

# A broadband optical cavity spectrometer for measuring weak near-ultraviolet absorption spectra of gases

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**Abstract.** Accurate absorption spectra of gases in the near-ultraviolet (300 to 400 nm) are essential in atmospheric observations and laboratory studies. This paper describes a novel incoherent broadband cavity-enhanced absorption spectroscopy (IBBCEAS) instrument for measuring very weak absorption spectra from 335 to 375 nm. The instrument performance was validated against the  ${}^3\text{B}_1\text{-X}^1\text{A}_1$  transition of  $\text{SO}_2$ . The measured absorption varied linearly with  $\text{SO}_2$  column density and the resulting spectrum agrees well with published spectra. Using the instrument, we report new absorption cross-sections of  $\text{O}_3$ , acetone, 2-butanone, and 2-pentanone in this spectral region, where literature data diverge considerably. In the absorption minimum between the Huggins and Chappuis bands, our absorption spectra fall at the lower range of reported ozone absorption cross-sections. The spectra of the ketones agree with prior spectra at moderate absorptions, but differ significantly at the limits of other instruments' sensitivity. The collision-induced absorption of the  $\text{O}_4$  dimer at 360.5 nm was also measured and found to have a maximum cross-section of ca.  $4.0 \times 10^{-46} \text{ cm}^5 \text{ molecule}^{-2}$ . We demonstrate the application of the instrument to quantifying low concentrations of the short-lived radical, BrO, in the presence of stronger absorptions from  $\text{Br}_2$  and  $\text{O}_3$ .

## 1 Introduction

Whereas large absorption cross-sections of atmospheric species are generally well established, weak absorption spectra are often poorly known. Accurate absorption cross-sections are important in atmospheric observations, labora-

tory work, and studies of fundamental molecular properties. The near-ultraviolet (300 nm to 400 nm) absorption spectra of gases are particularly relevant for understanding and modelling the chemistry of the atmosphere, since light in this spectral region drives most tropospheric photochemistry.

Weak absorptions, which occur when molecular absorption cross-sections are small or when the concentration of the species of interest is low, are a challenge to measure. In ozone, for instance, the absolute absorption cross-section drops to about  $10^{-23} \text{ cm}^2 \text{ molecule}^{-1}$  between the Chappuis and the Huggins bands (around 370 nm). Despite its prominent role in atmospheric chemistry, even recent  $\text{O}_3$  absorption cross-sections differ by almost an order of magnitude in this region (Brion et al., 1998; Bogumil et al., 2003; Voigt et al., 2001). The low concentrations associated with short-lived species also result in weak absorption. Bromine monoxide, BrO, exemplifies this case: although the absorption cross-section of BrO is large (about  $1.5 \times 10^{-17} \text{ cm}^2 \text{ molecule}^{-1}$  at 338.5 nm) (Fleischmann et al., 2004), the tropospheric mixing ratio is typically only a few ppt (Honninger and Platt, 2002; Peters et al., 2005; Wagner et al., 2007; Mahajan et al., 2009). At a high atmospheric concentration of 10 pptv, the maximum absorption of BrO would be only about  $4 \times 10^{-9} \text{ cm}^{-1}$ , making it extremely difficult to measure. Even in the laboratory, low concentrations of BrO are inevitable owing to its fast self-reaction (Fleischmann et al., 2004). Lastly, low gas concentrations are found also for substances with low vapour pressures; as a result, the ultraviolet spectra of many low volatility organic species of atmospheric interest have yet to be determined. This includes the large group of compounds with carbonyl or nitro functional groups, for which the  $n \rightarrow \pi$  transition occurs in the near-ultraviolet.

The pathlength of light through the sample primarily determines the spectral quality for weak absorptions. Long pathlengths result in larger, more easily measured absorptions



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according to the Beer-Lambert law. Extensive measurement times are required if the pathlength through the sample is relatively short (on the order of metres, say) which imposes stringent conditions on the stability of both the chemical source and the spectrometer during the measurement. Multi-pass cells can extend the pathlength from hundreds of metres to a few kilometres, but require large sample volumes (Steyert et al., 2001). Optical cavities provide an alternative means of obtaining a long effective pathlength of light (typically several kilometres) through the sample, and techniques exploiting this property include cavity ring-down spectroscopy and cavity-enhanced absorption spectroscopy (Mazurenka et al., 2005). Such laser-based techniques are exceptionally sensitive to weak absorptions ( $10^{-9} \text{ cm}^{-1}$  or better) but have some drawbacks for recording broad spectra. Because the laser must be scanned across the spectrum, the approach is limited by the stability of the absorber concentration over the duration of the measurement. Furthermore, the intrinsic high resolution of laser is compromised if measurements are taken at widely spaced intervals, as is sometimes the case (Xiang et al., 2009). The cost and experimental complexity of ultraviolet tuneable laser systems may also be prohibitive.

The simultaneous measurement of a broad spectral window across the long effective pathlength of an optical cavity would be convenient for measuring weak absorption spectra. A method combining these criteria is Incoherent Broadband Cavity-Enhanced Absorption Spectroscopy (IBBCEAS), first demonstrated by Fiedler and co-workers (Fiedler et al., 2003). IBBCEAS employs an optical cavity, a broadband light source (usually a Xe arc lamp or light emitting diode), and a spectrograph to cover a relatively broad spectral window with high sensitivity to weak absorptions (or other extinction processes) in the sample. The technique has proved valuable for quantifying trace gases and aerosol extinction in chamber and field measurements (Venables et al., 2006; Dixneuf et al., 2009; Ball et al., 2010; Varma et al., 2009; Thalman and Volkamer, 2010; Washenfelder et al., 2008) and good agreement with other techniques has been demonstrated in recent intercomparisons of  $\text{NO}_2$  and  $\text{NO}_3$  measurements (Dorn et al., 2011; Fuchs et al., 2010). Most IBBCEAS measurements thus far have been at visible wavelengths, with the shortest wavelength measurements between 360 nm and 380 nm to quantify HONO and  $\text{NO}_2$  (Gherman et al., 2008). To our knowledge, however, the technique has not been employed to record high quality reference spectra.

Our aim in this paper is to demonstrate an IBBCEAS instrument for measuring weak absorption spectra of gases over the wavelength range 335 to 375 nm. We do so by first showing the comparability of  $\text{SO}_2$  absorption spectra measured by the instrument with literature spectra. The technique is then applied to obtain moderate resolution spectra of several molecules for which the absorption cross-section is poorly constrained in this spectral region, namely ozone, the  $\text{O}_4$  dimer, and the small ketones acetone, 2-butanone, and 2-pentanone. Moreover, as weak absorption spectra are

often used to quantify sample constituents, we also demonstrate the retrieval of the short-lived species BrO from a background of  $\text{Br}_2$  and  $\text{O}_3$ .

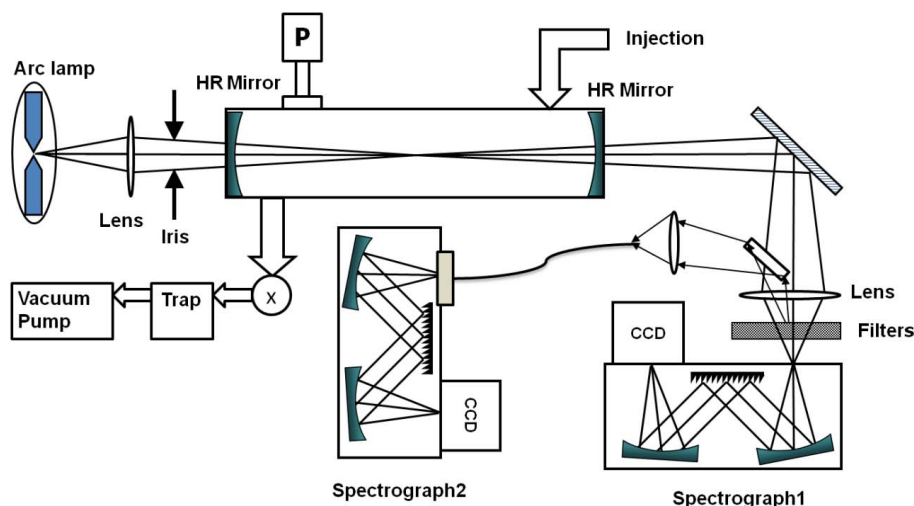
## 2 Experimental section

### 2.1 Spectrometer setup

The IBBCEAS system is shown schematically in Fig. 1. Light from a 75 W arc Xe lamp was focused into the centre of a 198 cm long optical cavity formed by two high reflectivity dielectric mirrors (Layertec). The mirrors had a maximum reflectivity of ca. 99.8% at 365 nm and a 5 m radius of curvature. Focusing the beam into the cavity centre improved the cavity quality and minimised dispersion losses (Fiedler et al., 2007; Wu et al., 2008). Light exiting the cavity was focussed into a Czerny-Turner spectrograph (Andor SR-163) with a 163 cm focal length and a  $12001 \text{ mm}^{-1}$ , 300 nm blaze grating. The spectrum was recorded using a back-illuminated CCD (Andor 420A-BU) optimized for the near-UV and cooled to  $-40^\circ \text{C}$  to reduce the dark current. The wavelength scale of the spectrograph was calibrated by fitting several emission lines of a Hg/Ne lamp to a second-order polynomial function. The uncertainty of the wavelength calibration was about 0.02 nm and the average resolution was 0.31 nm (25  $\mu\text{m}$  slit width) and 0.55 nm (100  $\mu\text{m}$  slit width) as determined from emission lines at 334 nm and 365 nm. The lamp output was highly stable: the short term intensity fluctuation of 1 s duration scans was below 0.1% and dimming of around 0.4% was observed over a 2 h period.

One of the practical challenges with IBBCEAS systems employing Xe arc lamps is to filter the extremely broadband light (200 nm to 2000 nm) to the high reflectivity region of the cavity mirrors. The cavity mirrors act as a band rejection filter but longer and shorter wavelengths are efficiently transmitted into the spectrograph, thereby adding to the background signal. A Schott UG11 band pass filter (270 to 380 nm) removed most visible and near-IR light in our system, while a heat-absorbing filter (Schott KG1) rejected wavelengths above 900 nm. A final band-pass filter (Semrock FF01-357/44-25) placed in front of the spectrograph provided a final, close match of the filter transmission to the high reflectivity range of the cavity mirrors and a usable spectral region from 335 nm to 375 nm.

The concentration of absorber must be known to calculate the absorption cross-section. Although any analytical method could be used to quantify the concentrations of  $\text{O}_3$  and the ketones, here we exploit the single-pass absorption at 320.1 nm to determine the concentration of the absorber. At this wavelength, which falls outside of the high reflectivity region of the optical cavity mirrors, most light traverses the sample once (a pathlength of 2 m). In contrast, the spectrometer is dramatically more sensitive between 335 and 375 nm



**Fig. 1.** Schematic diagram of the IBBCEAS instrument. Spectrograph 1 recorded the IBBCEAS spectra, while spectrograph 2 recorded single-pass spectra outside of the cavity mirror reflectivity range.

where the reflectivity is at a maximum and the effective path-length of light is about 0.5 km. (The effective pathlength in the absence of sample extinction is given by  $d/(1-R)$ , where  $d$  is the mirror separation and  $R$  is the mirror reflectivity). The IBBCEAS measurement is therefore about two orders of magnitude more sensitive than the 320 nm (single-pass) measurement. These species have moderate absorption cross-sections ( $10^{-20}$  cm<sup>2</sup> molecule<sup>-1</sup>) at 320 nm and literature spectra agree closely (Meyrahn et al., 1986; Hynes et al., 1992; Martinez et al., 1992; Brion et al., 1998; Gierczak et al., 1998; Burrows et al., 1999; Mu and Mellouki, 2000; Wollenhaupt et al., 2000; Voigt et al., 2001; Bogumil et al., 2003). The absorption was determined by measuring a small portion of the light reflected by the Semrock filter, transmitted through an Andover 320 nm band-pass filter, and recorded in another spectrograph (Andor SR-303). As the cavity mirrors have a reflectivity of 42% at 320 nm (as measured in a UV-visible spectrometer), about 20% of the light makes two or more passes through the sample. To account for multiple passes, the absorption coefficient was therefore determined from the general cavity equation, Eq. (4), in Fiedler et al. (2003).

The sample cell and optical cavity were enclosed in a vacuum tight system with the middle of the cell enclosed by a 100 cm glass tube. The outlet was connected to a foreline trap and rotary vane pump. The system pressure was monitored with a Leybold CTR 90 capacitance manometer and samples were studied at atmospheric pressure (ca. 1013 mbar) and under moderate vacuum. Acetone (Aldrich 99.9%), 2-pentanone (Aldrich 99%), 2-butanone (Aldrich 99%), bromine (Aldrich reagent grade) and SO<sub>2</sub> (BOC 99.9%) were used as received. Samples were introduced into the chamber at well-defined flow rates via MKS mass flow controllers. Br<sub>2</sub> and the ketones were introduced

into a stream of N<sub>2</sub> via a trap assembly. To reduce the Br<sub>2</sub> concentration to below  $10^{-13}$  cm<sup>-3</sup>, the trap was cooled to around  $-30$  °C in an alcohol/liquid nitrogen cooling bath and the Br<sub>2</sub>/N<sub>2</sub> stream diluted further in a higher flow-rate N<sub>2</sub> stream. Ozone was produced in a flow of oxygen by an ozone generator (OzoneLab OL80W).

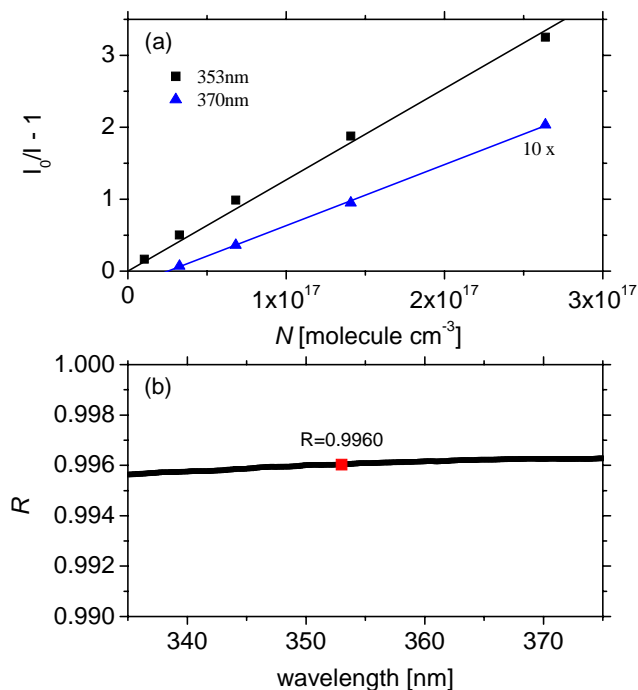
## 2.2 Mirror calibration

The extinction (or absorption) spectrum of a sample in the cavity is given by (Fiedler et al., 2003):

$$\varepsilon(\lambda) = \sum_i \sigma_i(\lambda) \cdot N_i = \frac{1}{d} \left( \frac{I_0(\lambda)}{I(\lambda)} - 1 \right) (1 - R(\lambda)) \quad (1)$$

where  $I$  and  $I_0$  are, respectively, the intensities transmitted through the optical cavity in the presence and absence of an absorber,  $d$  is the distance between the cavity mirrors,  $R$  is the mirror reflectivity, and  $\sigma_i$  (cm<sup>2</sup> molecule<sup>-1</sup>) and  $N_i$  (molecules cm<sup>-3</sup>) are the absolute absorption cross section and number density of species  $i$ . Equation (1) implies that the lamp intensity must be stable between the measurement of  $I_0$  and  $I$ , and that the mirror reflectivity is known. The reflectivity spectrum of the mirrors was calibrated in two steps. In the first step, the shape of the reflectivity spectrum was determined based on the difference in Rayleigh scattering of N<sub>2</sub> and CO<sub>2</sub>. The cavity was successively flushed with CO<sub>2</sub> and N<sub>2</sub> at about 1 L min<sup>-1</sup> for an hour each until the transmitted intensity attained a stable value. The sample was at 1013 mbar and 293 K. The observed extinction,  $\varepsilon_i$ , arising from Rayleigh scattering of gas  $i$  is related to the mirror reflectivity according to:

$$\alpha_i(\lambda) = N_i \sigma_i^{\text{Rayl}}(\lambda) = \left( \frac{I_0(\lambda)}{I_i(\lambda)} - 1 \right) \frac{(1 - R(\lambda))}{d} \quad (2)$$



**Fig. 2.** (a) Fractional absorption of O<sub>3</sub> at 353 nm and 370 nm. The slope of the graph,  $\sigma_{\text{O}_3}d/(1-R)$ , was used to determine  $R$  based on the literature absorption cross-section at 353 nm. (b) Mirror reflectivity scaled to give the literature cross-section of O<sub>3</sub> absorption at 353 nm.

In Eq. (2),  $I_0$  represents the (unmeasured) intensity of a non-scattering medium (that is, equivalent to the evacuated cavity) and  $\sigma^{\text{Rayl}}$  is the Rayleigh scattering cross-section of species  $i$ . Rayleigh scattering cross-sections were based on Naus and Ubachs (2000), Sneep and Ubachs (2005), Sutton and Driscoll (2004). The mirror reflectivity is thus related to the measured intensity through both CO<sub>2</sub> and N<sub>2</sub> by:

$$\frac{(1-R(\lambda))}{d} = \frac{I_{\text{N}_2}(\lambda)N_{\text{N}_2}\sigma_{\text{N}_2}^{\text{rayl}}(\lambda) - I_{\text{CO}_2}(\lambda)N_{\text{CO}_2}\sigma_{\text{CO}_2}^{\text{rayl}}(\lambda)}{I_{\text{CO}_2}(\lambda) - I_{\text{N}_2}(\lambda)} \quad (3)$$

Although Rayleigh scattering gives a smooth mirror reflectivity spectrum (Fig. 2b), the relatively small extinction can be susceptible to drifts in lamp intensity, thereby introducing a systematic error. The second step of the mirror calibration therefore scaled the reflectivity to match the absorption of O<sub>3</sub> in the cavity to the literature absorption cross-section at 353 nm. The cross-section used was the average of four literature cross-sections convoluted to our instrumental function (Brion et al., 1998; Burrows et al., 1999; Voigt et al., 2001; Bogumil et al., 2003). Figure 2a shows the linear dependence of the fractional absorption with O<sub>3</sub> concentration, where the slope equals  $\sigma_{\text{O}_3}d/(1-R)$ , as indicated by Eq. (1). The scaled mirror reflectivity is rather flat over this spectral range (Fig. 2b). The accuracy of the mirror calibration at 353 nm is limited by the 7% standard deviation of the

cross-sections at 353 nm, as well as the uncertainty in the individual cross-sections and the regression slope. The uncertainty in the reflectivity at other wavelengths is expected to be somewhat higher further away from 353 nm but cannot be explicitly quantified. Nevertheless, spectra generally agree closely with literature spectra, with some suggestion of a small overestimation of the mirror reflectivity below 340 nm in the ketone spectra.

### 2.3 Uncertainty of cross-sections

The uncertainty of the absorption cross-sections was estimated using standard uncertainty propagation. The uncertainty in the extinction coefficient,  $\Delta\varepsilon$ , was common to all spectra and was evaluated from Eq. (1):

$$\begin{aligned} \Delta\varepsilon^2 &= \left| \frac{\partial\varepsilon}{\partial d} \right|^2 \Delta d^2 + \left| \frac{\partial\varepsilon}{\partial(1-R)} \right|^2 \Delta(1-R)^2 + \left| \frac{\partial\varepsilon}{\partial I_0} \right|^2 \Delta I_0^2 \\ &= \varepsilon^2 \left( \frac{\Delta d}{d} \right)^2 + \varepsilon^2 \left( \frac{\Delta(1-R)}{1-R} \right)^2 + \left( \frac{1-R}{d} \right)^2 \left( \frac{\Delta I_0}{I} \right)^2 \end{aligned} \quad (4)$$

where  $\Delta d$ ,  $\Delta(1-R)$ , and  $\Delta I_0$  are the uncertainties in the mirror separation, the mirror reflectivity term, and the intensity. Here  $(1-R)$  and  $d$  are systematic sources of uncertainty, whereas the intensity term contributes randomly to the final uncertainty. The relative uncertainties in Eq. (4) are 7% for  $(1-R)$  at 353 nm, 0.5% for  $d$ , and less than 1% for short-term intensity fluctuations. The overall uncertainty in the extinction measurement at 353 nm is therefore around 7% and is clearly dominated by the uncertainty in the mirror, reflectivity. However, the characterisation of the lamp intensity appears to understate the longer-term drift in the intensity, as seen in the small offset at long wavelengths in the spectra of 2-butanone and 2-pentanone (where the absorption is expected to be effectively zero). On the basis of this evidence, the intensity contribution to the uncertainty is estimated to be around  $5 \times 10^{-7} \text{ cm}^{-1}$ .

The uncertainty in the absorption cross-sections must furthermore account for the uncertainty in the concentration of each species ( $\Delta N$ ):

$$\left( \frac{\Delta\sigma}{\sigma} \right)^2 = \left( \frac{\Delta\varepsilon}{\varepsilon} \right)^2 + \left( \frac{\Delta N}{N} \right)^2 \quad (5)$$

We have estimated  $\Delta N$  from the standard deviation of the literature cross-sections at 320 nm (Table 1), together with a 1.1% uncertainty in the single-pass absorption measurement. The uncertainties in the concentrations are 3% for O<sub>3</sub>, 4% for acetone, 3% for 2-butanone and 3% for 2-pentanone. The overall uncertainties in absorption cross-sections are thus 8% (O<sub>3</sub>), 8% (acetone), 10% (2-butanone) and 8% (2-pentanone).

We emphasize that this uncertainty applies to the absorption cross-section at 353 nm and is necessarily a lower limit. The uncertainty is larger at other wavelengths, and particularly at the long and short wavelength limits of the IBBCEAS spectrum. The uncertainty will therefore be revisited in the

**Table 1.** Instrumental parameters of literature absorption cross sections.

Author (year)	Temp. [K]	Wavelength [nm]	Resolution [nm]
Sulfur dioxide			
Bogumil et al. (2003)	293	239–395	0.26
Vandaele et al. (2009)	298	227–420	0.03
Ozone			
Bogumil et al. (2003)	293	230–1075	0.26
Voigt et al. (2001)	293	230–851	0.027
Burrows et al. (1999)	293	230–794	0.2–0.4
Brion et al. (1998)	295	350–830	0.01
Acetone			
Yujing et al. (2000)	298	240–350	0.04
Wollenhaupt et al. (2000)	296	220–346	0.16
Gierczak et al. (1998)	298	215–349	0.5
Martinez et al. (1992)	300	202–355	0.5
Hynes et al. (1992)	298	253.7–360	2
Meyrahn et al. (1986)	298	275–368	1.8
2-Butanone			
Yujing et al. (2000)	298	240–350	0.04
Martinez et al. (1992)	300	202–355	0.5
2-Pentanone			
Horowitz et al. (1999)	298	210–360	0.07
Martinez et al. (1992)	300	202–354	0.5

Discussion section to place a more conservative estimate on the overall uncertainty in the absorption cross-section.

### 3 Results

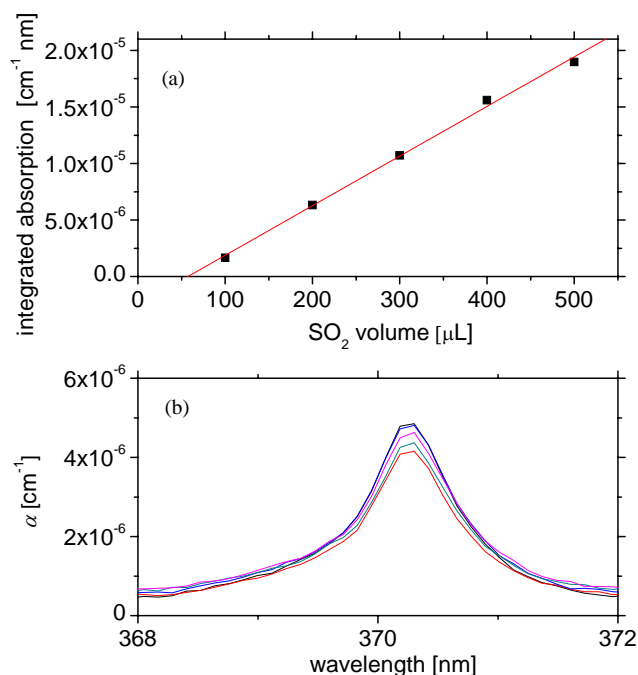
#### 3.1 Sulfur dioxide absorption

SO<sub>2</sub> has a relatively weak absorption band from 345 to 420 nm arising from the spin-forbidden  $a^3B_1-X^1A_1$  transition. Since recent spectra agree closely (Vandaele et al., 2009; Bogumil et al., 2003) and because SO<sub>2</sub> is stable and easily handled, the absorption of SO<sub>2</sub> in this region served as a convenient test of the spectrometer performance.

SO<sub>2</sub> was introduced into the spectrometer from a 500  $\mu$ L gas-tight syringe into a stream of N<sub>2</sub> (0.5 L min<sup>-1</sup>) at ambient temperature (296 K) and pressure (1013 mbar). The SO<sub>2</sub> flowed through the optical cavity as a short plug of gas with a relatively constant total column density. The sample absorption rose to a maximum for a short period before subsiding as the SO<sub>2</sub> was flushed from the chamber. Diffusion of gas into the dead volumes near the cavities mirrors resulted in a relatively long residence time of about 6 min. Measurements were based on the maximum absorption in the cell. The inte-

grated absorption from 350 to 375 nm increased linearly with volume of SO<sub>2</sub> (Fig. 3a). The small, negative intercept may have arisen from residual SO<sub>2</sub> in the chamber from one injection to another. Repeatability was good, with a standard deviation of 5% at the absorption peak at 370.2 nm found for five successive injections of 500  $\mu$ L SO<sub>2</sub> (Fig. 3b).

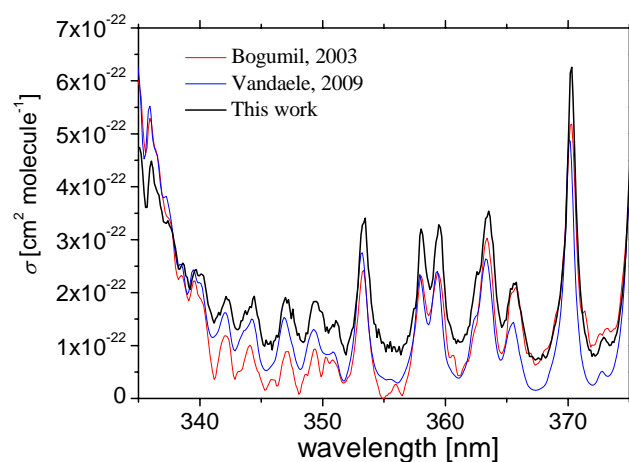
The average absorption cross-section of SO<sub>2</sub> from the five samples is shown in Fig. 4. The number density was calculated from the volume of the SO<sub>2</sub> injected and the volume of the sample cell. Figure 4 also displays the spectra of Bogumil et al. (2003) and Vandaele et al. (2009) at resolutions of 0.26 nm and 0.032 nm, respectively, as well as magnitudes of the absorption maxima convoluted to our instrument resolution. The magnitudes of the five largest absorption features above 350 nm were 28% larger than the average literature cross-sections, although around half of the difference ( $8.2 \times 10^{-23}$  cm<sup>2</sup> average) can be explained by our generally higher baseline ( $4.2 \times 10^{-23}$  cm<sup>2</sup> higher on average). Nevertheless, the agreement between the spectrum reported here and the literature spectra is generally good over this region, which includes changes in absorption coefficient of about an order of magnitude. Together with the absorption linearity and repeatability, these data thus provide evidence of the reliability of the spectra reported here.



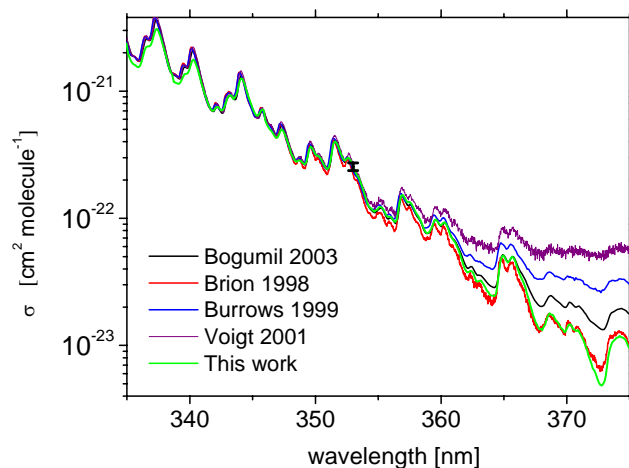
**Fig. 3.** (a) Dependence of the  $\text{SO}_2$  integrated absorption on the volume of  $\text{SO}_2$  introduced into the inlet stream. The correlation coefficient of the linear regression is  $R^2 = 0.9961$ . (b) Repeatability of five successive injections of  $500 \mu\text{L}$   $\text{SO}_2$  into the inlet stream. The standard deviation is 5% at the absorption peak at 370.2 nm.

A comparison of the experimental parameters illustrates the sensitivity of our instrument. The  $\text{SO}_2$  mole fraction of the sample gas is about 0.039% in our spectrometer compared to pure  $\text{SO}_2$  and 1%  $\text{SO}_2$  in  $\text{N}_2$  of Vandaele and Bogumil, respectively. Furthermore, the total measurement of 360 s for the averaged spectra (including the measurement of  $I_0$ ) was similar to that of the Bogumil spectra (320 s), but much shorter than the acquisition time of ca. 1 h that Vandaele and co-workers used for several pressure measurements. On the other hand, the lower resolution of our instrument would tend to improve our signal-to-noise over the much higher resolution of the Vandaele and Bogumil spectra. Despite these differences, the noise apparent in the absorption minima is comparable to or smaller than that in the literature spectra.

This absorption band of  $\text{SO}_2$  has some practical potential. The structured absorption features in this region have recently been used in DOAS measurements of  $\text{SO}_2$  column densities (Bobrowski et al., 2010). For a broadband optical cavity spectrometer like ours, injection of small volumes of  $\text{SO}_2$  would also be a convenient approach to calibrate mirror reflectivities since the small absorption cross-section results in moderate, easily measurable absorptions. Strongly absorbing gases such as  $\text{NO}_2$  must be highly diluted to low levels or no light will be transmitted through the optical cavity.



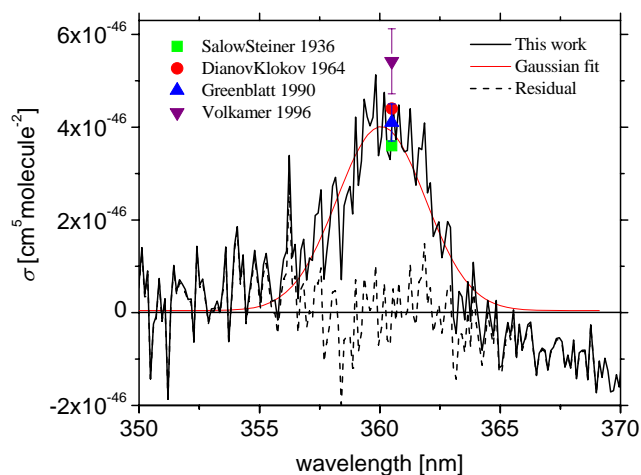
**Fig. 4.**  $\text{SO}_2$  absorption cross-sections spectra from this work (black), Bogumil (red) and Vandaele (blue). The literature spectra were convoluted to our spectrograph resolution.



**Fig. 5.**  $\text{O}_3$  absorption cross-sections from this work and recent literature spectra (Voigt et al., 2001; Burrows et al., 1999; Bogumil et al., 2003; Brion et al., 1998). The standard deviation of literature cross-sections at 353 nm is indicated by an error bar.

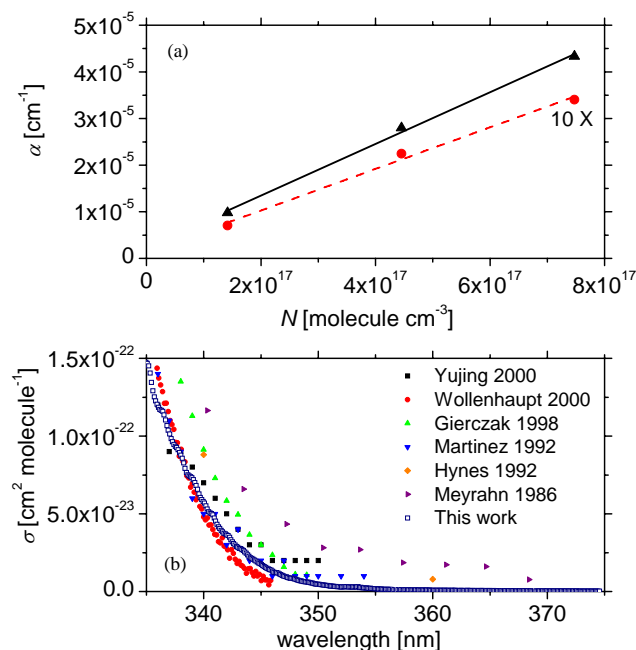
### 3.2 Ozone absorption

In contrast to the  $\text{SO}_2$  absorption cross-sections, relatively poor agreement exists between reported cross-sections of  $\text{O}_3$  in the minimum between the Huggins and Chappuis bands (Bogumil et al., 2003; Burrows et al., 1999; Voigt et al., 2001; Brion et al., 1998). Although various workers report similar absorption features, the absorption cross-sections differ enormously around 370 nm (Fig. 5). Agreement above 360 nm is poor owing to the very weak absorption and the relatively modest concentrations of  $\text{O}_3$  in samples. With the long effective pathlength of our instrument, we report an improved estimate of ozone absorption cross-section in this region.



**Fig. 6.** Experimental spectrum (black) and Gaussian fit (red) of the collision-induced absorption of the  $O_4$  dimer. Literature maxima shown are those of Salow and Steiner (green square), Greenblatt (blue triangle), Dianov-Klokov (red circle), and Volkamer (purple triangle). To facilitate comparison with the literature maxima, our measured spectrum was shifted downwards by the offset of the Gaussian fit.

Intensity spectra of around 1%  $O_3$  in  $O_2$  were recorded with a 5 min averaging time and at a resolution of 0.26 nm. The ozone concentration in the cavity was quantified from the absorption at 320 nm. The absorption cross-section and literature spectra are shown in Fig. 5. Of the literature spectra, Brion et al. (1998) and Voigt et al. (2001) employed FTS, while the data from Burrows et al. (1999) and from Bogumil et al. (2003) employed monochromators developed for satellite monitoring. The IBBCEAS spectrum was corrected for the weak collision-induced absorption of the  $O_4$  dimer at 360 nm (see below). (Note that the magnitude of the absorption cross-section at 353 nm is fixed to the average literature cross-section by the calibration of the mirror reflectivity.) Below 340 nm, our cross-sections are 10 to 20% smaller than the literature values, probably because our mirror reflectivity is overestimated at short wavelengths. At longer wavelengths, the spectra diverge, with the IBBCEAS spectrum agreeing closely with the Brion spectrum. Remarkably little noise is apparent in the IBBCEAS spectrum, and linear absorption behaviour with concentration is apparent even at 370 nm where the absorption is weak (Fig. 2a). Hence the spectrum reported here provides good evidence that the  $O_3$  absorption in this region is even weaker than most previous reports suggest.



**Fig. 7.** (a) Dependence of absorption coefficient on the acetone concentration at 340 nm (black triangles) and at 350 nm (red circles). Correlation coefficients are 0.996 and 0.987, respectively. (b) Acetone absorption cross-section of this work and of several literature spectra.

The collision-induced absorption of the  $O_4$  dimer ( $^1\Sigma_g^+ + ^1\Sigma_g^+ (v=1) \leftarrow ^3\Sigma_g^- + ^3\Sigma_g^+$ ) contributes marginally to the  $O_3$  absorption spectrum at 360.5 nm. The cross-section of the dimer was determined by comparing the sample extinction of pure  $O_2$  versus that of  $N_2$ . Differences between the Rayleigh scattering cross-sections of  $N_2$  and  $O_2$  are small and contribute only a small baseline offset to the extinction. The observed absorption feature was fitted to a Gaussian line shape on a sloping baseline (Fig. 6). The magnitude of the Gaussian maximum was  $4.0 \times 10^{-46} \text{ cm}^5 \text{ molecule}^{-2}$  at 360.1 nm with a FWHM of 4.3 nm, within 9% of the average literature value for the  $O_4$  cross-section peak, and in satisfactory agreement with the careful measurements of Greenblatt and co-workers ( $4.10 \times 10^{-46} \text{ cm}^5 \text{ molecule}^{-2}$  at 360.5 nm with a FWHM of 4.0 nm) and other work (Greenblatt et al., 1990; Dianov-Klokov, 1964; Volkamer, 1996; Salow and Steiner, 1936). Nevertheless, our signal-to-noise ratio is small despite the relatively long effective pathlength of our system ( $L_{\text{eff}} = d/(1-R) \approx 500 \text{ m}$ ) in pure  $O_2$ : the spectrum of Greenblatt taken at 55 atm  $O_2$  is to be preferred to our results. More precise estimates of the absorption cross-section would be feasible with higher mirror reflectivities or at higher  $O_2$  pressures.

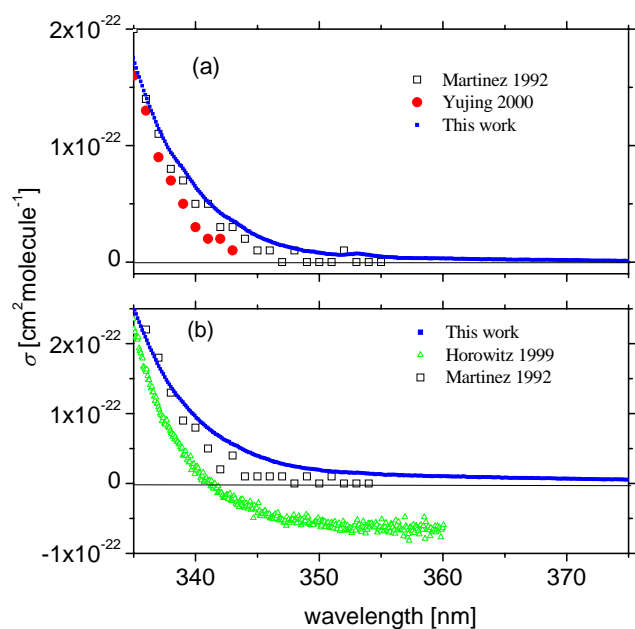


Fig. 8. Absorption cross-sections of (a) 2-butanone and (b) 2-pentanone.

### 3.3 Ketone spectra

Spectra of acetone, 2-pentanone, and 2-butanone were measured from 335 to 355 nm. The ketones were introduced via a  $N_2$  stream ( $0.5 \text{ L min}^{-1}$ ) into the cavity and quantified from the literature absorption cross-sections at 320 nm. The acetone cross-sections used were those of Gierczak et al. and the 2-pentanone and 2-butanone cross-sections were taken from Martinez and co-workers (Gierczak et al., 1998; Martinez et al., 1992). The acetone concentration was further diluted to check the linearity of the absorption measurement. The absorption of acetone increased linearly with concentration at 340 nm and 350 nm, where the absorption differs by over an order of magnitude (Fig. 7a). For acetone, the absorption cross-section was derived from the slope of the linear regression of the three concentrations; Fig. 7b shows the resulting spectrum reported to 355 nm (above which the correlation coefficient  $R^2$  fell below 0.9). The lamp stability would increasingly influence the measurement of progressively weaker absorption at longer wavelengths.

The resulting absorption cross-section of acetone compares favourably with the available literature spectra. Most literature spectra are evidently close to their detection limits for such weak absorptions and agreement is poor even at the shortest wavelengths (largest absorption) in Fig. 7b. Of the spectra, those of Mu and Mellouki, Wollenhaupt, Martinez, and Gierczak exhibit the lowest discernable measurement noise (Mu and Mellouki, 2000; Wollenhaupt et al., 2000; Martinez et al., 1992; Gierczak et al., 1998). Our spectrum is in reasonable agreement with these spectra. At wavelengths

shorter than 340 nm our measured absorption increases less strongly than other spectra, possibly indicating an overestimation of the mirror reflectivity at the short wavelength limit of the spectrometer. On the other hand, the linear dependence of the absorption with concentration at 340 nm, and particularly at 350 nm, provides strong evidence for the reliability of our cross-sections at these and similar wavelengths. Small undulations are visible in the IBBCEAS absorption spectra and probably correspond to the tail-end of small absorption structures that have been known for a long time (Noyes et al., 1956).

Spectra of 2-butanone and 2-pentanone are shown in Fig. 8. These spectra were determined from the absorption at a single concentration ( $2.07 \times 10^{17} \text{ molecule cm}^{-3}$  for 2-butanone and  $1.05 \times 10^{17} \text{ molecule cm}^{-3}$  for 2-pentanone). Cross-sections from this work agree with earlier spectra at the shortest wavelengths; again, lower noise is evident, especially at longer wavelengths. The spectrum of 2-butanone clearly agrees closely with that of Martinez and co-workers to 355 nm. For 2-pentanone, the IBBCEAS spectrum is higher than either the spectrum of Martinez or that of Horowitz. In the case of Horowitz et al., the negative values probably signify some measurement artefact (Horowitz, 1999). In like manner, the IBBCEAS spectrum does not fall to zero at longer wavelengths (as would be expected); we ascribe this result to a small decrease in the lamp output over the duration of the measurement, thereby giving rise to an apparent absorption.

### 3.4 Concentration measurement of BrO, Br<sub>2</sub>, and O<sub>3</sub>

With a broad spectral window, the concentrations of several species can be determined concurrently. We illustrate this application of the instrument by quantifying the composition of a sample comprising Br<sub>2</sub>, O<sub>3</sub>, and the short-lived species BrO. BrO was produced inside the cavity at reduced pressures (3 to 18 mbar) by photolysing Br<sub>2</sub> in the presence of O<sub>3</sub>:

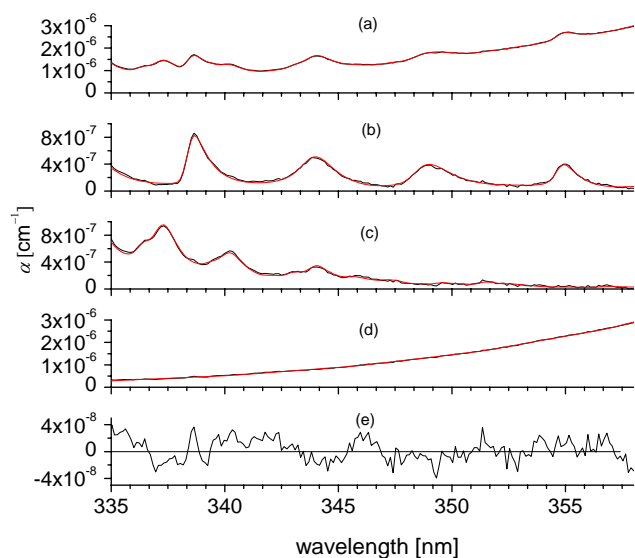


Molecular bromine, Br<sub>2</sub>, has a broad absorption band from 200 to 510 nm with a maximum at 420 nm and unity photolysis quantum efficiency (Atkinson et al., 2007). With filters placed behind the cavity, Br<sub>2</sub> is readily photolysed by residual light in the cavity. Under these conditions BrO formation is prompt. Nevertheless, rapid self-reaction suppresses the BrO concentration in the cavity:



BrO is readily identified by the  $A^2\Pi_{3/2} \leftarrow X^2\Pi_{3/2}$  transition, with the (7, 0), (6, 0), (5, 0) and (4, 0) vibrational





**Fig. 9.** Absorption and fitted spectra for the Br<sub>2</sub>, O<sub>3</sub>, and BrO system. (a) Measured spectrum (black) and fitted result (red). The individual components of the spectrum are (b) BrO, (c) O<sub>3</sub>, and (d) Br<sub>2</sub>. (e) Residual of the fit.

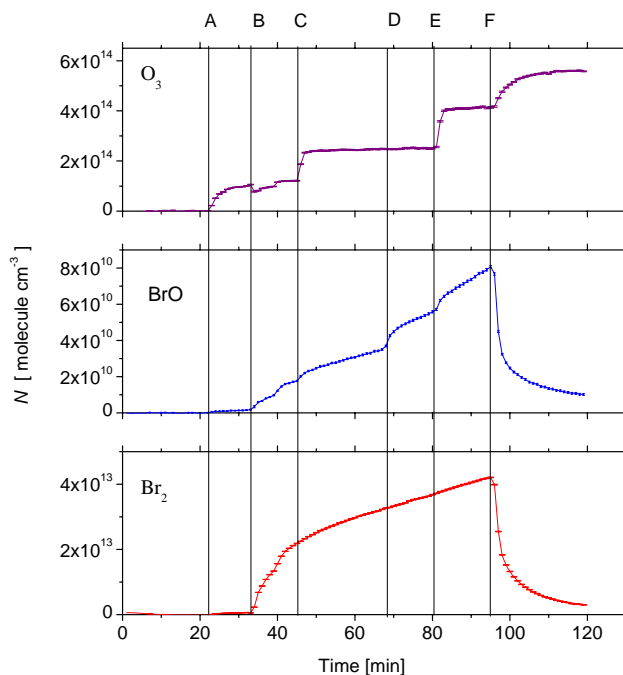
bands falling within the spectral range of our spectrometer (Fig. 9b). The large absorption cross-section of this transition permits sensitive retrieval of even very low concentrations of BrO.

The number densities of O<sub>3</sub>, Br<sub>2</sub>, and BrO were retrieved by least squares fitting their literature cross-sections to the absorption coefficient and including a linear offset:

$$\alpha(\lambda) = N_{\text{BrO}}\sigma_{\text{BrO}}(\lambda) + N_{\text{O}_3}\sigma_{\text{O}_3}(\lambda) + N_{\text{Br}_2}\sigma_{\text{Br}_2}(\lambda) + a + b\lambda \quad (10)$$

Literature cross-sections were those of Fleischman et al. (2004) for BrO, the IUPAC recommendation for Br<sub>2</sub> (Atkinson et al., 2007) and Voigt et al. (2001) for O<sub>3</sub> and were convoluted with our instrument function. As the composition of the sample is probably non-uniform in the sample cell, we note that the fitted number densities represent pathlength-integrated concentrations or column densities. A typical absorption spectrum of the sample and fit of the literature spectra over the range 336 nm to 358 nm is shown in Fig. 9. The absorption features of BrO and O<sub>3</sub> are readily evident against the monotonically increasing absorption of Br<sub>2</sub> with wavelength. The fitted curve closely matches the experimental absorption and the residual is small (RMS of  $1.8 \times 10^{-8} \text{ cm}^{-1}$ ). The fitted concentration values were little affected by extending the fitting window to 370 nm (a maximum change of 8% was found for  $N_{\text{Br}_2}$ ). Excluding parameters  $a$  and  $b$  influenced the fit result more, with the value of  $N_{\text{Br}_2}$  decreasing by 31%. In this case,  $a$  and  $b$  most likely compensate for a nominal offset in the absorption arising from a slow decline in the lamp intensity.

The concentration of BrO in the flow cell was explored as a function of reactant concentration (Fig. 10). The O<sub>3</sub> concen-



**Fig. 10.** Temporal dependence of the number density of O<sub>3</sub>, BrO, and Br<sub>2</sub> in the flow cell. Points A–F indicate changes in either the input concentrations or photolysis rate.

tration in the input stream was increased at points A, C, and E by ramping up the O<sub>3</sub> generator output. Br<sub>2</sub> was introduced at time B by flowing N<sub>2</sub> over Br<sub>2</sub> in a liquid nitrogen/alcohol cooling bath initially at  $-30^\circ\text{C}$ . As expected, photolysis of Br<sub>2</sub> in the presence of O<sub>3</sub> results in the immediate appearance of BrO; the BrO concentration increases with both Br<sub>2</sub> and O<sub>3</sub> concentration. The temperature of the cooling bath rose about  $10^\circ\text{C}$  over the course of the experiment, causing a gradual rise in the Br<sub>2</sub> and BrO concentrations. After the Br<sub>2</sub> source was stopped at time F, both Br<sub>2</sub> and BrO concentrations dropped off sharply, whereas the O<sub>3</sub> concentration rose by about a third as Br-catalysed O<sub>3</sub> loss is progressively choked off. The BrO concentration also rose at time D, when the Br<sub>2</sub> photolysis rate was increased by shining a high power 465 nm LED diagonally across the transparent centre of the flow tube.

The observed behaviour of the concentration of each species is in sensible agreement with known chemistry. Under stable inlet conditions, minimal scan-to-scan variation is observed. The precision of each concentration measurement was estimated based on the standard deviation from a seven-point running average between 49 and 65 min. Values of  $1.2 \times 10^{-12}$ ,  $4.5 \times 10^{-10}$ , and  $1.6 \times 10^{-8} \text{ molecules cm}^{-3}$  were found for O<sub>3</sub>, Br<sub>2</sub>, and BrO, respectively, which are comparable to the statistical uncertainties of the fit (respective averages of  $2.1 \times 10^{-12}$ ,  $9.0 \times 10^{-10}$ , and  $4.3 \times 10^{-8} \text{ molecules cm}^{-3}$ ). We note in passing that whereas our measurements were obtained under steady-state conditions

with low BrO concentrations, previous BrO measurements have recorded the absorption of higher instantaneous BrO concentrations, either by using a microwave discharge or with short-time observations following flash lamp or laser photolysis of Br<sub>2</sub> (Fleischmann et al., 2004; Wahner et al., 1988; Sander and Friedl, 1989; Basco and Dogra, 1971; Friedl and Sander, 1989; Gilles et al., 1997; Wilmouth et al., 1999; Laszlo et al., 1997; Cox et al., 1982; Orlando et al., 1991).

#### 4 Discussion

The results above show the potential of the IBBCEAS system for retrieving weak absorption spectra in the near-UV at moderate resolution. The instrument performance is affected by several factors which we discuss below.

Whereas earlier near-UV IBBCEAS systems employed a 365 nm LED, in this study a Xe arc lamp was used as the light source. The high radiance of LEDs at 365 nm is attractive for monitoring several species over a narrow spectral window, but the spectral output falls rapidly from the maximum and has a typical FWHM of only about 10 nm. High power LEDs are not yet available below 365 nm. In contrast, the Xe arc-lamp spectrum is relatively unstructured in this spectral range and better suited to retrieving reference spectra. Xe arc lamps also have useful output at even shorter wavelengths and can be used for relatively broad spectral coverage (greater than 100 nm with appropriate cavity mirrors), which we have demonstrated in recent chamber studies of low volatility organic species (Chen et al., 2011). Notwithstanding these advantages, arc-lamps (and other very broadband sources) require stringent filtering to minimise stray light levels.

The sensitive absorption of this spectrometer (ca.  $2 \times 10^{-8} \text{ cm}^{-1}$ ) is with mirrors of only modest reflectivity. Although adequate for the purposes of this study, the sensitivity of the instrument could be improved with higher reflectivity mirrors. Higher reflectivity mirrors also lower the cavity transmission and the signal-to-noise ratio, so a balance must be struck between these competing factors. The reflectivity of the dielectric mirrors must be calibrated. In this study, we used a combination of the Rayleigh scattering and the well-established absorption of O<sub>3</sub> at 353 nm to determine the overall reflectivity. The mirror reflectivity is mostly flat across the spectral window determined by the filters. Any uncertainty in the mirror reflectivity – or, more accurately, in the  $(1 - R)$  term in Eq. (1) – contributes linearly to the overall uncertainty in the absorption coefficient. In the results here, this uncertainty is 8 to 10% at 353 nm but increases towards the long and short-wavelength limits of the spectra owing to the uncertainty in the  $(1 - R)$  term. Below 340 nm, our spectra were 10 to 20% lower than previously reported absorption cross-sections, suggesting that the mirror reflectivity was overestimated at the short wavelength limit and

that our spectra are less accurate in this region. Taking the wavelength-dependence of the uncertainty into account, we estimate the uncertainty in our cross-sections above 340 nm to range from 10 to 15%, which is acceptable for many applications and in line with the uncertainty in prior IBBCEAS studies. This work has also demonstrated that transmission of the windows outside of the high reflectivity range can be exploited to quantify components of the sample based on absorption in other spectral regions. In the results above, the small uncertainty of the (larger) absorption cross-sections of O<sub>3</sub> and ketones allows direct in situ quantification of these species. Sample losses elsewhere in the system can therefore be ignored.

In comparison to numerous literature spectra, the IB-BCEAS spectra have significantly lower noise and also display linear absorption behaviour for weak absorptions. In these regions, where large variations between literature data are found, we believe that the spectra reported here are to be preferred to most prior spectra. Two caveats are in order, however. Firstly, even with long effective pathlengths, lamp intensity variations and gradual lamp dimming gives rise to apparent weak absorptions. This is evident in the small positive offset in the length wavelength limit of the 2-pentanone spectrum and implicitly in the linear offset of the fitting parameters of the BrO, O<sub>3</sub>, and Br<sub>2</sub> spectrum. Secondly, the accuracy of the absorption measurement tends to be worse at the limits of the high reflectivity region of the cavity mirrors, as we observe below 340 nm.

#### 5 Conclusions

A near-UV IBBCEAS instrument has been developed for measuring weak absorption spectra from 335 to 375 nm. Spectra of SO<sub>2</sub> measured with the instrument show good agreement with literature spectra. The instrument has been used to retrieve revised absorption cross-sections of O<sub>3</sub>, acetone, 2-butanone, and 2-pentanone, for which previously reported values diverge widely. The IBBCEAS spectra showed linear absorption behaviour and low spectral noise and provide revised estimates for the long-wavelength tail of these spectra. Concurrent retrieval of the concentrations of BrO, Br<sub>2</sub>, and O<sub>3</sub> in a steady-state flow system has been demonstrated.

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