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Tectonic overview of the West Gondwana margin

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Abstract

The oceanic southern margin of Gondwana, from southern South America through South Africa, West Antarctica, New Zealand (in its pre break-up position), and Victoria Land to Eastern Australia is one of the longest and longest-lived active continental margins known. It was the site of the 18,000 km Terra Australis orogen, which was initiated in Neoproterozoic times with the break-up of Rodinia, and evolved into the Mesozoic Australides. The Gondwana margin was completed, in Late Cambrian times, by closure of the Adamastor Ocean (between Brazilian and southwest African components) and the Mozambique Ocean (between East and West Gondwana), forming the Brasiliano–Pan-African mobile belts. During the Early Palaeozoic much of the southern margin was dominated by successive episodes of subduction-accretion. Eastern Australia, Northern Victoria Land and the Transantarctic Mountains were affected by one of the first of these events – the Late Cambrian Ross/Delamerian orogeny, remnants of which may be found in the Antarctic Peninsula – but also contain two accreted terranes of unknown age and origin. Similar events are recognized at the South American end of the margin, where the Cambrian Pampean orogeny occurred with dextral strike-slip along the western edge of the Río de la Plata craton, followed by an Ordovician active margin (Famatinian) associated with the collision of the Precordillera terrane. However, the

33 central part of the margin (the Sierra de la Ventana of eastern Argentina, the Cape
34 Fold Belt of South Africa and the Ellsworth Mountains of West Antarctica) seem to
35 represent a passive margin during the Early Palaeozoic, with the accumulation of
36 predominantly reworked continental sedimentary deposits (Du Toit's 'Samfrau
37 Geosyncline'). In many of the outer areas, accretion and intense granitic/rhyolitic
38 magmatism continued during the Late Palaeozoic, with collision of several small
39 continental terranes, many of which are nevertheless of Gondwana origin: e.g.,
40 southern Patagonia and (possibly) 'Chilenia' in the South American–South African
41 sectors, and the Western Province and Median Batholith terranes of New Zealand.
42 The rhyolitic Permo–Triassic LIP of southern South America represents a Permo-
43 Triassic switch to extensional tectonics, which continued into the early Jurassic, and
44 was followed by the establishment of the Andean subduction margin. Elsewhere at
45 this time the margin largely became passive, with terrane accretion continuing in New
46 Zealand. In the Mesozoic, the Terra Australis Orogen evolved into the accretionary
47 Australides, with episodic orogenesis in the New Zealand, West Antarctic and South
48 American sectors in Late Triassic–Early Jurassic and mid-Cretaceous times, even as
49 Gondwana was breaking up.

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51 Key words: Accretionary orogen, terrane, Palaeozoic, Laurentia, Rodinia

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1. Introduction

The oceanic margin of Gondwana was of the order of 40,000 km long (Fig. 1). Its northern boundary was the source of Avalonian and Cadomian terranes in the west and Cimmerian terranes in the east (Unrug, 1997). Its southern margin has been proposed as one of the largest and longest-lived accretionary orogens on Earth (Cawood, 2005; Vaughan et al., 2005b) - the Proterozoic and Palaeozoic Terra Australis orogen (Cawood, 2005), which evolved into the Australides (Vaughan et al., 2005b) during the Palaeozoic and Mesozoic. This orogen was over 18000 kilometres long, incorporating margins against the Iapetus and palaeo-Pacific oceans (Unrug, 1997) (Fig. 1), and is comparable in scale to the Late Palaeozoic Alleghenian/Hercynian/Uralian orogen of central Pangaea (Vaughan et al., 2005b). Today, the southern margin of Gondwana can be subdivided into Australian, Victoria Land, New Zealand, West Antarctic, South African and South American sectors (Figure 1). Apart from the West Antarctic and South African sectors, these have

91 recently been reviewed in a Geological Society Special Publication (Vaughan et al.,
92 2005a). The present paper focuses on the Iapetus and palaeo-Pacific margin of West
93 Gondwana (Fig. 1), i.e. the West Antarctic and South American sectors; it does not
94 deal with the collisional margin between East and West Gondwana, nor with the
95 Avalonian/Cadomian or Cimmerian margins (Fig. 1). However, it does touch on the
96 New Zealand and Victoria Land sectors (including the Transantarctic Mountains) of
97 the margin of East Gondwana, as these may have contributed detrital material and
98 terranes to the accretionary margin of West Gondwana from Palaeozoic times
99 onwards.

100

101 Moving clockwise along the southern margin of Gondwana, from modern-day east to
102 west (Figure 1), starting in East Gondwana, the Phanerozoic history of the Victoria
103 Land sector of the margin has recently been reviewed by Tessensohn and Henjes-
104 Kunst (2005) and the New Zealand sector has had recent and comprehensive reviews
105 by Mortimer (2004) and Wandres and Bradshaw (2005). Moving into West
106 Gondwana, aspects of the West Antarctic sector have been reviewed in the past 10
107 years by Pankhurst et al. (1998b) and Vaughan and Storey (2000), but is a sector of
108 the margin in need of an up-to-date treatment. Rapalini (2005) reviewed the southern
109 South American sector of the margin from the latest Proterozoic to the late Palaeozoic
110 on the basis of palaeomagnetic data, and a brief review of this sector was presented in
111 Vaughan et al. (2005b), but an up-to-date comprehensive review of the whole South
112 American sector is lacking. Given the pace of recent developments (e.g., Casquet et
113 al., 2006; Pankhurst et al., 2006), and the considerable controversy over the
114 Palaeozoic history of this sector of the margin, particularly regarding the origin of the

115 Precordillera or Cuyania terrane (e.g., Thomas and Astini, 2003; Finney et al., 2005),
116 a further review is appropriate.

117

118 **2. What is West Gondwana?**

119 In simple terms West Gondwana is that part of the supercontinent represented today in
120 South America, Arabia, Africa and West Antarctica. From a geological point of view,
121 however, this definition is over-simplified and it reflects a subdivision based on the
122 break-up rather than the amalgamation configuration of the supercontinent (e.g.,
123 Storey et al., 1996; Veevers, 2004). The earliest geologically-based separation of
124 Gondwana into eastern and western parts was made by Du Toit (1937) (Fig. 2). He
125 further separated Antarctica into eastern and western parts, as suggested by Suess
126 (1883–1901), assigning them to East and West Gondwana, respectively (see Thomson
127 and Vaughan (2005) for a brief discussion), but placed New Zealand in East
128 Gondwana, off the eastern coast of Australia (Fig. 2). More recently, West Gondwana
129 has been defined on the basis of the Archaean shields, cratons and cratonic fragments,
130 the intervening Mesoproterozoic and Neoproterozoic mobile belts, and the outer belts
131 of Proterozoic–Mesozoic terranes that make it up (e.g., Unrug, 1997; Pankhurst et al.,
132 1998b; Brito Neves et al., 1999; Vaughan and Storey, 2000; Murphy et al., 2004;
133 Tohver et al., 2006).

134

135 **2.1 Cratonic elements**

136

137 The major cratonic elements comprise the Amazonia–West Africa craton, Sao
138 Francisco–Congo craton, Kalahari–Grunehogna craton, Río de la Plata craton, and the
139 Arabian–Nubian shield (Tohver et al., 2006) (Fig. 3). Cordani et al. (2003) pointed

140 out that there are smaller cratonic fragments of considerable importance in
141 understanding the evolution of the supercontinent. These include the Central Goias
142 massif (Fischel et al., 2001), the Luiz Alves, Río Apa, Sao Luis and Paraná cratonic
143 fragments (e.g., Tohver et al., 2006) (Fig. 3). The Hoggar–Potiguar plate of Brito
144 Neves et al. (1999) is another potential cratonic fragment (e.g., Liegeois et al., 2003;
145 Ouzegane et al., 2003), although its limits are not currently delineated.

146

147 2.2 Mesoproterozoic and Neoproterozoic mobile belts

148

149 Brito Neves et al. (1999) used the term Brasiliano–Pan African collage for the
150 Mesoproterozoic and Neoproterozoic–Cambrian mobile belts involved in the final
151 amalgamation of West Gondwana. Tohver et al. (2006) listed 19 individual belts to
152 this collage, illustrated in Figure 3. Brito Neves et al. (1999) summarized them as the
153 Neoproterozoic Borborema/Trans-Saharan and Tocantins belts, and the
154 Neoproterozoic–Cambrian Pampean and Mantiqueira belts in modern-day South
155 America, and, in modern-day Africa, the Neoproterozoic Dahomeyide belt and the
156 Neoproterozoic–Cambrian Damara, and Zambesi belts. Other important parts of
157 Neoproterozoic–Cambrian West Gondwana include the Cariris-Velhos terrane (Brito
158 Neves et al., 1999) of northern South America–East Africa and the "Grenville"
159 Neoproterozoic rocks of the Haag Nunataks block of West Antarctica and the
160 Falklands Plateau (e.g., Storey et al., 1994; Wareham et al., 1998).

161

162 2.3 Palaeozoic–Mesozoic terranes

163

164 Accretion of new terrane material to Gondwana was active during amalgamation
165 (Cawood, 2005) and continued until the late stages of break-up of the supercontinent
166 (e.g., Vaughan et al., 2002b). In the Phanerozoic, these include the Cambrian rocks of
167 the Ellsworth–Whitmore Mountains block of West Antarctica (e.g., Curtis et al.,
168 1999), and the Cambrian rocks of the Western Province of New Zealand (Münker and
169 Cooper, 1995). Various Proterozoic fragments of West Gondwana also became part of
170 the margins of the Laurentia and Baltica cratons (Skehan, 1997). Murphy et al.
171 (2004) reviewed these and summarized them as being formed either of reworked
172 Neoproterozoic "juvenile crust within the Panthalassa-type ocean surrounding
173 Rodinia", the so-called Avalonian-type terranes, or of reworked West African
174 Palaeoproterozoic crust, the so-called Cadomian-type terranes. Following
175 amalgamation, the Gondwana margin continued to be active with addition of new
176 oceanic material (e.g., Cawood et al., 2002) and remobilization of existing parts of the
177 margin by strike-slip faulting (e.g., Cawood, 2005). Major episodes of terrane
178 addition and remobilization occurred during the Gondwanan Orogeny of the Permo-
179 Carboniferous (e.g., Cawood, 2005; Pankhurst et al., 2006) and during global
180 orogenesis in the Triassic–Jurassic and Cretaceous (e.g., Vaughan and Livermore,
181 2005).

182

183 2.4 Boundary with East Gondwana

184

185 The boundary with East Gondwana consists of a meandering zone of late
186 Neoproterozoic to earliest Cambrian orogenic and mobile belts, termed Pan-African,
187 extending from and including the Arabian–Nubian Shield in the north to Antarctica in
188 the south (e.g., Shackleton, 1996). Perhaps the most important of these belts is that of

189 the East African–Antarctic orogeny (Jacobs and Thomas, 2004). Unrug (1997) shows
190 a very broad zone of potential convergence in the northern segment, which include
191 eastern Africa and the Arabian–Nubian Shield. The southernmost extent of this
192 collision zone includes the Namaqua–Natal–Maud belt on the margin of the Kalahari–
193 Grunehogna craton in southern Africa and Dronning Maud Land in East Antarctica
194 (Jacobs et al., 2003). The essentially synchronous collision Brasiliano zone is the
195 subject of a new survey of geological links across the present South Atlantic region
196 (Pankhurst et al., in press).

197

198 **3. The formation and dispersal of West Gondwana**

199

200 Formation of the Gondwana supercontinent appears to have overlapped with the
201 break-up of Rodinia (a possible supercontinent built around Laurentia), which
202 occurred between 1000 and 750 million years ago (e.g. Cordani et al., 2003; Meert
203 and Torsvik, 2003). The series of accretionary and collisional events that formed
204 West Gondwana began 850 million years ago and were complete by the latest
205 Cambrian (490 million years ago) (e.g., Brito Neves et al., 1999). It is overly
206 simplistic to think of the final formation of Gondwana in terms of a collision between
207 the East and West parts (e.g., Meert, 2001). Recent palaeomagnetic data (Tohver et
208 al., 2006) suggest that prior to final amalgamation of Gondwana in the mid-Cambrian,
209 the Amazon–West Africa block of West Gondwana was still a separate entity from
210 Rodinia, and was separated from other blocks that constitute West Gondwana
211 (Congo–São Francisco–Kalahari–Arabia–Río de la Plata). Trindade et al. (2006)
212 provided palaeomagnetic support for this for Amazonia and proposed that
213 amalgamation involved successive suturing along three major orogenic belts, the

214 Mozambique, Kuunga and Pampean–Araguaia belts through closure of the
215 Mozambique, Adamastor and Clymene oceans. However, the associated complex
216 collisional processes produced deformation and magmatism throughout the late
217 Neoproterozoic and Early Cambrian in the East African–Antarctic belt and in the
218 Brasiliano belt between the Kalahari and Amazonia cratons. Jacobs & Thomas (2004)
219 suggest dispersal of smaller continental fragments by escape tectonics associated with
220 a Himalayan style and scale mountain range formed in the Mozambique belt. These
221 major orogenies, and their topographical and erosional consequences, are the most
222 probable explanation for the widespread occurrence of detrital zircons of this age span
223 in the subsequent sedimentary record of both East and West Gondwana margins (See
224 also Squire et al., 2006). According to Basei et al. (2005), a narrow band of
225 Neoproterozoic metasedimentary rocks on the Atlantic coast of South America is
226 equivalent to the southwest African sequences formed by erosion of the Kalahari and
227 Namaqua–Natal basement and was left behind on the Cretaceous opening of the South
228 Atlantic Ocean, so that the suture zone resulting from closure of the Adamastor ocean
229 now lies within southeastern Brazil and Uruguay.

230

231 During and subsequent to Late Cambrian times, West Gondwana continued to accrete
232 microcontinents and terrane fragments (e.g., Cawood, 2005; Vaughan and Livermore,
233 2005). The origin of some, such as the Precordillera terrane and its relationship to
234 Laurentia and the Pampia Terrane, continues to be extremely controversial (e.g.,
235 Thomas and Astini, 2003; Finney et al., 2005).

236

237 **4. The oceanic margin of West Gondwana**

238

239 4.1 South America

240

241 In the southern South American sector of the margin, the accretionary orogen model
242 has to take into account widely held ideas of collisional accretion of individual
243 terranes of pre-existing continental crust (Fig. 4). Many of these terranes were first
244 proposed and named by Ramos (1988) and, although many are accepted in general,
245 the essential details of their delineation, composition, and the timing of their accretion
246 to Gondwana continue to be controversial.

247

248 The best known of these is the Precordillera terrane (Astini et al., 1995), often equated
249 with and referred to as Cuyania (Ramos, 1988; 2004). This has an outcrop area at least
250 300 km from north-to-south and less than 100 km in width where the geology is
251 dominated by Cambrian to Middle Ordovician limestones, succeeded unconformably
252 by Silurian–Devonian clastic sediments that pass upwards into typical Gondwana
253 sequence lacustrine deposits and red beds of Carboniferous to Triassic age. Alonso et
254 al. (2008) present structural and sedimentological evidence for the passive margin
255 nature of this sequence. The most significant feature of the limestones is a change
256 from a Cambrian brachiopod and trilobite fauna of Laurentian affinity to a Middle to
257 Late Ordovician Gondwana fauna (Benedetto, 1998; Astini et al., 2004). For many,
258 this supports the idea that the Precordillera terrane was derived from Laurentia, but
259 approached Gondwana during the Early Ordovician, followed by accretion during a
260 Middle Ordovician collision. This idea is supported by a wide range of evidence, e.g.,
261 an Early-to-Middle magmatic arc including both I- and S-type granites developed on
262 the marginal continental crust of Gondwana – the Famatinian arc (Pankhurst et al.,
263 1998a; Pankhurst et al., 2000). Other aspects compatible with this scenario are

264 contemporaneous bentonite ash bands in the Precordillera limestones (Huff et al.,
265 1998; Fanning et al., 2004), and palaeomagnetic data (Rapalini, 2005). Middle
266 Ordovician metamorphism has been found in rocks east of the Precordillera (Casquet
267 et al., 2001; Vujovich et al., 2004) and equated with the collision stage, and Castro de
268 Machuca et al. (2008) ascribe an Early Silurian age to major post-collisional shear
269 zones. This is also the interpretation given in Chernicoff et al. (2007) who have
270 studied detrital zircon in a Late Ordovician–Devonian sedimentary sequence which
271 they regard as deposited in a post-collisional foreland basin. However, others (e.g.,
272 Aceñolaza et al., 2002) have proposed an alternative origin for the Precordillera
273 terrane in another part of West Gondwana, with Ordovician emplacement by massive
274 strike-slip movement along the margin. Attempts to resolve these opposing
275 hypotheses for the origin source of the Precordillera terrane continue without final
276 agreement, largely based on the patterns of detrital zircon provenance ages
277 determined by U–Pb geochronology (Thomas and Astini, 2003; Finney et al., 2005).

278

279 Another aspect of the Precordillera terrane hypothesis is the nature and origin of its
280 underlying crustal basement. Unfortunately, this is not unambiguously exposed. There
281 is indirect indication for it consisting of a high-grade metamorphic complex of
282 ‘Grenvillian’ age through the occurrence of ~1000 Ma amphibolite xenoliths brought
283 up in a Miocene dacite through the easternmost limestone outcrops (Kay et al., 1996).
284 High-grade rocks of 1200–1000 Ma have since been discovered throughout the
285 Western Sierras Pampeanas sequences to the east of the Precordillera (McDonough et
286 al., 1993; Varela et al., 1996; Pankhurst and Rapela, 1998; Casquet et al., 2001; 2005;
287 2006). Ordovician limestones are associated with high-grade granite gneiss of
288 ‘Grenville’ age as far south as Ponon Trehue (Fig. 4, Heredia, 2002; Cingolani et al.,

289 2005) and ‘Grenville’-age tonalites at Las Matras (Sato et al., 2000). Initially these
290 occurrences were mostly considered to be representative of the middle crustal
291 basement of the Precordillera terrane, consistent with a Laurentian origin, but more
292 recently (e.g., Galindo et al., 2004; Casquet et al., 2006) it has been suggested that the
293 ‘Grenville’-age rocks of the Western Sierras Pampeanas could be regarded as
294 autochthonous Gondwana basement during the Ordovician, and Casquet et al. (2006;
295 2007) have interpreted some at least as equivalent to the Arequipa-Antofalla block,
296 normally regarded as unambiguously autochthonous. The true nature of the
297 Precordillera basement thus remains questionable.

298

299 The Eastern Sierras Pampeanas constitute another putative continental terrane
300 accretion event (the Pampia terrane of Ramos, 1988, see Fig. 4). This is a belt of
301 migmatitic gneisses, low-grade metasediments, granites and metabasites which
302 underwent orogenic deformation, metamorphism and anatexis in Early-to-Middle
303 Cambrian times (Rapela et al., 1998a; Rapela et al., 1998b; Rapela et al., 2002),
304 although Guerreschi and Martino (2008) suggest that an even older migmatization
305 event may also have occurred. Their Early Palaeozoic history is thus incompatible
306 with the Palaeoproterozoic Río de la Plata craton to the east and the passive margin
307 limestones of the Precordillera sequence to the west, suggesting an exotic terrane. The
308 predominant Nd model age signature of these rocks is a Mesoproterozoic one (as is
309 that of the Famatinian rocks to the west). For this reason, Rapela et al. (1998b)
310 followed previous authors in thinking that the metasedimentary component must have
311 been derived from such a source to the east as a foreland sequence above an eastward
312 dipping subduction zone; however, no Mesoproterozoic source is exposed. They
313 suggested that the terrane was not allochthonous but had previously been rifted-off

314 from a similar position on the Gondwana margin in Neoproterozoic times, and was
315 similar to the Arequipa–Antofalla blocks of northern Chile and Peru. Simpson et al.
316 (2003) and Schwartz and Gromet (2004) proposed subduction of a spreading ridge in
317 Middle Cambrian times as an alternative to collision of a continental block. As a
318 recent development based on detrital zircon U–Pb and whole-rock Sm–Nd data,
319 Escayola et al. (2007) have proposed a radical model in which subduction towards the
320 west occurred in Neoproterozoic times, with sediments being deposited in a back-arc
321 basin from both the Grenville-age Western Sierras Pampeanas and the arc itself rather
322 than from the Río de la Plata craton to the east. The high-grade metamorphism of the
323 Pampean belt followed Early Cambrian closure of the back-arc basin. This could
324 explain the metabasites (as basin floor remnants) but there is no evidence for the arc
325 itself. The problem of the Pampean orogeny is ripe for new data to resolve these and
326 possibly other alternatives, and Rapela et al. (in press) present new evidence on the
327 extent of the craton, the origin of the Pampean belt metasedimentary rocks and the
328 Cambrian tectonic events leading to their juxtaposition.

329

330 The latest collisional event proposed by Ramos (1988) for the central part of this
331 sector is that of the hypothetical Chilenia terrane (Fig. 4). This is supposed to have
332 occurred in Devonian time, and was principally invoked in order to explain granite
333 magmatism of this age that occurs both within the Pampean belt and to the south. A
334 major unit in the former category is the Achala batholith in the southern Sierras de
335 Córdoba. This consists of evolved S-type granites (some with high U contents), of
336 generally post-orogenic characteristics (Lira and Kirschbaum, 1990). Geuna et al.
337 (2007) present palaeomagnetic data that support rapid cooling soon after
338 crystallization.

339

340 Finally, moving south to Patagonia, we arrive at a situation that has been a long-lived
341 puzzle. The source of the problem is the ?Cambrian to Permian Gondwanide
342 sedimentary sequence that forms a Late Permian fold and thrust belt in the Sierra de la
343 Ventana (*aka* Sierras Australes) of southernmost Buenos Aires province, Argentina
344 (Fig. 4). As emphasized by du Toit (1937), this has an obvious continuation in the
345 Cape Fold Belt of South Africa and the Ellsworth Mountains sequence of West
346 Antarctica – all of these must have been joined together as a single stratigraphical and
347 tectonic system during the Late Palaeozoic evolution of Gondwana.

348

349 Ramos (1984; 1986) proposed that an allochthonous (exotic) Patagonian terrane
350 collided with cratonic South America (supercontinental Gondwana) along the Río
351 Colorado zone (Fig. 4) in Carboniferous times. This was thought to have resulted
352 from southwest-dipping subduction beneath the North Patagonian Massif. Devonian–
353 Carboniferous penetrative deformation, southward-verging folds and southward-
354 directed thrusting of supracrustal rocks of the northeastern North Patagonian Massif
355 was described by Chernicoff and Caminos (1996) and elaborated in a detailed
356 structural study by von Gosen (2003), who argued for Permian rather than
357 Carboniferous crustal shortening, and possibly a northeastward-directed accretionary
358 process.

359

360 A major revision of the original collision model for Patagonia has been proposed by
361 Pankhurst et al. (2006). They claim that the majority of rocks in the North Patagonian
362 Massif are autochthonous to Gondwana. The basement to the immediate south of the
363 Sierra de la Ventana itself includes Late Neoproterozoic and Cambrian granites and

364 volcanic rocks of a similar age to those of the Pampean orogeny, albeit in a different
365 tectonic setting, and the northeastern part of the North Patagonian Massif has
366 Ordovician granite magmatism and metamorphism equivalent to the Famatinian
367 orogeny. There is no evidence of a Grenville-age belt similar to the Western Sierras
368 Pampeanas, but this could possibly be hidden beneath the deep Mesozoic and younger
369 sediments of the Río Colorado basin. Thus any collision must have occurred to the
370 south of this massif with its deformed Cambro-Ordovician cover. The discovery of
371 Early Carboniferous subduction-related magmatism followed by mid-Carboniferous
372 S-type granites in a belt that runs southeastwards from the western margin of the
373 North Patagonian Massif led to the proposal that this was essentially the zone of
374 collision, and that the distinctive crustal complexes of the Deseado Massif to the south
375 represents part of the colliding terrane (Pankhurst et al., 2006). The pre-Jurassic
376 geology of the Deseado Massif is very poorly exposed, but it includes Late
377 Neoproterozoic sedimentation, Cambrian plutonism, and both Silurian and Devonian
378 granite magmatism (Pankhurst et al., 2003).

379

380 Another prominent feature of the Palaeozoic geology of southern South America is
381 the enormously voluminous and extensive eruption of Permian and Triassic rhyolitic
382 rocks and the emplacement of associated granites (ca 290–220 Ma) – the Choiyoi
383 complex (Kay et al., 1989; Mpodozis and Kay, 1990). These are so far most closely
384 controlled in terms of their chronology in Patagonia, where they have a wide range
385 ages and isotopic characteristics. Initiation in Early Permian times was ascribed by
386 Pankhurst et al. (2006) to post-collisional break-off of the down-going slab, perhaps
387 with delamination of the crust beneath the North Patagonian Massif, leading to large-
388 scale access of heat to the middle crust. It was suggested that this could have lead to

389 promulgation of the slab break-off towards the north along the Gondwana margin,
390 where the magmatism of the Permo-Triassic Choiyoi Group may be more closely
391 related to east-directed subduction than to collision.

392

393 Some of the youngest rocks in this sector of West Gondwana are the accretionary
394 complexes forming the farthest outboard part of the margin (e.g., Vaughan and
395 Storey, 2000; Hervé and Fanning, 2003; Mortimer, 2004; Glen, 2005). These largely
396 formed after Gondwana was assembled and are semi-continuous from southern South
397 America to eastern Australia, ranging in age from Carboniferous to Cretaceous.
398 Detrital zircon studies show that the material within these complexes are of
399 Gondwanan origin (Hervé et al., 2003; Augustsson et al., 2006). Sepúlveda et al.
400 (2008) show that a relatively recent example, the Madre de Dios terrane (Fig. 4),
401 contains evidence of a Late Carboniferous–Early Permian mid-ocean ridge origin. The
402 terrane was accreted to the Gondwana margin during deformation in Late Triassic–
403 Early Jurassic times, called the Chonide orogeny (Hervé et al., 2003; Sepúlveda et al.,
404 2008) in Patagonia, but which was part of a global event (Vaughan and Livermore,
405 2005).

406

407 4.2 South Africa

408

409 The Cape Fold Belt of South Africa (e.g., Johnston, 2000) (including the Falkland
410 Islands block (Mitchell et al., 1986; Storey et al., 1999)), together with the Sierra de la
411 Ventana of eastern Argentina (e.g., Rapela et al., 2003) and the Ellsworth Mountains
412 of West Antarctica (e.g., Curtis, 2001), forms the central part of the margin of West
413 Gondwana. The basement consists of the 2000-1000 Ma metamorphic volcano-

414 sedimentary rocks of the Namaqua-Natal belt (e.g., Dewey et al., 2006; Eglington,
415 2006; McCourt et al., 2006), which was deformed during late Neoproterozoic to early
416 Palaeozoic Gondwana amalgamation (e.g., Jacobs et al., 2003). The Phanerozoic
417 continental margin sedimentary succession is represented by the 6–10 km thick,
418 siliciclastic Cape Supergroup (Broquet, 1992; Barnett et al., 1997) and subsequent
419 glacial, marine and terrestrial-fluvial successions of the Karoo Supergroup, which
420 includes the Dwyka, Ecca, Beaufort and Stormberg lithostratigraphic units
421 (Catuneanu et al., 2005). The sedimentary succession ranges in age from
422 Neoproterozoic to mid-Jurassic, terminated by basin-wide basaltic volcanism of the
423 Karoo Igneous Province (e.g., Duncan et al., 1997). This sector of the margin appears
424 to represent a passive margin during the Early Palaeozoic (Shone and Booth, 2005),
425 with the accumulation of predominantly reworked continental sedimentary deposits
426 (the ‘Samfrau Geosyncline’ (Du Toit, 1937)). It was deformed by the Gondwanide
427 Orogeny in the Late Permian-Early Triassic (e.g., Johnston, 2000). This major fold
428 belt is often modelled as an intraplate orogen representing far-field-deformation
429 related to distant subduction (e.g., Johnston, 2000), although Dalziel et al. (2000)
430 suggested that flattening of the subduction zone could have been driven by interaction
431 with mantle plume that was subsequently responsible for continental break-up.
432 However, a recent re-evaluation by Pankhurst et al. (2006), using data from the South
433 American, Sierra de la Ventana section of the fold belt, supports a possible collisional
434 origin.

435

436 4.3 West Antarctica

437

438 West Antarctica was originally split into four (Dalziel and Elliot, 1982), or five
439 (Storey et al., 1988), tectonic blocks. The innermost of these is the Ellsworth-
440 Whitmore mountains block, which has sedimentological affinities to the Cape Fold
441 Belt of South Africa (Curtis et al., 1999; Curtis, 2001). It preserves a passive margin
442 volcano-sedimentary succession that ranges from the Cambrian to the Permo-Triassic
443 and may have been derived from the Natal embayment (Randall and Mac Niocaill,
444 2004).

445

446 Recent reassessments of the large-scale structure of West Antarctica suggests that the
447 remaining blocks of West Antarctica can be subdivided into at least three main terrane
448 belts that appear to be continuous from the New Zealand sector of East Gondwana to
449 the Antarctic Peninsula (Pankhurst et al., 1998b; Vaughan and Storey, 2000). The
450 innermost and oldest of these is termed the Ross province in West Antarctica and
451 called the Eastern Domain in the Antarctic Peninsula (Vaughan and Storey, 2000).

452 The Hf-isotope composition of inherited zircons in Late Palaeozoic–Mesozoic
453 granites, migmatites and paragneisses from the Antarctic Peninsula show that they are
454 derived from Mesoproterozoic sources and have been taken to suggest that this
455 domain is underlain by crust of that age (e.g., Flowerdew et al., 2006). The oldest
456 rocks of this Palaeozoic ocean-marginal domain are the Ordovician turbidite
457 sequences of the Swanson Formation of Marie Byrd Land (Pankhurst et al., 1998b).

458 These have no equivalents elsewhere in West Antarctica although turbidites of similar
459 age are seen in the Robertson Bay terrane of Victoria Land in East Gondwana (Stump,
460 1995). These are intruded by the Ford Granodiorite in Marie Byrd Land, which are
461 equivalent in age to the older granitoids from Target Hill in the northern Antarctic
462 Peninsula (Millar et al., 2002). A suite of granitoids emplaced between 340 and 320

463 million years ago (Pankhurst et al., 1998b) are widely developed in Marie Byrd Land
464 and are also seen at Target Hill in the northern Antarctic Peninsula (Millar et al.,
465 2002). Although not developed in Marie Byrd Land, the Eastern Domain in the
466 Antarctic Peninsula contains a sequence of Middle Jurassic Gondwana break-up
467 rhyolite volcanic rocks, the Ellsworth Land Volcanic Group (Hunter et al., 2006b),
468 and an Early Jurassic to Cretaceous (Willan and Hunter, 2005; Hunter et al., 2006a)
469 sequence of deep and shallow marine clastic sedimentary rocks called the Latady
470 Group (Laudon et al., 1983; Hunter and Cantrill, 2006). The latest event seen in this
471 domain is the mid-Cretaceous emplacement of arc plutons of the voluminous Lassiter
472 Coast Intrusive Suite (e.g., Flowerdew et al., 2005).

473

474 Outboard of the Ross Province/Eastern Domain is a series of magmatic arc terranes
475 termed the Amundsen Province in Marie Byrd Land (Pankhurst et al., 1998b) and the
476 Central Domain in the Antarctic Peninsula (Vaughan and Storey, 2000). The
477 Amundsen Province and Central Domain are largely magmatic and show many
478 similarities in compositional types and in timing of magmatic emplacement (Vaughan
479 and Storey, 2000). Plutonism appears to have peaked in three discrete episodes in the
480 Late Triassic, mid-Jurassic, and Late Jurassic to Early Cretaceous (Leat et al., 1995;
481 Vaughan and Storey, 2000). Recent geophysical data from the Antarctic Peninsula
482 suggest that the Central Domain is composite and made up of smaller terranes
483 (Ferraccioli et al., 2006). So far, a mafic eastern Central Domain and a granitic
484 western Central Domain have been identified (Ferraccioli et al., 2006). Major
485 deformational episodes affected the Central Domain in Late Triassic-early Jurassic
486 and mid-Cretaceous times (Vaughan et al., 2002a; Vaughan et al., 2002b; Vaughan
487 and Livermore, 2005).

488

489 The outermost of the West Antarctic terrane belts is termed the Western Domain in
490 the Antarctic Peninsula (Vaughan and Storey, 2000). It has no equivalent in Marie
491 Byrd Land although similar accretionary complex terranes are developed in New
492 Zealand and in southern South America (Vaughan and Storey, 2000). Accretionary
493 complex rocks range in age from Late Carboniferous (Kelly et al., 2001) to Late
494 Cretaceous (Vaughan and Storey, 2000). The Western Domain in the Antarctic
495 Peninsula was affected by deformation in the Late Triassic-early Jurassic and in the
496 mid-Cretaceous (Vaughan and Livermore, 2005).

497

498 **5. Adjacent parts of the oceanic margin of East Gondwana**

499

500 5.1 New Zealand

501

502 The New Zealand sector of the eastern Gondwana margin (e.g. Mortimer, 2004;
503 Wandres and Bradshaw, 2005) is made up of a collage of terranes, composed of
504 basement rocks ranging in age from early Cambrian to late Early Cretaceous. These
505 can be grouped into three provinces, the Western Province, the Median Province, and
506 the Eastern Province (Coombs et al., 1976; Bishop et al., 1985; Bradshaw, 1989). The
507 Western Province is made up of two terranes that formed the Palaeozoic margin of
508 East Gondwana and largely consist of lower Palaeozoic metasedimentary rocks cut by
509 series of Devonian, Carboniferous and Early Cretaceous granite plutons (e.g., Cooper,
510 1989; Muir et al., 1996; Waight et al., 1998). In addition there are some minor
511 volcanic and metamorphic rocks of Cambrian age (e.g., Munker and Crawford, 2000).
512 The Median Province is largely magmatic and consists of suites of Carboniferous to

513 Early Cretaceous subduction-related arc plutons with subordinate volcanic and
514 sedimentary rocks (e.g., Muir et al., 1998; Mortimer et al., 1999). The Eastern
515 Province (e.g., Mortimer, 2004; Wandres and Bradshaw, 2005) consists of arc, fore-
516 arc and accretionary complex rocks that formed and accumulated during Permian to
517 Cretaceous plate convergence and subduction. These have been subdivided into up to
518 13 terranes, several of which are grouped into a Torlesse Superterrane (Campbell,
519 2000). As pointed out by Wandres and Bradshaw (2005) the bulk of New Zealand
520 continental crust is submerged by the sea. Adams (2008) examines the terrane
521 evidence from this hidden area by studying Rb-Sr metamorphic and U-Pb detrital
522 zircon ages from the emergent island parts of the submerged continental crust, called
523 "Zealandia". The data show that the Campbell Plateau segment of Zealandia has clear
524 affinities with the Western Province/Ross Province and the Median
525 Province/Amundsen Province, with little evidence for extension of the Eastern
526 Province.

527

528 5.2 Victoria Land and the Transantarctic Mountains

529

530 Although strictly part of East Gondwana, the Transantarctic Mountains are important
531 because they both acted as a source for sediments deposited in West Gondwana,
532 particularly in West Antarctica (e.g., Flowerdew et al., 2006), and were themselves a
533 sedimentary sink for sediments derived from West Gondwana in Late Palaeozoic and
534 Early Mesozoic times (e.g., Elliot and Fanning, 2007). At their most northerly extent,
535 in Northern Victoria Land, the Transantarctic Mountains are composed of Cambrian
536 and Ordovician terranes amalgamated during the Ross Orogeny (recently reviewed by
537 Tessensohn and Henjes-Kunst, 2005). The main part of the Transantarctic Mountains

538 is underlain by Neoproterozoic, and possibly older (e.g., Fanning et al., 1996;
539 Fitzsimons, 2003), basement, intruded by granitoid plutons of the Ross Orogeny
540 (Stump, 1995). This is unconformably overlain by the quartzose sandstones of the
541 Devonian Taylor Group (Isbell, 1999). The Taylor Group was deformed by the end-
542 Palaeozoic Gondwanan orogeny (Cawood, 2005) and is in turn unconformably
543 overlain by the Permo-Triassic glacial, marine, terrestrial and fluvial sedimentary
544 rocks of the Victoria Group (Collinson et al., 1994). This upper sedimentary
545 sequence was intruded in the Lower Jurassic by sills and dikes of Ferrar Dolerite (e.g.,
546 Hergt et al., 1991) with co-magmatic overlying basaltic pyroclastic rocks (e.g., Elliot
547 and Hanson, 2001) and Kirkpatrick Basalt flood lavas (e.g., Elliot et al., 1999).

548

549 **6. Concluding remarks**

550

551 The longevity and extent of the Gondwana margin has ensured that it has remained
552 the subject of intense study for over seventy years. It was one of the birthplaces of
553 terrane theory (e.g., Vaughan et al., 2005b) and it continues to be a proving ground for
554 theories of supercontinental amalgamation (e.g., Cawood, 2005) and break-up (e.g.,
555 Rapela et al., 2005; Veevers, 2005; Willan and Hunter, 2005).

556

557 An interesting question is the one of translation of terranes along the Gondwana
558 margin. Cawood et al.(2002) have shown evidence for translations of thousands of
559 kilometres along the Gondwana margin from the Permian to the Cretaceous, and this
560 idea has been inherent in some treatments of the older Palaeozoic tectonics. Structural
561 evidence suggests that large scale strike-slip faults exist (e.g. Vaughan and Storey,
562 2000). Some support for large-scale translation can be derived from zircon data

563 although the only way that these movements can be confirmed or quantified is by
564 multidisciplinary studies that include palaeomagnetic analysis and interpretation.

565

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569 **References**

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1058
 1059 Figure Captions

1060
 1061 Figure 1: Gondwana reconstruction after Unrug (1997) showing major terrane belts on
 1062 the margins of the supercontinent: NZ: New Zealand; TAM: Transantarctic
 1063 Mountains. Boundary zone between East and West Gondwana after Unrug
 1064 (1997) shown as overlay: ANS: Arabian–Nubian Shield; N–N–M: Namaqua–
 1065 Natal–Maud belt.

1066
 1067 Figure 2: Gondwana reconstruction after Du Toit (1937), showing earliest subdivision
 1068 of the supercontinent into eastern and western parts.

1069
 1070 Figure 3: Reconstruction of West Gondwana after Tohver et al.(2006) showing
 1071 cratonic and Brasiliano–Panafrican elements. Cratons shown in light grey:
 1072 Am, Amazonia; ANS, Arabian–Nubian Shield, C, Congo; GM, Goiás Massif;
 1073 K-G, Kalahari–Grunehogna; LA, Luis Alves, P, Paraná, RA, Río Apa, SF, São
 1074 Francisco; SL, São Lius; WA, West Africa. Brasiliano–Panafrican belts
 1075 (ringed): Ac, Araçuaí; Ag, Araguaia; Bo, Borborema; Br, Brasília; Da,
 1076 Damara; DF, Dom Feliciano; Dh/O, Dahomeides/Oubangides; G, Gariep; H,
 1077 Hoggar; Ka, Kaoko; K/Z, Katangan/Zambezi; LA, Lufilian Arc; M,
 1078 Mozambique; P, Paraguai; R/M, Ribeira/Mantequeira; Ro, Rockelides; Ta,
 1079 Tanzania; Tu, Tucavaca; WC, West Congo.

1080
 1081 Figure 4: Schematic representation of the tectonic elements of the margin of West
 1082 Gondwana, extensively modified after Rapalini (2005) and references therein,
 1083 using further information from Pankhurst et al. (2006) and personal
 1084 communications from C.W. Rapela and C. Casquet. Amazonia, Río Apa, Río
 1085 de la Plata (and in some schemes, Arequipa and Antofalla) are the cratonic
 1086 blocks of Palaeoproterozoic to Neoproterozoic age. The Pampean belt (which
 1087 encompasses the Eastern Sierras Pampeanas Pampia terrane of Ramos (1988),
 1088 is shown as continuous with the Araguaia belt of Brazil, following Trindade et
 1089 al. (2006), and the approximate form of the Patagonian plate is from Pankhurst
 1090 et al. (2006). The known extent of Grenville-age belts of Sunsas (S) and the
 1091 Western Sierras Pampeanas (W) is indicated, although the latter also occurs
 1092 beneath the Argentine Precordillera (Cy), as either stratigraphical or tectonic
 1093 basement. The Ordovician Famatinian orogenic belt (F) overprints the earlier
 1094 complexes, including those of the Antofalla block, where Lucassen et al.

1095 (2000) recognise Pampean metamorphism and magmatism as reflecting an
1096 accretionary orogeny.

1097

1098 **Alan P.M. Vaughan** is a Principal Investigator and head of the Palaeoenvironments
1099 Group at the British Antarctic survey, where he has worked for the past 16
1100 years. His work has taken to him to Antarctica six times. He is Earth sciences
1101 editor of Antarctic science and UK representative on the Gondwana
1102 Committee. He graduated from Trinity College Dublin in 1985 with a gold
1103 medal in natural sciences and he
1104 received his PhD from that institution
1105 in 1991. His main interests are on the
1106 long-term evolution of the Earth both
1107 tectonically and
1108 palaeoenvironmentally, with a
1109 particular focus on the influence of
1110 large scale tectonic and magmatic
1111 processes on global
1112 palaeoenvironments. He has also
1113 worked extensively on the tectonic
1114 evolution of the Antarctic sector of
1115 West Gondwana. He has collaborated
1116 with geoscientists in New Zealand,
1117 Brazil, Spain, Chile and the USA, and
1118 has been involved in several published
1119 papers and books.



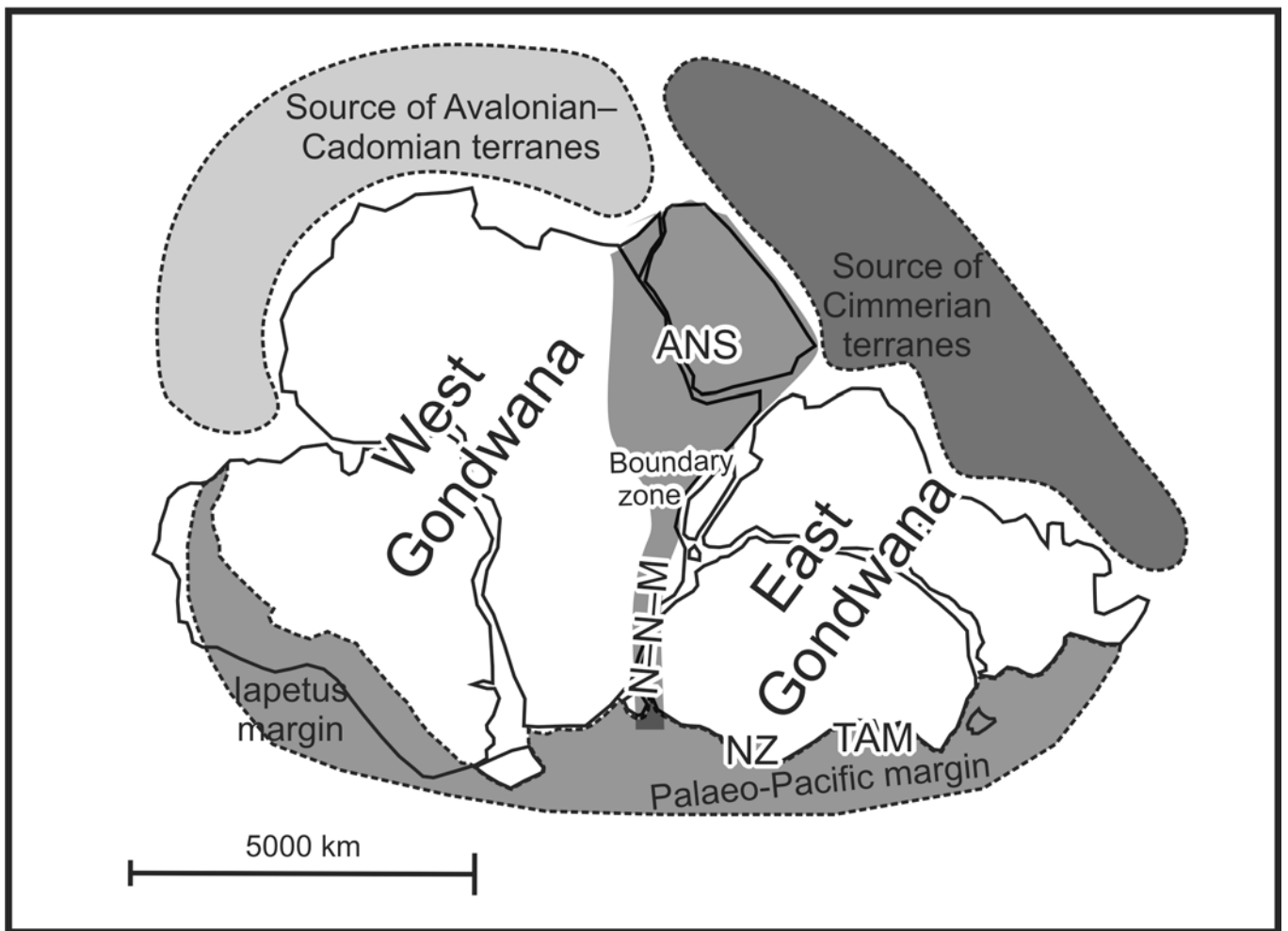
1120

1121 **Robert (Bob) Pankhurst** is Visiting Research Associate at the British Geological
1122 Survey, where he worked for 26 years in the NERC Isotope Geosciences
1123 Laboratory carrying out geochronological and isotope research on behalf of
1124 the British Antarctic Survey (BAS), during which he undertook fieldwork
1125 extensively in Antarctica (nine summer field seasons) and southern South
1126 America, concentrating on the latter since official retirement in 2002. He was
1127 awarded the Polar Medal in 1987 and has been elected corresponding member
1128 of both the Chilean and
1129 Argentine Academies of
1130 Science, as well as of the
1131 Argentine Geological
1132 Association. He is Chief
1133 Books Editor for the
1134 Geological Society, London,
1135 and Associate Editor for the
1136 Journal of South American
1137 Earth Sciences. He graduated
1138 from the University of
1139 Cambridge (B.A. 1964, M.A.
1140 1967) where he also holds the
1141 title of Doctor of Science
1142 (Sc.D. 1998). He received a
1143 Diploma in Geochemistry
1144 (1965) and then a D.Phil.

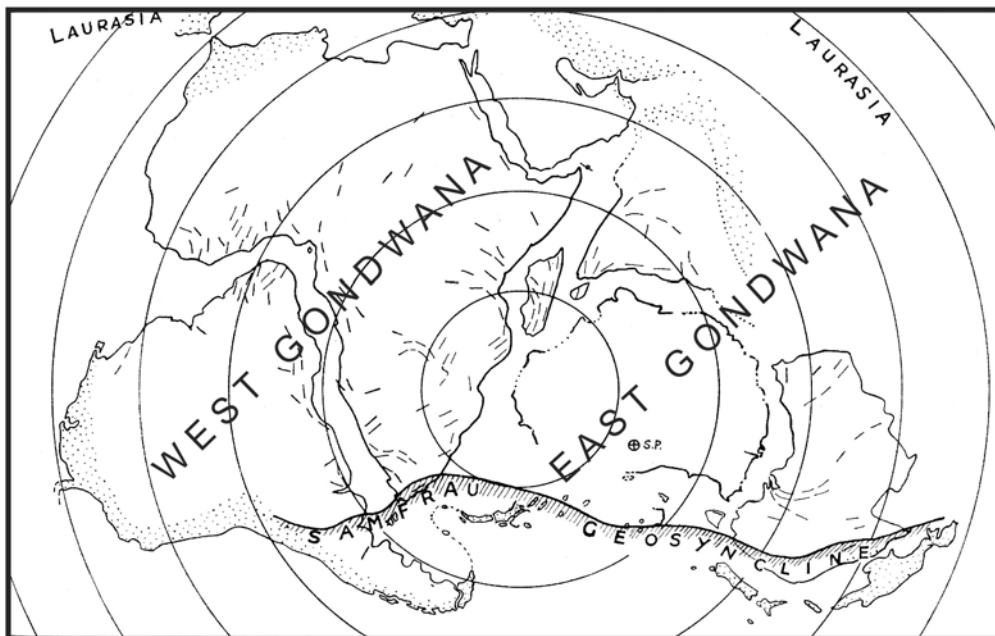


1145 (1968) from the University of Oxford, where he stayed as Research Fellow
1146 working on projects in Scotland, West Greenland and Iceland before joining
1147 BAS in 1976. His main interests are in isotope dating and geochemistry
1148 applied to igneous petrogenesis, metamorphism and sediment provenance in
1149 relation to the evolution of the continental crust of West Gondwana, and hence
1150 the tectonic processes involved. He has actively collaborated with
1151 geoscientists in Argentina, Australia, Brazil, Chile, New Zealand, Spain, and
1152 the USA, and has been involved in numerous published papers and several
1153 books.

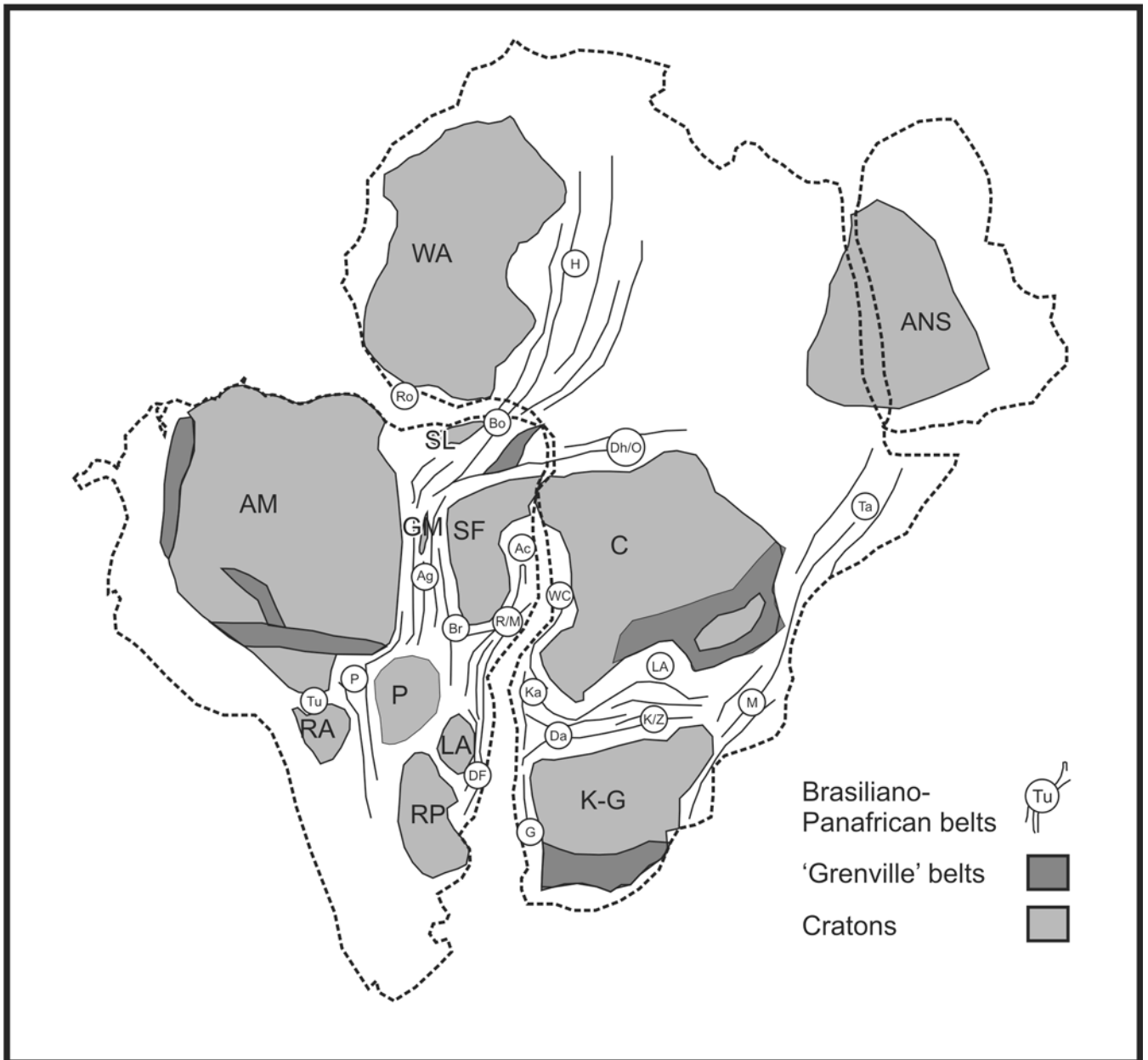
Vaughan & Pankhurst Figure 1



Vaughan & Pankhurst Figure 2



Vaughan & Pankhurst Figure 3



Vaughan & Pankhurst Fig. 4

