

1 **A seismic tomography, gravity, and flexure study of the crust and upper mantle**
2 **structure across the Hawaiian Ridge, Part 2 Ka'ena**

3 R. A. Dunn^{1,*}, A. B. Watts², C. Xu^{2,3}, and D. J. Shillington⁴

4 ¹Department of Earth Sciences, School of Ocean and Earth Science and Technology, University
5 of Hawaii at Manoa, Honolulu, HI, USA.

6 ²Department of Earth Sciences, University of Oxford, Oxford, UK.

7 ³Key Lab of Submarine Geosciences and Prospecting Techniques, Ministry of Education, and
8 College of Marine Geosciences, Ocean University of China, Qingdao, China.

9 ⁴School of Earth and Sustainability, Northern Arizona University, Flagstaff, AZ, USA.

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12 *Corresponding author. ORCID: orcid.org/0000-0003-1434-5512

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15 **Key words:** Hawaiian Ridge; Volcano structure; Plate loading; Lithospheric flexure; Seismic
16 tomography; Marine gravity

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18 **Key points:**

19 The Hawaiian Ridge at Ka'ena consists of volcanic edifices ~7 km thick, sitting on Pacific
20 oceanic crust ~6 km thick.

21 Seismic and gravity data reveal an inner volcanic core draped by layers of lava flows and
22 volcanic debris.

23 Broad lithospheric flexure (peak deflection ~3.5 km) suggests a flexural rigidity of the 90 Myr
24 old lithosphere equivalent to $T_e = \sim 25.6$ km.

25

26 **Abstract.** The Hawaiian Ridge, a classic intraplate volcanic chain in the Central Pacific Ocean,
27 has long attracted researchers due to its origin, eruption patterns, and impact on lithospheric
28 deformation. Thought to arise from pressure-release melting within a mantle plume, its mass-
29 induced deformation of Earth's surface depends on load distribution and lithospheric properties,
30 including elastic thickness (T_e). To investigate these features, a marine geophysical campaign
31 was carried out across the Hawaiian Ridge in 2018. Westward of the island of O‘ahu, a seismic
32 tomographic image, validated by gravity data, reveals a large mass of volcanic material
33 emplaced on the oceanic crust, flanked by an apron of volcanoclastic material filling the moat
34 created by plate flexure. The ridge adds ~7 km of material to pre-existing ~6-km-thick oceanic
35 crust. A high-velocity and high-density core resides within the volcanic edifice, draped by
36 alternating lava flows and mass wasting material. Beneath the edifice, upper mantle velocities
37 are slightly higher than that of the surrounding mantle, and there is no evidence of extensive
38 magmatic underplating of the crust. There is ~3.5 km of downward deflection of the sediment-
39 crust and crust-mantle boundaries due to flexure in response to the volcanic load. At Ka‘ena
40 Ridge, the volcanic edifice's height and cross-sectional area are no more than half as large as
41 those determined at Hawai‘i Island. Together, these studies confirm that volcanic loads to the
42 west of Hawai‘i are largely compensated by flexure. Comparisons to the Emperor Seamount
43 Chain confirm the Hawaiian Ridge's relatively stronger lithospheric rigidity.

44

45 **Plain Language Summary**

46 The Hawaiian Ridge, a seafloor volcanic chain in the Pacific, has long fascinated scientists for its
47 origin, eruptions, and impact on Earth's surface layers. It is believed to form from melted rock
48 originating in the Earth's mantle. As lava builds up each volcanic edifice, it presses down on the
49 Earth's surface, deforming it by a few kilometers vertically, with the degree of deformation
50 depending on its weight and the properties of the Earth's outer layers. Our geophysical study has
51 imaged the interior of the volcanic ridge and the underlying crust and mantle. The volcano
52 contains an interior core of highly crystalline rocks draped by volcanic material, and contributes
53 to a broad lithospheric flexural response of up to ~3.5 km, vertically. This study sheds light on
54 volcano formation, load-induced lithospheric flexure, and their implications for understanding
55 Earth's subsurface processes.

56 **1. Introduction**

57 Located in the Central Pacific Ocean, the Hawaiian Ridge is one of the most prominent
58 examples of an intraplate volcanic chain on the surface of the Earth. Long believed to have been
59 created by pressure-release melting within a rising mantle plume (e.g., Morgan, 1971; Wilson,
60 1963), the large mass of the Hawaiian Ridge has inspired many studies on the manner in which
61 the Earth's surface and upper layers yield to long-term surface loads. As each successive volcano
62 in the Hawaiian Ridge grows, its large mass deforms the Earth's surface, resulting in a deflection
63 or flexure of the lithosphere (e.g., Gunn, 1943; Walcott, 1970; Watts, 2001; Watts & Cochran,
64 1974; Watts & ten Brink, 1989; Wessel, 1993), the study of which can provide significant
65 information about the stress state, flexural rigidity and rheological properties of the oceanic
66 lithosphere (Calmant, 1987; Watts & Cochran, 1974). The degree of deformation is expected to
67 be controlled, in part, by the mass distribution of the load, and the effective elastic thickness (T_e)
68 of the lithosphere.

69 Along the Hawaiian Ridge, the volcanic load distribution is not well understood. Seismic,
70 gravity, petrologic, and other studies reveal that the ridge is composed of a series of massive
71 often-overlapping shield volcanoes, some of which appear to have cumulate cores. Early gravity
72 and seismic experiments found evidence for high-density or high-wave-speed bodies within the
73 islands of O'ahu and Hawai'i (e.g., Adams & Furumoto, 1965; Strange et al., 1965a; 1965b;
74 Woollard, 1951; Zucca et al., 1982) that were suggested to be each the result of up to a few
75 million years of accumulated mafic and ultramafic intrusive materials. More recently, an updated
76 residual gravity anomaly map of the Hawaiian Ridge (Flinders et al., 2013) and 2-D and 3-D
77 seismic tomographic imaging studies across the island of Hawai'i and its submarine flanks
78 clearly show that the summits and upper rift zones of volcanoes are characterized by high-
79 density and high-velocity materials (Lin et al., 2014; MacGregor et al., 2023; Park et al., 2007;
80 Park et al., 2009). Cumulate core features have been found by previous studies at other
81 seamounts as well, e.g., Jimmu guyot on the Emperor Seamounts (Watts et al., 2021; Xu et al.,
82 2022), Jasper seamount (Hammer et al., 1994), Louisville guyot (Contreras-Reyes et al., 2010),
83 Great Meteor seamount (Weigel & Grevemeyer, 1999) and La Reunion Island (Gallart et al.,
84 1999). Furthermore, not all of the added material may have been emplaced above the oceanic
85 crust. A seismic study by Watts et al. (1985) along the Hawaiian Ridge indicated the possibility

86 that magmatic material had been underplated to the bottom of the oceanic crust, setting off years
87 of investigation into the degree of underplating beneath seamounts around the world (e.g., Ali et
88 al., 2003; Caress et al., 1995; Dañobeitia and Canales, 2000 ; Leahy et al., 2010; Weigel and
89 Grevemeyer, 1999). Where present, it may have important implications for load and buoyancy
90 forces that control flexure (e.g., Wolfe et al., 1994).

91 The degree of deformation of the lithosphere is difficult to ascertain directly due to sediment
92 cover and the presence of the ridge itself. However, flexure is readily apparent in a regional
93 free-air gravity anomaly map (Figure 1), which reveals the Hawaiian Moat, the seafloor
94 depression around the Ridge that is partially filled by sediment. Previous seismic reflection and
95 refraction experiments presented a variable but obvious increase in depth of the sediment-crust
96 and crust-mantle boundary, or Moho, from the arch, beneath the moat, to the ridge near the
97 islands of Maui, O‘ahu and Hawai‘i (e.g., Furumoto & Woollard, 1965; Park et al., 2007; Shor &
98 Pollard, 1964; Zucca et al., 1982). In comparison to the Emperor Seamount Chain, where plate
99 age at the time of loading is relatively younger, studies of lithospheric flexure have suggested the
100 lithosphere underlying the Hawaiian Ridge has a relatively stronger flexural rigidity and hence
101 larger T_e values (Watts, 1978; Watts & Ribe, 1984; Watts et al., 2013). The first seismic
102 experiment directly aimed at investigating flexure along the Hawaiian Ridge, was carried out in
103 1982 aboard the R/V *Robert D. Conrad* and the R/V *Kana Keoki* (Watts et al., 1985). The
104 findings of the experiment revealed that the depth below sea level to the Moho, the boundary
105 between the crust and the mantle, increased from approximately 11 km below sea level beneath
106 the arch to over 14 km below sea level beneath the ridge.

107 The uncertainties in load distribution and corresponding degree and extent of flexure, leave
108 models of mantle properties without substantial constraints. To address this issue, in 2018 we
109 conducted a marine geophysical experiment across the Hawaiian Ridge, as part of a broader
110 study of the seismic structure of the Hawaiian-Emperor Seamount Chain (Boston et al., 2019;
111 Cilli et al., 2023; MacGregor et al., 2023; Watts et al., 2021; Xu et al., 2022). The experiment
112 consisted of two ocean bottom seismograph (OBS) lines for tomographic imaging (Lines 01 and
113 02 in Figure 1) and seven MCS lines, along with corresponding gravity and multibeam
114 bathymetry data collection (e.g., Watts et al, 2020). In this paper, we report results from the
115 seismic and gravity profile located to the western side of the island of O‘ahu (Line 02; Figures 1

116 and 2) across the Ka‘ena ridge. This paper is a companion paper to MacGregor et al., (2023),
117 which analyzed data collected along Line 01 located across the western flank of the island of
118 Hawaii. Our findings reveal that the Ka‘ena Ridge consists of a volcanic edifice of ~7 km thick,
119 which overlays pre-existing Pacific oceanic crust with a thickness of ~6 km. The edifice has a
120 high-velocity and high-density core, consistent with intrusive mafic and/or ultramafic rock
121 overlain by volcanic units and debris flows. The volcanic load produces a broad flexural
122 response of the Pacific lithosphere with a peak vertical deflection of ~3.5 km. The results
123 obtained from seismic tomography and gravity modeling are utilized to estimate the elastic
124 thickness of the Pacific oceanic lithosphere and are compared to the results for Line 01
125 (MacGregor et al., 2023), as well as to results of comparable studies located along the Emperor
126 Seamounts (Watts et al., 2022; Xu et al., 2022).

127 **2. Study area**

128 The oceanic crust that underlies the study area was formed at the Pacific-Farallon Ridge
129 during the Cretaceous Period, approximately 90 Ma, at an estimated spreading rate of roughly 80
130 km/Myr (Seton et al., 2020). Prior seismic studies have found that the crust has a thickness of
131 ~6-6.5 km and a seismic profile typical of oceanic crust formed along fast spreading ridges (e.g.,
132 Ohira et al., 2018; Watts et al., 1985). North of the islands, Ohira et al. (2018) found 9% mantle
133 *P*-wave anisotropy, presumably due to the alignment of olivine crystallographic *a*-axes during
134 seafloor spreading (Hess, 1964). Our seismic line is oriented ~58° from the paleo-spreading
135 direction, or roughly intermediate between the seismically fast and slow directions.

136 The seismic line intersects the Hawaiian Ridge at Ka‘ena Ridge (Figures 1 and 2), a
137 submarine topographic high extending approximately 100 km northwest from O‘ahu towards
138 Kauai. Ka‘ena Ridge, with its expansive spatial area and shallowly-submerged summits, has
139 frequently been proposed as an independent center of volcanism separate from the two volcanoes
140 that comprise the island of O‘ahu: Wai‘anae and Ko‘olau (e.g., Moore et al., 1989). Sinton et al.
141 (2014), who analyzed seafloor samples, bathymetry, and gravity data, present compelling
142 evidence that Ka‘ena Ridge formed as a distinct shield volcano with a significant shield building
143 phase occurring approximately 3.5 - 5 Ma, predating the major shield building stage of Wai‘anae
144 volcano by up to 1 My. The upper sections of the volcanic structure are expected to have
145 emerged above the water surface, reaching heights of approximately 1000 m, before subsiding

146 due to the flexural response of the lithosphere caused by the added weight of Wai‘anae and
147 Ko‘olau volcanoes to the southeast.

148 Cutting across the region is the Molokai Fracture Zone (Figure 1), which is composed of
149 several interacting tectonic strands oriented in an ENE-WSW direction, corresponding to the
150 presumed Cretaceous spreading direction (e.g., Normark et al., 1989). Where observable, the
151 abyssal hill fabric, which forms during the rifting process at the spreading ridge (e.g., Buck &
152 Poliakov, 1998), is perpendicular to the lineament trend. The seafloor to the north of the fracture
153 zone is younger compared to the southern side (e.g., Seton et al., 2020). As the fracture zone
154 nears the Hawaiian Ridge, its morphological features gradually fade as the strands become
155 buried beneath the sediments of the Hawaiian Moat and the Ridge itself, but it is largely
156 observable in the gravity field (Figure 1). Along the southern side of Ka‘ena, the seismic line
157 intersects two splays of the fracture zone, near seismic stations 209 and 217 (Figure 2a).

158 The seismic line intersects submarine landslides on both sides of the Ka‘ena Ridge. On the
159 northern flank, there is a significant landslide deposit known as the Ka‘ena Slide (Figure 2b),
160 which is believed to have formed predominantly from a debris avalanche that scattered blocks of
161 material into the Hawaiian Moat, covering a distance of over 90 km (Moore et al., 1989). The
162 composition of samples collected from the deposit aligns with origins from Ka‘ena Volcano,
163 indicating that the slide occurred during the later stages of the volcanic edifice's growth (Sinton
164 et al., 2014). Based on the difference in elevation between the deposit and the surrounding
165 seafloor, the upper portion, characterized by a more coherent slump feature, generally has a
166 thickness of less than 1 km, while the lower portion is typically less than a few hundred meters
167 thick. The spatial extent of the slide depicted in Figure 2b is derived from Sinton et al. (2014).
168 However, it is possible that this slide or a previous one extends further into the moat, as there
169 seems to be an additional ~100 m of infilling material to the NNE of the defined area.

170 On the southern flank of the ridge is the Wai‘anae Slump, an extensive submarine landslide
171 (Coombs et al., 2004; Hussong et al., 1987; Moore et al., 1989). Detailed seafloor mapping, on-
172 bottom visual observations, and petrologic analyses, indicate that the landslide occurred as a
173 complex series of slumps and smaller debris avalanches (Coombs et al., 2004; Sinton et al.,
174 2014). The thickness of the slump has been estimated to be around 1-2 km (Presley et al., 1997).
175 To the southwest of the Wai‘anae Slump lies a substantial debris field that stretches across the

176 abyssal seafloor, covering an area of at least 2700 km³ (Figure 2b). Coombs et al. (2004)
177 suggested that this debris field likely formed approximately concurrently with the activity of the
178 Wai‘anae Slump. When comparing the depth of the deposit to the surrounding seafloor, its
179 thickness along the seismic line measures up to 500 m. However, this estimation is complicated
180 by the overlapping volcanic features of the Southwest O‘ahu Volcanic Field.

181 The Southwest O‘ahu Volcanic Field (Figure 2b) was mapped in the late 1980s by GLORIA
182 sidescan imagery (e.g., Holcomb & Robinson, 2004) and its volcanics were later investigated by
183 Takahashi et al. (2001). It was found to be composed of numerous cones and thin lava flows. On
184 the basis of seafloor samples and relationships, Coombs et al. (2004) argued that the volcanic
185 field post-dates the debris field. Age estimates for the lava flows span 1.5-3 Ma (Coombs et al.,
186 2004; Noguchi & Nakagawa, 2003), postdating Ka‘ena’s main shield building stage.

187 The northern end of the seismic line intersects portions of the North Arch volcanic field
188 (Figure 2a), which consists of nearly 25,000 km² of ~0.5-1.15 Ma age lava flows (Clague et al.,
189 1990). The seismic line crosses the youngest of the flows (on the basis of backscatter imagery),
190 where the flows are thought to be thin, a few tens of meters thick at most, with only isolated
191 regions of thicker material near vents (Normark et al., 1989). Only a thin layer of sediment (~1 m
192 thick) is expected to cover these flows.

193 **3. Seismic experiment, data, and methods**

194 The seismic data were collected in 2018 using the R/V *Marcus G. Langseth* (cruise
195 MGL1806) and its 36-element 6,600-cubic-inch airgun array. The experiment consisted of two
196 ocean bottom seismograph (OBS) tomography lines and seven multi-channel seismic (MCS)
197 lines (Figure 1). This study analyzes data collected along the western OBS line, a ~550-km-long
198 line centered on the Ka‘ena Ridge (Figure 2). For 2-D refraction tomography, 35 OBS were
199 deployed along the line with an average spacing of ~15 km. Each OBS contained a hydrophone
200 and a 3-component geophone and recorded at a sampling rate of 200 Hz. The line was shot
201 twice, once with airgun pulses triggered every 390 m, which will be referred to as the “OBS
202 shots”, and once with pulses located every 63 m, which will be referred to as the “MCS shots”,
203 since that line was primarily for MCS analysis. Seismic waves recorded by the OBS from both
204 sets of shots are included in the analysis presented here. The MCS shots provide dense coverage
205 at short ranges, and a more limited amount of long-range data that varied by station. A collection

206 of records for each station and the details on the OBS relocation, time base, and travel time
207 picking approach is available in the online Supporting Information accompanying this paper.

208 A variety of seismic *P*-wave phases were identified and used in the analysis. The presence
209 and timing of a particular phase is largely dependent on a station's distance from the Hawaiian
210 Ridge. Most of the off-ridge stations recorded data to 300-350 km range, and *P* waves
211 undershoot the Ka'ena Ridge at crustal and mantle depths. Even seismic phases recorded from
212 the MCS shot line can be identified at distances >200 km for some instruments, since waterborne
213 shot noise is blocked by the ridge, facilitating the observation of long-range arrivals even if
214 short-range data is obscured by prior shot noise. The record in Figure 3 is typical for stations
215 located away from the ridge, showing a large delay in arrivals for shots located across the top of
216 Ka'ena that is presumably due to an increase in total crustal thickness. Figures 4 and 5 display
217 sets of records that emphasize the crustal and mantle phases dominating at shorter ranges in the
218 moat and across the Ka'ena Ridge, respectively. Figure 6 shows a plot of the travel times for all
219 stations and identified seismic phases, color-coded by type. There is very good data coverage and
220 correspondence of phases from station to station, meaning that shot-station pairs align well with
221 corresponding pairs in reversed positions. Figure 7 shows *P_n* travel times plotted with a reduced
222 time axis to illustrate the change in mantle wave speed along the seismic line. The next two
223 sections discuss the seismic phases in more detail.

224 **3.1 Stations located to the north and south of the Hawaiian Ridge**

225 Stations located away from the Ka'ena Ridge (Figures 3 and 4) exhibit typical seismic phases
226 for oceanic crust. The dominant seismic phases are *P*-wave refractions, which, in line with
227 standard practice, are interpreted as occurring in the sediment layer (*P1*), upper oceanic crust
228 (*P2*), lower oceanic crust (*P3*), and mantle (*P_n*). Additionally, reflections or triplications
229 interpreted to occur at the top of the oceanic crust (*P2P*) and crust-mantle interface, or Moho,
230 (*P_mP*) are present. *P2P* is easily recognizable across the moat areas, where the records indicate
231 the sediment is thickest (e.g., Figure 4g), compared to outside the moat areas (e.g., Figure 4h),
232 where sediment appears to be thinner, and identifying a *P2P* arrival can be more challenging.

233 The apparent velocities of phases in the record sections indicate a relatively simple oceanic
234 crustal structure with low-velocity upper layers, a ~7 km/s lower crust, and an upper mantle with
235 a wave speed of about 8 km/s on average. The northern side of the seismic line suggests laterally

236 homogeneous structures, while the area to the south, especially beneath the Southwest O‘ahu
237 Volcanic Field and the southernmost strand of the Moloka‘i Fracture Zone (near stations 208 and
238 209), wave arrival times and amplitude variations suggest that the crustal structure is somewhat
239 anomalous. Some records show clear *PmP* arrivals over large ranges (e.g., station 215 for shots
240 to the south; Figure 4), while others show more limited *PmP* arrivals, but robust *Pn* arrivals (e.g.,
241 stations 208 and 231 for shots to the north; Figures 3 and 4). This suggests that the crust-mantle
242 boundary is more transitional in some locations on the scale of a seismic wavelength, rather than
243 a sharp contact. The strong *Pn* arrivals indicate a positive velocity gradient in the upper mantle.
244 *PmP* was typically picked at ranges <60-80 km to avoid confusions with other waveform energy
245 in the records.

246 There is some evidence for a two-step Moho, primarily associated with the southern part of
247 the line where waves travel beneath the Southwest O‘ahu Volcanic Field. For instance, station
248 208 (Figure 3, inset) demonstrates evidence of a slightly slower *Pn* precursor phase to the
249 primary *Pn* phase. There is also evidence for a few sparse sub-Moho reflection events (e.g.,
250 Figures 4e, 4f, and 4i), indicating the possible existence of individual frozen melt sills. But these
251 do not occur with enough frequency to detectably reduce upper mantle velocities in a general
252 sense. Overall, there is no evidence for a broad low-velocity crust-mantle mixed layer that
253 underplates the oceanic crust, as has been suggested by past active-source (e.g., ten Brink &
254 Brocher, 1987; Watts et al, 1985) or passive-source (receiver function) seismic studies (e.g.,
255 Leahy et al., 2010).

256 Figure 7 shows *Pn* travel times using a reduced time axis to align the arrival times. The
257 reduction velocity that best aligns the refracted arrivals can be taken as an approximation of the
258 velocity for those ray paths, whereas the overall trends in the arrivals suggest broad structural
259 changes. The broad trend is a general delay for ray paths approaching the ridge, probably due to
260 the increase in sediment in the moats followed by increased crustal thickness at the ridge.
261 However, *Pn* arrivals do not align for any single reduction velocity (8 km/s in Figure 7a),
262 indicating lateral variations in mantle structure. Away from the ridge, the apparent velocity of the
263 travel times is close to 8 km/s (Figure 7a,c), but varies locally from values of ~7.8 km/s to 8.2
264 km/s. The sub-ridge area is discussed in Section 3.2.

265 3.2 Stations located on the volcanic edifice and its flanks

266 As stations approach the Ka‘ena Ridge, phase identification becomes more difficult due to
267 travel time delays and scattering caused by edifice topography and lateral velocity gradients. The
268 slopes of travel time curves are unreliable indicators of the seismic phase. Station-shot pairs were
269 examined in reverse configurations to ensure consistency. Several seismic phases, including *P1*,
270 *P2*, *P3*, *P2P*, *PmP*, and *Pn*, were observed, and an additional phase, *Pv*, was identified as waves
271 travelling through the volcanic edifice above the oceanic crust. Distinguishing between *P1* and
272 *Pv* could be challenging in some places, but it does not affect imaging since these phases are
273 treated similarly in the tomographic method (likewise for the *P2* and *P3* phase pair).

274 For stations located atop the ridge (Figure 5), the records exhibit a clear asymmetry in *Pv*
275 arrivals across a midpoint between stations 219 and 220. The apparent velocities for shots
276 moving away from that midpoint are slower than expected for intrusive rocks, indicating layers
277 of lava flows possibly mixed with debris from submarine landslides that occurred during the
278 shield building stage. For ray paths crossing the midpoint between stations 219 and 220, the *Pv*
279 phase shows a rapid change in slope with increasing range, indicating the presence of a central
280 high-velocity body within the edifice of the volcano. In Figures 5 and S4b-d, *P2P*, *P3*, and *PmP*
281 arrivals all indicate a deepening of the oceanic crust and Moho beneath the Ka‘ena Ridge. Some
282 possible sub-Moho reflection events are also observed (e.g., Figure 5f), suggesting the presence
283 of frozen melt sills. Secondary phases in Figure 5e indicate step-changes in velocity structure
284 within or at the base of the Wai‘anae Slump.

285 *Pn* data show progressive time delays toward the ridge, indicating thickening upper layers
286 and a deepening Moho (Figure 7a). In addition, for stations located on one side of the ridge and
287 shots on the other, the travel times of *Pn* waves that undershoot the edifice arrive much earlier
288 than expected when assuming an 8 km/s mantle (Figure 7a). These data are best aligned with a
289 reduction velocity of ~8.25 km/s (Figure 7b; green dots). To emphasize the sensitivity of the
290 phase alignment with changes in reduction velocity, a reduction velocity of 7.8 km/s is also
291 shown (Figure 7b; gray dots) and is strongly inconsistent with the data. On the other hand, 7.8
292 km/s provides a good fit to other *Pn* arrivals along the seismic line for rays that do not
293 undershoot the ridge axis (Figure 7c). From these considerations, we expect that mantle
294 velocities are variable and relatively high (8.2-8.3 km/s) beneath the ridge and lower elsewhere.

295 **3.3 Seismic methods**

296 Seismic imaging was carried out using an iterative tomographic technique that uses travel
297 time data to compute seismic velocity structure (Dunn et al., 2005). This implementation is
298 capable of handling both primary and secondary refracted and reflected phases. Internal
299 reflectors are defined on their own grid for ray-tracing accuracy and velocities defined on the
300 principal grid move vertically with changes in reflector position. The imaging procedure was
301 carried out by first solving for upper crustal structure with a shallow reflector representing the
302 top of the oceanic crust, and then solving for the full structure with a second reflector
303 representing the top of the mantle (the tomographic technique allows for only one reflector at a
304 time). In each case, a starting model was first determined and then 150 solutions were created by
305 varying that starting model. A single final solution was then determined by taking the mean of
306 the stack of solutions, the standard deviation of the stack provides a measure of the relative
307 uncertainty of the image. More details on the methodology and uncertainties are provided in the
308 Supporting Information file that accompanies this submission. The final tomographic image is
309 shown in Figure 8.

310 **4. Gravity data and processing**

311 Gravity data were acquired throughout cruise MGL1806, including along each seismic
312 reflection and refraction line. In this study, we analyze the gravity data collected along OBS
313 seismic Line 02 (Figure 2), coincident with the present seismic tomography analysis.

314 Gravity data were acquired with an axially-constrained BGM-3 sensor mounted on a
315 gyrostabilized platform. In June 2018, the sensor was replaced and the instrument recalibrated
316 with a new pulse rate count to mGal conversion factor of 5.096606269 mGal/count and bias of
317 852,513.49 mGal using tie-in data between the BGM-3 gravimeter and the Honolulu Alpha
318 absolute gravity station. Tie-in data since June 2018 indicate that the new sensor system has
319 performed well with a mistie at the start MGL1806 (some 80 days later) of -6.1 mGal, a mistie of
320 +1.5 mGal at the end of the cruise and a drift rate of -0.155 mGal/day during the cruise (Watts et
321 al., 2020).

322 Prior to correcting for latitude and the Eötvös effect, the 1 s converted count data were
323 filtered with a 120 s Gaussian filter in order to remove accelerations due to ship motions, as

324 recommended by the manufacturer (Bell Aerospace). While we found such a filter to be highly
325 effective at removing swell “noise” (Watts et al., 2020), noise remained at high frequencies
326 (short periods), albeit with significantly less power than swell “noise”. We therefore applied an
327 additional filtering step to the count data using a median filter (width = 1.0 km), which reduced
328 the high frequency noise and, importantly, extended the overall decrease in power with
329 increasing frequency seen in spectral data at low frequencies. The free-air anomaly was
330 calculated from the cleaned gravity data.

331 **5. Seismic structure**

332 **5.1 The Ka‘ena Ridge**

333 The volcanic edifice consists of a relatively high V_p inner “core”, surrounded by a highly
334 variable region of overall lower velocities (Figure 8). This structure is generally consistent with
335 previous studies that have detected localized high-density, high velocity cores within other
336 Hawaiian volcanoes (e.g., Adams & Furumoto, 1965; Strange et al., 1965a; 1965b; Woollard,
337 1951; Flinders et al., 2013). At Ka‘ena, the width of the core is ~30 km, and it stands ~5 km
338 high. V_p values reach a peak of ~7 km/s within the upper central part of the core. Overall, this
339 feature is consistent with a mix of intrusive and extrusive material. Immediately to the south of
340 this feature, there is a smaller, less-pronounced core-like structure. On the northern flank of the
341 edifice, near surface velocities beneath the Ka‘ena slide are low (<3 km/s), and the interior of the
342 flank is composed of large outward-dipping layers of alternating velocities. Likewise, the south
343 flank of the edifice, which includes the Wai'anae Slump, exhibits low near-surface velocities
344 followed by layers of alternating velocities within the interior. The small topographic depression
345 that formed between the edifice and the Wai'anae Slump appears to be filled with low-velocity
346 in-fill material. The maximum thickness of material added to the top of the oceanic crust is ~7
347 km and the width of the edifice is about 100 km at its base.

348 **5.2 Oceanic Crust and Mantle**

349 Away from the complexities of the Hawaiian Ridge, the crust has a typical V_p profile for
350 Pacific oceanic crust, with a thin sub-seafloor low-velocity layer (<4.5 km/s; seismic layer 1)
351 characteristic of accumulated volcanoclastic and oceanic sediments, a thicker low-velocity layer
352 (4-6.5 km/s; seismic layer 2), usually thought to be the upper oceanic crust and composed of

353 dikes and lava flows, and a deeper high-velocity layer (~6.5-7 km/s; seismic layer 3), usually
354 thought to be the lower oceanic crust and composed of intrusive gabbroic rocks (Figure 8) (e.g.,
355 Christeson et al., 2019). Using the distance between the two reflectors as a guide, the median
356 crustal thickness is 6 km, a typical value for oceanic crust formed at faster spreading rates (e.g.,
357 Christeson et al., 2019). The crust is thicker to the south (~6.3 km) than to the north (~5.5 km) of
358 the fracture zone. This asymmetry is similar to that found during a previous study of seismic
359 structure in this region (ten Brink & Brocher, 1988). Note that the nature of the upper reflector is
360 uncertain. While it appears to be the boundary between sediments and the oceanic crust, the
361 velocity contrast between volcanoclastic sediments and 90 Ma seafloor may not be large, and it is
362 possible that the reflector indicates a deeper feature, such as the seismic layer 2A/2B boundary.

363 The lower crust is fairly uniform beneath the northern part of the line, and more variable
364 beneath the southern part of the line, perhaps due to modification via the Southwest O‘ahu
365 Volcanic Field and the fracture zone splays that pass through this area (Figure 8). For example, a
366 distinct low velocity zone is observed in the lower oceanic crust beneath the southern strand of
367 the fracture zone (at -140 km of the profile) and a weaker one is observed beneath the northern
368 strand (at -30 km of the profile). Given their association with the fracture zone, they may be due
369 to crustal damage, increased porosity, and alteration (e.g., Roland et al., 2012).

370 Over wide areas along the seismic line (Figure 8; -250 to -180 km and 140 to 200 km), the
371 lower crust, at a depth of 4-5 km bsf, has slightly lower velocity than material just above. This is
372 corroborated by seismic records for these areas, which show the rapid loss of refracted *P3* wave
373 energy with increasing ray turning depth, indicative of the presence of a negative vertical
374 velocity gradient (Figures 4a and S1).

375 Approaching the Ka‘ena Ridge, the sediment layer thickens and the shallow seismic reflector
376 deepens at a greater rate than the seafloor (Figure 8). This indicates increasing downward
377 curvature of the top of the crust toward the edifice, presumably due to flexural loading of the
378 plate. The overall deflection of the top of the oceanic crust beneath the edifice (with respect to its
379 position at the ends of the seismic line) is ~3 km. The same measurement for the Moho yields a
380 value of ~3.5 km. In the upper oceanic crust, the lowest *V_p* values correspond with the location
381 of maximum curvature, just outboard of the flanks of the edifice.

382 Beneath the edifice, the oceanic crust has anomalously high velocities at all levels (Figure 8).
383 The low-velocity layer at the top of the crust is missing over a region roughly 75 km wide,
384 replaced with velocities exceeding 6 km/s with a peak central value of ~ 7 km/s. In the lower
385 crust, V_p generally exceeds 7 km/s and rises to ~ 7.5 km/s just above the Moho reflector over a
386 ~ 150 km wide area. The latter observation is difficult to verify with the travel time data alone.
387 Given the possible trade-off of V_p values with crustal thickness just above the Moho (a 1 km
388 Moho depth change relates to a change in V_p of almost ~ 0.4 km/s above the Moho; Figure S6),
389 we suggest caution when interpreting this.

390 Beneath the Hawaiian Ridge, a high V_p region extends downward into the mantle, widening
391 with depth. Here, mantle V_p exceeds 8.25 km/s over a wide area beneath the edifice, as
392 compared to ~ 7.75 -8 km/s outside of this area. Note that in the study area, the seismogenic layer
393 thickness is expected to be >40 km thick, on the basis of the depth limit of seismicity (e.g.,
394 McKenzie et al., 2005), and therefore the image is entirely contained within this layer. Overall,
395 the mantle exhibits a small positive gradient with depth of ~ 0.01 km/s per km of depth. There is
396 no evidence in the image nor in the data records themselves for a distinct sub-crustal low-
397 velocity layer (with velocities between crust and mantle values) either located beneath the edifice
398 (Watts et al., 1985) or underlying a broad area around the islands (Leahy et al., 2010).

399 **6. Gravity anomalies**

400 To verify the final seismic tomographic image, we computed the gravity effect of the
401 seismically constrained crust and mantle structure assuming different empirical relationships
402 between P wave velocity and density and compared it to the observed free-air gravity anomaly.
403 Figure 9 shows the observed and calculated gravity anomalies along Line 02, the iso-velocity
404 contours derived from Figure 8 and the density structure used in the gravity calculations. The
405 calculated gravity is based on a layered structure in which the density contrast between layers is
406 derived from the average P wave velocity above and below an individual iso-velocity contour.
407 Figure 9a (black solid line) shows the sum of the gravity effect of all the layers, which was
408 computed using a 3D Fast Fourier Transform modelling method for the seafloor bathymetry and
409 a 2D line-integral method for each sub-seafloor layer and an empirical relationship between P
410 wave velocity and density as defined by Nafe-Drake for $1.5 < V_p < 5.5$ km/s and Christensen
411 and Mooney for $V_p \geq 5.5$ km/s (Brocher, 2005).

412 Figure 9a shows a close agreement between the calculated gravity effect of the seismically
413 constrained crust and mantle structure in Figure 8 and the observed free-air gravity anomaly. The
414 Root Mean Square (RMS) difference between observed and calculated gravity is 8.7 mGal,
415 which is small when compared to the range of the observed free-air anomalies (~ -75 to $+225$
416 mGal). The main discrepancies are in the region of the flexural moat northeast of O‘ahu where
417 the calculated gravity effect is generally too low compared to the observed anomaly. The
418 discrepancy correlates with a 50-km wide region of relatively low density in the uppermost part
419 of the sub-crustal mantle and so we speculate that it is caused by a slightly denser mantle beneath
420 the moat northeast of O‘ahu than beneath the moat to the southwest of the island. A more
421 localized discrepancy occurs over the crest of the Ka‘ena Ridge which we attribute to the three-
422 dimensional structure of the high velocity core of the volcanic edifice.

423 The contributions of the individual sub-seafloor layers to the calculated gravity are shown in
424 Figure S8 of Supporting Information. That figure, together with Figure 9a, shows the observed
425 gravity high over the edifice is due to a combination of the gravity effect of seafloor bathymetry
426 and the shallowing of the sub-seafloor 3-7 km/s iso-velocity contours, with the amplitude and
427 wavelength of the high mainly controlled by the gravity effect of the 5.5-7 and 3-5 km/s iso-
428 velocity contours respectively. Seismic and gravity data therefore reflect the origin of the
429 volcanic edifice with its low velocity and density extrusive drape and its high velocity and
430 density intrusive core. Flanking the high are lows that are mainly caused by the gravity effect of
431 an increase in depth of the 5-7.5 km/s contours from the bulge region to beneath the edifice. We
432 attribute this increase to crustal flexure in response to long-term (> 1 Myr) volcano loading.
433 Competing with the gravity anomaly high and flanking lows associated with flexure is the mantle
434 structure, and our seismic and gravity modelling are consistent with the existence of a relatively
435 high P wave velocity and density mantle beneath the volcanic edifice and relatively low P wave
436 velocity and density mantle beneath its flanking moats.

437 **7. Discussion**

438 **7.1 Oceanic crustal and upper mantle structure**

439 In the moat areas, sediment wave speeds are < 3 km/s near the seafloor in places, and
440 generally < 5 km/s everywhere else (Figure 8). This is consistent with volcanoclastic sediments

441 (e.g., Hammer et al., 1994; Weigel & Grevemeyer, 1999), overlain by a thin layer of mixed
442 oceanic and volcanic sediment. There is no clear distinction between moat material and the
443 material that makes up the flanks of the edifice, and these materials are probably greatly
444 intermingled at the outer edges of the edifice (Figure 10).

445 At the intersections of the seismic line with strands of the fracture zone (Figures 8 and 10),
446 the low velocity regions in the crust may be a consequence of fault damage and alteration (e.g.,
447 Roland et al., 2012). Between the two strands, the crust is somewhat thinner (~5.5 km), and
448 thinner oceanic crust is often associated with fracture zone offsets (e.g., Detrick & Purdy, 1980;
449 Detrick et al., 1993; Minshull et al., 1991; White et al., 1984), and is usually attributed to
450 suppressed mantle melting near portions of the mid-ocean ridge where the fracture zone
451 originated (e.g., Detrick et al., 1993; Stroup & Fox, 1981; White et al., 1984).

452 In the oceanic crust, away from the edifice and fracture zones, there are wide low-velocity
453 regions at mid-crustal depth (Figure 8). We suggest that as the crust ages (e.g., Grevemeyer &
454 Weigel, 1996), fractures and pore space that formed nearer to the mid-ocean ridge will seal from
455 the top downward. This could result in relatively higher velocities overlying a region of
456 relatively lower values where this process has not yet taken place. In the upper oceanic crust, the
457 low V_p values located at maximum downward plate curvature, just outboard of the flanks of the
458 edifice, may be due to the formation of tensile cracks or larger fractures as a consequence of
459 flexure, which would lower seismic wave speed (e.g., Berge et al., 1992).

460 Across the Hawaiian Ridge at Ka'ena, the edifice is about 7 km thick, and the oceanic crust
461 plunges downward beneath the edifice, with a total deflection of ~3.5 km over a ~400-500-km-
462 wide region (Figures 8 and 10). Beneath the edifice, the fate of the top of the oceanic crust is
463 unknown. Presumably some combination of magmatic overprinting and compression of original
464 pore space has led to higher velocities. Here, the upper reflector beneath the edifice is at a depth
465 roughly consistent with the top of the oceanic crust (on the basis of its height above the Moho),
466 unlike what was found near the Island of Hawai'i (MacGregor et al., 2023), where it appears that
467 the dominant reflection marks the top of the lower oceanic crust and a reflection from the top of
468 the oceanic crust is weak and intermittent. In the lower oceanic crust, near the Moho, elevated
469 velocities, suggest magmatic overprinting, perhaps due to the crystallization of ultramafic
470 minerals as melts enter the crust. Furthermore, below this region, we see elevated mantle

471 velocities. This is well documented by the travel time data (e.g., Figure 7) and by the necessity
472 for its presence to fit the gravity data (Figure 9). The gravity constraint is important, because it
473 indicates a bulk mineralogical change, rather than a change in the orientation of anisotropic
474 minerals in the mantle (i.e., an alignment of the fast axes that is more parallel to the seismic line).
475 One could argue that the high wave speeds were inherited from processes in operation at the
476 spreading center during crustal formation, but the feature's close association with the ridge
477 suggests related processes rather than coincidence. The anomalous mantle wave speed suggests
478 that melts beneath the ridge rose through the mantle and chemically interacted with the cooler
479 lithosphere as they passed through, leaving behind a more ultramafic-rich material.

480 We find no evidence for significant underplating of melts beneath the Moho, as suggested by
481 prior studies of Hawai'i (e.g., Leahy et al., 2010; Watts et al., 1985) and some other marine
482 volcanoes (e.g., Gallart et al., 1999; Grevemeyer et al., 2001). We also rule out broad
483 serpentinization of the upper mantle, as suggested by Park and Rye (2019), though there may be
484 some minor local serpentinization such as indicated beneath the fracture zone. Instead, we find a
485 simple Moho transition, with some variability in its character, as well as evidence for isolated
486 sub-Moho frozen melt sills. Since the Ka'ena seismic line is not co-located with the previous
487 study of Watts et al. (1985), the differences in interpretation (no underplating versus
488 underplating) may simply indicate that underplating is variable along the ridge. However,
489 reprocessing of the earlier seismic data has not found evidence for underplating (Lindwall, 1988;
490 Cilli et al., 2023) and other recent studies of the Hawaiian-Emperor Seamount Chain have also
491 not found evidence for underplating (McGregor et al., 2023; Watts et al., 2022; Xu et al., 2022).

492 We do find some local variations in V_p that may be related to variable melt-rock reaction
493 effects and/or inefficient melt extraction, along with evidence of a few sub-Moho reflectors.
494 Ohira et al. (2018), using wide-angle seismic data from 750-km-long line located north of O'ahu,
495 also found evidence for several isolated sub-Moho frozen melt sills at depths ranging from 30-85
496 km, which they suggest formed during cooling of the plate. Imbedded within the high-velocity
497 mantle region beneath the Ka'ena Ridge is a smaller region of relatively lower velocities (~8
498 km/s) that lies directly beneath the ridge (Figure 8) and is associated with one of the possible
499 sub-Moho melt lenses (Figure 5f). This small region could be a consequence of frozen mafic
500 melts, associated with an incomplete separation of mafic and ultramafic material as the magmatic
501 phase of the seamount waned. Other studies have also suggested incomplete melt extraction from

502 the mantle, both in hot-spot affected areas (e.g., Gaherty & Dunn, 2007; Walther, 2003), and
503 near mid-ocean ridges (e.g., Conley & Dunn, 2011; Lizarralde et al., 2004). Nevertheless, it
504 would not represent a significant amount of melt as compared to the mass of the edifice or to
505 previous suggestions of underplating.

506 **7.2 Volcano structure**

507 Although evidence suggests that Ka'ena Ridge formed as a separate shield volcano from that
508 of O'ahu's Wai'anae volcano (Sinton et al., 2014), the location of its central vent and its
509 structural association with Wai'anae volcano and the Wai'alu Ridge are uncertain. Many
510 Hawaiian volcanoes are characterized by a central strong residual Bouguer gravity anomaly high
511 (e.g., Flinders et al., 2013; Strange et al., 1965a) and shallow high-velocity bodies (e.g.,
512 Furumoto et al., 1965; Park et al., 2009), which are often ascribed to the location of the principle
513 magmatic conduits through which the shield building stage of the volcano occurred, leaving
514 behind dense, possibly olivine-rich, cumulates. On the basis of the gravity structure and other
515 features of Ka'ena, Sinton et al. (2014) speculate that the center of Ka'ena Volcano lies buried
516 beneath lava flows from Wai'anae volcano at a position roughly centered just east of our seismic
517 line (Figure 2b). The gravity anomaly in this area is modest, only ~10 mgal as compared to >50
518 mgal for Wai'anae Volcano (Figure 2b). Sinton et al. (2014) suggest that the smaller gravity
519 anomaly and overall morphology of Ka'ena are consistent with Ka'ena being a smaller volcano
520 by volume, with more broadly disseminated volcanic activities. By this view, the portion of
521 Ka'ena Ridge extending westward from the gravity anomaly, and Wai'alu Ridge extending to the
522 northwest, are consistent with being rift zones of Ka'ena Volcano.

523 The seismic line is located just along the western edge of the estimated center of volcanism
524 of Ka'ena volcano (Figure 2b). Presumably then, the high-density and high-velocity core imaged
525 here (Figure 10) is the western edge of the center of volcanism of Ka'ena volcano. Alternatively,
526 the imaged core could be a slice of a more elongate feature that extends considerably along the
527 Ka'ena Ridge, such as is expected for the core-like feature detected beneath the Hāna Ridge
528 (MacGregor et al., 2023). The presence of the core suggests an edifice-level magma chamber or
529 mush zone, which would have evolved considerably in size and elevation during volcano growth.
530 The core properties are consistent with either intrusive gabbroic rocks, or a mixture of lavas,
531 dikes, gabbro, and ultramafic cumulates (i.e., a mixture of both high and low density/velocity

532 materials). The latter scenario seems more likely since magmatism would have built the core by
533 intruding previously formed dikes and lavas. The smaller core-like feature to the south (Figure
534 10) may have its origin in a small southeasterly trending rift zone located beneath the slump and
535 associated with a secondary center of volcanism of Ka'ena (Sinton et al., 2014).

536 Landslides of various types are ubiquitous features of the flanks of the Hawaiian Ridge. They
537 appear to form by both gravitational sector collapse and by lateral forces resulting from magma
538 injection into a volcano's rift zones (e.g., Moore et al, 1989). The shield stage of volcano growth
539 may generate the most landslides, due to active lava emplacement and oversteepening of the
540 volcano flanks (Coombs et al., 2004; Moore et al., 1989). The alternating layers of lower and
541 higher seismic velocity (Figure 8) that characterize the flanks of Ka'ena Ridge may represent
542 alternating lithologies (Figure 10), such as alternating layers of lava flows and debris avalanche
543 material. This 'layer-cake' structure may have implications for future landslides in that weaker
544 layers may allow for large-scale sector collapse of the volcano, such as has already occurred at
545 the Wai'anae Slump (Figure 2b). The slump is estimated to be ~1.5 km thick (Coombs et al.,
546 2004), and may thus have given way along a zone of weakness indicated by a band of low
547 velocities at that depth (Figure 8). Considering a trigger for the slump, we note that it overlies a
548 splay of the Molokai Fracture Zone. The age asymmetry across the fracture zone indicates that
549 the two halves of the plate have different thermomechanical properties and may have responded
550 differently to the volcanic load. We speculate that as the Hawaiian Ridge grew across the
551 fracture zone, there occurred enough differential flexure and vertical motion across it to trigger
552 the slow collapse of the volcano flank.

553 **7.3 Plate flexure**

554 The tomographic model can be used together with bathymetry and gravity data to constrain
555 the flexure caused by the volcanic load of the Hawaiian Ridge and the elastic parameters
556 associated with this flexure. The load was derived from a regional GEBCO bathymetric grid,
557 corrected for the effects of thermal age, assuming a cooling plate model (Parsons & Sclater,
558 1977) and the Müller et al. (1997) age grid, and for formation of the Hawaiian swell (e.g.,
559 Wessel, 2016), assuming it to be derived from a 500-km-wide median filter of the GEBCO
560 bathymetric grid. We compared the average depth to the top of oceanic crust and the Moho
561 derived from the tomographic model (red curves in Figure 8) to predictions based on a simple

562 three-dimensional elastic plate (flexure) model. The ‘best fit’ parameters for the model were
563 estimated from the RMS difference between the observed seismically constrained surfaces of top
564 of oceanic crust and Moho and the calculated surfaces for a range of densities for the material
565 infilling the flexure (ρ_{infill}) of 2100 to 2800 kg/m³ and T_e of 0-60 km, assuming a load density
566 (based on the P wave velocity structure of the volcanic edifice) of 2737 kg/m³. The results are
567 summarized in Figure S9 of Supporting Information which show a best fit parameter pair of ρ_{infill}
568 = 2450 kg/m³ and $T_e = 25.6$ km for the top of oceanic crust and $\rho_{infill} = 2450$ kg/m³ and a slightly
569 lower $T_e = 23.1$ km and the same $\rho_{infill} = 2450$ kg/m³ for the Moho. The minimum RMS values
570 for the two surfaces are 167.3 m and 325.9 m, respectively, indicating that the best overall fit (by
571 a factor of 2) is for the top of oceanic crust. Changing the load density within bounds determined
572 by the seismic constraints on the top of oceanic crust has little effect on the results: decreasing
573 the load density to 2637 kg/m³, for example, results in a best fit T_e , average ρ_{infill} and RMS of
574 23.1 km, 2500 kg/m³ and 166.2 m respectively, while increasing the load density to 2867 kg/m³
575 results in a best fit of 25.6 km, 2400 kg/m³ and 164.1 m respectively.

576 The T_e results here are in general accord with the results of Cilli et al. (2023) using
577 reprocessed R/V *Robert D. Conrad* and the R/V *Kana Keoki* multichannel seismic reflection
578 profile data. They used a similar range of load densities and derived a T_e in the range 25.7 to 27.7
579 km, slightly higher than the values derived here. However, the *Conrad* and *Kana Keoki* profiles
580 do not cross the Ka‘ena Ridge, intersecting instead the Hawaiian Ridge at the deep channel that
581 separates Kauai and O‘ahu and the insular shelf between O‘ahu and Moloka‘i. In addition, a
582 somewhat higher infill density was obtained (2701 kg/m³) and perhaps, most significantly, the
583 lowest RMS obtained for the top of the oceanic crust was 305.5 m, which is nearly a factor of 2
584 higher than that derived here.

585 The best-fit flexure model based on a load density to 2737 kg/m³ and a pre-existing oceanic
586 crustal thickness of 6.05 km (the mean difference between the depth to the two reflector surfaces
587 in Figure 8) is shown by the light blue shading in Figure 9b. There is an excellent agreement
588 between the seismically constrained surfaces of the top of oceanic crust and Moho and the
589 calculated flexure. The fit between the observed and calculated free-air gravity anomaly (Figure
590 9a) is not, however, as close. We attribute the misfit to our assumptions that the only loads that
591 act on the plate are surface loads as defined by the present-day bathymetry, and that the material
592 infilling the flexure is of uniform density. For example, the bathymetry is not the only

593 contributor to the positive part of the observed gravity anomaly. There is therefore likely to be a
594 significant contribution to the gravity anomaly of sub-surface loads, especially those associated
595 with the 5, 6 and 6.5 km/s iso-velocity contours. The effect of a non-uniform infill density is
596 more difficult to assess. The flexure model in Figure 9b assumes an average $\rho_{infill} = 2450 \text{ kg/m}^3$,
597 but it is likely that ρ_{infill} immediately beneath the loads is higher and of the order of the load
598 density and lower in the flanking flexural moats.

599 **7.4 Comparison of Lines 01 and 02**

600 In this study, we investigate the relationship between the Ka'ena Ridge's volcanic load
601 distribution, lithospheric deformation, and elastic properties. The results for this geophysical line
602 (Line 02) can be directly compared to those presented in our companion study (MacGregor et al.,
603 2023) located to the west of Hawai'i (Line 01). The two studies examine plate flexure for
604 situations in which the underlying plate age is similar, but the load size and structure are much
605 different.

606 At Ka'ena, the volcanic edifice's height and cross-sectional area are approximately half as
607 large as those at Hawai'i (Table 1). The presence of a large, high-density, and high-velocity core
608 suggests that the seismic line intersects a significant portion of the volcano's center of
609 volcanism. The core's properties are consistent with intrusive gabbroic rocks or a mixture of
610 dense mafic and ultramafic materials, indicating the previous existence of an edifice-level
611 magma chamber or mush zone. The proportion of intrusive to extrusive rocks is estimated at
612 approximately 7% (excluding materials that may have entered the moat). In contrast, the western
613 flank of the island of Hawaii lacks such a feature and appears to consist primarily of layer upon
614 layer of extrusive rocks. The relatively smaller cross-section of the Hāna Ridge may comprise as
615 much as ~20% dense intrusive material.

616 At Ka'ena, the amount of deflection is about half that measured westward of the island of
617 Hawai'i (MacGregor et al., 2023). Despite the large difference in load size and flexure between
618 the two locations, the calculated T_e values are similar (25.6 for Ka'ena versus 26.7 westward of
619 the island of Hawaii). These results are supported by Cilli et al. (2023), who used reprocessed
620 legacy MCS data as a basis for additional flexural modelling near O'ahu (T_e of 26.7), but largely
621 eastward of the Ka'ena seismic line. Together, these studies confirm that volcanic loads to the
622 west of Hawai'i are largely compensated by flexure. However, as discussed in MacGregor et al.

623 (2023), previous studies suggest on-going subsidence and a higher T_e value beneath of the
624 eastern flank of Hawai'i Island, suggesting that isostatic compensation may not yet be complete
625 at the youngest end of the ridge.

626 Magmatic underplating beneath the oceanic crust has been suggested for several intraplate
627 volcanic settings and affects loading and mass flux calculations. All else being equal,
628 underplating would be expected to lower the amplitude of the gravity anomaly over the crest of
629 the edifice and provide a positive buoyancy force that would make the plate appear more rigid
630 than it actually is. One hypothesis put forward to explain variations in the style of magmatic
631 emplacement at intraplate volcanoes hinges on the age of the lithosphere at the time of volcano
632 formation. In this model, shallow intrusion into the oceanic crust and overlying edifice is favored
633 for seamounts growing on young lithosphere, while magmatic underplating is favored for
634 seamounts growing on old lithosphere (Contreras-Reyes et al., 2010). However, the recent
635 seismic, gravity, and plate flexure studies conducted along the Hawaiian-Emperor Seamount
636 Chain (including this study, MacGregor et al., 2023, Watts et al., 2021, and Xu et al., 2022)
637 collectively provide clear evidence for shallow magmatic emplacement and contradict the notion
638 of significant magmatic underplating for ages at the time of loading of ~57 Ma (Emperor
639 Seamounts) and ~90 Ma (Hawaiian Ridge). Additionally, reprocessing of the legacy data of
640 Watts et al. (1985) did not find evidence for underplating (Lindwall, 1988; Cilli et al., 2023).
641 These findings invalidate a simple age-related hypothesis.

642 The recent seismic studies found only scant evidence for a mantle conduit that might be
643 ascribed to a narrow zone where melts passed upwards towards a volcano along the ridge.
644 Beneath the Ka'ena line we find a broad zone of high velocities that may be due to melt-rock
645 reactions, and an inner region near the Moho of lower velocities wherein melts may have frozen
646 before they could separate from the mantle. A similar structure may be present beneath Jimmu
647 Guyot (Xu et al., 2022), while Watts et al. (2021), using a ridge parallel seismic line, found
648 variations in mantle structure that are difficult to reconcile with any simple model of mantle
649 conduits feeding the cores of ridge volcanoes. In these studies, at mantle levels the width of the
650 minimum resolvable feature may be 20 km or more and perhaps regions of melt passage are too
651 isolated or too narrow to significantly influence the seismic wavefield and cannot be imaged
652 using these experiment geometries.

653 8. Conclusions

654 Using wide-angle seismic refraction and reflection data collected along a 550-km-long line
655 that intersects the Hawaiian Ridge at Ka‘ena volcano, we imaged the structure of the edifice and
656 the underlying Pacific oceanic crust and upper mantle using *P* wave tomography. By modeling
657 the density structure and gravity data and the degree of plate flexure, as determined by the
658 seismic results, we investigated the relationship between the ridge’s volcanic load distribution,
659 lithospheric deformation, and elastic properties and compare the results to other studies located
660 along the Hawaiian-Emperor Ridge. Our findings shed light on several key aspects of the
661 Ridge’s geological and geophysical characteristics.

662 *Volcanic Construction and Load Distribution.* Our seismic imaging and gravity data
663 collectively provide evidence for a complex internal structure within the Hawaiian Ridge at
664 Ka‘ena Ridge, characterized by a high velocity, high density, core overlain by volcanic layers
665 and debris flows. The presence of a cumulate core indicates that shallow intrusive processes
666 shape the Ridge’s volcanic evolution. The presence of this core is consistent with several other
667 studies of Hawaiian volcanoes that have detected the presence of such features and ascribed them
668 to central conduits of magmatic activity. Despite this interpretation, no corresponding features
669 have been detected in the mantle. In addition, no wide-spread crustal underplating of magmatic
670 material was detected here or beneath Hawaii island, but some scattered upper mantle intrusions
671 are indicated by the seismic data. Shallow layer-cake structures beneath the flanks of Ka‘ena
672 Ridge may indicate zones of weakness that promote flank collapse.

673 *Lithospheric Flexure.* Lithospheric deformation is prominently expressed through the
674 Hawaiian Moat and sub-ridge depression of the seismically determined sediment-crust and crust-
675 mantle boundaries. The added load is up to ~7 km in height, and the flexural response is broad
676 with a resultant peak vertical deflection of ~3.5 km, suggesting a flexural rigidity of the
677 lithosphere which is ~90 Myr old, equivalent to $T_e = \sim 25.6$ km. This is similar to other recent
678 estimates along the Hawaiian Ridge, despite the large differences in load size for these different
679 studies. Comparisons to the Emperor Seamount Chain confirm the Hawaiian Ridge’s relatively
680 stronger lithospheric rigidity, emphasizing the significance of lithospheric age at the time of
681 loading to the load-induced flexural response.

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689 to Repository archive (R²R, 2023). OBSIC data are archived at the EarthScope Consortium
690 facilities (HI-Emperor, 2019; SAGE, 2023). The seismic models for Lines 01 and 02 are
691 archived at Dunn et al. (2024) and Dunn (2024), respectively.

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954

955 **Tables and Figure Captions**

956

Table 1. Comparison of Approximate Edifice Size and Flexure

Lithospheric age at loading	Peak height of edifice above oceanic crust	Cross-sectional area of the edifice above the oceanic	Peak deflection of plate (km)	Calculated T_e value (km)

	(Myr)	(km)	crust* (km ²)		
Ka'ena Ridge	~82-95	7.3	800	3.5	25.6
Hawaii and Hāna	~90-95	12	1450	6-7	26.7
Jimmu Guyot	~55	8	550	3.8	14

* Not including moat fill, which also impacts the degree of flexure.

957

958 **Figure 1.** Free-air gravity anomaly map (Sandwell et al., 2019) of the area around the Hawaiian
 959 Islands and the seismic experiment. Heavy black lines indicate position of the seismic lines
 960 collected as part of the cruise MGL1806. The small white and gray circles indicate OBS
 961 positions. Line 02, the western line that runs approximately north-to-south is analyzed in this
 962 study. It crosses the Hawaiian Ridge at the topographic high called the Ka'ena Ridge, and spans
 963 the Hawaiian Trough or moat (the low gravity anomaly that surrounds the ridge), and part of the
 964 Hawaiian arch. Individual strands of the Molokai Fracture Zone appear to extend beneath the
 965 island chain and intersect the seismic line. To highlight the fracture zones, the gravity contour
 966 interval is 6 mgal up to 160 mgal, thereafter it is 48 mgal - the transition roughly coinciding with
 967 the shallow water and subaerial portions of the Hawaiian Ridge.

968 **Figure 2.** Maps of the O'ahu area showing location of the seismic line and local geographic
 969 features. Panel (a) shows locations of the OBS instruments (numbered circles) and co-located
 970 OBS and MCS seismic shot lines (heavy black line). Black dots indicate positions of the centers
 971 of ESP profiles from Watts et al. (1985). The irregular dotted curve indicates the location of the
 972 North Arch Volcanic Field. The black box indicates the area shown in (b). Panel (b) shows
 973 details of Ka'ena Ridge and other geological features discussed in the text. The Ka'ena Slide,
 974 Wai'anae Slump, a large debris field, and Southwest O'ahu Volcanic Field are labeled and their
 975 extents indicated by large-dashed lines. The dotted lines indicate the 10, 20, 30, 40, 50 mgal
 976 contours of residual Bouguer gravity (from Flinders et al., 2013). Proposed centers of volcanism
 977 for Ka'ena and Wai'anae volcanoes are indicated by the black and red stars, respectively. The
 978 heavy dashed lines indicate possible rift systems extending from Ka'ena and Wai'anae. The solid
 979 black line indicates the position of the seismic line.

980 **Figure 3.** Example of a common-receiver-gather record section for the complete seismic line
981 (station 208, vertical geophone channel). This record highlights the *Pn* phase and its observation
982 at distances greater than 300 km. On the north side of the instrument the *PmP* phase is faint,
983 whereas *Pn* is strong and exhibits a weak precursor (apparent velocity ~ 7.7 km/s). This may be
984 due to a stepped and more transitional nature of the crust-mantle boundary beneath the
985 Southwest O‘ahu Volcanic Field. Within the moat area, the asymmetry of the slope of the *Pn*
986 arrivals across the station indicates a dipping Moho (and thickening sediments) towards the
987 Ka‘ena Ridge, rather than actual differences in mantle velocity. Across the Ka‘ena Ridge, the
988 large travel time delay is probably the result of increased overall crustal thickness. Predictive
989 deconvolution was applied to the data shown in the inset.

990 **Figure 4.** (a-f) Common-receiver-gather records (OBS shot data) for stations located across the
991 Hawaiian moat, showing crustal and mantle phases. (g) and (h) are similar to (d) and (f),
992 respectively, but using MCS shot data and emphasizing *P2P* arrivals using a lower reduction
993 velocity. The station located deeper into the moat, plot (g), shows *P2P* reflections relatively
994 delayed, as compared to the station nearer the outer edge of the moat, plot (h), indicating
995 increased sediment thickness and down warping of the oceanic crust towards the Ka‘ena Ridge.
996 The box in (e) indicates the plot area in (i), which emphasizes the *PmP* arrival and possible sub-
997 Moho reflection to the north of station 227 (white curve in plots (e) and (f)). A companion sub-
998 Moho event can be seen on the south (left) side of station 231, as indicated by the white curve in
999 (f).

1000 **Figure 5.** Common-receiver-gather records for stations located across the Hawaiian Ridge. *P2P*
1001 and *PmP* phases are clearly delayed as compared with off-ridge stations. Arrows indicate *Pv*
1002 waves that passed through the center of the edifice and arrive earlier (~ 6.5 - 7.0 km/s), as
1003 compared to similar waves that travelled through the outer portions of the edifice (~ 5.0 - 5.5
1004 km/s). (e) shows an additional arrival from the interior of the landslide portion of the edifice. (f)
1005 Shows one of the possible sub-Moho arrivals that are observed intermittently along the seismic
1006 line.

1007 **Figure 6.** Travel times of the seismic phases used for tomographic imaging, color-coded by
1008 seismic phase. Station numbers and their positions are indicated by the labels. To better show the
1009 coherency of the seismic phases from station to station, the range to each pick has been divided

1010 by two before plotting. This is effective because the slope of each travel time curve is dependent
1011 on the wave speed at the turning depth of the ray, which is approximately halfway between the
1012 shot and station.

1013 **Figure 7.** Travel times of the *Pn* seismic phase using a reduced time axis to align the arrival
1014 times. (a) Shows all *Pn* travel time data; no single reduction velocity satisfies all regions of the
1015 data. In particular, rays that undershoot the ridge are greatly misaligned at this reduction
1016 velocity. (b) Shows data with ranges >100 km that undershoot the ridge axis recorded by off-axis
1017 stations (stations 203-212, 224-235) and off-axis shots, for two different reduction velocities. A
1018 velocity of 8.25 km/s (green values) provides an overall good alignment, despite local structural
1019 variations. A value of 7.8 km/s (gray values), provides poor alignment. (c) Shows data with
1020 ranges <100 km (omits ridge undershoot paths), and reveals an improvement in alignment over
1021 data shown in (a) for most areas away from the ridge (green dots) using a lower reduction
1022 velocity. As in Fig. 6, the range of each pick is divided by two.

1023 **Figure 8.** Final *P*-wave tomographic image (mean of all models), along with both the shallow
1024 reflector (upper red curve), whose position is controlled by the *P2P* phase and deeper refractions,
1025 and the deeper reflector (lower red curve), whose position is controlled by the *PmP* and *Pn*
1026 phases. The vertical exaggeration is a factor of 8.

1027 **Figure 9.** Comparison of the calculated gravity effect of the tomographic image with the
1028 observed free-air gravity anomaly. a) Observed and calculated free-air gravity anomaly. The
1029 observed gravity is based on the 1 s converted BGM-3 count data with a 120 s Gaussian filter
1030 (gray filled circles) and an additional 1.0 km median filter (red solid line). The calculated gravity
1031 is based on the tomographic image in Figure 8, the Nafe-Drake and Christensen and Mooney
1032 empirical relationships between *P* wave velocity and density (solid black line) and a continuous
1033 3D elastic plate (flexure) model based on a load density of 2737 kg/m³, an average infill density
1034 of 2450 kg/m³, an elastic thickness, T_e , of 25.6 km, and an assumed thickness of the pre-flexed
1035 oceanic crust of 6.04 km (dashed orange line). Light blue profile shows the calculated gravity
1036 effect of the bathymetry based on a 2D (solid line) and 3D (dashed line) model. b) *P* wave
1037 velocity contours (in km/s) and densities (in kg/m³) used in the calculation of the gravity
1038 anomalies. The light blue shaded area shows the calculated flexure of the oceanic crust based on
1039 the same parameters as used in the gravity calculation.

1040 **Figure 10:** Schematic model showing an interpretation of the seismic image. Melts rising
1041 through the cold lithosphere react with the host rock leaving behind higher density and wave
1042 speed minerals, such as olivine. As the melts move into the edifice, they enter a large mush zone
1043 that increases in elevation as the volcano grows. A small neighboring locus of melt injection is
1044 also observed on the south side of the main volcanic core. Some “off-ridge” melts may rise up to
1045 form the Southwest O‘ahu Volcanic field, the North Arch volcanics, and occasional mantle-level
1046 melt sills.

Figure 1.

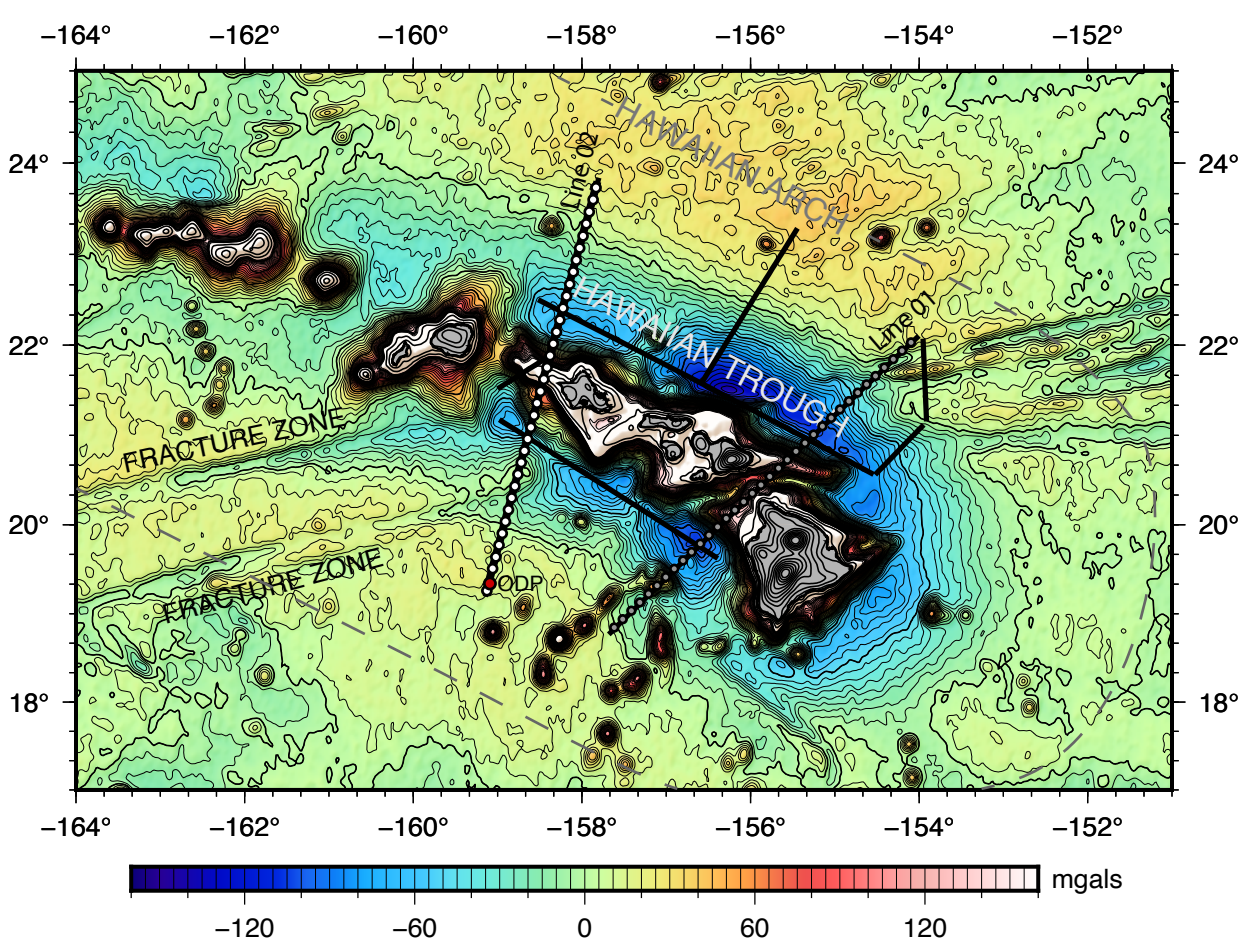


Figure 2.

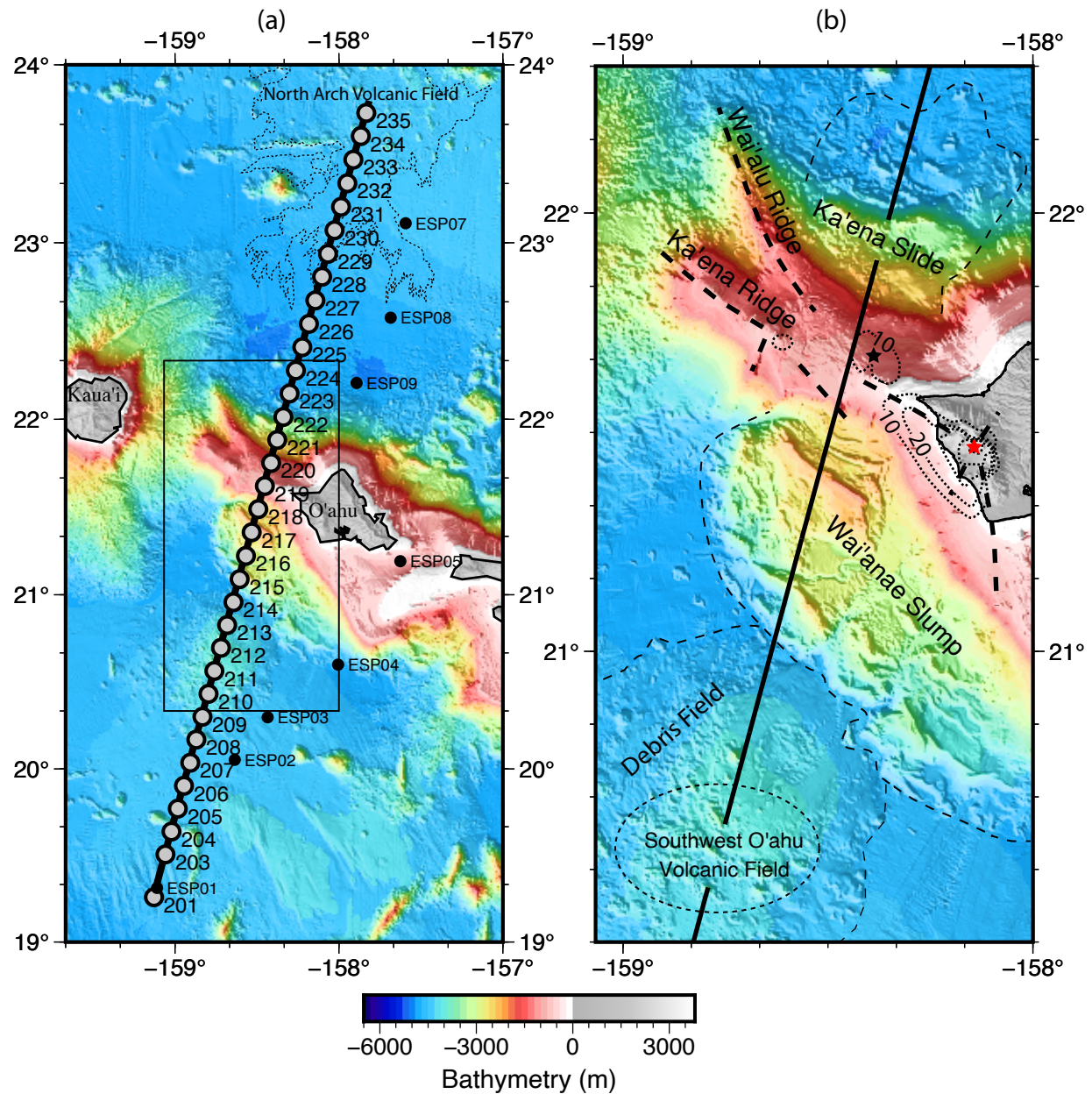


Figure 3.

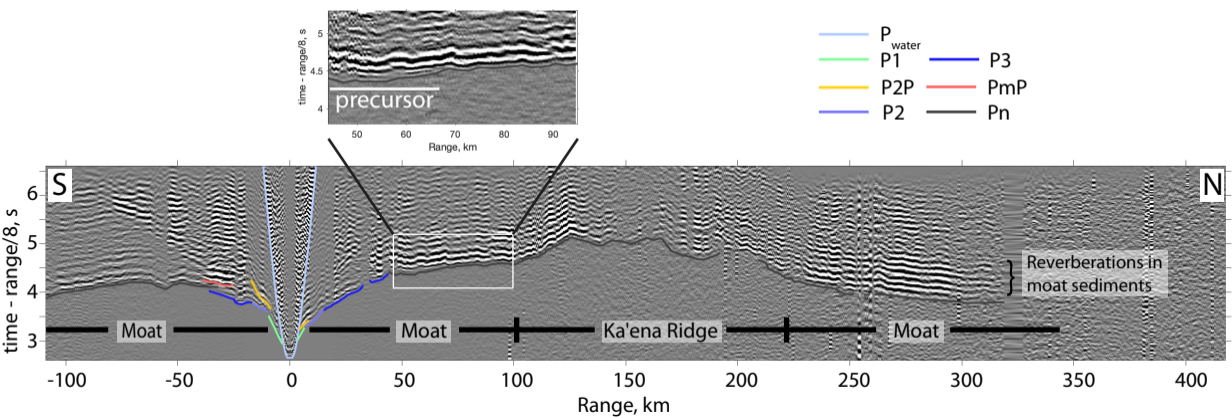


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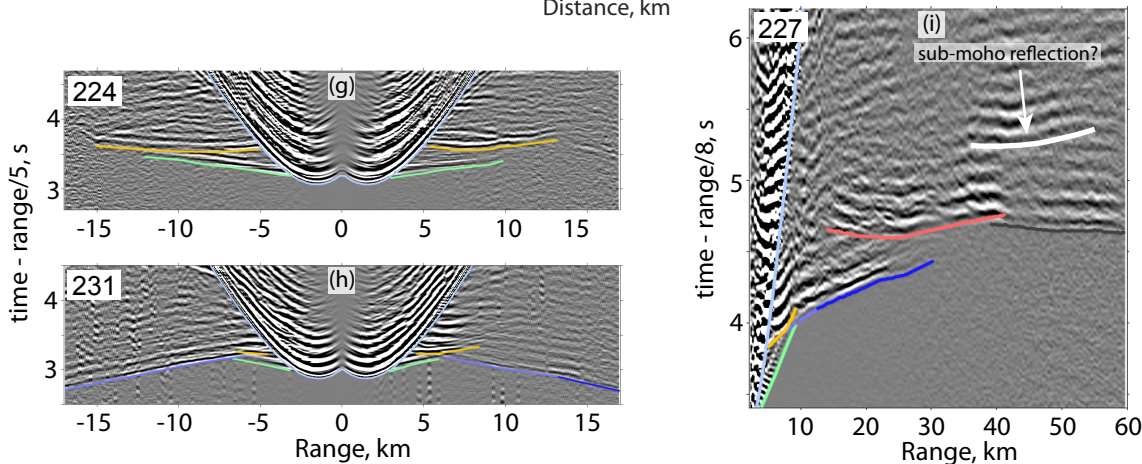
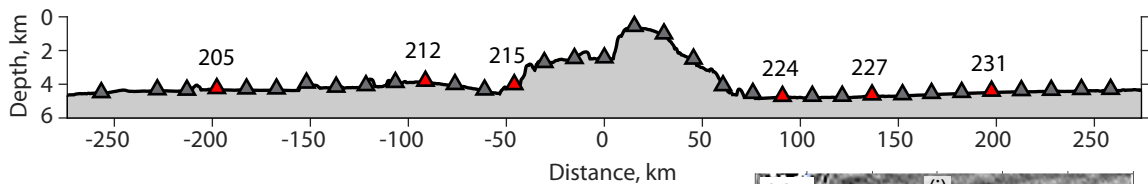
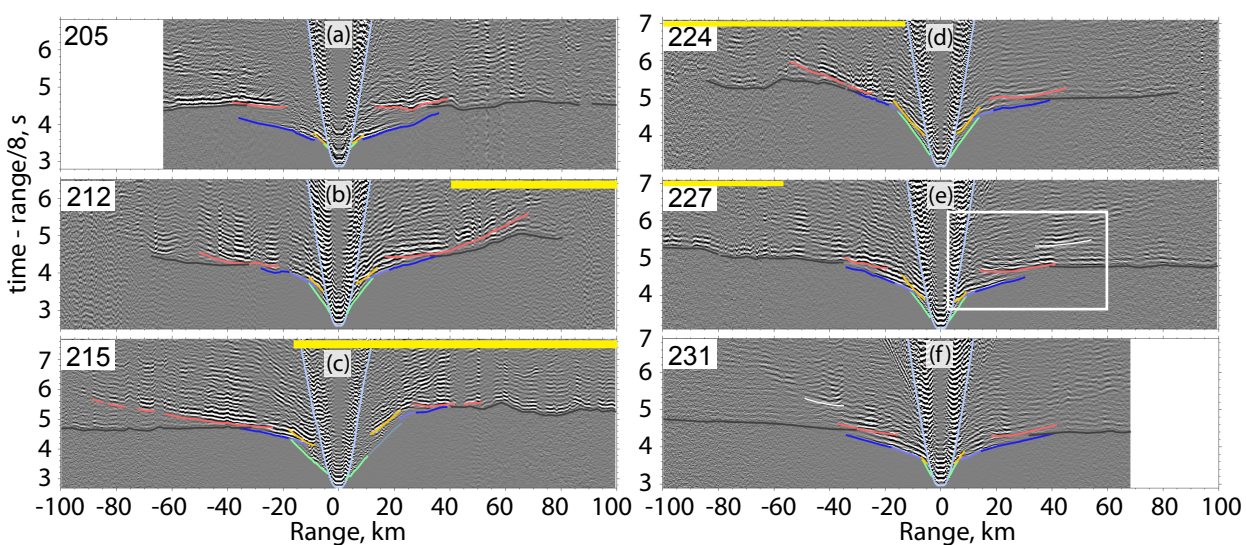


Figure 5.

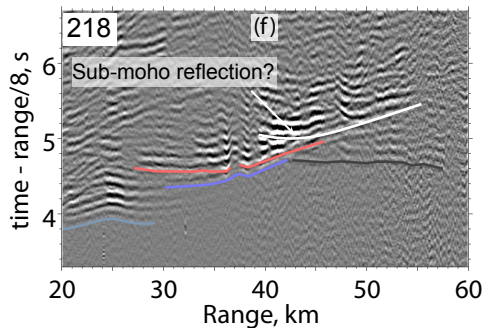
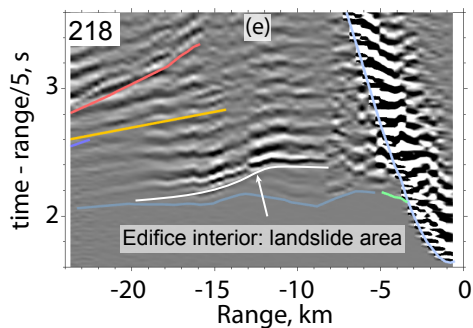
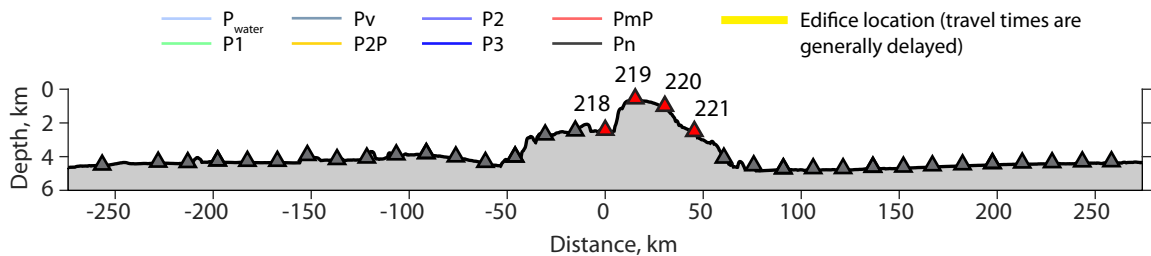
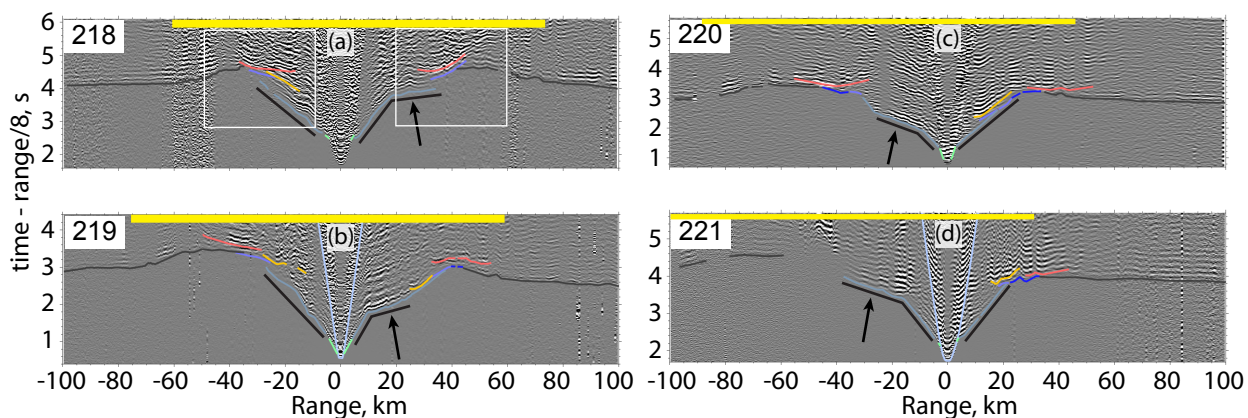


Figure 6.

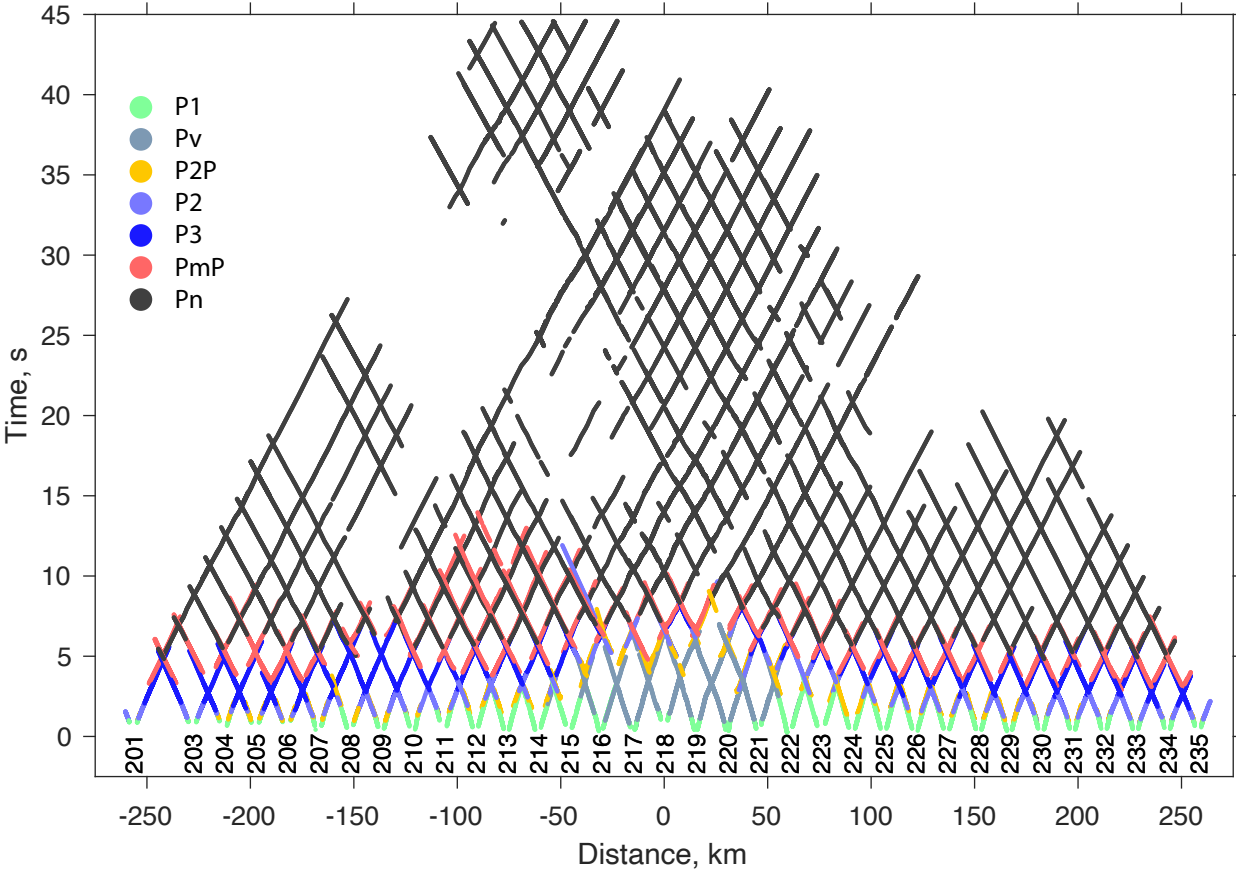


Figure 7.

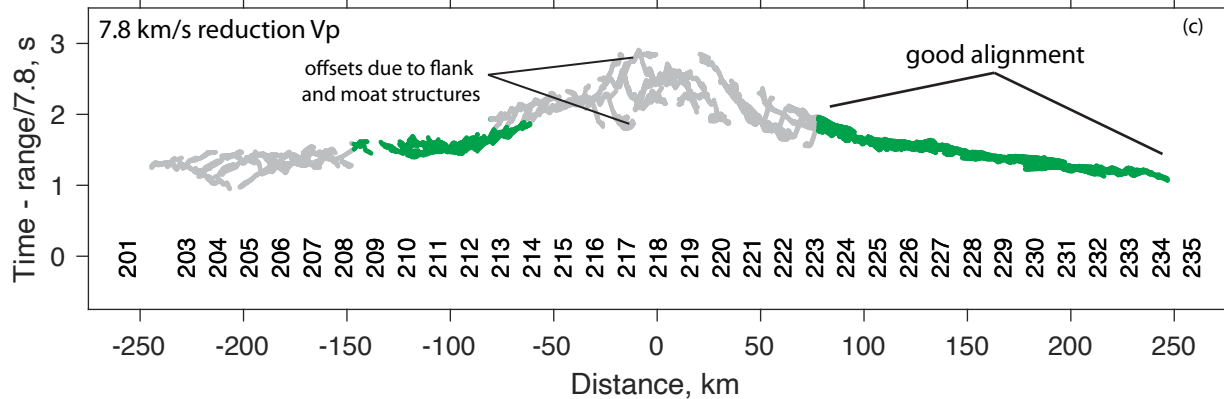
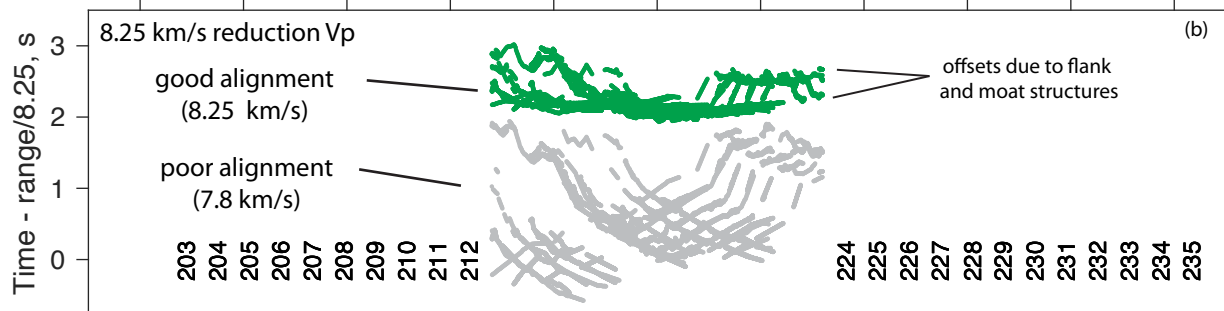
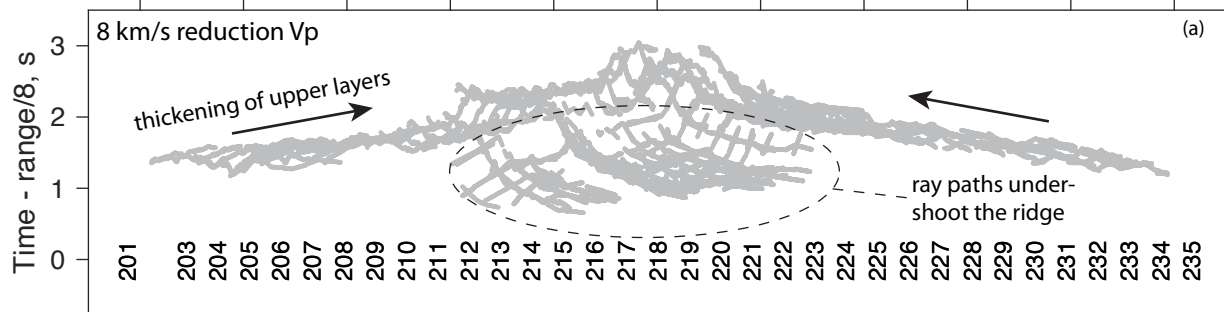


Figure 8.

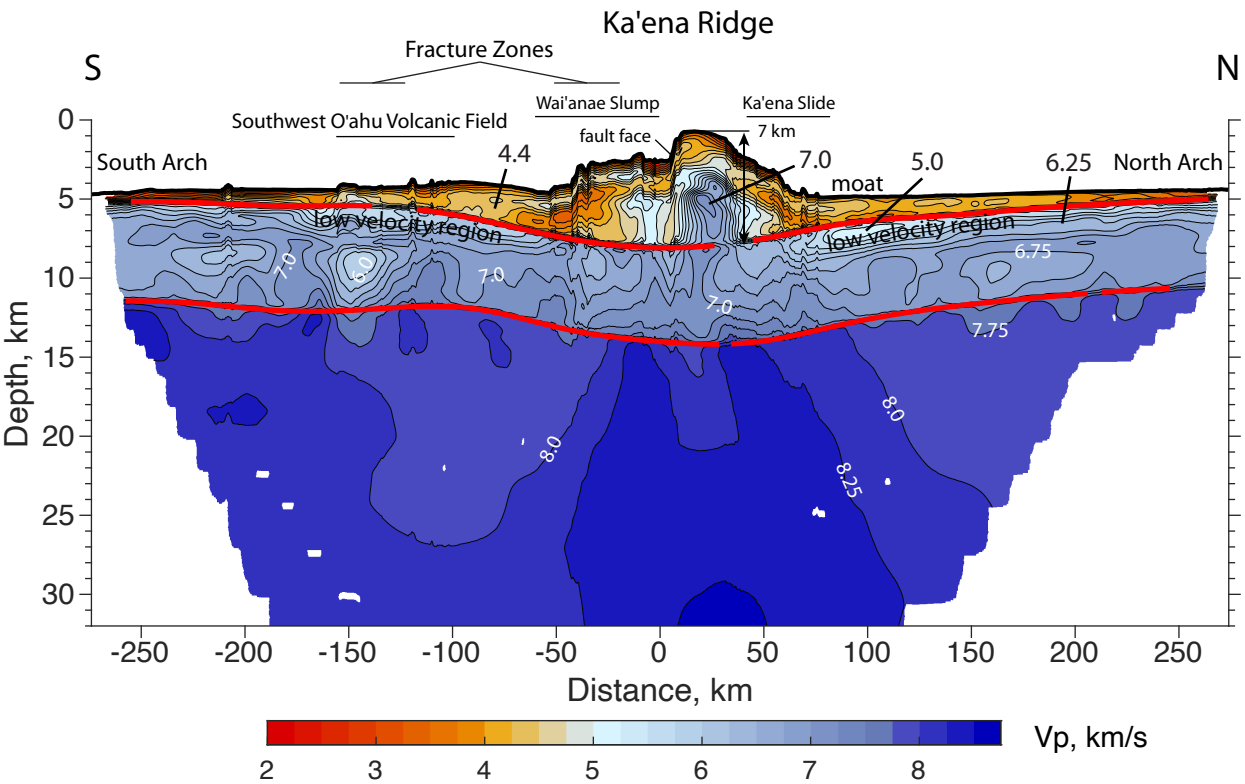
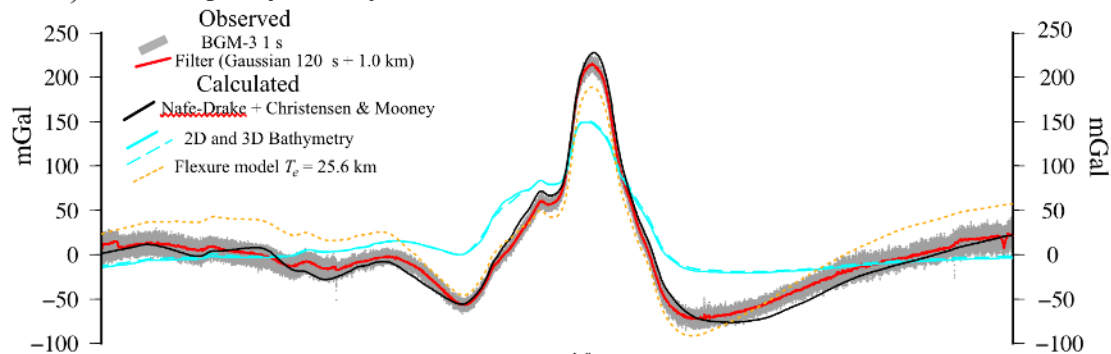


Figure 9.

a) Free-air gravity anomaly



b) Velocity (V_p) and Density model

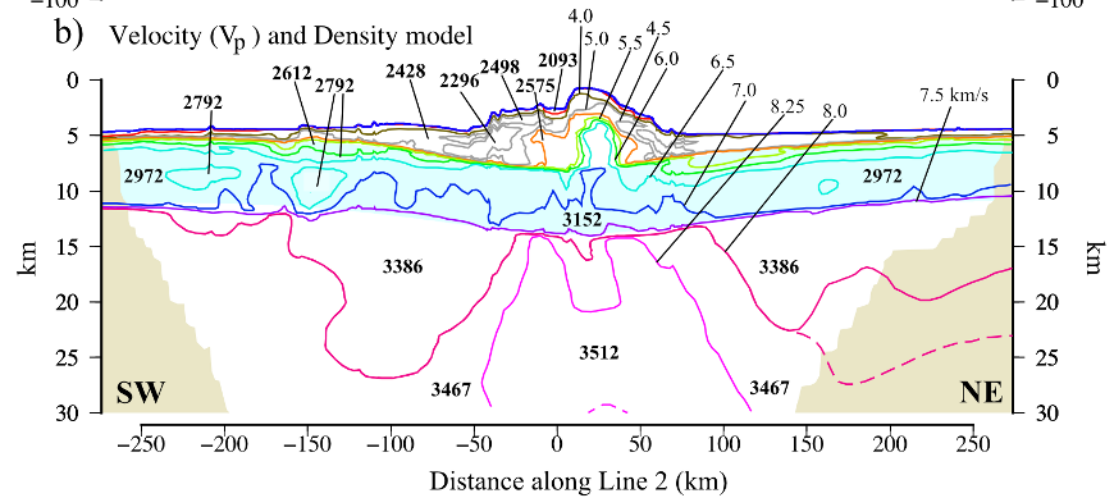


Figure 10.

Ka'ena Ridge

South

North

