- 1 Enhanced volcanic activity and long-term warmth in the middle Eocene revealed by
- 2 mercury and osmium isotopes from IODP Expedition 369 Site U1514
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### 16 Abstract

17 Rapid plate reorganization may have influenced global climate during the Eocene; however, its 18 linkage remains poorly constrained, particularly during the middle Eocene. To elucidate this tectonic-19 climatic relationship, here, we conducted a comprehensive analysis based on high-resolution mercury 20 (Hg) and osmium (Os) abundance and isotope data obtained from the complete Eocene sedimentary 21 sequence of Site U1514, drilled in the Mentelle Basin off southwest Australia. The Hg signals in this 22 sedimentary sequence, which are characterized by significantly high enrichment and insignificant mass-independent fractionation ( $\Delta^{199}$ Hg) signal, confirm that the middle Eocene (~45–38 Ma) was a 23 24 period of persistent, increased volcanism, accompanied by intense tectonic activity. In particular, a 25 remarkable seafloor volcanic eruption persisted for approximately 1.5 million years (~42.0-40.5 Ma), 26 immediately preceding the Middle Eocene Climate Optimum (MECO). Contemporaneously, the

trends toward a slightly more radiogenic seawater <sup>187</sup>Os/<sup>188</sup>Os (Osi) composition denote the 27 28 prevalence of intensified continental weathering under a warm, humid climate during the middle 29 Eocene, a phenomenon particularly evident during the MECO. Importantly, the Hg and Os records 30 from Site U1514 reveal the occurrence of a multi-million-year warming reversal amid the long-term 31 Eccene cooling trend, which likely contributed to significant CO<sub>2</sub> reduction during the late Eccene. 32 These findings significantly enhance our understanding of Eocene climate dynamics, which are 33 fundamentally linked to intensive tectonic-driven volcanic activity and associated continental 34 chemical weathering.

35 Keywords: Middle Eocene warmth; Hg and Os isotopes; volcanism; continental weathering; Site U1514

#### 36 **1. Introduction**

37 The Eocene was a globally significant period marked by extensive global plate reorganization 38 (Müller et al., 2016), as well as an Earth's climate shift from a warmhouse phase to a coolhouse 39 phase (Zachos et al., 2001). In the southern high latitudes, the early and middle Eocene was marked 40 by the extremely rapid separation of Australia from Antarctica through ridge extension and rifting on 41 the Kerguelen Plateau and Broken Ridge (Royer and Sandwell, 1989; Veevers, 2000; Borissova et al., 42 2010). This plate tectonic activity would have led to massive ridge/arc volcanism and prolonged CO<sub>2</sub> 43 outgassing (Jones and Fitzgerald, 1984; Rea et al., 1990 and references therein), which potentially 44 triggered global warming and associated perturbations in the land-ocean system (Wang et al., 2022). 45 However, most middle Eocene studies have concentrated primarily on specific warming periods, such 46 as the Middle Eocene Climate Optimum (MECO) (e.g., Bohaty and Zachos, 2003; Westerhold et al., 47 2018; van der Ploeg et al., 2018; Henehan et al., 2020; van der Boon et al., 2021), leaving gaps in our 48 understanding of the volcanic response to active tectonic movements during the middle Eocene and its environmental and climatic impacts. In addition, the mechanisms driving the MECO warming, 49

particularly the sudden rise in atmospheric CO<sub>2</sub>, are debated (Bohaty and Zachos, 2003; van der
Ploeg et al., 2018; Henehan et al., 2020).

52 A recent study proposed that the secular Cenozoic cooling trend in the southern mid-latitudes 53 was interrupted by a multi-million-year warmth, likely induced by large-scale volcanic eruptions 54 triggered by tectonic movements during the middle Eocene (Wang et al., 2022). The absence of 55 apparent cooling through the middle to late Eocene in subequatorial regions (Evans et al., 2018) also 56 raises the possibility of the global-scale middle Eocene warmth. The association of this long-term 57 warmth with intensified continental weathering may provide clues regarding the causal mechanism of 58 the drastic decline in atmospheric CO<sub>2</sub> across the late Eocene and the Eocene/Oligocene transition 59 (EOT) (DeConto and Pollard, 2003; Inglis et al., 2015). However, conclusive evidence for long-term 60 middle Eocene warmth in the form of sedimentary records linking active tectonic movements, 61 volcanism, and chemical weathering remains elusive. Therefore, it is necessary to re-examine the 62 long-term Eocene cooling trend that began at the Early Eocene Climatic Optimum (EECO). 63 Mercury (Hg) and osmium (Os) isotopes in sedimentary records are ideal indices of volcanic eruption and continental weathering intensity, respectively (e.g., Peucker-Ehrenbrink and Ravizza, 64 65 2000; Lu et al., 2017; Grasby et al., 2019). There is a growing agreement that Hg enrichment or 66 anomalies in sedimentary successions are potential signatures of large volcanic eruptions (Grasby et 67 al., 2016, 2019 and references therein), providing new insights into the relationship between volcanic activity and climate change (Sial et al., 2020 and references therein). Hg isotopes (e.g.,  $\Delta^{199}$ Hg), 68 which are unlikely to be altered by post-depositional processes, have also been adopted as indicators 69 70 of the volcanogenic origin of Hg enrichment (Blum et al., 2014; Grasby et al., 2019; Sial et al., 2020). The <sup>187</sup>Os/<sup>188</sup>Os isotope ratio of seawater reflects the mass balance between two Os sources: 71 72 unradiogenic Os ( $\sim 0.12$ ) from volcanic/hydrothermal inputs and radiogenic Os ( $\sim 1.4$ ) from

73 weathered continental crusts (Peucker-Ehrenbrink and Ravizza, 2000). Thus, records of age-corrected

74 seawater Os isotope composition have been utilized as an effective tool to identify the effects of 75 continental weathering feedback (Ravizza et al., 2001; Cohen et al., 2004; Dickson et al., 2015). 76 Nevertheless, the availability of combined Hg and Os isotope data covering the entire Eocene, which 77 could provide a broader understanding of the dramatic Eocene climate changes, is limited. In this 78 study, we reconstructed high-resolution Hg and Os abundance and isotope records from the Eocene 79 sedimentary sequence of Site U1514, drilled in the Mentelle Basin off southwestern Australia during 80 International Ocean Discovery Program (IODP) Expedition 369 (Fig. 1), to define the long-term 81 middle Eocene warmth and gain insights into the relationships among tectonics, volcanism, and 82 continental weathering during the middle Eocene.

#### 83 2. Samples and methods

84 During IODP Expedition 369, a continuous sedimentary sequence of Eocene was recovered at 85 Site U1514 (paleolatitude ~60~50°S, Fig. 1a) in the Mentelle Basin off southwest Australia (Fig. 1b). 86 The studied Eocene sedimentary interval at the site spans  $\sim 165$  m (70–235 m core composite depth 87 below sea floor, CCSF) of sediment and is dominated by light greenish gray clayey nannofossil ooze, 88 sponge spicule-rich clay, nannofossil-rich clay, and clayey nannofossil chalk (Fig. 2). In this study, a 89 total of 89 sediment samples from Site U1514 sequence, corresponding to Eocene-aged sediments of 90 ~34–50 Ma, were analyzed for total organic carbon (TOC), total sulfur (TS), total inorganic carbon 91 (TIC), and elemental abundances (i.e., Al, Fe, Mn, Mo, U, Hg, and Os), as well as Hg and Os isotopic 92 compositions.

93 2.1. Total contents of inorganic carbon, organic carbon, sulfur, and other elements

94 The total carbon (TC) and total sulfur (TS) contents were measured using an elemental analyzer

95 (FLASH 2000; Thermo Fisher Scientific, USA) and the total inorganic carbon (TIC) content was

96 measured using a CO<sub>2</sub> coulometer (CM5014; UIC, Joliet, IL, USA). The total organic carbon (TOC)

relative deviations between measured and certified values of standard reference materials (L-cysteine
for TC, BBOT for TS, and pure calcium carbonate for TIC), as well as their analytical reproducibility,
were less than 5%, indicating satisfactory recoveries. For the concentrations of elements (Al, Fe, Mn,
Mo, and U) for bulk sediments, each powdered sediment sample was fused with lithium metaborate
(LiBO<sub>2</sub>) flux and the molten beads were then poured into a volume of dilute nitric acid and stirred
until dissolved. The resultant solutions were then analyzed using a combination of a Thermo ICAP

content was calculated as the difference between the TC and TIC contents. For these components,

104 6500 radial inductively coupled plasma optical emission spectroscopy (ICP-OES) and Thermo

105 Elemental X Series II ICP mass spectrometry (ICP-MS). Calibration for both instruments was

106 achieved via matrix matched calibration standards produced from combinations of ICP-grade single

107 element standards. Standard reference material (SBC-1), was analyzed together with a batch of

sediment samples and the relative deviations between measured and certified values were less than 5-10 %.

# 110 2.2. Mercury concentration and a three-step sequential extraction

97

The mercury (Hg) concentrations of the sediment samples were determined using a mercury analyzer with thermal decomposition, amalgamation atomic absorption spectroscopy module (Hydra-C, a detection limit of 0.005 ng Hg; Teledyne Leeman Labs, Hudson, NH, USA) based on the US EPA method 7473. The analysis error, determined by repeated measurements of a standard reference material (MESS-3) together with a batch of sediment samples, was less than 3%, indicating satisfactory data acquisition. Hg enrichment factors (EF<sub>Hg</sub>) were calculated as:

117 
$$EF_{Hg} = (Hg/TOC)_{sample}/(Hg/TOC)_{bg}$$

118 where (Hg/TOC)<sub>sample</sub> and (Hg/TOC)<sub>bg</sub> represent the normalized Hg concentration of an individual

sample and a reference background (bg) interval (i.e., early and late Eocene sediments for Site

120 U1514). The median Hg/TOC ratio was assigned as its reference background value; note that the 121 median was used in preference to the mean owing to the sensitivity of the latter parameter to outlier 122 values.

123 To better understand factors controlling of THg concentrations, a three-step sequential 124 extraction scheme based on the procedures of Burt et al. (2003) was used to divide the carbonate- and 125 Fe-Mn hydroxide-bound fraction (Fe-Mn hydroxide-bound Hg), organic matter- and sulfide-bound 126 fraction (organic/sulfide-bound Hg), and residual faction (final residue-Hg). More detailed 127 description for the sequential extraction was addressed in Kim et al. (2022). 128 2.3. Hg isotopes 129 All analytical procedures and measurements for Hg isotopes were performed at the isotope 130 laboratory of Institute of Geochemistry, Chinese Academy of Sciences, Guiyang, China. A more 131 detailed description for the analytical procedure can be found in Yin et al. (2016). In brief, the 132 samples were prepared for Hg isotope analysis using a double-stage tube furnace coupled with 40% 133 anti aqua regia (HNO<sub>3</sub>/HCl = 2/1, v/v) trapping solutions. The solutions were diluted to 0.5 ng/mL

134 with an acid concentration of 10–20% prior to Hg isotope analysis using a Neptune Plus multi-

135 collector ICP-MS.

Mass dependent fractionation (MDF) of Hg isotopic results are expressed as delta (δ) values in
units of per mille (‰) variation relative to the bracketed NIST SRM 3133 Hg standard, as follows:

138 
$$\delta^{202} Hg = [({}^{202} Hg/{}^{198} Hg)_{sample}/({}^{202} Hg/{}^{198} Hg)_{standard} - 1] \times 1000$$

Hg-isotopic values that did not follow the theoretical MDF were considered to exhibit isotopic
anomalies caused by mass independent fractionation (MIF) (Bergquist and Blum, 2007). MIF values
were calculated for <sup>199</sup>Hg and ex-pressed as per mille deviations from the predicted values based on
the MDF law:

143  $\Delta^{199} Hg \approx \delta^{11}$ 

 $\Delta^{199} Hg \approx \delta^{199} Hg - \delta^{202} Hg \times 0.252$ 

The analytical uncertainties of Hg isotopic compositions were evaluated by repeated analysis of the isotopic compositions of NIST-3177 (n = 6) and GSS-4 (n = 6). The overall average and uncertainty of NIST-3177 ( $\delta^{202}$ Hg: -0.55 ± 0.09‰;  $\Delta^{199}$ Hg: -0.02 ± 0.05‰, 2SD), and GSS-4 ( $\delta^{202}$ Hg: -1.70 ± 0.07‰;  $\Delta^{199}$ Hg: -0.46 ± 0.02‰, 2SD) agree with previously published results (Yin et al., 2022; Chen et al., 2022).

## 149 2.4. Re-Os abundance and isotope analytical protocol

150 All rhenium (Re) and osmium (Os) isotope analyses were performed at the Durham 151 Geochemistry Centre, Durham University (UK). A more detailed description of the analytical 152 procedure can be found in van der Ploeg et al. (2018). In brief, powdered samples of approximately  $\sim$ 1 g were loaded into Carius tubes with a known amount of mixed <sup>185</sup>Re + <sup>190</sup>Os tracer solution 153 154 (spike), and 8 mL of CrO<sub>3</sub>-H<sub>2</sub>SO<sub>4</sub> solution, sealed and then heated in an oven at 220 °C for 48 hrs. 155 Osmium was isolated from the CrO<sub>3</sub>–H<sub>2</sub>SO<sub>4</sub> sample solution by using solvent extraction (CHCl<sub>3</sub>), 156 and then back extracted by hydrobromic acid (HBr), and then further purified through micro 157 distillation. Rhenium was isolated from the Os extracted CrO<sub>3</sub>-H<sub>2</sub>SO<sub>4</sub> sample solution by solvent 158 extraction using sodium hydroxide (NaOH) and acetone (C<sub>3</sub>H<sub>6</sub>O), and anion chromatography. 159 Following purification, the Re and Os fractions were loaded onto Ni and Pt filaments, respectively, 160 together with 0.5 µL BaNO<sub>3</sub> and BaOH activator solutions, respectively. Rhenium and Os isotope 161 ratios were determined by negative thermal ionization mass spectrometry, using Faraday cups for Re 162 and a Secondary Electron Multiplier for Os in peak-hopping mode on a ThermoScientific Triton mass 163 spectrometer.

164 The  ${}^{187}\text{Os}/{}^{188}\text{Os}_{initial}$  ratios (Os*i*) were calculated by correcting for post-depositional  ${}^{187}\text{Re}$  decay 165 over time with the following equation:

166 
$${}^{187}\text{Os}/{}^{188}\text{Os}_{\text{initial}} (\text{Os}_i) = {}^{187}\text{Os}/{}^{188}\text{Os}_{\text{measured}} - {}^{187}\text{Re}/{}^{188}\text{Os}_{\text{measured}} \times (e^{\lambda t} - 1)$$

where  $\lambda$  is the <sup>187</sup>Re decay constant (1.666  $\cdot$  10<sup>-11</sup> yr<sup>-1</sup>) and t is ages of the studied samples for Site 167 U1514. Procedural blanks were  $9.6 \pm 1.0$  pg for Re and  $0.04 \pm 0.05$  pg for Os, with a <sup>187</sup>Os/<sup>188</sup>Os 168 ratio of  $0.20 \pm 0.05$  (1SD; n = 4). The <sup>187</sup>Re/<sup>188</sup>Os and <sup>187</sup>Os/<sup>188</sup>Os uncertainties (2SE) include full 169 170 propagation of uncertainties in weighing, mass spectrometer measurements, spike calibrations, blank 171 corrections, and reproducibility of standards. Analytical precision for the lab was monitored through 172 repeated analysis of 50 pg DROsS and 125 pg ReSTD solution standards, which yield a running average of  $0.1608 \pm 0.0006$  (1SD; n = 916) for <sup>187</sup>Os/<sup>188</sup>Os and  $0.5986 \pm 0.0015$  (1SD; n = 784) for 173 <sup>185</sup>Re/<sup>187</sup>Re. 174

## 175 **3. Results**

176 The chronostratigraphic framework of the studied interval at IODP 369 Site U1514 was 177 constrained by linear interpolation between age points obtained from shipboard biostratigraphy and 178 magnetostratigraphy studies (Fig. 2). More information about age-depth model for Site U1514 179 section is given in Wang et al. (2022). Linear sedimentation rates were much higher (1.20–1.85 180 cm/kyr) during the middle Eocene (~45–38 Ma) than during the early and late Eocene (0.40–0.90 181 cm/kyr). All analytical results of this study are summarized in Supplementary Tables S1-S4. The 182 TOC and TS contents were less than ~0.30% and ~0.11%, respectively, with limited variation. The 183 Al, Fe, Mn, Mo, and U concentration ranges were 10.0–52.7 mg/g, 7.2–34.6 mg/g, 140–1403  $\mu$ g/g, 184 0.09–1.54 µg/g, and 0.18–1.45 µg/g, respectively. No distinct variation in Al, Fe, Mn, Mo, and U 185 concentrations was observed across the studied interval (Supplementary Table S2). Total Hg 186 concentrations in the Eocene sediments of Site U1514 ranged from  $\sim 1$  to 246 ng/g, with a significant 187 increase during the middle Eocene, featuring peaks exceeding 100 ng/g (Fig. 3a). Similar temporal 188 trends were observed in Al-, TIC-, TOC-, and TS-normalized Hg concentrations (Fig. 3b, c).

189 Sequential extraction data showed that sedimentary Hg was present primarily in organic matter

190 fraction (range: 46–89% of the total concentration, average:  $64 \pm 11\%$ ), followed by the Fe-Mn

191 hydroxide-bound (range: 6-39%, average:  $22 \pm 9\%$ ) and residual fractions (range: 4-23%, average:

192  $14 \pm 6\%$ ) (Supplementary Fig. S1).

Hg mass-dependent fractionation ( $\delta^{202}$ Hg) values from Site U1514 sediments ranged from 193 194 -1.74 to +0.86 ‰, with the majority falling between -1.15 and -0.11‰ (Supplementary Table S3). 195 In general, Hg undergoes a series of complex physical, chemical, and biological transformations in 196 the environment, potentially leading to ambiguous interpretations (Blum et al., 2014). Therefore, we 197 did not interpret  $\delta^{202}$ Hg record and instead focused on Hg mass-independent fractionation ( $\Delta^{199}$ Hg) 198 signals, providing clearer insights into the sources of Hg due to their conservation through postdepositional processes. The  $\Delta^{199}$ Hg values from Site U1514 sediments ranged from +0.04 to +0.25‰ 199 200 (Fig. 3d), indicating variation several times greater than the analytical uncertainty (2SD = 0.05). With 201 the exception of samples from specific periods with high Hg concentration peaks, exceeding ~100 ng/g, during the middle Eocene,  $\Delta^{199}$ Hg values exhibited limited variability, typically ranging 202 203 between +0.15 and +0.23‰, suggesting a relatively constant background level (+0.18  $\pm$  0.05‰) 204 across the studied sedimentary section. However, the middle Eocene strata exhibited abrupt and significant drops towards near-zero  $\Delta^{199}$ Hg values, particularly in samples associated with high Hg 205 206 concentration peaks (Fig. 3d).

207 Rhenium (Re) and Os abundance and isotopic composition varied considerably in the studied 208 samples. The Re and total Os, and <sup>192</sup>Os (the "common" Os component) concentration ranges were 209 0.03–1.28 ng/g, 22.61–167.77 pg/g, and 8.89–65.77 pg/g, respectively (Supplementary Table S4),

with a significant increase occurring during the middle Eocene (Fig. 3e). The  $^{187}$ Re/ $^{188}$ Os and

<sup>187</sup>Os/<sup>188</sup>Os ratios varied from 2.07 to 76.95 and 0.47 to 0.59, respectively (Supplementary Table S4).

212 Initial seawater <sup>187</sup>Os/<sup>188</sup>Os (Os<sub>i</sub>) values calculated using the assigned ages of each sample ranged

213	from 0.47 to 0.57, featuring a significant shift toward more radiogenic composition between ~43 and
214	38 Ma, with maximum values at ~40.5–40 Ma (Fig. 3f). The early and late Eocene samples exhibited
215	relatively less radiogenic characteristics, with $Os_i$ values of 0.47–0.50; however, these values
216	increased in the middle Eocene samples, ranging from 0.53 to 0.57. Although the magnitude of the
217	positive $Os_i$ shift at Site U1514 is small (~0.05), it exceeds the maximum analytical uncertainty (2SE
218	= 0.005) by a factor of ~10. The Os <sub>i</sub> values from Site U1514 were only moderately radiogenic, but
219	significantly lower than those of present-day seawater (i.e., ~1.04–1.07, Peucker-Ehrenbrink and
220	Ravizza, 2000). The Os <sub>i</sub> record from Site U1514 is in good agreement with data from CD29-2
221	(Klemm et al., 2005), Pacific manganese crust (Burton, 2006), as well as ODP Site 865, in terms of

values and broad temporal trends (Fig. 3f, Supplementary Fig. S2, see Fig. 1a for sample locations).

## 223 **4. Discussion**

#### 224 4.1. Enhanced middle Eocene volcanism

225 Large volcanic eruptions in geological history have often led to anomalously high Hg 226 enrichment or spikes in stratigraphic records, providing insights into their timing, duration, and 227 intensity (e.g., Sial et al., 2016; Percival et al., 2018; Grasby et al., 2019; Shen et al., 2019, 2023). 228 Accurate estimation of background Hg levels is a prerequisite for deepening our understanding of the 229 distinctive features and nuanced variations in Hg records from the geological past, particularly 230 because Hg anomalies may not always provide definitive signatures of enhanced volcanic emissions 231 (Grasby et al., 2019 and references therein). The middle Eocene interval (~45-38 Ma) in Site U1514 232 geological record exhibits significantly elevated Hg concentrations, reaching ~250 ng/g with peaks 233 exceeding 100 ng/g around 42–40.5 Ma (Fig. 3a). These concentrations markedly exceed the 234 background level of 11 ng/g (i.e., the average for early and late Eocene intervals) in the Site U1514 235 section (Fig. 3a), the average value of 50 ng/g for the upper continental crust (Rudnick and Gao,

2014), and the mean Hg value in sediments from several extinction and oceanic anoxic event zones
(62.4 ng/g, Grasby et al., 2019), suggesting substantial Hg inputs during the middle Eocene.

- 238 Furthermore, Hg concentrations during the middle Eocene, particularly in the period of 42–40.5 Ma,
- greatly surpass the organic-matter drawdown capacity limit (e.g.,  $Hg = 48.5TOC^{0.89}$  or  $150TOC^{0.89}$ ,
- 240 Grasby et al. 2019), potentially supporting increased Hg loading during this period.

241 To further clarify the cause of elevated Hg concentrations observed in Site U1514 sedimentary 242 records, it is necessary to normalize the data to the dominant Hg sequestration factor (Shen et al., 243 2020). Hg concentrations for Site U1514 sediments are primarily controlled by organic matter, as shown by the strong positive correlation between Hg and TOC ( $r^2 = -0.5$ ), but weak or no correlation 244 245  $(r^2 < 0.2)$  with TIC, TS, and Al concentrations, and Fe/Al and Mn/Al ratios (Supplementary Fig. S3). 246 This finding was further supported by the predominance of organic/sulfide-bound Hg ( $64 \pm 11\%$  of the total concentration), which was significantly correlated with TOC ( $r^2 = -0.5$ ), but not with TS 247 248 (Supplementary Fig. S1), indicating that organic matter played an essential role in the Hg 249 depositional pathway at Site U1514. Redox conditions played a limited role in Hg enrichment in the 250 Site U1514 sedimentary sequence, as indicated by low sulfur content (< 0.1%) and non-significant 251 correlations with redox elemental proxies (TS, Mo/Al, and U/Al) (Supplementary Fig. S3). Similar 252 temporal trends were observed in Al-, TIC-, TOC-, and TS-normalized Hg concentrations (Fig. 3b, c), 253 suggesting that local depositional environments (e.g., grain size and sedimentation rate) are unlikely 254 to be the primary factors controlling variation in Hg deposition at Site U1514. As such, the increased 255 environmental Hg loading at Site U1514 can be estimated by values normalized to organic matter 256 content (i.e., Hg/TOC). Site U1514 exhibited significantly elevated values of Hg/TOC (~200-760 ng/g/wt%), Hg enrichment factor (EF<sub>Hg</sub>, ~4–16), and Hg flux (5–25 ×  $10^2 \mu g/m^2/kyr$ ) throughout the 257 258 middle Eocene (Fig. 4a-c). The Hg/TOC values greatly exceeded the background level of 96 259 ng/g/wt% (i.e., the average of samples with TOC > 0.2% for early and late Eocene intervals) in Site

260 U1514 record and the anomalous enrichment threshold (72 ng/g/wt%) derived from sedimentary 261 records of major extinction and oceanic anoxic events (Grasby et al., 2019). This Hg increase during 262 the middle Eocene was further supported by elevated  $EF_{Hg}$  values (Fig. 4b), where  $EF_{Hg} > 2$  indicates 263 significant Hg enrichment, implying a volcanic source of Hg in geologic successions (e.g., Shen et 264 al., 2019). These results indicate that excess Hg inputs during the middle Eocene were likely driven by volcanism rather than increased organic matter flux (i.e., organic-matter Hg drawdown) and/or 265 266 increased runoff. Notably, we identified an intense volcanic episode lasting for 1.5 Myr (42.0-40.5 267 Ma) and three shorter periods of volcanic activity at ~45.5, ~44, and ~39 Ma (Fig. 4a-c). 268 Conclusive evidence for such a volcanic Hg source was provided by paired Hg enrichment and isotope measurement (Fig. 5). Site U1514 sediments yielded  $\Delta^{199}$ Hg values ranging from +0.04 to 269 270 +0.25% (0.16  $\pm$  0.06‰), with several peaks showing low values of up to +0.04% in the middle 271 Eccene (Fig. 4d), comparable to that (-0.30 to +0.27%) derived from extinction event records 272 (Grasby et al., 2019). Generally, geogenic Hg sources exhibit negligible Hg mass-independent fractionation signatures ( $\Delta^{199}$ Hg ~0%) (Sherman et al., 2009). However, recent studies proposed that 273 small, positive  $\Delta^{199}$ Hg values in deep water environments may result from the deposition of Hg<sup>2+</sup> 274 275 absorbed from the atmosphere by volcanic plume particles rather than direct atmospheric deposition 276 (Gong et al., 2017; Grasby et al., 2017, 2019). Such deposition is also likely in the deep-water basin of Site U1514, which displayed limited  $\Delta^{199}$ Hg variability, except for Hg isotope spikes, and an 277 278 overall slightly positive background value ( $+0.18 \pm 0.05\%$ ). Similar Hg signatures for dominant 279 volcanic Hg enrichments are reported for the latest Permian extinction and Ordovician-Silurian extinction successions (Gong et al., 2017; Grasby et al., 2017). Interestingly,  $\Delta^{199}$ Hg values at Hg 280 peaks (i.e., ~45.5, ~44, ~39, and ~42.0-40.5 Ma) of the middle Eocene showed a significant shift 281 282 toward values close to zero (Figs. 4, 5), indicating elevated Hg loading, likely attributable to 283 increased seafloor volcanic activity, such as ridge volcanism and/or hydrothermal activity, which

284 typically exhibit near-zero  $\Delta^{199}$ Hg values (Zambardi et al., 2009; Kim et al., 2022). Although such shifts in  $\Delta^{199}$ Hg can result from a mix of terrestrial (highly negative  $\Delta^{199}$ Hg) and seawater-sourced 285 Hg (positive  $\Delta^{199}$ Hg) (Sial et al., 2020; Shen et al., 2022), these sources are unlikely to explain the 286 identical trend between  $\Delta^{199}$ Hg and Hg enrichments and their negative correlation in this study (Fig. 287 288 5). Hg concentrations also showed no significant relationship with terrestrial-indicative Al content 289 (Supplementary Fig. S3). Therefore, the Hg isotopic signals at Site U1514 indicate a sustained 290 increase in widespread subaerial volcanism during the middle Eocene, with notable periods 291 (particularly 42.0–40.5 Ma) marked by intensive pulses of seafloor volcanic/hydrothermal eruptions. 292 This conclusion is further supported by the higher deposition of volcanic materials in the Mentelle 293 Basin during the middle Eocene (43–38 Ma) (Wang et al., 2022).

294 Although there are no candidates for a large igneous provenance event in the geological record 295 coincident with the middle Eocene (Sluijs et al., 2013; Ernst et al., 2021), increased volcanism 296 appears to be associated with active plate tectonic reorganization (e.g., the rifting of Australia from 297 Antarctica, the resumption of Pacific subduction, and the closing of Neotethys), and large igneous 298 activity in southeastern Australia as well (Fig. 4f). Significant increase in arc/mid-ocean ridge 299 volcanism and hydrothermal activity during the middle Eocene (~35-45 Ma) have been also 300 identified in the North Atlantic (Rea et al., 1990) and in various other regions (van der Ploeg et al., 301 2018). Both volcanism and hydrothermal activity would have been much more intense during 302 continental rifting and plate boundary rearrangements of middle Eocene (Jones and Fitzgerald, 1984; 303 Rea et al., 1990 and references therein).

304 *4.2. Volcanic climaxes and short-lived warming events* 

305 The Eocene epoch was distinguished by extensive global plate reorganization, including the306 abruptly accelerated separation of Australia and Antarctica at ~45–43 Ma and rifting of the Kerguelen

307	Plateau and Broken Ridge caused by the onset of seafloor spreading at the Southeast Indian Ridge at
308	~43-40 Ma (Veevers, 2000; Li et al., 2003). This tectonic activity would have been accompanied by
309	extensive seafloor volcanism/hydrothermal activity in the Mentelle Basin and on the Naturaliste
310	Plateau (Jones and Fitzgerald, 1984; Rea et al., 1990 and references therein; Borissova et al., 2010),
311	as well as at the Ninety East Ridge (Fleet and McKelvey, 1978). Igneous activity in southeastern
312	Australia peaked between 45 and 37 Ma (Fig. 4f) and around Antarctica at ~45 Ma (Wang et al.,
313	2009). Major volcanic climax records (i.e., 45.5, 44, and 42.0-40.5 Ma) interpreted from Site U1514
314	should accurately depict such a tectonic history in the Southern Ocean. Notably, the series of intense
315	volcanic eruptions that persisted for ~1.5 Myr (42.0-40.5 Ma) coincided with the period of greatest
316	separation between Australia and Antarctica. Additionally, these eruptions correspond to the warming
317	event known as the Late Lutetian Thermal Maximum (LLTM, ~41.5 Ma), which is mainly
318	documented in high-latitude Atlantic sites (Westerhold et al., 2018; Rivero-Cuesta et al., 2020). High
319	insolation forcing has been thought to underlie the LLTM (Westerhold et al., 2018 and reference
320	therein); however, our findings suggest that the greatly increased volcanic eruptions in the Southern
321	Ocean during this period are a plausible alternative mechanism for this transient warming.
322	The MECO, a transient global warming event that occurred at ~40.5–40 Ma in association with
323	an abrupt rise in atmospheric CO <sub>2</sub> (Sluijs et al., 2013), was first recognized through a ~1‰ negative
324	oxygen isotope excursion in bulk carbonate and benthic foraminiferal tests in Southern Ocean cores
325	(Bohaty and Zachos, 2003), and subsequently identified in the Atlantic Ocean and the central western
326	Tethys Ocean (Boscolo Galazzo et al., 2014 and references therein). Although a causal relationship
327	between volcanism and MECO warming has been suggested for some areas (e.g., van der Ploeg et
328	al., 2018; van der Boon et al., 2021 and references therein), a plausible source of excess volcanic CO <sub>2</sub>
329	remains to be identified, particularly in the high southern latitudes. In the Site U1514 record, the
330	potential MECO period was identified by pronounced negative excursions in the $\delta^{13}$ C value (Wang et

al., 2022) (Fig. 4e). Hg enrichment remained high throughout the MECO period; however, there were

332 no notable Hg peaks or isotope spikes corresponding to the period of negative carbon excursions,

implying the absence of catastrophic explosive volcanic eruptions (Fig. 4). Another likely

334 explanation for the absence of MECO Hg spikes in this area is dilution by the highest influx of

335 terrigenous materials during this period, revealed by Os<sub>i</sub> isotope signals with an abrupt increase

during the MECO at Site U1514 (discussed below). The intensified continental weathering suggests

337 significant rise in temperature and atmospheric CO<sub>2</sub> concentrations during this interval (Wang et al.,

338 2022). Nevertheless, the MECO was likely triggered by an abrupt, massive volcanic climax that

339 persisted for ~1.5 Myr (42.0 and 40.5 Ma) prior to its onset. These data support previous hypotheses

340 (e.g., Bohaty and Zachos, 2003) that enhanced volcanic activity associated with plate tectonics in the

341 Southern Ocean could explain contradictory global warming events (e.g., short-lived pulses of CO<sub>2</sub>)

342 during the middle Eocene (i.e., LLTM and MECO).

#### 343 4.3. Increased continental weathering and long-term middle Eocene warmth

344 The overall enhanced long-term volcanic activity during the middle Eocene, observed from Hg 345 proxy records from Site U1514 (Fig. 4), should have resulted in massive pCO<sub>2</sub> emissions, with 346 subsequent warming and environmental perturbations (e.g., Percival et al., 2016). This volcanism-347 climate warming feedback can be explained by the facilitation of continental chemical weathering by 348 elevated CO<sub>2</sub> levels and a warm and humid climate, i.e., the silicate weathering thermostat hypothesis (Cohen et al., 2004; Dickson et al., 2015; Penman et al., 2020). Initial seawater <sup>187</sup>Os/<sup>188</sup>Os ratio (Os<sub>i</sub>) 349 350 has been used as a powerful tool for reconstructing continental weathering changes in geological 351 successions involving hyperthermal climate events (e.g., ocean anoxic events, Paleocene-Eocene 352 Thermal Maximum, and Triassic-Jurassic transition) (Peucker-Ehrenbrink and Ravizza, 2000; Du 353 Vivier et al., 2014; Dickson et al., 2015 and references therein; Percival et al., 2016; Shen et al.,

2023). These studies are ultimately based on the proportional mixing of fluxes of two end-member Os isotope components: terrestrial ( $^{187}Os/^{188}Os$  of ~1.4) and volcanic–hydrothermal inputs ( $^{187}Os/^{188}Os$  of ~0.13). The Os<sub>i</sub> record from Site U1514 exhibited a significant positive shift toward more radiogenic values during the middle Eocene (~43.5–38 Ma), with enhanced volcanism (Figs. 3f, 6a). This overall Os<sub>i</sub> increase during the middle Eocene was also observed in sediment and manganese crust records from other sites, indicating that it is a global phenomenon (Supplementary Fig. S2).

361 Considering the relative invariability of both the Os<sub>i</sub> record and <sup>192</sup>Os abundance data (Fig. 3), 362 the balance of Os fluxes to the oceans and its uptake in organic matter did not change appreciably 363 over time. Thus, the Osi increase over the middle Eocene indicated in Site U1514 record may have 364 been caused by decreased volcanic/hydrothermal activity and/or increased weathering of felsic 365 radiogenic continental rocks. Given that Hg proxy signals at Site U1514 provide compelling evidence 366 of a significant increase in volcanism during the middle Eocene, the increase in Os<sub>i</sub> is more easily 367 reconciled with a notable increase in continental silicate weathering. This suggestion is corroborated 368 by a notable increase in terrigenous inputs from the Australian continent during the middle Eocene, 369 indicated by a higher mass accumulation rate of siliciclastic fractions (Fig. 6b). Therefore, this 370 definite Os<sub>i</sub> shift during the middle Eocene reflects the sequestration of hydrogenous Os derived from 371 the continent as a result of high weathering rates (Peucker-Ehrenbrink and Ravizza, 2000; Cohen, 372 2004; Percival et al., 2016). An analogous increase in radiogenic Os was also apparent in another 373 geological period (Oceanic Anoxic Event 1a) around the study area (ODP Site 763), which has been 374 interpreted as a sign of enhanced continental weathering triggered by global warming resulting from 375 outgassing during volcanic episodes at the Kerguelen Plateau (Matsumoto et al., 2022). 376 Based on a simple mixing model calculation (Dickson et al., 2015), radiogenic Os fluxes

377 increased by  $\sim 10-30\%$  during the middle Eocene, indicating a coeval increase in the supply of

378 terrigenous materials into the study area. The large increase in the mass accumulation rate of the 379 siliciclastic fraction during this period from  $\sim 0.4$  to 1.2 g/cm<sup>2</sup>/kyr supports our conclusion (Fig. 6b). 380 At Site U1514, strengthened chemical weathering and leaching of source rocks were also 381 corroborated by changes in indicators of the relative intensity of chemical weathering, i.e., an abrupt 382 increase in the kaolinite/smectite ratio and a decrease in the Mg/Al ratio between 43 and 38 Ma (Fig. 383 6c). The presence of subtropical to tropical rainforests in southwestern Australia during the middle 384 Eccene implies that there was considerable continental precipitation and runoff (Martin, 2006), which 385 likely facilitated sufficient hydrolysis for chemical weathering. Accordingly, the compelling evidence 386 in records of Os<sub>i</sub>, elemental indices, terrigenous input fluxes, and climate conditions indicates the 387 intensification of continental weathering during the middle Eocene on the Australian continent, 388 possibly driven by a warm, wet climate (Wang et al., 2022). Notably, the overall increase in Osi 389 values during the middle Eocene at other sites (Supplementary Fig. S2) indicates a global 390 enhancement in continental weathering during this period, despite slight differences in absolute Osi 391 values between sites, possibly due to differences in coastal proximity and oceanographic setting 392 (Paquay and Ravizza, 2012; Du Vivier et al., 2014). Globally enhanced chemical weathering during 393 the middle Eocene has also been inferred from lead ( $^{206}Pb/^{207}Pb$ ) and neodymium ( $\epsilon$ Nd) isotope 394 records (Burton, 2006).

Our analysis of Hg and Os isotope records, along with previously reported elemental and Sr isotope data (Wang et al., 2022), verify a distinct, prolonged middle Eocene warmth that persisted for ~5 Myr (~43–38 Ma, Fig. 6) in the southern mid-high latitudes. This long-term warmth period, possibly triggered by plate tectonics–volcanism–rising CO<sub>2</sub> feedback, challenges the prevailing notion of a continuous Eocene cooling phase following the early Eocene. Nevertheless, this middle Eocene warming reversal is reconciled with overall increasing trends in global deep-sea water and Austral high latitude surface water temperatures (Fig. 6d), as well as the absence of cooling 402 throughout the Eocene in tropical regions (Evans et al., 2018). This finding implies that long-term 403 Eocene cooling occurred at higher latitudes than previously thought, such as pole regions. The overall 404 intensified continental weathering during this global warming period may have contributed to the 405 prolonged reduction of atmospheric CO<sub>2</sub>, in turn potentially leading to late Eocene cooling and/or 406 EOT glaciation (DeConto and Pollard, 2003; Anagnostou et al., 2016; Cramwinckel et al., 2018). 407 Thus, our findings support tectonic-induced CO<sub>2</sub> forcing, rather than changes in ocean circulation 408 through ocean gateway reorganization or paleogeographic change, as the main driver of Eocene 409 climate (Inglis et al., 2015; Cramwinckel et al., 2018).

## 410 *4.4. The MECO pulse in continental weathering*

411 Close links among volcanism, continental chemical weathering, and climate warming were also 412 evident in the MECO interval from Site U1514, with the greatest shift toward more radiogenic Osi 413 compositions (Fig. 6a). The highest radiogenic Os signal during this interval signifies that the influx 414 of weathered terrestrial material in response to the inferred CO<sub>2</sub> rise and warming was most 415 prominent during the MECO period, which is in good agreement with predictions based on the 416 silicate weathering thermostat hypothesis. In particular, a coeval warming phase with increased sea 417 surface temperatures (~28°C during the peak of the MECO event) was identified on the Tasman 418 Plateau in eastern Australia (Bijl et al., 2013). This result supports the MECO as a period of 419 hydrological cycle intensification and more extensive flooding on the western Australian continent 420 (Wang et al., 2022). Furthermore, the rapid (10s of kyr) temperature drop after the MECO warming 421 peak can be explained by greatly strengthened silicate weathering during this warming period, 422 supporting the reactivation of the silicate weathering feedback after the hypothesized gradual decline 423 during the Eocene (Henehan et al., 2020). Interestingly, the MECO record of Os<sub>i</sub> at Site U1514 is in 424 conflict with the results from other MECO sedimentary successions (i.e., Site U1333, Site 959, and

425 Site 1263), exhibiting a slight Osi decrease (van der Ploeg et al., 2018) (Fig. 6a, Supplementary Fig. 426 S2). A relative increase in volcanic-hydrothermal input is required to explain this abrupt shift to 427 unradiogenic Osi value, (Peucker-Ehrenbrink and Ravizza, 2000; Du Vivier et al., 2014). However, 428 van der Ploeg et al. (2018) considered this decrease in  $Os_i$  across the MECO to be caused by 429 relatively weakened silicate weathering, which resulted in prolonged accumulation of volcanic CO2 430 in the oceans and atmosphere during this period. This scenario fails to account for a rapid drop in 431 temperatures following the peak of the MECO warming trend, potentially triggered by the 432 reactivation of silicate weathering feedback after million years of hypothesized gradual decline. 433 Although the regional disparity (i.e., heterogeneous seawater Os isotope composition) in Os<sub>i</sub> values 434 likely arises from differing environmental settings, including a restricted basin, and/or a more proximal depositional setting at the time of deposition (Du Vivier et al., 2014), this spatial 435 436 heterogeneity (i.e., contrasting pattern) of Os isotope signals in MECO sedimentary successions 437 requires further investigation with additional data from other locations. Nonetheless, the abrupt 438 increase to more radiogenic Osi at Site U1514 is difficult to explain without invoking enhanced 439 continental weathering rates (Peucker-Ehrenbrink and Ravizza, 2000; Percival et al., 2016). Although 440 the possibility of unradiogenic Os input from enhanced volcanism during the middle Eocene cannot 441 be excluded, a substantial amount of terrigenous detritus from the Australian continent may have 442 overwhelmed the seawater Os isotope composition, thereby masking volcanic-induced isotope 443 signals, and by inference of the Osi of the deposited sediments. This hypothesis may explain why the 444 MECO thermal maximum interval at Site U1514 was characterized by the largest increase in Os<sub>i</sub> 445 values. Thus, in some cases, silicate weathering feedback may obscure volcanic isotope signals, 446 limiting our understanding of the geographic distribution and global correlation of volcanic events.

## 447 **5.** Conclusions

448 We have documented the complete Eocene variations of Hg and Os concentrations and their 449 isotope signals from Site U1514, which provide compelling evidence of the operation of feedback 450 between plate tectonics, volcanism, continental chemical weathering, and climate change during the 451 middle Eocene. A comprehensive analysis of Hg and Os isotopes demonstrated that the middle 452 Eocene (~45–38 Ma) experienced extensive volcanic activity, evident from notably elevated Hg 453 enrichments and near-zero  $\Delta^{199}$ Hg values. This prolonged and enhanced volcanic activity was 454 intricately linked to widespread plate tectonic activity in the Southern Ocean, and likely played a 455 crucial role in triggering global warming incidents, including the LLTM and MECO. An analysis of seawater  ${}^{187}$ Os/ ${}^{188}$ Os (Os<sub>i</sub>) in conjunction with supporting data, such as influxes in terrigenous 456 457 materials and elemental indices, underscores that the middle Eocene was a critical period in the 458 intensification of the continental weathering response to CO2 rise and warming. This intensified 459 continental weathering was most prominent during the MECO warming, further supporting the 460 prevalence of warmer, more humid climate conditions. Notably, our findings suggest a long-term (~ 5 461 Myr) middle Eocene warming, instigated by a feedback loop involving plate tectonics and volcanic 462 activity in the southern high latitudes. This finding challenges the prevailing notion of a continuous 463 global cooling trend throughout the Eocene; however, the prolonged continental weathering during 464 this long-term warmth is reconcilable with the reduced atmospheric  $CO_2$  levels that contributed to 465 late Eocene cooling and/or EOT glaciation.

466

# **CRediT** authorship contribution statement

467 All authors have contributed to this work. **D. Lim:** Conceptualization, Methodology, Funding 468 acquisition, Investigation, Writing – original draft, Writing – review & editing; Z. Xu:

469 Conceptualization, Funding acquisition, Writing – review & editing; J. Kim: Investigation,

470 Methodology; Formal analysis; D. Selby, E. Ownsworth, and W. Wang: Investigation, Methodology,

- 471 Writing review & editing; **R. Yin:** Methodology, Writing review & editing; **T. Chang:** Writing –
- 472 review & editing. All authors have seen and approved manuscript.

473 <b>Declaration of competing int</b>	terest
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- 474 The authors declare that they have no known competing financial interests or personal
- 475 relationships that could have appeared to influence the work reported in this paper.

#### 476 Data availability

477 The original contributions presented in the study are given in the article/Supplementary478 Material.

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#### 486 Appendix A. Supplementary material

- 487 Supplementary material related to this article can be found on-line at
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#### 713 Figure captions

- Figure 1. (A) Paleo- (~40 Ma) and (B) present-location of IODP 369 Site U1514. Shown in a
- paleogeographic reconstruction are other drilled core sites (ODP Site 865, Rolewicz, 2013; ODP Site
- 716 959 and 1263, and IODP Site U1333, van der Ploeg et al., 2018) and manganese crust section (CD29-
- 717 2, Klemm et al., 2005) cited for comparison of osmium isotope data. The basic map of
- 718 paleogeographic reconstruction is from van der Ploeg et al. (2018).
- Figure 2. The lithology, linear sedimentation rate, and age-depth plot of sediments between 56 and
  260 m CCSF at Site U1514 (after Wang et al., 2022). Age determinations are based the geological
- time scale 2012 (GTS2012). E. Eoc.: early Eocene, L. Eoc.: late Eocene, Olig.: Oligocene.
- **Figure 3**. Vertical profiles of (A) Hg concentrations and their ratios normalized by (B) TOC and TIC, (C) Al and TS contents, (D) Hg isotope ( $\Delta^{199}$ Hg), (E) <sup>192</sup>Os concentrations, and (F) initial <sup>187</sup>Os/<sup>188</sup>Os (Os<sub>*i*</sub>) for bulk sediments of Site U1514. Vertical dashed lines indicate the baseline values (average values of early and late Eocene intervals). Four climax episodes of volcanic eruptions are shown as gray shaded vertical bars near y-axis. Os*i* data of CD29-2 (Klemm et al., 2005) is incorporated into Figure (F) for comparison.
- Figure 4. Vertical profiles of (A) Hg/TOC ratio, (B)  $EF_{Hg}$ , (C) Hg flux, (D) Hg isotope ( $\Delta^{199}$ Hg), (E)
- bulk carbon isotope ( $\delta^{13}$ C) for Site U1514, and (F) major tectonic and volcanic events. Bulk carbon
- isotope data for Site U1514 are from Edgar et al. (2022). Note the volcanic climax episode that lasted
- 731 ~1.5 Myr, just before the MECO. Horizontal dashed lines indicate submarine volcanic climax
- episodes coupled with low  $\Delta^{199}$ Hg values. 1) Cottin et al. (2011), 2) Borrissova et al. (2010), 3)
- 733 Veevers (2000), 4) Vilacis et al. (2022), 5) Niyazi et al. (2021), 6) van der Boon et al., (2021). LLTM:
- 734 Late Lutetian Thermal Maximum, MECO: Middle Eocene Climate Optimum, KP: Kerguelen
- 735 Plateau, BR: Broken Ridge, SE AUS: Southeastern Australian Continent, ANT: Antarctica
- Figure 5. Cross-plot of  $\Delta^{199}$ Hg versus Hg concentration and enrichment (Hg/TOC) for Site U1514.
- The values of  $\Delta^{199}$ Hg for mantle-derived volcanic emissions (0.00 ± 0.05‰, 2 $\sigma$ ) and seawater (+0.21
- $\pm 0.07\%$ ,  $2\sigma$ ) are from Moynier et al. (2021) and Štrok et al. (2015), respectively.

- 739 Figure 6. Temporal variations of (A) Os isotopic ratio (Os<sub>i</sub>), (B) mass accumulation rate
- 740 (MAR<sub>siliciclastic</sub>) of detrital materials, (C) chemical weathering indices (kaolinite/smectite ratio and
- 741 Mg/Al ratio) for Site U1514, and (D) the global deep-sea (Westerhold et al., 2020) and Austral high
- 742 latitude surface temperatures (Tremblin et al., 2016). Note synchronous increase in Os<sub>i</sub> and chemical
- 743 weathering indices during the middle Eocene (especially the MECO). The data of chemical
- 744 weathering indices for U1514 and Os<sub>i</sub> values for Site 1264 (see Fig. 1a for core site) are from Wang
- et al. (2022) and van der Ploeg et al. (2018), respectively. The middle Eocene warming reversal is
- shown as the shaded area.







# Figure 2.



Figure 3.



Figure 4.



Figure 5.



Figure 6.

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