

How to measure the spin of invisible states in $e^+e^- \rightarrow \gamma + X$

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(Received 30 August 2023; accepted 13 November 2023; published 12 December 2023)

We examine the production of an invisible state X together with a photon, $e^+e^- \rightarrow \gamma + X$, at electron positron colliders and present measurement strategies that can detect the spin of the invisible state as well as the underlying production mechanism, based on the angular distribution of the final-state photon, the cross sections for polarized initial states, and the photon polarization. Our measurement strategy can be used to identify whether the invisible state is a hidden photon or an axion. The results are compared with a detailed analysis of the Standard Model background, and we calculate the sensitivity reach for searches for axions and hidden photons at Belle II.

DOI: [10.1103/PhysRevD.108.115013](https://doi.org/10.1103/PhysRevD.108.115013)

I. INTRODUCTION

Searches for missing energy at electron-positron colliders with a final-state photon, $e^+e^- \rightarrow \gamma + X$, are one of the most sensitive probes for invisible states X with masses in the range 0.1–9 GeV. With unequal beam energies, $E_{e^+} = 4$ GeV and $E_{e^-} = 7$ GeV, and a center-of-mass energy of $\sqrt{s} = 10.58$ GeV, Belle II will provide a valuable opportunity to search for these signatures [1]. Here we discuss how we can take advantage of the clean environment at e^+e^- colliders and use the angular distribution of the final-state photons, polarized beams, and final-state photon polarization to identify the spin and the coupling structure of the invisible state. We introduce the most minimal model of a spin-1 hidden photon (also known as a dark photon) interacting with electrons, and a pseudoscalar spin-0 boson interacting with electrons and photons, which we call axions throughout the paper. For axions such a search has been proposed first by Wilczek [2] and searches have been performed with CLEO [3] and BABAR [4,5]. For Belle II, projections are available [6–8]. A search for hidden photons has been performed by BABAR [9] and a projection for Belle II can be found in [1]. The existing searches assume a model and production mechanism, and interpret searches for missing energy and a single photon as a signal in that model. We extend the search strategy proposed in [1] and identify observables that can distinguish between the different models in case a signal is observed.

First, we demonstrate that the angular distribution can be used to determine whether the invisible state has been produced in an s -channel or t/u -channel process. Second, we suggest extending the proposed “Chiral Belle program” that aims to use a 70% polarized electron beam in a future upgrade of Belle II [10]. We argue that a polarization of both the electron beam and the positron beam allows a measurement that would uniquely fix the quantum numbers of the invisible state. In addition, Standard Model background processes are largely suppressed and allow for more aggressive cuts on the photon energy and scattering angle. We discuss the different processes contributing to this background and propose new strategies that can lead to improved sensitivity. For the three models discussed, we compute the expected sensitivity reach at Belle II. An alternative, more speculative strategy is to explore the possibility to detect the polarization of the final-state photon. We discuss how such a measurement would give additional, complementary information on the production process and quantum numbers of the invisible state.

The rest of the paper is organized as follows. In Sec. II, we introduce the hidden photon and axion Lagrangians, and in Sec. III we discuss how the differential cross sections with and without polarized beams can be used to identify the production process. In Sec. IV, simulations and analyses for signal and background are presented, and we propose to improve the search strategy. We discuss the implications of measuring the photon polarization in Sec. V and in Sec. VI we present the sensitivity reach for future runs of Belle II.

II. AXION AND HIDDEN-PHOTON LAGRANGIAN

We compare two minimal models of invisible states that can be produced via $e^+e^- \rightarrow \gamma + X$ that carry either spin-1

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or spin-0. The relevant terms in the canonically normalized basis for the spin-1 field or hidden photon are given by

$$\mathcal{L} \ni -c_X \bar{\psi} \gamma_\mu \psi X^\mu - \frac{m_X^2}{2} X_\mu X^\mu, \quad (1)$$

where m_X is the mass of the hidden photon and c_X denotes its coupling strength to electrons. The interaction can be proportional to a new gauge charge in the case where X is the gauge boson of a new gauge group under which one of the SM global symmetries is charged, or generated via mixing with the SM photon. In the case of kinetic mixing one can write $c_X = \epsilon e Q_\psi$, where ϵ is the coefficient of the kinetic mixing term of the electromagnetic field strength tensor $F_{\mu\nu}$ and the hidden-photon field strength tensor, $\mathcal{L} \ni -\epsilon/2 F_{\mu\nu} X^{\mu\nu}$, and Q_ψ is the electric charge of the fermion ψ in units of the electron charge e [11–14]. Interaction due to mixing with the Z boson are suppressed for masses $m_X \leq 10$ GeV, but can play a role for hidden photons with masses closer to the Z -pole [15,16].

In the case of the spin-0 particle we consider an axion a with the interactions,

$$\mathcal{L} \ni c_e \frac{\partial_\mu a}{2f} \bar{\psi} \gamma^\mu \gamma_5 \psi - c_\gamma \frac{a}{f} F_{\mu\nu} \tilde{F}^{\mu\nu} - \frac{m_a^2}{2} a^2, \quad (2)$$

where c_e and c_γ are coupling constants, and m_a is the mass term, though we also use m_X to denote the mass of the new particle when referring to both the hidden photon and axions. Pseudoscalars with these interaction terms are also called axionlike particles (ALPs) in the literature. Since any axion will generically interact with photons and leptons due to renormalization group running we use the term *axion* here [17].

In all cases we consider the hidden photon or axion not to decay into SM particles on collider scales, e.g., by introducing a dominant decay into a set of invisible particles like dark matter. Otherwise for the mass range considered here with masses up to $m_X = 9$ GeV, both the hidden photon and axion decay back into electron-positron or photon pairs. Neglecting subleading terms suppressed by the electron mass, the corresponding decay lengths for the three models read,

$$\begin{aligned} \ell(c_X) &= \left(\frac{c_X^2 m_X}{12\pi} \right)^{-1} \approx 10^{-12} \left[\frac{1}{m_X c_X^2} \right] \text{ GeV m}, \\ \ell(c_e) &= \left(\frac{c_e^2 m_a m_e^2}{8\pi f^2} \right)^{-1} \approx 2 \times 10^{-8} \left[\frac{f^2}{m_a c_e^2} \right] \text{ GeV}^{-1} \text{ m}, \\ \ell(c_\gamma) &= \left(\frac{c_\gamma^2 m_a^3}{4\pi f^2} \right)^{-1} \approx 2 \times 10^{-15} \left[\frac{f^2}{m_a^3 c_\gamma^2} \right] \text{ GeV m}. \end{aligned} \quad (3)$$

The main differences between the models are their spin, the coupling structure to fermions, and the axion coupling to photons. The production of the hidden photon occurs only

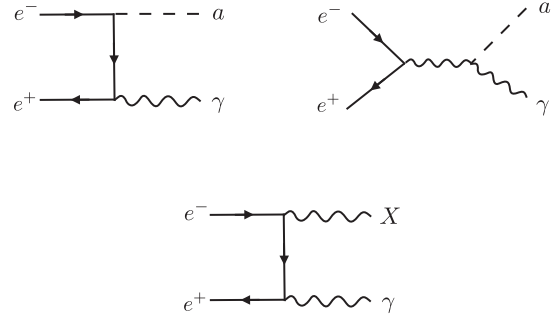


FIG. 1. Different contributions to the production of axions (top) and hidden photons (bottom) at an e^+e^- collider.

though couplings to the electron, whereas for the axion there are two contributions. The corresponding diagrams for are shown in Fig. 1.

III. ANGULAR DISTRIBUTIONS

In the following we discuss how the angular distribution of the final-state photon in the process $e^+e^- \rightarrow \gamma + X$ can be used to discriminate between the hidden photon and axion final states as well as between the axion coupling to electrons and photons. In general the amplitude for the process,

$$e^+(p_1, \lambda_{e^+}) + e^-(p_2, \lambda_{e^-}) \rightarrow \gamma(q_1, \lambda_\gamma) + X(q_2, \lambda_X), \quad (4)$$

can be written as $\mathcal{M} = \mathcal{M}_\mu e^\mu(q_1, \lambda_\gamma)$ with the photon polarization vector e^μ and helicities $|\lambda_{e^\pm}| = 1/2$, $|\lambda_\gamma| = 1$, $|\lambda_X| = 0$ for the axion, and $|\lambda_X| = 1$ for the hidden photon, respectively. At Belle II, the incoming beams are angled 83 mrad with respect to each other, with the z -axis defined with equal distance to the beams [1]. This distinction is assumed to have little effect on the results presented in this paper, and hence it will be assumed that the beams are antiparallel along the z -axis. We define the momenta of the incoming particles parallel to the z -axis. The differential cross section for the production of hidden photons in the center-of-mass frame reads up to corrections of order m_e^2/s ,

$$\frac{d\sigma}{d\Omega} = c_X^2 \frac{\alpha}{4\pi s} \frac{(1 + \tau_X)^2 + (1 - \tau_X)^2 \cos^2 \theta}{(1 - \tau_X)(1 - \cos^2 \theta)}, \quad (5)$$

where s is the center-of-mass energy squared, $\cos \theta$ is the angle between the photon and the beam axis in the lab frame, and $\tau_X = m_X^2/s$. For the production of axions the differential cross section reads [7],

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \frac{\alpha}{4\pi f^2} \left[c_e^2 \frac{m_e^2}{s} \frac{1 + \tau_X^2}{(1 - \tau_X)(1 - \cos^2 \theta)} \right. \\ &\quad \left. + c_e c_\gamma \frac{m_e^2}{2s} \frac{(1 - \tau_X)^2}{(1 - \cos^2 \theta)} + \frac{c_\gamma^2}{32} (1 + \cos^2 \theta)(1 - \tau_X)^3 \right], \end{aligned} \quad (6)$$

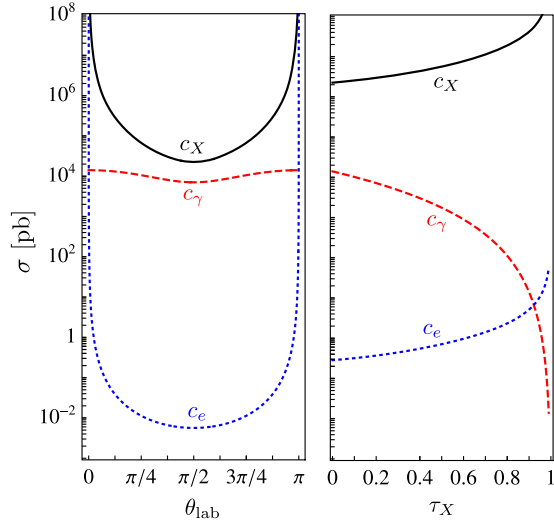


FIG. 2. Cross sections for the production of hidden photons with $c_X = 1$ (black), axions with couplings to photons for $c_\gamma/f = 1/\text{GeV}$ and $c_e = 0$ (red dashed), and electrons with $c_e/f = 1/\text{GeV}$ and $c_\gamma = 0$ (blue dotted) as a function of the scattering angle θ_{lab} (left) and τ_X (right). We set $s = 10 \text{ GeV}$, $\tau_X = 0$ (left), and $\theta_{\text{lab}} = \pi/4$ (right).

up to corrections of order m_e^2/s . Here the first term is the contribution from the diagram with the axion coupling to the electron, the second is the interference term, and the last term is the contribution from the diagram with the axion radiated from the photon (axion-strahlung). Appendix A describes the amplitude calculations in detail. The differential cross sections are shown in the left panel of Fig. 2. Both the hidden photon and the axion coupled to electrons are produced mostly for $\cos(\theta) \rightarrow 1$. The angular distribution can therefore distinguish a new particle produced in t/u -channel diagrams (axions coupling to electrons or hidden photons) from particles produced in the s -channel (axions coupling to photons), though it is not enough to distinguish within these categories.

We show the dependence of the differential cross sections on the mass of the invisible state m_X in the right panel of Fig. 2. The production via the s -channel is suppressed for large τ_X , but the production via t/u -channel is enhanced. Hence, independently of the background present, the expected signal from the s -channel contribution reduces for larger masses, decreasing the experimental sensitivity to the axion coupling to photons. We note that the t/u -channel contributions seem to diverge for both large angles and $\tau_X \rightarrow 1$, but are regularized by the electron mass.

The chiral Belle program is the proposal to use a polarized electron beam to collide with an unpolarized positron beam [10]. Since the amplitude for hidden photon and axion-strahlung production is dominated by electrons and positrons with opposite helicities, whereas the production of axions through coupling to electrons is

dominated by electrons and positrons with equal helicities, this will not have a qualitative effect on the differential production cross sections. If both the electron and positron beams are polarized, one can distinguish the hidden photon from an axion. In the case of a hidden photon, if electron and positron are polarized with equal helicity, the signal would be suppressed by m_e^2/s with respect to the unpolarized case. In contrast, the leading term for the cross section for axions produced via electron couplings remains unchanged for electron and positron beams with equal helicities. These two production channels have the same angular distribution, but with the polarization of both beams one can significantly reduce the contribution of either channel, distinguishing them from one another. Similarly, in the context of a dark vector boson Z_d , which differs from the hidden photon considered by the inclusion of an axial-vector coupling, the angular distribution of the polarized-differential cross section of $e^+e^- \rightarrow Z_d\gamma$ can be used to distinguish between the vector and axial-vector couplings [18]. We focus on longitudinal beam polarizations here, but for electron and positron beams with transversal polarization the information from the azimuthal angular distribution can be used to further discriminate between background and signal [19–21].

IV. EVENT GENERATION

The signature in the production of an invisible state together with a photon is a single photon recoiling against missing energy. Any search for New Physics with this signature needs to account for a number of SM processes that produce final states that are difficult to distinguish. The SM background is dominated by the process $e^+e^- \rightarrow e^+e^-\gamma$ in which both the electron and the positron escape the detector. Additional background processes are $e^+e^- \rightarrow \gamma\gamma(\gamma)$ in which one or two photons are lost, $e^+e^- \rightarrow e^+e^-\gamma(\gamma)$ in which the electron-positron pair is not detected, and the irreducible production of neutrinos $e^+e^- \rightarrow \nu\bar{\nu}\gamma(\gamma)$. The selection of signal events takes advantage of the signal and background event distribution, following the analysis in [1] with modifications.

A. Standard model background

Each SM background process has several contributions, where the corresponding Feynman diagrams are shown in Fig. 3. Implementing minimal cuts on the lab frame energy $E_\gamma > 0.01 \text{ GeV}$ and asymmetric angular coverage $12.4^\circ \leq \theta_{\text{lab}} \leq 155.1^\circ$ [22], we find that the fraction of the different processes contributing to the background shown in Table I. In Fig. 4, we show the normalized differential photon-energy distribution of the different background processes compared to the signal process for three different masses of the invisible state $m_X = 1, 6, 9 \text{ GeV}$ using the cuts described above. The shape of the differential energy distribution is independent, up to small corrections, of the

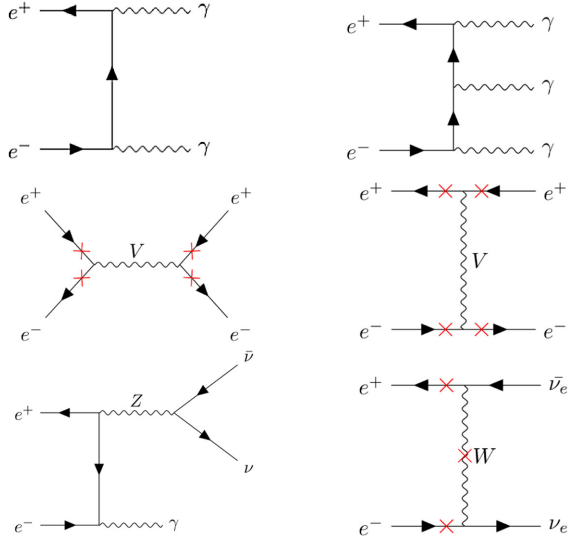


FIG. 3. Feynman diagrams for all background processes, $e^+e^- \rightarrow \gamma\gamma(\gamma)$ (top), $e^+e^- \rightarrow e^+e^-\gamma(\gamma)$ (middle), and $e^+e^- \rightarrow \nu\bar{\nu}(\gamma)$ (bottom), with $V \in [\gamma, Z]$.

spin or specific production process of the invisible state and shows how cuts on the energy can isolate the signal in particular for small masses m_χ . For photon with high energies or small mass m_χ , the background process $e^+e^- \rightarrow \gamma\gamma$ is increasingly important, whereas for low-photon energies or high-mass m_χ other background processes are more important. For this reason the *BABAR* search for hidden photons distinguishes two different mass regions, $-4 \text{ GeV}^2 < m_\chi^2 < 36 \text{ GeV}^2$ and $24 \text{ GeV}^2 < m_\chi^2 < 69 \text{ GeV}^2$,¹ in which the background from $e^+e^- \rightarrow \gamma\gamma$ and $e^+e^- \rightarrow e^+e^-\gamma$ dominate, respectively [4]. Similarly, Belle II anticipates two different signal regions for $m_\chi < 6 \text{ GeV}$ and $m_\chi = 6\text{--}8 \text{ GeV}$ [1]. In the following, we will analyze the proposed cuts for these analyses which are optimized to search for hidden photons and comment on the different optimization for axion searches and how polarized beams or a final-state photon polarization measurement could improve the analysis.

We simulate 5×10^6 background events using MadGraph [23] with a minimum transverse momentum cut applied to outgoing photons (which is also applied to the invisible states), $p_T \geq 0.01 \text{ GeV}$ in order to avoid divergences. The simulations are performed for tree-level, fixed-order QED contributions with a simplified detector setup. We, for example, do not consider the imperfections in the crystals in the detector, and the photon conversion probability. The beam polarization is specified within MadGraph for 0–100% polarized beams. We require only one photon within the angular acceptance of the detector; either in the end caps or the main barrel. All other particles must be undetected, and specifically for all charged particles we also require for outgoing fermions to have $p_T \geq 0.1 \text{ GeV}$.

¹The upper limit is $m_\chi^2 = 63.5 \text{ GeV}^2$ for the $\Upsilon(2S)$ dataset.

TABLE I. Fraction of the different SM backgrounds for simulations performed in this study.

SM process	Fraction (%)
$e^+e^-\gamma$	79.38
$e^+e^-\gamma\gamma$	10.39
$\gamma\gamma\gamma$	9.72
$\gamma\gamma$	0.51
$\bar{\nu}\nu\gamma(\gamma)$	< 0.01

The distribution of photons from the various SM backgrounds in the plane spanned by the scattering angle θ_{lab} and center-of-mass energy E_{CMS} is shown in Fig. 5. The left panel shows the total background, and the right panels show the distributions for the different background processes. Note that the color coding applies only to the left panel. In the right panels, the color coding merely indicates the distribution of events within each panel, but the scaling varies for each panel.

The panel with $\gamma\gamma$ has the same color coding as the total background on the left, whereas $\gamma\gamma\gamma$ and $e^+e^-\gamma(\gamma)$ backgrounds are rescaled by a factor 15–50, and $\bar{\nu}\nu\gamma(\gamma)$ by a factor of 5×10^7 with respect to the left panel in order to increase the visibility of the features. For larger displays and individual color coding (see Appendix B). The background from $\gamma\gamma$ final states has a fixed energy, but only contributes for certain angles because of two gaps between the end caps and barrel detectors in the forward and backward direction, and the asymmetric angular coverage of the detector where only one photon is lost along the beam pipe. As a consequence of the asymmetric beam energies and asymmetric angular coverage, processes in which photons are lost along the beam pipe in the other direction do not contribute to the background because the recoiling photon is not covered by the detector.

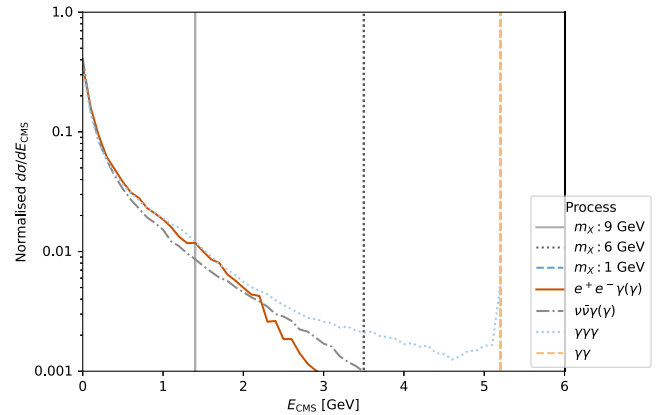


FIG. 4. Comparison of the differential cross section for the different SM background processes and the production of a new state with $m_\chi = 1, 6, 9 \text{ GeV}$ as a function of the lab frame energy of the final-state photon E_γ .

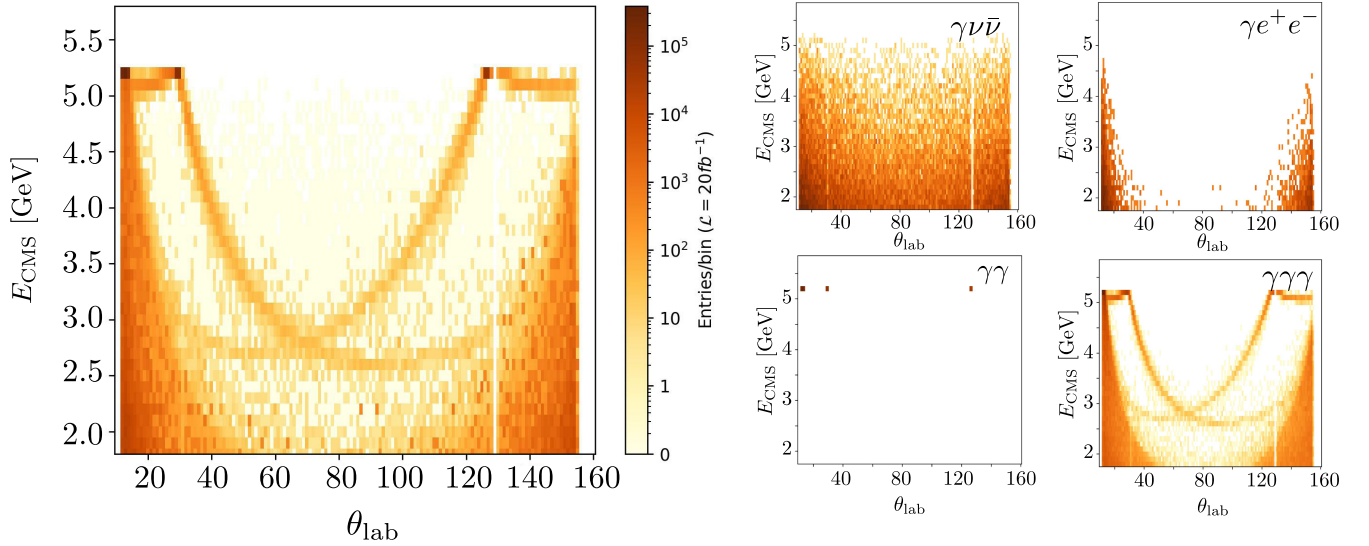


FIG. 5. Distribution of Standard Model background events for searches for final states with a photon and missing energy (left) and the separate contributions (right) displaying the shape of each contribution determined by the angular distribution and detector geometry, each with their own scaling (see text for further explanation).

The end cap gaps are also visible in the background from $\gamma\gamma\gamma$ final states, where the two bands crossing the central region of the plane correspond to photons lost in either forward or backward direction. The background from $e^+e^-\gamma(\gamma)$ final states is most pronounced for small angles with the beam pipe and there are fewer events for angles $40^\circ \lesssim \theta_{\text{lab}} \lesssim 120^\circ$, whereas background events from $\nu\bar{\nu}\gamma(\gamma)$ final states have an almost flat θ_{lab} distribution, but the contribution decreases for photon with high energies. Our background distribution agrees with the unpublished results found in [24] where a detailed analysis of the different background processes is performed. We find more background events in the for large θ_{lab} Belle physics paper [1], which includes higher-order QED corrections and performs a full detector analysis, though other cuts are applied.

For electron and positron beams polarized with equal helicities, the backgrounds are significantly reduced. Backgrounds from $\gamma\gamma$ and $\gamma\gamma\gamma$ final states are substantially suppressed with respect to the unpolarized case, and the remaining background is mostly peaked towards small angles with the beam axis. In the right panel of Fig. 6, we show the distribution of background events for beams polarized with the same helicities. Instead, for electron and positron beams polarized with opposite helicities the backgrounds are very similar to the unpolarized case. Hence in the following analysis, the case for oppositely polarized beams will not be considered.

B. Signal

We implement the signal processes into MadGraph using UFO models [25] based on modified FeynRules models for ALPs [26] and Z' models [27–29]. The energy of the photon recoiling against the invisible state is fixed by

the mass of the invisible state. The signal is therefore constant in E_{CMS} , and larger masses m_X correspond to lower photon energies. The angular distribution of the signal peaks towards small and large angles with respect to the beamline in the case of hidden photons or axions radiated off an electron, whereas it is approximately flat for axions coupled to photons. In case a signal is observed, and enough statistics are available, the angular distribution can be used to distinguish these models.

We simulate 10^6 events with the axion couplings fixed at $c_\gamma = 1$ and $c_e = 1 \times 10^4$ while the decay constant f is kept as a free parameter. We assume that all hidden photons and axions leave the detector before they decay, or that they decay into invisible particles, so that their width is considered to be zero for the remainder of this analysis. This would be a good assumption if they represent mediators that dominantly decay into dark matter. We vary the parameters within,

$$c_X \in [5 \times 10^{-6}, 1 \times 10^{-5}, 5 \times 10^{-5}, \dots, 5 \times 10^{-3}],$$

$$f \in [4 \times 10^5, 2 \times 10^5, 4 \times 10^4, \dots, 4 \times 10^2].$$

The mass of the invisible state is varied with changing step sizes in order to increase precision; $m_X \in [1.0, 8.0]$ GeV in steps of 1 GeV with additional $m_X = 0.1, 0.5, 8.5,$ and 9.0 GeV.

For polarized beams, the signal changes as discussed in Sec. III. If the beams have the same helicities the background is substantially suppressed. The signal, in the case of a hidden photon, is chirally suppressed and very small compared to the unpolarized case. Similarly, axion coupled only to photons are produced with a strongly suppressed cross section for beams with equal helicities. In contrast,

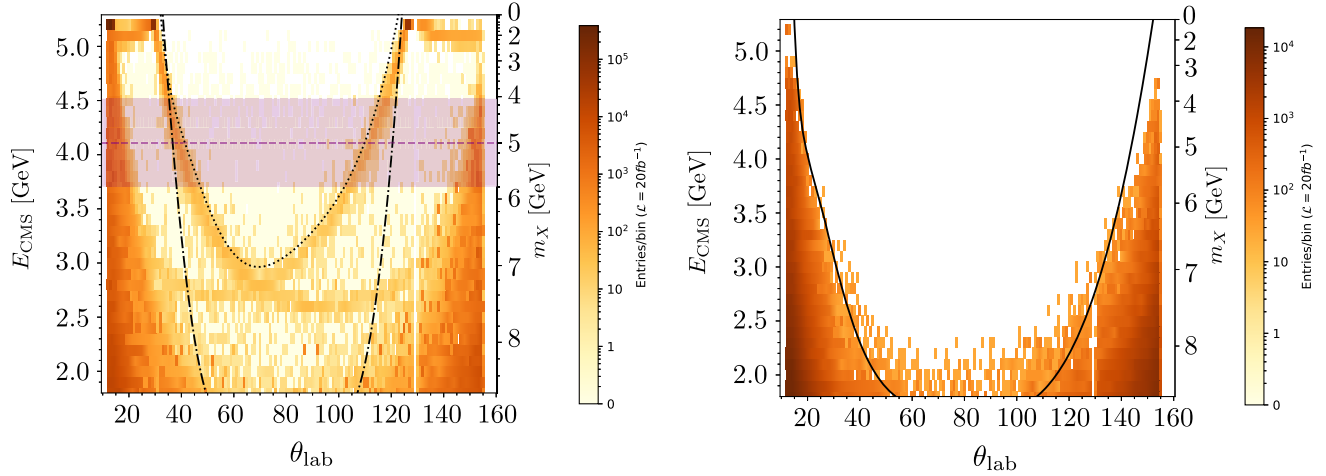


FIG. 6. Event selection for unpolarized beams (left) and equal helicities polarized beams (right). In the left panel we also show the E_{CMS} envelope used to define cuts in the search for a 5 GeV invisible state.

the cross section for the production of axions interacting with electrons is not suppressed and one could take advantage of the lower background for polarized beams.

C. Event selection

We distinguish searches for light new states $m_X < 6$ GeV and heavy new states $m_X \geq 6$ GeV, because there is more background for softer photons, and the photon trigger efficiency varies significantly with energy [1]. We use the Belle II angular acceptance regions as described in [1] and apply a cut on the energy of the detected photon, $E_{\text{CMS}} \geq 1.8$ GeV, which restricts the mass of the invisible state. The angular coverage consists of three regions; the forward end cap $12.4^\circ < \theta < 31.4^\circ$, the mail barrel $32.2^\circ < \theta < 128.7^\circ$, and the backwards end cap $130.7^\circ < \theta < 155.1^\circ$.

In a first step, we impose an energy-dependent cut in the $\theta_{\text{lab}} - E_{\text{CMS}}$ plane, taking advantage of the correlation between energy and scattering angle for the distribution of the background events. The cut functions are polynomials that were fitted using an algorithm designed to minimize background events. In contrast to the cut functions in [30], our cuts reduce the background from $\gamma\gamma\gamma$ and $\gamma\gamma$ final states. The details of the fit are explained in Appendix C. In the left panel of Fig. 6, the dotted black contour defines the cut for light invisible states, and the dash-dotted black contour defines the cut for heavy invisible states. The parameter space enclosed by these contours is the fiducial region. In the case of polarized beams with the same helicity, the energy-dependent θ_{lab} cut is shown by the black contour in the right panel of Fig. 6. The reduced background allows for a large fiducial region compared to the unpolarized case, so that the separation for different mass regions is unnecessary.

In a second step, we introduce a mass-dependent cut on E_{CMS} . The fixed relation between the final-state photon energy and the mass of the invisible state is shown by the

alternative y-axis in Fig. 6. It allows a more targeted search since signal events are predicted by the invisible state mass. The inclusion of higher-order corrections causes a smearing in the final-state photon energy, and hence in addition to the energy-dependent θ_{lab} cut we only select events in a window of $m_X \pm 0.4$ GeV. This window is shown in the left panel of Fig. 6 for the case of $m_X = 5$ GeV, which is wider than the one derived in the full Belle analysis based on the Novosibirsk function for higher masses m_X [1] and comparable to the energy window used in [30]. We checked that it captures more than 95% of the signal for the example given in [1] for $m_X = 7$ GeV.

The trigger efficiency ϵ_s is taken from [1] and interpolated for different masses as shown in Fig. 7. In the case of polarized beams, we make the conservative choice to use the trigger efficiency for the low-mass region, which is worse than the efficiency expected for the high-mass region as long as $m_X < 8$ GeV.

V. THE POLARIZATION OF THE FINAL-STATE PHOTON

We discussed how the angular distribution together with the polarization of the incoming beams can be used to

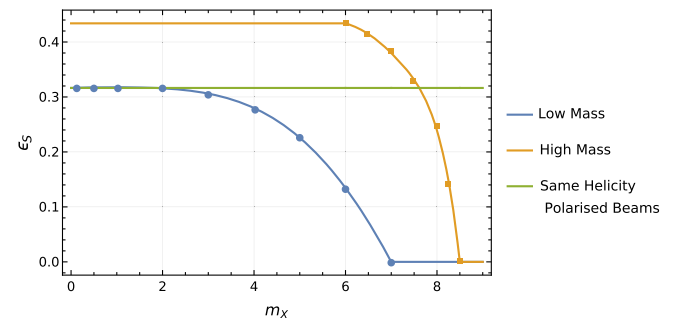


FIG. 7. Trigger efficiency for low- and high-mass region and the fixed efficiency used for the analysis with polarized beams.

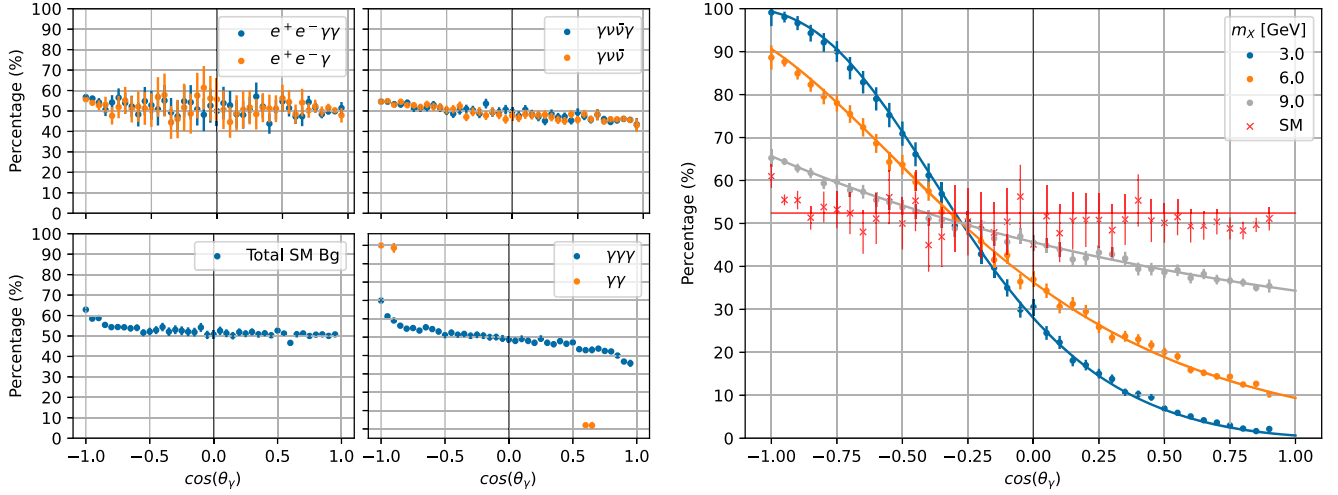


FIG. 8. Helicity fraction distribution of $2\lambda_{e^-} = \lambda_\gamma$ for background processes (left) and hidden photon (right) for $m_\chi = 3.0, 6.0,$ and 9.0 GeV as a function of the polar angle $\cos\theta_\gamma$.

distinguish between the hidden photon, axion coupling to photons, and axion coupling to electron contributions. The polarization of the beams can be used to significantly reduce the background for the axion coupling to electrons, but leaves the background for the other two dark matter contributions essentially unchanged. We therefore will now consider how the helicity of the outgoing photon can be used as a complementary discriminator.

Figure 8 displays the percentage of outgoing photons with helicity equivalent to that of the incoming electron for 100% polarized beams, a detailed description of the results can be found in Appendix D. The error bars shown are calculated using

$$D_i = \frac{N_i}{N_{\text{tot}}} \Rightarrow \delta D_i = \frac{N_i}{N_{\text{tot}}} \sqrt{\frac{1}{N_i} + \frac{1}{N_{\text{tot}}}}, \quad (7)$$

for N_i number of events for each $\cos\theta$ and total number of events N_{tot} . The distribution of photon helicity for the standard model background is centered around 50% with limited angular dependence as seen on the bottom-left panel. Any significant variation from this would indicate the presence of New Physics.

The hidden photon production has the photon helicity matching the helicity of the fermion traveling in the same direction, which is featured in the s-shape distribution seen on the panel on the right. For small m_χ , this correlation is more pronounced, but as m_χ increases, the direction of the photon becomes irregular and its helicity more independent. Therefore, the overall fraction tends towards 50% for increasing m_χ .

The helicity distribution for axion coupling to photons has the same s-shape, though the mass of the axion has no influence on the result as it factorizes out and becomes a part of the coupling constant for the process. The s-shape

helicity distribution is substantially different to the background and can therefore be used to distinguish the signal and provide improved exclusion limits. We will not carry this out due to the unknown detector setup needed to measure the helicity of the outgoing photon.

As the electron and positron helicities are the same for the axion coupling to electrons, the photon helicity will match the helicity of both or neither, and the resulting helicity fraction has no angular dependence. For small m_a , the photon helicity will match the incoming fermions, resulting in a constant percentage of 100%. But like the hidden photon process, as the axion mass increases, the distribution approaches 50% with increasing m_a . The axion coupling to electrons will cause a shift, and hence it can easily be distinguished from the other to dark matter contributions, but would be difficult to separate from the background.

A. Photon polarization

In addition to the beam polarization the polarization of the final-state photon can be used to determine the dirac structure in the production amplitude. This has been successfully used to examine of chiral structure of the operator responsible for $b \rightarrow s\gamma$ transitions at LHCb [31]. Measuring the final-state photon polarization is extremely challenging, but if it is possible it could provide an additional handle on the spin of the dark matter state. We define

$$\alpha_\gamma = \frac{\gamma_L - \gamma_R}{\gamma_L + \gamma_R}, \quad (8)$$

where $\gamma_{L/R}$ is the number of left- and right-handed photons respectively, and the photon polarization is therefore the ratio of the two polarization states. For the SM background, $\alpha_\gamma = 0.5$ is expected for random photon polarization.

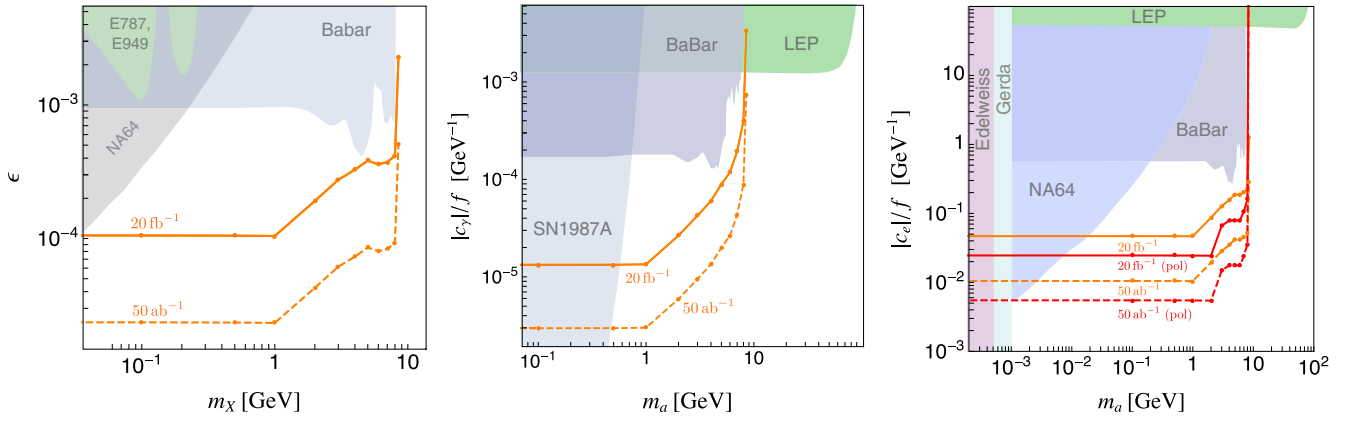


FIG. 9. Belle II sensitivity to hidden photons (left), axions coupling to photons (center), and axions coupling to electrons (right). Details are in the text.

If only one beam is polarized, all dark matter contributions will have the same photon polarization as the background, but this result can be significantly changed when polarizing both beams. When the incoming beams have opposite helicities, one can achieve $|\alpha_\gamma| \approx 0.8$ for axion-strahlung and hidden photon when implementing forwards/backwards angular cuts. Whereas for the axion coupling to electrons with incoming beams with equal helicity, for small axion masses regardless of angular cuts one finds $|\alpha_\gamma| \approx +1$. See Appendix A for further detail.

VI. SENSITIVITY TO NEW PHYSICS

In the following we discuss the sensitivity reach of Belle II for hidden photons and axions interacting with photons or electrons for unpolarized and polarized beams. We assume that the new states are stable on collider scales, for example because they decay into dark matter. We obtain the expected 90% confidence level upper limit on the observed number of signal events μ_S by demanding that the Poisson probability of observing not more than the number of background events μ_B if $\mu_B + \mu_S$ events are expected is $P(\mu_B, \mu_S + \mu_B) > 0.1$ as in [1].

Figure 9 shows the sensitivity reach of Belle II for 20 fb^{-1} (orange contour) and 50 ab^{-1} (dashed orange contour) for unpolarized beams and the improved sensitivity in the case of an axion coupling to electrons using polarized beams is shown by red contours.

For comparison, we show limits from a search for hidden photons by *BABAR* [9], limits from beam-dump experiments E787, E949 [32–35], and NA64 [36]. The center plots shows *BABAR* limits on axions from single-photon decays of $\Upsilon(1)$ [4] and monophoton searches by LEP [37,38], taken from [6], as well as constraints from the supernova SN1987A [39,40]. In the right panel we show the bounds from the neutrinoless double-beta decay experiment Gerda [41], the helioscope Edelweiss [42] taken from [43], together with the bounds by NA64 [44], *BABAR* [9], and LEP [38],

taken from [7]. Note that the bounds on astrophysical scales and therefore only constrain $m_a < 2m_e$.

The projected sensitivity reach exceeds the existing constraints for all three models by about an order of magnitude at least. In all models the sensitivity drops with larger mass m_X for which backgrounds are larger. For the case of axion-strahlung shown in the center panel, this drop is particularly steep, because the cross section drops for large masses as shown in the right panel of Fig. 2. The slight improvement of sensitivity for masses $m_X > 6 \text{ GeV}$ is a consequence of the better trigger sensitivity. Note that the projected sensitivity for masses $m_X < 2 \text{ GeV}$ at 50 ab^{-1} can only be achieved if backgrounds from cosmic rays are fully understood [8].

Our results can be compared with previous analyses of the sensitivity reach of Belle II. Our analysis largely follows [1], that also sets constraints on an invisible spin 1 state with the main differences that we implement different $\theta_{\text{lab}} - E_{\text{CMS}}$ cut functions and a constant mass-dependent window as opposed to previous estimates in [1] that consider higher-order effects and a more sophisticated detector simulation. As a consequence, we obtain projections roughly a factor three better. An analysis that translates this projection for the case of axions produced in axion-strahlung can be found in [6,7]. The improvement with the method we used is again roughly a factor of three in the projected sensitivity, consistent with the hidden photon case. A recast for axions interacting with electrons has been performed in [7], and we find again that our projections are roughly 3–4 times better using the improved cuts.

If a signal is observed at Belle II, it is possible in principle to determine whether the invisible state is a vector boson (c_X), an axion interacting with photons (c_γ), or an axion interacting with electrons (c_e). The angular distribution of the signal events can distinguish between c_γ and c_e , c_X , whereas the beam polarization suppresses the signal for c_X , c_γ and does not affect the signal for c_e .

VII. CONCLUSIONS

We propose measurement strategies that can be used at electron positron colliders to determine the spin, mass, and production mechanism of an invisible state produced in association with a photon $e^+e^- \rightarrow \gamma + X$. In particular, hidden photons and axions that interact with electrons or photons can be distinguished even in the absence of detecting any of their decay products. The angular distribution of the final-state photon is sensitive to the production mechanism and can distinguish s -channel production as in the case of axion-strahlung from t/u -channel production, e.g., of a hidden photon or an axion produced from interactions with e^+ or e^- directly. If both electron and positron beams can be polarized with equal helicities, the hidden-photon cross section is strongly suppressed with respect to the unpolarized cross section, and only the cross section for axions interacting with electrons remains unchanged. In combination, the angular distribution and the beam polarization can distinguish between these three models. Further, the dependence of the polarization of the final-state photon on the polar angle can be used to discriminate between the different models as well as the SM background if it can be reconstructed with future detectors. For beams with opposite helicities and small masses m_X , both the hidden-photon and axion coupling to photons helicity fraction distributions are significantly different from the SM background. While beams with equal helicities can be used to reduce backgrounds for axions interacting with electrons, the sensitivity for both axions interacting with photons as well as hidden photons can be increased if final-state photon polarization can be measured.

The SM background is also significantly reduced if both electron and positron beams are polarized. A careful analysis of different SM background processes shows that a combination of a universal cut in the plane spanned by the scattering angle θ_{lab} and center of mass energy E_{CMS} together with a mass-dependent cut could improve the projections for searches at Belle II, assuming the method described in this paper accounts for smearing effect from higher-order QED. If a run with polarized beams can be performed the sensitivity to axions interacting with electrons is further improved by a factor two. We compare the projections with constraints from other experiments and astrophysics and identify the parameter space that can be probed at Belle II using the proposed measurement strategy.

ACKNOWLEDGMENTS

We thank Christopher Hearty for helpful comments on the distribution of background events and the geometry of the Belle II detector. M. B. acknowledges support by the Future Leader Fellowship DARKMAP.

APPENDIX A: AMPLITUDE CALCULATIONS DETAILS

In the following we will describe how the amplitudes and analytical results were derived, for further details see [45]. The general helicity amplitude for the process is

$$e^+(p_1, \lambda_{e^+}) + e^-(p_2, \lambda_{e^-}) \rightarrow \gamma(q_1, \lambda_\gamma) + X(q_2, \lambda_X), \quad (\text{A1})$$

with

$$\begin{aligned} \mathcal{M} &= \mathcal{M}_\mu \epsilon^\mu(q_1, \lambda_\gamma) \\ &= \bar{v}(p_{e^+}, \lambda_{e^+}) \Gamma_\mu u(p_{e^-}, \lambda_{e^-}) \epsilon^\mu(q_1, \lambda_\gamma), \end{aligned} \quad (\text{A2})$$

where Γ_μ is the current for the interaction. Belle II has incoming electron and positron antiparallel along the z -axis with energies E_1 and E_2 respectively. Their momenta are defined,

$$p_1^\mu = \left(E_1, 0, 0, -\sqrt{E_1^2 - m_e^2} \right), \quad (\text{A3})$$

$$p_2^\mu = \left(E_2, 0, 0, \sqrt{E_2^2 - m_e^2} \right), \quad (\text{A4})$$

$$q_1^\mu = E_\gamma (1, \cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta), \quad (\text{A5})$$

$$q_2^\mu = p_1^\mu + p_2^\mu - q_1^\mu, \quad (\text{A6})$$

together with the two outgoing particles (photon and invisible state X) where E_γ is the photon energy in the lab frame, θ is the angle between the photon and incoming beams, and ϕ between the photon and the x -axis in the transverse plane.

1. Fermion spin vectors

Spin-1/2 particles with four-momentum $p^\mu = (E, \mathbf{p})$ have corresponding spin four-vector where λ is the helicity of the particle with $\lambda = \pm 1/2$ [46],

$$\text{Massive: } S^\mu = \frac{2\lambda}{m} (|\mathbf{p}|, E\hat{\mathbf{p}}), \quad (\text{A7})$$

$$\text{Massless: } S^\mu = 2\lambda(1, \hat{\mathbf{p}}). \quad (\text{A8})$$

Helicity projection expression can be written as

$$u(p, \lambda) \bar{u}(p, \lambda) = \frac{1}{2} (1 + \gamma^5 \not{S})(\not{p} + m), \quad (\text{A9})$$

$$\Rightarrow_{m \rightarrow 0} u(p, \lambda) \bar{u}(p, \lambda) = \frac{1}{2} (1 + 2\lambda \gamma^5) \not{p}, \quad (\text{A10})$$

$$v(p, \lambda) \bar{v}(p, \lambda) = \frac{1}{2} (1 + \gamma^5 \not{S})(\not{p} - m), \quad (\text{A11})$$

$$\Rightarrow_{m \rightarrow 0} v(p, \lambda) \bar{v}(p, \lambda) = \frac{1}{2} (1 - 2\lambda \gamma^5) \not{p}, \quad (\text{A12})$$

which are used to implement the fermion spin vector into the amplitude. Define amplitudes for right- and left-handed fermions,

$$|\mathcal{M}_{RR}|^2 = |\mathcal{M}|^2(\lambda_{e^-} = +1, \lambda_{e^+} = +1),$$

$$|\mathcal{M}_{LR}|^2 = |\mathcal{M}|^2(\lambda_{e^-} = -1, \lambda_{e^+} = +1),$$

similarly for $|\mathcal{M}_{LL}|^2$ and $|\mathcal{M}_{RL}|^2$. For longitudinally polarized fermion beams, the amplitude can be separated into parts proportional to the four combinations of fermion helicity,

$$\begin{aligned} |\mathcal{M}|^2 = & \frac{1}{4} \{ (1 + P_{e^-})(1 + P_{e^+}) |\mathcal{M}_{RR}|^2 \\ & + (1 - P_{e^-})(1 - P_{e^+}) |\mathcal{M}_{LL}|^2 \\ & + (1 + P_{e^-})(1 - P_{e^+}) |\mathcal{M}_{RL}|^2 \\ & + (1 - P_{e^-})(1 + P_{e^+}) |\mathcal{M}_{LR}|^2 \}, \end{aligned} \quad (\text{A13})$$

where P_{e^\pm} is the degree of electron and positron polarization,

$$P_{e^\pm} = \frac{n_{e_{R/L}^\pm} - n_{e_{L/R}^\pm}}{n_{e_{R/L}^\pm} + n_{e_{L/R}^\pm}}, \quad (\text{A14})$$

and $n_{e_{R/L}^\pm}$ denote the number of left- and right-handed electrons and positrons in each beam. An unpolarized beam has $P_{e^\pm} = 0$, and $P_{e^\pm} = \pm 1$ are 100% left- and right-handed polarized beams, respectively [47].

2. Photon polarization vectors

A spin-1 particle with four-momentum $k^\mu = (k_0, \mathbf{k})$ and helicity λ , moving in arbitrary direction has polarization four-vector basis,

$$\epsilon_1^\mu(k) = \frac{1}{|\mathbf{k}|k_T} (0, k_x k_z, k_y, k_z, -k_T^2), \quad (\text{A15})$$

$$\epsilon_2^\mu(k) = \frac{1}{k_T} (0, -k_y, k_x, 0), \quad (\text{A16})$$

$$\epsilon_3^\mu(k) = \frac{k_0}{|\mathbf{k}|\sqrt{k^2}} \left(\frac{\mathbf{k}^2}{k_0}, k_x, k_y, k_z \right), \quad (\text{A17})$$

$$\epsilon_4^\mu(k) = \frac{1}{\sqrt{k^2}} (k_0, k_x, k_y, k_z), \quad (\text{A18})$$

where $k_T = \sqrt{k_x^2 + k_y^2}$ [48].

The longitudinal polarization vector for a boson with $\lambda = 0$ is described by Eq. (A17) and helicity eigenvectors with $\lambda = \pm 1$ are

$$\epsilon^\mu(k, \lambda) = \frac{1}{\sqrt{2}} [-\lambda \epsilon_1^\mu(k) - i \epsilon_2^\mu(k)]. \quad (\text{A19})$$

The amplitude can be separated using the fraction of left- to right-handed outgoing photons P_γ ,

$$|\mathcal{M}|^2 = \frac{1}{2} \{ (1 + P_\gamma) \gamma_R + (1 - P_\gamma) \gamma_L \}, \quad (\text{A20})$$

where $\gamma_L = |\mathcal{M}|^2(\lambda_\gamma = -1)$ and $\gamma_R = |\mathcal{M}|^2(\lambda_\gamma = +1)$.

3. Hidden photon

The amplitude of the hidden photon production has two contributions,

$$\begin{aligned} \mathcal{M}_1 = & \frac{ec_X}{t - m_e^2} \epsilon_\mu(q_1, \lambda_\gamma) \epsilon_\nu(q_2, \lambda_X) \\ & \times \bar{v}(p_2, \lambda_{e^+}) \gamma^\mu (\not{p}_1 - \not{q}_1 + m_e) \gamma^\nu u(p_1, \lambda_{e^-}), \end{aligned} \quad (\text{A21})$$

$$\begin{aligned} \mathcal{M}_2 = & \frac{ec_X}{u - m_e^2} \epsilon_\nu(q_1, \lambda_\gamma) \epsilon_\mu(q_2, \lambda_X) \\ & \times \bar{v}(p_2, \lambda_{e^+}) \gamma^\mu (\not{p}_1 - \not{q}_2 + m_e) \gamma^\nu u(p_1, \lambda_{e^-}). \end{aligned} \quad (\text{A22})$$

Using FeynCalc [49], the matrix amplitude squared is calculated using the momenta, spin and polarization vectors described previously. As the hidden photon is undetectable, we will sum over its spin using its mass m_X . The leading term $\mathcal{O}(m_e^0)$ of the final total amplitude is given below in Eq. (A27), where

$$E_\pm = E_1 \pm E_2 \quad \text{and} \quad \beta_X^2 = 1 + \tau_X^2. \quad (\text{A23})$$

Two separate expressions containing the incoming fermion helicities emerge; $(\lambda_{e^-} - \lambda_{e^+})$ and $(\lambda_{e^-} \lambda_{e^+} - 1)$ which both go to zero for $\lambda_{e^+} = \lambda_{e^-}$. Hence, the helicities are required to be opposite and any dependence on them being equal are of order $\mathcal{O}(m_e^2)$. The effects of the polarization of the incoming beams on the amplitude is investigated using,

$$\mathcal{R}(P_{e^-}, P_{e^+}) = \frac{|\mathcal{M}|^2(P_\gamma = 0)}{|\mathcal{M}|^2(P_\gamma = P_{e^\pm} = 0)}, \quad (\text{A24})$$

where $\mathcal{R}(P_{e^-}, P_{e^+}) = (1 - P_{e^-} P_{e^+})$ at $\mathcal{O}(m_e^0)$ and the cross section is maximally enhanced for fully oppositely polarized beams. The photon polarization describes the ratio of right-handed to left-handed photons,

$$\alpha_\gamma = \frac{\gamma_L - \gamma_R}{\gamma_L + \gamma_R}, \quad (\text{A25})$$

which at order $\mathcal{O}(m_e^0)$ is found to be

$$\alpha_\gamma = \frac{(P_{e^-} - P_{e^+})}{(P_{e^-} P_{e^+} - 1)} f(E_1, E_2, m_X), \quad (\text{A26})$$

where largest α_γ is found for small m_X and fully oppositely polarized beams. As seen in the helicity fraction distributions for the hidden photon (see Sec. V) it is possible to significantly affect the photon polarization by applying angular cuts in the forwards or backwards directions with $|\alpha_\gamma| \geq 0.8$ for $m_X \leq 4$ GeV.

$$|\mathcal{M}|^2 = -\frac{c_X^2 e^2 \csc^2 \theta}{2s(\tau_X - 1)^2} \left\{ \beta_X^2 \lambda_\gamma (\lambda_{e^-} - \lambda_{e^+}) [E_+ E_- (\cos(2\theta) + 3) + (4E_+^2 - 2s) \cos \theta] \right. \\ \left. - 2(\lambda_{e^-} \lambda_{e^+} - 1) \left[-2\beta_X^2 E_+ E_- \cot \theta + \sin \theta \left(\beta_X^2 \left(E_+^2 - \frac{s}{2} \right) - s\tau_X \right) - \beta_X^2 (2E_+^2 - s) \csc \theta \right] \right\}. \quad (\text{A27})$$

4. Axion

The matrix amplitude of the axion production has two t -channel contributions together with a third s -channel contribution coming from the photon coupling,

$$\mathcal{M}_1 = \frac{ec_e}{f(t - m_e^2)} \epsilon_\beta^*(q_1, \lambda_\gamma) \\ \times \bar{v}(p_2, \lambda_{e^+}) \gamma^5 (\not{p}_1 - \not{q}_1 + m_e) \gamma^\beta u(p_1, \lambda_{e^-}), \quad (\text{A28})$$

$$\mathcal{M}_2 = \frac{ec_e}{f(u - m_e^2)} \epsilon_\beta^*(q_1, \lambda_\gamma) \\ \times \bar{v}(p_2, \lambda_{e^+}) \gamma^\beta (\not{p}_1 - \not{q}_2 + m_e) \gamma^5 u(p_1, \lambda_{e^-}), \quad (\text{A29})$$

$$\mathcal{M}_3 = \frac{ec_\gamma}{fs} g_{\mu\nu} \epsilon_\beta^*(q_1, \lambda_\gamma) \\ \times \bar{v}(p_2, \lambda_{e^+}) \gamma^\mu u(p_1, \lambda_{e^-}) \epsilon^{\nu\rho\sigma} (p_2 + p_1)_\rho (q_1)_\sigma. \quad (\text{A30})$$

The axion coupling to fermions c_e is proportional to m_e , hence its contribution will be proportional to m_e^2 , and the interference between the two channels is suppressed by a factor of m_e and is negligible. Following the same procedure as the hidden photon calculations, the amplitude is derived accounting for contributions from axion coupling to electrons, Eq. (A31), axion coupling to photons, Eq. (A32), and the interference between them, Eq. (A33). As described earlier, the two axion couplings have different dependencies on the axion mass. The s -channel contribution is directly proportional to $(m_a^2 - s)^2$, which for $m_a^2 \ll s$ simplifies to s^2 , and as m_a increases the expression tends to zero. The t -channel has a more complicated dependence, where for small axion masses an overall suppression by $1/s$ emerges, and the s -channel dominates over the t -channel, but as the mass increases and $m_a \sim \sqrt{s}$, the expression blows up and the t -channel starts to dominate.

$$|\mathcal{M}|_t^2 = \frac{c_e^2 e^2 \csc^2 \theta}{2f^2 s (\tau_X - 1)^2} (E_- \cos \theta + E_+)^2 [(\lambda_{e^-} \lambda_{e^+} + 1)(1 + \tau_X^2) + \lambda_\gamma (\lambda_{e^-} + \lambda_{e^+})(\tau_X^2 - 1)], \quad (\text{A31})$$

$$|\mathcal{M}|_s^2 = \frac{c_\gamma^2 e^2 s (\tau_X - 1)^2}{32f^2 (E_- \cos \theta + E_+)^2} \left\{ 2\lambda_\gamma (\lambda_{e^-} - \lambda_{e^+}) \left[4E_+^2 \cos \theta + \left(E_+^2 - \frac{s}{2} \right) (\cos \theta + 1)^2 \right] \right. \\ \left. + (\lambda_{e^-} \lambda_{e^+} - 1) \left[4E_+^2 \cos \theta - \left(E_+^2 - \frac{s}{2} \right) (\cos \theta + 1)^2 \right] \right\}, \quad (\text{A32})$$

$$|\mathcal{M}|_i^2 = \frac{c_\gamma c_e e^2 m_e \csc \theta}{2f^2 s^2} \left\{ \lambda_\gamma \left[s(\lambda_{e^-} - \lambda_{e^+}) \left(E_+^2 - \frac{s}{2} \right) (\cos \theta + 1)^2 (1 - \tau_X) - \frac{1}{2} s^2 (\lambda_{e^-} + \lambda_{e^+}) (\tau_X + 1) \sin^2(\theta) \right] \right. \\ \left. + s(\tau_X - 1) [\lambda_{e^-} \lambda_{e^+} (E_- + E_+ \cos \theta)^2 - (E_- \cos \theta + E_+)^2] - E_+^2 [\cos(2\theta) + 3] \right\}. \quad (\text{A33})$$

The leading-order terms for the contribution from the axion coupling to photons require the incoming fermions to have opposite helicity, and any contributions from equal helicities are of $\mathcal{O}(m_e^2)$. Unlike the hidden photon and s -channel contribution, the t -channel contribution from the axion coupling to fermions requires equal beam helicities whereas opposite helicities only give rise to terms $\mathcal{O}(m_e^4)$. Hence, with 100% polarized beams one can distinguish, to

leading-order approximation, between the two axion couplings. The s -channel contribution gives $\mathcal{R}(P_{e^-}, P_{e^+}) = (1 - P_{e^-} P_{e^+})$ at $\mathcal{O}(m_e^0)$, while the t -channel contribution gives rise to $\mathcal{R}(P_{e^-}, P_{e^+}) = (1 + P_{e^-} P_{e^+})$ at $\mathcal{O}(m_e^2)$, both with little to no corrections from higher orders. The full axion amplitude has a more complex structure with the two main components, from each axion coupling, having opposite polarization dependencies. As axion coupling to

photons dominates for small masses, for equal couplings, the maximum enhancement is found for opposite helicities, whereas the opposite becomes true for large mass.

For the axion coupling to photons the photon polarization is

$$\alpha_\gamma = \frac{(P_{e^-} - P_{e^+})}{(P_{e^-} P_{e^+} - 1)} f(E_1, E_2), \quad (\text{A34})$$

which, unlike the case of the hidden-photon expression (A26), does not depend on the axion mass as implied by the helicity fraction distribution. Values of $|\alpha_\gamma| \approx +1$ are achieved for all m_a when applying the forward angular cut ($10^\circ \leq \theta \leq 50^\circ$). For the axion coupling to electrons,

$$\alpha_\gamma = \frac{(P_{e^-} + P_{e^+})(s^2 - m_a^4)}{(P_{e^-} P_{e^+} + 1)(m_a^4 + s^2)}, \quad (\text{A35})$$

and the photon polarization is independent of $\cos\theta$, and maximum magnitude is found for small m_a and fully polarized beams. $|\alpha_\gamma| \approx +1$ for $m_a \leq 2$ GeV with no angular cuts applied. The full axion contribution containing both axion coupling to electron and photons has elements from both, having $(P_{e^-} \pm P_{e^+})$ term in the numerator and $(P_{e^-} P_{e^+} \pm 1)$ in the denominator. For opposite helicities the m_a dependence disappears, whereas for equal helicities the angular dependence goes away, and we recover the contribution from each coupling as expected.

APPENDIX B: STANDARD MODEL BACKGROUND

In Fig. 10, we display the individual $\theta_{\text{lab}} - E_{\text{CMS}}$ distributions for the four distinct contributions to the SM background considered. Each panel has an individual color coding.

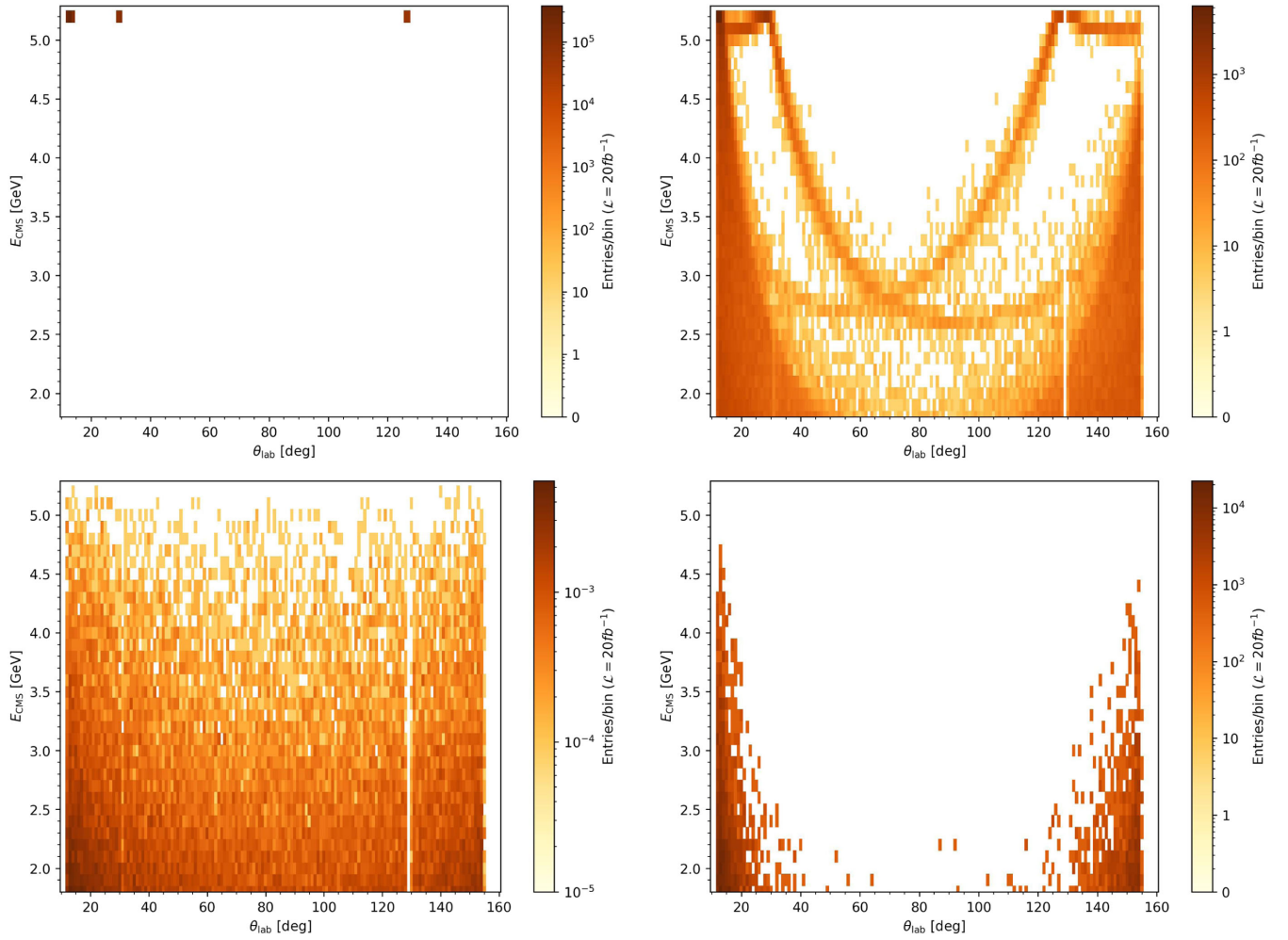


FIG. 10. SM background distributions for $e^+e^- \rightarrow \gamma\gamma$ (upper left), $\gamma\gamma\gamma$ (upper right), $\nu\bar{\nu}\gamma(\gamma)$ (bottom left), and $e^+e^-\gamma(\gamma)$ (bottom right).

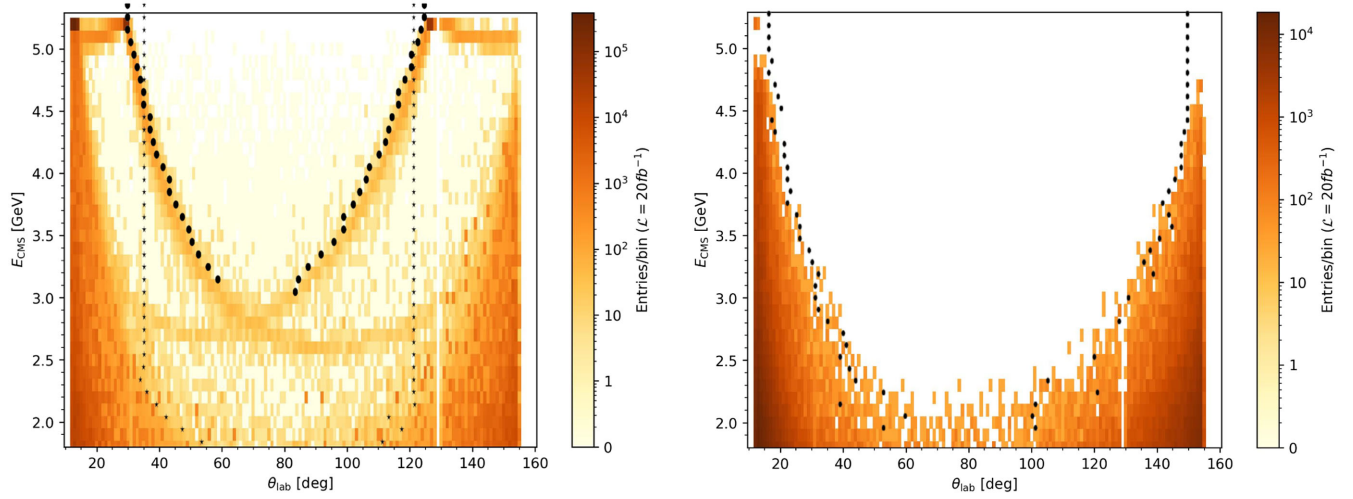


FIG. 11. Points used for unpolarized (left) with low mass (black circles) and high mass (black stars), and equal-beam polarization (right) background.

APPENDIX C: EXCLUSION LIMITS

In order to optimize the signal to background ratio, fits were performed, creating $\theta_{\text{lab}} - E_{\text{CMS}}$ cuts with the aim of selecting the regions with the least number of background events. In the following, we will present the work carried out to perform the fits. When observing the background in Fig. 5, two prominent regions stand out; the V-shape created by the $\gamma\gamma$ final-state bands, and the areas between the main $e^+e^-\gamma(\gamma)$ background and the bands. The signal for low-mass invisible states falls within the first region, but for increasing masses the second region has to be considered. Two fits are therefore performed. For the high mass fit, the points were found for each photon energy; going from small and large θ_{lab} values respectively, finding the first bin with less than 10 events. A similar procedure was done for the low mass fit, this time going from the middle ($\theta_{\text{lab}} \approx 80^\circ$) for both decreasing and increasing θ_{lab} , the last bin with fewer events than our threshold were found. Furthermore, the three peaks from the $\gamma\gamma$ final state were of particular interest to avoid. This was ensured by finding

the coordinates of the two center peaks and continuing these points downwards (constant θ_{lab} but decreasing E_{CMS}) until reaching the points found as described above. As the outgoing photon in the case of large masses has very low energy, the fit was extended downwards into the area of increased background by including the range of $[60, 100]^\circ$. The points found are displayed on of the relevant distributions in Fig. 11. The fits were found using *Mathematica* [50] for generic polynomial,

$$c_1 + c_2x + c_3x^2 + c_4\sqrt{x} + c_5x^{-1/2} + c_6x^{-1} + c_7x^{-2}.$$

The cut functions for the unpolarized SM background are described by Eqs. (C1) and (C2) for low and high m_X respectively. For the equal-beam helicity, we saw that the photon-only final states were not present, and therefore this opens a big area with little to no background (only from neutrino final states). This allows for a single, much wider fit which include more signal events. For equal-beam polarization, the cut function can be found in Eq. (C3).

$$E_{\text{CMS,low}}(\theta_{\text{lab}}) = -1.753 \times 10^4 - 61.42\theta_{\text{lab}} + 4.708 \times 10^{-2}\theta_{\text{lab}}^2 + 1.572 \times 10^3\sqrt{\theta_{\text{lab}}} + \frac{9.795 \times 10^4}{\sqrt{\theta_{\text{lab}}}} - \frac{2.379 \times 10^5}{\theta_{\text{lab}}} + \frac{6.994 \times 10^5}{\theta_{\text{lab}}^2}, \quad (\text{C1})$$

$$E_{\text{CMS,high}}(\theta_{\text{lab}}) = -2.601 \times 10^4 - 1.109 \times 10^2\theta_{\text{lab}} + 0.1001\theta_{\text{lab}}^2 + 2.583 \times 10^3\sqrt{\theta_{\text{lab}}} + \frac{1.304 \times 10^5}{\sqrt{\theta_{\text{lab}}}} - \frac{2.826 \times 10^5}{\theta_{\text{lab}}} + \frac{6.573 \times 10^5}{\theta_{\text{lab}}^2}, \quad (\text{C2})$$

$$E_{\text{CMS,equal}}(\theta_{\text{lab}}) = -2.929 \times 10^3 - 13.66\theta_{\text{lab}} + 1.293 \times 10^{-2}\theta_{\text{lab}}^2 + 3.061 \times 10^2\sqrt{\theta_{\text{lab}}} + \frac{1.379 \times 10^4}{\sqrt{\theta_{\text{lab}}}} - \frac{2.763 \times 10^4}{\theta_{\text{lab}}} + \frac{5.284 \times 10^4}{\theta_{\text{lab}}^2}. \quad (\text{C3})$$

APPENDIX D: HELICITY FRACTION DISTRIBUTIONS

In the following, we present explanations with visual aid for the helicity fraction distributions shown in this paper. The observable considered is the fraction of the final-state photon with helicity matching that of the incoming electron beam as a function of the angle of the photon,

$$P = \frac{|\mathcal{M}|^2(2\lambda_{e^-} = \lambda_\gamma = 1)}{|\mathcal{M}|^2(\lambda_{e^\pm} = \lambda_\gamma = 0)}, \quad (\text{D1})$$

where $|\lambda_\gamma| = 1$ is the helicity of the photon, and $|\lambda_{e^-}| = 1/2$ the helicity of the electron. Hence at each $\cos\theta_{\text{lab}}$ value, the distribution indicates that for X photons detected at this angle $P \times X$ of them will have the same helicity as the incoming electron beam, and $(P - 1) \times X$ will have the opposite helicity.

The representations below are 2D diagrams with a horizontal z -axis and arbitrary vertical axis in the x - y plane. Without loss of generality, the electron is assumed to have positive helicity, and two combinations of incoming beam helicities are possible; the electron and positron spin vectors are parallel ($\lambda_{e^+} = -\lambda_{e^-}$) or antiparallel ($\lambda_{e^+} = \lambda_{e^-}$).

1. Standard Model background

The helicity fraction distribution for the background has contributions from every SM background process, each with their own tendencies which will be described in this section. The simplest case is $e^+e^- \rightarrow \gamma\gamma$, where the incoming electron and positron must have parallel spin vectors (opposite helicities) due to the emission of two photons not changing the direction of the spin (see Fig. 12). The z -components of the spin vectors of the outgoing photons will have the same sign as the incoming fermions. The direction of the photon will determine its helicity; if emitted in the same direction as the electron, the helicity will match that of the electron, and vice versa if emitted in the positron direction. Hence the helicity fraction distribution has the s-shape seen in Fig. 13 due to the photon being more likely to travel in the same direction as the fermion which it was emitted from. At $\cos\theta_{\text{lab}} = -1$, all photons detected will have the same helicity as the incoming electron, whereas in the opposite direction,

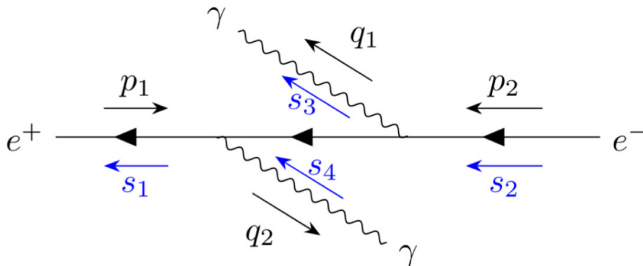


FIG. 12. Helicity visualization for $e^+e^- \rightarrow \gamma\gamma$.

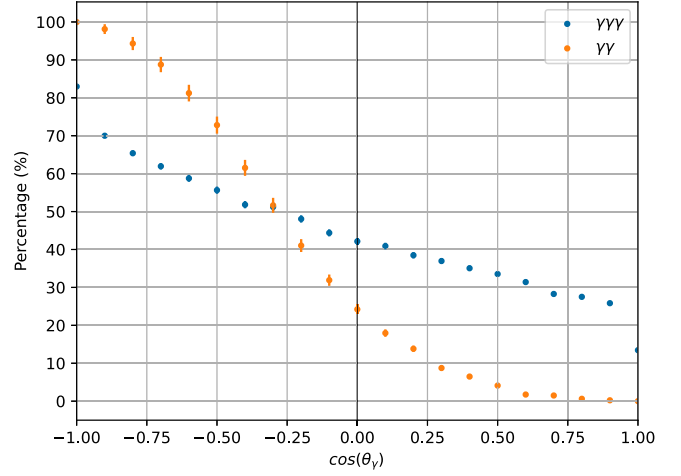


FIG. 13. Helicity fraction distributions for $e^+e^- \rightarrow \gamma\gamma(\gamma)$ with no cuts.

$\cos\theta_{\text{lab}} = +1$, none of them will and their helicity match the incoming positron. The skew away from the center line is due to the unequal beam energies at Belle II, for equal energies the shape will be centered around $(0, 50\%)$.

The spin vector diagram looks similar for $e^+e^- \rightarrow \gamma\gamma\gamma$, the fermions are still required to have opposite helicities, but the helicities of the outgoing photons are more random due to the 3-particle final state. Therefore the correlation between direction and helicity fraction reduces, and the distribution is closer to 50%. Figure 13 displays the helicity fraction distribution for $e^+e^- \rightarrow \gamma\gamma(\gamma)$ without any cuts applied. When implementing the criteria that only one photon can be detected, the few cases allowed for $e^+e^- \rightarrow \gamma\gamma$ are due to the asymmetric angular coverage of the detector and the gaps between the barrel and end-caps. As seen on the lower right of the left panel on Fig. 8, this corresponds to the two points around $\cos\theta \approx -1$ and another set at $\cos\theta \approx 0.6$.

The helicities involved in the main background process, $e^+e^- \rightarrow e^+e^-\gamma(\gamma)$, are much more complex as consists of many contributions and interferences between them, some of which require the incoming helicities to be opposite and others equal. The resulting angular distribution, upper left on the left panel of Fig. 8, does not show any directional dependence. Similar arguments can be made for $e^+e^- \rightarrow \nu\bar{\nu}\gamma(\gamma)$, and as seen on the upper right of Fig. 8, the process has a slight $\cos\theta$ dependence, but mostly resides around 50%.

As $e^+e^- \rightarrow e^+e^-\gamma$ is the dominating process, the full background results, displayed in the lower-left of Fig. 8, looks very similar. The slight upwards motion at $\cos\theta \approx -1$ and outlying point around $\cos\theta \approx 0.6$ are due to the large $e^+e^- \rightarrow \gamma\gamma$ contribution.

2. Hidden photon

The hidden photon interacts similarly to the SM photon, and hence the incoming fermions are required to have

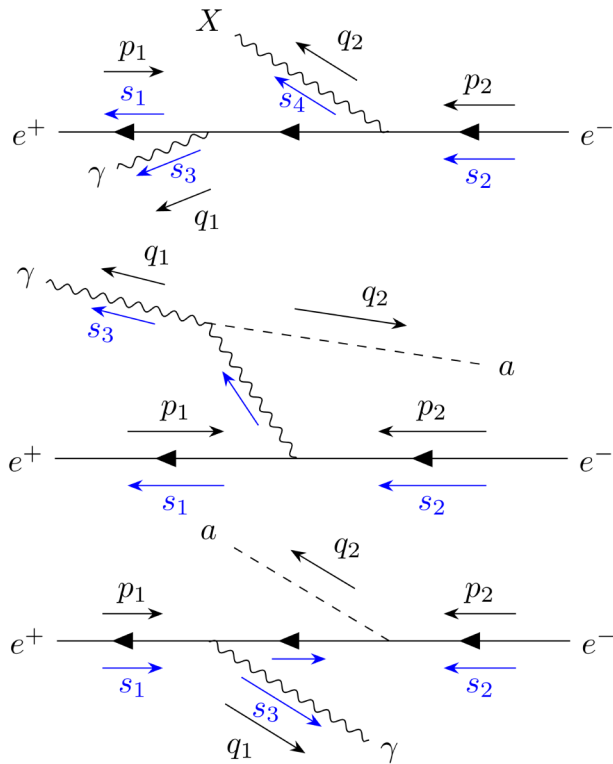


FIG. 14. Helicity visualization for the production of hidden photons for large m_X (top), axion-photon (middle), and axion-electrons (bottom) couplings.

opposite helicities. Two scenarios are possible; the photon and hidden photon travel in opposite directions along the z -axis (Fig. 12 with one γ replaced with X), or due to the unequal beam energies they travel in the same direction (see top panel on Fig. 14). The direction of the spin vectors involved remain the same, but as the momentum changes the photon helicity differs between the two scenarios. For small m_X effectively only scenario one happens, resulting in the photon direction and helicity being strongly coupled; the helicity of the photon will match the helicity of the fermion travelling in the same direction. But as the hidden photon mass increases, the photon becomes softer and its direction more random. This introduces occurrences of scenario two, pushing the helicity fraction distribution towards 50%. The results for the hidden photon can be seen on the right panel on Fig. 8.

3. Axions

As the two axion contributions have different spin structures, they will be described separately. For axion coupling to photons, the incoming fermions annihilate into a virtual photon which then emits an axion. The incoming fermions are required to have the same direction spin vector which the virtual photon inherits. The outgoing photon

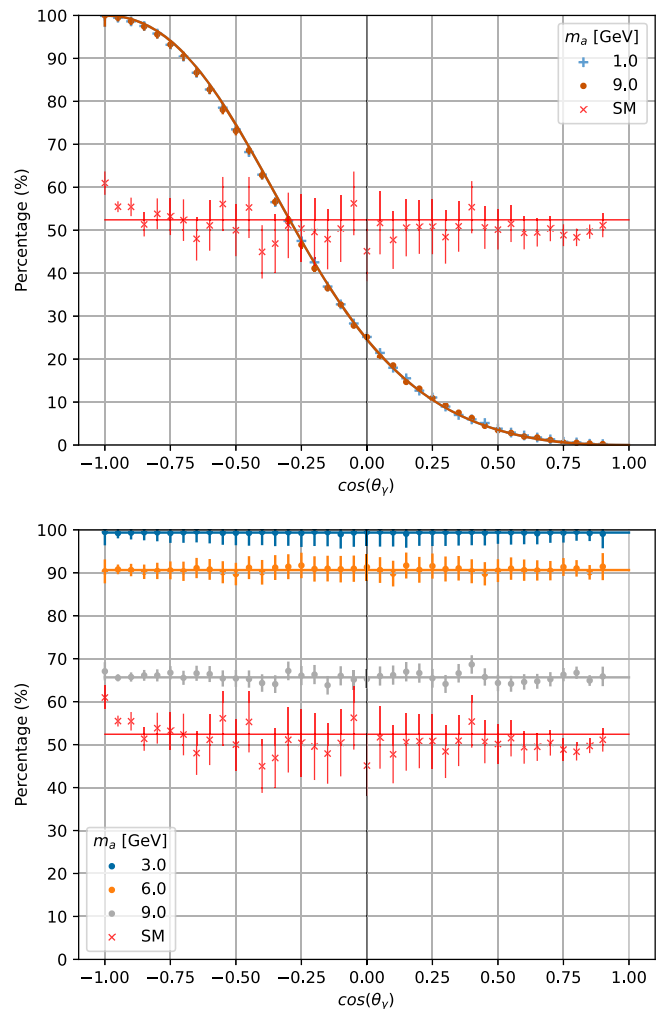


FIG. 15. Helicity fraction distributions for axion-photon (top) and axion-electron (bottom) couplings.

helicity is determined by its direction after the axion emission, see middle panel of Fig. 14, and depends on whether it is in the same direction as the electron or positron. Hence the helicity fraction distribution, top panel on Fig. 15, is the s-shape seen before. The distribution is not influenced by the mass of the axion, as it only changes the coupling strength. The axion coupling to electrons consists of t/u -channel diagrams like the hidden photon, but as the interaction introduces a helicity flip for the incoming fermions (bottom panel of Fig. 14), their helicities are equal. Therefore, the photon helicity will match both fermions or neither, resulting in no angular dependence as seen on bottom panel on Fig. 15. As described before, as the axion mass increases the photon direction becomes more random resulting in the helicity fraction approaching 50%.

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