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Observations of Typhoon Generated Gravity Waves

4

from the CIPS and AIRS instruments and comparison to the high-resolution ECMWF model.

5

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15

16 **Key Points**

17

(1) CIPS and AIRS measurements of concentric GWs generated by Typhoon Yutu are used to

18

verify ECMWF GWs in the altitude range of 30-55 km.

19

(2) Analysis of GW observations from CIPS and AIRS as well as ECMWF data provides detailed

20

characteristics of concentric GWs.

21

(3) Differences in GW amplitudes among five typhoon cases grow as GWs propagate upward,

22

indicating the importance of background winds.

23

24

25 **Abstract**

26 The satellite-based Cloud Imaging and Particle Size (CIPS) instrument and Atmospheric
27 Infrared Sounder (AIRS) observed concentric gravity waves (GWs) generated by Typhoon Yutu
28 in late October, 2018. This work compares CIPS and AIRS nadir viewing observations of GWs at
29 altitudes of 50-55 km and 30-40 km, respectively, to simulations from the high-resolution
30 European Centre for Medium-Range Weather Forecasting Integrated Forecasting System
31 (ECMWF-IFS) and ECMWF reanalysis v5 (ERA5). Both ECMWF-IFS with 9 km and ERA5 with
32 31 km horizontal resolution show concentric GWs at similar locations and timing as the AIRS and
33 CIPS observations. The GW wavelengths are ~225-236 km in ECMWF-IFS simulations, which
34 compares well with the wavelength inferred from the observations. After validation of ECMWF
35 GWs, five category 5 typhoon events during 2018 are analyzed using ECMWF to obtain
36 characteristics of concentric GWs in the Western Pacific regions. The amplitudes of GWs in the
37 stratosphere are not strongly correlated with the strength of typhoons, but are controlled by
38 background wind conditions. Our results confirm that amplitudes and shapes of concentric GWs
39 observed in the stratosphere and lowermost mesosphere are heavily influenced by the background
40 wind conditions.

41

42 **Plain Language Summary**

43 Atmospheric gravity waves (GWs) have an important role in coupling the different
44 atmospheric layers. One of the main sources of GWs is convection such as typhoons. In the
45 stratosphere, these GWs frequently appear as concentric (or ring-shaped) patterns in nadir-viewing
46 satellite measurements. In this work, data from two such nadir viewing satellite instruments, the

47 Cloud Imaging and Particle Size (CIPS) instrument and Atmospheric Infrared Sounder (AIRS),
48 are analyzed to study the GWs generated by typhoon Yutu, which occurred in late October 2018.
49 CIPS and AIRS observe at altitudes of 50-55 km and 30-40 km, respectively, providing us a unique
50 opportunity to study concentric GWs at two different altitudes. The satellite observations are then
51 used to validate the GW-resolving high-resolution European Centre for Medium-Range Weather
52 Forecasting (ECMWF) model. Utilizing ECMWF simulations, 4 more typhoon events were
53 analyzed. The results indicate that the amplitudes of concentric GWs in the stratosphere are not
54 correlated with the strength of typhoons. However, the amplitudes and shapes of concentric GWs
55 observed in the stratosphere and lower mesosphere are found to be influenced by the background
56 wind conditions. This work provides an understanding of the relative importance of GW source
57 strength and background wind conditions.

58

59 **1. Introduction**

60 Gravity waves (GWs) transport momentum and energy from the lower atmosphere to the
61 upper atmosphere and drive atmospheric circulations [e.g., Fritts & Alexander, 2003; Alexander
62 et al., 2010]. The main sources of GWs are frontal systems, convection, orography, and
63 spontaneous emission from unstable jets. Convectively generated GWs have been studied using
64 various satellite and ground-based observations and high-resolution models [e.g., Alexander &
65 Pfister, 1995; Kim & Chun, 2011; Yue et al., 2009]. Strong convective activity, including
66 hurricanes and typhoons, generates GWs that appear as concentric (or ring-shaped) patterns in
67 nadir-viewing satellite measurements, ground-based imager observations and high-resolution
68 simulations [e.g., Yue et al., 2009, 2013; Kim et al., 2009; Liu et al., 2014; Gong et al., 2015].
69 Convectively generated GWs can propagate up to the mesosphere and lower thermosphere (MLT),

70 and even higher into the thermosphere and ionosphere, and cause large disturbances in winds and
71 electron densities in the upper atmosphere [e.g., Vadas & Liu, 2013; Liu & Vadas, 2013; Azeem
72 et al., 2015]. Furthermore, typhoon-generated GWs can feed back to the development of the
73 typhoon itself [e.g., Kim & Chun, 2011; Hoffmann et al., 2018]; by affecting wind structure,
74 typhoon-generated GWs can intensify the typhoon during the development stage of the typhoon
75 [Kim & Chun, 2011]. Therefore, it is important to characterize these convectively generated GWs
76 in order to understand their impacts on the lower and also upper atmosphere.

77 The nadir-viewing Cloud Imaging and Particle Size (CIPS) instrument on the Aeronomy
78 of Ice in the Mesosphere (AIM) satellite measures perturbations in Rayleigh scattered ultraviolet
79 radiation that are indicative of GW-induced variations at altitudes of ~50-55 km [Randall et al.,
80 2017; Forbes et al., 2021; 2022]. These Rayleigh Albedo Anomaly (RAA) data show clear
81 concentric GWs in late October 2018, coincident with the occurrence of Typhoon Yutu. The nadir-
82 viewing Atmospheric Infrared Sounder (AIRS) instrument measures perturbations in brightness
83 temperature that are indicative of GW-induced variations at altitudes of ~30-40 km. [e.g.,
84 Hoffmann et al., 2013; Gong et al., 2015]. AIRS data also revealed concentric GWs in late October
85 2018. By combining analyses of data from both CIPS and AIRS, we can probe altitude variations
86 of concentric GWs in the middle stratosphere and lowermost mesosphere.

87 In this work, we combine CIPS and AIRS observations to study GWs generated by typhoon
88 Yutu. The high-resolution European Centre for Medium-Range Weather Forecasts Integrated
89 Forecasting System (ECMWF-IFS) dataset is also used to supplement the observational study.
90 Gong et al. [2015] conducted a study of concentric GWs with both AIRS and ECWFMF, using
91 automated detections of concentric GWs globally to investigate hemispheric differences of
92 concentric GWs in July and January. Their study showed that concentric GW phases and

93 wavelengths from ECMWF simulations and AIRS observations are comparable; however,
94 amplitudes of ECMWF GWs are weaker than those from AIRS observations. Since the work by
95 Gong et al. [2015], the ECMWF horizontal resolution has been improved from 16 km to 9 km, and
96 the number of vertical levels has been increased from 91 levels to 137 levels, with the same top
97 pressure level at 0.01 hPa. The hourly ECMWF Reanalysis v5 (ERA5) has ~31-km horizontal
98 resolution and also has been used for GW studies [e.g., Cullens & Thurairajah, 2021]. This work
99 evaluates GWs resolved by the updated version of ECMWF-IFS and ERA5 simulations via
100 comparisons against CIPS and AIRS observations. We also focus on individual strong typhoon
101 events in ECMWF, and obtain characteristics of typhoon-generated GWs including wavelengths,
102 altitude variations and time evolution of GWs.

103 In this paper, explanations of data sets are summarized in Section 2. Section 3 provides
104 results of GWs generated by Typhoon Yutu in October 2018 from both satellite observations and
105 ECMWF simulations. GW characteristics from five different typhoon events are summarized in
106 Section 4. Conclusions are presented in Section 5.

107

108 **2. Data**

109 **2.1. ECMWF**

110 ECMWF provides global analysis of atmospheric data from the ground to the lower
111 mesosphere. The cycle 41r2 version of the ECMWF-IFS high-resolution model is used here and
112 referred as ECMWF. ECMWF has ~9 km horizontal resolution with 137 vertical levels from the
113 ground to 0.01 hPa. This version of ECMWF is capable of resolving mesoscale GWs with
114 horizontal scales larger than ~70 km [Preusse et al., 2014]. Although ECMWF extends up to ~80
115 km, amplitudes of GWs are suppressed above ~45 km due to a sponge layer [Schroeder et al., 2009,

116 Ehard et al., 2017; Gisinger et al., 2022]. Earlier versions of ECMWF GWs have been validated
117 against AIRS, SABER, and COSMIC satellite observations, lidar and radiosonde ground-based
118 observations, and other high-resolution models including the Weather Research and Forecasting
119 (WRF) model [e.g., Alexander & Teitelbaum, 2007; Schroeder et al., 2009; Kim et al., 2009; Wu
120 & Eckermann, 2008; Yamashita et al., 2010]. The 6-hourly ECMWF is used in this study. The
121 hourly ECMWF Reanalysis v5 (ERA5) dataset provided by ECMWF, with ~31 km horizontal
122 resolution, is also used in this work for comparison [Hersbach et al., 2020]. Temporal variations
123 of GWs from ERA5 have been validated previously against COSMIC GW data [e.g., Cullens et
124 al., 2021].

125 Both ERA5 and ECMWF GW perturbations are estimated by removing large-scale
126 variations from the zonal wind (u) and vertical wind (w). Large-scale variations are estimated by
127 averaging over 15° longitude x 15° latitudes regions at each grid and each pressure level for both
128 ERA5 and ECMWF. Approximate altitudes are estimated using geopotential height. It should be
129 noted here that AIRS data is used in assimilation process for ECMWF data product. Yamashita et
130 al. [2010] compared ECMWF GW amplitudes between analysis and various forecast time to
131 examine assimilation influences on resolved GWs in ECMWF. Although Yamashita et al. [2010]
132 presented no clear sign of increase in GW amplitudes by assimilated process for climatology of
133 GWs, further analysis of assimilation influences on GW need to be conducted in the future.

134

135 **2.2. AIM-CIPS**

136 The CIPS instrument is onboard the AIM satellite, which was launched in April 2007
137 [Rusch et al., 2009; Russell et al., 2009]. CIPS observes Polar Mesospheric Clouds (PMCs) in the
138 summer by subtracting the background Rayleigh scattering from the ice particle scattering by

139 PMCs (if present) [Bailey et al., 2009; Lumpe et al., 2013]. In the absence of PMCs, perturbations
140 to the observed Rayleigh scattering signal, reported as RAA, are indicative of GW-induced
141 variations at 50-55 km altitude. The data are measured by four cameras with a total cross-track by
142 along-track field of view of $80^\circ \times 120^\circ$ ($\sim 1000 \text{ km} \times 2000 \text{ km}$), and a resolution of $7.5 \text{ km} \times 7.5$
143 km. The CIPS-RAA measurement spans $\sim 40^\circ$ – 85° latitude in the spring and summer hemisphere
144 during 2007-2015. Due to a change in operating mode in February 2016, the CIPS observations
145 after this change are global in sunlit regions year-round, except where PMCs are present [Randall
146 et al., 2017].

147 To quantify GW-induced perturbations to the Rayleigh scattering signal (controlled by
148 atmospheric neutral density and ozone), the RAA is calculated as the difference between the
149 observed Rayleigh scattering albedo and a baseline albedo that would be observed in the absence
150 of any small-scale atmospheric variations. More information on the retrieval process can be found
151 in Randall et al. [2017]. The RAA retrievals are sensitive to GW perturbations with horizontal
152 wavelengths of ~ 15 – 600 km and with vertical wavelengths $\geq 15 \text{ km}$. CIPS RAA data used in this
153 work are available from the CIPS website (<http://lasp.colorado.edu/aim/>) [Randall et al., 2017].

154

155 **2.3. AIRS**

156 AIRS is a nadir-sounding infrared radiometer on NASA's Aqua satellite. Data is available
157 from 2002 onwards [Aumann et al., 2003]. Stratospheric GWs are derived from emitted radiance
158 measurements in the $4.3 \mu\text{m}$ CO_2 fundamental band, which is most sensitive to 30-40 km altitude.
159 GW perturbations are extracted by removing background temperature and large-scale planetary
160 waves [Hoffmann et al., 2013] using a fourth-order cross-track polynomial. A detailed GW
161 analysis method can be found in Hoffmann et al. [2013], and all processed GWs can be found at

162 <https://data.fz-juelich.de/dataset.xhtml?persistentId=doi:10.26165/JUELICH-DATA/LQAAJA>.

163 AIRS observations are sensitive to GWs with vertical wavelengths longer than 10-15 km and
164 horizontal wavelength longer than 30-80 km.

165

166 **3. Results of Typhoon Yutu Generated Gravity Waves**

167 Typhoon Yutu occurred from October 22-31, 2018 in the western Pacific Ocean (10°N-
168 20°N latitudes and 115°E-155°E longitudes). The path of Typhoon Yutu is shown in Figure 1a and
169 the 10-minute averaged maximum sustained wind speed is shown in Figure 1b. Maximum
170 sustained wind speed is the maximum value of the average wind speed at the surface and is often
171 used to characterize typhoon activity according to the World Meteorological Organization (WMO).
172 Typhoon locations and 10-minute averaged maximum sustained wind speed data are obtained from
173 the Japan Meteorological Agency (JMA) ([http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-](http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-eg/besttrack.html)
174 [hp-pub-eg/besttrack.html](http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-eg/besttrack.html)). Typhoon Yutu traveled westward and reached maximum wind speed
175 on October 24. JMA categorized Typhoon Yutu as a “violent typhoon”, defined as having a
176 maximum sustained wind speed over 105 kt (54 m/s); specifically, the maximum wind speed for
177 Typhoon Yutu was 115 kt, which meets the WMO specification of a “Super Typhoon”. Such
178 typhoons are also referred to as “Category 5 Super Typhoons”.

179 Figure 2 shows GW observations from CIPS (50-55 km altitude) and AIRS (30-40 km) as
180 well as GWs in the simulated vertical wind perturbation at ~50 km and ~30 km altitude from ERA5
181 (31 km horizontal resolution) and high-resolution ECMWF (9 km horizontal resolution). AIRS
182 observations were made on October 27, 2018 at 13:30 LT (~4:30 UT) and CIPS observations were
183 made on October 26, 2018 around 22:25 UT. ECMWF and ERA5 simulations pertain to October
184 27, 2018 at 00 UT; the ECMWF archive their analyses only every 6 hours, so 00 UT is the closest

185 time to the CIPS and AIRS observations. Although ERA5 has an hourly data, to be consistent with
186 ECMWF data, ERA5 was used at the same time as ECMWF data in Figure 2. Both AIRS and
187 CIPS observations capture clear concentric GW structures associated with Typhoon Yutu. The
188 ERA5 and ECMWF simulations both show similar concentric GWs at the same location as the
189 AIRS and CIPS observations. As expected, the higher-resolution ECMWF simulation in Figures
190 2c and 2d shows finer-scale GW structures than the lower resolution ERA5 simulation in Figures
191 2e and 2f. To compare ERA5 and ECMWF GW amplitudes, GW amplitudes are averaged within
192 the red boxes indicated in Figures 2c-f. ECMWF and ERA5 GW amplitude at 10 hPa (~30 km)
193 are 3.2 cm/s and 2.0 cm/s, respectively. Averaged ERA5 GW amplitude is 38% smaller than
194 averaged ECMWF GW amplitude at 10 hPa. At 1 hPa, averaged GW amplitudes are 8.2 cm/s and
195 4.7 cm/s for ECMWF and ERA5, respectively. ERA5 GW amplitudes are 41% smaller than
196 ECMWF GWs.

197 Liu et al. [2014] showed clear convectively generated concentric GWs using a high-
198 resolution version of the Whole Atmosphere Community Climate Model (WACCM) with ~25 km
199 horizontal resolution, indicating that model resolutions of ~25 km can resolve somewhat
200 reasonable concentric GWs generated by typhoons. Gong et al. [2015] compared concentric GW
201 structures and amplitudes simulated with ECMWF using both 25 km and 16 km horizontal
202 resolution. Their results indicate that the structure of the concentric GWs is similar in the two
203 resolution datasets, but the GW amplitudes are larger, and closer to those observed, in the 16 km
204 resolution dataset. Our results are consistent with Gong et al. [2015], so the rest of this work was
205 conducted mainly using the 9 km horizontal resolution ECMWF output.

206 Figure 2 shows that the concentric GWs observed by CIPS at 50-55 km extend farther
207 horizontally than those observed by AIRS at 30-40 km, and this appears to be evident in the

208 ECMWF results as well. To confirm the latter, ECMWF simulations in ~10-km increments from
209 20 to 60 km are shown in Figure 3. Figures 3b and 3d are repeated from Figures 2d and 2c,
210 respectively. As expected, the horizontal extent of concentric GWs expands as the waves propagate
211 upward, consistent with the CIPS and AIRS observations shown in Figures 2a-b. Amplitudes of
212 concentric waves are getting weaker above 0.8 hPa level (~55 km) and wave structures are
213 significantly weakened and disappeared above 0.25 hPa level (~60 km), which we believe is due
214 to damping in ECMWF at higher altitudes by the sponge layers discussed by previous studies [e.g.,
215 Ehard et al., 2017; Gisinger et al., 2022]. Gisinger et al. [2022] observed that above ~45 km the
216 amplitudes of GWs in ECMWF are weaker than in lidar observations. Although GW amplitudes
217 might be reduced above ~45 km, our results indicate that GW structures are reasonably simulated
218 compared to CIPS around 50-55 km.

219 To further compare AIRS, CIPS, and ECMWF GWs, spectrum analysis of horizontal scale
220 is shown in Figure 4. Horizontal wavelength (λ_h) along red lines in Figures 4a, 4c, 4e are ~237 km,
221 247 km, 236 km for AIRS at 35-40 km altitudes, CIPS at 50-55 km altitudes, and ECMWF at ~50
222 km, respectively. It is not shown but ECMWF λ_h at ~30 km altitudes is 225 km. Horizontal
223 wavelengths are consistent between AIRS, CIPS and ECMWF. Averaged λ_h is 240 km.

224 One of the advantages of combining both observations and simulations is to supplement
225 information that these observations cannot obtain, such as the vertical wavelength, time-evolution,
226 and altitudinal variations of GWs. Figure 5 shows the time-evolution of GW signatures in the
227 ECMWF vertical wind at 3 hPa (~40 km) and at 300 hPa (~15 km) from October 23 to October
228 31 at 12 UT. From October 25 (Figure 5b) to October 29 (Figure 5d), concentric GW structures
229 move westward and almost disappear on October 31 (Figure 5e). At lower altitudes at 10 km, GW

230 structures are not as spread-out as GWs at 40 km and also disappear by October 31 (Figure 5k).
231 At 40 km, GW amplitudes maximize around October 27 (Figure 5c).

232 Figure 6a shows a concentric GW signature in the vertical wind at an altitude of ~35 km
233 from the ECMWF simulation. A vertical slice in Figure 6b is made along the white transect in
234 Figure 6a. Clear GW structures are seen from the lower altitudes up to ~60 km. The vertical
235 wavelength estimated from Figure 6b is ~10-14 km, depending on location. GW structures at the
236 higher altitudes are more spread out than those at lower altitudes, which is consistent with the
237 difference in GW structures from CIPS at 50-55 km and AIRS at 30-40 km shown in Figure 2.

238 To better understand changes in GW amplitudes and GW responses to the evolution of a
239 typhoon event, Figure 7 shows temporal variations of the averaged of GW amplitudes in the
240 ECMWF and ERA5 vertical wind (Figure 7a and 7b, respectively), along with the maximum
241 sustained wind speed during Typhoon Yutu (Figure 7c). GW amplitudes are averaged within +/-
242 15° longitude and latitude range from the center of a typhoon event at each time step to focus on
243 typhoon generated GWs. ERA5 provides hourly data in contrast to 6-hourly ECMWF data,
244 resulting in more data points in the ERA5 results in Figure 7b. Both ECMWF and ERA5 GW
245 amplitudes show the largest peak on October 27-28, which is a few days after the peak wind speed
246 on October 24-25. GW amplitudes significantly decay after October 30 as the typhoon weakened.
247 There are several smaller peaks on October 25 and October 29 in Figure 7a, and they are coincident
248 with peaks in wind speeds. To explain the temporal delay in GW amplitudes and typhoon activity,
249 vertical group velocity is calculated using the following equation,

$$250 \quad w_g = - \frac{Nmk}{(k^2 + m^2)^{\frac{3}{2}}} \quad (1)$$

251 where N is The Brunt–Väisälä frequency, m is vertical wavenumber, and k is horizontal
252 wavenumber [e.g., Yue et al., 2009]. Based on typhoon Yutu characteristics (wavelengths) from

253 Figure 6, estimated w_g is 0.9-1.2 m/s, and it takes ~10-12 hours to get to 40 km. Based on this
254 calculations, vertical velocity itself cannot explain the 1-2 days of delay in the peak of gravity
255 wave amplitudes compared to maximum wind speeds. However, small two peaks in GW vertical
256 wind amplitudes on October 25 and 29 can be explained by the strength of typhoon. Background
257 wind conditions that are suitable for wave propagation most likely contributed to the enhancements
258 on October 27.

259 Temporal variations of GW vertical wind amplitude from ERA5 and ECMWF show very
260 similar variations. However, ECMWF shows a peak on October 29 not seen in ERA5 GWs. ERA5
261 GW amplitudes are generally smaller than ECMWF GW amplitudes. Averaged GW amplitudes
262 from October 22 to November 1 at 40 km (30 km) altitudes shown in Figure 7 are 4.4 cm/s (2.6
263 cm/s) and 2.7 cm/s (1.7 cm/s) for ECMWF and ERA5, respectively. ERA5 amplitudes are 39%
264 and 35% smaller than ECMWF GWs at 40 km and 30 km altitudes, respectively, due to coarse
265 resolutions of ERA5.

266 Altitude variations of momentum fluxes are shown in Figure 8 at 2 hPa (~45 km), 5 hPa
267 (~35 km), 10 hPa (~30 km), and 50 hPa (~17 km). Momentum Flux (MF) is calculated using the
268 following equation.

$$269 \quad MF = \overline{\rho_0 u' w'} \quad (2)$$

270 where ρ_0 is atmospheric density, u' is zonal wind perturbation, and w' is vertical wind perturbation.

271 At higher altitudes (Figures 8a-b), the westward (negative) momentum flux is clear on the western
272 side of the concentric GW, whereas at lower altitudes (Figures 8c-d), both the eastern and western
273 sides reveal strong momentum fluxes with similar magnitude, but opposite direction. This anti-
274 symmetric GW pattern that develops at higher altitudes is caused by background wind conditions
275 that affect GW propagation between 17 km and 30 km [e.g., Piani et al.,2000; Yue et al., 2009;

276 Kim et al., 2009; Vadas et al., 2009]. There are strong south-eastward winds at 35 and 40 km in
277 Figure 8a and 8b. Under eastward wind conditions, it is easier for westward propagating GWs to
278 propagate upward because phases tilts vertically and vertical wavelengths get longer [Fritts &
279 Alexander, 2003]. Such background wind conditions can cause an asymmetric GW pattern at
280 higher altitudes.

281

282 **4. Five Typhoon Cases in 2018.**

283 Based on the validation of ECMWF GWs for Typhoon Yutu against CIPS and AIRS, we
284 further expanded our analysis to four additional typhoon events that occurred in 2018, specifically
285 Typhoon Jebi (August 26-September 4), Typhoon Mangkhut (September 6 – September 17),
286 Typhoon Kong-Rey (September 28 – October 6), and Typhoon Trami (September 20 to October
287 1). These typhoon events are all categorized as the strongest typhoon category, i.e. (“Violent
288 typhoon”) by JMA and “Category 5 typhoon” by NOAA. Table 1 summarizes the characteristics
289 of these typhoon events, including maximum wind speeds and duration of wind speeds over 100
290 kt to better understand typhoon activity.

291 Figure 9 shows snapshots of GW structures in the ECMWF vertical wind at an altitude of
292 ~35 km (5 hPa), vertical slices of the GW structures, and their power spectrum for each typhoon
293 event. Figure 9 shows results of Typhoon Jebi on September 2, 2018, Typhoon Mangkhut on
294 September 14, 2018, Typhoon Kong-Rey on October 2, 2018 and Typhoon Trami on September
295 28, 2018. These dates are selected based on the largest GW activity when typhoons are evident at
296 5 hPa, and dates are indicated in Table 1 as “GW analysis date”. Based on Figure 9, GW
297 characteristics from all cases are obtained and summarized in Table 1. The average horizontal
298 wavelength from the five typhoon-generated GWs is 236 km, with a range from 150 km to 400

299 km. GW vertical wavelengths range from 10 km to 17 km. Wu et al. [2015] derived a vertical
300 wavelength of 6-12 km from WRF simulations of Typhoon Mindulle in 2004, which is similar to
301 our study.

302 All four additional typhoon cases show concentric GWs at ~35 km altitude, though only
303 part of the concentric GWs propagate up to 35 km and appear in Figure 9. For example, concentric
304 GWs generated by Typhoon Kong-Rey (Figure 9c) show larger amplitudes in the southeast GW
305 region, but GWs generated by Typhoon Trami (Figure 9d) show larger amplitudes in the east GWs.
306 Figure 10 shows horizontal wind magnitudes and amplitudes along with vertical wind perturbation
307 at 35 km. Out of the 5 cases, typhoon Yutu is the only case where westward propagating waves
308 remained at 35 km, while the other four cases show that eastward propagating waves remained at
309 35 km. In Figure 10a for Typhoon Yutu case, it is clear that strong eastward winds exist, and that
310 the GWs are propagating against it. Typhoon Jebi (Figure 10b), Mangkhut (Figure 10c) and
311 Typhoon Trami (Figure 10e) all show relatively large westward winds, and GWs are propagating
312 against such winds as indicated by red arrows. For Typhoon Kong-Rey case, there are north-
313 westward winds and GWs are propagating south-east, opposite to the wind directions. When GWs
314 propagates against wind directions, all intrinsic phase speed, vertical wavelength and intrinsic
315 frequency become larger [Fritts & Alexander, 2003]. Large intrinsic frequency or shorter period
316 causes more vertical propagation, and GWs with longer vertical wavelength are less subject to
317 dissipation [Fritts & Alexander, 2003]. AIRS and CIPS observe more GWs with longer vertical
318 wavelength due to observational filtering. Such waves are therefore easier to be observed by AIRS
319 and CIPS nadir viewing instruments.

320 In addition to GW wavelengths, vertical wind amplitudes of GWs are also calculated and
321 summarized in Table 1. GW amplitudes at 3 hPa (~20 km) and 50 hPa (~40 km) in Table 1 are

322 calculated by the following steps. First, vertical wind GW perturbations are averaged within +/-
323 15° latitude and longitude from the center of the typhoon indicated by JMA's typhoon location
324 data at each time step. Then the largest GW amplitudes at 20 km and 40 km in Table 1 are selected
325 within 3 days of GW analysis date. Various grid sizes have been tested. Although absolute values
326 of GW amplitudes vary, conclusions from this analysis remain consistent. From Table 1, maximum
327 wind speeds are largest during typhoons Yutu and Kong-Rey; however, GWs during Typhoon
328 Mangkhut have the largest amplitudes at 20 km. Also, GW amplitudes for the three typhoons with
329 the shortest duration of winds > 100 kt vary from smallest to second largest. By comparing GW
330 amplitudes at 20 km and 40 km, the largest GW amplitudes changed from GWs generated by
331 Typhoon Mangkhut at 20 km to GWs generated by Typhoon Jebi at 40 km. For these five cases,
332 the strength of GWs in the stratosphere do not seem to correlate with the characteristics of typhoons.

333 Furthermore, at 20 km the maximum GW amplitude is 2.8 cm/s during Typhoon Mangkhut,
334 and the minimum GW amplitude is 2.2 cm/s during Typhoon Yutu. The difference between the
335 maximum and minimum amplitude is 21%. At 40 km, the maximum and minimum GW amplitudes
336 are 8.1 and 4.4 cm/s during typhoons Jebi and Kong-Rey, respectively. The difference here is 46%.
337 The GW amplitude differences grow as the GWs propagate upward. These results indicate that
338 differences in GW amplitudes between different typhoon cases at ~40 km are influenced by
339 background winds that affect GW propagation between 20 km and 40 km. Therefore, whether
340 typhoon generated GWs may impact the upper atmosphere is highly dependent on the seasonality
341 of the typhoons and background winds through which the waves subsequently propagate.

342

343 **5. Summary & Conclusion**

344 This work combines CIPS and AIRS nadir viewing observations at 50-55 km and 30-40
345 km, respectively, and the high-resolution ECMWF model data to study the altitudes and temporal
346 evolution of concentric GWs. Both AIRS and CIPS captured clear concentric GW structures
347 during Typhoon Yutu in late October 2018, with horizontal wavelength of ~ 237 - 247 km.
348 ECMWF-IFS (a horizontal resolution of 9 km) and ERA5 (a horizontal resolution of 31 km) are
349 able to generate concentric GWs at similar locations and timing as the AIRS and CIPS observations.
350 The GW horizontal wavelength simulated by the ECWMF-IFS was also ~ 225 - 236 km. Further
351 analysis of ECMWF data showed that the vertical wavelength of these concentric GWs was ~ 10 –
352 14 km, and GW amplitudes reached a maximum ~ 1 - 2 days after the peak wind speed of typhoon
353 Yutu.

354 After a validation case study of ECWMF GWs against CIPS and AIRS observations, four
355 additional category 5 typhoons were analyzed to obtain characteristics of concentric GWs
356 generated by these typhoons. ECMWF captured clear concentric GW structures generated by all
357 five cases. Analysis of all five cases showed an averaged horizontal wavelength of 237 km, ranging
358 from 150 km to 400 km, and vertical wavelength ranging from 10 km to 17 km. Differences in
359 maximum and minimum GW amplitudes among the five typhoon cases grow as GWs propagate
360 upward.

361 Our analysis also shows that the amplitudes of GWs in the stratosphere are not strongly
362 correlated with the strength of typhoons as indicated by wind. Our results confirm that typhoon-
363 generated concentric GWs in the stratosphere and lowermost mesosphere where AIRS and CIPS
364 observe have more influences from background wind conditions than strength of typhoons for our
365 cases. This work has shown that combining analyses of CIPS and AIRS data enables verification

366 of model simulations of GW propagation across the stratopause, a prerequisite to understanding
367 how GWs act to couple the lower and upper atmosphere and ionosphere.

368

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377

378 **Open Research**

379 **Data Availability Statement**

380 CIPS RAA used in this work is available from the CIPS website (<http://lasp.colorado.edu/aim/>)
381 [Randall et al., 2017]. AIRS GW data are produced by Dr. Lars Hoffmann, and obtained from
382 <https://data.fz-juelich.de/dataset.xhtml?persistentId=doi:10.26165/JUELICH-DATA/LQAAJA>
383 [[Hoffmann, 2021](#)]. Locations and maximum sustained wind speed data are obtained from Japan
384 Meteorological Agency ([http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-](http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-eg/besttrack.html)
385 [eg/besttrack.html](http://www.jma.go.jp/jma/jma-eng/jma-center/rsmc-hp-pub-eg/besttrack.html)) [Japan Meteorological Agency, 2018]. ECMWF and ERA5 were provided by
386 ECMWF Meteorological Archival and Retrieval System (MARS) at
387 <https://www.ecmwf.int/en/forecasts/datasets> [European Centre for Medium-Range Weather
388 Forecasts, 2016, 2019].

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507

508

509 Table

510

511 *Table 1. Characteristics of typhoon-generated GWs.*

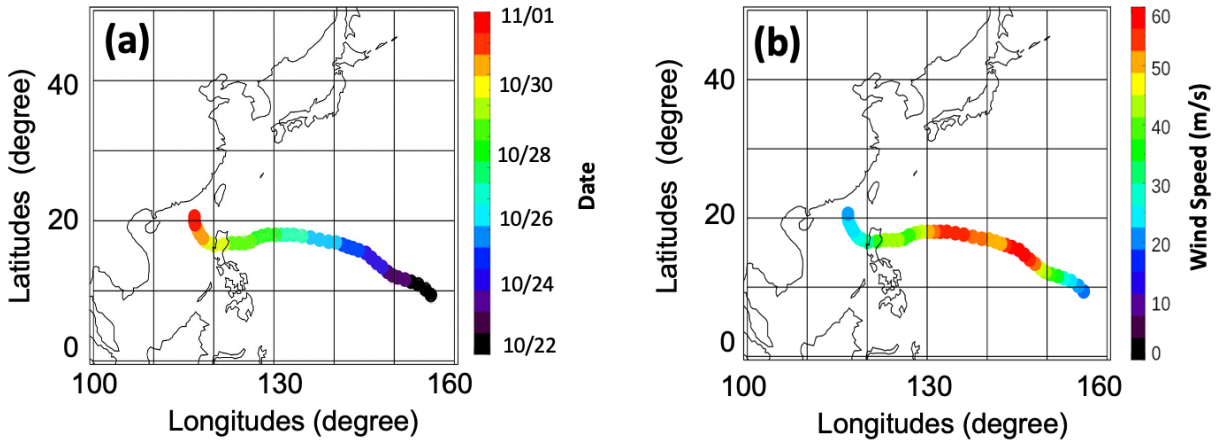
Name of Typhoon	Typhoon Date in 2018 (MM/DD)	Max. Wind Speed (Duration >100kt)	GW Analysis Date	λ_h (km)	λ_z (km)	w' (GW) Amplitude at 20 km	w' (GW) Amplitude at 40 km
Yutu	10/21 – 11/03	115 kt (78 h)	10/27	~ 250	~10-14	2.2 cm/s	6.4 cm/s
Jebi	08/26 – 09/04	105 kt (48 h)	09/02	~150	~15	2.6 cm/s	8.1 cm/s
Mangkhut	09/06 – 09/17	110 kt (90 h)	09/14	~150-400	~10-17	2.8 cm/s	6.3 cm/s
Kong-rey	09/28 – 10/06	115 kt (48 h)	10/02	~200	~10	2.3 cm/s	4.4 cm/s
Trami	09/20 – 10/01	105 kt (48 h)	09/28	~270	~10-12	2.5 cm/s	4.8 cm/s

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514 **Figures**

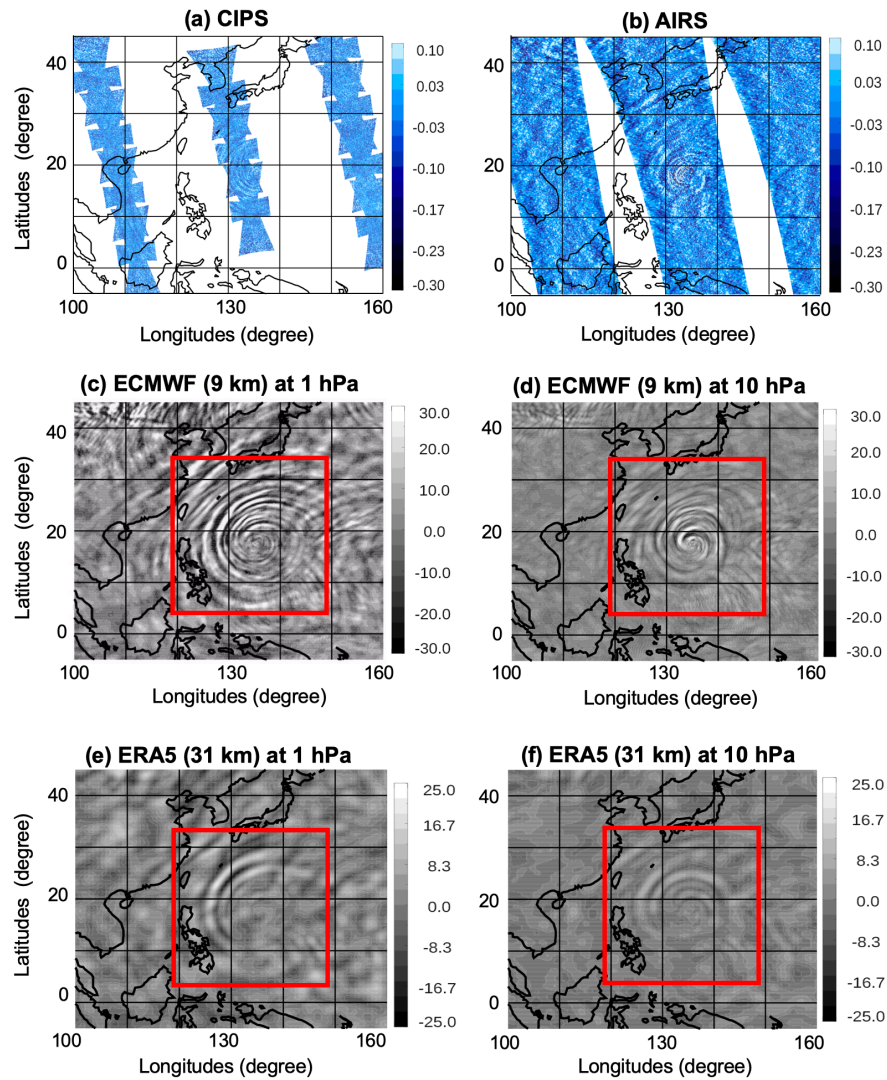
515 **Figure 1.** (a) Path of Typhoon Yutu from October 22 to November 1. (b) Maximum sustained
516 wind speed (10-minute average) along the path of Typhoon Yutu.



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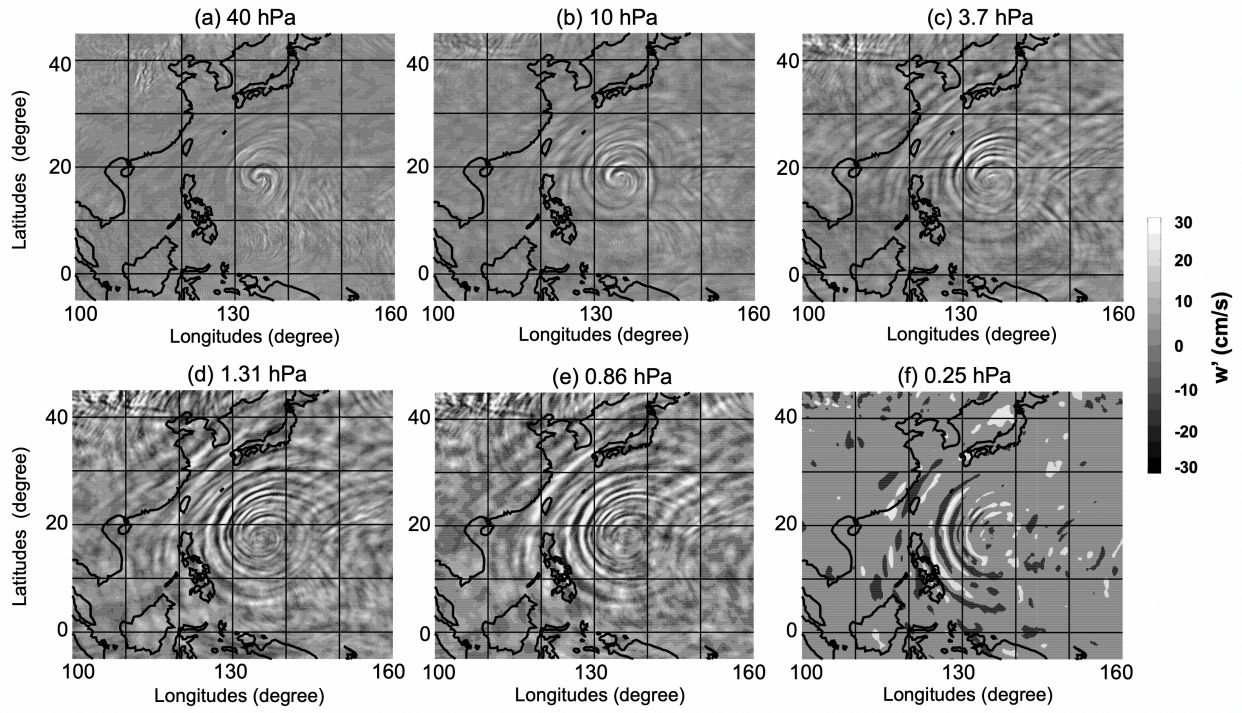
519 **Figure 2.** Snapshots of concentric GW observations from (a) CIPS RAA (%) at ~50-55 km and
 520 (b) AIRS brightness temperature perturbation (K) at ~30-40 km and vertical wind perturbation
 521 from (c, d) ECMWF and (e, f) ERA5 at 1 hPa (~50 km) and 10 hPa (~30 km). AIRS and CIPS
 522 observations were made on October 27, ~4:30 UT and October 26, ~22:25 UT, respectively.
 523 ERA5 and ECMWF plots are made with data from October 27, 00 UT. Red boxes indicate +/-
 524 15 degree from the center of typhoon location.
 525



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528 **Figure 3.** Vertical wind perturbation (cm/s) from ECMWF-IFS at 40, 10, 3.7, 1.31, 0.86, and 0.25
529 hPa for Typhoon Yutu on October 27 at 00 UT.



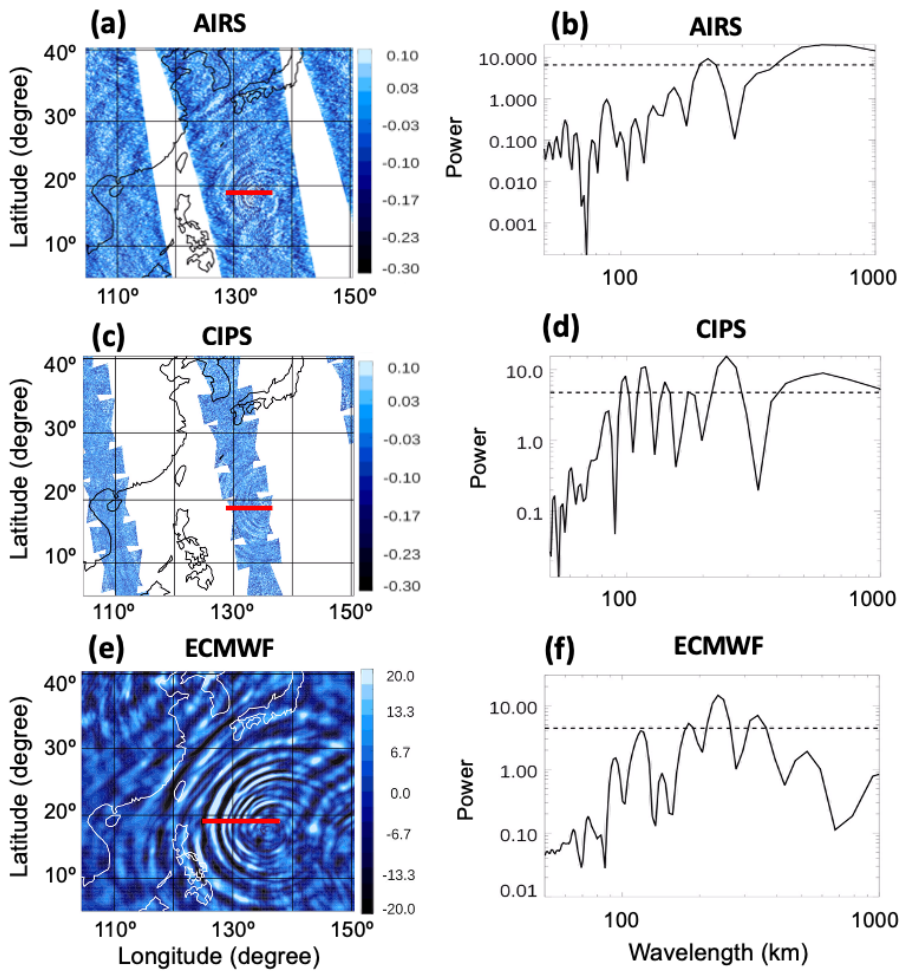
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533 **Figure 4.** Power spectrum of horizontal wavelength for (a, b) AIRS brightness temperature
 534 perturbations (K), (c, d) CIPS RAA (%), and (e, f) ECMWF vertical wind perturbation (cm/s)
 535 at 1 hPa (~50 km). Red thick lines indicate data used to calculate the power spectrum.

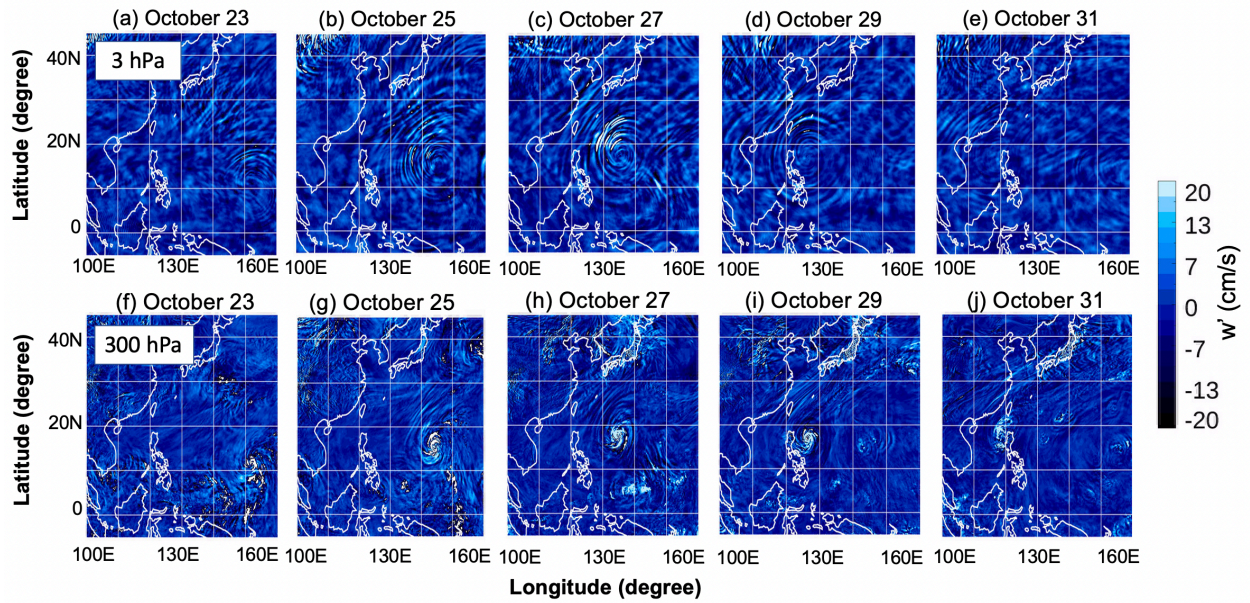
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540 **Figure 5.** Snapshots of vertical wind GW perturbation from October 23 to October 31 at 3 hPa
541 (~40 km) (top) and at 300 hPa (~10 km) (bottom). White lines indicate continents.

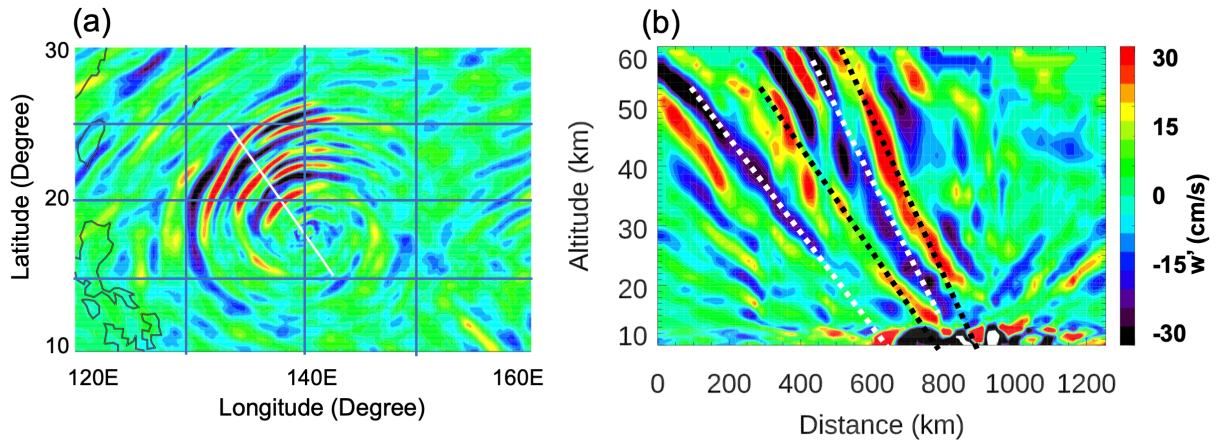
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545 **Figure 6.** (a) Snapshot of concentric vertical wind GW perturbation at 5 hPa (~35 km) on October
546 27, 2018 at 00 UT and (b) vertical slice at white transect in Figure 6a.

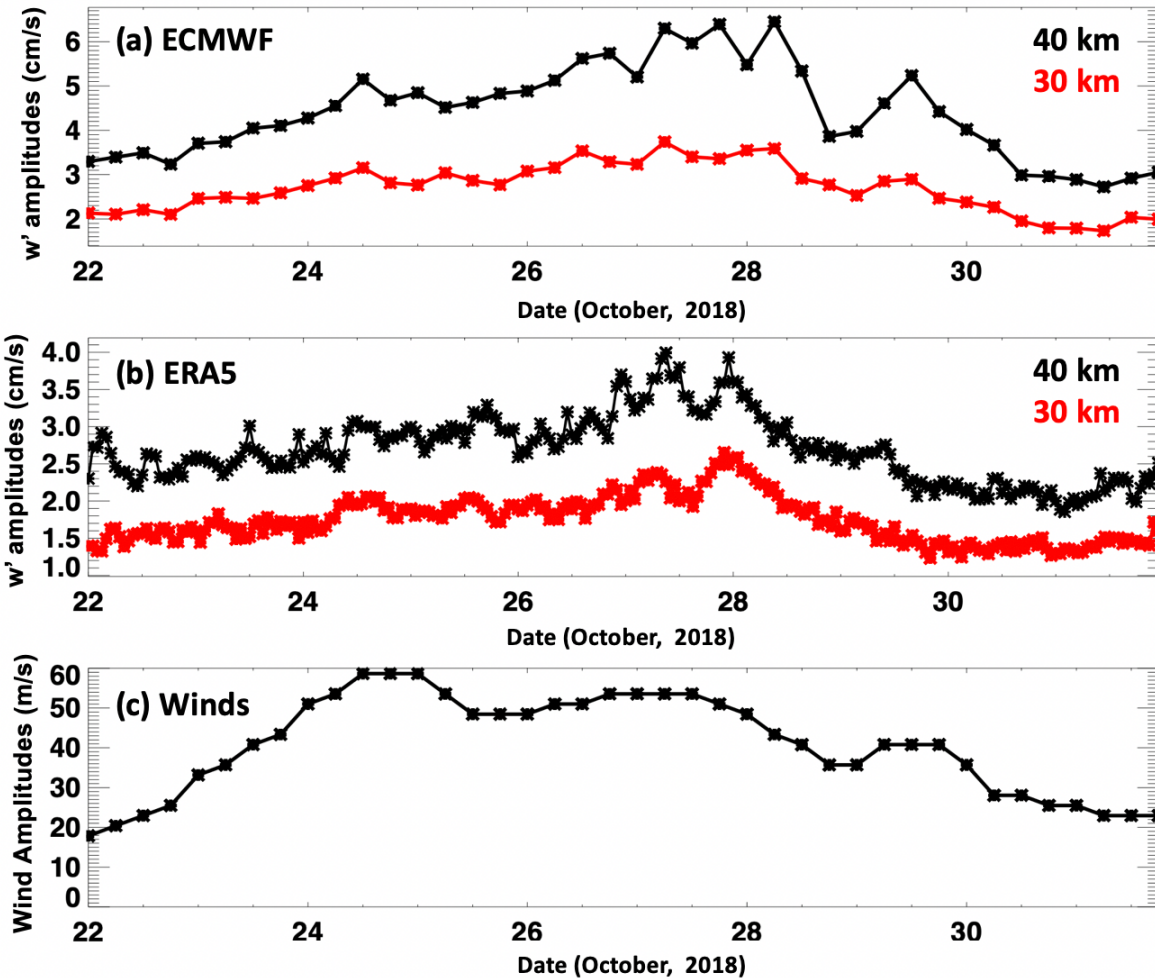


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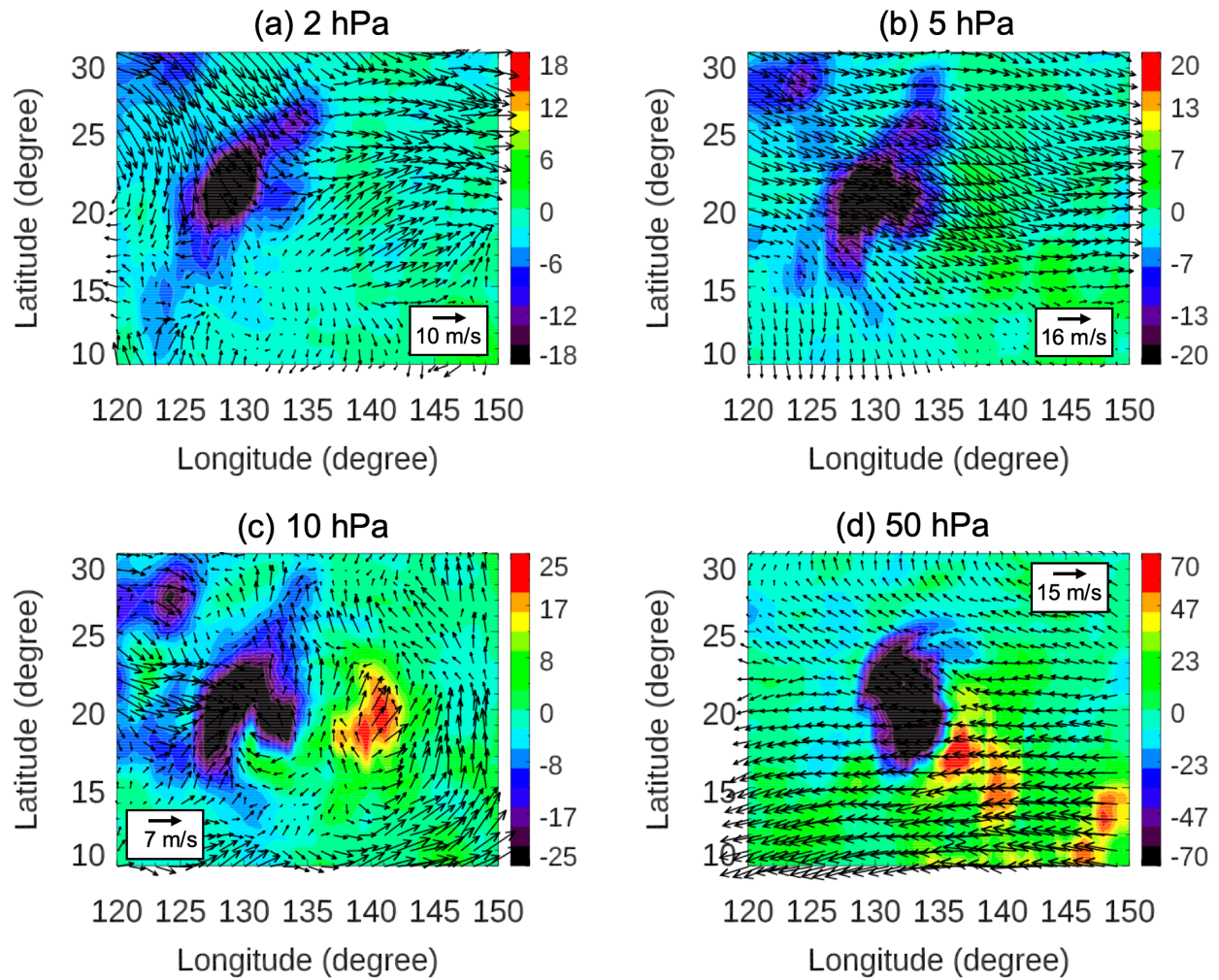
549 **Figure 7.** Averaged GW vertical wind (w) amplitudes at (black) 3 hPa (~ 40 km) and (red) 10 hPa
 550 (~ 30 km) from (a) 6-hourly ECMWF data and (b) hourly ERA5 data, and Typhoon Yutu maximum
 551 sustained wind speed provided by Japanese Meteorological Agency (JMA) from October 22 to
 552 November 1, 2018.

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557 **Figure 8.** Gravity wave momentum fluxes ($\times 10^{-4}$ Pa; color contours) on October 27, 2018 at 00
558 UT at (a) 2 hPa, (b) 5 hPa, (c) 10 hPa, and (d) 50 hPa. Arrows indicate horizontal wind
559 directions and magnitudes.



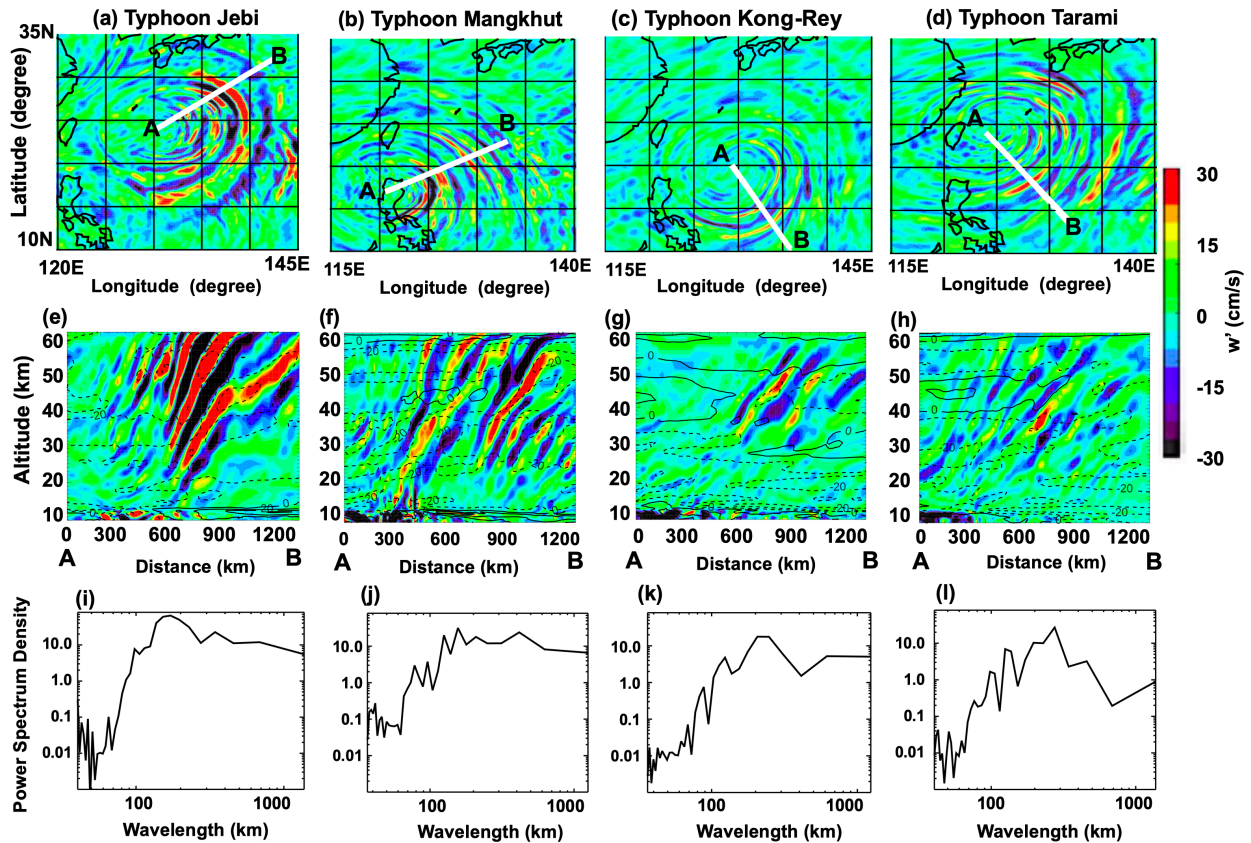
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564 **Figure 9.** (a-d) Concentric gravity wave structures in the ECMWF vertical wind at 5 hPa (~35 km)
 565 altitude; (e-h) vertical slices of vertical wind GW perturbation at the white transects in (a-d),
 566 respectively; and (i-l) corresponding power spectrum density at ~35 km for four each typhoon
 567 case.
 568

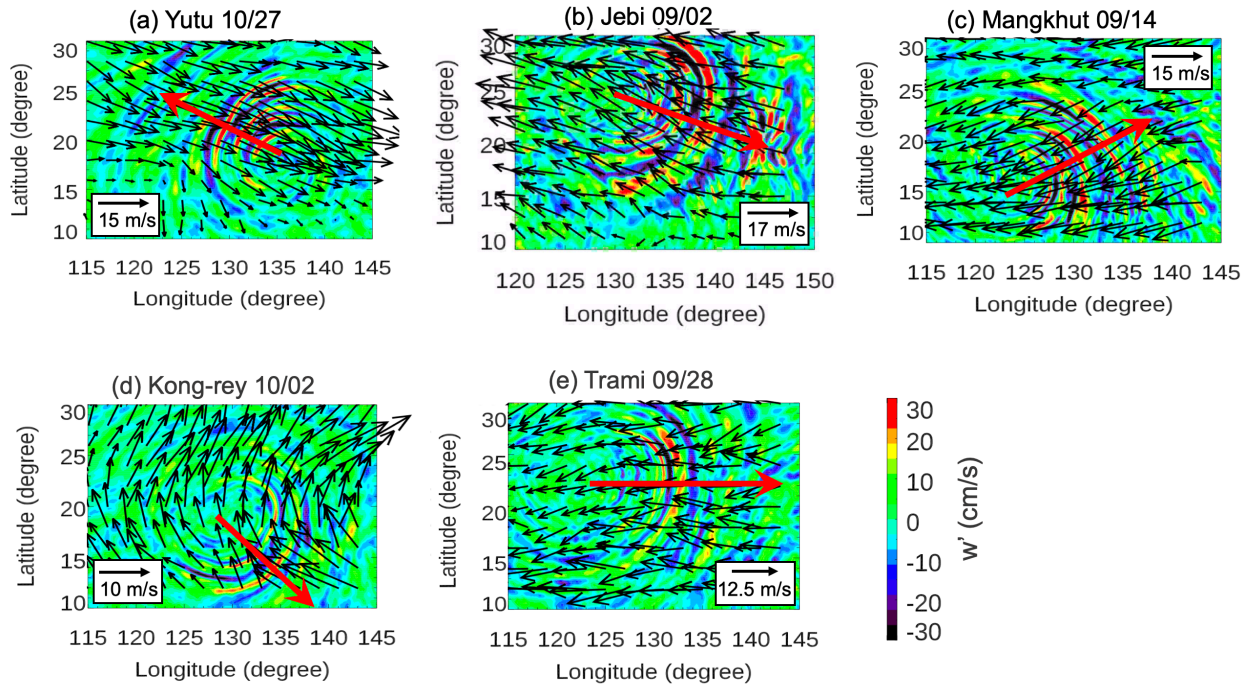


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571 **Figure 10.** (contours) Vertical wind perturbation and (black arrows) horizontal wind vectors at
 572 ~35 km for (a) Typhoon Yutu on October 27, (b) Typhoon Jebi on September 2, (c) Typhoon
 573 Mangkhut on September 14, (d) Typhoon Kong-Rey on October 2, and (e) Typhoon Trami on
 574 September 28. Red thick arrows indicate wave propagation directions.

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