

AN ABSTRACT OF THE THESIS OF

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Title: Design and Simulation of a Self-Powered Neutron Spectrometer.

Abstract approved:

Stephen E. Binney

A self-powered neutron detector (SPND) is a device that, coupled with a current meter, provides a readout proportional to neutron population. This thesis discusses the design parameters of an array of such devices, their characteristics, and the use of these devices as a self-powered neutron spectrometer (SPNS) to provide information about the energy distribution in a neutron radiation field.

Neutron absorption in an appropriate material produces subsequent beta emissions. In a SPND, some of these beta particles will cross a non-conducting region and stop in a collector material. A net exchange of charge between these regions can be read as a current flowing between the emission region and the collector region.

One potential SPNS design was modeled using a Monte Carlo simulation of the device's interaction with a radiation field. The Monte Carlo program used predicts the beta flux which is proportional to the current that would be produced by an actual device. Various beta emitting materials were considered for this device, and a sensitivity study of each was included.

The design considered is comprised of a concentric set of these cylindrical SPND detector elements which, in themselves, are currently available technology.

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Design and Simulation of a Self-Powered Neutron Spectrometer

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TABLE OF CONTENTS

	<u>Page</u>
1. INTRODUCTION	1
1.1 Problem Definition	2
1.2 Statement of Purpose.....	2
2. LITERATURE REVIEW	3
2.1 Bonner Spheres	3
2.2 Semiconductors	3
2.3 Self-Powered Neutron Detectors.....	4
3. SPECTROMETER DESIGN	7
3.1 Details of Self-Powered Neutron Detector Array.....	7
3.2 Emitter materials.....	7
3.3 Insulators.....	9
3.4 Conductors.....	10
3.5 A Complete Unit.....	10
3.6 Decoupling Signals For Different Energy Neutrons.....	11
4. EXPERIMENTAL DESIGN	12
4.1 Description Of Monte Carlo Technique.....	12
4.2 Advantages Of Monte Carlo Over Rules-Of-Thumb.....	12
4.3 MCNP4B Program.....	12
4.4 Neutron Source.....	13
4.5 Beta Dose/Counts For Current.....	13
4.6 Neutron, Gamma, and Beta Source Simulations.....	15

TABLE OF CONTENTS (CONTINUED)

	<u>Page</u>
4.7 Decoupling Matrix Derived From Model Response	16
4.8 Time Dependence.....	16
4.9 Final Design.....	17
5. RESULTS	19
6. DISCUSSION	21
7. SUMMARY AND CONCLUSIONS	22
BIBLIOGRAPHY	23
APPENDICES	24

LIST OF FIGURES

<u>Figure</u>		<u>Page</u>
1.	Neutron Spectrum	2
2.	A Typical SPND	4
3.	Concentric Detector Elements in a SPNS	5
4.	Beta Spectrum	9
5.	Effect of Beta Source Thickness on Relative Yield	14
6.	Effect of Insulator Thickness on Relative Yield	14

LIST OF TABLES

<u>Table</u>		<u>Page</u>
1.	Potential Nuclides for SPNS Emitters	8
2.	Beta Spectra of Emitter Progeny	8
3.	SPNS Dimensions	17
4.	Response of Three Material SPNS	18
5.	Response of Cadmium Shielded Rhodium SPNS	19

Chapter 1. INTRODUCTION

Neutrons are particulate radiation released in nuclear reactions under certain conditions such as a fission reaction. They may have various energies which are defined loosely as fast, epithermal, and thermal. A fast neutron has an energy on the order of the mass-defect of a nuclear reaction, from approximately 10 keV to 15 MeV. An epithermal neutron has a much lower energy and may react with various atomic nuclides in a non-linear fashion, with high interaction probability or cross section at specific energies, but a low interaction cross section at slightly above or below these resonance energies. Thermal neutrons, defined as below 0.5 eV, typically have average energies of about 1/40 eV [1], act much more like a gas, and follow a Maxwellian energy distribution. An application such as verification of a nuclear core model, evaluating a neutron beam port, or prediction of a dose in a neutron field may potentially require simultaneous measurement of neutrons in these three regions.

Figure 1 is plotted from theoretical distribution formulas and shows the three distinct neutron spectral regions. The thermal region follows a Maxwell-Boltzmann distribution which is dependent on the temperature of the surrounding media. The center region is called epithermal and follows an inverse kinetic energy ($1/E$) profile. Fast neutrons are distributed according to a Watt spectrum. Each distribution has a significantly different shape.

A self-powered neutron detector (SPND) can be used to measure the flux of an intense neutron field. This device utilizes beta emissions from materials, which absorb some fraction of the neutrons and provide a current that can be measured on a sensitive picoammeter. A series of these detector elements, nested within each other, could (in theory) provide readings of thermal, epithermal, and fast energy neutron fluxes.

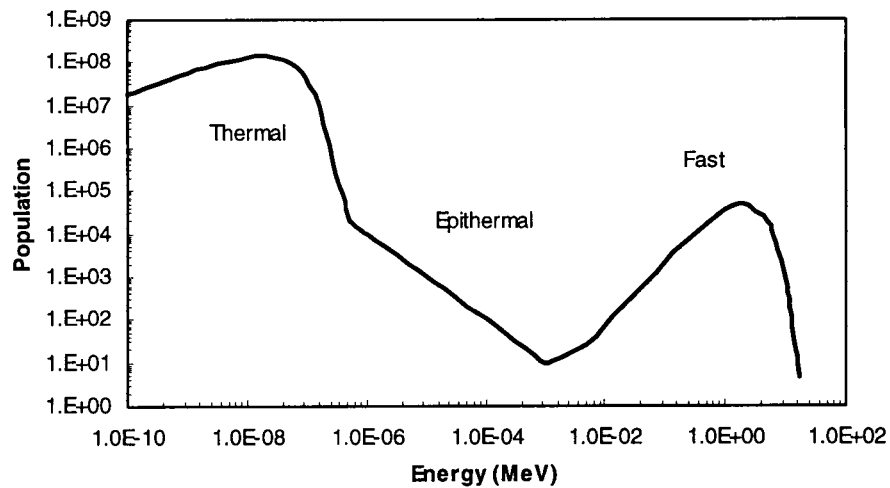


Figure 1. Neutron Spectrum

1.1 Problem Definition

Determine a suitable design for a SPND-based Self- Powered Neutron Spectrometer (SPNS) and use computer modeling to predict operating performance.

1.2 Statement of Purpose

Using established SPND technology, determine materials required for a SPNS. Use MCNP4B to determine the thicknesses of emitter and insulator regions in each detector element. Simulate actual operation of a complete SPNS using MCNP4B.

Chapter 2. LITERATURE REVIEW

A search of current literature shows many different types of neutron spectrometers, some with fine resolution or a very wide range. The actual list is quite long and includes devices that are based on neutron moderation, fast-neutron scattering, and fast neutron-induced reactions. Due to the large number of actual neutron spectrometers, only a select few will be mentioned here.

2.1 Bonner Spheres

Bonner spheres are a series of neutron detector elements, spherical in shape with radii varying from about 6 to 20 cm. An array of detectors, each with a distinct energy response, can be used to measure neutrons of various energies. Three detectors, for thermal, epithermal, and fast neutrons, are needed to characterize a three energy group neutron spectrum. More typically a much larger array of spheres is required to determine the shape of a spectrum.

Bonner spheres detect neutrons by first moderating the neutrons to thermal energies with various thicknesses of hydrocarbon shells [2] and then detecting some of these thermalized neutrons with a central LiI scintillation counter. A boron containing detector chamber or ^3He gas proportional counter may also be used. The spherical symmetry of this device eliminates any angular dependence. A drawback of this approach is that a set of these spheres can take up considerable volume, which means that they are hard to place in a reactor or other confined space and would certainly effect the criticality of any fissionable assembly. Hence they are more useful for external beams.

2.2 Semiconductors

Silicon semiconductor devices rely on electron-hole pair production to produce a signal to indicate the presence of neutrons. However, under a high flux

field, severe lattice damage may occur, limiting sensitivity and energy resolution, which results in a very short useful device life-span [3]. With various degrees of shielding, energy discrimination can be improved.

2.3 Self-Powered Neutron Detectors

A self-powered neutron detector, Figure 2 [4], uses an (n, γ) reaction with a delayed β which is read directly, without amplification, by a picoammeter. It involves three materials. The first is the beta source material which absorbs neutrons and emits beta particles (high energy electrons). This region may or may not need a conductive backing. The second is an insulator material, which is designed to electrically isolate the beta source material from the collector material. The most important consideration for this material is that it does not significantly attenuate the beta emissions. This is best done with a material composed of low atomic number atoms, and on the order of a millimeter thick, depending on the emitted beta energy. The last material is a conducting material that absorbs the emitted beta particles. Movement of charge, proportional to the neutron flux, can be read by connecting the source material to the collector material with a very sensitive current sensing instrument, a picoammeter.

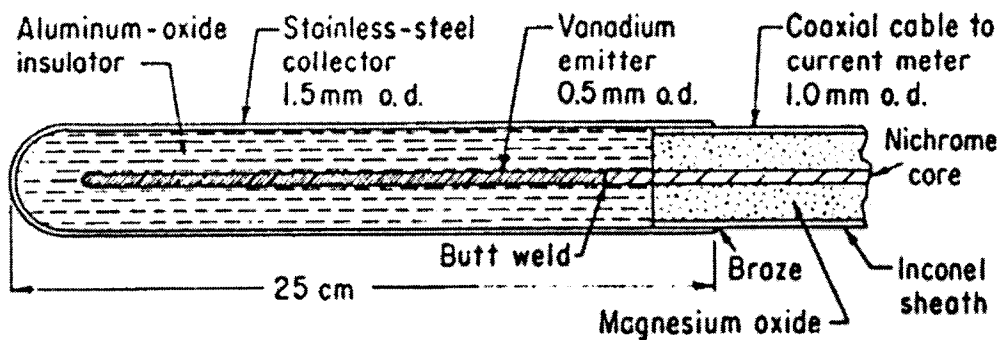


Figure 2. A Typical SPND

Several SPNDs can be nested inside each other provided they do not interfere with each other. These can be nested cylinders, such as shown in Figure 3. The dark lines show boundaries of detector units, and the light lines within signify surfaces between materials.

The sensitivity of a single SPND is dependent on several factors, including beta energy and geometry. An analytical formula for the sensitivity can be expressed as:

$$\text{Electrical Current} = \phi * \sigma * N * \% \beta * G * C/\beta$$

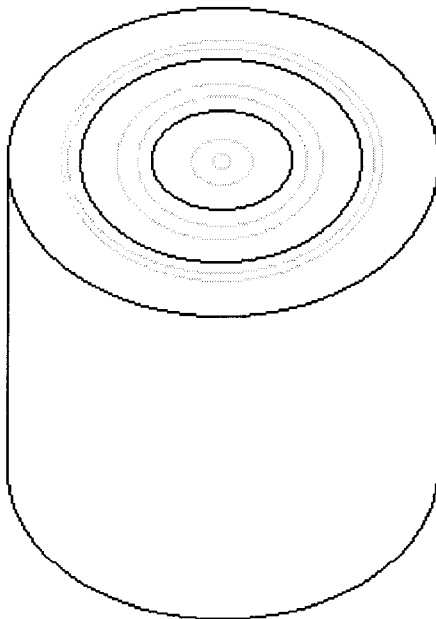


Figure 3. Concentric detector elements in a SPNS

The term ϕ , in units of neutrons $\text{cm}^{-2} \text{s}^{-1}$, is the neutron flux in the emitter material of the detector. The cross section σ represents the apparent size of one atom, or probability of interaction with a nearby free neutron, and has units of $\text{cm}^2 \text{atom}^{-1}$. N is merely the total number of atoms in the emitter region. The term $\% \beta$ refers to the

fraction of neutron captures that result in a beta decay and is 1.0 for most materials under consideration. The geometrical factor G is the most complicated term, as it depends on the combination of beta energy and material densities and arrangement. The last factor, C/β , is the charge of one electron, which causes the final result to have units of current (A).

Chapter 3. SPECTROMETER DESIGN

3.1 Details of Self-Powered Neutron Detector Array

The array of SPNDs, which is called a SPNS here, consists of three individual SPNDs which each have three regions, the neutron absorber/beta emitter region, the insulator, and the collector/shield. Wires attached to the absorber/emitter and collector regions deliver current to a remote meter. A typical SPND is cylindrical in design, so the spectrometer is also cylindrical, with concentric SPND units as shown in Figure 3.

3.2 Emitter Materials

The materials for the neutron absorbing, beta emitting region are subject to a few constraints, the first of which is that it be a metal or be present in some other solid, manufacturable form. A second consideration is half life, which should be reasonable as compared to measurement time requirements, with several minutes as an upper boundary. To approach steady state currents, the spectrometer would have to be exposed to the flux for at least four half lives of the longest half life beta emitter, and a half hour may be considered a very long time to wait for a reading. A third consideration is that the parent nuclide have a significant neutron cross section to aid in sensitivity of the device. A fourth concern is the beta spectrum emitted by the emitter material after neutron capture. The beta energies should be high enough so that a significant fraction will cross the insulator region, but not so high as to require heavy shielding. Table 1 shows a selection of nuclides that were considered as emitters [5].

Table 1. Potential Nuclides for SPNS Emitters

Parent Nuclide	Parent Abundance (%)	Parent Thermal Cross Section	Parent Epithermal Cross Section	Daughter Nuclide	Daughter Half life	Daughter beta-Max (MeV)
¹⁰³ Rh	100	11 b	13 b	¹⁰⁴ Rh	42.3 s	2.442
²⁷ Al	100	0.233 b	0.17 b	²⁸ Al	2.25 m	2.863
⁵¹ V	99.750	4.9 b	2.7 b	⁵² V	3.76 m	2.47
⁵⁵ Mn	100	13.3 b	14.0 b	⁵⁶ Mn	2.578 h	2.84

Table 2 shows the relative strength of emitted betas within several energy groups for the materials in Table 1. Each isotope has a finite upper limit to the beta energy. These are the beta energy spectra used as input in the MCNP4B model.

Table 2. Normalized Beta Spectra of Emitter Progeny

Energy Bin (MeV)	¹⁰⁴ Rh	⁵² V	⁵⁶ Mn	²⁸ Al
0 - 0.01	0.00241	0.00156	0.0069	0.00080
0.01 - 0.02	0.00247	0.00162	0.0070	0.00087
0.02 - 0.04	0.0051	0.00344	0.0143	0.00197
0.04 - 0.1	0.0170	0.0122	0.0466	0.0077
0.1 - 0.3	0.075	0.061	0.174	0.0425
0.3 - 0.6	0.159	0.141	0.230	0.108
0.6 - 1.3	0.446	0.434	0.270	0.380
1.3 - 2.5		0.345	0.245	0.448
1.3 - 2.442	0.289			
2.5 - 2.542		0.0000259		
2.5 - 2.849			0.0055	
2.5 - 2.863				0.0112

Figure 4 shows the same information graphically.

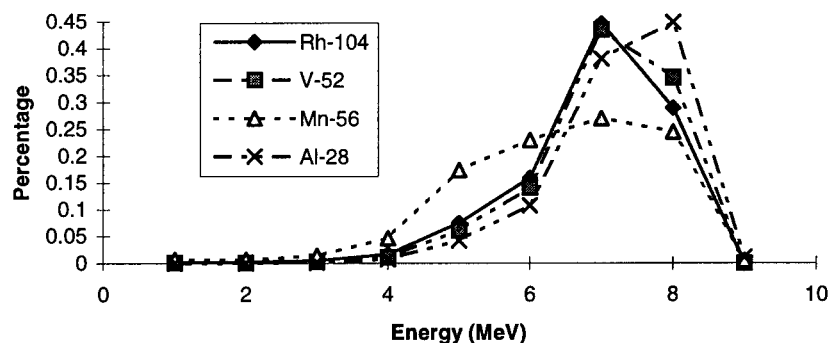


Figure 4. Beta Spectrum

3.3 Insulators

Many materials can be used to insulate the beta-emitting/neutron absorber material from the collector/conductor region. Ceramics make good insulators and are relatively stable under high temperatures. Physical shock or transients of various kinds can shatter them, however. Epoxy, however, can easily be applied in the form of coating, glue, or paint. Power coatings are formed with very low solvent fractions, are therefore harder than paints and epoxies, and have a minimum amount of void space. Hydrocarbon polymers make fair insulators, are very inexpensive, but are less stable than other alternatives. To choose between any of these options, principles of economy and manufacturability should be applied, but these are beyond the scope of this thesis. The exact formulation of any insulator compound, however, should be evaluated for long-term neutron activation products. Actual evaluation may require testing of different materials from different sources with their many impurities, and is beyond the scope of this paper.

3.4 Conductors

Lead is a fair conductor, and is suitable for the low currents in the SPNS. It is especially resistant to neutron bombardment, is ductile and corrosion resistant. Lead can anneal at fairly low temperatures, and has alloys with a broad range of melting points over the temperature expected in a reactor core. The closer a metal is to its melting temperature, the faster it will anneal. Copper is a very good conductor, second only to silver. It is ductile, a good plating agent, and can also anneal at relatively low temperatures.

Most other common metals are not good choices due to cost, difficulty in fabrication, or poor neutronic behavior. An isotope such as ^{59}Co can absorb a neutron and become ^{60}Co . ^{60}Co beta decays with a half life of 5.3 years but emits two penetrating gamma rays. These gammas can penetrate any light shielding, and the long half life means that they will be produced for some time after neutron bombardment ceases. This would make for a detector that is hazardous to service and inconvenient to dispose of. However, cobalt may be well suited for measuring higher energy neutrons due to the emission of prompt gammas during neutron capture. These gammas will produce Compton electrons which can be read as a signal in a SPND. Many materials must be avoided that could possibly contain this isotope or any similarly hazardous activating isotopes. Aluminum produces unreliable electric connections and therefore may not be a favorable choice.

Iron makes a fair conductor and can be used in this application, as signal strength losses would be minimal with the minute current expected. Iron and its alloys have been used in commercial SPNDs [4], can have low activation cross sections, and are therefore used in these simulations.

3.5 A Complete Unit

A complete unit consists of three (or more) SPNDs. Each beta emission region should have a different response to the neutron flux. This may be accomplished by requiring neutron absorbing/beta emitting materials to have different neutron cross

sections at the neutron energies of interest. An alternative is to shape the neutron flux between each region. Of key importance is a significant contrast in neutron response in each SPND element.

3.6 Decoupling Signals For Different Energy Neutrons

The neutrons of the three energy ranges will produce emitted betas, to some extent, in all the detector elements. This is an unavoidable aspect of the energy dependent cross section. As much as these elements will be cross-coupled to each other, the signals derived may still be used to properly register the different energies.

Signal data were generated, from known energy distributions, one energy or distribution region at a time. The responses from each energy region can be combined into a vector, and these vector responses used to form a response matrix. The inversion of this matrix can then be used to deconvolute signal data into an energy-dependent spectrum.

Chapter 4: EXPERIMENTAL DESIGN

4.1 Description Of Monte Carlo Technique

Monte Carlo simulations are a class of non-deterministic or stochastic computer programs. It is used here to emulate interaction of radiation fields with matter by tracking individual particles. This method is generally attributed to Fermi, von Neumann, and others and emerged from work done at Los Alamos during World War II [6]. Particles emanate from a source region, travel a distance randomly determined by their interaction cross sections, are eventually captured, and may also cause the creation of other particles which are then tracked later. The simulation requires the tracking of neutrons (the field particles of interest), and electrons (which will provide a current reading for a meter), and photons which can create additional electrons to be tracked.

4.2 Advantages Of Monte Carlo Over Rules-Of-Thumb

The Monte Carlo technique has several advantages over deterministic or rule-of-thumb procedures used for calculating interactions of radiation with matter. Particles follow the rules of the known experimental data through complicated geometries. In general, a large population for neutrons is tracked to estimate the effects of an extremely large neutron flux. Rare but important events can be tracked as well by biasing the input data to favor their occurrence. This bias can then be removed by a simple calculation at the end. Rule-of-thumb calculations usually involve very simple geometries and provide only an estimated result.

4.3 MCNP4B Program

The particular program used is called Monte Carlo, N-Particle, version 4B (MCNP4B) and is currently supported by the X-CI Code Integration Group at the Los

Alamos National Laboratory. MCNP4B is able to track particle histories for neutrons, photons, and electrons.

4.4 Neutron Source

Initial simulated particles can be placed in a source region or, as in these simulations, on a source surface with specified energy and direction distributions. They then proceed through the model, interact with the various materials until they are absorbed or leave the system. Particles other than neutrons may be eliminated if they fall below some kinetic energy cutoff, but neutrons are tracked down into thermal energies.

The neutrons in this experiment originate outside the detector and interact with materials to produce unstable nuclei. These nuclei are the source of beta particles which will eventually be read as a current with a picoammeter.

4.5 Beta Dose/Counts For Current

MCNP predicts current tallies on a surface that divides the insulator region in half. The reason for choosing this seemingly arbitrary surface is that many electrons will not penetrate the insulator region. They may later drift to one conductor or the other. However, there still will be an immediate effect on current as their presence will induce charge displacement in one region or the other.

The emitter region can attenuate some of its emitted betas in a process called self-shielding. Figure 5 shows the effect of increasing the beta emitting region's thickness while maintaining a constant specific activity which would be consistent with a fixed neutron flux. The yield scale on this chart indicates beta tallies (electrons that, for these purposes, have crossed the insulator junction) produced per source neutron. Aluminum has a slightly higher beta energy, but more significantly, a lower density, which greatly reduces the effect of this self-shielding. However, most of these materials produce little effective sensitivity increase beyond 0.1 cm thickness.

The effect of varying thickness of the insulator region is shown in Figure 6. The beta yield drops significantly with even a small increase in this thickness.

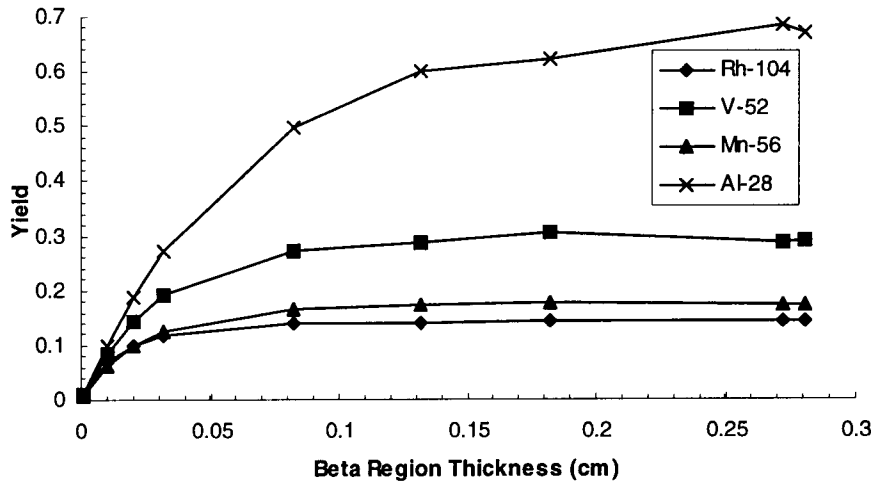


Figure 5. Effect of Beta Source Thickness on Relative Yield

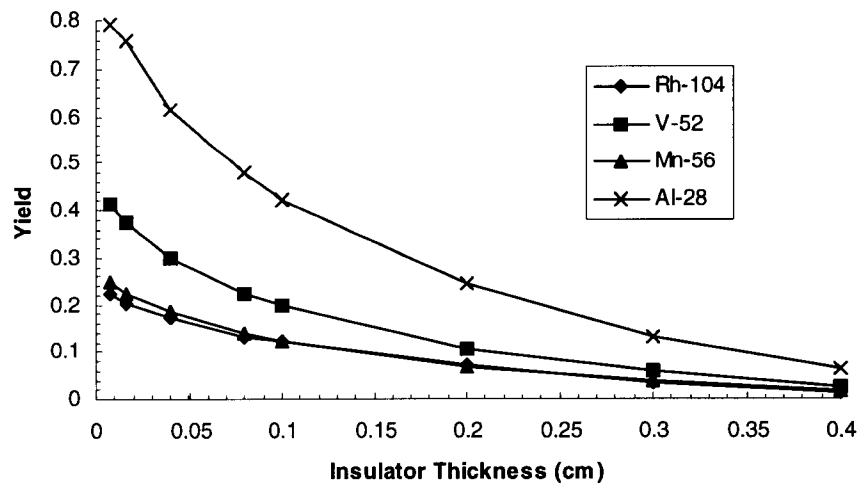


Figure 6. Effect of Insulator Thickness on Relative Yield

For arbitrary reasons, the insulator thickness used in the simulations was 0.008 cm, as this may be a reasonably manufacturable insulator thickness and has the lowest attenuation of the beta current.

Appendix A shows the MCNP4B input and output for one typical run which was used to generate data for the above figures. Not all runs are shown separately as each data point constitutes one MCNP4B run. Various beta spectra or variation of cell thickness were used to generate these graphs, but only one of the several dozen simulations is included for brevity.

Several factors have an effect on the beta yield and therefore the sensitivity of a detector element. Electrons produced by beta decay following (n, γ) reactions must first be able to leave the source material, and only half of these will be heading in the right direction, toward the collector. Some of these will be attenuated by the insulator material. As electrons are charged particles and have a finite range, this attenuation will be minimal if the insulator is thin (Figure 5). Betas can be detected farther away due to Compton production from gammas which are not as easily shielded. The probability of a given particle crossing the insulator region is dependent on the thickness of the region, the direction it is traveling, and its energy. Figure 3 shows the beta spectrum for the materials used in the beta-emitting regions. The remaining beta particles are then absorbed into the collector plate, producing a net exchange of charge. The ratio of the beta tally relative to incident neutrons determines the response of the device.

4.6 Neutron, Gamma, and Beta Source Simulations.

MCNP4B is capable of modeling multiple sources; however they must be of the same particle type. Also, all the cross section tables used in MCNP do not produce the delayed betas. This is a time dependent problem, as some betas may not be emitted for a significant time after absorption of the neutron. So the input deck should be run three times. The first run has a source that is an isotropic neutron flux

originating on the inside of a spherical surface. Some small fraction of these neutrons are absorbed in various cells to produce tallies. This was a neutron-only problem.

The results of the neutron absorption simulation could have been used as the source strength for the beta emission simulation, but this was not required as the results scale linearly, and only one beta simulation is needed for each combination of material type and thickness. The beta simulation tracks both betas and photons, but only betas were used as the source term.

A third simulation could have been run with photons as the source. This would have been a sum of the capture gammas, transition gammas and gammas associated with beta decay. This would also have been a photon/electron problem. Electrical current produced would be the product of the sum of the second two runs with appropriate conversion factors. A gamma production cross section is required for this type of calculation but was not available for rhodium, so this contribution was neglected.

Admittedly, this method leaves something to be desired as neutron capture in other regions, as well as other neutron-gamma production modes is ignored. This is unavoidable as a ^{103}Rh cross section table with gamma production is not available.

4.7 Decoupling Matrix Derived From Model Response

Since every detector element will be at least somewhat responsive to all the neutron energies, the signals derived from them will be coupled. By noting the response of each detector segment as it is irradiated with neutrons of known energies, a response matrix can be developed as a model for the spectrometer response. The successive measurements can be multiplied by the inverse of this response matrix, and the thermal, epithermal, and fast neutron spectrum components can be determined.

4.8 Time Dependence

There are two time-dependent considerations associated with SPNSs. The first is that there is a time delay between neutron absorption and beta emission. This can

easily be dealt with by waiting ten half lives of the daughter nuclide, so the device reaches approximate steady-state operation. In this case, there is no time dependence built into the MCNP model so all betas were assumed to be emitted at the same time as neutron capture.

The second time-dependent factor involves burning of the emitter atoms. Exposure of the SPNS while in operation will result in burnup or loss of the beta-emitting atoms as they are converted to another isotope with the capture of a neutron, and then to another element with the subsequent beta decay. The response of the detector will decrease with the loss of available material. Burnup is also a time-dependent problem beyond the scope of this thesis, but would have to be considered in an operational spectrometer.

4.9 Final design

The final form of the SPNS looks much like the form shown in Figure 3. The exact dimensions are shown in Table 3. The dimensions given produce a total

Table 3. SPNS Dimensions.

	Material	Density (g cm⁻³)	Outer radius (cm)	Thickness (cm)
Inner beta region	Rh	12.4	0.1	0.2
Inner insulator	MgO	3.97	0.108	0.008
Inner conductor	Fe	7.86	0.15	0.042
Middle beta region	V	6.1	0.25	0.1
Middle insulator	MgO	3.97	0.258	0.008
Middle conductor	Fe	7.86	0.3	0.042
Outer beta region	Mn	7.43	0.4	0.1
Outer insulator	MgO	3.97	0.408	0.008
Outer conductor	Fe	7.86	0.5	0.092

diameter of 1 cm. This device can be any length, but an arbitrary length of 6 cm was used for the model.

These dimensions were constrained by a 1 cm diameter. Actual thicknesses of the regions were compromises between size and device sensitivity as determined in Tables 4 and 5

Chapter 5. RESULTS

The spectrometer absorption response can be expressed as a vector, with each component corresponding to the neutron absorption per unit flux in each of the three detector regions. Normalizing (making unit length) is one way of comparing these vectors. The response vectors for the three energy regions (thermal, epithermal, fast) together become a response matrix which, when multiplied by a vector representing an incident flux, gives an approximate prediction of the system response.

The initial MCNP4B runs with different emitter materials in the spectrometer are shown in Table 4. The numbers in parentheses represent the relative (or fractional) error, instead of the absolute error, in the MCNP4B results.

Table 4. Response of Three Material SPNS

Raw Data	Inner SPND (Rh)	Middle SPNS (V)	Outer SPNS (Mn)
Thermal	0.087599 (0.04249)	0.024207 (0.0419)	0.13449 (0.0341)
Epithermal	0.18993 (0.0374)	0.051765 (0.0299)	0.28388 (0.0259)
Fast	0.51722 (0.0205)	0.131867 (0.0168)	0.729653 (0.0149)

These data show an almost complete indifference in the response ratios to the three neutron energy regions (see Appendix C).

Table 5 shows the response of the spectrometer with cadmium shielding added (in place of some of the conductor regions) to shape the neutron flux.

Table 5. Response of Cadmium Shielded Rhodium SPNS

Raw Data	Inner SPND	Middle SPNS	Outer SPNS
Thermal	5.12035E-5 (0.1698)	2.97292E-2 (0.0178)	1.60490E-1 (0.0107)
Epithermal	3.89807E-3 (0.0881)	8.85402E-2 (0.0148)	4.30689E-1 (0.0086)
Fast	1.00719E-2 (0.0205)	3.16939E-1 (0.0086)	1.62252E0 (0.0057)

While these vectors are not collinear (not in the same direction), taken together they still form a nearly singular matrix (one with a determinant of nearly zero, that does not span any significant volume.)

Chapter 6. DISCUSSION

Like the first spectrometer type using various neutron absorbers, the second spectrometer design has a nearly singular response with its associated non-invertable matrix. Unlike the first case which has collinear response vectors, the second arrangement is merely coplanar. This means that any pair of response vectors are linearly independent, but the system as a whole would not make a working three-region spectrometer. Two of the regions, thermal and epithermal, could be used for a two element, two-region spectrometer which could distinguish between thermal and higher energy neutrons. This is due to the fact that the fast and epithermal responses are in the same plane as the thermal response, but each has a different angle from the thermal response vector. Any linear combination of the vectors is in this same plane, which means a lack of linear independence among the response vectors. This means that the system cannot be inverted, so the incident neutron spectrum cannot be determined from the response of this SPNS.

Chapter 7. SUMMARY AND CONCLUSIONS

The self-powered neutron detector has many advantages. It can be formed using simple conduction and insulation materials, can withstand intense neutron bombardment, and can be made relatively small as compared to a series of Bonner spheres. In addition, it can be made in a cylindrical shape allowing easy insertion into small apertures, as may be required in a reactor or other heavily shielded structure.

The SPNS readout is simple and requires minimal instrumentation. A remote picoammeter is all that is needed as a display. A small number of parts, coupled with remote control electronics (away from any neutron flux) may provide high system reliability.

By reading several coaxial SPNDs with associated computer-interfaced meters, a neutron spectrum can be obtained with separate flux readings for thermal, epithermal, and fast neutron energies, if these responses are linearly independent.

Using various neutron absorbers/beta emitters in this design does not produce a workable spectrometer. Any emitter material seems to have much the same neutron absorption response to the various energy regions.

Modifying the neutron flux with cadmium shielding does produce a spectrometer that distinguishes between energies. This can be useful for a two-region spectrometer, but the response vectors for thermal, epithermal, and fast neutron fluxes are coplanar so this version of a SPNS also cannot be used to distinguish three regions.

Further work can be carried out by building a detector and operating it in a reactor to verify its operation against reactor models. Another idea for further work is to add more elements and make a spectrometer that acts more like a large number of Bonner spheres. One more important area to research is the time dependence of this spectrometer with varying flux and with “burnup” instead of assuming a steady-state calibrated system.

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APPENDICES

Appendix A: Beta Emitter MCNP4B Output

lmcnp version 4b ld=02/04/97 03/29/98 07:26:01
 ***** probid = 03/29/98 07:26:01

```

1- Sphere $ V sphere Unit area detector region, vary insulator
2- 10 5 -6.1 -200 IMP:N,P,E= 1. $ core (V)
3- 20 5 -6.1 200 -201 IMP:N,P,E= 1. $ beta emitter/neutron abs (V)
4- C 20 6 -7.43 200 -201 IMP:N,P,E= 1. $ beta emitter/neutron abs (Mn)
5- 30 3 -3.58 201 -202 IMP:N,P,E= 1. $ insulator (source side) (MnO)
6- 35 3 -3.58 202 -203 IMP:N,P,E= 1. $ insulator (collector side)
7- 40 2 -7.86 203 -204 IMP:N,P,E= 1. $ conductor (Fe)
8- 98 1 -1.00 204 -207 IMP:N,P,E= 1. $ water
9- 99 0 207 IMP:N,P,E= 0. $ outside universe (Void)
10-
11- C length units are in cm, mass densities in g/cu cm, cross sections in barns
12- C radius of sphere with surface area of 1 cu cm: 0.2820947917739 cm
13- C 0.02 cm thick beta region
14- 200 SO 0.2620947917739 $ vary 200 for plot (and source term below)
15- 201 SO 0.2820947917739
16- 202 SO 0.2860947917739 $ half way point in insulator
17- 203 SO 0.2900947917739 $ fixed insulator region (0.008 cm)
18- 204 SO 0.3020947917739
19- 207 SO 0.32
20-
21- M1 001001 2 008016 1 $ water {d=1.00}
22- M2 026000 1 $ iron (conductors/beta shield) {d=7.86}
23- M3 012000 1 008016 1 $ MnO magnesium oxide (insulator) {d=3.58}
24- C M4 45103.01e 1 $ rhodium (rh, 100 barn, 42s){.50c -> n} {Rh d=12.4}
25- M5 23000.01e 1 $ vanadium (5.0, 2.8) {V d=6.1}
26- C M6 025000.01e 1 $ Manganese {Mn d=7.43}
27- C neutron section *-*-*-*-*-*-*-*-*-*-*-*-*-*-*-*
28- C MODE N $ implied if missing
29- C SDEF SUR=207 NRM=-1 ERG=14 WGT=1 PAR=1
30- C PHYS:N 15 0
31- C F1:N 201
32- C F4:N 10 20 30
33- C associated beta emission *-*-*-*-*-*-*-*-*-*-*
34- C sdef: sphere defined by using pos and rad (not axs or ext -> cylinder)
35- MODE P E
36- SDEF POS=0 0 0 ERG=D1 RAD=D2 WGT=1 PAR=3
37- C beta spectra from Table of Radioactive Isotopes
38- C Edgardo Browne and Richard B. Firestone
39- C New York, John Wiley & Sons, 1986
40- C ==-== Rh-104 beta spectrum (p104-3): SpA: 2.544E9 Ci/g
41- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.442
42- C SP1 0 0.00241 0.00247 0.0051 0.0170 0.075 0.159 0.446 0.289
43- C ==-== V-52 beta spectrum(p52-1): 9.63E8 Ci/g
44- SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.5 2.542
  
```

```

45- SP1 0 0.00156 0.00162 0.00344 0.0122 0.061 0.141 0.434 0.345 0.0000259
46- C ---- Mn-56 beta spectrum(p56-2): SpA: 2.1702E7 Ci/g
47- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.5 2.849
48- C SP1 0 0.0069 0.0070 0.0143 0.0466 0.174 0.230 0.270 0.245 0.0055
49- SI2 0.2620947917739 0.2820947917739 $ <<----- variable source thickness
50- SP2 -21 2 $ -21 = sphere , 2 ==> r^2 source probability function/(default)
51- C Tally type 1: Current integrated over a surface (particles * weight / nps)
52- FC1 Tally comment == this should be a current, not flux, directional ==
53- F1:E 201 202 203 $ current over these spherical surfaces (particles)
54- FT1 ELC 2 $ special tally modifier, positron, electron, sum, separated
55- C common section *-*-*-*-*-*-*-*-*-*-*-*-*-*-*
56- NPS 20000 $ history limit
57- CTME 30 $ processor time (not absolute) limit
58-

```

warning. 2 materials had unnormalized fractions. print table 40.
lcells

print table 60

cell	mat	atom density	gram density	volume	mass	pieces	neutron importance	photon importance	electron importance	
1	10	5	7.21111E-02	6.10000E+00	7.54161E-02	4.60038E-01	1	1.0000E+00	1.0000E+00	1.0000E+00
2	20	5	7.21111E-02	6.10000E+00	1.86155E-02	1.13555E-01	1	1.0000E+00	1.0000E+00	1.0000E+00
3	30	3	1.06992E-01	3.58000E+00	4.05699E-03	1.45240E-02	1	1.0000E+00	1.0000E+00	1.0000E+00
4	35	3	1.06992E-01	3.58000E+00	4.17203E-03	1.49359E-02	1	1.0000E+00	1.0000E+00	1.0000E+00
5	40	2	8.47576E-02	7.86000E+00	1.32225E-02	1.03928E-01	1	1.0000E+00	1.0000E+00	1.0000E+00
6	98	1	1.00309E-01	1.00000E+00	2.17752E-02	2.17752E-02	1	1.0000E+00	1.0000E+00	1.0000E+00
7	99	0	.00000E+00	.00000E+00	.00000E+00	.00000E+00	0	.0000E+00	.0000E+00	.0000E+00

total 1.37258E-01 7.28756E-01

minimum source weight = 9.9984E-01 maximum source weight = 1.0000E+00

1 warning message so far.

lcross-section tables

print table 100

table length

tables from file mcplib02

1000.02p	623	01/15/93
8000.02p	623	01/15/93
12000.02p	643	01/15/93
23000.02p	651	01/15/93
26000.02p	651	01/15/93

total 3191

maximum photon energy set to 100.0 mev (maximum electron energy)

tables from file ell

1000.01e	478				11/16/88
8000.01e	478				11/16/88
12000.01e	478				11/16/88
23000.01e	478				11/16/88
26000.01e	478				11/16/88

decimal words of dynamically allocated storage

general	230520		
tallies	6608		
bank	6803		
cross sections	3191		
total	243258	=	973032 bytes

 dump no. 1 on file runtpe nps = 0 coll = 0 ctm = .00 nrn = 0

1 warning message so far.
 bank is full. bank backup file is being created. nps = 18

 dump no. 2 on file runtpe nps = 8198 coll = 56255636 ctm = 15.01 nrn = 347604307

1problem summary

run terminated when it had used 30 minutes of computer time.

+ Sphere \$ V sphere Unit area detector region, vary insulator probid = 03/29/98 07:56:51
 0 probid = 03/29/98 07:26:01

photon creation	tracks	weight (per source particle)	energy	photon loss	tracks	weight (per source particle)	energy
source	0	0.	0.	escape	1261	7.7321E-02	1.4938E-02
				energy cutoff	0	0.	0.
				time cutoff	0	0.	0.
weight window	0	0.	0.	weight window	0	0.	0.
cell importance	0	0.	0.	cell importance	0	0.	0.
weight cutoff	0	0.	0.	weight cutoff	0	0.	0.
energy importance	0	0.	0.	energy importance	0	0.	0.
dxtran	0	0.	0.	dxtran	0	0.	0.
forced collisions	0	0.	0.	forced collisions	0	0.	0.
exp. transform	0	0.	0.	exp. transform	0	0.	0.
from neutrons	0	0.	0.	compton scatter	0	0.	4.2913E-04
bremsstrahlung	6533	4.0059E-01	1.9328E-02	capture	13193	8.0896E-01	6.4447E-03

p-annihilation	0	0.	0.	pair production	0	0.	0.
electron x-rays	7796	4.7803E-01	2.4419E-03				
1st fluorescence	125	7.6647E-03	4.2196E-05				
2nd fluorescence	0	0.	0.				
total	14454	8.8628E-01	2.1812E-02	total	14454	8.8628E-01	2.1812E-02

number of photons banked	14329	average time of (shakes)	cutoffs
photon tracks per source particle	8.8642E-01	escape	tco 1.0000E+34
photon collisions per source particle	8.3773E-01	capture	eco 1.0000E-03
total photon collisions	13660	capture or escape	wc1 -5.0000E-01
		any termination	wc2 -2.5000E-01

0	electron creation	tracks	weight	energy	electron loss	tracks	weight	energy
			(per source particle)				(per source particle)	
	source	16306	9.9984E-01	1.1486E+00	escape	4601	2.8212E-01	2.8618E-01
					energy cutoff	1391595	8.5329E+01	8.4105E-02
					time cutoff	0	0.	0.
	weight window	0	0.	0.	weight window	0	0.	0.
	cell importance	0	0.	0.	cell importance	0	0.	0.
	weight cutoff	0	0.	0.	weight cutoff	0	0.	0.
	energy importance	0	0.	0.	energy importance	0	0.	0.
	pair production	0	0.	0.	scattering	0	0.	1.1390E+00
	compton recoil	189	1.1589E-02	4.2858E-04	bremsstrahlung	0	0.	1.9435E-02
	photo-electric	13153	8.0651E-01	6.2555E-03				
	photon auger	403	2.4711E-02	1.2990E-04				
	electron auger	47860	2.9347E+00	1.0618E-02				
	knock-on	1318285	8.0834E+01	3.6262E-01				
	total	1396196	8.5611E+01	1.5287E+00	total	1396196	8.5611E+01	1.5287E+00

number of electrons banked	1379890	cutoffs
electron tracks per source particle	8.5625E+01	tco 1.0000E+34
electron sub-steps per source particle	6.8982E+03	eco 1.0000E-03
total electron sub-steps	112482652	wc1 .0000E+00
		wc2 .0000E+00

computer time so far in this run	30.12 minutes	maximum number ever in bank	200
computer time in mcrun	30.01 minutes	bank overflows to backup file	2693
source particles per minute	5.4343E+02	dynamic storage	243262 words, 973048 bytes.
random numbers generated	694958440	most random numbers used was	195362 in history 8342

warning. random number stride 152917 exceeded 34 times.

range of sampled source weights = 9.9984E-01 to 9.9984E-01
 1photon activity in each cell

print table 126

cell	tracks entering	population	collisions	collisions * weight (per history)	number weighted energy	flux weighted energy	average track weight (relative)	average track mfp (cm)	
1	10	382	6453	5976	3.6643E-01	1.8211E-01	1.8211E-01	9.9984E-01	8.6588E-01

2	20	862	6078	4929	3.0223E-01	1.5851E-01	1.5851E-01	9.9984E-01	7.3425E-01
3	30	1237	1445	266	1.6310E-02	1.6834E-01	1.6834E-01	9.9984E-01	1.7186E+00
4	35	1265	1450	239	1.4655E-02	1.6611E-01	1.6611E-01	9.9984E-01	1.6796E+00
5	40	1165	3435	2223	1.3631E-01	1.7248E-01	1.7248E-01	9.9984E-01	5.5480E-01
6	98	1257	1283	27	1.6556E-03	1.9055E-01	1.9055E-01	9.9984E-01	6.5025E+00
total		6168	20144	13660	8.3760E-01				

electron activity in each cell print table 126

cell	tracks entering	population	substeps	substeps * weight (per history)	number weighted energy	flux weighted energy	average track weight (relative)	average track mfp (cm)
1	10	8453	555282	46734118	2.8656E+03	6.3605E-01	7.8123E-01	1.3256E-03
2	20	22585	514501	42433033	2.6019E+03	7.2212E-01	8.8685E-01	1.5067E-03
3	30	12337	64693	2646375	1.6227E+02	7.3125E-01	8.7549E-01	3.3505E-03
4	35	11975	63093	2586491	1.5860E+02	7.0545E-01	8.4995E-01	3.2483E-03
5	40	9090	212743	17383664	1.0659E+03	6.9208E-01	8.5015E-01	1.1031E-03
6	98	4878	23380	698971	4.2859E+01	8.4447E-01	9.5955E-01	1.4613E-02
total		69318	1433692	112482652	6.8972E+03			

1tally 1 nps = 16306
+ Talley comment == this should be a current, not flux, directional ==
tally type 1 number of particles crossing a surface.
tally for electrons
this tally is modified by ft elc

surface	201	.00000E+00	.0000
surface	201	7.57517E-01	.0106
surface	201	7.57517E-01	.0106
surface	202	.00000E+00	.0000
surface	202	7.25509E-01	.0109
surface	202	7.25509E-01	.0109
surface	203	.00000E+00	.0000
surface	203	7.12326E-01	.0116

surface 203 7.12326E-01 .0116
 lanalysis of the results in the tally fluctuation chart (tfc) for tally 1 with nps = 16306 print table 160

normed average tally per history = 7.57517E-01	unnormed average tally per history = 7.57517E-01
estimated tally relative error = .0106	estimated variance of the variance = .0007
relative error from zero tallies = .0076	relative error from nonzero scores = .0074
number of nonzero history tallies = 8361	efficiency for the nonzero tallies = .5128
history number of largest tally = 10496	largest unnormalized history tally = 1.19981E+01
(largest tally)/(average tally) = 1.58388E+01	(largest tally)/(avg nonzero tally) = 8.12142E+00
(confidence interval shift)/mean = .0001	shifted confidence interval center = 7.57594E-01

if the largest history score sampled so far were to occur on the next history, the tfc bin quantities would change as follows:

estimated quantities	value at nps	value at nps+1	value(nps+1)/value(nps)-1.
mean	7.57517E-01	7.58206E-01	.000910
relative error	1.06375E-02	1.06660E-02	.002678
variance of the variance	7.25050E-04	7.66049E-04	.056545
shifted center	7.57594E-01	7.57595E-01	.000002
figure of merit	2.94519E+02	2.92947E+02	-.005335

the estimated slope of the 84 largest tallies starting at 5.49914E+00 appears to be decreasing at least exponentially.
 the history score probability density function appears to have an unsampled region at the largest history scores: please examine.

=====

results of 10 statistical checks for the estimated answer for the tally fluctuation chart (tfc) bin of tally 1

tfc bin	--mean--	-----relative error-----			----variance of the variance----			--figure of merit--		-pdf-
behavior	behavior	value	decrease	decrease rate	value	decrease	decrease rate	value	behavior	slope
desired	random	< .10	yes	1/sqrt(nps)	<0.10	yes	1/nps	constant	random	>3.00
observed	random	.01	yes	yes	.00	yes	yes	constant	random	10.00
passed?	yes	yes	yes	yes	yes	yes	yes	yes	yes	yes

=====

this tally meets the statistical criteria used to form confidence intervals: check the tally fluctuation chart to verify.
 the results in other bins associated with this tally may not meet these statistical criteria.

estimated asymmetric confidence interval(1,2,3 sigma): 7.4953E-01 to 7.6565E-01; 7.4148E-01 to 7.7371E-01; 7.3342E-01 to 7.8177E-01
 estimated symmetric confidence interval(1,2,3 sigma): 7.4946E-01 to 7.6558E-01; 7.4140E-01 to 7.7363E-01; 7.3334E-01 to 7.8169E-01

fom = (histories/minute)*(f(x) signal-to-noise ratio)**2 = (5.434E+02)*(7.362E-01)**2 = (5.434E+02)*(5.420E-01) = 2.945E+02
 1status of the statistical checks used to form confidence intervals for the mean for each tally bin

tally result of statistical checks for the tfc bin (the first check not passed is listed) and error magnitude check for all bins
 1 passed the 10 statistical checks for the tally fluctuation chart bin result
 passed all bin error check: 9 tally bins had 3 bins with zeros and 0 bins with relative errors exceeding .10

the 10 statistical checks are only for the tally fluctuation chart bin and do not apply to other tally bins.

the tally bins with zeros may or may not be correct: compare the source, cutoffs, multipliers, et cetera with the tally bins.

1tally fluctuation charts

nps	mean	tally error	vov	slope	fom
1000	7.4488E-01	.0412	.0086	10.0	337
2000	7.6638E-01	.0295	.0051	10.0	316
3000	7.6621E-01	.0239	.0031	10.0	320
4000	7.5763E-01	.0208	.0023	10.0	319
5000	7.4848E-01	.0186	.0018	10.0	317
6000	7.5588E-01	.0171	.0015	10.0	311
7000	7.5888E-01	.0158	.0013	10.0	313
8000	7.5976E-01	.0148	.0011	10.0	311
9000	7.5755E-01	.0140	.0011	10.0	308
10000	7.5988E-01	.0134	.0011	10.0	303
11000	7.6143E-01	.0129	.0011	10.0	298
12000	7.6130E-01	.0123	.0010	10.0	299
13000	7.5796E-01	.0119	.0009	10.0	297
14000	7.5924E-01	.0114	.0008	10.0	297
15000	7.5602E-01	.0110	.0008	10.0	297
16000	7.5782E-01	.0107	.0007	10.0	295
16306	7.5752E-01	.0106	.0007	10.0	295

 dump no. 3 on file runtpe nps = 16306 coll = 112496312 ctm = 30.01 nrn = 694958440

2 warning messages so far.

run terminated when it had used 30 minutes of computer time.

computer time = 30.13 minutes

mcnp version 4b 02/04/97

03/29/98 07:56:52

probid = 03/29/98 07:26:01

Appendix B: Beta Emission MCNP4B Output

lmcnp version 4b ld=02/04/97 04/29/98 23:27:00

probid = 04/29/98 23:27:00

```

1- Self-Powered Neutron Spectrometer *Cd Shield* $Title card
2- C Edward K. Kropp, Oregon State University
3- C to execute this file (named inp), type: mcnp4b
4- C three detector spnd:
5- C Thermal E(n) < 0.5 eV
6- C Epithermal 0.5 eV < E(n) < 10 keV
7- C Fast 10 keV < E(n) (fission: 100 keV - 15 MeV)
8- C
9- C Cell Material Density Surfaces Particle importance
10- C V--- density: -value = g/cm^3
11- C INNER DETECTOR ELEMENT (Rh):
12- C inner core conductor/beta emitter material
13- 11 4 -12.4 -211 -250 260 $ Rh
14- C inner insulator material A
15- 12 3 -3.97 211 -212 -250 260 $ MgO
16- C inner insulator material B
17- 13 3 -3.97 212 -213 -250 260 $ MgO
18- C collector/conductor/beta shield
19- 14 2 -7.86 213 -214 -250 260 $ Fe
20- 15 30 -8.65 214 -215 -250 260 $ Cd
21- C
22- C MIDDLE DETECTOR ELEMENT (Rh):
23- C middle neutron absorber/beta emitter material
24- 21 4 -12.4 215 -221 -250 260 $ Rh
25- C 21 5 -6.1 215 -221 -250 260 $ V
26- C middle insulator material
27- 22 3 -3.97 221 -222 -250 260 $ MgO
28- C middle insulator material
29- 23 3 -3.97 222 -223 -250 260 $ MgO
30- C middle collector/conductor/beta shield
31- 24 2 -7.86 223 -224 -250 260 $ Fe
32- 25 2 -7.86 224 -225 -250 260 $ Fe (Rh is enough shield)
33- C 25 30 -8.65 224 -225 -250 260 $ Cd
34- C
35- C OUTER DETECTOR ELEMENT (Rh):
36- C middle neutron absorber/beta emitter material
37- 31 4 -12.4 225 -231 -250 260 $ Rh
38- C 31 6 -7.43 225 -231 -250 260 $ Mn
39- C outer insulator material A
40- 32 3 -3.97 231 -232 -250 260 $ MgO
41- C outer insulator material B
42- 33 3 -3.97 232 -233 -250 260 $ MgO
43- C outer detector collector material
44- 34 2 -7.86 233 -234 -250 260 $ Fe
45- C
  
```



```

46- C OTHER SPECTROMETER PARTS:
47- C endcaps (make material = water for no endcap) M30 = Cd
48- 40 30 -8.65 -225 250 -251 $ top
49- 41 30 -8.65 -234 250 -259 #40 $ endcaps split for IMP matching
50- 50 30 -8.65 -225 261 -260 $ bottom
51- 51 30 -8.65 -234 269 -260 #50
52- C
53- C OUTSIDE REGIONS:
54- 80 1 1.0 #(-234 -259 269) -299 -280 $ water (presume LLR coolant)
55- 81 1 1.0 280 -281 -299 $ water (presume LLR coolant)
56- 82 1 1.0 281 -282 -299 $ "
57- 83 1 1.0 282 -283 -299 $ "
58- 84 1 1.0 283 -299 $ "
59- 99 0 299 $ Void outside universe
60-
61- C Surface Geometry Dimension (radius, cm)
62- C cylindrical boundaries
63- 211 CX 0.1 $ boundary inner beta emitter & insulator (arbitrary) Rh
64- 212 CX 0.104 $ halfway point in insulator
65- 213 CX 0.108 $ boundary inner insulator & collector/conductor/shield
66- 214 CX 0.12 $ Cd shield
67- 215 CX 0.15 $ boundary inner and middle detectors (arbitrary)
68- 221 CX 0.25 $ outer surface of 2nd beta region V
69- 222 CX 0.254 $ halfway point in insulator
70- 223 CX 0.258 $ boundary middle beta insulator & collector/conductor/shield
71- 224 CX 0.29 $ Cd shield
72- 225 CX 0.3 $ boundary middle and outer detectors (arbitrary)
73- 231 CX 0.4 $ boundary outer beta emitter & insulator Mn
74- 232 CX 0.404 $ halfway point in insulator
75- 233 CX 0.408 $ boundary outer insulator $ collector/conductor/shield
76- 234 CX 0.5 $ outer radius of spectrometer outer boundaries
77- 250 PX 3.0 $ detector top
78- 251 PX 3.1 $ endcap mid
79- 259 PX 3.2 $ endcap top
80- 260 PX -3.0 $ detector bottom
81- 261 PX -3.1 $ endcap mid
82- 269 PX -3.2 $ endcap bottom
83- 280 CX 1.0 $ surfaces for neutron importances...
84- 281 CX 1.5 $ "
85- 282 CX 2.0 $ "
86- 283 CX 2.5 $ "
87- 299 SO 3.5 $ spherical system limit (and isotropic neutron flux source)
88-
89- C Instructions Section
90- C neutron cross sections in .50c and .60c
91- C common materials (Appendix G mcnp4a manual)
92- M1 001001 2 008016 1 $ water {d=1.00}
93- M2 026000 1 $ Fe (conductors/beta shield) {d=7.86}
94- M3 012000 1 008016 1 $ MgO magnesium oxide (insulator) {d=3.58}
95- M4 45103 1 $ rhodium (nb, 100 barn, 42s) (.50c -> n) {d=12.4}
96- C 45103 defaults to .50c, 45103.50c & 50d both lack gamma ray production

```

```

97-      C M5 23000.60c 1          $ vanadium (5.0, 2.8)                {d=6.1}
98-      C M6 25055.60c 1          $ manganese                      {d=7.43}
99-      C M4 45103.01e 1          $ rhodium (nb, 100 barn, 42s) (.50c -> n) {d=12.4}
100-     C M5 23000.01e 1          $ vanadium (5.0, 2.8)                {d=6.1}
101-     C M6 25000.01e 1          $ manganese                      {d=7.43}
102-     C M7 13027.01e 1          $ aluminum                       {d=2.70}
103-     C additional materials
104-     M30 048000 1              $ cadmium (thermal neutron shield) {d=8.65}
105-     C M40 027059 1            $ cobalt (nge, 35 barn, prompt) {d=8.90}
106-     C M41 078000 1            $ platinum (nge+ge, <10 barn, prompt) {d=21.4}
107-     C M50 82000 1             $ lead                          {d=11.4}
108-     C . . . . .
109-     C . step 1, neutron absorption problem . (with combinations)
110-     C . . . . .
111-     MODE N
112-     C . . . . . Neutron source term
113-     C TRIGA (@ Lazy Susan) t%:3E12 e%:1E11 f%:6E11 (@ Rabbit) t%:9E12 e%:4E11 f%:3E1
114-     C SDEF POS=0 5 0 CEL=98 ERG=14 WGT=1 $ neutron point source N:PAR=1
115-     SDEF SUR=299 NRM=-1 ERG=D1 WGT=154 PAR=1 $ inward/isotropic n
116-     C source on spherical surface, flux: 1 n cm^-2 s^-1 (area = 154 cm^2)
117-     SI1 S 2 3 4
118-     SP1 0 0 6 $ epithermal only!
119-     C SP1 30 1 6 $ Lazy Susan ratios
120-     C thermal: 0 - 0.5ev
121-     C ERG -2 a $ p(E)=C E^.5 exp(-E/a) a: temperature in MeV
122-     SP2 -2 2.53E-8 $ thermal/Maxwellian distribution
123-     C epithermal: 0.5 - 10 keV
124-     C $ epithermal/ 1/E distribution histogram
125-     SI3 H 0 5E-7 1E-6 2E-6 5E-6 1E-5 2E-5 5E-5 1E-4 2E-4 5E-4 1E-3 3E-3 5E-3 1E-2
126-     SP3 0 0.69 0.69 0.91 0.69 0.69 0.91 0.69 0.69 0.91 0.69 0.69 0.91 0.69 0.69
127-     C ERG -3 a b $ p(E) = C exp(-E/a) sinh(bE)^.5 a: width b: ave energy (p3-50)
128-     C defaults a = 0.965 MeV b = 2.29 MeV^-1
129-     C fast: > ???????????? MeV
130-     SP4 -3 $ fast/Watt spectrum
131-     PHYS:N 15 0 $ default p3-103
132-     C tallies.....
133-     C F4 flux averaged over a cell
134-     F4:N 11 $ straight tally (track length estimate??)
135-     FC4: Number of Rh-104 atoms created in cell 11
136-     C FM# C M# Xsect-type FMx: p3-75, ex p 4-34
137-     C where C = conversion Av*10^-24 * (mass of cell)/AW
138-     C AW = Rh:102.90660 V:50.9415 Mn:54.93805 Al:26.981539
139-     FM4 0.013678 4 102 $ Rh
140-     C 0.04786:conversion, atom/barn*cm) 4:material 4 (Rh) 102:radiative capture
141-     F14:N 21 $ straight tally
142-     FC14: Number of V-52 atoms created in cell 21
143-     FM14 0.054712 4 102 $ Rh
144-     C FM14 0.05437 5 102 $ V
145-     F24:N 31 $ straight tally
146-     FC24: Number of Mn-56 atoms created in cell 31
147-     FM24 0.095752 4 102 $ Rh

```

```

148- C FM24 0.107462 6 102 $Mn
149- C . . . . .
150- C . . step 2, beta emission/absorption problem . .
151- C . . . . .
152- C CUT:E 0.0001 $ MCNP4A low energy cutoff (100 Volts)
153- C SDEF CEL=D1 POS=0 0 0 AXS=0 0 1 EXT=D2 RAD=FCEL D3 ERG FCEL=D7 WGT=100 PAR=3
154- C SC1 beta sources (part II/after neutron section is run)
155- C SI1 L 11 21 31 $ beta source cells
156- C SP1 d 0.2 0.7 0.1 $ probability (strength) of each cell
157- C SB1 d 0.33 0.33 0.34 $ source bias (set equal for now)
158- C SI2 3. $ half height for each beta source cell
159- C DS3 s 4 5 6 $ set of set of cell radii
160- C SI4 0 0.3
161- C SI5 0.5 0.6
162- C SI6 0.8 0.9
163- C ---- Al-28 beta spectrum
164- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.5 2.863
165- C SP1 0 0.00080 0.00087 0.00197 0.0077 0.0425 0.108 0.380 0.448 0.0112
166- C ---- Rh-104 beta spectrum (p104-3): SpA: 2.544E9 Ci/g
167- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.442
168- C SP1 0 0.00241 0.00247 0.0051 0.0170 0.075 0.159 0.446 0.289
169- C ---- V-52 beta spectrum (p52-1): 9.63E8 Ci/g
170- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.5 2.542
171- C SP1 0 0.00156 0.00162 0.00344 0.0122 0.061 0.141 0.434 0.345 0.0000259
172- C ---- Mn-56 beta spectrum (p56-2): SpA: 2.1702E7 Ci/g
173- C SI1 H 0 0.01 0.02 0.04 0.1 0.3 0.6 1.3 2.5 2.849
174- C SP1 0 0.0069 0.0070 0.0143 0.0466 0.174 0.230 0.270 0.245 0.0055
175- C tallies.....
176- C . . . . .
177- C . step 3, associated photon emission .
178- C . . . . .
179- C need to look this up!!! Monday!!!
180- C . . . . .
181- C . . . . .
182- C beta current (directional) tallies
183- C F1: current integrated over a surface (unit: particles)
184- C F1:E 212 222 232 $ beta tallies for surfaces pg. 3-60
185- C FT1 ELC 2 $ special tally modifier, positron, electron, sum, seperated
186- C FC1 electron tallies 212:inner 222:middle 232:outer
187- C . . . . .
188- C . common commands for n, p and e runs: .
189- C . . . . .
190- IMP:N 64 64 64 64 64 32 32 32 32 16 16 16 16 16 32 16 32 16 8 4 2 1 1 0
191- IMP:P,E 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 0
192- C Print $ print tables....
193- NPS 3000000 $ run 50000 origin particles
194- CTME 180
195-

```

warning. 2 materials had unnormalized fractions. print table 40.
1cells

print table 60

cell	mat	atom density	gram density	volume	mass	pieces	neutron importance	photon importance	electron importance	
1	11	4	7.25650E-02	1.24000E+01	1.88496E-01	2.33734E+00	1	6.4000E+01	1.0000E+00	1.0000E+00
2	12	3	1.18648E-01	3.97000E+00	1.53812E-02	6.10635E-02	1	6.4000E+01	1.0000E+00	1.0000E+00
3	13	3	1.18648E-01	3.97000E+00	1.59844E-02	6.34582E-02	1	6.4000E+01	1.0000E+00	1.0000E+00
4	14	2	8.47576E-02	7.86000E+00	5.15724E-02	4.05359E-01	1	6.4000E+01	1.0000E+00	1.0000E+00
5	15	30	4.63393E-02	8.65000E+00	1.52681E-01	1.32069E+00	1	6.4000E+01	1.0000E+00	1.0000E+00
6	21	4	7.25650E-02	1.24000E+01	7.53982E-01	9.34938E+00	1	3.2000E+01	1.0000E+00	1.0000E+00
7	22	3	1.18648E-01	3.97000E+00	3.80007E-02	1.50863E-01	1	3.2000E+01	1.0000E+00	1.0000E+00
8	23	3	1.18648E-01	3.97000E+00	3.86039E-02	1.53257E-01	1	3.2000E+01	1.0000E+00	1.0000E+00
9	24	2	8.47576E-02	7.86000E+00	3.30546E-01	2.59809E+00	1	3.2000E+01	1.0000E+00	1.0000E+00
10	25	2	8.47576E-02	7.86000E+00	1.11212E-01	8.74129E-01	1	1.6000E+01	1.0000E+00	1.0000E+00
11	31	4	7.25650E-02	1.24000E+01	1.31947E+00	1.63614E+01	1	1.6000E+01	1.0000E+00	1.0000E+00
12	32	3	1.18648E-01	3.97000E+00	6.06202E-02	2.40662E-01	1	1.6000E+01	1.0000E+00	1.0000E+00
13	33	3	1.18648E-01	3.97000E+00	6.12234E-02	2.43057E-01	1	1.6000E+01	1.0000E+00	1.0000E+00
14	34	2	8.47576E-02	7.86000E+00	1.57462E+00	1.23765E+01	1	1.6000E+01	1.0000E+00	1.0000E+00
15	40	30	4.63393E-02	8.65000E+00	2.82743E-02	2.44573E-01	1	3.2000E+01	1.0000E+00	1.0000E+00
16	41	30	4.63393E-02	8.65000E+00	1.28805E-01	1.11417E+00	1	1.6000E+01	1.0000E+00	1.0000E+00
17	50	30	4.63393E-02	8.65000E+00	2.82743E-02	2.44573E-01	1	3.2000E+01	1.0000E+00	1.0000E+00
18	51	30	4.63393E-02	8.65000E+00	1.28805E-01	1.11417E+00	1	1.6000E+01	1.0000E+00	1.0000E+00
19	80	1	1.00000E+00	9.96924E+00	1.65095E+01	1.64587E+02	1	8.0000E+00	1.0000E+00	1.0000E+00
20	81	1	1.00000E+00	9.96924E+00	2.55972E+01	2.55184E+02	1	4.0000E+00	1.0000E+00	1.0000E+00
21	82	1	1.00000E+00	9.96924E+00	3.32023E+01	3.31001E+02	1	2.0000E+00	1.0000E+00	1.0000E+00
22	83	1	1.00000E+00	9.96924E+00	3.76965E+01	3.75806E+02	1	1.0000E+00	1.0000E+00	1.0000E+00
23	84	1	1.00000E+00	9.96924E+00	6.15624E+01	6.13730E+02	1	1.0000E+00	1.0000E+00	1.0000E+00
24	99	0	.00000E+00	.00000E+00	.00000E+00	.00000E+00	0	.0000E+00	.0000E+00	.0000E+00

total 1.79594E+02 1.78956E+03

minimum source weight = 1.5400E+02 maximum source weight = 1.5400E+02

1 warning message so far.

1cross-section tables

print table 100

table	length				
tables from file endf601					
1001.60c	1402	1-h-1 from endf-vi.1			
8016.60c	38367	8-o-16 from endf/b-vi	mat 8	25	mat 125 11/25/93
12000.60c	39099	12-mg-nat from endf/b-vi			mat1200 11/25/93
tables from file rmccs1					
26000.55c	76892	njoy		(260)	10/21/82
48000.51c	4762	njoy		(1281)	03/10/80
tables from file rmccsa1					

45103.50c 16241 njoy (1310) 11/01/85

total 176763

any neutrons with energy greater than $\text{emax} = 1.50000\text{E}+01$ from the source or from a collision will be resampled.
neutron cross sections outside the range from $.0000\text{E}+00$ to $1.5000\text{E}+01$ mev are expunged.

decimal words of dynamically allocated storage

general	9854		
tallies	19536		
bank	6803		
cross sections	176763		
total	201364	=	805456 bytes

```
*****
dump no. 1 on file runtpe nps = 0 coll = 0 ctm = .00 nrn = 0
```

1 warning message so far.

```
*****
dump no. 2 on file runtpe nps = 110603 coll = 15032720 ctm = 15.01 nrn = 256851376
```

```
warning. source energy > emax.
nps = 202651 ijk = 253436993943137 erg = 1.5013E+01
```

```
*****
dump no. 3 on file runtpe nps = 221298 coll = 30072519 ctm = 30.01 nrn = 513881005
```

```
*****
dump no. 4 on file runtpe nps = 330248 coll = 45116067 ctm = 45.02 nrn = 770934437
```

```
*****
dump no. 5 on file runtpe nps = 439752 coll = 60164968 ctm = 60.03 nrn = 1028060158
```

```
*****
dump no. 6 on file runtpe nps = 549490 coll = 75208187 ctm = 75.04 nrn = 1285052609
```

```
*****
dump no. 7 on file runtpe nps = 659154 coll = 90252217 ctm = 90.04 nrn = 1542100155
```

```

*****
dump no.   8 on file runtpe      nps =   768242   coll =   105288532   ctm =  105.04   nrn =   1799034046
*****
dump no.   9 on file runtpe      nps =   878101   coll =   120342437   ctm =  120.05   nrn =   2056221286
*****
dump no.  10 on file runtpe      nps =   988314   coll =   135386330   ctm =  135.06   nrn =   2313273594
*****
dump no.  11 on file runtpe      nps =  1098406   coll =   150430175   ctm =  150.07   nrn =   2570270320
*****
dump no.  12 on file runtpe      nps =  1208920   coll =   165472694   ctm =  165.07   nrn =   2827276837
*****

```

problem summary

run terminated when it had used 180 minutes of computer time.

```

+ Self-Powered Neutron Spectrometer *Cd Shield* $Title card      probid = 04/30/98 02:29:30
0                                     04/29/98 23:27:00

```

neutron creation	tracks	weight (per source particle)	energy	neutron loss	tracks	weight (per source particle)	energy
source	1318186	1.5400E+02	1.9819E+00	escape	991653	9.2417E+01	2.5376E-01
				energy cutoff	0	0.	0.
				time cutoff	0	0.	0.
weight window	0	0.	0.	weight window	0	0.	0.
cell importance	2997625	5.4635E+01	1.0488E-01	cell importance	2315232	5.4668E+01	1.0513E-01
weight cutoff	0	1.0092E+01	5.2979E-09	weight cutoff	1008937	1.0121E+01	5.5384E-09
energy importance	0	0.	0.	energy importance	0	0.	0.
dxtran	0	0.	0.	dxtran	0	0.	0.
forced collisions	0	0.	0.	forced collisions	0	0.	0.
exp. transform	0	0.	0.	exp. transform	0	0.	0.
upscattering	0	0.	4.7957E-07	downscattering	0	0.	1.7076E+00
				capture	0	6.1521E+01	2.0274E-02
(n,xn)	22	2.0954E-04	2.2688E-06	loss to (n,xn)	11	1.0477E-04	8.5299E-06
fission	0	0.	0.	loss to fission	0	0.	0.
total	4315833	2.1873E+02	2.0868E+00	total	4315833	2.1873E+02	2.0868E+00

number of neutrons banked	2997636	average time of (shakes)	cutoffs
neutron tracks per source particle	3.2741E+00	escape	tco 1.0000E+34
neutron collisions per source particle	1.3689E+02	capture	eco .0000E+00
total neutron collisions	180449845	capture or escape	wc1 -5.0000E-01
net multiplication	1.0000E+00 .0000	any termination	wc2 -2.5000E-01

```

computer time so far in this run 180.34 minutes      maximum number ever in bank      21
computer time in mcrun          180.01 minutes      bank overflows to backup file    0
source particles per minute      7.3227E+03    dynamic storage 201368 words,    805472 bytes.
random numbers generated         3083182460    most random numbers used was     91692 in history 46183

```

range of sampled source weights = 1.5400E+02 to 1.5400E+02

20 source particles had energies > emax.
1neutron activity in each cell

print table 126

cell	tracks entering	population	collisions	collisions * weight (per history)	number weighted energy	flux weighted energy	average track weight (relative)	average track mfp (cm)
1	11	141053	138508	14929	7.3912E-02	1.8049E-02	4.2448E+02	2.1698E+00
2	12	279665	144004	936	4.5489E-03	1.7880E-02	4.2212E+02	2.9015E+00
3	13	290731	149719	968	4.9208E-03	1.7899E-02	4.2148E+02	2.9047E+00
4	14	313434	167036	3632	1.7802E-02	1.7791E-02	4.2244E+02	3.7392E+00
5	15	425610	262594	60582	1.2758E-01	1.1547E-02	4.1924E+02	3.6982E+00
6	21	394205	336060	114365	5.8103E-01	2.5249E-03	3.7651E+02	2.0331E+00
7	22	463160	337187	1439	1.2355E-02	1.5311E-03	3.4991E+02	2.8063E+00
8	23	471057	341093	1567	1.3588E-02	1.5289E-03	3.5043E+02	2.8095E+00
9	24	510843	376254	20389	1.4108E-01	1.4882E-03	3.4944E+02	3.3690E+00
10	25	275093	225551	3262	4.4550E-02	1.5092E-03	3.5045E+02	3.3862E+00
11	31	544389	473194	259155	2.1067E+00	8.4309E-04	3.1755E+02	1.8128E+00
12	32	573132	469213	1983	2.5221E-02	3.8312E-04	2.7084E+02	2.6666E+00
13	33	581692	472352	2101	2.7098E-02	3.7536E-04	2.6935E+02	2.6757E+00
14	34	707740	556767	114700	1.1039E+00	3.0803E-04	2.5642E+02	2.6910E+00
15	40	22909	22467	1988	2.2536E-02	1.6545E-02	6.9933E+02	3.7025E+00
16	41	52407	47578	19371	3.3284E-01	8.7679E-03	7.1753E+02	3.6255E+00
17	50	22800	22313	2089	2.3466E-02	1.6366E-02	7.0537E+02	3.6915E+00
18	51	53027	48084	20030	3.4157E-01	8.1409E-03	7.1463E+02	3.6107E+00
19	80	1752324	1252212	28675198	4.5275E+02	1.1896E-04	2.0361E+02	1.3352E-01
20	81	2425602	1622766	41164056	9.5332E+02	7.5736E-05	1.4027E+02	1.1402E-01
21	82	2299975	1536000	38487087	1.4591E+03	7.4693E-05	1.1233E+02	1.1330E-01
22	83	1731916	888946	28064727	1.9747E+03	8.2130E-05	1.0281E+02	1.1701E-01
23	84	1723074	1032139	43415291	3.2975E+03	1.6137E-04	1.1356E+02	1.4922E-01
total	16055838	10922037	180449845	8.1424E+03				

```

tally 4 nps = 1318186
+ Number of Rh-104 atoms created in cell 11
tally type 4 track length estimate of particle flux.
tally for neutrons

```

```

volumes
cell: 11
1.88496E-01

```

```

cell 11
multiplier bin: 1.36780E-02 4 102

```



```

estimated tally relative error = .0086
relative error from zero tallies = .0037

number of nonzero history tallies = 67830
history number of largest tally = 982434
(largest tally)/(average tally) = 1.07916E+03

(confidence interval shift)/mean = .0001

estimated variance of the variance = .0008
relative error from nonzero scores = .0078

efficiency for the nonzero tallies = .0515
largest unnormalized history tally = 2.57884E+02
(largest tally)/(avg nonzero tally) = 5.55305E+01

shifted confidence interval center = 3.16969E-01

```

if the largest history score sampled so far were to occur on the next history, the tfc bin quantities would change as follows:

estimated quantities	value at nps	value at nps+1	value(nps+1)/value(nps)-1.
mean	3.16939E-01	3.17199E-01	.000818
relative error	8.64218E-03	8.67369E-03	.003647
variance of the variance	7.60800E-04	8.26143E-04	.085888
shifted center	3.16969E-01	3.16970E-01	.000003
figure of merit	7.43789E+01	7.38393E+01	-.007254

the estimated slope of the 200 largest tallies starting at 7.56632E+01 appears to be decreasing at least exponentially. the large score tail of the empirical history score probability density function appears to have no unsampled regions.

=====

results of 10 statistical checks for the estimated answer for the tally fluctuation chart (tfc) bin of tally 14

tfc bin	--mean--	-----relative error-----			----variance of the variance----			--figure of merit--		-pdf-
behavior	behavior	value	decrease	decrease rate	value	decrease	decrease rate	value	behavior	slope
desired	random	< .10	yes	1/sqrt(nps)	<0.10	yes	1/nps	constant	random	>3.00
observed	random	.01	yes	yes	.00	yes	yes	constant	random	10.00
passed?	yes	yes	yes	yes	yes	yes	yes	yes	yes	yes

=====

this tally meets the statistical criteria used to form confidence intervals: check the tally fluctuation chart to verify. the results in other bins associated with this tally may not meet these statistical criteria.

estimated asymmetric confidence interval(1,2,3 sigma): 3.1423E-01 to 3.1971E-01; 3.1149E-01 to 3.2245E-01; 3.0875E-01 to 3.2519E-01
estimated symmetric confidence interval(1,2,3 sigma): 3.1420E-01 to 3.1968E-01; 3.1146E-01 to 3.2242E-01; 3.0872E-01 to 3.2516E-01

fom = (histories/minute)*(f(x) signal-to-noise ratio)**2 = (7.323E+03)*(1.008E-01)**2 = (7.323E+03)*(1.016E-02) = 7.438E+01
ltally 24 nps = 1318186
+ Number of Mn-56 atoms created in cell 31
tally type 4 track length estimate of particle flux.
tally for neutrons

volumes
 cell: 31
 1.31947E+00

cell 31
 multiplier bin: 9.57520E-02 4 102
 1.62252E+00 .0057

analysis of the results in the tally fluctuation chart bin (tfc) for tally 24 with nps = 1318186 print table 160

normed average tally per history = 1.62252E+00	unnormed average tally per history = 2.14087E+00
estimated tally relative error = .0057	estimated variance of the variance = .0003
relative error from zero tallies = .0028	relative error from nonzero scores = .0049
number of nonzero history tallies = 118275	efficiency for the nonzero tallies = .0897
history number of largest tally = 177710	largest unnormalized history tally = 1.02632E+03
(largest tally)/(average tally) = 4.79396E+02	(largest tally)/(avg nonzero tally) = 4.30141E+01
(confidence interval shift)/mean = .0000	shifted confidence interval center = 1.62258E+00

if the largest history score sampled so far were to occur on the next history, the tfc bin quantities would change as follows:

estimated quantities	value at nps	value at nps+1	value(nps+1)/value(nps)-1.
mean	1.62252E+00	1.62311E+00	.000363
relative error	5.65047E-03	5.66005E-03	.001696
variance of the variance	2.93010E-04	3.07477E-04	.049373
shifted center	1.62258E+00	1.62258E+00	.000001
figure of merit	1.73991E+02	1.73402E+02	-.003384

the estimated slope of the 200 largest tallies starting at 3.56847E+02 appears to be decreasing at least exponentially. the large score tail of the empirical history score probability density function appears to have no unsampled regions.

=====

results of 10 statistical checks for the estimated answer for the tally fluctuation chart (tfc) bin of tally 24

tfc bin	--mean--	-----relative error-----			----variance of the variance----			--figure of merit--		-pdf-
behavior	behavior	value	decrease	decrease rate	value	decrease	decrease rate	value	behavior	slope
desired	random	< .10	yes	1/sqrt(nps)	<0.10	yes	1/nps	constant	random	>3.00
observed	random	.01	yes	yes	.00	yes	yes	constant	random	10.00
passed?	yes	yes	yes	yes	yes	yes	yes	yes	yes	yes

=====

this tally meets the statistical criteria used to form confidence intervals: check the tally fluctuation chart to verify. the results in other bins associated with this tally may not meet these statistical criteria.

estimated asymmetric confidence interval(1,2,3 sigma): 1.6134E+00 to 1.6318E+00; 1.6042E+00 to 1.6409E+00; 1.5951E+00 to 1.6501E+00
 estimated symmetric confidence interval(1,2,3 sigma): 1.6134E+00 to 1.6317E+00; 1.6042E+00 to 1.6409E+00; 1.5950E+00 to 1.6500E+00

fom = (histories/minute)*(f(x) signal-to-noise ratio)**2 = (7.323E+03)*(1.541E-01)**2 = (7.323E+03)*(2.376E-02) = 1.740E+02
 1status of the statistical checks used to form confidence intervals for the mean for each tally bin

tally result of statistical checks for the tfc bin (the first check not passed is listed) and error magnitude check for all bins

- 4 missed 1 of 10 tfc bin checks: the slope of decrease of largest tallies is less than the minimum acceptable value of 3.0
 passed all bin error check: 1 tally bins all have relative errors less than .10 with no zero bins
- 14 passed the 10 statistical checks for the tally fluctuation chart bin result
 passed all bin error check: 1 tally bins all have relative errors less than .10 with no zero bins
- 24 passed the 10 statistical checks for the tally fluctuation chart bin result
 passed all bin error check: 1 tally bins all have relative errors less than .10 with no zero bins

the 10 statistical checks are only for the tally fluctuation chart bin and do not apply to other tally bins.

warning. 1 of the 3 tally fluctuation chart bins did not pass all 10 statistical checks.
 1tally fluctuation charts

nps	tally 4					tally 14					tally 24				
	mean	error	vov	slope	fom	mean	error	vov	slope	fom	mean	error	vov	slope	fom
128000	9.3742E-03	.1029	.1178	2.7	5.4E+00	3.1126E-01	.0271	.0059	7.5	78	1.6316E+00	.0182	.0029	8.6	174
256000	9.3517E-03	.0707	.0491	2.9	5.8E+00	3.1276E-01	.0198	.0047	9.1	74	1.6308E+00	.0129	.0018	10.0	172
384000	9.3582E-03	.0577	.0316	3.0	5.7E+00	3.1644E-01	.0160	.0028	10.0	75	1.6458E+00	.0105	.0010	10.0	174
512000	1.0401E-02	.0674	.1525	2.6	3.1E+00	3.1979E-01	.0139	.0021	10.0	74	1.6366E+00	.0091	.0008	10.0	173
640000	1.0198E-02	.0621	.1104	2.6	3.0E+00	3.2158E-01	.0125	.0016	10.0	73	1.6276E+00	.0081	.0006	8.8	173
768000	9.9297E-03	.0547	.0988	2.7	3.2E+00	3.1999E-01	.0114	.0013	10.0	73	1.6244E+00	.0074	.0005	9.0	173
896000	9.9117E-03	.0513	.0860	2.8	3.1E+00	3.1848E-01	.0105	.0011	10.0	75	1.6226E+00	.0068	.0004	7.2	175
1024000	9.7946E-03	.0467	.0770	2.9	3.3E+00	3.1743E-01	.0098	.0010	10.0	74	1.6215E+00	.0064	.0004	9.9	175
1152000	9.9681E-03	.0451	.0605	2.9	3.1E+00	3.1676E-01	.0093	.0009	10.0	74	1.6239E+00	.0060	.0003	9.6	174
1280000	1.0088E-02	.0424	.0500	3.0	3.2E+00	3.1655E-01	.0088	.0008	10.0	75	1.6211E+00	.0057	.0003	8.3	175
1318186	1.0072E-02	.0418	.0478	3.0	3.2E+00	3.1694E-01	.0086	.0008	10.0	74	1.6225E+00	.0057	.0003	10.0	174

 dump no. 13 on file runtpe nps = 1318186 coll = 180449845 ctm = 180.01 nrn = 3083182460

4 warning messages so far.

run terminated when it had used 180 minutes of computer time.

computer time = 180.35 minutes

mcnp version 4b 02/04/97

04/30/98 02:29:30

probid = 04/29/98 23:27:00

Appendix C: Three-Emitter SPNS Spreadsheet

	Cell 11	Cell 21	Cell 31		A A ⁻¹	(Check)	
	Thermal run:				1	-7.1E-15	8.882E-16
response	8.76E-02	0.024207	0.13449		0	1	0
error	0.0449	0.0419	0.0341		0	0	1
	Epithermal run:						
response	1.90E-01	5.18E-02	0.28388				
error	0.0374	0.0299	0.0259				
	Fast run:						
response	5.17E-01	1.32E-01	7.30E-01				
error	0.0205	0.0168	0.0149	<--(relative or fractional error)			
Response Matrix (A)				Normalized Response matrix:			
8.76E-02	2.42E-02	1.34E-01		0.539675	0.149133	0.828559	
1.90E-01	5.18E-02	2.84E-01		0.549793	0.149845	0.821751	
5.17E-01	1.32E-01	7.30E-01		0.572117	0.145863	0.807097	
Error Matrix (dA)				Determinant of Normalized			
0.003933	0.001014	0.004586		Response Matrix:			
0.007103	0.001548	0.007352		-6.7E-05			
0.010603	0.002215	0.010872					
Calibration Matrix (A ⁻¹)							
-98.4354	-21.1123	26.35762					
-2415	1653.082	-198.017					
506.2285	-283.789	18.4735					
Intermediate Calculation (dA A ⁻¹)							
-0.51501	0.292152	-0.01245					
-0.71505	0.322074	0.016569					
-0.89018	0.353025	0.04163					
Intermediate Calculation (A ⁻¹ dA A ⁻¹)							
42.32878	-26.2529	1.973232					
237.989	-243.038	49.21848					
-74.2358	63.01654	-10.2367					
Associated Errors for Inverted Matrix (dA ⁻¹)							
-42.3288	26.25292	-1.97323					
-237.989	243.0377	-49.2185					
74.23579	-63.0165	10.23673					

Appendix D: Cadmium Shielded SPNS Spreadsheet

	Cell 11	Cell 21	Cell 31		A A ⁻¹	(Check)	
	Thermal run:				1	0	0
response	2.30E-01	1.71646	6.44764		8.88E-16	1	1.137E-13
error	0.0051	0.0028	0.0021		1.39E-17	1.11E-16	1
	Epithermal run:						
response	1.30E-01	6.71E-01	1.75473				
error	0.0089	0.0054	0.0048				
	Fast run:						
response	5.00E-03	1.99E-02	3.45E-02				
error	0.0073	0.005	0.0045				
Response Matrix (A)				Normalized Response matrix:			
2.30E-01	1.72E+00	6.45E+00		0.034411	0.257103	0.965771	
1.30E-01	6.71E-01	1.75E+00		0.069262	0.35647	0.931736	
5.00E-03	1.99E-02	3.45E-02		0.124608	0.496295	0.859165	
Error Matrix (dA)				Determinant of Normalized			
0.001172	0.004806	0.01354		Response Matrix:			
0.001161	0.003625	0.008423		-0.00052			
3.65E-05	9.96E-05	0.000155					
				Typical combined response vector			
Calibration Matrix (A ⁻¹)				(Lazy Susan ratios:)			
44.67875	-262.111	4985.044		0.191593	1.41382	5.28205	
-16.1932	92.07714	-1658.06		Associated errors (%):			
2.874042	-15.1731	263.7803		0.002	0.0011	0.0009	
				Experimental Error:			
Intermediate Calculation (dA A ⁻¹)				0.000383	0.001555	0.004754	
0.013436	-0.07001	1.443506					
0.017372	-0.09829	1.998171		Calculated incidence flux			
0.000464	-0.00275	0.057782		0.84667	-0.18339	4.201185	
Intermediate Calculation (A ⁻¹ dA A ⁻¹)				Associated derived errors:			
-1.63862	8.911983	-171.201		0.094497	-0.019	0.432891	
0.612182	-3.35222	64.80419					
-0.1025	0.564012	-10.928					
Associated Errors for Inverted Matrix (dA ⁻¹)							
1.63862	-8.91198	171.2012					
-0.61218	3.352215	-64.8042					
0.102503	-0.56401	10.92796					