MINIATURIZED LASER ALTIMETER FOR SMALL SATELLITE APPLICATIONS

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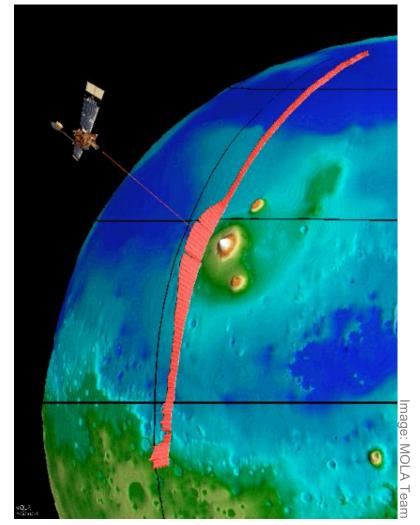
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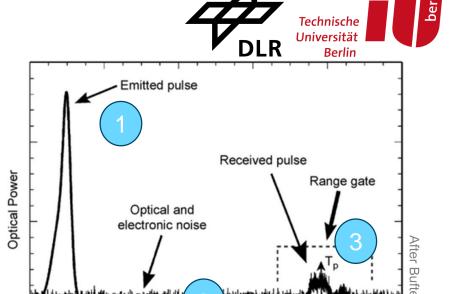


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Laser Altimeter Working Principle

- With a laser altimeter, one can measure the distance/altitude above an object or surface.
- The instrument emits a short laser pulse, detects the reflected pulse and calculates the time of flight resp. the distance.
- A laser altimeter consist of three main elements:
 - Transmitter: pulsed laser
 - Receiver: telescope, optics and detector
 - Electronics for analysis of the reflected pulse and instrument control
 - $N_{\rm ph}$ number of photons in the reflected pulse
 - E_T is energy of the emitted laser pulse,
 - ξ_T and ξ_R are the losses in the transmitter and receiver optics,
 - a is the fraction of emitted power reflected at the surface (the albedo),



Link budget equation:

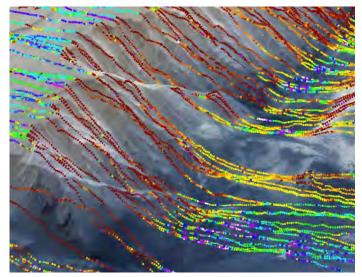
$$N_{\rm ph} = \frac{E_T \xi_T \xi_R a d_T^2 T_A^2 Q_e}{4R^2 h \lambda}$$

Travel Time $\Delta t_1 = 2z/c$

- d²_T is the diameter of the receiver optics,
- T_A is the fractional loss of optical power through attenuation in the atmosphere (if present),
- Q_e is the quantum efficiency of the sensor,
- R is the range to target,
- \(\lambda\) is wavelength and
- h is the Planck constant.

Laser Altimetry in Planetary Research

- Laser altimetry is used in planetary research for
 - Topography: craters, ridges, slopes, roughness, reflectance etc.
 - Surface deformation (tidal amplitudes)
 - Global reference model
- Laser altimeter data is combined with data from other instruments like cameras, spectrometers, magnetometer, gravimeter etc.



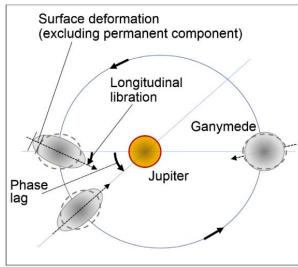
MOLA Team/DLR

Laser altimeter data (from MOLA) superposed with camera data (from HRSC)

Color codes: green & blue: small pulse

spread; red: large pulse spread





Enya et all, 2022

Surface deformation, phase lag and longitudinal libration

Laser altimetry helps to understand the evolution and the current state of a planetary body

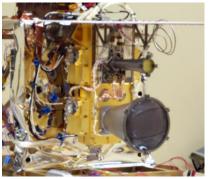
Classical Laser Altimeters



Numerous missions in our Solar System are equipped with laser altimeters

MOLA on Mars Orbiter LOLA on Lunar Orbiter MLA on MESSENGER





MOLA: NASA/JPL; LOLA: NASA MLA: Ott et al., 2008;

BELA on BepiColombo (to Mercury, launched in 2018)

with strong DLR contribution

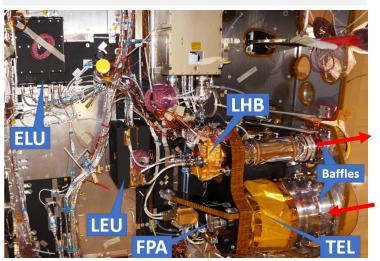
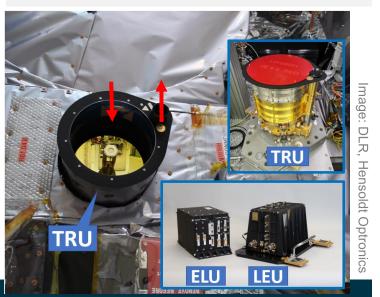


Image: DLR

GALA on JUICE (to Ganymede, launched in 2023)

with DLR as consortium lead



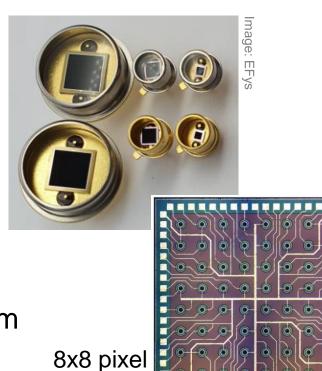
■ These laser altimeters need a lot of resources on the spacecraft: Mass (BELA: 15 kg, GALA: 23 kg) Power (30 W resp. 50 W)

Volume (100 l resp. 60 l)

Single-Photon Counting Laser Altimeters (1/3)



- The previously mentioned laser altimeters work with detectors which need a return signal consisting of a <u>few hundreds</u> photons at least.
- → Use a more sensitive detector, e. g. a <u>single</u> photon counting detector.
- Several single photon counting techniques do exist, we will focus on single-photon avalanche diodes (SPAD).
- The output signal of a SPAD is not an analogue waveform anymore, but already digital TTL-like (high/low).



SPAD array

 A new detector technology requires less strong return signal - and less strong transmitter laser pulses

Single-Photon Counting Laser Altimeters (2/3)



The main assemblies:

- Microchip laser or diode laser
- Combined transmitter-receiver optics
- Single-photon counting detector array (multi-pixel)
- Electronic: one FPGA for data processing, instrument control, data transfer

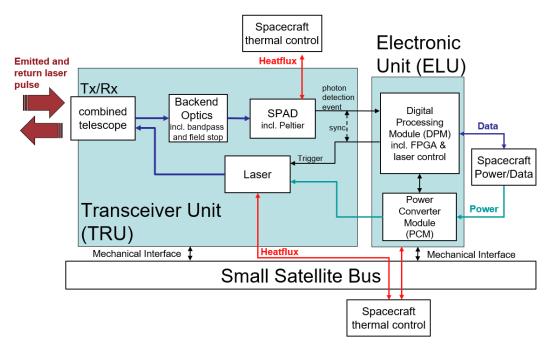


Image: DLR

Every part of the instrument can be miniaturized due to the more sensitive detector

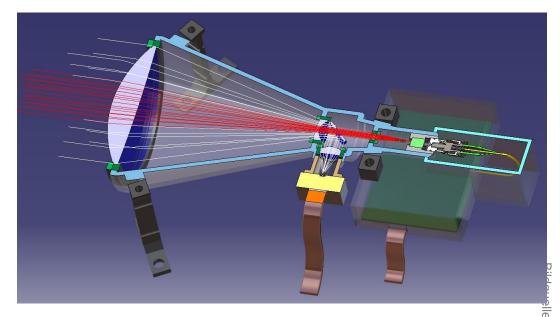
Single-Photon Counting Laser Altimeters (3/3)



For a potential mission to Moon, basic instrument parameters are set to...

- Orbit altitude: 70 km (max)
- Laser pulse energy: 2 mJ
- Laser pulse repetition rate: 30 to 50 Hz
- Telescope diameter: 8 cm
- Detector: 2x2 pixel array

... and a design and resource numbers are derived.



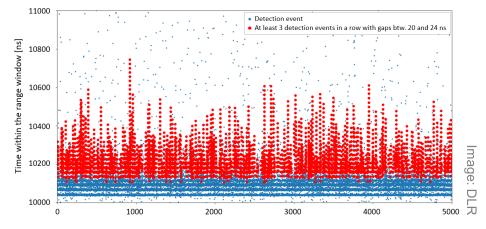
ParameterGALASPC LAMass23 kg5 kgPower50 W20 WVolume60 l5 l

The instrument design is more compact and much less resources are needed w.r.t to classical laser altimeters.

Experimental Results and Science Performance



- The instrument development is parallel on hardware and simulations. New findings are immediately verified and tested in the lab.
- In the lab, a SPC detector is fed with very weak laser pulses, e. g. generating less than 5 signal photons or detection events.
- Mathematical simulations concerning quantum efficiency, detection probabilities, time-of-flight measurement and detection schemes are verified.



A pulse of 80 ns length and approx. 5 signal photons is detected three times (blue lines) by a detector with 25 ns death

time.

ATTENUATION
FIBERS

OSCILLISCOPE

ENERGY
METER

ATTENUATION
FIBERS

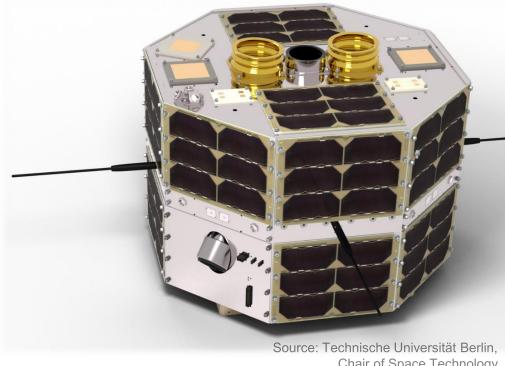
OPTICAL
SETUP

 We demonstrate the working principle of the single-photon counting laser altimeter and validate the relevant performance parameters.

Microsatellite Mission Studies (1/2) **Low Earth Orbit**



- In-Orbit Demonstration (IOD) utilizing TUBiX20 microsatellite platform
 - Highly modular, easy adaption to mission parameter
 - 5 to 25 kg payload mass
 - 25 to 60 W continuous payload power supply
- Flight proven platform
 - Low cost, low risk
 - high pointing accuracy of 0.1° (3-sigma)



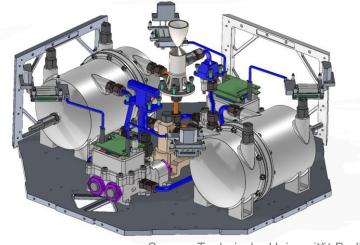
Chair of Space Technology

- Low cost, low risk, flight heritage
- Less favorable use case
 - → high noise outside eclipse, high altitude

Microsatellite Mission Studies (2/2) Lunar Orbit



- Much more favourable for instrument demonstration
 - Less signal noise & lower altitude
 - → better return signal and temporal & spatial coverage
- Utilizing ongoing TUBiX20 developments
 - Propulsion system and non-magnetic attitude control
 - Radiation hard (COTS) electronics
 - Higher ressource budgets for payloads
 - → Qualify TUBiX20 for missions beyond LEO



Source: Technische Universität Berlin, Chair of Space Technology

- Additional challenges to overcome for microsatellites
 - > Lunar orbit determination and control
 - → Communication with Earth

Future Work



- Improve design and simulations of the instrument
- Select the most relevant parts (laser, detector, FPGA, optics etc.)
- Establish representative (elegant) breadboards for spacequalification on ground
- Develop TUBiX20 platform for use beyond LEO
- Improve autonomous platform operation

Our goal is the space-qualification of the instrument and the platform.

Summary



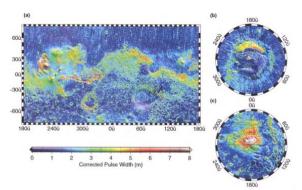
- Based on our heritage in planetary laser altimetry and small satellite systems, we develop a new instrument and satellite generation
- For the instrument, first resource budgets are available and science performance is verified by lab tests
- Combined with a small satellite platform with improved designs for propulsion system and attitude control, a technology demonstrator mission to the Moon is the obvious next step
- Lunar IOD of instrument and platform is more challenging but much more useful

- Single-photon counting laser altimetry is the next generation of laser altimeters in space
- A small satellite platform is the perfect fit for a such an miniaturized instrument



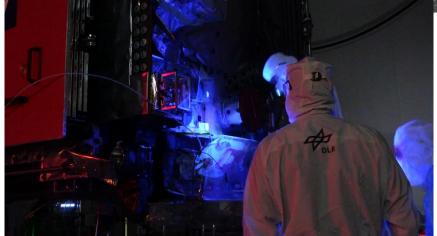
Thank you for your attention!







→ Smith et al. (1999)



Kay Lingenauber, Institut für Planetenforschung, 20.09.2023

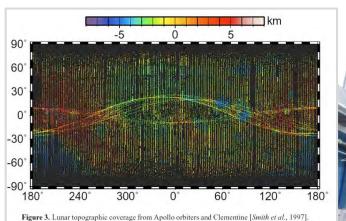


Figure 3. Lunar topographic coverage from Apollo orbiters and Clementine [Smith et al., 1997].

