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Unexpected ^{13}N concentrations in ISIS synchrotron room air

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Abstract

The specific activity of air in the large open room housing the 800-MeV proton synchrotron of the ISIS Spallation Neutron and Muon Source has been measured. Air from several positions within the ISIS synchrotron room was sucked through a long flexible tube, and run past a shielded HPGe gamma-ray detector outside the synchrotron room. In spite of an expectation that ^{13}N should be the largest component of the overall activity in the air, the results of the measurements are consistent with the presence in the air of ^{11}C and ^{41}Ar only, and suggest that the activity in the air is mostly created not in the synchrotron room itself but in the massive shielding monoliths around the neutron-producing targets, monoliths through which ventilation air is drawn into the synchrotron room. Typical specific activities of ^{11}C and ^{41}Ar in the air in the synchrotron room are ~ 0.10 and ~ 0.03 Bq cm^{-3} respectively, the upper limit for ^{13}N being at most ~ 0.01 Bq cm^{-3} .

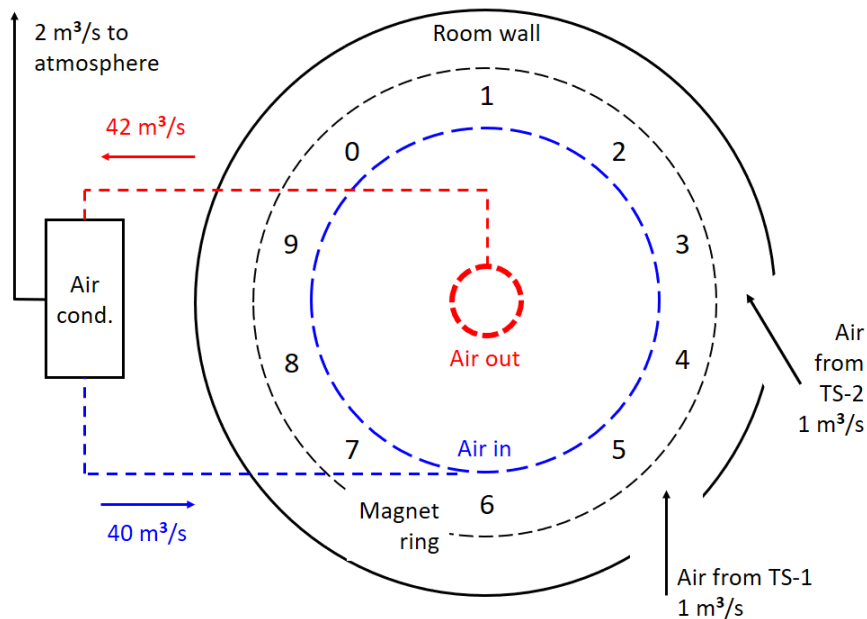
1. Introduction

The ISIS Spallation Neutron and Muon Source [1, 2] is driven by an 800-MeV $\sim 250\text{-}\mu\text{A}$ proton synchrotron running at 50 pulses per second (pps), the synchrotron in turn being driven by a 70-MeV H^- drift tube linac. The proton beam from the synchrotron is split and delivered at 40 pps to Target Station 1 (TS-1) and at 10 pps to Target Station 2 (TS-2). Accelerator, proton beam line and target areas are ventilated to avoid build-up of toxic and corrosive gaseous products, to remove heat, to maintain machine areas slightly below atmospheric pressure, and to facilitate the entry of personnel into machine areas after the beam has been switched off.

At ISIS, the synchrotron is located in an open room 61 metres in diameter and 9 metres in mean height, and this large room acts as a ~ 3 -hour delay/decay tank. Air is drawn into the $\sim 25000\text{-m}^3$ synchrotron room from ventilated areas around the neutron-producing targets and the proton beam lines through the two ~ 100 -metres-long tunnels for the extracted proton beam lines, and air is extracted from the synchrotron room at a rate of $2\text{ m}^3\text{ s}^{-1}$ and discharged to atmosphere (authorised by the UK Environment Agency). A schematic diagram is shown as Fig. 1.

Whilst previous calculations [3] (discussed in Sect. 4 below) of specific activities of air in the ISIS synchrotron room indicated that ^{13}N was likely to be the largest component of the overall activity, measurements made over several years using equipment for monitoring gaseous discharges have suggested that there is significantly more ^{11}C activity than ^{13}N activity in the air [4]. The work reported in this paper was carried out both to measure directly the specific activity of the air in the synchrotron room and to

42 try to resolve the question of the relative contributions of ^{11}C and ^{13}N activity. This
43 paper deals only with contributions to activity in the air from ^{11}C , ^{13}N , ^{15}O and ^{41}Ar .



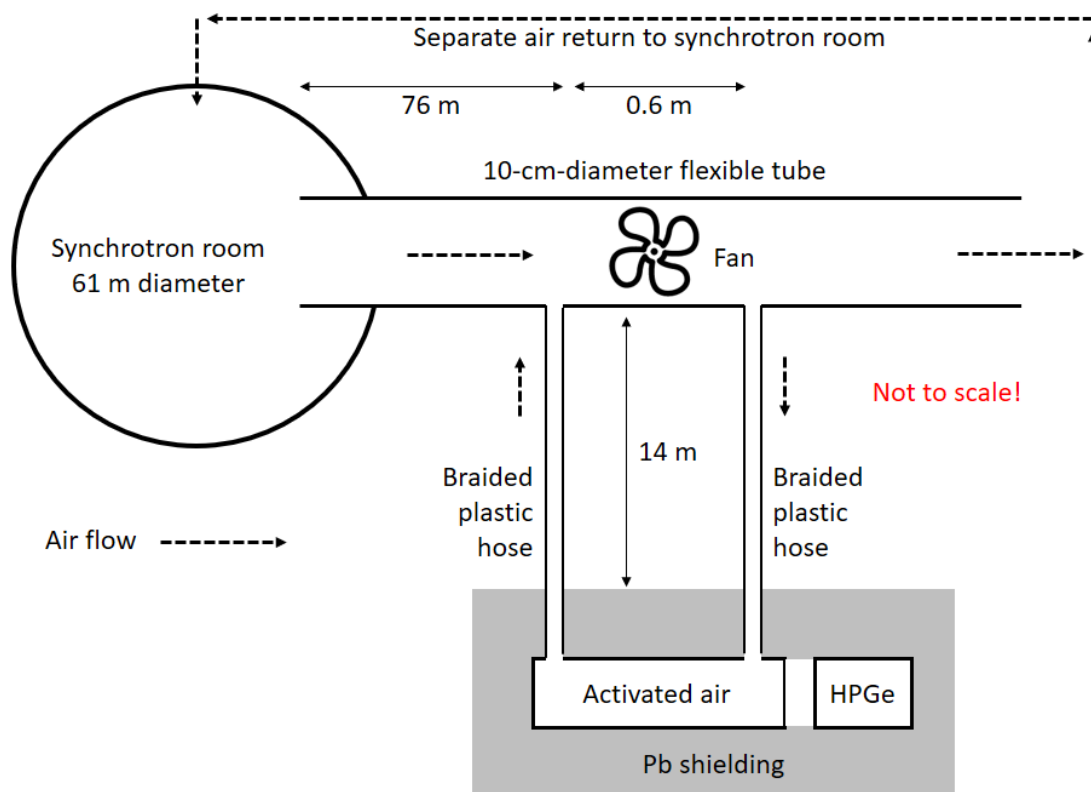
44

45 Fig. 1. Schematic diagram of air flows into and out of the ISIS synchrotron room. Air from
46 the full-flow-filtered air-conditioning system enters at floor level, and leaves at ceiling level.
47 The numbers 0–9 denote the positions of the ten superperiods (SPs) of the synchrotron
48 (injection takes place in SP0, and collimation and extraction take place in SP1). The air from
49 Target Stations TS-1 and TS-2 is essentially ventilation air from the shutter voids in the several-
50 thousands-of-tons monoliths of steel and concrete shielding around the neutron-producing
51 targets. In order to be able to switch on and off beams of neutrons to the research instruments
52 in the experimental halls independently, massive ~20-ton movable steel shutters are
53 incorporated in the monoliths, and the shutter voids are the voids in the steel monoliths
54 necessary to accommodate the movement of the shutters. Air is discharged to atmosphere at a
55 nominal rate of $2 \text{ m}^3 \text{ s}^{-1}$ through a stack on the roof of the air-conditioning building.

56 2. Measurements

57 Since the synchrotron and its surroundings become radioactive in use, it is not practical
58 to make measurements of activated air within the synchrotron room simply by using a
59 health physics monitor, since background from activated machine components and
60 support structures is always present. Consequently, a method was developed whereby
61 air from the synchrotron room is sucked out along a 76-metre length of 10-cm-diameter
62 flexible PVC air ducting hose, a fraction of this air is then circulated through a thin-
63 walled aluminium cylindrical vessel with an internal volume and diameter of 859 cm^3
64 and 9.3 cm respectively immediately in front of a shielded Canberra BE3825 HPGe
65 gamma-ray detector, and the air is then returned to the synchrotron room. The time for
66 activated air to travel from the input end of the long flexible tube to the HPGe was
67 measured using a smoke generator and found to be 69 ± 5 seconds [5]. A schematic
68 diagram is shown in Fig. 2.

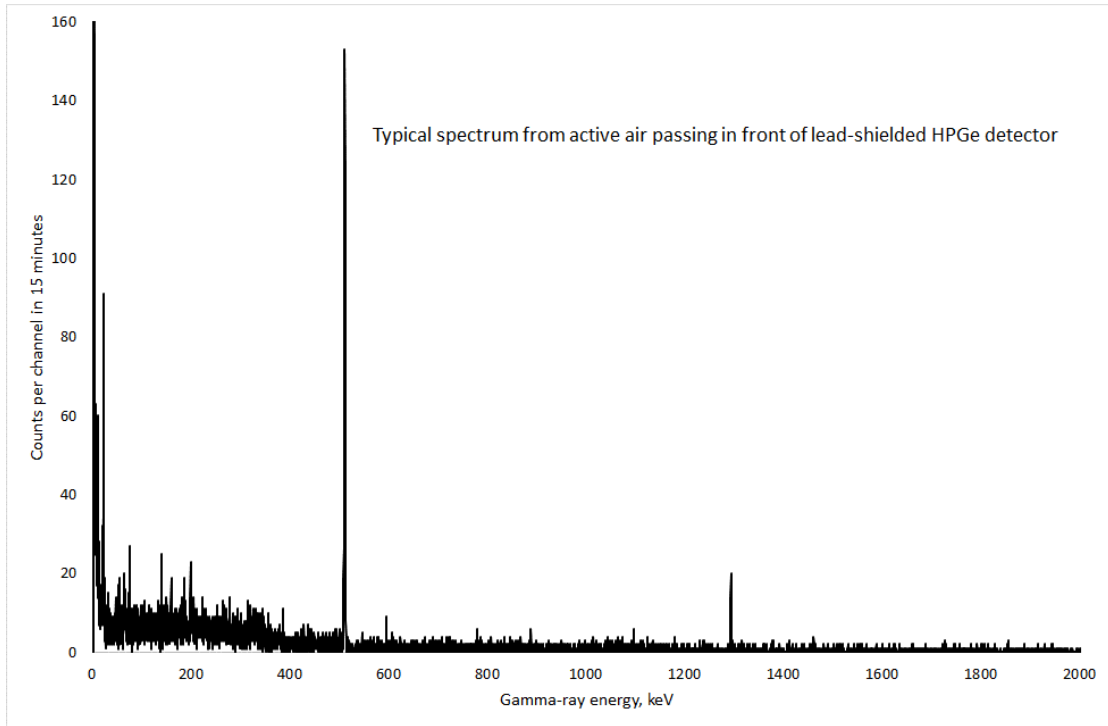
69 Successive 15-minute-long gamma-ray spectra¹ were recorded and stored over periods
 70 of two or three days — each period usually beginning just before the end of an
 71 irradiation campaign ('user cycle') and ending after a further ~1–2 days of accelerator
 72 physics work. A typical gamma-ray spectrum when the synchrotron was running
 73 steadily is shown in Fig. 3, in which the only gamma-ray lines of immediate
 74 significance are at 511 keV (from positron emitters such as ¹¹C, ¹³N and ¹⁵O) and at
 75 1294 keV (from ⁴¹Ar from neutron capture on the ~1% of argon naturally present in
 76 air). The beginnings and ends of counting periods were taken from the corresponding
 77 times recorded within the spectrum files themselves (correcting for two changes of local
 78 time from GMT (Greenwich Mean Time) to BST (British Summer Time)). Fig. 4
 79 shows a representative set of data collected: counts per 15 minutes in 511- and 1294-
 80 keV peaks, and beam current from the synchrotron measured at 4-minute intervals.



81

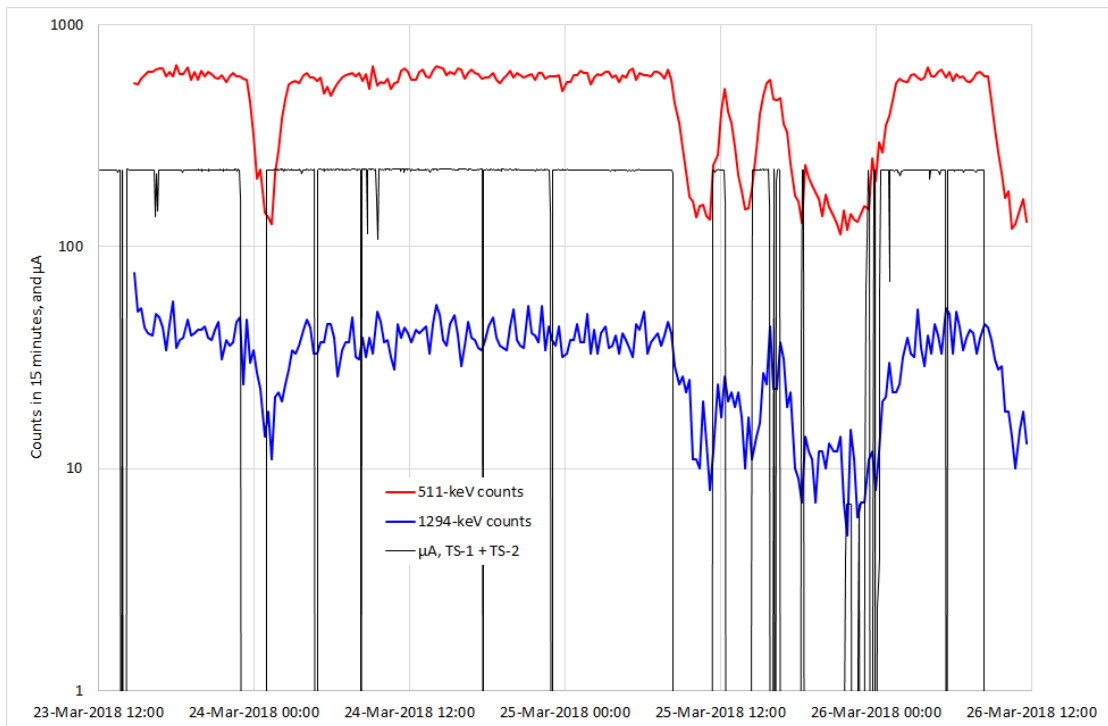
82 Fig. 2. Schematic representation of air from the ISIS synchrotron room being passed in front
 83 of the HPGe gamma-ray detector. Not to scale! The internal diameter and length of the
 84 activated air sample vessel in front of the HPGe detector were 9.3 and 12.7 cm respectively.
 85 The distance between the end of the activated air cylinder and the HPGe crystal was 1.0 ± 0.5
 86 cm. The activated air is moved through the activated air sample vessel in front of the HPGe
 87 detector by the pressure difference across the fan.

¹ The spectra for the December 2019 and February 2020 sets of data were each 30 minutes long, not 15 minutes long.



88

89 Fig. 3. Typical gamma-ray energy spectrum seen by the HPGe detector. The lines at 511 keV
 90 from positron-emitting radionuclides such as ^{11}C , ^{13}N and ^{15}O , and at 1294 keV from ^{41}Ar stand
 91 out very clearly.



92

93 Fig. 4. Representative data set covering three days: counts in 15 minutes in 511- and 1294-keV
 94 peaks in gamma-ray spectra, and the synchrotron proton beam current (which is also the sum
 95 of beam currents delivered to target stations TS-1 and TS-2).
 96

97 3. Analyses

98 3.1 Model

99 In order to extract information from sets of data such as the set illustrated in Fig. 4, the
100 following model was adopted. Assume atoms of radionuclides i are produced by the
101 proton beam at rates $r_i I$ where I is the proton beam current and r_i is a constant of
102 proportionality (atoms $\mu\text{A}^{-1} \text{s}^{-1}$). Suppose activated air takes time δ_s to get from
103 wherever it is produced to the synchrotron room. Then in time dt the number of atoms
104 of radionuclide i appearing at time t in the synchrotron room is

$$105 \quad dN_i = \exp(-\lambda_i \delta_s) r_i I(t - \delta_s) dt$$

106 where $\lambda_i = \ln(2)/t_{1/2,i}$ is the decay constant for radionuclide i and $t_{1/2,i}$ is its half-life.
107 Suppose there are N_i atoms of radionuclide i in the synchrotron room at time t , and
108 suppose the air exchange rate and synchrotron room volume are v ($\text{m}^3 \text{s}^{-1}$) and V (m^3)
109 respectively. Then, in time dt , $\lambda_i N_i dt$ atoms decay, and $(v/V) N_i dt$ atoms are
110 removed. So the net change in number of atoms of radionuclide i in the synchrotron
111 room is $dN_i = \exp(-\lambda_i \delta_s) r_i I(t - \delta_s) dt - \lambda_i N_i dt - (v/V) N_i dt$, which may be
112 re-written as

$$113 \quad dN_i = \exp(-\lambda_i \delta_s) r_i I(t - \delta_s) dt - \lambda'_i N_i dt$$

114 where $\lambda'_i = \lambda_i + v/V$. Since $I = I(t)$ is a known function of time ($I(t)$ is measured
115 and recorded in 4-minute steps), $N_i = N_i(t)$ may be obtained by numerical integration,
116 and thereby activities in the synchrotron room $\lambda_i N_i = \lambda_i N_i(t)$ may be obtained.

117 Suppose it takes time δ_c for the activated air to get from the synchrotron room to the
118 HPGe detector, and assume that the gamma-ray emission probability for radionuclide i
119 is α_i and that the HPGe detection efficiency for gamma-rays from radionuclide i is ε_i .
120 Then the number of HPGe counts $c_{i,j}$ from radionuclide i over a counting interval Δ
121 beginning at time t_j is

$$122 \quad c_{i,j} = c_i(t_j + \Delta) = \alpha_i \varepsilon_i \int_{t_j}^{t_j + \Delta} \exp(-\lambda_i \delta_c) \lambda_i N_i(t - \delta_c) dt.$$

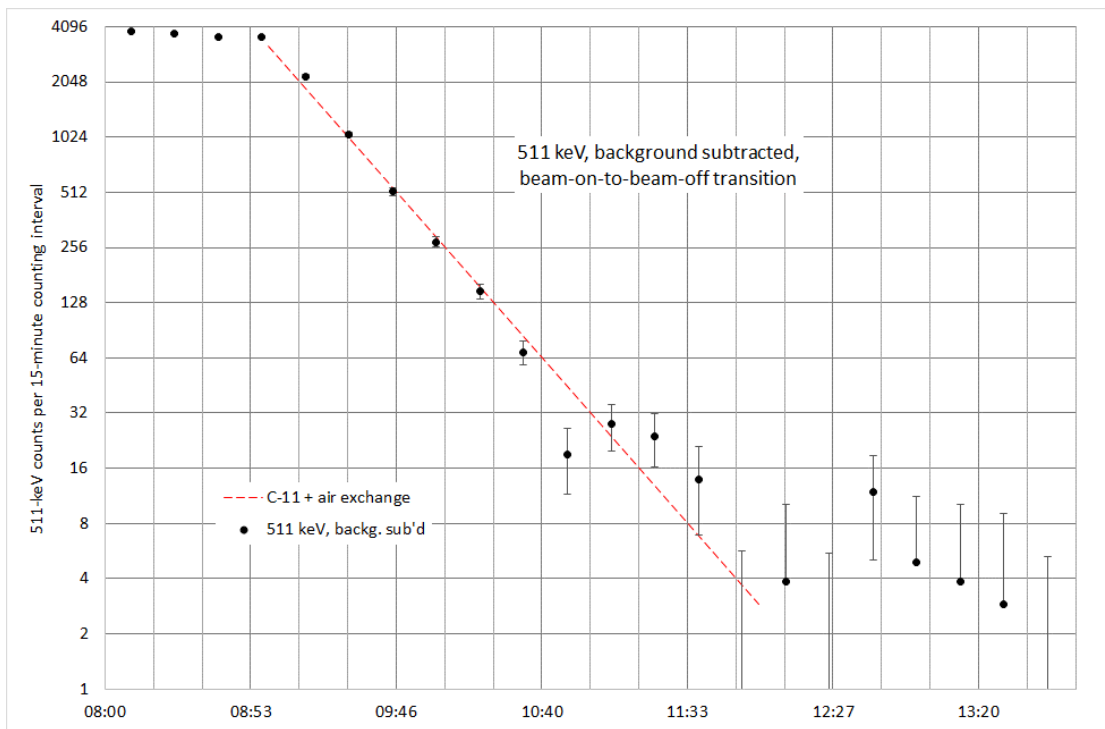
123 By fitting to the data by minimising the chi-squared per degree of freedom
124 $\chi_{\text{pdf}}^2 = \{\sum_i (\sum_j (C_{i,j} - B_i - c_{i,j})^2 / \delta C_{i,j}^2)\} / \{mn - (2m + 2)\}$ where m is the number of
125 radionuclides, parameters r_i , δ_s , δ_c and B_i may be extracted, where i runs from 1 up to
126 m , the C_i 's are the measured counts corresponding to radionuclide i , B_i is the
127 corresponding background, and j runs from 1 up to n where n is the number of counting
128 intervals.

129 If the passage of activated air from the source of production to the synchrotron room is
130 characterised not by a unique time but by a symmetrical distribution of times spanning
131 a finite range $2w$ described by a normalised function $s(\varepsilon) = s(\bar{t}, \varepsilon)$ where \bar{t} is the mean
132 time and ε is the deviation from the mean time (as will be seen to be the case in
133 Sect. 3.3), the resultant number of atoms $N_i'(t)$ of radionuclide i may be obtained from
134 $N_i'(t) = \int_{-w}^w N_i(t - \varepsilon) s(\varepsilon) d\varepsilon$.

135 3.2 Half-life of 511-keV component

136 It is clear that the half-life of the 1294-keV component should correspond to the
137 110-minute half-life of ^{41}Ar , but the 511-keV component could be due to some or all
138 of ^{11}C , ^{13}N and ^{15}O with half-lives of 20.3, 9.96 and 2.03 minutes respectively, since

139 these three radionuclides can all be produced by nuclear reactions in the air and they
 140 emit only positrons when they decay. Fig. 5 shows the 511-keV component for a
 141 typical beam-on-to-beam-off transition. Also shown is the rate at which ^{11}C alone
 142 would decay, taking into account the air extraction rate from the synchrotron room, *i.e.*
 143 the dashed line is proportional to $\exp(-\lambda'(t - t_0))$ where $\lambda' = \lambda_{^{11}\text{C}} + v/V$ is the
 144 effective ^{11}C decay constant, $\lambda_{^{11}\text{C}} = \ln(2)/t_{1/2,^{11}\text{C}}$ where $t_{1/2,^{11}\text{C}} = 20.3$ minutes, v
 145 and V are as set out in Sect. 3.1 above and have values of 2 m s^{-1} and 25000 m^3
 146 respectively, and t_0 is the beam-off time. Since the effective half-life $t'_{1/2} =$
 147 $\ln(2)/\lambda' = 17.8$ minutes, it seems clear that the measured data support the conclusion
 148 that the 511-keV activity is mostly due to ^{11}C alone.



149

150 Fig. 5. 511-keV component of counts in HPGe detector after background subtraction for a
 151 typical beam-on-to-beam-off transition, plotted on a base-2 logarithmic scale with 17.8-minute
 152 intervals (the effective ^{11}C half-life taking into account air exchange in the synchrotron room)
 153 marked on the horizontal axis. The dashed line is a decay proportional to $2^{-t/(17.8 \text{ minutes})}$. It
 154 is evident that the decay corresponds almost entirely to the decay of ^{11}C .

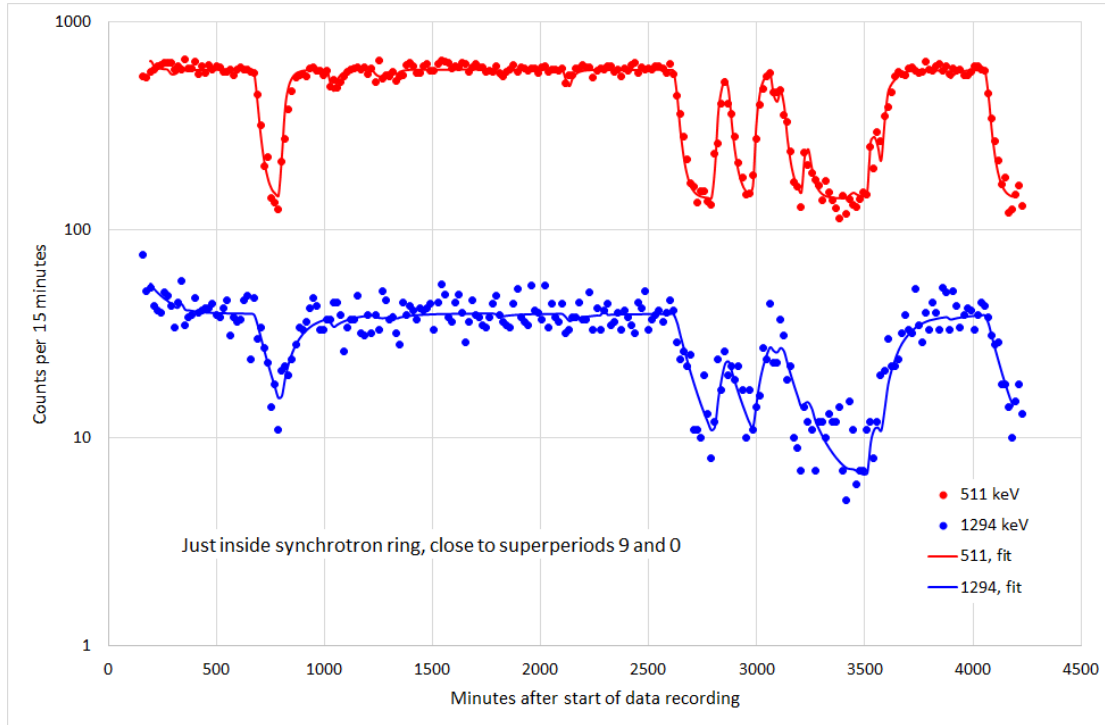
155 This conclusion, *viz.* that the 511-keV activity is mostly due to ^{11}C alone, is at variance
 156 with the results of Monte Carlo calculations [3] of saturation specific activities of air in
 157 the synchrotron room using MCNPX and CINDER-90 [6, 7] that were referred to in
 158 Sect. 1. Whilst these preliminary calculations involved some significant assumptions
 159 and approximations, the results suggested that saturation activities of ^{11}C , ^{13}N , ^{15}O and
 160 ^{41}Ar in the synchrotron room air should be in ratios of roughly
 161 $8 \pm 5 : 72 \pm 5 : 2 \pm 5 : 18 \pm 5$. How can the difference between these calculated results and
 162 the results of the present measurement be best explained? Consideration of this
 163 question is deferred until Sect. 3.5 below.
 164

165 **3.3 Delay times**

166 With the assumption that that the 511-keV component seen by the HPGe detector is
167 due to ^{11}C alone (so that in Sect. 3.1 $m = 2$), eight data sets (the one shown in Fig. 4
168 and seven others) were fitted using the model set out in Sect. 3.1. A typical fit is shown
169 in Fig. 6, and results are set out in Table 1. However, in view of the possibility that the
170 511-keV component does in fact contain a significant contribution from ^{13}N , Table 1
171 also lists results assuming that the 511-keV component seen by the HPGe detector is
172 due to ^{13}N alone. Whilst some of the fits in Table 1 are undoubtedly poor, the
173 explanation being that the fits include periods of time when the synchrotron was not
174 running normally but was being used for accelerator physics purposes when beam
175 delivery and beam conditions are often well outside normal operational envelopes, it is
176 evident that the fits are on average better assuming that the 511-keV component is due
177 to ^{11}C alone than assuming it to be due to ^{13}N alone, and that the overall time from
178 source to detector ranges between roughly 12 and 20 minutes depending on location
179 within the synchrotron room.

180 With what can these times be compared? At ISIS the most likely places for air to be
181 activated are the shutter voids in the massive shielding monoliths surrounding the
182 neutron-producing targets (see caption to Fig. 1) where high fluxes of neutrons pass
183 through the ventilating air. Elsewhere, little activity in air is likely to be produced, as
184 beam losses around the synchrotron and along the proton beam transport lines are low
185 [8] — confirmation being that when beam losses around the synchrotron were
186 deliberately doubled no increase in activation was observed. Using a smoke generator,
187 the mean time for air to be moved by the ventilation systems from the shutter voids to
188 the synchrotron room was measured [9] as 8 minutes, with a spread of ± 4 minutes. This
189 mean time of 8 minutes, with two additions, *viz* the few minutes for air to be moved by
190 the air-conditioning system within the synchrotron room from the point where the air
191 from the shutter voids enters the synchrotron room to the point where the air is sampled,
192 and the ~ 1 minute for air to travel from the sampling point to the HPGe gamma-ray
193 detector, is in satisfactory agreement with the ~ 12 – 20 minutes obtained from the fits.

194



195

196 Fig. 6. Fits to 511- and 1294-keV data measured during 23–26 March 2018 according to the
 197 model described in Sect. 3.1. In this figure (Fig. 6) the data points are the $C_{i,j}$'s in Sect. 3.1,
 198 and the fitted lines are the $c_{i,j}$'s. Values of fitted parameters are listed in Table 1.

Date of measurement	Location from which air being sampled is drawn	511 keV = ^{11}C		511 keV = ^{13}N	
		$\delta_s + \delta_c$, minutes	χ_{pdf}^2	$\delta_s + \delta_c$, minutes	χ_{pdf}^2
Feb. 2018	Foil change area, SP0	15.1 ± 0.3	3.99	22.6 ± 8.01	6.48
Mar. 2018	SP0/9 datum points	18.2 ± 2.1	2.05	26.6 ± 12.4	3.46
May 2018	SP9 datum and SP6 bridge	16.8 ± 1.0	7.49	23.6 ± 1.5	11.4
Jul. 2018	Inner side of SP5/6	15.0 ± 0.6	19.2	15.2 ± 0.5	21.2
Oct. 2018	Inner side of SP5/6	14.0 ± 0.5	31.5	14.0 ± 0.9	21.8
Dec. 2018	Inner side of SP5/6	14.2 ± 0.4	20.1	14.8 ± 0.4	30.6
Apr. 2019	Inner side of SP5/6	15.0 ± 1.4	23.7	17.1 ± 2.3	34.2
Dec. 2019	Inner side of SP5/6	12.5 ± 0.7	87.3	14.8 ± 8.6	93.6
Feb. 2020	[Data timings unavailable]				

199 Table 1. Sums of source-to-synchrotron-room and synchrotron-room-to-HPGe delay times
 200 from fitting eight sets of data (the individual delay times δ_s and δ_c are essentially completely
 201 anti-correlated) for two ‘extreme’ assumptions, *viz* assuming that the 511-keV activity is due
 202 entirely to ^{11}C , and assuming that the 511-keV activity is due entirely to ^{13}N . The fits include
 203 incorporation of a rectangular distribution s as described in Sect. 3.1 spanning a range of
 204 ± 4 minutes representing the spread in the nominal time δ_s for activated air to travel from the
 205 production source to the synchrotron room. The uncertainties quoted were obtained by
 206 perturbing one hundred times the counts $C_{i,j}$ per 15-minute counting period (see Sect. 3.1) by
 207 amounts chosen from random gaussian distributions with standard deviations $\delta C_{i,j}$ and then
 208 taking the standard deviations of the one hundred perturbed values of $\delta_s + \delta_c$, and finally, in
 209 accordance with ‘external consistency’, multiplying by χ_{pdf}^2 ^{1/2}.
 210

211 **3.4 Specific activity of air**

212 For both 511- and 1294-keV activity, the specific activity a of the air (Bq cm^{-3})
 213 sampled from the synchrotron room was calculated from

214
$$a = 2^{(\delta_c/t_{1/2})} \dot{c} / (\alpha \varepsilon V_{\text{vessel}})$$

215 where δ_c is the synchrotron-room-to-HPGe-detector delay time, $t_{1/2}$ is the half-life of
 216 the corresponding radionuclide, \dot{c} is the count rate of the 511- or 1294-keV component
 217 of the gamma-ray spectrum when the synchrotron is running steadily (*e.g.* the count
 218 rate corresponding to the ‘high flat’ parts of the 15-minute-counts data shown in Fig. 4),
 219 α is the emission probability (‘abundance’) of the gamma-rays being counted, and ε is
 220 the full-energy-peak efficiency for counting gamma-rays from the activated air vessel
 221 of volume V_{vessel} just in front of the HPGe detector. The full-energy-peak efficiency
 222 for the HPGe detector at 511 keV was calculated as $\varepsilon_{511} = 0.0122 \pm 0.0024$ by
 223 integrating the point-source full-energy-peak efficiency of the detector over the surface
 224 area of the inside of the activated air vessel using MORSE [10] with the DLC37F library
 225 [11], and the full-energy-peak efficiency at 1294 keV was integrated in a similar way
 226 over the volume of the vessel to give $\varepsilon_{1294} = 0.00388 \pm 0.00078$ (the range in air of the
 227 385-keV-mean-energy positrons from ^{11}C , ~ 100 cm, is much greater than the
 228 characteristic dimension of the vessel, and so most positrons annihilate on the inner
 229 surface of the vessel). In view of the conclusion of Sect. 3.1 and anticipating the
 230 conclusion of Sect. 3.5 that the 511-keV component is largely due to ^{11}C , from the half-
 231 lives of ^{11}C and ^{41}Ar of 20.3 and 110 minutes respectively, and from the gamma-ray
 232 emission probabilities per becquerel for ^{11}C and ^{41}Ar of 1.9952 and 0.9916 respectively,
 233 results for the specific activity a are presented in Table 2. Representative values for
 234 the specific activities of ^{11}C and ^{41}Ar in the air in the synchrotron room are ~ 0.10 and
 235 ~ 0.03 Bq cm^{-3} respectively.

236 **3.5 Search for ^{13}N component of 511-keV activity**

237 The extent to which the 511-keV activity is due to ^{11}C was investigated by
 238 simultaneously fitting the 511- and 1294-keV counts $C_{511,i}$ and $C_{1294,i}$ in the 15-minute
 239 bins i during beam-on-to-beam-off transitions with the expressions

240
$$C'_{511,i} = C'_{13\text{N},i} + C'_{11\text{C},i} + b_{511}(t_i - t_{i-1}) \text{ and}$$

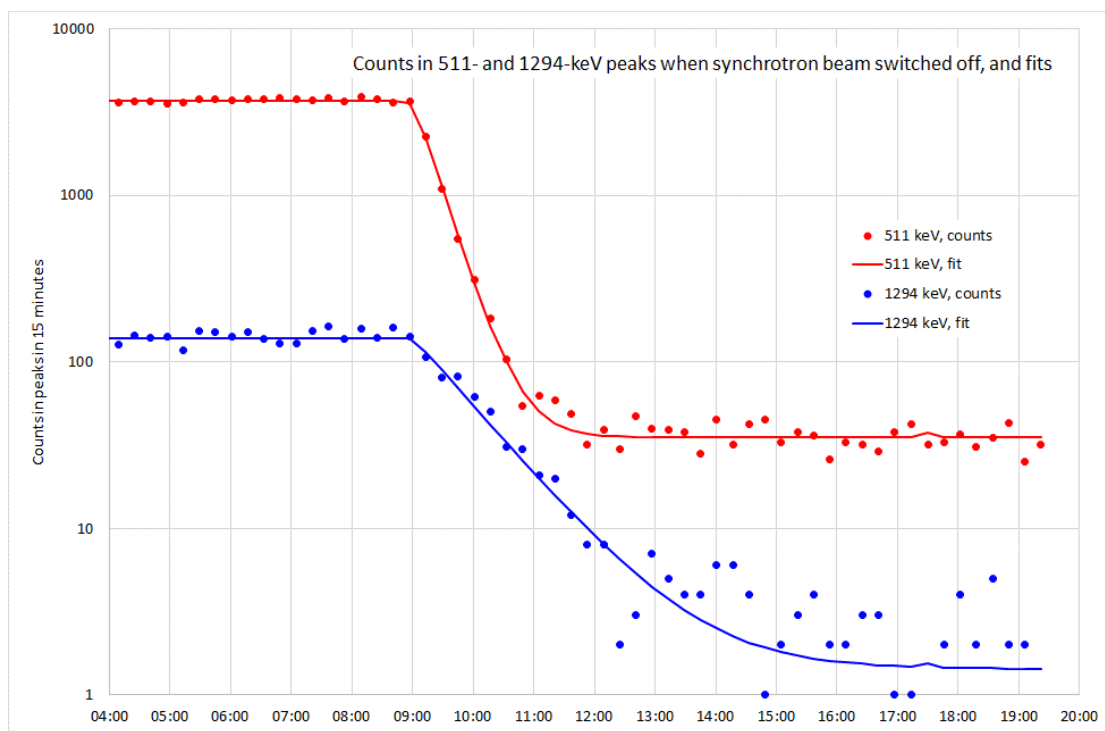
241
$$C'_{1294,i} = C'_{41\text{Ar},i} + b_{1294}(t_i - t_{i-1})$$

242 where the $C'_{AZ,i}$'s are the expressions $C'(t_1, t_2) = \int_{t_1}^{t_2} c'(t) dt$ given in the Appendix
 243 with $t_1 = t_{i-1}$, $t_2 = t_i$, $\lambda = \lambda_{AZ} + r_{\text{exch}}/V$ where $\lambda_{AZ} = \ln(2)/t_{1/2,ZA}$, $t_{1/2,ZA}$ is 9.96,
 244 20.3 and 110 minutes for ^{13}N , ^{11}C and ^{41}Ar respectively, V is the volume of the
 245 synchrotron room (25000 m^3) and $w = 4$ minutes. The expressions for $C'(t_1, t_2)$ were
 246 used because during the measurement of the time for air to be moved by the ventilation
 247 systems from the shutter voids to the synchrotron room (see Sect. 3.3) it was observed
 248 that by the time a short burst of smoke in the shutter void had reached the synchrotron
 249 room it had spread out in time over ± 4 minutes. The seven fitted parameters are $c_{0,13\text{N}}$,
 250 $c_{0,11\text{C}}$, $c_{0,41\text{Ar}}$, b_{511} , b_{1294} , r_{exch} and t_{off} where $c_{0,13\text{N}}$, $c_{0,11\text{C}}$ and $c_{0,41\text{Ar}}$, are the steady-
 251 state count rates c_0 in the expressions in the Appendix for ^{13}N , ^{11}C and ^{41}Ar
 252 respectively, b_{511} and b_{1294} are the corresponding background rates for the 511- and
 253 1294-keV signals, r_{exch} is the air exchange rate at the position where the measurements
 254 are being made, and t_{off} is the beam-off time (since beam-off times are not synchronised

255 with the ‘15-minute clock’ of the gamma-ray counting régime). Results are given in
 256 Table 3, and one of the fits is shown in Fig. 7. From Table 3, it is clear that the 511-
 257 keV activity is nearly all due to ^{11}C , in agreement with the earlier conclusion in Sect. 3.2
 258 — and, incidentally, in accordance with the suggestion [4] that on balance most of the
 259 511-keV air activity is more likely to be due to ^{11}C than to ^{13}N . From the numbers in
 260 Table 3, it is evident that the upper limit for the ^{13}N component of the 511-keV activity
 261 is at most one tenth of the ^{11}C component.

262 3.6 Comparison with results from discharge stack monitor

263 Air from the ISIS synchrotron room is exhausted to atmosphere through a filtered stack
 264 (the stack is included schematically in Fig. 1), and a drum-and-Geiger-tubes system [4]
 265 is used to monitor the activity in the air being discharged. The sensitivities of the
 266 monitoring system are respectively 5330 ± 690 and 3730 ± 340 $\text{Bq m}^{-3} \text{ cps}^{-1}$ for ^{11}C
 267 and ^{41}Ar in air, decay-curve measurements show [4] that essentially two-thirds of the
 268 activity is due to ^{11}C and one-third to ^{41}Ar , and the system typically counts at rates in
 269 the range of 25–30 cps. Consequently, including decay of ^{11}C and ^{41}Ar over the
 270 10 minutes it takes on average for air to be circulated once through the synchrotron
 271 room [4], the specific activity of the air in the synchrotron room deduced from the stack
 272 monitor is 0.174 ± 0.026 Bq cm^{-3} — a value consistent with the results from the HPGe
 273 detector in the present work.



274

275 Fig. 7. Typical 7-parameter simultaneous fits to 511- and 1294-keV lines when synchrotron
 276 beam is switched off. This fit is for the December 2018 data, where the activated air was
 277 sampled close to the inner side of the synchrotron ring between superperiods 5 and 6.

278 3.7 Comparison of measured and calculated activities

279 In ref. [3], saturation specific activities of the air in the ISIS synchrotron room were
 280 calculated using MCNPX and CINDER-90 [6, 7]. As already mentioned, these were

281 very preliminary calculations with a likely uncertainty of some $\pm 50\%$, but the result
282 was a total saturation specific activity of 0.56 Bq cm^{-3} , not so very different, given the
283 inevitable approximations and uncertainties involved, from the measured result.

284 4. Discussion

285 Why, contrary to the expectation from the Monte Carlo calculations [3] referred to in
286 Sect. 3.2 that ^{13}N should make the largest contribution to 511-keV activity, might the
287 511-keV activity in the air in the synchrotron room be largely due to ^{11}C ? A first
288 possible explanation could simply be that the Monte Carlo calculations did not include
289 all routes for the production of activity. The calculations included incident neutron
290 energies up to 25 MeV, which ought to be sufficiently high to include most of the
291 $^{14}\text{N}(n,2n)$ cross-section [12] for the most plentiful component (nitrogen) of air
292 especially in view of the $\sim E^{-3/2}$ energy dependence [13] of the spallation neutron
293 spectrum, but it may be that higher-energy neutrons that induce spallation reactions in
294 the air play a part. Neglecting the difference between neutron-induced and proton-
295 induced spallation, the cross-sections for the production of ^{11}C and ^{13}N from nitrogen
296 and oxygen by spallation lie in the ranges $\sim 10\text{--}20$ and $\sim 5\text{--}10$ millibarns respectively
297 [14]. In order to produce a saturation activity of $\sim 0.10\text{--}0.15 \text{ Bq cm}^{-3}$ (see Table 2) in
298 the air in the synchrotron room by spallation, the effective flux of neutrons inducing
299 spallation in the TS-1 shutter void would have to be $\sim 4\text{--}5 \times 10^5 \text{ cm}^{-2} \text{ s}^{-1}$, and such a flux
300 is not entirely inconsistent with the geometry and materials in the TS-1 monolith and
301 the TS-1 neutron source term of $2 \times 10^{16} \text{ s}^{-1}$ [15].

302 A second possible explanation could be that some chemical or filtering process
303 preferentially removes ^{13}N from the air (air from the shutter voids does pass through
304 HEPA filters). It could be that the nitrogen activity is accompanied by radiolytic
305 production of NO_x (nitrogen oxide), some of the NO_x being then effectively removed
306 chemically through combination with water vapour [16]. Such a possibility is not
307 inconsistent with corrosion of metal surfaces observed in practice in and around the
308 TS-1 target monolith.

309 A third possible explanation could be, in principle, that after its production the activated
310 air takes so long to reach the HPGe detector that the ^{13}N has mostly decayed. But it
311 was seen in Sect. 3.3 that the delay time between production of activity and detection
312 of activity in the HPGe gamma-ray detector is $\sim 12\text{--}20$ minutes, and this is consistent
313 with the delay time of 22 ± 2 minutes [4] observed between production of activity and
314 detection of activity in the drum-and-Geiger-tubes system monitoring the activity of air
315 discharged from the synchrotron room, since it takes several minutes for air from the
316 synchrotron to reach the drum-and-Geiger-tubes monitoring system whereas it takes
317 only ~ 1 minute for air from the synchrotron to reach the HPGe detector. But even if
318 the overall production-to-detection delay time were as much as 20 minutes, ^{13}N activity
319 would decrease relative to ^{11}C activity by a factor 2, and the application of such a factor
320 to the results of the Monte Carlo calculations would be far from sufficient to reproduce
321 the apparent paucity of ^{13}N in the measured data.

322 All things considered, the most plausible explanation is that spallation plays an
323 important part in activation of the air and that the chemical removal of ^{13}N is not
324 negligible. It may be noted that in the NuMI facility at Fermilab [17] and in spallation
325 at KEK [18] the measured values of ^{11}C and ^{13}N specific activity in activated air are in
326 the ratios of 1.6 : 1.0 and 2.3 : 1.0 respectively. It would be worthwhile to perform
327 Monte Carlo calculations with highly detailed models of the entire ISIS TS-1 and TS-2

328 monoliths (including detailed models of the shutters and detailed models of the target,
329 reflector and moderators (TRAM) assemblies) and tracking all particles with energies
330 up to 800 MeV, but the effort to perform such calculations would not be trivial.

331 It is noticeable in Table 1 that overall source-to-detection times $\delta_s + \delta_c$ are one or two
332 minutes greater for air sampled from near superperiods (SPs) 0 and 9 than for air
333 sampled from near SPs 5 and 6. This is not inconsistent with the facts that these two
334 air-sampling positions (near SPs 0 and 9 and near SPs 5 and 6) are on opposite sides of
335 the synchrotron ring and that activated air from the shutter voids enters the synchrotron
336 room nearer SPs 5 and 6 than SPs 0 and 9.

337 And it is noticeable in Table 2 that specific activities of both ^{11}C and ^{41}Ar in the air are
338 smaller for air sampled from near SPs 0 and 9 than for air sampled from near SPs 5 and
339 6, and that ratios of specific activity fall essentially into two sets, the ^{11}C -to- ^{41}Ar ratio
340 for air sampled from near SPs 0 and 9 SPs 5 and 6 being smaller than the ^{11}C -to- ^{41}Ar
341 ratio for air sampled from near SPs 5 and 6. These differences may be due to the
342 activated air ‘spreading out’ as it is circulated through the synchrotron room.

343 5. Summary and conclusions

344 Measurements of air activation in the synchrotron room of the ISIS Spallation Neutron
345 and Muon Source have been made over a period of time spanning two years. Typical
346 specific activities of ^{11}C and ^{41}Ar in the air in the synchrotron room were found to be
347 ~ 0.10 and ~ 0.03 Bq cm^{-3} respectively, but little or no ^{13}N was found. The total specific
348 activity of ~ 0.13 Bq cm^{-3} is consistent with specific activities found independently
349 using a drum-and-Geiger-tubes system [4] to monitor the air being discharged from the
350 synchrotron room.

351 There is good evidence that the air in the synchrotron is activated in the shutter voids
352 in the massive shielding monoliths surrounding the neutron-producing targets.

353 There is plausible evidence that measured specific activities and measured source-to-
354 detector times are consistent with expectations based on the known layout of the
355 synchrotron room.

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384

Date of measurement	Location	Counts / 15 minutes, 511 keV	Counts / 15 minutes, 1294 keV	Bq cm ⁻³ , ¹¹ C	Bq cm ⁻³ , ⁴¹ Ar	Bq cm ⁻³ , ¹¹ C ÷ Bq cm ⁻³ , ⁴¹ Ar
Feb. 2018	Foil change area, SP0	1500 ± 50	100 ± 5	0.083 ± 0.017	0.034 ± 0.007	2.5 ± 0.1
Mar. 2018	SP0/9 datum points	550 ± 25	40 ± 5	0.030 ± 0.006	0.014 ± 0.003	2.2 ± 0.3
May 2018	SP9 datum point and SP6 bridge	1550 ± 50	105 ± 10	0.086 ± 0.018	0.036 ± 0.008	2.4 ± 0.2
Jul. 2018	Inner side of SP5/6	4200 ± 200	155 ± 15	0.233 ± 0.049	0.053 ± 0.012	4.4 ± 0.5
Oct. 2018	Inner side of SP5/6	2900 ± 150	100 ± 10	0.161 ± 0.034	0.034 ± 0.008	4.7 ± 0.5
Dec. 2018	Inner side of SP5/6	3700 ± 150	140 ± 15	0.205 ± 0.043	0.047 ± 0.011	4.3 ± 0.5
Apr. 2019	Inner side of SP5/6	4000 ± 200	150 ± 15	0.222 ± 0.047	0.051 ± 0.012	4.4 ± 0.5
Dec. 2019	Inner side of SP5/6	4550 ± 200	155 ± 10	0.252 ± 0.053	0.053 ± 0.011	4.8 ± 0.4
Feb. 2020	Inner side of SP5/6	2000 ± 100	100 ± 10	0.111 ± 0.024	0.034 ± 0.008	3.3 ± 0.4
Means and standard deviations				0.099 ± 0.070	0.032 ± 0.013	0.131 ± 0.071

385 Table 2. Specific activity of air samples from the synchrotron room when the synchrotron was running steadily. Although there are nine sets of data, there are
386 data sets for only four distinct locations, and so the six measurements made in the same place have themselves been averaged before the means and standard
387 deviations are taken. The uncertainties have been obtained by the same method as that described in the caption to Table 1. In the rightmost column the
388 uncertainties are the statistical uncertainties from the counts in the peaks in the gamma-ray spectrum only, as the detection efficiency of the HPGe gamma-ray
389 detector is a common factor in the ¹¹C and ⁴¹Ar specific activities. Whilst in the February 2020 data set the timing data were unavailable, as indicated in Table 1,
390 count rates with the beam on and off were still perfectly visible. Although this table encompasses measurements made over a period of two years, it is perfectly
391 reasonable to form the averages set out in the table because internal ISIS physical configurations and ventilation arrangements for the synchrotron room and
392 target stations (fan speeds, air flow paths, *etc.*) were all always the same during these two years.
393

Location	Date	$c_{0,13\text{N}}$ min^{-1}	$c_{0,11\text{C}}$ min^{-1}	$c_{0,41\text{Ar}}$ min^{-1}	b_{511} min^{-1}	b_{1294} min^{-1}	r_{exch} $\text{m}^3 \text{s}^{-1}$	t_{off} mins.	χ_{pdf}^2
SP0	Feb. 2018	-67 ± 17	163 ± 17	6.4 ± 0.1	2.8 ± 0.1	0.08 ± 0.01	3.8 ± 0.3	± 2.0	1.20
SP0 (no -ve's)		$0^{+0.02}_{-0}$	96.0 ± 0.4	6.4 ± 0.1	2.9 ± 0.1	0.08 ± 0.01	3.7 ± 0.4	± 1.0	1.33
SP9, 6	May 2018	-89 ± 27	183 ± 27	6.9 ± 0.1	2.4 ± 0.1	0.08 ± 0.03	3.5 ± 0.5	± 3.8	1.93
SP9, 6 (no -ve's)		$0^{+0.15}_{-0}$	93.1 ± 0.5	6.7 ± 0.2	2.4 ± 0.1	0.07 ± 0.03	3.5 ± 0.5	± 1.5	2.12
SP5,6	Dec. 2018	6^{+36}_{-6}	223 ± 21	8.6 ± 0.1	2.2 ± 0.1	0.09 ± 0.03	4.1 ± 0.4	± 0.9	1.91

394 Table 3. Results of seven-parameter fits to beam-on-to-beam-off transitions, with inclusion of rectangular time-dispersion function of half-width 4 minutes as
395 described in the text (the February, May and December 2018 data sets are the only data sets for which the timing bins are 15 minutes wide and in which there
396 is at least twelve hours of steady running before beam-off and twelve hours after beam-off to establish background). The uncertainties have been obtained by
397 the same method as that described in the caption to Table 1. The values of the chi-squared of the fit per degree of freedom χ_{pdf}^2 are also shown. The three sets
398 of data were each first fitted allowing all seven parameters to have complete freedom; when unphysical negative numbers for the ^{13}N contribution to overall
399 activity were found for the 'SP0' and 'SP9, 6' data sets, these two data sets were re-fitted with very little resultant increase in χ_{pdf}^2 whilst constraining all seven
400 parameters to be ≥ 0 thereby showing the ^{13}N contribution to be minimal (these two re-fitted data sets are labelled 'no -ve's'). The absolute value of t_{off} is
401 irrelevant, as all that matters is the uncertainty with which it can be defined. The fitted value of r_{exch} , the rate of air exchange in the synchrotron room, is a little
402 greater than the nominal air extraction rate of $2 \text{ m}^3 \text{ s}^{-1}$, but this may be a consequence of likely non-uniform patterns of air movements within the synchrotron
403 room.
404

405 **Appendix**

406 Let the transition at time zero between a constant count rate and a decaying count rate
 407 be described by: $c(t) = c_0, t < 0$; $c(t) = c_0 \exp(-\lambda t), t \geq 0$.

408 Let the count rate $c(t)$ be smeared by the rectangular function:
 409 $s(t') = 0, t' < -w$; $s(t') = 1/(2w), -w \leq t' \leq w$; $s(t') = 0, t' > w$.

410 If the smeared count rate is $c'(t) = \int_{t-w}^{t+w} c(t') s(t-t') dt'$, then:

411 $c'(t) = c_0, t < -w$;

412 $c'(t) = \{c_0/(2w)\} \{w - t + (1 - \exp(-\lambda(t+w)))/\lambda\}, -w \leq t \leq w$; and

413 $c'(t) = \{c_0 \exp(-\lambda t)/(2w\lambda)\} \{\exp(\lambda w) - \exp(-\lambda w)\}, t > w$.

414 The integral of the smeared count rate between times t_1 and t_2 $C'(t) = \int_{t_1}^{t_2} c'(t) dt$
 415 is given by summing one or more of the three following integrals with appropriate
 416 choices of the limits τ_1 and τ_2 ($\tau_1 \leq \tau_2$):

417 $I_1(\tau_1, \tau_2) = c_0(\tau_2 - \tau_1), \tau_1$ and τ_2 both $< -w$;

418 $I_2(\tau_1, \tau_2) = \{c_0/(2w)\} \{w(\tau_2 - \tau_1) - (\tau_2^2 - \tau_1^2)/2 + (\tau_2 - \tau_1)/\lambda -$
 419 $(\exp(-\lambda w)/\lambda^2)(\exp(-\lambda\tau_1) - \exp(-\lambda\tau_2))\}, \tau_1$ and τ_2 both $\geq -w$ and $\leq w$; and

420 $I_3(\tau_1, \tau_2) = \{c_0/(2w\lambda^2)\} \{\exp(\lambda w) - \exp(-\lambda w)\} \{\exp(-\lambda\tau_1) - \exp(-\lambda\tau_2)\},$
 421 τ_1 and τ_2 both $> w$.

422 For example, if $t_1 < -w$ and $t_2 > w$,

423 $C'(t_1, t_2) = \int_{t_1}^{t_2} c'(t) dt = I_1(t_1, -w) + I_2(-w, w) + I_3(w, t_2)$.