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# Extreme Differentiation along Multiple Liquid Lines of Descent in Strongly Peralkaline Magma Series at Pantelleria (Italy)

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#### Abstract

The liquid line of descent from trachyte to pantellerite is controlled primarily by fractional crystallization of alkali feldspar, with wholerock compositions following a fractionation path along the 'thermal valley' in the peralkaline haplogranite system Qz-Ab-Or-Ac-Ns and terminating at a minimum on the feldspar-quartz cotectic. Although whole-rock compositions for different pantelleritic suites follow nearly identical paths in a Qz-Ab-Or projection that terminate near the experimental minimum (Qz<sub>40.5</sub>Or<sub>34.5</sub>Ab<sub>25</sub> at 100 MPa, projected from Ac-Ns), matrix glass from samples with near-minimum compositions record extreme differentiation and form a 'cotectic delta' beyond the terminus of the 'thermal valley'. Although each glass trend shows a continuing increase in Zr to >3000  $\mu$ g/g, the most evolved compositions in each suite differ in peralkalinity (mol [Na + K] / Al) and in the proportions of FeO<sup>T</sup>, Qz, Ab, Or, and other components, which are related to subtle variations in the mafic phases controlled mainly by differences in oxygen fugacity (fO<sub>2</sub>) and pressure (P). To determine the controls over mafic mineral crystallization in pantelleritic magmas and the various paths these suites take beyond the apparent (whole-rock) minimum, amphibole-phyric suites from the ~159 ka Cala dell'Altura and Cala Gadir volcanic centres and the ~8-10 ka Cuddia Mida volcanic centre on Pantelleria have been analyzed and compared with each other and with the well-characterized and amphibole-free, compositionally zoned Green Tuff, the ~46 ka caldera-forming ignimbrite of the Cinque Denti caldera. Differences between the extended fractionation trends may be ultimately attributed to variations in oxygen fugacity, depth of emplacement, and water saturation. Shallower (lower pressure) magma reservoirs such as the one for the Green Tuff are watersaturated and undergo degassing, which leads to an increase in relative oxygen fugacity. Deeper (higher pressure) magma reservoirs remain water-undersaturated and retain water in the melt, which both maintains lower relative oxygen fugacities and enables the crystallization of amphibole. Amphibole formation appears to require melt water contents >4 wt%, low oxygen fugacity (<ΔNNO-1.5), and low temperatures (<700°C), although fluorine may stabilize it at higher temperatures in some rocks.

Keywords: pantellerite, peralkaline, trachyte, oxygen fugacity, Extreme differentiation

# INTRODUCTION

In Petrogeny's Residua System, as represented by the plane Qz-Ab-Or, felsic melts evolve towards and along a so-called 'thermal valley' towards a low-temperature point ('minimum') on the quartz-feldspar cotectic (Tuttle & Bowen, 1958). As Bailey & Schairer (1964) recognized, Qz-Ab-Or is not the residua system for peralkaline (i.e. peralkalinity index, P.I. [mol (Na+K) / Al] > 1.0) silicic magmas; continued feldspar fractionation drives them out of the plane towards increasing peralkalinity and into a system better represented by Qz-Or-Ab-Ac-Ns (where Ac is acmite  $[NaFe^{3+}Si_2O_6]$  and Ns is sodium metasilicate  $[NaSiO_3]$ ), and terminating in a number of minima forming what might be termed a 'cotectic delta.' Attempts to describe these extended trends were made experimentally (Carmichael & MacKenzie, 1963; Bailey & Schairer, 1966; Thompson & MacKenzie, 1967) and by plotting compositional data into various projection schemes (Bailey & Macdonald, 1969; Roux & Varet, 1975). The form of the lowtemperature zone within the alkali feldspar primary phase region in peralkaline silicic systems and the existence of a unique lowtemperature point has remained, however, uncertain. For example, based on feldspar-rock relationships in peralkaline trachytes and rhyolites of the Menengai volcano, Kenya, Macdonald *et al.* (2011) suggested that several evolutionary paths are possible within the peralkaline low-temperature zone and that there is no unique endpoint.

Renewed interest in this question was raised by the discovery in peralkaline rhyolites of the Boseti volcano, Ethiopia, of very highly fractionated matrix glasses and melt inclusions in phenocrysts, with FeO<sup>T</sup> (total Fe as Fe<sup>2+</sup>) up to 17 wt%, Al<sub>2</sub>O<sub>3</sub> as low as 2.3 wt% and SiO<sub>2</sub> ~ 65 wt% (Macdonald *et al.*, 2012). Highly fractionated glasses had previously been reported from pantellerites of Palo Alto, Terceira (Mungall & Martin, 1995), and in an obsidian from the Fantale volcano, Ethiopia (Lacroix, 1930). Macdonald *et al.* (2012) used the information set to suggest that the effective minimum composition (EMC) for peralkaline silicic magmas has ~5 wt% Al<sub>2</sub>O<sub>3</sub>, 13 wt% FeO<sup>T</sup> and  $66 \pm 2$  wt% SiO<sub>2</sub>. The use of the

term 'effective' alluded to the fact that melts more evolved than this are unlikely to separate from their crystal mush hosts. The relatively high density due to the high Fe contents and their normal occurrence as low melt fractions is likely to inhibit crustal ascent and they remain trapped at depth. Possible exceptions are a pantellerite obsidian (lava?) from Fantale, Ethiopia (Lacroix, 1930) and a lava from Mt. Takahe, Marie Byrd Land, Antarctica (LeMasurier et al., 2018). It is important to note that at Boseti the highly evolved compositions occur in pantellerites where alkali feldspar + fayalite + hedenbergite + oxide  $\pm$  quartz is the dominant assemblage, whereas at Mt. Takahe alkali feldspar is the only reported phenocryst phase. In a majority of eruptive peralkaline silicic suites, the peralkaline varieties are derived by fractionation of metaluminous trachytic melts (White et al., 2009; Neave et al., 2012; Jeffery & Gertisser, 2018; Romano et al., 2018; Brenna et al., 2019; Macdonald et al., 2021). Protracted fractionation of the peralkaline melts may result in Fe-rich, Al-depleted melts which represent the most evolved stages of peralkaline systems.

In this report, we provide mineral chemical and glass data for three peralkaline suites from the island of Pantelleria (adj. Pantescan) to give further insights into how the magmas evolved within the low-temperature zone. On Pantelleria several eruptive sequences include rocks spanning the compositional range from trachyte to pantellerite. Some contain glass (melt) with extreme compositions (up to 13.2 wt% FeO<sup>T</sup>, as low as 4.7 wt% Al<sub>2</sub>O<sub>3</sub>, P.I. up to 3.5). Mineral and whole-rock chemical studies have revealed that each suite evolved along a different liquid-lineof-descent. This contribution focusses on the parameters (P, T,  $fO_2$ ,  $H_2O_m$ ) controlling the trends and the different assemblages formed. Particular attention is paid to the formation of amphibole which, in contrast to the other mafic phases such as olivine and pyroxene, shows considerable complexity in peralkaline silicic suites. For example, the composition of amphibole in comendites of the Olkaria complex, Kenya, changes with increasing wholerock peralkalinity from kataphorite to richterite to silica-poor riebeckite (Marshall et al., 2009). Macdonald et al. (2019) reported ferrorichterite in the Gold Flat pantelleritic ash flow tuff from southern Nevada. Ferrorichterite and arfvedsonite form small phenocrysts in pantellerites from the Eburru complex, Kenya (Ren et al., 2006). On the island of Pantelleria, Italy, amphibole phenocrysts are uncommon and absent from most eruptive units. To date, they have been described only in rocks from three smaller volcanic centres (Fig. 1), namely those being used in this report to examine the nature of the low-temperature zone.

These suites are compared with the results of relevant experiments on comenditic and pantelleritic samples (Scaillet & Macdonald, 2001, 2003, 2006; Di Carlo *et al.*, 2010; Romano *et al.*, 2020) and the Green Tuff, the best-understood unit on the island (Williams, 2010; Williams *et al.*, 2014; Liszewska *et al.*, 2018; Romano *et al.*, 2019). The pantelleritic portion of the Green Tuff is compositionally similar to the suites studied here but, critically, does not contain amphibole phenocrysts, although it was produced in low temperature experimental runs on the Green Tuff even at moderate crystallinities, allowing us to look for differences in the crystallization conditions.

# **GEOLOGICAL SETTING**

Pantelleria is located in the Strait of Sicily rift zone (SSRZ), a NW-SE trending transtensional rift system situated on the Pelagian Block, a northern promontory of the African plate (Catalano *et al.*, 2009; Fig. 1). The structure of the island is dominated by two nested calderas (Mahood & Hildreth, 1983, 1986): the La Vecchia

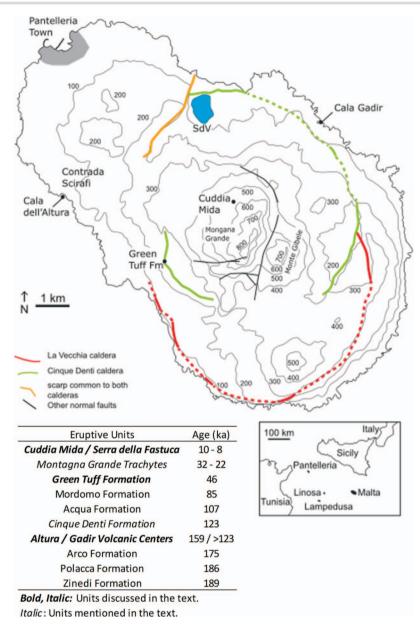
caldera, dated at 140–146 ka by <sup>40</sup>Ar/<sup>39</sup>Ar dating of feldspar (Rotolo et al., 2013, 2021), and the Cinque Denti (or Monastero) caldera, formed at 45.7 ± 1 ka (2 $\sigma$ ) (Scaillet *et al.*, 2013). The Cinque Denti caldera was created by the eruption of the compositionally zoned Green Tuff, which covered much of the island obfuscating the stratigraphy of the pre-caldera sequences. This was followed by a prolonged period of effusive activity from ~32–22 ka (Cornette et al., 1983; Civetta et al., 1984; Mahood & Hildreth, 1986; White et al., 2022) that formed the trachytic Montagna Grande-Monte Gibele shield volcano, erupted from deeper levels in the Green Tuff reservoir (Ferla & Meli, 2006; Neave et al., 2012), and several pantellerite shields and lava domes that erupted ~15–6 ka (Speranza et al., 2010; Scaillet et al., 2011) mostly from vents within the moat and along the rim of the caldera (Mahood & Hildreth, 1983, 1986; Civetta et al., 1988; Rotolo et al., 2007).

Pre-Green Tuff volcanic centres known to have erupted amphibole-phyric lavas and tuffs are limited to the Cala dell'Altura volcanic centre in Contrada Scirafi on the west coast and the Cala Gadir volcanic centre on the northeast coast. Gadir crops out along the coast as a volcanic shield buried by the Green Tuff and post-Green Tuff lavas; the Altura outcrop consists primarily of a sequence of welded and unwelded pumice fall deposits. Altura and Gadir occupy the same stratigraphic interval, situated below the 123 ka Cinque Denti Formation (based on <sup>40</sup>Ar/<sup>39</sup>Ar dating of feldspar; Jordan et al., 2018), which is consistent with the K-Ar age of 159±8 ka for Altura reported by Mahood & Hildreth (1986). Amphibole microphenocrysts have been reported from the Cinque Denti Formation (formerly called Unit P), a crystal-rich, welded and rheomorphic ignimbrite dated at  $123 \pm 1.6$  ka ( $^{40}$ Ar/ $^{39}$ Ar) by Rotolo et al. (2013); it is possible that the Altura, Cinque Denti Formation, and Gadir rocks are part of the same magmatic system, but this remains to be tested. Post-Green Tuff volcanic centres known to have erupted amphibolephyric lavas are restricted to two proximal vents, Cuddia Mida  $(8.2 \pm 2.2 \text{ ka K-Ar}; \text{Civetta et al., 1984, 1988})$  and Serra della Fastuca  $(9.7 \pm 0.6 \text{ ka} {}^{40}\text{Ar}/{}^{39}\text{Ar}; \text{ Scaillet et al., 2011})$ , associated with the 'Hingeline Vent System' located within the Cinque Denti caldera on the north-western slope of Montagna Grande and are likely part of the same magma system (Mahood & Hildreth, 1986; Orsi et al., 1989; Rotolo et al., 2007).

Although these magma series are clearly not consanguineous, they demonstrate similar incompatible trace element ratios (cf. Jordan *et al.*, 2021). They have, for example, Zr/Nb (4.9–6.6) and chondrite-normalized La/Yb ratios (Fig. 2a; La<sub>N</sub>/Yb<sub>N</sub> 9.5–11.3; McDonough & Sun, 1995) that overlap with the Green Tuff (4.9–5.6 and 8.8–11.6, respectively; Liszewska *et al.*, 2018) and the low-Ti and P basalts on Pantelleria (3.3–5.9 and 8.3–10.8, respectively; White *et al.*, 2020). They also show nearly identical patterns on a multi-element variation diagram normalized to depleted MORB mantle [DMM] (Fig. 2b; Salters & Stracke, 2004) with patterns similar to that of a representative sample of basalt, having elevated concentrations for all elements except Ba, Sr, P, and Ti. The comparisons suggest that they were derived from a similar source and evolved along broadly similar compositional trends.

# METHODS

Sixteen samples from amphibole-bearing units from the Cala dell'Altura ('Altura'), Cala Gadir ('Gadir'), and the Cuddia Mida / Serra della Fastuca ('Mida') volcanic centres were collected over the period 1998–2019 (Table 1). Data for four samples were presented in White *et al.* (2005, 2009), including two samples from Altura (98527 and 060522) that were erroneously identified as



**Fig. 1.** Map of Pantelleria showing topography (contour interval = 100 m), proposed calderas (after Mahood & Hildreth, 1986; Jordan *et al.*, 2018), post-caldera faults, the type locality of the Green Tuff (GT) Formation (Williams *et al.*, 2014), and the locations of Contrada Scirafi, Cala dell'Altura, Cala Gadir, and Cuddia Mida. SdV is Lago Specchio di Venere. Inset: Location of Pantelleria within the Strait of Sicily. Generalized stratigraphy of the eruptive units and the ages are adapted from Speranza *et al.* (2010), Scaillet *et al.* (2011, 2013), Rotolo *et al.* (2013, 2021); Jordan *et al.* (2018, 2021), and White *et al.* (2022).

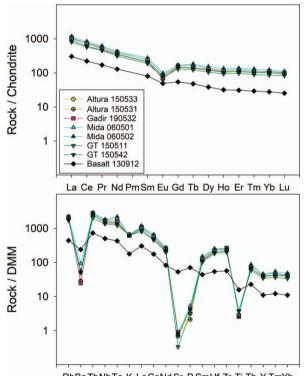
Green Tuff by those authors (Nina Jordan, personal communication, 2016). Samples 090532 (Altura) and 090511–14 (Mida) were analysed at Activation Laboratories, Canada. Major elements and Ba, Be, Sc, Sr, V, Y, and Zr were analysed by ICP-OES on a fused bead (Code WRA-ICP-4B). Cu, Ni, Pb, S, and Zn were analysed by ICP-OES following acid digestion (Code 4B1). All other trace elements plus Cl were analysed by INAA (Codes 4B-INAA and 4F-Cl); F was determined with ion selective electrodes (Code 4F-F). Samples 060501–02 (Mida) and 190532 (Gadir) were also analysed at Activation Laboratories using ICP-OES for major elements and ICP-MS for trace elements (Code 4Lithoresearch). Samples 150531–33 were analysed at Bureau Veritas Commodities Canada Ltd. Major elements and Cr were analysed by ICP-OES, and trace elements, including REE, by ICP-MS. Loss on ignition (LOI) was determined by weight difference before and after ignition at 1000°C. Mean detection limits on major elements were 0.01 wt%, and between 0.10 and 10  $\mu$ g/g for trace elements. Also included in our whole-rock dataset are WD-XRF analyses of three samples from Jordan (2014), used with permission: 171010–1 is a sample of Altura pantellerite sampled from the middle of the upper welded pyroclastic unit; Unk34A and Unk34B are samples of Gadir lava collected by Rebecca Williams.

New mineral and glass compositions (samples 090511–14, 150531–33, 190532) were determined by electron microprobe at the Inter-Institute Analytical Complex at the Department of Geochemistry, Mineralogy and Petrology, University of Warsaw, using a Cameca SX-100 microprobe equipped with four wavelength dispersive spectrometers. Minerals and glasses were analysed with a 15 kV accelerating voltage and 40 nA beam current, beam diameter from 5 to 20  $\mu$ m, with counting times of 20 s

Table 1:	Volcanic	centres,	sample	locations,	and	phenocr	yst assemblages

		Location				C	рх			Aen	Aı	mph		Ар	
Centre	Sample	Lat (N)	Long (E)	Matrix	Kfs	X-Fe <sub>tot</sub>	apfu Na	Ol	FeTiOx	Ti#	X-Fe <sub>tot</sub>	apfu Na(B)	Qtz	wt% F	Ро
Altu	r <b>:0</b> 60522	36.79000	11.94611	Glass	Or23–31	0.63–0.67	0.06	Fa84	Mgt	_	_	_	_	_	+
	98527	36.79000	11.94611	Glass	Or38	0.84-0.89	0.18	_	Ilm	0.16	0.75–0.83	0.94	+	-	_
	060521	36.79000	11.94611	Pumice	Or36	+	+	-	+	+	+	+	+	-	_
	090532	36.78942	11.94622	Glass	+	+	+	_	+	+	+	+	+	-	_
	150532	36.78847	11.94678	Glass	Or36	0.87	0.18	-	Ilm	0.17	0.73-0.82	0.88–0.94	+	+	+
	150533	36.78860	11.94666	Glass	Or37	0.82-0.87	0.13-0.20	_	Ilm	0.17	0.69–0.82	0.83-0.91	+	+	+
	150531	36.78847	11.94678	Pumice	+	+	+	-	-	+	+	+	-	-	-
Gadir	190532	36.81061	12.02731	Glass	Or34–37	0.87–0.90	0.16-0.22	-	Ilm	0.17	0.75–0.80	0.92-1.00	+	2.83	_
Mida	a 090512	36.78694	11.99506	Glass	Or37	0.92–0.95	0.15-0.22	Fa98	Ilm	0.17	+	+	+	2.12	+
	090511	36.78539	11.99389	Glass	Or32–36	0.94–0.95	0.20-0.23	_	Ilm	0.16	+	+	+	-	+
	060501	36.78250	11.98639	Glass	Or34–37	0.94–0.95	0.23	Fa95	Ilm	0.16	0.91–0.92	1.13	+	+	+
	090514	36.78269	11.98647	Glass	Or13–37	0.94–0.95	0.15-0.30	_	+	0.16	0.93	1.11	+	2.44	_
	090513	36.78353	11.98822	Glass	+	+	+	-	Mgt	+	0.93–0.94	1.20	+	2.21	_
	060502	36.78167	11.98750	Glass	Or38–44	0.94–0.96	0.21	-	-	0.15	0.92–0.94	1.18	+	3.61	-

Kfs, Alkali Feldspar; Cpx, Clinopyroxene (apfu = atoms per formula unit); Ol, Olivine; FeTiOx, Iron-Titanium Oxide(s); Aen, Aenigmatite ( $Ti#=Ti / [Ti + Fe^T]$ ); Amph, Amphibole (Na(B) = octahedral Na); Qtz, Quartz; Ap, Apatite; Po, Pyrrhotite. (+), observed but not analyzed; (-), not observed. X-Fe<sub>tot</sub> = mol Fe<sup>T</sup>/(Fe<sup>T</sup> + Mg). Geodetic datum: WGS84. Descriptions of 060522, 98527, 060521, 060501, and 060502 adapted from White *et al.* (2005, 2009).



RbBaThNbTa K LaCeNdSr P SmHf Zr Ti Tb Y TmYb

**Fig. 2.** Representative samples from each suite plotted on (a) a rare-earth element diagram (normalized to CI Chondrite; McDonough & Sun, 1995) and (b) a multi-element variation diagram (normalized to depleted MORB mantle [DMM]; Salters & Stracke, 2004). Included in each plot are representative samples of GT (Liszewska *et al.*, 2018) and a sample of low Ti-P basalt (130912; White *et al.*, 2020) for comparison.

on peak and 10 s on each of two background positions. The standards, crystals, X-ray lines used, and generalized detection limits are given in Supplementary Table 1. The 'PAP' $\varphi(\rho Z)$  program (Pouchou & Pichoir, 1991) was used for corrections. Apatite was analysed using the technique outlined in Macdonald *et al.* (2008a).

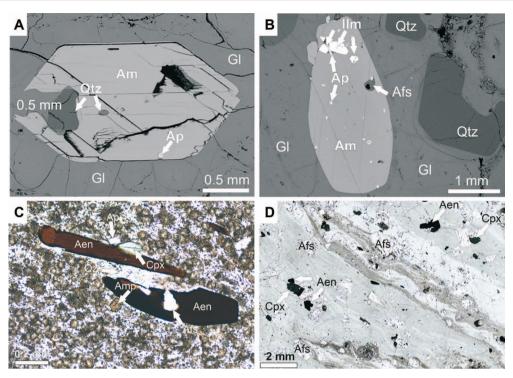
White *et al.* (2005, 2009) presented only representative analyses of minerals from four samples used in this study; the previously unpublished complete datasets of electron microprobe analyses are included in our supplementary tables and figures; the analytical methods for these are described in those papers.

# **RESULTS** Petrography

The petrography of the new analysed samples is here described in some detail, since this is the first detailed account of the amphibole-phyric rocks on Pantelleria and since the textural relationships provide important constraints on interpreting the preeruptive crystallization conditions. A summary of the petrography and mineral chemistry is presented in Table 1. Salient points to note at this stage are: (i) While the dominant phenocrysts are the same, there are small differences, e.g. the occasional presence of olivine and/or magnetite only in the Mida samples. (ii) Evidence from mutual inclusion relationships (Fig. 3) suggests that the phenocryst phases generally crystallized together, in some cases in the first 10% modally. (iii) The mafic phases occur in variable proportions but generally amphibole > aenigmatite > clinopyroxene. (iv) Quartz occurs as a macrophenocryst in these amphibolebearing rocks as opposed to the amphibole-free Green Tuff, in which it is absent except as heavily embayed microphenocrysts in the most evolved samples (Liszewska et al., 2018; Romano et al., 2019).

### Cala dell'Altura volcanic centre

Sample 060522 is a crystal-rich (20–25 vol%) unwelded tuff with a partially devitrified pumiceous groundmass from the upper part of the Altura section. The phenocryst assemblage in 060522 is dominated (~20 vol%) by seriate inequigranular (2– 10 mm) *anorthoclase*, with small amounts of olivine, augite, and magnetite. Samples 150532 and 150533 are gray-green, densely welded eutaxitic lapilli tuffs; 150532 was sampled from the welded unit above the unwelded pumice fall unit (150531) and 150533 was sampled from the welded unit below the pumice.



**Fig. 3.** (a) BSE image of sample 150532 (Altura) showing euhedral amphibole (Am) phenocryst with inclusions of quartz (Qtz) and apatite (Ap) in glass (Gl) matrix. (b) BSE image of sample 15033 (Altura) showing subhedral amphibole phenocryst with alkali feldspar (Afs), ilmenite (Ilm) and apatite inclusions. Note the partially resorbed, initially euhedral quartz. (c) Photomicrograph of sample 150531 (Altura) showing phenocrysts of aenigmatite (aen), clinopyroxene (cpx), and amphibole (amp). (d) Photomicrography of sample 090511 (Mida) showing streaky admixture of three glass types with phenocrysts of aenigmatite, clinopyroxene, and alkali feldspar.

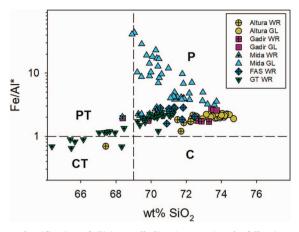
The fiamme are strongly flattened and comprise layers of colourless glass and largely devitrified brown material. Sample 150531 is a creamy brown pumice tuff, consisting of a mixture of flattened pumice lapilli and glassy, partially devitrified zones. The phenocryst assemblage, 10–15 vol%, is alkali feldspar + aenigmatite + amphibole  $\pm$  clinopyroxene  $\pm$  quartz  $\pm$  ilmenite  $\pm$  apatite (Table 1; Fig. 3c). Most abundant is anorthoclase, forming in 150532 and 150533 clear, subhedral to euhedral crystals up to  $\sim$ 1 mm across. In 150531, the feldspars, up to 2 mm across, are commonly fractured. Zonation within crystals is very small (<Or<sub>2</sub>). Contents of Fe<sub>2</sub>O<sub>3</sub> are low (<1.5 wt%). Amphibole (fluorian ferrorichterite) is the most abundant mafic phase, forming prisms up to 2 mm long, pleochroic from chestnut brown to deep brown (Fig. 3). Aenigmatite occurs as elongated plates up to 1 mm long with prismatic terminations. The clinopyroxene phenocrysts in 150532 and 150533 are green, prismatic, and up to 900  $\mu$ m in size. Quartz varies from hexagonal prims to rounded, partly embayed crystals up to 1.5 mm across. In 150531 it forms rare microphenocrysts. Occasional graphic intergrowths are found. Although quartz in Pantescan pantellerites has been considered to be a late-crystallizing phase, perhaps even formed during magma ascent (Di Carlo et al., 2010), it is clearly intratelluric in the Altura rocks. It forms, for example, inclusions in amphibole phenocrysts in 150532 and is of comparable size to amphibole in 150533. Ilmenite (Ilm<sub>97</sub>) is present as inclusions in hedenbergite and, less commonly, in amphibole. Fluorapatite forms prisms and subhedral plates, invariably associated with the mafic minerals. Pyrrhotite occurs as rare, small grains. Mutual inclusion relationships, e.g. ilmenite in amphibole (150533), amphibole in alkali feldspar (150532) and aenigmatite in alkali feldspar (150532), indicate that the phenocryst assemblage was forming in, or close to, equilibrium conditions.

### Cala Gadir volcanic centre

Sample 190532 contains the phenocryst assemblage alkali feldspar-quartz-clinopyroxene- aenigmatite-ilmenite (<10 modal %). The quartz is unusually large, comparable to the largest feldspars. Clinopyroxene including ilmenite is intergrown in one case with feldspar. Amphibole forms euhedral cross-sections and prisms up to 600 x 400  $\mu$ m in size.

#### Cuddia Mida volcanic centre

All samples from Mida belong to the 'hyalopantelleritic explosion breccia' unit of Rotolo et al. (2007). Samples 090511 and 090512 show a streaky admixture of glasses varying from colorless to pale grey to brown (Fig. 3d). Samples 090513 and 090514 are obsidians with pale homogeneous glass. All have the phenocryst assemblage (~10% modal) alkali feldspar + amphibole + aenigmatite + clinopyroxene + FeTi oxides + fluorapatite + quartz  $\pm$  olivine. The feldspar varies from euhedral plates up to 700  $\mu$ m long to small anhedra; some internal resorption is common; certain crystals in 090512 contain rounded melt inclusions. Aenigmatite forms thin rods and plates up to 0.5 mm across, with prismatic terminations. The amphibole, deep brown in color, also forms prisms, up to 0.6 mm long. Magnetite and ilmenite normally form small (~50  $\mu$ m) equant crystals. Deep green clinopyroxene is less abundant than amphibole and aenigmatite and ranges from prismatic crystals up to 300  $\mu$ m long to small ragged anhedra. Apatite is scarce, always associated with the mafic phases. Olivine has been positively identified only in sample 090512, forming



**Fig. 4.** Classification of silicic peralkaline (P.I. > 1.0) rocks following Le Maitre (2002) and Macdonald (1974), with Fe/Al\* =  $1.33 \cdot \text{FeO}^T$  / (Al<sub>2</sub>O<sub>3</sub>-4.4): CT, comenditic trachyte; C, comendite; PT, pantelleritic trachyte; P, pantellerite. WR, whole rock; GL, glass. Fastuca WR are WR analyses of the Serra della Fastuca pumice from Rotolo *et al.* (2007) and White *et al.* (2009). GT WR data are from Williams *et al.* (2014) and Liszewska *et al.* (2018).

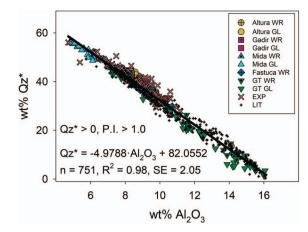
small anhedral platelets. Quartz phenocrysts, up to 700 microns across, are usually partly resorbed and may contain feldspars.

#### Whole-rock and glass compositions

Whole-rock analyses are presented in Table 2. Representative matrix glass analyses are presented in Table 3 and the complete data set is presented in Supplementary Table 2. Samples are classified using the methods of Le Maitre (2002) and Macdonald (1974); nearly all are pantellerites (viz., rhyolites [SiO $_2 > 69$  wt%], peralkaline [P.I. > 1.0], and pantelleritic  $[Fe/A]^* = 1.33 \cdot FeO^T/(Al_2O_3-4.4) > 1.0];$  Fig. 4). In silicaoversaturated peralkaline rocks, due to crystallization of alkali feldspar normative Qz will increase relative to Ab and Or. Here we employ the term  $Qz^* = Qz_n / (Qz_n + Or_n + Ab_n)$ , where  $Qz_n$ ,  $Or_n$ , and Ab<sub>n</sub> are the CIPW normative components reported in wt%] as a differentiation index (cf. Blundy & Cashman, 2008; Wilke et al., 2017; Jordan et al., 2021). CIPW norms were calculated using the method of Kelsey (1965) with iron oxides adjusted following Le Maitre (1976; FeO/FeO<sup>T</sup> = 0.87  $\pm$  0.02). For all the peralkaline analyses presented in this study plus all published analyses from the literature for both natural and experimental glasses for Pantelleria (n = 751) there is a strong negative correlation  $(R^2\!=\!0.98)$  between alumina (normalized anhydrous with total iron recalculated as  $FeO^{T}$ ) and  $Qz^{*}$  (Fig. 5). This observation enables us to slightly modify the Macdonald (1974) classification diagram for peralkaline rocks, allowing a more direct comparison between it and projections in Qz-Or-Ab-Ac-Ns for peralkaline compositions (Fig. 6). Compositional variations between Qz\* and selected components are presented in Fig. 7.

## Green tuff

The Green Tuff whole-rock analyses define a liquid line of descent (LLOD) confined within the 'thermal valley' of Carmichael & MacKenzie (1963) for peralkaline silicic melts. This LLOD extends from comenditic trachyte (63.9 wt% SiO<sub>2</sub>; 5.8 wt% FeO<sup>T</sup>) that plots near the feldspar join (1.27 wt% Qz<sup>\*</sup>) to pantellerite (71.3 wt% SiO<sub>2</sub>; 8.3 wt% FeO<sup>T</sup>) that terminates at 39.0 wt% Qz<sup>\*</sup>, near the Carmichael & MacKenzie (1963) experimental minimum of 40.5 wt% Qz<sup>\*</sup>, at approximately constant values of Or



**Fig. 5.** Wt% Al<sub>2</sub>O<sub>3</sub> plotted versus wt% Q*z*\* for WR and GL samples from this study plus GT WR and glass analyses (Williams, 2010; Liszewska *et al.*, 2018; Romano *et al.*, 2019), WR analyses of Serra della Fastuca samples (Rotolo *et al.*, 2007; White *et al.*, 2009), experimental glasses (EXP; Di Carlo *et al.*, 2001); Romano *et al.*, 2020), and all other pantellerite WR analyses from Pantelleria (LIT; Civetta *et al.*, 1994; Mahood & Stimac, 1990; Kovalenko *et al.*, 1994; Prosperini *et al.*, 2000; Avanzinelli *et al.*, 2004; White *et al.*, 2005, 2009; Ferla & Meli, 2006; Parker & White, 2008; Jordan *et al.*, 2021).

 $(35.1 \pm 1.5 \text{ wt\%})$ . These changes are accompanied by increases in P.I. (~1.0 to 1.8), Cl (~0.1 to  $0.8 \pm 0.2$  wt%), and Zr (291 to 2060  $\mu$ g/g) with decreases in TiO<sub>2</sub> ( $\sim$ 0.8 to 0.5 wt%) and CaO ( $\sim$ 1.2 to 0.4 wt%). A regression through the Green Tuff whole-rock data suggests that at  $Qz_{40.5}$  values would be P.I. =  $1.68 \pm 0.1$ ;  $0.49 \pm 0.05$  wt%  $TiO_2$ ; 71.69 ± 0.44 wt%  $SiO_2$ ; 0.32 ± 0.08 wt% CaO; 1986 ± 102  $\mu$ g/g Zr;  $8.34 \pm 0.38$  wt% FeO<sup>T</sup>; and  $8.23 \pm 0.23$  wt% Al<sub>2</sub>O<sub>3</sub>. These values are plotted in Fig. 7 to serve as a reference point for comparison with the other suites; the regression parameters are in Supplementary Table 3. Previous workers have modelled this LLOD as the result of ~85-90% fractional crystallization of an assemblage dominated by anorthoclase (>90% of the fractionating assemblage) with subordinate augite, olivine, ilmenite, and apatite (Liszewska et al., 2018; Romano et al., 2019). The majority (n = 112) of published Green Tuff glasses and melt inclusions more evolved than this (viz., Qz\* > 39 wt%) have a nearly constant composition of  $Qz_{42.8\pm1.3}Or_{36.6\pm1.7}Ab_{20.4\pm1.7}$ , with  $71.9\pm0.1$  wt%  $SiO_2$ , 8.2 ± 0.5 wt% FeO<sup>T</sup>, 0.9 ± 0.2 wt% Cl, and P.I. = 1.9 ± 0.1, although Zr continues to increase from  $\sim$ 2000 to 3183  $\mu$ g/g (Liszewska et al., 2018; Romano et al., 2019). The exceptions are matrix glass analyses from the precursor air-fall deposit of the Green Tuff (sample 150542; Liszewska et al., 2018) that overlap with the Altura glass trend discussed below, which forms a 'horizontal trend' towards lower Or (to 23) at fairly constant Qz\*  $(44.1 \pm 0.4 \text{ wt\%})$  along the feldspar-quartz cotectic, accompanied by decreases in P.I. (to 1.6),  $FeO^T$  (to 7.4 wt%),  $TiO_2$  (to 0.4 wt%), and Cl (to 0.2 wt%) along with slight increases in SiO<sub>2</sub> (to 73.5 wt%) and Zr (to 2318  $\mu$ g/g). Liszewska et al. (2018) attributed this trend to co-fractionation of aenigmatite with quartz on the cotectic.

# Altura

Whole-rock compositions overlap with the Green Tuff in the 'thermal valley' in the Qz-Or-Ab projection from 13.6 to 40.6 wt% Qz\* at nearly constant Or ( $33.8 \pm 2.2$  wt%), but occupy different locations in FeO<sup>T</sup>-Al<sub>2</sub>O<sub>3</sub> space. Altura samples follow a trend that runs parallel to the Green Tuff, but with FeO<sup>T</sup> concentrations ~1.5 wt% lower at a given value of Al<sub>2</sub>O<sub>3</sub> and Qz\*. The Altura suite is represented by a single sample of comenditic trachyte

Group: Cala dell'Altura	Cala dell'Altura	Altura			0				Cala Gadir	ir		Cuddia Mida	Aida				
Sample ID:	060522	150531	060521	150533	171010-1	090532	150532	98 527	Unk34B	190532	Unk34A	090511	090512	060501	090514	060502	090513
Class:	CT	പ	പ	Ь	Ч	Ь	Ч	Ь	Ч	പ	Ч	Ч	പ	പ	പ	Ч	പ
SiO <sub>2</sub> , wt%	66.43	68.51	66.33	71.35	69.48	72.56	71.35	72.04	70.86	71.37	71.88	67.46	68.64	68.58	69.78	69.95	70.35
TiO <sub>2</sub>	0.53	0.37	0.34	0.37	0.36	0.37	0.36	0.38	0.39	0.38	0.38	0.46	0.45	0.38	0.33	0.33	0.33
$Al_2O_3$	13.83	9.25	8.35	9.20	8.93	9.13	8.88	8.68	9.25	9.24	9.34	9.52	9.32	8.95	8.37	8.29	8.18
$Fe_2O_3^T$	5.52	6.93	6.54	6.91	6.73	7.08	7.10	7.78	7.15	7.28	7.16	8.78	8.80	8.71	8.89	8.81	9.12
MnO	0.23	0.26	0.23	0.25	0.25	0.26	0.26	0.26	0.27	0.26	0.27	0.29	0.29	0.30	0.30	0.31	0.31
MgO	0.28	0.32	0.68	0.21	0.16	0.13	0.09	0.10	0.46	0.28	0.42	0.19	0.11	0.06	0.07	0.04	0.10
CaO	0.83	0.19	0.28	0.31	0.29	0.34	0.27	0.32	0.10	0.30	lbdl	0.91	0.62	0.52	0.44	0.39	0.66
Na <sub>2</sub> O	6.08	4.65	5.50	5.65	6.23	5.67	5.86	5.72	5.13	4.62	4.88	6.52	6.32	6.92	6.56	6.98	6.62
$K_2O$	4.71	4.84	4.46	4.48	4.38	4.42	4.51	4.23	4.32	4.46	4.43	4.49	4.40	4.53	4.33	4.46	4.36
$P_2O_5$	0.09	0.02	0.04	0.03	0.02	0.02	0.01	0.03	0.03	< 0.01	0.02	0.06	0.03	0.05	0.02	0.04	0.02
SO <sub>3</sub>	n.a.	n.a.	0.09	n.a.	0.14	0.02	n.a.	n.a.	0.34	n.a.	lbdl	0.05	0.04	0.03	0.04	0.04	0.04
F	0.08	0.10	0.18	n.a.	n.a.	0.23	0.25	n.a.	n.a.	n.a.	n.a.	0.20	0.17	0.23	0.23	0.25	0.24
C]	0.06	n.a.	1.44	n.a.	0.97	0.61	n.a.	n.a.	n.a.	n.a.	n.a.	0.56	0.56	0.70	0.72	0.68	0.78
$H_2O^+$	n.a.	n.a.	n.a.	n.a.	n.a.	< 0.1	n.a.	n.a.	n.a.	0.70	n.a.	0.3	0.5	n.a.	0.1	n.a.	0.3
$H_2O^-$	n.a.	n.a.	n.a.	n.a.	n.a.	< 0.1	n.a.	n.a.	n.a.	0.70	n.a.	< 0.1	< 0.1	n.a.	< 0.1	n.a.	< 0.1
IOI	0.46	4.10	6.75	0.70	2.15	-0.19	0.80	n.a.	0.45	1.70	0.62	-0.33	-0.77	0.00	-1.14	00.00	-0.87
Sum	99.13	99.54	101.22	99.46	100.08	100.65	99.74	99.54	98.74	101.29	99.29	99.45	99.48	99.97	99.04	100.57	100.54
O=F, Cl	0.05	I	0.40	I	0.22	0.23	I	I		I	I	0.21	0.20	0.25	0.26	0.26	0.28
Total	99.08	99.54	100.82	99.46	99.86	100.41	99.74	99.54	98.74	101.29	99.29	99.24	99.28	99.71	98.78	100.31	100.26
P.I.	1.09	1.39	1.66	1.54	1.68	1.55	1.64	1.61	1.42	1.34	1.37	1.64	1.63	1.82	1.85	1.97	1.91
*20	13.06	35.28	37.36	37.38	37.51	38.62	38.78	40.09	37.47	37.59	38.04	29.76	32.45	33.89	38.22	38.54	39.22
ı Y				5												)	
Be, μg/g	4	7	11	11	n.a.	13	00	n.a.	n.a.	14	n.a.	12	12	11	14	12	14
Sc	S	2	4	2	< 1.1	2.86	2	4	2.9	2	3.6	2.05	2.04	ŝ	1.51	2	1.64
V	v V	n.a.	۸ آ	n.a.	< 0.8	۸ آ	n.a.	S	0.2	ہ د	2.3	۲ د	^ ℃	∧ ℃	^ ℃	∧ ℃	V U
Cr	< 20	n.a.	< 20	n.a.	< 0.7	4.4	n.a.	< 20	< 0.7	< 20	< 0.7	4	20.7	< 20	13.8	< 20	5.6
Co	۸ 1	0.4	۸ 1	0.7	3.8	< 0.1	0.3	ŝ	8.1	V V	7.4	< 0.1	< 0.1	< 1 1	< 0.1	~ 7	< 0.1
Ni	2	<20	2	<20	<0.7	с	<20	< S	< 0.7	< 20	< 0.7	2	4	<1	2	<	ю
Cu	4	n.a.	ŝ	n.a.	<0.6	2	n.a.	n.a.	2.3	< 10	2.4	Ŋ	9	7	Ŋ	9	9
Zn	172	n.a.	334	n.a.	380.1	403	n.a.	468	370	390	403	379	388	401	462	417	455
Ga	31	33.1	33	33.2	35.3	n.a.	31.1	35	33.3	35	32.8	n.a.	n.a.	38	n.a.	36	n.a.
Ge	1.5	n.a.	2	n.a.	n.a.	n.a.	n.a.	3.4	n.a.	2.3	n.a.	n.a.	n.a.	2.3	n.a.	2.2	n.a.
As	ر ک	n.a.	∞	n.a.	7.9	11	n.a.	∞	n.a.	< V	n.a.	6	7	6	10	6	12
Se	n.a.	n.a.	n.a.	n.a.	n.a.	8.5	n.a.	n.a.	n.a.	n.a.	n.a.	5.6	9.2	n.a.	5.8	n.a.	9.1
Br	n.a.	n.a.	n.a.	n.a.	n.a.	19.3	n.a.	n.a.	n.a.	n.a.	n.a.	18.6	18.5	n.a.	22.5	n.a.	23.6
Rb	90	174.6	177	194.1	214.1	190	190.8	201	182	187	169	170	190	192	200	198	200
Sr	44	6.8	13	7.2	8.0	7	3.7	9	25.0	б	12.2	12	б	7	<sup>∞</sup>	7	12
Y	79.4	173.5	208.0	164	195.3	168	168.9	162.8	151	159	183	155	153	188.0	182	183.0	185
Zr	917	2105.2	1926	2028	2292.0	1928	2067.4	1910	2346	2117	2345	1771	1640	1876	1958	2045	1986
Nb	167	369.8	339	361.9	426.8	n.a.	366.3	369.0	431	322	430	n.a.	n.a.	384	n.a.	384	n.a.
																	(Continued)

Operation         Obstant         Solution         Operation         O	Table 2: Continued	ontinued																	
Stanpelly         6652         1503         17001         6652         15033         171001         6653         15033         171001         6653         15033         15033         6653 <th>Group:</th> <th>Cala del</th> <th>l'Altura</th> <th></th> <th></th> <th></th> <th></th> <th></th> <th></th> <th>Cala Gadir</th> <th></th> <th></th> <th>Cuddia N</th> <th>lida</th> <th></th> <th></th> <th></th> <th></th>	Group:	Cala del	l'Altura							Cala Gadir			Cuddia N	lida					
00         3         1a         15         na         13 </th <th>Sample ID:</th> <th>060522</th> <th>150531</th> <th>060521</th> <th>150533</th> <th>171010-1</th> <th>090532</th> <th>150532</th> <th>98 527</th> <th>Unk34B</th> <th>190532</th> <th>Unk34A</th> <th>090511</th> <th>090512</th> <th>060501</th> <th>090514</th> <th>060502</th> <th>090513</th>	Sample ID:	060522	150531	060521	150533	171010-1	090532	150532	98 527	Unk34B	190532	Unk34A	090511	090512	060501	090514	060502	090513	
Mat         2.3         n.a         n.a         0.3         n.a         n.a <th n.a<="" th="" tr<=""><th>Mo</th><th>ŝ</th><th>n.a.</th><th>15</th><th>n.a.</th><th>18.3</th><th>17</th><th>n.a.</th><th>n.a.</th><th>5</th><th>4</th><th>3</th><th>14</th><th>17</th><th>13</th><th>15</th><th>13</th><th>15</th></th>	<th>Mo</th> <th>ŝ</th> <th>n.a.</th> <th>15</th> <th>n.a.</th> <th>18.3</th> <th>17</th> <th>n.a.</th> <th>n.a.</th> <th>5</th> <th>4</th> <th>3</th> <th>14</th> <th>17</th> <th>13</th> <th>15</th> <th>13</th> <th>15</th>	Mo	ŝ	n.a.	15	n.a.	18.3	17	n.a.	n.a.	5	4	3	14	17	13	15	13	15
CI         R.1         CO         R.1         CO         R.1         R.1         R.1         R.1         R.1         R.1         R.2         R.3         R.3 <thr.3< th=""> <thr.3< th=""> <thr.3< th=""></thr.3<></thr.3<></thr.3<>	Ag	2.3	n.a.	3.3	n.a.	n.a.	3.0	n.a.	n.a.	n.a.	6.2	n.a.	3.0	3.1	3.4	3.2	3.1	3.5	
50         6         12         13         13         13         13         13         13         13         13         14         13           50         c.02         n.2         13         13         0.3         1	Cd	< 0.5	n.a.	< 0.5	n.a.	n.a.	0.9	n.a.	n.a.	n.a.	n.a.	n.a.	1.2	1.1	0.5	1.2	< 0.5	1.2	
50         < 0.2         n.a         < 0.3         n.a         0.8         0.6         0.6         0.2         0.3         0.6         0.2         0.3         0.6         0.2         0.3         0.6         0.2         0.3         0.3         0.6         0.2         0.3<	Sn	9	22	13	15	15.5	n.a.	17	n.a.	9	14	n.a.	n.a.	n.a.	14	n.a.	14	n.a.	
(a)         (b)         (c)         (c) <td>Sb</td> <td>&lt; 0.2</td> <td>n.a.</td> <td>&lt; 0.2</td> <td>n.a.</td> <td>n.a.</td> <td>0.8</td> <td>n.a.</td> <td>n.a.</td> <td>n.a.</td> <td>0.2</td> <td>n.a.</td> <td>0.8</td> <td>0.6</td> <td>&lt; 0.2</td> <td>0.9</td> <td>&lt; 0.2</td> <td>1.1</td>	Sb	< 0.2	n.a.	< 0.2	n.a.	n.a.	0.8	n.a.	n.a.	n.a.	0.2	n.a.	0.8	0.6	< 0.2	0.9	< 0.2	1.1	
Ba         543         33         41         311         37         34         111         37         34         311         37         36         439         30         123         83         133         112         70         63           Cia         277         420         4409         4721         358         4504         4509         73         371         71         70         63           Fr         382         257         640         4699         4721         73         566         461         538         271         724         579         571         77         566         73         372         74         77         70         63         77         70         53         540         77         70         53         540         77         70         57         77         566         73         373         713         70         74<	Cs	0.2	2	2.1	2.0	15.5	2.4	2.3	2.1	13.0	0.8	16.6	2	2.1	2.7	2.5	2.7	2.8	
In         171         2653         281         2291         234         2758         450         281         210         216         225         234         278           F         382         52.08         663         65.2         181.4         113         157.5         185.6         183.6         135.7	Ba	548	33	43	34	11.1	37	34	36	43.9	30	12.9	89	83	112	70	69	80	
Circ         21/1         21/2         64/0         44/3         47/2         34/3         64/3         55/2         64/3         52/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3         51/3         53/3 <th< td=""><td>La</td><td>171</td><td>266.3</td><td>281</td><td>238</td><td>229.1</td><td>234</td><td>237.8</td><td>250.44</td><td>201.9</td><td>240</td><td>258.1</td><td>210</td><td>216</td><td>252</td><td>244</td><td>278</td><td>250</td></th<>	La	171	266.3	281	238	229.1	234	237.8	250.44	201.9	240	258.1	210	216	252	244	278	250	
PF 382 25/8 609 46.99 na 781 419 12/7 16/76 1543 162 na 489 na na 75 26 371 294 409 152 122 202 202 123 258 237 136 155 131, 149 1727 1577 1577 1577 1577 1577 1577 1577	Ce	247	427.9	440	449.9	472.1	358	435.9	425.65	489.6	464	528.8	321	330	463	373	511	382	
Nid 1300 1256 206 1552 1814 149 1727 1577 1575 1543 152 1395 128 134 179 152 139 Eu 4.38 4.29 5.38 3.71 na 2.65 3.276 3332 na 3.17 na 2.52 26 3.71 29.4 409 Eu 4.38 6.56 420 na 1.49 17.5 1.68 1.39 0.8 3.3 2.66 na 2.59 na 1.32 4.49 5.31 4.33 5.49 Th 3.77 5.66 5.5 7.44 6.14 na na 3.3 5.66 na 5.59 na na 1.8 na 3.3 2 4.09 5.31 4.33 5.49 Th 161 2.30 2.31 3.8 2.20 na na 1.8 7 194 na 1.75 na na 1.8 25 6.5 7.4 1.3 2.4 Th 161 2.30 2.31 1.9 1.79 na na 2.65 1.79 1.6 1.79 1.6 1.79 1.5 1.8 1.2 4.4 1.3 5.8 Th 161 0.30 184 194 15.6 na 1.8 7 194 na 1.8 7 175 na na 1.8 25 0.3 4.4 3 5.2 Th 161 0.30 184 194 15.6 na 1.6 1.792 16.9 na 175 na na 1.4 1.4 1.7 17.8 16.9 198 Hi 1.4 2.8 2.2 59 na na 2.25 na na 2.25 na na 2.25 na 1.4 1.3 1.7 178 16.9 198 Hi 21.9 4.7 3 4.5 4.7 na 2.25 1.79 1.8 16.9 198 Hi 21.9 4.7 3 4.5 2.47 na 2.23 2.53 na 1.1 1.1 1.2 1.7 8 na 1.3 2.66 2.03 2.29 2.78 Hi 21.9 4.2 5.7 1.7 na 2.23 2.23 7.8 na 1.3 193 2.05 2.62 2.24 2.55 2.29 2.78 Hi 21.9 4.2 5.7 1.7 na 2.23 2.53 na 2.73 na 2.76 2.4 4.3 5.5 1.7 1.7 1.8 16.9 198 Hi 21.9 4.7 3 4.5 1.4 1.4 1.4 1.4 1.4 1.4 1.4 1.2 1.7 1.8 16.9 198 Hi 21.9 4.1 2.4 2.4 1.4 1.3 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.2 1.1 5.6 84 6.5 1.0 2.0 13 2.0 2.0 13 2.0 10. Hi 16.9 3.8 3.5 3.9 4.0 4.7 3.3 5.6 1.0 2.0 13 2.0 2.0 13 2.0 10. Hi 16.9 3.8 3.5 3.9 4.0 4.7 3.3 5.6 3.0 5.0 1.3 2.2 1.2 8 1.5 4.1 1.1 1.2 1.1 1.1	Pr	38.2	52.78	60.9	46.99	n.a.	n.a.	50.14	46.12	n.a.	48.9	n.a.	n.a.	n.a.	54.4	n.a.	57.7	n.a.	
Sm         274         3569         426         3195         na         332         na         317         na         25         25         371         294         409           Lu         423         538         371         na         337         333         na         335         549         409           Th         371         586         686         490         na         na         337         143         533         566         na         3597         na         na         335         658         559         na         357         na         335         653         na         357         357         357         357         357         357         357<	PN	130.0	192.6	206	165.2	181.4	149	172.7	167.76	154.3	162	189.6	128	134	179	152	192	154	
EI         438         429         538         371         na         339         333         na         332         549         533         543         543         543         543         543         543         543         543         543         543         543         543         543         543         543         543         543         533         535         744         614         na         307         na         na         333         355         na         333	Sm	27.4	35.69	42.6	31.95	n.a.	26.5	32.76	33.32	n.a.	31.7	n.a.	25	26	37.1	29.4	40.9	29.6	
Gd2003313892920nana30792957na2935na335na335na335Tb3575.668.500.0na14.35.335.66na5.67na376.244.36.58Tb3526.557.446.14nana175nananana203na274Tr1612.83526.557.446.14nana175nanana203na274Tr1612.83.56.557.446.14nana175nana203na274Th1612.83.56.56na17.55na175nana203108129Th1642.9415.60.0na175na141147178na221Th1442542.762.47na2.66na20316.9019332Th1142.542.762.492.60na2.75na2.292.78Th13102.242.47117na2.352.492.60na2.9332Th13112.444.24.514.14.514.117.816.919.8Th910.110.111.112.14.510.32.74<	Eu	4.38	4.29	5.38	3.71	n.a.	3.9	3.87	3.93	n.a.	3.83	n.a.	3.92	4.09	5.31	4.53	5.49	4.65	
Tb         3.77         5.68         6.86         4.90         na.         4.3         5.33         5.66         na.         5.04         na.         3.8         3.2         6.24         4.3         6.58           Dy         1327         5.56         7.44         6.14         na.         na.         7.43         6.14         na.         na.         7.43         6.14         na.         7.43         6.14         na.         7.43         6.14         na.         7.43         6.14         na.         7.43         6.53         7.44         na.         7.44         na.         7.44         na.         7.44         na.         7.44         na.         7.44         na.         7.45         na.         7.45         na.         7.45         na.         7.44         na.         7.45         1.45         na.         2.75         na.         2.	Gd	20.0	33.1	38.9	29.20	n.a.	n.a.	30.79	29.67	n.a.	29.8	n.a.	n.a.	n.a.	33.5	n.a.	33.2	n.a.	
Dy19732.863852970na.na.31643109na.307na.na.34na.37.8Ho3526.557.446.14na.na.6.60na.7.9316.62na.7.44Tm1612.83.22.59na.na.6.05na.17.35na.7.37.44Tm1612.83.22.59na.na.16.717.9216.99na.17.93na.2.95na.2.03Tm1612.83.22.59na.na.2.6617.9216.99na.17.816.912.98Lu14.42.542.562.47na.2.562.47na.2.552.472.653.55Lu21.94.734.55na.2.561.431.67na.2.522.292.78Hf21.94.734.55na.2.551.69na.2.551.6910.82.552.651.691.78Lu21.94.744.734.55na.2.551.691.691.781.691.983.55Lu21.94.744.734.592.667.622.292.782.492.562.492.562.643.57W21.99.44.734.544.734.597.781.691.791.672.663.57W	Tb	3.77	5.68	6.86	4.90	n.a.	4.3	5.33	5.66	n.a.	5.04	n.a.	3.8	3.2	6.24	4.3	6.58	4.6	
Ho $352$ $655$ $7.44$ $6.14$ na. na. $6.39$ $6.05$ na. $5.97$ na. na. na. $6.62$ na. $7.44$ Tr. $10,0$ $1907$ $2.18$ $7792$ na. $175$ na. $1175$ na. $117$ $1147$ $178$ $163$ $123$ $123$ $127$ Tr. $1216$ $129$ $128$ $1792$ $169$ $128$ $1778$ $163$ $129$ $128$ $129$	Dy	19.7	32.86	38.5	29.70	n.a.	n.a.	31.64	31.09	n.a.	30.7	n.a.	n.a.	n.a.	34	n.a.	37.8	n.a.	
Er 1040 19.07 218 1792 na. na. 18.77 19.44 na. 17.5 na. na. 203 na. 22.1 Tm 161 28 3.2 259 na. na. 286 2.90 na. 2255 na. na. 14.1 14.7 17.8 16.9 19.8 11.2 1.1.2 1.1.2 14.7 17.8 16.9 19.8 11.2 1.1.2 1.1.2 14.7 17.8 16.9 19.8 11.2 1.1.2 1.1.2 14.7 17.8 16.9 19.8 11.2 1.1.2 14.4 12.5 12.9 2.0.2 2.0.2 2.0.2 2.0.2 2.0.2 2.0.2 16.9 10.8 12.1 14.4 2.54 2.47 na. 22.6 2.49 2.60 na. 2.55 na. 17.8 16.9 19.8 16.9 19.8 16.1 21.9 47.3 4.5 2.74 2.77 na. 22.6 2.49 2.60 na. 22.7 na. 17.8 16.9 19.9 2.1.2 2.0.2	Но	3.52	6.55	7.44	6.14	n.a.	n.a.	6.39	6.05	n.a.	5.97	n.a.	n.a.	n.a.	6.62	n.a.	7.44	n.a.	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Er	10.40	19.07	21.8	17.92	n.a.	n.a.	18.77	19.44	n.a.	17.5	n.a.	n.a.	n.a.	20.3	n.a.	22.1	n.a.	
Yb         10.30         18.4         19.4         16.76         n.a.         16         17.92         16.99         n.a.         17.8         n.a.         14.1         14.7         17.8         16.9         19.8           Lu         1144         2.54         2.76         2.47         n.a.         2.26         2.49         2.60         n.a.         1.93         2.05         2.62         2.29         2.78           Hf         21.9         47.3         46.2         47.4         n.a.         2.33         7         n.a.         19.3         2.05         2.62         2.29         2.78           Ta         13.0         22.4         2.45         n.a.         4.5         n.a.         4.5         n.a.         1.7         1.7         1.7         1.7         1.7         1.9         2.05         2.62         2.29         2.78           Ta         13.0         22.4         2.3         1.a.         n.a.         1.3         1.7         1.a.         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3         1.3 </td <td>Tm</td> <td>1.61</td> <td>2.8</td> <td>3.2</td> <td>2.59</td> <td>n.a.</td> <td>n.a.</td> <td>2.68</td> <td>2.90</td> <td>n.a.</td> <td>2.55</td> <td>n.a.</td> <td>n.a.</td> <td>n.a.</td> <td>2.95</td> <td>n.a.</td> <td>3.27</td> <td>n.a.</td>	Tm	1.61	2.8	3.2	2.59	n.a.	n.a.	2.68	2.90	n.a.	2.55	n.a.	n.a.	n.a.	2.95	n.a.	3.27	n.a.	
In       1.44       2.54       2.76       2.47       n.a.       2.26       2.49       2.60       n.a.       2.52       n.a.       1.93       2.05       2.62       2.29       2.78         Hf       21.9       47.3       46.2       47.4       n.a.       45.8       47       46.52       n.a.       43.7       n.a.       3.9       2.05       2.62       2.29       2.78         Ta       21.9       47.3       46.5       n.a.       45.8       47       46.52       n.a.       43.7       n.a.       39       40.4       47.3       45.5       51         W       2.5       4.4       2.1       n.a.       n.a.       15.3       n.a.       16       n.a.       0.1       n.a.       0.2       13.0       20.5       56.5       30.5         W       2.5       1.4       2.1       0.1       n.a.       16       n.a.       n.a.       17       11       13       19       11       10       20       39.5       20       39.2         Bi       <0.1	Yb	10.30	18.4	19.4	16.76	n.a.	16	17.92	16.99	n.a.	17.8	n.a.	14.1	14.7	17.8	16.9	19.8	17.1	
Hf         21.9         47.3         46.2         47.4         n.a.         45.8         47         46.52         n.a.         43.7         n.a.         39         40.4         47.3         45.9         51           Ta         13.0         22.4         24.7         21.7         n.a.         23.9         22.3         23.7         n.a.         20.8         n.a.         27.4         27.3         26.6         30.5           W         2.5         4.4         4.2         4.5         n.a.         n.a.         0.13         n.a.         23.5         n.a.         20.6         30.5         30.5           W         2.5         4.4         4.2         4.5         n.a.         n.a.         15.3         11         n.a.         0.15         n.a.         0.2         1.4         4.8         1         5.7           Pb         9         n.a.         0.4         n.a.         1.5         1.1         1.1         1.0         1.1         1.0         2.0         1.3         2.0         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5         3.5<	Lu	1.44	2.54	2.76	2.47	n.a.	2.26	2.49	2.60	n.a.	2.52	n.a.	1.93	2.05	2.62	2.29	2.78	2.35	
Ta       13.0       22.4       24.7       21.7       n.a.       23.9       22.3       23.7       n.a.       20.8       n.a.       22.6       22.4       27.3       26.6       30.5         W       2.5       4.4       4.2       4.5       n.a.       n.a.       n.a.       3.7       n.a.       2.1       4.8       <1	Hf	21.9	47.3	46.2	47.4	n.a.	45.8	47	46.52	n.a.	43.7	n.a.	39	40.4	47.3	45.9	51	47.4	
W       2.5       4.4       4.2       4.5       n.a.       -1       4.5       n.a.       n.a.       3.7       n.a.       -1       -1       4.8       <1	Та	13.0	22.4	24.7	21.7	n.a.	23.9	22.3	23.7	n.a.	20.8	n.a.	22.6	22.4	27.3	26.6	30.5	26.7	
T1         0.1         n.a.         0.13         n.a.         n.a.         n.a.         n.a.         n.a.         n.a.         n.a.         0.2         n.a.         0.26           Pb         9         n.a.         16         n.a.         15.3         11         n.a.         n.a.         18         13         19         11         10         20         13         20           Bi         < 0.1         n.a.         0.4         n.a.         n.a.         5.3         39.3         40.1         38.0         38.1         38.0         43.1         35.1         44.1         33.7         33.6         36.4         39.6         39.2           Th         16.9         38.7         35.3         39.3         43.1         35.1         44.1         33.7         33.6         36.4         39.6         39.2           U         3.65         11.6         11.1         12         11.1         12.1         5.6         8         6.5         10.2         10.2         13.2         13.2         12.8         15.4           La <sub>N</sub> /Yb <sub>N</sub> 11.3         9.8         9.6         5.1         12.1         10.1         10.2         10.2         10.2	W	2.5	4.4	4.2	4.5	n.a.	< 1	4.5	n.a.	n.a.	3.7	n.a.	< 1	< 1	4.8	< V	5.7	< 1	
Pb         9         n.a.         15.3         11         n.a.         n.a.         18         13         19         11         10         20         13         20           Bi         < 0.1         n.a.         0.4         n.a.         n.a. $< 2$ n.a. $< 33.7$ $35.1$ $41.1$ $33.7$ $33.6$ $36.4$ $39.6$ $39.2$ $0.4$ Th         16.9 $38.7$ $35.3$ $39.3$ $40.1$ $38.9$ $38.43$ $38.03$ $43.1$ $35.1$ $44.1$ $33.7$ $33.6$ $36.4$ $39.6$ $39.2$ U $3.65$ $11.6$ $11.4$ $11.6$ $11.1$ $12.1$ $12.1$ $5.6$ $8.6$ $5.1$ $10.2$ $10.2$ $13.2$ $12.8$ $15.4$ Law/Ybn $11.3$ $9.8$ $9.6$ $ 9.0$ $10.0$ $ 9.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$ $12.1$	T]	0.1	n.a.	0.13	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	0.15	n.a.	n.a.	n.a.	0.2	n.a.	0.26	n.a.	
Bi         < 0.1         n.a.         0.4         n.a.         n.a.         < 2         n.a.         < 2         0.4            Th         16.9         38.7         35.3         39.3         40.1         38.9         38.4         38.03         43.1         35.1         44.1         33.7         33.6         36.4         39.6         39.2           U         3.65         11.6         11.4         11.1         12         11.1         12.1         5.6         8         6.5         10.2         10.5         13.2         12.8         15.4           La <sub>N</sub> /Yb <sub>N</sub> 11.3         9.8         9.6         -         9.9         9.0         10.0         -         9.5         13.2         12.8         15.4           La <sub>N</sub> /Yb <sub>N</sub> 11.3         9.8         9.6         -         9.9         9.0         10.0         -         9.2         12.8         15.4           Zr/Nb         5.49         5.69         5.68         5.60         5.37         -         5.45         -         -         4.89         -         5.33           Another obsects1 sector 2.9 \$527,060521,060501, and 060501 were previously presented by White et al. (2009), except for Ag, Bi, TI, W, Sb, Sn, a	Pb	6	n.a.	16	n.a.	15.3	11	n.a.	n.a.	18	13	19	11	10	20	13	20	14	
Th       16.9       38.7       35.3       39.3       40.1       38.9       38.4       38.03       43.1       35.1       44.1       33.7       33.6       36.4       39.6       39.2         U       3.65       11.6       11.4       11.6       11.1       12       11.1       12.1       5.6       8       6.5       10.2       10.5       13.2       12.8       15.4         La <sub>N</sub> /Yb <sub>N</sub> 11.3       9.8       9.6       -       9.9       9.0       10.0       -       9.2       -       10.1       10.0       9.6       9.8       9.5         Zr/Nb       5.49       5.69       5.68       5.60       5.37       -       5.44       6.57       5.45       -       -       5.33         Samples 060522; 98527, 060521, and 060501, and 060501 were previously presented by White et al. (2009), except for Ag, Bi, TI, W, Sb, Sn, and As. Samples 17101-1, Unk34A, and Unk34B are WD-KRF analyses fro	Bi	< 0.1	n.a.	0.4	n.a.	n.a.	< 2	n.a.	n.a.	n.a.	< 0.1	n.a.	< 2	< 2	0.3	< 2	0.4	< 2	
U       3.65       11.6       11.4       11.6       11.1       12.1       12.1       5.6       8       6.5       10.2       10.5       13.2       12.8       15.4         Law/Yb <sub>N</sub> 11.3       9.8       9.6       -       9.9       9.0       10.0       -       9.6       9.8       9.5         Zr/Nb       5.49       5.69       5.68       5.60       5.37       -       5.44       6.57       5.45       -       -       4.89       -       5.33         Samples 060522; 98527; 060521; and 060501; and 060500; and 060501; and 060501; and 060501; and 060501; an	Th	16.9	38.7	35.3	39.3	40.1	38.9	38.4	38.03	43.1	35.1	44.1	33.7	33.6	36.4	39.6	39.2	39.8	
La <sub>N</sub> /Yb <sub>N</sub> 11.3       9.8       9.6       -       9.9       9.0       10.0       -       9.2       -       10.1       10.0       9.6       9.8       9.5         Zr/Nb       5.49       5.69       5.60       5.37       -       5.64       5.18       5.44       6.57       5.45       -       -       4.89       -       5.33         Samples 060522; 98527; 060521, 060501, and 060501 were previously presented by White et al. (2009), except for Ag, Bi, TI, W, Sb, Sn, and As. Samples 171010-1, Unk34A, and Unk34B are WD-XRF analyses from the state and	Ŋ	3.65	11.6	11.4	11.6	11.1	12	11.1	12.1	5.6	∞	6.5	10.2	10.5	13.2	12.8	15.4	12.7	
Zr/Nb         5.49         5.60         5.37         —         5.64         5.18         5.44         6.57         5.45         —         —         4.89         —         5.33           Samples 060522, 98527, 060521, and 060501 were previously presented by White et al. (2009), except for Ag, Bi, TI, W, Sb, Sn, and As. Samples 171010-1, Unk34A, and Unk34B are WD-XRF analyses fro         5.33	La <sub>N</sub> /Yb <sub>N</sub>	11.3	9.8	8.0	9.6		9.9	9.0	10.0		9.2		10.1	10.0	9.6	9.8	9.5	9.9	
Samples 060522, 98527, 060501, and 060501 were previously presented by White et al. (2009), except for Ag, Bi, Tl, W, Sb, Sn, and As. Samples 171010–1, Unk34A, and Unk34B are WD-XRF analyses fro	Zr/Nb	5.49	5.69	5.68	5.60	5.37		5.64	5.18	5.44	6.57	5.45		I	4.89		5.33		
(2014), UIK34A and UIK34B WERE COLECTED DY K. WILINAILIS. CLASS. CJ, CONTETINUEL ITACHYUE, T, FAINELILE (MACCUMIAIN, 12/4, LE MAULE, 2002). WIAI ILVII TEPUTEU AS FE203, LOVI, LOSS VII 18111011, 11.4, ILC	Samples 060: (2014); Unk34	522, 98527, ( A and Unk3	60521, 06050 4B were coll	11, and 0605( ected by R. V	01 were prev Villiams. Cla	//////////////////////////////////////	ted by White ditic Trachy	e et al. (2009) rte; P, Pantell	, except for erite (Macd	Ag, Bi, Tl, W lonald, 1974;	, Sb, Sn, and Le Maitre, 2	l As. Sample: 2002). Fe <sub>2</sub> O <sub>3</sub> T	s 171010–1, , total iron	Unk34A, an reported as I	d Unk34B a Fe2O3; LOI, J	re WD-XRF ( Loss on Igni	analyses fro ion; n.a., no	n Jordan t	

(2014), UnK34A and UnK34B were collected by R. Williams. Class: CI; Comenditic Trachyte, P. Pantellente (Macdonald, 1974; Le Mattre, 2002), Fe<sub>2</sub>O<sub>3</sub><sup>-1</sup>, total analyzed, bdl, below detection limit; P.I. (Peralkalinity Index) = mol Na + K/Al; La<sub>N</sub>/Yb<sub>N</sub>, chondrite-normalized (McDonough & Sun, 1995) La/Yb ratios. Qz\* = 100•Qz<sub>n</sub>/(Qz<sub>n</sub> + Or<sub>n</sub> + Ab<sub>n</sub>), CIPW norm calculated following Kelsey (1965) with iron recalculated following Le Maitre (1976).

	Table 3:	Representative	matrix g	glass	analyses
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Group:	Cala de	ll'Altura	a				Cala Ga	adir	Cuddia	Mida								
Sample ID:	150532		150533		150531		190532		090512								090514	
Point:	#6	#11	#6	#1	#1	#11	13 / 1.	18/1.	79/1.	11/1.	20/1.	21/1.	83/1.	77 / 1.	95 / 1.	94 / 1.	16/1	8/1
SiO <sub>2</sub> , wt%	73.51	74.11	72.40	75.68	72.48	72.77	67.61	72.34	68.90	70.87	71.00	71.39	69.73	69.38	68.82	67.67	72.56	70.71
TiO <sub>2</sub>	0.44	0.39	0.39	0.28	0.38	0.41	0.51	0.45	0.62	0.54	0.40	0.44	0.59	0.74	0.66	0.73	0.31	0.25
Al <sub>2</sub> O <sub>3</sub>	8.39	8.57	8.84	8.70	8.72	8.41	10.15	8.15	9.13	6.94	6.74	6.21	5.84	6.07	5.16	4.71	7.53	7.19
FeOT	6.17	6.93	6.79	5.89	6.41	6.61	7.83	7.80	8.00	9.57	8.64	9.39	10.76	11.00	11.68	12.98	8.54	8.21
MnO	0.25	0.29	0.27	0.00	0.32	0.23	0.00	0.34	0.33	0.37	0.33	0.37	0.44	0.43	0.48	0.51	0.32	0.31
MgO	0.09	0.09	0.09	0.07	0.08	0.07	0.55	0.13	0.06	0.04	0.04	0.05	0.06	0.07	0.10	0.08	0.06	0.04
CaO	0.28	0.14	0.07	0.16	0.31	0.26	0.70	0.35	0.39	0.32	0.41	0.43	0.51	0.53	0.60	0.59	0.37	0.32
Na <sub>2</sub> O	5.88	5.93	5.88	5.92	5.40	5.38	3.95	4.49	6.60	6.29	6.04	6.19	5.92	5.70	6.48	5.94	5.56	7.21
K <sub>2</sub> O	4.40	4.54	4.99	3.60	4.22	3.84	5.51	4.44	5.15	4.39	4.76	4.89	4.94	4.46	4.67	4.76	4.24	4.33
P <sub>2</sub> O <sub>5</sub>	bdl	bdl	bdl	bdl	bdl	bdl	0.00	0.03	0.02	0.00	0.02	0.02	0.00	0.09	0.03	0.04	0.03	0.01
SO <sub>3</sub>	0.04	bdl	0.03	bdl	0.03	0.05	0.01	0.04	0.10	0.09	0.07	0.08	0.03	0.12	0.07	0.07	0.06	0.00
ZrO <sub>2</sub>	n.a.	n.a.	n.a.	n.a.	0.30	0.40	0.97	0.46	0.26	0.42	0.38	0.34	0.46	0.31	0.50	0.48	0.34	0.43
F	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	0.40	n.a.	0.43	0.10	0.23	0.35	0.23	0.26	0.40	0.36	0.10	n.a.
Cl	0.31	0.60	0.49	0.57	0.68	0.81	0.05	0.07	0.89	0.90	0.89	0.94	1.11	1.02	1.15	1.28	0.87	0.88
Sum	99.76	101.59	100.24	100.87	99.03	98.84	97.26	99.10	100.88	100.85	99.58	100.74	100.14	99.86	100.30	99.71	100.88	99.44
O=F, Cl	0.07	0.14	0.11	0.13	0.15	0.18	0.18	0.02	0.38	0.25	0.30	0.36	0.35	0.34	0.43	0.44	0.24	0.20
Total	99.69	101.45	100.13	100.74	98.88	98.66	97.08	99.08	100.50	100.61	99.28	100.38	99.79	99.52	99.87	99.27	100.63	99.24
P.I.	1.72	1.71	1.71	1.57	1.54	1.55	1.23	1.50	1.81	2.18	2.24	2.49	2.58	2.34	3.05	3.17	1.82	2.30
Qz*	43.52	42.51	39.86	44.31	41.34	43.31	29.04	43.58	33.90	46.74	48.43	51.11	51.40	50.17	54.27	56.17	45.92	45.45
CIPW Norm																		
Qz	34.58	33.38	31.32	36.43	33.15	34.71	22.84	37.04	25.17	32.83	34.62	35.32	33.90	33.40	33.78	33.54	34.22	32.48
Or	26.13	26.54	29.54	21.19	25.33	23.13	33.58	31.17	30.64	26.09	28.53	29.04	29.49	26.75	27.93	26.18	25.11	25.91
Ab	18.75	18.60	17.72	24.60	21.70	22.30	22.25	22.24	18.43	11.32	8.33	4.75	2.57	6.43	0.53	0.00	15.19	13.08
Ac	3.04	3.33	2.95	3.36	3.41	3.74	3.79	2.01	2.62	4.10	3.68	3.83	4.34	4.94	4.32	5.06	4.39	3.03
Nms	6.47	6.34	6.70	5.00	4.85	4.62	1.85	2.06	8.11	8.74	9.17	10.13	10.04	8.58	11.66	10.58	6.29	10.54
Di	1.24	0.61	0.31	0.70	1.39	1.17	3.13	0.14	1.64	1.41	1.73	1.76	2.27	1.81	2.50	2.38	1.48	1.38
Hy	8.95	10.48	10.72	8.19	9.44	9.54	11.56	5.20	12.16	14.49	13.12	14.29	16.26	16.45	17.96	20.08	12.67	13.08
Il	0.84	0.73	0.74	0.53	0.73	0.79	1.00	0.15	1.18	1.02	0.78	0.83	1.13	1.42	1.26	1.40	0.59	0.47
Ар	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.04	0.01	0.04	0.05	0.00	0.20	0.07	0.10	0.06	0.02

FeO<sup>T</sup>, total iron reported as FeO; LOI, Loss on Ignition; n.a., not analyzed; bdl, below detection limit; P.I. (Peralkalinity Index) = mol Na + K/Al;  $Qz^* = 100 \cdot Qz_n/(Qz_n + Or_n + Ab_n)$ . CIPW norm calculated following Kelsey (1965) with iron oxides adjusted following Le Maitre (1976).

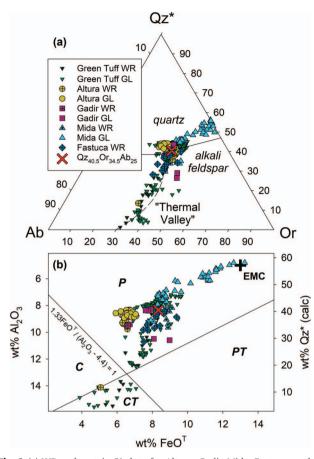
(67.8 wt% SiO<sub>2</sub>; 5.07 wt% FeO<sup>T</sup>) and several pantellerites (71.7-73.1 wt% SiO<sub>2</sub>; 6.1–7.1 wt% FeO<sup>T</sup>), with systematically higher SiO<sub>2</sub> and lower P.I. (1.1-1.7) than those seen in the Green Tuff for a given value of Qz\* over a similar range of Zr (917–2292  $\mu$ g/g). Altura matrix glasses have a nearly constant composition of  $Qz_{42.4\pm1.3} Or_{32.2\pm2.8} Ab_{25.4\pm2.1},$  with P.I. = 1.7  $\pm$  0.1, 73.7  $\pm$  0.5 wt%  $SiO_2$ , 6.6 ± 0.2 wt% FeO<sup>T</sup>, 0.39 ± 0.04 TiO<sub>2</sub>, and 0.22 ± 0.07 wt% CaO, although Zr continues to increase to 2961  $\mu$ g/g. Cl compositions range from 0.81 to 0.25 wt% and are uncorrelated with any other component, suggesting that this variation may be due to Cl loss during degassing or devitrification.

#### Gadir

The three samples from the Gadir volcanic centre are compositionally similar to Altura (~73.1 wt%  $SiO_2$ ; ~6.6 wt%  $FeO^T$ ) and plot near the experimental minimum in the Qz-Or-Ab projection (Qz<sub>38.1</sub>Or<sub>33.3</sub>Ab<sub>28.6</sub>). Considering that they occupy the same stratigraphic interval, albeit on different sides of the island, this suggests that there may be some petrogenetic relationship between the two suites. Glass in the Gadir dome is highly devitrified, which may account for its uniformly low Cl contents and should suggest some caution with this sample. Zr concentrations increase from  $\sim$ 2000 to 3500  $\mu$ g/g.

# Mida

Whole-rock samples from Mida plot along the 'thermal valley' projected on the Qz-Or-Ab plane, extending from 29.9 to 40.9 wt%  $Qz^*$  at nearly constant Or  $(35.1 \pm 1.5 \text{ wt\%})$ . The analyses are compositionally similar to the Green Tuff with respect to SiO<sub>2</sub> (69.0-70.4 wt%), Zr (1640-2550 µg/g), and CaO (0.40-0.93 wt%) but have slightly higher P.I. (1.2–2.0) and FeO<sup>T</sup> (7.8–8.2 wt%), with lower  $TiO_2$  (0.47-0.32 wt%) and Cl (0.67±0.09 wt%) at a given value of Qz\*. Glasses in Mida vary by sample. Glass compositions are fairly homogenous in sample 090514, with a composition of  $Qz_{45.9\pm1.1}Or_{35.8\pm0.9}Ab_{18.6\pm0.9}$ , 71.7±0.5 wt%  $SiO_2$ , 8.5 ± 0.2 wt% FeO<sup>T</sup>, P.I. = 2.3 ± 0.2, and 2505 ± 446  $\mu$ g/g Zr. Glass compositions in sample 090512, however, demonstrate an extreme differentiation trend: Qz\* increases from 45 to 56 wt% as Ab decreases to zero, P.I. increases from 2.0 to 3.5, Zr increases from  $\sim$ 2000 to 4000  $\mu$ g/g, CaO and SiO<sub>2</sub> both reverse their whole-rock trends and increase to 0.6 wt% and decrease to 69 wt%, respectively, and with the increase in  $Qz^*$ ,  $Al_2O_3$ decreases to 4.8 wt% and FeOT increases to 13.2 wt%. The most extreme compositions observed are very similar to those recorded in extremely fractionated magmas at Boseti, Ethiopia, which may, as noted earlier, represent the effective compositional minimum (EMC) composition for peralkaline silicic rocks



**Fig. 6.** (a) WR and matrix GL data for Altura, Gadir, Mida, Fastuca, and the GT plotted on an isobaric (100 MPa) projection from X (=  $H_2O$ +  $Na_2SiO_3$  + NaFeSi<sub>2</sub>O<sub>6</sub>) of the alkali feldspar-quartz cotectic (solid line), location of the 'thermal valley' (dashed line), and ternary minimum (**X**) of the 8.3 wt% Ns + 8.3 wt% Ac plane in the peralkaline haplogranite system Q-Or-Ab-X following Carmichael & MacKenzie (1963). Serra della Fastuca analyses are from Rotolo *et al.* (2007) and White *et al.* (2009). GT WR data are from Williams *et al.* (2014) and Liszewska *et al.* (2018). Normative components were calculated following Kelsey (1965) with iron oxides distributed following Le Maitre (1976) (see Supplementary Table S2). (b) Modified Macdonald (1974) classification scheme for peralkaline rocks. P. Pantellerite; PT, Pantelleritic Trachyte; C, Comendite; CT, Comenditic Trachyte; EMC, Effective Minimum Composition (Macdonald *et al.*, 2012).

(Macdonald et al., 2012). Cl in the glasses also shows a systematic increase from 0.8 to 1.3 wt as Qz\* increases, possibly due to its higher solubility in more peralkaline and iron-rich melts (Webster et al., 2015, 2020), the absence of a buffering exsolved vapor phase (contra Lowenstern, 1994), or both.

## Mineral compositions Feldspar

Representative feldspar analyses are presented in Table 4 and the complete data set is presented in Supplementary Table S4 and Fig. 8, which also includes analyses from the Green Tuff (Liszewska *et al.*, 2018). The total range is from *anorthoclase* (Or < 37 mol%) to *sanidine* (Or  $\geq$  37 mol%). Alkali feldspar compositions in all four suites are broadly similar. In the Green Tuff, alkali feldspars show a similarly continuous range of compositions with An decreasing from ~0.8 to 0.1 mol% as Or increases from ~20 to 35 mol%, after which it maintains a composition of 35.2 ± 1.0 mol% Or as An continues to decrease from 0.1 mol% to below the CaO detection limit (plotted as zero in Fig. 8b). Alkali feldspar in the sample of comenditic trachyte from Altura (060522) is compositionally zoned over a similar range from An<sub>5</sub>Or<sub>22</sub> to An<sub>0.5</sub>Or<sub>33</sub>, while alkali feldspars in the pantellerites have a nearly constant composition of <0.10 mol% An and  $37.3 \pm 1.1$  mol% Or. Anorthoclase from the Gadir pantellerite is similar to the Green Tuff, with a nearly constant composition of An <0.25 mol% and 35.5±1.3 mol% Or. Zonation within crystals is generally very small (< Or<sub>2-5</sub>) and contents of Fe<sub>2</sub>O<sub>3</sub> are low (<1.5 wt%). Although one sample from Mida (090514) has anorthoclase zoned from An<sub>20</sub>Or<sub>10</sub> to An<sub>5</sub>Or<sub>28</sub>, most samples also have rather monotonous compositions of <0.20 mol% An and  $37.2 \pm 1.1$  mol% Or. In the most extreme feldspars (<0.01 mol%) An), average Or compositions for the Green Tuff ( $35.2 \pm 1.0 \text{ mol}\%$ ) and Gadir ( $35.5 \pm 1.5$  mol%) are similar, Altura feldspars are similar but slightly higher  $(37.1 \pm 1.1 \text{ mol}\%)$ , and Mida are the highest  $(39.7 \pm 2.2 \text{ mol}\%)$ .

Arzilli et al. (2020) described a pressure- and H<sub>2</sub>O-independent relationship between temperature and Or content in feldspar in peralkaline silicic melts, with compositions of >37 mol% Or (i.e. the anorthoclase-sanidine boundary) crystallizing at <720°C and compositions between 31–37 mol% Or crystallizing at ≥720°C. A comparison with our data suggests higher temperatures for the Green Tuff, lower temperatures for Mida, and intermediate temperatures for Altura and Gadir.

#### Clinopyroxene

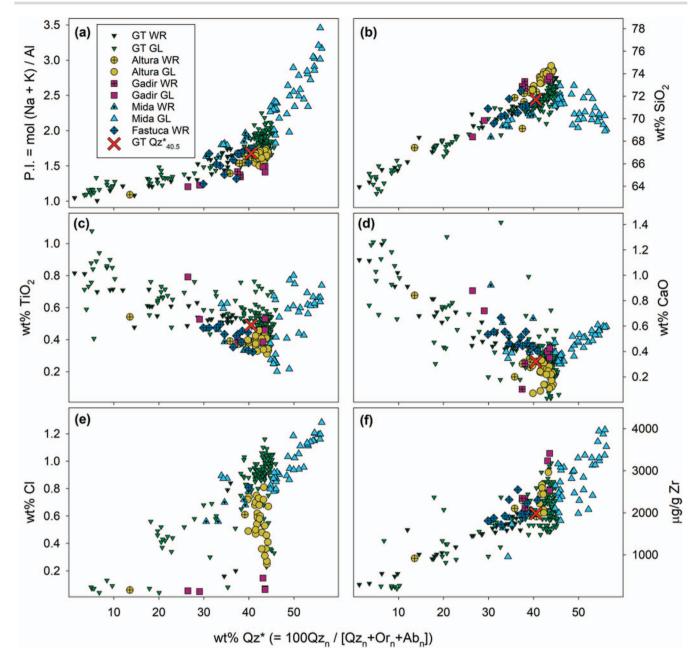
Representative clinopyroxene analyses are presented in Table 5 and the complete data set in Supplementary Table S5 and Fig 9. Clinopyroxene analyses from the comenditic trachyte sample from Altura are similar to those from the comenditic trachyte section of the Green Tuff and classify as *augites* (En<sub>19-21</sub>Fs<sub>38-41</sub>Wo<sub>39-41</sub>). Clinopyroxenes from the Altura and Gadir pantellerites have compositions similar to each other and to the most evolved samples from the Green Tuff  $(En_{7-11}Fs_{50-56}Wo_{36-41})$  with all points but one classifying as sodian augites (Morimoto, 1988), and with Na increasing from 0.13 to 0.22 atoms per formula unit (apfu) as X-Fe<sub>tot</sub>  $(= mol Fe^{T} / [Fe^{T} + Mg])$  increases from 0.82 to 0.90. In contrast, clinopyroxenes from Mida are distinctly Fe- and Na-rich, classifying as sodian augites and aegirine-augites (En<sub>3-4</sub>Fs<sub>53-59</sub>Wo<sub>39-43</sub>), with Na increasing from 0.11 to 0.30 apfu as X-Fetot increases from 0.92 to 0.95. Previous work (Larsen, 1976; Di Carlo et al., 2010; Romano et al., 2020) has shown that higher values of X-Fetot indicate both lower temperatures and lower relative oxygen fugacities.

## Olivine

Representative olivine analyses are presented in Table 6 and the complete data set is presented in Supplementary Table 6. Olivine was not observed in the Gadir pantellerite and at Altura is found only as a phenocryst (Fa<sub>84</sub>) in the comenditic trachyte. Unlike in other suites on Pantelleria such as the Green Tuff where it occurs in rocks with P.I. < 1.6 with compositions of Fa<sub>82-94</sub> (Liszewska *et al.*, 2018; Romano *et al.*, 2020), olivine at Mida is found in rocks with a P.I. of up to ~1.8 and are classified as *fayalites* with a composition of Fa<sub>95-98</sub>.

## Aenigmatite

Representative aenigmatite analyses are presented in Table 7 and the complete data set in Supplementary Table S7. Aenigmatite is found in all samples except the Altura comenditic trachyte. Aenigmatites show very little variation with respect to X-Ti (= mol Ti/[Ti + Fe<sup>T</sup>]), which has a nearly constant value of  $0.17 \pm 0.01$  typical in peralkaline silicic rocks (Nicholls & Carmichael, 1969). However, there is a small amount (~2%) of the substitution <sup>VIII</sup>Ca<sup>2+</sup>



**Fig. 7.** Compositional variation diagrams that plot (a) P.I., (b) wt% SiO<sub>2</sub>, (c) wt% TiO<sub>2</sub>, (d) wt% CaO, (e) wt% Cl, and (d)  $\mu$ g/g Zr versus wt% Qz\* (= 100·Qz<sub>n</sub> / [Qz<sub>n</sub> + Ab<sub>n</sub> + Or<sub>n</sub>]). GT data from Williams (2010), Liszewska et al. (2018) and Romano et al. (2019); Fastuca data are from Rotolo et al. (2007) and White et al. (2009). The red **X** symbol indicates the location of the projected intersection of the 'thermal valley' with the Carmichael & Mackenzie (1963) minimum on the alkali feldspar-quartz cotectic at Qz\* = 40.5 (see Supplementary Table S3 for details.)

+  $^{VI}Al^{3+} = ^{VII}Na^+ + {}^{IV}Si^{4+}$  within the individual series (Fig. 10), with aenigmatite from the Green Tuff generally being more calcic (Ca<sup>2+</sup> + Al<sup>3+</sup> = 0.15–0.25), which may have allowed it to stabilize at a slightly higher temperature and/or higher oxygen fugacity (Hodges & Barker, 1973; Larsen, 1977; Gaeta & Mottana, 1991).

# Fe-Ti oxides

Representative oxide analyses are presented in Table 8 and the complete set of analyses in Supplementary Table S8. Oxides were observed in all but one sample from Altura and one from Mida and consist dominantly of *ilmenite*, although *magnetite* is the sole oxide in the Altura comenditic trachyte (Usp<sub>66</sub>) and in a single sample from Mida (090513; Usp<sub>75</sub>). Ilmenite ranges in composition from

 $\rm Ilm_{95-98}$  at Altura,  $\rm Ilm_{96-97}$  at Gadir, and  $\rm Ilm_{96-99}$  at Mida, which all overlap with compositions in the Green Tuff ( $\rm Ilm_{95-98}).$ 

#### Amphibole

Representative amphibole analyses are presented in Table 9 and the complete data set in Supplementary Table S9 and Fig. 11. All amphiboles classify as *ferrorichterites* in the scheme of Hawthorne *et al.* (2012). Amphiboles from Altura and Gadir have similar compositions with respect to Fe and Na, but the Altura amphiboles have much higher fluorine concentrations (1.38–1.87 wt%; ~0.66–0.92 apfu) than Gadir (0.47–0.98 wt%; ~0.22–0.48 apfu) and Mida (0.47–0.57 wt%; ~0.24–0.29 apfu), although chlorine concentrations are similar in each series (0.05–0.07 wt%). As with the clinopyroxenes, amphiboles from Mida are higher in Na<sub>B</sub>

Table 4:	Representative	alkali feldspa	ar analyses
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Group:	Cala dell	'Altura				Gadir		Cuddi	a Mida							
Sample:	060522		150532	150533	98527	190532		090514	1	090511		090512	060502	1	060502	2
Point:	#18	#67	21 /1.	2/4.	#8	40 / 10.	38 / 14.	7/1.	52/1.	15 / 1.	5/1.	5/1.	#103	#19	#36	#38
Class:	Anr	Anr	Anr	Anr	San	Anr	Anr	Anr	Anr	Anr	Anr	San	Anr	San	San	San
SiO <sub>2</sub> , wt%	66.41	65.72	67.38	67.41	67.89	67.61	67.91	64.19	67.90	67.86	67.50	66.03	65.47	66.75	66.41	65.74
$Al_2O_3$	20.49	18.99	17.51	17.88	18.40	18.58	18.43	21.16	17.45	18.42	17.43	17.19	19.70	18.00	18.50	18.05
FeO <sup>T</sup>	bdl	bdl	1.04	0.95	0.87	0.79	0.78	0.30	0.93	0.23	1.32	1.00	0.36	1.08	0.86	1.14
CaO	0.92	0.09	bdl	bdl	0.00	0.04	0.01	3.64	0.004	0.50	0.01	0.02	0.39	0.01	0.01	0.00
Na <sub>2</sub> O	8.20	7.51	7.40	7.39	7.00	7.67	7.29	7.82	7.26	7.17	7.12	7.05	7.16	7.16	7.16	6.17
K <sub>2</sub> O	3.83	5.60	6.20	6.47	6.53	5.95	6.38	2.16	6.42	5.43	6.20	6.36	5.85	6.47	6.16	7.38
Total:	99.85	97.92	99.54	100.10	100.69	100.65	100.80	99.27	99.97	99.62	99.57	97.66	98.92	99.47	99.09	98.48
Formula	(MT4O8)															
Si	2.94	2.99	3.03	3.02	3.02	3.00	3.01	2.866	3.04	3.02	3.03	3.03	2.96	3.01	3.00	3.00
Al	1.09	1.02	0.93	0.94	0.96	0.97	0.96	1.113	0.92	0.97	0.92	0.93	1.05	0.96	0.98	0.97
Fe <sup>3+</sup>	0.00	0.00	0.04	0.04	0.03	0.03	0.03	0.011	0.03	0.01	0.05	0.04	0.01	0.04	0.04	0.04
$\Sigma^{\mathrm{T}}$	4.02	4.01	4.00	4.00	4.01	4.01	4.01	3.99	3.99	4.00	4.01	4.00	4.02	4.01	4.02	4.02
Ca	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.174	0.63	0.02	0.00	0.00	0.02	0.00	0.00	0.00
Na	0.71	0.66	0.65	0.64	0.60	0.66	0.63	0.677	0.37	0.62	0.62	0.63	0.63	0.63	0.60	0.55
К	0.22	0.32	0.36	0.37	0.37	0.34	0.36	0.123	0.00	0.31	0.36	0.37	0.34	0.37	0.36	0.43
$\Sigma M$	0.97	0.99	1.00	1.01	0.97	1.00	0.99	0.97	1.00	0.95	0.98	1.00	0.98	1.00	0.97	0.98
An,	4.54	0.45	0.00	0.00	0.00	0.17	0.03	17.85	0.02	2.52	0.03	0.09	1.90	0.06	0.05	0.00
mol%																
Ab	73.02	66.77	64.47	63.46	61.95	66.09	63.44	69.50	63.20	65.08	63.55	62.72	63.83	62.67	62.33	55.95
Or	22.45	32.78	35.53	36.54	38.04	33.74	36.53	12.65	36.79	32.40	36.42	37.19	34.27	37.27	37.62	44.05

Class: Anr, Anorthoclase; San, Sanidine. FeO<sup>T</sup>, total iron reported as FeO. bdl, below detection limit. Components: An, Anorthite; Ab, Albite; Or, Orthoclase.

(1.11–1.20 apfu) and X-Fe<sub>tot</sub> (0.93–0.94) than those at Altura and Gadir (0.70–1.03 apfu Na<sub>B</sub>; X-Fe<sub>tot</sub> = 0.69–0.82). A-site occupancy (Na<sub>A</sub> + K<sub>A</sub>) in the amphiboles is also lower in Altura and Gadir samples (0.80±0.03) than in the Mida samples (0.88±0.04) but does not correlate with X-Fe<sub>tot</sub>. The higher F contents in the Altura and Gadir samples compared to Mida may have allowed amphibole to stabilize at higher temperatures in the suites (Ernst, 1962; Giesting & Filiberto, 2014), and the greater range in compositions may suggest that amphibole crystallized continuously over a wider compositional interval during melt differentiation in these suites, whereas it may have only crystallized during the last stages of differentiation prior to eruption at Mida.

## Apatite

Representative analyses of *fluorapatite* are presented in Table 10 and the complete dataset in Supplementary Table S10. The sample from Gadir (190532) contains up to 8% britholite component, as determined by the substitution  $\Sigma REE^{3+} + Si^{4+} = Ca^{2+} + P^{5+}$ . Those from Mida and Altura have britholite components ranging between 1–7% and 5%, respectively. The analyses are similar to those reported from Pantelleria by Mahood & Stimac (1990) and Liszewska *et al.* (2018).

#### Summary

The phenocrysts in the Pantescan suites show compositional differences, some, at least, related to the degree of evolution of the host (as measured by wt%  $Qz^*$ ) and temperature. Mineral compositions imply low relative oxygen fugacities and lower temperatures (<720°C) for the amphibole-bearing suites compared to the Green Tuff. The next sections use petrographic, compositional,

and mineral chemical data to describe quantitatively the evolutionary paths of the various suites.

# DISCUSSION Liquid lines of descent

Mineralogical controls on liquid lines of descent (LLOD) can be constrained by mass balance and trace element modelling. Like most models, mass balance methods do not provide unique solutions; therefore, we present only those calculations that resulted in the lowest possible sum of the squares of the residuals ( $\Sigma r^2 < < 1$ ) with predicted mineral assemblages consistent with observed petrography and melt fractions (F) similar to those approximated by incompatible trace element geochemistry (F(Zr) = X(Zr)<sup>parent</sup> / X(Zr)<sup>daughter</sup>, taking D(Zr)  $\approx$  0). All calculations were performed with Microsoft Excel as described in White *et al.* (2009). Model results are summarized in Table 11 and detailed models are presented in Supplementary Table S11.

#### Green tuff ignimbrite

Four models are presented to provide a detailed description of the LLOD of the Green Tuff. The whole-rock trend that defines the 'thermal valley' is described in Model A1, which uses samples from the Green Tuff type section (Williams, 2010; Williams *et al.*, 2014) for the model parent (sample GT44F: 9.2 wt% Qz\*, 487  $\mu$ g/g Zr, P.I. = 1.15) and daughter (GT44B: 39.0 wt% Qz\*, 2021  $\mu$ g/g Zr, P.I. = 1.77). GT44F was sampled from 400 cm above the base of the section but was selected over higher samples because its lower crystal (<20 vol%) and Ba (<1500  $\mu$ g/g) contents make it less likely to have been affected by feldspar accumulation (White *et al.*, 2009).

Table 5:	Representative	clinopyroxene	analyses

Sample:															
oumpic.	060522	150532	150533		98527		190532		090511	090512		060501	090514	060502	
Point:	#85	30/1.	3/14.	3/1.	#8	#4	35 / 1.	37 / 6.	8/1.	28/1	60/1	#10	41/1	#29	#33
Class:	Aug	Na-Aug	Na-Aug	Na-Aug	Na-Aug	Ae-Aug	Na-Aug	Ae-Aug	Ae-Aug	Na-Aug	Ae-Aug	Na-Aug	Ae-Aug	Na-Aug	Ae-Aug
SiO <sub>2</sub> , wt%	49.75	49.31	49.76	49.55	49.23	49.09	49.98	49.54	47.92	47.54	48.03	48.58	48.28	47.90	48.78
TiO <sub>2</sub>	0.39	0.35	0.41	0.29	0.26	0.25	0.37	0.44	0.34	0.31	0.49	0.46	0.29	0.28	0.29
Al <sub>2</sub> O <sub>3</sub>	0.29	0.14	0.13	0.14	0.14	0.13	0.15	0.13	0.14	0.11	0.13	0.02	0.14	0.10	0.11
FeO <sup>T</sup>	21.26	26.69	26.62	26.78	27.09	27.22	26.86	27.78	28.38	28.31	28.76	29.16	28.19	28.84	29.24
MnO	1.80	1.54	1.53	1.56	1.79	1.53	1.57	1.54	1.16	1.38	1.30	1.36	1.08	1.21	1.12
MgO	6.36	2.31	3.10	2.15	2.95	2.04	2.96	1.69	0.90	1.30	0.82	0.80	0.80	1.09	0.76
CaO	18.56	16.75	16.83	16.34	16.71	16.57	16.60	16.09	16.10	17.90	16.52	14.90	15.19	16.51	14.91
Na <sub>2</sub> O	0.74	2.26	1.64	2.50	1.79	2.54	2.05	2.79	2.96	1.89	2.74	3.18	3.80	2.45	3.74
Sum	99.15	99.36	100.02	99.30	99.95	99.37	100.53	100.01	97.88	98.73	98.79	98.46	97.76	98.39	98.95
Recalculat	ed														
FeO <sup>T</sup>															
Fe <sub>2</sub> O <sub>3</sub>	1.99	5.60	3.31	5.79	5.28	7.20	4.91	7.00	8.74	7.33	8.37	8.13	10.90	7.79	10.64
FeO	19.48	21.65	23.64	21.57	22.34	20.74	22.44	21.48	20.51	21.72	21.23	21.85	18.38	21.83	19.67
Total	99.35	99.92	100.35	99.88	100.48	100.09	101.02	100.71	98.76	99.46	99.62	99.28	98.85	99.17	100.02
X-Fe <sub>tot</sub>	0.65	0.87	0.83	0.87	0.84	0.88	0.83	0.90	0.95	0.92	0.95	0.95	0.95	0.94	0.96
Formula															
(M2M1T <sub>2</sub> C	D <sub>6</sub> )														
Si	1.980	1.990	1.998	1.998	1.979	1.979	1.991	1.987	1.969	1.950	1.962	1.986	1.971	1.966	1.973
<sup>IV</sup> Al	0.014	0.007	0.002	0.002	0.006	0.006	0.007	0.006	0.007	0.005	0.006	0.001	0.007	0.005	0.005
VIFe <sup>3+</sup>	0.006	0.004	0.000	0.000	0.015	0.015	0.002	0.007	0.025	0.045	0.032	0.013	0.022	0.029	0.022
ΣΤ	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000
VIAl	0.000	0.000	0.004	0.004	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
VIFe <sup>3+</sup>	0.053	0.166	0.100	0.176	0.145	0.204	0.145	0.204	0.246	0.181	0.225	0.237	0.313	0.212	0.302
Ti	0.012	0.011	0.012	0.009	0.008	0.008	0.011	0.013	0.010	0.010	0.015	0.014	0.009	0.009	0.009
Mg	0.378	0.139	0.186	0.129	0.177	0.123	0.176	0.101	0.055	0.079	0.050	0.049	0.049	0.067	0.046
Fe <sup>2+</sup>	0.557	0.684	0.697	0.682	0.670	0.666	0.668	0.681	0.689	0.730	0.710	0.699	0.627	0.713	0.643
Mn	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.003	0.000	0.000
ΣM1	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000
Mg	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
Fe <sup>2+</sup>	0.091	0.046	0.096	0.045	0.080	0.033	0.080	0.039	0.016	0.015	0.015	0.048	0.000	0.037	0.022
Mn	0.061	0.053	0.052	0.053	0.061	0.052	0.053	0.052	0.040	0.048	0.045	0.047	0.035	0.042	0.038
Ca	0.792	0.724	0.724	0.706	0.720	0.716	0.709	0.691	0.709	0.787	0.723	0.653	0.664	0.726	0.646
Na	0.057	0.177	0.128	0.195	0.139	0.199	0.159	0.217	0.235	0.150	0.217	0.252	0.301	0.195	0.293
ΣM2	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000	1.000
X-En	0.2125	0.0884	0.1114	0.0844	0.1107	0.0824	0.1088	0.0676	0.0394	0.0549	0.0355	0.0338	0.0384	0.0461	0.0351
X-Fs	0.3647	0.4646	0.4765	0.4747	0.4702	0.4703	0.4632	0.4811	0.5060	0.5147	0.5146	0.5147	0.4959	0.5161	0.5105
X-13 X-Wo	0.4228	0.4470	0.4120	0.4409	0.4191	0.4473	0.4280	0.4513	0.4546	0.4305	0.4499	0.4516	0.4657	0.4378	0.4543

Class: Aug, Augite; Na-Aug, Sodian Augite; Ae-Aug, Aegirine-Augite. Iron oxides recalcualted following Papike *et al.* (1974). Formula:  $\Sigma$ T, sum of tetrahedral cations;  $\Sigma$ M1, sum of octahedral M1 site catiors;  $\Sigma$ M2, sum of octahedral M2 site cations. Components (Lindsley, 1983): Wo, Wollastonite; En, Enstatite; Fs, Ferrosilite; Wo, Wollastonite.

Model A2 also models crystal fractionation along the 'thermal valley' but uses a glass analysis from sample 150544 (Liszewska et al., 2018) with a composition most similar to the average value of highly evolved compositions (Qz\* > 40.5 wt%). Both models have excellent fits ( $\Sigma r^2 = 0.003$  and 0.010, respectively) with similar results. As would be expected, the melt fraction for the glass daughter (A2) is lower (0.254) than that calculated for the whole rock (0.279). Fractionating assemblages are also similar, both dominated by alkali feldspar (0.64–0.67, or 88–89% of the fractionating assemblage) with subordinate clinopyroxene, olivine, aenigmatite, ilmenite, and apatite. Crystallization model (B) tests the origin of this same glass analysis (150544) from its whole-rock host (30.8 wt% Qz\*). This model has similarly good results ( $\Sigma r^2 = 0.067$ ; F = 0.733) and predicted alkali feldspar fractions (0.22, or 88% of the fractionating assemblage) but requires a greater role for

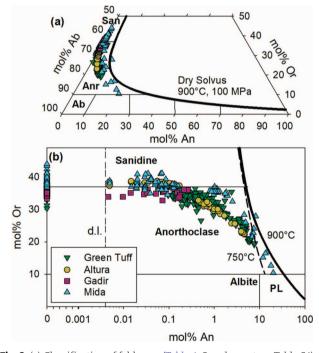
aenigmatite (0.03), a smaller amount of clinopyroxene (0.01), and no olivine. Mass balance models for the LLOD of the Green Tuff were also presented by Liszewska *et al.* (2018) for matrix glasses and Romano *et al.* (2019) for melt inclusions; both resulted in nearly identical fractionating assemblages, with the model of Romano *et al.* (2019) resulting in a similar melt fraction (0.274) and the model of Liszewska *et al.* (2018) resulting in a much lower melt fraction (0.073).

For differentiation beyond the average glass composition to the precursor air-fall tuff, an additional mass balance model was calculated that replicates the "horizontal trend" of Liszewska *et al.* (2018) of declining FeO<sup>T</sup> and Or at near-constant Qz\* and Al<sub>2</sub>O<sub>3</sub> (Model C,  $\Sigma r^2 = 0.613$ ). This model requires an additional 6% crystallization (F = 0.94) of an assemblage of aenigmatite (0.04) with alkali feldspar as the remainder. Liszewska *et al.*'s (2018)

Table 6:	Representative	olivine	analyses
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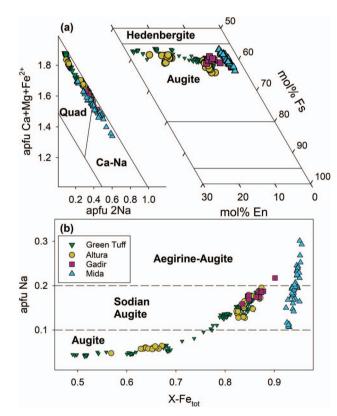
Group:	Altura	Cuddia Mida			
Sample ID:	060522	090512	060501		
Point:	#44	74 / 1.	#85	#110	
SiO <sub>2</sub> , wt%	30.25	28.92	28.88	28.74	
FeO <sup>T</sup>	58.51	64.80	64.60	67.18	
MnO	4.78	3.61	4.38	4.16	
MgO	6.32	0.65	2.33	1.03	
CaO	0.33	0.29	0.70	0.47	
Total	100.19	98.26	100.89	101.59	
Formula (M <sub>2</sub> TO <sub>4</sub> )					
Si (T)	0.982	0.994	0.965	0.964	
Fe <sup>2+</sup>	1.588	1.863	1.805	1.885	
Mn	0.132	0.105	0.124	0.118	
Mg	0.306	0.033	0.116	0.052	
Ca	0.011	0.011	0.025	0.017	
ΣM	2.037	2.012	2.070	2.072	
X-Fa	0.8335	0.9826	0.9276	0.9649	
X-Fo	0.1606	0.0174	0.0595	0.0264	
X-La	0.0060	0.0056	0.0129	0.0087	

Components: Fo, Forsterite; Fa, Fayalite; La, Larnite



**Fig. 8.** (a) Classification of feldspars (Table 4, Supplementary Table S4) shown with the 900°C, 100 MPa dry solvus calculated with SolvCalc (Wen & Nekvasil, 1994) using the solution model of Nekvasil & Burnahm (1987). Ab, Albite; Anr, Anorthoclase; San, Sanidine. (b) A semi-log plot of mol% Or versus mol% An (with the 900° and 750°C solvi) that displays anorthoclase compositions in greater detail. GT data are from Liszewska *et al.* (2018). The vertical dashed line indicates the approximate detection limit (d.l.); note that values with <0.001 mol% An are plotted on a linear scale and all samples with CaO below the detection limit are plotted at 0. PL, Plagioclase.

model for the horizontal trend requires a similar amount of additional crystallization (8%, F = 0.92) mostly of aenigmatite (0.04) and quartz (0.03), with trace amounts of clinopyroxene (0.01) and alkali feldspar (<0.01).



**Fig. 9.** (a) Classification of clinopyroxenes (Table 5, Supplementary Table S5) following Morimoto (1989). (b) Atoms per formula unit [apfu] Na plotted versus X-Fe<sub>tot</sub> [= apfu Fe<sup>T</sup> / (Fe<sup>T</sup> + Mg)] for the clinopyroxenes. in T- $\Delta$ NNO space.

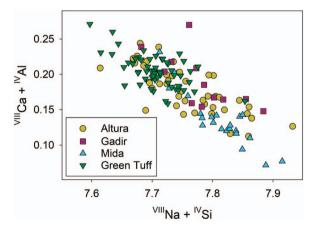
## Cala dell'Altura volcanic Centre

Two models are presented for Altura: Model D describes fractionation along the LLOD between comenditic trachyte (sample 060522; 13.1 wt% Qz\*, 917  $\mu$ g/g Zr, P.I. = 1.1) and pantellerite (sample

Group:	Cala dell'Altur	a	Gadir	Cuddia Mida		
Sample: Point:	150532 sub plate	150533 1	190532 7 / 1.	090512 67 / 1.	090511 1 / 1.	090514 42 / 1.
SiO <sub>2</sub> , wt%	41.27	40.92	40.503	40.83	40.10	40.30
TiO <sub>2</sub>	9.01	9.03	9.137	9.04	8.87	8.80
Al <sub>2</sub> O <sub>3</sub>	0.52	0.52	0.822	0.43	0.37	0.29
FeO <sup>T</sup>	39.50	39.85	40.716	41.12	41.23	40.97
MnO	1.41	1.40	1.343	1.23	1.09	1.21
MgO	0.95	0.90	0.918	0.42	0.40	0.38
CaO	0.53	0.43	0.657	0.38	0.41	0.28
Na <sub>2</sub> O	6.87	7.20	6.869	6.95	6.88	7.00
Total	100.06	100.25	100.97	100.39	99.34	99.23
Recalculated Fe	OT					
Fe <sub>2</sub> O <sub>3</sub>	1.03	3.24	3.18	2.33	3.29	3.36
FeO	38.57	36.94	37.86	39.02	38.27	37.94
Total	100.16	100.57	101.28	100.62	99.67	99.57
X-Ti	0.17	0.17	0.17	0.17	0.16	0.16
Formula (Na,Ca	)2M5(Ti,Fe <sup>3+</sup> )T6O20					
Si	5.884	5.810	5.732	5.830	5.789	5.817
<sup>IV</sup> Al	0.087	0.087	0.137	0.073	0.064	0.049
<sup>IV</sup> Fe <sup>3+</sup>	0.028	0.103	0.131	0.098	0.147	0.133
$\Sigma^{\mathrm{T}}$	6.000	6.000	6.000	6.000	6.000	6.000
Ti	0.966	0.964	0.972	0.970	0.963	0.955
<sup>VI</sup> Fe <sup>3+</sup>	0.034	0.036	0.028	0.030	0.037	0.045
$\sum$ (Ti, Fe <sup>3+</sup> )	1.000	1.000	1.000	1.000	1.000	1.000
VIAl	0.000	0.000	0.000	0.000	0.000	0.000
VIFe <sup>3+</sup>	0.049	0.208	0.180	0.123	0.173	0.187
Fe <sup>2+</sup>	4.599	4.386	4.481	4.659	4.621	4.579
Mn	0.170	0.168	0.161	0.148	0.133	0.148
Mg	0.202	0.191	0.194	0.089	0.085	0.081
<sup>VI</sup> Ca	0.000	0.048	0.000	0.000	0.000	0.004
$\sum M$	5.020	5.000	5.016	5.019	5.012	5.000
VIIICa	0.081	0.018	0.100	0.058	0.063	0.040
Na	1.899	1.982	1.885	1.923	1.925	1.960
$\sum$ (Na,Ca)	1.980	2.000	1.984	1.981	1.988	2.000

 Table 7: Representative aenigmatite analyses

X-Ti = mol Ti / (Ti + Fe<sup>T</sup>). Iron recalculated with charge balance. Formula recalculation following Kunzmann (1999) and Grew et al. (2008).



**Fig. 10.** Extent of the substitution  $Ca^{2+} + Al^{3+} = Na^+ + Si^{4+}$  in aenigmatite. Data from Table 7 and Supplementary Table S7. GT data are from Liszewska *et al.* (2018).

150 532; 38.8 wt% Qz\*, 2067  $\mu$ g/g Zr., P.I. = 1.6); Model E describes fractionation within that pantellerite sample from the whole-rock analysis to an analysis of evolved glass (42.3 wt% Qz\*, P.I. = 1.69) with slightly higher SiO<sub>2</sub> and lower FeO<sup>T</sup> and CaO. Model D has

a very good fit ( $\Sigma r^2 = 0.166$ ) with a melt fraction of 0.420 (58% crystallization) and suggests a LLOD that was the result of fractionation of an assemblage similar to the Green Tuff with respect to alkali feldspar (0.53, or 90.5% of the fractionating assemblage), but with a different mafic assemblage with more clinopyroxene (0.04) and magnetite (0.015%) and only trace amounts (<0.01) of olivine and apatite. The model of the extended LLOD seen in glass analyses in the daughter sample (Model E) has an excellent fit ( $\Sigma r^2 = 0.098$ ) and requires an additional 11% fractionation (F = 0.887) of alkali feldspar (0.04, and aenigmatite (<0.01).

# Cala Gadir volcanic centre

Due to its limited exposure, very few whole-rock samples are available for Gadir and they are similar enough to preclude any opportunity to calculate a meaningful LLOD for this centre. A single model (G) is presented to describe fractionation between the whole-rock composition (37.6 wt% Qz\*, 2117  $\mu$ g/g Zr, P.I. = 1.3) and evolved glass (43.6 wt% Qz\*, 3413  $\mu$ g/g Zr, P.I. = 1.5), which has an excellent fit ( $\Sigma$ r<sup>2</sup> = 0.019) and predicts a melt fraction of 0.772 after fractionation of a mineral assemblage dominated by alkali feldspar (0.17, or 73% of the fractionating assemblage) and quartz

Table 8:	Representative	analyses	of Fe-Ti	oxides
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Group:	Cala dell'Altı	ura		Gadir	Cuddia Mid	la		
Sample: Point: Class:	060522 Avg Mgt	150532 2 Ilm	150533 1 Ilm	190532 8 / 1. Ilm	090511 17 / 1. Ilm	090512 68 / 1. Ilm	090513 6 / 1. Mgt	060501 #23 Ilm
SiO <sub>2</sub> , wt%	0.07	0.06	1.16	0.02	0.03	0.29	0.12	bdl
TiO <sub>2</sub>	22.39	50.82	45.77	51.17	51.25	51.46	25.89	49.46
Al <sub>2</sub> O <sub>3</sub>	0.46	bdl	0.66	0.02	0.01	0.03	0.42	bdl
V2O3	n.a.	0.17	0.25	0.18	0.25	0.25	0.13	n.a.
Cr <sub>2</sub> O <sub>3</sub>	0.04	n.a.	n.a.	0.04	0.01	0.03	0.01	0.14
FeOT	69.96	45.73	44.07	46.02	45.31	46.30	67.70	46.25
MnO	1.94	2.54	0.62	2.41	1.91	1.98	1.70	2.12
MgO	0.40	0.14	0.21	0.16	0.08	0.10	0.25	0.15
CaO	0.06	0.04	0.64	0.03	0.01	0.03	0.01	n.a.
Nb <sub>2</sub> O <sub>5</sub>	n.a.	0.77	n.a.	0.90	0.52	0.27	0.09	n.a.
Total	95.32	100.27	93.38	100.95	99.38	100.73	96.31	98.11
Recalculated								
Fe <sub>2</sub> O <sub>3</sub>	23.67	3.21	4.90	3.03	1.43	2.48	16.80	4.62
FeO	48.66	42.84	39.66	43.29	44.03	44.07	52.59	42.09
Total	97.69	100.53	92.71	101.23	99.49	100.69	97.99	98.58
Formula (Ilm	1: M2O3, Mgt: M3C	04)						
0	4.000	3.000	3.000	3.000	3.000	3.000	4.000	3.000
Si	0.003	0.002	0.031	0.001	0.001	0.007	0.004	0.000
Ti	0.645	0.963	0.919	0.963	0.980	0.968	0.742	0.954
Al	0.021	0.000	0.021	0.001	0.000	0.001	0.019	0.000
V	_	0.003	0.005	0.004	0.005	0.005	0.004	_
Cr	0.001	_	_	0.001	0.000	0.001	0.000	0.003
Fe <sup>3+</sup>	0.683	0.042	0.074	0.036	0.015	0.034	0.479	0.089
Fe <sup>2+</sup>	1.560	0.921	0.909	0.927	0.949	0.935	1.680	0.902
Mn	0.063	0.054	0.014	0.051	0.041	0.042	0.014	0.046
Mg	0.023	0.005	0.008	0.006	0.003	0.004	0.000	0.006
Ca	0.002	0.001	0.018	0.001	0.000	0.001	0.001	_
Nb	_	0.009	0.000	0.010	0.006	0.003	0.000	_
$\Sigma^{M}$	3.000	2.000	2.000	2.000	2.000	2.000	2.945	2.000
X-Usp	0.6487	_	_	_	_	_	0.7491	_
X-Ilm	_	0.9678	0.9540	0.9696	0.9855	0.9754	_	0.9530

Class: Mgt, Magnetite; Ilm, Ilmenite. Oxides recalculated following Andersen *et al.* (1993) with total iron (FeO<sup>T</sup>) recalculated following Carmichael (1967). Formula:  $\Sigma$ M, Sum of Cations. Components: Usp, Ulvöspinel; Ilm, Ilmenite.

(0.05), with amphibole (0.02) as the main mafic phase along with trace ilmenite.

#### Cuddia Mida volcanic Centre

For whole-rock samples that lie in the 'thermal valley', an LLOD can be calculated only for a pantellerite with 32.5 wt% Qz\*, 1640  $\mu$ g/g Zr, and P.I. = 1.6 (090512) to pantellerite on the feldsparquartz cotectic with 39.2 wt% Qz\*, 1986  $\mu$ g/g Zr, and P.I. = 1.9 (090513). The resulting model (H) has an excellent fit ( $\Sigma r^2 = 0.094$ ) with a melt fraction of 0.820 and a fractionating assemblage of alkali feldspar (0.16, or 84% of the assemblage) with subequal amounts (~0.02 each) of amphibole and aenigmatite, similar to the observed modal assemblage. Mass balance models were also calculated for fractionation from the whole-rock (090512) composition to glass with 46.7 wt% Qz\*, 3109  $\mu$ g/g Zr, and P.I. = 2.5 (model J) and to one of the most evolved glasses with 56.2 wt% Qz\*, 3576  $\mu$ g/g Zr, and P.I. = 3.2 (model K). Both models have an acceptable fit ( $\Sigma r^2 = 0.072$  and 0.470, respectively) and similar results with varying mineral proportions, as would be expected from compositions lying on the same LLOD. with relatively low melt fractions (decreasing from 0.647 to 0.437) and a fractionating assemblage dominated by alkali feldspar (increasing from 0.29 to 0.43, 80-76% of the fractionating assemblage) and quartz (increasing from 0.02 to 0.08), with lesser amounts of amphibole (0.02-0.03), aenigmatite (0.02), trace (<0.01) apatite, and a negligible amount of clinopyroxene. The diminished role for clinopyroxene and greater role for quartz in this assemblage is consistent with the observed positive correlation of Qz\* with CaO and negative correlation of Qz\* and SiO<sub>2</sub> in this series. A model (L) of fractional crystallization from a glass analysis in sample 090512 with 45.6 wt% Qz\*, 2413  $\mu$ g/g Zr, and P.I. = 2.0 to the same highly evolved composition requires 36% crystallization (F = 0.65) of an assemblage only alkali feldspar (0.26, or 70% of the fractionating assemblage) and quartz (0.11), highlighting the that extreme fractionation to the EMC may require not just an assemblage of alkali feldspar ± quartz but also a dearth of mafic phases (cf. LeMasurier et al., 2018).

## Summary

Mass-balance modelling confirms that all suites evolved from trachyte to pantellerites with compositions that plot near the experimental (viz., Carmichael & MacKenzie, 1963) whole-rock

Table 9: Representative amphibole	analyses	
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Group:	Cala dell'A	Altura			Gadir		Cuddia M	ida		
Sample ID:	150532		150533		190 532		090514	090513	060501	060502
Point:	200 x 180 Core	$\mu$ m euh Rim	875 x 420 Core	$\mu$ m prism Rim	14 / 1.	21 / 1.	29/1.	1/1.	#28	#46
SiO <sub>2</sub> , wt%	48.87	48.89	48.63	48.80	49.23	48.90	47.68	47.54	48.45	47.78
TiO <sub>2</sub>	2.01	2.31	1.98	2.24	1.73	2.09	1.87	1.49	1.42	1.88
Al <sub>2</sub> O <sub>3</sub>	1.32	1.18	1.58	1.33	1.48	1.38	0.68	0.54	0.57	0.81
FeOT	27.89	29.36	26.13	29.12	26.83	29.00	32.58	33.50	34.21	33.18
MnO	1.22	1.26	1.10	1.21	1.18	1.23	1.13	1.05	1.27	1.21
MgO	5.21	4.27	6.41	4.31	5.95	4.38	1.44	1.26	1.70	1.41
CaO	6.07	5.60	6.46	5.89	5.94	5.55	4.78	4.23	4.08	4.61
Na <sub>2</sub> O	5.06	5.15	5.08	5.13	5.22	5.18	5.41	5.66	5.56	5.33
K <sub>2</sub> O	1.08	1.13	0.94	1.12	0.98	1.05	1.40	1.57	1.31	1.30
H <sub>2</sub> O <sup>+</sup>	n.a.	n.a.	n.a.	n.a.	1.43	1.58	1.57	n.a.	n.a.	n.a.
F	1.35	1.54	1.87	1.47	0.98	0.65	bdl	bdl	1.46	1.29
Cl	0.08	0.07	0.07	0.07	0.06	0.07	0.06	0.05	0.06	0.06
Sum	100.16	100.76	100.25	100.69	101.01	101.05	99.09	97.44	100.09	98.84
O=F,Cl	0.59	0.66	0.80	0.63	0.43	0.29	0.01	0.01	0.63	0.56
Total										
Recalculated	99.57 FeO <sup>T</sup>	100.10	99.45	100.06	100.58	100.76	98.86	97.20	99.46	98.28
Fe <sub>2</sub> O <sub>3</sub>	1.24	1.40	1.19	1.10	1.43	1.42	0.87	2.38	2.59	1.51
FeO	26.77	28.10	25.06	28.13	25.54	27.72	31.80	31.36	31.89	31.82
Total	99.70	100.24	99.57	100.17	100.72	100.90	98.95	98.98	99.72	98.44
X-Fe <sub>tot</sub>	0.75	0.79	0.69	0.79	0.72	0.79	0.93	0.94	0.92	0.93
Formula (A <sub>0-:</sub>	1B2C5T8O22X	( <sub>2</sub> )								
Si	7.606	7.620	7.544	7.608	7.621	7.631	7.761	7.756	7.747	7.729
<sup>IV</sup> Al	0.242	0.217	0.289	0.244	0.269	0.254	0.130	0.103	0.108	0.153
<sup>IV</sup> Ti	0.152	0.163	0.168	0.148	0.109	0.115	0.109	0.141	0.145	0.118
ΣT	8.000	8.000	8.001	8.000	7.999	8.000	8.000	8.000	8.000	8.000
<sup>VI</sup> Ti	0.083	0.108	0.063	0.115	0.092	0.130	0.120	0.042	0.026	0.111
<sup>VI</sup> Fe <sup>3+</sup>	0.145	0.163	0.140	0.128	0.166	0.168	0.106	0.293	0.310	0.184
Mn	0.078	0.073	0.064	0.087	0.060	0.067	0.095	0.081	0.000	0.060
Fe <sup>2+</sup>	3.485	3.664	3.249	3.669	3.307	3.617	4.330	4.278	4.259	4.305
Mg	1.209	0.992	1.482	1.002	1.374	1.018	0.350	0.306	0.405	0.340
ΣČ	5.000	5.000	4.998	5.001	4.999	5.000	5.001	5.000	5.000	5.000
Mn	0.083	0.093	0.080	0.073	0.095	0.096	0.060	0.064	0.172	0.105
Fe <sup>2+</sup>	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.006	0.000
Ca	1.012	0.935	1.074	0.984	0.985	0.928	0.834	0.740	0.700	0.799
Na	0.905	0.972	0.846	0.943	0.920	0.975	1.106	1.196	1.121	1.096
ΣB	2.000	2.000	2.000	2.000	2.000	1.999	2.000	2.000	1.999	2.000
Na	0.622	0.585	0.682	0.607	0.645	0.593	0.600	0.595	0.601	0.576
X	0.022	0.225	0.186	0.223	0.193	0.208	0.291	0.393	0.267	0.268
ΣΑ	0.214 0.836	0.225	0.186	0.225	0.195	0.208	0.291	0.922	0.267	0.268
ΣA ΣT,C,B,A	0.836 15.836	15.810	0.868 15.867	0.830 15.831	15.836	15.800	15.892	0.922 15.922	0.868 15.867	0.844 15.844
CH -	1.314	1.222	1.064	1.257	1.504	1.662	1.983	1.986	1.247	1.325
F	0.664	0.759	0.917	0.725	0.479	0.319			0.737	0.660
Cl	0.021	0.018	0.018	0.018	0.017	0.019	0.017	0.014	0.015	0.015
ΣΧ	2.000	2.000	2.000	2.000	2.000	2.000	2.000	2.000	1.999	2.000

Amphibole recalculated following Hawthorne *et al.* (2012) with the MS Excel spreadsheet of Locock (2014). FeO<sup>T</sup>, total iron reported as FeO; X-Fe<sub>tot</sub> = mol Fe<sup>T</sup>/(Mg + Fe<sup>T</sup>)

minimum via fractional crystallization of a similar assemblage dominated by alkali feldspar (~90% of the fractionating assemblage) with subordinate clinopyroxene, olivine, ilmenite, aenigmatite, and apatite. However, continued evolution to glass compositions beyond this point to the apparent minima for each suite require different assemblages that vary primarily by their mafic phases and the presence or absence of quartz. In the next section, we attempt to relate the paths to variable pre-eruptive conditions in the reservoirs.

# Previous Geobarometric estimates

The depth of the basaltic reservoir beneath Pantelleria has been estimated to be 8–6 km (200–150 MPa) based both on  $H_2O-CO_2$  equilibria in melt inclusions (Gioncada & Landi, 2010; Neave *et al.*, 2012) and thermodynamic modelling (White *et al.*, 2020). The application of various methodologies has consistently placed the silicic reservoirs at  $\leq 6$  km ( $\leq 150$  MPa; Mahood, 1984; Lowenstern, J.B., 1994; Civetta *et al.*, 1998; White *et al.*, 2009; Di Carlo *et al.*, 2010; Gioncada & Landi, 2010; Neave *et al.*, 2012;

Table 10: Representative apatite analyse
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Group:	Altura	Gadir	Cuddia Mida			
Sample ID: Point:	150533	190532 22 / 1.	090512 53 / 1.	090514 46 / 1.	090513 8/1.	
SiO <sub>2</sub> , wt%	2.39	2.25	2.09	1.06	0.45	
FeO	0.38	0.50	0.65	0.55	0.57	
MnO	0.00	0.13	0.09	0.18	0.13	
MgO	0.02	0.00	0.00	0.02	0.02	
CaO	49.75	47.90	51.98	53.07	55.39	
Na <sub>2</sub> O	0.62	0.59	0.56	0.32	0.17	
P <sub>2</sub> O <sub>5</sub>	37.91	35.72	33.66	40.14	41.88	
Y <sub>2</sub> O <sub>3</sub>	0.33	0.46	0.12	0.13	bdl	
La <sub>2</sub> O <sub>3</sub>	1.37	1.79	1.97	0.86	0.30	
Ce <sub>2</sub> O <sub>3</sub>	2.91	3.77	3.89	1.86	0.63	
Pr <sub>2</sub> O <sub>3</sub>	n.a.	0.47	1.67	0.95	0.36	
Nd <sub>2</sub> O <sub>3</sub>	n.a.	2.10	n.a.	n.a.	n.a.	
H <sub>2</sub> O	n.a.	0.24	0.58	0.55	n.a.	
F	2.02	2.83	2.12	2.44	2.21	
Cl	0.04	0.05	0.04	0.02	0.15	
Sum	97.74	98.80	99.40	102.17	102.25	
O=F, Cl	0.86	1.20	0.90	1.03	0.96	
Total	96.88	97.59	98.50	101.13	101.29	
Formula (M5(TO4)3	X)					
Р	2.875	2.756	2.612	2.891	2.992	
Si	0.214	0.205	0.191	0.091	0.038	
ΣT	3.089	2.961	2.803	2.981	3.030	
Ca	4.775	4.676	5.105	4.838	5.008	
Fe	0.028	0.038	0.049	0.039	0.040	
Mn	0.000	0.010	0.007	0.013	0.010	
Na	0.108	0.105	0.099	0.053	0.028	
Y	0.016	0.023	0.006	0.006	0.000	
La	0.045	0.060	0.066	0.027	0.009	
Ce	0.095	0.126	0.130	0.058	0.020	
Pr	_	0.015	0.056	0.030	0.011	
Nd	_	0.068	_	_	_	
ΣM	5.068	5.121	5.519	5.064	5.126	
F	0.572	0.816	0.615	0.657	0.590	
Cl	0.006	0.007	0.005	0.003	0.021	
OH	0.422	0.177	0.379	0.340	0.389	
ΣΧ	1.000	1.000	1.000	1.000	1.000	

Apatite analyses recalculated following Pasero et al. (2010)

Lanzo et al., 2013; Arzilli et al., 2020; Stabile et al., 2021; Romano et al., 2021). Many of these studies have, at least in part, relied on equivocal results that rely on volatile contents and calculated solubilities that have placed the felsic reservoir for the Green Tuff at  $\leq$ 4 km ( $\leq$ 100 MPa), based on maximum H<sub>2</sub>O concentrations in melt inclusions up to 4.2 wt% (Lanzo et al., 2013; Romano et al., 2019), and between 2 and 6 km (50-150 MPa) for post-Green Tuff pantellerites, based on H<sub>2</sub>O concentrations in melt inclusions between 2.5 and 4.9 wt% (Métrich et al., 2006; Gioncada & Landi, 2010; Neave et al., 2012). Neither Lowenstern & Mahood (1991) nor Gioncada & Landi (2010) detected any CO<sub>2</sub> by FTIR in melt inclusions by FTIR; however, Neave et al. (2012) reported concentrations (analyzed by SIMS) up to 150 ppm CO<sub>2</sub>, which they used with  $H_2O$  data to calculate a depth of ~5.5 km (~146 MPa) for post-Green Tuff pantellerites. Lanzo et al. (2013) compared Cl concentrations in the Green Tuff (~1 wt%) and post-Green Tuff pantellerites (~0.85 wt%) with the experimental results of Métrich & Rutherford (1992) and Signorelli & Carroll (2002) and concluded that the Green Tuff reservoir may have been emplaced at a much shallower level (2-3 km, or 50-100 MPa) than the post-Green Tuff

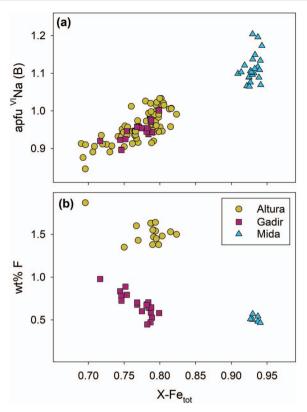
pantellerites ( $\geq$ 4.5 km, or  $\geq$  120 MPa). From their experimental results, Di Carlo *et al.* (2010) also placed the post-Green Tuff reservoir at 5 ± 1 km (120 ± 20 MPa). Based on this previous work, in the following section we will use values of 100 MPa and 150 MPa to calculate initial estimates of T and  $fO_2$ , and then use a new iterative technique to calculate a new estimate of P for fayalite-bearing rocks.

# Mineral Equilibria

Carmichael (1962) and Nicholls & Carmichael (1969) noted that ilmenite is frequently the sole oxide in pantelleritic rocks, from which they deduced the reduced nature of these melts. They also described an apparent antipathetic relationship between oxides and aenigmatite with increasing melt peralkalinity and suggested the presence of a 'no-oxide' field in melts with high peralkalinity and lower temperatures crystallizing at oxygen fugacities above the fayalite-magnetite-quartz (FMQ) buffer. White *et al.* (2005) proposed a generalized 'reaction series' for pantellerite that was demonstrated experimentally by Romano *et al.* (2020). With increasing silica activity, peralkalinity (P.I. ~1.0 to 2.0) and

	Model Parent	arent		Model d	Model daughter		Fraction	Fractionating assemblage	mblage							Melt Fraction	tion	
Model	wt% Qz* P.I.	* P.I.	µg/g Zr	Qz*	P.I.	Zr	Afs	Qz	Aen	Cpx	Amp	ol	Ap	Ilm	Mgt	F	F(Zr)	$\sum r^2$
Green Tuff Ignimbrite	ę																	
(A1)	9.16	1.15	487	38.97	1.77	2021	0.64		0.01	0.03		0.03	< 0.01	0.01		0.279	0.236	0.003
(A2)	9.16	1.15	487	44.53	1.84	2221	0.67	I	0.01	0.04	Ι	0.03		0.01	I	0.246	0.219	0.010
(B)	30.84	1.61	1634	44.53	1.84	2221	0.22		0.03	0.01						0.733	0.736	0.067
(C)	44.53	1.84	2221	43.78	1.79	2591	0.01		0.04							0.944	0.857	0.613
Cala dell'Altura Volcanic Centre	anic Cent	re																
(D)	13.06	1.09	917	38.78	1.64	2067	0.54			0.04		< 0.01	< 0.01		0.02	0.420	0.444	0.166
(E)	38.78	1.64	2067	42.34	1.69	n.a.	0.08	I	0.01		0.04					0.887	n.a.	0.098
Cala Gadir Lava Dome	ne																	
(G)	37.59	1.34	2117	43.58	1.50	3413	0.17	0.05			0.02			<0.01	I	0.772	0.620	0.019
Cuddia Mida Volcanic Centre	ic Centre																	
(H)	32.45	1.63	1640	39.22	1.91	1986	0.16		0.02		0.02		< 0.01			0.820	0.826	0.094
()	32.45	1.63	1640	46.74	2.45	3109	0.29	0.02	0.02	< 0.01	0.02		0.01	Ι	I	0.647	0.528	0.072
(K)	32.45	1.63	1640	56.17	3.17	3576	0.43	0.08	0.02	< 0.01	0.03		< 0.01			0.437	0.459	0.470
(T)	45.55	2.01	2413	56.17	3.17	3576	0.25	0.10	I	I			I	I		0.653	0.675	0.169
See text and Supplementary Table 10 for Model details. Qz* = 100Qzn/(Qzn + Orn + Abn), subscript n indicates normative values (in wt%) calculated following Kelsey (1965) with iron oxides recalculated following Le Maitre (1976). PI. (Peralkalinity Index) = mol (Na + K)/Al; Fractionating assemblage: Afs, alkali feldspar; Aen, aenigmatite; Cpx, clinpyroxene; Amp, amphibole; Ol, olivine; Qz, quartz; Ap, apatite; Ilm, ilmenite; Mgt, magnetite. F, calculated melt fraction. F(Zr) = Zr <sub>Parent</sub> /Zr <sub>Daughter</sub> $\Sigma$ r <sup>2</sup> , sum of squares of residuals.	nentary Tal alkalinity I: ed melt fra	ole 10 for ♪ ndex) = m( ction. F(Zr	Model details ol (Na + K)/Al )=Zr <sub>Parent</sub> /Zı	$Qz^* = 10C$ ; Fractions	)Qzn/(Qzn + ating assem `r <sup>2</sup> , sum of :	Or <sub>n</sub> + Ab <sub>n</sub> ), iblage: Afs, ? squares of r	subscript ı alkali felds <sub>l</sub> esiduals.	n indicates ı par; Aen, ae	normative v nigmatite;	values (in v Cpx, clinpy	vt%) calcula roxene; An	ated follow 1p, amphib	ing Kelsey ( ole; Ol, oliv	1965) with ine; Qz, qu	iron oxide lartz; Ap, a	s recalculat patite; 11m, 3	ed followin Imenite; M	g Le gt,

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**Fig. 11.** X-Fe<sub>tot</sub> [= apfu Fe<sup>T</sup> / (Fe<sup>T</sup> + Mg)] for the amphiboles plotted against (a) atoms per formula unit [apfu] Na occupying the octahedral (VI) B-site (on a 23 oxygen basis) and (b) wt% F. Data from Table 9 and Supplementary Table S9.

decreasing temperature (~900 to 700°C), the assemblages (each accompanied by alkali feldspar) are: (1) augite + olivine + ilmenite + magnetite, (2) sodian augite + fayalitic olivine + ilmenite; (3) sodian augite + fayalite + ilmenite + aenigmatite; (4) sodian augite + ilmenite + aenigmatite + quartz  $\pm$  amphibole, and (5) aegirine-augite + aenigmatite + quartz  $\pm$  amphibole. Rock samples from the magma series in this study have quartz-bearing assemblages (viz., types 4 and 5) that plot on the alkali feldsparquartz cotectic at or near the whole-rock minimum (or minima). Anomalous among these are two samples from Mida that have fayalitic (Fa<sub>>95</sub>) olivine, and one that crystallized magnetite instead of ilmenite.

Here we use mineral equilibria to provide constraints on the pressure (P), temperature (T), and oxygen fugacity  $(fO_2)$  for the pre-eruptive conditions of these samples. These assemblages are evaluated using equilibria in the Ca-QUIIF system (Frost et al., 1988; Frost & Lindsley, 1992; Lindsley & Frost, 1992) for reactions involving olivine, clinopyroxene, ilmenite, and quartz via use of the QUILF95 program (Andersen et al., 1993), along with reactions between aenigmatite, ilmenite, and quartz (White et al., 2005; Macdonald et al., 2011; Romano et al., 2020). For all calculations using the QUILF95 program, we use the results from Fe-Mg-Ca exchange for reactions involving silicates, but only Fe-Ti exchange for reactions involving silicates and oxides (viz., ignoring Mg-Mn in oxides). All oxygen fugacity values are reported as log units relative to the nickel-nickel oxide (NNO) buffer (O'Neil & Pownceby, 1993; Pownceby & O'Neil, 1994). The full results of these calculations are presented in Supplementary Tables 12 and Supplementary Table 13 and discussed below.

#### Clinopyroxene-Ilmenite-Aenigmatite-quartz Equilibria

For assemblages with quartz, clinopyroxene, aenigmatite, and ilmenite but neither olivine nor magnetite, White (2011) proposed that temperature and oxygen fugacity can be estimated by calculating the intersection of two curves in  $T-fO_2$  space, based on the assumption that all four phases equilibrated under the same conditions (cf. Berman, 1991):2Fe<sub>2</sub>Si<sub>2</sub>O<sub>6</sub> + O<sub>2</sub> = 2Fe<sub>2</sub>O<sub>3</sub> + 4SiO<sub>2</sub>,.

$$\label{eq:clinoferrosilite} \begin{split} & \text{clinoferrosilite} + \text{O}_2 = \text{hematite} + \text{quartz.Na}_2\text{TiFe}_5\text{Si}_6\text{O}_{20} \\ & + \text{O}_2 = 2\text{Fe}_2\text{O}_3 + \text{FeTiO}_3 + \text{Na}_2\text{Si}_2\text{O}_5 \ (\text{melt}) + 4\text{SiO}_2, \end{split}$$

aenigmatite +  $O_2$  = hematite + ilmenite + Nds + quartz.

Equation 1 is the displaced augite-hematite-quartz (AHQ) buffer, which is calculated using the QUILF95 program (Andersen et al., 1993) to solve for fO2 over a range of temperatures (660-760°C) for given values of X-Wo, X-En, and X-Hem at a stated pressure (with  $aSiO_2[Qtz] = 1.0$  for quartz saturation). Equation 2 defines the position of the so-called 'no oxide' boundary proposed by White et al. (2005) and hereafter referred to as the displaced aenigmatite-ilmenite-quartz buffer (AenIQ), which is calculated over the same range of temperatures and at the same pressure for given values of X-Hem and X-Ilm with activities for these components also calculated using QUILF95 (oxide solution model of Andersen et al., 1991), with unit values for aSiO<sub>2</sub>[Qtz] and aNa<sub>2</sub>Si<sub>2</sub>O<sub>5</sub> (for P.I. > 1.0; Carmichael et al., 1974, p. 288). The positions of these curves and their point of intersection are strongly pressure-dependent; therefore, we present and discuss results calculated at both 100 and 150 MPa (as described in the previous section). Temperatures calculated at 150 MPa versus 100 MPa are ~11°C lower; values for relative fO2 stated in terms of  $\Delta$ NNO are nearly identical. Likewise, rather than extrapolating conclusions based on calculations from arbitrary 'representative analyses' we use the entire range of ilmenite and clinopyroxene compositions to calculate parameters to define stability fields for these assemblages in  $T-fO_2$  space (Fig. 12).

Altura. At 100 MPa, calculated temperatures range from 754°C to 713°C, declining with increasing X-Fe<sub>tot</sub> and decreasing X-En in clinopyroxene (0.82–0.87 and 0.1174–0.0840, respectively). Calculated redox conditions are highly sensitive to ilmenite compositions (especially at values approaching end-member compositions) and range from  $\Delta$ NNO-1.2 to –2.1 as the X-Hem component in ilmenite varies from 0.0460 to 0.0244. At 150 MPa, calculated oxygen fugacities are similar ( $\Delta$ NNO-1.2 to –1.9) but temperatures are an average of ~42°C lower (711 to 671°C).

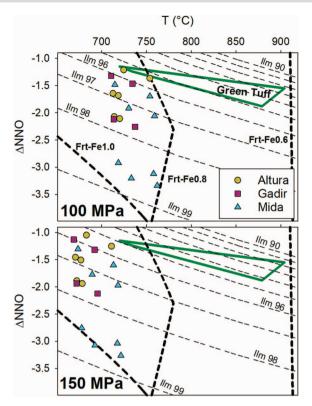
Gadir. Mineral compositions in the sample from Gadir are similar to Altura, and therefore result in similar T- $fO_2$  values. At 100 MPa, temperatures decrease from 737 to 711°C with increasing X-Fe<sub>tot</sub> and decreasing X-En in clinopyroxene (0.84–0.90 and 0.1088–0.0676, respectively). Relative oxygen fugacity is between  $\Delta$ NNO-1.3 (X-Hem = 0.0399) and – 2.3 (X-Hem = 0.0236). At 150 MPa, calculated T- $fO_2$  values range between 669–695°C and  $\Delta$ NNO-1.1 to –2.1.

Mida. Calculated temperatures from Mida are similar to those from Altura and Gadir (715–762°C at 100 MPa and 674–721°C at 150 MPa). However, the presence of some ilmenite grains with ~99 mol% Ilm extends the calculated range of oxygen fugacities from  $\Delta$ NNO-1.3 and – 1.5 (at X-Hem = 0.0363) to –3.3 (at X-Hem = 0.0128).

#### Geobarometry from

## olivine-Clinopyroxene-Ilmenite-Aenigmatite-quartz Equilibria

Olivine (92.6–97.6 mol% Fa) is found in two samples from Mida (090512 and 060501) in assemblages with clinopyroxene, ilmenite,



**Fig. 12.** Summary of the results of clinopyroxene-ilmenite-aenigmatitequartz geothermometry (T, °C) and oxygen barometry ( $fO_2$  plotted relative to the nickel-nickel oxide buffer ( $\Delta$ NNO; O'Neil & Pownceby, 1993, Pownceby & O'Neil, 1994) at (a) 100 MPa and (b) 150 MPa with isopleths for ilmenite in equilibrium with aenigmatite and quartz (Andersen *et al.*, 1991; White *et al.*, 2005; Romano *et al.*, 2020). See text and Supplementary Table S12 for details. Also plotted in both is the range of values for the GT (Liszewska *et al.*, 2018; Romano *et al.*, 2018, 2020) and the curves showing the experimental stability limits for ferrorichterite with compositions of X-Fe<sub>tot</sub> = 0.6, 0.8, and 1.0 (Charles, 1977), both at 100 MPa.

aenigmatite, and quartz. For the olivine-clinopyroxene-quartz sub-assemblage, the QUILF algorithm can be used to either: (1) calculate temperature and evaluate equilibrium between clinopyroxene and olivine as a function of pressure; (2) calculate pressure and evaluate equilibrium as a function of temperature; or (3) calculate pressure and temperature given an assumption of equilibrium (Andersen et al., 1993). The addition of ilmenite to this sub-assemblage can then be used to calculate  $fO_2$ . Because the olivine-clinopyroxene-ilmenite-quartz and aenigmatite-ilmenitequartz sub-assemblages are both quartz-saturated, it is possible to estimate the  $P-T-fO_2$  of equilibration of the entire assemblage (Nicholls et al., 1971). To do so, an initial temperature and oxygen fugacity are calculated from cpx-ol-q-ilm equilibrium at an arbitrary pressure (100 MPa.) These values are then applied to the AenIQ equation for silica activity (Romano et al., 2020) to calculate a new pressure:

$$P = \frac{-5666.7 + T \bullet \left(2.133 - \frac{1}{4}\log fO_2 + \frac{1}{4}\log a_{Ilm} + \frac{1}{2}\log a_{Hem}\right)}{0.55} + 0.1,$$
(Equation 3)

where P is the pressure in MPa, T is the temperature in K, and log  $fO_2$ ,  $a_{1lm}$ , and  $a_{Hem}$  are output values calculated by QUILF95. The newly estimated pressure is then returned to the QUILF95 program to calculate new T- $fO_2$  iteratively until pressure convergence was achieved; these results are presented

in Supplementary Table 13. For the two olivine-bearing samples from Mida, calculated pressures are 155 MPa (X-Hem = 0.0288;  $\Delta$ NNO-1.7) and 160 MPa (X-Hem = 0.0128;  $\Delta$ NNO-3.0) for sample 090512 and 159 MPa for sample 060501 (X-Hem = 0.470;  $\Delta$ NNO-1.0), with temperatures of 684–683° and 674°C, respectively. Therefore, our results from mineral equilibria agree with experimental results of Di Carlo *et al.* (2010) and the chlorine solubilities and provides additional evidence that the magma reservoir for the Mida/Fastuca system was likely deeper (~150 MPa, or ~ 5.7 km) than the reservoir for the Green Tuff ( $\leq$ 100 MPa, or  $\leq$  3.8 km).

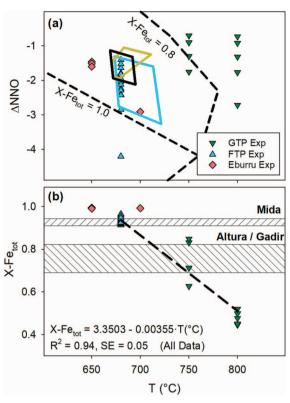
## Comparison with experimental results

In addition to the early experiments on synthetic silica-over saturated (viz., haplogranitic) peralkaline systems by Carmichael & MacKenzie (1963), Bailey & Schairer (1966), and Thompson & MacKenzie (1967), phase equilibria experiments have been performed on samples from natural peralkaline systems from Olkaria, Kenya (Scaillet & Macdonald, 2001, 2003), Eburru, Kenya (Scaillet & Macdonald, 2006), and Pantelleria (Di Carlo et al., 2010; Arzilli et al., 2020; Romano et al., 2020, 2021; Stabile et al., 2021). There have also been experimental studies on the stability of alkali amphiboles in general (Della Ventura et al., 2005) and the richterite-ferrorichterite series specifically (Charles, 1975, 1977). In this section, experimental and modelling results are integrated to provide further information on the conditions of amphibole formation and extreme differentiation in the Pantescan suites.

#### Amphibole stability

Phase equilibria studies on the solid-solution series between richterite (NaNaCaMg<sub>5</sub>Si<sub>8</sub>O<sub>22</sub>(OH)<sub>2</sub>; X-Fe<sub>tot</sub> = 0.0) and ferrorichterite  $(NaNaCaFe_5Si_8O_{22}(OH)_2; X-Fe_{tot} = 1.0)$  were performed by Charles (1975, 1977), who experimentally determined the stability of these amphiboles as a function of temperature and oxygen fugacity. These experiments showed that at 100 MPa the magnesian endmember is stable up to  $1025 \pm 5^{\circ}$ C at all values of fO<sub>2</sub> whereas the ferroan endmember has a maximum stability of  $760 \pm 10^{\circ}$ C at fO<sub>2</sub> equivalent to the wüstite-magnetite (WM) buffer ( $\sim \Delta NNO$ -4.2), falling to  $535 \pm 10^{\circ}$ C at fO<sub>2</sub> equivalent to the fayalite-magnetitequartz (FMQ) buffer (~∆NNO-0.8). For compositions more like the most ferrous amphiboles from Altura and Gadir (X-Fe<sub>tot</sub> = 0.8) maximum stability occurs at  $780 \pm 10^{\circ}$ C at  $fO_2$  equivalent to the C-CH<sub>4</sub> buffer ( $\sim \Delta$ NNO-2.3), falling to 730 ± 10°C at fO<sub>2</sub> equivalent to the FMQ buffer ( $\sim \Delta$ NNO-0.7). Temperatures and oxygen fugacities for the amphibole-bearing samples estimated from mineral equilibria at 150 and 100 MPa, along with the range of reported T-fO<sub>2</sub> for the Green Tuff (White et al., 2009; Liszewska et al., 2018; Romano et al., 2020), are plotted with the stability fields of Charles (1977) in Figs 12a and b. Calculated T-fO<sub>2</sub> values for these samples and the compositions of their amphiboles (X-Fe<sub>tot</sub> = 0.69– 0.83 for Altura and Gadir; X-Fetot = 0.91 to 0.97 for Mida) are consistent with the experimental results of Charles (1977) and therefore support the interpretation of equilibration at both low temperatures (<750°C) and oxygen fugacities (< $\Delta$ NNO-1.0). The amphibole-free Green Tuff equilibrated at temperature and redox conditions greater than those for these amphiboles, but still well within the range of richterites with X-Fe<sub>tot</sub> = 0.6-0.8 ( $<910 \pm 10^{\circ}$ C), which suggests that something other than favorable T and fO2 may be required to stabilize this phase under these conditions.

Experiments on pantellerites have crystallized amphiboles using starting material from the Serra della Fastuca pumice (Di Carlo et al., 2010; Romano et al., 2020), Eburru pantellerite (Scaillet & Macdonald, 2006), and curiously in Green Tuff pumice (Romano et al., 2020) despite its absence in natural samples.



**Fig. 13.** (a) The ranges of T-fO<sub>2</sub> (relative to  $\Delta$ NNO) for the amphibole-bearing pantellerites calculated at 150 MPa from mineral equilibrium (from Fig. 12) plotted with the experimental stability limits for ferrorichterite with compositions of X-Fe<sub>tot</sub> = 1.0 and 0.8 (Charles, 1977) and the experimental conditions that resulted in the crystallization of amphibole for the GT pumice (GTP; Romano et al., 2020) and the Fastuca pumice (FTP;Di Carlo et al., 2010; Romano et al., 2020). (b) Experimental amphibole compositions (as X-Fe<sub>tot</sub>) from Di Carlo et al. (2010) and Romano et al. (2020) plotted versus temperature (T, °C) along with the range of natural amphibole compositions from this study.

Scaillet & Macdonald (2006) synthesized arfvedsonite at 150 MPa under water-rich conditions and oxygen fugacities (fO<sub>2</sub>) equivalent to two log units below the nickel-nickel oxide buffer ( $\Delta$ NNO-2) from a pantellerite from Eburru centre, Kenya. Sodic  $(Na_B > 1.0)$  and ferroan  $(X-Fe_{tot} > 0.9)$  amphiboles similar in composition to Mida were synthesized by Scaillet & Macdonald (2001, 2006), Di Carlo et al. (2010), and Romano et al. (2020) over a wide range of pressures (50-150 MPa) and water contents (1.8-6.0 wt%  $H_2O_m$ ) but always at low oxygen fugacity (< $\Delta$ NNO-1.0) and temperature (<700°C), consistent with the results of Charles (1975, 1977). Amphiboles crystallized from the Green Tuff pumice range in composition from richterite (X-Fe<sub>tot</sub> = 0.45-0.52) to ferrorichterite (X-Fetot = 0.63-0.85) at temperatures of 800°C and 750°C, respectively, in experiments conducted at 100 MPa and between  $\Delta NNO\text{-}0.5$  and – 2.5 with 2.5–4.5 wt%  $H_2O_m.$  The compositions of the experimentally derived amphiboles from these experiments are consistent with the buffered experiments of Charles (1977) and overlap in  $T-fO_2$  space with the values for the natural samples calculated from mineral equilibria at 150 MPa (Fig 13a).

Romano *et al.* (2020) have drawn attention to the difficulties in determining the stability of amphibole in pantellerites, partly because of their scarce occurrence and wide compositional range. Factors which have been invoked include high  $fF_2$  (Ernst, 1962; Conrad, 1984; Scaillet & Macdonald, 2001, 2006; Di Carlo *et al.*, 2010; Giesting & Filiberto, 2014; Scaillet *et al.*, 2016), and relatively high melt Na contents (Romano et al., 2020). Fluorine contents in amphiboles that formed at 680°C and 100 MPa are 0.78–0.86 wt% (Romano et al., 2020), higher than natural ones from Mida (0.47–0.57 wt% F). In samples of the naturally amphibole-free Green Tuff pumice, Romano et al. (2020) synthesized amphiboles more similar in composition to Altura and Gadir (X-Fe<sub>tot</sub>  $\approx$  0.63–0.85, Na<sub>B</sub>  $\approx$  0.85–1.00) at temperatures of 750° and 800°C at 100 MPa under similar to slightly more oxidizing conditions (< $\Delta$ NNO-0.5) over a wide range of H<sub>2</sub>O<sub>m</sub> (1.3–4.5 wt%). These experimental amphiboles are higher in F (1.0–2.4 wt%), similar to the Altura and Gadir amphiboles (0.5–1.8 wt% F), reinforcing the proposition that high fluorine contents may stabilize amphibole at higher temperatures.

There is a strong  $(R^2 = 0.94)$  correlation between X-Fe<sub>tot</sub> and temperature in amphiboles crystallized in the experiments of Di Carlo et al. (2010), Romano et al. (2020), and Scaillet & Macdonald (2006) (Fig. 13b) that may be used to estimate the temperature of crystallization in the natural samples. Amphiboles in Altura and Gadir samples suggest temperatures declining from 750° to 710° and 743° to 720°C, respectively, both overlapping with the range of temperatures estimated from clinopyroxene-ilmeniteaenigmatite-quartz equilibrium at 100 MPa, which may indicate that these phases co-saturated over this temperature interval. Amphiboles from Mida, however, have much higher X-Fetot, with values nearly identical to their experimental counterparts from Fastuca, suggesting crystallization at temperatures between 680° and 690°C, in agreement with values determined from mineral equilibria above. The higher A-site occupancy in the Mida amphiboles relative to both Altura and Gadir is also consistent with equilibration at lower oxygen fugacities for that former compared to the latter (Della Ventura et al., 2005).

#### Fayalite, Aenigmatite, and quartz stability

White et al. (2005), Macdonald et al. (2011), and Romano et al. (2020) described an antipathetic relationship between fayalite and aenigmatite, with the two phases coexisting in equilibrium on a univariant curve in T-aSiO<sub>2</sub> space at constant P and  $\Delta$ NNO. In experiments on metaluminous and peralkaline (comenditic) trachyte (Romano et al., 2018), olivine (Fa<sub>57</sub> to Fa<sub>88</sub>) has been found to be stable over a wide range of pressures (50–150 MPa), temperatures (750–950°C), and oxygen fugacities (although under reducing conditions,  $<\Delta$ NNO-0.6), with fayalite contents that positively correlate with the P.I. of the coexisting glass. In experiments on pantellerites, olivine has been a more elusive phase that has only been synthesized in the experiments of Romano et al. (2020), which crystallized fayalitic olivine (Fa<sub>90-93</sub>) at 750°C, 100 MPa, and  $\Delta$ NNO-1.3 to -1.9; at  $\Delta$ NNO-1.9 fayalite co-precipitated with aenigmatite, and under more reducing conditions ( $\Delta$ NNO-2.4) aenigmatite crystallized without fayalite. In these experiments (Di Carlo et al., 2010; Romano et al., 2020), aenigmatite crystallized under reducing conditions ( $<\Delta$ NNO-1.0) at temperatures between 725 and 680°C and pressures between 100 to 150 MPa. Under more reducing conditions (≤△NNO-2.4) aenigmatite crystallized at temperatures as high as 750°C at pressures of 50 to 150 MPa.

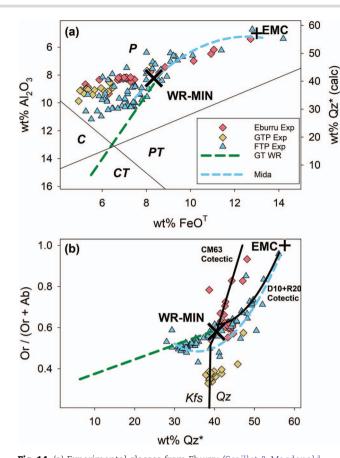
Fayalite is stable in pantelleritic melts until the silica activity relative to quartz saturation ( $aSiO_2[Qtz]$ ) is near unity, at which point aenigmatite ( $\pm$ quartz) begins to crystallize instead; at pressures of 100 MPa this occurs at ~750°C, when the P.I. of the melt is ~1.6 (White *et al.*, 2005; Macdonald *et al.*, 2011; Romano *et al.*, 2020). However, in some pantellerites—such as our samples from Mida and pantellerites from Eburru (Ren *et al.*, 2006)—fayalite and aenigmatite are in equilibrium to ~680°C and the P.I. of the melt is ~2.0. It has long been understood that olivine stability in silicic magmas is favored at lower silica activities (Carmichael et al., 1970; White et al., 2005; Macdonald et al., 2011) and that silica activity is inversely proportional to total pressure in magma (Nicholls et al., 1971), so it stands to reason that the stability field of fayalite would be expanded to lower temperatures at higher pressures as our geothermobarometric calculations suggest. Additional evidence in support of this conclusion comes from experimental relations in comendites, in which fayalite crystallized only at P  $\geq$  150 MPa and  $\Delta$ NNO-1.4 to -3.2 at temperatures as low as 661°C (Scaillet & Macdonald, 2001, 2003).

#### Glass compositions

The compositions of glasses from experiments conducted on the pantelleritic Green Tuff Pumice (GTP; Romano et al., 2020), the pantelleritic Fastuca della Serra Pumice (FTP; Di Carlo et al., 2010; Romano et al., 2020), and the Eburru pantellerite (EBU; Scaillet & Macdonald, 2006) are plotted in Fig. 13 along with LLOD curves regressed through the natural data for the Green Tuff and Mida datasets (see Fig. 6). Experimental glasses from GTP runs have low FeO<sup>T</sup> (Figure 14a) and are displaced along the feldspar-quartz cotectic to lower values of Or/(Or + Ab) (Figure 14b) than the main Green Tuff trend. This is certainly due to the nature of the starting material, a sample of the precursor air fall deposit of the Green Tuff that has a whole-rock composition similar to these glasses, which Liszewska et al. (2018) suggested resulted from late fractionation of aenigmatite > quartz > clinopyroxene > alkali feldspar (summarized in our Table 11). Therefore, although the experimental results from GTP can inform our understanding of the stability of amphibole and other phases, the whole-rock composition appears to represent an endpoint of a particular LLOD and will not be discussed further. FTP and EBU, on the other hand, have starting compositions that lie in the thermal valley and near the whole-rock minimum, respectively. The Mida glasses lie plot along the FTP trend, reaching the effective minimum composition. The phenocryst assemblage in the pumice is alkali feldspar, aenigmatite and clinopyroxene, with fayalite, amphibole and quartz occurring in minor amounts (Di Carlo et al., 2010; Romano et al., 2020) comparable to that in the Cuddia Mida rocks. Experimental FTP glasses with compositions approaching the EMC were performed at a wide range of pressures (50-150 MPa) and temperatures (725-800°C) but all required low relative oxygen fugacities ( $<\Delta$ NNO-3). Proportions of glass (melt), alkali feldspar, and quartz in all runs vary similarly with Qz\* regardless of P-T $fO_2$ , with values that agree with those predicted by mass balance models K and L for the Mida glass trend (Figure 15a). With respect to FeO<sup>T</sup>, however, only runs conducted at lower oxygen fugacities generally crystallized lesser (<4%) proportions of mafic minerals, which enabled iron enrichment also predicted by the mass balance models (Figure 15b). One notable exception to this is a single run (9–7; Di Carlo et al., 2010) at 680°C, 150 MPa, and  $\leq \Delta$ NNO-2.2 that has 21% amphibole and a very low melt fraction (0.03) that drove the melt composition to lower values of both Qz\* and FeO<sup>T</sup> and may represent near-solidus conditions.

### Summary

Evidence from equilibria in the natural mineral assemblages and the experimental results suggests similarly low (670–700°C) minimum temperatures for all three amphibole-bearing units, lower than those determined for the amphibole-free Green Tuff (~720°C). Oxygen fugacity is also systematically lower (between  $\Delta$ NNO-1.2 and -3.3) for amphibole-bearing units compared to amphibole-free systems, with the most reduced oxygen fugacities also associated with the most extreme glass compositions.

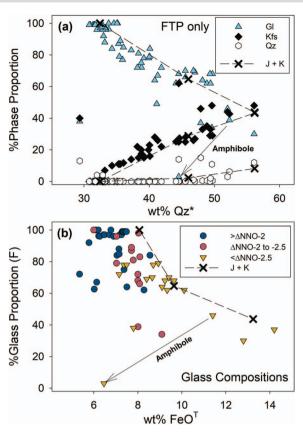


**Fig. 14.** (a) Experimental glasses from Eburru (Scaillet & Macdonald, 2006), GTP (Romano et al., 2020), and FTP (Di Carlo et al., 2010; Romano et al., 2020) plotted on a modified Macdonald (1974) diagram with the natural GT and Mida trends (regressed through the data in Fig. 6) shown with the whole-rock minimum (WR-MIN) and effective minimum composition (EMC). (b) The isobaric projection of Fig. 6a re-plotted in a cartesian system (cf. Wilke et al., 2017) with the ternary minimum and alkali feldspar-quartz cotectic from Carmichael & MacKenzie (1963) [CM63] and a new cotectic calculated from the results of Di Carlo et al. (2010) and Romano et al. (2020) [D10 + R20].

# Implications for silicic peralkaline systems

Numerous studies have concluded that silicic peralkaline rocks form by protracted fractional crystallization of alkali basalts, along LLOD that include mugearite, benmoreite, and metaluminous trachyte (Macdonald et al., 2021, and references therein). Whether metaluminous trachyte continues to evolve via fractional crystallization to comenditic (P.I. > 1.0, Fe/Al\* < 1.0) or to pantelleritic (P.I. >1.0, Fe/Al\* > 1.0) compositions depends on when iron oxides (magnetite, ilmenite) and Na-Fe silicates (aenigmatite, aegirine-augite, ferrorichterite) crystallize, which is primarily controlled by  $fO_2$ ,  $H_2O_m$ , and  $fF_2$  (Scaillet & Macdonald, 2001, 2003, 2006; Markl et al., 2010; Di Carlo et al., 2010; Marks et al., 2011; Romano et al., 2020.) Oxygen fugacity in particular is widely understood to be a critical variable in controlling the evolution of alkaline igneous rocks in general, with lower values associated with oxide-free assemblages, the crystallization of aegirine-rich pyroxenes, and the driving of residual melts to more strongly peralkaline compositions. The transition from comenditic to pantelleritic compositions is mirrored in intrusive sequences by the transition from miaskitic to agpaitic compositions (Mann et al., 2006; Markl et al., 2010; Marks et al., 2011; Giehl et al., 2013).

Although Pantelleria is the type locality for pantellerite, the pre-Green Tuff eruptive history of the island has been dominated



**Fig. 15.** (a) Percent phase proportions of glass (Gl), alkali feldspar (Kfs), and quartz (Qz) from experiments on FTP (Di Carlo *et al.*, 2010; Romano *et al.*, 2020) plotted versus the Qz\* of the glass compositions and with the results of mass balance models J and K from Mida (see Table 11 and Supplementary Table 11). (b) Percent phase proportions of glass only sorted by experimental oxygen fugacity values (relative to NNO) plotted versus wt% FeO<sup>T</sup> of the glass and with the results of mass balance models J and K from Mida. 'Amphibole' shows the effect of amphibole crystallization on melt composition under crystal-rich conditions (from 725°C [Run 4–7] to 680°C [Run 9–7] at 150 MPa; Di Carlo *et al.*, 2010).

by eruptions of comenditic trachyte to comenditic ignimbrites, with pantellerites occurring mostly as small inter-caldera volcanic centres, all of which have nearly identical trace-element enrichment patterns and incompatible trace-element ratios that strongly suggest derivation from a common parental magma along broadly similar LLOD (Rotolo et al., 2013, 2021; Jordan et al., 2018, 2021; Romano et al., 2022). The occurrence of both basalttrachyte-comendite and basalt-trachyte-pantellerite suites in the same volcanic centre has been documented in many peralkaline systems. In some instances, comendite and pantellerite occur as part of the same magma series, with the latter interpreted as the result of continued fractional crystallization of the former; in these series, pantellerite eruptive products are subordinate in volume to the comendites and are, relative to other pantellerites, only mildly peralkaline (P.I.  $\leq$  1.3). Examples of these include Torfajökull, Iceland (McGarvie et al., 2006), Olkaria, Kenya (Macdonald et al., 2008b; Marshall et al., 2009), and numerous volcanic centres in the Basin and Range Province, USA (Wallace et al., 1980; Conrad, 1984; Mathis, 1993; Henry et al., 2017). An exception is the Black Mountain Volcanic Centre in Nevada where a long sequence of comenditic eruptive systems was terminated by a strongly peralkaline ash flow (Macdonald et al., 2019). In other instances, comendite and pantellerite form the end-member of distinct magma series. In addition to Pantelleria, these include

Boseti, Ethiopia (Ronga et al., 2009), Nemrut, Turkey (Macdonald et al., 2015), Changbaishan, China-North Korea (Andreeva et al., 2019), and Terceira, Azores (Mungall & Martin, 1995, 1996).

On Pantelleria, ilmenite is the dominant oxide phase in pantellerites whereas ulvöspinel-rich magnetite is the dominant phase in comendites (Jordan, 2014; Jordan et al., 2021). Mungall & Martin (1995) observed the same for the comenditic Santa Barbara and pantelleritic Pico Alto volcanic suites at Terceira, Azores, which they attributed to differences in oxygen fugacity (higher for trachyte-comendite and lower for trachyte-pantellerite) that they postulated were the result of early degassing of water and H<sub>2</sub> from Santa Barbara trachytes. A salient feature of the Pico Alto pantellerites reported by Mungall & Martin (1996) are matrix glasses and melt inclusions with extreme compositions of up to 13.3 wt% FeO<sup>T</sup>, 3600 ppm Zr, and P.I.  $\approx$  4.3, which they also ascribed to very low oxygen fugacity and described as geochemically analogous to the hypersolvus (miaskitic) phase of the rare-metal rich Strange Lake granite in Québec-Labrador, Canada (Siegel et al., 2018).

The relative oxygen fugacity of silicic magmas typically overlaps with, and is inherited from, the parental mafic magmas (Carmichael, 1991). Major and trace element data strongly suggest that the four suites discussed in this study were derived from a similar parent along similar differentiation paths from transitional basalt to trachyte and through a 'thermal valley' dominated by the fractionation of alkali feldspar to pantelleritic compositions that plot near a 'whole-rock minimum' in a Q-Or-Ab-Ac-Ns projection. Compositional differences between the suites can be attributed to variability in the fractionating mafic phases, which is mainly controlled by the oxygen fugacity during crystallization.

Although pressure has negligible effect on relative oxygen fugacity, we posit that the differences in oxygen fugacity between the suites-and most markedly between the Green Tuff and Mida suites-are most likely due to differences between the depths of the magma reservoirs and the hydrous nature of these magmas, inherited from their mafic parents. Increases in relative oxygen fugacity with decreasing temperature in pantelleritic systems have been described by Mungall & Martin (1995, 1996) for Terceira and Liszewska et al. (2018) for the Green Tuff, who each attributed the change to a reaction between Fe-rich silicate melt and H<sub>2</sub>O in the magma via thermal dissociation of H<sub>2</sub>O. For this mechanism to be viable, the magma must be degassing and  $H_2$  (reduced from  $H^+$  by  $Fe^{2+}$  by the reaction  $2FeO + H_2O = Fe_2O_3 + H_2$ ) able to diffuse (Sato, 1978). Diffusion of H<sub>2</sub> in silicate glasses has been experimentally shown to be rapid (Zhang & Ni, 2010) and in order to degas the magma must be saturated in a fluid phase. Water saturation in pantelleritic melts is primarily a function of pressure (Romano et al., 2021), with lower temperatures and higher sodium contents also contributing to increased solubility (Stabile et al., 2018; Allabar et al., 2022). We therefore suggest that although all these pantelleritic suites were sourced from a similar mafic parent along similar liquid lines of descent, differences in the evolutionary paths beyond the 'minimum' are primarily controlled by the depth of emplacement. Deeper magma reservoirs, such as Mida, had high water contents but remained undersaturated relative to a fluid phase due to higher pressures, which also kept the water dissolved in the melt, oxygen fugacity low, and enabled the crystallization of amphibole. Shallower reservoirs, such as the one that supplied the Green Tuff, had similarly high water contents but became saturated at lower pressures that allowed the water to degas, which ultimately increased the relative oxygen fugacity

and inhibited the crystallization of amphibole. The depth-control will also have important implications for the eruptive potential and style of the system. Furthermore, whether or not the magma degases strongly influences whether soluble rare-metals are retained in the system or lost.

While this discussion is aimed primarily at peralkaline systems, it would be instructive to assess non-peralkaline systems for evidence of a relationship between oxygen fugacity and pressure, for example, the high-Fe rhyolites associated with tholeiitic suites in Iceland.

# CONCLUSIONS

- Despite a common basalt-trachyte origin and evolution through a common 'thermal valley', several evolutionary paths exist and there is no unique end point for peralkaline silicic magmas, forming a 'cotectic delta' towards multiple glass minima beyond the whole-rock minimum on the alkali feldspar-quartz cotectic.
- Experiments on both comendites and pantellerite and shown that fractionation at higher  $fO_2$  expands the thermal stability of oxides, which results in residual liquids that are less, or similarly, peralkaline and lower in FeO<sup>T</sup> than the bulk compositions, whereas at very low  $fO_2$  (< $\Delta$ NNO-2.5) residual liquids have compositions with P.I. up to ~4.3 and up to 14.3 wt% FeO<sup>T</sup> (Scaillet & Macdonald, 2003, 2006; Di Carlo *et al.*, 2010; Romano *et al.*, 2020). The close correlation between the compositions of the natural Mida and experimental glasses convincingly demonstrates the validity of this conclusion and suggests that our lower estimate for oxygen fugacity in the Mida system (~ $\Delta$ NNO-3.0) is accurate.
- A key control is pressure. At high pressures, water stays dissolved in the melt and is available to form amphiboles; at lower pressures, water will exsolve and thermally dissociate resulting in an increase in relative oxygen fugacity. Thus, at higher pressures water will remain dissolved and maintain lower  $fO_2$  in the melt, resulting in a melt that is both waterrich and reduced (cf. Di Carlo *et al.*, 2010), providing ideal conditions for amphibole crystallization. Amphibole formed in the Mida, Altura, and Gadir suites as a result of low temperatures, low  $fO_2$  and high  $H_2O_m$ . Its formation was inhibited in the Green Tuff by higher  $fO_2$  and lower  $H_2O_m$ .
- Evidence is accumulating on how the depth and nature of the magma reservoirs at Pantelleria have changed over time (Neave, 2020; Jordan *et al.*, 2021; Rotolo *et al.*, 2021). Here we have described a mechanism by which the depth of the reservoir can affect the ultimate fractionation history of the magma.

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# DATA AVAILABILITY

The data underlying this article are available in the article and in its online supplementary material.

# Supplementary data

Supplementary data are available at Journal of Petrology online.

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