

9-26-2023

On the Propagation of Whistler-Mode Waves in the 2 Magnetic Ducts

Salman A. Nejad
Embry-Riddle Aeronautical University

Anatoly V. Streltsov
Embry Riddle Aeronautical University, streltsa@erau.edu

Follow this and additional works at: <https://commons.erau.edu/publication>



Part of the [Atmospheric Sciences Commons](#)

Scholarly Commons Citation

Nejad, S. A., & Streltsov, A. V. (2023). On the Propagation of Whistler-Mode Waves in the 2 Magnetic Ducts. *JGR-Space Physics*, (). <https://doi.org/10.1029/2023JA031839>

This Article is brought to you for free and open access by Scholarly Commons. It has been accepted for inclusion in Publications by an authorized administrator of Scholarly Commons. For more information, please contact commons@erau.edu.

On the Propagation of Whistler-Mode Waves in the Magnetic Ducts

Salman A. Nejad¹, Anatoly V. Streltsov¹

¹Department of Physical Sciences, Embry-Riddle Aeronautical University, Daytona Beach, FL 32114,
USA

Key Points:

- MMS satellites observe whistler-mode waves inside the low-magnetic ducts.
- We provide an analytical explanation for the occurrence of leaking ducts and show that the low-magnetic ducts are prone to leak.
- Simulations confirm our analytical prediction of the most trapped whistler-mode waves in the low-magnetic ducts, provided by the least degree of leakage.

Corresponding author: S. A. Nejad, abarghos@my.erau.edu

This article has been accepted for publication and undergone full peer review but has not been through the copyediting, typesetting, pagination and proofreading process, which may lead to differences between this version and the [Version of Record](#). Please cite this article as [doi: 10.1029/2023JA031839](https://doi.org/10.1029/2023JA031839).

This article is protected by copyright. All rights reserved.

Abstract

This paper studies extremely-low frequency (ELF) whistler-mode waves' behavior within small-scale magnetic field irregularities in the Earth's magnetosphere, known as magnetic ducts. Based on the magnetic fields' magnitude inside and outside these ducts, they are categorized as high-magnetic ducts (HBD) and low-magnetic ducts (LBD). Using the whistler-mode dispersion relation analysis, our primary focus is to show that LBDs are prone to leak electromagnetic energy outside the duct. We further investigate the hypothesis that whistlers can propagate within LBDs without any signal loss when the width of the duct corresponds to an integer multiple of the perpendicular wavelengths of the waves inside it. This condition offers a straightforward and effective method for identifying non-leaking eigenmodes of LBDs. Our analysis of this non-leaking condition reveals that every LBD possesses a finite number of non-leaking eigenmodes directly proportional to the duct's width and the magnitude of the ambient magnetic field within it. The analytical results are then validated using two-dimensional, time-dependent simulations of the electron-Magnetohydrodynamics (EMHD) model. Also, we model the non-leaking propagation of an ELF whistler-mode wave observed inside the LBD by the NASA Magnetospheric Multiscale mission (MMS) satellites.

1 Introduction

Whistler-mode waves have attracted attention from the space plasma community due to their ability to interact with high-energy electrons in the Earth's radiation belt through cyclotron resonance. This interaction changes the pitch angle of energetic particles, causing them to precipitate out of the magnetosphere. By intentionally introducing controlled injections of whistler-mode waves from Earth or space into the magnetosphere, the presence of energetic particles in the Earth's radiation belt can be reduced, thereby creating a safer environment for electronics and crews on space platforms (Inan et al., 1985, 2003).

As a notable characteristic, whistler-mode waves in the magnetosphere propagate along the ambient magnetic field, guided by field-aligned plasma density inhomogeneities referred to as the density ducts, which can be formed as the low-density ducts (LDD) and the high-density ducts (HDD). The presence of these ducts enables the propagation of whistler-mode waves over considerable distances in the Earth's magnetosphere, including from one hemisphere to another, with minimal attenuation.

44 Recently, (Streltsov & Nejad, 2023) have pointed out the occurrence of small-scale,
45 localized packages of ELF waves within the inhomogeneities of the ambient magnetic field
46 in the equatorial magnetosphere in the data gathered by the NASA Magnetospheric Mul-
47 tiscule Mission (MMS) satellites, which are interpreted as another mechanism of ELF
48 guiding that happens through the magnetic structures, referred to as high-magnetic duct
49 (high- B duct or HBD) and low-magnetic duct (low- B duct or LBD).

50 For almost 70 years, an extensive study has been done on different density struc-
51 tures, guiding whistler-mode waves in space and laboratory plasmas (Storey, 1954; Nunn,
52 1974; Karpman et al., 1974; Stenzel, 1976; Omura et al., 1991; Trakhtengerts et al., 1996;
53 Nunn & Smith, 1996; Stenzel, 1999; Hobara et al., 2000; Kostrov et al., 2000; Helliwell,
54 1965; Sazhin, 1993; Kondrat'ev et al., 1999).

55 An important feature of density ducts is the ability to guide ELF waves without
56 any leakage for specific conditions. These conditions have been studied in different works.
57 Notably, one of the key findings from these studies is that the leakage of whistler-mode
58 waves from the HDD with very narrow walls can be minimized when the HDD channel
59 encompasses an integer number of the smallest perpendicular wavelengths and the ra-
60 tio of the largest to the smallest internal perpendicular wavelengths is large (Zaboronkova
61 et al., 1992).

62 Further studies conducted by (Streltsov et al., 2007; Streltsov, 2007) shed light on
63 the complete elimination of leakage from HDDs with narrow walls, given certain condi-
64 tions. Specifically, the leakage is entirely eliminated when the width of the duct chan-
65 nel corresponds to an integer multiple of both the smallest and largest perpendicular wave-
66 lengths of the internal waves.

67 Despite the presence of magnetic structures in the inner magnetosphere, the trap-
68 ping mechanism of ELF whistler-mode waves within them has not been extensively stud-
69 ied. (Yu & Yuan, 2022) investigated the ducting of whistler-mode waves by magnetic dips
70 and peaks by analyzing the wave refractive index. In a recent study, (Streltsov & Ne-
71 jad, 2023) utilized the whistler-mode dispersion relation to derive analytical criteria for
72 wave trapping in HBD and LBD and to identify thresholds in the magnitude of the am-
73 bient magnetic field that facilitates the trapping of waves with specific frequencies. They
74 also determined the ranges of parallel and perpendicular wavelengths of the waves trapped
75 in LBDs and HBDs.

76 In this study, we examine the leakage from LBDs with narrow walls and investi-
 77 gate the conditions under which the leakage can be entirely eliminated. Specifically, we
 78 explore the scenarios where the width of the duct channel corresponds to an integer mul-
 79 tiple of both the smallest and largest perpendicular wavelengths of the internal waves.
 80 This condition will be further validated through numerical simulations utilizing the full
 81 wave model. By conducting these investigations, we will gain a comprehensive under-
 82 standing of the mechanisms involved in eliminating leakage from LBDs with narrow walls.

83 2 Observation

84 Figure 1 shows an event demonstrating a correlation between the wave and the mag-
 85 netic structure. It consists of the electric field, magnetic field, and plasma density mea-
 86 sured by the MMS1 satellite from 19:09:20 to 19:09:40 UT on March 06, 2016. The data
 87 are obtained by the MMS1 FPI/DES electron number density, and MMS1 DSP from FIELDS
 88 Instrument Suite (Torbert et al., 2016), and magnetic field vector from MMS1 Flux-Gate
 89 Magnetometer (FGM) instrument (Russell et al., 2016).

90 Figure 1a shows with a color pallet the power spectral density (PSD) of the x -component
 91 of the electric field (in the GSE coordinate system) in the frequency range 150–330 Hz
 92 and the magnitude of the background magnetic field (white line). Figure 1b shows the
 93 corresponding electron density. The minimum magnitude of the ambient magnetic field
 94 inside the duct is 33.37 nT (here the electron cyclotron frequency $\omega_{ce} = 5.87 \times 10^3$ rad/s
 95 and the lower hybrid frequency $\omega_{LH} = 1.36 \times 10^2$ rad/s), and it changes during this
 96 time interval by ≈ 16.91 nT or 50.6% compare to the maximum values outside the duct
 97 of 50.28 nT. For comparison, during this time interval, the electron density changes from
 98 83.4 cm^{-3} to 81.9 cm^{-3} or by 1.8%. The density inside the duct is 81.9 cm^{-3} ($\omega_{pe} =$
 99 5.1×10^5 rad/s). The frequency of the wave in the duct is 248 Hz, and the width of the
 100 duct is ≈ 13.483 km.

101 Figure 1c shows the trajectory and the location of MMS satellites in the GSE X-
 102 Y plane on March 6, 2016. The red dot indicates the location of MMS satellites at 19:00:00
 103 UT.

104 In the next section, we present an analytical investigation of the conditions where
 105 any LBD with arbitrary values of the duct width and ambient magnetic field guides some
 106 whistlers without leakage. Our approach is based on the analysis of the dispersion re-

107 lation for whistler-mode waves, and it is similar to the analysis of the waveguiding via
 108 density ducts developed earlier by (Streltsov et al., 2007; Streltsov, 2007). We also use
 109 the same approach and simulations to model an event observed by the MMS1 satellite
 110 in Figure 1.

111 3 Model

The ELF waves ducting in the plasma density inhomogeneities and the magnetic field are investigated by the analysis of the dispersion relation derived from the linearized equations of electron-MHD (EMHD), also called quasi-longitudinal approximation of EMHD, where the displacement current in the Ampere's law is omitted (Helliwell, 1965; Gordeev et al., 1994; Sazhin, 1993). This model is valid for the cases which satisfy the following conditions

$$\omega_{LH} < \omega < \omega_{ce} \ll \omega_{pe}, \quad (1)$$

112 where, ω_{LH} is the lower hybrid frequency, ω is the angular frequency of the wave; ω_{ce}
 113 is the electron gyrofrequency, and ω_{pe} is the electron plasma frequency.

The dispersion relation of ELF-mode waves derived from the quasi-longitudinal EMHD in the homogeneous media is

$$k^2 - \frac{\omega_{ce}}{\omega} k_{\parallel} k + \frac{1}{\lambda_e^2} = 0 \quad (2)$$

114 where k_{\parallel} and k_{\perp} are parallel and perpendicular to \mathbf{B}_0 wavenumber, with $k^2 = k_{\parallel}^2 +$
 115 k_{\perp}^2 , and $\lambda_e = c/\omega_{pe}$ is the electron plasma skin depth.

The relation (2) can be used to express B in terms of ω , k_{\perp} , k_{\parallel} , and ω_{pe} in the form of the following function.

$$B = \frac{m_e \omega}{e k_{\parallel}} \left[\frac{1}{k} \left(k^2 + \frac{1}{\lambda_e^2} \right) \right]. \quad (3)$$

116 Figure 2 illustrates the meaning of B_1 and B_2 , where the plot of B , i.e. equation (3),
 117 is depicted as a function of k_{\perp} , k_{\parallel} , and ω_{pe} . It shows that B_1 is the value of the B when
 118 $k_{\perp} = 0$, and B_2 is the minimum value of the B when $k_{\perp 1} = k_{\perp 2}$. Here,

$$B_1 = \frac{m_e}{e} \omega \left(1 + \frac{1}{k_{\parallel}^2 \lambda_e^2} \right), \quad (4)$$

$$B_2 = \frac{m_e}{e} \omega \left(\frac{2}{k_{\parallel} \lambda_e} \right). \quad (5)$$

119 and,

$$k_{\perp 1,2} = k_{\parallel} \left[\frac{\omega_{ce}^2}{4\omega^2} \left(1 \mp \sqrt{1 - \frac{B_2^2}{B^2}} \right)^2 - 1 \right]^{\frac{1}{2}}. \quad (6)$$

120 Figure 2 reveals another important feature of the LBDs. It shows that for the partic-
 121 ular value of ω , there are two different real solutions, $k_{\perp 1,2}$, for any B in the range $B_2 <$
 122 $B < B_1$; one real solution outside the duct for $B > B_1$, denoted by $k_{\perp 3}$; and no real
 123 solution exists for $B < B_2$.

With all that said above, the ELF wave can propagate inside the LBD if B_I and B_L , the magnetic fields inside and outside the duct respectively, satisfy the following condition.

$$B_2 < B_I < B_1 < B_L. \quad (7)$$

124 If the magnitude of B inside the channel allows propagation of at least one ELF wave
 125 with $B_2 < B$, then for any $B_L > B_1$, there always exists a propagating wave with the
 126 same ω and k_{\parallel} outside the duct. Therefore, the waves inside the duct may couple to the
 127 propagating wave outside and carry energy away from the duct, resulting in the leak-
 128 age of electromagnetic energy from low- B ducts.

129 The results by (Streltsov et al., 2007) lead to an important condition where the leak-
 130 age from high-density ducts with thin walls vanishes. Hence, for a duct with a width of
 131 $2L$, the leakage is identically zero if

$$\begin{cases} 2L = l\lambda_{\perp 1} \\ 2L = m\lambda_{\perp 2} \end{cases} \quad \text{or} \quad \begin{cases} k_{\perp 1}L = l\pi \\ k_{\perp 2}L = m\pi \end{cases} \quad (8)$$

132 with both l and m as integer numbers.

133 Due to the same leaking features, this condition remains valid for the low- B ducts.
 134 The following section will study the discrete non-leaking eigenmodes of LBDs.

135 3.1 Non-leaking Eigenmodes of Low- B Ducts

136 Following (Streltsov et al., 2007), we introduce dimensionless variables $\Omega = \omega/\omega_{ce}$,
 137 $\Omega_p = \omega_{pe}/\omega_{ce}$, $K_{\perp} = k_{\perp}c/\omega_{ce}$, and $K_{\parallel} = k_{\parallel}c/\omega_{ce}$. In these variables, relation (2) has
 138 two solutions for K_{\perp} as the function of K_{\parallel} , Ω , Ω_p as

$$K_{\perp 1,2} = K_{\parallel} \left[\frac{1}{4\Omega^2} \left(1 \mp \sqrt{1 - \left(\frac{2\Omega\Omega_p}{K_{\parallel}} \right)^2} \right)^2 - 1 \right]^{\frac{1}{2}}. \quad (9)$$

139 These two perpendicular wave numbers represent solutions propagating inside the LBD,
 140 where the relation (7) is satisfied. Hence, $K_{\perp 1,2}$ are real when K_{\parallel} , Ω_p , and Ω satisfy in-

141 equalities

$$2\Omega < \frac{K_{\parallel}}{\Omega_p} < \sqrt{\frac{\Omega}{1-\Omega}}. \quad (10)$$

142 These inequality relations are the same for HDDs as in (Streltsov, 2007). Figure 3 shows
 143 the area satisfying these inequalities.

144 Applying conditions $K_{\perp_1}L = l\pi$, and $K_{\perp_2}L = m\pi$ to the expressions (9), one
 145 gets

$$1 - \sqrt{1 - \left(\frac{2\Omega\Omega_p}{K_{\parallel}}\right)^2} = 2\Omega\sqrt{\left(\frac{l\pi}{K_{\parallel}L}\right)^2 + 1}, \quad (11)$$

$$1 + \sqrt{1 - \left(\frac{2\Omega\Omega_p}{K_{\parallel}}\right)^2} = 2\Omega\sqrt{\left(\frac{m\pi}{K_{\parallel}L}\right)^2 + 1}. \quad (12)$$

146 Solutions of the relations (11) and (12) determine families of curves, referred to as l -curves
 147 and m -curves, inside the region that is defined by (10). The intersections of these curves
 148 provide a finite set of discrete non-leaking eigenmodes of an LBD. Figure 3 shows some
 149 of these intersections for specific values of l and m .

150 4 Simulation

151 The simulations are based on a quasi-longitudinal electron-MHD model, consist-
 152 ing of equations for the wave electric field, \mathbf{E} , magnetic field, \mathbf{B} , and the electron veloc-
 153 ity, \mathbf{v} (Streltsov et al., 2006)

$$\frac{m_e}{\mu_0 n_e e^2} \nabla \times \nabla \times \mathbf{E} + \mathbf{E} = -\frac{m_e}{e} (\mathbf{v} \cdot \nabla) \mathbf{v} - \mathbf{v} \times \mathbf{B}, \quad (13)$$

$$\frac{\partial \mathbf{v}}{\partial t} = \frac{1}{\mu_0 n_e e} \nabla \times \nabla \times \mathbf{E}, \quad (14)$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}. \quad (15)$$

154 The finite-difference, time-domain (FDTD) technique is employed to numerically imple-
 155 ment these equations within a rectangular domain (x, z) (Streltsov et al., 2006). The z -
 156 direction aligns with the background magnetic field, while the plasma density and back-
 157 ground magnetic fields exhibit homogeneity in the z -direction but inhomogeneity in the
 158 x -direction. The domain extends from $-l_z/2$ to $l_z/2$ in the z -direction and from $-l_x/2$
 159 to $l_x/2$ in the x -direction.

160 The boundary conditions in the z -direction are set to be periodic, with l_z being equiv-
 161 alent to one λ_{\parallel} . On the boundaries in the x -direction, the values of \mathbf{E} , \mathbf{B} , and \mathbf{v} are all

162 set to zero, following the Dirichlet boundary conditions. The simulations begin with the
 163 initial conditions for \mathbf{E} , \mathbf{B} , and \mathbf{v} as described in (Streltsov et al., 2006).

164 To solve the Helmholtz equation (13) for \mathbf{E} at every time step, the method of suc-
 165 cessive overrelaxation (SOR) is used. A third-order predictor-corrector algorithm is used
 166 to advance the equations in time, with the Adamse-Bashforth method as a predictor and
 167 the Adamse-Moulton method as a corrector (Streltsov et al., 2006).

168 4.1 Leaking/Non-leaking Low- B Duct Model

169 Several numerical experiments were conducted to verify the accuracy of the pre-
 170 diction, which states that leakage from LBD is minimal under conditions (8).

171 To achieve this goal, we first simulate the propagation of whistler-mode waves within
 172 a model LBD with parameters that satisfy the relations (8). Then, we purposely disrupt
 173 the non-leaking conditions by modifying the width of the duct in two distinct schemes,
 174 e.g., by shrinking and extending the duct size.

175 In these simulations, we assume that the plasma density remains uniform in the
 176 x and z directions, with $n = 60 \text{ cm}^{-3}$. On the other hand, the magnetic field maintains
 177 homogeneous in the z direction while exhibiting inhomogeneity in the x direction. Within
 178 the duct, the magnetic field has a magnitude of $B_I = 37.35 \text{ nT}$, while it measures $B_L =$
 179 60 nT outside the duct.

180 The wave frequency is set to $f = 302 \text{ Hz}$, accompanied by a chosen parallel wave-
 181 length of $\lambda_{\parallel} = 7.298 \text{ km}$. For these specific wave and plasma parameters, equations (4),
 182 (5), and (6) yield the following values: $B_1 = 41.64 \text{ nT}$, $B_2 = 36.5 \text{ nT}$, $k_{\perp 1} = 0.79$
 183 rad/km (with $\lambda_{\perp 1} = 7.92 \text{ km}$), and $k_{\perp 2} = 1.59 \text{ rad/km}$ (with $\lambda_{\perp 2} = 3.94 \text{ km}$).

184 To model a non-leaking duct, the initial width of the duct is chosen as $2L = 7.94$
 185 km . This size and the wave parameters calculated above satisfy non-leaking relations $Lk_{\perp 1} =$
 186 π (or $2L = \lambda_{\perp 1}$), and $Lk_{\perp 2} = 2\pi$ (or $2L = 4\lambda_{\perp 2}$) inside the channel. To smash the
 187 non-leaking case, we repeat the simulation of the same wave two more times by only chang-
 188 ing the width of the duct to 60% and 140% of its primary size.

189 Figure 4 shows the results of the simulations of the three situations mentioned above,
 190 accompanied by the propagation of the same wave in the homogeneous background mag-

191 netic field. These results show the wave simulations during the time interval of 331 ms
 192 or 100 wave periods.

193 Based on the results presented in Figures 4a, 4b, and 4c, it can be concluded that
 194 the wave characterized by the specific values of ω and λ_{\parallel} is confined within the LBD,
 195 aligning closely with the prediction made using the analytical criteria (7). Figure 4d pro-
 196 vides another confirmation for the occurrence of ducting due to the presence of inhom-
 197 ogeneity in the background magnetic field, where in the absence of that inhomogeneity,
 198 the same wave propagates at a significant angle relative to the magnetic field.

199 Figure 4a shows the simulation of the propagation of the whistler-mode wave in
 200 the shrunken duct with $2L = 0.6\lambda_{\perp 1}$. The propagation of the whistler-mode wave in
 201 the initial duct with $2L = \lambda_{\perp 1}$ is shown in Figure 4b, while the extended duct, for which
 202 $2L = 1.4\lambda_{\perp 1}$, is shown in the Figure 4c.

203 Figures 4a', 4b', and 4c' show profiles of B_0 across \mathbf{B}_0 for the shrunken, original,
 204 and extended ducts, respectively. Also, the dynamics of $E_x(x, z = 0, t)$ for the shrunken,
 205 original, and extended ducts are illustrated respectively in Figures 4a'', 4b'', and 4c''.

206 The simulations' results in Figure 4 validate the theoretical prediction that leak-
 207 age from the LBD can be entirely eliminated by ensuring the width of the duct chan-
 208 nel aligns with an integer multiple of perpendicular wavelengths of internal waves. This
 209 result is evident by comparing the amplitude of E_x component of the propagating wave
 210 in Figures 4a'', 4b'', and 4c'', whilst the width of the duct deviates (reduced or extended)
 211 from the initial width in Figure 4b', the amplitude of E_x gets smaller, showing that the
 212 energy is leaking from the duct.

213 These simulations highlight the importance of maintaining precise alignment be-
 214 tween the channel width and the integer number of perpendicular wavelengths to guide
 215 the most trapped wave, as any mismatch between them results in substantial leakage from
 216 the duct.

217 4.2 Leaking/Non-leaking Observed LBD Event

218 As illustrated in Figure 1, the observed event showcases the confinement of whistler-
 219 mode waves within the LBD. In our analysis, we assume that the plasma density and
 220 the background magnetic field exhibit homogeneity along the z direction but display in-

221 homogeneity along the x direction. The plasma parameters within the duct are obtained
 222 from the observations presented in Figure 1. The wave frequency is 248 Hz. Also, $B_0 =$
 223 33.37 nT, and $n_0 = 81.92$ cm $^{-3}$, representing the minimum value of the magnetic field
 224 inside the duct and the corresponding electron density.

225 To simulate the observed propagating ELF wave inside the LBD, we need to es-
 226 tablish initial conditions by calculating k_{\parallel} in a way that the wave can be trapped inside
 227 the duct. (Williams & Streltsov, 2021) showed that the whistler-mode waves cannot prop-
 228 agate along a high-density duct for every parallel wavelength. Due to the existing sim-
 229 ilarity between HDDs and LBDs, not all parallel wavelengths will be trapped inside LBDs.
 230 For this particular event, $0.903 < k_{\parallel} < 1.023$ provides two real $k_{\perp s}$, meaning that the
 231 wave may be guided by the duct.

232 It is also important to note that the approach used to set the simulation param-
 233 eters for the non-leaking and leaking cases of the event differs from the one employed
 234 in simulating the LBD model. In the LBD model, the wave parameters, e.g., λ_{\parallel} and $\lambda_{\perp s}$,
 235 were kept fixed while altering the duct width to create various non-leaking and leaking
 236 scenarios using relations (8). However, when dealing with an observed event, it is not
 237 admissible to modify the width of the duct. Consequently, the wave parameters are ad-
 238 justed for the fixed duct width to generate different scenarios for non-leaking and leak-
 239 ing conditions.

240 • Wave I

241 To simulate the non-leaking propagation, k_{\parallel} should be chosen in a way that sat-
 242 isfies the condition (8). This goal will be achieved by choosing $k_{\parallel} = 0.98$ rad/km
 243 ($\lambda_{\parallel} = 6.89$ km), resulting to $Lk_{\perp 1} = 2\pi$ (or $2L = 2\lambda_{\perp 1}$), and $Lk_{\perp 2} = 3\pi$ (or
 244 $2L = 3\lambda_{\perp 2}$) inside the channel.

245 Based on this wave and plasma parameters of the event, relations (4), (5), and (6)
 246 yield the following values: $B_1 = 39.69$ nT, $B_2 = 33.05$ nT, $k_{\perp 1} = 1.168$ rad/km
 247 (with $\lambda_{\perp 1} = 5.38$ km), and $k_{\perp 2} = 1.726$ rad/km (with $\lambda_{\perp 2} = 3.64$ km).

248 • Wave II

249 To investigate the leaking situation, we will simulate the wave with the $k_{\parallel} = 0.925$
 250 rad/km ($\lambda_{\parallel} = 6.792$ km), resulting in $Lk_{\perp 1} = 1.8\pi$ (or $2L = 1.8\lambda_{\perp 1}$), and $Lk_{\perp 2} =$
 251 3.4π (or $2L = 3.4\lambda_{\perp 2}$) inside the channel. Based on the wave and plasma param-

eters, $B_1 = 38.83$ nT, $B_2 = 32.59$ nT, $k_{\perp 1} = 1.008$ rad/km (with $\lambda_{\perp 1} = 6.232$ km), and $k_{\perp 2} = 1.901$ rad/km (with $\lambda_{\perp 2} = 3.303$ km).

The outcomes of these simulations, alongside the wave propagation in a homogeneous background magnetic field, are presented in Figure 5. All these results exhibit wave propagation within a time interval of 403 ms or 100 wave periods.

Figure 5 shows that the waves I and II, characterized by the above-mentioned values of ω and λ_{\parallel} s, are trapped inside the LBD, following the analytical prediction made by (7). Figure 5c illustrates the dispersion of the whistler-mode wave I in the absence of the magnetic duct.

Figure 5a illustrates the simulation's result of the propagation of a whistler-mode wave I in the duct, satisfying the relations $2L = 2\lambda_{\perp 1}$, and $2L = 3\lambda_{\perp 2}$. On the other hand, Figure 5b presents the simulation's outcome of the propagation of whistler-mode wave II in the same duct, where the relations $2L = 1.8\lambda_{\perp 1}$, and $2L = 3.4\lambda_{\perp 2}$ are satisfied. The Profiles of B_0 across \mathbf{B}_0 for these cases are presented in Figures 5a' and 5b', respectively.

By comparing the dynamics of $E_y(x, z = 0, t)$ for the ducts in Figure 5, one can observe the difference between the propagation of the ELF waves in three different situations, i.e., in the presence of the LBD ducts where its width corresponds to an integer multiple of perpendicular wavelengths of the internal wave shown in Figure 5a", in the presence of the LBD ducts where its width does not correspond to an integer multiple of perpendicular wavelengths of the internal wave shown in Figure 5b", and in the absence of the LBD shown in Figure 5c".

5 Conclusions

In this paper, we study the property of LBD to guide whistler-mode waves with minimum attenuation. This is conducted by analyzing the dispersion relation obtained from the linearized equations of the quasi-longitudinal EMHD (Electromagnetic Hydrodynamics) model. Our analysis demonstrates that the low-magnetic duct can leak energy outside due to the potential coupling between the waves propagating inside and outside the duct channel. This property of magnetic ducts is similar to the property of high-

281 density ducts, where they can leak electromagnetic energy due to coupling with the waves
282 outside the channel.

283 To expand on the earlier studies conducted by (Zaboronkova et al., 1992; Streltsov,
284 2007) on high-density ducts, our research offers new findings for LBDs. Specifically, we
285 discover that an LBD of a specific width and plasma density can guide a finite number
286 of discrete eigenmodes without experiencing any attenuation.

287 The central concept explored in this study is the complete elimination of whistler
288 leakage from the LBD by ensuring that the width of the duct aligns with an integer mul-
289 tiple of the perpendicular wavelengths of internal waves. This condition provides a straight-
290 forward and efficient method for identifying non-leaking eigenmodes of LBDs.

291 Finally, by conducting two-dimensional, time-dependent simulations using the com-
292 plete set of EMHD equations, accounting for the inhomogeneity of the magnetic field of
293 a modeled LBD, we demonstrate that the leak-free condition which is established as an
294 integer ratio between the channel width and the perpendicular wavelength of the wave,
295 allows for the propagation of whistler-mode waves inside the LBD with the least atten-
296 uation. Furthermore, we extend our investigation to the propagation of an ELF whistler-
297 mode wave propagating in an observed LBD event.

298 **Acknowledgments**

299 This research was supported by the Air Force Office of Sponsored Research Grants FA9453-
300 322 21-2-0039.

301 **Data Availability**

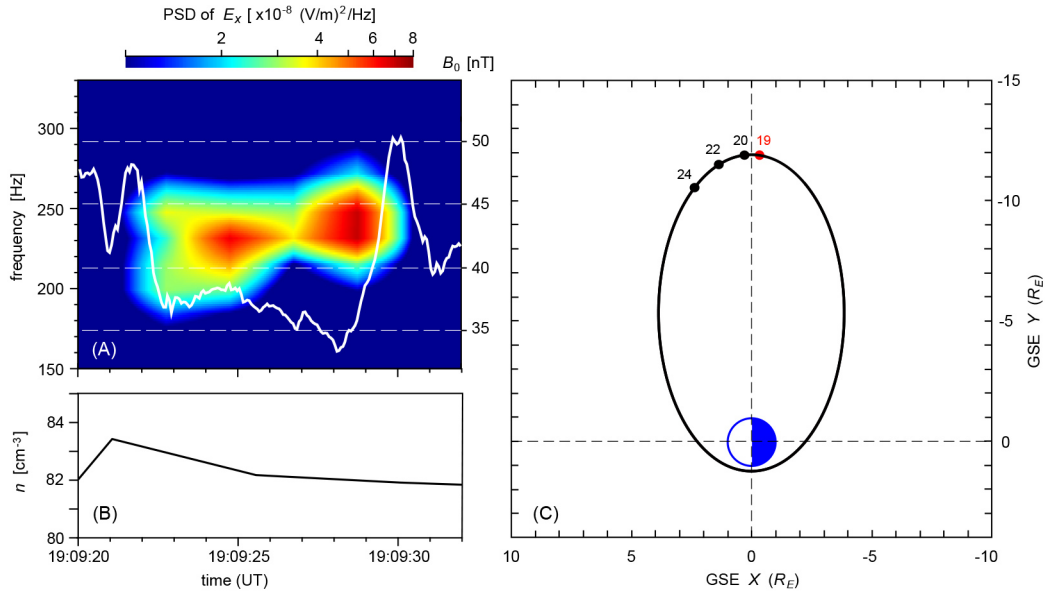
302 The MMS data used in this study are available from <https://lasp.colorado.edu/mms/sdc/public/datasets/>
303 and <https://cdaweb.gsfc.nasa.gov/>.

304 The Linux executable code (r1000), data files used to run the code (rbsp.dat_newN.dat
305 and Dens_Bfield_In.dat), and the results from the simulations (ExfieldS.dat) shown in
306 Figures 4 and 5 are available from <https://doi.org/10.6084/m9.figshare.c.6835749.v1>.

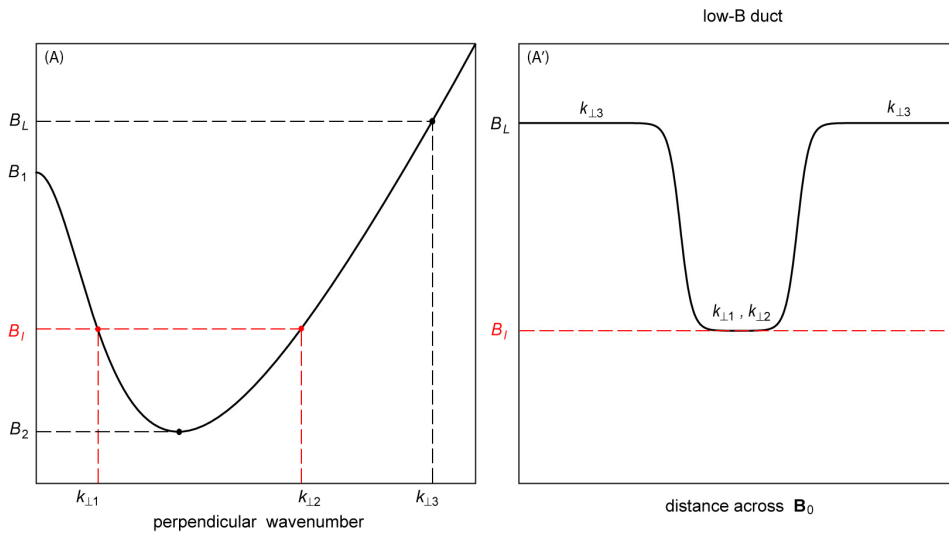
References

- 307
- 308 Gordeev, A., Kingsep, A., & Rudakov, L. (1994). Electron magnetohydrodynamics.
- 309 *Physics Reports*, *243*, 215-315.
- 310 Helliwell, R. (1965). *Whistlers and related ionospheric phenomena*. Stanford: Stan-
- 311 ford University Press.
- 312 Hobara, Y., Trakhtengerts, V. Y., Demekhov, A. G., & Hayakawa, M. (2000).
- 313 Formation of electron beams by interaction of a whistler wave packet with
- 314 radiation belt electrons. *J. Atmos. Sol.-Terr. Phys.*, *62*, 541.
- 315 Inan, U., Bell, T., Bortnik, J., & Albert, J. (2003). Controlled precipitation of radia-
- 316 tion belt electrons. *J. Geophys. Res.*, *108*, 1186. doi: 10.1029/2002JA009580
- 317 Inan, U., Chang, H., Helliwell, R., Imhof, W., Reagan, J., & Walt, M. (1985).
- 318 Precipitation of radiation belt electrons by man-made waves: A comparison
- 319 between theory and measurement. *J. Geophys. Res.*, *90*, 359.
- 320 Karpman, V. I., Istomin, Y. N., & Shklyar, D. R. (1974). Nonlinear theory of a
- 321 quasimonochromatic whistler mode packet in inhomogeneous plasma. *Plasma*
- 322 *Phys.*, *16*, 685.
- 323 Kondrat'ev, I. G., Kudrin, A. V., & Zaboronkova, T. M. (1999). *Electrodynamics of*
- 324 *density ducts in magnetized plasmas*. Amsterdam: Gordon and Breach.
- 325 Kostrov, A. V., Kudrin, A. V., Kurina, L. E., Luchinin, G. A., Shaykin, A. A., &
- 326 Zaboronkova, T. M. (2000). Whistlers in thermally generated ducts with
- 327 enhanced plasma density: excitation and propagation. *Phys. Scripta*, *62*, 51.
- 328 Nunn, D. (1974). A self-consistent theory of triggered vlf emissions. *Planet. Space*
- 329 *Sci.*, *22*, 349.
- 330 Nunn, D., & Smith, A. J. (1996). Numerical simulations of whistler-triggered vlf
- 331 emissions observed in antarctica. *J. Geophys. Res.*, *101*, 5261.
- 332 Omura, Y., Nunn, D., Matsumoto, H., & Rycroft, M. J. (1991). A review of obser-
- 333 vational, theoretical and numerical studies of vlf triggered emissions. *J. Atmos.*
- 334 *Terr. Phys.*, *53*, 351.
- 335 Russell, C. T., Anderson, B. J., Baumjohann, W., & et al. (2016). The magneto-
- 336 spheric multiscale magnetometers. *Space Sci. Rev.*, *199*, 189.
- 337 Sazhin, S. (1993). *Whistler-mode waves in a hot plasma*. Cambridge: Cambridge
- 338 University Press.
- 339 Stenzel, R. (1976). Whistler wave propagation in a large magnetoplasma. *Phys. Flu-*

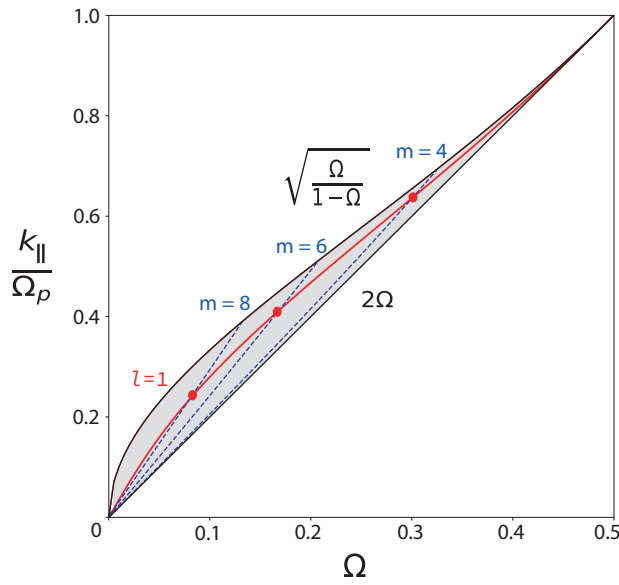
- 340 *ids*, 19, 857.
- 341 Stenzel, R. (1999). Whistler waves in space and laboratory plasma. *L. Geophys.*
342 *Res.*, 104, 14379.
- 343 Storey, L. (1954). An investigation of whistling atmospheres. *Phil. Trans. Roy. Soc.*
344 *London, A*, 246, 113.
- 345 Streltsov, A. (2007). Spectral properties of high-density ducts. *J. Geophys. Res.*,
346 112, A12218. doi: 10.1029/2007JA012710
- 347 Streltsov, A., Lampe, M., & Ganguli, G. (2007). Whistler propagation in non-
348 symmetrical density channels. *J. Geophys. Res.*, 112. doi: 10.1029/
349 2006JAa012093
- 350 Streltsov, A., Lampe, M., Manheimer, W., Ganguli, G., & Joyce, G. (2006).
351 Whistler propagation in inhomogeneous plasma. *J. Geophys. Res.*, 111. doi:
352 10.1029/2005JA011357
- 353 Streltsov, A., & Nejad, S. (2023). Whistler-mode waves in magnetic ducts. *J. Geo-*
354 *phys. Res.: Space Phys.*, 128, e2023JA031716. doi: 10.1029/2023JA031716
- 355 Torbert, R.B., Russell, C.T., Magnes, W., & et al. (2016). The fields instrument
356 suite on mms: Scientific objectives, measurements, and data products. *Space*
357 *Sci. Rev.*, 199, 105. doi: 10.1007/s11214-014-0109-8
- 358 Trakhtengerts, V., Rycroft, M., & Demekhov, A. (1996). Interaction of noise-like
359 and discrete ELF/VLF emissions generated by cyclotron interactions. *J. Geo-*
360 *phys. Res.*, 101, 13293.
- 361 Williams, D., & Streltsov, A. (2021). Determining parameters of whistler
362 waves trapped in high-density ducts. *J. Geophys. Res.: Space Phys.*, 126,
363 e2021JA029228. doi: 10.1029/2021JA029228
- 364 Yu, X., & Yuan, Z. (2022). Duct effect of magnetic structures on whistler waves. *J.*
365 *Geophys. Res.*, 127. doi: 10.1029/2022JA031013
- 366 Zaboronkova, T., Kostrov, A., Kudrin, A., Tikhonov, S., Tronin, A. V., & Shaikin,
367 A. I. (1992). Channeling of waves in the whistler frequency range within
368 nonuniform plasma structures. *Sov. Phys. JETP*, 75, 625.



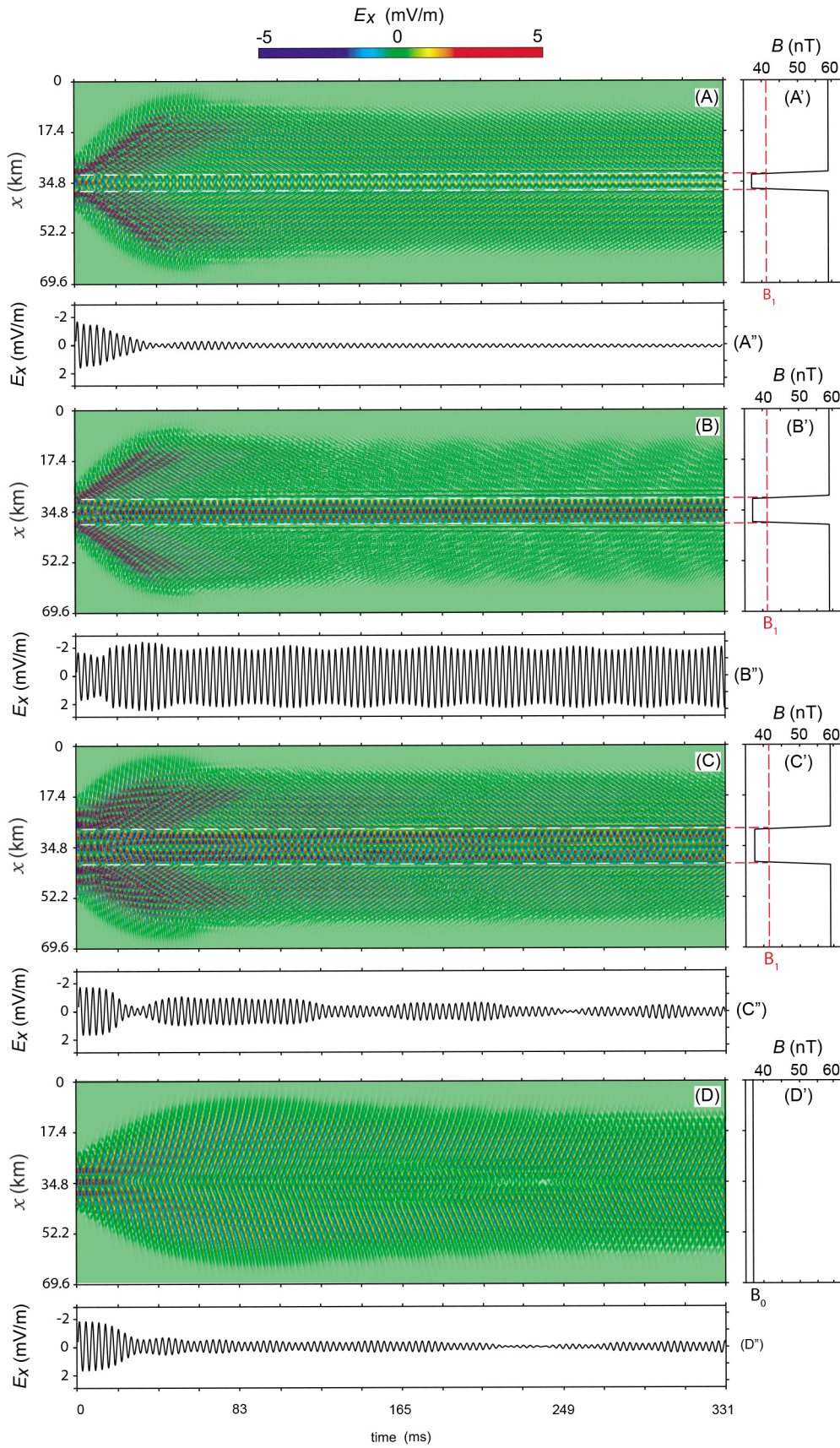
369 **Figure 1.** ELF wave packets localized in the low- B duct (LBD) observed by MMS1 satellite
 370 on March 06, 2016, from 19:09:20 to 19:09:40 UT. Panel (A) shows the power spectral density
 371 (PSD) of E_x component of the electric field in the GSE coordinate system of the satellite (shown
 372 with a color pallet) and background magnetic field (shown with the white line). Panel (B) shows
 373 the electron density measured by the MMS1 satellite. Panel (C) shows the trajectory and the
 374 location of MMS satellites in the GSE X-Y plane on March 6, 2016. The red dot marks the
 375 location of the satellites at 19:00 UT.



376 **Figure 2.** Panel (A): Schematic plot of B as a function of k_{\perp} , ω , k_{\parallel} , and ω_{pe} . Panel (A'):
 377 Magnitude of the magnetic field in the direction perpendicular to \mathbf{B}_0 corresponding to the low- B
 378 duct.



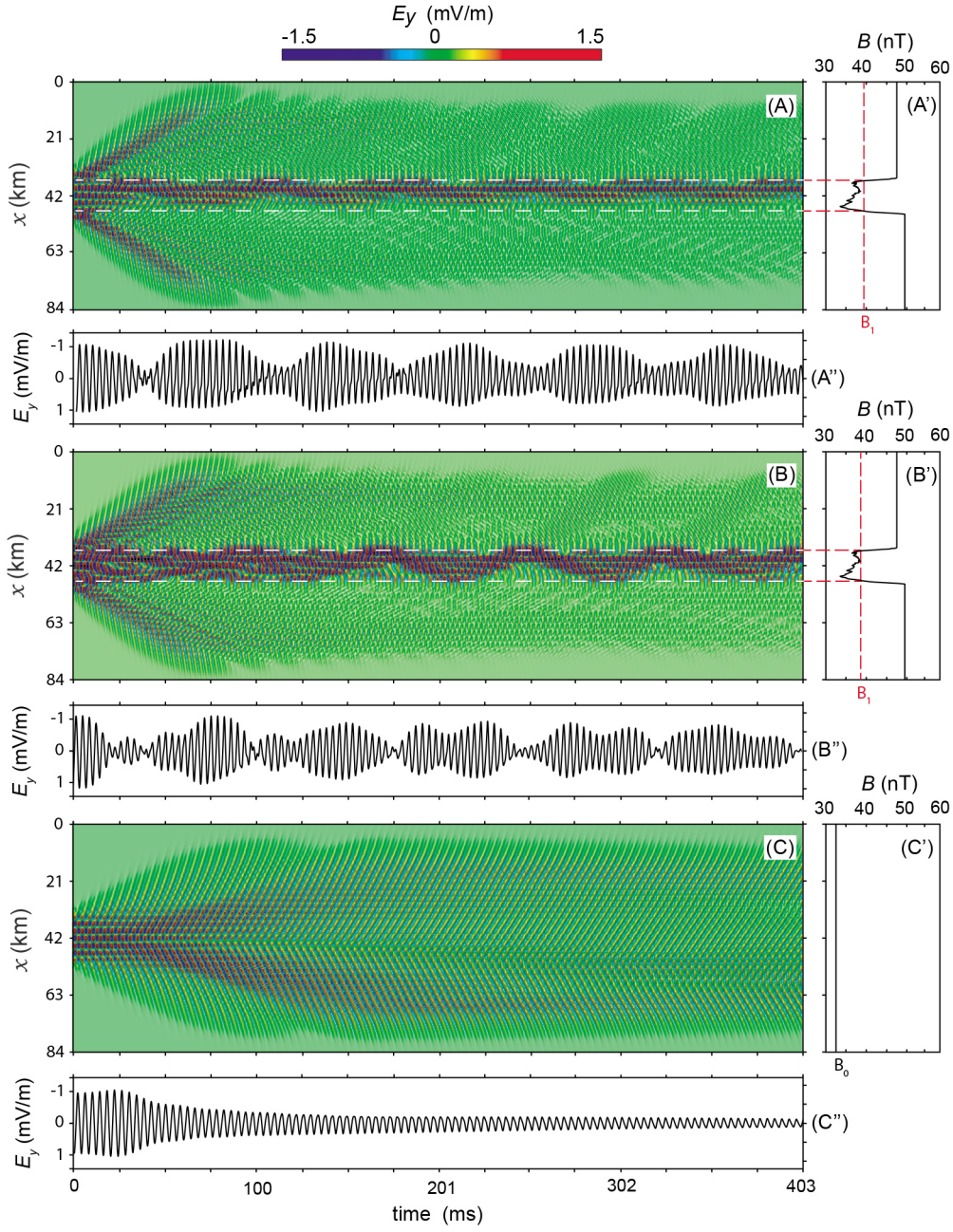
379 **Figure 3.** The region in the $(\Omega, K_{\parallel}, \Omega_p)$ space where real K_{\perp} s exist. The solution of $K_{\perp 1} =$
 380 $l\pi$ for $l = 1$ is plotted by the solid red line, and the solutions of $K_{\perp 2} = m\pi$ are plotted by dashed
 381 blue lines for $m = 4, 6$ and 8 , when $l = 1$ and $\Omega_p = 10$. Red dots show the intersection of $l = 1$
 382 and $m = (4, 6, 8)$ curves, indicating a set of discrete non-leaking eigenmodes of a low- B duct.



383

Figure 4. Continued on the following page.

384 **Figure 4.** Panel (A): Dynamics of $E_x(x, z = 0, t)$ in the simulation of the model low- B duct.
 385 The wave frequency is $f = 302$ Hz and $\lambda_{\parallel} = 7.298$ km. Panel (A'): Profile of the magnetic field
 386 across \mathbf{B}_0 . The duct width satisfies $2L = 0.6 \lambda_{\perp 1}$. Panel (A''): Dynamics of E_x at the center of
 387 the computational domain, $E_x(x = 0, z = 0, t)$. Panels (B), (B'), and (B'') show the results from
 388 the simulations of the same wave with the width satisfying $2L = \lambda_{\perp 1}$. Panels (C), (C'), and (C'')
 389 show results from the simulations of the same wave with the width satisfying $2L = 1.4 \lambda_{\perp 1}$. Pan-
 390 els (D), (D'), and (D'') show results from the simulations of the same wave with the homogeneous
 391 magnetic field, $B_0 = 37.35$ nT.



392 **Figure 5.** Panel (A): Dynamics of $E_y(x, z = 0, t)$ in the simulation of the observed low- B duct
 393 event. The wave frequency is $f = 248$ Hz and $\lambda_{\parallel} = 6.89$ km. Panel (A'): Profile of the magnetic
 394 field across \mathbf{B}_0 . The width of the duct satisfies $2L = 2 \lambda_{\perp 1}$. Panel (A''): Dynamics of E_y at the
 395 center of the computational domain, $E_y(x = 0, z = 0, t)$. Panels (B), (B'), and (B'') show results
 396 from the simulations of the wave with the same frequency and $\lambda_{\parallel} = 6.792$ km, satisfying $2L = 1.8$
 397 $\lambda_{\perp 1}$. Panels (C), (C'), and (C'') show results from the simulations of the same wave within the
 398 panel (A) with the homogeneous magnetic field, $B_0 = 33.37$ nT.