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A stochastic approach to detect fragmentation epoch from a single fragment orbit determination

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Abstract

In the last decades, the growing in-orbit population of resident space objects has become one of the main concerns for space agencies and institutions worldwide. In this context, fragmentations further contribute to increase the number of space debris and, operationally, it is fundamental to identify the event epoch as soon as possible, even when just a single fragment orbital state, resulting from an Initial Orbit Determination (IOD) process, is available.

This work illustrates the Fragmentation Epoch Detector (FRED) algorithm, which deals with the problem through a stochastic approach, starting from a single fragment IOD result (expressed through mean state and covariance) and parent ephemeris (assumed as deterministic). The process populates the fragment ephemeris with a multivariate normal distribution and, for each couple sample-parent, the epochs of parent transit through the Minimum Orbital Intersection Distance (MOID) are first computed on a time window and then clustered in time. For each cluster, both the three-dimensional MOID and the three-dimensional relative distance distributions are derived, and their similarity is statistically assessed. Given that, at the actual fragmentation epoch, MOID and relative distance were equal, the cluster featuring the best matching between the two distributions is considered as the optimal candidate, and the related fragmentation epoch is returned from the time of parent transit through the MOID, in terms of mean and standard deviation.

FRED algorithm performance is assessed through a numerical analysis. The algorithm robustness decreases when parent and fragment orbits share a similar geometry, and results get deteriorated if the perturbations and, moreover, the IOD errors are included in the process, but the correct fragmentation epoch is always present among candidates. Overall, FRED algorithm turns out to be a valid choice in operational scenarios, and a sensitivity analysis tests the algorithm out of the nominal conditions.

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Keywords: Fragmentations; Initial Orbit Determination ; Space Surveillance and Tracking ; Minimum Orbital Intersection Distance

1 1. Introduction

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In the last decades, the number of man-made objects orbiting the Earth has dramatically increased. In around 65 years of space activities, more than 6340 successful launches have taken place, which turned out in about 14710 objects placed in Earth orbit (ESA, 2023). Among these, 9780 are still orbiting, but only 7000 are active. In addition, about 640 break-ups, explosions, collisions, or anomalous events resulting in fragmentation have been recorded, which have further contributed to the increase in the orbiting population of man-made objects. In this context, space debris are considered as all the artificial objects including fragments and elements thereof, in Earth orbit or reentering the atmosphere, that are non functional (IADC, 2002).

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Nowadays, 32500 debris objects are regularly tracked by space 14 surveillance networks and maintained in their catalogue (ESA, 15 2023). In addition to them, statistical models estimate that there 16 are 36500 objects greater than 10 cm, one million objects be-17 tween 1 cm and 10 cm, and 130 million objects between 1 mm 18 and 1 cm. Their presence may jeopardise the operative mission 19 of active satellites, given that the possible impact with a space 20 debris ranges from cumulative erosion of satellite surface, for 21 debris smaller than 0.1 mm, to the possible satellite destruction, 22 with the generation of thousands of additional pieces of debris 23 and inevitable environmental drawbacks and possible cascade 24 effects (Kessler & Cour-Palais, 1978).

To mitigate mission-related risks, specific Space Surveillance 26 and Tracking (SST) programs were started to build the exper-27 tise required to manage the challenges posed by the Space Traf-28 fic Management (STM). To prevent the above-mentioned pro-29 liferation of space debris, particular attention is devoted to frag-30 mentation events, which may further contribute to increase the 31 number of space debris objects (McKnight et al., 2021). There-32 fore, it is fundamental to apply models predicting the frag-33 ments cloud evolution, like the ones in (Letizia et al., 2015) 34 35 and (Letizia et al., 2016), in order to assess possible collisions, and, for this purpose, the time when the break-up occurred shall 36 be identified to set the proper initial conditions. 37

In (Andrisan et al., 2016) the fragmentation epoch is evaluated 38 as the point of minimum distance of all the fragments with re-39 spect to the cloud centre of mass. In (Frey et al., 2018) the 40 break-up epoch is determined by detecting a convergence of 41 fragments in the space of inclination and right ascension of 42 the ascending node. In (Di Mare et al., 2019) a critical study 43 is conducted to identify the best criterion to assess the event 44 epoch from the fragments ephemerides, and a sensitivity anal-45 ysis on the cloud orbital position is conducted. In (Romano 46 et al., 2021) a process is proposed, which screens a catalogue of 47 ephemerides, detects possible break-ups of satellites and iden-48 tifies those related to fragments, through the filters presented in 49 (Hoots et al., 1984). After the filtering phase, the same cri-50 teria are applied combined with SGP4 propagation (Vallado 51 et al., 2006) and, by comparing the algorithm outputs among all the fragments, the fragmentation epoch is identified. All 53 these works need many fragments ephemerides, and use them 54 as a deterministic information. 55

The numerous accurate ephemerides availability of the space debris originated by the fragmentation event is a quite opti-57 mistic assumption, as, from an operational point of view, it 58 could be necessary to estimate the fragmentation epoch just few 59 hours after the event, and very few ephemerides (even only one) 60 could be available. Indeed, it may take days and even months 61 to have a large number of ephemerides. In addition, when a 62 fragments cloud is observed, the correlation of measurements 63 to a single fragment is a very challenging task, and this further decreases the number of ephemerides which can be used 65 in a reliable way. Next, such ephemerides could be inaccurate, 66 because of the noise of the observation measurements and the 67 error introduced by the Initial Orbit Determination (IOD) algo-68 rithm exploited, and their uncertainty cannot be neglected dur-69 ing the event characterisation. Nevertheless, a prompt knowl-70

edge of the fragmentation epoch would be fundamental to plan 71 additional observations of the fragments cloud, e.g. by task-72 ing the sensors to point at the right ascension and declination 73 where the parent was when broke up. Indeed, all the fragments 74 are expected to transit close to that inertial region in the first 75 hours after the event, before that their orbit modification due 76 to orbital perturbations becomes too relevant. Also, knowing 77 the fragmentation epoch would allow to model the break-up 78 event, which may be used to task sensors for early detection. 79 In addition, the knowledge of the fragmentation epoch would be important to refine the processing of the observation mea-81 surements, aiming at obtaining more and more accurate orbit determination results. This would lead to also refine the es-83 timation of the fragmentation epoch and, so, a virtuous cycle 84 would be generated. 85

The aim of the present work is to provide an operational procedure to estimate the fragmentation epoch starting from the last available ephemeris of the parent object (assumed as a deterministic quantity) and a single fragment orbital state provided with uncertainty. The latter is considered as determined, in the hours right after the fragmentation alert, by a IOD process from a single observation with no transit prediction. Such an approach would support operators to characterise fragmentations when a satellite break-up is detected and a measurements track (sufficient to provide an orbit determination result) is acquired few hours later, and it is associated to the event..

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To accomplish the purpose of the work, the FRagmentation Epoch Detector (FRED) algorithm, implementing a stochastic approach, is described in Sec. 2 and its performance are assessed in Sec. 3 through numerical simulations.

2. FRagmentation Epoch Detector - FRED

Let's consider the fragmentation of a space object whose last 103 available ephemeris x^p is dated to t_{eph} , and is considered as a 104 deterministic information. The event has occurred at $t_0 > t_{eph}$ 105 and the related alert has been notified at $t_a > t_0$. Some hours 106 later, one fragment is detected by an on-ground sensor at t_{obs} 107 (with $t_{obs} > t_a$) and its orbital state $\{\mathbf{x}^{fg}, \mathbf{P}^{fg}\}$ is first determined, 108 where the mean x^{fg} and covariance P^{fg} are directly derived 109 from the IOD process. 110

If the orbit determination were very accurate and both the phys-111 ical parameters and the dynamical model were well known, it 112 would be theoretically possible to propagate both the fragment 113 and the parent object in the time window $|t_{eph}, t_a|$ and search for 114 the epoch of the minimum relative distance, which would cor-115 respond to the fragmentation epoch t_0 . However, in real appli-116 cations, both the measurements accuracy and the IOD process 117 introduce an error in the reconstruction of the observed frag-118 ment state vector, and the above-mentioned method turns out 119 to be unreliable. As an example, Fig. 1 represents the relative 120 distance trend on an analysis time window between the par-121 ent object last available ephemeris and an observed fragment 122 mean state to which an IOD error of 1.85e-02 km in position 123 and 4.99e-04 km/s in velocity is attributed (continuous line). 124



Fig. 1: Relative distance between the parent object and the mean state of one observed fragment. Their state vectors are propagated on a time window ranging from the last available ephemeris of the parent object to the event alert. The dashed curve line shows the theoretical trend and the dashed straight line corresponds to the epoch of minimum value, that is the fragmentation epoch. On the contrary, the continuous black line shows the relative distance trend when an IOD error is attributed to the fragment mean state, and the dashed dense line corresponds to the minimum value, that is the estimated fragmentation epoch. It is possible to see that the estimated fragmentation epoch is completely different from the correct value.

Such an error is retrieved from a synthetic IOD process based 125 on the method presented in (Siminski, 2016) and starting from 126 angular track and slant range to which Gaussian noises of 0.01 127 deg and 30 m are added, respectively. It can be observed that 128 the epoch of the minimum relative distance between fragment 129 and parent mean states (dashed dense line) is completely dif-130 ferent from the correct fragmentation epoch (dashed line), that 131 is the epoch corresponding to the theoretical minimum rela-132 tive distance (dashed black line). A further source of error is 133 represented by the mismatching between the actual fragment 134 trajectory and the propagation model used, due, for instance, 135 to the fact that the actual physical parameters of the observed 136 fragment are not known. For all these reasons, assessing the 137 fragmentation epoch by just searching for the minimum rela-138 tive distance between \mathbf{x}^{p} and \mathbf{x}^{fg} in the time window $|t_{eph}, t_{a}|$ is 139 an unreliable methodology. 140

The considerations above imply that the orbit determination un-141 certainty cannot be a-priori neglected. For this reason FRED al-142 gorithm deals with the fragmentation epoch identification prob-143 lem through a stochastic approach, starting from a Monte Carlo 144 distribution of the orbit determination result. Ideally, at the 145 fragmentation epoch, both the Minimum Orbital Intersection 146 Distance (MOID) (Gronchi, 2005) and the relative distance be-147 tween the parent and the fragment are expected to be zero. Due 148 to the considerations above, in practical cases neither MOID 149 nor relative distance turn out to be null, but they should statis-150 tically match each other. Therefore, the correct fragmentation 151 epoch is expected to feature a matching between the MOID and 152 the relative distance distributions. 153

FRED algorithm flowchart is reported in Fig. 2, and is structured as follows.

156 1. In order to include the fragment state uncertainty in the 157 event epoch identification, N_s samples x^s are generated from the orbital state $\{x^{fg}, P^{fg}\}$ according to a multinormal distribution (Kotz et al., 2000). The parameter N_s can be selected by the user to guarantee a trade-off between a proper uncertainty sampling and the computational demand of the algorithm (which is directly proportional to the number of samples used).

- 2. The time window $[t_{eph}, t_a]$ is sampled with frequency $1/T^p$ (where T^p is the parent orbital period). This results in the epochs t_i , whose number is n_{orb} .
- 3. Both parent and fragment samples orbital states are propagated to each *t_i*.
- 4. For each t_i and for each *j*-th fragment sample, the epochs of transit through the MOID of both the parent and the fragment *j*-th sample are computed analytically, according to (Gronchi, 2005), and indicated as t_j^p and t_j^s . The parent and the *j*-th sample state vectors are propagated up to t_j^p and t_j^s respectively, resulting in the orbital states $\mathbf{x}^p(t_j^p)$ and $\mathbf{x}^s(t_j^s)$, and the analytical computations of t_j^p and t_j^s are updated. The epochs t_j^p and t_j^s are iteratively modified in this manner until, between two consecutive steps, they do not change anymore (according to a tolerance set equal to 1e-03 s).

This iterative process results in $N_s \times n_{orb}$ couples of (t_j^p, t_j^s) and $(\mathbf{x}^p(t_j^p), \mathbf{x}^s(t_j^s))$. It is important to observe that the difference between $\mathbf{p}^s(t_j^s)$ and $\mathbf{p}^p(t_j^p)$ (the $\mathbf{x}^s(t_j^s)$ and $\mathbf{x}^p(t_j^p)$ positions) allows to compute the MOID (usually described in a scalar way (Gronchi, 2005)) in 3 dimensions: $\mathbf{m}_j = \mathbf{p}^s(t_j^s) - \mathbf{p}^p(t_j^p)$.

- 5. The fragment *j*-th sample state vector $\mathbf{x}^{s}(t_{j}^{s})$ is propagated up to the epoch of parent transit through the MOID, resulting in $\mathbf{x}^{s}(t_{j}^{p})$. It is worth to observe that the difference between the $\mathbf{p}^{s}(t_{j}^{p})$ (the $\mathbf{x}^{s}(t_{j}^{p})$ position) and $\mathbf{p}^{p}(t_{j}^{p})$ provides the three-dimensional relative distance between the *j*-th sample and the parent, at the epoch of parent transit through the MOID: $\rho_{j} = \mathbf{p}^{s}(t_{j}^{p}) - \mathbf{p}^{p}(t_{j}^{p})$. Figure 3 provides a two-dimensional sketch of the parent and fragment sample orbits, with the involved quantities.
- 6. To exclude unfeasible solutions, the $N_s \times n_{orb}$ couples enter a filtering phase, which is based on the epoch of parent transit through the MOID t_j^p . Being related to the parent ephemeris, that is the information considered more reliable (and so assumed as deterministic), it is selected instead of the time of the fragment *j*-th sample transit through the MOID t_j^s . The filtering phase is structured as follows:
 - 6..1 First, the couples for which t_j^p is not included in the boundaries $[t_{eph}, t_a]$ are filtered out.
 - 6.2 Then, the couples computed from the state vectors 204 propagated at epoch t_i and for which $t_i^p < (t_i - T^p/2)$ 205 or $t_i^p > (t_i + T^p/2)$ are removed from the data 206 set. This operation is done because the MOID data 207 $(t_i^p, t_j^s, \boldsymbol{p}^p(t_j^p), \boldsymbol{p}^s(t_j^s), \boldsymbol{p}^s(t_j^p))$ are computed for each 208 periodicity. Thus, if t_i^p is computed from orbital 209 states at t_i , it must belong to the *i*-th periodicity, that 210 is the time difference $|t_i - t_i^p|$ shall be smaller than half 211 of the parent orbital period T^p . 212

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Fragmentation epoch $\{\mu_t, \sigma_t\}$ from *F* mean and standard deviation

Fig. 2: FRED algorithm flowchart.



Fig. 3: Sketch of the parent and fragment sample orbits, with the quantities involved in FRED algorithm process.

7. All the remaining n_{filter} epochs t_i^p are clustered accord-213 ing to a Density-Based Spatial Clustering of Applications 214 with Noise (DBSCAN) (Ester et al., 1996). From this op-215 eration, n_{orb} are expected to be identified. However, for 216 those situations in which parent and fragment orbits are 217 similar (especially in inclination and right ascension of the 218 ascending node), multiple clusters are possibly identified 219 for each *i*-th periodicity, as the epochs t_i^p change signifi-220 cantly from a *j*-th sample to another one. So, more gen-221 erally, n_{cl} clusters are considered to be identified. Figure 222 4a presents the obtained clusters, in the plane t_i^p (in Co-223 ordinated Universal Time, UTC) versus scalar MOID. It is 224 worth to remark that the MOID values are equal from a pe-225 riodicity to the other, as the graph is related to a Keplerian 226 scenario, in which, for a single parent *i*-th sample couple, 227 the MOID does not change. 228

8. For each *n*-th cluster, the candidate fragmentation epoch 229 t_n^{fg} can be computed (in terms of mean and standard devi-230 ation) from the distribution of the epoch of parent transit 231 through the MOID, which is indicated as F, and which is 232 represented in Fig. 4b (for the correct cluster). In addition, 233 M and R distributions (grouping the m_i and ρ_i respec-234 tively) are associated to each cluster. Figure 5 shows the 235 two distributions in Earth-Central-Inertial (ECI) reference 236 frame, both for the correct candidate and for a non-correct 237 one. It is possible to observe that the three-dimensional 238 MOID distribution M is much more concentrated than 239 the relative distance one R. This is due to the fact that, 240 from sample to sample, the change in t_i^p causes a remark-241 able modification in the relative distance ρ_i (as it is time-242 dependent), but not in the MOID m_i , which is the geo-243 244 metrical difference between the parent and the *j*-th sample orbits and, so, does not vary remarkably from a sample to 245 another. 246

²⁴⁷ 9. Afterwards, for each cluster:

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9..1 All the m_i and ρ_j are rotated in the Modified Equidis-

tant Cylindrical (EQCM) reference frame (Vallado & Alfano, 2014). This operation results in MOID and relative distance distributions like in Fig. 6. The MOID distribution M is almost two-dimensional, as, in all the m_j , the y-component, expressing the along orbit curvature relative distance, is negligible.

- 9..2 The statistical distance between *M* and *R* distributions is computed according to one of the metrics discussed below.
- 10. Repeating the operations above for each cluster results in Fig. 7, which shows the statistical distance computed through the Earth Mover's Distance (EMD) (Levina & Bickel, 2001) (discussed below) in function of the *F* distribution mean. Finally, the cluster featuring the minimum statistical distance between the *M* and *R* distributions is selected, and the fragmentation epoch is returned from the related distribution *F*, in terms of mean μ_t and standard deviation σ_t .

As mentioned above, this process provides a pattern to derive the fragmentation epoch (in terms of mean and standard deviation) through a stochastic approach, starting from the last available parent ephemeris and the fragment IOD result. However, there are two theoretical sources of failure:

- The MOID computation turns out to be very sensitive 273 when the orbital planes of the fragment and parent orbits 274 are very close each other (that is, they have similar in-275 clination and right ascension of the ascending node). In 276 this case, the change in the fragment orbit, occurring from 277 sample to sample, may provoke a remarkable variation in 278 the MOID data computation. As result, F distribution 279 expand, and, for the correct candidate, it may not clus-280 ter around the actual fragmentation epoch, but around an 281 epoch distant up to tens of minutes. 282
- The relative distance distribution R does not change from 283 a cluster to another when the fragment and parent orbital 284 periods are very close each other (that is, they have similar 285 semi-major axis). In this case, for a *j*-th sample, from a *i*-286 th periodicity to the following one, the relative distance ρ_i 287 does not change significantly. As result, it is not straight-288 forward to recognise the correct cluster from the statisti-289 cal distance metrics, and the wrong fragmentation epoch 290 is possibly returned by the process. 291

As introduced above, FRED needs a statistical distance metrics to assess the best epoch candidate. Expressing M and R distributions through their mean and covariance as $\{\mu_M, P_M\}$ and $\{\mu_R, P_R\}$ respectively, a possible choice is represented by the Mahalanobis Distance (Mahalanobis, 1936):

$$\xi = \sqrt{\{\mu_M - \mu_R\}^T \{P_M + P_R\}^{-1} \{\mu_M - \mu_R\}}$$
(1)

However such a metrics applies to Gaussian distributions only. Even if supported by the rotation to EQCM reference frame, assuming Gaussian distributions would be a particularly strong

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(a) Distribution of the t_i^p epochs in the time window of the analysis.

(b) Distribution of the t_i^p epochs for the cluster related to the correct solution.

Fig. 4: Results of the clustering phase. The epochs are reported in UTC.



Fig. 5: *M* and *R* distributions in ECI reference frame, for the correct cluster and a non-correct one.



Fig. 6: *M* and *R* distributions in EQCM reference frame, for the correct cluster and a non-correct one.



Fig. 7: EMD statistical distance computed for each cluster.

assumption for M and R distributions. To be as generic and 295 agnostic as possible regarding the distributions characteristics, 296 metrics suitable both for Gaussian and no Gaussian distribu-297 tions are investigated. 298

A first choice is represented by the Earth Mover's Distance 299 (EMD) (Levina & Bickel, 2001), which measures the flow to 300 pass from a distribution to the other one. Such a flow can be 301 evaluated based on different distance metrics, and the Euclidean 302 distance weighted on the distribution variance is chosen to bet-303 ter account for M and R distributions shape and elongation. 304 The implementation provided in (SciPy, 2022) is used. 305

A third metrics is investigated, which has been developed 306 specifically for FRED algorithm. It is addressed as the quan-307 *tile* metrics given its workflow, which is described as follows. 308

1. For both M and R distributions a Principal Component 309 Analysis (PCA) is performed to rotate them in their re-310 spective principal coordinate reference frame (Jolliffe, 311 2011). Then, for each distribution, the quantiles 10%, 312 25%, 50%, 75% and 90% are computed for the three coor-313 dinates separately. This operation does not account for the 314 correlations among coordinates, but these have been min-315 imised thanks to the rotation to the principal coordinate 316 reference frame. This results in two sequences (for M and 317 \mathbf{R}) of three-dimensional points, expressed in two different 318 principal coordinate reference frames. 319

2. The two sequences of three-dimensional points (express-320 ing the quantiles) are rotated back to the original EQCM 321 reference frame, in order to have them in a common coor-322 dinate system. Figure 8 shows the two sequences of three-323 dimensional points, for the correct and for a wrong epoch. 324 Then, the five quantile-to-quantile Euclidean distances are 325 computed and summed together in a weighted manner ac-326 cording to the quantile percentage (that is, by advantaging 327 more the central quantiles with respect to the side ones). 328 This weighted sum provides the statistical distance which 329 accounts for the similarity between the two non-Gaussian 330 distributions *M* and *R*. 331

A critical comparison among the metrics presented above is 332 proposed during the numerical analysis in Sec. 3.2. 333

Analogies and differences with conjunction analysis

From the FRED description, the reader may easily notice that 335 dealing with the fragmentation detection problem in such a stochastic way presents analogies with the conjunction analy-337 sis. In particular, the process involves the MOID and the rel-338 ative distance, which are quantities usually exploited also in 339 the screening part of the conjunction assessment (Hoots et al., 1984), as well as in other fragmentation epoch identification al-341 gorithms (like in (Di Mare et al., 2019) and (Romano et al., 342 2021)) which use the availability of many fragments orbital 343 states, then processed in a deterministic way. However, at this 344 level a first difference arises. Indeed, in FRED, the screening is 345 fully stochastic and is only based on the time of parent transit 346 through the MOID. In addition, the FRED screening phase does not aim at identifying possible conjunctions, as the fragmenta-348 tion is already known to have occurred, but to rank conjunction 349 (that is fragmentation epoch) candidates. Thus, the MOID and 350 the relative distance are not quantities used to search for a possible conjunction in a deterministic way, but they are stochastically represented at the fragmentation epoch candidates, and 353 then their statistical distance is computed.

At this point, a second analogy may be noticed, as in both cases a stochastic quantity is expressed at the time of closest approach: the Probability of Collision (PoC) in the conjunction analysis and the statistical distance between MOID and relative distance distributions in FRED. However, besides the two metrics differently defined, a remarkable difference arises: while in conjunction analysis the PoC is a quantity assessing the danger associated to a single conjunction and, so, expressing an absolute meaning, in FRED the statistical distance is used to rank the fragmentation epoch candidates previously identified, and so it has a relative meaning.

3. Numerical simulations

3.1. Data set generation

A numerical simulation is here conducted to test FRED al-368 gorithm. The fragmentation scenario is the one which involved 369 the Russian satellite COSMOS 1408 during the kinetic anti-370 satellite (ASAT) test which occurred around 02:47 UTC of 371 November 15th, 2021 (EUSST, 2021). The ASAT test took 372 place when the satellite was flying over the north-west Rus-373 sia and the sensors of the EUSST consortium (European Space 374 Surveillance and Tracking, 2021) observed the fragments gen-375 erated by such an event. 376

The data set to test FRED algorithm is generated as follows:

1. The last available COSMOS 1408 ephemeris before the 378 event are retrieved from the last TLE (Two-Line Elements) 379 available on Spacetrack, which are dated to 00:55 UTC of 380 November 15th (Space-track, 2022) (Hoots & Roehrich, 381 1980). To make the analysis time window more symmetri-382 cal with respect to the break-up epoch, they are propagated 383 one orbital period back to the 23:20 UTC of November 384 14th, and the orbital state at this epoch is considered as 385

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(a) Cluster related to the correct epoch.

(b) Cluster related to a wrong epoch

Fig. 8: Quantile sequences for *M* and *R* distributions in EQCM reference frame.

 x^{p} . This operation has been taken to assess algorithm behaviour when the fragmentation epoch candidates distribution is as symmetric as possible with respect to the correct epoch.

2. The state vector x^p is propagated up to 02:47:00 UTC of 390 November 15th. Table 1 reports COSMOS 1408 orbital 391 parameters, simulated at the fragmentation epoch. 392

3. The fragmentation event is modelled as a set of impulses 393 applied to the satellite orbital state at 02:47:00 UTC. These 394 impulses, generating one single fragment each, are retrieved from the NASA standard break-up model (NASA, 396 2011). By this way, a data set of 231 fragments is created 397 by setting the parent object mass equal to 200 kg and the 398 fragments characteristic length ranging from 0.01 m to 2.1 399 m. These values were selected to obtain a manageable and 400 complete impulse data set size rather than to model the 401 event in a realistic way. The simulated fragments cloud 402 characteristics are described in Fig. 9, both in terms of 403 impulse magnitude distribution of the fragmentation event 404 and Gabbard diagram. 405

<i>a</i> [km]	е	<i>i</i> [deg]	Ω [deg]	ω [deg]	θ [deg]
6862.2	2.9e-03	82.7	123.4	91.9	341.8

Table 1: COSMOS 1408 orbital parameters simulated on November 15th 2022, at 02:47:00 UTC.

The obtained ephemerides, representing the fragments, are 406 propagated until the epoch t_{obs} , when they are detected by an 407 on-ground sensor, and the orbital states $\{x^{fg}, P^{fg}\}$ are deter-408 mined. The propagation model used depends on the analysis 409 conducted, as detailed throughout the rest of Sec. 3. 410

In this way all the inputs for the process described in Sec. 2 are 411 obtained, and FRED algorithm can be tested, considering an 412 analysis time window ranging from 23:20 UTC of November 413 14th (epoch of the simulated last available ephemeris of the 414 parent object) to 06:00 UTC of November 15th, retracing the 415

fact that the COSMOS 1408 fragmentation alert was provided 416 in the early morning (considering UTC time coordinates). 417 These two epochs correspond to t_{eph} and t_a introduced in Sec. 418 2. Instead, the t_{obs} changes from an analysis to the other, as discussed below.

Based on this data set, FRED is run on each fragment IOD result $\{x^{fg}, P^{fg}\}$ separately, considering $N_s=1e+03$ samples for the multi-normal distribution.

3.2. Unperturbed scenario with no IOD error

First, the unperturbed scenario, considering a two-body dy-426 namics and with no IOD orbital state error is tested to assess 427 the theoretical characteristics of FRED algorithm in ideal con-428 ditions. For this purpose, an analytic propagator with no orbital 429 perturbations is exploited. This simulation just associates a co-430 variance P^{fg} (with standard deviations 2.6e-02 km and 7.0e-04 431 km/s, for inertial position and velocity respectively, computed 432 simulating an IOD with the method presented in (Siminski, 433 2016)) to the nominal value x^{fg} , that is the fragments propa-434 gated state vectors. Thus, the fragment mean state μ^{fg} is the 435 actual fragment position and velocity at t_{obs} . The parent last 436 available ephemeris x^p is the same used above to generate the 437 fragmentation, and the observation epoch t_{obs} is set 13 h after 438 the event, as the method aims at reconstructing the fragmenta-439 tion epoch from a single fragment observation conducted in the 440 hours right after the event. 441

For a single fragment analysis, the result is considered success-442 ful if the difference between the epoch estimation and the cor-443 rect value (t_{err}) is below a threshold quantity, which is set equal 444 to 1 min in the analysis, coherently with the time uncertainty 445 associated to the estimated fragmentation epoch in (Muciaccia 446 et al., 2022). As introduced in Sec. 2, possible FRED failures 447 can be linked to either the MOID computation or to the distri-448 butions comparison performed through the statistical metrics, 449 and for this reason they are classified as follows: 450

• MOID failures - compliant: 1 min < t_{err} and $t_{err} < 3\sigma_t$. 451 These are cases for which the fragment orbit orientation 452

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Table 2: Unperturbed scenario results for the different statistical distance metrics.

is so similar to the parent one that a slight change in the 453 fragment orbit, occurring from fragment mean state to its 454 samples, causes a remarkable variation in the MOID data 455 computation. This leads to an erroneous estimation of 456 the mean epoch of parent transit through the MOID, but 457 the distribution is wide enough to include such an error. 458 Therefore, the resulting epoch estimation is wrong, but sta-459 tistically compliant. 460

• MOID failures - uncompliant: 1 min < t_{err} and $3\sigma_t$ < $t_{err} < T^p/2$. In these cases, the erroneous estimation of the epoch is not mitigated by its uncertainty. The epoch estimation is wrong, but the error is smaller than the half of the parent orbital period.

• Periodicity failures: $t_{err} > T^p/2$. In these cases, the statistical comparison among clusters identifies a wrong candidate and, so, a wrong result is returned. It is worth to remark that MOID failures may occur also when a wrong candidate is identified. Nevertheless, also this situation is addressed as a periodicity failure, as the time error is anyways larger than half of the parent orbital period.

The results are reported on Tab. 2, for each statistical distance
metrics introduced in Sec. 2. It can be observed that all the
metrics feature comparable results, but the EMD ones are the
most appreciable.

An analysis is also conducted to assess the Gaussianity of the
problem, in order to evaluate whether the Mahalanobis Distance
metrics, which needs the Gaussian assumption of the involved

distributions, is a suitable choice. For each fragment, the Maha-480 lanobis Distances between each ρ_i and each m_i and the distri-481 butions **R** and **M** respectively is computed, and a χ^2 test is con-482 ducted to check how many Mahalanobis Distances are smaller 483 than the 3σ level, for all the n_{cl} clusters. To fulfil the Gaussian 484 assumption, this condition shall be matched in the 99% of cases. 485 Figure 10 shows the number of fragments (in logarithmic scale) 486 in function of the mean percentage of samples (across the clus-487 ters) satisfying the 3σ level, both for the MOID distribution M 488 and for the relative distance distribution R, by also focusing on 489 the portion of the diagram closest to the expected value of 99%. 490 It can be observed that no fragment satisfies the 99% require-491 ment in the MOID distribution M, with lot of cases showing a 492 low percentage of samples within the 3σ level. For some frag-493 ments the relative distance distribution R features Gaussianity, 494 but the 99% requirement is not fulfilled in most cases. 495

This analysis proves that a non-Gaussian metrics shall be considered and, so, the Mahalanobis Distance is rejected. Furthermore, given the results in Tab. 2, the Earth Mover Distance metrics is selected, as it features the best performance. Therefore, next analyses always apply EMD to identify the best epoch candidates.

EMD results and failures assessment

Figure 11 shows, for each fragment analysed, the relationship between the magnitude of the impulse which generated it (in logarithmic scale) and the time error between the estimated and the correct fragmentation epochs. It is possible to notice that, over the 231 fragments analysed, 12 MOID failures occur, out of which 11 are compliant and 1 is not. Then, 5 periodicity

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Fig. 10: Number of fragments (in logarithmic scale) in function of the mean percentage of samples (across the clusters) satisfying the 3σ level, both for the MOID distribution M and for the relative distance distribution R, by also focusing on the portion of the diagram closest to the expected value of 99%.



Fig. 11: Results of the numerical analysis on the unperturbed scenario with no orbital state error, by using the EMD metrics. The graph represents, for each fragment analysed, the relationship between the magnitude of the impulse which generated it (in logarithmic scale) and the time error between the estimated and the correct fragmentation epochs. The fragments for which a failure occurs are highlighted according to the legend.

failures are present, and they are cases for which the EMD met-509 rics returns similar values across the candidates, among which 510 the correct solution is always present, and the process returns 511 a wrong epoch. It may be noticed that, as general trend, the 512 larger the impulse, the more robust FRED algorithm is. Indeed, 513 a fragment originated by a large impulse magnitude is expected 514 to feature an orbit remarkably different from the parent one both 515 in terms of orbital plane (inclination and right ascension of the 516 517 ascending node) and of shape (semi-major axis and eccentricity). Thus, it does not run into the theoretical failure sources 518 mentioned in Sec. 1. 519

To further assess the problem, it is useful to relate the time 520 standard deviation of the computed fragmentation epoch to the 521 difference between parent and fragments orbital parameters, as 522 523 represented in Fig. 12. The closer the fragment orbit to the parent one, the larger the time standard deviation associated to 524 the FRED solution, especially for what concerns the inclination 525 and the right ascension of the ascending node (Fig. 12c and 526 Fig. 12d respectively). This behaviour is linked to the fact that 527 the closer the fragment orbit orientation to the parent one, the 528 larger the excursion of the MOID data from a sample to another 529 (as commented in Sec. 2) and, so, the larger the uncertainty of 530 the time of parent transit through the MOID, that is of the frag-531 mentation epoch candidates. On the contrary, the smallest time 532 uncertainty is related to those fragments with an orbit signifi-533 cantly different from the parent one, as the MOID data do not 534 vary much from a sample to another. Focusing on the the fail-535 ures characteristics, from Fig. 12a and Fig. 12b it is possible to 536 observe that the periodicity failures regard cases in which the 537 fragment orbit semi-major axis and eccentricity are very close 538 to the parent values. Indeed, in this situation, the two orbits 539 have a similar period and shape, and, from a *i*-th periodicity to 540 the following one, there is not a remarkable difference in the 541 relative distribution R (the MOID distribution M is al-542 ways the same, being the scenario Keplerian). This weakens 543

the statistical comparison result, as the EMD is similar across 544 multiple clusters, and the algorithm possibly converges to an 545 erroneous solution. Instead, from Fig. 12c and Fig. 12d it 546 is worth to notice that both compliant and uncompliant MOID 547 failures regard cases in which fragment and parent inclination 548 and right ascension of the ascending node are very close each 549 other, as the similar orientation provokes a remarkable excur-550 sion of MOID data from a sample to another, and the samples 551 cluster around a quantity corresponding to an epoch which is 552 not the correct value. Overall, this practically confirms the two 553 theoretical sources of failure mentioned in Sec. 1. 554

A detailed computational demand study is not carried out, given the current prototype implementation in MATLAB (MATLAB, 2020), but it can be quantified in about 30 s per fragment by using a single core with the same Intel(R) Core(TM) i7-8700 CPU @ 3.20 GHz - 3.19 GHz processor. This low computational demand is linked to the analytical propagation exploited in the unperturbed scenario.

Sensitivity analysis on the number of samples used

As described in Sec. 2, FRED algorithm starts from the IOD result (expressed in terms of mean state and covariance), and populate it by samples according to a multi-normal distribution. Thus, the larger the number of samples used, the more accurate the IOD uncertainty representation. The number of samples used is a key point in assessing FRED performance and, for this reason, a sensitivity analysis is here conducted by modifying the nominal value of $N_s = 1000$ to 100, 500, 2000 and 10000. It must be pointed out that the larger the number of samples used, the larger the computational demand, as more conjunctions for each fragment are to be computed (both in terms of MOID and relative distance evaluation). In addition, also the computational demand of the EMD metrics is proportional to the number of samples.

The results are reported in Tab. 3. It is possible to notice that 577 the performance are stable across the different values of N_s , and remain similar to the EMD metrics results reported in Tab. 2. 579 In particular, it is to point out that the convergence rate to the correct solution does not improve for a larger number of samples used in a monotonic way. This confirms that the failure cases are not related to an uncertainty representation which is not dense enough, but to the mutual geometry between parent and fragment orbits, as discussed above regarding Fig. 12. On the one hand, this is an important result, as the method computational demand can be reduced by using a lower number of samples, without a performance degradation. On the other hand, the larger the number of samples, the better the representation of the IOD uncertainty. Therefore, a trade-off choice must be conducted. For these reasons, the nominal value of $N_s = 1000$ samples is kept in the following analyses.

3.3. Perturbed scenario with no IOD error

The same analysis as above is conducted on a perturbed sce-594 nario in which SGP4 (Vallado et al., 2006) is used both to derive 595 the fragments actual trajectory, and in FRED algorithm. The 596 data set is created as follows: 597

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Fig. 12: Unperturbed scenario: relationship between the standard deviation associated to the computed fragmentation epoch and the fragment semi-major axis, eccentricity, inclination and the right ascension of the ascending node. The fragments for which a failure occurs are highlighted according to the legend, and the dashed line shows the parent orbital parameters.

	Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
100	92.4 %	3.9 %	0.4 %	3.4 %
500	92.0 %	4.2 %	0.4 %	3.4 %
2000	92.4 %	4.2 %	0.4 %	3.0 %
10000	92.8 %	4.2 %	0.4 %	2.6 %

Table 3: Unperturbed scenario: sensitivity analysis on the number of samples used.

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Fig. 13: B* distribution and FRED clusters in the perturbed scenario.

- 1. The last available TLE of the parent object is propagated 598 up to the fragmentation epoch, which is always set at 599 02:47:00 UTC of November 15th, 2021, and converted in 600 Cartesian coordinates. 601
 - 2. The fragmentation impulses are applied, again according to the NASA standard break-up model (NASA, 2011).

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- 3. Each fragment state is converted in SGP4 elements 604 through a fixed-point iteration loop based on a Non-linear 605 Least Squares (Coleman & Li, 1996). In particular, the 606 B*, which accounts for the physical characteristics of the 607 object, is computed by: 608
 - 3..1 Propagating the fragment orbital state through the high-fidelity propagator described in (Cipollone et al., 2022). To this end, the ballistic coefficient was provided by the NASA break-up model.
- 3..2 Searching for the B* which allows the SGP4 prop-613 agation to best match the high-fidelity propagation, 614 through a Non-linear Least Squares filter. Out of the 615 237 fragments of the original data set, for 28 the pro-616 cess does not converge to a solution. Thus, a data set of 209 fragments is considered from now on. 618
- The computed B* distribution is reported in Fig. 13a. 619
- 4. Similarly to the analysis in Sec. 3.2, each fragment ele-620 ments are propagated through SGP4 for 13 h, when the ob-621 servation is simulated by computing the fragment orbital 622 state in Cartesian coordinates and associating the same co-623 variance used in Sec. 3.2. 624
- Then, in FRED algorithm, each fragment sample is propagated 625 through SGP4. This operation implies a first conversion from 626 Cartesian coordinates to SGP4 elements (at the IOD epoch), 627 and then from SGP4 elements to Cartesian coordinates at the 628 end of the propagation (that is at the epochs t_i defined in Sec. 629 2) to compute the MOID and the relative distance. 630

Both in data set generation and inside FRED algorithm, the 631 conversion from Cartesian coordinates to SGP4 elements in-632 troduces an error which, although negligible at the considered 633 epoch, increases with the propagation and may affect results at 634

the epochs t_i . On the contrary, the presence of perturbations in 635 the propagation introduces an additional difference among clusters, besides the one related to the phasing effect between parent 637 and fragment samples orbital states. This can be observed in Fig. 13b., which reports the clusters in the plane time of transit 639 of parent through the MOID versus MOID magnitude, for the 640 same case as the one reported in Fig. 4a for the Keplerian sce-641 nario. Comparing the two figures, it can be appreciated how the 642 perturbations introduce a difference among the clusters.

FRED results for the perturbed scenario are reported in Tab. 4 644 considering the Earth Mover Distance metrics, and represented 645 in Fig. 14. A deterioration in performance may be noticed, due 646 to the fact that the number of fragments in data set decreases, 647 as mentioned above, and both the uncompliant MOID and the 648 periodicity failures increase, passing from 1 and 5 to 2 and 8 649 respectively. Similarly to Fig. 11, Fig. 14 confirms that FRED 650 algorithm is more prone to fail for those fragments originated 651 by a small impulse magnitude.

Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
90.0 %	5.3 %	0.9 %	3.8 %

Table 4: Perturbed scenario results for EMD metrics.

As in Sec. 3.2, it is interesting to study the relationship between the time standard deviation associated to the solution and the orbital parameters, as represented in Fig. 15. All the considerations as in Sec. 3.2 are valid, to testify that the most failure prone situations (similar orbital period and orientation) do not change when perturbations are considered in the dynamics. The computational demand increases with respect to the unperturbed scenario (under the same conditions), resulting in about 5 min per fragment analysed. This is due both to SGP4, which 661 requires more computational time than the unperturbed analytical propagation, and to the fact that, for each *j*-th fragment 663 sample, the MOID data are recursively refined until the flying time to the MOID falls below 1e-03 s (as described in Sec. 2).

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Fig. 14: Results of the numerical analysis on the perturbed scenario with no orbital state error. The graph represents, for each fragment analysed, the relationship between the magnitude of the impulse which generated it (in logarithmic scale) and the time error between the estimated and the correct fragmentation epochs. The fragments for which a failure occurs are highlighted according to the legend.

To assess the general applicability of FRED algorithm, the same
simulation as in Sec. 3.3 is reported in Appendix considering a
Medium Earth Orbit (MEO) and a Geostationary Orbit (GEO)
fragmentation.

670 3.4. Perturbed scenario with IOD error

The analyses in Sec. 3.2 and Sec. 3.3 are conducted with no 671 error associated to IOD, that is starting from an orbital state ob-672 tained by simply propagating the fragment nominal ephemeris 673 up to a certain epoch, considering it as the mean state and as-674 sociating a covariance to it. However, in real applications, at 675 the orbit determination epoch a mismatching between the or-676 bital state mean and the ground truth is introduced by the IOD 677 process, and its effects on FRED algorithm must be assessed. 678 For this purpose, an analysis is carried out by starting from an 679 orbital state generated through a surveillance radar observation, 680 which allows to run a IOD from the measurements acquired 681 during a single observation, also if this lasts few tens of sec-682 onds (Bianchi et al., 2022). 683

- The ground truth of the fragment orbital state is generated in the same manner as in Sec. 3.3, that is propagating the fragment ephemeris for 13 h from the event through SGP4 (Vallado et al., 2006) and with the estimated B*.
- Geodetic latitude and longitude are computed from the
 fragment position, and a monostatic radar station is simulated at 0 km altitude and with a small variation of +1
 deg from the fragment coordinates. Such a variation prevents the target from exactly transiting through the station
 zenith direction.
- Azimuth, elevation and slant range are simulated for the following 30 s. A Gaussian noise is added of 0.01 deg (on angular coordinates) and 30 m (on slant range), coherently with the real data analyses presented in (Montaruli et al., 2022a).

- Azimuth, elevation and slant range are simulated for the following 30 s. A Gaussian noise is added of 0.01 deg (on angular coordinates) and 30 m (on slant range), coherently with the real data analyses presented in (Montaruli et al., 2022a).
- The orbital state is computed at the initial observation 704 epoch, through the IOD procedure presented in (Simin-705 ski, 2016), which computes the orbital state at the first 706 observation epoch through an unperturbed analytic prop-707 agation. By this way, a dynamical model error is included 708 in the IOD process, as the measurements were simulated 709 through a propagation based on SGP4. No further refine-710 ment is done in the simulations, to test the procedure for a 711 coarse IOD result. 712

In this way, the fragment orbital state $\{x^{fg}, P^{fg}\}$ is obtained, and FRED algorithm is run. It is worth to stress that an error between x^{fg} and the fragment actual position and velocity is now present, and the covariance P^{fg} is computed from the measurements through the IOD procedure, that is differently from what done in Sec. 3.2 and in Sec. 3.3.

Correct	MOID failures	MOID failures	Periodicity
solutions	compliant	uncompliant	failures
68.9 %	9.6 %	0.5 %	21.0 %

Table 5: Results for the perturbed scenario and accounting for the orbital state error introduced by the IOD process. The EMD metrics is used.

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Results are reported in Tab. 5 and represented in Fig. 16. It is 719 worth to observe that in most cases the algorithm converges to 720 the correct solution. However, comparing Tab. 5 to Tab. 2 and 721 Tab. 4, it can be noticed that the IOD mismatching remarkably 722 affects the algorithm performance, especially for what concerns 723 the metrics to select the correct candidate. This can be visu-724 alised also by comparing Fig. 16 with Fig. 11 and Fig. 14. 725 Concerning the relationship between the time standard devia-726 tion associated to the solution and the orbital parameters, rep-727 resented in Fig. 17, it may be noticed that the more similar 728 the fragment and the parent orbits are, the larger the time un-729 certainty associated to the FRED solution, as already discussed 730 about Fig. 12 and Fig. 15. This relationship is more evident for 731 the inclination (Fig. 17c) and the right ascension of the ascend-732 ing node (Fig. 17d). The relationships between orbital parame-733 ters and failures are analogous to those in Fig. 12 and Fig. 15, 734 but they are less clear because of the orbit determination error. 735 Overall, the computational time is similar to the one in Sec. 3.3. 736 To further appreciate FRED results, an alternative analysis, 737 analogous to the method described at the beginning of Sec. 2, 738 is carried out. Such an approach assesses the fragmentation 739 epoch as the time of the minimum relative distance between 740 parent and fragment mean states (both assumed as determinis-741 tic), propagated on the analysis time window. This would allow 742 a lower computational demand. The results are reported in Tab. 743 6, where a much smaller convergence to the correct solution 744 can be observed. Therefore, besides providing statistical in-745 formation and the correct solution among fragmentation epoch 746



Fig. 15: Perturbed scenario: relationship between the standard deviation associated to the computed fragmentation epoch and the fragment semi-major axis, eccentricity, inclination and right ascension of the ascending node. The fragments for which a failure occurs are highlighted according to the legend, and the dashed line shows the parent orbital parameters.



Fig. 16: Results for the perturbed scenario and accounting for the orbital state error introduced by the IOD process. The graph represents, for each fragment analysed, the relationship between the magnitude of the impulse which generated it (in logarithmic scale) and the time error between the estimated and the correct fragmentation epochs. The fragments for which a failure occurs are highlighted according to the legend.



Fig. 17: Perturbed scenario and accounting for the orbital state error introduced by the IOD process: relationship between the standard deviation associated to the computed fragmentation epoch and the fragment semi-major axis, eccentricity, inclination and right ascension of the ascending node. The fragments for which a failure occurs are highlighted according to the legend, and the dashed line shows the parent orbital parameters.

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	Correct solutions	$1 \min < t_{err} < T^p/2$	$t_{err} > T^p/2$
Relative distance	12.4 %	67.0 %	20.6 %

Table 6: Results for the perturbed scenario and accounting for the orbital state error introduced by the IOD process. A deterministic metrics is used, according to which the fragmentation epoch is assessed as the time of the minimum relative distance between the parent and the fragment mean state (both assumed as deterministic), propagated on the analysis time window.

candidates, FRED convergence to the correct solution turns outto be more robust.

749 3.5. Sensitivity Analysis

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A sensitivity analysis regarding the scenario in Sec. 3.4 is
 conducted to test FRED robustness. Operationally, three as pects may negatively affect the results:

- A larger time elapsed between the event and the IOD:
 given the IOD error, the larger the propagation time, the
 larger the mismatching at the fragmentation epoch.
- A wrong evaluation of the physical parameter of the fragment: the physical characteristics of the fragment can be either assumed or reconstructed during the IOD process, and this likely create an additional source of mismatching.
 - A larger measurements noise: this generally induces a more noisy IOD result, with larger mismatching between IOD mean state and larger covariance.

For all these aspects a sensitivity analysis is carried out as follows, by also comparing the FRED results with the ones obtained through the relative distance metrics introduced in Sec.
3.4.

767 3.5.1. Sensitivity Analysis on the IOD epoch

In Sec. 3.2, Sec. 3.3 and Sec. 3.4, the IOD epoch is always 768 set 13 h after the event, as FRED algorithm aims at provid-769 770 ing a method to identify the fragmentation epoch from a fragment orbital state determined in the first hours right after the 771 event. However, in real case scenarios, the algorithm may need 772 to be applied starting from an orbital state resulting from an 773 IOD conducted later. For this reason, it is fundamental to as-774 sess the FRED performance by considering larger time elapsed 775 between the fragmentation and the IOD epochs. Three cases 776 are investigated: 24 h, 48 h and 72 h from the event to the 777 first observation epoch. As above, the IOD method presented 778 in (Siminski, 2016) is applied. Results are reported in Tab. 7 779 and show a deterioration in performance, and this confirms that 780 the longer the time elapsed, the less robust the algorithm is. 781 Furthermore, a longer time elapsed implies a longer fragment 782 samples propagation, which increases the computational cost. 783

The FRED results are compared to those which could be obtained with the deterministic relative distance metrics, which are reported in Tab. 8. There is an oscillating behaviour of the correct solution, but the general trend confirms that the longer the time elapsed, the less performing the deterministic metrics. Moreover, the results are always much worse than the FRED ones.

3.5.2. Sensitivity Analysis on the B* mismatching

In the above analyses, the same B* (expressing the physical 792 parameter in the SGP4 propagator (Vallado et al., 2006)) is used 793 to generate the ground truth and inside FRED algorithm. This is 794 a strong assumption, as operationally no physical information 795 about the observed fragment is known. Generally, during an 796 OD process, the physical parameters can be estimated as well, 797 but accurate measurements are needed, as well as a long ob-798 servation arc (possibly obtained by linking more measurements 799 tracks). This is not the case for a single observation right after 800 a fragmentation event, and the physical parameters are either 801 roughly estimated or not estimated at all and, so, assumed. In 802 addition, the IOD procedure used (Siminski, 2016) estimates 803 the orbital state only, which is voluntarily not refined through 804 additional filters, as stated in Sec. 3.4. 805

To test FRED algorithm robustness to the physical parameter mismatching, a sensitivity analysis is carried out considering, inside the FRED algorithm, B^* values different from the one used to generate the ground truth. This modification is obtained by multiplying the correct B^* times: 1e+01, 1e-01, 1e-02, 1e-03, 0.

The results are reported in Tab. 9. FRED performance turns out to be robust to erroneous physical parameter estimation, and, for the 1e+01, the 1e-01, the 1e-03 and the 0 cases, the percentages are exactly the same as the nominal scenario ones (Tab. 5). Moreover, in the 1e-02 case the result for one fragment passes from being a compliant MOID failure to a correct solution. Overall, these results cannot be considered as a general behaviour, as the algorithm sensitivity on the physical parameters always depends on the perturbations experienced by the fragment and, so, on its orbital regimen. This is even more true considering the short propagation period of the simulation. For the scenario analysed, also the distribution of the relative distance metrics result does not change, as visible in Tab. 10.

3.5.3. Sensitivity Analysis on the measurements noise

As mentioned above, the performance of FRED algorithm 826 in operational scenarios strongly depends on the IOD accuracy, 827 which in turn depends on the algorithm used, the observation 828 geometry and length, and on the measurements quality. Indeed, 829 the deterioration of measurements can lead to two effects on 830 the IOD result and, so, on FRED performance: an erroneous 831 orbital mean state and a larger uncertainty. For this reason, it is 832 fundamental to assess FRED algorithm sensitivity to the mea-833 surements noise. In particular, since in surveillance radars (the 834 on-ground sensors of the nominal analysis) the angular track is 835 the less accurate measurement, the noise associated to the range 836 is kept fixed to the nominal value of 30 m, while the angular 837

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Time from the event	Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
24 h	60.8 %	5.7 %	0.0~%	33.5 %
48 h	43.1 %	3.8 %	1.0 %	52.1 %
72 h	31.6 %	2.4 %	0.5 %	65.5 %

Table 7: Perturbed scenario with orbital state error introduced by the IOD process: FRED results for the sensitivity analysis on the time elapsed between the fragmentation and the IOD epoch.

Time from the event	Correct solutions	$1 \min < t_{err} < T^p/2$	$t_{err} > T^p/2$
24 h	8.1 %	57.9 %	34.0 %
48 h	3.4 %	42.1 %	54.5 %
72 h	4.3 %	31.6 %	64.1 %

Table 8: Perturbed scenario with orbital state error introduced by the IOD process: deterministic relative distance metrics results for the sensitivity analysis on the time elapsed between the fragmentation and the IOD epoch.

noise is made varying from the nominal value of 1e-02 deg to:
2e-02 deg, 5e-02 deg and 1e-01 deg.

The results are reported in Tab. 11. It is possible to notice that 840 the larger the noise associated to the angular track, the lower 841 the convergence to the correct solution and the larger the pe-842 riodicity failures percentage. There is a slight increase also 843 in the MOID compliant failures, while the uncompliant ones 844 tend to zero. These results depend on the IOD result deteriora-845 tion, which introduces a mismatching affecting the estimation 846 of MOID data. On the one hand this may lead to a wrong eval-847 uation through the EMD metrics, with still the correct epoch 848 among candidates. On the other hand the IOD result may in-849 duce a wrong computation of time of parent transit through the 850 MOID and, so, the epoch candidates may be wrongly estimated, 851 and this may result in the absence of the correct solution among 852 candidates. In any case, FRED is always better performing than 853 the relative distance metrics, whose results are reported in Tab. 854 12. Also in this case there is a performance deterioration with 855 the angular noise increase. 856

857 4. Conclusions

The paper described FRED algorithm, which deals with 858 the fragmentation epoch identification problem focusing on the 859 case in which, besides the last available ephemeris of the par-860 ent object (assumed as a deterministic quantity), just one sin-861 gle fragment stochastic orbital state is available and already 862 linked to the event. The algorithm computes the fragmentation epoch candidates, which are ranked according to the matching 864 between MOID and relative distance distributions, given that, 865 at the actual fragmentation epoch, the MOID and the relative 866 distance were equal. To compute the statistical matching, three 867 metrics are discussed: the Mahalanobis distance, a tailored pro-868 cedure based on the quantiles coupled with a Principal Com-869 ponent Analysis and the Earth Mover's Distance. The latter is 870

eventually selected as the most performing and the most suitable for the problem, given the non-Gaussian distributions involved.

The numerical simulations highlighted that the algorithm reli-874 ability decreases when the observed fragment orbit has either 875 the period or the orbital plane similar to the parent object one, and a sensitivity analysis showed that there is no remarkable 877 dependence on the number of samples used in representing the 878 fragment orbital state. The inclusion of the perturbations and, 879 moreover, of the orbit determination error deteriorates the performance, but the correct fragmentation epoch can still be iden-881 tified among candidates. In addition the algorithm always fea-882 tures much better results with respect to an alternative determin-883 istic metrics based on the minimum relative distance between 884 the parent ephemeris and the fragment mean state propagated 885 on the analysis time window. A further sensitivity analysis 886 shows a deterioration proportional to the angular noise asso-887 ciated to the solution and to the time elapsed between the event 888 and the observation, but FRED is always much more perform-889 ing than the relative distance metrics. Instead, no remarkable 890 change occurs considering a mismatching between the actual 891 value of the fragment physical parameter and the one used in 892 the algorithm, but this depends on the fragment orbital regimen 893 and on the elapsed time from the event to the observation, and 894 so it is not possible to consider it as a general result.

In operational applications, FRED performance may be im-896 proved through multiple sensors contributions and by refining 897 the fragment orbital state with a smarter orbit determination 808 process (Montaruli et al., 2022b), by possibly exploiting the parent orbital state prediction as first guess for those fragments 900 generated by small magnitude impulses. Furthermore, the plau-901 sibility of FRED fragmentation epoch candidates can be exam-902 ined by tasking the sensors to point at the right ascensions and 903 declinations where the parent was at those epochs and retain 904 only candidates featuring a sufficient number of fragments de-905 tected. This action cannot be decisive, as periodicity failures 906 may share the same right ascension and declination as the correct solution, but it can support to shrink the candidates set. Fi-908 nally, the parallel use of FRED algorithm on different fragments 909 would allow to reach a higher level of confidence and precision 910 in the provided results, both in terms of the epoch candidates 911 set and of the one eventually returned by the algorithm. This 912 could be beneficial when multiple fragments are detected and 913 associated to the event, and they provide an orbit determination 914 result, but they are too few to be used in the deterministic ap-915 proaches mentioned in Sec. 1. 916

Concerning possible upgrades, the multivariate normal distribu-

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Factor multiplying B*	Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
1e+01	68.9 %	9.6 %	0.5 %	21.0 %
1e-01	68.9 %	9.6 %	0.5 %	21.0 %
1e-02	69.4 %	9.1 %	0.5 %	21.0 %
1e-03	68.9 %	9.6 %	0.5 %	21.0 %
0	68.9 %	9.6 %	0.5 %	21.0 %

Table 9: Perturbed scenario with orbital state error introduced by the IOD process: FRED results for the sensitivity analysis on the B*.

Factor multiplying B*	Correct solutions	$1 \min < t_{err} < T^p/2$	$t_{err} > T^p/2$
1e+01	12.4 %	67.0 %	20.6 %
1e-01	12.4 %	67.0 %	20.6 %
1e-02	12.4 %	67.0 %	20.6 %
1e-03	12.4 %	67.0 %	20.6 %
0	12.4 %	67.0 %	20.6 %

Table 10: Perturbed scenario with orbital state error introduced by the IOD process: deterministic relative distance metrics results for the sensitivity analysis on the B*.

Angular noise [deg]	Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
2e-02	66.5 %	11.5 %	0.5 %	21.5 %
5e-02	53.1 %	20.6 %	0.0 %	26.3 %
1e-01	33.5 %	29.7 %	0.0 %	36.8 %

Table 11: Perturbed scenario with orbital state error introduced by the IOD process: FRED results for the sensitivity analysis on the angular track noise.

Angular noise [deg]	Correct solutions	$1 \min < t_{err} < T^p/2$	$t_{err} > T^p/2$
2e-02	13.4 %	57.4 %	29.2 %
5e-02	12.0 %	59.8 %	28.2 %
16-01	11.3 %	51.9%	50.0 %

Table 12: Perturbed scenario with orbital state error introduced by the IOD process: deterministic relative distance metrics results for the sensitivity analysis on the angular track noise.

tion used represents the most generic approach, but an alterna-918 tive and less computational demanding way of covariance prop-919 agation may be integrated in the process. In addition, the algo-920 rithm considers the last available ephemeris as a deterministic 921 information, while an uncertainty is associated also to it and 922 may be included in the overall process. Another aspect which 923 may be further studied is the fragmentation epoch candidates 924 ranking strategy, which is currently performed based on the sta-925 tistical matching between the relative distance and the MOID 926 distributions, but which may profit from other conjunction anal-927 928 ysis tools, like the long-term risk assessment. Finally, it would be interesting to deal with the fragmentation epoch identifica-929 tion problem in the case that it is not possible to determine the 930 fragment orbital state, with a tailored procedure conducted in 931 the measurements space. To this end, developing an approach 932

to solve a track to track association problem to link multiple measurements referred to a same fragment would allow to derive an orbit determination result and to exploit FRED algorithm. Overall, all these possible algorithm improvements and developments should be carried out together with test on real data and the final operational implementation shall include a detailed computational demand assessment and minimisation.

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Appendix

The same analysis as in Sec. 3.3, that is a perturbed scenario with no IOD error, is here conducted simulating the fragmentations of two objects: the COSMOS 1490, flying in Medium Earth Orbit (MEO), and the EDRS-C, flying in geostationary

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	<i>a</i> [km]	е	i [deg]	Ω [deg]	ω [deg]	θ [deg]
COSMOS 1490	2.6e+04	1.5e-03	64.2	130.2	18.9	182.1
EDRS-C	4.2e+04	6.2e-05	0.0	89.9	165.5	104.3

Table 13: COSMOS 1490 and EDRS-C orbital parameters simulated on November 15th 2022, at 02:47:00 UTC.

	Last available ephemeris	Event alert	Orbit determination
	epoch (UTC)	epoch (UTC)	epoch (UTC)
COSMOS 1490	November 14 th , 02:07:03	November 16 th , 00:00:00	November 16 th , 07:35:00
EDRS-C	November 12 th , 18:26:49	November 18 th , 06:00:00	November 18 th , 10:00:20

Table 14: Epochs of COSMOS 1490 and EDRS-C last available ephemerides and event alert

	Correct solutions	MOID failures compliant	MOID failures uncompliant	Periodicity failures
COSMOS 1490	89.9 %	9.1 %	0.0 %	1.0 %
EDRS-C	86.5 %	11.5 %	0.5 %	1.9 %

Table 15: FRED results for the COSMOS 1490 and EDRS-C simulated fragmentations. The perturbed scenario with no orbital state error is assessed, and the EMD metrics is used.

orbit (GEO). Analogously to Sec. 3.1, the fragmentation event 954 is simulated at 02:47:00 UTC of November 15th, 2021, and 955 modelled through the same set of impulses. The orbital parame-956 ters of the two parent objects at the break-up epoch are reported 957 in Tab. 13. 958

The epochs of the last available ephemeris, of the considered 959 event alert and of the orbit determination result are reported in 960 Tab. 14. These epochs were selected to set an analysis time 961 window which includes the same number of periodicities as the 962 one in Sec. 3. Similarly to Sec. 3.3, at the orbit determination 963 epoch a covariance is associated, with inertial position and velocity standard deviations of 1.4e+00 km and 2.5e-04 km/s, for 965 the COSMOS 1490, and 3.1e+00 km and 4.9e-04 km/s, for the 966 EDRS-C. These quantities were derived from an orbit determi-967 nation process. 968

The results are reported in Tab. 15. It can be noticed that both 969 for COSMOS 1490 and EDRS-C the convergence to the cor-970 rect solution is similar to the one in Sec. 3.3. The increase in 971 compliant MOID failures is motivated by the larger propagation 972 time window, which makes the samples more spread, resulting 973 in possible wrong fragmentation epoch estimates provided with 974 a time error smaller than the associated uncertainty. Overall, 975 this analysis confirms FRED general applicability, as the algo-976 rithm behaviour does not depend on the orbital regimen of the 977 fragmentation event.

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Declaration of interests

 \boxtimes The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

□The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

