1	Hf isotope systematics of seamounts near the East Pacific Rise
2	(EPR) and geodynamic implications
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23	May 20, 2016 version

## 24 Abstract

We report new Hf isotopic data for basaltic glasses from seamounts flanking the East Pacific Rise (EPR) between 5° and 15°N that have been previously analyzed for Sr-Nd-Pb isotopes as well as major and trace elements. The Hf isotopic data offer new perspectives on the petrogenesis of these samples in a broader context on mantle dynamics. The Hf isotope compositions show significant correlations with Sr-Nd-Pb isotopes and with both abundances and ratios of incompatible elements. The seamount lavas are thus best interpreted as products of melting-induced mixing in a two-component mantle.

32 The range in composition of EPR seamount lavas cannot be generated by simple mixing of melt and melting of variably heterogeneous mantle in which enriched and 33 depleted materials contribute equally to melting (source mixing). Instead, the trace element 34 35 and isotope compositions of seamount lavas can be reproduced by melting models in which more enriched, fertile mantle component are preferentially melted during mantle upwelling. 36 At progressively lower degrees of melting, erupted lavas are thus more enriched in 37 incompatible trace elements, have higher <sup>87</sup>Sr/<sup>86</sup>Sr, <sup>208</sup>Pb/<sup>204</sup>Pb ratios and lower <sup>143</sup>Nd/<sup>144</sup>Nd, 38 <sup>176</sup>Hf/<sup>177</sup>Hf ratios. The "EM1" and "pyroxenite" endmember might be the suitable enriched 39 component. 40

The Hf-Nd isotopic variations on global scale might result from the variations in amounts of residual continental lithospheric mantle that detached into upper mantle during continental rifting. The significant correlations of Rb/Sr vs <sup>87</sup>Sr/<sup>86</sup>Sr, Sm/Nd vs <sup>143</sup>Nd/<sup>144</sup>Nd and Lu/Hf vs <sup>176</sup>Hf/<sup>177</sup>Hf give pseudochron ages of 182±33 Ma, 276±50 Ma and 387±93 Ma, respectively. These different "ages" have no significance, but result from meltinginduced mixing with the pseudochron slopes controlled by the compositions of enriched 47 component and depleted end-member.

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Keywords: East Pacific Rise; seamounts; melting-induced mixing; Hf isotopes; Hf-Nd
correlation.

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# 52 **1. Introduction**

53 Isotopic studies of oceanic basalts provide constraints on models of Earth's chemical differentiation and convection processes. One of the major advances in the solid Earth 54 science over the past 50 years is the recognition of mantle compositional heterogeneity and 55 the identification of several isotopically distinct mantle end-members through these 56 isotopic studies (e.g., Gast et al., 1964; White, 1985; Zinder and Hart, 1986). Numerous 57 58 geochemical studies of MORB (e.g. Allègre and Turcotte, 1986; Arevalo and McDonough, 2010; Donnelly et al., 2004; Hirschmann and Stolper, 1996; Niu et al., 1999; Phipps 59 Morgan and Morgan, 1999; Schilling et al., 1983; Zindler and Hart, 1986; Haase et al., 60 2011; Brandl et al., 2012) have shown that even in the absence of nearby hotspots, the 61 upper mantle is chemically and isotopically heterogeneous. The origin of these 62 heterogeneities is debated, but they are likely results from recycled material (Donnelly et 63 64 al., 2004; Niu and O'Hara, 2003; Pilet et al., 2005; White and Hofmann, 1982). The Nd and Hf isotope ratios (i.e., <sup>143</sup>Nd/<sup>144</sup>Nd and <sup>176</sup>Hf/<sup>177</sup>Hf) are shown to be well correlated in ocean 65 island basalts (OIB) due to the similar behavior during mantle melting the parent-daughter 66 (P/D) pairs (i.e.,  $D_{Sm} > D_{Nd}$  and  $D_{Lu} > D_{Hf}$ ) (Blichert-Toft, 2001; Patchett and Tatsumoto, 67 1980; Salters and Hart, 1991; Vervoort et al., 1996). In contrast to OIB, mid-ocean ridge 68 basalts (MORB) on a global scale have large variations in Hf isotopic composition at a 69

given Nd isotopic composition (e.g., Blichert-Toft et al., 2005; Debaille et al., 2006; Johnson and Beard, 1993; Patchett, 1983; Patchett and Tastumoto, 1980; Salters and Hart, 1991; Salters and White, 1998; Salters and Zindler, 1995). A more recent study, however, shows that Hf and Nd isotopes in MORB are in fact well correlated on some ridge segment scales (Salters et al., 2011). This is an important observation that points to large regional scale differences in mantle sources and source histories.

76 To help address issues on MORB genesis and the variable MORB Hf-Nd isotopic correlations at different ridges (Salters et al., 2011), we study Hf isotope compositions of 77 near-ridge seamounts flanking the East Pacific Rise (EPR) (see Fig. 1 and Fig. A1). We 78 79 choose these seamount lavas because (1) they have already been well-characterized for major elements, trace elements, and Sr-Nd-Pb isotope compositions (Niu and Batiza, 1997; 80 Niu et al, 2002); (2) they represent an integral part of EPR ridge magmatism as they were 81 82 derived from the same upper mantle source of EPR axial MORB; (3) they have avoided 83 melt homogenization during melt aggregation in the mantle and in the long-lived axial magma chambers, and thus more faithfully record the nature of the MORB mantle source 84 beneath the EPR than axial lavas; and importantly, (4) they display a large compositional 85 spectrum that encompasses much of the global MORB compositional variability from 86 ridges unaffected by mantle plumes/hotspots, and can thus help answer first-order 87 questions on MORB mantle sources and processes, especially beneath the fast-spreading 88 89 EPR.

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91 **2. Samples and analytical methods** 

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The samples were dredged during the 1988 Raitt 02 expedition aboard R/V Thomas

Washington (Batiza and Niu, 1992; Batiza et al., 1990). All the samples studied are basalt
glasses. They were collected from near-ridge seamounts within the 1 Ma isochron (<~ 60</li>
km) of the EPR axis between 5° and 15°N on both the Pacific and Cocos Plates (Fig. 1).
These 36 samples have been studied for major and trace elements (Batiza and Niu, 1992;
Batiza et al., 1990; Niu and Batiza, 1997) and Sr-Nd-Pb isotopes (all samples were
analyzed for Sr isotope, 28 of them were analyzed for Nd isotope and 34 of them were
analyzed for Pb isotope, see Niu et al., 2002).

100 All the samples were carefully hand-picked under a binocular microscope before they were leached at room temperature in 10% H<sub>2</sub>O<sub>2</sub> for a few minutes to remove Mn oxides in 101 102 possible micro-fractures and other potential labile contaminants. The samples were then repeatedly washed ultrasonically in 18 megohm Milli-Q water before digestion. The 103 method of rock digestion and Hf separation follows those of Yang et al. (2010). Hf isotopic 104 analysis was done using a Neptune MC-ICP-MS in the Institute of Geology and 105 106 Geophysics, Chinese Academy of Sciences (IGGCAS). The analytical details for Hf isotopic measurements are given by Li et al. (2005). Hf isotopic compositions were 107 normalized to <sup>179</sup>Hf/<sup>177</sup>Hf=0.7325. In our analysis, we measured <sup>176</sup>Hf/<sup>177</sup>Hf values for 108 USGS reference rock standards W-2 (0.282711±0.000013 vs. reference value: 109 0.282715±0.000030, Le Fèvre and Pin, 2001), BHVO-2 (<sup>176</sup>Hf/<sup>177</sup>Hf=0.283082±0.000010 110 vs. reference value: 0.283096±0.000020, Weis et al., 2005) and BCR-2 111 (<sup>176</sup>Hf/<sup>177</sup>Hf=0.282858±0.000009 vs. reference value: 0.282870±0.000008, Weis et al., 112 2007). We also measured the in-house standard Alfa Aesar hafnium solution produced by 113 Johnson Matthey Company (i.e. Alfa Hf, 10000 µg/ml, stock No. 14374, plasma standard 114 115 solution) (Wu et al., 2006, 2007; Yang et al., 2007). The measured value for Alfa Hf gave an average  ${}^{176}$ Hf/ ${}^{177}$ Hf of 0.282181±0.000010 (2 $\sigma$ , n=7) (reference value: 0.282189±0.000019, Wu et al., 2006). All of these are indistinguishable from their reference values within analytical errors.

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120 **3. Results** 

The Hf isotope data are given in Table 1 and presented in Fig. 2. The  $\varepsilon_{Hf}$  are calculated 121 using ([<sup>176</sup>Hf/<sup>177</sup>Hf]<sub>sample</sub>/[<sup>176</sup>Hf/<sup>177</sup>Hf]<sub>CHUR</sub>-1)×10000 (the subscript sample and CHUR 122 denote the value of samples and chondrite), in which,  $[^{176}Hf/^{177}Hf]_{CHUR} = 0.282772$ 123 (Blichert-Toft and Albarède, 1997). The ENd values are calculated in the same manner, and 124 the [<sup>143</sup>Nd/<sup>144</sup>Nd]<sub>CHUR</sub> =0.512638 (Dickin, 1997). The ε<sub>Hf</sub> values for seamount lavas vary 125 from 15.61 to 6.86. The  $\varepsilon_{Hf}$  and  $\varepsilon_{Nd}$  correlate well (see Fig. 2), giving a linear expression 126 of  $\varepsilon_{Hf}$ =1.72 $\varepsilon_{Nd}$ -2.83 with R<sup>2</sup>=0.91. This slope is significantly similar to, although slightly 127 shallower than the mantle array defined by MORB and OIB (i.e.,  $\epsilon_{Hf}=1.59\epsilon_{Nd}+1.28$ , 128 129 Chauvel et al., 2008). The Hf and Nd isotope analyzed in this study cover the range defined by previous analyses of northern EPR MORB (Salters et al., 2011 and reference therein), 130 131 and more enriched than MORB from southern EPR (Fig. 2). The large range of isotopic variability of the near-EPR seamount lavas suggests that the scale of mantle source 132 heterogeneities may be as small as a few kilometers or even smaller, as recognized 133 134 previously (e.g., Batiza and Vanko, 1984; Niu and Batiza, 1997). Large variety in compositions for lavas from seamount located at 12°45'N of EPR also implied that the 135 upper mantle is extremely heterogeneous on the scale of a single seamount (Brandl et al., 136 2012). It is worthy to note that the HIMU-like samples (R78-5-1 and R83-2, shown in green 137 138 circles in Fig. 2), are offset from the other samples. It is most likely that some mantle source 139 heterogeneities exist on a small scale and the enriched component may have different origins and, thus, possess different isotopic signatures (Niu et al., 1999, 2002). 140

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#### 4. Discussion 142

#### 4.1 The effects of fractional crystallization 143

144 The major element compositions had been published previously (Niu et al., 2002). 145 The low Mg# values (70 to 52) indicate that none are likely to represent primary melts in equilibrium with mantle olivine (Fo = 89), and all have probably undergone some degree 146 147 of crystal fractionation. However, the large variations in major and trace element are difficult to explain by fractional crystallization. Additionally, the correlations between 148 isotope compositions and incompatible trace element ratios, and major and trace element 149 150 concentrations (Fig. 3, and 4) are also inconsistent with closed system crystal fractionation. Other processes other than fractional crystallization must be responsible for variation in 151 composition of most seamount lavas. Considering the composition of seamount lavas is 152 similar to, and more primitive than axial lavas, the crustal contamination is also negligible. 153

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# 4.2 Melting of two component mantle

4.2.1 Evidence for melting of a component source 155

156 Our new Hf isotopic data correlate with Sr, Nd, and Pb isotopes, with ratios of 157 incompatible elements (Fig. 4) and with the abundances of major-trace elements (Fig. 3). Previous studies also showed that the Sr-Nd-Pb isotopic data of these samples correlate 158 with each other, with ratios of incompatible elements and with the abundances and ratios 159 of major elements (Niu and Batiza, 1997; Niu et al., 1999, 2002). The large variation in 160 161 major-trace element and isotope and the correlations are result from source heterogeneity

and easily explained by mixing between enriched melt (lower MgO, <sup>143</sup>Nd/<sup>144</sup>Nd, 162 <sup>176</sup>Hf/<sup>177</sup>Hf, higher incompatible element concentrations and <sup>87</sup>Sr/<sup>86</sup>Sr ratios) and relatively 163 depleted melt. Simple mixing of melts derived from lithologically distinct enriched and 164 depleted mantle in variable proportions will result in linear arrays in element-element 165 diagrams, and in all ratio-ratio diagrams in which the denominator element is the same (i.e., 166 A/X vs. B/X diagram) (Langmuir et al., 1978). For the same reason, melting of a source 167 168 composed of two different lithologies, which are mixed in variable proportions (mixing of 169 sources) but melt to the same extent, also give a straight mixing line. However, the seamount data in this study define curved arrays (Fig. 5), which is inconsistent with the 170 171 above mentioned mixing. Similar curved arrays have been reported for lavas from the fossil Galapagos Rise (Haase et al., 2011) and seamounts east of the EPR at 12°45'N (Brandl et 172 al., 2012). Variable degrees melting of a two-component mantle in which different lithologies 173 have different trace element and isotope compositions but also different melting behavior may 174 account for the chemical and isotopic variation (Niu and Batiza, 1997; Niu et al., 1999, 2002; 175 Paulick et al., 2010; Haase et al., 2011; Brandl et al., 2012). This concept emphasizes the 176 177 physical state of a mantle with enriched component as easily melted domains in a more 178 depleted matrix. Increasing evidence suggested that the upper mantle source of oceanic 179 basalts consists of relatively fertile and more refractory lithologies. They have different solidus temperatures and thus begin melting at different depths during adiabatic upwelling. 180 The most important assumption is that the enriched mantle lithologies have lower solidus 181 temperature than the more depleted matrix, thus are preferentially tapped at low degrees of 182 melting (Sleep, 1984; Prinzhofer et al., 1989; Ito and Mahoney, 2005). It is evidenced by 183 184 the clinopyroxenes in residual abyssal peridotites tend to have more radiogenic Nd isotope 185 than lavas from the same section of ridge, indicating a preferential melting of eclogite or pyroxenite with lower Sm/Nd and <sup>143</sup>Nd/<sup>144</sup>Nd ratios during decompression melting
(Salters and Dick, 2002). Melting of the fertile lithologies would extract heat from the more
refractory matrix, and inhibit the melting of the latter; once the more refractory lithologies
begin to melt, the remaining fertile materials may stop melting (Phipps Morgan, 2001), and
the melting paths may be curved.

There have been several attempts to quantify the range in melt compositions produced by melting of two-component mantle (Ito and Mahoney, 2005; Phipps Morgan, 2001; Phipps Morgan and Morgan, 1999; Stracke and Bourdon, 2009). The contribution of fertile lithologies decreased with the increasing degrees of melting, and lavas produced at lower degree of melting are expected to have higher incompatible element concentrations, higher <sup>87</sup>Sr/<sup>86</sup>Sr, and lower <sup>143</sup>Nd/<sup>144</sup>Nd ratios (e.g., Stracke and Bourdon, 2009).

In the following section, the melting model of Stracke and Bourdon (2009) was used to quantitatively model the melt compositions derived from melting of a two-component mantle, in order to investigating the effects of varying solidus temperature, initial proportion of each component, and compositions f of both endmembers, and constraining the mantle heterogeneity beneath the EPR.

4.2.2 Quantitative modeling of the melting process

The trace-elemental and isotopic compositions, melting behavior, and initial proportion ratio of each component all play an important role in determining melt compositions. Here we assume the simplest case of a two-component mantle, consisting of a volumetrically minor enriched component dispersed in the depleted peridotite matrix. The initial proportion of enriched component is assumed to be no more than 10%. These two components have different solidus temperatures and thus begin melting at different depths during upwelling. Peridotite melting is assumed to start at ~2.5 GPa, and the beginning depth of the enriched component melting is significantly deeper than to similar to that of peridotite melting (Stracke and Bourdon, 2009). Mineral modes and partition coefficients are given in Appendix 1.

The origin of enriched component is unclear, but likely result from subduction and recycling of oceanic crust, sediments, metasomatised oceanic lithosphere or mantle wedge material (e.g., Hofmann and White, 1982; Niu and O'Hara, 2003; Donnelly et al., 2004).

216 Recycled ancient oceanic crust might be the enriched component (i.e., ROC model) (e.g., Hofmann and White, 1982), although Niu and O'Hara (2003) suggested that oceanic 217 218 crust is isotopically too depleted. As a variant of the ROC model, Donnelly et al. (2004) suggested that low-degree melts of subducted oceanic crust could metasomatize the 219 overlying mantle wedge, and the metamomatized mantle wedge could recycle by mantle 220 221 convection to the ridge supplying enriched MORB (E-MORB) source. One difficulty with 222 the model is that the cold oceanic crust that has undergone dehydration may be too refractory to melt at subduction zone. Another problem is that mantle convection model 223 (Hofmann, 1997; Tolstikhin et al., 2006) may not support this scenario, and the mantle 224 wedge "corner flow" (Niu, 2005) is blocked by the subducting slab, thus, it is difficult for 225 the metasomatized mantle wedge to reach the ocean ridge. Terrigenous sediments may 226 provide enriched composition to depleted mantle (DM) source because few OIB source 227 may have a component of subducted terrigenous sediments (Jackson et al., 2007; Mahoney 228 et al., 1995). However, continental crust (CC) and global subducting sediments (GLOSS) 229 230 are characterized by depletions in Nb, Ta, P and Ti and enrichment in Pb (Plank and 231 Langmuir, 1998; Rudnick and Gao, 2003). Such a "CC-like signature" is so strong that it 232 should be inherited in the incompatible element characteristics of OIB and E-MORB if 233 terrigenous sediments were indeed the enriched component of DM (Niu et al., 2012). However, it is not observed in global oceanic basalts except for two OIB suites (Jackson et 234 al., 2007; Mahoney et al., 1995). Furthermore, there is no convincing evidence that the 235 terrigenous sediments are the geochemically enriched components of DM on global scale. 236 Recycled metasomatized oceanic lithospheric mantle (ROLM) had also been proposed as 237 238 a possible enriched component (Niu, 2008, 2009; Niu and O'Hara, 2003, 2009; Niu et al., 239 2011, 2012). The deep portion of oceanic lithospheric mantle can be enriched in terms of incompatible elements throughout its long histories through mantle metasomatism by low-240 241 degree melt. The melt collects atop the seismic low velocity zone (LVZ), and then percolates through and metasomatizes the growing lithosphere via basal accretion. The 242 percolating/metasomatizing Low-degree melt can precipitate mineral phases (e.g., 243 amphiboles; 244 "macroscopic metasomatism") before finally absorbed ("cryptic 245 metasomatism") in the lithosphere, generating metasomatically veined mantle (e.g., Niu, 2008). 246

No matter what the enriched component is (recycled crust or lithospheric mantle), the 247 complex process (i.e., melting, metasomatism, and dehydration during subduction) that 248 they had undergone would increase difficulties in determining their initial trace-element 249 250 and isotopic compositions. Nevertheless, previous research suggested that the enriched 251 endmember of MORBs array in Hf-Nd isotopic diagram tends to plot either within or in 252 the extension of the OIB array, and none of them towards the HIMU component (Salters et al., 2011). Thus, the previous published trace element and isotope compositions of 253 pyroxenite (recycled MORB), EM1 and EM2 were used to represent that of the enriched 254

component in the modeling. The isotopic compositions of pyroxenite, EM1 and EM2 are
from Ito and Mahoney (2005). And the trace element compositions of pyroxenite, EM1 and
EM2 are from Stracke and Bourdon (2009), Willbold and Stracke (2006), and Workman
et al. (2004), respectively.

The compositions of depleted endmember are generally estimated from the 259 compositions of oceanic basalts and abyssal peridotites. Actually, the estimated 260 261 compositions of depleted MORB mantle (DMM) based on that of abyssal peridotites might 262 more faithfully reflect the characteristics of depleted upper mantle due to that the enriched component had been exhausted. And the MORBs, including the most depleted Garrett 263 264 Transform lavas, are enriched than depleted endmember (discussed below). It is supported by the abyssal peridotites that components of the upper mantle beneath spreading ridges, 265 which have more radiogenic Nd and Hf isotope compositions than those of lavas from the 266 267 ridge (Salters and Dick, 2002; Stracke et al., 2011). Thus, the estimated composition of DMM mantle based on numerous abyssal peridotites following Workman and Hart (2005) 268 were used here as the depleted endmember in the modeling. The assumption may or may 269 not be realistic, however, it is reasonable. Further researches are needed to better constrain 270 271 the compositions of depleted/enriched endmembers.

The results of fractional melting model following Stracke and Bourdon (2009) are shown in Figures 5-7. The enriched components are assumed to preferentially melting than the more depleted matrix. The parameters used in the modeling are summarized in Appendix 1. The compositions, mineralogy, melting behavior, and proportion of both components must be more various than that used in the modeling, however, the melting curves still provide important clues on melting of a heterogeneous mantle that composed of at least two component (i.e., enriched and depleted component). Then, the effects of
different variables on melt compositions would be discussed below.

The melting modeling can broadly reproduce the range in compositions of the 280 seamount lavas in this study (Fig. 5, 6, 7). In the La/Nd versus Yb/Nd and La/Yb versus 281 Sm/Yb diagram (Fig. 5), variable degrees of melting of a heterogeneous mantle fits the 282 curved array defined by EPR seamount lavas that mentioned above, which is contrast with 283 284 the linear arrays of simple mixing. The melting curves have similar trend, however, the 285 "pyroxenite" endmember fits best to the seamount data in this study and also the previous published axial MORB from EPR 5-15°N. The deviation of "EMI" trajectory from the 286 287 seamount data, especially in Fig. 5b, is most likely result from the extremely higher incompatible element concentration and the higher degree of fractionation between LREEs 288 and HREEs in "EMI" source. Meanwhile, the degree of enrichment in incompatible 289 element increased with the increasing incompatibility. Consequently, the "EMI" 290 291 endmember is too enriched to fits the seamount lava in Fig. 5b, although the incompatible element concentration decreased dramatically with the increasing degrees of melting. 292 Previous researches suggested that the curvature of melt curves increased with the 293 increasing proportion of enriched endmember relative to depleted endmember (See Figure 294 8c in Brandl et al., 2012). And increasing in enriched component/depleted component ratio 295 may lead the EMI endmember better fits the seamount data in Fig. 5a, however, it is invalid 296 297 in Fig. 5b.

The effects of difference in solidus temperatures between the two endmembers on melt compositions were shown in Fig. 6. Peridotite melting is assumed to start at ~2.5 GPa (Stracke and Bourdon, 2009). And the melt evolutionary paths that the beginning depth of

301 the enriched component melting is assumed to be similar to and significantly deeper than 302 that of peridotite melting were exhibited in solid lines and dashed lines, respectively. The largest melt heterogeneity in the ratios of trace element is observed for small differences in 303 304 solidus temperature between the two components (Fig. 6). If the contrast in solidus temperature is too large, the enriched component would be completely exhausted before 305 peridotite melting, and the clear correlations between trace element and isotope ratios of 306 307 the seamount lavas in this study would not be observed in the pooled melts (Fig. 6). The 308 seamount data lies between the melt curves at different solidus temperatures, and "EM1 /EM2" endmember, especially in condition that the solidus temperatures of the two 309 310 endmembers are similar, could better reproduce the seamount data. The "pyroxenite" endmember is significantly depleted in Sr isotope than the enriched seamount lavas (Fig. 311 6b), indicating that "pyroxenite" endmember is not a reasonable enriched component or 312 313 other enriched components must also involve in the source. The melting modeling in correlations between Sr-Nd-Pb isotopes further supported that the enriched component 314 might be complex, and it is most likely composed of EM1 and/or pyroxenite (Fig. 7). Melt 315 compositions in assumption that the initial proportion of EM1 and pyroxenite is 0.06 and 316 317 0.04, respectively, can match well with the isotopic compositions of seamount lavas (Fig. 7b, c). 318

In summary, the large variation in trace element ratios (e.g., La/Sm and Nb/Zr) and isotope of seamount lavas in this study and the correlations between them is the result of melting of a heterogeneous mantle. And "EM1" and/or "pyroxenite" as the enriched component could better fits the seamount data and axial MORB from EPR at 5-15°N both in trace element and isotope compositions in the modeling. Thus, we infer that the metasomatized lithospheric mantle and/or oceanic crust might be the reasonable enriched component in mantle source beneath northern EPR. The parameters used in the modeling should be further constrained.

Furthermore, the most depleted seamount lavas in this study and previous published 327 328 axial MORB from northern EPR approximate to the depleted endmember, however, it is enriched than the mantle source. Previous researches also suggested that the average 329 330 incompatible trace element and isotope composition of mantle source is significantly more 331 depleted than mantle-derived magmas (e.g., Stracke and Bourdon, 2009; Brandl et al., 2012). Oceanic basalts were often used to infer the composition of the upper mantle. 332 333 However, if melting of two-component mantle is common and an important process, then the studies of oceanic basalts that assume that the isotope compositions of basalts are 334 identical to those of their mantle source might be wrong. 335

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# **4.3 Implications for Hf-Nd isotopic variations on a global scale**

Previous studies show that there is a good correlation between Hf and Nd isotope for 338 global MORBs (e.g., Chauvel et al., 2008; Vervoort et al., 2011). However, a later 339 comprehensive study of MORBs (Salters et al., 2011) shows that the EHF values of global 340 341 MORBs varies widely at given  $\varepsilon_{Nd}$  value. Here we use  $\Delta \varepsilon_{Hf}$  values to describe the derivation of the measured  $\varepsilon_{Hf}$  value from that of the mantle array defined by Chauvel et 342 al. (2008), i.e.,  $\epsilon_{Hf}=1.59\epsilon_{Nd}+1.28$ . It is obvious that the  $\Delta\epsilon_{Hf}$  values of Pacific MORBs, 343 Atlantic MORBs and Indian MORBs are different (Fig. 8b). In general, MORBs and 344 345 seamount lavas of Pacific spreading centers have the lowest and relatively constant  $\Delta \epsilon_{\rm Hf}$ values (Fig. 8). However, Atlantic and Indian MORBs range widely in  $\Delta \varepsilon_{\rm Hf}$  values, and the 346

347 Atlantic MORBs, especially the North Atlantic MORBs, have the highest  $\Delta \epsilon_{\rm Hf}$  values (Fig. 8). The fundamental reason must be the global scale heterogeneity of DMM, i.e., the large 348 349 variations in  $\Delta \epsilon_{\rm Hf}$  values are most likely dominated by local variations of source compositions. The above discussed fractional melting modeling of a heterogeneous mantle 350 also supported that the isotopic composition of melt is mainly controlled by source 351 compositions, melting behavior and initial proportion of each component. Addition of a 352 353 component with high  $\Delta \epsilon_{\rm Hf}$  values to the mantle source might lead the melt towards high  $\Delta \epsilon_{\rm Hf}$  values on different extent. The mantle component with high  $\epsilon_{\rm Hf}$  value is most likely to 354 355 be the residues of ancient melting events with garnet present as a residual phase. The partition coefficients of Lu between garnet and basaltic liquids is much higher than that of 356 Hf (i.e.,  $D_{Lu}/D_{Hf} >> 1$ ), and the residues produced by melt extraction at great depth where 357 garnet is stable and as a residual phase of the source would impart an elevated Lu/Hf ratios 358 and evolve to high <sup>176</sup>Hf/<sup>177</sup>Hf in the residues if the partial melting event is ancient (e.g., 359 360 Schmidberger et al. 2002). Salters et al. (2011) proposed that various amounts of residual oceanic lithosphere (ReLish) can account for the variation in  $\varepsilon_{Hf}$  value at constant Nd 361 isotope. And the amount of ReLish that contributes to the Pacific MORB was suggested to 362 363 be relatively constant and that various widely as to Atlantic and Indian MORBs (Salters et al., 2011). The more uniform  $\Delta \epsilon_{\rm Hf}$  value of Pacific MORB may also simply reflect the fact 364 365 that melt mixing in magma chambers at fast-spreading ridges is more efficient (Rubin and Sinton, 2007). 366

Alternatively, the sub-continental lithospheric mantle (SCLM) was also proposed to be a suitable component to produce the high Hf isotope of lavas from Lucky Strike Ridge that along the northern Mid-Atlantic Ridge (Hamelin et al., 2013). The Pacific spreading 370 center is always suggested to be ancient and not associated with the continental rifting, thus, 371 the continental lithospheric mantle is absent or exhausted. In contrast, the opening time of Atlantic and Indian spreading center is no longer than 200 Ma, and the continental 372 lithospheric mantle that detached into the upper mantle can feed the MORB source. It had 373 been approved that the continental materials could introduce into the Atlantic-Indian upper 374 mantle during continental rifting (e.g., Torsvik et al., 2013). Furthermore, ancient depleted 375 376 continental lithospheric mantle with extremely radiogenic Hf isotopes had been observed 377 (Schmidberger et al., 2002; Bedini et al., 2004; Simon et al., 2007; Chu et al., 2009). Although the Hf and Nd isotopic compositions of SCLM range widely from negative to 378 379 positive values, most of them have extremely high  $\varepsilon_{Hf}$  value (even to +200) (Fig. 8) that can interpret the lowest and relatively constant  $\Delta \varepsilon_{\rm Hf}$  values for Pacific MORBs and largely 380 various and higher  $\Delta \epsilon_{\rm Hf}$  values for Atlantic and Indian MORBs. 381

### 382 **4.4 Pseudochron ages**

Fig. 9 shows that our samples give Rb-Sr, Sm-Nd and Lu-Hf pseudochron ages of 383 182±33 Ma, 276±50 Ma and 387±93 Ma, respectively. The ages have been reported at 384 385 other locations of ocean ridge and nearby seamounts (Donnelly et al., 2004; Niu et al., 1996; Zindler et al., 1984). Zindler et al. (1984) suggested that the Sm-Nd pseudochrons of near-386 EPR seamounts have age significance and interpreted that to be controlled by (1) the 387 position of depleted and enriched end-member in <sup>143</sup>Nd/<sup>144</sup>Nd vs. Sm/Nd space and (2) 388 389 independent magma source of two end-members. The Rb-Sr and Sm-Nd pseudochron ages 390 of near-EPR seamounts (187 Ma and 238 Ma, respectively) are similar to our results, which 391 were interpreted to provide constraints on the origin and timing of E-MORB source enrichment (Donnelly et al., 2004). However, Niu et al. (1996) suggested that these 392

393 pseudochrons are statistically significant mixing trends without chronological significance. 394 These ages shown by different radioactive decay systems are non-conformable, suggesting indeed that the "ages" are of no chronological significance, but are best 395 explained as results of melting-induced mixing with the pseudochron slopes controlled by 396 the compositions of the enriched component and the depleted end-member. The fractional 397 melting modeling can reproduce the arrays to a first order (Fig. 9) with the "EM1" 398 399 endmember fits the seamount data better. It is consistent with the above conclusion that "EM1" is a possible origin of enriched component. The source mantle contains a 400 volumetrically minor enriched component, and with increasing degree of melting, the 401 402 contribution to the pooled melt from the more fertile component with higher Rb/Sr, <sup>87</sup>Sr/<sup>86</sup>Sr, and lower Sm/Nd, Lu/Hf, <sup>143</sup>Nd/<sup>144</sup>Nd, <sup>176</sup>Hf/<sup>177</sup>Hf ratios progressively decreases. 403 404

# 405 **5. Conclusions**

1. New Hf isotopic data of basaltic glasses from seamounts flanking the EPR between
5° and 15°N demonstrate that Hf and Nd isotope of these lavas are correlated and form a
well-defined trend that parallel to the mantle array.

2. The correlated Hf-Nd isotopic variations, together with their correlations with Sr-Pb isotope as well as with abundances and ratios of incompatible elements, suggest that the geochemistry of EPR and near-EPR seamount lavas is most consistent with meltinginduced mixing of a two-component mantle (i.e., an easily-melted enriched component dispersed in the more depleted refractory matrix).

414 3. Fractional melting modeling of a heterogeneous mantle can reproduce the
415 variations in trace-element and isotope compositions of seamount lavas. "Pyroxenite"

and/or EM1 endmember better fits the data, indicating that the enriched component is most
likely to be recycled oceanic crust or metasomatized oceanic lithospheric mantle.

- 418 4. The contribution of residual continental lithospheric mantle may interpret the 419 difference in  $\Delta \epsilon_{\rm Hf}$  values between Pacific, Atlantic and Indian MORBs.
- 5. The Rb-Sr, Sm-Nd and Lu-Hf pseudochrons have no age significance, but are
  mixing lines constrained by the composition of the enriched component and the depleted
  end-member.
- 423

# 424 Acknowledgements

425 Y.H. Yang and J.Y. Chen are thanked for their help during the chemical separation 426 and analyses. We thank C. Chauvel, P. Mueller, J. Blichert-Toft, C.L. Waters, A. Stracke and V.G.M. Salters for their constructive suggestions for this paper. A. Stracke and J. 427 Phipps Morgan are appreciated for help in quantitatively modeling. This study is supported 428 by the National Natural Science Foundation of China (Grants 41273013, 41130314, 429 91014003), China Postdoctoral Science Foundation, Specialized Research Fund for the 430 431 Doctoral Program of Higher Education of China (2012021110021), Scientific Research Foundation of Shandong University of Science and Technology for Recruited Talents 432 (2016RCJJ008). 433

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# 734 **Figure captions**

Fig. 1 (a) Tectonic framework of the northern (5°-15°N) EPR and vicinity; (b) Simplified
map of the study area showing the locations of near-ridge seamounts. The size of the circles
(sample locations) are not to scale (Niu and Batiza, 1997; Niu et al., 2002).

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Fig. 2 Hf and Nd isotopic variations of seamount lavas on the flanks of the EPR between 739 740 5° and 15°N. Nd isotope data are from Niu et al. (2002). The oceanic basalts field, defined by MORBs, OIBs and island arc volcanic rocks (IAVs), is from Chauvel et al. (2008) and 741 742 Vervoort et al. (1999). The mantle array is according to Chauvel et al. (2008). The field of abyssal peridotite and marine sediments is from Stracke et al. (2011), Vervoort et al. (2011), 743 respectively. The Pacific, Atlantic and Indian MORB data are from Salters et al. (2011), 744 745 Sims et al. (2002), Blichert-Toft et al. (2005), and Janney et al. (2005). The blue and green circles are alkali basalts and HIMU-like samples, respectively. All the samples give a linear 746 expression of  $\varepsilon_{Hf}=1.72\varepsilon_{Nd}-2.83$ . 747

748

Fig. 3 Diagram of correlations between Hf isotopic data in this study and the abundances
of major and trace elements. Major and trace elements are from Niu and Batiza (1997).

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Fig. 4 Correlations of Hf isotopes with Sr-Nd-Pb isotopes and with ratios of incompatible elements for near-EPR seamount lavas. The correlations are best interpreted as resulting from melting-induced mixing of a two-component mantle with the enriched component dispersed as physically distinct domains in the more depleted matrix. The correlations also indicate that both the enriched component and the depleted matrix are ancient and have developed their isotopic signatures independently. The Sr-Nd-Pb isotope data are from Niu
et al. (2002) and the trace-elements are from Niu and Batiza (1997). Data for ridge axis
MORB from EPR 5-15°N are from Sims et al. (2002), Salters et al. (2011), and Gale et al.
(2013).

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Fig. 5 Variation of (a) Yb/Nd with La/Nd and (b) La/Yb with Sm/Yb for lavas from 762 seamounts flanking the East Pacific Rise (EPR) between 5° and 15°N. Curves show range 763 764 in pooled melt compositions produced by variable degrees of fractional melting of a twocomponent mantle source consisting of 10% enriched component and 90% depleted matrix, 765 766 calculated using the method of Stracke and Bourdon (2009). The mass fractions of enriched endmembers are labelled. The enriched component is assumed to have a lower solidus 767 temperature and therefore contributes more to melting at low melt fractions, compared to 768 769 the more refractory matrix, which begins melting at a slightly lower pressure. Thus, unlike 770 two-component mixing of melts or sources that result in linear arrays on these diagrams, the resulting melt evolution paths are curved because with increasing degree of melting. 771 the contribution to the pooled melt from the more fertile component progressively 772 decreases, which reproduce the data array defined by lavas from seamounts in this study. 773 Detailed parameters used in modeling are given in Appendix 1. Data for northern EPR 774 775 axial MORBs are from Salters et al. (2011), Donnelly (2002), Chauvel and Blichert-Toft 776 (2001) and references therein. "Pyroxenite" endmember better fits the seamount data. 777

Fig. 6 Variation of (a) La/Sm vs. <sup>143</sup>Nd/<sup>144</sup>Nd and (b) Nb/Zr vs. <sup>87</sup>Sr/<sup>86</sup>Sr for lavas from
seamounts flanking the East Pacific Rise (EPR) between 5° and 15°N. The symbols are the

780 same as in Fig. 5. The different trends reflect mixtures of melts from peridotite matrix and 781 enriched component with different solidus temperatures. Peridotite melting is assumed to start at about 2.5 GPa in the spinel stability field. The dashed lines and solid lines represent 782 the onset of enriched component melting from depths significantly deeper (i.e., great 783 difference in solidus temperature between the two components) and depths similar to the 784 onset of peridotite melting (i.e., small difference in solidus temperature), respectively. The 785 786 compositions of depleted mantle is from Workman and Hart (2005). The pyroxenite 787 represents a 2 Ga old recycled MORB that has undergone chemical modification during subduction and is isotopically similar to FOZO (Stracke and Bourdon, 2009). The isotopic 788 789 composition of EM1 and EM2 is from Ito and Mahoney (2005). The large variation in La/Sm and Nb/Zr ratios and their good correlations with isotopes indicating a small 790 difference in solidus temperature between the two components. "EM1" endmember seems 791 792 to be better reproduce the array.

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Fig. 7 Isotope ratios of lavas from seamounts flanking the East Pacific Rise (EPR) between 794 5° and 15°N. The initial depleted:enriched component ratio Is assumed to be 90:10. Large 795 796 open circles show assumed isotope ratios of each component. The dashed curves in figure (b) and (c) using a mixture of 0.06:0.04 EM1:pyroxenite as the enriched component. 797 798 Px=pyroxenite,  $\phi_m$ = initial proportion of EM1/EM2, and initial proportion of pyroxenite=0.1-  $\phi_{m}$ . Data for ridge axis MORB from EPR 5-15°N are from Sims et al. 799 800 (2002), Salters et al. (2011), and Gale et al. (2013). The EM1 and/or pyroxenite as the enriched component can reproduce the seamount arrays. 801

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803 Fig. 8 (a) EHF versus ENd and (b)  $\Delta$ EHF versus ENd diagram of global MORBs. In general, Atlantic MORBs and Pacific MORBs have the highest and the lowest  $\varepsilon_{Hf}$  value at given 804 805 Nd isotopes, respectively, and Indian MORBs lies between them. In order to show more clearly, the average values are also plotted. The mantle array is  $\varepsilon_{Hf}=1.59\varepsilon_{Nd}+1.28$  (Chauvel 806 et al., 2008). The published data is from Agranier et al. (2005), Blichert-Toft et al. (2005), 807 808 Chauvel and Blichert-Toft (2001) and refs therein, Debaille et al. (2006), Dosso et al. 809 (1993), Douglass and Schilling (1999), Fontignie and Schilling (1996), Graham et al. 810 (2006), Hamelin et al. (2011), Hanan et al. (2004), Janney et al. (2005), Kempton et al. 811 (2000, 2002), Klein et al. (1988), Mahoney et al. (2002), Meyzen et al. (2007), Nowell et al. (1998), Patchett (1983), Patchett and Tatsumoto (1980), Pyle et al. (1992), Salters 812 (1996), Salters and White (1998), Salters et al. (2011) and refs therein, Schilling et al. (1994, 813 814 1999, 2003), Sims et al. (2002), Yu et al. (1997). The difference in  $\Delta \varepsilon_{\rm Hf}$  values between Pacific, Atlantic and Indian MORBs is likely result from various contribution of residual 815 continental lithospheric mantle with high  $\Delta \varepsilon_{Hf}$  values. Data for SCLM are from Bianchini 816 et al. (2007), Carlson et al. (2004), Chu et al. (2009), Lapen et al. (2005), Schmidberger et 817 al. (2002), Simon et al. (2007), and Witting et al. (2007). 818

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Fig. 9 The correlations between parent-daughter and isotopes ratios for Rb-Sr, Sm-Nd and Lu-Hf give pseudochron ages of 182±33 Ma, 276±50 Ma and 387±93Ma, respectively. These different "ages" have no age significance, but are best explained as resulting from melting-induced mixing with the pseudochron slopes controlled by the compositions of the enriched component and the depleted end-member. The two end-members have independently developed their distinct chemical and isotopic signatures. The symbols are

- the same as in Fig. 4. The fractional melting modeling can reproduce the arrays. The trace-
- 827 element and Sr-Nd isotope compositions are from Niu and Batiza (1997) and Niu et al.
- 828 (2002). Data for ridge axis MORB from EPR 5-15°N are from Sims et al. (2002), Salters
- et al. (2011), and Gale et al. (2013).
- 830
- 831 Fig. A1 The locations of EPR 5-15°N and Garrett Transform in the Pacific (According to
- 832 Amante and Eakins, 2009).

Sample	Туре	Latitude/°N	Longitude/°W	Depth/m	<sup>176</sup> Hf/ <sup>177</sup> Hf	2σ	ε <sub>Hf</sub>	
<b>R74-5</b>	Ν	10.62	103.84	2320.00	0.283174	$\pm 0.000012$	14.22	
<b>R7-13</b>	Ν	8.14	103.19	2020.00	0.283189	$\pm 0.000015$	14.76	
<b>R3-4</b>	Ε	5.78	102.21	1773.00	0.283212	$\pm 0.000013$	15.56	
<b>R3-1</b>	Ν	5.78	102.21	1773.00	0.283194	$\pm 0.000014$	14.93	
<b>R3-3</b>	Ν	5.78	102.21	1773.00	0.283201	$\pm 0.000017$	15.18	
<b>R1-14</b>	Ν	5.77	102.18	1834.00	0.283187	$\pm 0.000012$	14.66	
R66-1	Ν	10.14	103.34	2600.00	0.283189	$\pm 0.000013$	14.75	
<b>R28-8</b>	Ν	8.81	103.90	1984.00	0.283156	$\pm 0.000010$	13.57	
R103-13	Ν	13.84	103.80	2870.00	0.283201	$\pm 0.000010$	15.16	
<b>R22-1</b>	Ν	8.90	104.10	2749.00	0.283161	$\pm 0.000011$	13.77	
R25-1	Ν	8.88	103.79	1980.00	0.283158	$\pm 0.000013$	13.64	
R65-1	Ν	10.13	103.41	2074.00	0.283186	$\pm 0.000013$	14.63	
<b>R8-8</b>	Ν	8.34	103.06	3180.00	0.283153	$\pm 0.000013$	13.49	
R19-4	Ν	8.94	104.41	2267.00	0.283158	$\pm 0.000013$	13.65	
R60-1	Ν	10.00	104.91	2640.00	0.283213	$\pm 0.000014$	15.61	
<b>R71-2</b>	Ν	10.26	103.74	3380.00	0.283173	$\pm \ 0.000006$	14.18	
R21-6	Ε	8.89	104.14	2657.00	0.283152	$\pm 0.000011$	13.46	
R16-2	Ν	8.84	104.57	2985.00	0.283155	$\pm 0.000013$	13.56	
R62-4	Ε	10.03	104.19	2320.00	0.283167	$\pm 0.000015$	13.95	
R24-3	Ν	8.97	103.87	3054.00	0.283142	$\pm \ 0.00008$	13.10	
R102-1	Ε	13.22	102.68	2350.00	0.283178	$\pm 0.000013$	14.37	
<b>R96-24</b>	Ε	13.07	103.45	2577.00	0.283209	$\pm 0.000011$	15.46	
<b>R72-2</b>	Ε	10.38	103.93	2748.00	0.283120	$\pm 0.000010$	12.29	
<b>R4-2</b>	Ε	5.60	103.02	2263.00	0.283110	$\pm 0.000009$	11.94	
<b>R73-1</b>	Ε	10.38	103.92	2547.00	0.283124	$\pm 0.000010$	12.43	
R32-1	Ε	9.09	104.92	3025.00	0.283084	$\pm 0.000012$	11.03	
R110-4	Ε	14.14	104.36	2760.00	0.283136	$\pm 0.000009$	12.86	
R18-3	Ε	8.93	104.46	2720.00	0.283061	$\pm 0.000011$	10.21	
<b>R80-6</b>	Ε	11.80	103.25	1619.00	0.283144	$\pm \ 0.00008$	13.14	
R17-1&2	Ε	8.91	104.57	2715.00	0.283027	$\pm \ 0.000007$	9.01	
R109-5	Ε	14.15	104.30	2610.00	0.283110	$\pm 0.000011$	11.96	
<b>R83-2</b>	Ε	11.24	103.59	2900.00	0.283084	$\pm 0.00008$	11.04	
<b>R79-2</b>	Ε	11.79	103.25	1620.00	0.283047	$\pm 0.000011$	9.72	
<b>R78-5</b>	Ε	11.22	103.58	2450.00	0.283043	$\pm 0.000009$	9.58	
R13-1	Alkali	8.40	104.07	2140.00	0.282974	$\pm \ 0.000007$	7.15	
R15-1	Alkali	8.76	104.54	1682.00	0.282966	$\pm 0.000008$	6.86	

Table 1 Hf isotopic data of glass samples from near-ridge seamounts between 5 and 15°N EPR





Fig. 2















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Fig. 9



# Fig. A1

