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# Highlights

- Equation of state (EOS) modelling is a powerful tool to estimate mineral properties at conditions not accessed by high pressure and temperature experiments.
- Experimental errors, both random and systematic (e.g. pressure scale, functional forms), data consistency and sparsity all contribute to the uncertainties in mineral seismic properties.
- Conventional explicit EOSs which are assumed to follo <sup>7</sup> certain form provide a priori information by fixing their functional form or pressure scale, thereby providing a biased estimate of uncertainties.
- Neural networks based approach can implicitly cal ture full uncertainties together with highlighting data gaps and identifying data inconsistencies.



# Inferring material properties of the lower mantle minerals using Mixture Density Networks

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#### Abstract

Interpretation of information available from seismic  $\dot{a}$  a in terms of temperature and composition requires an understanding of the physical properties of minerals, in particular, the elastic properties of candidate Earth minerals at the relevant (here, lower mantle) pressure and temperature. A common practise for the bulk elastic properties is to measure volume at a range of pressures and temperature using experiments or computational methods. These datasets are then typically fit to a pre-determined functional form, or equation of state to allow computation of elastic properties at any other pressure or temperature. However, errors, both random and systematic, limitations in the number of data and choice of pressure marker and scale, as where different functional forms of equations of state, all contribute to the uncertainties in mineral seismic properties. In an attempt to present a more comprehensive view of these uncertainties, we use neural-network based techniques to infer the relationship among: pressure, temperature, volume, bulk modulus, and thermal expansivity of MgO. We illustrate our approach on experimental data, but an extension to ab initio data is straightforward. The type of neural network used is called a Mixture Density Network (MDN) which is a combination of a conventional feed-forward neural network and a mixture

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model that consists of Gaussian functions. MDNs are capable of approximating arbitrary probability density functions, which allows us to compute the uncertainties in the predicted equations of state. Since the networks interpolate locally between input samples, pressurevolume-temperature relations are implicitly learned from data without imposing any explicit thermodynamic assumptions or ad-hoc relationships. We use the partial derivatives of the mapping between inputs (pressure and temperature) and output (volume) to compute the isothermal bulk modulus and thermal expansivity. Flexibility of the MDNs allows us to investigate the uncertainty due to certain data in one region of pressure-temperature space without influencing the posterior probability density everywhere. In general, we find that the elastic properties of MgO are well-constrained by experime. \* data. However, our study highlights regions in which sparse or inconsistent data hand to poorly constrained elastic properties, namely: at low pressure and high temperature (<25GPa and >1500 K), and temperatures above 2700 K. While the former conditions are likely not important for the Earth's lower mantle, they are relevant in the planetary bodies such as the Moon and Mars. Comparison with conventional  $\epsilon_{\rm cu}$  tion of state forms shows that assuming a certain functional form of the pressure-volume-temperature relationship leads to potential bias in uncertainty quantification, because the uncertainties are then specific to the underlying form. In combination with data orts of other lower mantle minerals, this technique should improve uncertainty quantification in interpretations of seismic data.

Keywords: equations of state; lower mantle; neural networks; periclase; MgO

#### 1 1. Introduction

Information such as variation of wave speeds (e.g. Dziewonski and Anderson 1981, Kennett
et al. 1995), obtained by studying seismic data is crucial for understanding the internal
structure of the Earth. Various studies have reported the presence of seismically distinct
structures at multiple scales in the Earth's mantle (e.g. Garnero and Helmberger 1998,
Ritsema et al. 1999, Romanowicz 2008, Hernlund and Houser 2008, Deschamps et al. 2012,

Garnero et al. 2016). In order to relate those observed seismic structures to appropriate 7 temperature and composition, constraints from mineral physics on the sensitivity of seismic 8 wave speeds to these parameters are required (e.g. Jackson 1998, Trampert et al. 2001). 9 The sensitivities have been used to infer the probable existence of chemical heterogeneities 10 within the mantle (e.g. Trampert et al. 2004, Dobrosavljevic et al. 2019, Jackson and Thomas 11 2021). Other studies have tried to constrain the (average) mantle geotherm and composition 12 by combining seismic data and mineral seismic properties (e.g. Cammarano et al. 2003, 13 2005a,b, Deschamps and Trampert 2004, Stixrude and Lithgow-Bertelloni 2005, Matas et al. 14 2007, Cobden et al. 2008, 2009, Simmons et al. 2010, Khan et al. 2009, 2011, 2013). Mantle 15 convection simulations (e.g. Nakagawa et al. 2009, 2010, 2012) Schuberth et al. 2009, 2012) 16 have also incorporated mineral properties to illustrate *+* 1e ....portance of joint geodynamical-17 mineralogical approaches to explain the seismic and values in the mantle. Mineral properties 18 can be derived from experimental or theoretical methods. In particular, information on the 19 density (or volume V), incompressibility and ignate are required to obtain the seismic wave 20 speeds in a material. Since it is not practical or feasible yet to perform experiments at each 21 pressure (P) and temperature (T) that may exist within the Earth, the convention is to use 22 equations of state (EOSs) to define the relationship among the thermodynamic variables P, 23 and T (e.g. Duffy and Wang 1998), and hence be able to estimate mineral properties at V 24 the conditions not accessed by `xperiments. 25

However, a number of the containties are associated with this procedure. Experimental measurements contain random and systematic errors. The choice of pressure scale as well as different functional forms of the EOS (e.g. Vinet EOS, third/fourth order finite strain equations, also called Birch-Murnaghan EOSs, as well as the choice of Grüneisen models) all contribute to the uncertainties in mineral seismic properties. As a result, it becomes challenging to determine realistic uncertainties for the interpretations which relate seismic observations to temperature and composition.

<sup>33</sup> In this study, we present an Artificial Neural Network (ANN) based approach to infer the

pressure-volume-temperature (P-V-T) relationship of MgO, with a view to extend the appli-34 cation to other major lower mantle minerals. We collate experimental P-V-T data for MgO 35 together with reported uncertainties, regardless of pressure scale or functional form used. 36 By applying ANN techniques, P-V-T relationships are implicitly learned from data without 37 any prior assumption on the functional form (or thermodynamic model) of the relationship. 38 Specifically, we use Mixture Density Networks to infer material properties and assess their 39 uncertainties. We compute the partial derivatives of inferred volume with respect to pres-40 sure and temperature to extract the bulk modulus and thermal expansivity, respectively. 41 In order to test the feasibility of this approach, we train the notworks only on experimen-42 tal data, although a combination of theoretical and experimental data is also possible and 43 straightforward. 44

#### 45 2. Equations of state: Uncertainties

Experimental approaches (e.g. Vassiliou er d Ahrens 1981, Yoneda 1990, Utsumi et al. 1998, 46 Duffy and Ahrens 1995, Fei 1999, Sinospikin and Bass 2000, Sinogeikin et al. 2000, Dewaele 47 et al. 2000, Speziale et al. 2001, Li et al. 2006, Dorogokupets and Dewaele 2007, Hirose et al. 48 2008, Murakami et al. 2009, Kon, et al. 2010, Dorfman et al. 2012, Ye et al. 2017) have 49 been used to establish the P-V-T relationship of MgO. Experiments using a diamond anvil 50 cell (DAC), a multi-anvil press (MAP) and shock compression have provided a huge number 51 of data covering a wide ringe of pressure and temperature. Laboratory measurements of 52 volume are done at a discrete set of pressure and temperature points. To cover the en-53 tire pressure and temperature range of lower mantle requires pressure extrapolation and/or 54 interpolation of the measurements using a thermal equation of state. The most common 55 procedure (e.g. Matas et al. 2007, Cobden et al. 2009) is to use an isothermal equation of 56 state with a Mie-Grüneisen model for thermal pressure. In this approach, the total pressure 57 is considered to be the sum of a static pressure and a quasiharmonic thermal pressure. The 58 static pressure term describes the pressure-volume relationship at a reference temperature 59

(usually 300 K). Different functional forms, such as third/fourth order finite strain and Vinet, 60 have been widely used to model isothermal compression curves often leading to different esti-61 mates of fitting parameters or ambient mineral properties such as volume  $(V_0)$ , bulk modulus 62  $(K_{0T})$  and pressure derivative of bulk modulus  $(K'_{0T})$  at 0 GPa pressure (e.g. Speziale et al. 63 2001, Dorogokupets and Dewaele 2007, Tange et al. 2009). To compute temperature effects 64 (more precisely, thermal pressure) this framework uses a Grüneisen parameter whose volume 65 dependence is uncertain (Ye et al. 2017). Although anharmonic effects are very small com-66 pared to the harmonic contribution to thermal pressure, some authors (e.g. Dorogokupets 67 and Dewaele 2007) use models to account for this term as well. 68

Additionally, the exact determination of pressure using a reliable pressure scale in static high 69 pressure and temperature experiments is still a challenging task. The ruby pressure scale of 70 Forman et al. 1972 used in DAC experiments has bee. largely calibrated (Liu and Bi 2016) 71 using both static and dynamic compression d. t., but still suffers from large experimental 72 uncertainties. Dynamic shock compression experiments provide an absolute pressure scale. 73 But the correction for thermal effects can be very uncertain (e.g. Dorfman et al. 2012, Duffy 74 and Wang 1998), especially at high sinck temperatures because the corresponding thermal 75 contribution also increases. Other vial used pressure scales are gold, platinum and MgO 76 scales. A recent study by Ye et al. 2017 shows the inter-comparison of those scales up to 140 77 GPa and 2500 K. They report 2 1 to 4 GPa (sometimes systematic) differences in pressure 78 among those pressure real. Although their study optimized different Au, Pt and MgO 79 pressure scales to make onem agree within  $\pm 1$  GPa, it concludes that the most preferred 80 form of EOS (and the pressure standard itself) remains uncertain. 81

Measurement errors, lack of an absolute pressure scale, and a variety of functional forms of EOSs all contribute to the uncertainties in mineral seismic properties. Assuming one particular EOS or pressure scale has the potential to produce biased uncertainty estimates that are specific to the underlying functional form. In this study we train neural networks to learn the implicit relation between pressure and temperature (as inputs) and volume,

<sup>87</sup> bulk modulus and thermal expansivity (as outputs). The results are entirely data-driven <sup>88</sup> without a priori selection of experiments or a functional form to explain the data. In this <sup>89</sup> way, we can infer the relative contributions of data sparsity versus prior conditioning to the <sup>90</sup> uncertainties. We can also map the level of certainty of the elastic parameters in pressure-<sup>91</sup> temperature space, which can be propagated into seismic interpretation.

#### 92 3. The Mixture Density Network (MDN)

#### 93 3.1. Background

Conventional neural networks (Hornik et al. 1989) are general function approximators, which 94 can be used to infer an (arbitrary nonlinear) relationship (Cybenko 1989) between inputs 95 and targets/outputs. However, the conditional everage (i.e. the mean value of output 96 conditioned on input data) given by such net ve ks only provides limited information about 97 that relationship (Bishop 1994). Since experimental P-V-T data contain measurement errors, 98 and inferring P-V-T relationship using those data is an inverse problem which can have 99 multiple solutions, naturally we seek to treat the problem in a probabilistic framework. 100 Hence, instead of having only the average volume output, we want to find the posterior 101 probability density function (pdi) for volume. The pdf for volume at a given pressure and 102 temperature can be denoted as 103

104

$$\sigma(V|P,T).\tag{1}$$

We can represent a general pdf by combining a conventional feed-forward neural network with a Gaussian Mixture Model (GMM), which is then called a Mixture Density Network (MDN) (Bishop 1994 and Bishop 1995). The architecture of the MDN used in this study is shown in Figure 1, and consists of a two layer feed-forward neural network and a GMM. The GMM contains a mixture of a finite number of Gaussian kernels which are then weighted to give the posterior pdf. The mean, standard deviation and weight of each Gaussian kernel are parameterized by weights and biases of the feed-forward neural network, also known as

<sup>112</sup> network parameters ( $\alpha$ ).

Application of MDNs in Earth Sciences ranges from inversion of surface wave data for global 113 crustal thickness (Meier et al. 2007a,b), temperature and water content variations within 114 the transition zone (Meier et al. 2009), inference of Earth's radial seismic structure (de Wit 115 et al. 2013), inversion of free oscillations (de Wit et al. 2014), constraints on lower mantle 116 anisotropy (de Wit and Trampert 2015), nonlinear petrophysical inversion (Shahraeeni and 117 Curtis 2011), source inversion of strong-motion data (Käufl et al. 2016b), inferring parame-118 ters governing mantle convection (Atkins et al. 2016) to travel time tomography (Earp and 119 Curtis 2020). In our case, based on some experimental P-V  $\Gamma$  Gata, we seek to approximate 120 the true posterior pdf (Equation 1) by a parameterized post rior 121

$$p(V|P,T;\boldsymbol{\alpha}) \approx \sigma(V_{\perp}^{D},\Gamma).$$
(2)

In other words, for a given pressure and ten, erature, the posterior probability density for volume is given by the pdf in expression 2 which is parameterized by the weights and biases ( $\alpha$ ) of the feed-forward neural network. These parameters are learned during the network training process (see Sub-section 3.2). The posterior pdf (Equation 2) can be expressed as a linear combination of a fixed number of Gaussian kernels (also see Figure 1) as

$$p V|_{\boldsymbol{\mathcal{I}}}, T; \boldsymbol{\alpha}) = \sum_{n=1}^{M} \pi_n(P, T; \boldsymbol{\alpha}) \phi_n(V|P, T; \boldsymbol{\alpha})$$
(3)

<sup>129</sup> where M denotes the number of kernels used, and  $\pi_n$  are mixing coefficients which satisfy

130 
$$\sum_{n=1}^{M} \pi_n(P, T; \alpha) = 1.$$
 (4)

If the number of Gaussian kernels is M, then the total number of outputs from the feedforward network is K= 3M because each kernel is parameterized by its weight  $(\pi_n)$ , mean  $(\mu_n)$  and standard deviation  $(\sigma_n)$ . Equation 4 ensures that the posterior integrates to 1



two layer feed-forward neural network

135

mixture model posterior pdf

Figure 1: Architecture of the Mixture Density Network (MDN). A two layer feed-forward neural network (left) is combined with a GMM (centre) to get the posterior pdf (right). P & T denote the network inputs,  $h_j$  are the hidden nodes, and  $y_k$  are the outputs of  $\sum i$ -forward network. Indices J and K represent the number of hidden and output nodes, respectively. Except for the input nodes, each circle represents a computational node. Hidden layer nodes take a weighted sum. (with weights  $\alpha_{ij}$ , where  $i \neq 0$ ) of input data (P & T) plus a bias term ( $\alpha_{0j}$ ) as inputs and apply a sign of activation function. The output layer nodes take a weighted sum (weighted by  $\alpha_{jk}$ , where  $j \neq 0$ ) of the output's  $y_k$ . These outputs are related to the mean, standard deviation and weight of each Gaussian in the CM.<sup>4</sup> (see Appendix A for details). Each Gaussian in the GMM is then weighted to give the final posterior  $\mathbf{p}_i^{\mathsf{re}}$ 

making it a valid probability density.  $\phi_n$  in equation 3 are Gaussian kernels of the form

$$\phi_n(V|P,T;\boldsymbol{\alpha}) = \frac{1}{\sqrt{2\pi}\sigma_n(P,T;\boldsymbol{\alpha})} exp\left\{-\frac{(V-\mu_n(P,T;\boldsymbol{\alpha}))^2}{2\sigma_n(P,T;\boldsymbol{\alpha})^2}\right\}$$
(5)

where  $\mu_n$  and  $\sigma_n$  are the mean and standard deviation of Gaussian kernels in the GMM. These parameters of the GMM are related to the outputs  $(y_k)$  of the feed-forward network (see details in Appendix A).

#### 139 3.2. MDN initialization and training

In order to find the appropriate weights and biases of the feed-forward neural network, 140 we train the MDN using a sub-set of the experimental P-V-T data. In fact, the total 141 experimental P-V-T data, shown in Figure 2 (Fei 1999, Jacobsen et al. 2008, Fei et al. 142 2004a, Fei et al. 2004b, Dewaele et al. 2000, Speziale et al. 2001, Utsumi et al. 1998, Figuet 143 et al. 1999, Ye et al. 2017, Kono et al. 2010, Dorfman et al. 2012, Zhang 2000, Fiquet et al. 144 1996, Dubrovinsky and Saxena 1997, Hirose et al. 2008, Litasov et al. 2005, Murakami et al. 145 2012, Sinogeikin and Bass 2000, Li et al. 2006 and Fan et U 2019), is divided into three 146 sets: training (70%), monitoring (20%) and test (10%) sets. Duing training, the MDN takes 147 pressure and temperature from the training data and cutputs a pdf for volume according 148 to Equation 3. However, we need to decide on the initial values of the network parameters 149 of the feed-forward neural network to compute the list output. We randomly draw the 150 input layer and hidden layer weights (Bishop 1.95) according to Gaussian distributions (see 151 Appendix B for details). Once the MDN is 1 littlelized and training has started, the difference 152 between the output and the target can be computed according to an error function defined 153 in Appendix B. This function is also select the loss function which is minimized iteratively 154 using the ADAM optimization method (see detailed algorithm in Kingma and Ba 2014). We 155 use TensorFlow (1.13.1) (Abad. et al. 2015) to construct, train and evaluate the MDN. 156

Overfitting is a general property of the maximum likelihood technique (Bishop 1995). We use a separate monitoring data set to monitor the error decay during training. We evaluate the monitoring set error at the end of each iteration; if the monitoring error starts to increase (i.e. the network starts to over-fit the training data) then we stop the training procedure and save the last best trained model. This technique is also called the early-stopping technique.

It is known that the inverse problem can have multiple solutions (i.e. a range of network parameters can possibly provide equally likely solutions). We train a number of independent MDNs, and combine them by a weighted sum (e.g. Käufl et al. 2016a). The weight of each network is based on how well it performs on the test data which is not used during training.



Figure 2: Experimental P-V-T data for MgO used in bis study (Fei 1999, Jacobsen et al. 2008, Fei et al. 2004a, Fei et al. 2004b, Dewaele et al. 2000, Speziale et al. 2001, Utsumi et al. 1998, Fiquet et al. 1999, Ye et al. 2017, Kono et al. 2010, Dorfman et al. 2012, Thang 2000, Fiquet et al. 1996, Dubrovinsky and Saxena 1997, Hirose et al. 2008, Litasov et al. 2005, Maral ami et al. 2012, Sinogeikin and Bass 2000, Li et al. 2006 and Fan et al. 2019) to train the MDNs. Data with uncertainties from X-ray diffraction experiments (in static high P-T, Brillouin spectroscopy and Arab not plotted because the scale would be inappropriate to visualize them.

The performance is measured by the same error function that we use to calculate training 166 and monitoring errors (for details see Appendix B). In this way, the explicit dependence of 167 the posterior on the network parameters can be avoided. The choice of the number of MDNs 168 depends on the problem t hand. A rough estimate for a relatively simple problem (e.g. a 169 few inputs and a target/output) may lie in the range 10-20 (Käufl et al. 2016a). However, 170 in order to compute the uncertainties in bulk modulus and thermal expansivity (details in 171 Section 5) we train a large number of MDNs  $(10^3)$ . The number of hidden nodes to use in 172 each MDN are randomly selected from a pre-defined range which is 16-32. We conducted a 173 separate test (not shown here) to find the range that provides the lowest errors for the test 174 set. Similarly, we propagate the uncertainties in experimental data through the MDNs by 175 randomly perturbing the thermodynamic variables within the reported uncertainty range. 176

#### 177 3.3. Network performance

We use the test data set to examine how well the trained MDNs perform when a new datum is presented. Since the test data are not used in network training, we can use them to predict the output and subsequently compare with target data. In Figure 3 (top panel) the predicted volume is compared with the target data. The MDNs predict pdfs for volume, and for this comparison we compute the conditional mean volume (conditioned on inputs P & T), instead of using the full posterior pdfs on volume, as

$$< V|P,T; \boldsymbol{\alpha} > = \sum_{n=1}^{M} \pi_n(P,T;\boldsymbol{\alpha}) \mu_n(P,T;\boldsymbol{\alpha}).$$
(6)

This special case of MDN corresponds to the standard neural network output (Bishop 1994), i.e. only the feed-forward network with one volume catput. Equation 6 shows the mean volume output for one MDN, and we calculate the weighted sum (weights are chosen according to the test set error as mentioned previously) of mean volumes from all MDNs. One alternative to the conditional posterior mean could be the posterior mode. However, the posterior mode may be biased toward: contain pressure scales which contain relatively more data in the training set compared to other scales.

In the region of high temperatures and low pressures (Figure 3, top panel) the trained MDNs 192 show lower resolving cap condition, providing more uncertain volume predictions. We found that 193 this discrepancy in network predictions comes from the inclusion of specific training data 194 points (high temperature data of Figuet et al. 1996) in those ranges. We note that Figuet 195 et al. 1996 did not include a thermal pressure term in their experiments and so it is likely 196 that the total pressure is underestimated. Moreover, the reported temperatures are likely 197 overestimated by about 20 to 50%. We trained another network excluding these data in our 198 training set and access the prediction performance (Figure 3, bottom panel). In doing so, 199 MgO volumes are resolved within the prior range of experimental data, also in the region 200 of low pressure and high temperature. This shows the networks' ability to capture the 203



Figure 3: Performance of MDNs. Target volu.  $\infty$  from the test data set are compared with mean volumes (Equation 6) predicted by the MDNs. Top panel shows mean volumes predicted by the MDNs trained with all experimental data while bottom shows insults with high temperature data of Fiquet et al. 1996 and Murakami et al. 2012 excluded (also see Srippection 4.2). The pressure (left) and temperature (right) range of the test data set is shown by colourbars or both panels. We note that the solid red line in the Figure refers to a perfectly resolved network prediction. Foints located near this line are well resolved and those located away represent more uncertain volume  $e_1$  redictions. The MDNs best predict the volumes in low temperature regions and at simultaneous high to preature and pressure. However, including high temperature data of Fiquet et al. 1996 into training provides more uncertain volume predictions in the low pressure, high temperature region. For two data, points marked with "+" in both left and right plots in the top panel, we plot posterior pdfs for volume in Figure 4. One datum is located in the low pressure, high temperature region where the effect of high imperature data from Figure et al. 1996 is significant and another away from it.

<sup>202</sup> underlying data consistency.

Low pressure data (approximately less than 30 GPa) are relatively dense up to about 1400 K compared to higher temperatures. Similarly, most of the high pressure data, i.e. extending to the lower mantle environment, come either from approximately between 1500 K to 2700 K or from ambient temperature measurements. Besides that, the experimental data doesn't cover simultaneous high temperature and high pressure regions, for example temperatures greater

than  $\sim 2700$  K at pressures expected near the bottom of lower mantle. Hence, we expect wider posterior probability density functions for volume in regions of sparse experimental data coverage.

So far we have only shown the mean of the posterior pdf for volume. To illustrate more 211 clearly the effect of the high temperature data of Figuet et al. 1996 on the posterior pdf 212 at low pressure, high temperature, we take two data points from the test set (denoted by 213 '+' in Figure 3, top panel). Both points are drawn at low pressures, but one is at high 214 temperature and located away from the solid line and another at low temperature is close 215 to it. In Figure 4 posterior pdfs at those points are show. They show a more uncertain 216 prediction for the high temperature, low pressure input. Once we remove Figuet et al. 217 1996 data from training (see Sub-section 4.2), the network predicts narrow posterior pdfs 218 showing less uncertainty (cf. including those in training) in volume. Although excluding 219 Fiquet et al. 1996 provides less uncertain volu per redictions, due to limited availability of 220 experimental data at high temperature  $\epsilon_{10}$  ow pressure (approximately >1500 K and < 25 22 GPa) the predicted posterior pdfs are still lightly wider than at similar temperatures and 222 high pressures (also see Sub-section ±.? and Appendix C.1). 223



Figure 4: Posterior pdfs for MgO volume (solid curves) for two data points from top panel of Figure 3 together with their target values (red dashed line) and conditional mean volume (black dashed line). Left: inputs are 24.86 GPa and 300.19 K. The posterior pdf is narrow and uni-modal with the posterior mode located close to the target value. Right: inputs are 1.36 GPa and 2116.03 K. The posterior pdf is broad and multi-modal with target volume located away from the posterior modes. The smaller peak is the due to experimental P-V-T data of Figure et al. 1996.

#### 224 4. MDN predicted material properties

#### 225 4.1. P-V relationship at 300 K

The predicted pdfs for volume along a 300 K isotherm are presented in Figure 5. A subset of the training data (i.e. only around 300 K temperature) is also shown along with the MDN predictions. The uncertainty in volume increases with pressure as shown by the increasing width of pdfs. This is expected as the training data (around 300 K) are more consistent with each other at lower pressures.



Figure 5: The predicted pdf by the MDNs for volume of MgO along a 300 K isotherm. Left: pdf for volume up to lower mantle pressures is shown as a continuous function of pressure. The colour scale shows the value of the probability density function. Right: pdfs on volume are shown at 5 GPa pressure intervals together with training data around 300 K (shoorn as circles in the background). The training data show less variation at low pressures which results in nanower pdfs compared to high pressures.

In Figure 6 we compare has for the volume of MgO along a 300 K isotherm with EOSs 231 of Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2005, 2011 and 232 Dorogokupets and Dewaele 2007 (denoted as T09, S01, SLB0511 and DD07, respectively). In 233 this study, we use MINUTI (Sturhahn 2020) to compute volume, bulk modulus and thermal 234 expansivity as a function of pressure (and temperature) from these EOSs. For ambient 235 temperature comparisons, static equations (i.e. third-order finite strain or Vinet) together 236 with respective fitting parameters  $(V_0, K_{0T} \text{ and } K'_{0T})$  as reported in the literature are used. 237 We show the pdfs for volume (Figure 6, left panel) at every 5 GPa. The EOSs diverge as 238 the pressure increases. At 135 GPa, the difference in volume between the equations of state 239

of Stixrude and Lithgow-Bertelloni 2005, 2011 and Tange et al. 2009 is  $\sim 0.68 \text{ Å}^3$ , whereas 240 one standard deviation predicted by the neural networks is  $\pm 0.54 \text{ Å}^3$ . Moreover, the slope of 241 each individual EOS differs. This can best be visualized by computing  $\frac{\partial P}{\partial V}$  for all EOSs (see 242 Figure 6, right panel). Although Speziale et al. 2001 and Stixrude and Lithgow-Bertelloni 243 2005, 2011 are based on third order Birch-Murnaghan EOSs, their fitting parameters are 244 different. Comparisons between different EOSs and their fitting parameters are given by 245 other studies (e.g. Dorogokupets and Dewaele 2007, Tange et al. 2009, Ye et al. 2017, etc.). 246 The mean slope predicted by the neural network shows a slightly stiffer EOS compared to 247 the "standard" EOSs from the literature. This may be due to the fact that our training data 248 include experiments which make use of different pressure surgards (e.g. Ruby, NaCl, Pt, 249 Au) than the EOSs considered for comparison (which  $z \in b$  and on MgO). Nevertheless, such 250 a difference in slope together with the volume difference will inevitably lead to a significant 251 divergence in the inferred compressibility and thermal expansivity (see Section 5). 252



Figure 6: Left: our predicted pdfs for volume of MgO along a 300 K isotherm (black lines) compared with previously published EOSs (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2005, 2011 and Dorogokupets and Dewaele 2007) (coloured lines). Pdfs for volume are shown at 5 GPa pressure intervals. Right:  $\frac{\partial P}{\partial V}$  of MgO EOSs from the left panel. For this computation, we take the mean (Equation 6) of the output posterior on volume at every 0.1 GPa interval. The divergence between different EOSs increases with pressure.

#### 253 4.2. High temperature P-V-T relationships

We use the trained MDNs to predict volumes of MgO at different temperatures. As an 254 example, we plot the predicted pdfs for volume along a 2500 K isotherm in Figure 7, left panel 255 (other isotherms are provided in Appendix C.1). Similar to the ambient temperature (Sub-256 section 4.1), the 2500 K isotherm shows a well-constrained volume prediction at lower mantle 257 pressures. However, the high temperature pdfs show more uncertain volume predictions at 258 low pressures (except at 0 GPa). For example, at 5 GPa the pdf is relatively wide and 259 bimodal compared to that at high pressures (e.g. 100 GPa) which is unimodal. As discussed 260 earlier in Section 3.3, high temperature experimental data of Fig. et et al. 1996 do not include 261 a thermal pressure term, and it is likely the total pressure is underestimated. This can be 262 visualised in Figure 7, left panel, where training data points located approximately between 263 5-15 GPa have a smaller volume compared to data around 20 GPa and  $\sim$ 2500 K. We train 264 another network without the high temperature <sup>4</sup> at  $\varepsilon$  of Figuet et al. 1996 and plot the results 265 on the right panel of Figure 7. The posterior pdf for volume at 5 GPa now shows a unimodal 266 peak and the width is decreased by approximately a factor of 2 (cf. left panel at 5 GPa). 267 Although removing Figuet et al. 1966, duces the uncertainties in volume, the posterior pdf 268 is still wider than at high pressures for the same temperature. This region of low pressure, 269 high temperature is known to be dominated by anharmonic effects. Although these effects 270 are implicitly represented in the volume pdfs, there are limited experimental data in this 271 region (temperature  $>_150_0$  K and pressure <25 GPa) to further constrain them. 272

We compare the MDN predicted pdfs along a 2500 K isotherm (Figure 7) with some conventional EOSs (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2005, 2011 and Dorogokupets and Dewaele 2007). The variation in volume between these EOSs at high pressures is similar to that observed at 300 K. It has been noted in earlier studies (e.g. Ye et al. 2017) that the discrepancies in high temperature EOSs are partly due to persistence of the disagreement between them at 300 K (reference isotherm). Furthermore, at low pressure (<25 GPa) Speziale et al. 2001 diverges from other EOSs. This deviation is



Figure 7: Left: pdfs for volume of MgO along a 2500 K isotherm predicted by MDNs trained with all data. Right: same as left but Fiquet et al. 1996 and Murakami et al. 2012 c. (a are excluded. For comparison, volumes along the high temperature isotherm for some previously published L OSs (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2005, 2011 and Dorogokupets and Dewaele 2007) are computed using MINUTI (Sturhahn 2020). On both panels we plot a sub-set of the total training data, namely those data at temperatures between 2100 and 2600 K. Excluding Figure et al. 1996 data from neural network training significantly reduces the width of the pdfs at high temp vature and low pressure.

likely due to different values of fitting parameters to gether with distinct Grüneisen models to compute the thermal behavior. For example,  $S_1 \circ i$  is le et al. 2001 do not consider anharmonic effects, and their ambient Grüneisen parameters are also different than other studies (see e.g. Ye et al. 2017, Dorogokupets and Dewacle 2007). Besides that, as with the case of the



Figure 8: Left: pdfs for volume of MgO along a 2700 K isotherm predicted by the MDNs trained with all data. We also plot a sub-set of the training data, namely those whose temperatures lie between 2600 and 2800 K. Note: the large uncertainty in volume in the low pressure region (approximately below 25 GPa) is due to inclusion of data from Fiquet et al. 1996 as discussed in the text. Right: Comparison of posterior pdfs for volume predicted by MDNs trained with and without Murakami et al. 2012 (M12) and Fiquet et al. 1996 (F96) data at 2700 K and 60 GPa. The small peak at around 66  $Å^3$  is due to Murakami et al. 2012 data.

<sup>284</sup> 300 K isotherm, all explicit EOSs lie within the uncertainty range predicted by our MDNs,
<sup>285</sup> which is expected because some training data come from the MgO pressure scales described
<sup>286</sup> by these EOSs.

At 2700 K, the MDN predicted pdfs (Figure 8) show bimodal volumes in the pressure range 287 of approximately 45-90 GPa. Once we plot the associated training data on top, it becomes 288 clear that the smaller peaks in the pdfs are the representation of experimental data points 289 of Murakami et al. 2012. Surprisingly, for the same reported volume and temperature they 290 report pressures which are different from each other by alout 36 GPa. However, their 29 reported densities appear to be physically reasonable. Never new ss, we train another network 292 to discriminate how much uncertainty is coming from those specific data points. In doing 293 so, the posterior becomes unimodal. At 60 GPa, including Murakami et al. 2012 data leads 294 to a factor of approximately 3.5 wider pdfs for volum. (Figure 8, right panel) compared to 295 results without those data. However, the effect of those data points seems to be local in 296 P-V-T space and their influence decreases r recample, at higher pressures. This is because 29 MDNs interpolate locally in between samp. s, and data in one region of P-T space doesn't 298 influence uncertainties everywhere. 299

#### <sup>300</sup> 5. Bulk modulus and thermal expansivity

Since the training data do not contain explicit values for the volume derivatives with respect 301 to the inputs (P and T), getting constraints on bulk modulus  $(-V\frac{\partial P}{\partial V})$  and thermal expansivity 302  $\left(\frac{1}{V}\frac{\partial V}{\partial T}\right)$  is less straightforward than constraining the volumes. Hence, we follow a slightly 303 different approach compared to volume. We calculate the mean volume using Equation 6 for 304 any given P and T from each earlier obtained MDN. Then we perturb pressure  $(P+\delta P)$  while 305 keeping the temperature fixed and compute the mean volume ( $\langle V(P + \delta P, T) \rangle$ ) for that 306 pressure from the same MDN. This way, we can compute the mean isothermal bulk modulus 307 (K) as shown in Equation 7. Similarly, we evaluate mean volumes for two slightly different 308 temperatures but at a fixed pressure, and use that to compute the thermal expansivity,  $\alpha$ 309

(Equation 8). For numerical differentiation, we use  $\delta P = 0.1$  GPa and  $\delta T = 1$  K. Using a different value for  $\partial P$  or  $\partial T$  provides similar results.

$$\langle K|P,T; \boldsymbol{\alpha} \rangle = \langle -V(P,T) \rangle \frac{\delta P}{\langle V(P+\delta P,T) \rangle - \langle V(P,T) \rangle}$$
(7)

313

312

$$_{314} \qquad \qquad <\alpha|P,T;\boldsymbol{\alpha}> = \frac{1}{\langle V(P,T)\rangle} \frac{\langle V(P,T+\delta T)\rangle - \langle V(P,T)\rangle}{\delta T} \tag{8}$$

Hence, in this approach, we take the derivatives of the P-V (or T-V) curve defined by the 315 mean of the posterior pdfs from each neural network rath r than fitting P-V-T data to a 316 predefined EOS to get fitting parameters (such as  $K_{0T}$  and  $K'_{0T}$ ). Since we have trained a 317 large number of MDNs  $(10^3)$  to predict the posterior 21 for volume, we get the same number 318 of mean isothermal bulk modulus and thermal expansivity values. This way, each neural 319 network approximates a slightly different map ing and its derivatives, and the distribution 320 on the mean bulk modulus and therma' expansivity can approximate the uncertainties on 321 them. Moreover, we use the same networks to compute the pdfs for volume and the mean 322 volumes; the volume that goes into 'ne calculation of bulk modulus and thermal expansivity 323 is therefore consistent. 324

As an example, Figure 9 shows bulk modulus as a function of pressure along two selected 325 isotherms (refer to App ndi. C.2 for other isotherms). The bulk modulus predicted by 326 neural networks shows a jigher value at high pressure along the 300 K isotherm compared 327 to conventional EOSs. As mentioned earlier, this is likely due to the fact that the training 328 data come from experiments which make use of different EOSs and pressure standards than 329 those (MgO based) EOSs considered for comparison. Moreover, the fitting parameters  $(V_0,$ 330  $K_{0T}$  and  $K'_{0T}$ ) are different for different EOSs. Hence, although these EOSs predict volume 331 within the uncertainty range predicted by MDNs (Figure 6, left panel), their derivatives 332 (Figure 6, right panel) differ significantly from each other and also from the MDN prediction, 333 leading to different values of bulk modulus. 334



Figure 9: Comparison of the mean bulk modulus (a, b, c and d) and thermal expansivity (e and f) predicted by the neural networks with previously published equations of state for MgO (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lungow-Bertelloni 2011 and Dorogokupets and Dewaele 2007) as a function of pressure. The output from the neural networks is shown with greyscale- the darker the region of the plot, the greater the number of MDNs which predict the bulk modulus (or thermal expansivity) has that value. Frequency counts for output from the MDNs are at intervals of 1 GPa for pressure and bulk modulus, and  $10^{-7} \text{ K}^{-1}$  for thermal expansivity. For (a), (c) and (e) neural networks are trained with all collected data, whereas for (b), (d) and (f) data from Fiquet et al. 1996 and Murakami et al. 2012 have been excluded. Due to the inclusion of Fiquet et al. 1996 data we obtain large uncertainties in bulk modulus and thermal expansivity in low pressure, high temperature regions. Note: the overlapping of different EOSs makes the background histogram difficult to visualise.

One high temperature (2000 K) comparison between the neural network predicted mineral properties and other studies is shown in Figure 9- c, d, e and f. In general, bulk modulus



Figure 10: Comparison of the MDN predict d mean bulk modulus at (a) 2700 K, 60 GPa, (b) 2700 K, 135 GPa, (c) 300 K, 135 GPa and (d) thermal expansivity at 2700 K, 135 GPa of MgO trained with and without Murakami et al. 2012 (M12) and Fiquet  $\epsilon$  71. 596 (F96). The effect of Murakami et al. 2012 data on bulk modulus and thermal expansivity is main by around 2700 K, and it gradually reduces as pressure decreases or increases outside the interval approx imately 45-90 GPa.

values predicted by the neural networks agree well with explicit EOSs, although Tange 337 et al. 2009 shows slightly higher values at moderate pressures (e.g. 60 GPa). The mean 338 bulk modulus predicted  $r_y$  the neural networks shows a large uncertainty at low pressures 339 (below  $\sim 25$  GPa) when high temperature data by Figuet et al. 1996 are included. In 340 Figure 9- d, we show the bulk modulus predicted by the neural network trained without 341 Fiquet et al. 1996 (and Murakami et al. 2012). Here, the uncertainties at low pressure 342 are significantly decreased. Similarly, neural networks trained without those two data sets 343 predict physically reasonable thermal expansivities (Figure 9-f) compared to those trained 344 with all data sets (Figure 9- e). At high temperatures, we still see a sharp bend around 20 345 GPa (also see Appendix C) which we suggest may be related to anharmonic effects. As the 346

experimental data is relatively sparse in this region, one would need additional measurements 347 (or theoretical studies) to confirm this. Furthermore, the thermal expansivity of Speziale 348 et al. 2001 deviates from other EOSs. As mentioned in earlier studies (e.g. Dorogokupets 349 and Dewaele 2007), this may be improved by including anharmonic terms in the EOS. In 350 equation of state formalisms, one can add an anharmonic term to the total free energy. This 35 additional term has a  $T^2$  dependence, rather than simply a linear temperature term. The 352 effect of adding this term is most significant at low pressures, and can potentially capture 353 more accurately the volume dependence at high temperatures compared with the standard 354 thermal models without anharmonicity (for temperatures less than or equal to 2700 K in 355 this meta dataset). 356

Besides low pressure, including Murakami et al. 2012 date during network training provides mean bulk modulus uncertainties that are more that 4 times larger (Figure 10- a) than excluding them together with Fiquet et al. 1.95, and this discrepancy reduces at higher pressures (Figure 10- b). Moreover, as expected, neither Fiquet et al. 1996 nor Murakami et al. 2012 data influence bulk modulus at the temperatures, as shown in Figures 10- c and 9- a, b.

#### 363 6. Discussion

Fitting parameters (such as  $K_{0T}$  and  $K'_{0T}$ ) are inherent to explicit global EOSs, and a 364 correlation between them tells us how one parameter changes with another providing optimal 365 global fit. We do not estimate the uncertainties on fit parameters of EOSs which are specific 366 to the underlying global functional form. Instead, we directly provide the uncertainties on 367 volumes which are local in P-T space. The MDN is a kernel based method where we fit 368 (a mixture of Gaussian) kernels to the experimental data and get an arbitrary probability 369 density function on volume at any given P and T. The neural networks are flexible and 370 interpolate locally; the uncertainties in one region of P-T space don't impact the posterior 37 pdf everywhere. For example, Figure 7 shows no change in high pressure pdfs while removing 372

Fiquet et al. 1996 data in the region of low pressures. Our approach is also very powerful at
identifying data inconsistencies when using different data sources.

The posterior pdfs given by the MDNs represent uncertainties in volume due to experi-375 mental errors, data gaps and data inconsistencies from different studies. Together with the 376 uncertainties in mean isothermal bulk modulus and thermal expansivity, these results can 377 be used by, for example, seismologists working on thermochemical interpretation of seismic 378 data. Although uncertainties in volume, bulk modulus and thermal expansivity vary locally 379 depending on sparsity and consistency of the experimental *lata*, using these outputs from 380 MDNs, one can directly compute bulk wave speed ( $\phi^2 = K_{\xi}/\rho$ ) and density ( $\rho$ ) at any given 381 pressure and temperature. However, in order to compute buck wave speeds at temperatures 382 applicable to the lower mantle, we need the adiabatic bulk modulus  $(K_S = K_T(1 + \alpha \gamma T))$ , 383 where  $\gamma$  is Grüneisen parameter and  $\alpha$  is the thermal expansivity. Nevertheless, assuming 384 that the difference between isothermal  $(K_T)$  a  $\nu_{\star}$  is diabatic  $(K_S)$  bulk moduli, at 300 K is 385 roughly within  $\pm 1.0\%$  (Marquardt et al 2018), the bulk wave speed of MgO is  $11.14\pm0.07$ 386 km/s at 135 GPa. At the same condition, the relative uncertainty (one standard deviation 387 around mean) in density predicted  $b_1$  the MDNs is about  $\pm 1.0\%$ . This is larger than or 388 comparable to the relative density ranations in lower mantle estimated by previous studies 389 (e.g. Ishii and Tromp 1999, Trompert et al. 2004, Koelemeijer et al. 2017). Although the 390 Grüneisen parameter varies is a function of volume that ultimately depends on pressure 391 (and temperature), we assume it to be approximately  $1.1\pm0.3$  (e.g. Stixrude and Lithgow-392 Bertelloni 2011, Ye et al. 2017) at 2700 K and 135 GPa to give an estimate of uncertainties 393 in bulk wave speed. In doing so, the relative uncertainty in bulk wave speed is about  $\pm 1.77\%$ 394 which is larger than the reported bulk sound speed variation in the lower mantle (e.g. Tram-395 pert et al. 2004). 396

Estimation of mineral properties beyond the range of experimental data requires extrapolation. The standard EOSs can easily be used for extrapolation provided that the assumptions of the functional form hold in the region of no data. In general, it has been observed that



Figure 11: Probability density function for volume of MgO along a 2700 K isotherm (a) and 100 GPa isobar (b). Training data belonging to temperature be owe in 2400 and 3000 K (a), and pressure range from 96 to 103 GPa (b) are also shown. Magenta (SLB051.) and red (T09) curves are Stixrude and Lithgow-Bertelloni 2005, 2011 and Tange et al. 2009 EOS, respectively. They follow the volume trend predicted by the network. In the region outside the prior data, the train. d MDNs provide wider pdfs as they are forced to extrapolate the volume. To illustrate this more clear'y, "olume pdfs at a fixed temperature (and pressure) and three different pressure (and temperature) are also shown in c (and d).

MDNs provide a wider estimate of uncertainties in the region of little to no training data 400 (Käufl et al. 2016a). Here is as shown by the wider pdfs in Figure 11, the uncertainty 401 in predicted mineral p. portles increases when the network has to extrapolate from distant 402 training data. We note that EOSs of Stixrude and Lithgow-Bertelloni 2005, 2011 and Tange 403 et al. 2009 closely follow the pdf predicted by the network indicating that it learns a func-404 tional form present in the data, but errs on the cautious side by returning larger uncertainties. 405 From a Bayesian perspective, we would advise against extrapolation as this covers a region 406 outside the prior. Figure 11, however, demonstrates some capability of neural networks to 407 extrapolate beyond the ranges of the data, although we would need to establish how far this 408 is related to the precise network architecture. 409

The shear modulus is required to calculate compressional and shear wave speeds. There is no 410 thermodynamic expression for the shear modulus, but functional forms are often assumed, 411 for example third order finite-strain and shear counterpart of the Keane EOSs (Keane 1954) 412 by Kennett 2017, to compute the shear modulus which are based on the bulk modulus 413 calculation. One can also use the linear relationship among shear modulus, adiabatic bulk 414 modulus and pressure given by Stacey 1995. However, the uncertainties in shear modulus 415 would then be dependent on those in bulk modulus, and the assumption that shear properties 416 can be constrained from the bulk properties. An alternative is to use data from experiments 417 such as Brillouin Spectroscopy that provide shear wave speed information. Together with 418 unit-cell volume, as measured by X-ray diffraction on the same sample (e.g. Murakami et al. 419 2012, Kurnosov et al. 2017) and known sample composition, the density and thus shear 420 moduli can be determined. However, these data set: do lot cover simultaneous high pressure 421 and temperature regions that are expected in the Earth's lowermost mantle. For example, 422 the highest temperature and pressure data for MgO reported in Murakami et al. 2012 are 423 six measurements at 2700 K and between 32.5-68.4 GPa. Nevertheless, a combination of 424 wave speed data from ultrasonic tech. ques and Brillouin Spectroscopy together with high 425 P-V-T data from x-ray diffraction be highlighted has the potential to exhaustively sample the 426 lower mantle geotherm in the read future (Marquardt and Thomson 2020). 427

We note that, in principle, a combination of experimental data and theoretical calculations 428 (e.g. Karki et al. 1999, Og. )v and Dorogokupets 2003, Wu et al. 2008) is possible. This may 429 provide additional constraints on the predicted mineral properties covering a wider range of 430 pressure and temperature. Since our approach implicitly identifies the consistency between 431 different data sources, a proper rationale can be developed to mix data and uncertainties from 432 theory with experiments. Furthermore, the MDN based approach can easily be extended to 433 the upper mantle and the core. Since MDNs are flexible, they can be employed to model 434 multi-mode targets/outputs. This would be helpful to model for example volume anomalies 435 induced by the iron spin transition (e.g. Marquardt et al. 2009, Speziale et al. 2007, Lin 436 et al. 2006, Crowhurst et al. 2008, Solomatova et al. 2016). A natural progression of this 437

work is to extend it for solid solution. It is straightforward to include composition, e.g. the
Mg/Fe ratio, by including it as an extra dimension in the input data (i.e. P, T and mol% Fe
in ferropericlase) provided there is enough training data.

#### 441 7. Conclusions

This study demonstrates the feasibility of a neural network based approach to infer the 442 material properties of lower mantle minerals. In our approach, we learn the underlying P-V-443 T relationship providing a reasonable approximation of the  $P-v \cdot T$  data of MgO. This allows 444 us to compute the uncertainties in density, thermal expansivity and bulk modulus without 445 prescribing an explicit EOS. Once the networks are traned, it is a simple function that can 446 be evaluated at any given pressure and temperature to get volume, mean bulk modulus and 447 thermal expansivity with uncertainties. In order to train the networks, we collect data from 448 high P-V-T experiments without prior selection of data (e.g. based on pressure scale or 449 functional form used). Hence, our uncervities are not biased towards a subjective selection 450 of experimental data. Furthermore, ur approach identifies inconsistencies between data 451 from different sources. The assumpt of that an EOS follows a particular form provides a 452 priori information by fixing then form (or thermodynamic model) and/or pressure scale. 453 It remains to be determined which EOS form best describes the thermodynamic behaviour 454 of MgO at wide range of p essures and temperatures. In this study, we compare a few 455 "standard" EOSs with the material properties inferred from neural networks and show that 456 choosing one particular explicit form provides a biased estimate of uncertainties. 457

<sup>458</sup> Based on the prediction performance of the MDNs and comparison with conventional EOSs <sup>459</sup> (such as Figures 3, 7, 9, and Appendix C), we can be most confident about physical inter-<sup>460</sup> pretation of seismic data in the lower mantle within the prior range of experimental data <sup>461</sup> (Figure 2). In the regions where there exists little evidence about how the P-V-T relationship <sup>462</sup> behaves, such as at low pressure, high temperature (<25 GPa, >1500 K), and temperatures <sup>463</sup> approximately >2700 K at pressures expected towards the core-mantle boundary, neural

<sup>464</sup> networks show increasingly uncertain predictions. Although for the Earth's lower mantle, <sup>465</sup> low pressure and high temperature environments may not be relevant, they are expected in <sup>466</sup> other planetary bodies such as the Moon and Mars (e.g. Khan et al. 2014, 2018). With <sup>467</sup> currently available data, it likely provides meaningful uncertainties that could be used by <sup>468</sup> seismologists within certain ranges of pressure and temperature, while highlighting the P, <sup>469</sup> T regions in which more experimental (or theoretical) data is needed before we can draw <sup>470</sup> robust conclusions on temperature and composition.

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#### 477 Appendices

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#### 478 Appendix A. Generalised theory of the MDN

Let,  $\boldsymbol{x} = \{x_1, x_2, ..., x_I\}$  be the input data to the feed-forward part of the MDN. Please note, to generalise this section, we write inputs as  $\boldsymbol{x}$  and targets as  $m_k$  instead of P & T and V, respectively. The feed-forward network outputs  $y_k$  are computed as a weighted sum of the outputs from the hidden nodes plus a bias

$$y_k = f_2 \left( \sum_{j=1}^J \alpha_{jk} h_j + \alpha_{0k} \right) \tag{A.1}$$

where the function  $f_2$  is an identity function such that  $f_2(p) = p$ ,  $\alpha_{jk}$  is the hidden layer weight matrix and  $\alpha_{0k}$  represents a bias term of each output node. Now, the hidden node

486 outputs  $h_j$  are computed as

487

$$h_j = f_1 \left( \sum_{i=1}^{I} \alpha_{ij} x_i + \alpha_{0j} \right) \tag{A.2}$$

where the function  $f_1$  is a logistic sigmoid function  $f_1(p) = \frac{1}{1 + exp(-p)}$ ,  $\alpha_{ij}$  is the input layer weight matrix,  $\alpha_{0j}$  are the biases of hidden nodes and  $x_i$  are input data.  $y_k$  are related to the parameters, namely weights  $(\pi_n)$ , means  $(\mu_n)$  and standard deviations  $(\sigma_n)$  of Gaussians in the Gaussian Mixture Model (GMM) by the following relationship (for details see e.g. Bishop 1994, de Wit et al. 2013)

493 
$$\pi_n(\boldsymbol{x};\boldsymbol{\alpha}) = \frac{exp\left(y_k^{(\pi)}(\boldsymbol{x};\boldsymbol{c})\right)}{\sum_{n=1}^M exp\left(y_{\perp}^{(\pi)}(\boldsymbol{\omega};\boldsymbol{\alpha})\right)},$$
(A.3)

$$\mu_n(\boldsymbol{x};\boldsymbol{\alpha}) = \underline{\boldsymbol{y}}_k^{(\mu)}(\boldsymbol{x},\boldsymbol{\alpha}) \text{ and }$$
(A.4)

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 $\sigma_n(\boldsymbol{x};\boldsymbol{\alpha}_j) = exp\Big(y_k^{(\sigma)}(\boldsymbol{x};\boldsymbol{\alpha})\Big).$ (A.5)

#### 498 Appendix B. MDN initial zation and training details

The total data  $(\boldsymbol{x})$  is divided into three sets- training (70%), monitoring (20%) and test (10%) sets such that

$$\boldsymbol{x}^{train} \subset \boldsymbol{x}, \ \boldsymbol{x}^{monitor} \subset \boldsymbol{x} \ and \ \boldsymbol{x}^{test} \subset \boldsymbol{x}$$
 (B.1)

with  $x^{train} \cap x^{monitor} = \emptyset$ ,  $x^{train} \cap x^{test} = \emptyset$  and  $x^{monitor} \cap x^{test} = \emptyset$ . Using the training data  $(x^{train})$  we train the MDN. However, before we train the MDN we need to decide on initial values of the network parameters. We randomly draw the input layer and hidden layer weights (Bishop 1995) according to the following Gaussian distributions

506 
$$\alpha_{ij} \sim \mathcal{N}\left(0, \frac{1}{I+1}\right) \tag{B.2}$$

507 and

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524

$$\alpha_{jk} \sim \mathcal{N}\Big(0, \frac{1}{J+1}\Big),$$
(B.3)

respectively. Where I and J are number of input and hidden nodes, respectively. Similarly, the output layer biases are initialized by a K-means clustering algorithm (i.e. fitting a GMM to the training data set). Once the initialization is done and the training begins, the difference between the output and the target can be computed according to the error function

$$E^{train} = \sum_{train} -\ln\left(p(m_k | \boldsymbol{x}^{train}; \boldsymbol{\alpha})\right)$$
(B.4)

which is summed over all training data providing the average error. This function is also called the loss function which is minimized iteratively using the ADAM optimization method (see detailed algorithm in Kingma and Ba 2014). The explicit dependence of output posterior on the network parameters (see Käufl et al. 2016a and references therein) can be avoided by using multiple MDNs and combining them by weighted sum. The weight of each MDN is determined by the test set error as

$$w_i = e z_F \left( -\frac{E^{test}(\boldsymbol{x}^{test}, \boldsymbol{\alpha}_i)}{N} \right)$$
(B.5)

where index i denotes the i-th MDN (C MDNs in total) and N is the size of the test data set, and the MDNs are combined according to

$$\sum p(m_k | \boldsymbol{x}; \boldsymbol{\alpha}) = \sum_{i=1}^{C} \frac{w_i}{\sum_j w_j} p_i(m_k | \boldsymbol{x}; \boldsymbol{\alpha}_i).$$
(B.6)

#### 525 Appendix C. Mineral properties

526 Appendix C.1. P-V-T EOS



Figure Appendix C.1: P-V relationship of MgO predicted by MDNs trained with (left) all data and (right) excluding Murakami et al. 2012 and Fiquet et al. 1996. Comparison with previously published EOSs (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2005, 2011 and Dorogokupets and Dewaele 2007) along 1500 K (top), 2000 K (middle) and 2700 K (bottom) isotherms also shown.





Figure Appendix C.2: Comparison of the bulk into this of MgO predicted by the neural network along 1500 K (top) and 2700 K (bottom) isotherms with only studies (Tange et al. 2009, Speziale et al. 2001, Stixrude and Lithgow-Bertelloni 2011 and Dorogokupets and Dewaele 2007) as a function of pressure. Left panel shows results from MDNs trained with all on the right panel shows results from MDNs excluding Murakami et al. 2012 and Fiquet et al. 19% and the right.



#### <sup>528</sup> Appendix C.3. Thermal expansivity

Figure Appendix C.3: Con., anson of the thermal expansivity of MgO predicted by neural networks with Tange et al. 2009, Stixrude ar 1 Lithgow-Bertelloni 2011 and Dorogokupets and Dewaele 2007 along 300 K (top), 1500 K (middle) and 2700 K (bottom) isotherms as a function of pressure. Left panel: MDNs trained with all data. Right: MDNs trained without Murakami et al. 2012 and Fiquet et al. 1996 data.

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#### **Declaration of interests**

 $\boxtimes$  The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

□The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

#### Author statement

**Manuscript title**: Inferring material properties of the lower mantle minerals using Mixture Density Networks.

Ashim Rijal: methodology, software, validation, formal analysis, investigation, visualisation, writing - original draft

Laura Cobden: conceptualisation, supervision, funding acquisition, writing - review and editing

Jeannot Trampert: supervision, writing - review and editing

Jennifer M. Jackson: validation, formal analysis, writing - review and editing

Andrew Valentine: methodology, writing - review and editing

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