

The non-linear break-up of an inviscid liquid jet using the spatial-instability method

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THE NON-LINEAR BREAK-UP OF AN INVISCID LIQUID JET USING THE SPATIAL-INSTABILITY METHOD

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Abstract—A liquid jet originating from a nozzle with radius r_0^* breaks up into droplets in consequence of disturbances of certain frequencies, depending on the fluid properties and the nozzle geometry. A theoretical model is developed to describe the growth of these disturbances at the jet surface. The model is based on the inviscid and irrotational flow governed by the Laplace equation together with the kinematical and dynamical conditions at the free surface of the jet. A comparison is made between the model and experimental data from literature. The model predicts a dependence on the disturbance-amplitude of the break-off mode. Contrary to other experimental results, the model predicts satellites (i.e. smaller droplets between the main larger ones) at wavelengths exceeding a critical value of $(10/7) 2\Pi r_0^*$. The disturbances grow at wavelengths more than the theoretical bound of $2\Pi r_0^*$. Discrepancies with experimental data are possible because of the effect of viscosity in the theory. It is shown that the effect of viscosity on the jet can be neglected under certain conditions.

INTRODUCTION

A liquid jet originating from a nozzle is sensitive to disturbances. Disturbances of certain frequencies cause the jet to break up into a series of successive droplets. There are two theoretical methods to investigate the behaviour of a disturbed liquid jet. The "spatial-instability" method describes the disturbance of the jet surface as a travelling wave in axial direction. The "temporal-instability" method describes the surface disturbance of the jet as a standing wave on an infinitely long cylinder with the nozzle at infinity.

The model presented in this paper is based on spatial instability and describes the jet form close to the nozzle in the form of travelling waves with harmonic influences. The Laplace equation together with the dynamical and kinematical boundary condition for the free surface are used to describe the radius of the jet and the velocity. The model is mathematically simplified by neglecting both the effects of viscosity and the surroundings. It is possible to approximate the solution of the equations for the radius of the jet by the solution of a simplified fluid dynamical theory, the Cosserat theory, which is a simplified one-dimensional theory. In this article we present an approximation for radius and velocity by Taylor series expansions with respect to the disturbance-amplitude.

The break-up process depends on the surface tension σ^* , density ρ^* , nozzle radius r_0^* and initial disturbance-amplitude δ_0^* . The disturbances grow with time and distance from the nozzle. Piezo-crystals or mechanical vibrators can be used as sources when applied to the jet surface, the velocity or the pressure distribution in the jet. Previously published theoretical models are of a more limited use as the model presented in this paper, either because the stability analysis ignores higher harmonic effects or because these models are based on a simplified fluid dynamical theory. The spatial-instability method describes the physical reality better than the temporal-instability method does, whereas the first method does not impose periodic axial demands on the jet. The model based on spatial instability shows that satellites can break up before or after a main drop, dependent on the disturbanceamplitude. A condition is derived by which the influences of the viscosity can be estimated and eventually neglected.

LITERATURE

The first mathematical model was published by Rayleigh (1878). This model was a stability analysis for infinitely small disturbances of an inviscid jet based on the temporal-instability method. Weber (1931) and recently Sterling and Sleicher (1975) extended this analysis to an aerodynamically influenced viscous jet. According to these linearized models the only dependent variable for the break-up process is the break-up time.

By approximately 1960 these linearized theories were found inadequate to describe the phenomenon accurately. Only the break-up time could be reasonably well compared with experiments. At present the formation of satellites is considered to be more characteristic for the phenomenon than the break-up length or break-up time. Yuen (1968) and Lafrance (1974, 1975) have obtained analytical approximations for an

inviscid jet ignoring influences of the environment on the basis of temporal instability and mass conservation for higher harmonic disturbances. They arrived at establishing the formation and the existence of satellites. Rutland and Jameson (1970) compared experimental data with numerical results based on Yuen's model. By coupling the initial and final volume of the droplets they succeeded in calculating diameters of representative satellites and main drops. They concluded that in case of wavelengths $\lambda^* \leq (10/7)2\Pi r_0^*$ no satellites could be formed. With a modified model, Lafrance (1974) showed that no satellites were formed for $\lambda^* \leq (10/8) 2 \Pi r_0^*$. These conclusions are certainly not correct in the case of viscous jets. Chaudhary and Maxworthy (1980a, b) and Chaudhary and Redekopp (1980) used a comparable model and performed experiments on the jet behaviour and satellite drop formation.

Keller et al. (1973) used spatial instability in combination with a stability analysis to describe the breakup process of a jet. As a result, the description of the break-up process is improved. From a mathematical viewpoint the disturbances behave like waves progressing on the jet surface. Bogy (1978, 1979a, b) expanded the Cosserat theory of Green (1976) to develop a new analytical model for higher harmonics based on spatial instability. The Cosserat theory is a one-dimensional theory simplifying the flow in the jet, especially in the radial direction.

The model presented in this paper is an extension of the models referred to above. It describes the break-up process of a jet with the spatial-instability method and mass conservation including higher harmonic effects. The jet form is approximated by a Taylor series expansion with respect to the disturbance-amplitude. The geometry of the jet, the velocity components in it, the break-up length and the break-up mode of satellites before or behind a main drop are calculated with our model. Computation of the last mentioned phenomenon is not found in earlier published articles. The model still has physical limitations on account of the neglect of viscosity and environment effects such as gravity, mass or heat transfer.

MATHEMATICAL MODEL

General description

A semi-infinite axisymmetric jet with liquid density ρ^* and surface tension σ^* emerges from a nozzle of

radius r_0^* . The uniform velocity v_0^* at the nozzle is harmonically disturbed. As a result the radius R^* of the jet is a function of the axial coordinate z^* and time t^* . Dimensional variables are indicated with the superscript *. Figure 1 shows the geometry of the jet. The characteristic length and time are r_0^* and t_0^* $= r_0^*/v_0^*$, respectively. The break-up process of the jet has a characteristic time-scale $t_1^* = (\rho^* r_0^* 3/\sigma^*)^{1/2}$. The characteristic dimensionless number is the Weber number which is defined as the quadratic ratio of the characteristic time scales t_1^* and t_0^* :

$$We = (t_1^*/t_0^*)^2 = \rho r_0^* (v_0^*)^2 / \sigma^*.$$
(1)

The influence of the surroundings, viscosity and gravity are neglected. Rotational symmetry and irrotational flow in the jet are assumed. The flow in the jet is described by means of the Laplace equation, a pressure continuity condition at the jet boundary and a containment boundary condition, in the sense that all liquid remains within the jet surface. The nozzle diameter is constant and the axial jet velocity at the nozzle is sinusoidally disturbed.

The local jet radius R^* and the velocity potential Φ^* are approximated by series expansions. The nonlinear differential equations are reduced to a set of linear differential equations; these are solved up to the third order to enable the description of satellites. From the homogeneous differential equations the characteristic dispersion relation is derived. The dispersion relation is solved numerically on a mainframe Burroughs. The break-up length z_B^* and the local jet radius $R^*(z^*, t^*)$ are calculated by the computer. These results are compared with the experimental data taken from the literature (Chaudhary and Maxworthy, 1980a, b).

Formulation

A velocity potential Φ^* with $u^* = \frac{\partial \Phi^*}{\partial z^*}$ and $v^* = \frac{\partial \Phi^*}{\partial r^*}$ is introduced. All variables are made dimensionless with respect to the characteristic time $t_0^* = r_0^*/v_0^*$ and the characteristic length r_0^* . The Laplace equation in terms of dimensionless Φ is

$$\nabla^2 \Phi = \Phi_{rr} + \frac{1}{r} \Phi_r + \Phi_{zz} = 0 \quad (0 \le r \le R, \ 0 \le z \le z_B).$$
 (2)

The kinematical boundary condition for free surface of



Fig. 1. Sketch of the geometry of a liquid jet emerging from a nozzle at $z^*=0$. The entrance velocity is disturbed sinusoidally. The local jet radius R^* is a function of the axial coordinate and time. Under certain conditions the jet breaks up into main drops, spaced at wavelength λ^* , and satellites at the break-up length z_B^* .

the jet is

$$\Phi_r = R_t + \Phi_z R_z \quad (r = R). \tag{3}$$

The dynamical boundary condition is given by Busker (1983):

$$\Phi_{r} + \frac{1}{2}(\Phi_{r}^{2} + \Phi_{z}^{2}) + \frac{1}{We} \{1/[R(1 + R_{z}^{2})^{1/2}] - R_{zz}/(1 + R_{z}^{2})^{3/2}\} = \frac{1}{2} + \frac{1}{We} \quad (r = R).$$
(4)

The nozzle diameter is constant:

$$R = 1 \qquad (z = 0) \tag{5}$$

and the jet velocity is disturbed sinusoidally according to

 $\Phi_z = 1 + \delta_0 \cos \omega_1 t \qquad (z = 0, \ 0 \le r \le R).$ (6)

Chaudhary and Redekopp (1980) started from the same eqs [(2)-(4)]. Our approach, however, is to seek to find solutions of the form

$$R = 1 + R' = 1 + \delta_0 \eta_1 + \delta_0^2 \eta_2 + \delta_0^3 \eta_3 + O(\delta_0^4)$$
 (7)

and

$$\Phi(r, z, t) = z + \delta_0 \Phi_1 + \delta_0^2 \Phi_2 + \delta_0^3 \Phi_3 + O(\delta_0^4).$$
 (8)

Here η_i is the *i*th-order harmonic surface distortion, and Φ_i is the *i*th-order harmonic velocity potential of the main disturbance with i=1. The derivative of the velocity potential on the boundary is transferred to the undisturbed jet geometry R=1 by a re-expansion of the velocity potential Φ into a series of the boundary transformation R'. After substitution of the boundary transformation R' the terms of eqs (2)-(6) can be expressed in terms of δ_0^0 , δ_0^1 , δ_0^2 and δ_0^3 , respectively. After that, the equations are separated into powers of δ_0 . The zeroth-order solution gives the undisturbed cylindrical jet.

The first-order equations are given by

$$\nabla^2 \Phi_1 = 0 \qquad (0 \le r \le 1) \tag{9}$$

$$\Phi_{1,r} - \eta_{1,t} - \eta_{1,z} = 0 \qquad (r = 1) \tag{10}$$

$$-\Phi_{1,t} - \Phi_{1,z} + \frac{1}{We}(\eta_1 + \eta_{1,zz}) = 0 \quad (r = 1) \quad (11)$$

with the following nozzle conditions:

$$\eta_1 = 0$$
 (z=0) (12)

$$\Phi_{1,z} = \cos \omega_1 t \qquad (z = 0, \ 0 \le r \le 1). \tag{13}$$

From these first-order equations a stability rule, the dispersion relation, will be derived in the next paragraph. The second-order (δ_0^2) and the third-order (δ_0^3) equations are given by Busker (1983).

Dispersion relation

Because of the first-order nozzle conditions [see eqs (12) and (13)] and eq. (9) which is defined at r=0, a first-order solution is assumed which has the form

$$\eta_1 = C e^{i(\omega t - kz)} \qquad (z, t \ge 0) \tag{14}$$

$$\Phi_1 = DI_0(kr)e^{i(\omega t - kz)} \qquad (z, t \ge 0, r \le 1).$$
(15)

Here, $I_n(x)$ is the modified Bessel function of the first kind and *n*th order. Φ_1 is a solution of the Laplace equation.

Substitution of η_1 and Φ_1 into eqs (10) and (11) produces the following characteristic equation:

$$(\omega - k)^2 = \frac{k(k^2 - 1)}{We} \frac{I_1(k)}{I_0(k)}.$$
 (16)

This equation is called the dispersion relation, because it connects the frequency ω to the wave-number k at a given Weber number. The real part of the wavenumber is related to the wavelength λ by

$$Re(k) = \frac{2\Pi}{\lambda}.$$
 (17)

The imaginary part of k determines the rate of growth of the disturbance. The dispersion relation has an infinite number of solutions for k per ω . The solutions for k are functions of ω and We. Keller et al. (1973) gives the zeroth-order solutions for this equation:

$$(\omega - k)^2 = O(We^{-1})$$
(18)

leading to $k = \omega + O(We^{-1/2})$, and

$$\frac{I_0(k)}{I_1(k)} = O(We^{-1})$$
(19)

leading to $k = \pm i j_{on} + O(We^{-1})$ with $I_0(\pm i, j_{on}) = 0$.

Keller *et al.* (1973) gives an approximation for the solution of eqs (18) and (19), respectively:

$$k_{1,2} = \omega \pm \left[\frac{\omega(\omega^2 - 1)}{We} \frac{I_1(\omega)}{I_0(\omega)} \right]^{1/2} + O(We^{-1}) \quad (20)$$

$$k_{3,4n} = \pm i j_{on} + \frac{j_{on}(j_{on} j_{on} + 1)}{We(\omega^2 + j_{on}^2)^2} [2\omega j_{on} \mp i(\omega^2 - j_{on}^2)] + O(We^{-2}). \quad (21)$$

Another solution is (private communication of J. Boersma, Department of Mathematics, Eindhoven University of Technology)

$$k_{5} = We - \left(2\omega - \frac{1}{2}\right) + \frac{(9/8 + \omega - 3\omega^{2})}{We} + O(We^{-2}).$$
(22)

Only those solutions of k which have a positive phasevelocity $c = \omega/Re(k) > 0$ and a positive group-velocity $c_g = \frac{\partial \omega}{\partial [Re(k)]} > 0$ are valid solutions because only then energy is transported downstream and the waves are travelling in the same direction. The wave-numbers $k_{1,2}$ come up to these conditions. Here it is assumed that the disturbance excitator is at the nozzle and therefore energy goes downstream in the same manner as the waves. Other assumptions are possible depending on the place of the disturbance source along the jet but they are beyond the scope of this article. A singularity-point exists for $\omega = \omega_s$, because there is an abrupt change over from complex to real wave-number values. The dispersion relation has now two equal solutions. For $\omega < \omega_s$ the two solutions are conjugate complex (Fig. 2). By approximation we find for the singularity-point ω_s that

$$\omega_s \approx 1 - \frac{I_1(1)}{2We I_0(1)} \approx 1 - \frac{0.221}{We}.$$
 (23)

Substitution of eq. (23) in eq. (20) gives for $k_{1,2}$ the expression

$$k_{1,2} \approx 1 + \frac{I_1(1)}{2We I_0(1)} \approx 1 + \frac{0.221}{We}.$$
 (24)

Moreover the following relations hold:

$$\lim_{\omega \to \omega_s} \operatorname{Im}(k_{1,2}) = 0; \quad \lim_{\omega \uparrow \omega_s} \frac{\partial \operatorname{Im}(k_{1,2})}{\partial \omega} = \pm \infty;$$
$$\lim_{\omega \downarrow \omega_s} \frac{\partial \operatorname{Im}(k_{1,2})}{\partial \omega} = 0. \quad (25)$$

The maximum growth rate occurs at $\omega_{opt} \approx 0.4858^{1/2} \approx 0.6969$ for large Weber numbers.

Effect of viscosity

The dispersion relation for a jet with viscosity μ under the same conditions as an inviscid jet is given by Busker (1983):

$$(\omega - k)^{2} = + \frac{k(k^{2} - 1)}{We} \frac{I_{1}(k)}{I_{0}(k)} - \frac{2i(\omega - k)k}{Re} \left[2k - \frac{I_{1}(k)}{I_{0}(k)} \right] - \frac{4k^{3}}{Re^{2}} \left[k - \frac{lI_{0}(l)I_{1}(k)}{I_{1}(l)I_{0}(k)} \right]$$
(26)

with $l^2 = k^2 + i(\omega - k)Re$.

In the case of $\omega < 1$ and by approximating $\frac{I_1(k)}{I_0(k)} \approx \frac{k}{2} + O(k^2)$ eq. (26) can be rewritten analogously to Weber (1931) as

$$(\omega - k)^{2} = + \frac{k(k^{2} - 1)}{We} \frac{I_{1}(k)}{I_{0}(k)} - \frac{3i(\omega - k)k^{2}}{Re}.$$
 (27)

The solution for this equation is

$$k_{1,2} = \omega \pm \left[\omega(\omega^2 - 1) \frac{I_1(\omega)}{I_0(\omega)} \right]^{1/2} \left[\frac{1}{We^{1/2}} + \frac{(3i\omega^2 + 2)}{(3i\omega^2 - 1)Re} \right].$$
 (28)

The influence of the viscosity might be neglected if the influence of the viscous terms are less than 2% of the surface tension term, thus

$$\left|\frac{(3i\omega^2+2)We^{1/2}}{(3i\omega^2-1)Re}\right| < 0.02.$$
(29)

For the examples of the experiment of Chaudhary and Maxworthy (1980a, b) this gives in the case of ω_1 = 0.4313, We=922.6 and Re=1486 a value about 3.7% and in case of ω_1 =0.720, We=330.9 and Re=889.9 a value about 2.8% (see Application).

First-order solution

The dispersion relation provides the wave-number values $k_{1,2}$ for a given dimensionless frequency ω_1 and Weber number which are used in the first-order solutions (14) and (15):

$$\eta_1 = P_1 \cos \theta_1 + P_2 \cos \theta_2 \tag{30}$$

$$\Phi_1 = Q_1 I_0(k_1 r) \sin \theta_1 + Q_2 I_0(k_2 r) \sin \theta_2 \qquad (31)$$

where $\theta_1 = \omega_1 t - k_1 z$, $\theta_2 = \omega_1 t - k_2 z$, and P_1 , P_2 , Q_1 and Q_2 are constant. Substitution into the kinematical boundary condition and the nozzle conditions produces the following first-order solution:

$$\eta_1 = \frac{P}{V} (\cos \theta_1 - \cos \theta_2) \tag{32}$$

$$\Phi_{1} = \frac{P}{V} \left[\frac{-(\omega_{1} - k_{1})I_{0}(k_{1}r)}{k_{1}I_{1}(k_{1})} \sin \theta_{1} + \frac{(\omega_{1} - k_{2})I_{0}(k_{2}r)}{k_{2}I_{1}(k_{2})} \sin \theta_{2} \right].$$
(33)



Fig. 2. Solutions of the characteristic dispersion relation at We = 10. The path of wave-number k as a function of the frequency ω is given. A singularity-point exists at $\omega = \omega_s$. In case of $\omega < \omega_s$ the solutions of k are complex conjugated.

Here
$$P = I_1(k_1)I_1(k_2)$$
 and $V = (\omega_1 - k_1)I_1(k_2) - (\omega_1 - k_2)I_1(k_1)$.

Second- and third-order solutions

The second- and third-order solutions are derived from the second- and third-order differential equations together with the nozzle conditions. They consist of a particular part and a homogeneous part. The particular parts of the second-order boundary differential equations are composed of terms with η_1 and Φ_1 . In the case of the third-order equations it is composed of terms with η_1 , Φ_1 , η_2 and Φ_2 . By substituting the lower-order solutions in the particular part of the second- and third-order differential equations a general form of the particular second- and third-order solution is derived. These general forms for the second- and third-order solutions are substituted in the homogeneous parts of the boundary differential equations. The particular solutions are then determined. As homogeneous solutions of the ithorder (i = 2 or 3) are taken:

$$\eta_i^H = C_1 \cos \theta_{i1} + C_2 \cos \theta_{i2} \tag{34}$$

$$\Phi_i^H = D_1 I_0(k_{1i}) \sin \theta_{i1} + D_2 I_0(k_{2i}) \sin \theta_{i2} \qquad (35)$$

with $\theta_{2i} = 2\omega_1 t - k_{2i}z$ and $\theta_{3i} = 3\omega_1 t - k_{3i}z$.

 k_{2i} and k_{3i} come from the dispersion relation with $\omega = 2\omega_1$ and $3\omega_1$, respectively. The homogeneous solutions are substituted in the boundary differential equations. A relationship between η_i^H and Φ_i^H is deduced. The complete solution is found by substitution of the homogeneous and particular parts of the solution in the nozzle conditions. The equations with which all constants necessary for the complete solution can be calculated are given by Busker (1983). In the case of $\omega_1 < \omega_s$ and $k_{1,2}$ values are conjugate complex and it is possible to make simplifications in R and Φ . The third-order solution is given in Appendix 1.

If necessary the velocity components, pressure distribution and volume in the jet could be calculated up to the break-up point. The field of interest in this paper is restricted to the unstable growth modes of the disturbances, i.e. $\omega_1 < \omega_s$. Because R(z) is continuous, R cannot describe the discontinuous behaviour of the jet beyond the break-up point. Therefore only $z \leq z_B$ is considered.

APPLICATION

Comparing the model with a particular experiment

Chaudhary and Maxworthy (1980a, b) investigated experimentally the break-up process with a piezoelement as a disturbance source. This was driven by a sinusoidal wave at a frequency $f^* = 100$ kHz with a peak voltage V_e^* as amplitude. The jet velocity v_0^* was fitted to reach the desired wavelength λ^* . It resulted in the dimensionless wave-number

$$k = \frac{2\pi r_0^*}{\lambda^*}.$$
 (36)

$$v_0^* \approx \lambda^* f^* \tag{37}$$

$$We \approx \frac{2\pi^2 \rho^* r_0^* (f^*)^2}{\sigma^* (k^*)^2}$$
 (38)

and

$$z_B = (We T_B)^{1/2}.$$
 (39)

The theoretical value of ω_1 is found by a variation of ω in the dispersion relationship at the given Weber number to the effect that $Re(k_{1,2}) = k$.

values of ω_1 , We and z_B for the spatial-instability

model with the relationships

The Reynolds number in the jet is given by

$$Re = \frac{\rho^* v_0^* r_0^*}{\mu^*} \approx \frac{2\pi \rho^* (r_0^*)^2 f^*}{\mu^* k}.$$
 (40)

Chaudhary and Maxworthy (1980a) used a water jet and varied k between 0.3 and 1. The values of the physical quantities are given in Table 1. This resulted in values of We in the range of 170–1910 and values of Re in the range of 640-2140. A series of experimental data have been compared with a series of numerical results of our model. The wave-number k = 0.4312is used which gives We = 922.6, Re = 1486 and $\omega_1 = 0.4313$. Some pictures are given by Chaudhary (1980a). In the case of We = 922.6 the experimental amplitude V_{e}^{*} varied between 1 and 80 V and the corresponding minimum break-up length z_B between 860 and 305. A break-up length smaller than 586 causes a satellite break-up after the main drop, as opposed to the break-up before the main drop for higher values of z_B . The corresponding calculated values of z_B lead to a range of dimensionless disturbance-amplitude values δ_0 between 3×10^{-5} and 4 $\times 10^{-2}$ (Table 2).

 V_e^* and δ_0 can be written in a logarithmic linear relation. The constants are calculated by linear regres-

Table 1. Values of the physical quantities usedwith the experiments of Chaudhary and
Maxworthy (1980a, b)

Physical quantity	Value	Dimension	
r*	3.048×10^{-5}	m	
σ*	65.3×10^{-3}	N/m	
ρ*	1002	kg/m ³	
μ *	9.128×10^{-4}	Pa s	

Table 2. Experimental data of Chaudhary and Maxworthy for the peak voltage V_*^* and the break-up length z_B compared with the calculated values of z_B and the disturbance-amplitude δ_0 at the dimensionless frequency $\omega_1 = 0.4313$ and Weber number We = 922.6

Experimen and N	ts of Chaudhary Maxworthy	Calculated results of the model		
V *	Z _B	Z _B	δ_0	
80	305	298	5×10^{-3}	
70	321	324	4×10^{-3}	
40	397	399	2×10^{-3}	
30	437	434	15×10^{-4}	
20	497	493	9×10^{-4}	
15	533	537	6×10^{-4}	
10	586	584	4×10^{-4}	
4	703	694	15×10^{-5}	
2	784	779	7×10^{-5}	
1	860	876	3×10^{-5}	

sion leading to

$$\ln (V_e^*) = 9.26 + 0.896 \ln (\delta_0). \tag{41}$$

A second series of experimental data with k = 0.720and We = 330.9 is used as a comparison. The experimental values of z_B were between 417 and 169. The calculated values of δ_0 with $\omega_1 = 0.720$ are in the range between 3×10^{-5} and 3×10^{-3} (Table 3).

The same log-linear relation for V_e^* and δ_0 does not apply when break-up lengths are smaller than 300 (Fig. 3).

Break-up mode of main drops and satellites

Another test criterion for the validity of the model is the break-off mode of main drops and satellites at the minimum break-up length z_{Bmin} . For small amplitudes

Table	3.	Experi	mental	data	of	Chaud	hary	and
Maxwo	orth	y for th	ne peak	voltage	V*	and th	e brea	k-up
length	Z _B C	ompare	d with	calculat	ed va	alues of	z_{R} and	d the
disturb	anc	e-ampli	itude δ_0	at the d	imer	isionles	s frequ	ency
1.	$\omega_1 =$	= 0.720	and We	eber nur	nber	We = 3	30.9	-

Experimen and N	ts of Chaudhary Maxworthy	Calculated results of the model		
V .	ZB	Z _B	δ_0	
70	175	173	3×10^{-3}	
50	196	193	2×10^{-3}	
30	228	231	1×10^{-3}	
20	256	257	6×10^{-4}	
15	270	268	5×10^{-4}	
10	291	295	3×10^{-4}	
7	311	314	2×10^{-4}	
2	379	379	6×10^{-5}	
1	417	417	3×10^{-5}	

 δ_0 the main drop will break off first of all at the breakup length $z_B = z_{Bmin}$ followed by the satellite at $z_B > z_{Bmin}$. For large amplitudes δ_0 the satellite will break off first of all at $z_B = z_{Bmin}$ followed by the main drop. The jet shape in the surroundings of the breakup length z_B for small and large δ_0 values are shown in Figs 4 and 5.

In the case of $\omega_1 = 0.4313$ and We = 922.6 the experimental transition value of the different modes occurs for V_e^* between 4 and 10 V (Chaudhary and Maxworthy, 1980a) corresponding with δ_0 between 15×10^{-6} and 4×10^{-4} for the fixed log-linear relation between V_e^* and δ_0 . The analytical model predicts that the transition value for δ_0 lies between 5×10^{-3} and 6×10^{-3} . This discrepancy is probably due to the neglect of viscous effects in the jet (*Re*



Fig. 3. Comparison between calculated values of the break-up length $z_B(+)$ with experimental data from Chaudhary and Maxworthy (1980a, b) (\bullet). Curve 1 represents values of z_B in case of We=992.6 and ω_1 = 0.4313. Curve 2 in case of We=330.9 and $\omega_1=0.720$ under the conditions of eq. (41).



Fig. 4. Calculated jet shape just before the break-up point z_B for small values of the disturbance-amplitude δ_0 . The calculated diameters and their lengths of the main drop and satellite are given.



Fig. 5. Calculated jet shape just before the break-up point z_B for large values of the disturbance-amplitude δ_0 . The calculated diameters and the lengths of the main drop and satellite are given.

= 1486). The reason for this is the slow specific velocity of the break-up process $(\sigma^*/\rho^* r_0^*)^{1/2}$, in which the viscosity effects cannot be neglected. The model of Chaudhary and Maxworthy (1980a) seems to predict the break-off mode.

The forming of satellites

Sterling and Sleicher (1975) demonstrated the unstable break-up for viscous liquid jets with aerodynamic effects. Their model and the model described in this article do not predict the forming of satellites for $\omega_1 \ge 0.7$. Lafrance (1974) and Rutland and Jameson (1970) did not succeed in finding satellites either. Figure 6 shows the divergence between the model and experiments.

Consistent with the model of Yuen (1968) our model predicts the formation of two satellites for $\omega_1 \leq 0.3$ as a result of the conjugate complex k values for the second- and third-order solutions of the dispersion relation (Fig. 7). For long wavelengths more than one satellite will be formed. When the analytical model is extended to higher orders (larger than three) more satellites are formed at suitable values of ω_1 .

Effect of the disturbance-amplitude on the dimensions of satellites and the main drop

The volume of satellites and main drops is dependent on the disturbance-amplitude δ_0 . The ratio of the maximum diameters of satellites and main drops changes the same way as the ratio of their lengths changes with δ_0 . These phenomena can be seen in Figs 4 and 5. The ratio of the diameters of the satellite to the main drop for $\delta_0 = 1 \times 10^{-2}$ is 0.579 while the ratio of the lengths is 1.05. In case of $\delta_0 = 1 \times 10^{-6}$ the ratio of the diameters becomes 0.479 and the ratio of the lengths 0.728. This means that satellites are larger for greater values of δ_0 . About this no experimental data are available. Because of the limited use of this model, this point will not be pursued further.

Break-up length as function of the frequency

The dependence of the minimum break-up length z_B on the frequency for several Weber and amplitudenumbers is shown in Fig. 8. The plot shows that the minimum break-up length over the whole range $0 < \omega_1 < \omega_s$ occurs somewhere between $\omega_1 = 0.7$ and $\omega_1 = 0.8$ for small δ_0 , and moves towards ω_s for large



Fig. 6. Calculated jet shape at We = 300, $\omega = 0.7$ and $\delta_0 = 1 \times 10^{-4}$ shows no forming of satellites.



Fig. 7. Forming of two satellites under the conditions of We = 300 and $\delta_0 = 1 \times 10^{-2}$ in the case of the dimensionless frequency $\omega_1 = 0.3$. The calculated wavelength is given.



Fig. 8. Values of the break-up length z_B as a function of the dimensionless frequency ω_1 with the disturbance-amplitude δ_0 and We as parameters. (1) Lines with We = 100, (2) lines with We = 300, and (3) lines with We = 1000.

 δ_0 . It can be expected that an undisturbed jet picking up a disturbance from its surroundings will not break up at the most unstable mode (just below $\omega_1 = 0.7$) but that it breaks up at frequencies in an interval $0.65 < \omega_1 < 0.8$.

DISCUSSION AND CONCLUSIONS

The theoretical model of spatial instability provides a good insight into the break-up process in a jet and may be extended. In some respects the model is in accordance with the results of experiments published

We

z

in literature, in other respects it shows discrepancies therewith. The model describes the physical reality better than models based on temporal instability. Our model and Yuen's (1968) are analogous and have as common features the forming of two satellites for $\omega_1 < \frac{1}{3}$ and the absence of satellites for $\omega_1 > \frac{7}{10}$.

The model is more extensive than the one-dimensional Cosserat theory because we made use of twodimensional fluid mechanical equations taking account of the flow in radial direction. However, by comparing the theoretical predictions with available experimental data, it follows that the discrepancies exist due to the neglect of the viscosity or higher-order terms of the solution. Further research could take into account the effect of the nozzle shape and the disturbances on the break-up mode of the jet. In addition, the effects of gravity, aerodynamical, mass and heat transfer phenomena could be investigated.

NOTATION

с	dimensionless phase velocity
C _q	dimensionless group velocity
Č	constant of eq. (14)
C_{1}, C_{2}	constants of eq. (34)
D	constant of eq. (15)
D_{1}, D_{2}	constants of eq. (35)
f*	frequency, 1/s
$I_n(x)$	modified Bessel function of the first kind and
	the nth-order
Im ()	imaginary part of an indicated variable
j _{on}	roots of $I_0(i, j_{on}) = 0$
k	dimensionless wave-number in eqs (14)-(16),
	(26) and (27)
$k_{1,2}$	dimensionless wave-numbers in eqs (20), (24)
	and (28)
k _{3.4n}	dimensionless wave-numbers in eq. (21)
k5	dimensionless wave-number in eq. (22)
<i>O</i> ()	order of terms of the indicated variable
P	numerator of eqs (32) and (33)
$P_{i,i}$	constants of eq. (42); subscript i indicates the
	ith order and subscript j the number of the
	constant
P_{1}, P_{2}	constants of eq. (30)
Q_1, Q_2	constants of eq. (31)

- dimensionless radial distance
- r* radial distance, m

- nozzle radius, characteristic length, m r_0^*
- R dimensionless local jet radius
- R* local jet radius, m
- R' boundary transformation

Re Reynolds number
$$\left(Re = \frac{\rho^* v_0^* r_0^*}{\mu^*}\right)$$

- real part of the indicated variable Re(...) t^* time, s
- t_0^* characteristic time, s
- t_1^* characteristic time-scale, s
- dimensionless minimum break-up time T_E
- velocity component in axial direction, m/s u*
- v^* velocity component in radial direction, m/s
- v_0^* uniform velocity at the nozzle, m/s

- $\boldsymbol{\nu}$ denominator in eqs (32) and (33)
- V^* experimental voltage, V

Weber number
$$\left(We=\frac{\rho}{W}\right)$$

- dimensionless axial distance
- z* axial distance, m
- Z_B^* break-up length, m
- Z_{Bmin} dimensionless minimum break-up length

δ_0	dimensionless disturbance-amplitude
η_i	th-order harmonic surface distortion
θ_1, θ_2	constants of eqs (30), (31), (32), (33) and (42)
θ_{2i}, θ_{3i}	constants of eqs (34), (35) and (42)
λ*	wavelength, m
μ*	viscosity, Pa s
ρ*	density, kg/m ³
σ*	surface tension, N/m
Φ	dimensionless velocity potential
Ф,	ith-order harmonic velocity potential
ω, ω1	dimensionless frequency
	dimonsionlass fragmanon with maximum

- ω_{opt} dimensionless maximum Irequency with growth rate
- singular value of the dimensionless fre- ω_{s} quency

Subscripts

- B break-up value
- derivative with respect to the radial variable r r
- t derivative with respect to the time t
- z derivative with respect to the axial variable z

Superscripts

*

- H homogeneous part of the solution
- I imaginary part
- R real part
 - dimensional quantity

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APPENDIX 1

The third-order solution for the jet surface is

 $R = 1 + 2\delta_0 [P_{1,1} \operatorname{Im}(\cos \theta_1)]$

$$+\delta_0^2 \operatorname{Re}(\{P_{2,1}\cos\theta_{21}+P_{2,2}\cos\theta_{22}+2P_{2,3}\cos2\theta_1\})$$

$$+P_{2,5}\cos[2(\omega_{1}t-k^{R}z)]+P_{2,6}[\cosh(2k^{I}z)-1]\})$$

+ $\delta_{0}^{3} \operatorname{Im}\left(\left\{\sum_{i=1}^{2}P_{3,i}\cos\theta_{3i}+2\sum_{i=1}^{2}[P_{3,4i}\cos(\theta_{2i}+\theta_{1})+P_{3,(4i+1)}\cos(\theta_{2i}-\theta_{1})]+2P_{3,12}\cos(3\theta_{1})+2P_{3,14}\cos(2\theta_{1}+\theta_{2})+2P_{3,15}\cos(2\theta_{1}-\theta_{2})+(2P_{3,18}+P_{3,20})\cos\theta_{1}+P_{3,21}\cos\theta_{2}\right\}\right)$ (42)

with $\theta_1 = \omega_1 t - k_1 z$ and $\theta_2 = \omega_1 t - k_2 z$.

The subscripts *i* and *j* of the constants $P_{i,j}$ indicate the *i*th order and the number of the constant, respectively.

To obtain numerical results for the analytical solution of the jet-surface it is necessary to calculate the k values for the first-, second- and third-order equations using ω_1 and the Weber number as parameters. The calculation of the various constants $P_{i,j}$ is the next step. After that the jet surface R can be determined for different amplitudes δ_0 and time t in a given distance interval Δz . From these results computer-plots can be made. A description of the software is given by Busker (1983).