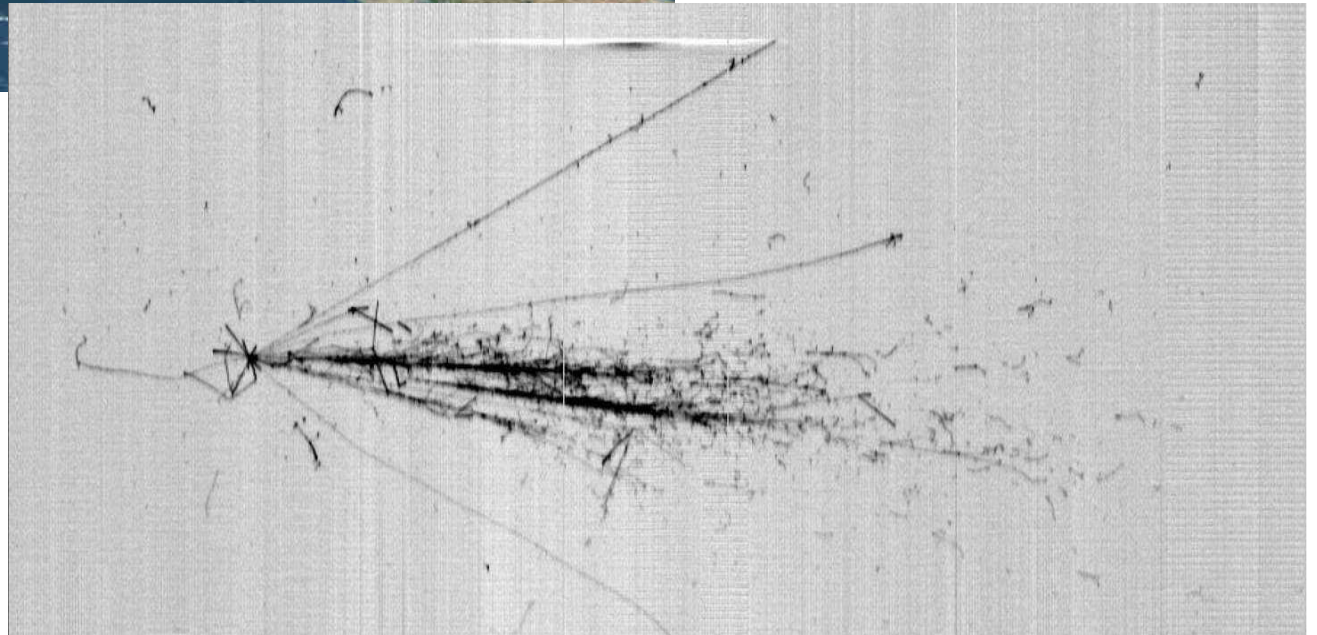


# NEUTRINOS FROM ICARUS



C. Farnese  
For the ICARUS Collaboration

Vulcano Workshop 2012  
1 June 2012



# The ICARUS Collaboration

M. Antonello<sup>a</sup>, P. Aprili<sup>a</sup>, B. Baibussinov<sup>b</sup>, P. Benetti<sup>c</sup>, E. Calligarich<sup>c</sup>, N. Canci<sup>a</sup>, S. Centro<sup>b</sup>, A. Cesana<sup>f</sup>, K. Cieslik<sup>g</sup>, D. B. Cline<sup>h</sup>, A.G. Cocco<sup>d</sup>, A. Dabrowska<sup>g</sup>, D. Dequal<sup>b</sup>, A. Dermenev<sup>i</sup>, R. Dolfini<sup>c</sup>, C. Farnese<sup>b</sup>, A. Fava<sup>b</sup>, A. Ferrari<sup>j</sup>, G. Fiorillo<sup>d</sup>, D. Gibin<sup>b</sup>, S. Gninenko<sup>i</sup>, A. Guglielmi<sup>b</sup>, M. Haranczyk<sup>g</sup>, J. Holeczek<sup>l</sup>, A. Ivashkin<sup>i</sup>, J. Kisiel<sup>l</sup>, I. Kochanek<sup>l</sup>, J. Lagoda<sup>m</sup>, S. Mania<sup>l</sup>, G. Mannocchi<sup>n</sup>, A. Menegolli<sup>c</sup>, G. Meng<sup>b</sup>, C. Montanari<sup>c</sup>, S. Otwinowski<sup>h</sup>, L. Periale<sup>n</sup>, A. Piazzoli<sup>c</sup>, P. Picchi<sup>n</sup>, F. Pietropaolo<sup>b</sup>, P. Plonski<sup>o</sup>, A. Rappoldi<sup>c</sup>, G.L. Raselli<sup>c</sup>, M. Rossella<sup>c</sup>, C. Rubbia<sup>a,j</sup>, P. Sala<sup>f</sup>, E. Scantamburlo<sup>e</sup>, A. Scaramelli<sup>f</sup>, E. Segreto<sup>a</sup>, F. Sergiampietri<sup>p</sup>, D. Stefan<sup>a</sup>, J. Stepaniak<sup>m</sup>, R. Sulej<sup>m,a</sup>, M. Szarska<sup>g</sup>, M. Terrani<sup>f</sup>, F. Varanini<sup>b</sup>, S. Ventura<sup>b</sup>, C. Vignoli<sup>a</sup>, H. Wang<sup>h</sup>, X. Yang<sup>h</sup>, A. Zalewska<sup>g</sup>, K. Zarembo<sup>o</sup>.

*a Laboratori Nazionali del Gran Sasso dell'INFN, Assergi (AQ), Italy*

*b Dipartimento di Fisica e INFN, Università di Padova, Via Marzolo 8, I-35131 Padova, Italy*

*c Dipartimento di Fisica Nucleare e Teorica e INFN, Università di Pavia, Via Bassi 6, I-27100 Pavia, Italy*

*d Dipartimento di Scienze Fisiche, INFN e Università Federico II, Napoli, Italy*

*e Dipartimento di Fisica, Università di L'Aquila, via Vetoio Località Coppito, I-67100 L'Aquila, Italy*

*f INFN, Sezione di Milano e Politecnico, Via Celoria 16, I-20133 Milano, Italy*

*g Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Science, Krakow, Poland*

*h Department of Physics and Astronomy, University of California, Los Angeles, USA*

*i INR RAS, prospekt 60-letiya Oktyabrya 7a, Moscow 117312, Russia*

*j CERN, CH-1211 Geneve 23, Switzerland*

*k Institute of Theoretical Physics, Wroclaw University, Wroclaw, Poland*

*l Institute of Physics, University of Silesia, 4 Uniwersytecka st., 40-007 Katowice, Poland*

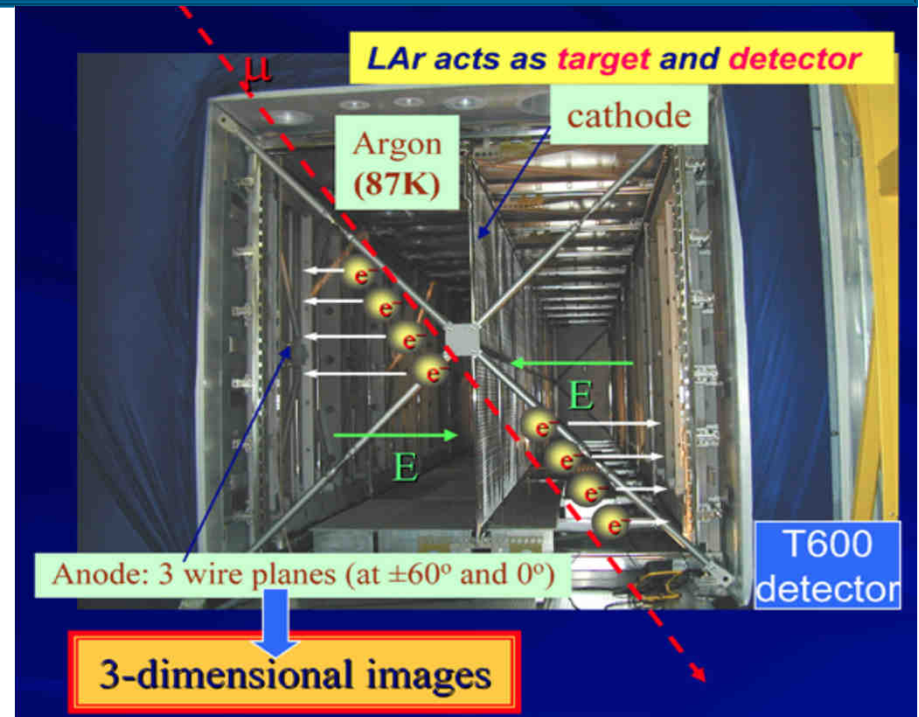
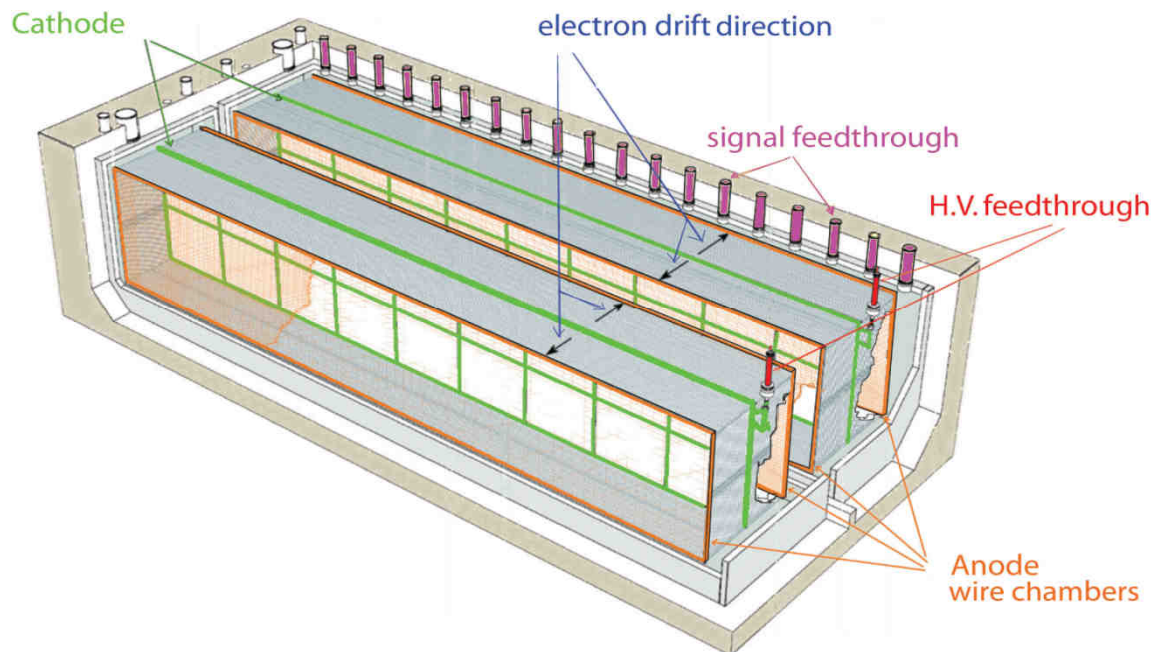
*m National Centre for Nuclear Research, A. Soltana 7, 05-400 Otwock/Swierk, Poland*

*n Laboratori Nazionali di Frascati (INFN), Via Fermi 40, I-00044 Frascati, Italy*

*o Institute of Radioelectronics, Warsaw University of Technology, Nowowiejska, 00665 Warsaw, Poland*

*p INFN, Sezione di Pisa. Largo B. Pontecorvo, 3, I-56127 Pisa, Italy*

# The ICARUS T600 detector

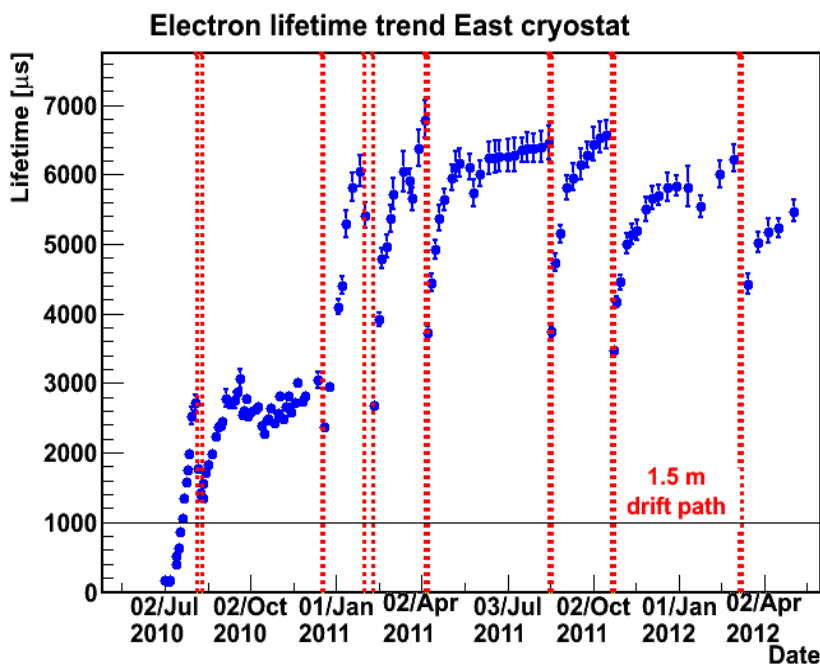
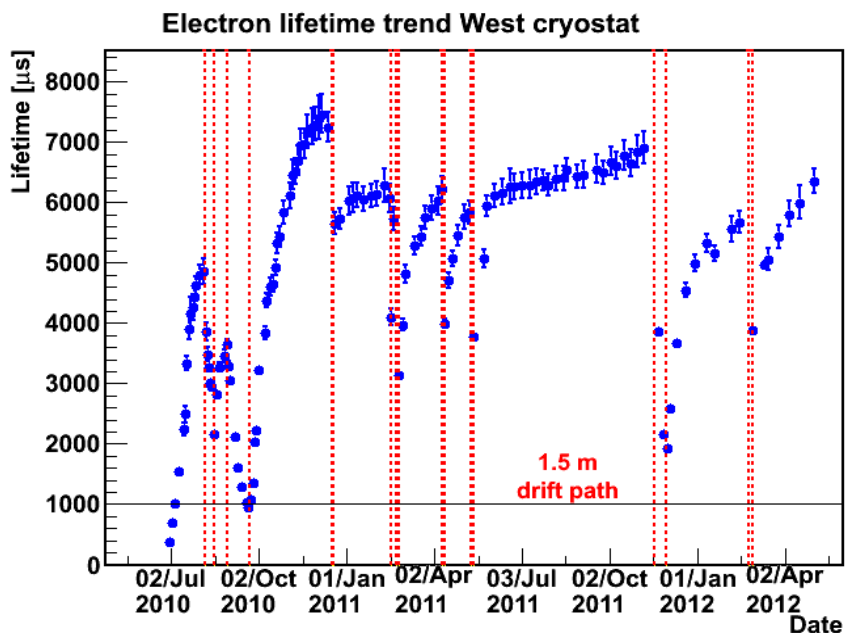


- Two identical modules
  - $3.6 \times 3.9 \times 19.6 \approx 275 \text{ m}^3$  each
  - Liquid Ar active mass:  $\approx 476 \text{ t}$
  - Drift length = 1.5 m (1 ms)
  - HV = -75 kV     $E = 0.5 \text{ kV/cm}$
  - v-drift = 1.55 mm/ $\mu\text{s}$
- 4 wire chambers:
  - 2 chambers per module
  - 3 readout wire planes per chamber, wires at  $0, \pm 60^\circ$
  - $\approx 54000$  wires, 3 mm pitch, 3 mm plane spacing
- 20+54 PMTs , 8"  $\varnothing$  , for scintillation light detection: VUV sensitive (128nm) with wave shifter (TPB)

Taking data in LNGS hall B

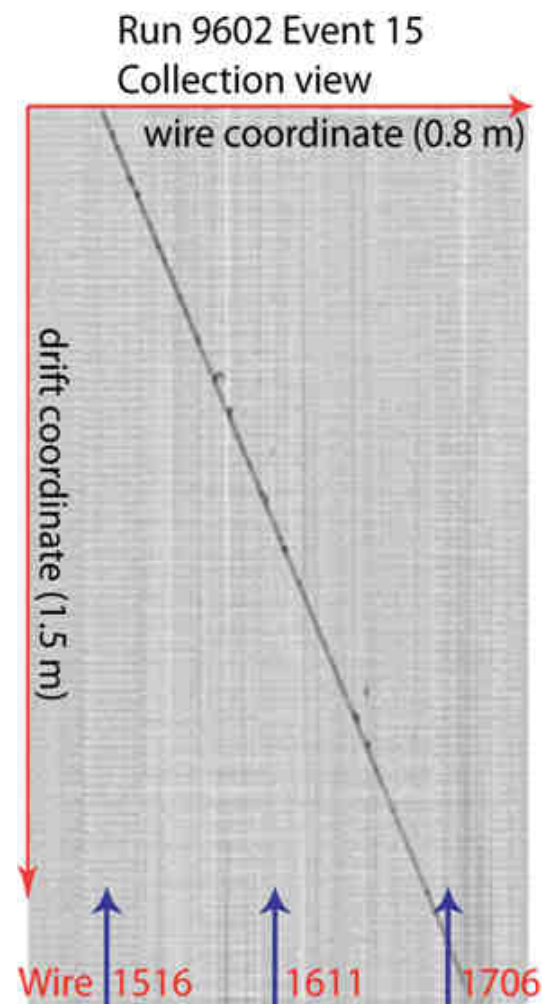
# LAr purification

Key feature: LAr purity from electro-negative molecules ( $O_2$ ,  $H_2O$ ,  $CO_2$ ).  
LAr continuously filtered,  $e^-$  life-time measured by charge attenuation study on cosmic  $\mu$  track



$\tau_{ele} > 5ms$   
( $\sim 60$  ppt  $[O_2]_{eq}$ )

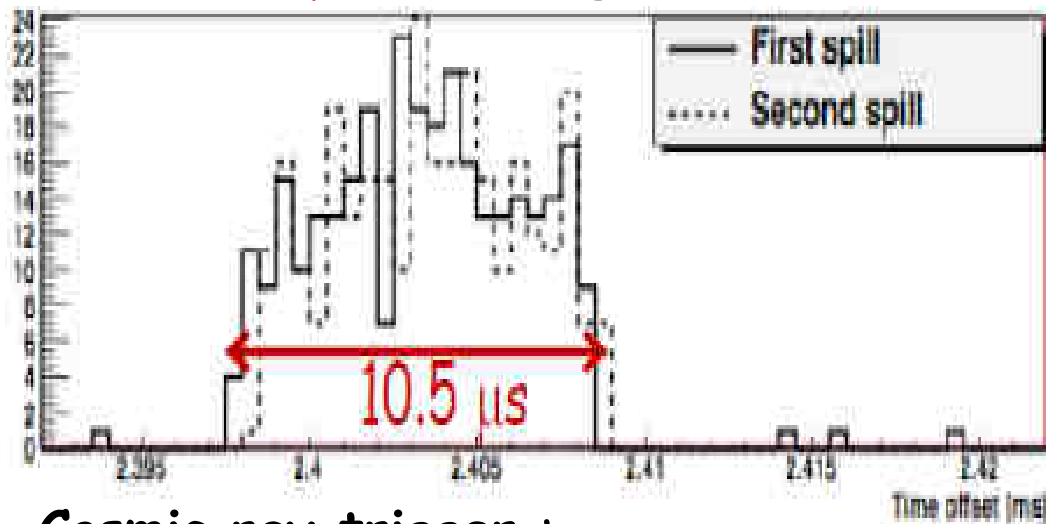
corresponding to 17% max. charge attenuation at 1.5m  
These results would allow operation at larger drift distances.



# Trigger

## CNGS trigger :

- CNGS "Early Warning" signal sent 80 ms before the SPS p extraction. It contains the predicted extraction time of the 2 spills → opens a 60  $\mu$ s wide gate.
- **Photomultiplier** sum signal for each chamber in coincidence with the beam gate



- 2.40 ms offset value in agreement with 2.44 ms  $v$  tof (40  $\mu$ s fiber transit time from external lab to Hall B).
- Spill duration reproduced (10.5 $\mu$ s)
- 1 MHz event rate ,  $\approx$  80 events/day

## Cosmic ray trigger :

- **Photomultiplier** sum signal, requiring coincidence of two adjacent chambers (50% cathode transparency)
- Globally 36 MHz trigger rate achieved:  $\sim$ 130 cosmic events/h

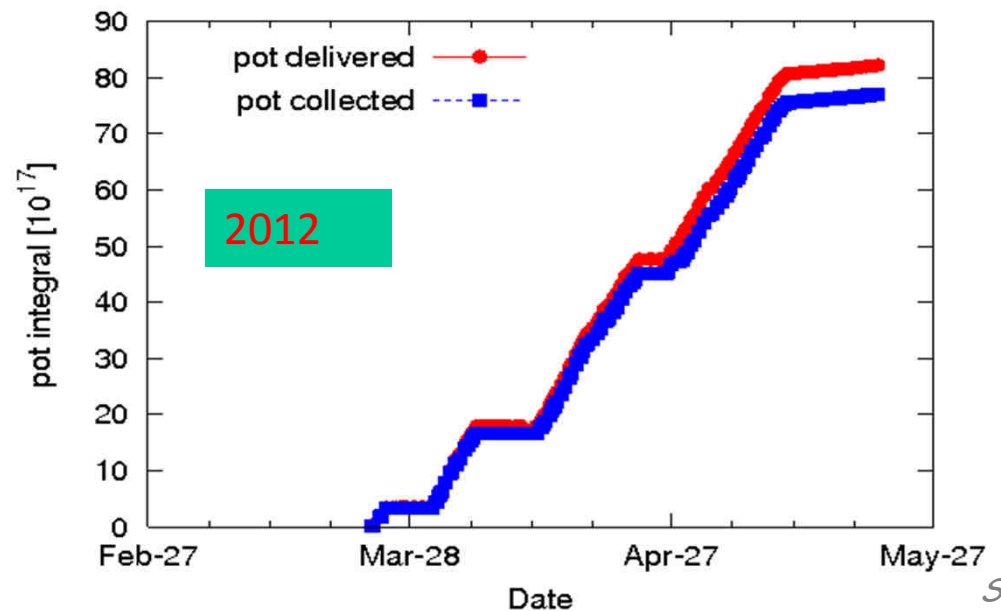
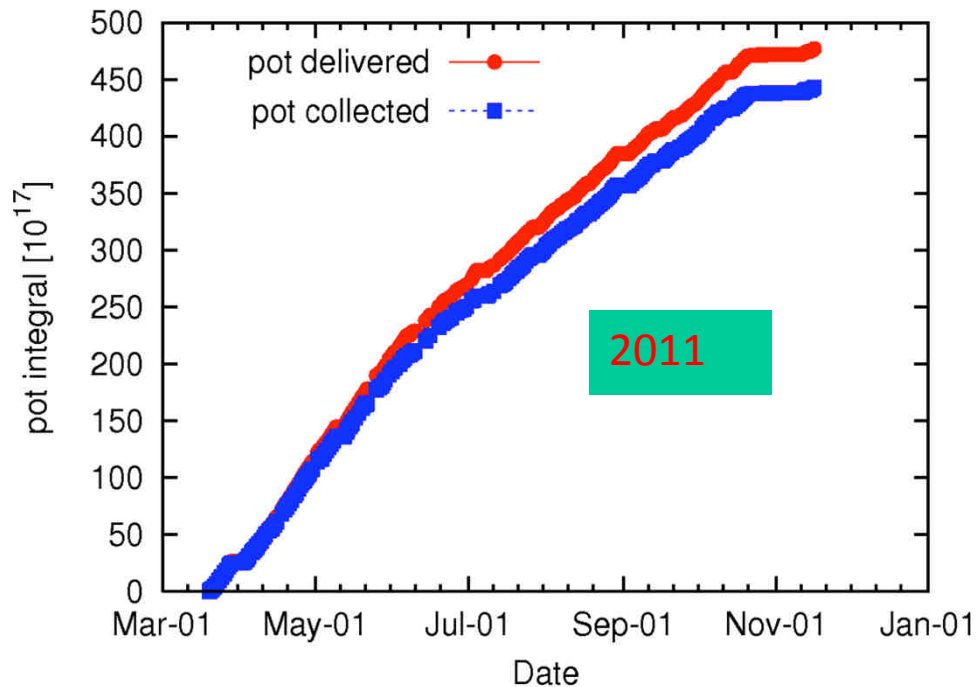
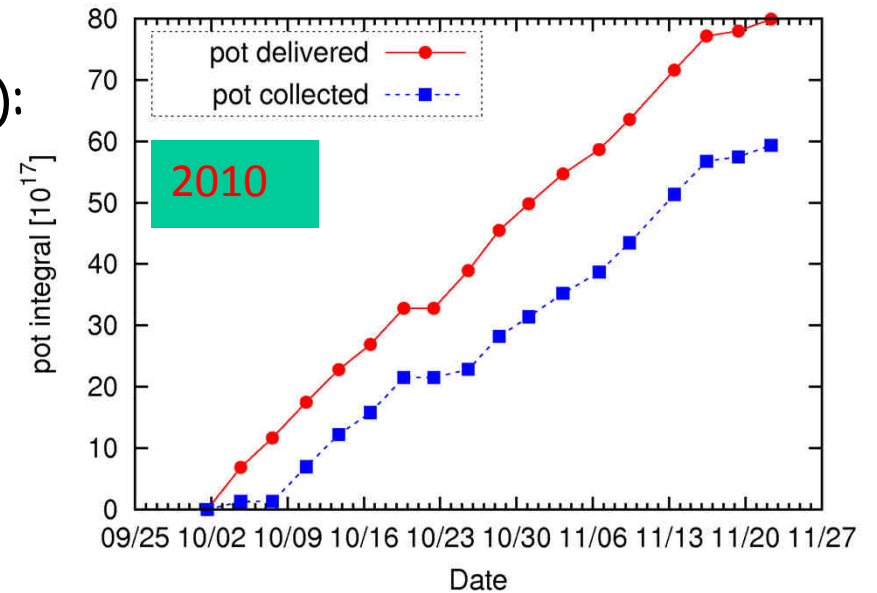
**SuperDedalus** : New trigger system based on charge deposition on TPC wires: DR-slw algorithm implemented in a new SuperDedalus chip (FPGA), installed and used to improve trigger efficiency for CNGS events at low energy (i.e. below 500 MeV)

# ICARUS T600 physics potential

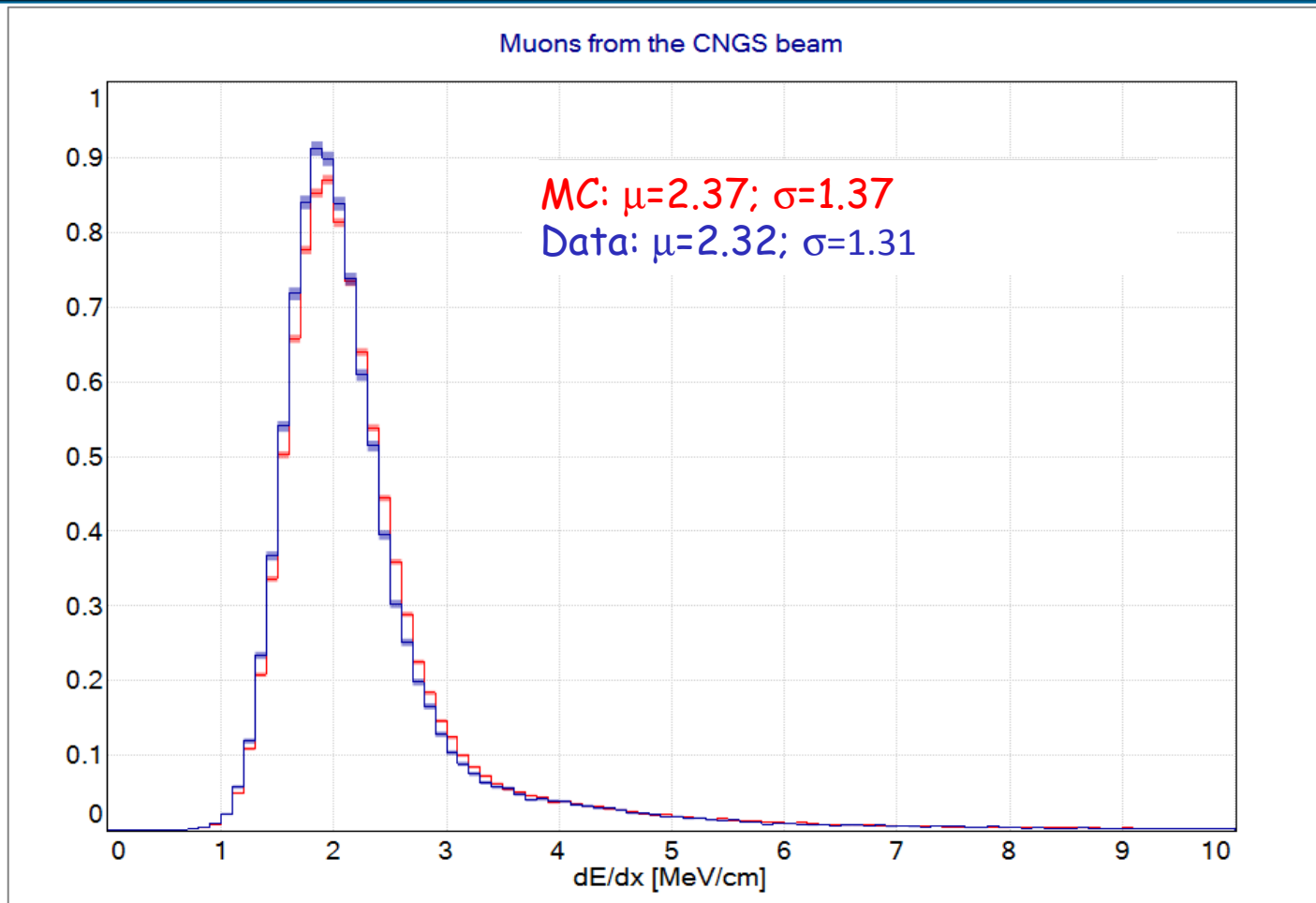
- T600 is a major milestone towards the realization of a much more massive multikton LAr detector, **but it offers also some interesting physics in itself.** The unique imaging capability of ICARUS, its spatial/calorimetric resolutions, and  $e/\pi^0$  separation allow "to see" events in a new way
- The detector is collecting "bubble chamber like" CNGS events: for  $10^{20}$  pot
  - CC event expected  $\approx 2800$  ev
  - NC event expected  $\approx 900$  ev
  - Muons from upstream GS rock  $\approx 12000$  ev ( $\approx 8200$  on TPC front face)
  - Intrinsic beam  $\nu_e$  CC  $\approx 26$  ev
  - $\nu_\mu \Rightarrow \nu_\tau$  detecting  $\tau$  decay with kinematical criteria ( $\sim 2$  event  $\tau \rightarrow e$ ).
  - $\nu_\mu \Rightarrow \nu_e$  ( $\theta_{13}$ ) from e-like CC events excess at  $E < 20\text{GeV}$  ( $\sim 5$  events CC)
  - Search for sterile neutrinos in LSND parameter space, with e-like CC events excess at  $E > 10\text{GeV}$ .
- The T600 is also collecting simultaneously "self triggered" events:
  - $\approx 100$  ev/year of atmospheric  $\nu$  CC interactions.
  - Proton decay with  $3 \times 10^{32}$  nucleons, zero bckg. in some of the channels

# CNGS neutrino runs

- ICARUS fully operational since Oct. 1<sup>st</sup> 2010:  $5.8 \times 10^{18}$  pot collected in 2010.
- 2011 CNGS run: Mar. 19<sup>th</sup> to Nov. 14<sup>th</sup>
  - $4.44(4.78) \times 10^{19}$  pot collected (delivered):  
**93% detector live-time**
  - Expected  $\sim 1280$  CC and  $\sim 395$  NC events
- March 23<sup>rd</sup> 2012: new CNGS events,  
 $7.7 \times 10^{18}$  pot collected.



# Calibration with CNGS muons



$dE/dx$  distribution  
for real and MC  
muon tracks from  
CNGS events

Tracks reconstructed in 3D.  $\delta$  rays and showers rejected.

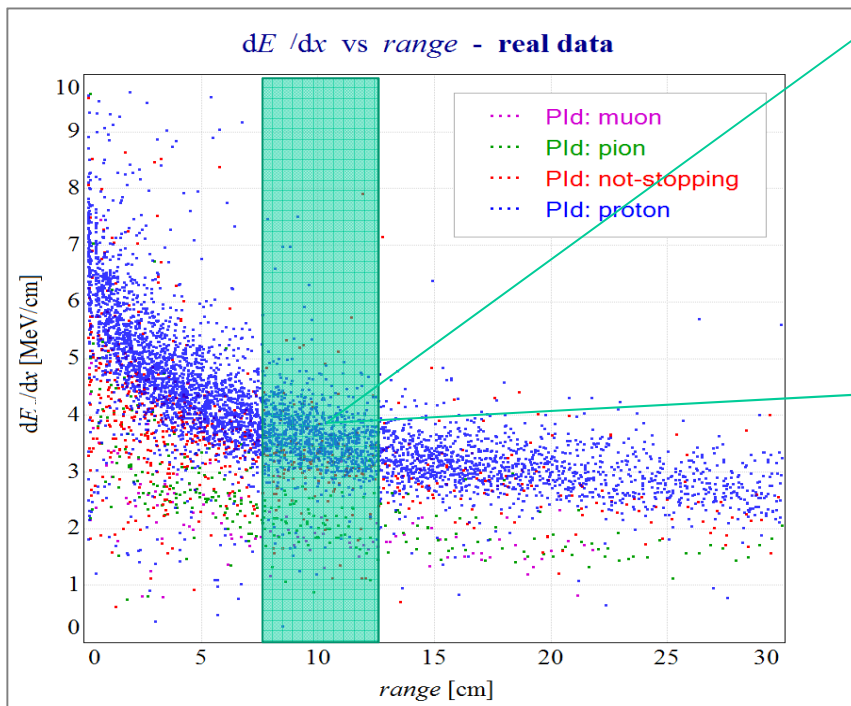
Same reconstruction on MC muons with CNGS spectrum.

Very nice agreement ( $\sim 2-3\%$ ) - still possible small different conditions of data and MC (noise patterns and their effects on  $\delta$  ray selection).

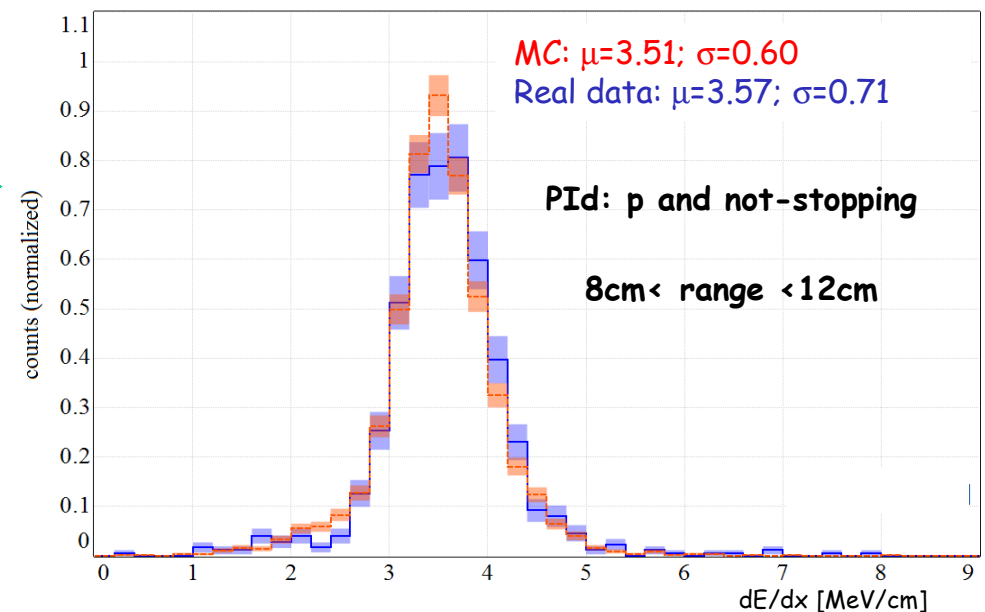
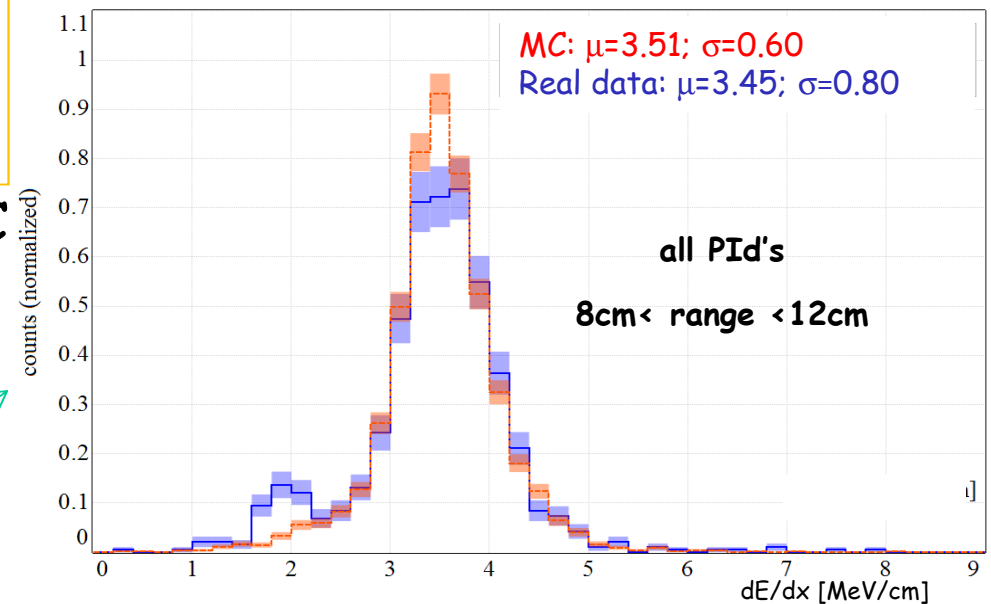


# Study of stopping tracks

- Deposited  $dE/dx$  vs residual range
- No quenching corr. for  $dE$  estimation
- Residual range between 8 and 12 cm
- Good agreement between Data and MC
- $\pi$  clearly separated from protons
- MC: only protons are considered

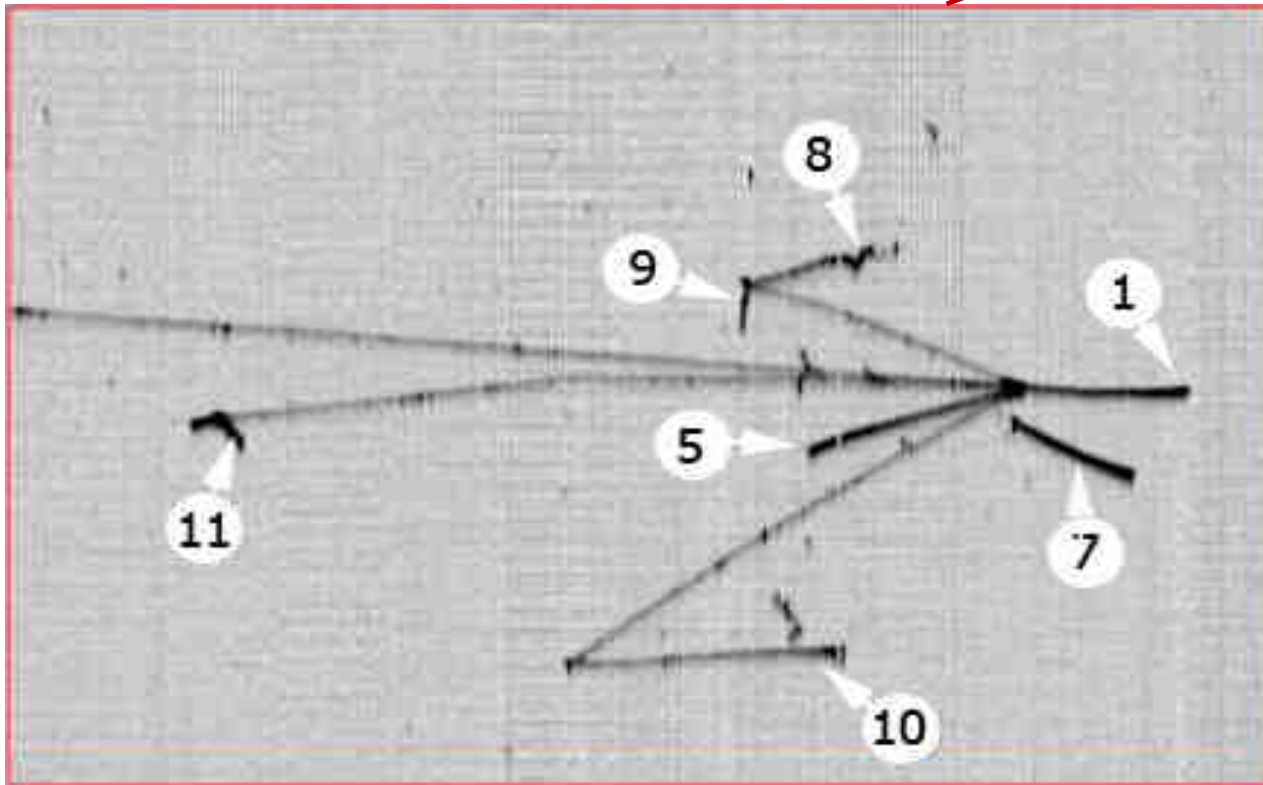
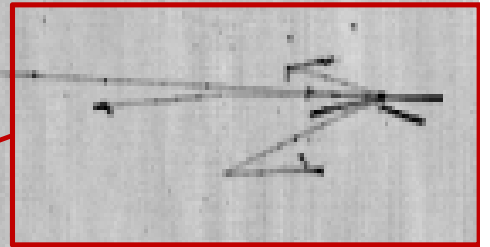


Methods for identification of non-stopping particles are under development



# $\nu_\mu$ CC CNGS event: reconstruction of stopping tracks

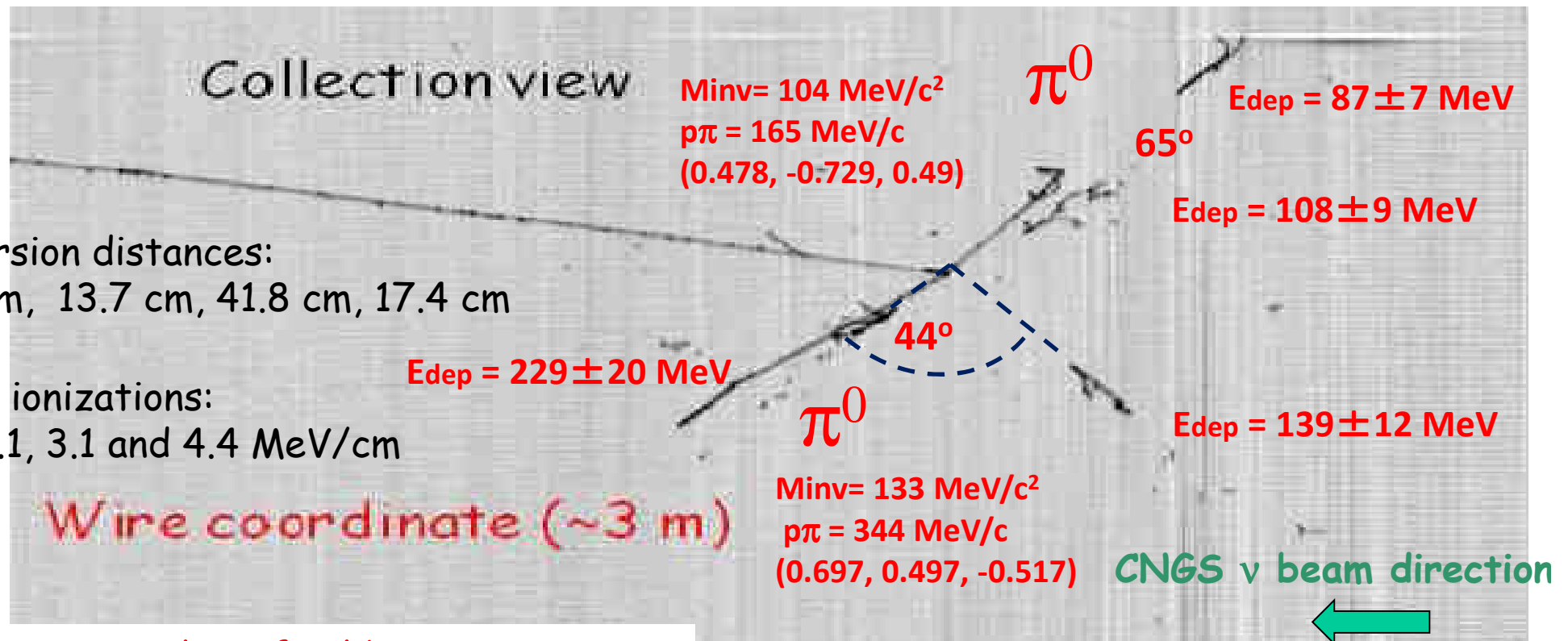
Run 9809 Event 651



Track	$E_{\text{dep}}$ [MeV]	range [cm]
1(p)	$185 \pm 16$	15
5(p)	$192 \pm 16$	20
7(p)	$142 \pm 12$	17
8( $\pi$ )	$94 \pm 8$	12
9(p)	$26 \pm 2$	4
10(p)	$141 \pm 12$	23
11(p)	$123 \pm 10$	6

6 protons, 1 pion decays at rest

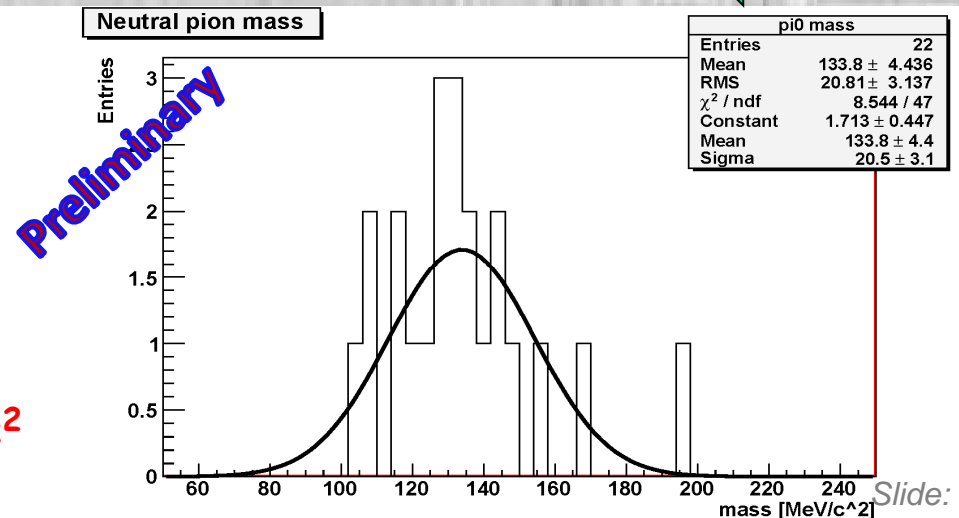
# $\pi^0$ reconstruction in CNGS event



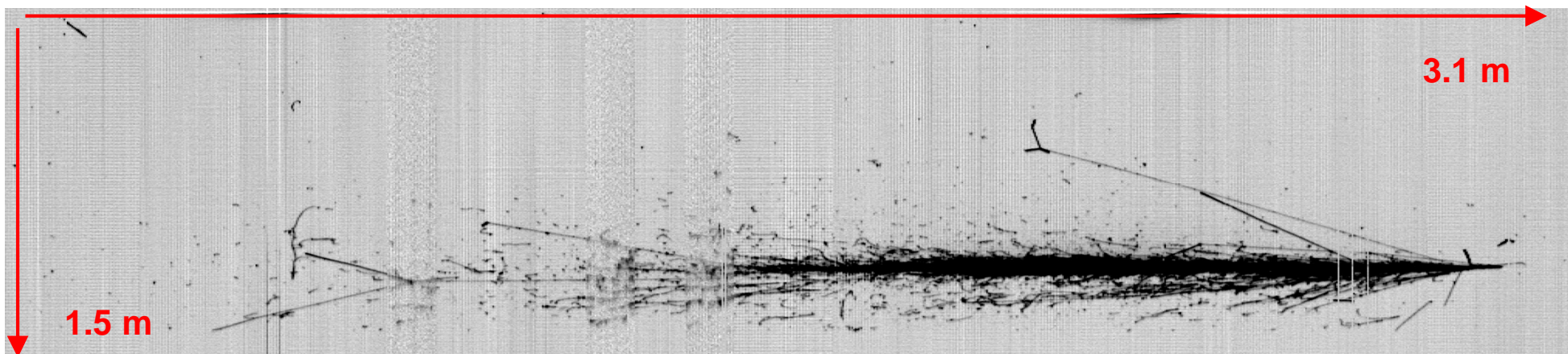
$\pi^0$ -showers identified by

- 2 $\gamma$  conversion separated from primary vertex
- Reconstruction of  $\gamma\gamma$  invariant mass
- Ionization in the first segment of showers (1 mip or 2 mips)

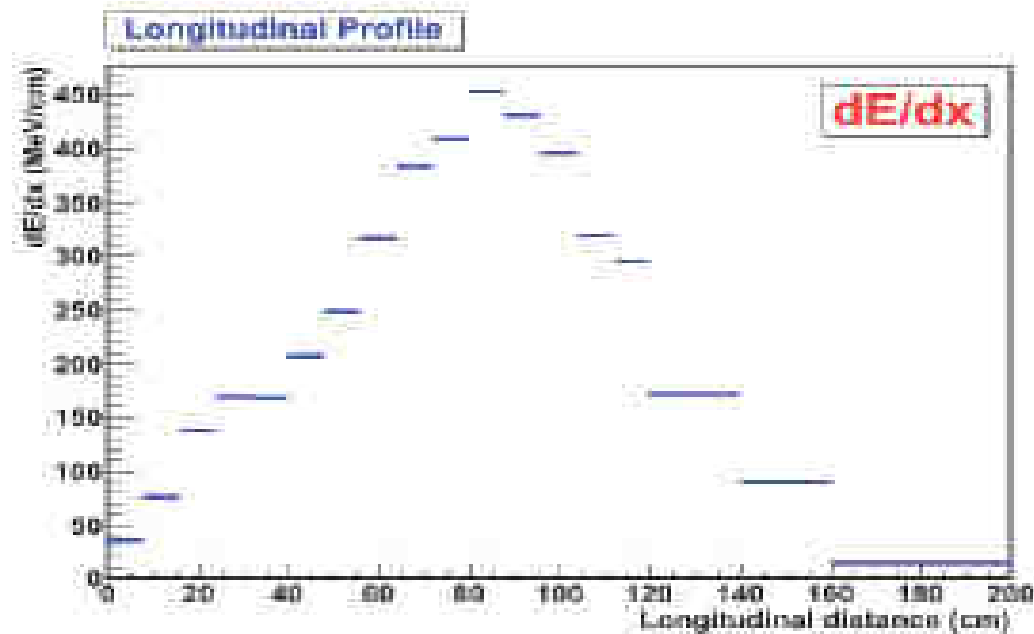
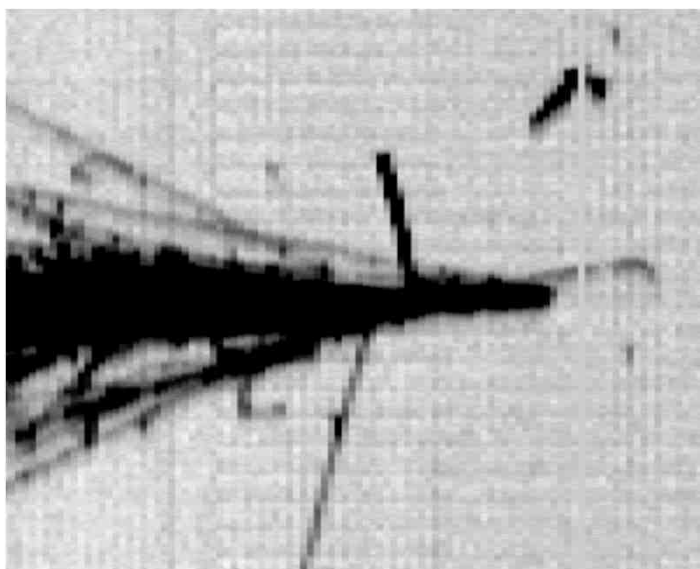
Mean:  $133.8 \pm 4.4(\text{stat}) \pm 4(\text{syst})$  MeV/c<sup>2</sup>  
 $\sigma = 20.5$  MeV



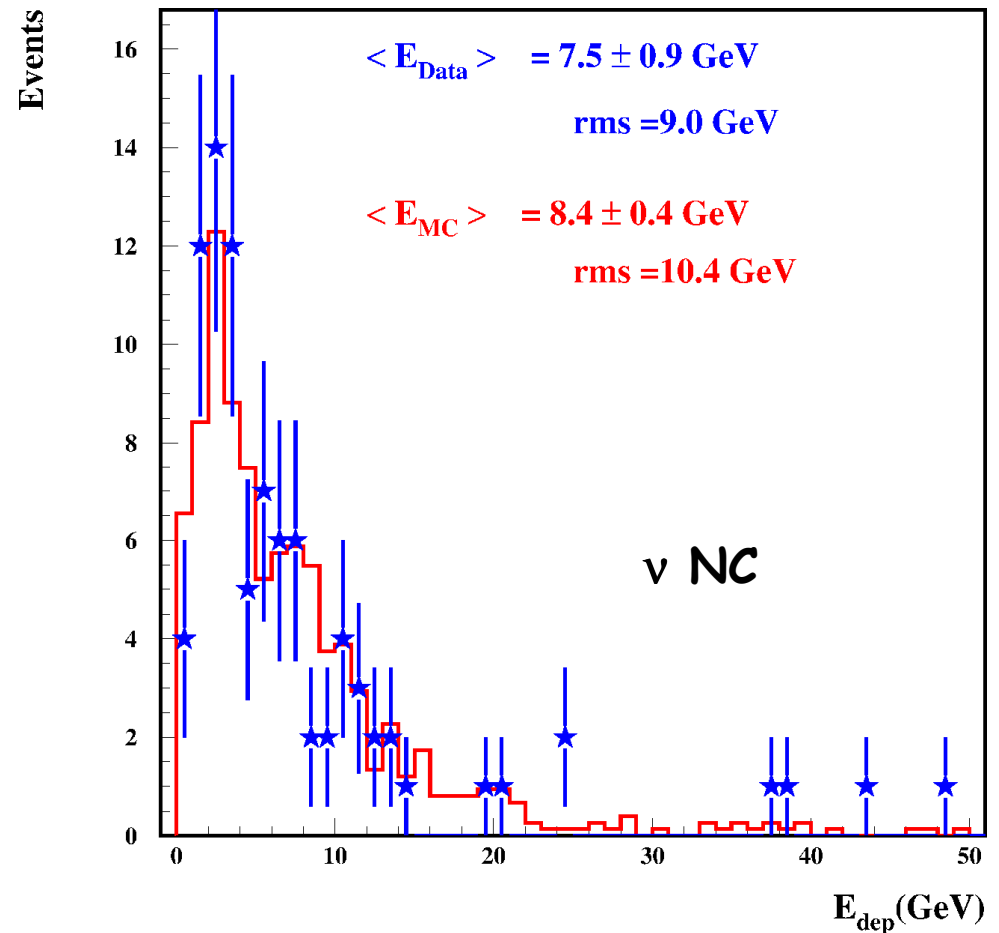
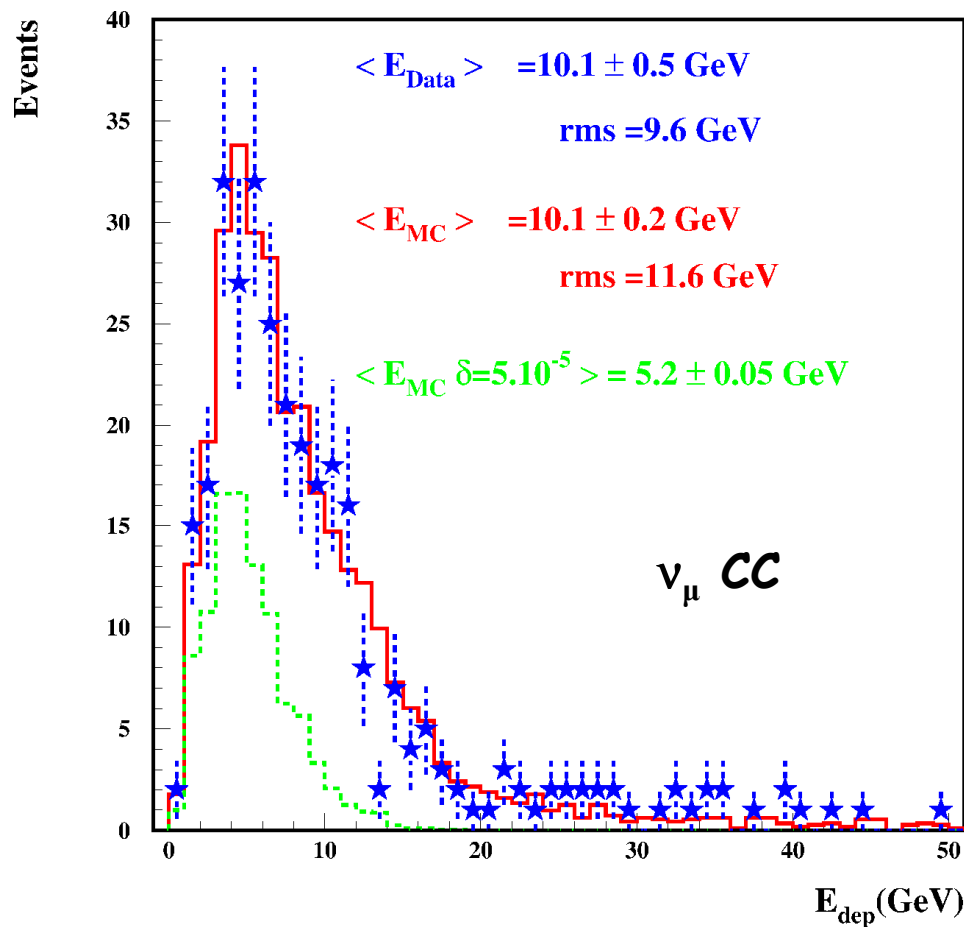
# Electron event candidate



- A  $\nu_e$ CC candidate from 2010 run. This event has  $45 \text{ GeV}$  energy with a single powerful  $37 \text{ GeV}$  e.m. shower at vertex with a longitudinal profile peaking at the expected position ( $\sim 88 \text{ cm}$ ).



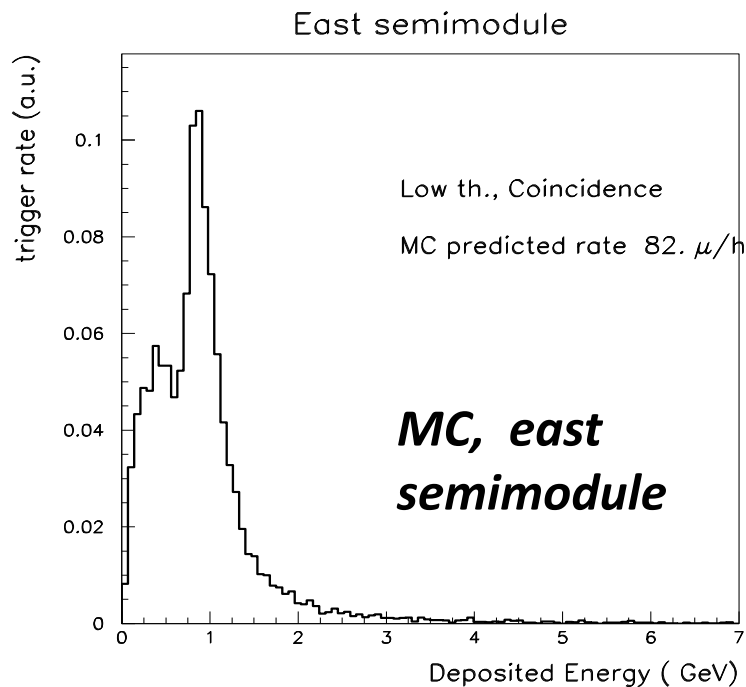
# Total energy deposition in CNGS $\nu$ events



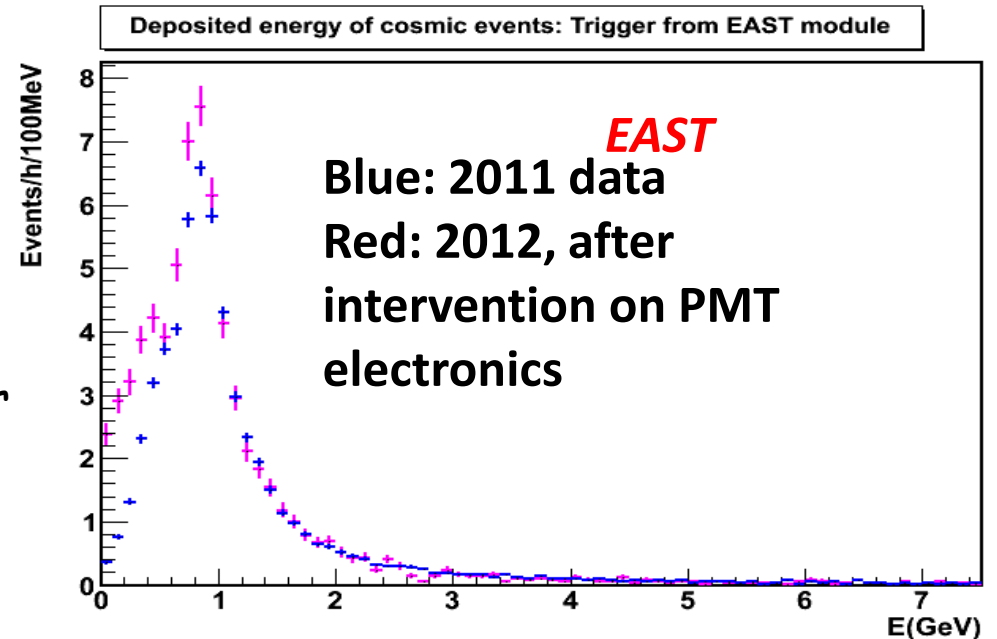
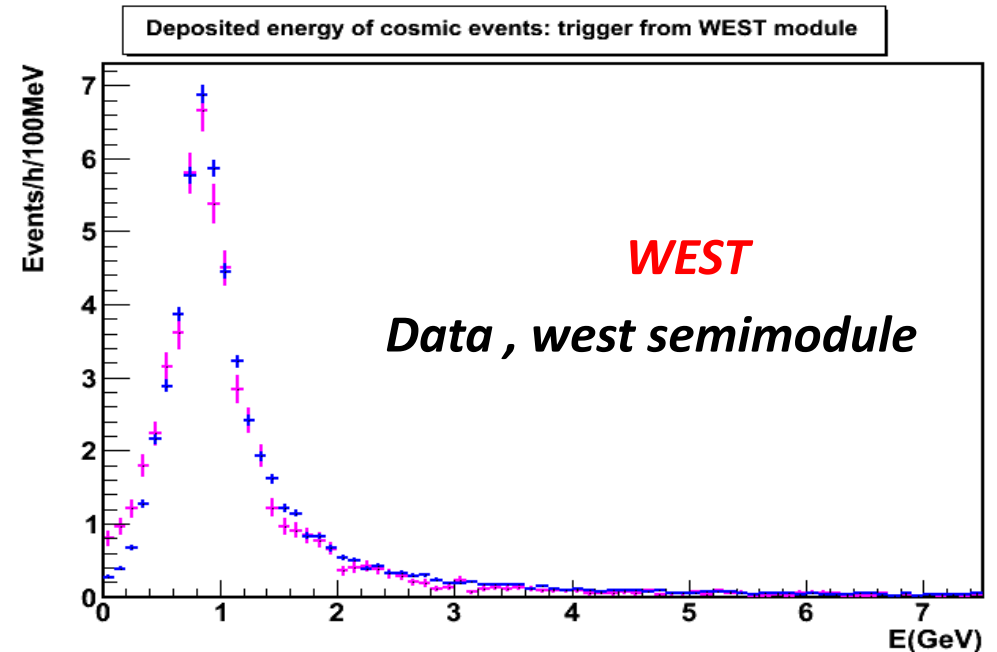
- Comparison of the predicted ( full MC) and detected deposited energy spectrum from NC and CC events on 2010 statistics and a subset of the 2011 statistics.
- Used for the "superluminal" neutrino searches.

# C-ray spectrum

- CR data automatically filtered
- Good agreement of energy spectrum with MC expectation. (MC simulation includes light collection/trigger conditions)

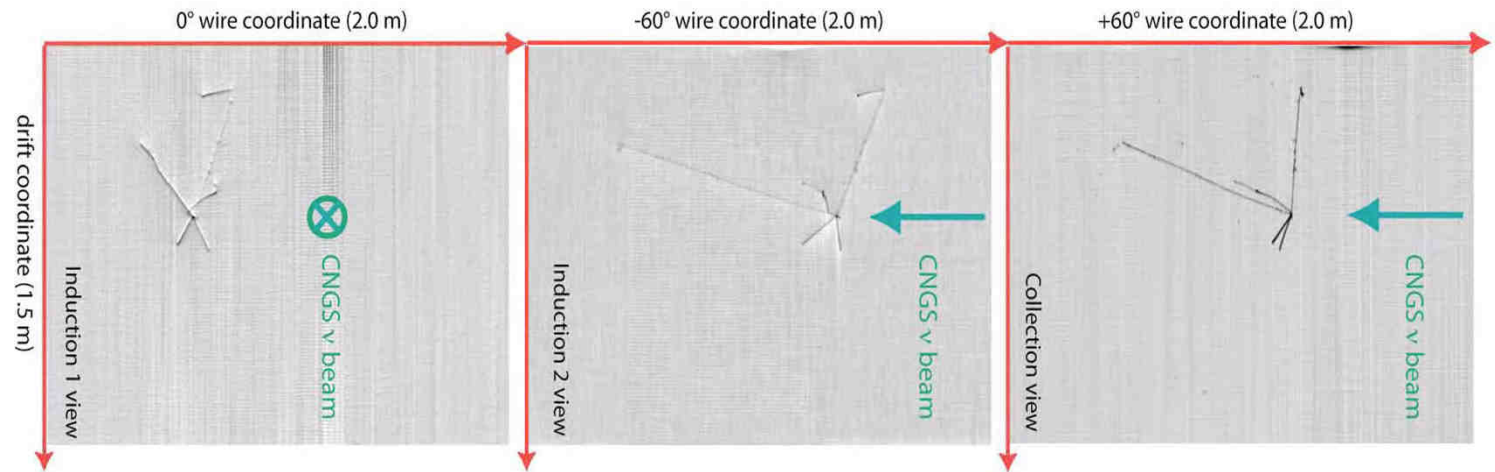


- 2012: 30% improvement w.r.t. 2011 trigger rate thanks to new PMT's HV biasing signal readout: signal amplitude increased  $\rightarrow$  *efficiency at low energy increased*



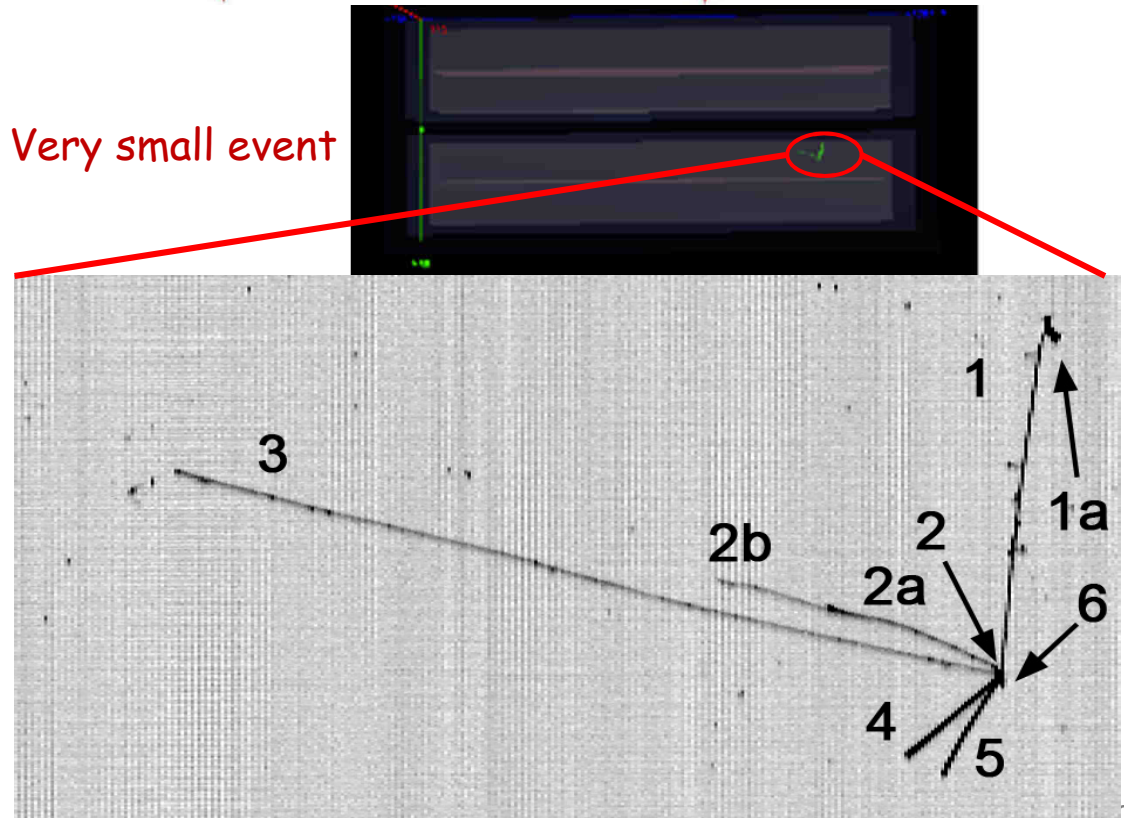
# Atmospheric $\nu$ candidate

- Total visible energy: 887 MeV
- Out-of-time wrt CNGS spill, 35° angle w.r.t. beam direction.



Track	$E_k$ [MeV]	Range [cm]
1 ( $\pi$ , decays in flight)	136.1	55.77
2 ( $\pi$ )	26	3.3
2a ( $\mu$ )	79.1	17.8
2b (e)	24.1	10.4
3 ( $\mu$ )	231.6	99.1
4 (p)	168	19.2
5 (p)	152	16.3
6 (?) (merged with vtx)		2.9

Very small event



# Search for superluminal $\nu$ 's radiative processes in ICARUS

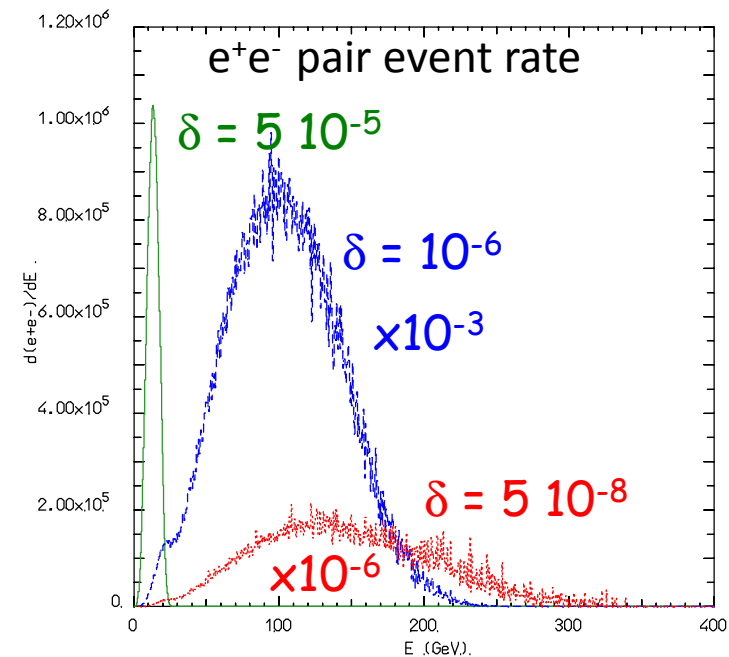
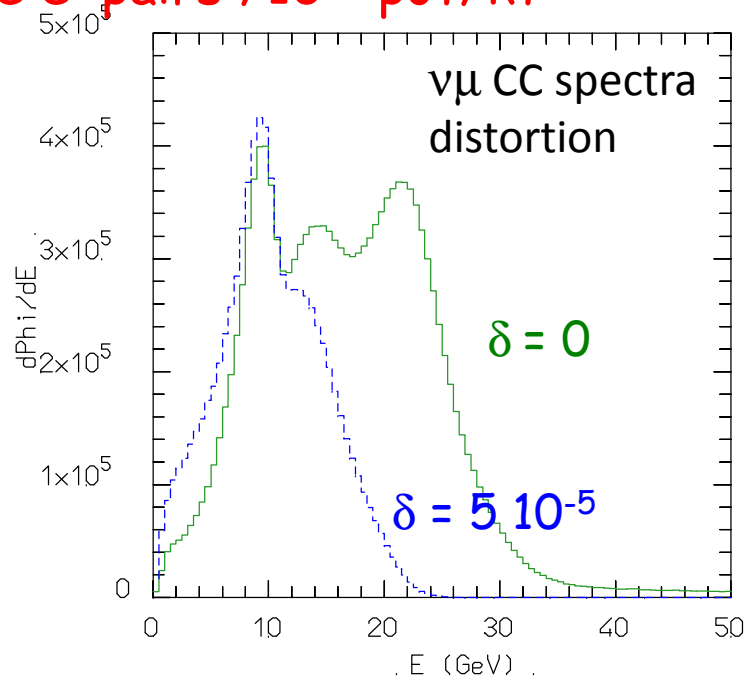
<http://dx.doi.org/10.1016/j.physletb.2012.04.014>

Phys.Lett.B711. (2012) 270-275

- Cohen and Glashow [Phys. Rev. Lett., 107 (2011) 181803] argued that superluminal  $\nu$  should lose energy mainly via  $e^+e^-$  bremsstrahlung, on average  $0.78 \cdot E_\nu$  energy loss/emission
- Full FLUKA simulation of the process kinematics, folded in the CNGS beam, studied as a function of  $\delta = (v_\nu^2 - c^2)/c^2$

For  $\delta = 5 \cdot 10^{-5}$  (OPERA first claim):

- full  $\nu$  event suppression for  $E > 30$  GeV
- $\sim 10^7$   $e^+e^-$  pairs /  $10^{19}$  pot/kt





# Search for superluminal $\nu$ 's radiative processes in ICARUS

Expected  $\nu$  event rate and  $e^+e^-$  pair production spectra for  $10^{19}$  pot $\cdot$ kt of ICARUS exposure and different  $\delta$  values

$\delta$	CC (all flavours)	NC (all flavours)	CC > 60 GeV ( $\nu_\mu + \bar{\nu}_\mu$ )	$e^+e^-$
0	644	203	57	0
$5 \cdot 10^{-8}$	644	203	57	27
$5 \cdot 10^{-7}$	643	203	56	$2.1 \cdot 10^4$
$5 \cdot 10^{-6}$	594	188	8.5	$7.2 \cdot 10^5$
$5 \cdot 10^{-5}$	203	85	$< 10^{-6}$	$1.1 \cdot 10^7$

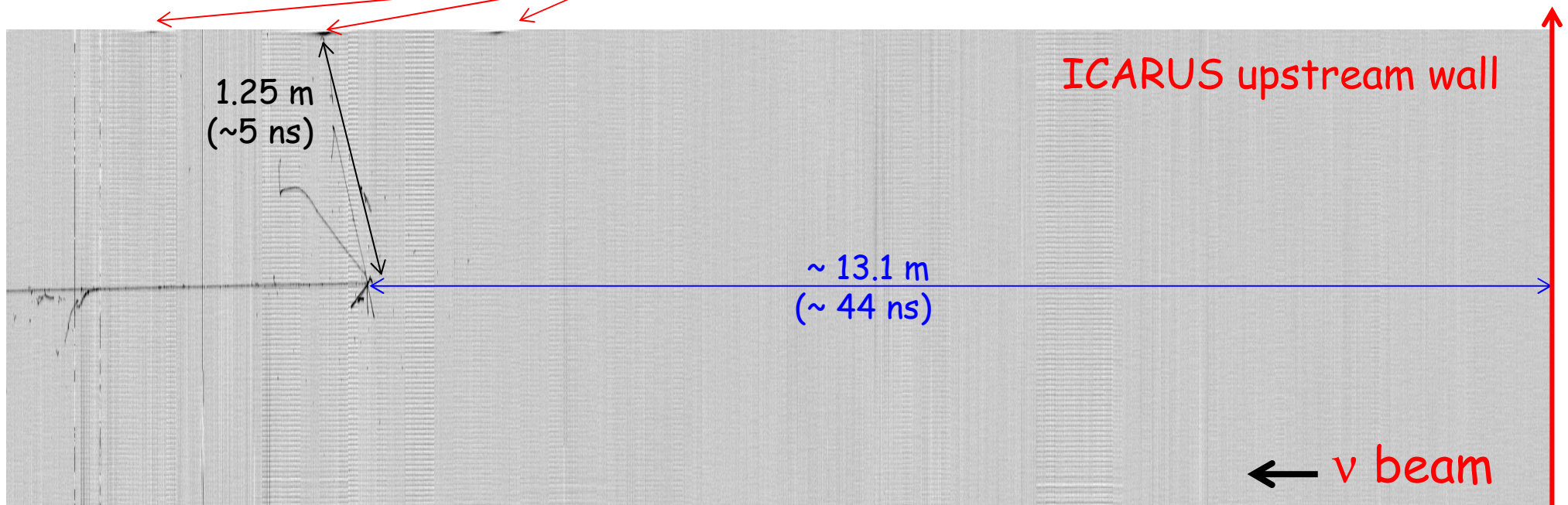
- Effects searched in  $6.7 \cdot 10^{18}$  pot $\cdot$ kt ICARUS exposure (2010/11) to CNGS
  - No spectrum suppression found in both NC, CC data ( $\sim 400$  events)
  - No  $e^+e^-$  pair bremsstrahlung event candidate found
- The lack of pair in CNGS ICARUS 2010/2011 data, sets the limit:

$$\delta = (v_\nu^2 - c^2)/c^2 < 2.5 \cdot 10^{-8} \text{ 90\% CL}$$

- comparable to the SuperK limit  $\delta < 1.4 \cdot 10^{-8}$ , somewhat larger than the lower energy velocity constraint  $\delta < 4 \cdot 10^{-9}$  from SN1987A.

# Neutrino time of flight with 2011 bunched beam

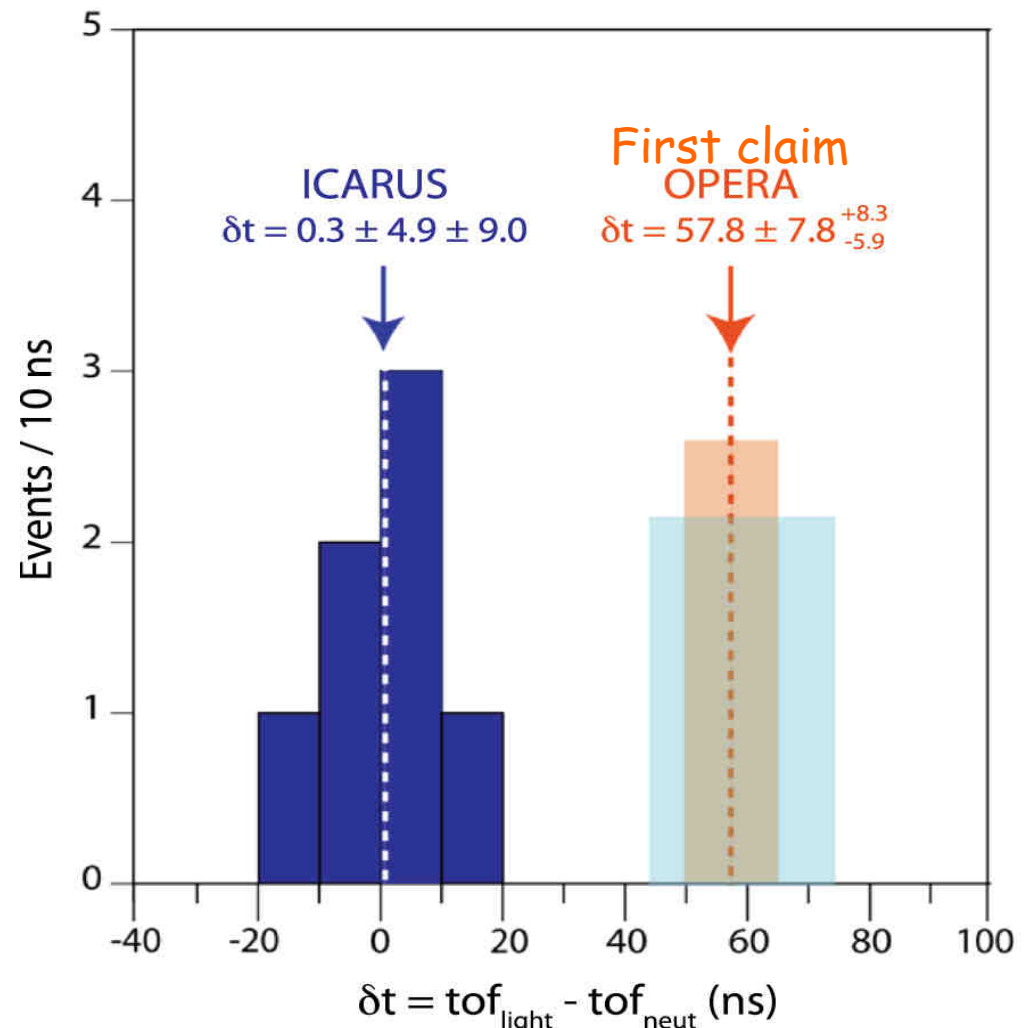
- Bunched beam: 4 bunches/spill, 3 ns FWHM, 524 ns separation
- From October 31<sup>st</sup> to Nov. 5<sup>th</sup> ICARUS observed 7 bunched-beam events
- Timing from ICARUS PMT readout equipped with an independent DAQ
- Reference point : upstream wall of active volume -> corrections needed:
  - the position of interaction vertex along 18 m of detector length
  - the distance of event vertex from closest PMT
- Both corrections precisely ( $\sim 1\text{ns}$ ) deduced from event topology in LAr-TPC through visual scanning.



# Neutrino time of flight result

[http://dx.doi.org/10.1016/j.physletb.2012.05.033\(PhysLettB\)](http://dx.doi.org/10.1016/j.physletb.2012.05.033(PhysLettB))

- All fixed delays/propagation times calibrated
  - Baseline estimation relies on existing available geodesy data (OPERA/LNGS)
  - Variable corrections to GPS from OPERA/CERN recipe
- ↓
- The average  $\delta t = \text{tof}_c - \text{tof}_v$  for the 7 events is  $+0.3 \text{ ns} \pm 4.9 \text{ ns (stat)} \pm 9 \text{ ns (syst)}$
  - $v$  velocity compatible with speed of light



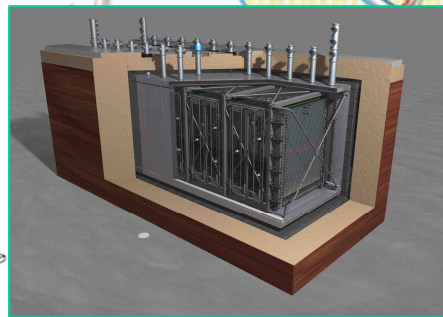
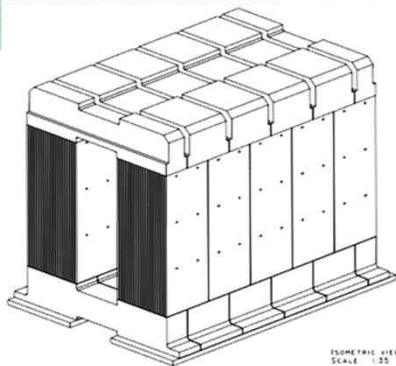
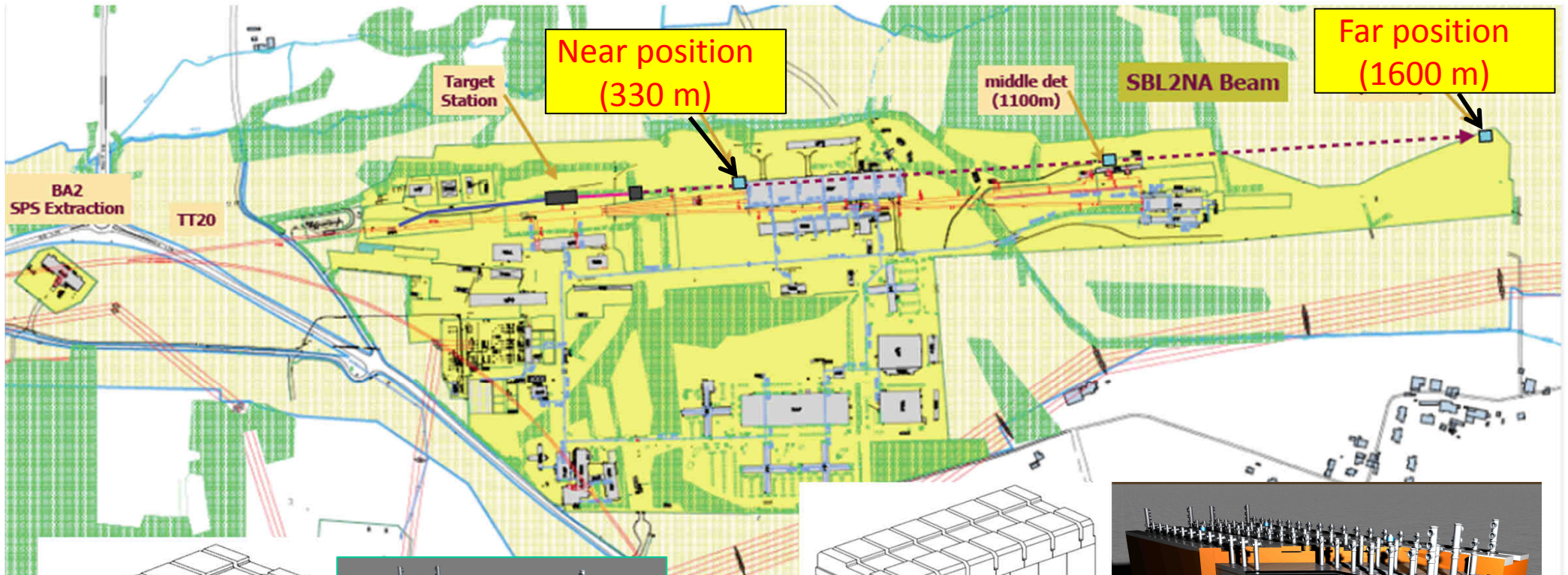
Presently analyzing data with the new bunched beam run,  
Common effort LNGS and CERN , involving Borexino, LVD, Opera, Icarus

# ICARUS after CNGS2: a new approach to sterile $\nu$ at CERN/SPS

The experimental "anomalies" found by LSND/Miniboone (observation of electron *excess* in a anti- $\nu_\mu$  beam from accelerators) and by the reactor neutrino experiments (apparent *disappearance signal* in the anti- $\nu_e$  events) might be due to the presence of "sterile" neutrino

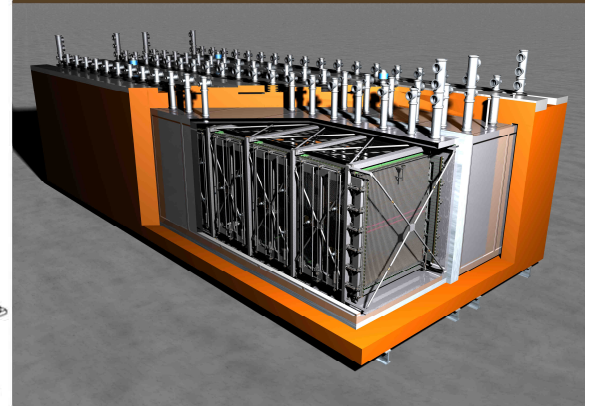
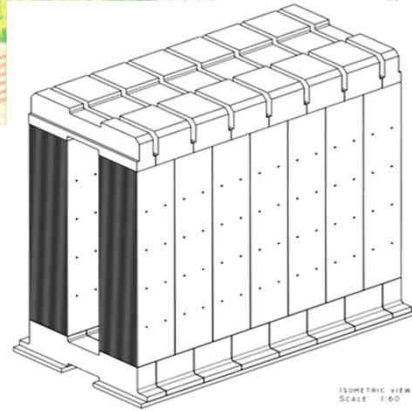
- ❑ The LAr-TPC is the viable device to solve these "anomalies" thanks to
  - detection capability of genuine  $\nu_e$  events
  - energy resolution/detector granularity largely adequate for  $E < 3\text{GeV}$
  - high level of rejection of associated background events ( $\pi^0$ ).
- ❑ A novel experimental search based on two strictly identical LAr-TPC detector + 2 magnetic spectrometers at 330 m and 1600 m from the p target is proposed at CERN - SPS
- ❑ Neutrino beam produced by a 100 GeV proton beam fast extracted from SPS will be centred at  $\sim 2\text{ GeV}$
- ❑ Anti-neutrino beam by inverting the current of the horn
- *Technical proposal: "Search for "anomalies" from neutrino and anti-neutrino oscillations at  $\Delta m^2 \approx 1\text{eV}^2$  with muon spectrometers and large LAr-TPC imaging detectors" (SPSC-P-347) of March 15<sup>th</sup>, 2012.*

# New Neutrino Facility in the CERN North Area



**NEAR**

New detector T150 identical to ICARUS but of smaller size

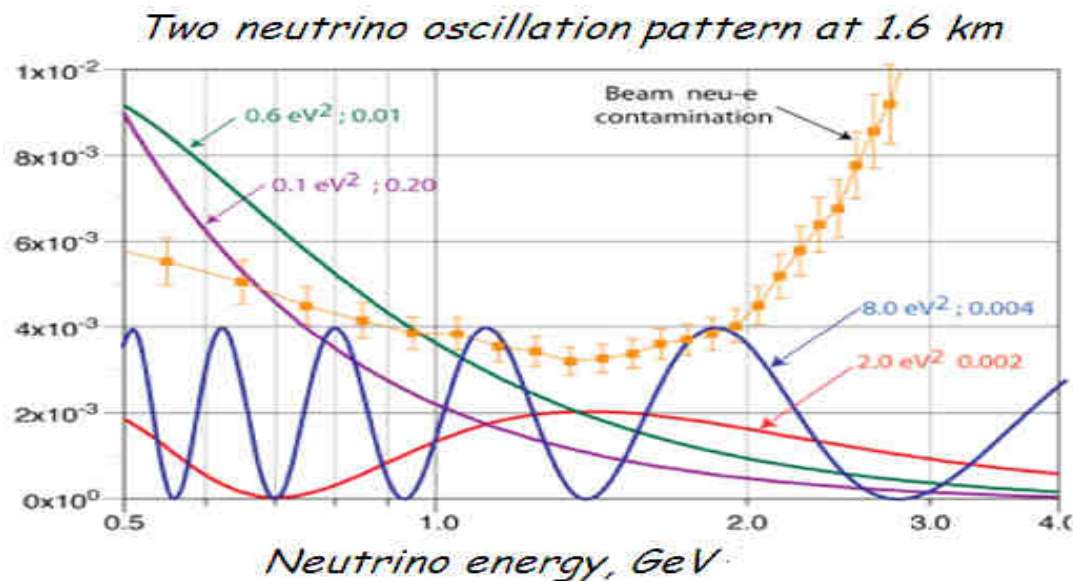


**FAR**

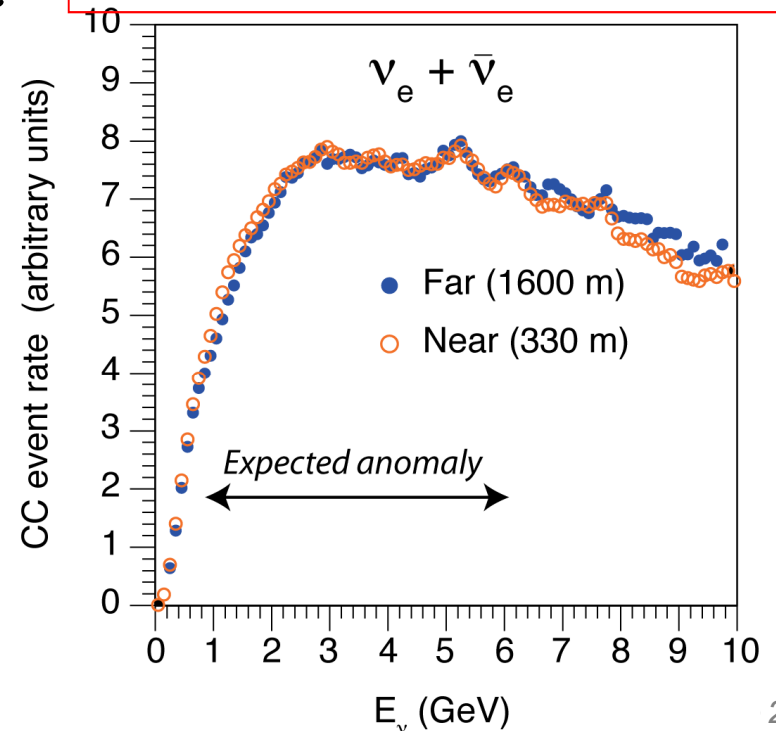
T600 moved from LNGS to CERN with new insulation

# Unique features of the CERN beam

- The present proposal is a search for spectral differences of electron like specific signatures in **two identical detectors** but at two different neutrino decay distances.
- *In absence of oscillations*, apart some beam related small spatial corrections, the two  $\nu_e$  intrinsic spectra are a precise copy of each other, independently of the specific experimental event signatures and without any Monte Carlo comparison.
- Therefore an exact, observed proportionality between the two  $\nu_e$  spectra implies directly the absence of neutrino oscillations over the measured interval of  $L/E$ .



Precise identity of intrinsic  $\nu_e$  events in the near and far positions

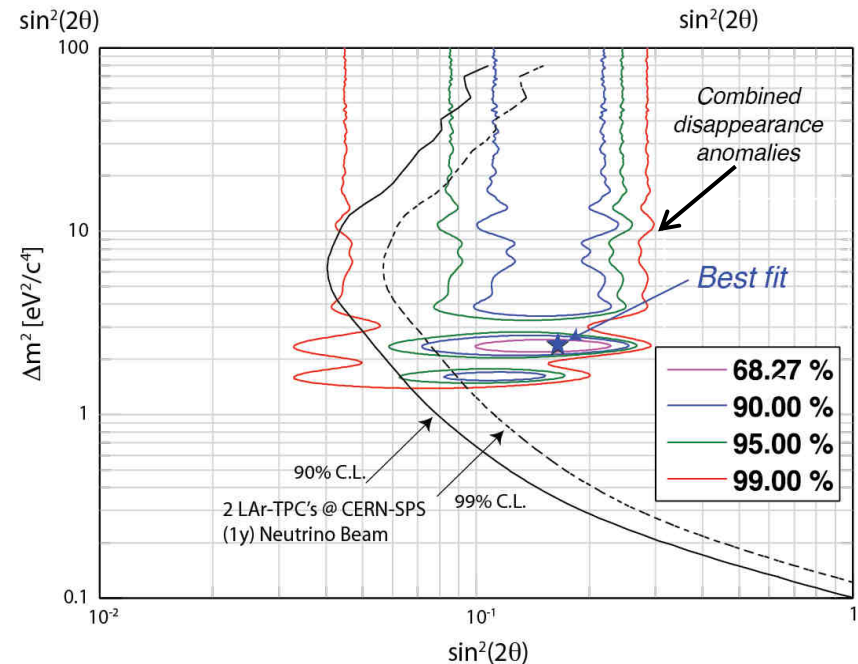
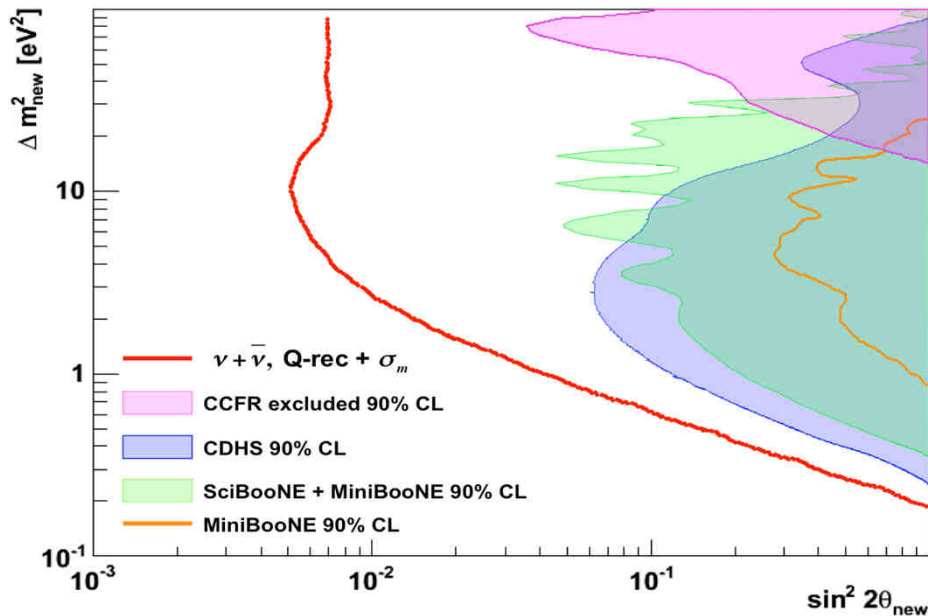
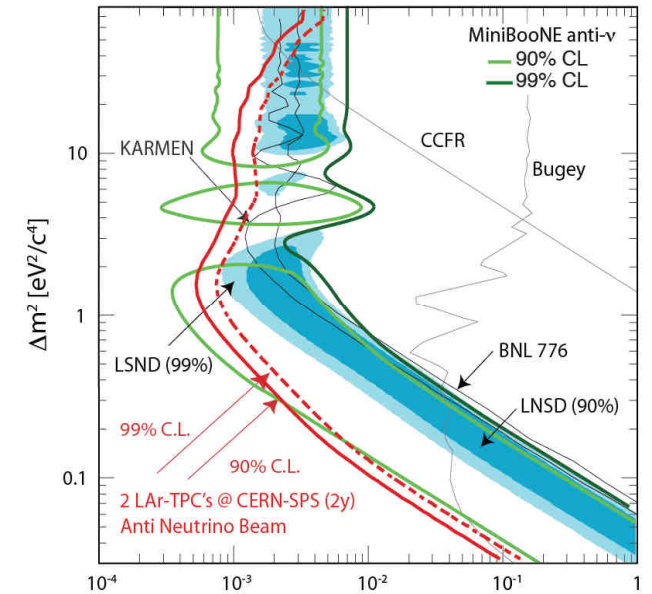
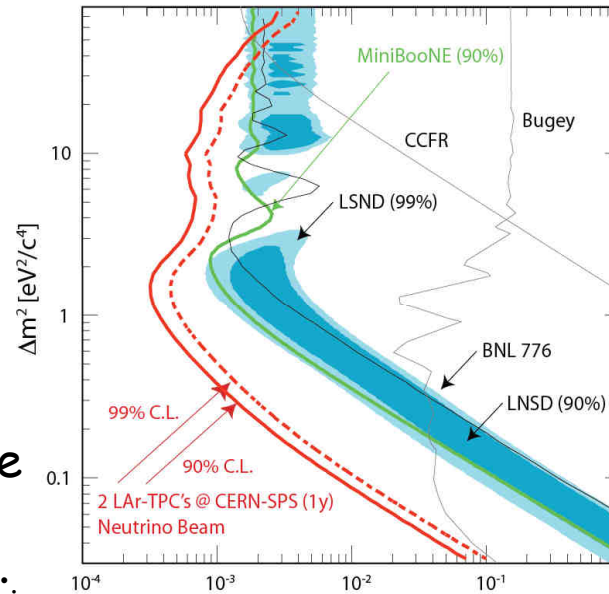


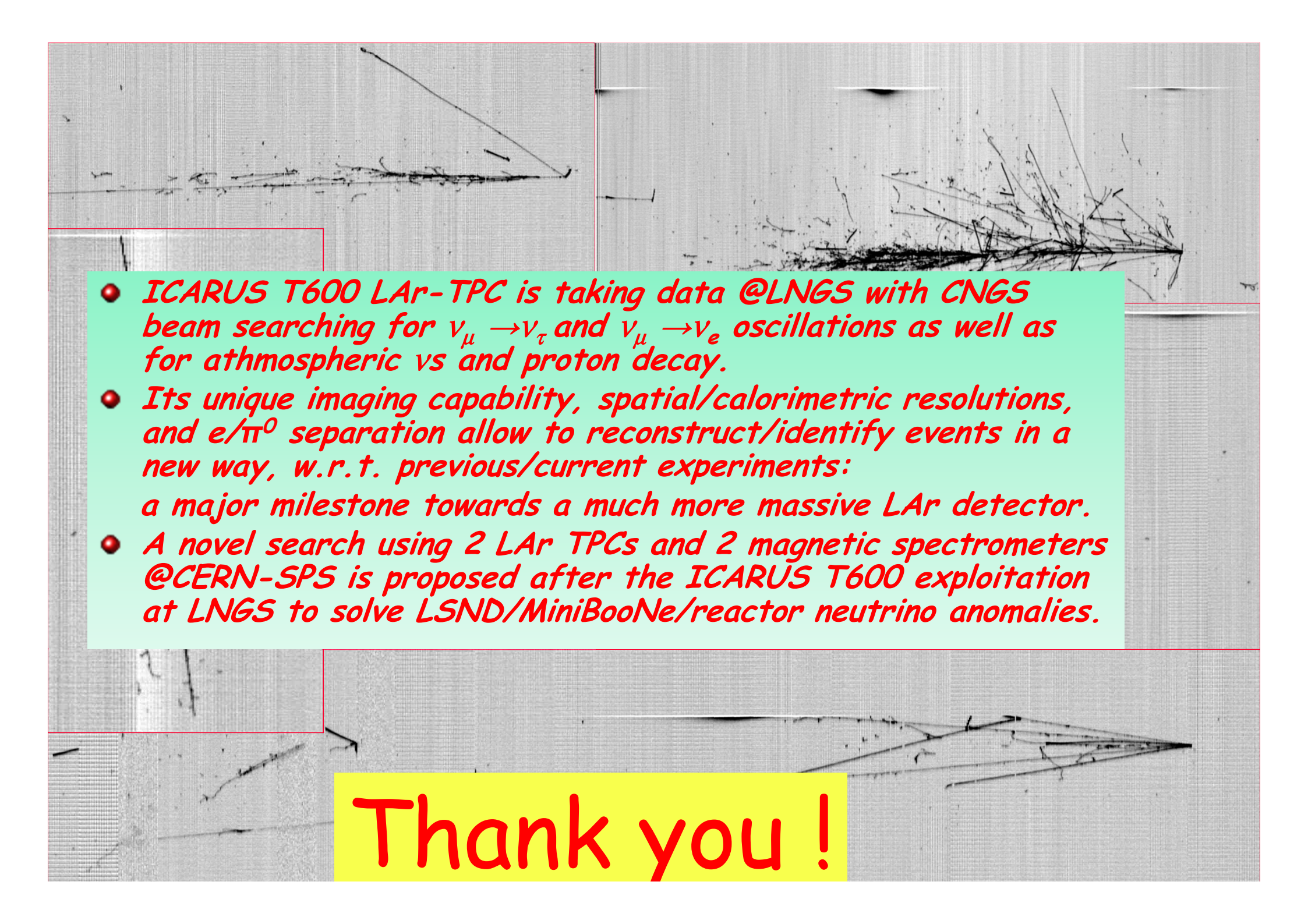
# Physics program

Full exploration of LNSD  $\nu_\mu \rightarrow \nu_e$  allowed region both with  $\nu_\mu$  and anti- $\nu_\mu$  beam.

Expected sensitivity at neutrino beam (top left) for  $4.5 \times 10^{19}$  pot and twice as much for anti-neutrino (top right).

- Search for a possible oscillatory disappearance in the  $\nu_\mu$  (bottom left) and in the initial  $\nu_e$  signals (bottom right).



- 
- *ICARUS T600 LAr-TPC is taking data @LNGS with CNGS beam searching for  $\nu_\mu \rightarrow \nu_\tau$  and  $\nu_\mu \rightarrow \nu_e$  oscillations as well as for atmospheric  $\nu$ s and proton decay.*
  - *Its unique imaging capability, spatial/calorimetric resolutions, and  $e/\pi^0$  separation allow to reconstruct/identify events in a new way, w.r.t. previous/current experiments:  
a major milestone towards a much more massive LAr detector.*
  - *A novel search using 2 LAr TPCs and 2 magnetic spectrometers @CERN-SPS is proposed after the ICARUS T600 exploitation at LNGS to solve LSND/MiniBooNe/reactor neutrino anomalies.*

Thank you !