# Spin chirality fluctuation in two-dimensional ferromagnets with perpendicular magnetic anisotropy

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Non-coplanar spin textures with scalar spin chirality can generate an effective magnetic field that deflects the motion of charge carriers, resulting in a topological Hall effect (THE)<sup>1-3</sup>. However, spin chirality fluctuations in two-dimensional ferromagnets with perpendicular magnetic anisotropy have not been considered so far. Here, we report evidence of spin chirality fluctuations by probing the THE above the Curie temperature in two different ferromagnetic ultra-thin films, SrRuO<sub>3</sub> and V-doped Sb<sub>2</sub>Te<sub>3</sub>. The temperature, magnetic field, thickness and carrier-type dependence of the THE signal, along with Monte Carlo simulations, suggest that spin chirality fluctuations are a common phenomenon in two-dimensional ferromagnets with perpendicular magnetic anisotropy. Our results open a path for exploring spin chirality with topological Hall transport in two-dimensional magnets and beyond<sup>4-7</sup>.

Understanding quantum transport of electrons in magnets is a fundamental issue in strongly correlated systems and spintronics<sup>1-3</sup>. If the magnetic moments form a non-coplanar spin texture as in many frustrated and/or chiral magnets, an electron travelling through the system will experience an effective magnetic field originated from the real-space Berry phase as it hops along a loop of three neighbouring magnetic moments (called a triad)2. Thus, this effective field is proportional to the scalar spin chirality of the triad and gives rise to a transverse response known as the topological Hall effect (THE). The THE is distinguished from both the ordinary Hall effect (OHE), which requires the application of an external magnetic field, and the anomalous Hall effect (AHE) found in ferromagnets with a uniform magnetization, in which the dominant contribution comes from the Berry phase in the momentum space in moderately conducting samples8. Because of its close relation to non-coplanar spin textures, the THE is a powerful probe to detect unconventional magnetic orders. Indeed, the THE has been instrumental in the electric detection of skyrmion phase in chiral magnets and heterostructures<sup>3,9-11</sup>.

While the THE is often interpreted as a signature of static spin textures with spin chirality in two-dimensional (2D) magnetic thin films, it has been proposed that thermal fluctuations of topological excitations could also result in a substantial Hall effect in three-dimensional magnets such as manganites<sup>12–14</sup>. However, the temperature dependence of the observed Hall resistivity cannot exclude the conventional mechanisms<sup>2</sup>. Herein, we report evidence of spin chirality fluctuation in 2D ferromagnets with perpendicular

magnetic anisotropy (PMA) by probing the THE. A substantial THE signal was observed above the ferromagnetic transition temperatures  $T_c$  in two different itinerant ferromagnets, SrRuO<sub>3</sub> (SRO) and V (5%)-doped Sb<sub>2</sub>Te<sub>3</sub> (VST) thin films in the 2D limit. The persistence of THE in the paramagnetic phase suggests a thermal-fluctuation origin. SRO is a metallic ferromagnet with n-type carriers<sup>15</sup>, while VST is a magnetically doped topological insulator with p-type carriers-type dependence of the THE are in agreement with our Monte Carlo (MC) simulations, supporting our assertion of spin chirality fluctuations in 2D ferromagnets with PMA. Given the existence of chiral spin orders and fluctuations in correlated systems, our finding opens the door to exploring the spin chirality with topological Hall transport in 2D ferromagnets<sup>4-6</sup> or quantum spin liquids<sup>1,7</sup>.

The SRO thin films were grown on SrTiO<sub>3</sub> (STO) (001) substrates using pulsed laser deposition, while the VST films were grown on STO (111) substrates with molecular beam epitaxy (see Methods and Supplementary section A) We first focus on SRO films to illustrate the chiral-fluctuation-driven THE. A cartoon schematic of the SRO thin film structure and device configuration is shown in Fig. 1a. Electric contacts were fabricated with wire bonding and silver paint. The STO capping layer helps to enhance the ferromagnetic ordering in the ultra-thin limit <sup>17</sup>. Metallic behaviour was observed in all STO capped SRO films with thickness  $t \ge 3$  unit cells (u.c.), while the 2 u.c. film is a non-magnetic insulator (see Supplementary Fig. 2 for transport data). The Curie temperature  $T_c$  is characterized by the peak anomaly of the slope of the longitudinal resistance  $(dR_{xx}/dT)^{15,18}$ . Thinner SRO films are more resistive with lower  $T_c$ , which is consistent with the effect of reduced dimensionality <sup>15</sup>.

In general, there are three contributions to the Hall resistivity in a magnetic metal: the OHE proportional to H, the conventional AHE proportional to the magnetization M and the THE due to the real-space Berry phase<sup>3,10,11</sup>. Therefore, we can express the Hall resistivity as

$$\rho_{vx}(H) = R_0 H + R_S M + \rho_{vx}^T \tag{1}$$

where  $R_0$  is the OHE coefficient,  $R_S$  is the conventional AHE coefficient, and  $\rho_{yx}^T$  is the THE contribution. Here we will focus on the anomalous part,  $\widetilde{\rho}_{yx} \equiv \rho_{yx} - R_0 H$ , by removing the OHE contribution. Well below  $T_c$ , a hysteresis loop of  $\widetilde{\rho}_{yx}$  was observed for all SRO

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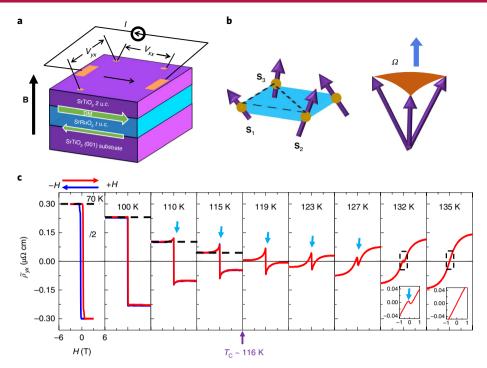


Fig. 1 | Scalar spin chirality and topological Hall effect in SrRuO<sub>3</sub> film. a, A schematic of t u.c. SRO film capped with 2 u.c. STO with electric contacts (orange patches) for transport measurements. The (STO)<sub>2</sub>/(SRO), bilayers are grown on STO (001) substrates. t is the thickness of the SRO film. A current (I) was passed through the film and longitudinal ( $V_{xx}$ ) and transverse ( $V_{yx}$ ) voltages were measured simultaneously. The Dzyaloshinsky-Moriya (DM) vectors (green arrows) on the two interfaces point in opposite directions.  $\mathbf{b}$ , a snapshot of neighbouring non-coplanar spins on a square lattice (cyan).  $S_1$ ,  $S_2$ and S<sub>3</sub> are three neighbouring spins (purple arrows) of the Ru atoms (yellow spheres) in SRO films. These three non-coplanar spins subtend a solid angle  $\Omega$ , resulting in an effective magnetic field (blue arrow).  $\mathbf{c}$ , the Hall resistivity  $\widetilde{\rho}_{uv}$  of SRO film (6 u.c.) with OHE subtracted as a function of magnetic field Hat various temperatures from 100 K to 135 K. Blue (red) curves were taken as the magnetic field from -6 to 6T (6 to -6T). The cyan arrows indicate the THE humps emerge above 100 K. The insets show enlarged curves of 132 and 135 K data around zero field, suggesting THE peaks are absent above 135 K.

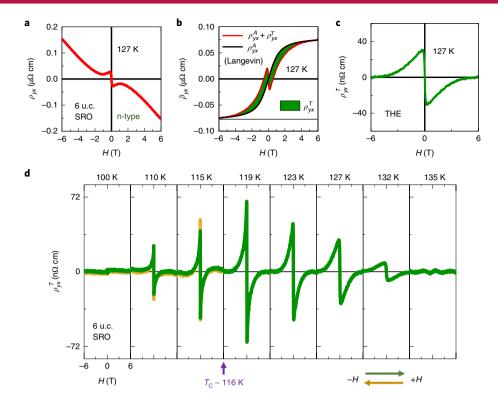
(and VST) films (see Supplementary sections I-K). This loop has the same shape as the magnetization hysteresis loop, indicating that below  $T_c$  the main contribution to  $\widetilde{\rho}_{vx}$  comes from the conventional AHE ( $\propto M$ ). This is similar to previous reports of  $\widetilde{\rho}_{vx}$  on magnetic thin films with PMA<sup>16,19-21</sup>.

Surprisingly, at higher temperatures a sharp anomaly in  $\widetilde{\rho}_{vr}(H)$ develops around  $T_c$ . Figure 1c shows  $\widetilde{\rho}_{yx}(H)$  of the STO-capped SRO film (6 u.c.). Below  $T_c \approx 116 \,\mathrm{K}$ ,  $\widetilde{\rho}_{vx}(H)$  shows a pronounced square shape, indicating a robust ferromagnetic ordering with a strong uniaxial anisotropy<sup>15</sup>. As T approaches  $T_c$ , a prominent antisymmetric peak near zero field emerges. This emergent feature indicates an additional contribution to the Hall signal, which we attribute to the THE. In particular, the AHE changes its sign around 119 K as evidenced from the high-field values of  $\widetilde{\rho}_{vv}(H)$ . This is due to the strong energy dependence of the AHE near the Fermi energy<sup>22</sup>. Therefore,  $\widetilde{\rho}_{vx}(H)$  data at 119 K are dominated by the THE signal (see Supplementary section I.4 for the complete dataset). The THE data at 119 K are different from the conventional AHE, showing a sharp antisymmetric peak near zero field followed by a smooth suppression at high fields. Furthermore, this antisymmetric peak feature becomes strongest slightly above  $T_c$ , then gradually decreases and disappears around 132 K, well into the paramagnetic phase. Note that the THE feature is visible in the raw Hall data (see Supplementary section C). The persistence of THE in the paramagnetic phase above  $T_c$  demonstrates a thermally driven spin chirality fluctuation mechanism<sup>23,24</sup>, which is distinct from the THE due to static spin texture emerging below  $T_c$  in previous studies<sup>19–21</sup>.

To extract the T evolution of the THE signal, the AHE contribution  $(\propto M)$  at each T needs to be properly subtracted. It is technically challenging to directly measure the uniform out-of-plane magnetization of thin films in the 2D limit because of diminishing stray field due to the demagnetization factor<sup>25</sup>. For our SRO films, the M(H) curves of the paramagnetic phase  $(T > T_c)$  can be described by a Langevin function (or self-consistent equation using a Brillouin function, see Supplementary section D). Since the the antisymmetric peak feature becomes negligible for  $T \ge 135 \,\mathrm{K}$ , these high-temperature AHE data are used for the Langevin fitting. For  $T < T_c$ , a step function is used to approximate the AHE contribution. Figure 2 shows an example of AHE background subtraction. The raw data  $\rho_{vx}(H)$  at 127 K are shown in Fig. 2a. The negative slope at high field indicates n-type charge carriers 15. Figure 2b shows  $\widetilde{\rho}_{vr}(H)$ after removing the OHE and the Langevin fitting of the AHE background. The difference (green area) between them is the THE signal  $\rho_{yx}^{T}$ , which is plotted in Fig. 2c. Using this procedure, we extracted  $\rho_{vx}^{T}$  of the 6 u.c. SRO film at various T as shown in Fig. 2d. The

antisymmetric THE peaks are visible from 110 K to 132 K and are most pronounced around  $T_c$ , which suggests that the observed THE originates from the thermal fluctuation of spin chirality. Similar T-H dependence of THE was also observed in thin-

ner STO-capped SRO films with t=3, 4, 5, consistent with a THE induced by spin chirality fluctuations (see Supplementary sections I and J). Furthermore, the THE is absent in thicker SRO films (t > 7), supporting the 2D nature of the spin chirality fluctuation due to the spatial confinement effect. Figure 3a shows the maximum values of  $\rho_{yx}^T$  as a function of reduced temperature  $T/T_c$ . The THE signal peaks around  $T_c$ , then diminishes approximately 20% above  $T_c$ . The non-monotonic thickness dependence of the THE magnitudes might come from the combined effects of the effective exchange (J) and other parameters (for example, anisotropy, magnetic disLETTERS NATURE MATERIALS



**Fig. 2 | Extracting the THE signal**  $(\rho_{yx}^T)$  **in SRO films. a,** Raw data  $\rho_{yx}$  of SRO film (6 u.c.) at 127 K. The negative slope at high field indicate n-type carriers. **b,**  $\hat{\rho}_{yx}$  (black) of the SRO film at 127 K was fitted with a Langevin function (red). The green area is the THE signal, which is plotted in **c. d,**  $\rho_{yx}^T$  of SRO film as a function of magnetic field H at various temperature from 100 K to 135 K. Green (yellow) curves are taken as the magnetic field from -6 to 6 T (6 to -6 T). The Curie temperature  $T_x \sim 116$  K is indicated by the purple arrow.

orders). The qualitative T-H dependences of the THE signal are consistent with the prior theoretical studies of thermal fluctuations in 2D chiral magnets<sup>23,24</sup>. However, spin chirality fluctuations in 2D ferromagnets with strong PMA have not been addressed.

To gain insight into the emergent spin chirality fluctuation in 2D ferromagnets with PMA, we carried out MC simulations with the following Hamiltonian (see Methods and Supplementary section N for details):

$$H = \sum_{\langle ij \rangle} \left[ -J(\mathbf{S}_i \cdot \mathbf{S}_j) + \mathbf{D}_{ij} \cdot (\mathbf{S}_i \times \mathbf{S}_j) \right] - K \sum_i (S_i^z)^2 - B_z \sum_i S_i^z$$

where  $S_i = Sn_i$  is the spin on the ith lattice site. The first term (J>0) describes ferromagnetic Heisenberg exchange coupling. The second term describes the chiral interaction, namely the Dzyaloshinsky–Moriya interaction (DMI)<sup>26,27</sup>, arising from the inversion symmetry breaking (due to interfaces, for example). The uniaxial anisotropy K is included in the third term. With  $D < D_c \equiv 2\sqrt{2JK} / \pi$  ( $D_c$  is the critical DMI value that separates uniform ferromagnet and chiral magnet ground states), the ground state is a ferromagnet with uniform out-of-plane (OOP) magnetization despite of the presence of DMI<sup>28</sup>. The last term is the Zeeman energy due to the external magnetic field  $B_z$ .

It appears that the DMI is forbidden by the symmetric structure (STO/SRO/STO) of our films. However, it has been shown that the STO film grown on SRO is slightly different from the STO substrate<sup>29</sup>. This slight difference breaks the OOP inversion symmetry, allowing a small but non-zero chiral interaction (DMI). Thus, non-zero spin chirality emerges when the ferromagnetic order is 'melted' by thermal fluctuations around  $T_c$ . The emergent spin chirality is characterized by the topological charge Q defined as:

$$Q = \frac{1}{4\pi} \int d^2 \mathbf{r} \left[ \mathbf{n} \cdot (\partial_x \mathbf{n} \times \partial_y \mathbf{n}) \right]$$
 (2)

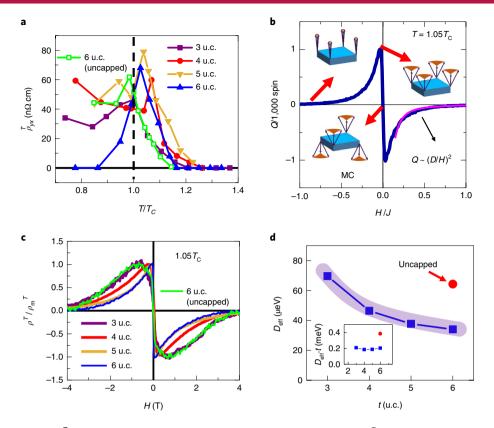
where **n** is a unit vector describing the local spin direction. On a lattice, the integral is replaced by the sum of the solid angle  $\Omega$  subtended by three neighbouring spins (a triad) as shown in Fig. 1b<sup>30</sup>. As shown in Fig. 3b, our MC simulations demonstrate that the topological charge density Q(H) is non-zero around  $T_{c}$ , showing qualitatively the same antisymmetric peak structure as that of the observed THE (see Supplementary section N). The agreement suggests that the observed THE originates from the effective magnetic field generated by the thermal-fluctuation-driven spin chirality.

The H dependence of the THE can be understood in the following physical picture. At the limit of  $H \rightarrow 0$ , the induced Q is proportional to H because each triad carries a net moment as shown in Fig. 3b. Therefore,  $Q \propto H \propto M$  in the paramagnetic phase at  $H \rightarrow 0$ . At the high-field limit  $(H \rightarrow \infty)$ , the Zeeman energy competes with the DMI, thus suppressing the solid angles  $\Omega$  of all triads, resulting in a vanishing Q (and THE). The leading order in the high-field expansion is proportional to  $D^2$  because Q respects the spatial inversion symmetry<sup>23</sup>, so  $Q \propto (D/H)^2$  as demonstrated by the cyan line in Fig. 3b. Thus, the overall H dependence can be described by a phenomenological function that interpolate these two limits:

$$Q(H) \propto \frac{M(H)}{1 + (H/D_{\text{eff}})^2} \tag{3}$$

where  $D_{\rm eff}$  is the effective DMI. This function reasonably describes both MC simulations and experimental data, supporting the simple intuitive picture (see Supplementary section L for detailed analysis). As shown in Fig. 3c, the width of THE peak decreases systematically as t increases, indicating decreasing  $D_{\rm eff}$ . This behaviour is also supported by our MC simulations with varying DMI (see Supplementary section L). Using equation (3), we extracted the values of  $D_{\rm eff}$  (up to a numerical factor) of the STO-capped (3, 4, 5 and

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**Fig. 3** | T-H dependence of THE ( $\rho_{yx}^T$ ) in SRO films and the effective DMIs ( $D_{eff}$ ). **a**, Maximum values of  $\rho_{yx}^T$  as a function of temperature T of SRO films with thickness t = 3-6 u.c. capped with STO and the 6 u.c. uncapped SRO film. **b**, MC simulations (uniaxial anisotropy K = 0.5J and DMIs  $D = 0.25J < D_c$ ) of topological charges Q at  $T = 1.05T_c$  show an antisymmetric profile similar to the experimental data of THE. The high-field tail can be described by  $(D/H)^2$  (the magenta curve). The inset cartoons show the behaviour of spin triads at various magnetic fields, as indicated by the red arrows. **c**, H dependence of normalized  $\rho_{yx}^T$  at the reduced temperature  $1.05T_c$ .  $\rho_m^T$  is the maximum THE signal. The 6 u.c. uncapped SRO film is indicated by the green curve. **d**, t dependence of effective DMIs ( $D_{eff}$ ). The  $D_{eff}$  was extracted from fitting the H dependence of  $\rho_{yx}^T$  shown in **c**. The 6 u.c. uncapped SRO film is indicated by the red circle. The inset shows that the product  $D_{eff} \times t$  is approximately a constant, suggesting an interface origin.

6 u.c.) and uncapped (6 u.c.) SRO films.  $D_{\rm eff} \times t$  values are shown in the inset of Fig. 3d. Essentially  $D_{\rm eff}$  of the STO-capped films is inversely proportional to t, in agreement with the interface origin of  $D_{\rm eff}$ . Furthermore, the  $D_{\rm eff}$  of uncapped SRO film (6 u.c.) is about twice as large as that of the capped one, which is consistent with the stronger inversion symmetry breaking of the uncapped SRO film. The estimated interface DMI value is  $D\approx 0.2\,{\rm meV}$ , which is approximately one to two orders of magnitude smaller than that of the SrIrO<sub>3</sub>/SRO interface<sup>20</sup>, in good agreement with the expectation of weak DMI in our SRO films.

To demonstrate the common nature of spin chirality fluctuation, we present Hall data of five-quintuple (QL) VST thin film capped with three-QL Sb<sub>2</sub>Te<sub>3</sub> (ST). Because the carrier density of our ST/ VST film is approximately 1,000 times smaller than that of SRO, the Fermi level of the ST/VST film can be shifted substantially by a gate voltage. More interestingly, the AHE can be tuned to zero at every T, which allows an unambiguous isolation of the THE signal. The schematic of the heterostructure is shown in Fig. 4a. Figure 4b shows the raw data  $(\rho_{vx})$  of the ST/VST film. The positive slope at high field indicates p-type carriers. An example of gate tuning AHE at 30 K is shown in Fig. 4c (see Supplementary section K for the complete dataset). Figure 4d shows the THE data at T above and below  $T_c \approx 27.5$  K, showing qualitatively the same T-H dependences as that of SRO films. Furthermore, the sign of the THE is consistent with that of the carrier type, in agreement with the same effective magnetic field due to the same spin chirality fluctuations in these two different 2D ferromagnets with PMA.

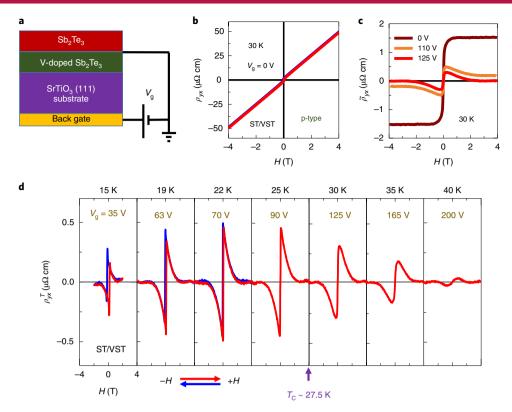
This agreement suggests that the THE observed in SRO and VST films originates from the 2D spin chirality fluctuations. For  $T > T_c$ , the suppression of the THE signal approximately follows a power law, in reasonable agreement with that of MC simulations (see Supplementary section M). This would inspire future studies of the critical scaling behaviour of the spin chirality fluctuation. Although the ground states of both systems are simple ferromagnets where the static skyrmion phase is unstable because  $D \ll D_c$ , the thermal fluctuations around  $T_c$  are chiral due to the presence of the (weak) chiral interaction D. How D influences the critical behaviour of 2D ferromagnets with PMA is an interesting subject to explore in future studies.

Although the THE diminishes at  $T \ll T_c$  in STO-capped 6 u.c. SRO film, the THE persists to low temperatures in the 3 u.c. SRO film and the ST/VST film (see Supplementary sections I and K). It is unclear whether the low-temperature THE originates from static topological spin texture (for example, skrymions), which will inspire future studies of the THE in Berry phase engineered magnetic thin films such as the ST/VST films. Our report of spin chirality fluctuations opens the door to further exploration of topological Hall transport in other itinerant 2D ferromagnets and beyond 4-6, and may help to identify chiral spin liquid states with quantum entanglements in 2D correlated systems using topological Hall transport 1-7.

## Online content

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**Fig. 4 | THE**  $(\rho_{yx}^T)$  of ST-capped VST films. **a**, A schematic of the  $(ST)_3/(VST)_5$  heterostructure grown on STO (111) substrate. A back-gate voltage  $(V_g)$  was applied to tune the AHE. **b**, Raw data  $\rho_{yx}$  of ST/VST film at 30 K. The positive slope at high field indicates p-type carriers. **c**,  $\widetilde{\rho}_{yx}$  of VST film at 30 K with different gate voltages. The AHE is tuned to zero at 125 V so that only THE is visible. **d**,  $\rho_{yx}^T$  of VST film as a function of magnetic field H at various temperature from 15 K to 40 K. Red (blue) curves were taken as the magnetic field from -6 to 6 T (6 to -6 T). The Curie temperature  $T_c \sim 27.5$  K is indicated by the purple arrow. The overall behaviour is qualitatively the same as that of SRO films.

associated accession codes are available at https://doi.org/10.1038/s41563-019-0454-9.

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# **Author contributions**

W.Wu conceived and supervised the project. Z.L., J.W., G.K. and G.R. synthesized the SRO samples and performed X-ray diffraction. Y.Z. and C.-Z.C. synthesized the VST samples. W.Wang performed the magnetotransport experiments and analysed the data. M.W.D. and D.X. performed the MC simulations. W.Wang, W.Wu, M.W.D. and D.X. wrote the manuscript. All authors discussed the data and contributed to the manuscript.

# **Competing interests**

The authors declare no competing interests.

#### Additional information

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LETTERS NATURE MATERIALS

#### Methods

Sample growth and characterization. Both SRO and STO layers were epitaxially grown on  ${\rm TiO_2}$ -terminated (001) STO substrates using a pulsed laser deposition technique. The single  ${\rm TiO_2}$ -terminated STO substrates were achieved by standard buffer hydrofluoric acid (HF) etching for 30 s and subsequent annealing at 950 °C for 90 min. During the growth of either SRO or STO, the laser fluence and repetition rate were  $2\,{\rm J\,cm^2}$  and 1 Hz, respectively. The substrate temperature and oxygen partial pressure during growth were 650 °C and 0.25 mbar, respectively. In-situ reflection high-energy electron diffraction was used to monitor the growth and confirmed a layer-by-layer growth fashion. All the films showed an atomic flat surface with similar terrace structure with substrates. X-ray diffraction was carried out using a PANalytical XPert Materials Research Diffractometer in high-resolution mode.

The Sb<sub>2</sub>Te<sub>3</sub>/Sb<sub>2-x</sub>V<sub>x</sub>Te<sub>3</sub> (x = 0.1) heterostructure was grown on SrTiO<sub>3</sub>(111) substrate with a molecular beam epitaxy (MBE) method. The insulating SrTiO<sub>3</sub> (111) substrates used for the growth were first soaked in 90°C dejonized water for 1.5 h, and then annealed at 985 °C for 3 h in a tube furnace with flowing pure oxygen gas. Through the above heat treatment, the STO (111) substrate surface become passivated and atomically flat; the topological insulator (TI) heterostructure growth was carried out using a commercial EPI-620 MBE system with a vacuum that was better than  $2 \times 10^{-10}$  mbar. The heat-treated insulating STO (111) substrates were outgassed at ~530 °C for 1 h before the growth of the TI heterostructures. High-purity Sb (99.999%) and Te (99.999%) were evaporated from Knudsen effusion cells and V (99.995%) was evaporated from an e-gun. During growth of the TI, the substrate was maintained at 240 °C. The flux ratio of Te to Sb was set to be >10 to prevent Te deficiency in the samples. The pure or magnetic TI growth rate was at ~0.25 QL min<sup>-1</sup>. Following the growth, the TI films were annealed at ~240 °C for 30 min to improve the crystal quality before being cooled down to room temperature. Finally, to avoid possible contamination, an 18-nm-thick Te layer was deposited at room temperature on top of the sandwich heterostructures before their removal from the MBE chamber for transport measurements.

**Transport measurements.** The Hall resistance and longitudinal resistance were measured by standard lock-in techniques with an alternating current of  $40\,\mu\mathrm{A}$  modulated at 314 Hz. All Hall data are antisymmetrized.

**Monte Carlo simulations.** The MC simulations were conducted on a  $N^2 = 32 \times 32$  square lattice with periodic boundary conditions. The Hamiltonian was used as the energy function underlying the standard Metropolis algorithm. The expression of Q is as follows:

$$\exp\left(\frac{i\Omega}{2}\right) = r^{-1}(1 + \mathbf{S}_1 \cdot \mathbf{S}_2 + \mathbf{S}_2 \cdot \mathbf{S}_3 + \mathbf{S}_3 \cdot \mathbf{S}_1 + i\chi) \tag{4}$$

where  $\chi \equiv S_1 \cdot (S_2 \times S_3)$  is the spin chirality of the three neighbouring spins¹ and  $r = \sqrt{2(1 + S_1 \cdot S_2)(1 + S_2 \cdot S_3)(1 + S_3 \cdot S_1)}$  is the normalization factor. The sign of chirality is fixed by the chiral interaction (DMI). The total topological charge Q is then given by summing  $\Omega$  over all the right-handed spin triangles in the lattice.

$$Q = \frac{1}{4\pi} \sum_{\Delta_{ijk}} \Omega(\mathbf{S}_{ij}, \mathbf{S}_{j}, \mathbf{S}_{k})$$
 (5)

Q (or  $\chi$ ) is an odd function with respect to time reversal symmetry, so Q(H) = -Q(-H). Thus, for the paramagnetic phase with H = 0, the net topological charge is zero, Q(0) = 0, corresponding to equal populations of 'up' and 'down' triads.

To illustrate chiral fluctuations, we present representative MC simulation results with parameters D=0.25J and K=0.5J. The simulation of  $32\times32$  sites shows a ferromagnetic ordering with  $T_c\approx0.85J$ . The extrapolated thermodynamic  $T_c\approx0.71J$ . At  $T=J>T_c$ , the M(H) curve shows the expected paramagnetic behaviour without hysteresis.

## Data availability

The data that support the findings of this study are available from the corresponding author upon reasonable request.

# Code availability

The code for the MC simulations is available from M.W.D. and D.X. on reasonable request.