

## Assessing the Internal Uppermost Crustal Structure of the Central Pyrenees by Gravity-Constrained Cross Sections



### Key Points:

- Gravity, petrophysical and geological data are integrated to constrain the geometry of the upper crust of the Central Pyrenees
- Bouguer and residual anomalies and six serial gravity-constrained cross sections of the study area are obtained
- Gravity variations respond to changes in the volume of Triassic evaporites, the geometry of the igneous bodies, and lateral structural variations

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**Abstract** The Pyrenees constitutes an exceptional example of an Alpine orogenic belt characterized by basement thrust sheets involving Paleozoic rocks and Mesozoic and Cenozoic cover units detached on the Triassic evaporites, the main décollement level. This work is located in the Central Pyrenees, where gravity data help to better constrain the internal architecture of the upper crust of the southern half of the Axial Zone and the northern part of the South Pyrenean Zone, a key area to understand the orogenic evolution of the chain. Previous and new gravity, petrophysical and geological data have been used to obtain the Bouguer and residual anomaly maps of the study area and six serial gravity-constrained cross sections perpendicular to the main structural trend. The residual anomaly map shows a good correlation between basement units involved in thrust sheets of the study area and gravity highs whereas negative anomalies are interpreted to correspond with Mesozoic/Cenozoic basins, Triassic evaporites, Late Variscan igneous bodies, and Ordovician gneisses. The six gravity-constrained cross sections highlight strong along-strike variations on the gravity signal due to lateral differences of the superficial and subsurface occurrence of Triassic evaporites, different geometry at depth of the Late Variscan igneous bodies outcropping in the study area, and geometric lateral variations of the basement thrust sheets and their relationship with the Mesozoic-Cenozoic units.

## 1. Introduction

The integration of geophysical modeling to classical structural studies constitutes a powerful approach to inferring the Earth's crust architecture (e.g., Calcagno et al., 2008; Crombez et al., 2020; Godin & Harris, 2014; Jessel, 2001). Focusing on the upper crust, the characterization of the crustal-basement structure and its relationships with the overlying sedimentary cover has deserved much attention during the last decades due to its implications for heat-flow prediction and hydrocarbon system modeling (e.g., Forster & Smith, 1989; Ten Brink et al., 1993). Obtaining an accurate interpretation of the upper crustal structure is not easy, especially in areas where data arising from wells and/or seismic sections are absent or scarce. Gravity data, due to their relative low acquisition cost, wide availability, and well-established processing and interpretation techniques, constitute a valid and widespread geophysical method used in the upper crust to study basement geometries (e.g., Ayala et al., 2019; Goleby et al., 1989; Izquierdo-Llavall et al., 2019), assess the internal structure and thickness of allochthonous complexes (Martínez Catalán et al., 2012), delineate the distribution and geometry of evaporitic bodies and/or evaporitic detachment levels (e.g., Ali et al., 2014; Calvín et al., 2018; Santolaria et al., 2016, 2020; Sarsar-Naouali et al., 2011) and/or study geometry at depth of granitic bodies (e.g., Aranguren et al., 2003; Terrinha et al., 2018; Vigneresse & Clemens, 2000).

In the last decades, considerable progress has been made in understanding the crustal and lithospheric structure of the Pyrenees based on several geophysical methods; wide-angle reflection/refraction profiles (Daignières et al., 1981; Gallart et al., 1981), vertical incidence reflection profiles (e.g., ECORS-Pyrenees Team, 1988; Muñoz, 1992), analyzing seismic and teleseismic data (e.g., Chevrot et al., 2018; Díaz et al., 2012; Wang et al., 2016), magnetotelluric data (e.g., Campanyà et al., 2011, 2012; Ledo et al., 2000), magnetic anomalies (e.g., Canva et al., 2020), seismic tomography (e.g., Chevrot et al., 2014), heat flow (e.g., Fernández et al., 1990)

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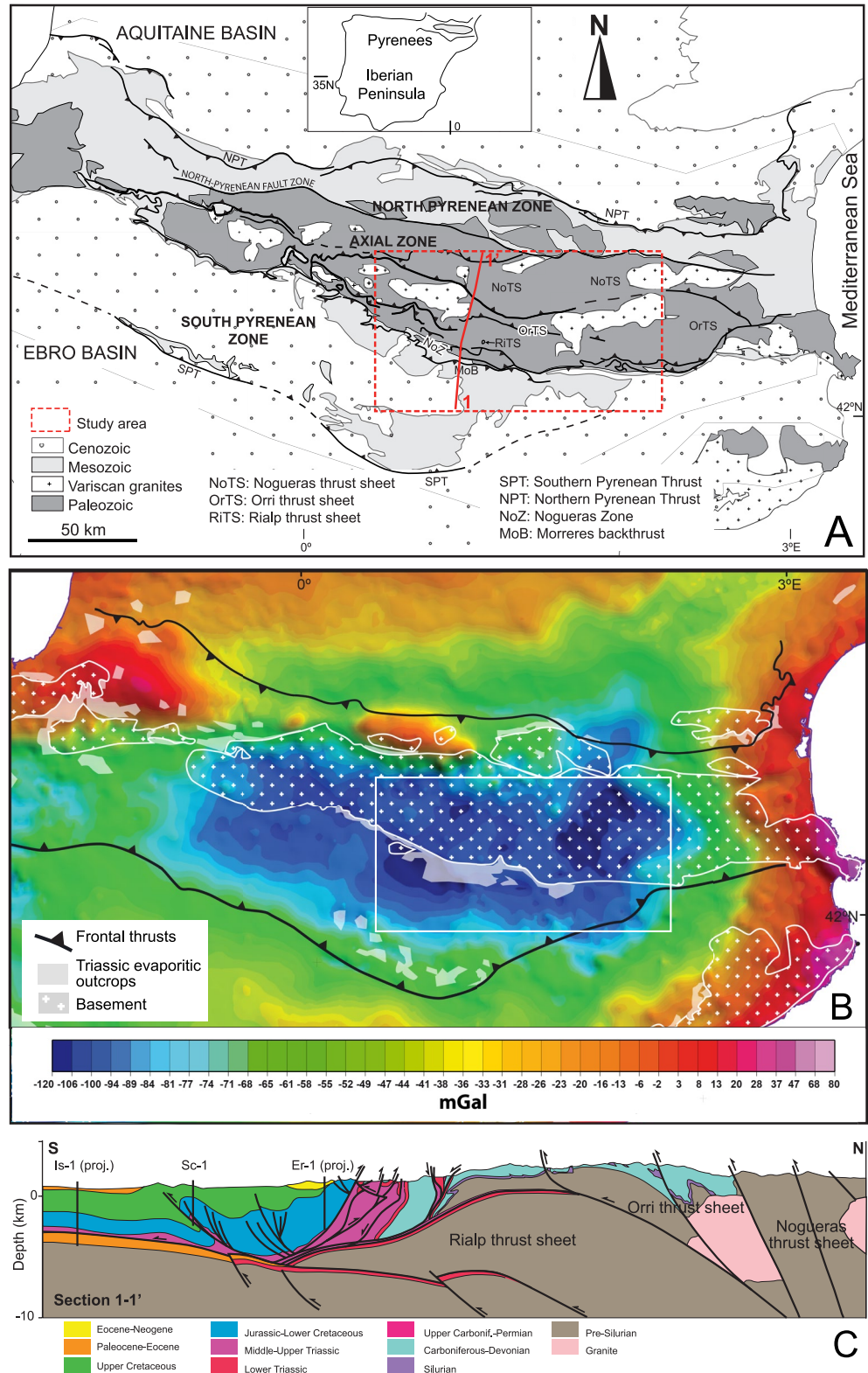
and gravimetric data (Casas et al., 1997; Pedrera et al., 2017; Rivero et al., 2002; Torné et al., 1989, 2015; Vacher & Souriau, 2001). As numerous works demonstrate, the crustal deep structure of the Pyrenees constitutes a first-order element controlling the evolution of the Pyrenees and also its uppermost structure (e.g., Beaumont et al., 2000; Soto et al., 2006; Jammes et al., 2009; Saspiturry et al., 2019; Muñoz, 2019 and references therein). A rapid glance at the crustal structure of the Pyrenees can be done by looking at its Bouguer anomaly map (e.g., Ayala et al., 2016; Casas et al., 1997; Torné et al., 1989) (Figure 1). It shows from North to South a high-amplitude positive-negative-positive anomaly pattern across the WNW-ESE Pyrenean direction. The large negative anomaly reaching c.  $-100$  mGal coincides with the core of the chain and is caused by the thick Pyrenean crustal roots (Torné et al., 2015). In the Northern Pyrenees, strong positive anomalies have been related to uplifted dense bodies (from the either lower crust or mantle) (e.g., Casas et al., 1997; Pedrera et al., 2017; Vacher & Souriau, 2001). To the North and South, a higher but laterally variable  $-60$  to  $5$  mGal gravity response corresponds to the sedimentary filling at its foreland basins, the Aquitanian and Ebro basins (Figure 1). The area studied in this work lies on the long-wavelength large negative anomaly at the core of the chain (up to  $400$  km). This area also shows, at a shorter wavelength (c.  $2$  km), different relative gravity maxima and minima whose detailed analysis adding a dense gravity prospecting constitute one of the main targets of this work (Figure 1).

Interpretations deduced from gravity data are non-unique and different models can fit the same observations (e.g., Blakely, 1996; Crombez et al., 2020; LaFehr & Nabighian, 2013). Decreasing the range of possible solutions represents one of the major challenges when modeling subsurface geology (e.g., Crombez et al., 2020; Saltus & Blakely, 2011). The relatively well-known geology with exceptional outcrops and availability of geophysical data make the Central Pyrenees the ideal candidate for attempting to better understand their upper crustal structure and analyze the gravity response of their different geological crustal bodies. This approach can be applied to other orogenic areas with less subsurface information where gravity data are available.

## 2. Geological Setting

The Pyrenees are a WNW-ESE asymmetric mountain range with a mainly southwards vergence, developed from the Late Cretaceous to the Early Miocene (e.g., Muñoz, 1992). It resulted from the convergence between the Iberian and European plates which was accommodated by underthrusting of the Iberian lithosphere underneath Europe and thrusting and stacking of the upper crust (e.g., Choukroune and ECORS team, 1989; Roest & Srivastava, 1991; Robenbaum et al., 2002). The present-day crustal architecture of the Pyrenees results from a complex evolution influenced by several factors such as the re-activation of previous Variscan structures and/or inversion of inherited Mesozoic extensional basins (e.g., Bond & McClay, 1995; Martínez-Peña & Casas-Sainz, 2003; Muñoz, 1992; Vergés et al., 1995), weakness of the inherited Variscan crust and the lithospheric thermal state inherited from Early Mesozoic rifting stages (e.g., Clerc & Lagabriele, 2014; Jammes et al., 2014; Jammes & Huisman, 2012), and the distribution and position of mechanically weak horizons (e.g., Beaumont et al., 2000; Muñoz, 1992). Regarding the last factor, in the Pyrenees three main different weak horizons have been recognized to control its evolution: (a) a mid-crustal weak zone decoupling the Iberian crust in the Central Pyrenees (e.g., Beaumont et al., 2000; Muñoz, 1992), (b) the Silurian shales acting as an effective décollement in Variscan and Alpine basement thrust sheets (e.g., Casas et al., 2019; Cochelin et al., 2017; García-Sansegundo, 1990; García-Sansegundo et al., 2011; Marcén et al., 2018; Matte, 1969), and (c) the Mid-Upper Triassic evaporites which exert a major control on the tectonic style of the North and South Pyrenean Zones (e.g., Ford & Vergés, 2021; Muñoz, 1992).

The overall geometry and structural architecture of the mountain range have already been established based on numerous geological and geophysical data obtained during several decades. The regional cross sections constructed in different sectors of the mountain belt point to an along-strike variability of the basement and sedimentary cover structure (e.g., Casas et al., 1997; Casas and Faccenna, 2001; Chevrot et al., 2018; Espurt et al., 2019; Martínez-Peña & Casas-Sainz, 2003; Muñoz, 1992; Teixell, 1998; Vergés et al., 1995). Across the central part of the chain,  $167$  km of shortening have been estimated (Muñoz, 1992). Geophysical data coming from a series of receiver-functions profiles crossing the Pyrenees also show an along-strike variability on its deep crustal architecture (Chevrot et al., 2015, 2018).



**Figure 1.** (a) Tectonic sketch map of the Pyrenees, showing the main basement and cover units and the study area (in dashed red rectangle). (b) Bouguer anomaly map of the Pyrenees showing the study area (modified from Ayala et al., 2016). The location of the Paleozoic basement and Triassic evaporitic rocks is also shown. (c) Geological cross section along the ECORS profile (modified from Muñoz et al. [2018]). Location on Figure 1A.

## 2.1. Structural Overview

The study area is located in the central sector of the Central Pyrenees and comprises the southern half of the Axial Zone and the northernmost part of the South Pyrenean Zone (Figures 1 and 2a).

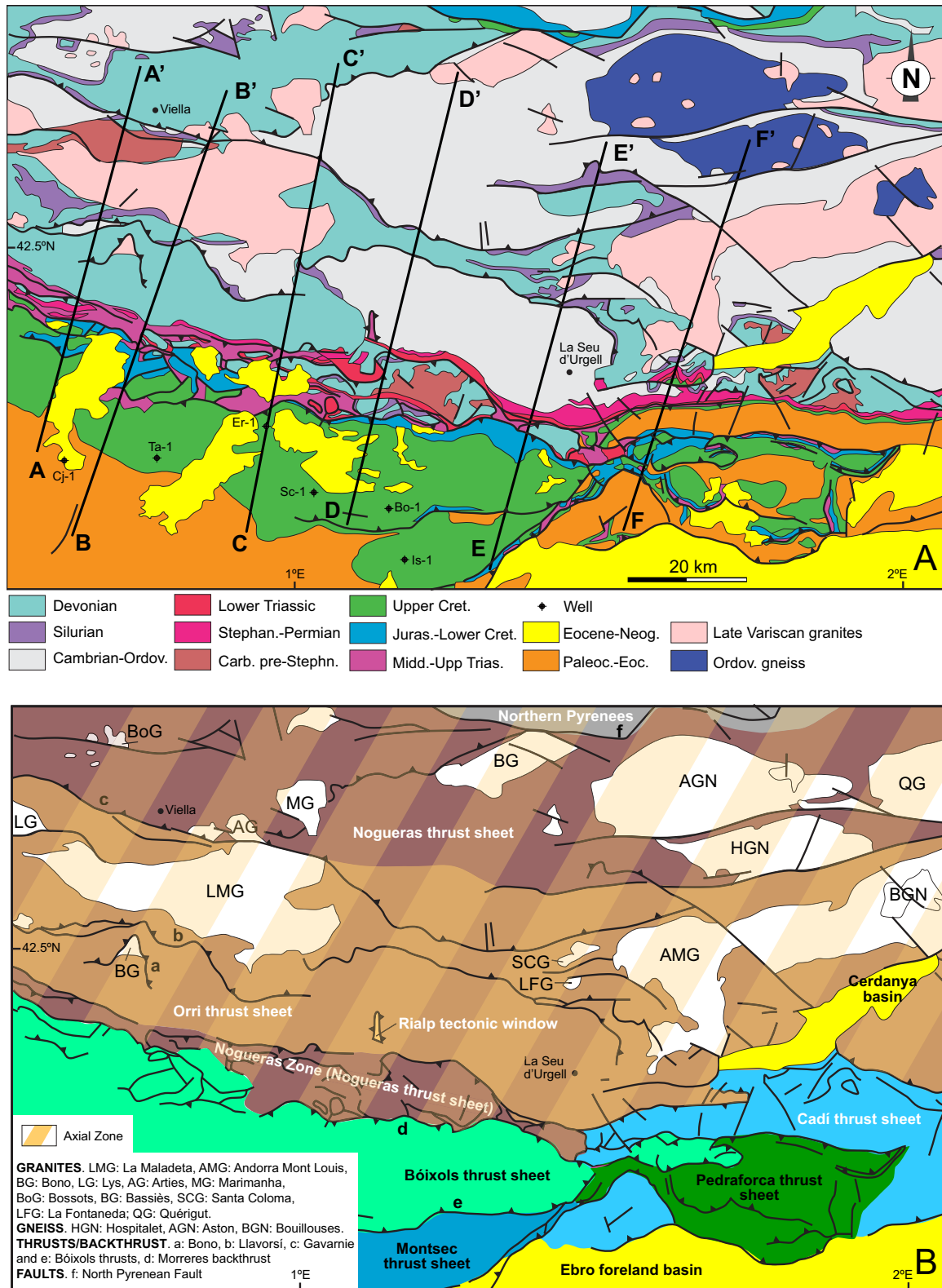
In the Axial Zone of the Central Pyrenees, Alpine deformation in the Paleozoic basement was accommodated by several south-verging thrust sheets varying along-strike in number and distribution (e.g., Espurt et al., 2019; García-Sansegundo, 1992; García-Senz et al., 2020; Izquierdo-Llavall et al., 2018; Labaume et al., 2016; Martínez-Peña & Casas-Sainz, 2003; Muñoz, 1992; Parish, 1984; Teixell, 1996; Teixell et al., 2018; Williams, 1985). In the western part of the study area four basement thrust sheets can be differentiated, which from top to bottom are (e.g., Muñoz et al., 2018): Nogueras, Orri, Rialp, and Ribagorza (not outcropping) (Figure 2b). Toward the east, only the first three basement units, Nogueras, Orri, and Rialp thrust sheets, are differentiated (Muñoz, 1992; Vergés, 1993; see Figure 1c). These basement thrust units form an antiformal stack whose sequential pattern is characterized by underthrusting of the lower and younger basement-involved thrust sheets (e.g., Muñoz, 1992). Consequently, the structures of the uppermost unit (the Nogueras thrust sheet) have been tilted in the southern leading edge of the antiformal stack showing downward facing structures and involving upper Paleozoic rocks and Permian to Lower Triassic red beds (Figure 2; Nogueras Zone; e.g., Seguret [1972]). Toward the west and east of the study area, these basement thrust sheets show a smaller vertical overlap and downward facing structures are not found (Martínez-Peña & Casas-Sainz, 2003; Saura & Teixell, 2006; Teixell, 1996; Vergés, 1993).

South of the Axial Zone, the Mesozoic-Cenozoic cover was decoupled from the basement along the Triassic evaporites and deformed in a thin-skinned style following a forward propagating thrust sequence from the Late Cretaceous to the Oligocene (e.g., Muñoz, 1992; Seguret, 1972; Vergés, 1993). The study area includes the Bóixols thrust sheet, which continues to the west into the Cotiella thrust sheet (e.g., García-Senz, 2002) and to the Upper Pedraforca thrust sheet to the east (e.g., Muñoz et al., 2018), a small portion of the Montsec thrust unit and its eastern continuation, the Middle Pedraforca thrust sheet (Muñoz et al., 2018) and the Lower Pedraforca thrust sheet (Figure 2b). The Bóixols and Montsec thrust sheets belong to a thrust salient developed as a progressive arc due both to a thrust displacement gradient and changes in the mechanical stratigraphy (Muñoz et al., 2013; Sussman et al., 2004). The Ebro foreland basin appears in the southeastern portion of the study area (Figure 2). These sediments continue in the foot-wall, below the allochthonous Pyrenean thrust sheets, as shown by seismic and well data (e.g., Muñoz et al., 2018; Vergés, 1993). Located in the Axial Zone of the study area, the Cerdanya basin (see Figure 2b) is filled with Neogene rocks related to an ENE-WSW oriented half-graben (e.g., Cabrera et al., 1988).

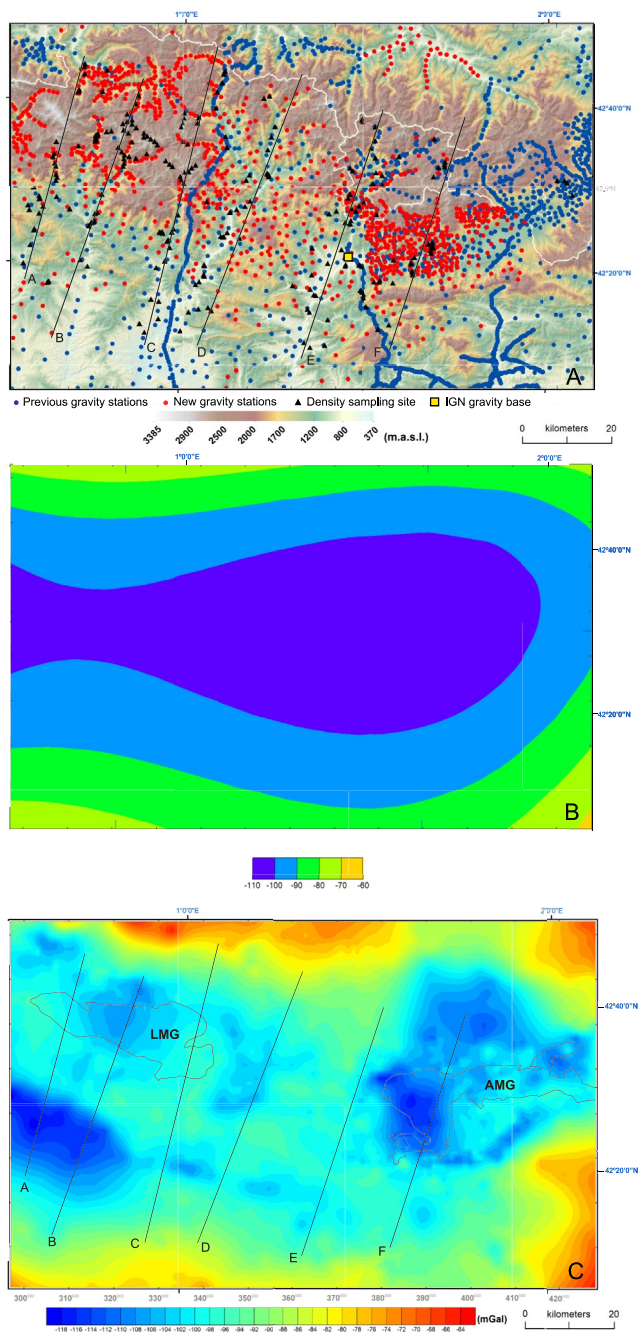
## 2.2. Stratigraphic Units

The Axial Zone is formed, in the study area, by sedimentary, igneous, and high-, medium- and low-grade metamorphic rocks of Cambrian to Carboniferous age deformed during the Variscan Orogeny (Middle-Late Carboniferous; Capellà, 1991; Clariana, 2015; Denèle, 2007; García-Sansegundo, 1992; Margalef, 2015) unconformably overlain by Stephanian, Permian, and Triassic rocks (Zwart, 1979) (Figure 2a). The oldest outcropping sedimentary record corresponds to Cambrian and Ordovician siliciclastic rocks with some intercalations of limestones and microconglomerates. Above, the Silurian black shales constitute an important detachment level (e.g., García-Sansegundo, 1990; Matte, 1969), as mentioned previously, below the Devonian slates and limestones (Zwart, 1979). The Carboniferous series, older than the Stephanian succession, mainly consist of siliciclastic rocks (Devolvé, 1987). The Late Carboniferous (Stephanian)-Permian succession is mainly formed by andesitic lava flows and pyroclastic units with some interbedded coal beds and red sandstones, shales, and microconglomerates (e.g., Gisbert, 1983; Mey et al., 1968). This succession is unconformably covered by the Triassic sediments, which show the typical Germanic facies (Calvet et al., 2004): Lower Triassic red beds (Buntsandstein facies), Middle Triassic dolostones and limestones (Muschelkalk facies) and Upper Triassic evaporites and shales (Keuper facies). Tholeiitic subvolcanic rocks (i.e., dolerites) are relatively common within the Upper Triassic series (Lago et al., 2000).

Intrusive rocks are represented by syn-to late-Variscan granitoids (Late Carboniferous-Early Permian in age) (Figure 2), which intrude the pre-Late Carboniferous succession (e.g., Evans, 1993; López-Sánchez et al., 2019; Solé et al., 1997) (Figure 2) and produced local contact metamorphism (e.g., Autran et al., 1970). The metamorphic rocks showing the highest metamorphic grade are located in the northeastern sector of the study area where



**Figure 2.** (a) Geological map with the location of the modeled sections A-A' to F-F' and Tamurcia-1, Cajigar-1, Erinya-1, San Corneli-1, Bóixols-1 and Isona-1 wells (Comiols-1 well is located outside the study area) (well data; see Lanaja [1987]; Mencos et al. [2015]). (b) Sketch showing the main structural units of the study area.



**Figure 3.** (a) Relief map of the study area showing the previous and new acquired gravity stations considered in this work and the density sampling sites. In yellow square IGN (Instituto Geográfico Nacional) gravity base of La Seu d’Urgell. The color scale refers to the elevation in meters above sea level. (b) Regional anomaly map of the study area. (c) Bouguer anomaly map of the study area. Contour interval is 2.50 mGal. The location of the modeled sections A-A’ to F-F’ is shown. LMG = La Maladeta granite and AMG = Andorra Mont-Louis granite.

three Ordovician gneiss domes crop out (Figure 2). These gneiss domes are characterized by a core of ortho- to paragneisses (Denèle et al., 2014) and an envelope of Cambrian and Ordovician metasedimentary rocks affected by a high temperature-low pressure metamorphic gradient from the biotite zone to the garnet, andalusite-staurolite-cordierite, staurolite and sillimanite zones (Alías, 1995; Clariana, 2015; Guitard et al., 1996).

The Mesozoic-Cenozoic cover in the study area consists of thick sequences of Jurassic and Lower Cretaceous, mainly marine limestones and marls, deposited in relation to previous extensional basins (e.g., García-Senz, 2002). The Upper Cretaceous series is characterized by a succession of limestones, calcarenites, and marls (Simó, 2004). The Paleocene-Eocene series mainly consists of limestones, sandstones, and marls of marine origin, together with non-marine sandstones and conglomerates; whereas the youngest deposits considered in the study area are Oligocene and Neogene continental units, mainly conglomerates and sandstones (e.g., Barnolas & Gil-Peña, 2001; Martínez et al., 1988; Puigdefàbregas & Souquet, 1986).

### 3. Gravity Data Acquisition and Processing

Gravity data used in this study is a compilation of public gravity datasets (ICC, 1987; BRGM, 2009; IGC-ICC-ICTJA-CSIC, 2012), previous works (Ayala et al., 2020; Casas et al., 1987, 1997; Rivero Marginedas, 1994; Torné et al., 1989) and new gravity surveys. In total, 3,590 gravity stations have been used in an area of ~10,000 km<sup>2</sup>, 2,449 previous and 1,141 new stations. Gravity station distribution is not homogeneous showing a higher number of measurements on La Maladeta and Andorra-Mont Louis granites only along accessible tracks (Figure 3). Elevation and spatial positioning of the new gravity data were obtained using a differential GNSS instrument (Triumph from Javad). New gravity measurements were taken with a Scintrex CG5 (63% stations), a Lacoste & Romberg G582 (25% stations) and a Scintrex CG6 (12% stations). Data were linked to La Seu d’Urgell base that belongs to the IGN’s Fundamental Network directly linked to the IGSN71 network and other stations that coincide with the ones of the IGN REDNAP high precision leveling network. Gravity measurements were corrected for effects caused by instrumental drift and tides to obtain the observed gravity.

A complete Bouguer anomaly was recalculated for the entire dataset using the Geodetic Reference System GRS80 (Moritz, 1980), and a density of reduction of 2.67 g/cm<sup>3</sup> (Figure 3). The terrain correction was applied up to 167 km using the Oasis Montaj module based on Kane (1962) and Nagy (1969). We use a combined 5 m DTM from Spain, France, and Andorra that was regridded to 20 × 20 m as a local grid considering a radius of 2 km, and a combined grid from the SRTM database (e.g., Farr et al., 2007) and bathymetry data from GEBCO database (Sandwell et al., 2002) regridded to 100 × 100 m as regional grid considering a radius of 167 km.

Bouguer anomalies result from the combination of sources at different depths related to density contrasts; the deep-seated geological bodies or huge shallow sedimentary basins generate the long-wavelength components of the anomaly and the small and near-surface geological bodies originate the short-wavelength components.

In order to investigate gravity anomalies related to the uppermost crust, a regional-residual separation was obtained. The selected regional gravity anomaly (assumed as the contribution of deeper sources) consists of a third order polynomial surface (Figure 3b). The regional anomaly was chosen after doing a careful analysis of

different regional anomalies (i.e., isostatic calculation based on the Airy hypothesis, assuming different polynomial degrees and using a different type of filters) and estimating the radially averaged power spectrum. The regional anomaly was subtracted from the Bouguer anomaly (Figure 3c) in order to obtain the residual Bouguer anomaly (Figure 4), which in our work we assume allows modeling down to  $\sim 10,000$  m depth, where the structures that are the target of this study are located. Additionally, vertical and horizontal gradients of the residual anomaly were calculated (Figures 5a–b5c). This enhances the edges between geological bodies of different densities and defines particular structural alignments that can be associated with faults or lithological contacts at depth. Taking into account the main structural trend of the Pyrenees, a N110 E direction, we also calculated the horizontal gradients both in the N110 E and N200 E (i.e., parallel and perpendicular to the Pyrenean trend respectively) directions (Figures 5b and 5c).

#### 4. Bouguer and Residual Gravity Anomaly Maps

In the entire study area, the Bouguer anomaly shows negative values between approximately  $-120$  and  $-67.5$  mGal (Figure 3c). The central part of the study area is characterized by a long-wavelength elongated relative minimum associated with the crustal Pyrenean root (see Torné et al., 1989) which follows a roughly N110 E direction. The orientation of this long minimum is parallel to the southern limit of the Axial Zone and most Alpine structures (Figures 1 and 3). Superimposed to this minimum, several relative maxima, and minima of medium and short wavelengths with variable amplitudes can be observed (Figure 3c). The most prominent of these relative minima are (a) a N110 E oriented relative minimum located to the South of the southern limit of the Axial Zone (i.e., located on the Nogueras Zone and Bóixols thrust sheet; see Figures 1 and 2) which reaches more than  $-115$  mGal and amplitude of  $\sim 25$  mGal and (b) a minimum formed by two rounded gravity lows aligned in an NNE direction reaching more than  $-110$  mGal and an amplitude of 15–20 mGal located inside the Axial Zone and coinciding at the South with the western part of the Andorra-Mont Louis granite (Figures 2 and 3). The first of these relative minima can be continued to the East following a roughly N110 E orientation and along a total of  $\sim 125$  km before it is closed to the East by a N-S positive gradient. The gravity response of the La Maladeta granite coincides with a relative minimum showing variations in the amplitude of only 5 to 10 mGal, smaller than that shown by the Andorra-Mont Louis granite (Figures 2 and 3).

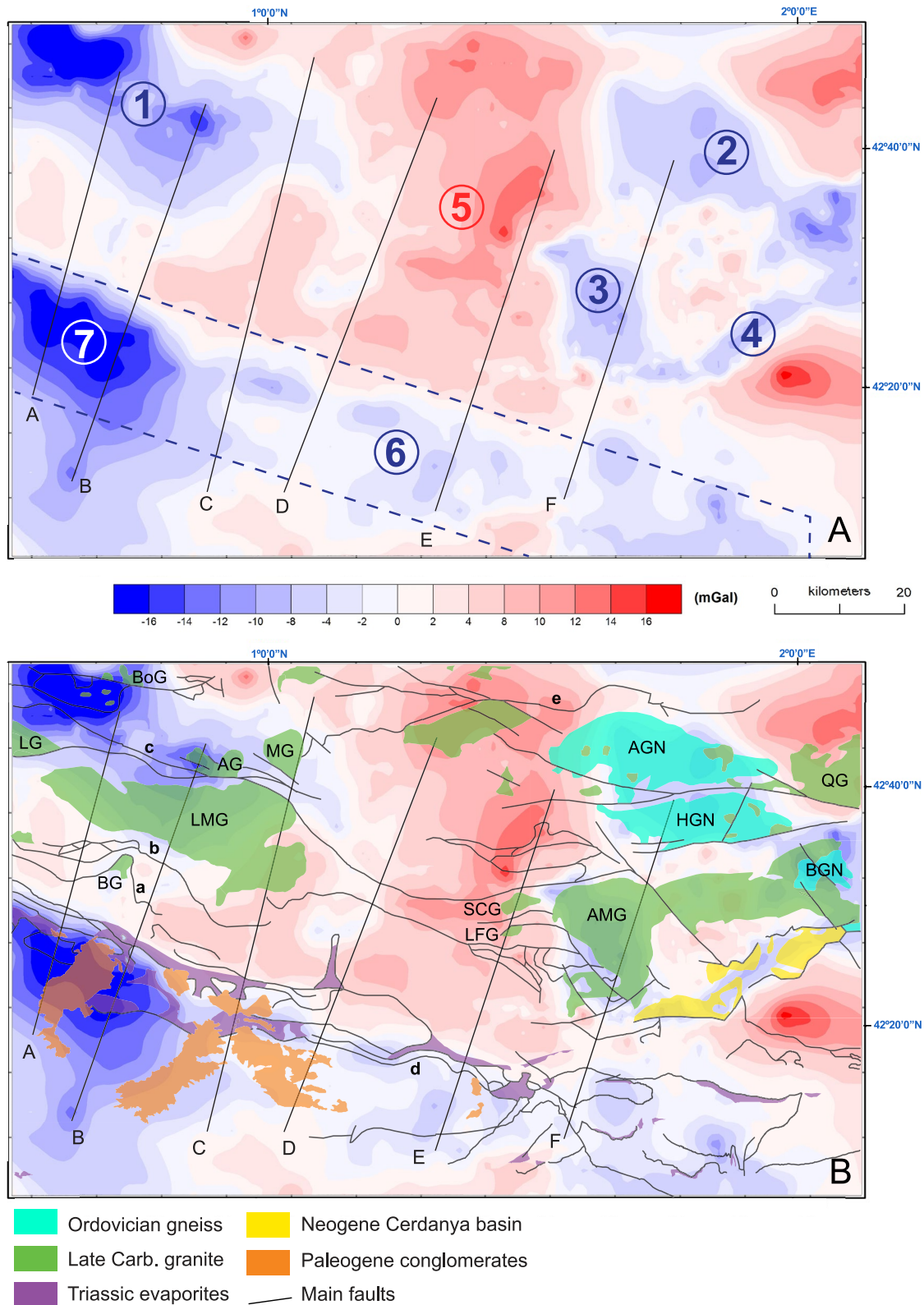
The residual gravity anomalies range from  $-18$  to 18 mGal (Figure 4). The residual gravity anomaly map shows two areas with a different residual anomaly pattern; the Axial Zone and the South Pyrenean Zone (Figure 4). Both areas are separated by a Pyrenean-oriented limit, which roughly coincides with the southern termination of the Axial Zone with values around 0 mGal. In the Axial Zone, the residual anomaly pattern is complex (Figure 4a). Gravity lows labeled as 1 and 2 in Figure 4a show a Pyrenean-oriented trend. Gravity low 1 reaches  $-18$  mGal and coincides at the surface with the outcrops of the Arties and Bossost granites (Figure 4b). Gravity low 2 shows lesser amplitude than the previous one ( $\sim -4$  mGal) and coincides with the Aston and Hospitalet gneisses (Figure 4b). To the East, the gravity minima 3 and 4 coincide with the Andorra-Mont Louis granite and the Neogene Cerdanya basin, respectively (Figure 4). The gravity maximum 5 displays a noteworthy NNE-SSW in the central sector of the studied Axial Zone (Figure 4). In the South Pyrenean Zone, the large Pyrenean-oriented gravity minimum labeled as 6 coincides with the Bóixols and Pedraforca thrust sheets and partly with the Nogueras Zone (Figure 4). It shows its lowest gravity values ( $-18$  mGal) in a large minimum labeled as 7 in Figure 4a. Positive gravity values surround the gravity low six except at its SW corner (Figure 4).

The vertical gradient of the residual anomaly map enhances the negative gravity anomaly of the Late Carboniferous-Early Permian granites, Ordovician gneisses, and the Cerdanya Cenozoic basin in the Axial Zone and the Triassic evaporites at depth in the South Pyrenean Zone (Figure 5a). Whereas the horizontal gradient map along the N200 E direction enhances the negative gravity anomaly 6, which coincides with the WNW-ESE structural band comprising the Nogueras Zone (Figure 5b). The equivalent map along the N110 E direction magnifies the positive gravity anomaly 5 (Figure 5c).

#### 5. Gravity Modeling

##### 5.1. Density Properties

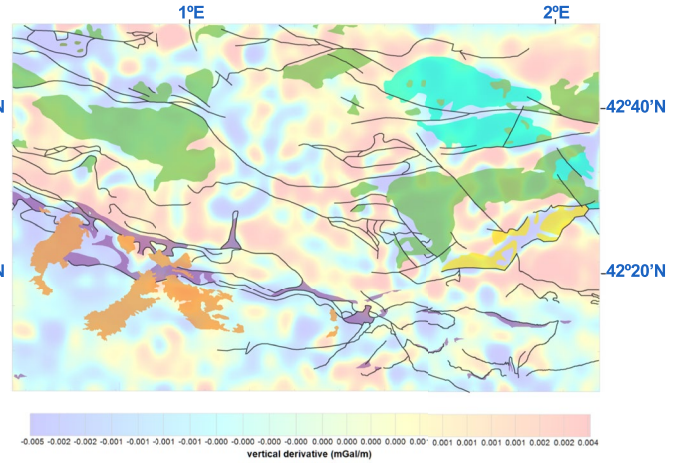
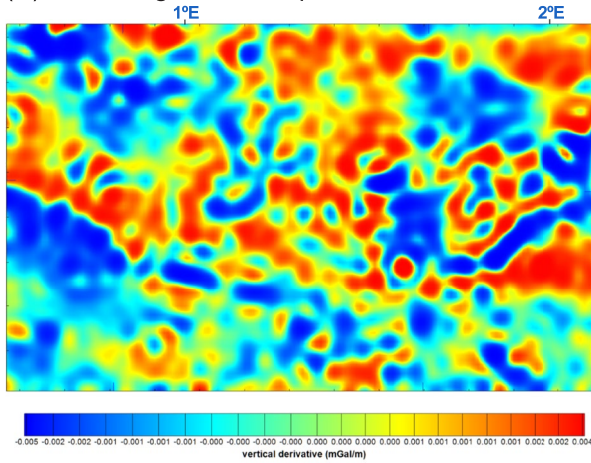
Table 1 shows the range and main statistics (arithmetic mean, standard deviation, and mode) of the measured density of representative rocks. They have been grouped into 13 rock units ranging between the Cambrian and the



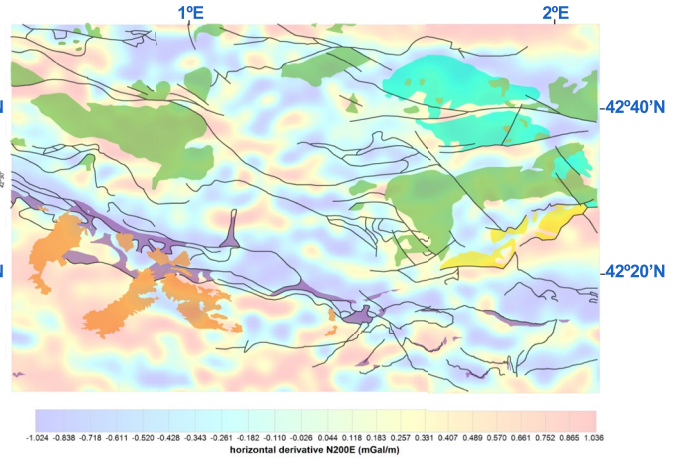
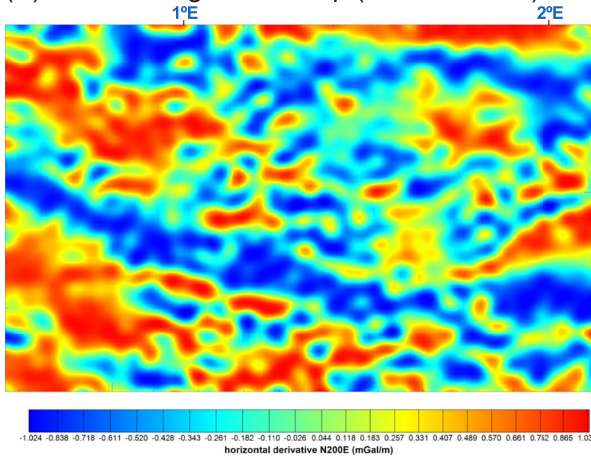
**Figure 4.** (a) Residual anomaly map of the study area. Main gravity minima and maximum are labeled to facilitate their description through the text. Contour interval is 2.0 mGal. (b) Residual anomaly map of the study area superposed with the main structures (labeled as in Figure 2B), main igneous bodies (labeled as in Figure 2B) and Triassic evaporitic outcrops of the study area. The location of the modeled sections A-A' to F-F' is shown.



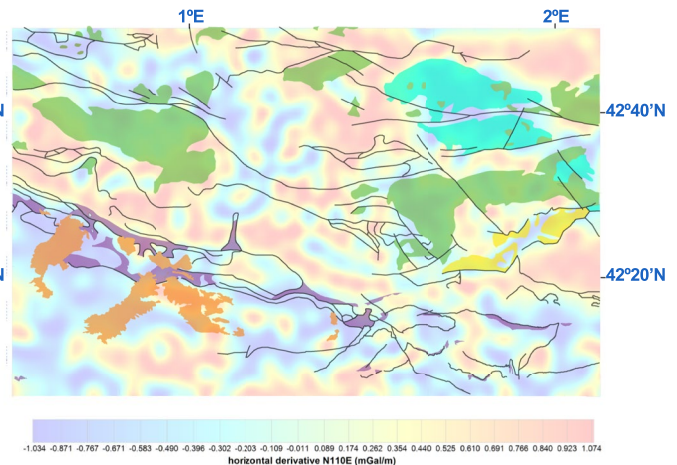
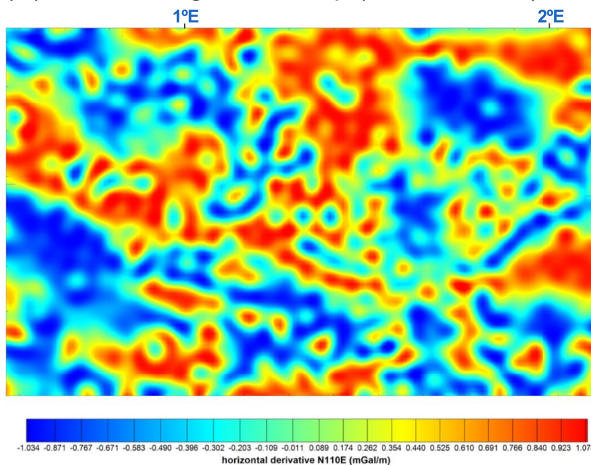
(A). Vertical gradient map



(B). Horizontal gradient map (N200E direct.)



(C). Horizontal gradient map (N110E direct.)



**Figure 5.** (a, b, c) Vertical, N200 E and N110 E horizontal gravity residual gradients. To the right, same maps superposed with the main structures (labeled as in Figure 2B), main igneous bodies (labeled as in Figure 2B) and Triassic evaporitic outcrops of the study area. Maps covering the same area as maps of Figures 3 and 4.

Oligocene-Neogene. Density was measured on 526 irregular non-weathered samples using the double weighting method (standard UNE-EN, 1936, p. 2007) obtaining the bulk density from each of them. All density values range between 1.85 and 3.03 g/cm<sup>3</sup>. The highest density mean values of 2.69, 2.68 and 2.67 g/cm<sup>3</sup> correspond to the Devonian, pre-Stephanian Carboniferous, and Cambro-Ordovician rocks, respectively (Table 1). Upper Triassic

**Table 1**  
Density Values ( $\text{g/cm}^3$ ) of Sampled Rocks and Statistics

Units	N	Mean ( $\text{g/cm}^3$ )	Standard deviation ( $\text{g/cm}^3$ )	Mode ( $\text{g/cm}^3$ )	Range ( $\text{g/cm}^3$ )
<i>Cambro-Ordovician</i>	75	2.67	0.08	2.68	2.35–2.81
<i>Silurian</i>	29	2.45	0.19	2.44	2.04–2.74
<i>Devonian</i>	86	2.69	0.102	2.69	2.41–3.03
<i>Pre-Stephanian Carboniferous</i>	5	2.68	0.013	2.70	2.67–2.7
<i>Late Carboniferous-Permian</i>	52	2.52	0.182	2.57	1.85–2.82
<i>Early Triassic</i>	24	2.58	0.110	2.69	2.31–2.69
<i>Middle-Late Triassic</i>	39	2.40	0.313	2.47	1.93–3.00
<i>Jurassic-Late Cretaceous</i>	17	2.66	0.059	2.65	2.51–2.78
<i>Late Cretaceous</i>	32	2.59	0.094	2.69	2.34–2.7
<i>La Maladeta granite</i>	48	2.65	0.093	2.64	2.29–2.89
<i>Andorra-Mont Louis granite</i>	58	2.63	0.051	-	2.50–2.74
<i>Cenozoic</i>	7	2.65	0.024	-	2.62–2.68
<i>Ordovician gneiss</i>	54	2.62	0.064	-	2.39–2.78

and Silurian rocks are the less dense ones, with mean density values of 2.40 and 2.45  $\text{g/cm}^3$ . The ranges of each rock unit vary widely, being the Middle-Upper Triassic rocks the unit showing the largest range (i.e., between 1.93 and 3.00  $\text{g/cm}^3$ ). This wide range responds to the high variety of lithologies included in the Middle-Upper Triassic unit; evaporites, mudstones, lutites, limestones, dolostones, and dolerites. Pyrenean pre-Cambrian rocks crop out to the East of the study area (e.g., Padel et al., 2018), but no samples have been collected in this work. We interpret that they could be present at depth in the upper crust in our sections and have been assigned a density value of 2.70  $\text{g/cm}^3$ . The density of collected samples indicates density mean values of 2.65 and 2.63  $\text{g/cm}^3$  for the La Maladeta and Andorra-Mont Louis granites, respectively, and 2.62  $\text{g/cm}^3$  for the Ordovician gneisses (Table 1).

## 5.2. Geological Cross Section Construction

Six almost parallel sections (A-A' to F-F') have been selected for modeling in order to cover the entire study area. Sections strike N010 E and N018 E, thus being perpendicular to most structures, are ~68–80 km long and keep a distance of ~20 km. Sections A and B were already created and modeled in a previous work (Clariana et al., 2021). The same geological units shown in the geological map of Figure 2 are considered in the cross sections.

Sections consist of composite geological cross sections constructed taking into account field data, surface geological data coming from geological maps (1:50.000 scale, MAGNA series, Instituto Geológico y Minero de España, and the geologic digital continuous map of Catalunya from Institut Cartogràfic i Geològic de Catalunya), well data (Tamura-1, Cajigar-1, Erinyà-1, San Corneli-1, Bóixols-1, Comiols-1, and Isona-1; Lanaja, 1987) (see well location in Figure 2) and previous interpreted cross sections from several authors (Table 2).

## 5.3. Processing and Interpretation of Modeled Sections

Forward gravity modeling was carried out using the GM-SYS module (implemented in the Oasis Montaj® software by Geosoft), which calculates the gravimetric response of given density models with an algorithm based on Talwani et al. (1959) and Won and Bevis (1987). The modeling has been carried out through a feedback process between geology and gravity that ended when the gravimetric response of each cross section (i.e., calculated gravity data) fitted the observations (i.e., observed gravity data). The different layers of each model were extended far enough at the end of the profiles to avoid edge effects. Since the density values are well constrained by laboratory measurements, in order to adjust the observed anomalies, we kept the mean densities fixed and modified

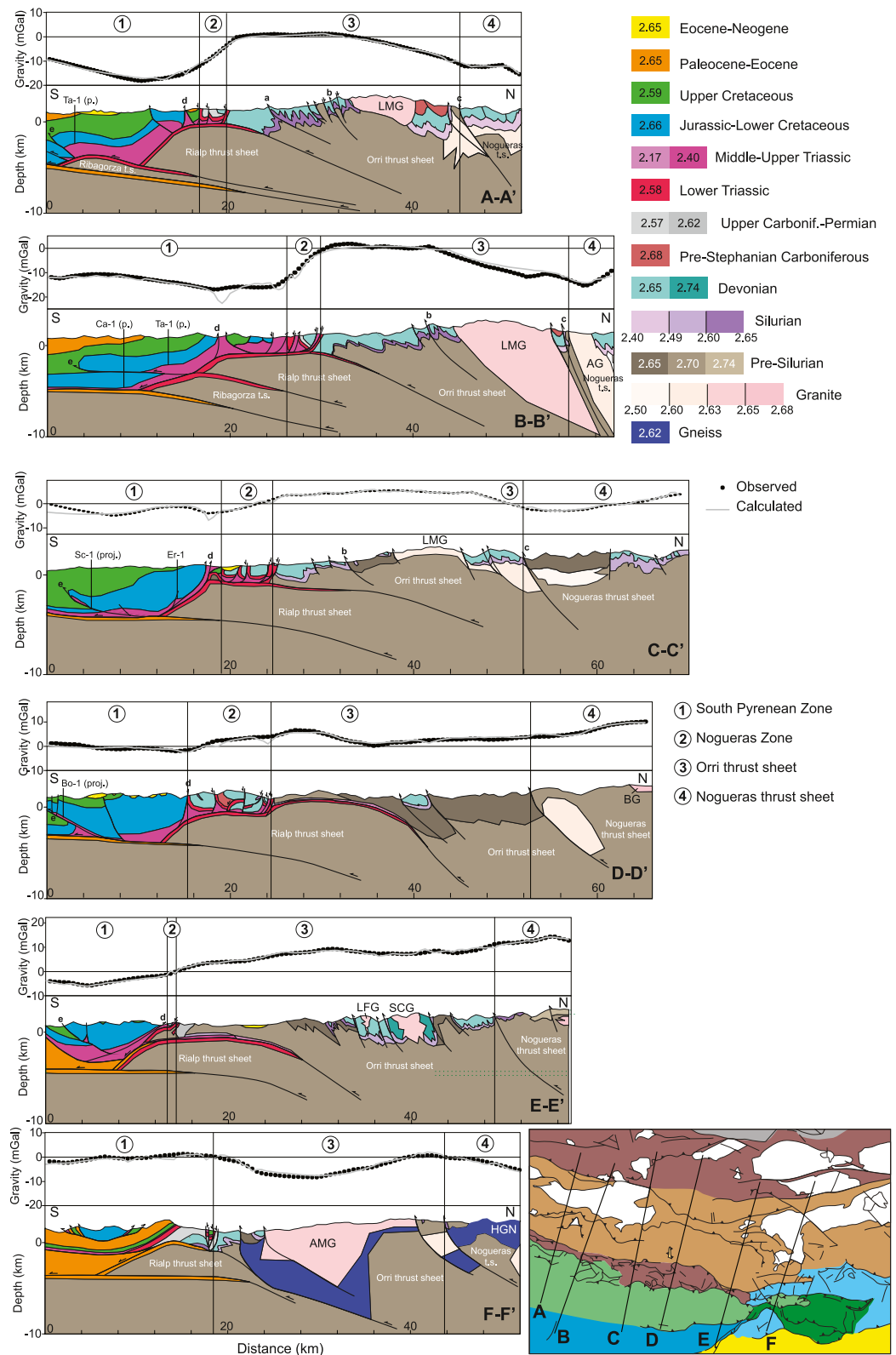
**Table 2**

*Data Considered for the Construction of All Sections: Coordinates, Geological Maps, Wells and Previous Interpreted Cross Sections*

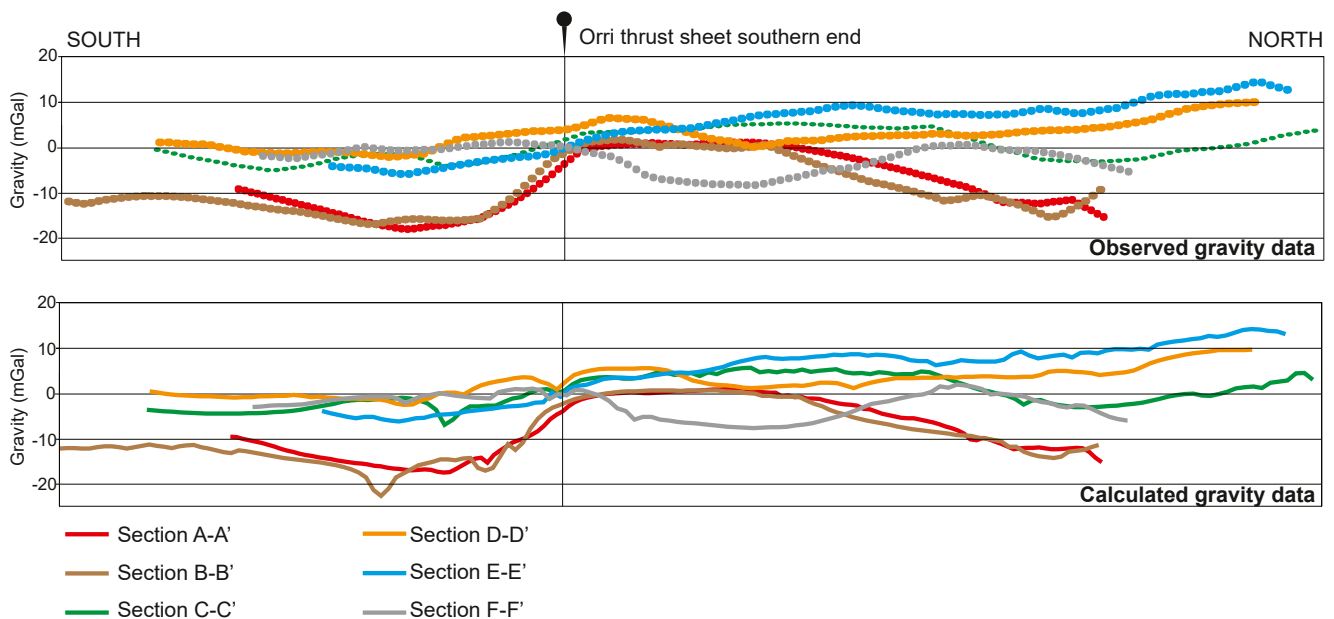
Section	Coordinates ETRS89 H31	Geological maps	Well data	Previous cross-sections
Section A-A'	North X: 313,500,014 Y: 4,735,852,718 South X: 300,077,947 Y: 4,685,835,61	-García-Sansegundo et al. (2013)-MAGNA 118-bis148 -Ríos-Aragües et al. (2002)-MAGNA 180. -García-Senz and Ramirez-Merino (2009)-MAGNA 213. -López-Olmedo and Ardevol (2016)-MAGNA 251. -Alta Ribagorça-ICGC (2007).	-Tamurcia-1 (projection)	-García-Sansegundo (1992). -García-Senz (2002). -Teixell and Muñoz (2000). -Izquierdo-Llavall et al. (2013).
Section B-B'	North X: 326,764,021 Y: 4,730,763,27 South X: 306,125,282 Y: 4,672,572,645	-Sanz-López and Palau (2013)-MAGNA 149. -Martín-Parra et al. (2016)-MAGNA 181. -Rosell (1994)-MAGNA 252. -Alta Ribagorça-ICGC (2007).	-Cajigar-1 (projection) -Tamurcia-1 (projection)	-Poblet (1991). -Soler et al. (1998). -Gil-Peña and Barnolas (2001). -García-Senz (2002). -Teixell and Muñoz (2000). -Saura et al. (2016).
Section C-C'	North X: 343,484,853 Y: 4,737,960,244 South X: 327,022,759 Y: 4,670,932,804	-Sanz-López and Palau (2013)-MAGNA 149. -Martín-Parra et al. (2016)-MAGNA 181. -Rosell (1994)-MAGNA 252. -Muñoz et al. (2009). -Muñoz et al. (2010). -Pallars-Sobirà-ICGC (2007). -Pallars-Jussà-ICGC (2007).	-Erinyà-1 -San Corneli-1 (projection)	-García-Senz (2002). -Mencos et al. (2015). -Muñoz et al. (2018). -Beamud et al. (2010). -Saura (2004).
Section D-D'	North X: 362,146,385 Y: 4,731,097,02 South X: 338,873,82 Y: 4,670,849,247	-Clariana (2015). -Muñoz et al. (2010). -Pallars-Sobirà-ICGC (2007). -Pallars-Jussà-ICGC (2007).	-Bóixols-1 (projection)	-Poblet (1991). -Muñoz (1992). -Mencos et al. (2015). -Gil-Peña and Barnolas (2004). -Saura (2004).
Section E-E'	North X: 380,598,66; Y: 4,723,660,49 South X: 362,674,16 Y: 4,669,353,45	-Clariana (2015). -Margalef (2015). -Alt-Urgell-ICGC (2007).		-Poblet (1991). -Vergés (1993). -García-Senz (2002). -Saura (2004). -Muñoz et al. (2018). -Clariana (2015). -Margalef (2015). -Casas et al. (2019).
Section F-F'	North X: 399,003 Y: 4,722,092 South X: 382,067 Y: 4,669,801	-Cirés et al. (1994), MAGNA 217. -Martínez et al. (2020), MAGNA 254. -Alt-Urgell-ICGC (2007).		-Cirés et al. (1990). -Soler (1990). -Poblet (1991). -Vergés (1993). -Margalef (2015).

the depth geometries in areas with scarce subsurface data. In some cases, when using the mean density was incompatible with a geological geometry, we modified the densities within the range of the petrophysical data (Table 1). The RMS fitting error (Root Mean Square of the difference between calculated and observed anomalies) in the six sections ranged between 0.55 and 1.3 mGal, less than 7% of the observed anomaly amplitudes. The most prominent misfits are small and correspond to the lack of a denser coverage of gravity stations in the Middle-Upper Triassic diapirs.

All sections, except the easternmost section F-F', cross the same structural units. From South to North these units are (Figure 6): the Montsec and Bóixols thrust sheets included in the South Pyrenean Zone and the Nogueras Zone and the Orri and Nogueras thrust sheets belonging to the Axial Zone. Section F-F' crosses from South to North, the Cadí, Pedraforca, and Bóixols thrust sheets in the South Pyrenean Zone and the Orri and Noguera's thrust sheets in the Axial Zone. All sections show at their southern sector an autochthonous succession of Paleogene sediments in continuation with the Ebro foreland basin underneath the basal Pyrenean southern thrust as



**Figure 6.** A-A' to F-F' modeled sections. Both the observed (points) and calculated (gray line) residual gravity anomalies are shown on top of each model section. See location of sections on Figure 2. Density values are in g/cm<sup>3</sup>. When different density values are used in the same section, they are indicated on the caption. A schematic geologic map is also shown to help following results. Outcropping granites are labeled according to Figure 2.



**Figure 7.** Calculated and observed gravity data of all sections together. All curves have been fixed in the southern end of the Orri thrust sheet.

evidenced by well (data projected from Comiols-1 and Isona-1 wells; see Figure 2 for location) and seismic data (e.g., Muñoz et al., 2018). The depth of this autochthonous succession ranges between 3,500 and 5,000 m showing a general slight westwards deepening following data and results obtained by Muñoz et al. (2018). The cut-off line between the Paleogene autochthonous succession and the south Pyrenean basal thrust cannot be delineated exactly from the available seismic data, but it can be mapped northwards of the branch line between the southernmost basement-involved thrust sheets and the south pyrenean basal thrust (see Muñoz et al., 2018).

In general, gravity data fit well with the proposed geometry interpreted by previous authors at the different parts of all composite sections. Most sectors showed the observed residual gravity data coinciding with the calculated one (Figure 6). The changes done in the sections to fit the calculated with the observed gravity data can be grouped into: (a) areal changes in the Middle-Upper Triassic evaporites at depth in the South Pyrenean Zone in areas without proximal well data, (b) subsurface areal changes in the outcropping granites and gneisses, (c) addition of buried granites in areas with surface evidences of their presence (e.g., northern sector of section A-A'), (d) addition of buried granites without surface evidences of their presence (e.g., northern sector of sections D-D' and F-F'), (e) presence of a thin layer of Middle-Upper Triassic evaporites at depth in the Orri thrust sheet in sections D-D' and E-E' inferred from the outcrops of Middle-Upper Triassic evaporites in the Rialp tectonic window (see Figure 2), (f) addition of a gneiss under the Andorra-Mont Louis granite (section F-F') already suggested by surface evidences (Denèle et al., 2014; Guitard et al., 1996), and (g) density value changes in the Middle-Upper Triassic (between 2.17 and 2.40 g/cm<sup>3</sup>), Upper Carboniferous-Permian (between 2.57 and 2.62 g/cm<sup>3</sup>), Silurian (between 2.40, 2.49, 2.60, and 2.65 g/cm<sup>3</sup>), pre-Silurian (between 2.65, 2.70, and 2.74 g/cm<sup>3</sup>) and granite (between 2.50, 2.60, 2.63, 2.65, and 2.68 g/cm<sup>3</sup>) rocks considering the ranges obtained from collected samples (Table 1).

### 5.3.1. South Pyrenean Zone

The residual anomaly map shows a very strong along-strike variation of gravity residual values in the South Pyrenean Zone (Figure 4). The gravity low 1 shows a similarly observed gravity curve in sections A-A' and B-B' reaching values of −18 and −16 mGal, respectively. In section A-A', most of this gravity low does not correspond to Triassic evaporites at the surface, but its high amplitude suggests a large volume of this lithology at depth (Figure 6). The observed residual gravity curves of sections C-C', D-D' and F-F' show similar behavior, with data ranging around 0 mGal (Figure 7) which have been interpreted as linked to smaller volumes of Triassic evaporites at depth with respect to sections A-A' and B-B' (Figure 6). Section E-E', which shows lower gravity values ranging roughly between −5 and −2 mGal, has been modeled with an intermediate volume of Triassic

evaporites at depth (Figure 6). The modeled cross sections showing higher thicknesses of Triassic evaporites in the South Pyrenean Zone coincide with thinner Jurassic-Lower Cretaceous succession. The northern part in all curves, except that corresponding to section F-F', shows an upward trend toward higher gravity values. These upward trends match at depth with the southernmost leading edges of the youngest basement-involved thrust sheets; the Rialp and Ribagorza thrust sheets in sections A-A' and B-B' and the Rialp thrust sheet in sections C-C', D-D', and E-E' (Figure 6).

### 5.3.2. Nogueras Zone

Sections show a lateral variation on the geometry and width of the Nogueras Zone as previous authors have pointed out (Izquierdo-Llavall et al., 2018; Muñoz, 1992; Saura, 2004; Saura & Teixell, 2006; Seguret, 1972). Modeled sections also show that, in order to match the observed and calculated gravity data, an along-strike variation on the depth reached by the different thrust slices forming the Nogueras Zone is necessary. Sections C-C' and D-D' show a wider and deeper set of structures belonging to the Nogueras Zone with respect to sections A-A' and B-B'. Section E-E' only crosses the easternmost end of a small thrust slice of this unit (Figure 6). Modeled sections also show that the position of the Nogueras Zone must coincide, with a flat floor thrust (Figure 6). The depth to this flat of the floor thrust varies along-strike in parallel with the size of the thrust slices of the Nogueras Zone between 2,200 and 1,400 m-depth in sections D-D' and A-A', respectively.

The observed and calculated gravity curves of modeled sections crossing the Nogueras Zone show in all cases an increasing trend toward its limit with the Orri basement thrust sheet to the North. It is worth noting that the gravity minimum 7 (located in the South Pyrenean Zone) coincides with the western slices of the close Nogueras Zone that form a narrower band (sections A-A' and B-B'). Whereas the eastern half of the Nogueras Zone, outcropping as a much wider band, is not accompanied by such a gravity minimum (Figure 4).

### 5.3.3. Orri and Nogueras Thrust Sheet

Regarding the Orri unit, all sections, except F-F', coincide at their southern half with positive gravity residual values. These positive values show a discontinuous roughly Pyrenean WNW-ESE trend interrupted by several perpendicular minima of negative values (Figure 4). Sections A-A', B-B' and C-C' cross the gravity low 1 which coincides with the outcrops of the La Maladeta, Arties and Marimán granites. The eastern end of the La Maladeta granite corresponds with a slight relative gravity high (section C-C'). In order to adjust the observed and calculated gravity curves a lower density value was chosen for the La Maladeta granite there. Section D-D' also shows a relative gravity minimum of positive values at the northern half of the Orri thrust sheet, whereas section E-E' displays higher positive values in equivalent positions (Figure 6). Section F-F' shows a different behavior; it traverses the gravity minimum 3 coinciding with the outcrop of the Andorra-Mont Louis granite and a gravity relative maximum between gravity lows 2 and 3 (Figure 6). In general, outcrops of Cambro-Ordovician rocks in the Orri thrust sheet are more abundant in the studied area toward the East, and the comparison of the observed and calculated gravity curves roughly highlights this tendency (i.e., higher gravity values toward the East) (Figure 7). In the Nogueras thrust sheet, the observed and calculated gravity curves of all sections also show the same tendency. Sections A-A' and B-B' display the lowest negative gravity values; in section B-B' they coincide with the outcrop of Arties granite and in section A-A' they have been interpreted to be linked to a buried granite at depth possibly related to the Bossost granite located to the north (Figure 6). Except for values of section F-F', sections C-C', D-D' and E-E' show a gradual increment of gravity values in the Nogueras thrust sheet (Figure 7).

## 6. Implications and Discussion

### 6.1. Along-Strike Gravity Differences in the Bóixols Thrust Sheet

One of the most remarkable features that can be observed in the residual anomaly map of the study area is the presence of the gravity low 7 located in the Bóixols thrust sheet (Figure 4). Toward the East, the Bóixols thrust sheet still coincides with gravity negative values, but much higher. According to the modeled cross sections, a gravity minimum 7 is interpreted as related to an important volume of Middle-Upper Triassic rocks at depth (Figure 6). Exploration wells located in that area (i.e., Tamurcia-1 and Cajigar-1; Lanaja, 1987) reached the Triassic evaporites but they did not cross the whole unit to infer its total thickness. It is worth noting that the modeled cross sections show an inverse proportion between the thickness of the Middle-Upper Triassic rocks and that of the Jurassic-Lower Cretaceous rocks. The westernmost sections (A-A' and B-B') show the

higher thickness of Middle-Upper Triassic rock and lower thickness of Jurassic-Lower Cretaceous rocks when compared with sections that traverse the eastern half of the Bóixols thrust sheet, which show the lower thickness of Middle-Upper Triassic rock and higher thickness of Jurassic-Lower Cretaceous rocks. This relationship could point to a lateral westward migration of the Triassic evaporites due to differential loading prior to the Cenozoic Alpine compression and probably to an increase in early diapirism to the West (e.g., Saura et al., 2016). Unraveling the pre-compressional thickness variations of the Triassic evaporites inside the Pyrenean range constitutes a difficult task due to the subsequent development of compressional structures and possible across- and/or along-strike evaporitic flows during its syn-rift, post-rift, and compressional stages (e.g., Burrel & Teixell, 2021; García-Senz, 2002; Santolaria et al., 2016; Storti et al., 2007). A visualization of the volumes of Triassic evaporites at depth based on gravity data could help to solve the pre-compression scenario.

At depth, this gravity low 7 also coincides with the presence of the Ribagorza basement thrust sheet in the modeled cross sections A-A' and B-B' (Figure 6). The Ribagorza basement thrust sheet constitutes the youngest basement thrust sheet in the footwall of the Rialp thrust sheet at this sector (e.g., Muñoz et al., 2018). Its formation could be favored by the presence of an important volume of Middle-Upper Triassic evaporites there, as the Bóixols and Montsec cover thrust sheets would have been already emplaced, and/or the occurrence of a thinnest Jurassic-Lower Cretaceous succession at this portion of the Bóixols thrust sheet. To the East, the thicker Jurassic-Lower Cretaceous deposits could prevent the southwards propagation of basement thrusts and the formation of the Ribagorza basement thrust sheet. The emplacement of the Bóixols and Montsec thrust sheets is dated from Late Santonian to Maastrichtian time and from Ilerdian to Cuisian time, respectively (Bond & McClay, 1995; Farrell et al., 1987; García-Senz, 2002; Garrido-Megías, 1973; Nijman & Nio, 1976; Simó, 1986; Teixell & Muñoz, 2000; Verges & Munoz, 1990), whereas the Rialp and Ribagorza basement thrust sheets have been interpreted to be emplaced after the Rupelian (Muñoz, 2002).

Another remarkable feature arises from the coincidence between the location of the northern boundary of the gravity minimum 7 with the sector of the Nogueras Zone where it is narrower and characterized by strongly stacked and rotated minor thrust units (Figure 4b). Toward the East, the Nogueras Zone is much wider and its thrust slices are larger and show lower dips. The emplacement of the Nogueras basement thrust sheet has been dated as Late Cretaceous to Middle Eocene (Muñoz, 1992), but the thrust slices belonging to the Nogueras Zone were emplaced and were tilted later, between the Late Eocene and Early Oligocene (Muñoz, 2002; Saura & Teixell, 2006). When the small units of the Nogueras Zone started to emplace at the Late Eocene, the Bóixols and Montsec thrust sheets were already emplaced and probably, the western half of the Bóixols thrust sheet already presented higher volumes of Triassic evaporites at depth due to evaporitic flow during the syn-rift, post-rift stage and compressional stage previous to the Late Eocene. The fact that the western half of the Bóixols thrust sheet coincides with higher volumes of Triassic evaporites at depth and the presence of steeper and narrower thrust slices in the Nogueras Zone suggests that the Triassic salt distribution could influence the formation and geometry of the Nogueras Zone structures. García-Senz and Muñoz (2019) already suggested the presence of Late Jurassic-Early Cretaceous evaporitic salt walls coinciding with the location of the Nogueras Zone. An along-strike variation on the geometry of these salt walls derived from variable volumes of Triassic evaporites prior to their emplacement could have controlled the geometry of the thrust slices of the Nogueras Zone.

Finally, another feature derived from the modeled sections is that thrust slices belonging to the Nogueras Zone arrange above a flat floor thrust in all modeled sections. This geometry is similar to the geometry interpreted along the Ribagorzana valley (to the West) by Muñoz et al. (2018) and contrasts with that interpreted along the Pallaresa valley (to the East) located these units on a ramp of the floor thrust (Muñoz et al., 2018).

## 6.2. Variability of Gravity Values in the Orri and Nogueras Thrust Sheets

The residual anomaly map of the Orri and Nogueras thrust sheets of the study area reflects two different patterns. (a) Pyrenean WNW-ESE elongated gravity lows and highs more evident at the western half and easternmost sector of the study area (gravity lows 1 and 2) and (b) a complex pattern defined by WNW-ESE and NNE-SSW to N-S oriented gravity maxima and minima (gravity lows 3 and 4 and gravity high 5) in the central sector of the study area (Figure 4). The gravity minima 1, 2, and 3 coincide at the surface with outcrops of plutonic rocks (Figure 4). All these plutonic bodies with exception of Andorra-Mont Louis granite give a gravity response coinciding with WNW-ESE elongated negative anomalies, parallel to the main Pyrenean trend. It is worth noting that the La Maladeta and Andorra-Mont Louis granites point to a different gravity response probably reflecting differences

in their density values (see Table 1) and geometry at depth. The results obtained from gravity modeling show that the La Maladeta granite displays a laccolithic shape at its ends (sections A-A' and C-C') with its basal contact deepening to the North (Figure 6) and the Andorra Mont-Louis granite has a more batholithic shape (section F-F'; Figure 6). Section B-B' shows an elongated vertical shape for the Arties granite to fit the observed and calculated gravity curves. Below the Andorra Mont-Louis granite, the presence of a gneiss dome of lower density has been interpreted to match both gravity curves (see section F-F'; Figure 6). This feature has already been suggested by previous authors (Debon et al., 1996; Denèle et al., 2014) based on geological evidence. Ortogneisses as those surrounding the Andorra Mont-Louis granite display slightly lower densities than the density mean value obtained for this granite (see Table 1). The presence of granitic bodies both coinciding with surface outcrops (i.e., Marimanha, Bassiès, La Fontaneda, and Santa Coloma granites; Figures 2 and 6) and with buried granites of lesser volume compared to the modeled volume of La Maladeta and Andorra Mont-Louis granites, has been interpreted for sections C-C', D-D' and E-E' (Figure 6). Also the Bono granite (Figure 2) coincides with part of a small N-S gravity low suggesting its higher volume at depth (Figure 4). The gravity high 5 does not coincide with any structural trend at surface and could reflect inherited Variscan or pre-Variscan compositional crustal variations following this trend. Previous authors have suggested the important role played by lateral variations of the rheological inheritance and/or inherited faults on the structural architecture of the Pyrenees (e.g., Chevrot et al., 2014; Chevrot et al., 2018; Jammes et al., 2014). To the South of the Orri thrust sheet, Triassic evaporites crop out in the Rialp window (e.g., Muñoz, 1992). A gravity low showing an NNE-SSW orientation appears to the North of the Rialp window (Figures 2 and 4) pointing out a possible relationship.

Figure 7 shows all observed and calculated gravity data along the modeled sections together to facilitate their comparison. In the southern half of the Orri thrust sheet, all curves with exception of section F-F' show gravity values ranging between 0 and 6 mGal, a relatively narrow range. This range grows toward the North (i.e., between -13 and 14 mGal; Figure 7). The highest gravity values are shown in the northernmost part of section E-E'. This area coincides with outcrops of high-grade metamorphic rocks; biotite zone, garnet zone, andalusite-staurolite-cordierite zone, staurolite out zone, and sillimanite zone (Alfías, 1995; Clariana, 2015; Guitard et al., 1996). Cambro-Ordovician samples of this area show density values ranging between 2,610 and 2,744 kg/m<sup>3</sup>, similar to the mean density value obtained for these rocks in the whole study area (Table 1). The density of metamorphic rocks does not always show high values because they strongly depend on the composition, density of the protolith, the degree of metamorphism, and thermodynamic conditions and processes (Schön, 2011). The residual gravity maximum 5 (Figure 4b) might be linked to higher density rocks located at depth.

## 7. Conclusions

In this work we use gravity, petrophysical and geological data to constrain the geometry of the upper crust of the Central Pyrenees. The study area lies on a long-wavelength negative Bouguer anomaly ranging between -120 and -67.5 mGal together with several relative gravity maxima and minima following a complex pattern. The residual anomaly map shows values ranging between -18 and 18 mGal and highlights two areas separated by a Pyrenean-oriented limit coinciding with the Axial Zone southern termination.

In general, gravity data fit well with the geometry proposed by previous authors at the different parts of the six composite sections done for this work, except for; (a) different thicknesses and/or presence of Triassic evaporites at depth in the South Pyrenean Zone and Axial Zone, (b) different geometry at depth of the Late Variscan granites and Ordovician gneisses, (c) buried Late Variscan granites, with or without evidence of their presence, at the surface, and (d) the existence of a gneiss under the Andorra-Mont Louis granite. Our results highlight strong along-strike variations both (a) in the South Pyrenean Zone mainly due to different along-strike variations in the thicknesses of the Triassic evaporites and Mesozoic rocks and (b) in the Axial Zone linked to lithological and rheological variations in the Paleozoic rock and presence of igneous bodies. The across- and along-strike differences in the superficial and subsurface occurrence of Triassic evaporites have important implications for controlling the structural style in the Pyrenees and in the Axial Zone, the distribution of basement thrust sheets and igneous bodies are relevant to improving the understanding of the orogenic building process.



## Data Availability Statement

Data supporting the conclusions can be obtained from the Digital.CSIC repository (<http://hdl.handle.net/10261/254934> or <https://doi.org/10.20350/digitalCSIC/14026>) and in the Application SIGEOF (<https://info.igme.es/SIGEOF/>).

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