

# **Time-variable gravity field determination from GRACE Follow-On data using the Celestial Mechanics Approach extended by empirical noise modelling**

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GRACE Follow-On Science Team Meeting 2022

18 October 2022

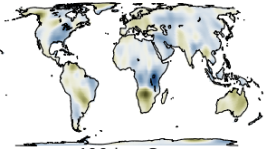
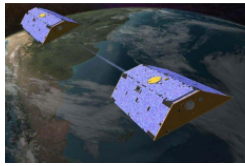
Potsdam, Germany



Astronomical Institute, University of Bern

**AIUB**

# Operational processing



## Basic parametrisation

- initial conditions 2x[6]
- accelerometer bias 2x[3]
- accelerometer scaling 2x[3]

parameters per arc 24

## Additional parameters

- 15 min PCA per satellite in
  - radial 2x[96]
  - along-track 2x[96]
  - cross-track 2x[96]

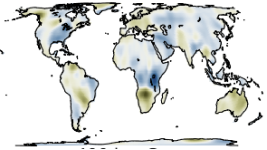
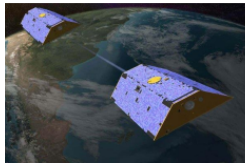
parameters per arc 576

## in daily arcs (30 days):

- 18000 parameters,
- 17280 for the noise model
- + gravity field

→ frequently used in the  
Celestial Mechanics Approach  
[Beutler et al., 2010]

# Operational processing



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## Force models

Gravity field	Internal AIUB static GRACE field
Astromomic bodies	JPL DE421 (all planets + Pluto)
Mean pole	Linear
Solid Earth tides	IERS2010
Solid Earth pole tides	IERS2010
Ocean tides	FES2014b (+ admittances from TUG)
Ocean pole tides	Desai
Atmospheric tides	AOD RL06
Atmospheric & oeanic dealiasing	AOD RL06
Relativistic effects	IERS2010

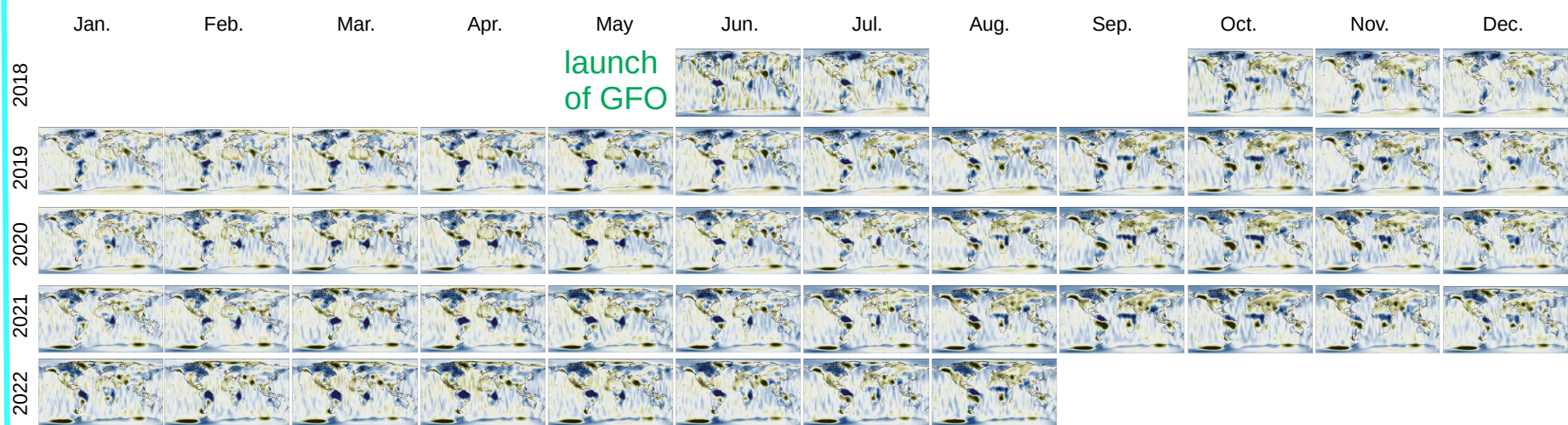
Non-conservative forces:  
**ACT from TUG**

# Processing – operational solution

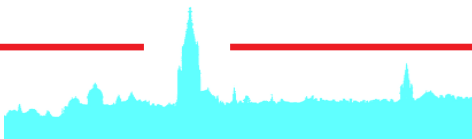
$$\mathbf{N} = (\mathbf{A}^T \mathbf{P} \mathbf{A})$$

$$\mathbf{b} = \mathbf{A}^T \mathbf{P} \mathbf{l}$$

$$\hat{\mathbf{x}} = \mathbf{N}^{-1} \mathbf{b}$$



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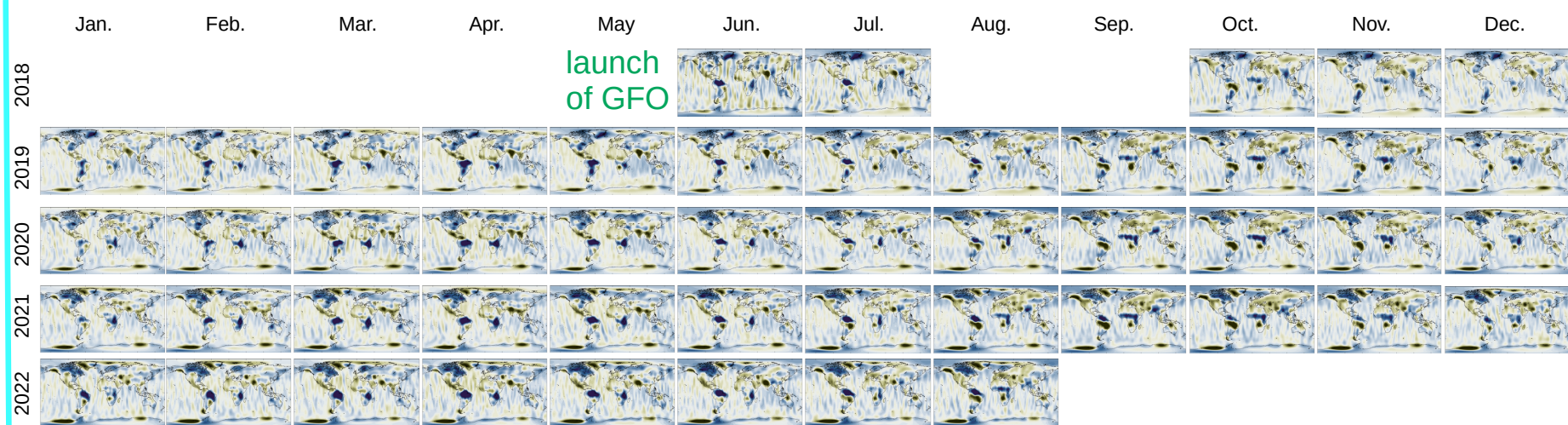


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## Facts

- operational since September 2020
- available at **ICGEM** as [AIUB-GRACE-FO\\_operational](#)
- continuation of AIUB-RL02
- current status: 49 months from June 2018 until August 2022

## Features

- updated background models
- data screening with variance component estimation
- use of alternative GF2 Level 1B ACT product from TUG

## Empirical Noise Modelling



# Empirical noise modelling based on post-fit residuals

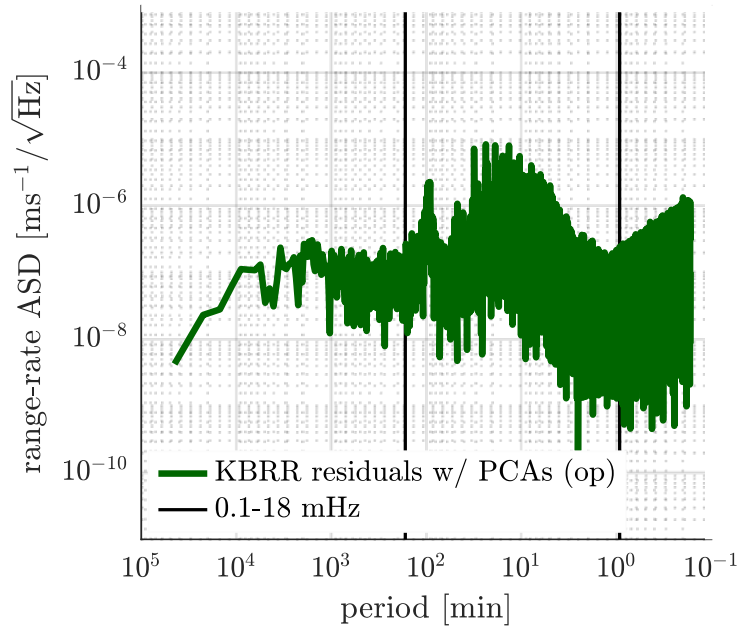
Least-squares

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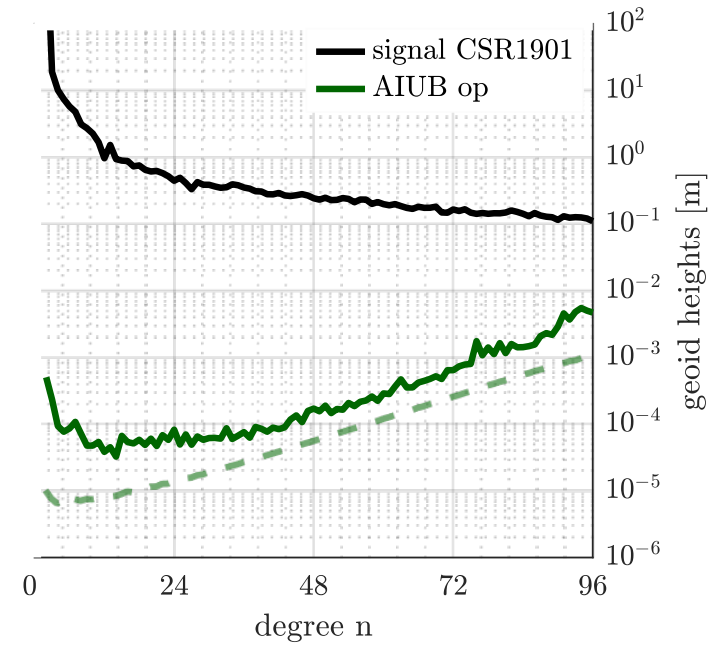
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$$\hat{\mathbf{e}} = \mathbf{l} - \mathbf{A} \hat{\mathbf{x}}$$



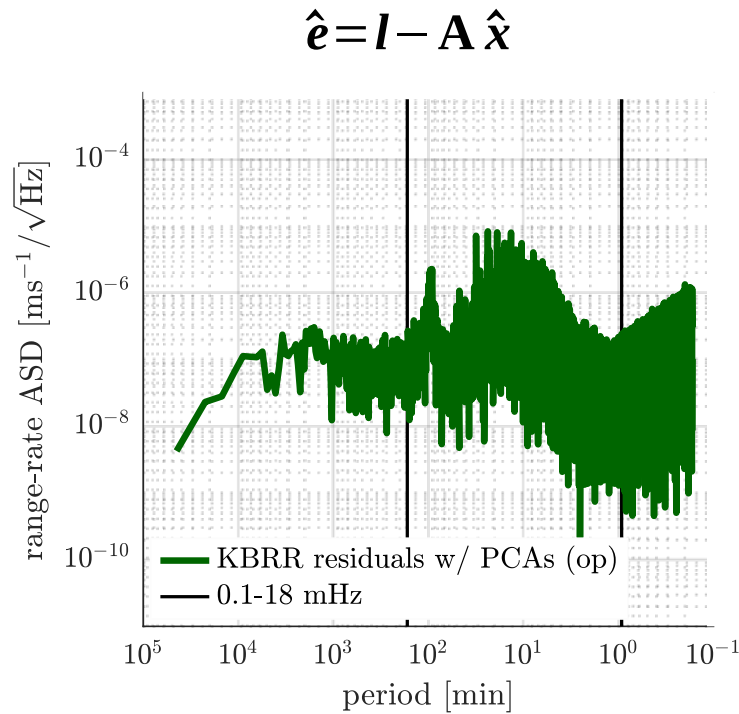
- information in the residuals (noise?)

$$\hat{\mathbf{x}} = \mathbf{N}^{-1} \mathbf{b}$$



- information in the parameters (signal?)

# What do we expect from residuals?

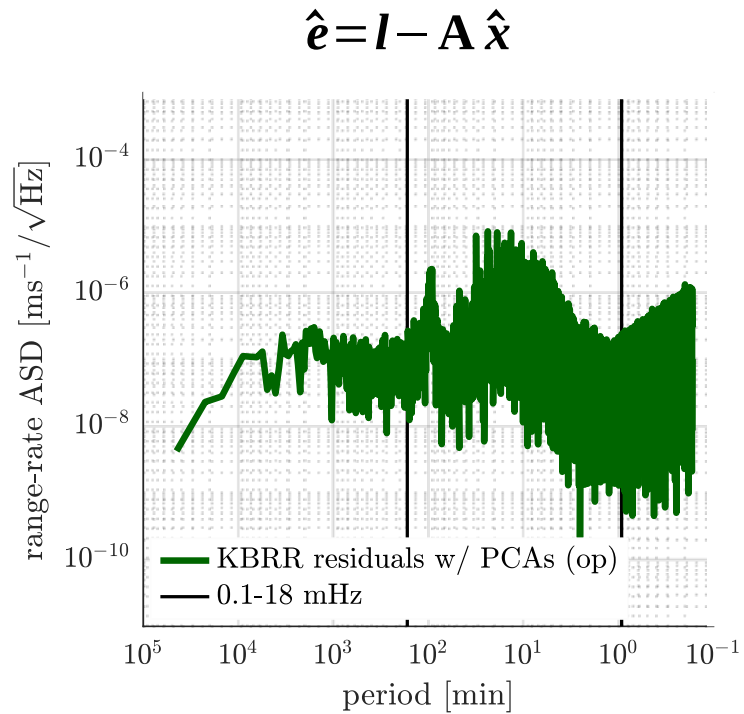


The estimator is BLUE  
(best – linear – unbiased) if

- $E(\hat{\mathbf{x}}) = \mathbf{x}$
- $E(\mathbf{e}|\mathbf{x}) = E(\hat{\mathbf{e}}) = \mathbf{0}$
- $D(\mathbf{e}|\mathbf{x}, \sigma_0^2) = D(\mathbf{l}|\sigma_0^2) = \sigma_0^2 \mathbf{P}^{-1}$   
 $= \sigma_0^2 \mathbf{Q}_{ee} = \sigma_0^2 \mathbf{Q}_{ll}$



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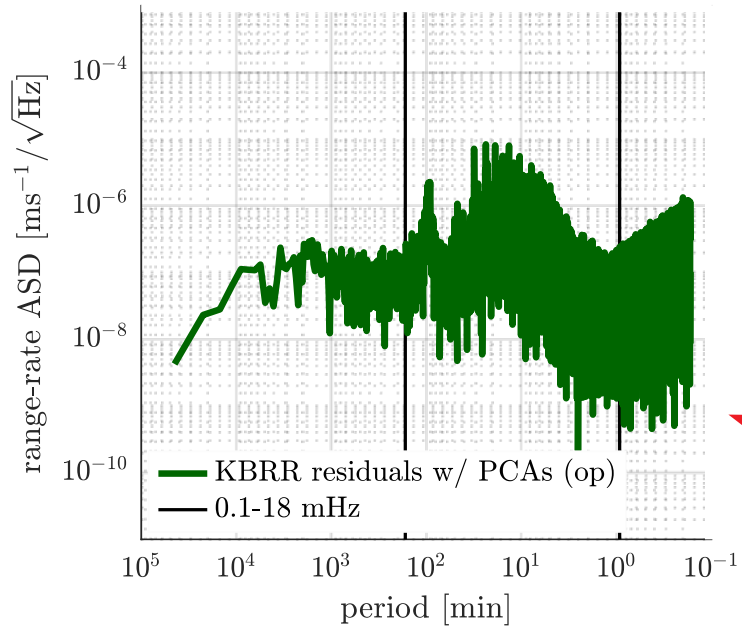
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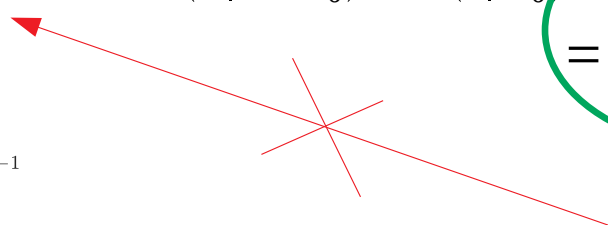
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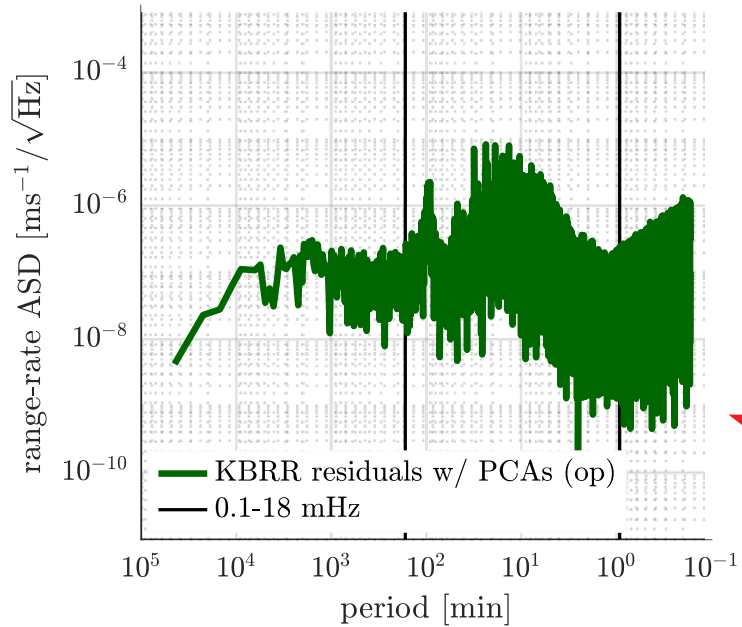
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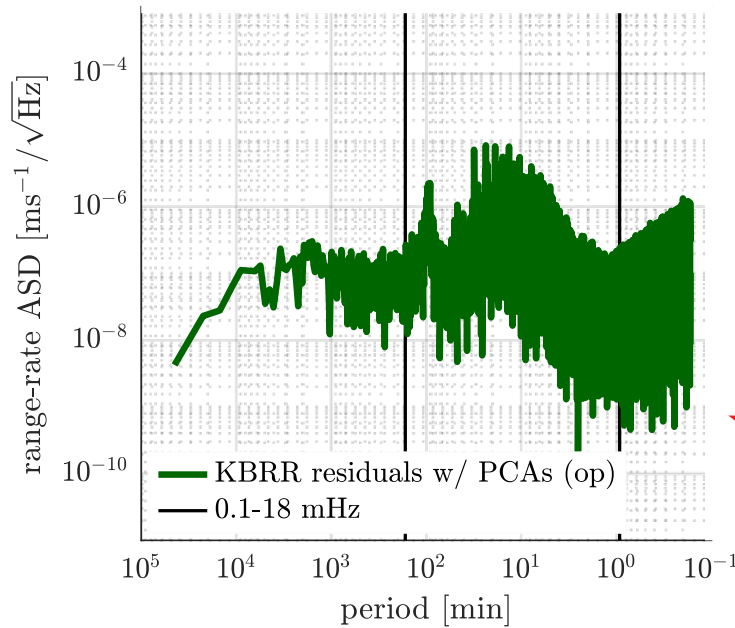
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$\hat{\mathbf{x}} = \mathbf{N}^{-1} \mathbf{b}$	unbiased
$\mathbf{C}_{\hat{\mathbf{x}}\hat{\mathbf{x}}} = \sigma_0^2 (\mathbf{A}^T \mathbf{P} \mathbf{A})^{-1}$	biased

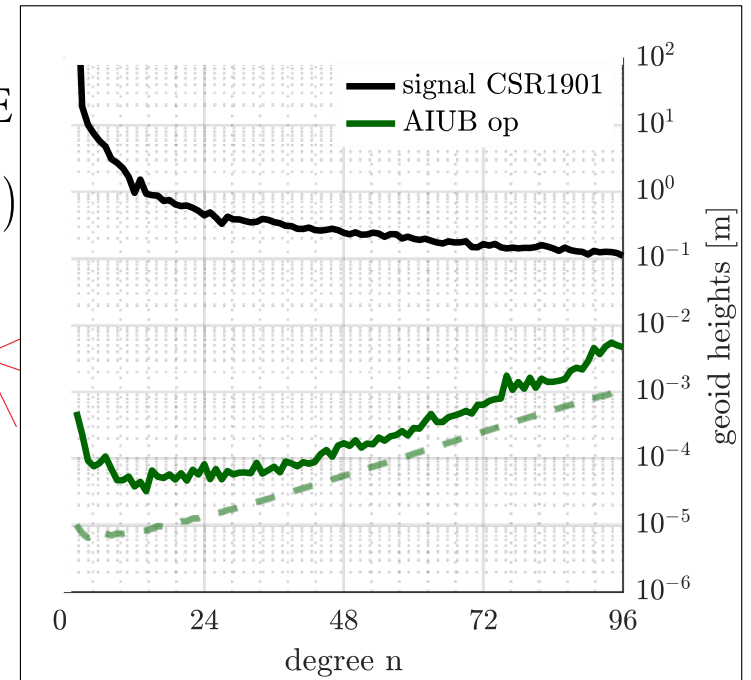
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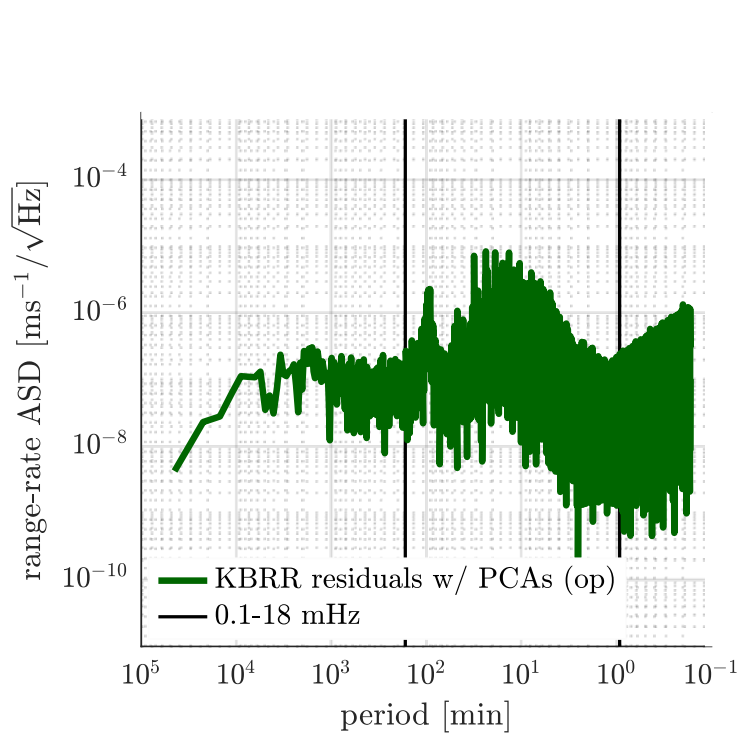
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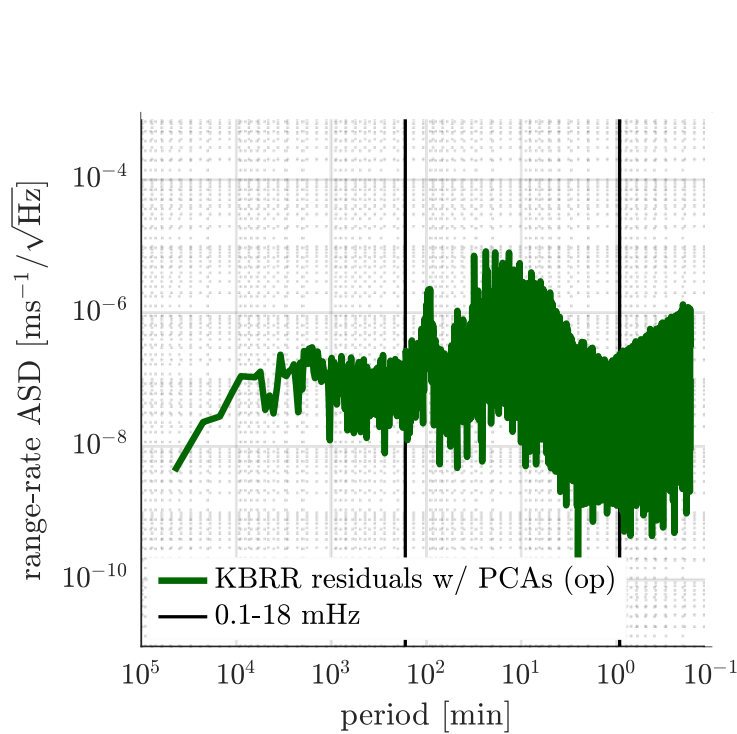
# Auto-covariance function



$$\hat{\mathbf{e}} = \mathbf{l} - \mathbf{A} \hat{\mathbf{x}} \quad (\text{post-fit residuals})$$

$$\text{cov}(\Delta t_k) = \frac{1}{N} \sum_{i=0}^N \hat{\mathbf{e}}(t_i) \hat{\mathbf{e}}(t_i + \Delta t_k)$$

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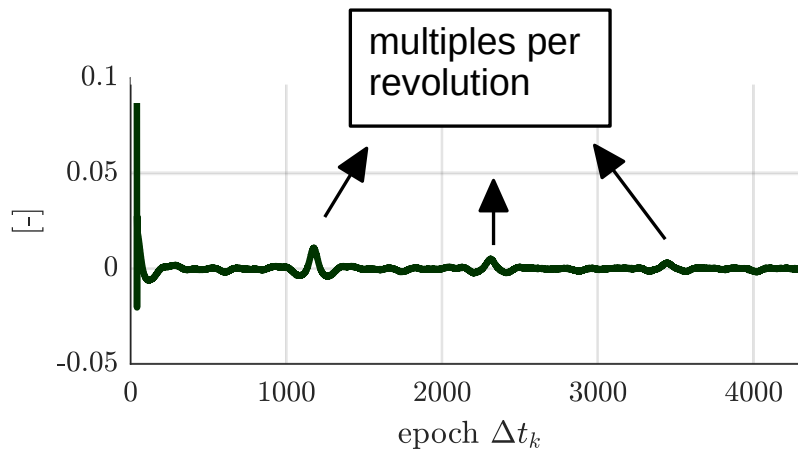


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- stationarity assumed
- biased estimation of auto-covariance  
→ covariance matrix nondegenerate

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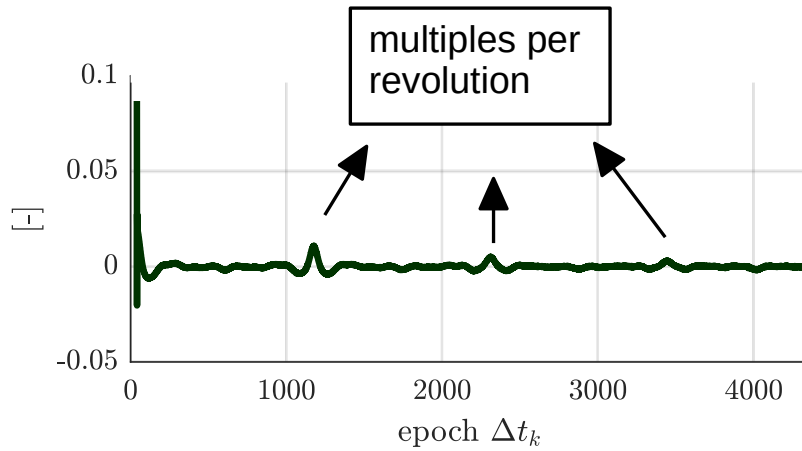


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# Auto-covariance function and weight matrix

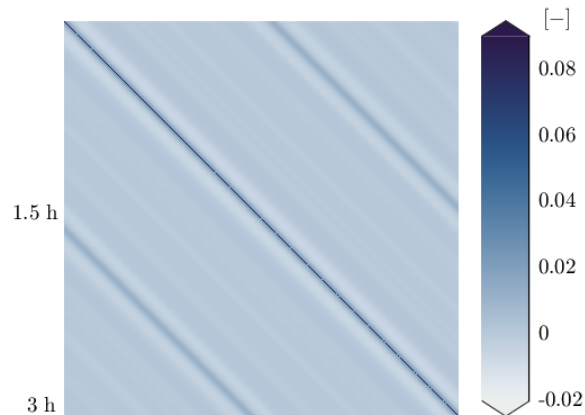


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block  
Toeplitz  
matrix

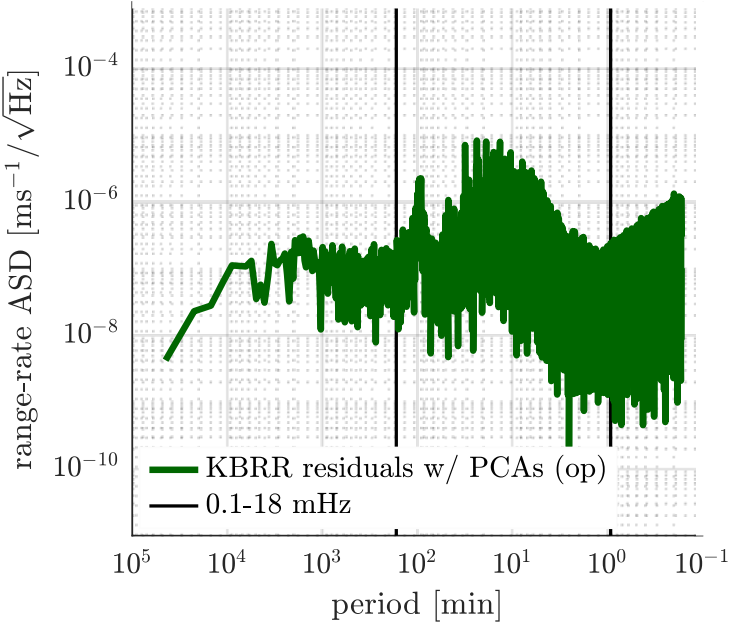


**P**

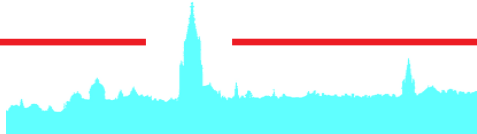


# Auto-covariance function – examples

$$\hat{e} = I - A \hat{x}$$



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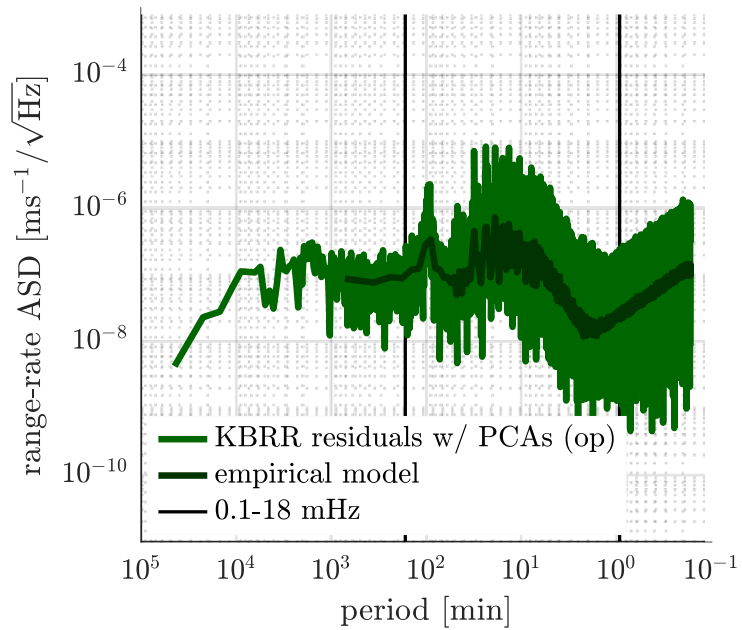


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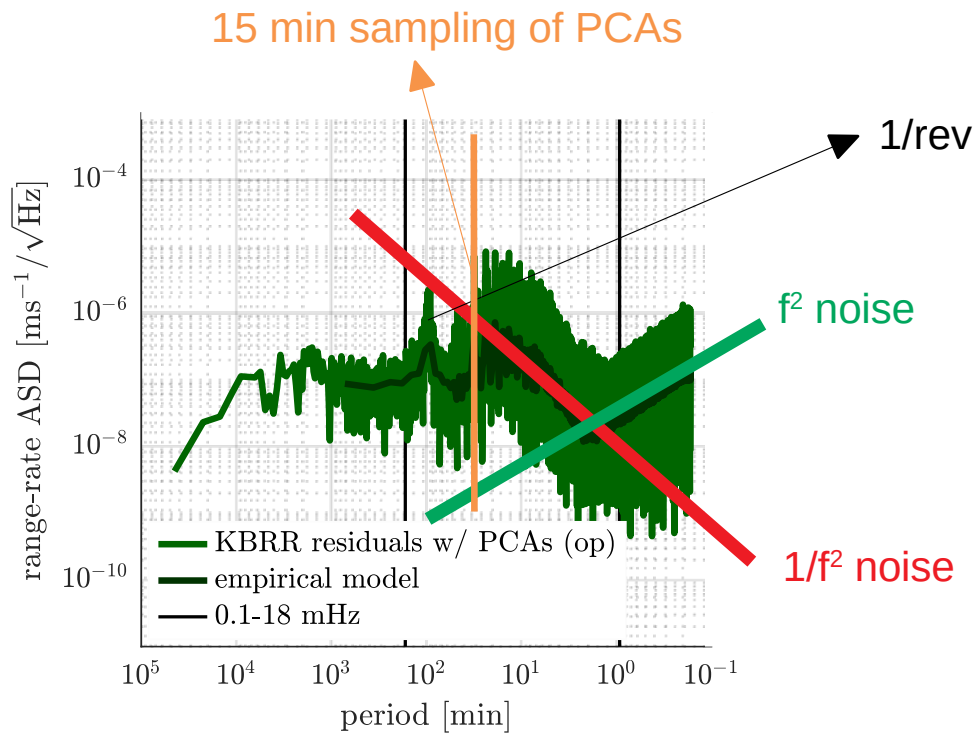


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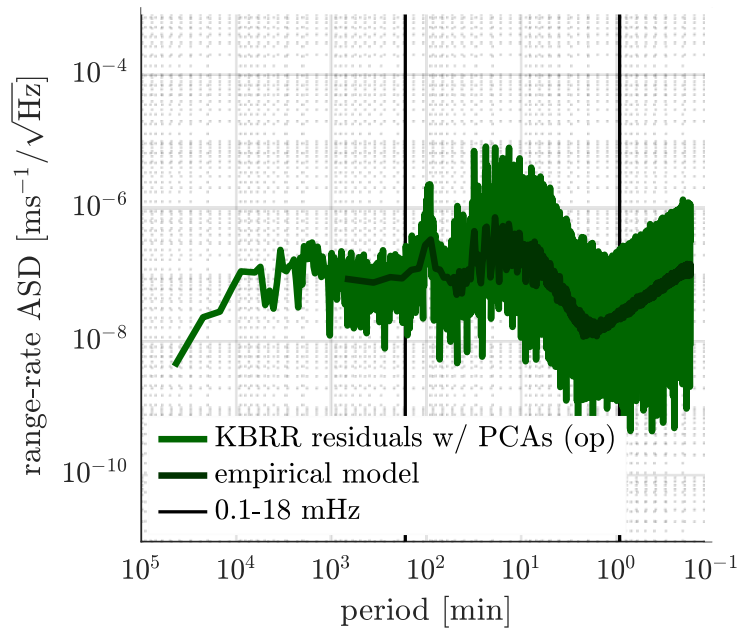


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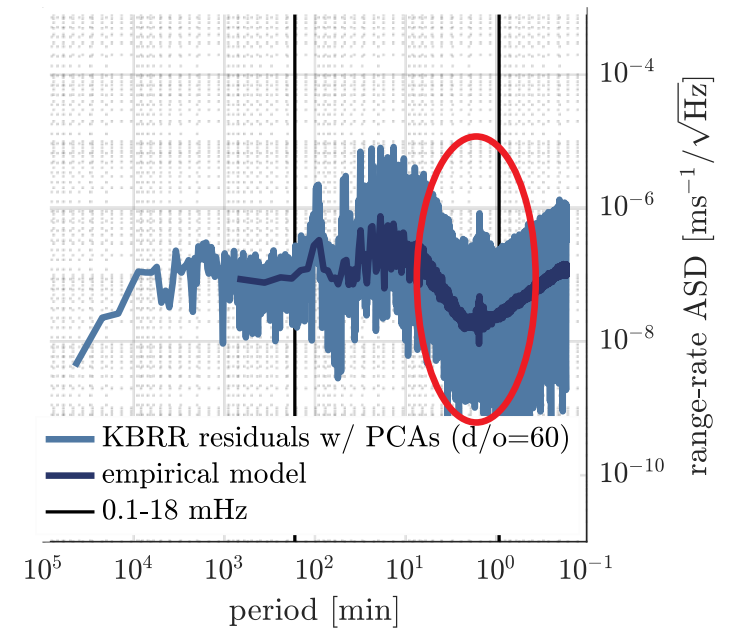
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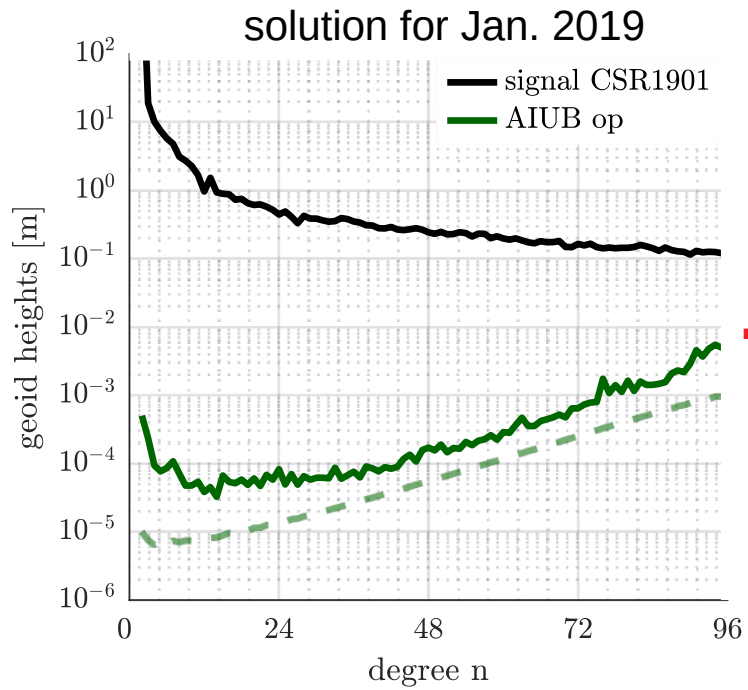
$$\sum_{i=0}^N \hat{e}(t_i) \hat{e}(t_i + \Delta t_k)$$



truncation at d/o=60  
(artefact  $\rightarrow$  bias)  
is reflected

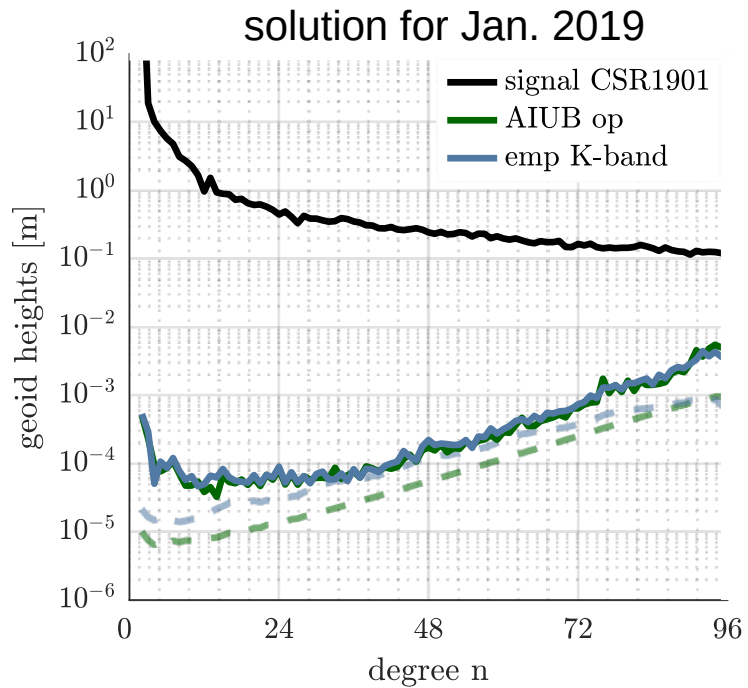


# Results of empirical modelling – gravity field



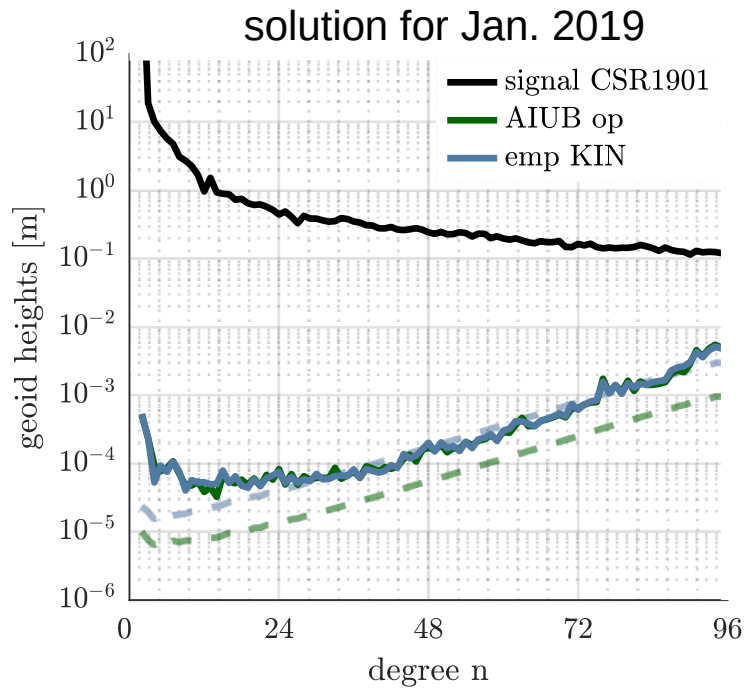
■ formal errors too optimistic compared to the assessed noise

# Results of empirical modelling (K-band) – gravity field



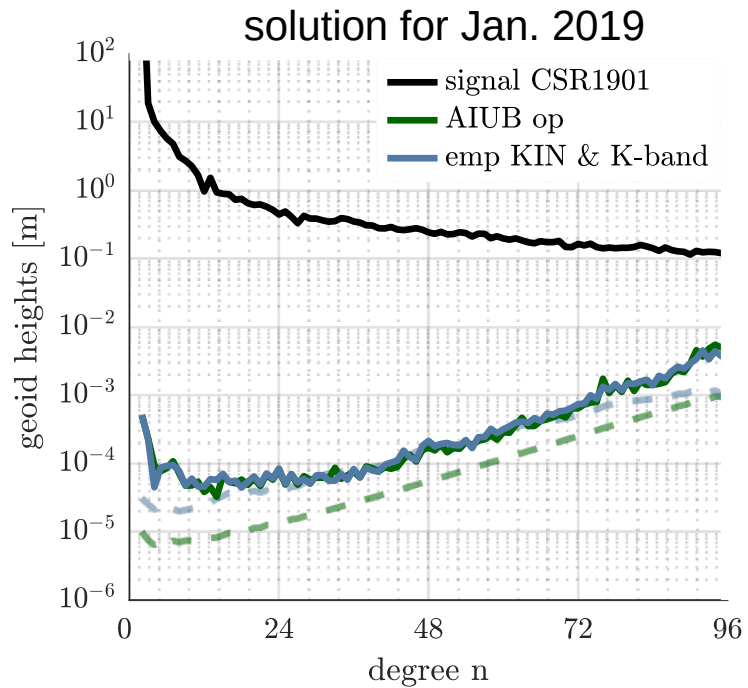
- formal errors reflect assessed noise very well
- including features of resonance orders

# Results of empirical modelling (KIN) – gravity field



- formal errors reflect assessed noise very well
- basically the formal error curve is shifted

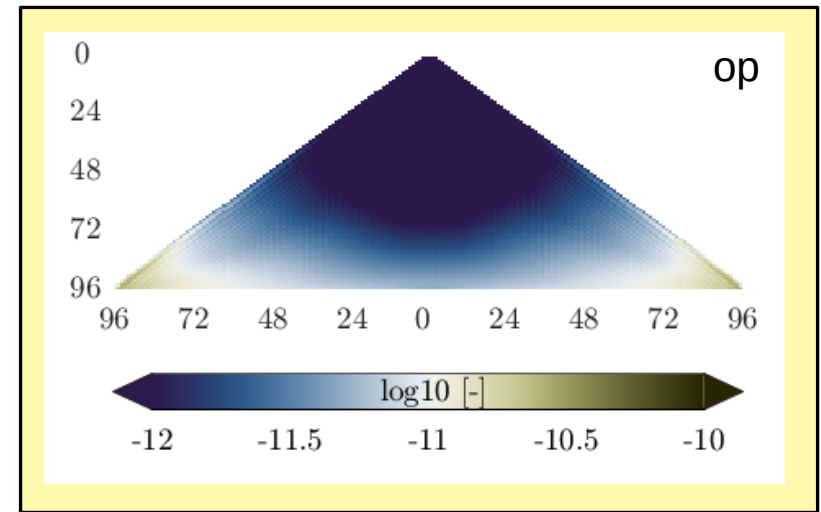
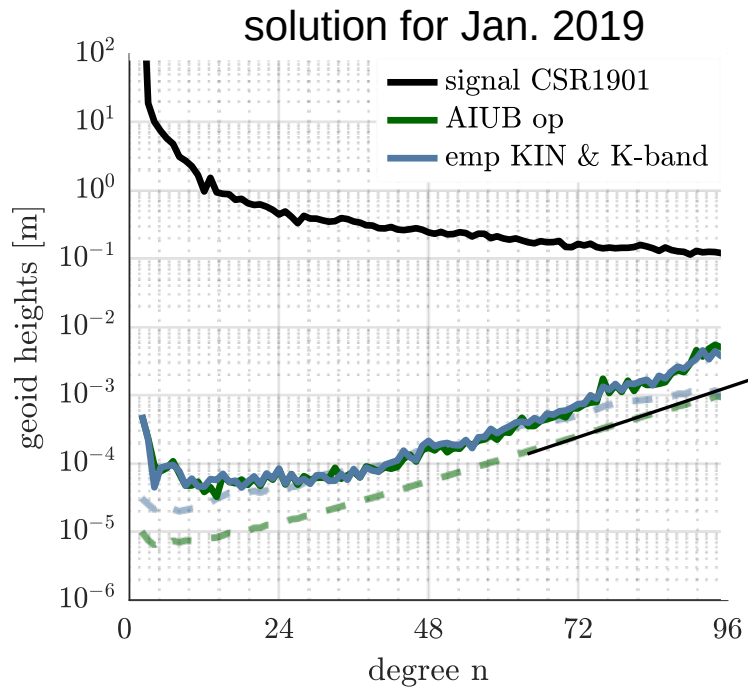
# Results of empirical modelling (KIN & K-band) – gravity field



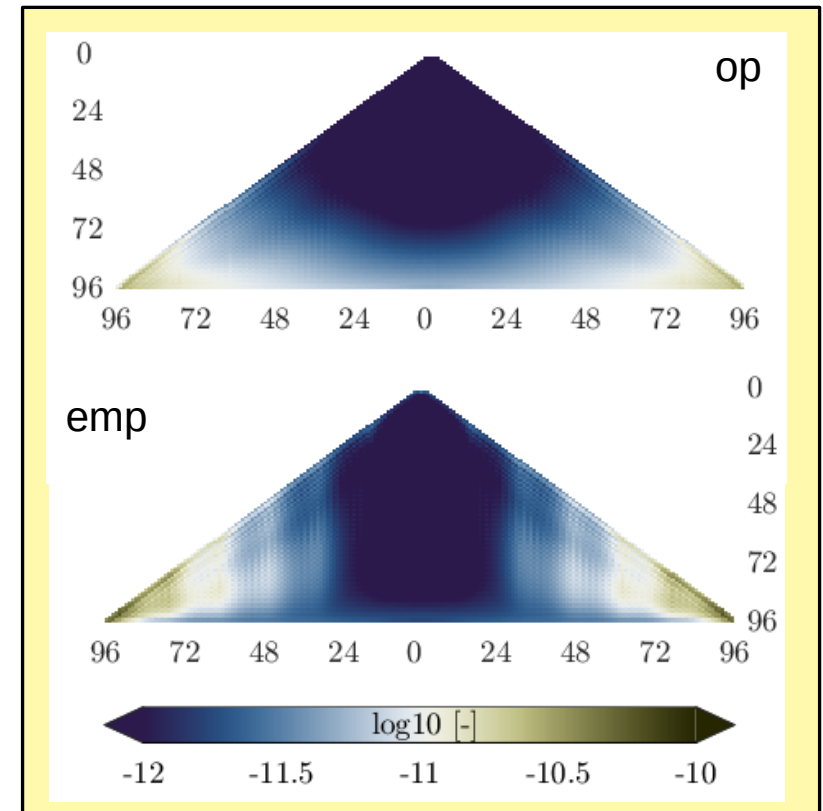
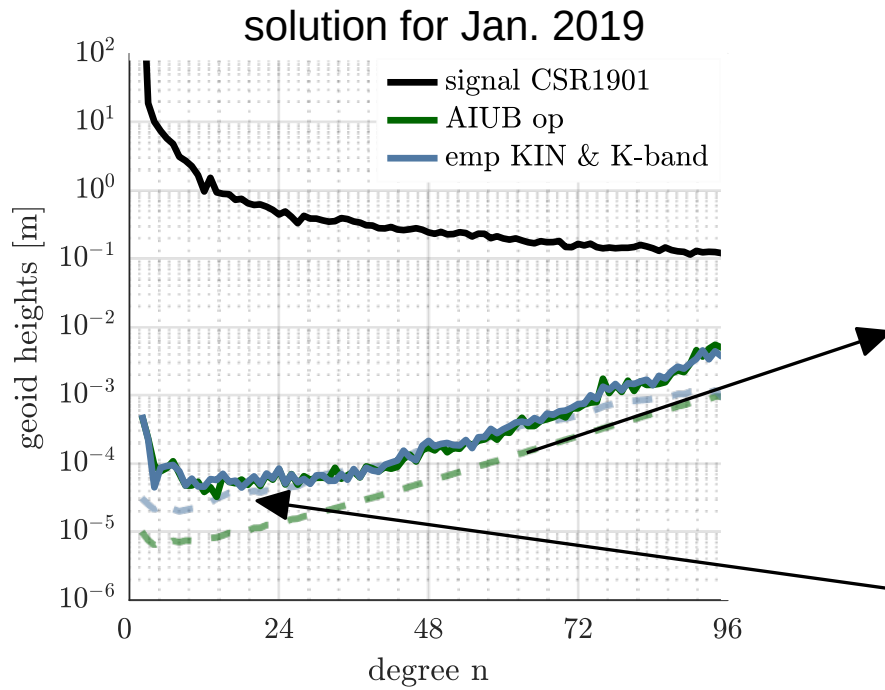
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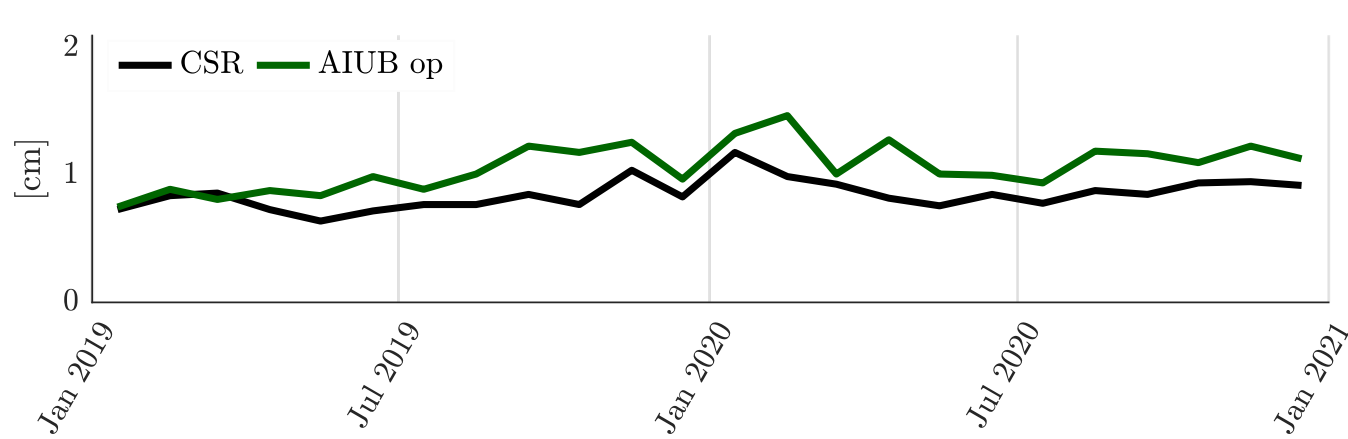
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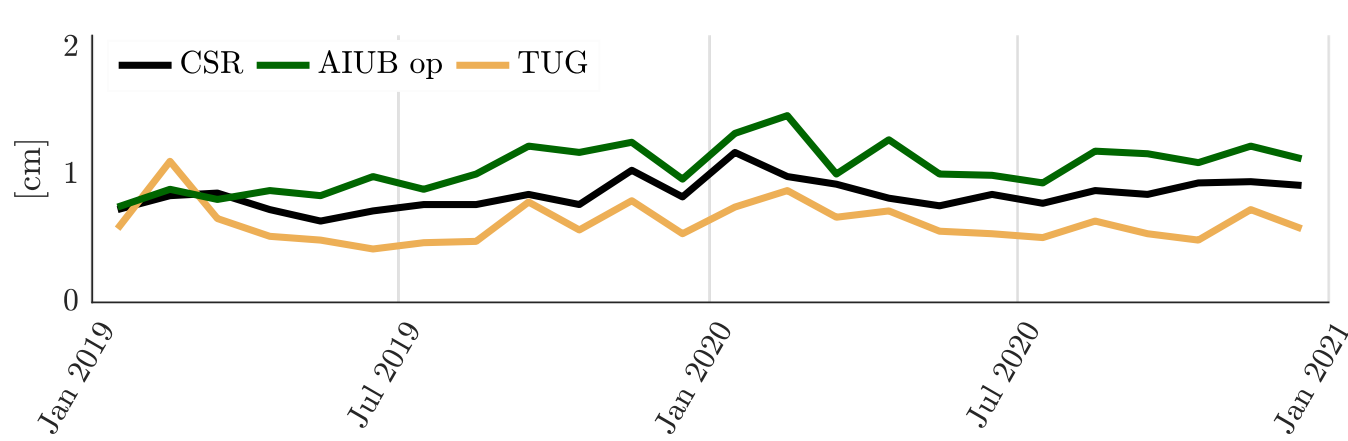
## Results of empirical modelling – RMS over the ocean



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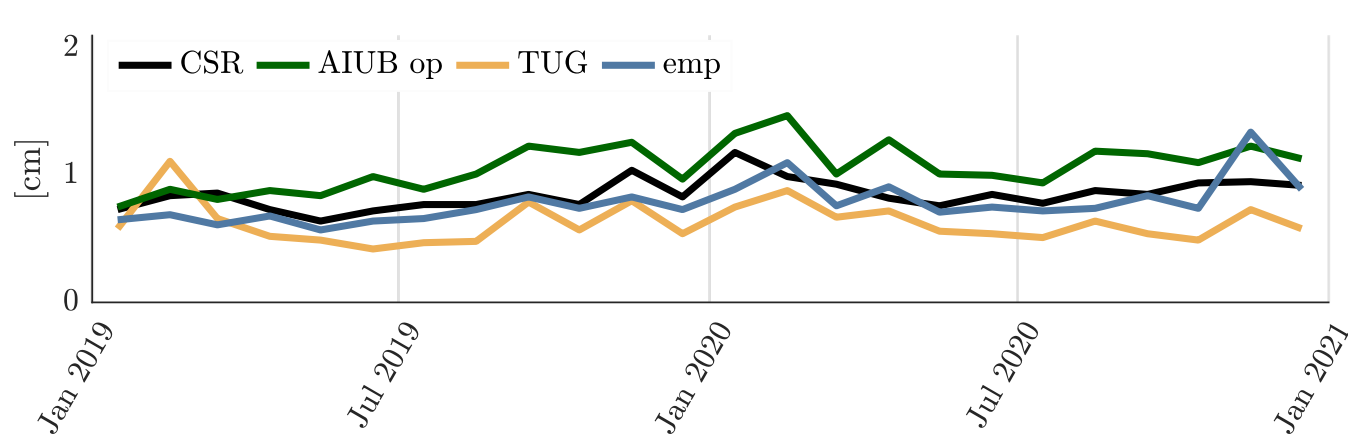
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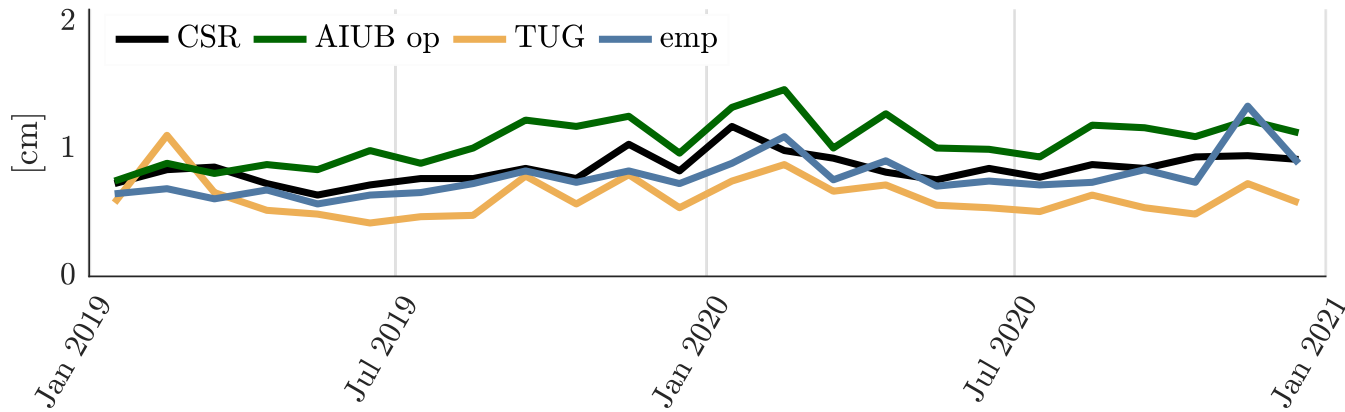
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## Results of empirical modelling – summary

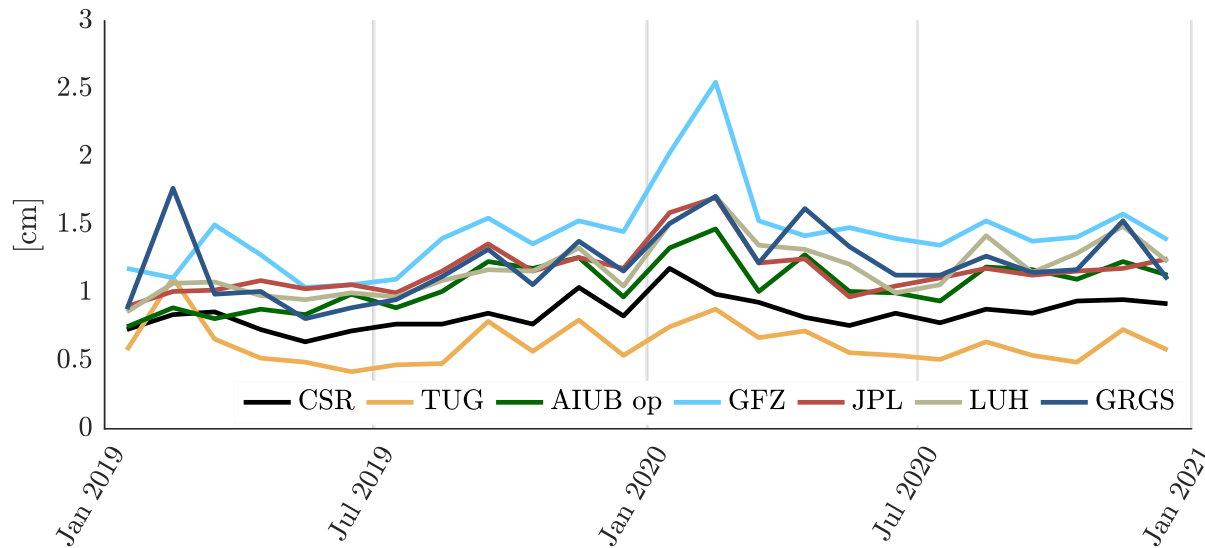


- possible on any (stationary) residuals time series
  - additional parameters can be reduced as stationary behaviour can be absorbed
  - formal errors much more realistic and show resonance orders (if correlation length  $> 3$  h)
  - no constraints needed
  - no/few a priori knowledge needed
- 
- iterations required (might be time consuming)
  - memory consumption and inversion time dependent on length of auto-covariance function

# Performance in COST-G – RMS over the ocean



Combining time-variable gravity field solutions to provide for a product of improved quality, robustness and reliability

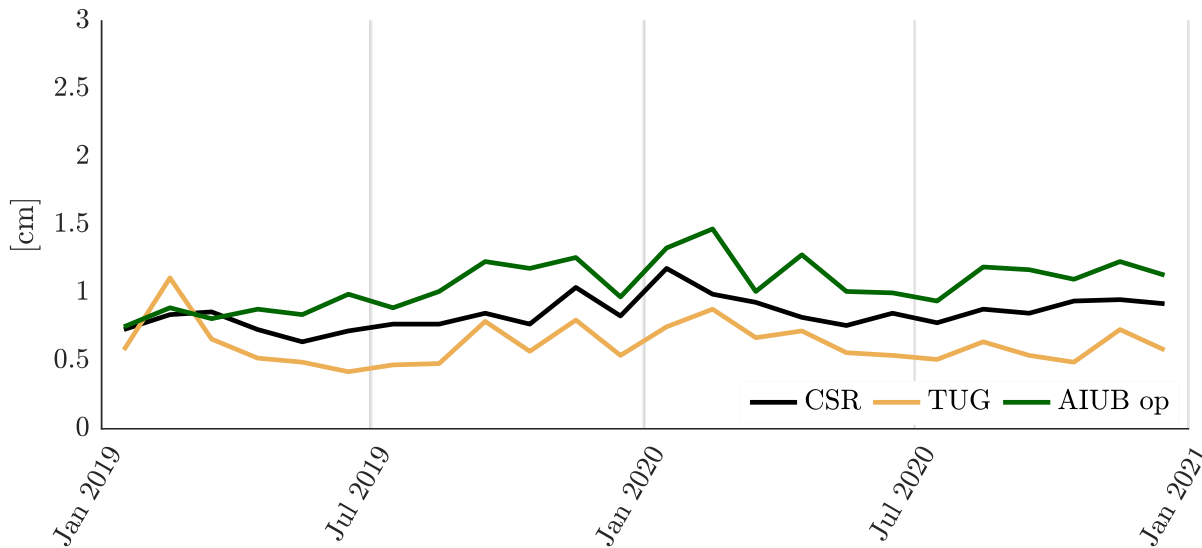


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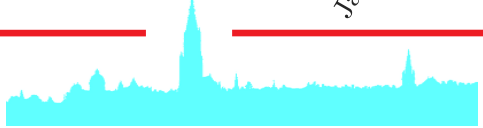
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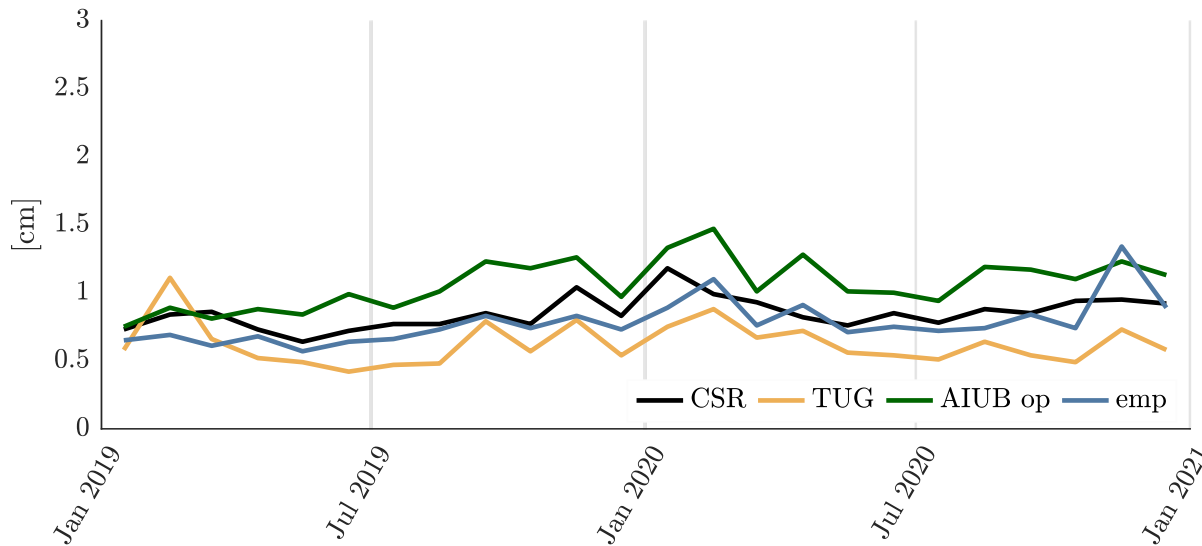




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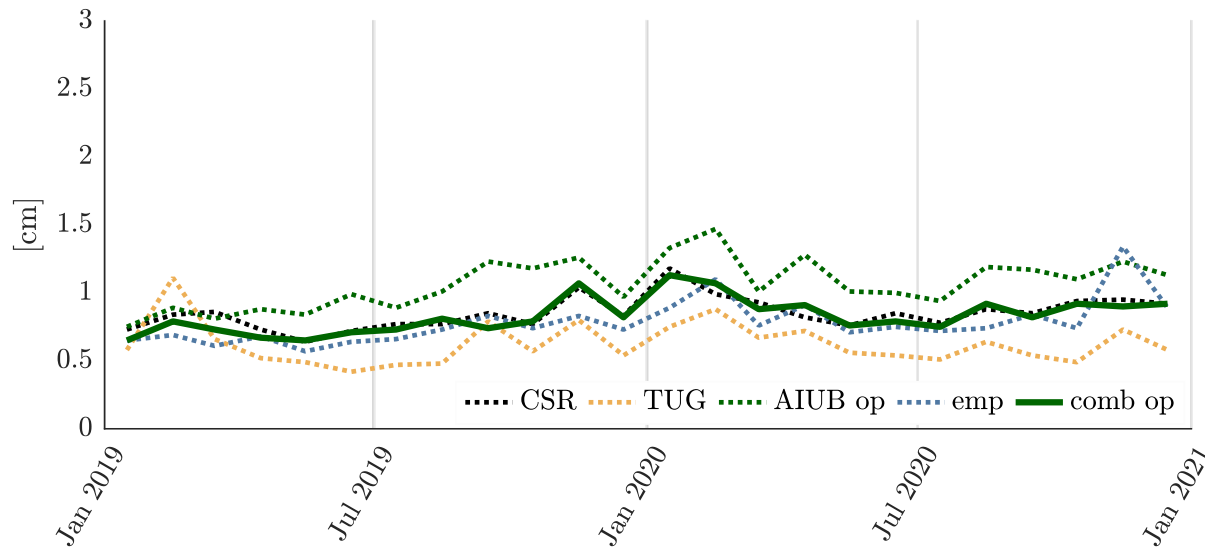


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Combining time-variable gravity field solutions to provide for a product of improved quality, robustness and reliability

- Combination not at the level of the best individual solutions

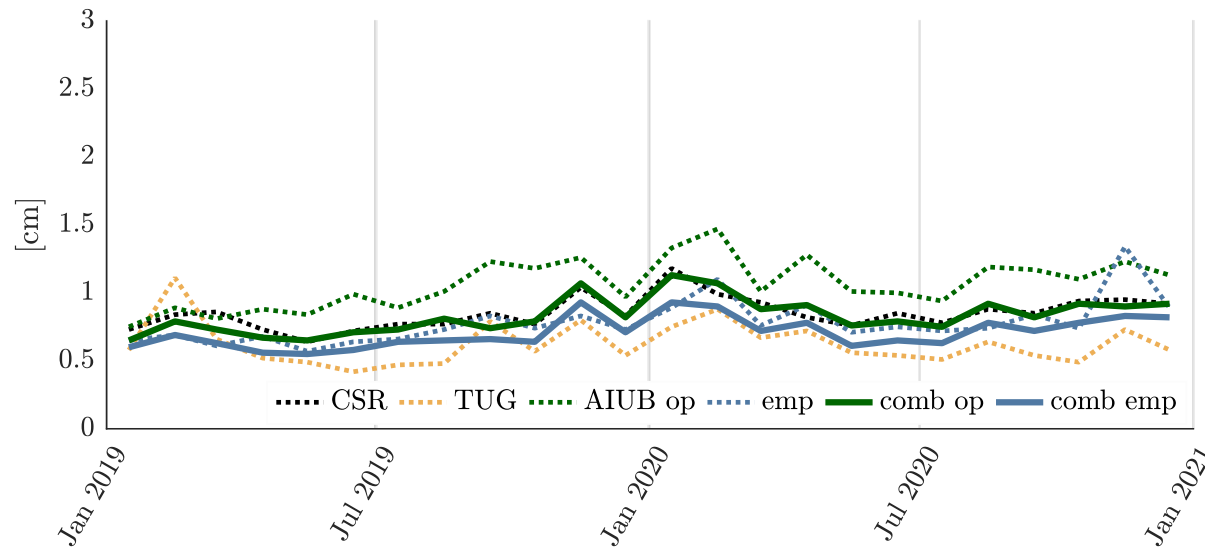


# Performance in COST-G

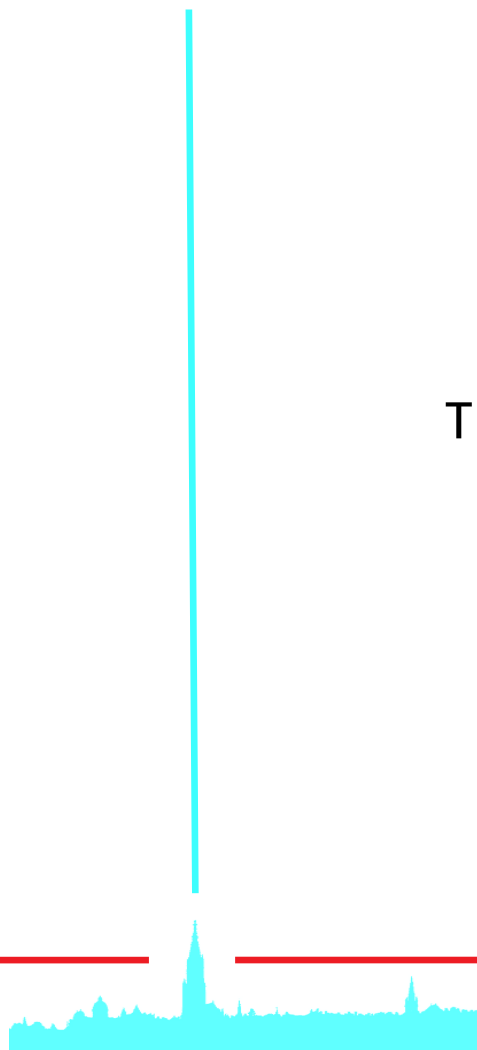


Combining time-variable gravity field solutions to provide for a product of improved quality, robustness and reliability

- (Combination not at the level of the best individual solutions)
- Combination significantly improved by the new processing scheme of one analysis centre



Thank you for your attention



# References

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