

1 **Title**

2 The Palmer ice core as a candidate Global boundary Stratotype Section and Point for the  
3 Anthropocene series.

4

5 **Authorship**

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18

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## 36 **Abstract**

37 The remote Antarctic continent, distant from human industrial activity, should be one of the last  
38 places on Earth to capture Anthropogenic change. Hence, stratigraphic evidence of pollution and  
39 nuclear activity in the Antarctic provides proof of the global nature of the Anthropocene epoch. We  
40 propose an Antarctic Peninsula ice core candidate for the Global boundary Stratotype Section and  
41 Point (GSSP) to the onset of the Anthropocene. The Palmer ice core captures the first evidence of  
42 spheroidal carbonaceous fly ash particles (SCPs), resulting from high temperature combustion  
43 deposited in Antarctic ice. SCPs first appear in 1936 CE, preceding the rise in plutonium ( $^{239+240}\text{Pu}$ )  
44 concentrations from 1945 CE onward. GSSP 1952 CE occurs at a depth of 34.9 m, coincident with the  
45 peak in  $^{239+240}\text{Pu}$  the primary marker for this site.

## 46 **Keywords**

47 Antarctic Peninsula, Spheroidal carbonaceous particles, ice core, methane, plutonium.

## 48 **Introduction**

49 Antarctica's remote location, distant from industrial sources and with no permanent or indigenous  
50 human population, makes it an ideal candidate for identifying the onset of the Anthropocene. Often  
51 referred to as the "last great wilderness", Antarctica is assumed to be a pristine environment free  
52 from the influence of human activity. However, the presence of atmospheric pollutants, soot and  
53 black carbon from fossil fuel and biomass burning (Cordero et al., 2022), heavy metals from mining  
54 (Potocki et al., 2016; Bargagli, 2008) and microplastics (Aves et al., 2022) are a powerful reminder  
55 that the Anthropocene is a global phenomenon and even the most remote region on Earth bears its  
56 mark.

57 The Antarctic Peninsula was chosen as a location to identify the Global boundary Stratotype Section  
58 and Point (GSSP) for the Anthropocene Series due to its potential suitability as a GSSP and sensitivity  
59 to anthropogenic climate change. The preparatory activities of the Anthropocene Working Group  
60 (AWG), including events leading to the submission of GSSP proposals and the binding decision that  
61 the base of the Anthropocene should align with stratigraphic signals dating to the mid-20th century,  
62 are detailed in the introductory article to this special issue (Waters, 2022). In this paper we focus on  
63 the Palmer ice core from the Antarctic Peninsula, as an invited contribution to the AWG to propose a  
64 suitable Antarctic ice core GSSP site.

65 Here we present the Palmer ice core as a candidate for the Anthropocene series. We provide  
66 evidence from two Anthropogenic proxies, spheroidal carbonaceous fly ash particles (SCPs) and  
67 plutonium ( $^{239+240}\text{Pu}$ ). We also include the stable water isotopes ( $\delta^{18}\text{O}$ ), snow accumulation and  
68 atmospheric methane ( $\text{CH}_4$ ) concentrations from this site as supporting material relating to climate  
69 and anthropogenic change. The Palmer ice core was selected from a transect of ice cores drilled  
70 along the spine of the Antarctic Peninsula since the mid-1990s (Thomas. and Tetzner., 2018).  
71 Although arguably many other ice cores could have been proposed to represent the Antarctic GSSP,  
72 the Palmer ice core was selected based on the following criteria:

### 73 ***The exceptional chronology and sample resolution***

74 The Antarctic Peninsula receives the highest amount of snowfall of all Antarctic regions, exceeding  
75 ~4 m per year in some areas (van Wessem et al., 2016). Snowfall is highest around the coast and  
76 areas of orographic uplift, such as the Palmer site. Deposition of anthropogenic proxies is likely  
77 related to snowfall (wet deposition) and thus the concentration of proxies is expected to be highest

78 at coastal and high snow accumulation sites. Evaluation of reanalysis data (ERA-Interim ,1979–2010)  
79 confirms that the ice core site receives an estimated annual average precipitation minus evaporation  
80 of 49 cm (water equivalent). While lower elevation Peninsula ice cores have higher snow  
81 accumulation (Thomas et al., 2017), few extend beyond the past ~150 years (Thomas. and Tetzner.,  
82 2018; Emanuelsson et al., 2022b). The Palmer ice core was selected to provide sufficient layer  
83 thickness to capture changes at sub-annual resolution, while ensuring a record that extends over the  
84 past ~300 years.

### 85 ***The sensitivity to Anthropogenic climate change***

86 In contrast to much of the Antarctic continent, the Antarctic Peninsula is the only region to reveal  
87 strong and statistically significant warming trends since observations began in 1957 (Jones et al.,  
88 2016; Gonzalez and Fortuny, 2018). Paleoclimate observations from ice cores and moss peat banks  
89 suggest that the warming from 1957 onwards is part of a 100-year trend (Thomas et al., 2013; Royles  
90 et al., 2013) that began in the ~1920s. The rise in surface temperatures is accompanied by significant  
91 increases in snow accumulation (Thomas et al., 2017; Thomas et al., 2008; Thomas et al., 2015;  
92 Medley and Thomas, 2019), which has been linked changes in atmospheric circulation, sea ice  
93 conditions and the increased moisture content associated with a warming atmosphere (Medley and  
94 Thomas, 2019). The increased surface temperatures, together with the increased occurrence of  
95 warm and dry fohn winds descending over the Antarctic Peninsula, have been attributed to the  
96 collapse of Antarctic Peninsula ice shelves (e.g. Banwell et al., 2013).

### 97 ***The site has excellent sample preservation, not impacted by melt***

98 Despite the evidence of increased surface temperatures during the 20<sup>th</sup> century (e.g. Jones et al.,  
99 2016), the occurrence of surface melt on the Antarctic Peninsula is limited to the floating ice shelves  
100 and low-elevation coastal locations (van Wessem at al., 2016). While the James Ross Island ice core,  
101 from the northern tip of the Antarctic Peninsula, reveals an acceleration in surface melting during  
102 the 20<sup>th</sup> century (Abram et al., 2013), there are very few visible melt features in the Palmer core.  
103 Therefore, the impact of melt on the climate and anthropogenic proxies at Palmer is assumed to be  
104 negligible. Indeed, a recent study revealed that the stable water isotope record ( $\delta^{18}\text{O}$ ) from Palmer  
105 has the potential to reconstruct melt over the Larsen ice shelves (Emanuelsson et al., 2022a), even  
106 though the  $\delta^{18}\text{O}$  itself is not directly impacted by melting. Thus, the strategic location of the Palmer  
107 core makes it sensitive to climate variability, without being adversely impacted by it.

### 108 ***The link to long-range transport for anthropogenic proxies***

109 Back-trajectory analysis confirms that the Amundsen- Bellingshausen Sea is the dominant source  
110 region for air-masses reaching the Palmer ice core site (Thomas and Bracegirdle, 2015). The  
111 geographical setting of the Antarctic Peninsula, which forms a barrier for the strong southern  
112 hemisphere westerly winds, deflects onshore (northerly) air-masses to the Palmer site. An observed  
113 seasonal migration in the trajectories (1979-2010) is related to changes in the Amundsen Sea Low, a  
114 climatological low-pressure system that is driven by large-scale modes of climate variability (Thomas  
115 and Bracegirdle, 2015). These include the Southern Annular Mode (SAM), largest mode of variability  
116 in the Southern Hemisphere, and El Nino Southern Oscillation (ENSO). The shift to the positive phase  
117 of the SAM, in response to stratospheric ozone depletion (Lubin et al., 2008), constitutes one of the  
118 largest trends in the past 200 years (Jones et al., 2016). Antarctic Peninsula ice cores capture  
119 changes in the SAM, ENSO and tropical teleconnections in the western Pacific (Thomas et al., 2008;  
120 2013; 2015; 2017; Medley and Thomas, 2019). The Palmer core specifically displays a strong  
121 relationship with the Inter-decadal Pacific oscillation (IPO) (Emanuelsson et al., 2022a). Thus, Palmer

122 provides an optimal location to capture the deposition of anthropogenic proxies, due to the  
123 continent's remote location, involves long-range atmospheric transport and a clear pathway of  
124 northerly air-mass incursions.

### 125 ***The availability of uncontaminated frozen sample with sufficient volume for new analysis***

126 An advantage of the Palmer ice core, over many deep ice cores, is that it was drilled without drilling  
127 fluid. The high-density drilling fluid, used routinely to keep the ice core borehole open, is one of the  
128 main sources of chemical and biological contamination in ice cores (Alekhina et al., 2018). While  
129 most Antarctic ice cores which span the mid-20<sup>th</sup> century were drilled in the 1990s (e.g., Thomas et  
130 al., 2017), the Palmer core was drilled in 2012 and has therefore not experienced any potential  
131 degradation or damage associated with long-term storage. A key objective of the AWG was to  
132 undertake new analysis of anthropogenic proxies, requiring substantial ice volumes. Unlike more  
133 established and heavily sampled Antarctic ice cores, the Palmer ice core has sufficient sample  
134 volume for new analysis whilst meeting the GSSP requirements to maintain an archive.

135

## 136 **Materials and methods**

### 137 ***Geographic setting of core sites***

138 The Palmer ice core was drilled in the Palmer Land region of the Antarctic Peninsula [73.86° S, 65.46°  
139 W] (Figure 1). The Antarctic Peninsula ice sheet sits within a 2000 m high mountain chain that  
140 extends ~1,300-km northward from continental Antarctica toward the southern tip of South  
141 America. The ice core was drilled at 1897 m above sea level, close to the ice divide, where horizontal  
142 ice flow is expected to be minimal. The estimated ice thickness at the drill site is ~1000 m (Fretwell  
143 et al., 2013). With average 2 m temperatures of -28.5°C, and summer averages of -16.2°C, (ERA5  
144 reanalysis data) this cold site is unlikely to be affected by surface melting. As one of the highest  
145 points on the ice sheet, the site is distant from any orographic disturbances, such as mountain peaks  
146 or areas of exposed rock. The closest rock outcrop is a small nunatak approximately 35 km from the  
147 drill site. The closest areas of human activity are the summer only stations of Sky Blu (1435 m a.s.l.),  
148 ~210 km to the southwest, which began operations in 1995 CE, and Fossil Bluff (92 m a.s.l.), ~300 km  
149 to the northwest which has been used intermittently since 1961 CE. The closest year-round research  
150 station, Rothera (4 m a.s.l.), is over 700 km to the north, which began operation in 1975 CE. Thus,  
151 the influence of local human activity or light industry is expected to be minimal at this remote  
152 location, especially prior to the 1960s.



153

154 **Figure 1.** Map of the Antarctic Peninsula showing the Palmer ice core location (red star), and the  
 155 location of research stations. Insert map of the Antarctic continent, highlighting the Antarctica  
 156 Peninsula (black rectangle). Copyright Laura Gerrish, BAS.

157 ***Field collection of core and sampling***

158 The ice core was drilled over six days between the 24<sup>th</sup> – 29<sup>th</sup> December 2012. It was drilled using the  
 159 British Antarctic Survey (BAS) electromechanical dry drill, which operates in a dry borehole without  
 160 drill fluid. The final drilling depth was 133.45m, drilled in approximate 1 m sections. The cores were  
 161 cut to 85 cm length, using a cross-saw, and wrapped in pre-labelled layflat tubing. The cores were  
 162 weighed and measured in the field before being loaded into insulated boxes, with polystyrene  
 163 inserts to prevent core damage during transport. The boxes were stored in a trench in the field,  
 164 where the atmospheric temperature remained below -20°C. The ice was transported via Twin Otter  
 165 aircraft to Rothera research station, loaded into a -25°C refrigerated container and shipped to BAS  
 166 for analysis and long-term storage. The ice samples, and remaining archive, are kept at -25°C in a  
 167 monitored, temperature-controlled freezer facility in the dedicated BAS ice core laboratories in  
 168 Cambridge, UK.

169 ***Chronological controls***

170 A suite of chemical and isotopic species analysed using continuous flow analysis (CFA), using  
 171 longitudinal ice sections (32 x 32 mm). Ice samples are loaded onto a heated melt head (in a -25°C  
 172 freezer) and the sample stream of water and gas, from trapped air bubbles, are separated in a

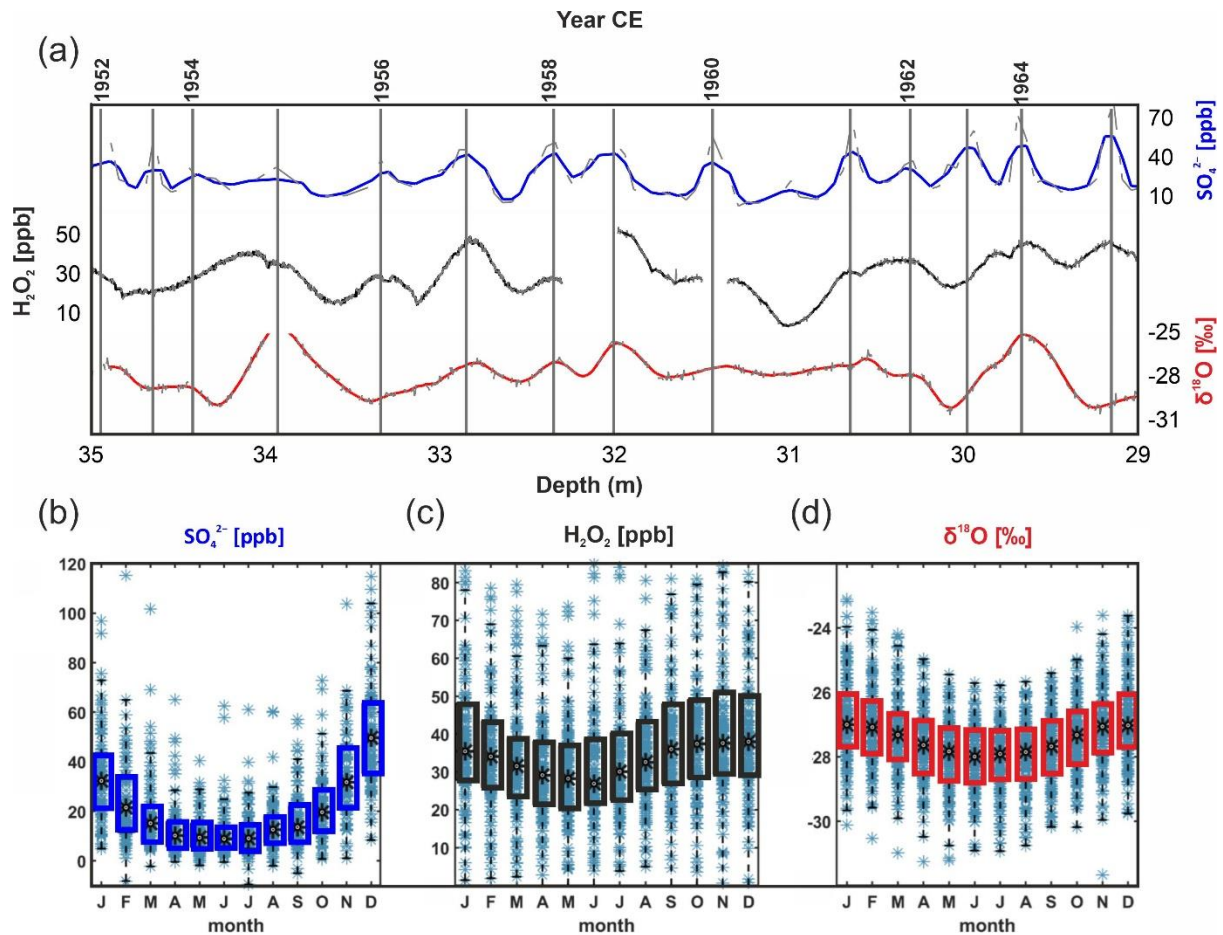
173 debubbler before being pumped into a series of analytical instruments (Grieman et al., 2022). The  
174 bespoke CFA system comprises Fast Ion Chromatography (FIC), Dionex ICS-3000, to measure major  
175 anions sulphate [ $\text{SO}_4^{2-}$ ], nitrate [ $\text{NO}_3^-$ ], chloride [ $\text{Cl}^-$ ], and methane sulphonic acid [MSA-]. The FIC  
176 data acquisition is one 2.63 cm sample every 4 cm, thus providing an average depth resolution of 4  
177 cm (Grieman et al., 2022). Hydrogen peroxide [ $\text{H}_2\text{O}_2$ ] was measured using a FIALab photomultiplier  
178 tube-fluorescence detector (PMT-FL), connected to the CFA. The estimated resolution for the  
179 fluorescence detection is  $\sim 1.4 \pm 0.5$  cm (Grieman et al., 2022). Water isotopes ( $\delta^{18}\text{O}$ ) were  
180 measured using a Picarro analyzer (L2130-i) operating in a custom-made continuous-flow setup.  
181 Internal standards, calibrated to the V-SMOW2/Standard Light Antarctic Precipitation 2 (V-  
182 SMOW2/SLAP2) scale, were used to calibrate the isotope records. The average depth resolution for  
183  $\delta^{18}\text{O}$  is  $\sim 4$  cm.

184 The number of data points per year is dependent on the ice depth and annual snow accumulation  
185 rate. During the 30-year interval surrounding the proposed GSSP (1940-1970), the average snow  
186 accumulation was  $43 \pm 27$  cm snow per year. Providing an average of  $10 (\pm 6)$  data points per year for  
187 major ions and  $\delta^{18}\text{O}$ , and  $30 (\pm 10)$  data points per year for  $\text{H}_2\text{O}_2$ .

188 The ice core was dated using annual layer counting, based on seasonal deposition of chemical and  
189 isotopic species (Fig. 2). Clear seasonal cycles are observed in sulphate [ $\text{SO}_4^{2-}$ ], which peaks during  
190 the austral summer (November – January) when marine productivity is highest, and hydrogen  
191 peroxide ( $\text{H}_2\text{O}_2$ ), a photochemical species which peaks during the summer solstice.  $\delta^{18}\text{O}$  also displays  
192 a seasonal cycle, although with a weaker amplitude than [ $\text{SO}_4^{2-}$ ] and  $\text{H}_2\text{O}_2$  (Fig. 2d) (Emanuelsson et  
193 al., 2022b). The annual layers were counted manually, assigning years where a clear peak was  
194 observed in at least two of the chronological markers  $\text{SO}_4^{2-}$  and  $\text{H}_2\text{O}_2$ , or  $\delta^{18}\text{O}$ . Years with only one  
195 clear marker, or during unavoidable data gaps (e.g., core breaks), were marked as uncertain.

196 The age-scale is independently verified by identifying volcanic [ $\text{SO}_4^{2-}$ ] peaks, classified as [ $\text{SO}_4^{2-}$ ]  
197 values which exceed two standard deviations ( $2\sigma$ ) above the mean (Fig. 3a). To account for the  
198 natural variability in background [ $\text{SO}_4^{2-}$ ] we apply a running 200 data-point (approx. 10-years) mean  
199 and standard deviation (Fig. 3b).

200



201  
 202 **Figure 2.** Annual layer counting. (a) Seasonal cycles in sulphate [ $SO_4^{2-}$ ] (blue), hydrogen peroxide  
 203 [ $H_2O_2$ ] (black), and stable water isotopes [ $\delta^{18}O$ ] (red) from 1952-1965 CE. High resolution data for all  
 204 species shown in grey dashed curves, solid lines represent a 3-point running mean (7-point running  
 205 mean for [ $H_2O_2$ ]). Seasonality plots modified from Emanuelsson et al., (2022b) for (b) [ $SO_4^{2-}$ ] (blue),  
 206 (c) [ $H_2O_2$ ] (black), and (d) [ $\delta^{18}O$ ] (red). Monthly averaged values for each month, the coloured boxed  
 207 indicating the 25<sup>th</sup> and 75<sup>th</sup> percentiles.

208 **Anthropocene proxies**

209 **Radioisotopes**

210 Discrete samples were cut to annual resolution, spanning the period 1930 to 1998 CE. The surface  
 211 area of each annual sample was  $11.4 \text{ cm}^2$ , and the sample volumes ranged between 64 mL - 430 mL.  
 212 This reflects the varying snow accumulation during this period (1930-1998 CE), which can range  
 213 between 0.11 and 0.57 m of water equivalent per year. Subsequently, the annual samples were  
 214 combined to two-yearly resolution due to the extremely low sample concentrations detected. The  
 215 average sample volume (1930-1998 CE) was  $459 \pm 366 \text{ ml}$ .

216 All samples were analysed in the GAU-Radioanalytical Laboratories at the University of  
 217 Southampton. Samples were spiked with  $Pu-^{242}$  recovery tracer, transferred into PTFE beakers, and  
 218 evaporated to dryness. Samples underwent a series of acid treatments, with the residues  
 219 evaporated to dryness after each treatment. First residues were dissolved using concentrated  
 220 hydrofluoric acid (HF) solution (2x10 ml), followed by nitric acid ( $HNO_3$ ) (3x20ml), and finally boiled  
 221 undercover in 40ml concentrated hydrochloric acid (HCl) for 2 hours. After this 2g of boric acid was  
 222 added and the samples were boiled again until complete boric acid dissolution (2hrs).

223 Solutions were evaporated to dryness, re-dissolved in 30ml 1M HNO<sub>3</sub> and boiled again under cover  
224 until a completely clear solution was obtained. 10mg of iron (Fe) carrier was added and <sup>239+240</sup>Pu was  
225 co-precipitated with Fe(OH)<sub>3</sub> using concentrated ammonia solution. Obtained precipitates were  
226 dissolved in HCl and acid molarity was adjusted to 9M. Samples were loaded onto anion exchange  
227 columns (1x5cm), previously conditioned with 9M HCl. After the load solution passed the columns,  
228 columns were washed with further 30ml of 9M HCl followed by 50ml of 8M HNO<sub>3</sub> and again 10 ml of  
229 9M HCl.

230 Finally, <sup>239+240</sup>Pu was stripped from the columns using 9MHCl/NH<sub>4</sub>I (ammonium iodide) solution into  
231 clean beakers. Solutions were evaporated to dryness with 5ml conc HNO<sub>3</sub> added to remove excess  
232 iodide. Thin alpha-spectrometric sources were prepared using electrodeposition from diluted  
233 HCl/oxalate solution. Prepared sources were counted for <sup>238</sup>Pu and <sup>239+240</sup>Pu using Alpha Octete  
234 spectrometers with PIPS detectors and spectra analysed using Maestro 32 software. Detection limits  
235 were ca. 0.5 uBq/g.

### 236 ***Spheroidal carbonaceous fly-ash particles (SCPs)***

237 SCPs are only produced by the high temperature, industrial combustion of coal-series and oil fuels  
238 (Rose 2015). They are morphologically distinct under the light and scanning electron microscope  
239 making them unambiguous indicators of deposition from these sources (Rose 2008). Two methods  
240 were applied to identify the presence of SCPs in the ice. The first approach used discrete samples cut  
241 from the archive section of the Palmer ice core. The second approach used scanning electron  
242 microscopy of filters of melt water collected as part of the CFA analysis.

243 **Method 1.** Annual samples were provided between 1900 to 1998 CE. The ice core sample volumes  
244 provided ranged between 64 mL to 430 mL, and SCP analysis was conducted on every sample (total  
245 of 98) down-core. The surface area of each annual sample was 11.4 cm<sup>2</sup>.

246 Each ice core water sample, representing a single year, was filtered through a glass microfibre (GF/C)  
247 filter. The filtering apparatus was rinsed with deionised water between samples. The filter papers  
248 were then analysed for SCPs following a method adapted from Rose (1994), involving dissolution of  
249 the filter paper using hydrofluoric acid followed by a treatment with hydrochloric acid to remove any  
250 precipitate formed from the filter (Rose, 1994). Between acid treatments, the samples were washed  
251 with distilled water, centrifuged, and decanted. This was repeated twice as a final washing step to  
252 remove any remaining acid. SCPs are composed mostly of elemental carbon and are chemically  
253 robust and undamaged by the acid process. A known fraction of the final suspension was evaporated  
254 onto multiple coverslips and mounted onto microscope slides. The number and sizes of SCPs on the  
255 coverslips were counted using a light microscope at x400 magnification using criteria for SCP  
256 identification (Rose, 2008). Analytical blanks were included in duplicate for each sample batch, and  
257 no SCPs were observed in these blanks.

258 **Method 2.** The second approach used the melt water produced while processing ice samples on the  
259 BAS CFA system (Grieman et al., 2022). Melt water from the CFA waste lines was collected in new  
260 and sealed low-density polyethylene (LDPE; Nalgene™) bottles. Melt water represents the effective  
261 deposition of snow and particles over a 6.53 cm<sup>2</sup> surface. Samples were collected at annual  
262 resolution between 2011 and 1980 CE (33 samples), and between 5 and 8-year resolution between  
263 1900 and 1979 CE (16 samples). Sample volumes ranged between 44 mL and 927 mL. Meltwater  
264 from each sample was then filtered through 13 mm diameter, 1.0 µm pore size Whatman™  
265 Polycarbonate membrane filters, inside clean polypropylene Swinnex™ filter holders. Each filter  
266 was then mounted onto an aluminium stub for analyses on a scanning electron microscope (SEM) in  
267 the Earth Sciences Department at the University of Cambridge. Filters were imaged on a Quanta-



268 650F using back scattered electrons (BSE) on a low-pressure mode. Each filter was imaged at x800  
269 magnification for SCPs identification, following the analysis strategy presented in (Tetzner, 2021).  
270 SCPs were identified, based on their characteristic morphological and textural features. Analytical  
271 blanks were included every 15 samples in the filtration process. No SCPs were observed in the  
272 blanks.

### 273 ***Methane***

274 The Palmer methane (CH<sub>4</sub>) record was measured using a continuous flow analysis (CFA) (Stowasser  
275 et al., 2012; Grieman et al., 2022). The melted water stream (with an average melt rate of 3-5  
276 cm/min) is pumped through three steps of gas extraction; 1) a debubbler, where gravity takes away  
277 excess water, 2) a hydrophobic membrane, which separates gas bubbles from the water completely,  
278 and 3) a Nafion dryer, which absorbs water vapor on a molecular level. The dry gas sample stream is  
279 measured continuously with the commercially available laser spectrometer Picarro G2301 using  
280 wavelength-scanned cavity ring-down spectroscopy. The effective CH<sub>4</sub> stratigraphic resolution is 1.5-  
281 17.5 cm. The instrumental CFA CH<sub>4</sub> uncertainty is 10 ppb. The raw CH<sub>4</sub> mixing ratio was calibrated to  
282 NOAA gas standards (<https://gml.noaa.gov/ccl/airstandard.html>) to obtain absolute values and  
283 corrected for solubility in the meltwater stream. NOAA primary air standards (405.8 ppb and 869.1  
284 ppb [CH<sub>4</sub>] calibrated against the WMO (World Meteorological Organization) X2004A reference scale  
285 Dlugokencky et al., 2005) were measured before and after the CFA analysis providing a slope of  
286 1.034 and intercept of 3.708. The percentage of dissolved gas was calculated daily based on an  
287 experiment mimicking a melted water stream by mixing a working standard gas ([CH<sub>4</sub>] = 606.97 ppb)  
288 and deionized water. An average factor to increase the CH<sub>4</sub> mixing ratio is 8.75%.

289 The age of the CH<sub>4</sub> (or any other gas) is not equal to the age of the ice at a certain depth. The so-  
290 called gas-ice age difference (or ΔAge) (Lemieux-Dudon et al., 2010) relates to the atmospheric air  
291 trapping when the glacier ice forms. Pores between grains close at a certain depth (typically  
292 between 50 to 70 m in ice sheets but can be down to 120 m – (Buizert et al., 2021)), based on the  
293 snow and firn (granular snow that has not yet been compressed into ice) densification process.

294 The gas age at the depth that the pores lock-in (56.8 m for Palmer) is 0 years, because the air is  
295 exchanged in the opened pores with the atmospheric air on the surface. Throughout the lock-in zone  
296 bubbles close steadily, and the air defuses until the close-off depth. This occurs at 62.8 m, where  
297 the ΔAge is 17 years. The ΔAge is unique to each ice core site and often assumed to be constant for  
298 calculation convenience. One can identify ΔAge either by firn densification modelling (see e.g.  
299 Buizert et al., 2021) or by synchronizing with existing atmospheric gas records (Blunier et al., 2007).  
300 In this study, synchronisation to the Law Dome CH<sub>4</sub> record (Rubino et al., 2019) is adopted, matching  
301 the absolute values and slopes of the two records. The ΔAge was verified using the community firn  
302 model (Stevens et al., 2020) implemented using a varying snow accumulation rate, ranging from 0.17  
303 to 0.33 m w.e. yr<sup>-1</sup> based on measured density, annual layer depths (Emanuelsson et al. 2022a) and  
304 an annual average surface air temperature estimate of -28.5°C (ERA 5 reanalysis). The ΔAge from the  
305 community firn model approach ranges from 112 to 162 years, in agreement with the ΔAge estimate  
306 based on gas synchronisation of 90 to 140 years.

### 307 ***Numerical analysis***

308 Change points in CH<sub>4</sub> detected using a Bayesian Estimator of Abrupt change, Seasonal change, and  
309 Trend (BEAST) (Zhao et al., 2019). It is a decomposition algorithm with built-in Bayesian model  
310 averages. It allows identification of a time stamp of a long-term trend change in a noisy time series.

311 For CH<sub>4</sub>, the change-point analysis was run using 5-yearly averages (to reflect the mixing in the firn  
312 column) and presented with an estimated error of +/- 5 years.

## 313 **Results**

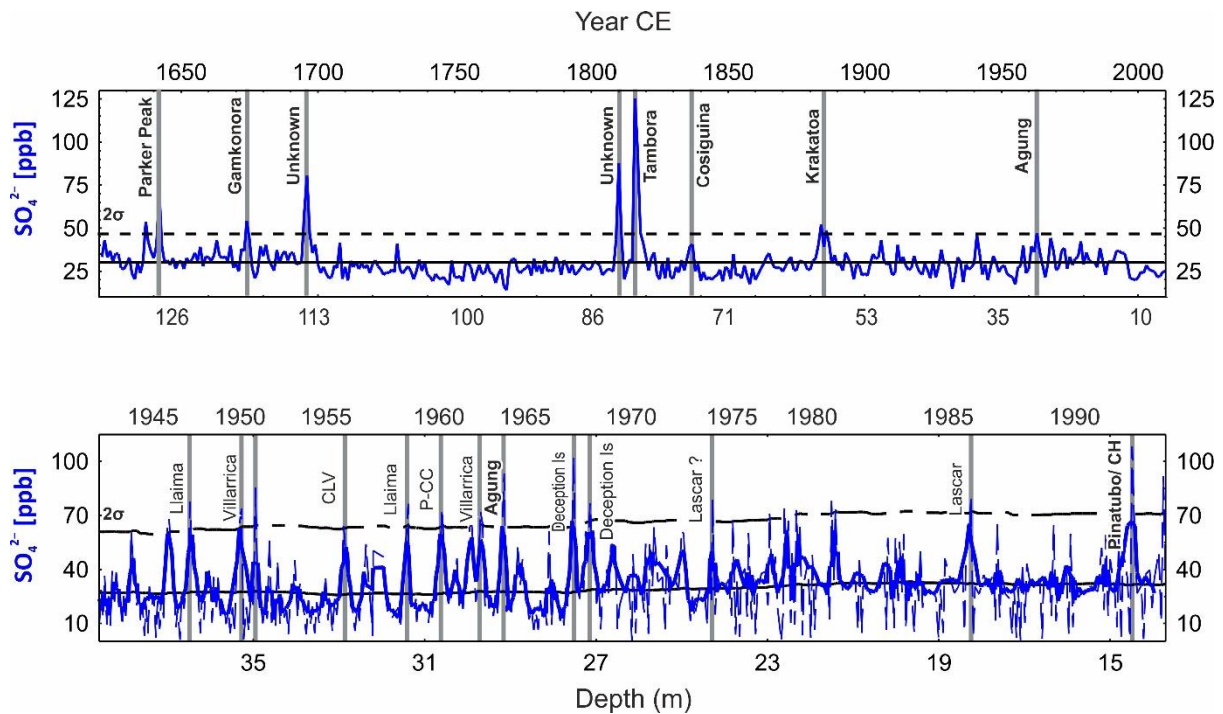
### 314 ***Chronology***

315 The Palmer ice core covers 391 years, 1621–2011 CE, and is one of the oldest records from the  
316 Antarctic Peninsula. The annual layer counted age scale is verified using the dates of known volcanic  
317 eruptions that have been detected in ice cores from both Antarctic and Arctic ice cores (Sigl et al.,  
318 2013). Six distinct peaks are observed in the annual averaged [SO<sub>4</sub><sup>2-</sup>] (Fig 3a) corresponding to  
319 volcanic eruptions documented in the WAIS divide ice core (Sigl et al., 2013). This includes Parker  
320 peak (1640), Gamkonora (1673), Tambora (1815), Cosiguina (1835), Krakatoa (1883), and Agung  
321 (1963). Two additional peaks are assigned to eruptions of unknown sources (1694 and 1809), which  
322 have been widely documented and identified in other Antarctic and Arctic sites (Emanuelsson et al.,  
323 2022b). The chronological error is established by evaluating the absolute age differences of volcanic  
324 horizons in Palmer, compared with documented ages in the WAIS divide ice core (Sigl et al., 2013).  
325 The maximum age difference is 6-months and thus our chronological error is ± 0.5-year.

326 The 1963 eruption of Agung, Indonesia, provides the closest tie-point for the proposed GSSP. While  
327 the annual average [SO<sub>4</sub><sup>2-</sup>] does not exceed the threshold of 2σ above the mean, the threshold is  
328 exceeded in the raw (~4 cm) data, at a depth of 29.2 m corresponding to 1964 (Fig 3b). The 1991  
329 eruption of Pinatubo, Philippines, is another well-documented eruption visible in many polar ice  
330 cores (Sigl et al., 2013). The threshold of 2σ is also only exceeded in the raw data, as a distinct [SO<sub>4</sub><sup>2-</sup>  
331 ] peak at 14.5 m corresponding to 1992 (Fig 3b). A delay of between one and three years is expected  
332 between the volcanic eruption and the deposition of [SO<sub>4</sub><sup>2-</sup>] at an Antarctic ice core site (Cole-Dai et  
333 al., 1997).

334 A number of smaller, more proximal, eruptions provide additional age constrains during the mid-20<sup>th</sup>  
335 century (3b) (Global Volcanism Program., 2022). Deception island, located in the South Shetland  
336 islands to the north of the Antarctic Peninsula is known to have been active between 1968 and 1970  
337 with a volcanic explosivity index (VEI) of 3. This is coincident with multiple [SO<sub>4</sub><sup>2-</sup>] peaks that exceed  
338 2σ (and 3σ) above the running 200-point average between 29.5 and 27.09 m (1968-1969 CE). Two  
339 peaks in [SO<sub>4</sub><sup>2-</sup>] at 36.5 m and 34.9 m most likely correspond to documented eruptions of Llaima  
340 (1945 CE) and Villarrica (1948/1949 CE), two stratovolcanoes from southern Chile with documented  
341 VEI 3 eruptions (Global Volcanism Program, 2022). These peaks are used to constrain the lower  
342 portion of the GSSP period.

343



344  
 345 **Figure 3.** Volcanic reference horizons in the Palmer chronology plotted on a depth (bottom axis) and  
 346 age-scale (top axis). (a) Annual average  $[SO_4^{2-}]$ , highlighting volcanic peaks (grey vertical lines)  
 347 identified in other Antarctic and bi-polar ice cores. Solid horizontal line indicates record average and  
 348 dashed horizontal line indicates  $2\sigma$  above the record average. (b) Raw ( $\sim 4$  cm resolution)  $[SO_4^{2-}]$   
 349 between 1942 and 1993 CE, highlighting both mid-latitude (Agung and Pinatubo, bold) and Southern  
 350 Hemisphere eruptions (Global Volcanism Program., 2022). Solid vertical lines indicate peaks  
 351 exceeding  $2\sigma$  above a 200-point running average (solid black curve), with corresponding dated  
 352 volcanic eruptions exceeding VEI 3. Cerro Hudson (CH), Carran-Los Venados (CLV) and Puyehue-  
 353 Cordon Caulle (P-CC).

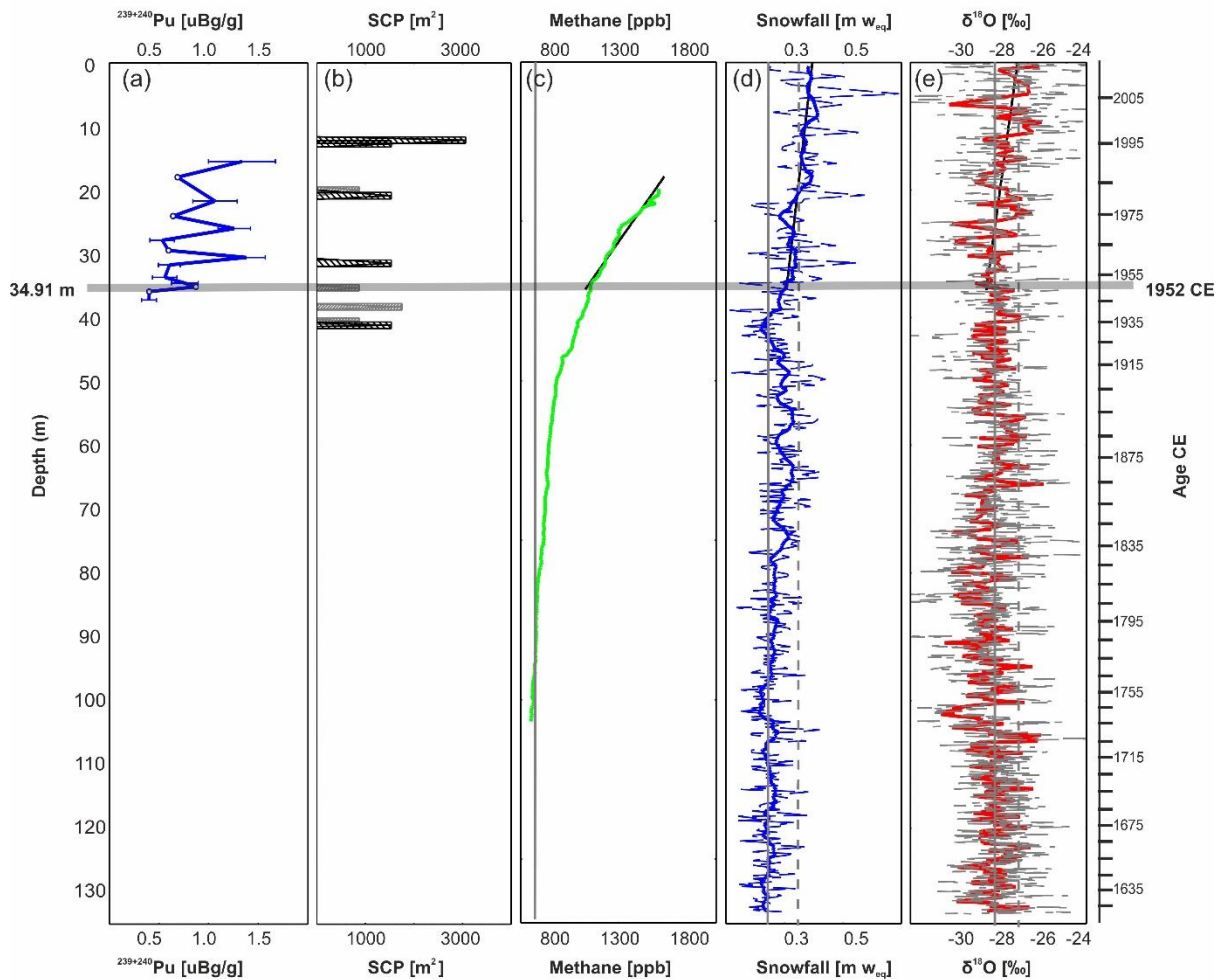
354

### 355 **Radioisotopes**

356 The  $^{239+240}\text{Pu}$  activity concentrations are presented in Figure 4 and 5. The first evidence of  $^{239+240}\text{Pu}$  in  
 357 the Palmer ice core is at 37.16 m depth, corresponding to years 1945-1946 CE. This constitutes the  
 358 lowest concentration (0.39  $\mu\text{B/g}$ ) in the Palmer record. The concentrations of  $^{239+240}\text{Pu}$  at Palmer  
 359 increase after 1945, through a subsidiary maximum in the early 1950s, reaching an absolute  
 360 maximum (1.38  $\mu\text{B/g}$ ) at 30.48 m by 1960-1961.

361 The low concentrations are consistent with previous  $^{239+240}\text{Pu}$  measured in Antarctic ice cores  
 362 (Arienzo et al., 2016), which are considerably lower than concentrations measured in Greenland ice  
 363 cores or Alpine ice cores (Gabrieli et al., 2013) or other mid-latitude archives (Waters et al., 2015).  
 364 Temporal trends are also broadly similar to those reported previously (e.g., Arienzo et al. 2016), with  
 365 two exceptions, in the date of first detection of Pu in the ice core, and post-1980 concentration  
 366 trends. In previous studies, the first detected Pu in Antarctic ice cores was 1952 CE, while in Alpine  
 367 cores it was as late as 1954 CE. The earlier detection in the Palmer ice core may reflect the different  
 368 analytical approaches, which has previously been measured using Inductively Coupled Plasma –  
 369 Sector Field Mass Spectrometer (ICP-SFMS) (e.g., (Gabrieli et al., 2013). However, the proximity to  
 370 the test sites (surrounding the Pacific region) may also explain the earlier detection on the Antarctic  
 371 Peninsula. Back-trajectory analysis indicates that the Pacific sector of the Southern Ocean is the

372 dominant source region for the 5-day air-parcels reaching Palmer (Thomas and Bracegirdle, 2015).  
 373 An apparent secondary increase in  $^{239+240}\text{Pu}$  in 1990 (albeit with low precision, with concentrations  
 374 close to detection limits), if real, may be a result of increased aerosol or dust supply, and likely  
 375 reflects local processes rather than any regional fallout process.

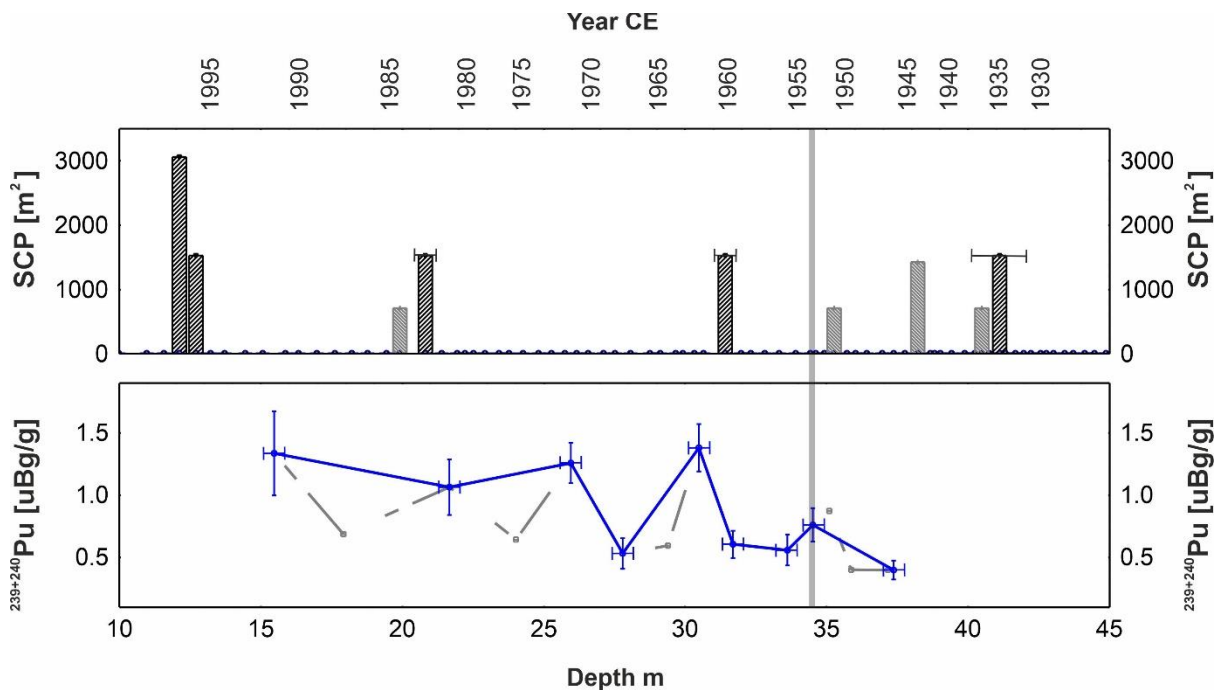


376  
 377 **Figure 4.** Anthropogenic proxies in the Palmer ice core in the context of the last 391 years. (a)  
 378 Plutonium ( $^{239+240}\text{Pu}$ ) [ $\mu\text{Bq/g}$ ] (solid), with error bars, presented as 2-yearly resolution. Open circles  
 379 indicate data below detection limit. (b) SCPs [ $\text{m}^2$ ] determined using discrete sampling (black bars,  
 380 annual resolution) and continuous sampling (grey bars, 1–5-year resolution). Auxiliary climatological  
 381 records (c) ( $\text{CH}_4$  [ppb]), (d) annual snowfall (meters of water equivalent) (dashed line) with running  
 382 decadal mean (thick solid curve) and (e)  $\delta^{18}\text{O}$  (‰) at 4cm resolution (grey dashed curve) and annual  
 383 averages (solid red curve). All data plotted on a depth scale (left axis), with age-scale shown on right  
 384 axis. The proposed GSSP is highlighted in the grey line corresponding to depth 34.91 m (1952 CE).

385  
 386 **Novel materials**

387 **Fly-ash (SCPs).** A total of five SCPs were identified using method 1 (discrete sampling), with a further  
 388 six SCPs identified using method 2 (continuous meltwater). These eleven particles constitute the first  
 389 evidence that SCPs have been deposited in Antarctica (Rose, 2015), and to our knowledge are the  
 390 first dated record of deposition in an Antarctic ice core. The SCP record is presented as a  
 391 concentration ( $\text{m}^2$ ) in Figure 4 & 5. The first SCP observed using method 1 is in the ice layer

392 corresponding to 1936 CE (40.48 m). For the continuous method (methods 2) the first SCP is  
 393 identified in the sample centred on 41.12 m depth, which encapsulates years 1930-1937 CE.  
 394



395  
 396 **Figure 5.** Anthropogenic proxies in the Palmer ice core between 1930-1998 CE. (a) SCPs [m<sup>2</sup>]  
 397 determined using discrete sampling (black bars, annual resolution) and continuous sampling (grey  
 398 bars, 1–5-year resolution). Blue dots indicate samples analysed where no SCPs were found. (b)  
 399 Plutonium (<sup>239+240</sup>Pu) [µBg/g], measured at bi-annual resolution from 1946 to 1991 CE (grey dashed  
 400 curve). Values which exceeded the limit of detection shown in blue (presented in Fig. 4). Error bars in  
 401 the x-axis denote depth resolution and error in the y-axis denote analytical uncertainty. Vertical grey  
 402 line denotes the proposed GSSP horizon at 34.91 m (1952 CE).

403 **Geochemical proxies for climate**

404 **Stable water isotopes.** δ<sup>18</sup>O are commonly used to reconstruct past surface temperature from  
 405 Antarctic ice cores (e.g., Stenni et al., 2017). The stable water isotope record from Palmer displays  
 406 evidence of seasonal cycles, ranging from -33.12 ‰ (winter) to -20.51‰ (summer). When converted  
 407 to decadal averages, the values from the most recent decade (2001-2010 CE) are the most  
 408 isotopically enriched (-26.82‰), thus suggesting that temperatures during the first decade of the  
 409 21<sup>st</sup> century were warmer than any decade since 1621 CE. δ<sup>18</sup>O has been increasing at a rate of 0.22  
 410 ‰ per decade since 1950 CE (Figure 4). This is consistent with other ice cores from the Antarctic  
 411 Peninsula, which display an isotopic warming during the late 20<sup>th</sup> century (Thomas et al., 2013;  
 412 Stenni et al 2017; Thomas and Tetzner, 2017), consistent with observations.

413 **Snow accumulation.** The amount of snowfall (snow accumulation) per year is determined as the sum  
 414 of precipitation, sublimation, evaporation, and melt (Thomas et al., 2017). Snow accumulation is  
 415 measured using the distance between chronological peaks (SO<sub>4</sub><sup>2-</sup> and H<sub>2</sub>O<sub>2</sub>) and converted to metres  
 416 of water equivalent per year (m w<sub>eq</sub> yr<sup>-1</sup>) based on the measured density. The record is corrected for  
 417 ice thinning and horizontal flow using the Nye model, which assumes thinning is proportional to  
 418 burial and is appropriate to use in the upper 10-15% of the ice sheet (Nye, 1963).

419 The snow accumulation increases from a decadal average of  $0.24 \text{ m w}_{\text{eq}} \text{ yr}^{-1}$  at the beginning of the  
420 record (1621-1630 CE) to a decadal average of  $0.38 \text{ m w}_{\text{eq}} \text{ yr}^{-1}$  in the final decade (2001-2010 CE).  
421 Snowfall has been increasing at a rate of 1.5 cm per decade since 1950 CE (Figure 4). This increase is  
422 consistent with other Antarctic Peninsula ice cores (Thomas et al., 2015, Porter et al., 2016) and is  
423 part of a regional trend, which began in the ~1920s (Thomas et al., 2015; 2017). Snow accumulation  
424 variability in the Antarctica Peninsula is driven by changes in tropical sea surface temperatures,  
425 regional sea ice conditions (e.g., (Thomas et al., 2015; Porter et al., 2016; Thomas et al., 2017)) and  
426 large scale modes of atmospheric circulation, most notably the SAM (Medley and Thomas, 2019).  
427 However, the underlying trend in snow accumulation can also be explained by the observed  
428 atmospheric warming (Medley and Thomas, 2019).

#### 429 ***Methane.***

430 The oldest  $\text{CH}_4$  measurement corresponds to approximately 1719 CE (+/- 5 years) with a  
431 concentration of  $645 \pm 10$  ppb. Concentrations of  $\text{CH}_4$  increase throughout the record, with a marked  
432 acceleration during the 20<sup>th</sup> century (Figure 4). The concentrations of  $\text{CH}_4$  have doubled by 1969 CE.  
433 The concentration at the lock-in depth (56.8 m) is  $1839 \pm 10$  ppb, which corresponds with modern  
434  $\text{CH}_4$  concentration in the atmosphere. At the pore close-off depth (62.8 m) the concentration is  $1511$   
435  $\pm 10$  ppb, which in terms of gas age is approximately 1977 CE. The corresponding concentration in  
436 the Law Dome ice core in 1977 CE is  $1476 \pm 5$  ppb (Rubino et al., 2019). The earliest observational  
437 monthly data, from air collected in glass flasks at the Palmer Station in Antarctica (Lan et al., 2022)  
438 from NOAA Global Monitoring Laboratory, began in January 1983 with a  $\text{CH}_4$  concentration of 1557  
439 ppb.

440 Change point analysis identifies two significant changes in the 5-yearly averaged  $\text{CH}_4$  below the  
441 critical depth (1977 CE). The change-points correspond to 1958 and 1883 CE, with probabilities of  
442 99% and 82% respectively. We emphasize that due to the gradual gas trapping and residence time in  
443 the atmosphere ( $9.1 \pm 0.9$  years, (Prather et al., 2012)), the  $\text{CH}_4$  change points cannot be interpreted  
444 as an exact year. However, the record suggest that the most significant decadal change detected in  
445 the 264-year  $\text{CH}_4$  record is centred on 1958 (+/- 5 years) CE.

#### 446 **Discussion**

447 The Palmer ice core meets the requirements of a GSSP site. The ice layers are of adequate thickness  
448 to support global correlation, with an accurate chronology independently verified using well-dated  
449 volcanic horizons. The Agung eruption of 1963 provides a well-dated reference horizon observed in  
450 other Antarctic and Greenland ice cores (Sigl et al., 2013) in close proximity to the GSSP depth. The  
451 Palmer location is also sensitive to volcanic  $[\text{SO}_4^{2-}]$  deposition from South American volcanos, with  
452 six  $[\text{SO}_4^{2-}]$  peaks in the 20-year window surrounding the GSSP (1952) coincident with dated Chilean  
453 eruptions (Fig. 3)(Global Volcanism Program, 2022). Snowfall is continuous, with small variations in  
454 seasonal deposition (Thomas and Bracegirdle, 2015), ensuring the record is not seasonally biased.  
455 The risk of reworking is small, with redistribution of surface snow estimated at just 5% (van Lipzig et  
456 al., 2004). The ice layers, and anthropogenic proxy data they contain, are continuous and unaffected  
457 by tectonic and sedimentary movements, or metamorphism. For this ice core specifically, the high  
458 elevation (1897 m) and southerly latitude ( $74^\circ\text{S}$ ) ensure that the site is not at risk of surface melting.  
459 Even under extreme future warming scenarios, the proposed marker for the onset of the  
460 Anthropocene is 34.91 m below the surface.

461 Antarctica is a unique location, governed by the Antarctic Treaty as a continent for peaceful  
462 purposes, freedom of scientific investigation and scientific observations and results that shall be

463 exchanged and made freely available. The Protocol on Environmental Protection designated  
464 Antarctica as a “natural reserve” in 1998, ensuring that future scientific exploration of the Palmer  
465 site is not at risk of exploitation or political changes. The Palmer ice core site is easily accessible by  
466 Twin Otter aircraft from several international research stations on the Antarctic Peninsula, and from  
467 other Antarctic gateways (e.g., South America, South Africa, New Zealand, and Australia). Thus,  
468 despite its remote location, future access for repeat drilling is feasible, achievable, and uniquely  
469 protected. The Palmer ice core used in this study is stored at a government funded research facility  
470 in the United Kingdom, ensuring safe, long-term storage and accessibility for the public and  
471 academics.

472 The Palmer ice core provides clear evidence that anthropogenic proxies, in the form of SCPs and  
473 radionuclides (Pu), have been deposited in Antarctica. Both proxies have no known natural source  
474 but are evident in the Palmer ice core from the mid-twentieth century onwards. SCPs have been  
475 recorded in the records of coastal mainland and sub-Antarctic lakes sediments (Rose et al., 2012),  
476 and in near-shore marine sediments (Martins et al., 2010). However, the SCP data presented here  
477 are the first for Antarctic ice and demonstrate for the first time that these markers of heavy industry  
478 (high temperature combustion) reached this more remote part of the Antarctic continent as early as  
479 1936 CE. The most likely source region of these SCPs is the South American or Australian continents,  
480 where coal powered energy industry was delayed relative to the United Kingdom and other nations  
481 (Rose and Appleby, 2005). However, the low analytical detection and transport mechanisms may  
482 also in part explain the delayed appearance relative to other global locations.

483 The presence of  $^{239+240}\text{Pu}$  in the samples corresponding to 1945/46 CE indicate that even the earliest  
484 bomb tests, which began in July 1945 CE, are detected in the Palmer ice core. The earlier tests were  
485 fission weapons, with fallout in the lowest layers of the atmosphere (Aarkrog, 2003). Despite the  
486 extremely low concentrations, the detection of  $^{239+240}\text{Pu}$  cannot be attributed to any natural source  
487 or emission scenario. Tropospheric transport of volcanic ash from South American and mid-latitude  
488 volcanoes has been observed reaching West Antarctica just 2-3 weeks following an eruption  
489 (Koffman et al., 2017). Thus, we hypothesize that Palmer’s location in the Western Antarctic  
490 Peninsula, with strong teleconnections with the tropical Pacific (e.g. (Thomas et al., 2013)),  
491 facilitated the tropospheric transport of  $^{239+240}\text{Pu}$  to the ice core. Making it possible to detect the fall-  
492 out from the first explosions in 1945/46 CE.

493 It was not until after 1952 CE, and the first thermonuclear tests, that stratospheric transport was  
494 sufficient to produce a global signal detectable at other polar (Arienzo et al., 2016) and mid-latitude  
495 ice core sites (Gabrieli et al., 2011; Gabrieli et al., 2013). While ash layers in ice cores can be used to  
496 demonstrate rapid tropospheric transport, the sulphate emitted from explosive volcanic eruptions  
497 can often take several years to be detected in Antarctic ice (Koffman et al., 2017). This may explain  
498 the delayed deposition in of  $^{239+240}\text{Pu}$  at other polar and mid-latitude locations. While the first  
499 detection of  $^{239+240}\text{Pu}$  occurs in 1945/1946 CE, the first recorded peak in values (significantly above  
500 the detection limit) corresponds to 1952/53 CE. This is consistent with the first thermonuclear tests  
501 and is a globally identifiable feature (Waters et al., 2015).

502 The highest concentrations of  $^{239+240}\text{Pu}$  (1960/61 CE) are coincident with the largest nuclear test (Tsar  
503 Bomba) in October 1961 CE. The radioactive debris from such large thermonuclear explosions can  
504 remain in the stratosphere for between 15-18-months (Zander and Araskog, 1973). Thus, despite the  
505 extremely distant source region for this test (Novaya Zemlya, Arctic Russia), the sufficiently long  
506 time-window for global stratospheric transport is feasible. This is also supported by other well  
507 documented stratospheric transport and deposition of particles and chemical species in the  
508 Antarctic ice core record (e.g., volcanic eruptions) (Sigl et al., 2013; Koffman et al., 2017). The brief

509 drop to lower concentrations during years 1963-1967 CE, may reflect the reduction in nuclear testing  
 510 following the partial test ban treaty in 1963 CE. A sharp decrease in deposition corresponding to  
 511 1967 CE has previously been observed in an ice core from the Alps (Gabrieli et al., 2011),  
 512 demonstrating the sensitivity of the ice core record to detect changes in  $^{239+240}\text{Pu}$  emissions. By 1970  
 513 CE, the concentrations of  $^{239+240}\text{Pu}$  returned to the 1960/61 CE levels, despite most of the testing  
 514 being moved underground. This may reflect the delayed and ongoing stratospheric  $^{239+240}\text{Pu}$   
 515 deposition. The analogous stratospheric transport of sulphate aerosols has been documented in the  
 516 Antarctic ice core record several years after the eruption took place (Koffman et al., 2017).  $^{239+240}\text{Pu}$   
 517 is still detected as recently as 1990/1991 CE.

518 Both SCPs and  $^{239+240}\text{Pu}$  are distinct markers of the Anthropocene, with no known natural source. The  
 519 long half-life of  $^{239}\text{Pu}$  (24,110 years), together with the high weathering resistance of the SCPs, mean  
 520 they will both be preserved in the ice as a *near*-perpetual marker. Thus, the presence of both could  
 521 be used as a primary stratigraphic marker that defines the lower boundary of the Anthropocene.  
 522 However, based on the globally synchronous detection of  $^{239+240}\text{Pu}$  corresponding to the first  
 523 thermonuclear tests, we propose that  $^{239+240}\text{Pu}$  be the primary marker at this site. The SCPs provide a  
 524 clear secondary marker for the Palmer GSSP, visible under magnification.

525 In addition to the clear evidence of Anthropogenic proxies (SCPs and  $^{239+240}\text{Pu}$ ), the auxiliary  
 526 climatological records confirm that the Palmer site has experienced significant climate and  
 527 environmental change since the 1950s. The observed isotopic warming, and positive trend in snow  
 528 accumulation, since 1950 is consistent with regional observations (e.g., Thomas et al., 2017; Stenni  
 529 et al., 2017), suggesting warmer atmospheric surface temperatures and increased precipitation.  $\text{CH}_4$   
 530 provides a unique record of past atmospheric conditions that place this proposed GSSP site in a  
 531 global context. Although produced by a variety of natural processes, the acceleration of  $\text{CH}_4$  in the  
 532 atmosphere in recent centuries has been linked to changes in land use, landfills, agricultural activity,  
 533 industrialisation, and coal mining (Kirschke et al., 2013), the latter directly linked to the SCP record,  
 534 as a marker for coal combustion. As a more potent greenhouse gas than carbon dioxide,  $\text{CH}_4$   
 535 accounts for more than one-quarter of the radiative imbalance (IPCC, 2013) and plays a major role in  
 536 driving present and future climate.  $\text{CH}_4$  does not represent a spike, or a single stratigraphic point,  
 537 and is thus not proposed as a GSSP marker. However, change point analysis identifies a statistically  
 538 significant marker centred around 1958 CE (+/- 5 years), which is broadly coincident with the first  
 539 recorded  $^{239+240}\text{Pu}$  peak (1952/1953 CE). Thus, we have included  $\text{CH}_4$  in this study to provide a unique  
 540 global context to support this GSSP proposal.

541 **Table 1.** Summary of anthropogenic proxies detected in the Palmer ice core and their ranking as a  
 542 GSSP marker. Presented with the year they are first detected ( $^{239+240}\text{Pu}$  and SCPs), their peak  
 543 concentration ( $^{239+240}\text{Pu}$  and SCPs) and the proposed year to mark the onset of the Anthropocene.

Proxy/ record	First detected (CE)	Peak (CE)	GSSP (CE)	Depth (m)	GSSP ranking
$^{239+240}\text{Pu}$	1946	1960	1952	34.91	Primary
SCPs	1936	1996	1936	40.73	Secondary

544

545 To conclude, we propose the Palmer ice core as a candidate for the Global Boundary Stratotype  
 546 Section and Point (GSSP) to mark the onset of the Anthropocene. Based on the first peak in  $^{239+240}\text{Pu}$ ,  
 547 our primary marker, we propose that the onset of the Anthropocene was 1952 CE observed at a  
 548 depth of 34.91 m.



549 While ice cores have been used as previous GSSP sites (Walker et al., 2009), to date, there are no  
550 GSSP sites in Antarctica. Thus, if this site is selected as the GSSP for the Anthropocene we propose it  
551 be named the “Antarctican”. The anthropogenically driven mass loss from the Antarctic ice sheet,  
552 observed since the mid-20<sup>th</sup> century, is predicted to produce 5 cm of global mean sea level rise by  
553 2100 (DeConto et al., 2021). Thus, it seems especially fitting that this new epoch be named for the  
554 continent that will arguably play the largest role in governing the planet’s future habitability.

555

#### 556 **Data availability**

557 The Palmer data is available from the UK Polar Data Centre (UKPDC)  
558 <https://doi.org/10.5285/b3eca350-79aa-49b2-bd6b-ffee86ad6559>, and  
559 <https://doi.org/10.1017/jog.2021.75>.

560

561

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