
#### Abstract

\title{ of Document: \\ CONSTRAINTS ON THE DEPOSITIONAL AGES OF LESSER HIMALAYAN ROCKS IN CENTRAL NEPAL AND THEIR STRUCTURAL IMPLICATIONS }


Katherine Burgy, Master of Science, 2009

Directed by:
Dr. Aaron Martin, Department of Geology

The lack of good exposures and paucity of datable horizons in central Nepal has hindered the ability of geologists to piece together a relatively cohesive and straightforward stratigraphic succession within the Lesser Himalaya. U-Pb isotopic analyses of detrital zircons from the Modi Khola valley indicates maximum depositional ages of $\sim 1875$ Ma for the Kuncha Formation, $\sim 1800 \mathrm{Ma}$ for the Fagfog Formation, and $\sim 1780$ Ma for the Kushma Formation. The intrusive $1831 \pm 17$ Ma Ulleri augen gneiss provides a minimum depositional age bound for the Kuncha. Combined, these data suggest the Kuncha Formation is the oldest member of the Lesser Himalayan series in central Nepal. Additionally, $\delta^{13} \mathrm{C}$ data suggest the Malekhu Formation of the Lakharpata Group was deposited before ca. 1250 Ma . A field mapping comparison based on the redefined stratigraphy indicates the Ramgarh thrust is located $>10 \mathrm{~km}$ farther south than previously mapped, potentially reducing regional shortening estimates.

CONSTRAINTS ON THE DEPOSITIONAL AGES OF LESSER HIMALAYAN ROCKS IN CENTRAL NEPAL AND THEIR STRUCTURAL IMPLICATIONS

By

Katherine Burgy

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Advisory Committee:
Assistant Professor Aaron Martin, Chair
Associate Professor Alan J. Kaufman
Professor Roberta Rudnick
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## Chapter 1 - Introduction to the Himalaya

### 1.1 Orogenesis

"Mountains are the beginning and end of all scenery" - John Ruskin, art critic and author
To geologists mountains are more than just majestic scenery; they represent a remarkable result of plate tectonics. Orogenesis, or mountain building, offers insight into the workings of the Earth; including past and present plate motions, crustal strength and deformation, and related magmatism and metamorphism. In addition, mountains influence regional and global climate by shifting air circulation patterns, leading to rain shadows and orographic effects. Mountains isolate plant and animal populations, leading to the development of new species and act as barriers to migration. Numerous human civilizations have also benefitted from the strategic advantages conferred by occupation of the topographic high ground.

### 1.2 The Himalaya

The word Himalaya comes from the Sanskrit hima or "snow" and ālaya or "abode", in other words "place where there is snow". Bounded to the north by the Tibetan Plateau and to the south by the Indian subcontinent, the range extends from $\sim 76^{\circ}$ to $\sim 91^{\circ} \mathrm{E}$ longitude, and is home to eight of the ten highest peaks in the world including Mt. Everest and K2 (Fig. 1.1; Hodges, 2000). Additionally, the range partially or entirely encompasses the countries of Nepal, Tibet, China, India, Pakistan, and Bhutan.

The Himalaya are considered the "type" example of a continent-continent collision. Beginning around 55-50 Ma ago and continuing to the present day, the collision of the Indian and Eurasian plates has accommodated a minimum of 2500 km of convergence (Achache et al., 1984; Patriat and Achache, 1984; Besse et al., 1984; Besse and Courtillot, 1988; Rowley, 1996; Guillot et al., 2003; DeCelles et al., 2004; Leech et al., 2005; Leech et al., 2007). Of this, approximately $800-1200 \mathrm{~km}$ are thought to be accommodated by the Himalayan fold and thrust belt (Lyon-Caen and Molnar, 1985; Lillie et al., 1987; Srivastava and Mitra, 1994; Bilham et al., 1997; Powers et al., 1998; Larson et al., 1999; Lavé and

Avouac, 2000; DeCelles et al., 2002). The fold and thrust belt includes mostly south-vergent thrusts and associated folds, which consist of lithified sediments accumulated on the northern margin of India prior to collision (DeCelles et al., 2001).

Regional balanced cross sections across the fold and thrust belt reveal minimum shortening estimates, however they do not include the shortening contributions of penetrative strain or small scale folds and faults, which would increase total shortening estimates considerably (DeCelles et al., 2002). The western portion of the Himalaya, including Pakistan, western Nepal, and northwestern India, appear to have accommodated the most shortening, on the order of 353-743 km from the South Tibetan Detachment System (STDS) to the Main Frontal thrust (MFT) (Coward and Butler, 1985; Srivastava and Mitra, 1994; DeCelles et al., 1998; DeCelles et al., 2001; DeCelles et al., 2002; Robinson et al., 2006). In central Nepal, Schelling (1992) estimates 210-280 km of shortening between the STDS and the Main Frontal thrust. In the eastern portion of the Himalaya, estimates of shortening range from 185-245 km between the STDS and MFT, and up to 323 km from the Indus-Yalu suture to the MFT (Schelling and Arita, 1991; Hauck et al., 1998).


Figure 1.1- The Himalaya from space. Image from NASA World Wind.

In Nepal, where the focus of this research is located, the fold and thrust belt is divided into four tectonostratigraphic units from south to north, the Subhimalayan (SH), Lesser Himalayan (LH), Greater Himalayan (GH), and Tibetan Himalayan (TH) zones, all of which are bounded by faults (Heim and Gansser, 1939). These tectonostratigraphic, or geologic, units should not be confused with the physiographic provinces of the Himalaya by the same name (Fig. 1.2). This paper will refer only to tectonostratigraphic divisions. The fold and thrust belt region has received a great deal of attention because it can potentially answer many of the questions geologists have about the Himalaya and mountain building in general.


Figure 1.2 - Geologic cross section of the Himalaya with the tectonostratigraphic units marked on the cross section and physiographic provinces delineated below. TH - Tibetan Himalaya, GH - Greater Himalaya, LH- Lesser Himalaya, LHD - Lesser Himalayan Duplex, SH - Subhimalaya, STD - South Tibetan Detachment, MCT - Main Central thrust, MBT - Main Boundary thrust, MFT - Main Frontal thrust. Modified from Robinson et al., 2006.

### 1.3 Fold and Thrust Belt Formation

"It is a truism in structural geology that broad, thin-skinned fold and thrust belts are easy to find, but difficult to explain mechanically" - William Chapple, Geologist

Early workers recognized fold and thrust belts as areas of extensive crustal shortening and their geometries and kinematics became the focus of extensive research (Armstrong and Oriel, 1965; Bally et al., 1966; Dahlstrom, 1970; Elliott 1976 a, b; Chapple, 1978; Bombolakis, 1986; Butler, 1987). Chapple
was one of the first to enumerate several defining characteristics of fold and thrust belts as we know them today. He described a wedge shaped region of deformation, thicker at the back (hinterland) and thinner at the toe (foreland), which was bounded at its base by a relatively weak geologic layer. He noted that shortening and thickening occurred throughout the wedge, primarily towards the back-end of the wedge. Chapple posited that material within the wedge maintained cohesion and that displacement on the thrust faults was relayed to successive, adjacent faults. The driving force responsible for the development of the fold and thrust belt was a "compressive flow" or "push from behind", as opposed to gravity driven "gliding" or other previously proposed mechanisms (Elliott 1976 a,b).

Additional work revealed other common characteristics of fold and thrust belts including a positive, linear relationship between fault displacement and fault length, and foreland propagation of insequence thrusting, although there are cases of out-of-sequence thrusting as well (Armstrong and Oriel, 1965; Bally et al., 1966; Elliott 1976 a, b; Goff and Wiltschko, 1992; Pearson and DeCelles, 2005; Robinson et al., 2006; Robinson, 2008). A regular decrease in spacing between thrust ramps toward the foreland has also been noted (Bombolakis, 1986; Goff and Wiltschko, 1992; Panian and Wiltschko, 2007).

To explain these observations, the critical taper theory was proposed by Davis et al. in 1983, and has subsequently been elaborated upon by numerous workers (Dahlen et al., 1984; Goff and Wiltschko, 1992; Makel and Walters, 1993; Horton, 1999; DeCelles and Mitra, 1995; Strayer, 2001; Bollinger et al., 2006). The theory accepts the basic geometry described by Chapple and explains the mechanics of wedge formation as analogous to a wedge of snow or dirt pushed along by a plow. At first, the material deforms internally until a wedge shape forms, and sliding begins along the base. At this point the wedge has attained a critical taper, where all points within the wedge are on the verge of failure under horizontal compression. As material is added or lost, or as internal factors such as pore fluid pressure change, the wedge geometry ( $\alpha$ or $\beta$ ) responds through internal deformation or faulting to maintain its' critical state (Fig. 1.3; Davis et al., 1983; Horton et al., 1999).

How and where the first thrust fault in a thrust sheet forms is currently debated. Most modelers agree that maintenance of wedge geometry is likely the main controlling factor in thrust fault formation. However, a variety of factors can influence where the break occurs. Work by Goff and Wiltschko (1992) suggests that the weight of overlying thrust sheets stabilizes the footwall of the incipient thrust, i.e. that the length and mass of overlying thrust sheets have a direct effect on the location of the next thrust sheet. Therefore, rates of thrust fault emplacement and erosion might also contribute to variation in the spacing of thrust faults. It is easy to understand why the prediction of thrust sheet dimensions is difficult! Nevertheless, it is reasonable to assume there is some minimum thickness of a thrust sheet necessary for the rocks to remain cohesive during movement and to avoid complete erosion during the time it takes a thrust sheet to be emplaced.


Figure 1.3 - Original model developed by Davis et al. (1983) that used sand and a moving mylar sheet to approximate wedge formation. Lower schematic shows an idealized critical taper (Horton, 1999).

### 1.4 Enduring Questions

By developing an extensive and cohesive understanding of the Himalayan fold and thrust belt, geologists can gain insight into the processes that shape the face of our entire planet. However, many questions remain incompletely answered, including: (1) How do mountain ranges fully accommodate shortening? Possible mechanisms include faulting, internal deformation, and extrusion of crust, but may likely be a combination of these and other mechanisms. (2) What is the order and timing of shortening
events? Thrusting may occur in- or out- of- sequence. (3) How much shortening occurred and where? Regional shortening estimates vary and are dependent upon correct mapping of faults and geologic units. (4) Where are the major faults located, how much movement occurred on them, and are they still active? This information has ramifications for the area's geologic history.

In order to answer questions such as the ones listed above, fundamental information about the geology of the fold and thrust belt is needed. What are the ages of the geological units? How are the units related? The development of regional stratigraphy and subsequent field mapping provide the foundation of knowledge necessary to make interpretations of the larger "big picture" structures involved in mountain building. In this study, I examine Lesser Himalayan rocks in central Nepal to obtain depositional ages of the units. From these ages I evaluate the accuracy of the current stratigraphy and subsequently modify regional geologic maps, as necessary. This work relates to the broader questions about fold and thrust belts because the location of faults, hence the thicknesses of the overlying thrust sheets and the development of accurate cross- sections and regional shortening estimates, are dependent upon correct mapping based on accurate stratigraphy.

Of the two units that make up the Lesser Himalayan series, only the younger unit, the Tansen, is fossiliferous. The older Nawakot Unit lacks fossils that might indicate relative ages of its members, so other methods of dating are necessary. To this end, I use U-Pb isotopic analyses of detrital zircons sampled from several siliciclastic units to obtain maximum ages of deposition. These ages are further constrained by a cross cutting igneous unit of known age. I also use chemostratigraphic data (time series $\delta^{13} \mathrm{C}$ variations), sampled from the stratigraphically highest carbonate within the Lesser Himalaya, to place a lower limit on deposition of the Lesser Himalayan sequence. The rearrangement of stratigraphy based on these analyses lead to a revision of current maps and fault placement, thus changing regional shortening estimates based on balanced cross-sections.

## Chapter 2 - Geochemical Background of Isotopic Systems Utilized

### 2.1 Perspective

This work examines several geologic units from central Nepal whose lithologies range from clean quartzite and sandy phyllite to crystalline carbonate and calcareous shale. In order to constrain depositional ages of these units several methods were applied. A brief description of each isotopic system used in this research is covered below.

### 2.2 Carbon

The carbon cycle is a series of complex interactions between the biosphere, atmosphere, hydrosphere, and lithosphere. Of all the carbon available on Earth, $98.89 \%$ is found in the form of carbon-12, while only $1.11 \%$ consists of the carbon-13 isotope. The ratio of the two isotopes in the Earth reflects global-scale oceanographic and climatic changes based upon isotopic fractionation (Prothero and Schawb, 2004).

Isotope fractionation includes equilibrium isotope effects, where the heavier isotope ( in this case carbon-13) preferentially goes into the compound in which it is most strongly bound (Bigeleisen, 1965), and kinetic isotope effects, where the rate of a chemical reaction is sensitive to the atomic mass at a particular position of the reaction species (Hayes, 1983).

### 2.2.1 Delta Notation

Due to the fact that isotopic abundances may differ at or beyond the third significant figure, geologists have developed delta notation expressed in parts per mil or parts per thousand (\%), as the conventional way to communicate findings. Delta notation is given by the following equation (1):

$$
\begin{equation*}
\delta^{13} \mathrm{C}=\frac{\left.\left[\left({ }^{13} \mathrm{C} /{ }^{12} \mathrm{C}\right)_{\text {sample }}-\left({ }^{13} \mathrm{C} /{ }^{12} \mathrm{C}\right)_{\text {standard }}\right)\right] * 1000}{\left({ }^{13} \mathrm{C} /{ }^{12} \mathrm{C}\right)_{\text {standard }}} \tag{1}
\end{equation*}
$$

The standard used for calibration has traditionally been the Pee Dee Belemnite (PDB). As there is no longer a supply of this material, alternative standards such as NBS-19 (this study) are used and then referenced to Vienna-PDB (V-PDB).

### 2.2.2 Carbon Cycling

While critical to life on the planet, the majority of carbon on the Earth's surface is actually sequestered in sedimentary rocks, namely carbonates and solid organic compounds (Fig. 2.1; Ripperdan, 2001). In ocean water carbon is largely present in the form of the bicarbonate ion $\left(\mathrm{HCO}_{3}{ }^{-}\right)$. Carbonate minerals are formed by the combination of divalent cations, predominantly $\mathrm{Ca}^{2+}$, with the carbonate ion (Hoefs, 2004):
$\mathrm{Ca}^{2+}+\mathrm{CO}_{3}{ }^{2-}=\mathrm{CaCO}_{3}$


Figure 2.1 - Diagram of carbon cycling pathways. Modified from Ripperdan, 2001.

Short term cycling and fractionation of carbon isotopes is dominated by photosynthetic fractionation of $\mathrm{CO}_{2}$, which preferentially incorporates ${ }^{12} \mathrm{C}$ into plant tissues (O'Leary, 1981; Farquhar et
al., 1989). When photosynthesis occurs in the oceans, it may enrich the surrounding ocean water in ${ }^{13} \mathrm{C}$. This kinetic isotope effect is the primary basis for the fractionation of carbon isotopes between the $\mathrm{C}_{\text {organic }}$ ( $\mathrm{C}_{\text {org }}$ ) and $\mathrm{C}_{\text {carbonate }}\left(\mathrm{C}_{\text {carb }}\right)$ reservoirs (Kaufman and Knoll, 1995). The magnitude of fractionation between two compounds, A and B, can be described by a fractionation factor $(\alpha)$ in which:
$\alpha_{A-B}=R_{A} / R_{B}$
where R is the ratio of numbers of any two isotopes in a compound A divided by the corresponding ratio in compound B (Hoefs, 2004, Kaufman, Pers. Comm.).

Long term cycling of carbon is based on the flux of carbon into the oceans from weathering and out gassing, and the corresponding outward flux through sedimentation and burial (Broeker, 1970; Hayes, 1983; Buick et al., 1995; Ripperdan, 2001). Most importantly, secular changes in the carbon isotope record reflect the relative proportions of $\mathrm{C}_{\text {org }}$ and $\mathrm{C}_{\text {carb }}$ being buried in marine sediments (Broeker, 1970; Hayes, 1983; Kaufman and Knoll, 1995; Veizer et al., 1999; Anbar and Knoll, 2002).


Figure 2.2 - Schematic of the open system ocean as a reactant chamber with input/reactant R, and outputs/ products Q and P , which in this diagram represent $\mathrm{C}_{\text {carb }}$ and $\mathrm{C}_{\text {org }}$ respectively. Bottom diagram demonstrates the isotopic fractionation (A) between Q , and P .

The open system behavior of the ocean can be described mathematically in the terms of carbon inputs and carbon outputs (Fig. 2.2; Hayes, 1983). If a steady state of inputs (reactant R in Fig. 2.2) and
outputs (products Q and P in Fig. 2.2) is assumed, fractionation between outputs is controlled by the kinetic isotope effect:
$\mathrm{R}_{\mathrm{pe}}=\alpha_{\mathrm{P} / \mathrm{R}} * \mathrm{R}_{\mathrm{re}}$
$\mathrm{R}_{\mathrm{qe}}=\alpha_{\mathrm{Q} / \mathrm{R}} * \mathrm{R}_{\mathrm{re}}$
which when combined yield:

$$
\begin{equation*}
\mathrm{R}_{\mathrm{pe}} / \mathrm{R}_{\mathrm{qe}}=\alpha_{\mathrm{P} / \mathrm{R}} / \alpha_{\mathrm{Q} / \mathrm{R}} \tag{6}
\end{equation*}
$$

where $\mathrm{R}_{\mathrm{pe}}$ is the ratio of P at equilibrium, $\mathrm{R}_{\mathrm{qe}}$ is the ratio of Q at equilibrium, $\alpha_{\mathrm{P} / \mathrm{R}}$ is the fractionation factor between $P$ and $R, \alpha_{Q / R}$ is the fractionation factor between $Q$ and $R$, and $R_{r e}$ is the ratio of reactant at equilibrium. Considering the law of conservation of mass:
$n_{i} \mathrm{~F}_{\mathrm{i}}=\mathrm{n}_{\mathrm{a}} \mathrm{F}_{\mathrm{a}}+\mathrm{n}_{\mathrm{b}} \mathrm{F}_{\mathrm{b}}$
where n is the molar quantity of carbon in i , a , and b , and F is the fractional isotopic abundance or $\mathrm{F}=$ $\left[{ }^{13} \mathrm{C} /\left({ }^{13} \mathrm{C}+{ }^{12} \mathrm{C}\right)\right]$. Therefore:
$F_{i}=f_{a} F_{a}+\left(1-f_{a}\right) F_{b}$
where $f_{a}$ is the fraction of input $i$ that is going to product $A\left(n_{a} / n_{i}\right)$, and $\left(1-f_{a}\right)$ is equal to the fraction of input going to product $B\left(n_{b} / n_{i}\right)$. Simultaneous solving of equations 6 and 8 yields:
$\delta_{\mathrm{p}}=\delta_{\mathrm{r}}+\left(1-\mathrm{f}_{\mathrm{p}}\right) \varepsilon$
and
$\delta_{\mathrm{q}}=\delta_{\mathrm{r}}-\mathrm{f}_{\mathrm{p}} \varepsilon$
where $\delta_{\mathrm{r}}$ is the isotopic composition of carbon entering the ocean, $\delta_{\mathrm{p}}$ is the isotopic composition of organic carbon, $\delta_{q}$ is the isotopic composition of carbonate, $\mathrm{f}_{\mathrm{p}}$ is the fraction of organic carbon being buried in sediments, and $\left.\varepsilon=\left[\left(\alpha_{\mathrm{P} / \mathrm{R}} / \alpha_{\mathrm{Q} / \mathrm{R}}\right)-1\right)^{*} 10^{3}\right]$. The typical values of carbon entering and leaving the
ocean are $\delta_{\mathrm{r}} \sim 5.5 \%$ and $\varepsilon \sim 25 \%$, respectively (Kaufman, Pers. Comm.). Ultimately, if burial of $\mathrm{C}_{\mathrm{org}}$ increases, the long term carbon isotope record preserved in carbonates will record more positive values. If relatively less $\mathrm{C}_{\text {org }}$ is being buried in sediment, long term carbon isotope values of carbonates will be more negative.

Therefore, variation in the carbon isotope record is enhanced or attenuated by a variety of processes including: changes in biological productivity, changes in geologic preservation rates of organic carbon, atmospheric $\mathrm{CO}_{2}$ sorption by weathering of continental materials, changes in $\mathrm{CO}_{2}$ content in ocean water based on temperature changes, oceanic circulation, and volcanic out gassing (Ripperdan, 2001). Typical values of some carbon reservoirs can be seen in Figure 2.3.

Additional variations in the carbon record are caused by complicating factors such as: nonequilibrium processes at the depositional interface, preservational bias, local ${ }^{13} \mathrm{C}$ heterogeneities from organic matter, and post-depositional diagenesis, including the liberation of volatiles with increasing temperature and the infiltration of externally derived fluids (Valley, 1986; Kohn and Valley, 1994; Ripperdan, 2001).


Figure 2.3 - Typical $\delta^{13} \mathrm{C}$ values of some carbon reservoirs in \%ov-PDB (Hoefs, 2001).

Many researchers, working in sedimentary basins around the world, have developed a composite record of carbon isotopes through time. It appears that throughout Earth history, with one or two notable exceptions, variable and fluctuating carbon isotope values have been the norm. Though prior to 2.6 Ga the carbon isotope curve is essentially flat, workers note that significant positive and negative excursions similar to modern day values were seen (Schidlowski et al., 1983; Karhu and Holland, 1996; Lindsay and Brasier, 2002). From $\sim 1.8$ to 1.0 Ga another relatively quiescent interval in the carbon record occurs, with isotopic values typically near $0 \pm 2 \%$ (Knoll et al. 1995, Buick et al. 1995; Kaufman, 1997; Des Marais, 1997). Towards the end of the Mesoproterozoic and the beginning of the Neoproterozoic, researchers observe a general trend of increasing fluctuations in carbon isotopes (Kah et al., 1999; Kumar et al., 2002), which gradually increase in amplitude up to $10 \%$ and greater, most often associated with widespread glaciations (Kaufman and Knoll, 1995; Kaufman et al., 2006; Kaufman et al., 2007; Tewari and Sial, 2007). The wildly swinging values of the Neoproterozoic eventually give way to more moderate fluctuations of the Phanerozoic (Zachos et al., 2001; Tewari and Sial, 2007). Several authors attribute the large fluctuations in the isotopic carbon record to tectonism and the formation or break up of supercontinents and a web of closely linked phenomena including carbon sequestration, rising atmospheric oxygen levels, and glaciations (Derry et al., 1992; Kaufman and Knoll, 1995; Berner, 2001; Ripperdan, 2001; Anbar and Knoll, 2002; Lindsay and Brasier, 2002; Kaufman et al., 2007).

### 2.3 Oxygen

Typically measured simultaneously with carbon isotopes and expressed in the same delta notation, oxygen isotopes provide important chemostratigraphic information to geologists. Changes in atmospheric oxygen are intimately related to the carbon cycle, including the biologic and tectonic processes of photosynthesis, sedimentation/burial of organics, weathering, and erosion (Berner et al., 2001). The formation of supercontinents leads to the burial of large amounts of organic carbon, which in turn leads to an increase in atmospheric $\mathrm{O}_{2}$ levels (Lindsay and Brasier, 2002). Increases in atmospheric oxygen, in turn, affect the oxygen content of ocean waters.

In addition, Urey (1947) and Emiliani (1955) determined that oxygen isotopes $\left({ }^{16} \mathrm{O}=99.756 \%\right.$ and ${ }^{18} \mathrm{O}=0.205 \%$ ) fractionate based on temperature. Later work indicated that ice volume (glaciation) plays a larger role in changing the isotopic ratios of oxygen in seawater (Prothero and Schwab, 2004). In this system, ${ }^{16} \mathrm{O}$ rich water is preferentially evaporated, which is then precipitated from clouds. Due to global atmospheric circulation patterns, these clouds move out from the equator towards the poles. During glaciated periods, this water is captured on land and fails to return to the oceans. The oceans thus become progressively enriched in ${ }^{18} \mathrm{O}$ over time (Fig. 2.4; Alley and Cuffey, 2001). These findings make oxygen isotopes indicators of paleo-ice volume.


Figure 2.4 - Schematic of oxygen isotope fractionation during evaporation and precipitation. Values given are only illustrative. Modified from Alley and Cuffey, 2001.

A globally correlated curve of isotopic oxygen values throughout time has been developed. However, while $\delta^{13} \mathrm{C}$ values have been shown to be relatively resistant to diagenetic changes, oxygen isotopes are much more easily altered (Hudson, 1977; Tucker, 1983; Burdett et al., 1990; Kaufman et al., 1991). This is likely due to the nature of most meteoric and metamorphic fluids, which have relatively little carbon to exchange with the rocks, but plenty of oxygen. The oxygen atoms are supplied from the fluid. Work by Baumgartner and Valley (2003) reveal that as metamorphism increases, the $\delta^{18} \mathrm{O}$ values of carbonates decrease due to a variety of volatization reactions. Similarly, Alley and Cuffey (2003) found that most meteoric fluids are relatively depleted in ${ }^{18} \mathrm{O}$, which would in turn result in lower measured $\delta^{18} \mathrm{O}$ values. Furthermore, Jacobsen and Kaufman (1999) demonstrate that a much larger (factor of 10 or
greater) water to rock ratio is necessary in order to reset ${ }^{13} \mathrm{C}$ values versus ${ }^{18} \mathrm{O}$ values (Fig. 2.5). Typical values of some oxygen reservoirs are seen in Fig. 2.6.


Figure $2.5-\delta^{18} \mathrm{O}$ and $\delta^{13} \mathrm{C}$ values with increasing water to rock ratio for both closed and open systems. Modified from Jacobsen and Kaufman, 1999.


Figure 2.6 - Typical $\delta^{18} \mathrm{O}$ values of some oxygen reservoirs in \%ov-PDB (Hoefs, 2001).

## 2. 4 Uranium-Lead

The uranium-thorium-lead system is one of the more complicated systems in radiogenic isotopes. The system includes three radiogenic parent isotopes which decay to three different daughter products, as
well as emit alpha ( $\alpha$ or ${ }^{4} \mathrm{He}$ ), beta $\left(\beta^{-}\right)$, and antineutrino ( $\overline{\mathrm{v}}$ ) particles (Fiorentini et al., 2007). For U-Pb dating, the ${ }^{238} \mathrm{U}$ and ${ }^{235} \mathrm{U}$ pathways are used:
${ }^{238} \mathrm{U}=8{ }^{4} \mathrm{He}+6 \beta^{-}+6 \overline{\mathrm{v}}+{ }^{206} \mathrm{~Pb}$
${ }^{235} \mathrm{U}=7{ }^{4} \mathrm{He}+4 \beta^{-}+4 \bar{v}+{ }^{207} \mathrm{~Pb}$
The system also has one stable isotope of lead, ${ }^{204} \mathrm{~Pb}$. This isotope is used as the denominator in the decay equations written below as (Hole, 1998):
${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}={ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}_{\mathrm{i}}+{ }^{238} \mathrm{U} /{ }^{204} \mathrm{~Pb} *\left(\mathrm{e}^{\lambda \mathrm{t}}-1\right)$
${ }^{207} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}={ }^{207} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb} b_{\mathrm{i}}+{ }^{235} \mathrm{U} /{ }^{204} \mathrm{~Pb} *\left(\mathrm{e}^{\text {ג2t }}-1\right)$
where ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ and ${ }^{207} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}$ are the isotope ratios at the present, $i$ denotes the initial lead isotope ratio at the time of sample formation, ${ }^{238} \mathrm{U} /{ }^{204} \mathrm{~Pb}$ and ${ }^{235} \mathrm{U} /{ }^{204} \mathrm{~Pb}$ are the isotope ratios at the present time, $\lambda 1$ and $\lambda 2$ are decay constants, and $t$ is the time elapsed since the closure temperature of the mineral was reached (Hole, 1998). The decay constants used are:
$\lambda 1=1.551 * 10^{-10} \quad$ half life $=4.468 \mathrm{Ga}$
$\lambda 2=9.848 * 10^{-10} \quad$ half life $=704 \mathrm{Ma}$
These independent decay systems provide two independent geochronometers, which should provide two independent ages for a given sample. Ideally, these two ages coincide, which is termed concordant. If you rearrange equations 5 and 6 to:
${ }^{206} \mathrm{~Pb} *{ }^{238} \mathrm{U}=\mathrm{e}^{\lambda 1 \mathrm{t}}-1$
${ }^{207} \mathrm{~Pb} *{ }^{235} \mathrm{U}=\mathrm{e}^{\lambda 2 \mathrm{t}}-1$
where the * denotes the amount of radiogenic lead that has been produced by the sample, for a specified value of $t$, the amount of radiogenic lead of a sample can be calculated. If plotted on a ${ }^{207} \mathrm{~Pb}^{*}{ }^{235} \mathrm{U}$ versus ${ }^{206} \mathrm{~Pb} *{ }^{238} \mathrm{U}$ graph, this would produce a curve known as the concordia (Fig. 2.7). However, because the elements are radioactive, crystal damage to the zircon containing the U and Pb is not uncommon.

Radiation damage can allow lead to be lost from the mineral, as the zircon crystal lattice preferentially incorporates uranium and excludes lead. Lead loss by diffusion during metamorphism or incorporation into igneous rocks can result in different ages for the ${ }^{238} \mathrm{U}$ and ${ }^{235} \mathrm{U}$ systems, which is called discordance (Lee et al., 1997). Several studies have also shown re-crystallization of zircons during metamorphism can lead to redistribution of Pb within the mineral, possibly to zones of lower uranium concentration (Mattinson et al., 1996; Hawkins and Bowring, 1997; Mezger and Krogstad, 1997; Pidgeon and Wilde, 1998; Carson et al., 2002a, b; Romer, 2003). When re-crystallized zones of high Pb are sampled, they plot above the concordia and are said to be reverse discordant.

Zircon $\left(\mathrm{ZrSiO}_{4}\right)$ is one the most common minerals dated using the U-Pb system. Zircon is a stable, tetragonal mineral at the Earth's surface. Its crystal lattice accepts uranium but excludes lead, and it has a closure temperature of $\sim 1,000^{\circ} \mathrm{C}$ (Dahl, 1997; Lee et al., 1997; Hole, 1998; Cherniak and Watson, 2001; Finch and Hanchar, 2003; Schmitz and Bowring, 2003). Zircon is quite common in felsic and intermediate rocks, but its abundance is low in mafic and ultramafic rocks. In this study detrital zircons, those that are not primary but have been weathered from the original rocks in which they formed and reincorporated into younger sedimentary rocks, are dated using the $\mathrm{U}-\mathrm{Pb}$ system to render populations of ages for a given formation. Based on the geologic principle of inclusions, the sedimentary unit containing detrital zircons of interest must be younger than the youngest zircon found within the rock. Therefore, the youngest robust ages of detrital zircons are interpreted to be the maximum depositional age of a given unit (Fedo et al., 2003).


Figure 2.7 - Sample concordia plot (Hole, 1998).

## Chapter 3 - Lesser Himalayan Rocks

### 3.1 Introduction

The Himalaya has been a natural focal point in the study of continent-continent collisional orogenies. Numerous theories and models have been developed to elucidate the processes responsible for the ongoing uplift and thickening of crust in the world's largest mountain range. While grand in scope, explanations ranging from the classic wedge shaped, forward-propagating thrust systems (Davis et al., 1983; DeCelles et al., 2001; Robinson et al., 2006; Kohn, 2008) to more avant-garde theories of crustal extrusion (Beaumont et al., 2001; Jamieson et al., 2004) are undeniably dependent upon accurate geometric and kinematic information to make correct predictions. Conversely, geometric and kinematic information gathered act as tests of the theoretical models.

The most basic and perhaps, therefore, the most important step towards creating these models is the development of an understanding of regional stratigraphy. In the Lesser Himalayan series of central Nepal (Fig. 3.1), the rocks have few fossils and are difficult to date and correlate regionally (DeCelles et al., 2001). Workers have thus conceptualized a variety of stratigraphic columns in order to evaluate field relations (Heim and Gansser, 1939; Pêcher, 1978; Hodges et al., 1996; Upreti, 1996; DeCelles et al., 2001; Martin et al., 2005; Pearson and DeCelles, 2005; Robinson et al., 2006). Through the use of U-Pb isotopic analyses of detrital zircons and time series $\delta^{13} \mathrm{C}$ variations, this study aims to constrain depositional ages for several units in the Lesser Himalayan sequence, and as a result, will inform the regional stratigraphic succession and structural organization.

### 3.2 Geologic Setting

Although the Himalaya offers a spectacular opportunity to study the processes of mountain building, it also presents great challenges. Steep topography, abundant vegetation, and a paucity of datable horizons throughout much of the region are several obstacles to the identification and correlation of stratigraphic units across the range. This is evidenced by the multitude of monikers for individual
formations and the proliferation of stratigraphic columns for the region that are not in agreement. Most problematic are the lack of definitive ages and relationships between units, which necessarily makes it difficult to establish links between units of eastern and central Nepal with that of the better exposed and documented units in western Nepal, and along the rest of the Himalayan arc (DeCelles et al., 2001;

Robinson et al., 2006).


Figure 3.1 -Geologic map of Nepal with study areas outlined in red (Martin et al., 2005). K Kathmandu, P- Pokhara, L- Langtang, IYSZ - Indus Yarling Suture Zone, STDS - South Tibetan Detachment System, MCT- Main Central thrust, RT - Ramgarh thrust, MT- Mahabarat thrust, DT Dadeldhura thrust, MBT - Main Boundary thrust, MFT - Main Frontal thrust

The four main tectonostratigraphic units that constitute the Himalaya are, from south to north, the Subhimalayan (SH), Lesser Himalayan (LH), Greater Himalayan (GH), and Tibetan Himalayan (TH) zones, all of which are divided by faults (Fig. 3.1; Heim and Gansser, 1939). This work focuses solely on the Lesser Himalayan rocks, using the stratigraphic scheme of Upreti (Fig. 3.2a, b; 1996) as a basis for mapping.


Figure 3.2a, b - a) Proposed stratigraphy of Lesser Himalayan rocks by (r) Martin et al. (2005) and (l) Upreti (1996). b) Schematic of stratigraphy used in this study, with approximate thicknesses labeled. Symbols (Kn) Kuncha Formation, (Ks) Kushma Formation, (Fg) Fagfog Formation, (Da) Dandagon Formation, (N) Norpul Formation, (D) Dhading Formation, (B) Benighat Formation, (M) Malekhu Formation, (GW_FB) Gondwana and Foreland Basin units, (MBT) Main Boundary thrust, (MCT) Main Central thrust. The Dandagon and Norpul formations are not recognized in the Modi Khola.

The Lesser Himalayan series is structurally bounded below by the Main Boundary thrust (MBT) and above by the Main Central thrust (MCT) (DeCelles et al., 2001). While there are several faults associated with the Lesser Himalayan series with displacements on the order of tens of kilometers, only the Main Central thrust and the Ramgarh thrust (RT) are considered in this study. The Ramgarh thrust is located within the Lesser Himalayan series, generally placing the oldest Proterozoic rocks against younger LH rocks or Miocene foreland basin deposits (Pearson and DeCelles, 2005). The MCT and RT are thought to accommodate over 140 km and 120 km of slip, respectively, making them the largest factors in regional shortening estimates.

Upreti (1996) designates the basal member of the Nawakot unit of the Lesser Himalayan sequence in central Nepal the Kuncha Formation. However, the base of this formation has not been described or located within Nepal. Largely composed of grey-green phyllites and phyllitic quartzites, the Kuncha Formation extends westward from central Nepal for over 300 km , and may be the lateral
equivalent of the Ranimata Formation of western Nepal (DeCelles et al., 2001; Pearson and DeCelles, 2005). The Kuncha Formation has a thickness of several thousand meters and is intruded by the Ulleri augen gneiss, a granitic intrusion containing zircons that are $1831 \pm 17 \mathrm{Ma}$, providing a minimum age constraint for the formation (DeCelles et al., 2000). Not recognized as a separate unit by Upreti in central Nepal, the Kushma Formation is a several hundred meter thick unit of white to grey, fine to very fine grained, well sorted quartzite (DeCelles et al., 2001). In western Nepal, the Kushma Formation is stratigraphically below the Ranimata (Kuncha?) Formation (DeCelles et al., 2000; Pearson and DeCelles, 2005; Robinson et al., 2006) and the exact nature of the relationship between the Kuncha and Kushma formations in central Nepal, as well as the similarity in names, has caused considerable confusion. Martin et al. (Fig. 3.2; 2005) included the Kushma quartzite in the Kuncha Formation, however, a depositional age for either unit in central Nepal has not been determined and is a focus point of this research.

Moving upward in the section, the Fagfog Formation is an off-white to cream colored quartzite, with a large amount of ripples and trough cross bedding. In western Nepal it has a thickness of $\sim 500 \mathrm{~m}$ and detrital zircons indicate an age younger than 1.68 Ga (DeCelles et al., 2000). The Dandagon Formation (or Galyang in western Nepal) is a green to grey phyllite that lacks any particularly useful diagnostic characteristics. This unit is stratigraphically below the much more distinctive Norpul (or Syangia) Formation. The Norpul Formation consists of pink/white quartzites and stunning reddish purple phyllites and slates. Moving upwards again are the three members of the Lakharpata Group. In stratigraphic order; the ridge-forming, blue-gray Dhading dolomite, the black-gray Benighat slate, and the blue-gray Malekhu limestone. The lower portion of the Dhading Formation is described as a collection of gray slates and stromatolitic limestones, while the upper portion is characterized as more than $80 \%$ stromatolitic dolomite, interspersed with a few quartz sandstones and oolitic dolomites (Upreti, 1996). The Malekhu Formation consists of thinly bedded dolomites and shales, moving upward into thickly bedded dolomite, with the upper portion including intraformational pebble conglomerates and chert nodules as well. Stromatolite structures are not as pervasive in the Malekhu Formation.

Unconformably overlying the Lakharpata Group (and the Nawakot unit) is the Tansen Unit, consisting of the Gondwanan unit and the Bhainskati and Dumri formations. The Carboniferous to Paleocene Gondwanan sequence consists of sandstone, black shale, coal, lignite, and quartz pebble conglomerate (DeCelles et al., 2001). Bryozoa from the genera Fenestrella, Polypora, and Acanthocladia and spores of the genus Vittatiana have been used to ascertain relative ages of this sequence (Upreti, 1996). The Bhainskati and Dumri formations, which make up the Lower Foreland Basin Unit of Martin et al. (2005), consist of Eocene to Miocene black shale, fossiliferous (Nummulites and Assilina) limestone, and sandstone (Upreti, 1996).

The petrography and metamorphic history of the Lesser Himalayan rocks in central Nepal have been studied by numerous workers (Arita 1983; Upreti, 1996; Paudel and Arita, 2000; Catlos et al., 2001; Beyssac et al., 2004; Bollinger et al., 2004; Kohn, 2008). Work by Martin (2005) in the Modi Khola valley indicates a peak equilibrium assemblage of quartz + muscovite + biotite $\pm$ garnet $\pm$ plagioclase feldspar in the Lesser Himalayan series. Exceptions include one sample that lacks muscovite and one which has chlorite, but no biotite. Accessory minerals present in the rocks include allanite, apatite, hematite/magnetite, ilmenite, monazite, pyrite, thorianite, tourmaline, xenotime, and zircon. Therefore, the pelites in the hanging wall of the Ramgarh thrust in central Nepal are garnet zone.

Garnet-biotite and garnet-ilmenite cation exchange thermometry and GMBP barometry by Martin (2005) indicates that the Lesser Himalayan rocks between the Ramgarh thrust and the Main Central thrust underwent a peak temperature of $575^{\circ} \mathrm{C}$ and pressure at the peak temperature of $9 \mathrm{kbar}(900 \mathrm{MPa})$. Work by Beyssac et al. (2004), using Raman spectroscopy of carbonaceous matter, yielded peak temperatures of $330^{\circ}-540^{\circ} \mathrm{C}$ which agree within the uncertainty with the finding of Martin (2005). In the Marsyangdi river valley, several kilometers east of the Modi Khola valley, Catlos et al. (2001) found peak conditions of $450^{\circ}-550^{\circ} \mathrm{C}$ and $6-8 \mathrm{kbar}(600-800 \mathrm{MPa})$ using muscovite ${ }^{40} \mathrm{Ar} /{ }^{39} \mathrm{Ar}$ analyses and thermobarometric analyses of garnet bearing assemblages. Research thus suggests that the Lesser Himalayan rocks in central Nepal experienced metamorphism from greenschist to lower amphibolite facies.

The focus of this research is in the previously mapped Modi Khola valley in the Annapurna range of central Nepal (Fig. 3.1). The area has relatively good access and exposure of Lesser Himalayan rocks, and the structural top of the section has been well defined (Martin et al., 2005). A second study is located in the well-mapped Galcchi Bajar area of Trishuli River Valley, approximately 40 km west of Kathmandu, where the uppermost member (Malekhu Formation) of the Lakharpata Group is exposed (Fig. 3.3; Pearson and DeCelles, 2005). The lower carbonate member of the Lakharpata Group, the Dhading Formation, is poorly exposed in both areas and was thus not sampled.

With the aim of resolving some of the aforementioned stratigraphic ambiguities, dating of detrital zircons in quartzites from the Modi Khola valley and $\delta^{13} \mathrm{C}$ stratigraphy of the Malekhu Formation in Galcchi Bajar are used to constrain the depositional ages of the units of the Lesser Himalayan sequence and subsequently the position of the Ramgarh thrust and other structures. As such, the three hypotheses that guide this research are as follows:

1) In the Lesser Himalayan rocks of the Modi Khola, the Kushma quartzite is older than the Kuncha Formation.
2) In the Lesser Himalayan rocks of the Modi Khola, the quartzite stratigraphically below the carbonates of the Lakharpata Group is the Fagfog quartzite.
3) The Malekhu Formation in the Lakharpata Group north of Galcchi Bajar was likely deposited before $\sim 1250 \mathrm{Ma}$.

### 3.3 Methods

### 3.3.1 Field Mapping

Field mapping and sampling in the Modi Khola valley was conducted during May and June 2007. Carbonate samples were also collected in 2007 from the Trishuli River valley north of the town of Galcchi Bajar (Fig. 3.3). Detrital zircon samples were collected from type localities in central Nepal by Dr. Martin during the 2006 field season (Fig. 3.4). All field stations were marked using the Global Positioning System and recorded. Thin section samples were oriented using a Brunton compass. Samples were then stored individually to prevent cross-contamination.


Figure 3.3 - Geologic map and cross section of the Trishuli river valley north of Galcchi Bajar (Pearson and DeCelles, 2005).


Figure 3.4 - Map of type localities in central Nepal where samples were taken during the 2006 and 2007 field seasons. The unknown sample is located within the Modi Khola valley, which is one of two study areas in this research. Inset map modified from Figure 3.1. Base map courtesy of Lisa Walsh (2008).

### 3.3.2 Malekhu Formation

Previously mapped by Pearson and DeCelles (Fig. 3.3; 2005), the Trishuli River north of Galcchi Bajar has relatively good rock exposure and is easily accessible, providing an ideal location for measuring and sampling the Malekhu Formation. Carbonate samples were collected at 3 m intervals, with additional samples taken at notable lithologic changes, to prevent sampling bias for specific rock types or exposures. Total section thickness measured was 443 m , including covered intervals at the base and top of the section.

In the lab, the samples were cut using an MK tile saw, then ground and polished with a Struers Labopol-21 two-wheel grinding apparatus. Drilling sites were determined by identifying homogeneous areas within the rocks that were not located close to veins or weathered surfaces. Analytical procedures followed are described in Kaufman et al. (1991) and Kaufman and Knoll (1995). Once drilled, $100 \mu \mathrm{~g}$ of material were acidified with $102 \%$ phosphoric acid $\left(\mathrm{H}_{3} \mathrm{PO}_{4}\right)$ for 10 minutes at $90^{\circ} \mathrm{C}$ under vacuum. The resulting $\mathrm{CO}_{2}$ samples were analyzed using the GV IsoPrime dual inlet gas source mass spectrometer at the University of Maryland. Eight to ten NBS-19 standards were analyzed per run and all data were subsequently reported relative to V-PDB. Uncertainties for both carbon and oxygen isotopes are less than 0.05\%。.

### 3.3.3 Zircons

Samples of quartzite from the type localities in central Nepal were collected by Dr. Aaron Martin during the 2006 field season. Additional samples were collected from the Modi Khola valley during the 2007 field season by Dr. Martin, as well. Detrital zircon samples include type Kuncha (Sample \#506078), Kushma (Sample \#406102), Fagfog (Sample \#506079), and Norpul/Syangia (Sample \#406154) formations, as well as one sample taken from an unidentified unit in the Modi Khola (Sample \#406020). One igneous sample was also collected, the Ulleri augen gneiss (Sample \#507056).

Samples were crushed by hand using a stainless steel mortar and pestle, and sifted through 2 mm and 0.25 mm sieves. The resulting material was washed and panned by hand and allowed to dry overnight. Magnetic separation was accomplished using the Frantz LB-1 magnetic barrier separator, set with a $17.5^{\circ}$ front angle and a $20^{\circ}$ side angle. Samples were run through the Frantz at $0.5,1,1.5$, and 2.25 (maximum) amps. Due to machine variance, maximum amp settings were as following; for sample number 406154 (2.25A), 406020 (1.97A), 507056 and 506079 (1.85A), and 406102 (1.5A). Dense liquid separation was accomplished using methylene iodide (MEI) with a density of $3.325 \mathrm{~g} / \mathrm{ml}$. The samples were then hand-picked using a stereomicroscope with both transmitted and reflected light. Sample 406154, the Norpul (Syangia) type locality, did not yield any zircons and will therefore not be further evaluated. During picking, a random selection of zircon grains were chosen to prevent bias in the zircon populations of each sample based on grain morphology. With the addition of zircon standard SL-2, the zircons were mounted, polished, and taken to the University of Maryland Electron Probe Microanalyzer Laboratory. Backscatter electron and cathodoluminescent images were collected using the JEOL JXA8900 SuperProbe (Figure 3.5a, b).

Figure 3.5a, b - Backscatter electron images (left) and cathodoluminescent images (right) of zircons from the type Kuncha Formation (Sample 506078). Red arrows indicate candidate spots for laser ablation and the red circle encompasses an inclusion which was avoided when firing the laser.

$\mathrm{U}-\mathrm{Pb}$ isotopic data were collected in-situ from the laser ablation multicollector inductively coupled plasma mass spectrometer (LA-MC-ICPMS) at the University of Arizona LaserChron Center. The LA-ICPMS uses a New Wave/Lambda Physik DUV 193 Excimer laser, with a 193 nm wavelength. The laser is set with an 8 Hz repetition rate and a fluence of $\sim 4 \mathrm{~J} / \mathrm{cm}^{2}$. Spot diameter for detrital samples was $30 \mu \mathrm{~m}$, and for the igneous sample a spot size of $20 \mu \mathrm{~m}$ was used to allow for the collection of rim and core data from individual grains. Depth of the ablation pit is $\sim 12-15 \mu \mathrm{~m}$. Care was taken to avoid hitting inclusions or multiple zones within a single zircon with the laser (Fig. 3.5a, b). Ablated material was carried in a stream of helium to the plasma source and $\mathrm{U}, \mathrm{Th}$, and Pb were measured simultaneously in static mode on 10 E 11 ohm Faraday detectors for ${ }^{238} \mathrm{U},{ }^{232} \mathrm{Th},{ }^{208} \mathrm{~Pb}$, and ${ }^{206} \mathrm{~Pb}$, a 10 E 12 ohm Faraday collector for ${ }^{207} \mathrm{~Pb}$, and an ion counting channel for ${ }^{204} \mathrm{~Pb}$. Examples of systematic uranium and lead changes during laser ablation can be seen in Figure 3.6a-c.




Figure 3.6a-c - Graphs of (a) ${ }^{206} \mathrm{~Pb}$ intensity, (b) ${ }^{238} \mathrm{U}$ intensity, and (c) ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ intensity changes during laser ablation. Data points are integrated over 1 second intervals.

Reduction of detrital zircon analyses was accomplished using the procedures of Gehrels et al. (2006, 2008). Three corrections were made prior to age calculation using an Excel spreadsheet supplied by the University of Arizona LaserChron Center. Depth-related fractionation was accommodated by discarding the first three seconds of data, which removes the early, typically rapidly fluctuating signal (Figure 3.7), and then extracting the nine seconds of remaining analyses as nine 1 -second integrations which yield isotope ratios from the integrated intensities. A least squares regression through the remaining data to the initial ratio (fourth second of signal acquisition) was made to account for ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ and ${ }^{208} \mathrm{~Pb} /{ }^{232} \mathrm{Th}$ depth dependent changes. Correction for inter-element fractionation of ${ }^{206} \mathrm{~Pb} /^{238} \mathrm{U}$ and fractionation of ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ was based on in-run analyses of standard SL-2 (Sri Lanka), which has an IDTIMS age of $564 \pm 4 \mathrm{Ma}(2 \sigma)$ (Gehrels et al., 2008). Inter-element fractionation of $\mathrm{Pb} / \mathrm{U}$ is typically $20 \%$ and fractionation between Pb isotopes is $\sim 2 \%$. Uncertainties on the calibration corrections are 1-2\% (2б). The standard was analyzed after every five to seven unknowns. Unknowns were corrected using a sliding window average of the six closest standards, excluding the maximum and minimum fractionation factor values.

Common lead correction was made by measuring ${ }^{204} \mathrm{~Pb}$ and assuming an initial lead composition based on the work of Stacey and Kramers (1975). Background ${ }^{204} \mathrm{Hg}$ and ${ }^{204} \mathrm{~Pb}$ were measured on peaks. Uranium and thorium concentrations were also determined by comparison to the standard zircon (SL-2)
which has concentrations of 518 ppm and 68 ppm , respectively, calculating the average intensity of the ${ }^{238} \mathrm{U}$ and ${ }^{232} \mathrm{Th}$ for standards used between samples and adjusting the unknowns by that factor according to their intensities. To remove detrital zircon ages without crystallization significance, e.g. an age resulting from an analyzed zircon with demonstrable lead loss, analyses with greater than $10 \%{ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ error, $5 \%$ ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ error, $25 \%$ discordance, or $5 \%$ reverse discordance were discarded. For zircon grains older than $1,000 \mathrm{Ma}$, interpreted ages are based on ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$, as opposed to ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$, with an uncertainty of $1-2 \%(2 \sigma)$.

Although the formation of "new" metamorphic zircon is rare at temperatures and pressures less than upper amphibolite or granulite facies, recrystallization of inherited zircons and growth of metamorphic rims is possible (Hoskin and Schaltegger, 2003). Therefore, in order to avoid including analyses of detrital zircons that do not reflect igneous crystallization ages, samples were evaluated for the presence of metamorphic zircon in several ways. Backscatter electron and cathodoluminescent images, which have been demonstrated to be effective methods for revealing internal crystal structure (Hanchar and Miller, 1993; Hanchar and Rudnick, 1995), were taken of all samples. Locations of cores and rims were noted for better placement of the laser during ablation, and only cores were sampled. Oscillatory zoning, which is a common characteristic of igneous zircons (Zhao et al., 2002; Hoskin and Schaltegger, 2003), was also noted in many of the zircons. Several researchers have noted that metamorphic zircons typically have higher U concentrations and higher $\mathrm{U} / \mathrm{Th}$ ratios than igneous zircons (Belousova et al., 2002; Rubatto, 2002; Xian et al., 2004; Dziggel et al., 2005). Thus for each sample, outlying analyses with relatively high $U$ concentrations or $U / T h$ ratios compared to the sample population as a whole were discarded. No more than 5 analyses were excluded from any given sample based on the $U$ concentration or U/Th ratio. Lastly, because only clusters of three or more analyses with overlapping ages are interpreted to have significance, the likelihood of any $\mathrm{n}>3$ remaining metamorphic zircons yielding similar ages is significantly diminished.

The number of analyses discarded varied by sample. For the Kuncha type (Sample 506078) there were 39 analyses discarded out of 166 or $23 \%$ of the total analyses. From the Kushma type (Sample 406102), 53 analyses out of 178 analyses or $30 \%$ were of the total analyses were discarded. In the Fagfog type (Sample 506079) 35 analyses out of 211 analyses or $17 \%$ of the total analyses were discarded. From the unidentified quartzite in the Modi Khola (Sample 406020), 19 of 53 analyses were discarded or 36\% of the total analyses.


Figure 3.7 - Example of signal from ablation of a zircon using the LA-ICP-MS at the University of Arizona LaserChron Center. Laser is fired for 12 seconds, with first 3 seconds of data discarded based on rapidly fluctuating signal. Figure from Gehrels et al., 2008.

### 3.3.4 Justification of Detrital Zircon Method

Although accurate and precise analyses are important, the use of detrital zircons to determine a maximum depositional age for a given geologic unit is somewhat less dependent on the precision of any one individual analysis, and more focused on the number of analyses and the care with which individual zircons are selected (Anderson, 2005). Link et al. (2005) recognized detrital zircons could create a unique "barcode" of age spectra that could be used to describe a geologic unit or terrane. However, researchers need to be aware of several biases that could influence the representativeness of a given sample and take care to avoid them when possible. The abundance of (or lack of) zircon in source rocks will ultimately lead to variability in the amount of zircon produced from a given rock. Zircon is found in sedimentary, igneous, and metamorphic rocks, but its abundance is much greater in rocks with felsic or intermediate compositions than those that are mafic or ultramafic in nature (Yamashita et al., 2000, Belousova et al., 2002).

Furthermore, zircons formed in mafic rocks typically have low U concentrations, which when sampled by laser ablation can lead to a large error on the ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ measurement (Rubatto, 2002; Hoskin and Schaltegger, 2003). The high error can result in the exclusion of those zircons and an introduction of bias into the sample. Understanding the source rocks of the detrital zircon sample and the relative quantity of mafic or ultramafic rocks that might have contributed to the sample is essential. Research on the Lesser Himalayan series indicates that the rocks consist of sediment weathered from the Indian craton and deposited on the passive margin of northern India (Gansser, 1964; DeCelles et al., 1998; Sharma, 1998; Hodges, 2000; Bollinger et al., 2004; Robinson et al., 2001, Myrow et al., 2003; Najman, 2006). The Indian subcontinent consists of six Precambrian terrains. The terrains include the Dharwar, Bastar, and Singhbhum cratons, the Southern Granulite Terrain, the Eastern Ghat Mobile Belt, and the Aravalli-Delhi Mobile Belt (including the Bundelkhand craton). The major lithologies consist of sandstones, limestones, shales, quartzites, felsic and mafic metavolcanics, granites, and gneisses along with Archean aged pillow lavas, amphibolites, and ultramafics (Divakara Rao et al., 1998; Sharma, 1998; Mishra et al., 2000;

Menon et al., 2003; Myrow et al., 2003; Leelanandam et al., 2006; Pati et al., 2007; Mall et al., 2008). Importantly, the amount of mafic or ultramafic rock of mid-Paleoproterozoic age or older in the Indian shield appears to be quite small. Therefore, it is inferred that the potential contribution of zircon from mafic or ultramafic rocks is likely restricted and the detrital samples in this study should not be unduly biased.

Surficial processes can also create natural biases in the detrital zircon record. Weathering has been shown to more easily destroy zircons or portions of zircon grains that are older, have higher U concentrations, or are metamict (Fedo et al., 2003). Distance traveled from the source will affect the concentration of zircon that is incorporated into downstream sedimentary rocks. Upriver sources may be masked by downstream inputs, especially if the lower reaches of the rivers transporting material are incised and contribute zircons from underlying units (Cawood et al., 2003; Link et al, 2005; Moecher and Samson, 2006).

While researchers have little to no control over the natural processes that bias detrital zircon populations, there are introduced biases in sampling that can be avoided. Larson and Poldervaart (1957) demonstrated that zircon breaking during crushing does not appear to be significant. However, a common grain size fraction (ex. $<0.25 \mathrm{~mm}$ ) should be used in order to avoid a grain size bias between samples (Morton et al., 1996). There is a lack of evidence to support any biases introduced by water separation on the Wilfley table or in this research, hand panning, or from the dense liquid separation of zircons (Fedo et al., 2003). Magnetic separation has the potential to cause bias in detrital zircon samples because of the positive correlations between Pb loss, U content, discordance, and magnetic susceptibility (Silver, 1963; Sircombe and Stern, 2002). It is noted that the most paramagnetic fractions of detrital zircons are typically the most unreliable (discordant) and it is suggested that a Frantz setting of 1.8 A with a $10^{\circ}$ side angle should be an acceptable compromise between a representative sampling and analytical reliability. However, this is perhaps open to some interpretation (Sircombe and Stern, 2002).

Hand picking of the detrital zircon population has the potential to introduce a large bias into the final age population of a sample and has thus been a focus of several studies. Several researchers argue for the importance of a random sampling of grains (Dodson, 1988; Morton et al., 1996; Moecher and Samson, 2006), while others suggest picking non-random samples based on grain morphology (Ross and Villenuve, 2003), or both random and non-random samples to compare results (Anderson, 2005). Link and Fanning (2003) sampled from the the same units as Ross and Villenuve (2003) and found the same populations of zircons using a random picking method versus the previous author's selective method. They therefore maintain that picking grains based on a specific morphology is unnecessary. These arguments about random and non-random picking lead to the question: How many zircons need to be picked and analyzed to yield a statistically robust answer?

Dodson (1988) determined a sample of $\mathrm{n}=60$ randomly chosen detrital zircons should detect an age population as low as $5 \%$ of a given sample at a $95 \%$ confidence interval. This assumes that any zircon in the sample has an equal opportunity of being picked. Link and Fanning (2003) found in sample sizes $n>60$, the smallest age populations became better defined but no new populations were detected. They also noted that one grain analysis does not constitute a population, but that a multi-grain population ( $\mathrm{n} \geq$ 3) was important to demonstrate the presence of any specific age group. Several other researchers use sample sizes ranging from $\mathrm{n}=35$ to $\mathrm{n}=75$ grains based on the work of Dodson (Morton et al., 1996; Link and Fanning, 2003; Moecher and Samson, 2006).

However, Vermeesch (2004) posits that Dodson's work does not sufficiently account for a "worst case" scenario in which there is an equal abundance of all age populations. If this were the case a sample size of $\mathrm{n}=117$ would be necessary to ensure at the $95 \%$ confidence interval that an age population of $5 \%$ is not missed. If research suggests that there is not an equal abundance of all age populations, Anderson (2005) suggests the ideal number of analyses needed is likely somewhere between Dodson's $\mathrm{n}=60$ and Vermeesch's $\mathrm{n}=117$. In this study, three of the four detrital zircon samples analyzed have $\mathrm{n}>117$. The remaining sample has $\mathrm{n}=34$, after much of the sample was lost during final polishing.

### 3.4 Results

### 3.4.1 Mapping

Work in the Modi Khola valley (Fig. 3.8) reveals the Kushma Formation is exposed in the southernmost and topographically lowest portion of the map area (Fig. 3.9). Outcrop near the town of Birethati (Fig. 3.10) demonstrates that the Kushma quartzite lies above the phyllitic Kuncha Formation. Moving topographically higher and northward, the Kuncha Formation is again exposed, but this time above the Kushma. The Kuncha Formation is intruded by the Ulleri augen gneiss in the valley. The repetitive sequence of Kushma and Kuncha formations is followed by exposure of another clean quartzite, also identified as the Kushma Formation.

Structurally above the quartzite are, in order, the Dhading, Benighat, and Malekhu formations of the Lakharpata Group, followed by the Gondwana and Foreland Basin sediments. Thin section photographs of the units sampled from type localities in central Nepal and from the Modi Khola can be seen in Figures 3.11a-j. These are some of the first pictures of type locality thin section to be illustrated and are included here for reference. Petrographic descriptions of the type location detrital zircon samples can be seen in Table 1. The Main Central thrust represents the structural termination of the Lesser Himalayan series and is located at the northernmost portion of the map area. Measured foliations and bedding in the Modi Khola are predominantly north (hinterland) dipping.


Figure 3.8 - View of the Modi
Khola valley looking south from Ghandruk.



Figure 3.10 - Birethati exposure with quartzitic Kushma Formation $\left(\mathrm{pC}_{\mathrm{ks}}\right)$ above phyllitic Kuncha Formation ( $\mathrm{pC}_{\mathrm{kn}}$ ). Red dashed line denotes contact. Outcrop location in the Modi Khola valley marked by black star in Figure 3.9.

### 3.4.2 Malekhu Formation

Results from the analysis of the Malekhu Formation from the Trishuli River valley can be seen in Figure 3.12 and Table 2.The $\delta^{13} \mathrm{C}$ values of the section range from -1.7 to $+0.2 \%$, with a mean value of $0.9 \% \pm 0.4 \%$ ( 1 s.d.), while $\delta^{18} \mathrm{O}$ values range from -17 to $-9 \%$, with a mean value of $-12.4 \% \pm 1.5$ $\%$. A remarkably long stratigraphic interval with limited variation about the mean can be seen in the carbon isotope data throughout the entire section, with the exceptions of a slight negative trend noted from approximately 133 m to 147 m and a somewhat larger variation is seen in the data above 387.5 m . Otherwise, there is remarkable consistency in the data over the entire 309.5 m section and across a wide variety of lithologies, including limestone, dolomite, and calcareous shale (Fig. 3.13a-d). The data for $\delta^{18} \mathrm{O}$ reveal slightly more variability, but no significant excursions are seen. Thin sections from the Trishuli River valley reveal recrystallization and a lack of primary structures, as well as an increasing siliciclastic input near the top of the section.

A sample of the Malekhu Formation (Sample \#07033) located in the Modi Khola valley (Fig. 3.9) reveals a $\delta^{13} \mathrm{C}$ value of $-0.3 \%$ and a $\delta^{18} \mathrm{O}$ value of $-12 \%$, which is similar to findings in the Trishuli River valley. Thin sections of the sample (Fig. $3.11 \mathrm{i}, \mathrm{j}$ ) demonstrate that this unit is predominantly made up of carbonate minerals and is not a sandstone as suggested by previous workers (Pearson and DeCelles, 2005). Of note, the thin sections from samples in the Modi Khola valley also demonstrate a consistently larger grain size than the thin sections from the Trishuli River locality. This is expected since the Malekhu Formation in the Modi Khola valley is located within the Ramgarh thrust sheet, and experienced higher temperatures and pressures than in the Trishuli River valley.

Additionally, several samples dispersed throughout the section were drilled in multiple locations to ascertain the variability of carbon and oxygen isotope measurements on the millimeter scale. The results of this investigation demonstrate very little variability of $\delta^{13} \mathrm{C}$ and $\delta^{18} \mathrm{O}$ values on this scale, although there are some minor differences beyond analytical uncertainty (Fig. 3.14). The limited variation in carbon and oxygen isotopic values at the millimeter scale is also reflected in the larger, meter scale findings which demonstrate a similar tendency.


Figure 3.11a-j (and on next page). Thin sections of type localities in central Nepal and samples from the Modi Khola. Pictures (a-h) taken in cross-polar light, scale the same as in (b). (a) Kuncha Formation type locality - grey-green phyllite and phyllitic quartzite, (b) Kushma Formation type locality - coarse grained, well sorted quartzite, (c) Ulleri augen gneiss type locality, (d) Fagfog Formation type locality medium grain, well sorted quartzite, (e) Dhading Formation - carbonate- from Modi Khola, (f) Benighat Formation - slate- from Modi Khola ,(g) Malekhu Formation - carbonate-from type locality, (h) diorite mapped in the Modi Khola, ( $\mathrm{i}, \mathrm{j}$ ) Scale as seen in ( j ). Cross-polar and plane polar light respectively. Malekhu Formation from Modi Khola, which was interpreted as a sandstone by previous workers. Symbols (q) quartz, (bt) biotite, (ms) muscovite, (amp) amphibole, (ca) calcite, (pl) plagioclase feldspar.


Figure 3.12 (next page) - Stratigraphic column with carbon and oxygen isotopic data for the Malekhu Limestone. Interval between base of section and 133.5 m is covered, column starts at 133.5 m . All values are referenced to V-PDB. Errors less than $0.05 \%$ and contained within symbols. (ca) calcite, (dol) dolomite.



Figure 3.13a-d. - Thin sections of carbonate lithofacies present in the Malekhu Formation in the Trishuli River valley. Photos taken in cross-polar light, field of view 2 mm . (a) Crystalline carbonate - calcite with quartz vein, (b) Crystalline carbonate - dolomite with re-crystallized calcite in vein (c) Calcareous shale with quartz vein, (d) Micaceous calcareous shale.


Figure 3.14 - Carbon and oxygen cross-plot of selected samples from the stratigraphic column in Fig. 3.12 demonstrate limited variablitiy of $\delta^{13} \mathrm{C}$ and $\delta^{18} \mathrm{O}$ values with samples on a millimeter scale. Errors contained within symbols.

### 3.4.3 Interpretation of Detrital Zircon Chronology

In this study detrital zircon data are presented with a combination of histograms and probability density plots. The histograms plot the number or frequency of analyses or samples on the y -axis and ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ ages of samples in binned groups along the x -axis. Overlying the histogram is a probability density plot. This plot represents the ages and frequencies, as well as associated uncertainties (Vermeesch, 2004). The probability density plot is constructed by calculating a normal distribution for each analysis based on its measured age and uncertainty, then summing the probability distributions into a single curve (Gehrels et al., 2006).

Based on the premise that a geologic unit is younger than the youngest detrital zircon contained within it (law of inclusions); the youngest portion of the detrital zircon sample is of greatest interest for interpreting maximum depositional ages. While the entire age spectra of an analysis may be useful for studies of provenance, no interpretations are made regarding these ages. All interpretations focus on the youngest zircons within a given sample.

Although Vermeesch (2004) suggests that a single age analysis can be used to determine the presence of an age population, other researchers prefer to interpret only clusters of three or more analyses with overlapping ages as defining robust age populations (DeCelles et al., 2000, 2004; Gehrels et al., 2006, 2008). A single analysis is more likely to have been compromised in some way (e.g. Pb loss) and yield an erroneous age, than three or more analyses with ages that overlap. Therefore, this study does not interpret clusters of ages with less than three analyses. Furthermore, a range of ages are given for each sample and the size of the analyzed cluster is also indicated. From the range of ages, a conservative maximum depositional age, as well as a less conservative, but still robust maximum depositional age is determined for each unit.

### 3.4.4 Zircons

Reduction of detrital zircon analyses yield several distinct maximum depositional ages for units in the Modi Khola. For zircon grains older than 1,000 Ma, interpreted ages are based on ${ }^{207} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$, as opposed to ${ }^{207} \mathrm{~Pb} /{ }^{235} \mathrm{U}$, with an uncertainty of $1-2 \%(2 \sigma)($ Table 3$)$. Zircons ( $\mathrm{n}=127$ ) from the type Kuncha Formation have a youngest peak on the probability density plot (Fig. 3.15a, b) of 1896 Ma, however Table 2 shows a cluster of 10 analyses with ages ranging from 1871 to 1881Ma. Therefore a conservative maximum depositional age for the Kuncha Formation is $\sim 1896$ Ma. A slightly less conservative, but still robust age for the Kuncha Formation is $\sim 1875 \mathrm{Ma}$. The unit's depositional age is further constrained by the $1831 \mathrm{Ma} \pm 17 \mathrm{Ma}$ age of the Ulleri augen gneiss, which intrudes the Kuncha Formation and thus constrains the youngest possible age of deposition (DeCelles et al., 2000). Sample (507056), the Ulleri augen gneiss collected in 2007, did not yield useable ages for either cores or rims. As such, it was not included in this analysis and the published age of $1831 \mathrm{Ma} \pm 17 \mathrm{Ma}$ is used instead (DeCelles et al., 2000).

The maximum depositional age indicated by the youngest peak on the probability density plot for zircons ( $\mathrm{n}=125$ ) in the type Kushma Formation is 1795 Ma (Fig. 3.16a, b). Nine analyses ranging from 1771-1785 Ma suggest a less conservative, but still robust maximum depositional age of $\sim 1780 \mathrm{Ma}$ for the Kushma Formation. The maximum age of deposition peak on the probability density plot for the type Fagfog Formation ( $\mathrm{n}=176$ ) is $\sim 1800 \mathrm{Ma}$ (Fig. 3.17a, b). Seven analyses with ages ranging from 1791 to 1810 Ma also suggest a maximum depositional age of 1800 Ma for this unit. The unknown sample ( $\mathrm{n}=34$ ) taken from the Modi Khola indicates a likely maximum depositional age of $\sim 1900$ Ma, with the youngest reliable zircon population ranging between 1895 and 1908 Ma , although the sample size is much smaller than the preferred $n>60$, as the majority of the sample was lost during final polishing (Fig. 3.18a, b).

### 3.5 Implications

### 3.5.1 Constraints on depositional ages

Broadly consistent with the findings of DeCelles et al. (2000), the detrital zircon analyses of this study have several interesting implications for the depositional ages of Lesser Himalayan units in central Nepal. First, the $\sim 1875$ Ma maximum depositional age of the Kuncha Formation coupled with the $1831 \pm$ 17 Ma age of the intruded and cross-cutting Ulleri augen gneiss, provide upper and lower depositional


Figure 3.15a -b. - Relative probability plot of age distributions and concordia diagram of detrital zircon ages in the type Kuncha Formation (506078). Inferred depositional age is $\sim 1875 \mathrm{Ma}$.


Figure 3.16a -b. - Relative probability plot of age distributions and concordia diagram of detrital zircon ages in the type Kushma Formation (406102). Inferred depositional age is $\sim 1780 \mathrm{Ma}$.


Figure 3.17a -b. - Relative probability plot of age distributions and concordia diagram of detrital zircon ages in the type Fagfog Formation (506079). Inferred depositional age is $\sim 1800 \mathrm{Ma}$.



Figure 3.18a-b. - Relative probability plot of age distributions and concordia diagram of detrital zircon ages in the unknown formation in the Modi Khola (406020). Inferred depositional age is $\sim 1900 \mathrm{Ma}$.
bounds for the unit (DeCelles et al., 2000). These data support the interpretation that the Kuncha Formation is older than the Kushma and Fagfog formations, which have maximum depositional ages that are younger than the crystallization age of the Ulleri gneiss. This finding does not support my earlier hypothesis (1), in which the Kushma Formation was expected to be older than the Kuncha Formation.

Second, the $\sim 1800$ Ma maximum age of deposition for the type Fagfog Formation allows the unit to have been deposited before the type Kushma Formation, but it cannot completely rule out that the unit may actually have been deposited after the deposition of the Kushma Formation. Due to the restricted sample size, the zircon population from the quartzite (406020) stratigraphically below the Lakharpata Group in the Modi Khola does not match exactly any of the type localities. However, there are similarities between the histogram profiles of the unknown and the Fagfog type locality, as well as that of the Kushma type locality. If the northernmost quartzite is the Kushma Formation, then a depositional contact between it and the underlying Kuncha Formation would be inferred. In contrast, if the northernmost quartzite is the Fagfog Formation a fault contact is required. The simplest and therefore the preferred explanation is that the northernmost quartzite is actually a repetition of the Kushma Formation and is mapped as such. Since the data cannot distinguish between the Fagfog and Kushma formations, I am unable to falsify hypothesis (2) that the northernmost quartzite is the Fagfog Formation.

Third, the trend of near $0 \% \delta^{13} \mathrm{C}$ values throughout the Malekhu section in Trishuli River valley is consistent with hypothesis (3) of a Mesoproterozoic (>1250 Ma) age of deposition. Unfortunately, the data cannot differentiate between an early Mesoproterozoic depositional age and that of a younger, relatively quiescent interval between the larger $\delta^{13} \mathrm{C}$ shifts of the later Mesoproterozoic or Neoproterozoic. However, in practical terms, there appears to be no evidence for carbonates from a thick, Neoproterozoic aged succession with near $0 \%{ }_{0} \delta^{13} \mathrm{C}$ values continuously throughout the section. Buick et
al. (1995) found a comparably limited range of $\delta{ }^{13} \mathrm{C}$ values in the Mesoproterozoic Bangemall Group in Australia and noted that the "flat" carbon values also existed over an approximately 2500 m thickness of strata. Knoll et al. (1995) examined Proterozoic carbonates of the 2200 m thick Anbar Massif in northwestern Siberia which span the Meso-Neoproterozoic boundary. They found $\delta^{13} \mathrm{C}$ values of 0 to $1.9 \%$ in the older units ( $1600-1200 \mathrm{Ma}$ ), and slightly greater variation, -2.7 to $4.6 \%$, in the younger units (1200-850 Ma). Spanning a similar time rage, Kumar et al. (2002) investigated the 4300 m thick Proterozoic Vindhyan Basin in central India, in which the older, Mesoproterozoic, unit ( $\sim 2100 \mathrm{~m}$ thick) also shows $\delta^{13} \mathrm{C}$ values of $0 \%$ and the younger Neoproterozoic unit values show a range from -7.5 to $2 \%$. Kaufman et al. (2006) found significantly more negative and positive $\delta^{13} \mathrm{C}$ values in the $\sim 1400 \mathrm{~m}$ thick Neoproterozoic (Ediacaran Period) Krol platform in the Lesser Himalaya of northern India, as did Tewari and Sial (2007). These studies lead to a preference for an early ( $>1250 \mathrm{Ma}$ ) Mesoproterozoic age of deposition for the Malekhu Formation (Fig. 3.19).

The $\delta^{13} \mathrm{C}$ age of the Malekhu Formation is significant because it confirms that the Lakharpata Group of central Nepal is not correlative with other carbonate successions in the Lesser Himalayan physiographic province in India, e.g., the Neoproterozoic Infra Krol Formation and the Krol group (Aharon et al., 1987; Kaufman et al., 2006; Tewari and Sial, 2007).


Figure 3.19- $\delta^{13} \mathrm{C}$ curve for Meso-and Neo-Proterozoic rocks (Kah et al., 1999).

Post-depositional diagenesis has been shown to cause $\delta^{18} \mathrm{O}$ values to become more negative through aqueous fluid interactions, coupled with temperature increases after burial, while leaving $\delta^{13} \mathrm{C}$ values relatively unchanged (Tucker, 1983; Banner and Hanson, 1990; Wickham and Peters, 1993; Jacobsen and Kaufman, 1999; Kah et al., 1999; Kaufman et al., 2006). The wide range of negative $\delta^{18} \mathrm{O}$ values suggests the influence of an isotopically depleted metamorphic or meteoric fluid interaction with the rocks (Jacobsen and Kaufman, 1999; Baumgartner and Valley, 2003). The variation in oxygen values appears to be unrelated to or uncoupled from the carbon isotopes which remain unchanging throughout the section. The significant amount of veining in parts of the measured section, which can be seen in thin section (Fig. 3.13a-d), is physical evidence consistent with fluids moving through the system.

### 3.5.2 Stratigraphy in the Modi Khola

Based on depositional age constraints of the detrital zircons, a reorganization of the stratigraphy of central Nepal is necessary. Detrital zircons and Ulleri augen gneiss magmatic zircons constrain the Kuncha Formation as older than the Kushma, therefore the Kuncha Formation should be considered the basal unit of the Lesser Himalayan package. This is supported by field evidence in the geologic exposure south of the town of Birethati (Fig. 3.10), which clearly demonstrates the relationship of Kushma Formation overlying Kuncha Formation, as well as the larger map scale contact between the units. Furthermore, it has been noted by previous workers that exposures of the Ulleri augen gneiss exclusively intrude the Kuncha Formation, but never the Kushma Formation or younger units. If the Kushma Formation were the older unit, evidence of the Ulleri augen gneiss intruding the Kushma would be expected.

While mapping, no evidence was found to indicate the existence of the Dandagon or Syangia formations in the Modi Khola. The northernmost quartzite in the Modi Khola is likely the Kushma Formation, based on detrital zircon analyses and evaluation of field relationships. Based on this identification and from the absence of the Dandagon and Syangia formations, a significant amount of the

Lesser Himalayan series is missing/ omitted from the region, either through normal faulting or a stratigraphic hiatus. A normal fault is inferred based on the presence of the omitted units in nearby regions.

The package of rocks topographically and stratigraphically above the northernmost quartzite are likely the Dhading, Benighat, and Malekhu formations (the Lakharpata Group as recognized in western Nepal). This interpretation is supported by thin section petrography, which confirm the carbonate-shalecarbonate lithologies of the units (Fig. 3.11a-j). Furthermore, thin sections of the Malekhu Formation from the Modi Khola demonstrate a larger grain size than that of the measured section in the Trishuli River valley. This finding suggests that the rocks have experienced more recrystallization in the Modi Khola valley, which is consistent with its metamorphic history. However, the similarity in $\delta^{13} \mathrm{C}$ and $\delta^{18} \mathrm{O}$ values between the Modi Khola and Trishuli river valleys suggests limited diagenetic differences between the two.

Topographically above the Malekhu Formation, the Gondwana/Foreland Basin units are exposed. The Main Central thrust is structurally above the Gondwana and Foreland Basin and represents the top of the Lesser Himalayan series.

### 3.5.3 Structural Significance

The presence of a thrust fault is inferred from the repetition of Kuncha and Kushma units in the Modi Khola. Since the Ramgarh thrust is defined as the first major thrust fault below the MCT, the northernmost exposure of the Kuncha Formation marks the location of the Ramgarh thrust. Based on this interpretation, the location of the Ramgarh thrust is $\geq 10 \mathrm{~km}$ further south than indicated by previous mapping (Fig.3.9 and 3.20; Martin et al., 2005; Pearson and DeCelles, 2005). DeCelles et al. (2001) note that the Ramgarh thrust sheet, where currently located, is unusually thin for such a regionally extensive feature. This study's findings imply that the Ramgarh thrust does in fact carry a thick package of rock similar to, but thinner, than that of the Main Central thrust, and supports the assertion of Pearson and

DeCelles (2005) that the thrust is a kinematically important structure in the Himalayan fold and thrust belt. Furthermore, this interpretation reduces the number of thrust faults mapped below the MCT within the Modi Khola from four to one (Fig. 3.9), and would thus reduce the estimates of shortening accommodated in the region.

### 3.6 Conclusions

Field observations, $\delta^{13} \mathrm{C}$ analyses, and $\mathrm{U}-\mathrm{Pb}$ detrital zircon ages support the following conclusions in the central Nepal Himalaya:
(1) The Kuncha Formation is older than the Kushma Formation in central Nepal. Detrital zircons indicate a conservative maximum depositional age for the Kuncha Formation of $\sim 1896 \mathrm{Ma}$. A slightly less conservative, but still robust age for the Kuncha Formation is $\sim 1875$ Ma. The Ulleri augen gneiss, which intrudes the Kuncha Formation, is dated at $1831 \pm 17$ Ma age and constrains the youngest possible age of deposition for the unit. Detrital zircons from the Kushma Formation, which is not intruded by the Ulleri augen gneiss, indicate a maximum depositional age of the $\sim 1795 \mathrm{Ma}$. A less conservative, but still robust maximum depositional age of $\sim 1780 \mathrm{Ma}$ for the Kushma Formation is still consistent with this conclusion.
(2) The Malekhu Limestone was likely deposited before 1250 Ma . There appears to be no evidence for a thick succession of Neoproterozoic aged carbonate rocks with consistently near $0 \% \delta^{13} \mathrm{C}$ values throughout the section, especially those of the Infra Krol and Krol formations in the physiographic Lesser Himalaya (Buick et al., 1995; Knoll et al., 1995; Kumar et al., 2002).
(3) Upper and lower age bounds for Lesser Himalayan rocks in central Nepal between the Kuncha and Malekhu formations are $\sim 1875$ Ma to $\sim 1250 \mathrm{Ma}$ ( likely lower bound on the Malekhu Formation), indicating deposition occurred over no more than a 625 Ma time span.
(4) Based on the depositional ages determined in this study, the Ramgarh thrust is likely located farther south than previously mapped. The new location of the Ramgarh would imply that the thrust does in fact carry a thick package of rock similar to that of the Main Central thrust, and the
reduction in the number of faults mapped in the Modi Khola may reduce shortening estimates in the fold and thrust belt of central Nepal.
(5) Reevaluation of fault locations in central Nepal is called for based on the stratigraphy defined in the Modi Khola. Confirmation of this finding in other parts of Nepal is necessary, and if supported, should be reflected in balanced cross sections of the region.


Figure 3.20 - Comparison of mapping by Pearson and DeCelles (2005) and mapping completed in this study in 2007. Lithologic units are the same as in Figure 3.9. The Ranimata Formation in Pearson and DeCelles (2005) may be the lateral equivalent of the Kuncha Formation (orange) in central Nepal.
Contour interval is 200 m for both maps.

Table 1 - Petrographic descriptions of detrital zircon samples in this study.

| Sample Number | Sample Name | Predominant <br> Minerals | Accessory Minerals |
| :---: | :---: | :---: | :---: |
| 406020 | Unknown in Modi <br> Khola | quartz | muscovite, zircon |
| 406102 | Kushma Formation | quartz | chlorite, muscovite, <br> sulfides, tourmaline, <br> zircon |
| 506078 | Kuncha Formation | quartz, muscovite | hematite/magnetite, <br> sulfides, tourmaline, <br> zircon |
| 506079 | Fagfog Formation | quartz | muscovite, rutile, <br> tourmaline, zircon |

Table 2 - Carbon and oxygen values of the Malekhu formation. Data highlighted in color correspond to Figure 3.9.

| Height above base (m) | $\delta^{13} \mathrm{C}$ | $\delta^{18} \mathrm{O}$ | Height above base (m) | $\delta^{13} \mathrm{C}$ | $\delta^{18} \mathrm{O}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 133.5 | -0.1 | -11 | 292.5 | -0.9 | -14 |
| 136.5 | 0.0 | -11 | 292.5 | -1.0 | -14 |
| 139.5 | -0.1 | -12 | 292.5 | -1.0 | -14 |
| 139.5 | -0.1 | -12 | 292.5 | -1.0 | -14 |
| 139.5 | -0.2 | -13 | 292.5 | -1.0 | -14 |
| 139.5 | -0.2 | -13 | 295.5 | -0.9 | -13 |
| 139.5 | -0.2 | -13 | 298.5 | -0.7 | -13 |
| 139.5 | -0.1 | -13 | 301.5 | -0.9 | -12 |
| 145.5 | -0.4 | -11 | 304.5 | -1.1 | -13 |
| 148.5 | -0.6 | -12 | 304.5 | -1.2 | -13 |
| 148.5 | -0.6 | -12 | 310.5 | -1.0 | -13 |
| 148.5 | -0.6 | -12 | 310.5 | -1.1 | -14 |
| 148.5 | -0.6 | -12 | 313.5 | -1.2 | -14 |
| 148.5 | -0.6 | -12 | 313.5 | -1.2 | -14 |
| 151.5 | -0.8 | -13 | 319.5 | -1.2 | -12 |
| 154.5 | -1.0 | -13 | 322.5 | -1.4 | -12 |
| 163.5 | -1.0 | -13 | 325.5 | -1.1 | -12 |
| 166.5 | -1.0 | -14 | 331.5 | -1.1 | -11 |
| 172.5 | -0.8 | -11 | 334.5 | -1.2 | -11 |
| 175.5 | -0.7 | -10 | 337.5 | -0.9 | -12 |
| 178.5 | -0.8 | -11 | 339.5 | -1.0 | -12 |
| 178.5 | -0.8 | -11 | 342.5 | -1.1 | -11 |
| 181.5 | -1.1 | -11 | 345.5 | -1.1 | -12 |
| 184.5 | -0.9 | -9 | 345.5 | -1.1 | -13 |
| 187.5 | -0.9 | -11 | 357.5 | -0.8 | -14 |
| 190.5 | -0.9 | -11 | 360.5 | -0.7 | -14 |
| 196.5 | -0.8 | -10 | 363.5 | -1.1 | -14 |
| 199.5 | -1.0 | -11 | 363.5 | -1.1 | -14 |
| 202.5 | -1.3 | -11 | 366.5 | -0.9 | -12 |
| 208.5 | -0.7 | -10 | 366.5 | -0.9 | -12 |
| 211.5 | -1.1 | -11 | 369.5 | -1.1 | -12 |
| 214.5 | -1.0 | -11 | 369.5 | -1.1 | -12 |
| 217.5 | -1.1 | -10 | 372.5 | -0.9 | -12 |
| 220.5 | -1.6 | -13 | 372.5 | -0.9 | -12 |
| 223.5 | -1.5 | -13 | 376.5 | -0.6 | -12 |
| 225.5 | -0.8 | -11 | 379.5 | -1.1 | -13 |
| 229.5 | -0.6 | -9 | 382.5 | -1.1 | -14 |
| 232.5 | -1.0 | -10 | 384.5 | -1.3 | -14 |
| 235.5 | -1.3 | -11 | 387.5 | -0.0047 | -12 |
| 238.5 | -1.1 | -10 | 387.5 | 0.1180 | -12 |
| 241.5 | -0.8 | -9 | 387.5 | 0.1874 | -11 |
| 244.5 | -1.0 | -12 | 387.5 | -0.0112 | -12 |
| 250.5 | -1.0 | -12 | 393.5 | -1.0 | -14 |
| 250.5 | -0.9 | -12 | 393.5 | -1.1 | -14 |
| 253.5 | -1.2 | -13 | 396.5 | -0.9 | -13 |
| 259.5 | -0.8 | -13 | 396.5 | -0.8 | -14 |
| 265.5 | -1.3 | -13 | 408.5 | -0.6 | -13 |
| 274.5 | -1.1 | -15 | 408.5 | -0.7 | -14 |
| 280.5 | -1.7 | -17 | 414.5 | -0.7 | -14 |
| 280.5 | -1.2 | -15 | 414.5 | -0.6 | -14 |
| 283.5 | -1.4 | -15 | 420.5 | -1.7 | -16 |
| 289.5 | -0.9 | -14 | 429.5 | -1.6 | -15 |

Table 3- U-Pb detrital zircon analyses.


|  | 506 | uncha | , | 28.126 | 4.3 |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 799 | 75325 | 5.14828 | 1.67 | 0.33217 | 1.00 | 0.60 | 1849 | 16 | 1844 | 14 | 1839 | 24 | 1.01 | 1839 | 24 |
| 290 | 14050 | 5.11838 | 1.90 | 0.32830 | 1.00 | 0.52 | 1830 | 16 | 1839 | 16 | 1849 | 29 | 0.99 | 1849 | 29 |
| 321 | 30368 | 5.39134 | 1.58 | 0.34242 | 1.22 | 0.77 | 1898 | 20 | 1883 | 14 | 1867 | 18 | 1.02 | 1867 | 18 |
| 324 | 32306 | 5.41025 | 2.53 | 0.34355 | 1.32 | 0.52 | 1904 | 22 | 1886 | 22 | 1868 | 39 | 1.02 | 1868 | 39 |
| 198 | 15979 | 5.43952 | 1.50 | 0.34517 | 1.00 | 0.67 | 1912 | 17 | 1891 | 13 | 1869 | 20 | 1.02 | 1869 | 20 |
| 836 | 33214 | 5.08251 | 1.42 | 0.32226 | 1.00 | 0.71 | 1801 | 16 | 1833 | 12 | 1870 | 18 | 0.96 | 1870 | 18 |
| 248 | 10120 | 5.53482 | 1.82 | 0.35062 | 1.00 | 0.55 | 1938 | 17 | 1906 | 16 | 1872 | 27 | 1.04 | 1872 | 27 |
| 264 | 19544 | 5.39347 | 1.81 | 0.34147 | 1.00 | 0.55 | 1894 | 16 | 1884 | 16 | 1873 | 27 | 1.01 | 1873 | 27 |
| 376 | 37490 | 5.40051 | 1.41 | 0.34159 | 1.00 | 0.71 | 1894 | 16 | 1885 | 12 | 1875 | 18 | 1.01 | 1875 | 18 |
| 256 | 18555 | 5.39259 | 1.41 | 0.34091 | 1.00 | 0.71 | 1891 | 16 | 1884 | 12 | 1876 | 18 | 1.01 | 1876 | 18 |
| 893 | 69830 | 5.33576 | 1.56 | 0.33687 | 1.00 | 0.64 | 1872 | 16 | 1875 | 13 | 1878 | 22 | 1.00 | 1878 | 22 |
| 294 | 21963 | 5.48941 | 1.66 | 0.34638 | 1.00 | 0.60 | 1917 | 17 | 1899 | 14 | 1879 | 24 | 1.02 | 1879 | 24 |
| 223 | 22910 | 5.23708 | 1.92 | 0.33042 | 1.00 | 0.52 | 1840 | 16 | 1859 | 16 | 1879 | 30 | 0.98 | 1879 | 30 |
| 653 | 66194 | 5.21299 | 1.59 | 0.32889 | 1.00 | 0.63 | 1833 | 16 | 1855 | 14 | 1879 | 22 | 0.98 | 1879 | 22 |
| 442 | 34453 | 5.27139 | 1.41 | 0.33219 | 1.00 | 0.71 | 1849 | 16 | 1864 | 12 | 1881 | 18 | 0.98 | 1881 | 18 |
| 921 | 75827 | 5.54070 | 1.95 | 0.34904 | 1.19 | 0.61 | 1930 | 20 | 1907 | 17 | 1882 | 28 | 1.03 | 1882 | 28 |
| 291 | 20731 | 5.36995 | 1.42 | 0.33824 | 1.00 | 0.71 | 1878 | 16 | 1880 | 12 | 1882 | 18 | 1.00 | 1882 | 18 |
| 207 | 19472 | 5.29985 | 1.41 | 0.33357 | 1.00 | 0.71 | 1856 | 16 | 1869 | 12 | 1884 | 18 | 0.99 | 1884 | 18 |
| 797 | 40867 | 4.95667 | 1.81 | 0.31190 | 1.00 | 0.55 | 1750 | 15 | 1812 | 15 | 1884 | 27 | 0.93 | 1884 | 27 |
| 248 | 18072 | 5.39717 | 1.41 | 0.33953 | 1.00 | 0.71 | 1884 | 16 | 1884 | 12 | 1884 | 18 | 1.00 | 1884 | 18 |
| 184 | 20669 | 5.15907 | 1.60 | 0.32415 | 1.00 | 0.62 | 1810 | 16 | 1846 | 14 | 1887 | 23 | 0.96 | 1887 | 23 |
| 364 | 25208 | 5.42230 | 1.87 | 0.34058 | 1.21 | 0.65 | 1889 | 20 | 1888 | 16 | 1887 | 26 | 1.00 | 1887 | 26 |
| 1021 | 47042 | 4.88344 | 2.01 | 0.30646 | 1.00 | 0.50 | 1723 | 15 | 1799 | 17 | 1889 | 31 | 0.91 | 1889 | 31 |
| 598 | 40312 | 4.79973 | 4.97 | 0.30105 | 4.83 | 0.97 | 1697 | 72 | 1785 | 42 | 1890 | 21 | 0.90 | 1890 | 21 |
| 137 | 8882 | 4.99931 | 4.67 | 0.31354 | 4.50 | 0.96 | 1758 | 69 | 1819 | 40 | 1890 | 22 | 0.93 | 1890 | 22 |
| 230 | 22802 | 5.22026 | 3.27 | 0.32735 | 1.00 | 0.31 | 1826 | 16 | 1856 | 28 | 1890 | 56 | 0.97 | 1890 | 56 |
| 220 | 21475 | 5.17269 | 2.39 | 0.32434 | 1.71 | 0.72 | 1811 | 27 | 1848 | 20 | 1890 | 30 | 0.96 | 1890 | 30 |
| 204 | 19858 | 5.52960 | 2.30 | 0.34672 | 1.74 | 0.76 | 1919 | 29 | 1905 | 20 | 1890 | 27 | 1.02 | 1890 | 27 |
| 364 | 22118 | 5.47920 | 2.14 | 0.34340 | 1.84 | 0.86 | 1903 | 30 | 1897 | 18 | 1891 | 20 | 1.01 | 1891 | 20 |
| 183 | 13296 | 5.12942 | 1.49 | 0.32147 | 1.00 | 0.67 | 1797 | 16 | 1841 | 13 | 1891 | 20 | 0.95 | 1891 | 20 |
| 382 | 22829 | 5.37802 | 1.49 | 0.33693 | 1.11 | 0.74 | 1872 | 18 | 1881 | 13 | 1892 | 18 | 0.99 | 1892 | 18 |
| 175 | 15954 | 5.33255 | 2.58 | 0.33383 | 2.38 | 0.92 | 1857 | 38 | 1874 | 22 | 1893 | 18 | 0.98 | 1893 | 18 |
| 201 | 22723 | 5.21034 | 1.41 | 0.32605 | 1.00 | 0.71 | 1819 | 16 | 1854 | 12 | 1894 | 18 | 0.96 | 1894 | 18 |
| 524 | 47243 | 5.12015 | 1.52 | 0.32032 | 1.14 | 0.75 | 1791 | 18 | 1839 | 13 | 1894 | 18 | 0.95 | 1894 | 18 |
| 555 | 74574 | 5.25434 | 1.41 | 0.32865 | 1.00 | 0.71 | 1832 | 16 | 1861 | 12 | 1895 | 18 | 0.97 | 1895 | 18 |
| 430 | 35768 | 5.56044 | 2.17 | 0.34766 | 1.00 | 0.46 | 1923 | 17 | 1910 | 19 | 1895 | 35 | 1.01 | 1895 | 35 |


| 562 | 41138 | 5.43393 | 1.96 | 0.33971 | 1.04 | 0.53 | 1885 | 17 | 1890 | 17 | 1896 | 30 | 0.99 | 1896 | 30 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 265 | 25997 | 5.21799 | 2.31 | 0.32617 | 1.99 | 0.86 | 1820 | 32 | 1856 | 20 | 1896 | 21 | 0.96 | 1896 | 21 |
| 565 | 45610 | 5.15678 | 1.84 | 0.32185 | 1.00 | 0.54 | 1799 | 16 | 1846 | 16 | 1899 | 28 | 0.95 | 1899 | 28 |
| 468 | 43635 | 5.46276 | 1.47 | 0.34093 | 1.00 | 0.68 | 1891 | 16 | 1895 | 13 | 1899 | 19 | 1.00 | 1899 | 19 |
| 192 | 20518 | 5.53399 | 1.97 | 0.34523 | 1.70 | 0.86 | 1912 | 28 | 1906 | 17 | 1899 | 18 | 1.01 | 1899 | 18 |
| 687 | 54206 | 5.37924 | 1.41 | 0.33523 | 1.00 | 0.71 | 1864 | 16 | 1882 | 12 | 1901 | 18 | 0.98 | 1901 | 18 |
| 185 | 12530 | 5.37782 | 5.40 | 0.33497 | 1.41 | 0.26 | 1862 | 23 | 1881 | 46 | 1902 | 94 | 0.98 | 1902 | 94 |
| 479 | 34461 | 5.49145 | 1.41 | 0.34205 | 1.00 | 0.71 | 1897 | 16 | 1899 | 12 | 1902 | 18 | 1.00 | 1902 | 18 |
| 497 | 35611 | 5.47237 | 1.44 | 0.34055 | 1.03 | 0.72 | 1889 | 17 | 1896 | 12 | 1904 | 18 | 0.99 | 1904 | 18 |
| 752 | 77232 | 5.43694 | 1.41 | 0.33829 | 1.00 | 0.71 | 1878 | 16 | 1891 | 12 | 1904 | 18 | 0.99 | 1904 | 18 |
| 824 | 51174 | 4.69642 | 2.03 | 0.29216 | 1.17 | 0.58 | 1652 | 17 | 1767 | 17 | 1904 | 30 | 0.87 | 1904 | 30 |
| 1035 | 76872 | 5.26111 | 2.11 | 0.32687 | 1.86 | 0.88 | 1823 | 30 | 1863 | 18 | 1907 | 18 | 0.96 | 1907 | 18 |
| 341 | 33024 | 5.61773 | 1.41 | 0.34901 | 1.00 | 0.71 | 1930 | 17 | 1919 | 12 | 1907 | 18 | 1.01 | 1907 | 18 |
| 935 | 95878 | 5.11073 | 2.27 | 0.31749 | 1.45 | 0.64 | 1777 | 23 | 1838 | 19 | 1907 | 31 | 0.93 | 1907 | 31 |
| 249 | 20051 | 5.73419 | 1.41 | 0.35613 | 1.00 | 0.71 | 1964 | 17 | 1937 | 12 | 1907 | 18 | 1.03 | 1907 | 18 |
| 258 | 19237 | 5.72667 | 1.53 | 0.35556 | 1.00 | 0.65 | 1961 | 17 | 1935 | 13 | 1908 | 21 | 1.03 | 1908 | 21 |
| 422 | 31010 | 5.04518 | 2.41 | 0.31322 | 1.67 | 0.69 | 1757 | 26 | 1827 | 20 | 1908 | 31 | 0.92 | 1908 | 31 |
| 438 | 29870 | 5.38184 | 1.87 | 0.33411 | 1.00 | 0.53 | 1858 | 16 | 1882 | 16 | 1908 | 28 | 0.97 | 1908 | 28 |
| 390 | 23517 | 5.18912 | 1.57 | 0.32183 | 1.00 | 0.64 | 1799 | 16 | 1851 | 13 | 1910 | 22 | 0.94 | 1910 | 22 |
| 294 | 44950 | 5.29166 | 1.41 | 0.32807 | 1.00 | 0.71 | 1829 | 16 | 1868 | 12 | 1911 | 18 | 0.96 | 1911 | 18 |
| 196 | 10216 | 5.40277 | 1.64 | 0.33494 | 1.10 | 0.67 | 1862 | 18 | 1885 | 14 | 1911 | 22 | 0.97 | 1911 | 22 |
| 642 | 35288 | 5.41618 | 1.41 | 0.33567 | 1.00 | 0.71 | 1866 | 16 | 1887 | 12 | 1911 | 18 | 0.98 | 1911 | 18 |
| 276 | 25714 | 5.53956 | 1.41 | 0.34318 | 1.00 | 0.71 | 1902 | 16 | 1907 | 12 | 1912 | 18 | 0.99 | 1912 | 18 |
| 541 | 43454 | 5.35886 | 1.41 | 0.33196 | 1.00 | 0.71 | 1848 | 16 | 1878 | 12 | 1912 | 18 | 0.97 | 1912 | 18 |
| 359 | 15048 | 5.47800 | 1.41 | 0.33927 | 1.00 | 0.71 | 1883 | 16 | 1897 | 12 | 1912 | 18 | 0.98 | 1912 | 18 |
| 312 | 18354 | 5.72687 | 2.28 | 0.35461 | 1.61 | 0.71 | 1957 | 27 | 1935 | 20 | 1913 | 29 | 1.02 | 1913 | 29 |
| 347 | 23579 | 5.29138 | 1.95 | 0.32747 | 1.67 | 0.86 | 1826 | 27 | 1867 | 17 | 1914 | 18 | 0.95 | 1914 | 18 |
| 226 | 9290 | 5.13442 | 1.77 | 0.31699 | 1.33 | 0.75 | 1775 | 21 | 1842 | 15 | 1918 | 21 | 0.93 | 1918 | 21 |
| 391 | 36253 | 5.10155 | 1.96 | 0.31491 | 1.00 | 0.51 | 1765 | 15 | 1836 | 17 | 1918 | 30 | 0.92 | 1918 | 30 |
| 325 | 13546 | 4.90064 | 1.70 | 0.30238 | 1.34 | 0.79 | 1703 | 20 | 1802 | 14 | 1919 | 19 | 0.89 | 1919 | 19 |
| 348 | 20288 | 5.34281 | 1.68 | 0.32942 | 1.14 | 0.68 | 1836 | 18 | 1876 | 14 | 1921 | 22 | 0.96 | 1921 | 22 |
| 552 | 24251 | 4.55454 | 4.92 | 0.28050 | 3.62 | 0.74 | 1594 | 51 | 1741 | 41 | 1923 | 60 | 0.83 | 1923 | 60 |
| 467 | 43282 | 5.30852 | 2.62 | 0.32687 | 2.42 | 0.92 | 1823 | 38 | 1870 | 22 | 1923 | 18 | 0.95 | 1923 | 18 |
| 587 | 47168 | 5.12682 | 1.62 | 0.31528 | 1.00 | 0.62 | 1767 | 15 | 1841 | 14 | 1925 | 23 | 0.92 | 1925 | 23 |
| 174 | 15392 | 5.46316 | 1.42 | 0.33532 | 1.00 | 0.70 | 1864 | 16 | 1895 | 12 | 1929 | 18 | 0.97 | 1929 | 18 |
| 505 | 24696 | 5.16365 | 2.16 | 0.31671 | 1.60 | 0.74 | 1774 | 25 | 1847 | 18 | 1930 | 26 | 0.92 | 1930 | 26 |
| 442 | 46963 | 5.65633 | 3.18 | 0.34681 | 2.64 | 0.83 | 1919 | 44 | 1925 | 27 | 1930 | 32 | 0.99 | 1930 | 32 |
| 242 | 26659 | 5.68954 | 1.61 | 0.34585 | 1.24 | 0.77 | 1915 | 21 | 1930 | 14 | 1946 | 18 | 0.98 | 1946 | 18 |
| 308 | 33397 | 5.74878 | 1.56 | 0.34869 | 1.00 | 0.64 | 1928 | 17 | 1939 | 14 | 1950 | 21 | 0.99 | 1950 | 21 |
| 942 | 64760 | 5.31715 | 2.04 | 0.32062 | 1.00 | 0.49 | 1793 | 16 | 1872 | 17 | 1960 | 32 | 0.91 | 1960 | 32 |
| 145 | 18019 | 6.19633 | 1.41 | 0.37196 | 1.00 | 0.71 | 2039 | 17 | 2004 | 12 | 1968 | 18 | 1.04 | 1968 | 18 |
| 901 | 48270 | 5.84359 | 2.16 | 0.35014 | 1.80 | 0.83 | 1935 | 30 | 1953 | 19 | 1972 | 21 | 0.98 | 1972 | 21 |
| 304 | 33614 | 5.88007 | 2.80 | 0.35074 | 1.00 | 0.36 | 1938 | 17 | 1958 | 24 | 1980 | 46 | 0.98 | 1980 | 46 |


| 431 | 45909 | 6.16332 | 1.80 | 0.36513 | 1.33 | 0.74 | 2006 | 23 | 1999 | 16 | 1992 | 22 | 1.01 | 1992 | 22 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 354 | 36179 | 6.19520 | 1.67 | 0.36524 | 1.00 | 0.60 | 2007 | 17 | 2004 | 15 | 2000 | 24 | 1.00 | 2000 | 24 |
| 224 | 20781 | 6.07491 | 1.67 | 0.35599 | 1.00 | 0.60 | 1963 | 17 | 1987 | 15 | 2011 | 24 | 0.98 | 2011 | 24 |
| 491 | 47683 | 5.94776 | 2.46 | 0.34667 | 1.46 | 0.59 | 1919 | 24 | 1968 | 21 | 2021 | 35 | 0.95 | 2021 | 35 |
| 389 | 20933 | 6.04941 | 3.93 | 0.35220 | 3.68 | 0.94 | 1945 | 62 | 1983 | 34 | 2023 | 25 | 0.96 | 2023 | 25 |
| 317 | 16150 | 6.38643 | 1.41 | 0.36774 | 1.00 | 0.71 | 2019 | 17 | 2030 | 12 | 2042 | 18 | 0.99 | 2042 | 18 |
| 108 | 9554 | 6.52287 | 1.79 | 0.37436 | 1.00 | 0.56 | 2050 | 18 | 2049 | 16 | 2048 | 26 | 1.00 | 2048 | 26 |
| 365 | 18342 | 5.96098 | 1.82 | 0.34196 | 1.00 | 0.55 | 1896 | 16 | 1970 | 16 | 2049 | 27 | 0.93 | 2049 | 27 |
| 225 | 17742 | 6.35210 | 2.78 | 0.36225 | 2.44 | 0.88 | 1993 | 42 | 2026 | 24 | 2059 | 23 | 0.97 | 2059 | 23 |
| 473 | 45482 | 6.36356 | 1.59 | 0.35797 | 1.00 | 0.63 | 1973 | 17 | 2027 | 14 | 2083 | 22 | 0.95 | 2083 | 22 |
| 322 | 27184 | 6.54882 | 2.98 | 0.36694 | 2.43 | 0.81 | 2015 | 42 | 2052 | 26 | 2090 | 30 | 0.96 | 2090 | 30 |
| 393 | 41451 | 7.11704 | 1.41 | 0.39844 | 1.00 | 0.71 | 2162 | 18 | 2126 | 13 | 2092 | 18 | 1.03 | 2092 | 18 |
| 1014 | 58632 | 7.33811 | 2.21 | 0.39602 | 1.24 | 0.56 | 2151 | 23 | 2153 | 20 | 2156 | 32 | 1.00 | 2156 | 32 |
| 684 | 43603 | 5.47820 | 5.63 | 0.29214 | 4.19 | 0.74 | 1652 | 61 | 1897 | 48 | 2177 | 65 | 0.76 | 2177 | 65 |
| 120 | 14979 | 7.46065 | 1.42 | 0.39745 | 1.00 | 0.71 | 2157 | 18 | 2168 | 13 | 2179 | 17 | 0.99 | 2179 | 17 |
| 105 | 10285 | 7.58167 | 2.92 | 0.40133 | 2.74 | 0.94 | 2175 | 51 | 2183 | 26 | 2190 | 17 | 0.99 | 2190 | 17 |
| 241 | 34075 | 7.43213 | 2.02 | 0.39267 | 1.00 | 0.50 | 2135 | 18 | 2165 | 18 | 2193 | 30 | 0.97 | 2193 | 30 |
| 175 | 22298 | 7.32228 | 1.56 | 0.38473 | 1.00 | 0.64 | 2098 | 18 | 2152 | 14 | 2203 | 21 | 0.95 | 2203 | 21 |
| 238 | 24782 | 7.45694 | 1.69 | 0.39070 | 1.36 | 0.81 | 2126 | 25 | 2168 | 15 | 2208 | 17 | 0.96 | 2208 | 17 |
| 210 | 24323 | 7.60750 | 1.84 | 0.39823 | 1.00 | 0.54 | 2161 | 18 | 2186 | 17 | 2209 | 27 | 0.98 | 2209 | 27 |
| 281 | 27386 | 7.67173 | 1.55 | 0.39254 | 1.00 | 0.64 | 2135 | 18 | 2193 | 14 | 2249 | 21 | 0.95 | 2249 | 21 |
| 434 | 25074 | 6.86762 | 2.33 | 0.34672 | 1.41 | 0.61 | 1919 | 23 | 2094 | 21 | 2272 | 32 | 0.84 | 2272 | 32 |
| 261 | 27843 | 8.59628 | 3.63 | 0.42922 | 2.71 | 0.75 | 2302 | 52 | 2296 | 33 | 2291 | 42 | 1.00 | 2291 | 42 |
| 768 | 33091 | 6.25552 | 2.35 | 0.31153 | 1.08 | 0.46 | 1748 | 17 | 2012 | 21 | 2295 | 36 | 0.76 | 2295 | 36 |
| 381 | 31978 | 8.19314 | 1.41 | 0.40779 | 1.00 | 0.71 | 2205 | 19 | 2253 | 13 | 2296 | 17 | 0.96 | 2296 | 17 |
| 670 | 77469 | 9.14145 | 1.41 | 0.43214 | 1.00 | 0.71 | 2315 | 19 | 2352 | 13 | 2384 | 17 | 0.97 | 2384 | 17 |
| 491 | 75883 | 8.94157 | 2.60 | 0.41168 | 2.40 | 0.92 | 2223 | 45 | 2332 | 24 | 2429 | 17 | 0.91 | 2429 | 17 |
| 236 | 18829 | 9.62632 | 1.84 | 0.43921 | 1.00 | 0.54 | 2347 | 20 | 2400 | 17 | 2445 | 26 | 0.96 | 2445 | 26 |
| 489 | 18634 | 7.97869 | 3.42 | 0.36395 | 2.86 | 0.84 | 2001 | 49 | 2229 | 31 | 2445 | 32 | 0.82 | 2445 | 32 |
| 120 | 18869 | 10.18803 | 1.41 | 0.45361 | 1.00 | 0.71 | 2411 | 20 | 2452 | 13 | 2486 | 17 | 0.97 | 2486 | 17 |
| 221 | 34902 | 9.93674 | 1.41 | 0.43974 | 1.00 | 0.71 | 2349 | 20 | 2429 | 13 | 2496 | 17 | 0.94 | 2496 | 17 |
| 268 | 20950 | 10.15060 | 1.54 | 0.44900 | 1.17 | 0.76 | 2391 | 23 | 2449 | 14 | 2497 | 17 | 0.96 | 2497 | 17 |
| 302 | 35691 | 9.75769 | 1.75 | 0.42977 | 1.00 | 0.57 | 2305 | 19 | 2412 | 16 | 2504 | 24 | 0.92 | 2504 | 24 |
| 67 | 10498 | 10.45692 | 1.59 | 0.45968 | 1.24 | 0.78 | 2438 | 25 | 2476 | 15 | 2507 | 17 | 0.97 | 2507 | 17 |
| 347 | 36030 | 10.79749 | 1.59 | 0.47443 | 1.00 | 0.63 | 2503 | 21 | 2506 | 15 | 2508 | 21 | 1.00 | 2508 | 21 |
| 312 | 30238 | 10.06373 | 1.52 | 0.44213 | 1.00 | 0.66 | 2360 | 20 | 2441 | 14 | 2508 | 19 | 0.94 | 2508 | 19 |
| 384 | 33354 | 11.36050 | 3.23 | 0.49887 | 2.55 | 0.79 | 2609 | 55 | 2553 | 30 | 2509 | 33 | 1.04 | 2509 | 33 |
| 401 | 31728 | 10.82332 | 1.41 | 0.47394 | 1.00 | 0.71 | 2501 | 21 | 2508 | 13 | 2514 | 17 | 0.99 | 2514 | 17 |
| 114 | 12357 | 11.16286 | 1.56 | 0.48739 | 1.00 | 0.64 | 2559 | 21 | 2537 | 15 | 2519 | 20 | 1.02 | 2519 | 20 |
| 92 | 13022 | 10.50956 | 2.59 | 0.45865 | 1.49 | 0.58 | 2434 | 30 | 2481 | 24 | 2520 | 36 | 0.97 | 2520 | 36 |
| 356 | 44131 | 10.74324 | 2.66 | 0.46624 | 2.46 | 0.93 | 2467 | 50 | 2501 | 25 | 2529 | 17 | 0.98 | 2529 | 17 |
| 204 | 21637 | 11.44159 | 2.77 | 0.49429 | 1.14 | 0.41 | 2589 | 24 | 2560 | 26 | 2537 | 42 | 1.02 | 2537 | 42 |
| 880 | 110760 | 11.17538 | 2.39 | 0.47892 | 1.02 | 0.43 | 2523 | 21 | 2538 | 22 | 2550 | 36 | 0.99 | 2550 | 36 |


| 336 | 58186 | 10.75940 | 1.53 | 0.45499 | 1.00 | 0.65 | 2417 | 20 | 2503 | 14 | 2572 | 19 | 0.94 | 2572 | 19 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 137 | 13726 | 10.19257 | 2.74 | 0.42924 | 2.52 | 0.92 | 2302 | 49 | 2452 | 25 | 2579 | 18 | 0.89 | 2579 | 18 |
| 82 | 4293 | 15.04288 | 3.96 | 0.56565 | 2.00 | 0.51 | 2890 | 47 | 2818 | 38 | 2767 | 56 | 1.04 | 2767 | 56 |
| 262 | 35130 | 16.06230 | 2.21 | 0.54393 | 1.00 | 0.45 | 2800 | 23 | 2881 | 21 | 2937 | 32 | 0.95 | 2937 | 32 |
| 118 | 23405 | 24.37904 | 1.56 | 0.64886 | 1.00 | 0.64 | 3224 | 25 | 3284 | 15 | 3320 | 19 | 0.97 | 3320 | 19 |

Sample 506079 - Fagfog Fm (Lat, Long = $\mathbf{2 7 . 8 5 1 6 5}{ }^{\circ}$, 84.85856${ }^{\circ}$ )

| 453 | 21239 | 3.94169 | 9.21 | 0.26830 | 8.87 | 0.96 | 1532 | 121 | 1622 | 75 | 1741 | 46 | 0.88 | 1741 | 46 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 208 | 17373 | 4.99994 | 1.41 | 0.33379 | 1.00 | 0.71 | 1857 | 16 | 1819 | 12 | 1777 | 18 | 1.05 | 1777 | 18 |
| 403 | 24298 | 3.92498 | 1.70 | 0.25998 | 1.37 | 0.81 | 1490 | 18 | 1619 | 14 | 1791 | 18 | 0.83 | 1791 | 18 |
| 242 | 7933 | 3.88763 | 2.36 | 0.25697 | 2.13 | 0.90 | 1474 | 28 | 1611 | 19 | 1795 | 18 | 0.82 | 1795 | 18 |
| 1011 | 39800 | 4.51018 | 1.41 | 0.29751 | 1.00 | 0.71 | 1679 | 15 | 1733 | 12 | 1799 | 18 | 0.93 | 1799 | 18 |
| 665 | 34879 | 3.92561 | 4.82 | 0.25850 | 4.72 | 0.98 | 1482 | 63 | 1619 | 39 | 1802 | 18 | 0.82 | 1802 | 18 |
| 614 | 44412 | 5.03546 | 1.55 | 0.33097 | 1.19 | 0.77 | 1843 | 19 | 1825 | 13 | 1805 | 18 | 1.02 | 1805 | 18 |
| 250 | 13593 | 4.49646 | 1.42 | 0.29545 | 1.00 | 0.71 | 1669 | 15 | 1730 | 12 | 1806 | 18 | 0.92 | 1806 | 18 |
| 447 | 43076 | 4.77328 | 2.34 | 0.31291 | 1.84 | 0.79 | 1755 | 28 | 1780 | 20 | 1810 | 26 | 0.97 | 1810 | 26 |
| 509 | 18253 | 3.75043 | 1.90 | 0.24536 | 1.41 | 0.74 | 1414 | 18 | 1582 | 15 | 1814 | 23 | 0.78 | 1814 | 23 |
| 599 | 32209 | 5.01587 | 1.41 | 0.32763 | 1.00 | 0.71 | 1827 | 16 | 1822 | 12 | 1816 | 18 | 1.01 | 1816 | 18 |
| 436 | 36687 | 5.19961 | 1.87 | 0.33861 | 1.00 | 0.53 | 1880 | 16 | 1853 | 16 | 1822 | 29 | 1.03 | 1822 | 29 |
| 100 | 9582 | 5.05454 | 2.38 | 0.32837 | 1.00 | 0.42 | 1830 | 16 | 1829 | 20 | 1826 | 39 | 1.00 | 1826 | 39 |
| 189 | 13936 | 5.12215 | 1.93 | 0.33270 | 1.18 | 0.61 | 1851 | 19 | 1840 | 16 | 1827 | 28 | 1.01 | 1827 | 28 |
| 342 | 21958 | 4.48376 | 2.83 | 0.29123 | 1.00 | 0.35 | 1648 | 15 | 1728 | 23 | 1827 | 48 | 0.90 | 1827 | 48 |
| 298 | 16526 | 5.09035 | 1.58 | 0.32945 | 1.03 | 0.65 | 1836 | 16 | 1834 | 13 | 1833 | 22 | 1.00 | 1833 | 22 |
| 463 | 25384 | 4.59584 | 1.59 | 0.29744 | 1.00 | 0.63 | 1679 | 15 | 1749 | 13 | 1833 | 22 | 0.92 | 1833 | 22 |
| 455 | 28384 | 4.94158 | 1.42 | 0.31819 | 1.00 | 0.71 | 1781 | 16 | 1809 | 12 | 1842 | 18 | 0.97 | 1842 | 18 |
| 132 | 12571 | 5.21251 | 1.88 | 0.33461 | 1.23 | 0.66 | 1861 | 20 | 1855 | 16 | 1848 | 26 | 1.01 | 1848 | 26 |
| 146 | 12224 | 5.17455 | 1.72 | 0.33111 | 1.10 | 0.64 | 1844 | 18 | 1848 | 15 | 1854 | 24 | 0.99 | 1854 | 24 |
| 340 | 23228 | 4.80266 | 1.42 | 0.30722 | 1.01 | 0.71 | 1727 | 15 | 1785 | 12 | 1854 | 18 | 0.93 | 1854 | 18 |
| 688 | 43823 | 5.29792 | 1.56 | 0.33700 | 1.00 | 0.64 | 1872 | 16 | 1869 | 13 | 1864 | 22 | 1.00 | 1864 | 22 |
| 414 | 33141 | 4.97424 | 1.94 | 0.31496 | 1.54 | 0.79 | 1765 | 24 | 1815 | 16 | 1873 | 21 | 0.94 | 1873 | 21 |
| 634 | 64177 | 5.35703 | 1.41 | 0.33896 | 1.00 | 0.71 | 1882 | 16 | 1878 | 12 | 1874 | 18 | 1.00 | 1874 | 18 |
| 315 | 30235 | 5.60571 | 2.02 | 0.35445 | 1.00 | 0.49 | 1956 | 17 | 1917 | 17 | 1875 | 32 | 1.04 | 1875 | 32 |
| 175 | 15893 | 5.35631 | 1.59 | 0.33856 | 1.00 | 0.63 | 1880 | 16 | 1878 | 14 | 1876 | 22 | 1.00 | 1876 | 22 |
| 175 | 17476 | 5.43296 | 2.05 | 0.34237 | 1.00 | 0.49 | 1898 | 16 | 1890 | 18 | 1881 | 32 | 1.01 | 1881 | 32 |
| 176 | 14951 | 5.47082 | 1.61 | 0.34411 | 1.00 | 0.62 | 1906 | 17 | 1896 | 14 | 1885 | 23 | 1.01 | 1885 | 23 |
| 609 | 45214 | 5.52377 | 1.41 | 0.34715 | 1.00 | 0.71 | 1921 | 17 | 1904 | 12 | 1886 | 18 | 1.02 | 1886 | 18 |
| 278 | 16505 | 5.60621 | 1.85 | 0.35122 | 1.00 | 0.54 | 1940 | 17 | 1917 | 16 | 1892 | 28 | 1.03 | 1892 | 28 |
| 374 | 24768 | 5.67712 | 1.61 | 0.35556 | 1.00 | 0.62 | 1961 | 17 | 1928 | 14 | 1892 | 23 | 1.04 | 1892 | 23 |
| 420 | 16985 | 5.17608 | 1.51 | 0.32396 | 1.00 | 0.66 | 1809 | 16 | 1849 | 13 | 1894 | 20 | 0.96 | 1894 | 20 |
| 97 | 9760 | 5.54698 | 1.69 | 0.34715 | 1.36 | 0.81 | 1921 | 23 | 1908 | 15 | 1894 | 18 | 1.01 | 1894 | 18 |
| 228 | 18859 | 5.66053 | 1.61 | 0.35384 | 1.00 | 0.62 | 1953 | 17 | 1925 | 14 | 1896 | 23 | 1.03 | 1896 | 23 |
| 334 | 21219 | 5.17283 | 4.22 | 0.32323 | 3.91 | 0.93 | 1805 | 62 | 1848 | 36 | 1897 | 28 | 0.95 | 1897 | 28 |
| 286 | 21168 | 5.44763 | 2.55 | 0.33996 | 1.29 | 0.51 | 1887 | 21 | 1892 | 22 | 1899 | 40 | 0.99 | 1899 | 40 |


| 476 | 35737 | 5.38823 | 1.41 | 0.33609 | 1.00 | 0.71 | 1868 | 16 | 1883 | 12 | 1900 | 18 | 0.98 | 1900 | 18 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 194 | 14825 | 5.56936 | 2.20 | 0.34738 | 1.60 | 0.73 | 1922 | 27 | 1911 | 19 | 1900 | 27 | 1.01 | 1900 | 27 |
| 142 | 14385 | 5.04758 | 7.44 | 0.31479 | 7.29 | 0.98 | 1764 | 113 | 1827 | 63 | 1900 | 26 | 0.93 | 1900 | 26 |
| 399 | 35083 | 5.37783 | 2.04 | 0.33537 | 1.49 | 0.73 | 1864 | 24 | 1881 | 18 | 1900 | 25 | 0.98 | 1900 | 25 |
| 221 | 9742 | 5.31334 | 1.42 | 0.33118 | 1.00 | 0.71 | 1844 | 16 | 1871 | 12 | 1901 | 18 | 0.97 | 1901 | 18 |
| 216 | 20311 | 5.75861 | 1.65 | 0.35890 | 1.00 | 0.61 | 1977 | 17 | 1940 | 14 | 1901 | 24 | 1.04 | 1901 | 24 |
| 274 | 19166 | 5.49059 | 1.66 | 0.34213 | 1.33 | 0.80 | 1897 | 22 | 1899 | 14 | 1902 | 18 | 1.00 | 1902 | 18 |
| 152 | 12780 | 5.67512 | 1.41 | 0.35346 | 1.00 | 0.71 | 1951 | 17 | 1928 | 12 | 1902 | 18 | 1.03 | 1902 | 18 |
| 312 | 29445 | 5.65000 | 1.98 | 0.35187 | 1.26 | 0.64 | 1944 | 21 | 1924 | 17 | 1902 | 28 | 1.02 | 1902 | 28 |
| 278 | 23695 | 5.74208 | 1.64 | 0.35759 | 1.09 | 0.67 | 1971 | 19 | 1938 | 14 | 1903 | 22 | 1.04 | 1903 | 22 |
| 416 | 32624 | 5.47085 | 1.41 | 0.34055 | 1.00 | 0.71 | 1889 | 16 | 1896 | 12 | 1903 | 18 | 0.99 | 1903 | 18 |
| 328 | 34676 | 5.45839 | 1.57 | 0.33948 | 1.21 | 0.77 | 1884 | 20 | 1894 | 13 | 1905 | 18 | 0.99 | 1905 | 18 |
| 403 | 35290 | 5.50084 | 1.43 | 0.34202 | 1.00 | 0.70 | 1896 | 16 | 1901 | 12 | 1905 | 18 | 1.00 | 1905 | 18 |
| 286 | 25640 | 5.51823 | 1.41 | 0.34307 | 1.00 | 0.71 | 1901 | 16 | 1903 | 12 | 1906 | 18 | 1.00 | 1906 | 18 |
| 745 | 43661 | 4.81380 | 1.61 | 0.29906 | 1.26 | 0.78 | 1687 | 19 | 1787 | 14 | 1907 | 18 | 0.88 | 1907 | 18 |
| 339 | 30022 | 5.66664 | 1.41 | 0.35202 | 1.00 | 0.71 | 1944 | 17 | 1926 | 12 | 1907 | 18 | 1.02 | 1907 | 18 |
| 227 | 22437 | 5.57425 | 1.42 | 0.34621 | 1.00 | 0.71 | 1916 | 17 | 1912 | 12 | 1907 | 18 | 1.00 | 1907 | 18 |
| 501 | 33006 | 5.25583 | 1.44 | 0.32637 | 1.00 | 0.70 | 1821 | 16 | 1862 | 12 | 1908 | 19 | 0.95 | 1908 | 19 |
| 519 | 45834 | 5.64544 | 1.43 | 0.35053 | 1.00 | 0.70 | 1937 | 17 | 1923 | 12 | 1908 | 18 | 1.02 | 1908 | 18 |
| 189 | 18283 | 5.60327 | 1.58 | 0.34790 | 1.00 | 0.63 | 1925 | 17 | 1917 | 14 | 1908 | 22 | 1.01 | 1908 | 22 |
| 504 | 45074 | 5.72532 | 1.50 | 0.35531 | 1.00 | 0.67 | 1960 | 17 | 1935 | 13 | 1909 | 20 | 1.03 | 1909 | 20 |
| 571 | 39502 | 5.12461 | 1.46 | 0.31775 | 1.00 | 0.69 | 1779 | 16 | 1840 | 12 | 1910 | 19 | 0.93 | 1910 | 19 |
| 200 | 19858 | 5.52905 | 1.64 | 0.34278 | 1.19 | 0.72 | 1900 | 20 | 1905 | 14 | 1911 | 20 | 0.99 | 1911 | 20 |
| 341 | 29656 | 5.58306 | 1.41 | 0.34612 | 1.00 | 0.71 | 1916 | 17 | 1913 | 12 | 1911 | 18 | 1.00 | 1911 | 18 |
| 199 | 19646 | 5.83880 | 1.42 | 0.36194 | 1.00 | 0.71 | 1991 | 17 | 1952 | 12 | 1911 | 18 | 1.04 | 1911 | 18 |
| 457 | 40317 | 5.61243 | 1.41 | 0.34774 | 1.00 | 0.71 | 1924 | 17 | 1918 | 12 | 1912 | 18 | 1.01 | 1912 | 18 |
| 246 | 16782 | 5.48689 | 1.85 | 0.33990 | 1.00 | 0.54 | 1886 | 16 | 1899 | 16 | 1912 | 28 | 0.99 | 1912 | 28 |
| 407 | 27628 | 5.68313 | 1.41 | 0.35197 | 1.00 | 0.71 | 1944 | 17 | 1929 | 12 | 1912 | 18 | 1.02 | 1912 | 18 |
| 115 | 9166 | 5.78962 | 1.42 | 0.35838 | 1.00 | 0.70 | 1974 | 17 | 1945 | 12 | 1913 | 18 | 1.03 | 1913 | 18 |
| 212 | 18860 | 5.74132 | 1.67 | 0.35534 | 1.34 | 0.80 | 1960 | 23 | 1938 | 14 | 1914 | 18 | 1.02 | 1914 | 18 |
| 473 | 35637 | 5.67523 | 1.69 | 0.35124 | 1.00 | 0.59 | 1941 | 17 | 1928 | 15 | 1914 | 24 | 1.01 | 1914 | 24 |
| 585 | 35144 | 5.46652 | 2.08 | 0.33819 | 1.24 | 0.60 | 1878 | 20 | 1895 | 18 | 1914 | 30 | 0.98 | 1914 | 30 |
| 975 | 80405 | 5.55378 | 2.11 | 0.34355 | 1.73 | 0.82 | 1904 | 29 | 1909 | 18 | 1915 | 22 | 0.99 | 1915 | 22 |
| 828 | 62572 | 5.67947 | 1.41 | 0.35125 | 1.00 | 0.71 | 1941 | 17 | 1928 | 12 | 1915 | 18 | 1.01 | 1915 | 18 |
| 341 | 38412 | 5.71902 | 2.51 | 0.35367 | 1.76 | 0.70 | 1952 | 30 | 1934 | 22 | 1915 | 32 | 1.02 | 1915 | 32 |
| 407 | 26937 | 5.45092 | 1.47 | 0.33707 | 1.00 | 0.68 | 1873 | 16 | 1893 | 13 | 1915 | 19 | 0.98 | 1915 | 19 |
| 705 | 51992 | 5.58856 | 1.41 | 0.34554 | 1.00 | 0.71 | 1913 | 17 | 1914 | 12 | 1915 | 18 | 1.00 | 1915 | 18 |
| 276 | 19230 | 5.75485 | 2.17 | 0.35560 | 1.81 | 0.83 | 1961 | 31 | 1940 | 19 | 1917 | 22 | 1.02 | 1917 | 22 |
| 609 | 32048 | 5.48230 | 3.21 | 0.33873 | 2.87 | 0.89 | 1881 | 47 | 1898 | 28 | 1917 | 26 | 0.98 | 1917 | 26 |
| 214 | 18233 | 5.65798 | 1.45 | 0.34938 | 1.00 | 0.69 | 1932 | 17 | 1925 | 13 | 1918 | 19 | 1.01 | 1918 | 19 |
| 637 | 36945 | 5.43625 | 1.41 | 0.33558 | 1.00 | 0.71 | 1865 | 16 | 1891 | 12 | 1918 | 18 | 0.97 | 1918 | 18 |
| 598 | 32869 | 5.43971 | 1.57 | 0.33567 | 1.17 | 0.74 | 1866 | 19 | 1891 | 13 | 1919 | 19 | 0.97 | 1919 | 19 |
| 289 | 29840 | 5.68174 | 1.71 | 0.35058 | 1.39 | 0.81 | 1937 | 23 | 1929 | 15 | 1919 | 18 | 1.01 | 1919 | 18 |


| 1023 | 46134 | 4.34357 | 11.00 | 0.26791 | 10.82 | 0.98 | 1530 | 147 | 1702 | 91 | 1920 | 36 | 0.80 | 1920 | 36 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 270 | 26641 | 5.94655 | 1.78 | 0.36677 | 1.00 | 0.56 | 2014 | 17 | 1968 | 15 | 1920 | 26 | 1.05 | 1920 | 26 |
| 580 | 49404 | 5.76550 | 1.41 | 0.35558 | 1.00 | 0.71 | 1961 | 17 | 1941 | 12 | 1920 | 18 | 1.02 | 1920 | 18 |
| 972 | 52736 | 5.05853 | 2.57 | 0.31196 | 2.20 | 0.86 | 1750 | 34 | 1829 | 22 | 1920 | 24 | 0.91 | 1920 | 24 |
| 123 | 9252 | 5.46362 | 1.42 | 0.33690 | 1.00 | 0.71 | 1872 | 16 | 1895 | 12 | 1920 | 18 | 0.97 | 1920 | 18 |
| 451 | 29658 | 5.72789 | 1.93 | 0.35313 | 1.00 | 0.52 | 1950 | 17 | 1936 | 17 | 1921 | 30 | 1.02 | 1921 | 30 |
| 521 | 38051 | 5.13031 | 2.24 | 0.31614 | 1.66 | 0.74 | 1771 | 26 | 1841 | 19 | 1922 | 27 | 0.92 | 1922 | 27 |
| 521 | 35654 | 5.29774 | 1.61 | 0.32630 | 1.26 | 0.78 | 1820 | 20 | 1868 | 14 | 1922 | 18 | 0.95 | 1922 | 18 |
| 350 | 20311 | 5.48871 | 1.56 | 0.33801 | 1.00 | 0.64 | 1877 | 16 | 1899 | 13 | 1923 | 22 | 0.98 | 1923 | 22 |
| 694 | 39403 | 5.52498 | 2.11 | 0.34006 | 1.00 | 0.47 | 1887 | 16 | 1904 | 18 | 1924 | 33 | 0.98 | 1924 | 33 |
| 288 | 24658 | 5.78060 | 1.71 | 0.35567 | 1.27 | 0.74 | 1962 | 21 | 1944 | 15 | 1924 | 21 | 1.02 | 1924 | 21 |
| 470 | 38098 | 5.44224 | 1.55 | 0.33438 | 1.00 | 0.65 | 1860 | 16 | 1892 | 13 | 1927 | 21 | 0.97 | 1927 | 21 |
| 254 | 21652 | 5.84415 | 1.64 | 0.35889 | 1.30 | 0.79 | 1977 | 22 | 1953 | 14 | 1928 | 18 | 1.03 | 1928 | 18 |
| 476 | 29063 | 5.64058 | 2.51 | 0.34579 | 2.20 | 0.88 | 1914 | 36 | 1922 | 22 | 1931 | 22 | 0.99 | 1931 | 22 |
| 334 | 24153 | 5.76147 | 3.22 | 0.35313 | 2.94 | 0.91 | 1950 | 49 | 1941 | 28 | 1931 | 24 | 1.01 | 1931 | 24 |
| 929 | 71581 | 5.90616 | 1.47 | 0.36189 | 1.00 | 0.68 | 1991 | 17 | 1962 | 13 | 1932 | 19 | 1.03 | 1932 | 19 |
| 386 | 17267 | 5.82923 | 1.58 | 0.35666 | 1.00 | 0.63 | 1966 | 17 | 1951 | 14 | 1934 | 22 | 1.02 | 1934 | 22 |
| 346 | 27937 | 5.88072 | 1.56 | 0.35977 | 1.00 | 0.64 | 1981 | 17 | 1958 | 14 | 1934 | 21 | 1.02 | 1934 | 21 |
| 316 | 18551 | 5.47917 | 2.96 | 0.33512 | 2.79 | 0.94 | 1863 | 45 | 1897 | 25 | 1935 | 18 | 0.96 | 1935 | 18 |
| 272 | 14026 | 5.39750 | 4.48 | 0.32994 | 4.32 | 0.97 | 1838 | 69 | 1884 | 38 | 1936 | 21 | 0.95 | 1936 | 21 |
| 653 | 42195 | 5.52216 | 1.90 | 0.33731 | 1.00 | 0.53 | 1874 | 16 | 1904 | 16 | 1937 | 29 | 0.97 | 1937 | 29 |
| 732 | 44321 | 5.42583 | 1.55 | 0.33135 | 1.00 | 0.64 | 1845 | 16 | 1889 | 13 | 1938 | 21 | 0.95 | 1938 | 21 |
| 485 | 42949 | 5.73825 | 1.41 | 0.35025 | 1.00 | 0.71 | 1936 | 17 | 1937 | 12 | 1939 | 18 | 1.00 | 1939 | 18 |
| 191 | 13256 | 5.77188 | 1.48 | 0.35143 | 1.09 | 0.74 | 1941 | 18 | 1942 | 13 | 1943 | 18 | 1.00 | 1943 | 18 |
| 200 | 11424 | 4.55435 | 5.13 | 0.27720 | 4.61 | 0.90 | 1577 | 65 | 1741 | 43 | 1944 | 40 | 0.81 | 1944 | 40 |
| 302 | 25399 | 5.73173 | 2.13 | 0.34641 | 1.00 | 0.47 | 1917 | 17 | 1936 | 18 | 1956 | 34 | 0.98 | 1956 | 34 |
| 454 | 45746 | 6.01775 | 1.41 | 0.36174 | 1.00 | 0.71 | 1990 | 17 | 1978 | 12 | 1966 | 18 | 1.01 | 1966 | 18 |
| 729 | 25483 | 5.16099 | 4.68 | 0.30907 | 4.42 | 0.94 | 1736 | 67 | 1846 | 40 | 1973 | 27 | 0.88 | 1973 | 27 |
| 449 | 45616 | 6.24842 | 1.46 | 0.37230 | 1.00 | 0.69 | 2040 | 17 | 2011 | 13 | 1982 | 19 | 1.03 | 1982 | 19 |
| 540 | 39741 | 5.74426 | 1.41 | 0.34092 | 1.00 | 0.71 | 1891 | 16 | 1938 | 12 | 1989 | 18 | 0.95 | 1989 | 18 |
| 221 | 20071 | 5.83401 | 3.40 | 0.34477 | 1.81 | 0.53 | 1910 | 30 | 1951 | 30 | 1996 | 51 | 0.96 | 1996 | 51 |
| 562 | 45033 | 6.20229 | 1.41 | 0.36511 | 1.00 | 0.71 | 2006 | 17 | 2005 | 12 | 2003 | 18 | 1.00 | 2003 | 18 |
| 108 | 6978 | 5.74012 | 1.47 | 0.33639 | 1.06 | 0.72 | 1869 | 17 | 1937 | 13 | 2011 | 18 | 0.93 | 2011 | 18 |
| 227 | 15618 | 6.29533 | 1.52 | 0.36547 | 1.00 | 0.66 | 2008 | 17 | 2018 | 13 | 2028 | 20 | 0.99 | 2028 | 20 |
| 316 | 32921 | 6.50273 | 1.41 | 0.37689 | 1.00 | 0.71 | 2062 | 18 | 2046 | 12 | 2031 | 18 | 1.02 | 2031 | 18 |
| 315 | 26069 | 6.64897 | 1.93 | 0.38150 | 1.00 | 0.52 | 2083 | 18 | 2066 | 17 | 2049 | 29 | 1.02 | 2049 | 29 |
| 196 | 19604 | 6.64987 | 1.42 | 0.37914 | 1.00 | 0.71 | 2072 | 18 | 2066 | 12 | 2060 | 18 | 1.01 | 2060 | 18 |
| 486 | 33279 | 6.19572 | 1.70 | 0.34854 | 1.00 | 0.59 | 1928 | 17 | 2004 | 15 | 2083 | 24 | 0.93 | 2083 | 24 |
| 589 | 34343 | 5.92447 | 6.52 | 0.33250 | 4.01 | 0.62 | 1851 | 65 | 1965 | 57 | 2087 | 90 | 0.89 | 2087 | 90 |
| 275 | 18532 | 6.36796 | 2.18 | 0.35294 | 1.67 | 0.77 | 1949 | 28 | 2028 | 19 | 2109 | 25 | 0.92 | 2109 | 25 |
| 316 | 26592 | 6.94728 | 2.80 | 0.38177 | 1.00 | 0.36 | 2085 | 18 | 2105 | 25 | 2124 | 46 | 0.98 | 2124 | 46 |
| 736 | 44712 | 5.70417 | 5.96 | 0.31305 | 5.51 | 0.93 | 1756 | 85 | 1932 | 51 | 2127 | 40 | 0.83 | 2127 | 40 |
| 357 | 28324 | 6.90652 | 2.21 | 0.37850 | 1.00 | 0.45 | 2069 | 18 | 2099 | 20 | 2129 | 34 | 0.97 | 2129 | 34 |


| 361 | 40498 | 6.93259 | 2.07 | 0.37870 | 1.00 | 0.48 | 2070 | 18 | 2103 | 18 | 2135 | 32 | 0.97 | 2135 | 32 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 434 | 25260 | 5.85780 | 2.21 | 0.31443 | 1.60 | 0.72 | 1762 | 25 | 1955 | 19 | 2166 | 27 | 0.81 | 2166 | 27 |
| 257 | 26767 | 7.57592 | 3.15 | 0.40208 | 1.55 | 0.49 | 2179 | 29 | 2182 | 28 | 2185 | 48 | 1.00 | 2185 | 48 |
| 433 | 22221 | 7.85162 | 2.38 | 0.41221 | 1.34 | 0.56 | 2225 | 25 | 2214 | 21 | 2204 | 34 | 1.01 | 2204 | 34 |
| 338 | 20833 | 7.05794 | 2.41 | 0.36943 | 1.06 | 0.44 | 2027 | 18 | 2119 | 21 | 2209 | 38 | 0.92 | 2209 | 38 |
| 124 | 14381 | 8.10867 | 1.49 | 0.42197 | 1.00 | 0.67 | 2269 | 19 | 2243 | 14 | 2219 | 19 | 1.02 | 2219 | 19 |
| 196 | 17758 | 7.68770 | 2.58 | 0.39757 | 2.08 | 0.81 | 2158 | 38 | 2195 | 23 | 2230 | 27 | 0.97 | 2230 | 27 |
| 102 | 13091 | 8.30100 | 1.58 | 0.42493 | 1.00 | 0.63 | 2283 | 19 | 2264 | 14 | 2248 | 21 | 1.02 | 2248 | 21 |
| 402 | 35361 | 7.93988 | 3.12 | 0.40425 | 1.03 | 0.33 | 2189 | 19 | 2224 | 28 | 2257 | 51 | 0.97 | 2257 | 51 |
| 747 | 64328 | 8.88887 | 3.93 | 0.44503 | 3.11 | 0.79 | 2373 | 62 | 2327 | 36 | 2286 | 41 | 1.04 | 2286 | 41 |
| 483 | 41119 | 8.49440 | 1.88 | 0.41864 | 1.38 | 0.74 | 2254 | 26 | 2285 | 17 | 2313 | 22 | 0.97 | 2313 | 22 |
| 236 | 22065 | 7.80004 | 1.43 | 0.37986 | 1.02 | 0.71 | 2076 | 18 | 2208 | 13 | 2334 | 17 | 0.89 | 2334 | 17 |
| 336 | 35156 | 8.89805 | 1.85 | 0.43115 | 1.43 | 0.77 | 2311 | 28 | 2328 | 17 | 2342 | 20 | 0.99 | 2342 | 20 |
| 241 | 22653 | 9.35810 | 1.41 | 0.44950 | 1.00 | 0.71 | 2393 | 20 | 2374 | 13 | 2357 | 17 | 1.02 | 2357 | 17 |
| 467 | 49695 | 8.91830 | 1.46 | 0.42440 | 1.00 | 0.68 | 2280 | 19 | 2330 | 13 | 2373 | 18 | 0.96 | 2373 | 18 |
| 538 | 47593 | 9.24269 | 3.90 | 0.43896 | 2.02 | 0.52 | 2346 | 40 | 2362 | 36 | 2377 | 57 | 0.99 | 2377 | 57 |
| 124 | 17176 | 9.78670 | 1.41 | 0.46172 | 1.00 | 0.71 | 2447 | 20 | 2415 | 13 | 2388 | 17 | 1.02 | 2388 | 17 |
| 104 | 13087 | 9.96281 | 1.94 | 0.46925 | 1.66 | 0.86 | 2480 | 34 | 2431 | 18 | 2391 | 17 | 1.04 | 2391 | 17 |
| 783 | 88452 | 9.84823 | 1.80 | 0.45874 | 1.00 | 0.55 | 2434 | 20 | 2421 | 17 | 2409 | 25 | 1.01 | 2409 | 25 |
| 693 | 42007 | 8.33071 | 4.81 | 0.38040 | 1.00 | 0.21 | 2078 | 18 | 2268 | 44 | 2443 | 80 | 0.85 | 2443 | 80 |
| 76 | 7890 | 9.93253 | 1.42 | 0.44787 | 1.00 | 0.70 | 2386 | 20 | 2429 | 13 | 2465 | 17 | 0.97 | 2465 | 17 |
| 477 | 61435 | 10.59911 | 2.41 | 0.47146 | 1.43 | 0.59 | 2490 | 30 | 2489 | 22 | 2488 | 33 | 1.00 | 2488 | 33 |
| 356 | 38856 | 10.39074 | 2.51 | 0.45951 | 1.00 | 0.40 | 2437 | 20 | 2470 | 23 | 2497 | 39 | 0.98 | 2497 | 39 |
| 589 | 49758 | 9.87460 | 2.76 | 0.43636 | 2.08 | 0.75 | 2334 | 41 | 2423 | 25 | 2499 | 31 | 0.93 | 2499 | 31 |
| 516 | 54977 | 10.55048 | 1.82 | 0.46492 | 1.52 | 0.84 | 2461 | 31 | 2484 | 17 | 2503 | 17 | 0.98 | 2503 | 17 |
| 438 | 50835 | 11.02207 | 1.62 | 0.48569 | 1.16 | 0.72 | 2552 | 24 | 2525 | 15 | 2503 | 19 | 1.02 | 2503 | 19 |
| 228 | 20199 | 10.70599 | 1.63 | 0.46745 | 1.11 | 0.68 | 2472 | 23 | 2498 | 15 | 2519 | 20 | 0.98 | 2519 | 20 |
| 782 | 52763 | 10.69395 | 3.31 | 0.46552 | 3.01 | 0.91 | 2464 | 62 | 2497 | 31 | 2524 | 23 | 0.98 | 2524 | 23 |
| 932 | 80239 | 11.05779 | 1.92 | 0.48108 | 1.32 | 0.69 | 2532 | 28 | 2528 | 18 | 2525 | 24 | 1.00 | 2525 | 24 |
| 280 | 25492 | 10.46292 | 1.44 | 0.45403 | 1.00 | 0.70 | 2413 | 20 | 2477 | 13 | 2529 | 17 | 0.95 | 2529 | 17 |
| 396 | 37176 | 10.58262 | 1.72 | 0.45770 | 1.00 | 0.58 | 2429 | 20 | 2487 | 16 | 2535 | 23 | 0.96 | 2535 | 23 |
| 307 | 29780 | 10.86399 | 2.53 | 0.46884 | 1.68 | 0.66 | 2478 | 35 | 2512 | 24 | 2538 | 32 | 0.98 | 2538 | 32 |
| 246 | 30945 | 11.70167 | 1.73 | 0.50399 | 1.41 | 0.82 | 2631 | 30 | 2581 | 16 | 2542 | 17 | 1.04 | 2542 | 17 |
| 110 | 11753 | 10.91751 | 5.65 | 0.46672 | 5.50 | 0.97 | 2469 | 113 | 2516 | 53 | 2554 | 22 | 0.97 | 2554 | 22 |
| 76 | 12191 | 11.88819 | 1.75 | 0.50679 | 1.24 | 0.71 | 2643 | 27 | 2596 | 16 | 2559 | 21 | 1.03 | 2559 | 21 |
| 141 | 19370 | 11.88210 | 1.41 | 0.50537 | 1.00 | 0.71 | 2637 | 22 | 2595 | 13 | 2563 | 17 | 1.03 | 2563 | 17 |
| 565 | 45260 | 10.38622 | 1.91 | 0.44148 | 1.39 | 0.73 | 2357 | 27 | 2470 | 18 | 2564 | 22 | 0.92 | 2564 | 22 |
| 374 | 43758 | 11.66867 | 1.49 | 0.49444 | 1.00 | 0.67 | 2590 | 21 | 2578 | 14 | 2569 | 19 | 1.01 | 2569 | 19 |
| 398 | 43350 | 11.01575 | 3.38 | 0.46526 | 2.98 | 0.88 | 2463 | 61 | 2524 | 31 | 2574 | 27 | 0.96 | 2574 | 27 |
| 333 | 25457 | 11.38063 | 1.63 | 0.48009 | 1.15 | 0.70 | 2528 | 24 | 2555 | 15 | 2576 | 19 | 0.98 | 2576 | 19 |
| 380 | 34169 | 11.32346 | 1.65 | 0.47754 | 1.31 | 0.79 | 2517 | 27 | 2550 | 15 | 2577 | 17 | 0.98 | 2577 | 17 |
| 155 | 14182 | 9.80210 | 2.24 | 0.41315 | 1.57 | 0.70 | 2229 | 30 | 2416 | 21 | 2578 | 27 | 0.86 | 2578 | 27 |
| 72 | 9435 | 11.69598 | 1.49 | 0.49041 | 1.10 | 0.74 | 2572 | 23 | 2580 | 14 | 2587 | 17 | 0.99 | 2587 | 17 |


| 404 | 41781 | 12.33962 | 3.05 | 0.51658 | 1.76 | 0.58 | 2685 | 39 | 2631 | 29 | 2589 | 42 | 1.04 | 2589 | 42 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :--- | :--- | :--- | :--- |
| 418 | 42983 | 12.28759 | 1.54 | 0.51117 | 1.00 | 0.65 | 2662 | 22 | 2627 | 14 | 2600 | 20 | 1.02 | 2600 | 20 |
| 213 | 29871 | 12.03874 | 1.68 | 0.50026 | 1.00 | 0.60 | 2615 | 21 | 2607 | 16 | 2602 | 23 | 1.01 | 2602 | 23 |
| 97 | 8106 | 10.51285 | 1.50 | 0.43679 | 1.04 | 0.69 | 2336 | 20 | 2481 | 14 | 2602 | 18 | 0.90 | 2602 | 18 |
| 135 | 19859 | 13.18117 | 2.41 | 0.52587 | 2.19 | 0.91 | 2724 | 49 | 2693 | 23 | 2669 | 17 | 1.02 | 2669 | 17 |
| 524 | 70357 | 13.87237 | 2.03 | 0.53719 | 1.56 | 0.77 | 2772 | 35 | 2741 | 19 | 2719 | 21 | 1.02 | 2719 | 21 |
| 241 | 34540 | 13.75199 | 1.80 | 0.53245 | 1.50 | 0.83 | 2752 | 34 | 2733 | 17 | 2719 | 16 | 1.01 | 2719 | 16 |
| 504 | 53437 | 13.62567 | 1.81 | 0.52655 | 1.00 | 0.55 | 2727 | 22 | 2724 | 17 | 2722 | 25 | 1.00 | 2722 | 25 |
| 258 | 45679 | 17.91657 | 1.81 | 0.60354 | 1.00 | 0.55 | 3044 | 24 | 2985 | 17 | 2946 | 24 | 1.03 | 2946 | 24 |
| 464 | 55045 | 22.86169 | 1.97 | 0.61891 | 1.70 | 0.86 | 3106 | 42 | 3221 | 19 | 3294 | 16 | 0.94 | 3294 | 16 |
| 518 | 82235 | 26.87348 | 1.48 | 0.68445 | 1.00 | 0.68 | 3361 | 26 | 3379 | 14 | 3389 | 17 | 0.99 | 3389 | 17 |

Sample 406102 - Kushma Fm (Lat, Long $=\mathbf{2 8 . 2 3 0 8 2}^{\circ}$, $\mathbf{8 3 . 6 8 1 1 8}^{\circ}$ )

| 185 | 7504 | 3.99509 | 1.77 | 0.27293 | 1.45 | 0.82 | 1556 | 20 | 1633 | 14 | 1735 | 19 | 0.90 | 1735 | 19 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 702 | 18520 | 3.57136 | 2.44 | 0.24204 | 2.18 | 0.89 | 1397 | 27 | 1543 | 19 | 1749 | 20 | 0.80 | 1749 | 20 |
| 183 | 14152 | 4.07971 | 1.77 | 0.27609 | 1.00 | 0.56 | 1572 | 14 | 1650 | 14 | 1752 | 27 | 0.90 | 1752 | 27 |
| 351 | 32252 | 4.25295 | 2.66 | 0.28736 | 2.10 | 0.79 | 1628 | 30 | 1684 | 22 | 1755 | 30 | 0.93 | 1755 | 30 |
| 356 | 12787 | 4.16873 | 1.62 | 0.28147 | 1.01 | 0.63 | 1599 | 14 | 1668 | 13 | 1756 | 23 | 0.91 | 1756 | 23 |
| 563 | 37554 | 3.94270 | 3.52 | 0.26512 | 3.25 | 0.92 | 1516 | 44 | 1622 | 29 | 1764 | 25 | 0.86 | 1764 | 25 |
| 963 | 49448 | 3.69448 | 1.41 | 0.24809 | 1.00 | 0.71 | 1429 | 13 | 1570 | 11 | 1766 | 18 | 0.81 | 1766 | 18 |
| 740 | 27759 | 3.67196 | 5.09 | 0.24590 | 4.95 | 0.97 | 1417 | 63 | 1565 | 41 | 1771 | 22 | 0.80 | 1771 | 22 |
| 758 | 23460 | 3.90341 | 4.21 | 0.26133 | 3.92 | 0.93 | 1497 | 52 | 1614 | 34 | 1772 | 28 | 0.84 | 1772 | 28 |
| 345 | 28018 | 4.32199 | 1.43 | 0.28908 | 1.00 | 0.70 | 1637 | 14 | 1698 | 12 | 1773 | 19 | 0.92 | 1773 | 19 |
| 279 | 9603 | 4.63653 | 2.08 | 0.30936 | 1.00 | 0.48 | 1738 | 15 | 1756 | 17 | 1778 | 33 | 0.98 | 1778 | 33 |
| 801 | 37514 | 4.05965 | 2.11 | 0.27054 | 1.12 | 0.53 | 1544 | 15 | 1646 | 17 | 1780 | 33 | 0.87 | 1780 | 33 |
| 463 | 41638 | 4.40820 | 1.75 | 0.29376 | 1.18 | 0.67 | 1660 | 17 | 1714 | 14 | 1780 | 24 | 0.93 | 1780 | 24 |
| 819 | 49993 | 4.30500 | 1.58 | 0.28672 | 1.00 | 0.63 | 1625 | 14 | 1694 | 13 | 1781 | 22 | 0.91 | 1781 | 22 |
| 576 | 26173 | 4.31714 | 2.57 | 0.28727 | 2.37 | 0.92 | 1628 | 34 | 1697 | 21 | 1783 | 18 | 0.91 | 1783 | 18 |
| 329 | 16693 | 4.74969 | 2.02 | 0.31558 | 1.00 | 0.50 | 1768 | 15 | 1776 | 17 | 1785 | 32 | 0.99 | 1785 | 32 |
| 230 | 23540 | 4.73214 | 1.69 | 0.31422 | 1.00 | 0.59 | 1761 | 15 | 1773 | 14 | 1786 | 25 | 0.99 | 1786 | 25 |
| 213 | 23420 | 4.73452 | 1.73 | 0.31426 | 1.25 | 0.72 | 1762 | 19 | 1773 | 15 | 1787 | 22 | 0.99 | 1787 | 22 |
| 260 | 14316 | 4.56332 | 2.64 | 0.30289 | 1.00 | 0.38 | 1706 | 15 | 1743 | 22 | 1787 | 44 | 0.95 | 1787 | 44 |
| 125 | 14424 | 4.79172 | 1.55 | 0.31802 | 1.00 | 0.64 | 1780 | 16 | 1783 | 13 | 1787 | 22 | 1.00 | 1787 | 22 |
| 247 | 18381 | 4.64407 | 2.63 | 0.30804 | 2.19 | 0.83 | 1731 | 33 | 1757 | 22 | 1788 | 26 | 0.97 | 1788 | 26 |
| 380 | 19305 | 4.61506 | 1.70 | 0.30603 | 1.00 | 0.59 | 1721 | 15 | 1752 | 14 | 1789 | 25 | 0.96 | 1789 | 25 |
| 191 | 7725 | 4.41570 | 1.64 | 0.29274 | 1.00 | 0.61 | 1655 | 15 | 1715 | 14 | 1789 | 24 | 0.93 | 1789 | 24 |
| 192 | 14886 | 4.96671 | 1.45 | 0.32910 | 1.00 | 0.69 | 1834 | 16 | 1814 | 12 | 1790 | 19 | 1.02 | 1790 | 19 |
| 445 | 18194 | 3.83488 | 2.99 | 0.25367 | 2.82 | 0.94 | 1457 | 37 | 1600 | 24 | 1793 | 18 | 0.81 | 1793 | 18 |
| 683 | 39891 | 3.88028 | 7.05 | 0.25655 | 6.98 | 0.99 | 1472 | 92 | 1610 | 57 | 1794 | 18 | 0.82 | 1794 | 18 |
| 358 | 8850 | 3.92576 | 1.42 | 0.25944 | 1.00 | 0.70 | 1487 | 13 | 1619 | 11 | 1795 | 18 | 0.83 | 1795 | 18 |
| 169 | 14933 | 4.70962 | 1.42 | 0.31121 | 1.00 | 0.71 | 1747 | 15 | 1769 | 12 | 1795 | 18 | 0.97 | 1795 | 18 |
| 310 | 13558 | 4.56422 | 1.88 | 0.30159 | 1.00 | 0.53 | 1699 | 15 | 1743 | 16 | 1795 | 29 | 0.95 | 1795 | 29 |
| 265 | 19902 | 4.52185 | 2.25 | 0.29811 | 1.00 | 0.45 | 1682 | 15 | 1735 | 19 | 1800 | 37 | 0.93 | 1800 | 37 |


| 329 | 29875 | 4.70442 | 2.79 | 0.30948 | 1.85 | 0.66 | 1738 | 28 | 1768 | 23 | 1803 | 38 | 0.96 | 1803 | 38 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 827 | 39597 | 3.67812 | 2.82 | 0.24145 | 2.64 | 0.94 | 1394 | 33 | 1567 | 23 | 1807 | 18 | 0.77 | 1807 | 18 |
| 601 | 25329 | 4.32514 | 2.15 | 0.28346 | 1.90 | 0.88 | 1609 | 27 | 1698 | 18 | 1810 | 18 | 0.89 | 1810 | 18 |
| 609 | 45994 | 4.23812 | 4.60 | 0.27762 | 4.49 | 0.98 | 1579 | 63 | 1681 | 38 | 1811 | 18 | 0.87 | 1811 | 18 |
| 233 | 18602 | 4.54823 | 1.91 | 0.29780 | 1.00 | 0.52 | 1680 | 15 | 1740 | 16 | 1812 | 30 | 0.93 | 1812 | 30 |
| 280 | 13880 | 4.69912 | 1.55 | 0.30763 | 1.00 | 0.64 | 1729 | 15 | 1767 | 13 | 1812 | 22 | 0.95 | 1812 | 22 |
| 166 | 19265 | 4.89824 | 1.93 | 0.32056 | 1.03 | 0.53 | 1792 | 16 | 1802 | 16 | 1813 | 30 | 0.99 | 1813 | 30 |
| 363 | 16283 | 4.34767 | 1.77 | 0.28449 | 1.10 | 0.62 | 1614 | 16 | 1702 | 15 | 1813 | 25 | 0.89 | 1813 | 25 |
| 951 | 28356 | 3.93803 | 1.41 | 0.25766 | 1.00 | 0.71 | 1478 | 13 | 1622 | 11 | 1813 | 18 | 0.81 | 1813 | 18 |
| 277 | 15622 | 4.30911 | 1.78 | 0.28193 | 1.00 | 0.56 | 1601 | 14 | 1695 | 15 | 1813 | 27 | 0.88 | 1813 | 27 |
| 422 | 22577 | 4.60584 | 1.71 | 0.30118 | 1.00 | 0.58 | 1697 | 15 | 1750 | 14 | 1814 | 25 | 0.94 | 1814 | 25 |
| 406 | 34021 | 4.71752 | 1.55 | 0.30822 | 1.00 | 0.65 | 1732 | 15 | 1770 | 13 | 1816 | 21 | 0.95 | 1816 | 21 |
| 211 | 13491 | 4.74295 | 2.06 | 0.30977 | 1.36 | 0.66 | 1740 | 21 | 1775 | 17 | 1817 | 28 | 0.96 | 1817 | 28 |
| 628 | 20422 | 4.55078 | 1.55 | 0.29718 | 1.00 | 0.65 | 1677 | 15 | 1740 | 13 | 1817 | 21 | 0.92 | 1817 | 21 |
| 582 | 46600 | 4.71254 | 2.80 | 0.30746 | 2.33 | 0.83 | 1728 | 35 | 1769 | 23 | 1819 | 28 | 0.95 | 1819 | 28 |
| 517 | 48208 | 4.81391 | 1.59 | 0.31370 | 1.24 | 0.78 | 1759 | 19 | 1787 | 13 | 1821 | 18 | 0.97 | 1821 | 18 |
| 357 | 20458 | 4.68253 | 1.90 | 0.30508 | 1.00 | 0.53 | 1716 | 15 | 1764 | 16 | 1821 | 29 | 0.94 | 1821 | 29 |
| 790 | 28313 | 3.95354 | 3.05 | 0.25659 | 2.82 | 0.92 | 1472 | 37 | 1625 | 25 | 1828 | 21 | 0.81 | 1828 | 21 |
| 391 | 12893 | 4.75162 | 2.36 | 0.30767 | 1.82 | 0.77 | 1729 | 28 | 1776 | 20 | 1832 | 27 | 0.94 | 1832 | 27 |
| 455 | 30306 | 4.92036 | 2.42 | 0.31842 | 1.00 | 0.41 | 1782 | 16 | 1806 | 20 | 1833 | 40 | 0.97 | 1833 | 40 |
| 336 | 13517 | 5.06443 | 2.13 | 0.32772 | 1.00 | 0.47 | 1827 | 16 | 1830 | 18 | 1833 | 34 | 1.00 | 1833 | 34 |
| 1022 | 32716 | 4.42412 | 1.97 | 0.28585 | 1.09 | 0.55 | 1621 | 16 | 1717 | 16 | 1836 | 30 | 0.88 | 1836 | 30 |
| 617 | 15517 | 3.91601 | 3.35 | 0.25290 | 1.95 | 0.58 | 1453 | 25 | 1617 | 27 | 1837 | 49 | 0.79 | 1837 | 49 |
| 199 | 17659 | 4.68096 | 2.52 | 0.30226 | 1.02 | 0.40 | 1702 | 15 | 1764 | 21 | 1837 | 42 | 0.93 | 1837 | 42 |
| 496 | 18161 | 4.63598 | 1.96 | 0.29896 | 1.00 | 0.51 | 1686 | 15 | 1756 | 16 | 1840 | 30 | 0.92 | 1840 | 30 |
| 802 | 26874 | 4.77141 | 6.57 | 0.30735 | 5.08 | 0.77 | 1728 | 77 | 1780 | 55 | 1842 | 75 | 0.94 | 1842 | 75 |
| 231 | 7164 | 4.48075 | 2.82 | 0.28853 | 1.71 | 0.61 | 1634 | 25 | 1727 | 23 | 1842 | 41 | 0.89 | 1842 | 41 |
| 179 | 22196 | 5.00367 | 1.88 | 0.32193 | 1.00 | 0.53 | 1799 | 16 | 1820 | 16 | 1844 | 29 | 0.98 | 1844 | 29 |
| 414 | 25411 | 4.97921 | 2.27 | 0.31897 | 1.16 | 0.51 | 1785 | 18 | 1816 | 19 | 1852 | 35 | 0.96 | 1852 | 35 |
| 792 | 37687 | 4.43254 | 2.49 | 0.28284 | 1.55 | 0.62 | 1606 | 22 | 1718 | 21 | 1859 | 35 | 0.86 | 1859 | 35 |
| 543 | 18915 | 4.75875 | 1.47 | 0.30155 | 1.00 | 0.68 | 1699 | 15 | 1778 | 12 | 1871 | 19 | 0.91 | 1871 | 19 |
| 244 | 12013 | 4.65777 | 1.88 | 0.29469 | 1.00 | 0.53 | 1665 | 15 | 1760 | 16 | 1874 | 29 | 0.89 | 1874 | 29 |
| 283 | 11929 | 4.84100 | 1.67 | 0.30544 | 1.00 | 0.60 | 1718 | 15 | 1792 | 14 | 1879 | 24 | 0.91 | 1879 | 24 |
| 248 | 18222 | 5.59661 | 1.41 | 0.35181 | 1.00 | 0.71 | 1943 | 17 | 1916 | 12 | 1886 | 18 | 1.03 | 1886 | 18 |
| 608 | 58236 | 4.91793 | 1.53 | 0.30873 | 1.00 | 0.65 | 1734 | 15 | 1805 | 13 | 1888 | 21 | 0.92 | 1888 | 21 |
| 310 | 23863 | 5.26897 | 1.54 | 0.33056 | 1.17 | 0.76 | 1841 | 19 | 1864 | 13 | 1889 | 18 | 0.97 | 1889 | 18 |
| 387 | 47605 | 5.25551 | 1.41 | 0.32946 | 1.00 | 0.71 | 1836 | 16 | 1862 | 12 | 1891 | 18 | 0.97 | 1891 | 18 |
| 274 | 13449 | 4.41425 | 2.01 | 0.27639 | 1.74 | 0.87 | 1573 | 24 | 1715 | 17 | 1893 | 18 | 0.83 | 1893 | 18 |
| 130 | 13088 | 5.13824 | 1.71 | 0.32126 | 1.00 | 0.58 | 1796 | 16 | 1842 | 15 | 1895 | 25 | 0.95 | 1895 | 25 |
| 284 | 13672 | 5.25798 | 1.42 | 0.32796 | 1.00 | 0.71 | 1828 | 16 | 1862 | 12 | 1900 | 18 | 0.96 | 1900 | 18 |
| 583 | 17701 | 4.96598 | 1.41 | 0.30875 | 1.00 | 0.71 | 1735 | 15 | 1814 | 12 | 1906 | 18 | 0.91 | 1906 | 18 |
| 229 | 12301 | 4.85388 | 2.94 | 0.30156 | 2.76 | 0.94 | 1699 | 41 | 1794 | 25 | 1907 | 18 | 0.89 | 1907 | 18 |
| 692 | 32868 | 4.68083 | 2.05 | 0.29013 | 1.00 | 0.49 | 1642 | 14 | 1764 | 17 | 1911 | 32 | 0.86 | 1911 | 32 |


| 386 | 37021 | 5.20243 | 1.85 | 0.32138 | 1.29 | 0.70 | 1796 | 20 | 1853 | 16 | 1917 | 24 | 0.94 | 1917 | 24 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 388 | 27421 | 5.15771 | 1.41 | 0.31625 | 1.00 | 0.71 | 1771 | 15 | 1846 | 12 | 1930 | 18 | 0.92 | 1930 | 18 |
| 659 | 31468 | 5.24594 | 2.48 | 0.32093 | 2.27 | 0.92 | 1794 | 36 | 1860 | 21 | 1935 | 18 | 0.93 | 1935 | 18 |
| 245 | 24480 | 5.52233 | 2.60 | 0.33719 | 1.96 | 0.75 | 1873 | 32 | 1904 | 22 | 1938 | 30 | 0.97 | 1938 | 30 |
| 351 | 13978 | 4.97112 | 1.89 | 0.30279 | 1.45 | 0.77 | 1705 | 22 | 1814 | 16 | 1942 | 22 | 0.88 | 1942 | 22 |
| 583 | 24448 | 4.80202 | 1.59 | 0.29244 | 1.00 | 0.63 | 1654 | 15 | 1785 | 13 | 1943 | 22 | 0.85 | 1943 | 22 |
| 336 | 11124 | 4.57250 | 2.93 | 0.27689 | 2.53 | 0.86 | 1576 | 35 | 1744 | 24 | 1953 | 26 | 0.81 | 1953 | 26 |
| 329 | 18402 | 5.84084 | 1.74 | 0.34823 | 1.00 | 0.58 | 1926 | 17 | 1952 | 15 | 1981 | 25 | 0.97 | 1981 | 25 |
| 642 | 17007 | 4.60406 | 2.99 | 0.27366 | 2.82 | 0.94 | 1559 | 39 | 1750 | 25 | 1986 | 18 | 0.79 | 1986 | 18 |
| 425 | 42199 | 5.67492 | 2.49 | 0.33267 | 1.50 | 0.60 | 1851 | 24 | 1928 | 22 | 2011 | 35 | 0.92 | 2011 | 35 |
| 881 | 43557 | 5.07487 | 10.08 | 0.29484 | 10.00 | 0.99 | 1666 | 147 | 1832 | 86 | 2026 | 23 | 0.82 | 2026 | 23 |
| 148 | 8115 | 6.05948 | 2.03 | 0.34939 | 1.24 | 0.61 | 1932 | 21 | 1984 | 18 | 2040 | 28 | 0.95 | 2040 | 28 |
| 237 | 10304 | 6.47044 | 1.44 | 0.36429 | 1.00 | 0.69 | 2002 | 17 | 2042 | 13 | 2082 | 18 | 0.96 | 2082 | 18 |
| 295 | 10830 | 5.80291 | 1.65 | 0.31874 | 1.00 | 0.61 | 1784 | 16 | 1947 | 14 | 2125 | 23 | 0.84 | 2125 | 23 |
| 302 | 20350 | 7.10673 | 1.98 | 0.38919 | 1.31 | 0.66 | 2119 | 24 | 2125 | 18 | 2131 | 26 | 0.99 | 2131 | 26 |
| 368 | 27076 | 6.25070 | 2.47 | 0.32568 | 1.00 | 0.40 | 1817 | 16 | 2012 | 22 | 2217 | 39 | 0.82 | 2217 | 39 |
| 387 | 37889 | 7.33722 | 1.93 | 0.38204 | 1.47 | 0.76 | 2086 | 26 | 2153 | 17 | 2218 | 22 | 0.94 | 2218 | 22 |
| 123 | 9053 | 7.29063 | 2.05 | 0.37255 | 1.61 | 0.79 | 2041 | 28 | 2148 | 18 | 2251 | 22 | 0.91 | 2251 | 22 |
| 911 | 52374 | 7.94134 | 1.77 | 0.39906 | 1.00 | 0.57 | 2165 | 18 | 2224 | 16 | 2280 | 25 | 0.95 | 2280 | 25 |
| 343 | 31158 | 8.18101 | 1.64 | 0.39463 | 1.00 | 0.61 | 2144 | 18 | 2251 | 15 | 2350 | 22 | 0.91 | 2350 | 22 |
| 595 | 39802 | 9.44799 | 1.84 | 0.43936 | 1.21 | 0.66 | 2348 | 24 | 2383 | 17 | 2412 | 23 | 0.97 | 2412 | 23 |
| 465 | 33900 | 8.76030 | 2.01 | 0.39807 | 1.24 | 0.62 | 2160 | 23 | 2313 | 18 | 2451 | 27 | 0.88 | 2451 | 27 |
| 270 | 24911 | 10.10751 | 2.02 | 0.44799 | 1.47 | 0.73 | 2386 | 29 | 2445 | 19 | 2494 | 23 | 0.96 | 2494 | 23 |
| 559 | 35729 | 9.43780 | 1.90 | 0.41786 | 1.00 | 0.53 | 2251 | 19 | 2382 | 17 | 2495 | 27 | 0.90 | 2495 | 27 |
| 156 | 11343 | 9.23105 | 2.86 | 0.40652 | 2.40 | 0.84 | 2199 | 45 | 2361 | 26 | 2504 | 26 | 0.88 | 2504 | 26 |
| 420 | 27023 | 9.87075 | 3.48 | 0.43466 | 3.30 | 0.95 | 2327 | 64 | 2423 | 32 | 2505 | 18 | 0.93 | 2505 | 18 |
| 689 | 32722 | 8.46825 | 1.94 | 0.37056 | 1.39 | 0.72 | 2032 | 24 | 2283 | 18 | 2515 | 23 | 0.81 | 2515 | 23 |
| 547 | 24740 | 8.86558 | 3.53 | 0.38653 | 1.48 | 0.42 | 2107 | 27 | 2324 | 32 | 2521 | 54 | 0.84 | 2521 | 54 |
| 243 | 28930 | 11.13003 | 2.00 | 0.48477 | 1.73 | 0.87 | 2548 | 36 | 2534 | 19 | 2523 | 17 | 1.01 | 2523 | 17 |
| 335 | 43585 | 11.13759 | 1.90 | 0.47902 | 1.00 | 0.53 | 2523 | 21 | 2535 | 18 | 2544 | 27 | 0.99 | 2544 | 27 |
| 181 | 10794 | 9.95865 | 1.98 | 0.42814 | 1.00 | 0.50 | 2297 | 19 | 2431 | 18 | 2545 | 29 | 0.90 | 2545 | 29 |
| 159 | 18792 | 11.01834 | 2.31 | 0.47284 | 1.55 | 0.67 | 2496 | 32 | 2525 | 21 | 2548 | 29 | 0.98 | 2548 | 29 |
| 113 | 4921 | 9.97476 | 2.45 | 0.42745 | 1.00 | 0.41 | 2294 | 19 | 2432 | 23 | 2550 | 38 | 0.90 | 2550 | 38 |
| 218 | 22871 | 11.32644 | 1.74 | 0.48322 | 1.00 | 0.58 | 2541 | 21 | 2550 | 16 | 2558 | 24 | 0.99 | 2558 | 24 |
| 616 | 31558 | 10.87973 | 1.90 | 0.46169 | 1.03 | 0.54 | 2447 | 21 | 2513 | 18 | 2567 | 27 | 0.95 | 2567 | 27 |
| 359 | 36683 | 11.54088 | 1.47 | 0.48074 | 1.07 | 0.73 | 2530 | 22 | 2568 | 14 | 2598 | 17 | 0.97 | 2598 | 17 |
| 289 | 18468 | 11.01828 | 1.41 | 0.45848 | 1.00 | 0.71 | 2433 | 20 | 2525 | 13 | 2599 | 17 | 0.94 | 2599 | 17 |
| 63 | 12414 | 12.27508 | 1.42 | 0.49298 | 1.00 | 0.71 | 2584 | 21 | 2626 | 13 | 2658 | 17 | 0.97 | 2658 | 17 |
| 258 | 17974 | 11.69858 | 1.67 | 0.46248 | 1.00 | 0.60 | 2451 | 20 | 2581 | 16 | 2684 | 22 | 0.91 | 2684 | 22 |
| 678 | 33847 | 11.11985 | 5.24 | 0.43116 | 1.00 | 0.19 | 2311 | 19 | 2533 | 49 | 2716 | 85 | 0.85 | 2716 | 85 |
| 341 | 22083 | 12.92678 | 1.47 | 0.49243 | 1.00 | 0.68 | 2581 | 21 | 2674 | 14 | 2746 | 18 | 0.94 | 2746 | 18 |
| 278 | 41158 | 12.47589 | 3.33 | 0.46278 | 2.31 | 0.69 | 2452 | 47 | 2641 | 31 | 2789 | 39 | 0.88 | 2789 | 39 |
| 173 | 24233 | 13.62468 | 2.95 | 0.50499 | 1.00 | 0.34 | 2635 | 22 | 2724 | 28 | 2790 | 45 | 0.94 | 2790 | 45 |


| 303 | 37463 | 14.63241 | 2.76 | 0.52206 | 1.00 | 0.36 | 2708 | 22 | 2792 | 26 | 2853 | 42 | 0.95 | 2853 | 42 |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| 221 | 11445 | 12.86812 | 4.12 | 0.45530 | 1.03 | 0.25 | 2419 | 21 | 2670 | 39 | 2866 | 65 | 0.84 | 2866 | 65 |
| 202 | 21364 | 15.04269 | 4.40 | 0.52224 | 2.21 | 0.50 | 2709 | 49 | 2818 | 42 | 2897 | 62 | 0.93 | 2897 | 62 |
| 799 | 123872 | 16.49168 | 1.43 | 0.55313 | 1.00 | 0.70 | 2838 | 23 | 2906 | 14 | 2953 | 16 | 0.96 | 2953 | 16 |
| 217 | 13346 | 15.86391 | 2.42 | 0.51988 | 1.00 | 0.41 | 2699 | 22 | 2869 | 23 | 2990 | 35 | 0.90 | 2990 | 35 |
| 402 | 22745 | 14.60165 | 7.85 | 0.44958 | 7.28 | 0.93 | 2393 | 146 | 2790 | 75 | 3090 | 47 | 0.77 | 3090 | 47 |
| 272 | 35260 | 19.81526 | 1.41 | 0.60890 | 1.00 | 0.71 | 3066 | 24 | 3082 | 14 | 3093 | 16 | 0.99 | 3093 | 16 |
| 187 | 25411 | 20.59891 | 4.35 | 0.60533 | 3.32 | 0.76 | 3051 | 81 | 3120 | 42 | 3164 | 45 | 0.96 | 3164 | 45 |
| 541 | 79400 | 23.01154 | 2.39 | 0.63930 | 2.08 | 0.87 | 3186 | 52 | 3227 | 23 | 3253 | 19 | 0.98 | 3253 | 19 |

## Sample 406020 - Unidentified sample (Lat, Long $=\mathbf{2 8 . 3 8 1 8 7}{ }^{\circ}$, $83.80073^{\circ}$ )

| 617 | 23248 | 4.28844 | 3.15 | 0.28279 | 1.00 | 0.32 | 1605 | 14 | 1691 | 26 | 1799 | 54 | 0.89 | 1799 | 54 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 63 | 6768 | 4.85940 | 5.58 | 0.31303 | 4.67 | 0.84 | 1756 | 72 | 1795 | 47 | 1842 | 55 | 0.95 | 1842 | 55 |
| 334 | 22828 | 5.13569 | 2.14 | 0.32500 | 1.00 | 0.47 | 1814 | 16 | 1842 | 18 | 1874 | 34 | 0.97 | 1874 | 34 |
| 316 | 17533 | 5.02149 | 2.25 | 0.31757 | 1.00 | 0.45 | 1778 | 16 | 1823 | 19 | 1875 | 36 | 0.95 | 1875 | 36 |
| 1191 | 30436 | 3.98662 | 4.47 | 0.25088 | 2.19 | 0.49 | 1443 | 28 | 1631 | 36 | 1884 | 70 | 0.77 | 1884 | 70 |
| 1005 | 26602 | 5.00331 | 1.57 | 0.31286 | 1.00 | 0.64 | 1755 | 15 | 1820 | 13 | 1895 | 22 | 0.93 | 1895 | 22 |
| 586 | 32183 | 5.18701 | 1.66 | 0.32298 | 1.00 | 0.60 | 1804 | 16 | 1850 | 14 | 1903 | 24 | 0.95 | 1903 | 24 |
| 427 | 29994 | 5.28227 | 1.66 | 0.32877 | 1.00 | 0.60 | 1832 | 16 | 1866 | 14 | 1904 | 24 | 0.96 | 1904 | 24 |
| 481 | 20472 | 4.72198 | 1.65 | 0.29350 | 1.00 | 0.61 | 1659 | 15 | 1771 | 14 | 1906 | 24 | 0.87 | 1906 | 24 |
| 771 | 46531 | 5.17165 | 1.75 | 0.32116 | 1.00 | 0.57 | 1795 | 16 | 1848 | 15 | 1908 | 26 | 0.94 | 1908 | 26 |
| 375 | 22937 | 5.28125 | 1.67 | 0.32277 | 1.00 | 0.60 | 1803 | 16 | 1866 | 14 | 1936 | 24 | 0.93 | 1936 | 24 |
| 387 | 15720 | 5.11488 | 2.42 | 0.30936 | 1.98 | 0.82 | 1738 | 30 | 1839 | 21 | 1955 | 25 | 0.89 | 1955 | 25 |
| 502 | 25032 | 5.54236 | 2.39 | 0.33140 | 1.64 | 0.69 | 1845 | 26 | 1907 | 21 | 1975 | 31 | 0.93 | 1975 | 31 |
| 184 | 8168 | 6.09649 | 2.68 | 0.35818 | 1.00 | 0.37 | 1974 | 17 | 1990 | 23 | 2007 | 44 | 0.98 | 2007 | 44 |
| 469 | 30798 | 6.39960 | 1.96 | 0.35855 | 1.00 | 0.51 | 1975 | 17 | 2032 | 17 | 2090 | 30 | 0.94 | 2090 | 30 |
| 43 | 6973 | 7.19554 | 2.47 | 0.40135 | 1.93 | 0.78 | 2175 | 36 | 2136 | 22 | 2098 | 27 | 1.04 | 2098 | 27 |
| 425 | 27362 | 6.54321 | 1.75 | 0.36382 | 1.00 | 0.57 | 2000 | 17 | 2052 | 15 | 2104 | 25 | 0.95 | 2104 | 25 |
| 436 | 24482 | 6.68588 | 5.49 | 0.36305 | 1.07 | 0.20 | 1997 | 18 | 2071 | 48 | 2145 | 94 | 0.93 | 2145 | 94 |
| 468 | 21485 | 6.37485 | 4.81 | 0.34496 | 4.12 | 0.86 | 1910 | 68 | 2029 | 42 | 2151 | 43 | 0.89 | 2151 | 43 |
| 398 | 21951 | 6.03807 | 3.66 | 0.32162 | 3.52 | 0.96 | 1798 | 55 | 1981 | 32 | 2179 | 17 | 0.82 | 2179 | 17 |
| 157 | 17146 | 7.70224 | 1.63 | 0.40649 | 1.00 | 0.61 | 2199 | 19 | 2197 | 15 | 2195 | 22 | 1.00 | 2195 | 22 |
| 206 | 10299 | 7.51050 | 1.47 | 0.39293 | 1.00 | 0.68 | 2136 | 18 | 2174 | 13 | 2210 | 19 | 0.97 | 2210 | 19 |
| 411 | 29052 | 9.00424 | 1.67 | 0.42619 | 1.00 | 0.60 | 2289 | 19 | 2338 | 15 | 2382 | 23 | 0.96 | 2382 | 23 |
| 314 | 14477 | 6.94502 | 4.01 | 0.32781 | 3.75 | 0.93 | 1828 | 60 | 2104 | 36 | 2387 | 24 | 0.77 | 2387 | 24 |
| 1353 | 75914 | 8.56405 | 3.28 | 0.38792 | 2.87 | 0.87 | 2113 | 52 | 2293 | 30 | 2457 | 27 | 0.86 | 2457 | 27 |
| 785 | 59391 | 9.83780 | 1.53 | 0.44446 | 1.00 | 0.65 | 2371 | 20 | 2420 | 14 | 2461 | 20 | 0.96 | 2461 | 20 |
| 457 | 38351 | 9.78517 | 2.39 | 0.44013 | 1.00 | 0.42 | 2351 | 20 | 2415 | 22 | 2469 | 37 | 0.95 | 2469 | 37 |
| 70 | 8122 | 10.96234 | 3.13 | 0.48103 | 2.29 | 0.73 | 2532 | 48 | 2520 | 29 | 2510 | 36 | 1.01 | 2510 | 36 |
| 214 | 18883 | 10.86886 | 1.50 | 0.47416 | 1.00 | 0.67 | 2502 | 21 | 2512 | 14 | 2520 | 19 | 0.99 | 2520 | 19 |
| 470 | 30679 | 10.22232 | 1.67 | 0.44488 | 1.00 | 0.60 | 2372 | 20 | 2455 | 15 | 2524 | 23 | 0.94 | 2524 | 23 |
| 504 | 44632 | 11.68696 | 1.44 | 0.48998 | 1.00 | 0.69 | 2571 | 21 | 2580 | 13 | 2587 | 17 | 0.99 | 2587 | 17 |
| 235 | 25125 | 12.65721 | 2.74 | 0.49362 | 1.18 | 0.43 | 2586 | 25 | 2654 | 26 | 2707 | 41 | 0.96 | 2707 | 41 |


| 400 | 30642 | 13.27228 | 2.31 | 0.51182 | 1.45 | 0.63 | 2664 | 32 | 2699 | 22 | 2725 | 30 | 0.98 | 2725 | 30 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 364 | 36858 | 13.87242 | 1.77 | 0.52393 | 1.00 | 0.57 | 2716 | 22 | 2741 | 17 | 2760 | 24 | 0.98 | 2760 | 24 |
| $*=$ radiogenic Pb. |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |

$U$ concentration has an uncertainty of $\sim 25 \%$.

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