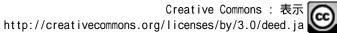
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Effects of bootstrap current on magnetic configuration in Chinese first quasi-axisymmetric stellarator

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Abstract

Neoclassical properties in quasi-axisymmetric (QA) stellarators are analogous to these in tokamaks. Consequently, a substantial bootstrap current could significantly modify the MHD equilibrium properties of a QA stellarator, which is an important characteristic in this type of stellarator. This paper is dedicated to systemically investigate the effects of bootstrap current on the magnetic configuration in Chinese first quasi-axisymmetric stellarator (CFQS). For the first time, self-consistent bootstrap currents in free-boundary equilibria are calculated with an accurate Fokker-Planck neoclassical numerical mode in CFQS. Several important results are achieved: (a) as the bootstrap current grows with increasing volume-averaged normalized pressure β , magnetic shear develops in the bulk plasma and meanwhile, a deep magnetic well is robustly sustained, which leads to improved stabilization of interchange modes up to $\beta \sim 2.0\%$. (b) In the analysis of global ideal MHD instability, as the bootstrap current rises to 39 kA $(\beta \sim 1.3\%)$, external kink modes become destabilized and the unstable mode with m/n = 2/1 is dominant. (c) From $\beta = 0$ to 1.5%, the bootstrap current hardly changes the OA property and a low neoclassical transport is maintained. However, as β is enhanced beyond 2.0%, the substantial bootstrap current gives rise to an increase of non-QA magnetic field components, which weakens the neoclassical transport properties. (d) An increase of the negative magnetic shear at the core region by the bootstrap currents has a favorable effect on the properties of J(second adiabatic invariant). The maximum-J region can be extended by raising bootstrap currents.

Keywords: quasi-axisymmetric stellarator, bootstrap current, MHD stabilities, neoclassical transport, micro-instability

(Some figures may appear in colour only in the online journal)

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1. Introduction

An essential optimization strategy for a stellarator is to achieve omnigenity [1], meaning the particle radial drift vanishes on a time average [2]. In quasi-symmetric configurations the omnigenious property can be realized. The quasi-symmetric configurations include quasi-axisymmetric (QA), quasi-helical symmetric and quasi-poloidal symmetric magnetic field topologies [3]. To date, the QA configuration has been accomplished in many physics designs e.g. CHS-qa [4], NCSX [5], ESTELL [6], etc, none of which, however, has been successfully fabricated and operated. The Chinese first quasi-axisymmetric stellarator (CFQS) will represent the first operational quasiaxially symmetric stellarator in the world, which has been constructed as a collaborative project between the National Institute for Fusion Science and Southwest Jiaotong University to prove the intrinsic advantages of quasi-axisymmetry [7–11]. The principal parameters of CFQS are as follows: the major radius is 1.0 m, the magnetic field strength is 1.0 T, the aspect ratio is 4.0, and the toroidal periodic number is 2 [7]. A magnetic field coil system consists of 16 modular coils (MCs), 12 toroidal field coils (TFCs), and 4 poloidal field coils (PFCs). It has been designed to possess a number of advanced features in fixed and free-boundary equilibria. The MHD equilibrium of the CFQS configuration is stable up to $\beta \sim 1.1\%$ in the plasma-current-free scenario [9]. Rotational transform profiles with weak magnetic shear were always sustained, which did not contribute to stabilization of Mercier modes. The resistive ballooning modes were simulated in CFQS equilibria with magnetic islands [10]. For the high-resistivity regimes, the modes were unstable, whereas in the low-resistivity regimes, the modes were stabilized. With bootstrap current effects, in $\beta \sim 1.0\%$ plasma the low-order islands (2/4) were generated, which may lead to a stochastization of magnetic field lines [10]. Moreover, the effects of the bootstrap current on neoclassical transport were investigated and a satisfactory neoclassical transport maintained up to $\beta = 1.5\%$ [11]. Now the fabrication of the CFQS device is steadily progressing to achieve its first plasma [12]. It is noted in previous work [9-12], that the bootstrap current was estimated with the BOOTSJ code, which uses a semi-analytic formula in the collision-less limit [13]. The bootstrap current density predicted with this code is usually noisy or has 'spikes' near the rational surfaces [10, 14], which results from the shortcoming of the low-collisionality bootstrap analytical formula. Even though these spikes can be smoothed over with a damping parameter, the current profile still does not closely match the accurate Fokker-Planck neoclassical numerical results in high β plasma [14].

The bootstrap current is a net toroidal current driven by the temperature and density gradients, which was predicted and calculated in tokamak geometry five decades ago by Galeev [15] and Bickerton, Connor, and Taylor [16]. A proper estimation of the bootstrap current in a stellarator configuration was first done by Shaing and Callen [17], using the Hirshman-Sigmar moment method, which was improved by a more direct calculation by Boozer and Gardner [18]. Both in tokamaks and stellarators the bootstrap current has been observed experimentally [19, 20]. Since then, it has proven to be of great importance to both tokamaks and stellarators [21, 22], especially in QA stellarators [10, 23]. A new analytic expression of bootstrap current in a low-collisionality stellarator plasma was given by Helander [24], which was suitable for all low-collisionality regimes. The QA configuration has neoclassical properties that are similar to those in tokamaks [25], since the physics of the bootstrap current is the same in the Boozer coordinates, which indicates this current may significantly change the equilibrium properties in a QA configuration. Previous studies in NCSX [26, 27] and CHS-qa [28, 29] had shown that large bootstrap currents dramatically impacted the rotational transform profile, magnetic shear, MHD modes, etc. Recently, a magnetic configuration (a 1 Tesla mean field) with precise QA for plasma confinement has been achieved in which the axisymmetry-breaking fields throughout the torus of aspect ratio 6 are as small as the 50 μ T geomagnetic field [30].

In addition, a discrepancy between accurate Fokker–Planck neoclassical numerical results and the analytical model, for higher collisionality, was identified in the studies by Koh *et al* [31] and Landreman and Ernst [32]. In W7-X experiments, the neoclassical effects were frequently observed and studied [33], which could be precisely predicted with the code SFINCS (the Stellarator Fokker–Planck Iterative Neoclassical Conservative Solver). This code solves the steady-state drift-kinetic equation for multiple species, allowing arbitrary collisionality, and using the full linearized Fokker–Planck–Landau collision operator. In this paper, we use the SFINCS code to accurately calculate bootstrap currents in CFQS and systematically study bootstrap current effects on MHD instabilities and plasma transport behaviors.

2. MHD equilibria with bootstrap currents

The simulation of 3D ideal MHD equilibria is executed using the free-boundary VMEC code [34]. The code approaches an equilibrium state by minimization of the plasma energy. The bootstrap current is estimated with the SFINCS code. A convergent iteration operation between VMEC and SFINCS codes is performed to self-consistently calculate the bootstrap current in CFQS. After several iterations a self-consistent equilibrium state is obtained. The plasma β is scanned from 0% to 2.03%. Problems of convergence towards an equilibrium state with VMEC arise partially due to the large Shafranov shift for $\beta > 2.03\%$. The plasma temperature and density profiles are assumed as $T = T_0 (1 - \rho^2)$ and $n_{\rm e} = n_{\rm i} = n_0 \ (1 - \rho^2)$, respectively and set $T_{\rm e0} \ / T_{\rm i0} = 3/2$. ρ represents the normalized minor radius. n_0 is scanned form 0 to $4.5 \times 10^{19} \text{m}^{-3}$ and T_{i0} is fixed at 1.3 keV. The dependence of rotational transform (ι) on bootstrap currents and bootstrap current density as a function of the normalized minor radius are shown in figures 1(a) and (b), respectively. As the plasma β increases to 2.03% the bootstrap current is enhanced

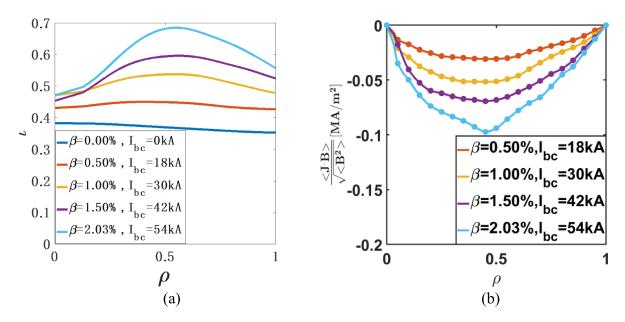


Figure 1. Profiles of (*a*) rotational transform and (*b*) bootstrap current density for $\beta = 0.00\%$, 0.50 %, 1.00 %, 1.50% and 2.03% with plasma bootstrap current = 0, 15, 30, 42 and 54 kA, respectively. These results are obtained with the self-consistent iterations between SFINCS and VMEC codes.

to 54 kA. In the regions with ρ from 0.4 to 0.6 a peak exists in the profile of the bootstrap current density, which results in generation of magnetic shear in the bulk plasma. This could be beneficial for the stabilization of MHD modes, e.g. interchange modes. Meanwhile, the rotational transform crosses the low-order rational surfaces with n/m = 1/2 as β increases to 1.0%, which may lead to the generation of magnetic islands.

The impact of increasing β and bootstrap currents on the equilibria is shown in figure 2 which displays cross sections of magnetic flux surfaces with the bean shape in equilibrium states with various bootstrap currents and corresponding magnetic field strength spectra $B_{m,n}$ normalized by $B_{0,0}$ in the Boozer coordinates. In order to demonstrate clearly the emergence of non-axisymmetric magnetic field components, the dominant component $B_{1,0}$ with a large amplitude is not shown. Figures 2(a)-(c) show as β is increased, that the magnetic axis is shifted horizontally by the net toroidal current because the effective vertical field is changed. Meanwhile, the flux surfaces become deformed and compressed towards the outside of the torus. The core region develops a $B_{0,1}$ symmetry breaking component from $\beta > 1\%$ in figure 2(*e*) and from $\beta > 2\%$ the $B_{2,1}$ term becomes significant in figure 2(f), which indicates as β increases beyond 2.0%, the deformation of magnetic flux surfaces may degrade the quasi-axisymmetry of the magnetic configuration in the core. This phenomenon is further studied in section 4.

3. Effects of bootstrap current on MHD stability

The Mercier criterion and global linear MHD stability are estimated with the VMEC and TERPSICHORE codes[34, 35]. The Mercier criterion for the stable interchange modes is expressed by $D_{\text{merc}} = D_{\text{shear}} + D_{\text{well}} + D_{\text{curr}} + D_{\text{geod}} > 0$ where they represent magnetic shear term, magnetic well term, current term and geodesic curvature term, respectively [36]. At three different radial positions, the variations of Mercier stabilities with β are given in figure 3(a), which illustrates that the interchange modes are stable up to $\beta = 2.03\%$ at least. What is interesting is that in the equilibria with bootstrap currents the stabilization of interchange modes is attributed to the enhancement of magnetic shear and the magnetic well, as shown in figure 3(b). These two stabilization effects enable to robustly suppress the destabilization effect of the geodesic curvature term. The previous study [10] found that free boundary equilibria without bootstrap currents, they were also stable when $\beta \leq 2.0\%$ in CFOS. However in that case, a flat ι profile was well sustained as β increased. Thus, interchange modes were merely stabilized by magnetic well structures. In addition, the stabilization effect of a net toroidal current on interchange modes in heliotron plasmas was also observed [37]. The resistive ballooning mode was also studied in CFQS [10] and its stability is dependent on the MHD resistivity. If the resistivity is not large, the ballooning modes can be stabilized well.

A global linear MHD stability analysis of CFQS equilibria is computed with respect to n = 1 and n = 2 family of modes with the TERPSICHORE code, which evaluates the variational energy principle. Figure 4 displays bootstrap currents as a function of β and it is approximately proportional to β . The ranges of Fourier modes for the global MHD stability analysis are m < 49 and n < 14 for the poloidal and toroidal modes, respectively. The typical marginal stability is represented by a threshold value of $\lambda = -1 \times 10^{-4}$ and more negative eigenvalues than this crucial number are thought to be more unstable [38]. The simulation results reveal that a global MHD stability is obtained for β up to 1.35% ($I_{bc} = 38$ kA).

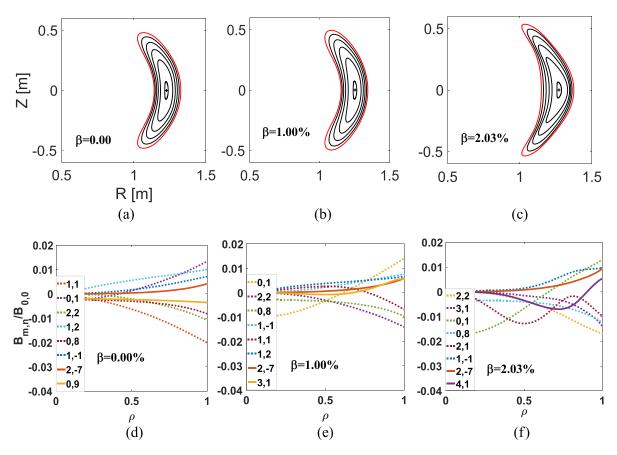


Figure 2. Cross sections of magnetic flux surfaces with the bean shape and magnetic field strength spectra with $n \neq 0$ normalized by $B_{0,0}$ in the Boozer coordinates for three of the cases: (a) and (d) for $\beta = 0.00\%$, (b) and (d) for $\beta = 1.00\%$ with a 30 kA bootstrap current, and (c) and (f) for $\beta = 2.03\%$ with a 54 kA bootstrap current.

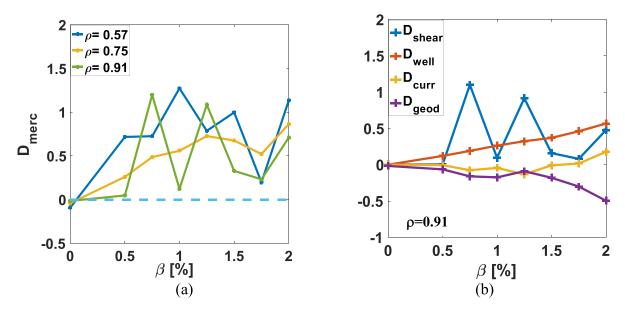


Figure 3. (a) Mercier stabilities versus β at three different radial positions $\rho = 0.57$, 0.75 and 0.91 respectively; (b) magnetic shear, magnetic well, current and geodesic curvature terms versus β at $\rho = 0.91$ in various free-boundary equilibria with bootstrap currents.

The stability is determined by the rotational transform at the edge, i.e. the iota value higher than 0.5 is unstable. As β increases beyond 1.35%, global instabilities arise. As an instance, the Fourier amplitudes of the radial component of the

displacement vector (ξ^{s}) profiles are displayed in figure 5(*a*). In this equilibrium the plasma β and bootstrap current are 1.75% and 46 kA, respectively. The unstable mode structure is dominated by the m/n = 2/1 term with the eigenvalue

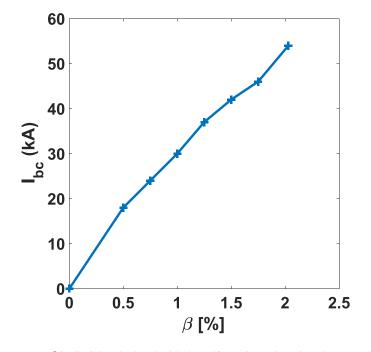


Figure 4. Bootstrap currents versus β in CFQS, calculated with the self-consistent iterations between SFINCS and VMEC codes.

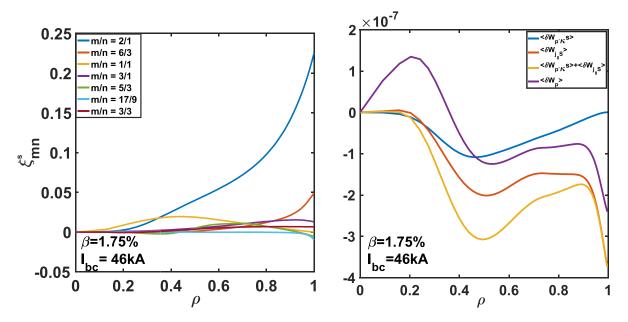


Figure 5. Fourier amplitudes of the radial component of the displacement vector $\xi^{s}(a)$ and surface-averaged plasma potential energy δW_{p} , ballooning mode driven energy $\delta W_{p_{1}\kappa^{s}}(b)$ as a function of ρ for $\beta = 1.75\%$ and $I_{bc} = 46$ kA. These results are calculated with TERPSICHORE.

 $\lambda = -3.8724 \times 10^{-4}$. In order to identify the category of these unstable modes, figure 5(*b*) shows radial profiles of surfaceaveraged plasma potential energy δW_p which mainly includes two instability drive terms, $\delta W_{p',\kappa^s}$ and $\delta W_{j_{\parallel}\kappa^s}$ [38, 39]. The former drive term associates with ballooning modes in which the plasma pressure gradient interacts with the curvature of magnetic field lines. The latter describes kink modes which are caused by the parallel current density interacting with the local magnetic shear. In the unstable regions with $\delta W_p < 0$, the second term completely exceeds the first term. Hence, the 2/1 mode is mainly current-driven external kink mode.

4. Effects of bootstrap current on neoclassical transport and J

The neoclassical transport study of the quasiaxisymmetric stellarator is implemented to evaluate how close it is to *a* axisymmetric configuration during increasing plasma bootstrap currents. The estimation of the neoclassical transport can be characterized by the effective helical ripple strength $\varepsilon_{\rm eff}[40]$ and neoclassical transport matrix [41]. In the $1/\nu$ regime the neoclassical diffusion coefficient is proportional to $\varepsilon_{\rm eff}^{3/2}$ which is investigated with the NEO code [40]. The NEO code solves

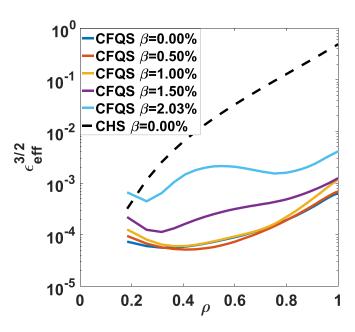


Figure 6. Radial profiles of effective helical ripple calculated with NEO for various β in the free-boundary equilibria with bootstrap currents. The CHS result is shown as a reference, plotted by a black dashed curve.

for the radial profile of $\varepsilon_{\rm eff}^{3/2}$ using an analytic solution of the banana kinetic equation. Figure 6 shows the radial profiles of $\varepsilon_{\rm eff}^{3/2}$ in the free boundary equilibria with various β . As a comparison with CFQS, the $\varepsilon_{\rm eff}^{3/2}$ profile in the fixed boundary equilibrium of CHS [42] is given by a black dashed curve. From $\beta = 0$ to 1.50%, the $\varepsilon_{\rm eff}^{3/2}$ in the CFQS configurations is approximately two orders lower than that in CHS, which is in good agreement with A.Shimizu's, et.al results [11]. In their study, they used a flatter density profile to estimate the bootstrap current. However, as β increases beyond 2.0% the $\varepsilon_{\rm eff}^{3/2}$ gets enhanced notably in the core region due to an increase of non-axisymmetric components of *B*, e.g. $B_{0,1}, B_{2,1}$, etc, which results from the deformation of magnetic flux surfaces by high β and bootstrap currents, as displayed in figures 2(*c*) and (*f*).

Furthermore, in order to reveal impacts of bootstrap currents on neoclassical transport in various plasma collisional regimes, the transport matrix is computed with the SFINCS code with the full linearized Fokker-Planck-Landau collision operator. In this collision operator, the momentum conservation property is adequately sustained, which can more accurately describe the collisional effects than the pitch-angle scattering collision operator [43]. Figure 7 shows two transport matrix elements $L_{1,1}$ and $L_{1,2}$ as a function of ν^* for the CFQS and CHS geometries at half-radius. The quantities $L_{1,1}$, $L_{1,2}$ and ν^* are defined by formula 40 and 41 in [43], which represent the radial particle and heat transport coefficients and normalized collisionality, respectively. Here the ambipolar electric field on transport coefficients is not considered. At high collisionality $\nu^* > 10^\circ$, $L_{1,1}$ and $L_{1,2}$ are quite low and basically independent of the toroidal current in CFQS. The neoclassical transport in CHS is comparable with that in CFQS. In the rare collision regime $\nu^* < 10^{-1}$, however, as the current increases to 54 kA ($\beta \sim 2.0\%$) these two coefficients are approximately one order of magnitude higher than that in low bootstrap current equilibria in CFQS, which indicate a degradation of neoclassical transport properties in high bootstrap current equilibria. Overall, when $\beta 1.5\%$ the neoclassical transport in CFQS is much lower than that in CHS and the quasi-axisymmetry is robustly sustained. This is in good agreement with the results from NEO in figure 6. The operational parameters of CFQS were estimated using the ISS95 scaling law [9, 12]. With a heating power of 1.0 MW, the β ranged from 1.0% to 2.0% in high-density plasmas. Our results are also consistent with the expected operation.

Quasi-axisymmetry makes a stellarator as analogous as possible to a tokamak, in some sense, which is not an essential condition for achieving good plasma confinement. Mathematically, the necessary condition is to require parallel adiabatic invariant $J = \oint m v_{\parallel} dl$ to be (approximately) constant on flux surfaces for trapped particle orbits [2]. If the magnetic field has been written as $\boldsymbol{B} = \nabla \psi \times \nabla \alpha$ and ψ measures the toroidal flux and $\alpha = \theta - \iota \zeta$ labels the different field lines on each flux surface, a configuration with vanishing bounceaveraged drift, $\partial J/\partial \alpha = 0$, is called omnigenous. Meanwhile, J decreases away from the center, i.e. $\partial J/\partial \psi < 0$, it is the so-called the maximum-J condition [44], which is very beneficial for stabilizing trapped-particle-driven micro-instability modes [45, 46]. The perfectly omnigenous configurations with the maximum-J property cannot be exactly achieved. Alternatively, the wave vector perpendicular to the magnetic field is expressed by $\boldsymbol{k}_{\perp} = k_{\psi} \nabla \psi + k_{\alpha} \nabla \alpha$ and if k_{ψ} is very small, $\partial J/\partial \psi < 0$ is able to suppress these micro instabilities in incompletely omnigenous configurations at least [46, 47]. The maximum-J capability has been broadly investigated in QA and quasi-isodynamic configurations [29, 45, 48]. It is of importance to evaluate the maximum-J capability in CFQS.

To obtain the distribution of J, the guiding center trajectories of tracer particles are calculated in Boozer coordinates, which are reflected at the same B. The motion equations used in the simulation were expressed in [49]. The tracing particle energy is W = 10 eV and set the magnetic field strength at the bounce point equal to 0.95 T. The initial positions of these particles are distributed in the $(\rho, \zeta, \theta = 0)$ plane. The J contours on the $(\rho, \zeta/\pi)$ plane are calculated in various equilibria, as shown in figure 8. The J contour values are normalized to the largest J. The color bar represents the normalized J values. The areas on the left side of the red solid curve represent the maximum-J criterion satisfied. In low bootstrap current equilibria with $\beta = 0.3\%$, the maximum-J region can hardly be generated in figure 8(a). As the bootstrap current increases, the negative magnetic shear in the core region is enhanced in figure 1(a), leading to a decrease of J away from the center, which is to extend maximum-J regions in figures 8(b)and (c). A similar result was also observed in reversed shear tokamaks [50] that the maximum-J region coincides with the region of the reversed q profile (negative shear). Meanwhile,

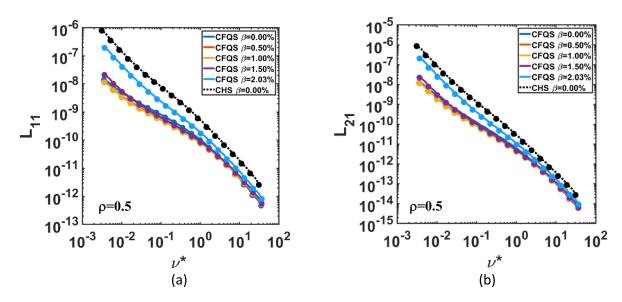


Figure 7. The radial particle and heat transport coefficients $L_{1,1}$ and $L_{2,1}$ as a function of the normalized collisionality in CFQS for various β in the free-boundary equilibria with bootstrap current. The CHS result is shown as a reference, plotted by a black curve. These results are calculated with SFINCS.

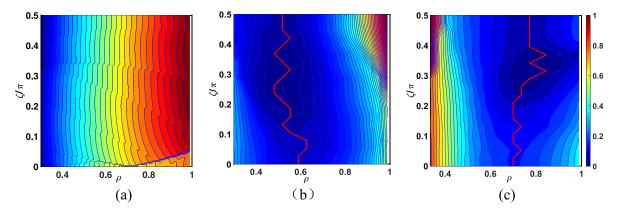


Figure 8. Contours of normalized J on the $(\rho, \zeta/\pi)$ plane for $(\beta, I_{bc}) = (0.30\%, 11 \text{ kA})$ (a), (0.75%, 24 kA) (b) and (1.50%, 42 kA) (c) in CFQS configurations, respectively. The J is normalized to the largest J. The areas on the left side of the red solid curve represent the maximum-J criterion satisfied.

with increasing β , it is found that $dB_{00}/d\psi$ decreases monotonically with radius, which could make poloidal drift decrease with radius. This process is also beneficial for obtaining $\partial J/\partial \psi < 0$. Notably, an inhomogeneity of maximum-*J* regions in the ζ direction arises due to the existence of the non-axisymmetry of B.

To assess the capability of the maximum-*J* in CFQS, linear gyrokinetic simulations using the GENE code [51] are performed to explore electrostatic trapped electron modes (TEMs) in these equilibria. The code GENE solves the gyrokinetic equation along with Maxwell's equations. In this study, both electrons and ions are kinetically treated with a resolution of $n_x \times n_{ky} \times n_z \times n_v \times n_w = 30 \times 1 \times 128 \times 64 \times 16$. The wave number $k_y \rho_i$ ($k_y = k_\alpha B_0 (d\rho/d\psi)/\sqrt{2}$) is scanned from 0.5 to 2.0 with an interval = 0.1. The local flux-tube is fixed at the position at (ρ , α) = (0.5,0). The stability study is simulated with the same temperature for ions and electrons $T_i = T_e$, while setting the ion-temperature normalized gradient a/L_{Ti} and the density normalized gradient a/L_n to zero and the electron-temperature normalized gradient $a/L_{\text{Te}} = 7$. Figure 9 shows the growth rate spectrum of electrostatic TEMs in three equilibria with $\beta = 0.0\%$, 0.75%, and 1.50%, respectively. The propagating direction of these modes is the electron diamagnetic drift direction. In the plasma-pressure-free equilibrium, the TEMs are significantly excited, displayed by the blue curve. As β increases to 0.75%, weak maximum-J regions are produced in the core, which slightly mitigate the TEMs, denoted by the red curve. In the equilibrium with $\beta = 1.50\%$, the development of the core maximum-J becomes prominent, which remarkably suppresses the TEMs, as shown by the black curve. This numerical result is consistent with the above analytical analysis that even though the perfect omnigeneity is not achieved in CFQS, the TEMs in maximum-J regions can be effectively mitigated. In addition, in low β plasmas a stability-valley-like structure was found along the transition boundary of the ion temperature

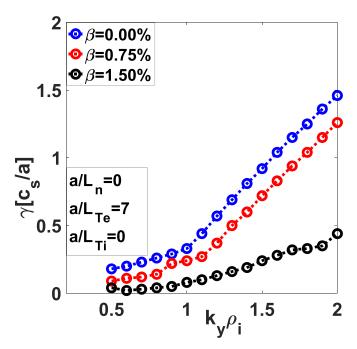


Figure 9. The growth rates of electrostatic TEM modes as a function of $k_y \rho_i$ in three equilibra with $\beta = 0.0\%$, $\beta = 0.75\%$ and 1.50%, respectively.

gradient mode and TEM, where the growth rate is quite low [52].

The magnetic islands and chaos could modify the properties of MHD equilibria, which may result in enhanced radial transport [53]. In CFQS the magnetic islands emerge in edge plasma region at $\beta \sim 1.0\%$. This is due to the rotational transform crossing the rational surface with $\iota = 2/4 = 1/2$. An equilibrium with 2/4 magnetic islands was found in the simulation results with the HINT code [10]. 12 planar TFC have been designed to adjust the rotational transform profile [54], healing magnetic islands. Moreover, active control of islands is also available by electron cyclotron heating and tailoring the pressure profile [55, 56]. In the appendix we give a case that in the equilibrium with $\beta = 1.0\%$, the 2/4 magnetic islands are fully suppressed by TFC. How to control magnetic islands is an interesting and important research topic, which will be further studied in future work.

5. Conclusion

Characteristics of MHD activities and plasma transport affected by self-consistent bootstrap currents are studied in free boundary equilibria of CFQS. The bootstrap currents are calculated accurately with the SFINCS code. The volumeaveraged normalized pressure is scanned from 0.00% to 2.03% and the bootstrap current is enhanced to 54 kA.

The interchange modes are stabilized by the magnetic shear and magnetic well effects, which are able to suppress the unfavorable effect of the geodesic curvature till $\beta = 2.03\%$. In the free boundary equilibria without bootstrap currents, interchange modes were also stable up to $\beta = 2.0\%$ in CFQS [9]. However, the magnetic shear has not been produced as β increased and these modes were merely stabilized by a magnetic well structure. Furthermore, with β exceeding 1.35%, the current-driven kink modes become unstable and the dominant mode is m/n = 2/1. This instability is sensitive to the rotational transform at the edge and as the iota value is higher than 0.5, this instability arises. Such a dependence was also found in the CHS-qa configuration [29].

From $\beta = 0$ to 1.5%, the $\varepsilon_{eff}^{3/2}$, $L_{1,1}$ and $L_{1,2}$ in the CFQS configurations are quite low and the QA property is sustained well. Whereas, as plasma β increases to 2.0%, these neoclassical transport coefficients grow significantly in the core regions, which result from an increase of non-axisymmetric magnetic field components due to unfavorable deformations of the QA configuration.

Additionally, the extension of maximum-*J* regions can be achieved by increasing bootstrap currents, which is beneficial for mitigating trapped-electron-driven microinstabilities. The numerical result is consistent with the analytical studies.

In conclusion, a stable plasma with good confinement properties in $\beta \sim 1.35\%$ equilibria is expected to be experimentally accomplishable in CFQS.

Acknowledgments

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Appendix

In order to assess the possible formation of magnetic islands and chaotic magnetic fields in CFQS, we apply the multiregion relaxed MHD (MRXMHD) framework based on the Taylor's relaxation theory [57] to develop MHD equilibria with the stepped-pressure equilibrium code (SPEC) [58]. SPEC allows the plasma to undergo possible magnetic reconnection events that would minimize the plasma

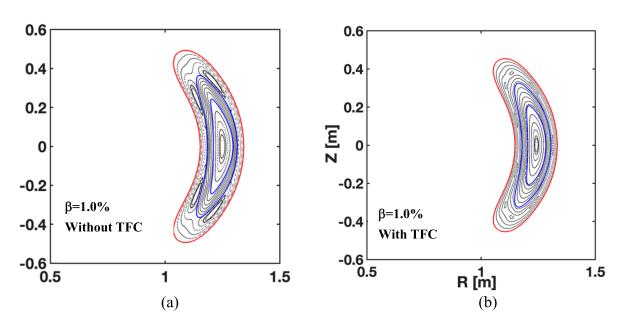


Figure A1. (a) Without and (b) with considering TFC, Poincaré plots of the poloidal cross section with the bean shape in the MHD equilibria with $\beta = 1.0\%$. The ratio of the current in TFC to that in MC $I_{\text{TFC}}/I_{\text{MC}} = 15\%$.

potential energy. In this work, we consider two interfaces so that the entire plasma volume is divided into three distinct sub-volumes in CFQS equilibria. Given that both relaxed regions and ideal interfaces exist in MRXMHD, two types of toroidal currents coexist, namely currents flowing in sub-volumes and current sheets in ideal interfaces [59]. Meanwhile, the spatially distributed current density in each sub-volume can change self-consistently with variations of magnetic islands and stochasticities. The stepped pressure profile is interpolated from the VMEC output. Figure A1 displays Poincaré sections of the equilibrium magnetic fields calculated from the SPEC code with increasing values of β and self-consistent bootstrap currents. The magnetic islands emerge in the edge plasma domain at around $\beta = 1.0\%$ in figure A1(a). This is due to the rotational transform crossing the rational surface with $\iota = 2/4 = 1/2$ (see figure 1). In the CFQS device, 12 planar TFC have been designed to adjust the rotational transform profile, healing magnetic islands. With the ratio of the current in TFC to that in MC, $I_{TFC}/I_{MC} = 15\%$, a strong resilience against magnetic islands and chaos is produced and the 2/4 magnetic islands are well suppressed in figure A1(b).

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References

- Nührenberg J., Lotz W. and Gori S. 1994 Quasi-axisymmetric tokamaks theory of fusion plasmas, Varenna, 1994 (Bologna: Editrice Compositori) p 3
- [2] Hall L.S. and McNamara B. 1975 Phys. Fluids 18 552
- [3] Landreman M. and Catto P.J. 2012 Phys. Plasmas 19 056103
- [4] Okamura S. et al 2001 Nucl. Fusion 41 1865
- [5] Nelson B.E. et al 2003 Fusion Eng. Des. 66 169-74
- [6] Drevlak M., Brochard F., Helander P., Kisslinger J., Mikhailov M., Nührenberg C., Nührenberg J. and Turkin Y. 2013 Contrib. Plasma Phys. 53 459–68
- [7] Liu H. et al CFQS Team 2018 Plasma Fusion Res. 13 3405067
- [8] Isobe M. et al CFQS Team 2019 Plasma Fusion Res.
- **14** 3402074
- [9] Liu H. et al 2021 Nucl. Fusion **61** 016014
- [10] Wang X. et al 2021 Nucl. Fusion **61** 036021
- [11] Shimizu A. *et al* 2018 *Plasma Fusion Res.* **13** 3403123
- [12] Shimizu A. *et al* 2022 *Nucl. Fusion* **62** 016010
- [13] Shaing K.C., Crume E.C., Tolliver J.S., Hirshman S.P. and van Rij W.I. 1989 *Phys. Fluids* B 1 148
- [14] Landreman M., Buller S. and Drevlak M. 2022 Phys. Plasmas 29 082501
- [15] Galeev A.A. 1971 Sov. Phys.-JETP 32 752
- [16] Bickerton R., Connor J.W. and Taylor J.B. 1971 Nat. Phys. Sci. 229 110
- [17] Shaing K.C. and Callen J.D. 1983 Phys. Fluids 26 3315
- [18] Boozer A.H. and Gardner H.J. 1990 Phys. Fluids B 2 2408
- [19] Zarnstorff M.C. and Prager S. 1984 Phys. Rev. Lett. 53 454
- [20] Murakami M. et al 1991 Phys. Rev. Lett. 66 707–10
- [21] Neuner U. et al 2021 Nucl. Fusion 61 036024
- [22] Helander P., Geiger J. and Maaßberg H. 2011 Phys. Plasmas 18 092505
- [23] Xu Y. 2016 Matter Radiat. Extremes 1 192–00
- [24] Helander P., Parra F.I. and Newton S.L. 2017 J. Plasma Phys. 83 905830206
- [25] Boozer A.H. 1983 Phys. Fluids 26 496
- [26] Zarnstorff M.C. et al 2001 Plasma Phys. Control. Fusion 43 A237
- [27] Neilson G.H. et al 2000 Phys. Plasmas 7 1911
- [28] Isobe M. et al 2001 28th EPS Conf. on Contr. Fusion and Plasma Physics

- [29] Okamura S. et al 2004 Nucl. Fusion 44 575
- [30] Landreman M. and Paul E. 2022 Phys. Rev. Lett. 128 035001
- [31] Koh S., Chang C.S., Ku S., Menard J.E., Weitzner H. and Choe W. 2012 Phys. Plasmas 19 072505
- [32] Landreman M. and Ernst D.R. 2012 Plasma Phys. Control. Fusion 54 115006
- [33] Pablant N. et al 2020 Nucl. Fusion 60 036021
- [34] Hirshman S., van Rij W.I. and Merkel P. 1986 Comput. Phys. Commun. 43 143–55
- [35] Anderson D.V., Cooper W.A., Gruber R., Merazzi S. and Schwenn U. 1990 Int. J. Supercomput. Appl. 4 34
- [36] Ichiguchi K., Nakajima N., Okamoto M., Nakamura Y. and Wakatani M. 1993 Nucl. Fusion 33 481
- [37] Ichiguchi K., Nishimura T., Nakajima N., Okamoto M., Oikawa S. and Itagaki M. 2002 Nucl. Fusion 42 557–67
- [38] Cooper W.A., Graves J.P., Sauter O., Terranova D., Gobbin M., Marrelli L., Martin P. and Predebon I. 2011 *Plasma Phys. Control. Fusion* 53 084001
- [39] Cooper W.A. 1992 Plasma Phys. Control. Fusion 34 1011
- [40] Nemov V.V., Kasilov S.V., Kernbichler W. and Heyn M.F. 1999 Phys. Plasmas 6 4622
- [41] Beidler C.D. and Isaev M.Y. 2007 Proc. 17th Int. Toki Conf. and 16th Int. Stellarator/Heliotron Workshop Toki
- [42] Matsuoka K. et al 1988 12th Int. Conf. Plasma Phys. and Control. Nucl. Fusion Research, Nice, 1988 vol 2 (IAEA: Vienna) p 411
- [43] Landreman M., Smith H.M., Mollén A. and Helander P. 2014 Phys. Plasmas 21 042503
- [44] Rosenbluth M.N. 1968 Phys. Fluids 11 869

- [45] Proll J.H.E., Helander P., Connor J.W. and Plunk G.G. 2012 *Phys. Rev. Lett.* **108** 245002
- [46] Helander P. 2014 Rep. Prog. Phys. 77 087001
- [47] Alcusón J.A., Xanthopoulos P., Plunk G.G., Helander P., Wilms F., Turkin Y., Stechow A.V. and Grulke O. 2020 *Plasma Phys. Control. Fusion* 62 035005
- [48] Helander P., Proll J.H.E. and Plunk G.G. 2013 Phys. Plasmas 20 122505
- [49] Garren D.A. and Boozer A.H. 1991 Phys. Fluids B 3 2822
- [50] Masayuki Y., Kimitaka I., Shoichi O. and Keisuke M. 2002 Nucl. Fusion 42 1094–101
- [51] Jenko F., Dorland W., Kotschenreuther M. and Rogers B.N. 2000 Phys. Plasmas 7 1904
- [52] Jie H. et al 2022 Phys. Plasmas 29 052505
- [53] Horton C.W. and Ichikawa Y.H. 1996 *Chaos and Structures in Nonlinear Plasmas* (Singapore: World Scientific)
- [54] Okamura S., Liu H., Shimizu A., Kinoshita S., Isobe M., Xiong G. and Xu Y. 2020 J. Plasma Phys. 86 815860402
- [55] Kanno R., Hayashi T. and Okamoto M. 2005 Nucl. Fusion 45 588
- [56] Sinha P., Böckenhoff D., Endler M., Geiger J., Hölbe H., Smith H.M., Pedersen T.S. and Turkin Y. 2019 Nucl. Fusion 59 126012
- [57] Taylor J.B. 1986 Rev. Mod. Phys. 58 741
- [58] Hudson S.R., Dewar R.L., Dennis G., Hole M.J., McGann M., von Nessi G. and Lazerson S. 2012 Phys. Plasmas 19 112502
- [59] Baillod A., Loizu J., Qu Z.S., Kumar A. and Graves J.P. 2021 J. Plasma Phys. 87 905870403