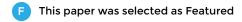
# Epitaxial growth of the first two members of the $Ba_{n+1}In_nO_{2.5n+1}$ Ruddlesden-Popper homologous series •

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### **ABSTRACT**

We demonstrate the epitaxial growth of the first two members, and the  $n = \infty$  member of the homologous Ruddlesden-Popper series of  $Ba_{n+1}In_nO_{2.5n+1}$  of which the n=1 member was previously unknown. The films were grown by suboxide molecular-beam epitaxy where the indium is provided by a molecular beam of indium-suboxide  $[In_2O(g)]$ . To facilitate *ex situ* characterization of the highly hygroscopic barium indate films, a capping layer of amorphous  $SiO_2$  was deposited prior to air exposure. The structural quality of the films was assessed by x-ray diffraction, reflective high-energy electron diffraction, and scanning transmission electron microscopy.

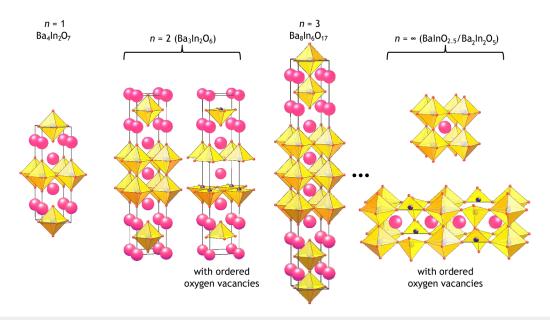
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Oxide Ruddlesden–Popper (RP) phases describe the homologous series of the perovskite-related  $A_{n+1}B_nO_{3n+1}$  phases, where n represents the number of perovskite unit cells sandwiched between double rock salt layers as illustrated in Fig. 1.  $^{1-3}$  The interest in RP phases is based on the intriguing and often fundamentally different physical properties that the distinct members of the series exhibit. A famous example is the  $Sr_{n+1}Ru_nO_{3n+1}$  series where with an increase of n, the phase changes from superconducting  $(n=1)^4$  to ferromagnetic (n=3-5). Consequently RP phases are highly pertinent to a wide variety of applications, e.g., tunable dielectrics, resistive switching, and photovoltaics.  $^{9,10}$ 

The trend toward the formation energy of phase n being degenerate with phase n+1 (or n-1) and an increasingly close cation stoichiometry with an increasing n make RP phases challenging to synthesize by most methods, especially for large n.

Molecular-beam epitaxy (MBE) with its unique flux control paired with recent developments in an improved *in situ* growth control of RP phases can, however, overcome these challenges. <sup>11,54</sup>

The  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series differs from most RP series as the perovskite end member  $BaInO_{2.5}$  usually referred to as  $n=\infty$  contains a huge concentration of oxygen vacancies. If these oxygen vacancies arrange in an ordered rather than a random manner, the  $n=\infty$  member transforms into the brownmillerite  $Ba_2In_2O_5$ . The literature is equivocal over the parameters promoting this transformation.  $I^{2-14}$  In addition to the  $n=\infty$  phase, two other members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series have been synthesized in bulk:  $n=2^{15}$  and n=3. The lack of evidence in the literature for an n=1 phase is a result of another phase forming for the same cation ratio,  $Ba_4In_2O_7$  with a unique structure sometimes misidentified as  $Ba_5In_2O_8$ , not related to the RP series.  $I^{7-20}$ 



**FIG. 1.** General structure of the homologous  $A_{n+1}B_nO_{3n+1}$  RP phases with increasing n from left to right. The oxygen octahedra around the violet B-site atoms are depicted in yellow, the A-site atoms in magenta, and the oxygen atoms in red. The specific structures of the  $Ba_{n+1}In_nO_{2.5n+1}$  series with oxygen vacancy ordering are included for the structures where the oxygen vacancy order is known, namely, the observed oxygen vacancy order for the n=2 and  $n=\infty$  phase. For clarity, only a quarter of the ordered  $n=\infty$  unit cell is depicted.

Intriguing properties for the known members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series are either predicted or already known. The  $n = \infty$  perovskite BaInO<sub>2.5</sub> and brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub> both combine high proton and oxygen conductivity and are, thus, interesting for applications such as solid oxide fuel When it was first prepared as a high-quality single crystal, Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> was considered a potential high transition temperature  $(T_c)$  superconductor based on its structure which is closely related to the n = 2 RP phase  $Sr_3Ti_2O_7$  and isostructural to the known high- $T_c$  superconductors  $La_{(2-x)}Sr_xCaCu_2O_6$  and  $La_2CaCu_2O_{(6+\delta)}$ . More recently, machine learning efforts to map out potential high  $T_c$  superconductors have predicted it to have a  $T_c$  of 45.9 K.<sup>26</sup> Despite this to our knowledge, there are no reports of the epitaxial growth of any members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series. A possible reason for this is the rather challenging growth. The large a-axis bulk lattice constants of the known members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series  $(Ba_3In_2O_6 \ a = 0.419 \text{ nm},^{27} \ Ba_8In_6O_{17} \ a = 0.417 \text{ nm})^{16}$ suggest a limited choice of substrates and most barium-containing indates are highly hygroscopic.<sup>20</sup> The hygroscopy of Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> is also the reason why its potential superconductivity was not investigated for single crystals.

In this letter, we describe the growth and structural characterization of the first two members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series by MBE of which the n=1 member was previously unknown. The films were grown using a molecular beam of  $In_2O$  emanating from a suboxide source; such growth is referred to as suboxide molecular-beam epitaxy (S-MBE). S-MBE has previously been demonstrated as a powerful tool for the growth of III-VI

materials at high growth rates and low substrate temperature with high structural perfection.  $^{28-30}$  This S-MBE approach for the growth of  $Ba_{n+1}In_nO_{2.5n+1}$  phases produces films with high crystalline quality as revealed by x-ray diffraction (XRD) and scanning transmission electron microscopy (STEM). The successful growth of RP phases by codeposition proves the applicability of S-MBE for a system requiring precise flux calibrations. The stability of the hygroscopic films is substantially increased by an *in situ* deposited layer of amorphous SiO<sub>2</sub>. We see indications for the formation of both  $n=\infty$  members (the perovskite and the brownmillerite member) being dependent on the choice of substrate.

All films are grown in a Veeco Gen10 MBE. The growth temperature is measured by an optical pyrometer at a measurement wavelength of 980 nm. A substrate temperature of 950 °C is employed as lower temperatures promote the formation of undesired compounds according to the Ba-In-O phase diagram. <sup>14</sup> The barium-containing indate films are grown in a mixture of approximately 10% distilled ozone and 90% oxygen at a background pressure of 10<sup>-6</sup> Torr. Before the deposition of the amorphous SiO<sub>2</sub> capping layer, the substrate is cooled to  $\leq 350$  °C and the oxidant background pressure is decreased to  $5 \times 10^{-7}$  Torr. The growth is monitored in real time by reflection high-energy electron diffraction (RHEED). Barium is evaporated from an elemental source. The In<sub>2</sub>O (g) suboxide molecular beam is obtained from heating a mixture of indium and In2O3 powders with a mole fraction of oxygen x(O) = 0.4 (provided by the  $In_2O_3$  powder) in a SUMO medium-temperature MBE effusion cell (Sumo\* is a registered trademark of Veeco Instruments, Inc.). An indium + In<sub>2</sub>O<sub>3</sub> mixture with x(O) = 0.4 has proven to give the most stable and pure

In<sub>2</sub>O flux,<sup>30</sup> note that a gallium + Ga<sub>2</sub>O<sub>3</sub> mixture with x(O) = 0.4 is used to produce a Ga<sub>2</sub>O (g) suboxide molecular beam in the growth of Ga<sub>2</sub>O<sub>3</sub> films by S-MBE.<sup>28,29</sup> The silicon species for the SiO<sub>2</sub> capping is supplied by a molecular beam of SiO (g) emanating from a standard medium-temperature MBE source containing chunks of amorphous SiO.<sup>31–33</sup> The fluxes of barium and In<sub>2</sub>O are roughly established by a quartz-crystal microbalance (QCM) and refined by the subsequent growth of binary oxide calibration films.<sup>34</sup> The thickness of the amorphous SiO<sub>2</sub> capping layer is calibrated by growing an amorphous SiO<sub>2</sub> film directly on a (001) MgO substrate and subsequently measuring the thickness by x-ray reflectivity (XRR) obtained with a Panalytical Empyrean with Cu-K<sub>α1</sub> radiation. (001) MgO substrates are supplied by CrysTec GmbH, and (001) SrTiO<sub>3</sub> substrates are supplied by Shinkosha Co. Ltd and prepared as described by Koster *et al.*<sup>35,36</sup>

Cross-sectional STEM specimens were prepared using a Thermo Fisher Scientific Helios G4 UX focused ion beam (FIB) using standard lift out and thinning methods. HAADF-STEM data were acquired using a Thermo Fisher Scientific Spectra 300 X-CFEG operating at 300 kV with a convergence angle of 21.4 mrad. Inner and outer collection angles were approximately 60 and 200 mrad, respectively. To obtain high signal-to-noise-ratio (SNR), many image frames were acquired at fast acquisition rates and then aligned using a rigid registration process optimized for noisy images.<sup>37</sup> STEM specimens were stored in a vacuum chamber between FIB preparation and imaging to minimize air exposure.

For any homologous RP series, the n=1 member has the highest A/B ratio (A/B=2) for n=1) and as n increases A/B approaches one. Once it reaches A/B=1, it is called the  $n=\infty$  member which is usually a perovskite, but in the case of  $\mathrm{Ba}_{n+1}\mathrm{In}_n\mathrm{O}_{2.5n+1}$  can also be a brownmillerite if the oxygen vacancies order into planes. An approach for achieving the different members of a homologous RP series in codeposition is to start with an A/B ratio of two and to subsequently decrease the ratio systematically. Considering the aforementioned large a-axis lattice constant of all known members of the  $\mathrm{Ba}_{n+1}\mathrm{In}_n\mathrm{O}_{2.5n+1}$  RP series, we choose (001) MgO as a substrate. In addition to its large a-axis lattice constant  $(0.42\,\mathrm{nm})$ ,  $^{39}$  it has the advantage of industrial relevance.  $^{40-44}$  Its obvious drawback is the nonperovskite structure. All discussed films have a thickness of  $\approx 160\,\mathrm{nm}$ .

Following the strategy of starting with Ba/In = 2, the resulting film under the aforementioned conditions shows the green diffraction pattern shown in Fig. 2(a). The pattern is indicative of a phase-pure film. The peak positions do not, however, correspond to the non-RP phase previously reported for this cation ratio 19 and correspond to a c-lattice constant of 1.34 nm. If we compare this pattern to the diffraction pattern of epitaxial n = 1 RP phases, e.g.,  $Sr_2TiO_4^{45}$  and Sr<sub>2</sub>RuO<sub>4</sub>, a strong resemblance is noticeable. This is the first hint of the successful growth of the previously unknown n = 1 member,  $Ba_4In_2O_7$ , of the  $Ba_{n+1}In_nO_{2.5n+1}$  RP series. This hypothesis is further substantiated by the RHEED image along the [100] in-plane axis of the film in Fig. 2(b). It shows the typical pattern for the expected in-plane perovskite-related structure and is sharp with slightly visible Kikuchi lines, indicating a smooth surface and high surface crystallinity. The bulk crystallinity is equally of high quality as evident from the full width half maximum (FWHM) of the rocking

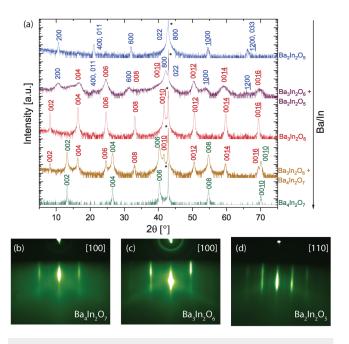


FIG. 2. (a)  $\theta$ -2 $\theta$  XRD scans of epitaxial films with varying Ba/In content offset along the y-axis for clarity. All films are about 160 nm thick and are grown on (001) MgO substrates. The 002 MgO substrate peaks are marked by an asterisk (\*). Cation ratios not fitting an exact phase stoichiometry result in a phase mixture (brown and purple). The diffraction pattern shown in green best agrees with the c-axis of the previously unknown n = 1 RP phase Ba<sub>4</sub>In<sub>2</sub>O<sub>7</sub> and we indexed it accordingly. The red diffraction pattern exhibits sharp peaks of high intensity at positions expected for c-axis oriented Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> based on its bulk lattice constant as reported by Antipov et al. (Ref. 27). The blue diffraction pattern is in agreement with epitaxial (100)-oriented brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub> based on the bulk lattice constant. The peak positions expected for an (011)-oriented brownmillerite as reported by Fischer et al. are also marked behind the comma (Ref. 38). RHEED patterns taken during growth along the [100] azimuth of the 160 nm thick epitaxial films of (b)  $Ba_4In_2O_7$  and (c) Ba<sub>3</sub>ln<sub>2</sub>O<sub>6</sub>. The Kikuchi lines and sharp features are a sign of the good crystal and surface quality. (d) RHEED pattern along the [110] azimuth of a 160 nm thick film of Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub>.

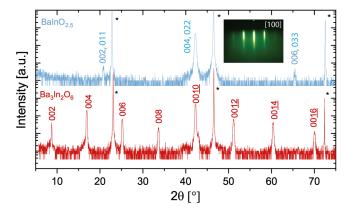
curve of the 006 peak of this phase being dominated by the substrate quality:  $0.03^{\circ}$  (see Fig. S1 in the supplementary material).<sup>53</sup>

As per our strategy in the next step, we increase the  $In_2O$  flux and, thus, decrease the Ba/In-ratio. For the resulting film, we observe the brown diffraction pattern in Fig. 2(a). It shows a mix of  $Ba_4In_2O_7$  and a second phase. This second phase is easily identified considering the diffraction pattern resulting for a further increased Ba/In-ratio of 3/2 (red). At a Ba/In ratio of 3/2, the  $\theta$ -2 $\theta$  scan in XRD displays phase-pure narrow peaks of exceptional intensity in agreement with the bulk *c*-axis lattice constant of  $Ba_3In_2O_6$ . The RHEED pattern along the [100] azimuth of the  $Ba_3In_2O_6$  film shows defined streaks and clear Kikuchi lines, again indicating a smooth surface and high surface crystallinity [Fig. 2(c)]. The narrow FWHM of the rocking curve of the  $00\underline{10}$  peak of the  $Ba_3In_2O_6$  film (0.11°, see Fig. S1 in the supplementary material<sup>53</sup>) indicates high crystalline quality, however, less so than the n=1 film. A possible reason is a

rotation of adjacent perovskite layers with respect to each other as shown in Fig. S2 (a) in the supplementary material.<sup>53</sup>

An even further decrease of the Ba/In-ratio again results in a film with a diffraction pattern [purple Fig. 2(a)] indicative of a mixture of phases. It shows broadened peaks at the same positions as the Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> film (red) plus additional peaks. These peaks are at the same positions as the peaks of the diffraction pattern for the film grown with a Ba/In-ratio of 1:1 [blue Fig. 2(a)] and are in agreement with the bulk *a*-lattice constant of the brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub> (space group Ima2; a = 1.672 nm, b = 0.608 nm, c = 0.596 nm). For growth on (001) MgO, the diffraction pattern, thus, suggests the epitaxial growth of a (100)-oriented brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub> film.

The phase observed is different when this same 1:1 flux ratio of Ba/In is incident on a (001) SrTiO<sub>3</sub> substrate. Figure 3 shows the diffraction for a film grown with a Ba/In-ratio of 1:1 (light blue) on a (001) SrTiO<sub>3</sub> substrate. The diffraction pattern does not show the peaks expected for the (100) brownmillerite phase. Instead peaks and peak intensity ratios in agreement with the bulk c-lattice constant of the perovskite BaInO<sub>2.5</sub> are seen. 12 The difference is also obvious from the RHEED images of both  $n = \infty$  films. The epitaxial brownmillerite grown on (001) MgO shows pronounced additional reflections at the half-order position of the substrate [Fig. 2(d)] suggesting a doubled unit cell. The epitaxial perovskite phase of BaInO<sub>2.5</sub> grown on (001) SrTiO<sub>3</sub> exhibits the RHEED pattern shown in the inset of Fig. 3, which is typical for a perovskite and shows no reflections at the half-order position of the substrate (see Fig. S3 in the supplementary material<sup>53</sup>). An explanation for this substrate-dependent observation of two different phasepure  $n = \infty$  films is that at the given growth parameters the formation of the brownmillerite phase is energetically favorable; the perovskite structure of the SrTiO<sub>3</sub> substrate, however, epitaxially



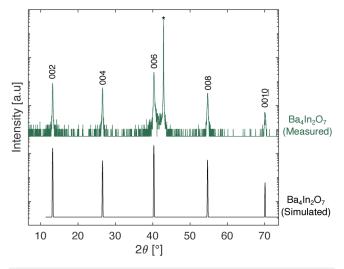
**FIG. 3.**  $\theta$ –2 $\theta$  scans of a 160 nm thick Ba $_3$ In $_2$ O $_6$  film (red) and a 160 nm thick perovskite BalnO $_2$ 5 film (light blue) grown on (001) SrTiO $_3$  with the film peaks indexed accordingly. The peak positions expected for an (011)-oriented brown-millerite are also marked following the comma (Ref. 38). The peaks marked by asterisks (\*) are the SrTiO $_3$  substrate peaks. The diffraction peaks are in good agreement with the respective values expected from the bulk c-lattice constants reported previously (Refs. 12 and 27). The inset shows the RHEED pattern along the [100] azimuth of the BalnO $_2$ .5 film at the end of the growth.

stabilizes<sup>46–50</sup> the perovskite phase. We cannot completely exclude the formation of (011)-oriented brownmillerite formation on (001) SrTiO<sub>3</sub> as its peaks coincide with the perovskite peaks indicated in Fig. 3. Nonetheless, the RHEED pattern and peak intensity ratios of the diffraction pattern suggest a perovskite film.

The red diffraction pattern in Fig. 3 belongs to the n=2 RP phase grown on (001) SrTiO<sub>3</sub>. The diffraction pattern of the n=2 Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> film is, despite the much higher lattice mismatch on (001) SrTiO<sub>3</sub> of 7.4% compared with -0.7% on (001) MgO, of similar structural quality. The much higher FWHM of the rocking curve (0.48°, see Fig. S1 in the supplementary material<sup>53</sup>), despite the isostructural substrate, means that the increased mismatch has a stronger negative effect on the crystalline quality than the isostructural substrate has a positive effect.

Comparing our x-ray diffraction data to the general structures in Fig. 1, we are able to epitaxially synthesize the  $n=\infty$  members, the n=2 member,  $\mathrm{Ba_3In_2O_6}$ , and the previously unknown n=1 member,  $\mathrm{Ba_4In_2O_7}$  (shown in Fig. 1), we compare a simulated diffraction pattern of it to the measured diffraction pattern from Fig. 2(a) in Fig. 4. We find good agreement between the measured and calculated diffraction patterns. This strengthens our claim of having grown the previously unknown n=1 RP phase,  $\mathrm{Ba_4In_2O_7}$ .

While the x-ray diffraction data suggest we are able to grow high-quality, phase-pure epitaxial films of the first two members of the  $Ba_{n+1}In_nO_{2.5n+1}$  homologous RP series, we have further assessed the crystalline structure and quality by STEM. Figure 5 shows cross-sectional HAADF-STEM images of the films and corroborates the identities of the phases hypothesized above. The n=1 and the n=2 RP phases are shown in Figs. 5(a) and 5(b), respectively, and the  $Ba_2In_2O_5$  brownmillerite phase is shown in Figs. 5(c) and 5(d). The depicted images are representative off all observed sample areas. Figure 5(a) confirms that we have indeed



**FIG. 4.** Comparison of a simulated  $\theta$ – $2\theta$  scan for the proposed Ba<sub>4</sub>In<sub>2</sub>O<sub>7</sub> structure to the measured  $\theta$ – $2\theta$  scan from Fig. 2(a).



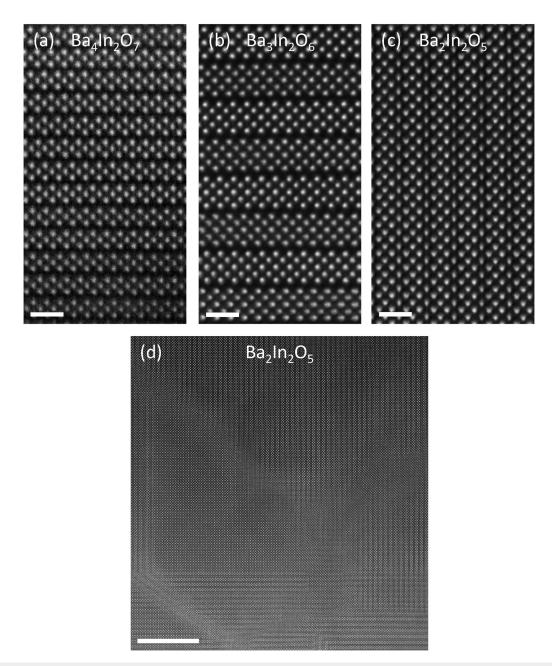


FIG. 5. Cross-sectional HAADF-STEM images of (a) n=1 Ba<sub>4</sub>In<sub>2</sub>O<sub>7</sub>, (b) n=2 Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub>, (c) the brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub>, and (d) an overview scan of the brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub>. The expected ordering for the respective n=1 and n=2 RP phases is clearly visible with the dark lines indicative of the double rock salt BaO layers. The brownmillerite in (c) is (011)-oriented. (d) reveals that the (011) and the (100) orientation coexist throughout the film. Scale bars for (a)–(c) are 1 nm. Scale bar for (d) is

epitaxially grown the previously unknown n = 1 RP phase, Ba<sub>4</sub>In<sub>2</sub>O<sub>7</sub>. Figure 5(b) verifies the high quality of the epitaxial Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> film grown on (001) MgO with the expected oxygen vacancy ordering. Nonetheless, some defects can be found within the Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> film and are discussed in Fig. S2 in the supplementary material.<sup>53</sup> Most interestingly, an on- and off-axis rotation of adjacent perovskite layers is commonly observed throughout the film which, as evident by the diffraction data shown in Fig. 4(a), has no effect on the c-lattice constant. Figure 5(c) reveals a (011) orientation of the epitaxial brownmillerite which is

in agreement with the RHEED image in Fig. 2(d), but not expected from the  $\theta$ -2 $\theta$  scan in Fig. 2(a). The broader overview HAADF-STEM image in Fig. 5(d) reveals that both orientations (011) and (100) coexist throughout the film. Some areas also show a contrast consistent with either the perovskite phase or a projection through stacking faults of the (100)- and/or (011)-oriented brownmillerite phase; latter being more likely considering the  $\theta$ -2 $\theta$ scan in Fig. 2(a). The lack of a (011) related diffraction pattern in Fig. 2(a) is easily explained by the peak positions. The 011 peak is hidden by the 400 peak of the (100)-oriented brownmillerite and the 022 peak is at the same position as the 002 substrate peak.

At this point, we can comfortably claim the ability to grow the first two members of the  $Ba_{n+1}In_nO_{2.5n+1}$  RP homologous series as well as the  $n = \infty$  members. We overcame the general challenge of RP phases needing precise stoichiometry control and did so utilizing an indium suboxide source, demonstrating its applicability to the growth of these layered indates. We overcame the specific challenge of the large a-axis lattice parameters of the barium-containing indates by using (001) MgO substrates, but our results suggest that Ba<sub>n+1</sub>In<sub>n</sub>O<sub>2.5n+1</sub> phases even grow epitaxially on highly mismatched substrates such as (001) SrTiO<sub>3</sub>. The remaining challenge to the growth of barium-containing indates is hygroscopy.

All of the films discussed thus far were capped with a 50 nm thick amorphous SiO2 layer. The importance of this capping layer will be discussed for the specific example of Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> films grown on (001) MgO. Figure 6(a) shows  $\theta$ -2 $\theta$  scans of two stoichiometric 160 nm thick Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> films. One of them was not capped. It was exposed to air and measured by XRD (blue). The other was capped with 50 nm of amorphous SiO2 before exposing it to air for the XRD measurement (red). The capped film is the same film as shown in Fig. 2(a). Both films were measured immediately after growth (bottom), after five days of storage in a desiccator (center), and for the capped sample after 45 days of storage in a desiccator (top). The  $\theta$ -2 $\theta$  scans of the uncapped and capped films are offset along the y-axis for clarity. It is obvious that the uncapped sample, despite storage in a desiccator, is not stable. The only peak remaining after five days is the 002 peak of the MgO substrate. In stark contrast to this behavior, all film peaks of the capped film remain after five days. Only the formation of a shoulder on the lower  $2\theta$  angle side of the lower-angle  $00\ell$  peaks (up to 0010) is observed. After 45 days, there is slightly more change. Instead of shoulders additional broad peaks can be observed at the former shoulder positions. These broad peaks coincide with the bariumpoor phases discussed above, making a loss of barium by hydroxylation probable. The observation at lower diffraction angles which are more surface sensitive means that the hydroxylation likely happens at the surface. Nonetheless, the distinct Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> diffraction pattern stays intact. A scan of the 0010 peak over a narrow  $\theta$ –2 $\theta$  range that includes the 002 MgO substrate peak [Fig. 6(b)] reveals that even the thickness fringes are still visible after 45 days of storage, a sign of preserved good crystallinity and a smooth surface.

Figure 6(c) shows a cross-sectional HAADF-STEM image of a similar capped sample with a 50 nm thick amorphous SiO<sub>2</sub> layer after being stored for over 120 days in a desiccator. The typical 1/5 ordering of n = 2 RP phases is maintained, which further demonstrates the preservation of the crystal quality of the sample by capping.

Prolonging the stability of epitaxial barium-containing indate films in general, and Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub>, in particular, enables their properties to be investigated. The electrical properties of Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> will be reported elsewhere.<sup>51</sup> It will also enable the investigation of the previously unknown n = 1 RP member Ba<sub>4</sub>In<sub>2</sub>O<sub>7</sub>, which we demonstrated in this paper for the first time. We expect that higher-nmembers of the  $Ba_{n+1}In_nO_{2.5n+1}$  RP homologous series, e.g., the

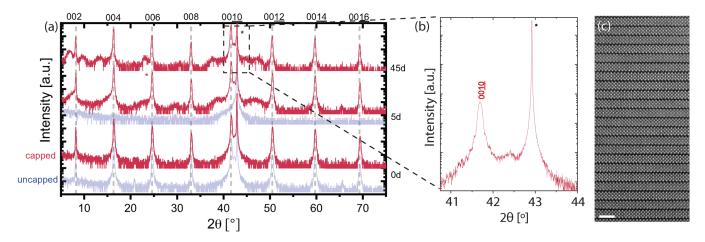


FIG. 6. (a)  $\theta$ -2 $\theta$  scans of an uncapped (blue) and a capped (red) Ba<sub>3</sub>In<sub>2</sub>O<sub>6</sub> film as a function of time after air exposure. Both films were stoichiometric as grown and ≈ 160 nm thick. The scans are offset along the y-axis for clarity. For the uncapped film, complete loss of crystallinity (no peaks remain) is seen after five days of being stored in a desiccator (center). For the capped film, the crystalline structure remains mostly intact for 45 days (top). This is further indicated by the thickness fringes of the 0010 film peak remaining visible after 45 days (b). (c) Representative HAADF-STEM image showing the characteristic 1/5 ordering of an n=2 RP phase even after an SiO<sub>2</sub> capped film is stored for over 120 days in a desiccator. Scale bar is 2 nm.



n=3 member,  ${\rm Ba_8In_6O_{17}}$ , that has already been demonstrated in bulk, <sup>16</sup> or yet to be demonstrated higher n can be grown by changing from codeposition to more advanced shuttered growth strategies. <sup>11,54</sup> Our ability to grow the n=1 and n=2 members by codeposition with high crystalline quality demonstrates the applicability of the newly developed  ${\rm In_2O}$  suboxide source to growths requiring a precise flux calibration.

The potential ability to select between the formation of the  $n=\infty$  phases of perovskite BaInO<sub>2.5</sub> and brownmillerite Ba<sub>2</sub>In<sub>2</sub>O<sub>5</sub> is of interest for the solid oxide fuel cell community as mentioned in the Introduction. Epitaxial films have recently been investigated by this community as model materials as they allow the influence of film microstructure, crystal orientation, strain, and other factors on the material's performance to be investigated.<sup>52</sup>

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# **AUTHOR DECLARATIONS**

### Conflict of Interest

The authors have no conflicts to disclose.

### **Author Contributions**

Felix V. E. Hensling: Conceptualization (lead); Data curation (lead); Formal analysis (lead); Funding acquisition (supporting); Investigation (lead); Methodology (equal); Writing – original draft (lead); Writing – review & editing (equal). Michelle A. Smeaton: Data curation (equal); Formal analysis (equal); Investigation (equal); Methodology (equal); Writing – original draft (supporting); Writing – review & editing (equal). Veronica Show: Data curation (supporting); Formal analysis (supporting); Visualization (supporting); Writing – review & editing (supporting). Kathy Azizie: Data curation (supporting); Methodology (supporting). Matthew R. Barone: Data curation (equal); Formal analysis (equal); Investigation (equal); Methodology (equal); Writing – original draft (supporting); Writing – review & editing (equal). Lena F. Kourkoutis: Funding acquisition (equal); Project

administration (equal); Supervision (equal); Writing – review & editing (equal). **Darrell G. Schlom:** Conceptualization (equal); Funding acquisition (lead); Project administration (lead); Supervision (lead); Writing – review & editing (lead).

### **DATA AVAILABILITY**

The data supporting the findings of this study are available within the paper. Additional data related to the growth and structural characterization are available in the PARADIM data repository at <a href="https://doi.org/10.34863/e7qx-7487">https://doi.org/10.34863/e7qx-7487</a>, Ref. 55. Any additional data connected to the study are available from the corresponding author upon reasonable request.

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