

A mathematical model for the energy stored in green roofs

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Abstract

A simple mathematical model to estimate the energy stored in a green roof is developed. Analytical solutions are derived corresponding to extensive (shallow) and intensive (deep) substrates. Results are presented for the surface temperature and energy stored in both green roofs and concrete during a typical day. Within the restrictions of the model assumptions the analytical solution demonstrates that both energy and surface temperature vary linearly with fractional leaf coverage, albedo and irradiance, while the effect of evaporation rate and convective heat transfer is non-linear. It is shown that a typical green roof is significantly cooler and stores less energy than a concrete one even when the concrete has a high albedo coating. Evaporation of even a few millimetres per day from the soil layer can reduce the stored energy by a factor of more than three when compared to an equivalent thickness concrete roof.

Keywords: Green roof, Heat island, Climate change, Mathematical model
2020 MSC: 93A30, 35K05

1. Introduction

2 With 55% of the world's population currently living in urban areas the
3 transition from rural living to urbanisation is a global issue. Antrop [1]

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4 predicts that by 2030, 85% of Europe's population will live in urban areas
5 whilst two-thirds of the global population will be in urban areas by 2050 [2].
6 Increasing levels of urbanisation lead to problems such as poor air and water
7 quality, growing demands on water availability, high energy consumption and
8 a deterioration of the natural environment. Urban areas are also significantly
9 warmer than rural areas, a phenomenon known as the urban heat island
10 effect. Heat islands arise in densely populated regions due to several factors.
11 The albedo of a surface indicates how well it reflects solar energy. Relative
12 to their surroundings, urban areas have high concentrations of dark surfaces
13 such as buildings, roofs, paved surfaces and car parks. These have low albedos
14 and absorb a high percentage of the incoming solar energy. Paved surfaces are
15 typically impermeable, causing rainwater to be directed to drainage systems
16 instead of being absorbed by vegetation which could later cool the area by
17 evapotranspiration. Buildings have high thermal masses, and thus store heat
18 during the day and slowly release it at night whereas natural landscapes such
19 as forest and other green areas consist largely of shaded, air filled regions with
20 a much lower thermal mass, thus storing less heat. Human activities, such as
21 factory and vehicle emissions, heating etc. also impact urban temperatures.

22 Early heat island studies include Howard's report of the London climate
23 in the late 1800s [3] and Schmidt's meteorological descriptions of urban ar-
24 eas in the early 1900s [4]. Since then heat islands have been observed and
25 examined globally [5, 6]. The resulting higher temperatures are a concern for
26 local authorities and residents. Heat islands can have public health implica-
27 tions such as high temperature ailments, hospitalisation and mortality [7–9].
28 Circumstances become worse with climate change and heat waves. For ex-
29 ample, the 2003 European heat wave resulted in thousands of excess deaths
30 [10]. From an economic perspective, heat islands lead to an increased demand
31 for air conditioning and other cooling equipment, and thus higher electricity
32 costs. Higher temperatures may also lead to a reduction in tourism over the
33 summer months. Thus, there is pressing need for a clear understanding of
34 heat islands, and practical, effective solutions.

35 To counteract heat islands, local authorities and government bodies are
36 constantly seeking cost-effective, environmentally friendly solutions [11]. One
37 option is a green roof which consists of vegetation planted over a waterproofed
38 system installed on top of a flat or slightly inclined roof. The two main cate-
39 gories of green roofs are extensive and intensive. Extensive green roofs have
40 thin soil with little or no irrigation whereas intensive green roofs have deep
41 soil, an irrigation system and more favourable growing conditions for vege-
42 tation. Intensive green roofs typically require a lot of maintenance, whereas
43 extensive green roofs are left to grow naturally with minimal maintenance.
44 Green roofs are made up of various layers including an upper vegetation

45 layer, a growing medium of organic and aggregate materials, a filter mem-
46 brane to prevent clogging of drains, a drainage layer to inhibit build-up of
47 excess water, a root barrier to stop root penetration and some form of roof
48 support. Green roofs have been shown to reduce the heat island effect as
49 they provide shade, remove heat from the air, and decrease temperatures of
50 roofs and the surrounding air [12–14]. Green roofs have other advantages
51 including rainwater management [15–18], and improvement of air and water
52 quality [19, 20]. They also greatly enhance an urban area’s aesthetic value
53 by increasing the level of urban fauna and wildlife habitat.

54 Modelling heat transfer in green roofs is complicated due to the various
55 layers involved and the movement of moisture [21]. Consequently researchers
56 apply a variety of simplifying hypotheses. Even so the standard approach re-
57 sults in a complex system which must be solved numerically. This is further
58 complicated by the typically 20 different parameter values, many of which
59 are difficult to determine experimentally [22] and therefore require the use
60 of multi-parameter fitting techniques (see [23] for a summary of such mod-
61 els). An attempt to develop a relatively simple model with an analytical
62 solution is contained in the workshop report [24]. In order to produce some
63 preliminary conclusions regarding green roof design this focuses primarily on
64 the energy stored in different substrates subject to constant ambient bound-
65 ary conditions and a semi-infinite layer. Here we deal primarily with more
66 realistic conditions, where the solar energy and ambient temperature vary
67 during the day, and also a finite thickness substrate. We obtain solutions via
68 separation of variables (verified against numerical solutions) to estimate the
69 energy stored in both extensive and intensive roofs. The ultimate aim being
70 to provide a set of guidelines to aid in the design of green roofs.

71 Theoretical models dealing with moisture dynamics and heat transport
72 in soil are typically based on those developed by Philip and De Vries in the
73 1950s [25]. Examples for models extending these descriptions by accounting
74 for energy transport in the soil support and the canopy on top of it include the
75 models presented by Palomo del Barrio [26] or Kondo and Saigusa [27]. Sailor
76 [28] presented a linearized model for determining the temperature evolution
77 at either side of the soil surface, which was later extended by Ouldboukhitine
78 et al. [29]. An extensive review of existing models may be found in Quezada-
79 Garcia et al. [21]. The majority of models incorporate evaporation as a
80 sink term in the soil heat equation, sometimes combined with a term in the
81 boundary condition at the soil surface. Since evaporation is known to be
82 a surface phenomenon, it is more realistic to include it exclusively in the
83 boundary condition.

84 A number of studies have shown that plants have some ability to regulate
85 their temperature over a wide range of ambient temperatures, see [30–32].

86 Citing these studies, and many others, Michaletz et al. [33] point out that the
87 difference between the plant and air temperature is generally small and the
88 correspondence is particularly close for ambient temperatures in the range 20-
89 30°C. The correspondence is achieved primarily through evapotranspiration.
90 Air has a very low viscosity (approximately 1.8×10^{-5} Pa s) and so is easily
91 mixed, for example by the motion of air above the canopy, motion of the
92 plants or convective currents due to temperature differences between the air
93 and the soil. Palomo [26] discusses the turbulent nature of the air both in
94 and above the canopy (and states that as a consequence of this the energy
95 and mass fluxes cannot be exactly predicted). It is well known from fluid
96 mechanics studies that turbulent diffusion is orders of magnitude greater than
97 thermal diffusion: as soon as turbulent diffusion occurs the air temperature
98 may be assumed approximately constant (and so set to the ambient value).

99 Given this dynamic mechanism for maintaining the air and plant tem-
100 perature at a similar level as well as the tendency for turbulent mixing of
101 the air, when developing a mathematical model the thermal response of the
102 canopy may be partially decoupled from that of the soil. Specifically the
103 canopy affects the soil in that the canopy air removes or adds heat by con-
104 vection while the leaves provide shade. However, any energy that the soil
105 gives to the air is rapidly transported away through turbulent diffusion and
106 has a negligible effect on the surroundings. Plants also remove water from
107 the soil, which affects the thermophysical parameter values. Consequently,
108 if the layers are to be decoupled it must be under the assumption that water
109 removal is relatively small, which may come through investigating situations
110 with a low evapotranspiration rate or for short times.

111 As a consequence of the above arguments, in this paper we will focus
112 on the development of a simple mathematical model to determine the effect
113 of different substrates on energy stored during daylight hours (which here
114 we will define as when the solar irradiance is above zero). This energy can
115 then go to heating the surface or be released at night, both with detrimental
116 effects on the comfort and energy footprint in the local area. The model
117 requires a number of assumptions. With regard to evaporation we assume
118 that the mass loss is proportional to the (varying) surface temperature. We
119 also assume an approximately constant moisture content in the soil. This
120 requires studying situations involving the removal of a few millimetres of wa-
121 ter per day, which is not a strong restriction, and only carry out calculations
122 for a single day. Subsequent days may be dealt with by imposing new initial
123 conditions. The resulting mathematical model is simple enough to be solved
124 analytically and the thermal performance of the layer can be ultimately de-
125 scribed by a single expression. The ultimate goal being to provide a set of
126 guidelines to demonstrate the effect of substrate type and physical character-

127 istics (depth, composition etc) on the energy storage and to provide simple
128 explicit formulae to compute this energy.

129 The paper is organised as follows. In Section 2 we specify the model as-
130 sumptions to obtain the governing heat equation in the roof layer as well as
131 the surface boundary condition which accounts for heat transfer due to con-
132 vection, radiation, solar irradiance and surface evaporation. In Section 3 ex-
133 act and approximate analytical solutions are obtained for both intensive and
134 extensive green roofs. In Section 4 we analyse different roof configurations
135 and discuss the range of validity of the different approximations obtained in
136 the previous section. Finally, Section 5 is devoted to the conclusions.

137 2. Derivation of governing equations

138 The mathematical model will be developed subject to the following as-
139 sumptions:

- 140 1. The plant temperature is close to the air temperature [33].
- 141 2. The air forms an infinite sink and is well mixed, so the air temperature
142 may be assumed independent of the plant cover and substrate.
- 143 3. The plants have two main effects on the soil, firstly they provide shade
144 (so reducing the amount of solar energy available to be absorbed),
145 secondly they remove water through transpiration.
- 146 4. The variation in the soil moisture is small such that it does not produce
147 any significant variation in the thermal parameters.
- 148 5. The temperature flow is predominantly one-dimensional.

149 The first four assumptions were discussed in the introduction. The as-
150 sumption on the moisture content obviously imposes restrictions on the cal-
151 culation time, which must be sufficiently small that only small quantities of
152 water evaporate, or that the substrate is carefully watered at regular inter-
153 vals. This is perhaps the most restrictive assumption, as well as the condi-
154 tions it imposes on the time the moisture content affects the albedo and the
155 thermal properties (such as density and thermal conductivity). Our results
156 may therefore be viewed as providing bounds on the energy storage capacity
157 for different moisture levels.

158 Regarding the final assumption, our aim is to produce a simple model
159 capable of predicting and providing insight into a green roof's performance.
160 Given that the main drivers for energy change occur at the top and bottom
161 boundaries it seems apparent that the heat flow is predominantly perpendic-
162 ular to the surface. Small local changes, for example due to variations in leaf
163 coverage, would require a detailed analysis for each individual green roof. By
164 dealing with the one-dimensional problem we are effectively considering the
165 average temperature over a cross-sectional area.

Table 1: Summary of the symbols used in this work.

Symbols

A	Albedo (-)
H_{as}	Air-soil heat transfer coefficient ($\text{W}/\text{m}^2\text{°C}$)
ρ	Density (kg/m^3)
f, g	Deviation from average irradiance, temperature ($\text{W}/\text{m}^2, \text{°C}$)
z	Distance from surface (m)
E	Energy absorbed (J/m^2)
λ_n	Eigenvalue (-)
$\varphi_n, \phi_n, \psi_n$	Eigenfunction (-)
d, \dot{d}	Evaporated amount, Evaporation rate (m, m/s)
α	Evaporation proportionality constant ($\text{m}/\text{°C}$ s)
F_{VC}	Fractional vegetative coverage (-)
q	Heat flux (W/m^2)
ϵ	Heat loss parameter (-)
L_e	Latent heat of evaporation (J/kg)
p	period (s)
L	Soil thickness (m)
C_1, C_2	Specific energy terms ($\text{W}/\text{m}^2, \text{W}/\text{m}^2\text{°C}$)
c	Specific heat capacity ($\text{J}/\text{kg}\text{°C}$)
Q	Sun irradiance (W/m^2)
T	Temperature (°C)
k	Thermal conductivity ($\text{W}/\text{m}\text{°C}$)
D	Thermal diffusivity (m^2/s)
t	Time (s)

Superscripts

0 Constant environmental conditions

Subscripts

s	Soil	c	Concrete	a	Air
w	Water	rad	Radiation	conv	Convection
evap	Evaporation	lap	Laplace	tot	Total
max	Maximum value	min	Minimum value	mn	Mean value
av	Average value	num	Numerical	app	Approximate
net	Net value				

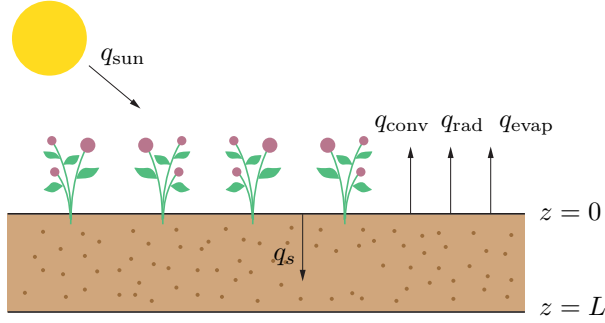


Figure 1: Illustration of the green roof model.

166 *2.1. Mathematical model*

167 The one dimensional model that we will consider is represented in Fig-
 168 ure 1. Subject to the above restrictions the substrate satisfies the one-
 169 dimensional heat equation

$$\rho_s c_s \frac{\partial T}{\partial t} = \frac{\partial}{\partial z} \left(k_s \frac{\partial T}{\partial z} \right), \quad (1)$$

where the notation is described in Table 1. The thermal parameters may depend on the water content (here assumed to be constant) and the substrate occupies the region $z \in [0, L]$, with $z = 0$ being open to the atmosphere. With constant coefficients we denote the thermal diffusivity as $D_s = k_s / \rho_s c_s$. At the surface, $z = 0$, the energy conducted into or out of the substrate must balance the incoming heat from the sun with that lost through convection, radiation and evaporation, hence

$$q_s = -k_s \frac{\partial T}{\partial z} \Big|_{z=0} = q_{\text{sun}} + q_{\text{conv}} + q_{\text{rad}} + q_{\text{evap}}. \quad (2)$$

170 As for the lower boundary conditions, we assume the roof has a perfect
 171 thermal insulation, so there is no heat flowing out of the roof and into the
 172 house, which can be written in terms of a homogeneous Neumann boundary
 173 condition like $\partial T(L, t) / \partial z = 0$. However, if the soil is deep enough the
 174 domain can be assumed to be semi-infinite where the temperature far from
 175 the surface remains constant, which can be written like $\partial T(z, t) / \partial z \rightarrow 0$
 176 as $z \rightarrow \infty$. We shall therefore consider one of the two following boundary
 177 conditions:

$$\text{i) } \lim_{z \rightarrow \infty} \frac{\partial T}{\partial z} = 0, \quad \text{ii) } \frac{\partial T}{\partial z} \Big|_{z=L} = 0. \quad (3)$$

178 The first applies to thick soil layers, the second to thinner layers. Clearly,
 179 due to weight considerations, green roofs are far from semi-infinite: later

180 we will discuss what constitutes thick and thin and demonstrate that, at
 181 least mathematically, roofs greater than 15cm thick may be treated as semi-
 182 infinite.

183 As initial condition we shall take the daily mean temperature value,

$$T(z, 0) = T_{\text{mn}} = \frac{T_{\text{max}} + T_{\text{min}}}{2}. \quad (4)$$

184 2.2. Specific energy terms

185 The most important part of the above energy balance is the surface
 186 boundary condition, equation (2) which requires expressions for the con-
 187 vection, radiation and evaporation terms.

Convection and radiation. Firstly we note that the convective and radiation terms are usually combined to make a single effective heat transfer term. The sum of the convective and radiation terms is

$$E_{\text{conv}} + E_{\text{rad}} = -H'_{as}(T - T_a) - \epsilon\sigma(T^4 - T_a^4), \quad (5)$$

188 where H'_{as} is the soil-air heat transfer coefficient, T_a is the ambient air tem-
 189 perature, ϵ is an emissivity coefficient and σ is the Stefan–Boltzmann con-
 190 stant. Assuming the soil surface temperature is not greatly different to the
 191 air temperature we may write $T = T_a + f$, where $f = (T - T_a) \ll T_a$, then

$$(T^4 - T_a^4) = (T_a^4 + 4fT_a^3 + 6f^2T_a^2 \dots) - T_a^4 = 4fT_a^3 \left(1 + \frac{3}{2}\frac{f}{T_a} + \dots\right). \quad (6)$$

Provided $f/T_a = (T - T_a)/T_a \ll 1$, then

$$E_{\text{conv}} + E_{\text{rad}} \approx -H_{as}(T - T_a) \quad (7)$$

192 where $H_{as} = H'_{as} + 4\epsilon\sigma T_a^3$ is a combined soil-air heat transfer coefficient.

Energy flux related to the sun irradiance. The solar energy term in (2) may be written as

$$q_{\text{sun}} = (1 - F_{VC})(1 - A_s)Q_{\text{net}}, \quad (8)$$

193 where A_s is the substrate albedo and Q_{net} the net irradiance received from
 194 the sun. The parameter $F_{VC} \in [0, 1]$ is the fractional vegetative coverage,
 195 a dimensionless quantity for whose definition we may use the Leaf Area
 196 Index (LAI), see [21, 26, 34], which is used to characterise plant canopies.
 197 The LAI is defined as the one-sided green leaf area per unit ground surface
 198 area (see for instance [35]). Since dense plant canopies can have several
 199 layers of leafs, values of LAI > 1 are permissible. Plant leaves very rapidly
 200 prevent the sun from reaching the soil surface, so a common definition has
 201 $F_{VC} = 1 - \exp(-\text{LAI})$.

Energy loss due to evaporation. Evaporation is a form of phase change which occurs below the standard phase change temperature. For fluid molecules to escape from the liquid surface they require sufficient energy. At low temperatures the fluid molecules have low energy, however they may gain above average energy through collisions with other molecules. The closer to the phase change temperature the more likely it is that a molecule will gain sufficient energy to break free from the liquid and join the vapour layer above the surface. For this reason we will define an evaporation rate proportional to the surface temperature $\dot{d} = \alpha T(0, t)$, that is, the evaporation rate increases with the surface temperature and so

$$q_{\text{evap}} = -\rho_w L_e \dot{d}(t) = -\rho_w L_e \alpha T(0, t), \quad (9)$$

202 where \dot{d} the rate at which water is evaporated per unit area. The constant of
 203 proportionality α in fact depends on quantities such as the relative humidity,
 204 air flow, surface roughness which all influence the evaporation but will here
 205 be taken as a constant.

The soil surface boundary condition may now be written as

$$\begin{aligned} -k_s \frac{\partial T}{\partial z} \Big|_{z=0} &= (1 - F_{VC})(1 - A_s) Q_{\text{net}} \\ &- H_{as}(T(0, t) - T_a) - \rho_w L_e \alpha T(0, t). \end{aligned} \quad (10)$$

206 The relation between the boundary condition (10) and the Penman-Monteith
 207 (PM) equation is discussed in [24]. The PM equation is a widely used formula
 208 primarily aimed at estimating the amount of evapotranspiration. It includes
 209 a variety of empirical quantities, some, such as the ‘ground heat flux’, are dif-
 210 ficult to measure. The ground heat flux is identical to $-k_s T_z(0, t)$ which here
 211 will be determined during the solution process. As discussed earlier trans-
 212 spiration plays a limited role in the substrate energy balance. The surface
 213 boundary condition depends on the surface evaporation whereas transpira-
 214 tion acts to remove water from within the soil which is then evaporated in the
 215 plant layer and the associated energy passed to the air. The effect of tran-
 216 spiration on the soil is then purely through the amount of water removed.
 217 Experimentally this may lead to an issue in distinguishing where the water
 218 has been removed, since this affects the measurement of d . In which case we
 219 could make an estimate via an empirical formula. However, our goal here
 220 is to provide a comparative study of different forms and depths of roof so
 221 we will simply specify a value for d and then see how it affects the different
 222 substrates.

223 *Time-dependent environment conditions.* Both the solar energy and ambient
 224 temperature vary throughout the day, typical forms are shown in Figure 2.

225 The solid curve in Figure 2(a) depicts the variation of solar radiation. We
 226 take the origin of our time axis as the moment when the solar radiation
 227 first reaches the surface and, in this example, the radiation returns to zero
 228 at the end time $t_f = 14$ hours (note, the calculations are in SI units so
 229 in all equations we work in seconds, $t_f = 14 \times 3600$ s). Consequently we
 230 approximate the irradiance by

$$Q_{\text{net}}(t) = Q_{\text{max}} \sin\left(\frac{2\pi t}{p_1}\right), \quad (11)$$

231 where the period is $p_1 = 2t_f$ s. This holds for $t \in [0, t_f]$ s, which is how we
 232 define our daytime and is the time period over which we will calculate the
 233 energy absorbed by the substrate. In Figure 2(b) the solid curve represents
 234 the ambient temperature. This reaches a minimum a short time after the
 235 solar radiation begins, while the maximum occurs shortly after the peak of
 236 radiation. If we choose a minimum at $t = t_{\text{min}} = 1$ hour and a maximum at
 237 $t = t_{\text{max}} = 9$ hours then the sine wave representation takes the form

$$T_a(t) = T_{\text{min}} - \frac{\Delta T}{2} \cos\left(\frac{2\pi(t - t_{\text{min}})}{p_2}\right), \quad (12)$$

238 where $\Delta T = T_{\text{max}} - T_{\text{min}}$ and the period p_2 is twice the distance between the
 239 maximum and minimum, $p_2 = 2(t_{\text{max}} - t_{\text{min}}) = 2 \times 8 \times 3600$ s. Given that
 240 neither period p_1 or p_2 coincide with the day length these functions cannot
 241 be applied for a full 24 hours. However, our goal is to calculate the energy
 242 absorbed by the substrate over daylight hours in which case the definitions
 243 (11, 12) provide a realistic approximation.

244 In the following sections we will calculate analytical and numerical solu-
 245 tions for the temperature and energy in the substrate. Some of the analytical
 246 approximations employed will require constant values for irradiance and am-
 247 bient temperature and thus some form of average value must be determined.
 248 Consequently we write

$$Q_{\text{net}}(t) = Q_{\text{av}} + f(t), \quad T_a(t) = T_{\text{av}} + g(t). \quad (13)$$

249 To determine the average solar irradiance we first calculate the daily total
 250 (i.e., the solar exposure or insolation)

$$Q_{\text{tot}} = Q_{\text{max}} \int_0^{t_f} \sin\left(\frac{2\pi t}{p_1}\right) dt = \frac{2}{\pi} Q_{\text{max}} t_f, \quad (14)$$

251 where we have used $p_1 = 2t_f$. Similarly, the integral of the temperature

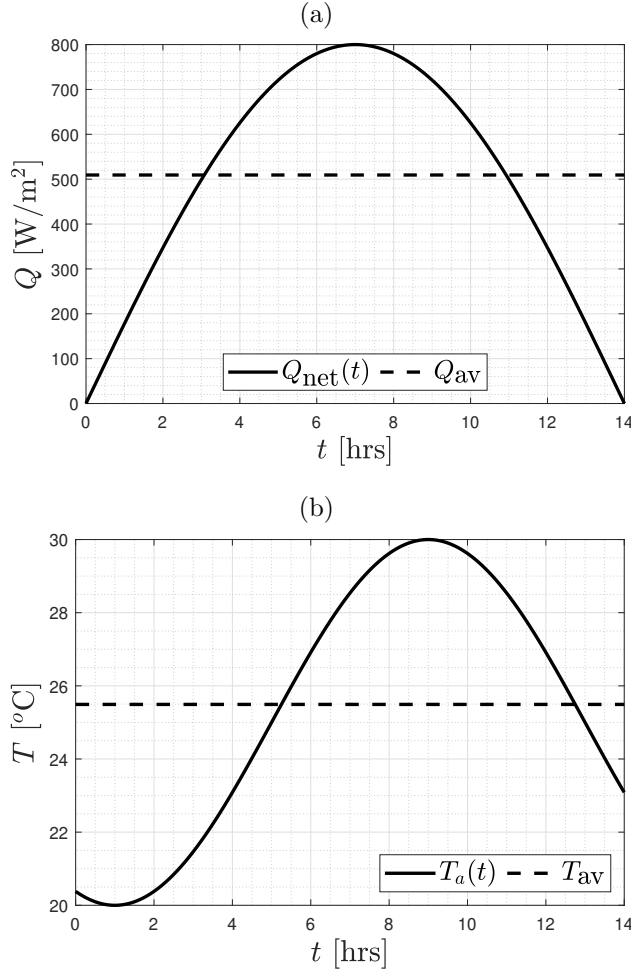


Figure 2: (a) Evolution of the solar energy (irradiance) and (b) the air temperature during solar radiation hours. The dashed lines represent the daily averages.

252 during the time period $[0, t_f]$ is

$$\begin{aligned}
 T_{\text{tot}} &= \int_0^{t_f} T_a(t) dt \\
 &= T_{\text{mn}} t_f - \frac{\Delta T p_2}{4\pi} \left[\sin \left(\frac{2\pi (t_f - t_{\text{min}})}{p_2} \right) + \sin \left(\frac{2\pi t_{\text{min}}}{p_2} \right) \right]. \quad (15)
 \end{aligned}$$

The daily averages are then

$$Q_{\text{av}} = \frac{Q_{\text{tot}}}{t_f}, \quad T_{\text{av}} = \frac{T_{\text{tot}}}{t_f}. \quad (16)$$

The daily average temperature, T_{av} , and the mean temperature, T_{mn} , are only equal if $t_f = p_2$, this would occur with exactly 12 hours of daylight.

Substituting for $T_{\text{av}}, Q_{\text{av}}$ in equations (13) defines $f(t), g(t)$

$$f(t) = Q_{\text{max}} \left(\sin \left(\frac{2\pi t}{p_1} \right) - \frac{2}{\pi} \right), \quad (17)$$

$$g(t) = \frac{\Delta T}{2} \left(\frac{p_2}{2\pi t_f} \left[\sin \left(\frac{2\pi(t_f - t_{\text{min}})}{p_2} \right) + \sin \left(\frac{2\pi t_{\text{min}}}{p_2} \right) \right] - \cos \left(\frac{2\pi(t - t_{\text{min}})}{p_2} \right) \right). \quad (18)$$

253 The final form for the surface boundary condition is then

$$-k_s \frac{\partial T}{\partial z} \Big|_{z=0} = C_1 + G(t) - C_2 T(0, t), \quad (19)$$

254 where C_1, C_2 are constant and $G(t)$ incorporates the $f(t), g(t)$ terms:

$$C_1 = (1 - F_{VC})(1 - A_s)Q_{\text{av}} + H_{as}T_{\text{av}}, \quad C_2 = H_{as} + \rho_w L_e \alpha, \quad (20)$$

$$G(t) = (1 - F_{VC})(1 - A_s)f(t) + H_{as}g(t). \quad (21)$$

255 *Energy flux due to evaporation.* With regard to the evaporation rate we first
 256 note that the value of the constant α must be consistent with the specified
 257 daily evaporation rate. If we observe that d metres have evaporated in time
 258 t_f then

$$d = \int_0^{t_f} \dot{d} dt = \alpha \int_0^{t_f} T(0, t) dt. \quad (22)$$

259 In the case where \dot{d} is constant the above equation states that $\dot{d} = d/t_f$ and α
 260 is redundant. In the variable case the method to obtain α from this relation
 261 will be discussed later.

262 2.3. Absorbed energy

To determine the energy absorbed (above the initial energy) in a layer of
 thickness L we must evaluate

$$E(t) = \int_0^L \rho_s c_s (T(z, t) - T_{\text{mn}}) dz. \quad (23)$$

263 We note that the energy is related to the surface boundary condition and, in
 264 turn, to the evaporation rate. Indeed, integrating (1) with respect to time
 265 and using (19) one obtains

$$\frac{dE}{dt} = k_s \frac{\partial T}{\partial z} \Big|_{z=L} - k_s \frac{\partial T}{\partial z} \Big|_{z=0}. \quad (24)$$

266 With an insulated layer $\partial T/\partial z = 0$ at $z = L$, whereas for a deep layer
 267 $T \rightarrow T_\infty$ as $z \rightarrow \infty$, hence $\partial T/\partial z \rightarrow 0$. In either case the first temperature
 268 derivative in (24) may be neglected while the second is simply the energy
 269 flux at the surface which is defined by the surface boundary condition

$$\frac{dE}{dt} = C_1 + G(t) - C_2 T(0, t), \quad (25)$$

270 and therefore the energy absorbed in the layer is

$$E(t) = C_1 t + \int_0^t (G(\tau) - C_2 T(0, \tau)) dt. \quad (26)$$

271 According to the definition of $G(t)$ given in (21), $\int_0^{t_f} G(t) dt = 0$. Therefore,
 272 the energy absorbed at the end of the day, and thus available for release
 273 through the night, is

$$E(t_f) = C_1 t_f - C_2 \int_0^{t_f} T(0, t) dt = C_1 t_f - C_2 \frac{d}{\alpha} \quad (27)$$

$$= (1 - F_{VC})(1 - A_s) Q_{av} t_f - \rho_w L_e d + H_{as} \left(T_{av} t_f - \frac{d}{\alpha} \right). \quad (28)$$

274 It would appear that we can finish the calculation here given that we have
 275 an expression for the energy absorbed through the day and hence the en-
 276 ergy available for release at night. The simplicity of the above expression
 277 clearly shows how the absorbed energy depends on the problem parameters.
 278 However, α is *a priori* unknown and depends on the surface temperature.
 279 Consequently in the following sections we will investigate the temperature
 280 flow more carefully to determine the evaporation rate and specifically the
 281 form of the function $d(\alpha)$. Further, this will determine the surface tem-
 282 perature which is an important quantity in the comfort of the population.
 283 Importantly the analysis will show that due to the interplay between evapo-
 284 ration and surface temperature the absorbed energy does not depend linearly
 285 on t_f as suggested by the above result.

286 2.4. Numerical solution

287 The mathematical problem defined by the heat equation (1), boundary
 288 conditions (3i) and 10 and the initial condition (4) can be solved numerically
 289 by using Matlab's built-in function `pdepe`, which performs a discretization of
 290 the spatial variable to obtain a system of ODEs which is then solved using
 291 the function `ode15s`. It is an excellent way to demonstrate the accuracy of
 292 the analytical solutions but does not provide any insight into the role of the
 293 model parameters.

294 **3. Analytical solutions**

295 In this section we seek analytical solutions for both infinite and finite
 296 depth layers. Specifically we will calculate the temperature profile through-
 297 out the layer and from this determine the surface temperature which is re-
 298 quired to calculate the energy absorbed throughout the day. We note that
 299 in the case of a substrate with infinite depth the analytical method requires
 300 constant average values for the ambient temperature and solar irradiance (as
 301 studied in [24]).

302 *3.1. Deep substrate solution for averaged input values*

303 As discussed above, for thick layers the surface temperature variation may
 304 not affect the lower boundary $z = L$. Consequently we may seek an analytical
 305 solution based on the assumption $L \rightarrow \infty$ (and subsequently determine a
 306 lower bound for this approximation).

Taking the average values for the ambient temperature and solar radiation
 discussed in Section 2.2, that is $G = 0$, we may use Laplace transforms (see
 Appendix A) to obtain

$$T_{\text{lap}}(z, t) = T_{\text{mn}} + \frac{C_1 - C_2 T_{\text{mn}}}{C_2} \left[\operatorname{erfc} \left(\frac{z}{2\sqrt{D_s t}} \right) - \exp \left(\frac{C_2}{k_s} \left(\frac{C_2 D_s t}{k_s} + z \right) \right) \operatorname{erfc} \left(\frac{z}{2\sqrt{D_s t}} + \frac{C_2 \sqrt{D_s t}}{k_s} \right) \right], \quad (29)$$

where $\operatorname{erfc}(z)$ is the well known complementary error function,

$$\operatorname{erfc}(z) = \frac{2}{\sqrt{\pi}} \int_z^\infty e^{-t^2} dt.$$

In the limit $z \rightarrow \infty$ this gives $T \rightarrow T_{\text{mn}}$ which satisfies $\partial T / \partial z \rightarrow 0$ so this applies for both boundary conditions (3) for sufficiently large values of L . The surface temperature is simply

$$T_{\text{lap}}(0, t) = T_{\text{mn}} + \frac{C_1 - C_2 T_{\text{mn}}}{C_2} \left[1 - \exp \left(\frac{C_2^2 D_s t}{k_s^2} \right) \operatorname{erfc} \left(\frac{C_2 \sqrt{D_s t}}{k_s} \right) \right]. \quad (30)$$

Using expression (23) the energy absorbed throughout the layer after time t is

$$E_{\text{lap}}(t) = \frac{k_s^2}{D_s C_2^2} (C_1 - C_2 T_{\text{mn}}) \left[\exp \left(\frac{C_2^2 D_s t}{k_s^2} \right) \times \operatorname{erfc} \left(\frac{C_2}{k_s} \sqrt{D_s t} \right) + \frac{2}{\sqrt{\pi}} \frac{C_2}{k_s} \sqrt{D_s t} - 1 \right]. \quad (31)$$

We note that $E_{\text{lap}}(t)$ is linear in C_1 (defined in (20)), this is in keeping with equation (27) although the time dependence is different. The dependence of the energy on the evaporation rate, which appears in C_2 , is not so clear. We can find a simpler expression by considering large time solutions, i.e. approaching the end of the day. For the soils considered in this work $C_2\sqrt{D_s t_f/k_s}$ takes values between 2 and 5. Noting that for large y , $e^{y^2}\text{erfc}(y) = 1/(y\sqrt{\pi}) + \mathcal{O}(1/y^3)$. For sufficiently large times the energy may then be approximated by

$$E_{\text{lap}}(t) \sim \frac{k_s^2}{D_s C_2^2} (C_1 - C_2 T_{\text{mn}}) \left[\frac{k_s}{C_2 \sqrt{D_s t}} + \frac{2}{\sqrt{\pi}} \frac{C_2}{k_s} \sqrt{D_s t} - 1 \right]. \quad (32)$$

For t close to t_f the second term in the square brackets is dominant indicating

$$E_{\text{lap}}(t_f) \rightarrow 2 \left(\frac{C_1}{C_2} - T_{\text{mn}} \right) \frac{k_s}{\sqrt{\pi D_s}} \sqrt{t_f} \quad (33)$$

$$= 2 \left(\frac{(1 - F_{VC})(1 - A_s)Q_{\text{av}} + H_{as}T_{\text{av}}}{H_{as} + \rho_w L_e \alpha} - \frac{T_{\text{max}} + T_{\text{min}}}{2} \right) \sqrt{\frac{\rho_s c_s k_s t_f}{\pi}}. \quad (34)$$

307 This simple expression provides an estimate of the energy for a deep substrate
 308 layer when the average of irradiance and ambient temperature are imposed.
 309 It clearly indicates the effect of each system parameter. There is a linear
 310 dependence on the fractional vegetative coverage, albedo and irradiance. The
 311 thermophysical parameters, ρ_s, c_s, k_s appear as a square root.

312 3.2. Finite substrate solution

Here we consider a finite substrate on top of an insulating layer. In this case we apply the boundary condition (3 ii). The heat equation (1) is now defined over the finite domain $z \in [0, L]$, which suggests a solution by separation of variables. This leads to an approximate solution of the form

$$T(z, t) = T_{\text{mn}} + \frac{C_1 L}{k_s} \sum_{n=1}^{\infty} \frac{\cos(\lambda_n)}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)} \cos\left(\lambda_n \left(1 - \frac{z}{L}\right)\right) \psi_n(t). \quad (35)$$

313 The derivation is provided in Appendix B. The eigenvalues satisfy

$$\tan(\lambda_n) = \frac{\beta_3}{\lambda_n}. \quad (36)$$

and

$$\psi_n(t) = \beta_2 \left[1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \right] + \lambda_n^2 \mathcal{G}_n(t), \quad (37)$$

where $\mathcal{G}_n(t) = \beta_1 I_{1,n}(t) + (1 - \beta_1) I_{2,n}(t)$ and

$$I_{1,n}(t) = \frac{D_s L^2 p_1 \pi}{8\pi^2 L^4 + 2D_s^2 p_1^2 \lambda_n^4} \left[\frac{D_s p_1 \lambda_n^2}{L^2} \sin\left(\frac{2\pi t}{p_1}\right) - 2\pi \cos\left(\frac{2\pi t}{p_1}\right) + 2\pi \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \right] - \frac{1}{\lambda_n^2} \left[1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \right], \quad (38)$$

$$I_{2,n}(t) = \frac{1}{\lambda_n^2} \left(\frac{T_{\text{mn}}}{T_{\text{av}}} - 1 \right) \left[1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \right] + \frac{\Delta T D_s^2 p_2^2 \lambda_n^2}{2T_{\text{av}} (4\pi^2 L^4 + D_s^2 p_2^2 \lambda_n^4)} \left[\exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \cos\left(\frac{2\pi t_{\text{min}}}{p_2}\right) - \cos\left(\frac{2\pi (t - t_{\text{min}})}{p_2}\right) \right] - \frac{\Delta T D_s L^2 p_2 \pi}{T_{\text{av}} (4\pi^2 L^4 + D_s^2 p_2^2 \lambda_n^4)} \times \left[\exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \sin\left(\frac{2\pi t_{\text{min}}}{p_2}\right) + \sin\left(\frac{2\pi (t - t_{\text{min}})}{p_2}\right) \right], \quad (39)$$

314 where $p_1 = 2t_f$ and $p_2 = 2(t_{\text{max}} - t_{\text{min}})$. These expressions involve three
315 non-dimensional parameters

$$\beta_1 = \frac{(1 - F_{VC})(1 - A_s)Q_{\text{av}}}{C_1}, \quad \beta_2 = 1 - \frac{C_2 T_{\text{mn}}}{C_1}, \quad \beta_3 = \frac{C_2 L}{k_s}. \quad (40)$$

Setting $z = 0$ in (36) gives the surface temperature

$$T_{\text{N}}(0, t) = T_{\text{mn}} + \frac{C_1 L}{k} \sum_{n=1}^{\infty} \frac{\psi_n(t)}{\lambda_n^2 + \beta_3^2 + \beta_3}. \quad (41)$$

316 The eigenvalues, λ_n , are found through the numerical solution of (36). In
317 the following we will also employ an approximation to the first eigenvalue

$$\lambda_1 \approx \lambda_{\text{app}} = \left(\frac{5^{1/3}}{6} y - \frac{5}{6} - \frac{13 \cdot 5^{2/3}}{6y} \right)^{1/2} \quad (42)$$

where

$$y = \left(110 + 162\beta_3 + 9\sqrt{285 + 440\beta_3 + 324\beta_3^2} \right)^{1/3}, \quad (43)$$

318 see Appendix B.

319 In Figure 3 we compare the exact and approximate values of λ_1 for varying
320 L and parameter values for a typical soil (parameter values used here and in
321 the next three figures are discussed in detail in §4). As may be seen, the error

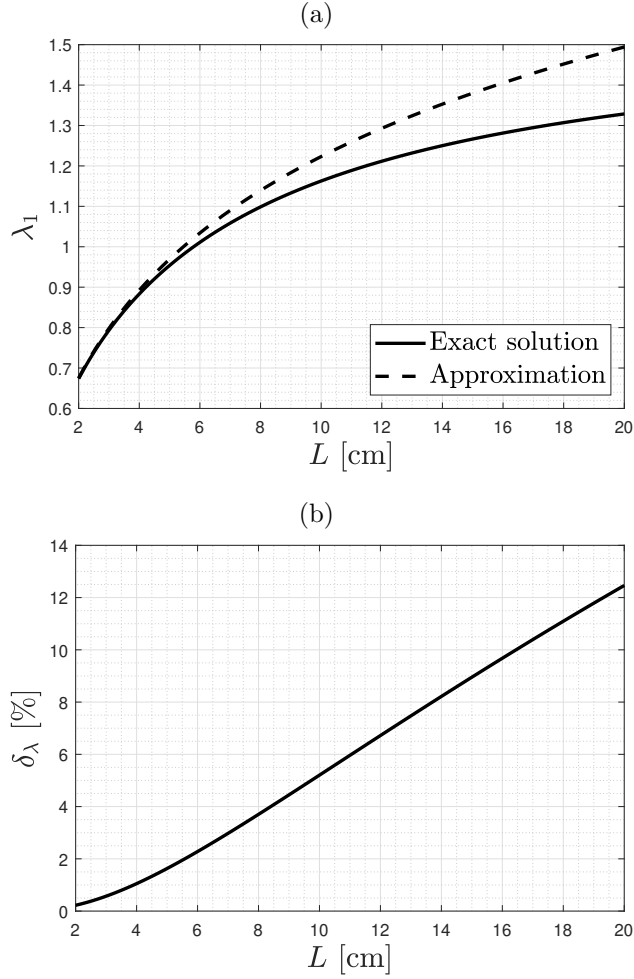


Figure 3: (a) Value of λ_1 according to the numerical solution of (36) and the approximation (42). (b) Relative error of the approximation (42), defined by $\delta_\lambda = 100 \times |\lambda_{1,\text{num}} - \lambda_{1,\text{approx}}| / \lambda_{1,\text{num}}$.

322 grows with L . When $L = 16\text{cm}$ the error is around 10%, beyond this depth
 323 we should expect solutions using this approximation to exhibit significant
 324 errors.

In practice, we can only use a finite amount of terms, hence we call T_N the approximation that uses N terms of the series (35). The energy (to N terms) absorbed within the layer is found by integrating (35)

$$E_N(t) = \frac{L^2 C_1}{D_s} \sum_{n=1}^N \frac{\beta_3 \psi_n(t)}{\lambda_n^4 + \lambda_n^2 \beta_3 (1 + \beta_3)}. \quad (44)$$

325 Again we may observe a strong dependence on C_1 , which represents the

326 energy gains. However, both C_1 and the energy loss term C_2 appear in
 327 ψ_n while $\beta_3 = C_2L/k_s$, so the dependence is not so clear as for the thick
 328 substrate.

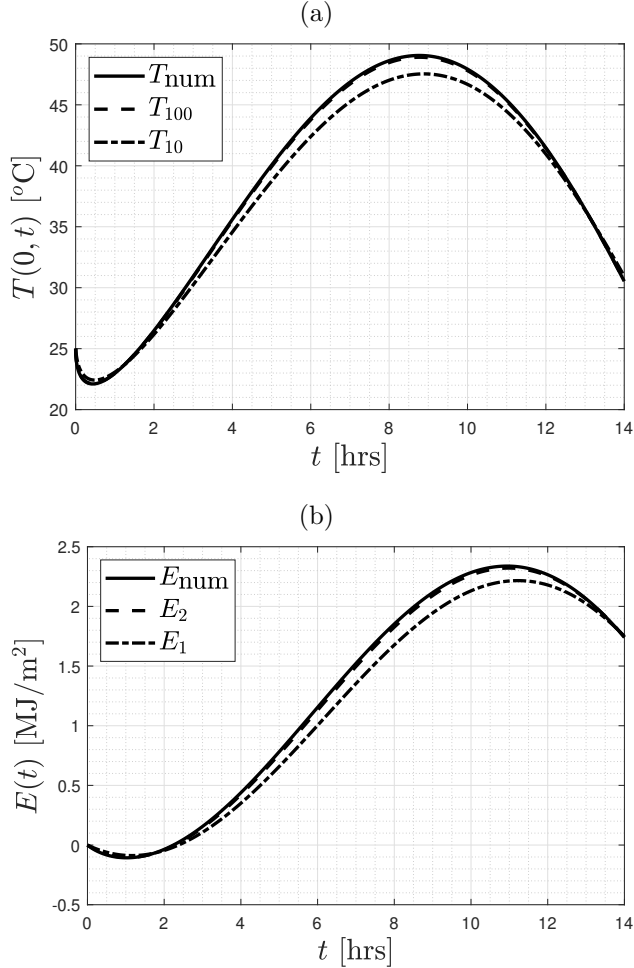


Figure 4: (a) Evolution of the surface temperature using the numerical solution and equation (41) retaining 10 and 100 terms in the expansion. (b) Evolution of the stored energy using the numerical solution and expression (44) with $N = 1, 2$.

329 In Figure 4 the series solution is compared with a numerical solution ob-
 330 tained using the Matlab *pdepe* function (parameter values are given in §4, the
 331 value $\alpha = 10^{-9}\text{m}/^\circ\text{Cs}$). In the case of the surface temperature the conver-
 332 gence rate of the series is slow such that only as $N \rightarrow 100$ do the numerical
 333 and analytical results coincide. The integration of the temperature results in
 334 the energy series being $\mathcal{O}(1/\lambda_n^2)$ smaller than that of the temperature, this
 335 leads to a much more rapid convergence rate: with $N = 1$ the error is small

336 while with $N = 2$ the series is almost indistinguishable from the numerical
 337 result.

If we define

$$E_1(t) = \frac{L^2 C_1 \rho_s c_s}{k_s} \frac{\beta_3 \psi_1(t)}{\lambda_1^4 + \lambda_1^2 \beta_3 (1 + \beta_3)} \quad (45)$$

338 and $E_{\text{app}}(t)$ as the value of E_1 after replacing λ_1 with λ_{app} and then define
 339 the relative error

$$\delta(t) = 100 \times \frac{|E_{\text{num}}(t) - E_{\text{app}}(t)|}{|E_{\text{num}}(t)|}, \quad (46)$$

340 we find $\delta(t_f) \approx 1\%$. The error with E_1 is even smaller. Hence we conclude
 341 that either equation (45) or E_{app} are sufficiently accurate to describe the
 342 energy in a shallow substrate (at least for large times, i.e, close to $t = t_f$).
 343 All terms of E_{app} may be written explicitly so this solution provides a signif-
 344 icantly more tractable form than the full series which requires the numerical
 345 calculation of the eigenvalues. We will verify the accuracy further in the
 346 results section.

For the case where we choose constant averaged values to represent the irradiance and ambient temperature, so that $G = 0$ and hence $\mathcal{G}_n = 0$, then

$$T_N^0(z, t) = T_{\text{mn}} + \frac{C_1 L \beta_2}{k} \sum_{n=1}^N \left\{ \frac{\cos(\lambda_n)}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)} \right. \\ \left. \times \cos\left(\lambda_n \left(1 - \frac{z}{L}\right)\right) \left(1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right)\right) \right\}, \quad (47)$$

which results in a surface temperature and energy

$$T_N^0(0, t) = T_{\text{mn}} + \frac{C_1 L \beta_2}{k} \sum_{n=1}^N \frac{1}{\lambda_n^2 + \beta_3^2 + \beta_3} \left(1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right)\right), \quad (48)$$

$$E_N^0(t) = \frac{L^2 (C_1 - C_2 T_{\text{mn}})}{D_s} \sum_{n=1}^N \left\{ \frac{\beta_3}{\lambda_n^4 + \lambda_n^2 \beta_3 (1 + \beta_3)} \right. \\ \left. \times \left(1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right)\right) \right\}. \quad (49)$$

347 This solution will be compared later with the thick substrate solution which
 348 requires $G = 0$. It is tempting to permit $L \rightarrow \infty$ in the above such that
 349 the term $1 - \exp(-\lambda_n^2 D_s t / L^2) \rightarrow \lambda_n^2 D_s t / L^2$, giving the appearance that
 350 for thick layers the surface temperature and energy grow proportional to

351 time, so contradicting the previous deep substrate results. However this
 352 naive expansion does not capture the correct behaviour since $\beta_3 \rightarrow \infty$ and
 353 $\lambda_n \rightarrow \sqrt{\beta_3}$ (independent of n) as $L \rightarrow \infty$.

354 In Figure 5 (a) we show a comparison of the surface temperature pre-
 355 dicted by the separable solution, (48), with $N = 100$, $L = 10, 15$ cm, and
 356 that obtained for an infinite depth substrate (30). Here it becomes clear that
 357 as L increases the separable solution approaches the infinite depth solution.
 358 All solutions in the figure have $G(t) = 0$, that is the temperature and irradi-
 359 ance are set to their average values. For this reason the surface temperature
 360 steadily increases during the day, in contrast to the result presented earlier
 361 in Figure 4, which has a variable energy input. For $L = 15$ cm the results
 362 are virtually identical for the whole day whereas for $L = 10$ cm they start
 363 to diverge after around 6 hours (when $\lambda_n^2 D_s t / L^2$ becomes $\mathcal{O}(1)$). The cor-
 364 responding energy stored in the layer is presented in Figure 5 (b), but now
 365 we only take two terms in the series solution. The nonlinear behaviour is
 366 apparent. Again, as the depth of the substrate increases the series solution
 367 approaches that of the Laplace solution, with an error of approximately 4%
 368 at the end of the day when $L = 15$ cm.

369 3.3. Method to determine the amount evaporated, $d(\alpha)$

370 Evaporation plays a key role in the absorption and storage of energy. It
 371 is one of the main reasons why a dark soil surface can remain cooler than a
 372 lighter concrete one: evaporation requires a lot of energy. Here we have taken
 373 a standard approximation by setting the evaporation rate proportional to the
 374 surface temperature, $\dot{d} = \alpha T(0, t)$. However, the constant of proportionality,
 375 which determines the energy loss term C_2 , is *a priori* unknown and must be
 376 calculated during the solution process.

377 In the case where the evaporation rate is assumed to be constant the
 378 solution is simple. Writing $\dot{d} = d/t_f$, where it is assumed that the liquid
 379 evaporated during the day, d , is a known depth, then α may be removed from
 380 the problem. The constants C_1, C_2 take on a slightly different definition

$$C_1 = (1 - F_{VC})(1 - A_s)Q_{av} + H_{as}T_{av} - \rho_w L_e \frac{d}{t_f}, \quad C_2 = H_{as}. \quad (50)$$

381 These different definitions do not affect the previous calculations, so the
 382 expressions derived in the preceding sections for temperature and energy
 383 absorbed still hold but with C_1, C_2 defined by (50).

In the more physically realistic situation, where evaporation varies through
 the day the constants are as defined in (20) and α may be determined by
 integrating the derived expression for $T(0, t)$ and employing equation (22) or

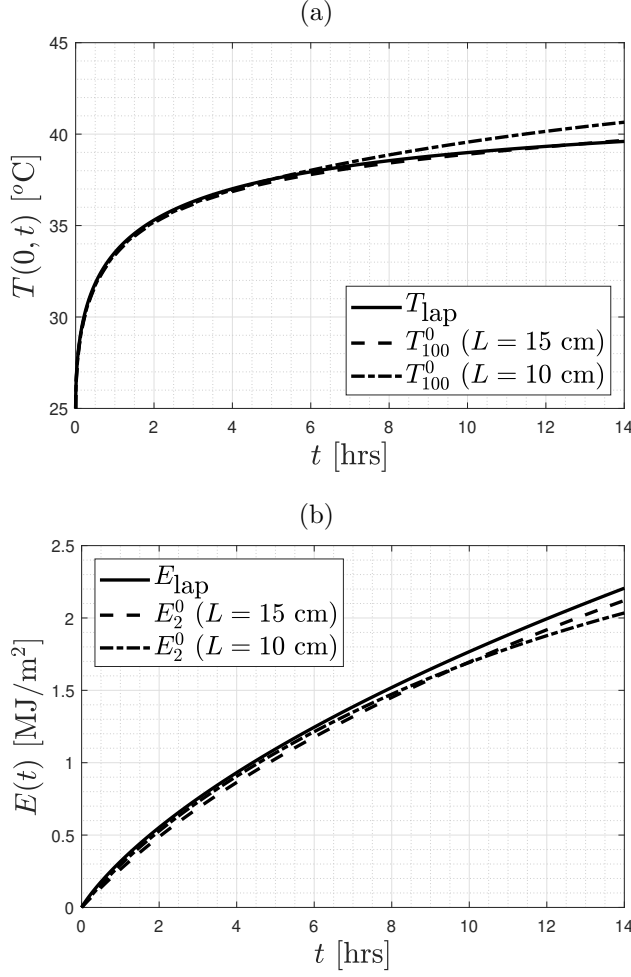


Figure 5: Comparison of separable (48) and Laplace (30) solutions for the surface temperature and energy when $G(t) = 0$ for $L = 10, 15$ cm: (a) Surface temperature taking $N = 100$ (b) Stored energy taking $N = 2$.

by equating the energy expression with the result of equation (27). For the case of a thick layer setting $t = t_f$ in (31) and equating the result with (27) we obtain

$$d = \frac{\alpha}{C_2} \left[C_1 t_f - \frac{k_s \rho_s c_s}{C_2^2} (C_1 - C_2 T_{\text{mn}}) \right. \\ \left. \times \left(\exp \left(\frac{C_2^2 D_s}{k_s^2} t_f \right) \operatorname{erfc} \left(\frac{C_2}{k_s} \sqrt{D_s t_f} \right) + \frac{2}{\sqrt{\pi}} \frac{C_2}{k_s} \sqrt{D_s t_f} - 1 \right) \right], \quad (51)$$

For thinner layers, setting $t = t_f$ in (44) and equating with (27) we obtain

$$d = \frac{\alpha C_1}{C_2} \left[t_f - \frac{L^2 \rho_s c_s}{k_s} \sum_{n=1}^{\infty} \frac{\beta_3 \psi_n(t_f)}{\lambda_n^4 + \lambda_n^2 \beta_3 (1 + \beta_3)} \right], \quad (52)$$

384 where λ_n is obtained by solving (36) and $\psi_n(t_f)$ given explicitly by (37).

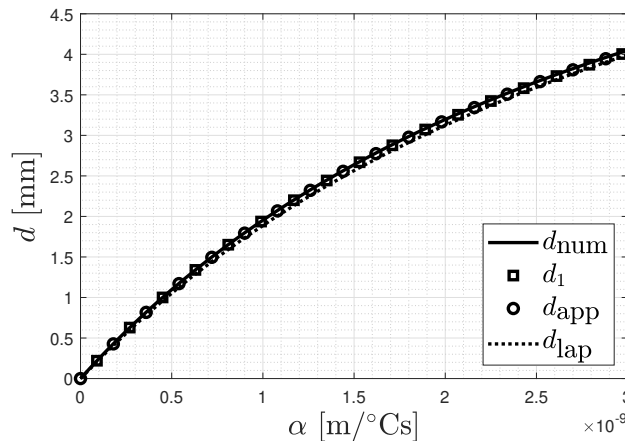


Figure 6: Depth evaporated as a function of α , for $L = 8\text{cm}$, obtained by numerical integration, d_{num} , the solution of (52) using only the first term with the exact value of λ_1 , denoted d_1 , and using the first term with the approximate value of λ_1 , denoted d_{app} , and the solution of (51), denoted d_{lap}

In Figure 6 we show the depth of water evaporated as a function of α using infinite and finite depth solutions, (51, 52). Having demonstrated that the energy converges rapidly and that $E_1(t_f)$ agrees very well with the full numerical solution for the finite depth we only use the first term of the finite depth solution which defines

$$d_1 = \frac{\alpha C_1}{C_2} \left[t_f - \frac{L^2 \rho_s c_s}{k_s} \frac{\beta_3 \psi_1(t_f)}{\lambda_1^4 + \lambda_1^2 \beta_3 (1 + \beta_3)} \right]. \quad (53)$$

385 Taking the exact value of λ_1 , determined by solving $\lambda_1 \tan \lambda_1 = \beta_3$ numeri-
 386 cally, we obtain the curve labelled d_1 . Using the approximate expression for
 387 the first eigenvalue, (42), gives d_{app} . Also shown is the numerical solution of
 388 the heat equation. In all cases we take $L = 8\text{cm}$ and parameter values spec-
 389 ified in the following section. Clearly the agreement between the four forms
 390 is excellent. In each case the depth increases non-linearly with α . Although
 391 calculated with $L = 8\text{cm}$ we found virtually identical results for larger values,
 392 such that for sufficiently thick layers we may assume that $d(\alpha)$ is independent
 393 of depth. This may be attributed to the fact that evaporation is a surface

394 effect. Here affecting a maximum of approximately 4.5mm into the soil layer.
 395 The soil far below this, such that $L \gg 4.5\text{mm}$ will have little influence on
 396 this process. From the deep substrate result (51) we know that d depends
 397 linearly on C_1 . From Fig. 6 it is clear that the computed value of d is almost
 398 identical for both shallow and deep layers. As a consequence the energy absor-
 399 bed, expressed by (27), is linear in C_1 for any depth of substrate (greater
 400 than a few millimetres). This means that the energy decreases linearly with
 401 fractional leaf coverage F_{VC} , albedo, A_s , and increases linearly with irradi-
 402 ance Q_{net} and average daily temperature. The nonlinear dependence of both
 403 expressions on C_2 shows that $E(t_f)$ depends nonlinearly on the evaporation
 404 rate and heat transfer coefficient. This will be demonstrated in §4.3.

405 4. Results

406 After developing a set of solutions appropriate for thick and thin sub-
 407 strates the goal is now to see how they behave under realistic conditions and
 408 in doing so determine guidelines or advice concerning the design of green
 409 roofs.

410 The thermophysical properties of soils vary significantly depending on the
 411 composition and moisture content. Coma *et al.* [36] analysed the properties
 412 of five different dried soils for extensive green roofs, changing the percentage
 413 of compost, coco peat, crushed building wastes, coarse grained sand and
 414 pozzolana, to obtain $k_s \in [0.1, 0.19]$ W/m°C, $c_s \in [724, 873]$ J/kg°C, $\rho_s \in$
 415 $[375, 1360]$ kg/m³. In [37] the soil layer has the following properties: $k_s =$
 416 0.27 W/m°C, $c_s = 1307$ J/kg°C, $\rho_s = 1210$ kg/m³, the layer thickness was
 417 $L = 8$ cm and the albedo $A_s = 0.26$. We will use this latter soil as a reference
 418 case and choose a fractional vegetative cover of $F_{VC} = 0.5$.

419 The typical properties of concrete used in construction may also vary
 420 with composition. For example, the density of regular concrete is around
 421 2400 kg/m³ while lightweight concrete has $\rho_c \approx 1750$ kg/m³ [38]. In our
 422 subsequent analysis, we will use parameter values consistent with HSC con-
 423 crete as presented in Table 2 [39, 40]. With regard to the ambient conditions,
 424 unless otherwise specified, we will use the parameter values given in Table
 425 3, which are consistent with a roof in a city with warm but not extreme
 426 temperatures and with a medium irradiance.

427 4.1. Temperature and energy profiles

428 In this section we focus on a green roof using the data provided in Tables 2,
 429 3 and $A_s = 0.26$, $L = 8\text{cm}$ unless specified otherwise. The numerical solution
 430 of the heat equation (1) subject to (19) is used to determine the temperature
 431 at the soil-air interface and the energy stored as a function of time in Figure

Table 2: Representative values for the thermophysical properties of soil and concrete [37, 38, 40, 41].

Material	ρ_i (kg/m ³)	c_i (J/kg°C)	k_i (W/m°C)
soil ($i=s$)	1210	1307	0.27
concrete ($i=c$)	2300	1000	1.5

Table 3: Parameters used to represent ambient conditions, values for H_{as} and Q_{\max} are taken from [42], [43] respectively.

Maximum temperature	T_{\max}	30	°C
Minimum temperature	T_{\min}	20	°C
Maximum irradiance	Q_{\max}	800	W/m ²
Density of water	ρ_w	1000	kg/m ³
Soil-air heat transfer coefficient	H_{as}	5	W/m ² °C
Latent heat of evaporation	L_e	2.26×10^6	J/kg

432 7. The left panel shows the evolution of the surface temperature for different
433 thicknesses while the right one corresponds to the evolution of the absorbed
434 energy. The surface temperature is always highest with the thinnest layer,
435 reaching a maximum of approximately 52°C soon after midday (which occurs
436 at $t = 7$ hours). As the layer thickness increases the surface temperature
437 decreases, since the heat can diffuse further into the layer. The two plots for
438 $L = 13, 15$ cm are difficult to distinguish, suggesting that for $L > 15$ cm the
439 surface temperature is independent of depth (and hence the infinite depth
440 solution will be accurate). The higher surface temperature results in a higher
441 evaporation rate and convective cooling and for this reason less energy is
442 retained in the thinner soil layer. Consequently, in contrast to the surface
443 temperature, the energy stored is greater in the thicker layers, attaining a
444 maximum of around 2.4 MJ/m² when $L = 15$ cm while for the 5cm layer the
445 maximum is close to 2 MJ/m². Thicker layers clearly have the ability to
446 store more energy whilst maintaining a lower surface temperature.

447 4.2. Effect of evaporation

448 In Figure 8 we plot the energy at the end of the day as a function of
449 the evaporated depth as predicted by integrating the numerical solution for
450 the full problem via equation (23) with $G \neq 0$, denoted $E_{\text{num}}(t_f)$, and with
451 the average values (hence $G = 0$), $E_{\text{num}}^0(t_f)$. Also shown are the analytical
452 solutions, the Laplace solution, (31), $E_{\text{lap}}(t_f)$, which has $G = 0$ and the
453 approximate separable solution E_{app} , which has $N = 1$ and $\lambda_1 = \lambda_{\text{app}}$. As
454 discussed earlier the energy decreases linearly with d , due to the increasing

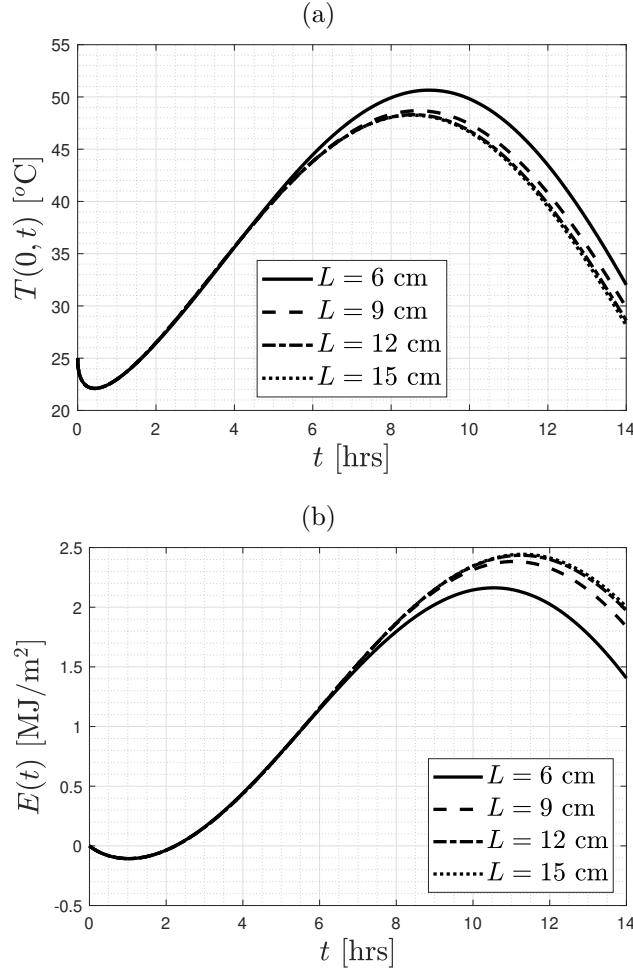


Figure 7: Numerical solutions for (a) Evolution of the temperature at the air-soil surface for different soil thicknesses, (b) Evolution of the energy for different soil thicknesses. In both cases $\alpha = 1 \text{ nm}/^\circ\text{Cs}$.

455 energy removed by evaporation. In Figure 8(a) the Laplace solution provides
 456 a poor approximation, indicating that this substrate is too shallow for the
 457 infinite layer analysis. The approximate solution E_{app} shows good agreement
 458 with the full numerical solution, demonstrating that the behaviour is well
 459 captured by this simple explicit representation. The numerical solution with
 460 $G = 0$ shows an error of approximately 3% which results from neglecting the
 461 variation in ambient temperature and irradiance. Increasing the substrate
 462 depth to 15cm on Figure 8(b) it may be seen that the Laplace solution now
 463 becomes accurate, within the assumption of $G = 0$: the agreement with
 464 E_{num}^0 is excellent. Surprisingly E_{app} also appears to closely match these

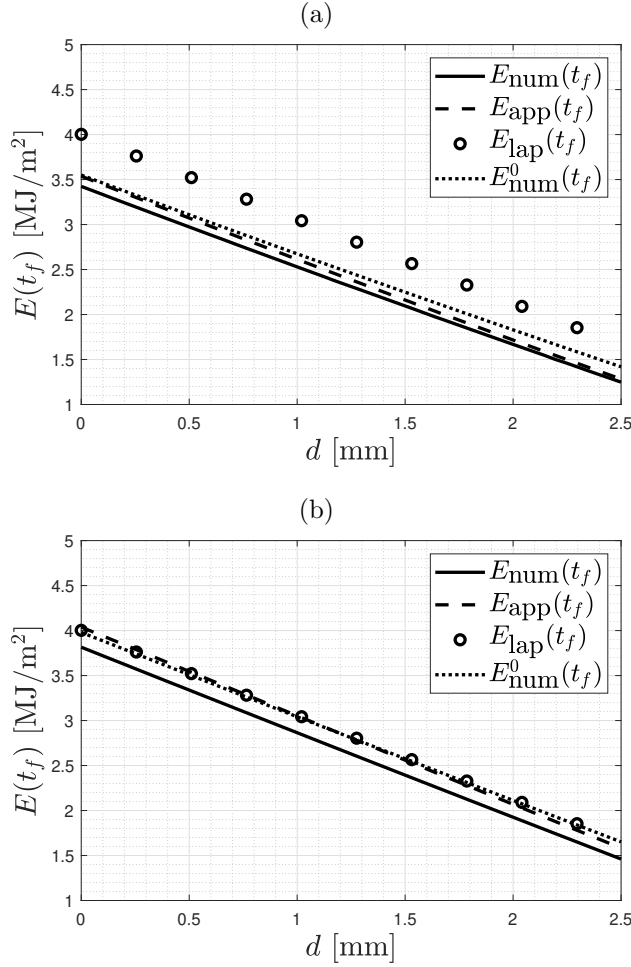


Figure 8: Energy absorbed at the end of the day using the numerical solution of the heat equation and expression (23) with $G \neq 0$, denoted $E_{\text{num}}(t_f)$, and with $G = 0$, denoted $E_{\text{num}}^0(t_f)$, also shown are (45) for E_{app} and (31) for $E_{\text{lap}}(t_f)$: (a) Soil thickness of $L = 8$ cm. (b) Soil thickness of $L = 15$ cm.

465 two solutions. The distance from the full numerical solution is a result of
 466 the approximation to λ_1 which becomes worse with increasing depth, for
 467 $L = 15$ cm the error in λ_1 is of the order 9% but this only results in an error
 468 for the energy prediction of approximately 5%.

469 4.3. Albedo and vegetative cover effects

470 The albedo and the fractional vegetative cover are properties which di-
 471 rectly affect the amount of solar irradiance reaching the surface and conse-
 472 quently must play a strong role in the energy storage. As discussed in §3.3
 473 for a thick layer the energy absorbed varies linearly with C_1 and so must

474 decrease linearly with increasing F_{VC} and A_s . For a thin layer the depen-
475 dence is not clear from the analytical solution but the fact that the deep and
476 shallow results coincide for the depth $d(\alpha)$ suggests a linear decrease with C_1 .
477 In Figure 9 we show the dependence of the total energy as a function of the
478 fractional vegetative cover F_{VC} when $L = 8\text{cm}$ and with various evaporation
479 rates. The linear decrease is apparent in all cases, even though the layer is
480 not sufficiently thick to warrant applying the infinite thickness solution. As
481 we increase the vegetative cover to unity, with $\alpha = 0$ (that is no evaporation),
482 the stored energy becomes very small. This is to be expected, there is no in-
483 coming irradiance so energy only passes to the substrate through convective
484 heat transfer which is much less efficient. Increasing the amount evaporated
485 the energy at the end of the day can become negative, that is there is less
486 energy in the substrate than at the beginning of the day. This demonstrates
487 the strong effect evaporation can have on the process. Since the incoming
488 energy depends on $(1 - F_{VC})(1 - A_s)Q_{\text{net}}$ exactly the same behaviour can be
489 expected by fixing F_{VC} and varying the albedo. Consequently Figure 9 also
490 holds for varying $A_s \in [0, 1]$ with fixed F_{VC} .

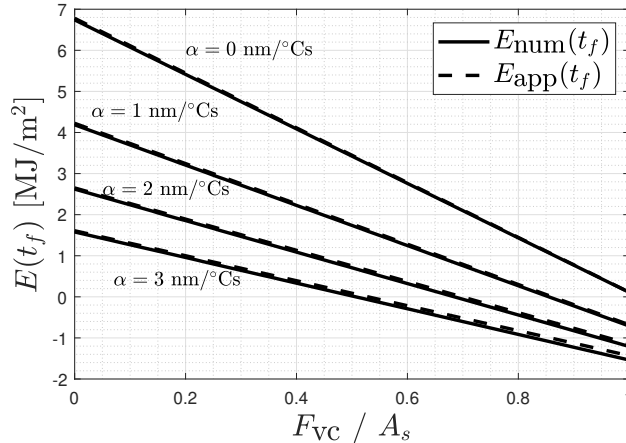


Figure 9: Total energy stored as a function of F_{VC} or A_s using direct numerical simulation and the approximate expression (45) for different values of α .

491 4.4. Surface temperature and energy stored in concrete layers

492 The model is easily applied to a concrete or other man-made surface by
493 setting $F_{VC} = d = 0$ in all expressions. Thermophysical parameters are
494 shown in Table 2 and typical values for the albedo are shown in Table 4:
495 unpainted concrete has values between 0.2 and 0.45. In Figure 10(a) we
496 show the surface temperature with $A_c = 0.3$ for three different concrete
497 layer thicknesses. As before the thinner layers show a higher value. In

Table 4: Typical albedo values for different types of surfaces.

Material	Albedo	Reference
concrete	0.2-0.45	[44–46]
asphalt	0.05-0.20	[45, 46]
brick, stone	0.20–0.40	[46]
sandy soil	0.25-0.45	[47]
bare fields	0.1-0.25	[47]
grass, bushes	0.16-0.27	[37, 43, 47]
trees	0.15-0.18	[46]
white paint	0.5-0.9	[46]
black paint	0.05	[48]

498 order to compare with a green roof, we also show the equivalent result taken
 499 from Figure 7(a). For the 6cm layer the maximum difference in surface
 500 temperatures between green and concrete roofs is close to 40°C. Even for
 501 the 12cm layer the difference is of the order 30°C. This clearly demonstrates
 502 how green roofs can affect comfort levels. Figure 10(b) shows the energy
 503 stored in a concrete roof for $L \in [6, 16]$ cm when $A_c = 0.3$, we also present a
 504 second result corresponding to using a reflective paint, with $A_c = 0.7$. The
 505 corresponding green roof result is shown as a solid line. With the unpainted
 506 surface the energy stored ranges between four to six times that in a typical
 507 green roof. Even with the reflective coating the energy stored is typically
 508 twice that of the green roof. All this extra energy is available for release
 509 during the night and so will add to the heat island effect.

510 4.5. Comparison between intensive and extensive green roofs

511 The key difference between intensive and extensive green roofs is the
 512 depth of the substrate. Extensive roofs typically have a substrate between
 513 approximately 5-12cm while intensive roofs are deeper.

514 The deeper growing medium of an intensive roof permits a wide range of
 515 plant types, ranging from grasses to shrubs or even trees. These large plants
 516 provide a lot of shade on the soil, so their mean leaf area index (LAI) ranges
 517 from 1 to 6 or even 7 (see [49]), and they have been found to have albedo
 518 values ranging from 0.16 to 0.28 (see [50]).

519 Extensive green roofs typically allow for hardy plants which require min-
 520 imum maintenance. Since these plants are typically smaller, so are their
 521 leaves and they usually require little watering. Extensive roofs are usually
 522 “brownier”, so their albedo is low. Since plants of extensive green roofs are
 523 typically grass or have small leaves, their LAI is also small, ranging from 0.8
 524 to 2.

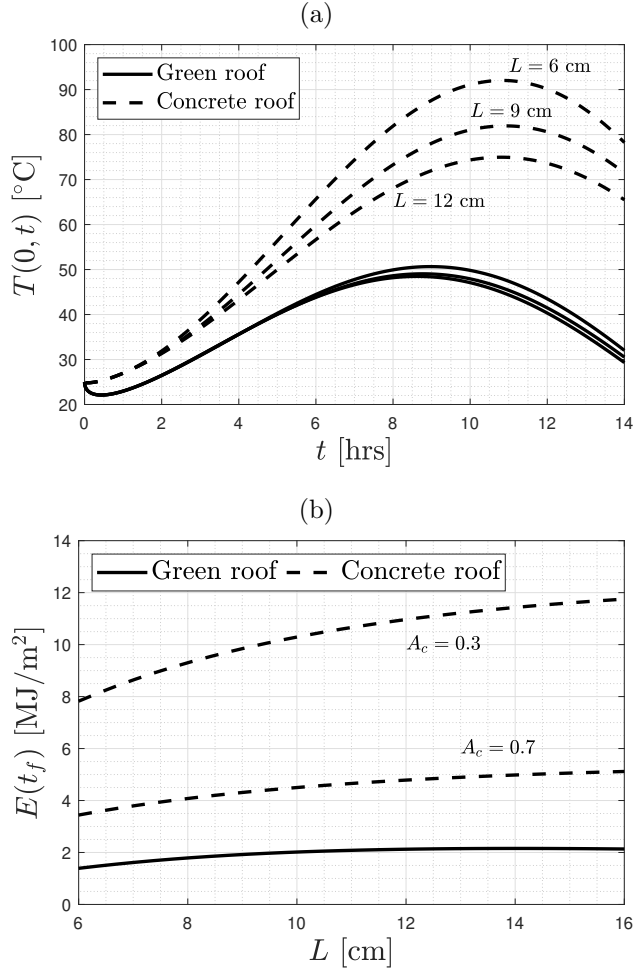


Figure 10: (a) Surface temperature for a concrete and green roofs using direct numerical simulation. Albedo value for concrete used here is $A_c = 0.3$. Curves relative to green roofs are plotted for comparative purposes, we suggest observing Figure 7(a) for a better reference. (b) Total energy stored at the end of the day. The different albedo values correspond to unpainted ($A_c = 0.3$) and painted ($A_c = 0.7$) concrete.

525 These two types of green roofs are dramatically different in terms of
 526 maintenance, leading to the natural question whether it is worth investing
 527 in a more sophisticated “rooftop garden” rather than simply providing a
 528 shallow layer of soil with small vegetation. We therefore now use our model
 529 and consider two different set of values for the albedo and the vegetative
 530 coverage and compare the total energy stored in the two cases as a function
 531 of the soil thickness. We will consider a range of values in the two cases.

532 Table 5 shows the values used in the numerical simulations. The rest of the
 533 parameters are kept the same as in the previous sections.

Table 5: Representative values for intensive and extensive green roofs [49, 50].

Type of roof	A	LAI	F_{VC}	L (cm)
Intensive	0.26	2.5	0.92	12-20
Extensive	0.20	1	0.63	5-12

534 Figure 11 shows the evolution of $E(t)$ over the day for the two roof types.
 535 We recall that our models are valid for small evapotranspiration rates and
 536 root uptake. The predictions should therefore be treated with caution when
 537 dealing with large plants, for this reason we use a moderate value for the LAI
 538 in the intensive case. The curves were obtained by numerically computing the
 539 evolution of the energy absorbed during the day for a range of soil thicknesses
 540 and then calculating the mean energy for each range. For extensive roofs we
 541 used 5-10.5cm, whereas for the intensive ones we used 12-20cm (as shown in
 542 Fig. 10b) the energy shows only a weak depth dependence in these ranges).
 543 In this case the total stored energy at the end of the day for extensive green
 544 roofs is around 1.5 MJ/m². For the intensive green roof a negative value,
 545 around -0.4 MJ/m², is obtained. The negative value indicates that energy
 546 has been lost during the day, so highlighting the significant advantage of
 547 intensive over extensive roofs.

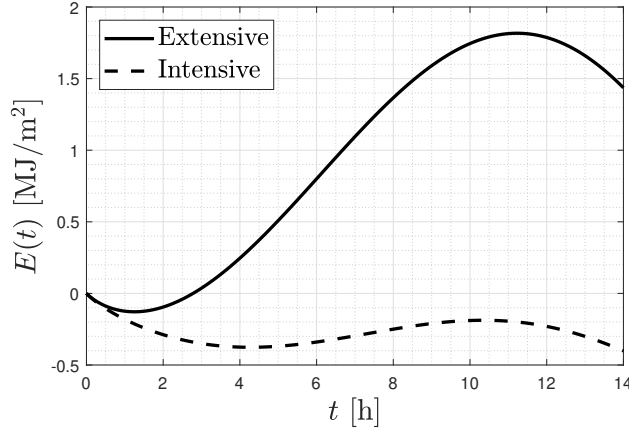


Figure 11: Time evolution of the absorbed energy for intensive and extensive green roofs.

548 4.6. Non-insulated soil bottom

549 In the original model, the boundary conditions at the bottom of the soil
 550 represent two possible scenarios: either the temperature there remains at the

551 initial temperature during the whole day or there is no heat transfer from the
 552 soil to the building. In practice, we may wish to extend the second case and
 553 assume that a certain amount of heat is lost into the building. This requires
 554 modifying the boundary condition at $z = L$ to

$$-k_s \left. \frac{\partial T}{\partial z} \right|_{z=L} = \epsilon C_1, \quad \epsilon > 0. \quad (54)$$

This form indicates that the heat lost at the base is proportional to that input at the surface, represented by C_1 . The constant of proportionality, $\epsilon > 0$, is expected to be small and in the limit $\epsilon \rightarrow 0$ the insulated boundary condition is retrieved. Imposing (54) the temperature can be expressed in the form of equation (35), although now the expression for ψ_n is

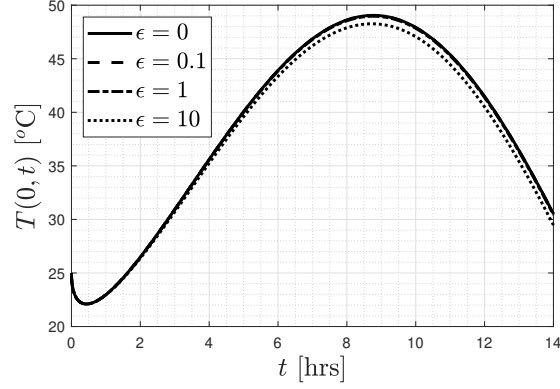
$$\psi_n(t; \epsilon) = \left(\beta_2 - \frac{\epsilon}{\cos(\lambda_n)} \right) \left[1 - \exp\left(-\frac{\lambda_n^2 D_s t}{L^2}\right) \right] + \lambda_n^2 \mathcal{G}_n(t), \quad (55)$$

555 see Appendix A for the detailed procedure. In particular, we note that
 556 equation (55) reduces to (37) in the limit $\epsilon \rightarrow 0$.

557 In Fig. 12 we show the effect of the parameter ϵ on the surface temper-
 558 ature and the amount of energy stored during the day for a green roof with
 559 $L = 8$ cm. The temperature is truncated after 100 terms, the energy after
 560 2 terms (these results are almost identical to the numerical solution). From
 561 Fig. 12a) it is apparent that the substrate heat loss has little impact on the
 562 surface temperature, which is dominated by surface effects. Since energy is
 563 now being lost at the bottom of the substrate ϵ has a much stronger influence
 564 on the absorbed energy, which decreases by around 15% at the end of the
 565 day as ϵ increases from 0 to 10, as shown in Fig. 12b).

566 To estimate a typical value for ϵ we consider a situation where there is a
 567 room with temperature $T_{room} = 25$ °C located at $z = L$, i.e. directly below
 568 the green roof. The right hand side of Eq. (54) represents the heat flux
 569 flowing into the room through the ceiling, which can be expressed as $q_{room} =$
 570 $H_{cr}(T_{ceiling} - T_{room})$ where H_{cr} is the ceiling-room heat transfer coefficient and
 571 $T_{ceiling}$ is the temperature of the ceiling. Heat transfer coefficients between
 572 indoor air and ceilings in regular buildings have values around $5 \text{ W/m}^2 \text{ °C}$
 573 or smaller [51, 52]. To obtain an upper bound for ϵ we assume that the
 574 temperature of the ceiling takes a very high value $T_{ceiling} = \max[T(z = 0, t)]$
 575 where $T(z = 0, t)$ is the temperature at the roof surface for the insulated
 576 case (in practice $T_{ceiling}$ would be much lower than the surface temperature).
 577 Combining the expression for q_{room} with the right hand side of Eq. (54)
 578 shows $\epsilon = H_{cr}(T_{ceiling} - T_{room})/C_1$. Using the results for both green and
 579 concrete roofs with $L = 6$ cm as shown in Fig. 10, we have $T_{ceiling}^{GR} = 50.65$ °C,

(a)



(b)

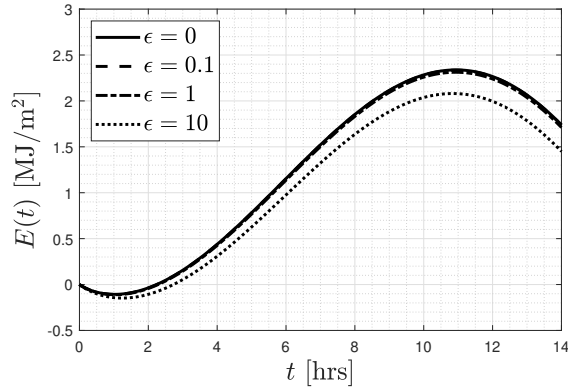


Figure 12: Evolution of (a) surface temperature and (b) stored energy for different values of the heat loss parameter ϵ . In both cases $L = 8\text{cm}$, $\alpha = 1\text{nm}/^\circ\text{Cs}$.

580 $C_1^{GR} = 315.9$ and $T_{ceiling}^{concrete} = 92.03^\circ\text{C}$, $C_1^{concrete} = 483.97$. These values lead
 581 to $\epsilon^{GR} = 0.41$ and $\epsilon^{concrete} = 0.69$, respectively. So, even in the $L = 6\text{cm}$
 582 case which has the highest temperature of our calculations and with no heat
 583 loss through the layer, the upper bound has $\epsilon < 1$. Referring to Fig. 12
 584 which shows negligible differences between the $\epsilon = 0$ and $\epsilon = 1$ cases we may
 585 conclude that the insulating boundary condition is sufficient for all practical
 586 purposes.

587 5. Conclusions

588 The primary aims of this paper were to develop a mathematical model
589 for heat flow in a green roof, to develop analytical solutions and to use these
590 to gain a better understanding of the process and finally, to evaluate the
591 differences in energy storage between green and concrete or man-made roofs.
592 The intention being that such a model could provide simple guidelines with
593 regard to the design of green roofs.

594 Various analytical solutions were presented, which explicitly define the
595 role of system parameters. A numerical solution was also developed this per-
596 mitted verification of the analytical solutions but could not clarify the role of
597 individual parameters. For an infinitely thick substrate an analytical expres-
598 sion was presented, provided that the daily variation of ambient temperature
599 and irradiance is averaged. For large time, close to the end of the day, this
600 solution demonstrated a simple linear dependence of the energy on many of
601 the ambient conditions, while the energy varied with the square root of time
602 and thermophysical properties of the substrate. For finite thickness layers a
603 separable solution was found. The energy expression converged rapidly at
604 large times and so could be written in an explicit form, involving only the
605 first term, which again permits the role of ambient conditions to be clearly
606 understood.

607 Comparison between the analytical solutions and the numerical solution
608 demonstrated that:

- 609 1. The separable solution for the energy rapidly converges such that by the
610 end of the day only a single term is required. The surface temperature
611 converges much more slowly.
- 612 2. The infinite depth solution holds for substrates deeper than approxi-
613 mately 15cm.
- 614 3. The averaging of the ambient temperature and irradiance results in
615 slightly higher predictions for energy absorption, but the trend is the
616 same as the numerical solution of the full, time-dependent system.

617 Given that the majority of green roofs have a depth greater than 15cm this
618 final point suggests that the Laplace solution provides a simple way to quan-
619 tify the energy storage.

620 Significant conclusions from the model include:

- 621 1. Subject to identical conditions, e.g. the same albedo, solar radiation,
622 vegetative coverage etc, an extensive (thin layer) green roof will ex-
623 hibit higher surface temperatures than an intensive (thick) roof but
624 will absorb less energy. However, in practice an intensive roof typically
625 has the higher vegetative coverage and albedo. In which case intensive

626 roofs absorb significantly less energy and so, whenever possible, are
627 preferable to extensive roofs.

628 2. Concrete roofs are significantly hotter and store significantly more en-
629 ergy than a standard green roof. Our example showed a maximum
630 surface temperature almost 40°C higher and a factor three more en-
631 ergy stored between concrete and green roofs.

632 If the energy stored in a green roof is a factor x less than that in an equivalent
633 concrete layer then we may easily deduce that a 10% increase in an area's
634 green coverage corresponds to a $10(x-1)/x$ % decrease in stored energy. For
635 the present example $x \approx 3$ which corresponds to a 6.7% decrease in energy
636 stored for every extra 10% of green area.

637 With regards to guidelines to aid in the design of green roofs, within the
638 restrictions of the present model, the results show that the energy absorbed:

- 639 1. Decreases linearly with: depth of water evaporated; fractional vegeta-
640 tive coverage; albedo.
- 641 2. Increases linearly with average ambient temperature.
- 642 3. Varies non-linearly with surface temperature (related to the evapora-
643 tion rate) and the surface heat transfer coefficient.
- 644 4. Increases with the square root of the thermophysical properties such as
645 density of the soil, thermal conductivity and specific heat capacity.

646 Obviously not all of these may be controlled but ensuring good leaf coverage,
647 a lighter coloured surface and a wet surface layer are possible. These are the
648 most important factors. Choosing a substrate with a low density, conduc-
649 tivity and heat capacity will, to a lesser extent, reduce the energy storage
650 capacity.

651 While there exist many more complex descriptions of heat flow in urban
652 landscapes, including detailed descriptions of air flow and energy input from
653 human activity it is imperative that these models are founded on the correct
654 building block. The current model is designed to achieve this. It may be
655 adapted to incorporate empirical modifications, such as those employed in
656 the Penman-Monteith equation where, for example, the surface heat transfer
657 accounts for the air speed some distance above the ground, or the evaporation
658 is split into evaporation and transpiration. However, swapping the heat
659 transfer coefficient for an air speed expression or removing liquid from inside
660 the soil layer as well as at the surface will not change the basic conclusions.
661 Other modifications, which can add to the generality of the model include
662 permitting heat transfer between the green roof and the building surface, so
663 paving the way to study the effect inside the building, and also permitting
664 moisture movement within the soil.

665 **Author Statement**

666 The work has not been published previously and it is not under consider-
667 ation for publication elsewhere. Its publication is approved by all authors. If
668 accepted, it will not be published elsewhere in the same form, in English or
669 in any other language, including electronically without the written consent
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671 **Declaration of Competing Interest**

672 All authors have participated in (a) conception and design, or analysis
673 and interpretation of the data; (b) drafting the article or revising it criti-
674 cally for important intellectual content; and (c) approval of the final version.
675 This manuscript has not been submitted to, nor is under review at, another
676 journal or other publishing venue. The authors have no affiliation with any
677 organization with a direct or indirect financial interest in the subject matter
678 discussed in the manuscript.

679 **CRedit authorship contribution statement**

680 **Maria Aguares:** Conceptualization, Methodology, Formal analysis,
681 Writing - original draft, Software. **Marc Calvo-Schwarzwalder:** Formal
682 analysis, Writing - review & editing. **Francesc Font:** Conceptualization,
683 Methodology, Formal analysis, Writing - original draft, Software. **Timothy**
684 **G. Myers:** Supervision, Conceptualization, Writing - review & editing.

685 **Acknowledgements**

686 M. Aguares acknowledges grant no. MTM2017-84214-C2-2-P. M. Calvo-
687 Schwarzwalder acknowledges financial support from Zayed University through
688 the grant no. R21033 and the Policy Research Incentive Program 2022. F.
689 Font acknowledges the support of the Serra-Hunter Programme of the Gen-
690 eralitat de Catalunya. T.G. Myers thanks the CERCA Programme of the
691 Generalitat de Catalunya. All authors acknowledge the support of Ministerio
692 de Ciencia e Innovación Grant No. PID2020-115023RB-I00.

693 **Appendix A. Temperature profile for infinite depth substrate**

694 The model presented in this work can be solved analytically if we assume
695 that

- 696 1. the air temperature and sun irradiation remain constant along the day;

697 2. the thickness of the soil can be assumed to be infinite.

Under these conditions, the governing equations are

$$\frac{\partial T}{\partial t} = D_s \frac{\partial^2 T}{\partial z^2}, \quad T(z, 0) = T_{\text{mn}}, \quad (\text{A.1a})$$

$$-k_s \frac{\partial T}{\partial z} \Big|_{z=0} = C_1 - C_2 T(0, t), \quad \frac{\partial T}{\partial z} \Big|_{z \rightarrow \infty} = 0. \quad (\text{A.1b})$$

698 To simplify notation and to understand the contributions and relative
 699 importance of the different terms in the equations, it is helpful to formulate
 700 the problem first in terms of scaled variables. The arising dimensionless
 701 numbers can then yield some interesting behaviour of the solutions.

702 *Appendix A.1. Non-dimensional formulation*

703 For the problem defined by (A.1) there is a clearly defined time scale t_f .
 704 The length and temperature scales can be determined by writing $z = \mathcal{L}\hat{z}$,
 705 $t = t_f \hat{t}$ and $T = T_{\text{mn}} + \mathcal{T}\hat{T}$, which leads to

$$\frac{\partial \hat{T}}{\partial \hat{t}} = \frac{\partial^2 \hat{T}}{\partial \hat{z}^2}, \quad \hat{T}(\hat{z}, 0) = 0, \quad \frac{\partial \hat{T}}{\partial \hat{z}} \Big|_{\hat{z} \rightarrow \infty} = 0, \quad (\text{A.2})$$

706 provided we choose $\mathcal{L} = \sqrt{t_f/D_s}$. The boundary condition at the surface
 707 reads

$$-\frac{k_s \mathcal{T}}{\mathcal{L}} \frac{\partial \hat{T}}{\partial \hat{z}} \Big|_{\hat{z}=0} = C_1 - C_2 T_{\text{mn}} - C_2 \mathcal{T} \hat{T}(0, \hat{t}). \quad (\text{A.3})$$

708 Upon choosing the temperature scale $\mathcal{T} = C_1 \mathcal{L}/k_s$, this condition reduces to

$$-\frac{\partial \hat{T}}{\partial \hat{z}} \Big|_{\hat{z}=0} = \beta_2 - \beta_3 \hat{T}(0, \hat{t}), \quad (\text{A.4})$$

709 where

$$\beta_2 = 1 - \frac{C_2 T_{\text{mn}}}{C_1}, \quad \beta_3 = \frac{C_2 \mathcal{L}}{k_s}. \quad (\text{A.5})$$

710 *Appendix A.2. Analytical solution*

Taking the Laplace transform of the system defined by Eqs. (A.2) and (A.4) we obtain

$$\hat{s} \hat{u} = \frac{\partial^2 \hat{u}}{\partial \hat{z}^2}, \quad \hat{u}(\hat{z}, 0) = 0, \quad (\text{A.6a})$$

$$-\frac{\partial \hat{u}}{\partial \hat{z}} \Big|_{\hat{z}=0} = \frac{\beta_2}{s} - \beta_3 \hat{u}(0, \hat{s}), \quad \frac{\partial \hat{u}}{\partial \hat{z}} \Big|_{\hat{z} \rightarrow \infty} = 0, \quad (\text{A.6b})$$

711 where $\hat{u}(\hat{z}, \hat{s})$ is the Laplace transform of $\hat{T}(\hat{z}, \hat{t})$. The solution to the ordi-
 712 nary differential equation is

$$\hat{u}(\hat{z}, \hat{s}) = A \exp\left(-\sqrt{\hat{s}}\hat{z}\right), \quad (\text{A.7})$$

where the positive branch has been neglected to satisfy the far-field condition. Applying the boundary condition at $\hat{z} = 0$ determines

$$A = \frac{\beta_2}{\hat{s}(\beta_3 + \sqrt{\hat{s}})}, \quad (\text{A.8})$$

and hence we can transform back to obtain

$$\begin{aligned} \hat{T}(\hat{z}, \hat{t}) = & \frac{\beta_2}{\beta_3} \left[\operatorname{erfc}\left(\frac{\hat{z}}{2\sqrt{\hat{t}}}\right) \right. \\ & \left. - \exp\left(\beta_3(\beta_3\hat{t} + \hat{z})\right) \operatorname{erfc}\left(\beta_3\sqrt{\hat{t}} + \frac{\hat{z}}{2\sqrt{\hat{t}}}\right) \right], \end{aligned} \quad (\text{A.9})$$

where $\operatorname{erfc}(z)$ is the complementary error function,

$$\operatorname{erfc}(z) = \frac{2}{\sqrt{\pi}} \int_z^\infty e^{-t^2} dt.$$

In the original variables, the temperature reads

$$\begin{aligned} T(z, t) = & T_{\text{mn}} + \frac{C_1 - C_2 T_{\text{mn}}}{C_2} \left[\operatorname{erfc}\left(\frac{z}{2\sqrt{D_s t}}\right) \right. \\ & \left. - \exp\left(\frac{C_2}{k_s} \left(\frac{C_2 D_s t}{k_s} + z\right)\right) \operatorname{erfc}\left(\frac{z}{2\sqrt{D_s t}} + \frac{C_2 \sqrt{D_s t}}{k_s}\right) \right]. \end{aligned} \quad (\text{A.10})$$

713 *Appendix A.3. Total energy absorbed*

714 We can rewrite the energy in terms of the non-dimensional variables,

$$E(t) = \rho_s c_s \mathcal{L}\mathcal{T} \int_0^\infty \hat{T}(\hat{z}, t/t_f) d\hat{z}, \quad (\text{A.11})$$

which requires the results

$$\begin{aligned} \int_0^\infty \operatorname{erfc}\left(\frac{z}{a}\right) dz = & a \left[z \cdot \operatorname{erfc}\left(\frac{z}{a}\right) - \frac{1}{\sqrt{\pi}} \exp\left(-\frac{z^2}{a^2}\right) \right]_0^\infty \\ = & \frac{a}{\sqrt{\pi}}, \end{aligned} \quad (\text{A.12})$$

$$\begin{aligned} \int_0^\infty \exp(bz) \operatorname{erfc}\left(\frac{ab}{2} + \frac{z}{a}\right) dz = & \frac{1}{b} \left[\exp(bz) \operatorname{erfc}\left(\frac{ab}{2} + \frac{z}{a}\right) \right. \\ & \left. + \exp\left(-\frac{a^2}{4b^2}\right) \operatorname{erf}\left(\frac{z}{a}\right) \right]_0^\infty \\ = & \frac{1}{b} \left[\exp\left(-\frac{a^2}{4b^2}\right) - \operatorname{erfc}\left(\frac{ab}{2}\right) \right], \end{aligned} \quad (\text{A.13})$$

where $a = 2\sqrt{D_s t}$ and $b = C_2/k_s$. In the original dimensional variables, the absorbed energy is

$$E(t) = \frac{k_s^2}{D_s C_2^2} (C_1 - C_2 T_{\text{mn}}) \left[\exp\left(\frac{C_2^2 D_s t}{k_s^2}\right) \operatorname{erfc}\left(\frac{C_2}{k_s} \sqrt{D_s t}\right) + \frac{2}{\sqrt{\pi}} \frac{C_2}{k_s} \sqrt{D_s t} - 1 \right]. \quad (\text{A.14})$$

715 Appendix B. Derivation of the temperature for a finite substrate 716 with varying Q_{net} and T_a

717 In the case where the environmental conditions are not assumed con-
718 stant, the problem is impossible to solve analytically and numerical and
719 approximate solutions are required. Whereas numerical solutions are useful
720 to rapidly visualise the solution to specific problems, explicit solutions allow
721 a better understanding of the role of the different parameters of the problem.
722 Here we derive a solution based on a generalised eigenfunction expansion,
723 which is always a valid method when the equations are linear.

In a soil with a finite length L and with varying environmental conditions, the governing equations are

$$\frac{\partial T}{\partial t} = D_s \frac{\partial^2 T}{\partial z^2}, \quad T(z, 0) = T_{\text{mn}}, \quad (\text{B.1a})$$

$$-k_s \frac{\partial T}{\partial z} \Big|_{z=0} = C_1 + G(t) - C_2 T(0, t), \quad \frac{\partial T}{\partial z} \Big|_{z=L} = 0, \quad (\text{B.1b})$$

724 where $C_1 = (1 - F_{VC})(1 - A_s)Q_{\text{av}} + H_{as}T_{\text{av}}$, $C_2 = H_{as} + \rho_w L_e \alpha$ and $G(t)$ is
725 defined by

$$G(t) = (1 - F_{VC})(1 - A_s)Q_{\text{net}}(t) + H_{as}T_a(t). \quad (\text{B.2})$$

726 Similarly to the case of infinite depth solved in Appendix A, we will first
727 reformulate the problem in terms of scaled variables. Additionally to the
728 non-dimensional numbers arising in the case of the infinitely deep soil, in this
729 case we have an additional parameter relative to the varying environmental
730 conditions.

731 Appendix B.1. Non-dimensional formulation

The thickness of the soil L provides the length scale to this problem, which has an associated diffusion time scale L^2/D_s . We introduce the rescaled

variables defined by $z = L\bar{z}$, $t = L^2\bar{t}/D_s$ and $T = T_{\text{mn}} + \mathcal{T}\bar{T}$, so that heat equation and initial condition become

$$\frac{\partial \bar{T}}{\partial \bar{t}} = \frac{\partial^2 \bar{T}}{\partial \bar{z}^2}, \quad (\text{B.3})$$

$$\bar{T}(\bar{z}, 0) = 0. \quad (\text{B.4})$$

The value of \mathcal{T} will be determined by the boundary condition at the soil surface. Firstly, let us define the variable quantities Q_{net} and T_{a} in terms of the relative changes with respect to the average values rather than absolute changes, i.e., we write

$$\begin{aligned} Q_{\text{net}}(t) &= Q_{\text{av}} (1 + \bar{f}(t)), \\ T_{\text{a}}(t) &= T_{\text{av}} (1 + \bar{g}(t)). \end{aligned}$$

732 Using the same approach as for obtaining $f(t)$ and $g(t)$ in section 2, obtaining

$$\bar{f}(t) = \frac{\pi}{2} \sin\left(\frac{2\pi t}{p_1}\right) - 1, \quad (\text{B.5})$$

$$\bar{g}(t) = \frac{T_{\text{mn}}}{T_{\text{av}}} - \frac{\Delta T}{2T_{\text{av}}} \cos\left(\frac{2\pi(t - t_{\text{min}})}{p_2}\right) - 1, \quad (\text{B.6})$$

733 The definition of G in (B.2) can be rearranged to be expressed in terms of a
734 non-dimensional function \bar{G}

$$\begin{aligned} G(t) &= (1 - F_{VC})(1 - A_s)Q_{\text{net}}(t) + H_{as}T_{\text{a}}(t) \\ &= C_1 + (1 - F_{VC})(1 - A_s)Q_{\text{av}}\bar{f}(t) + H_{as}T_{\text{av}}\bar{g}(t) \\ &= C_1 (1 + \beta_1\bar{f}(t) + (1 - \beta_1)\bar{g}(t)) \\ &= C_1 (1 + \bar{G}(\bar{t})), \end{aligned} \quad (\text{B.7})$$

735 where we have introduced the non-dimensional parameter

$$\beta_1 = \frac{(1 - F_{VC})(1 - A_s)Q_{\text{av}}}{C_1}. \quad (\text{B.8})$$

736 Upon using the non-dimensional quantities, the boundary condition at
737 the surface becomes

$$-\frac{k_s \mathcal{T}}{C_1 L} \frac{\partial \bar{T}}{\partial \bar{z}} \Big|_{\bar{z}=0} = 1 + \bar{G}(\bar{t}) - \frac{C_2}{C_1} (T_{\text{mn}} + \mathcal{T}\bar{T}(0, \bar{t})). \quad (\text{B.9})$$

738 Thus, if one takes $\mathcal{T} = LC_1/k_s$, then

$$-\frac{\partial \bar{T}}{\partial \bar{z}} \Big|_{\bar{z}=0} = \beta_2 + \bar{G}(\bar{t}) - \beta_3 \bar{T}(0, \bar{t}). \quad (\text{B.10})$$

739 where we have introduced the dimensionless parameters

$$\beta_2 = 1 - \frac{C_2 T_{mn}}{C_1}, \quad \beta_3 = \frac{C_2 \mathcal{T}}{C_1}, \quad (\text{B.11})$$

740 As for the boundary condition at the end of the soil, it simply reads

$$\left. \frac{\partial \bar{T}}{\partial \bar{z}} \right|_{\bar{z}=1} = 0. \quad (\text{B.12})$$

741 *Appendix B.2. Eigenfunction Expansion*

742 The problem to be solved is defined by the heat equation (B.3) subject to
743 the boundary conditions (B.10) and (B.12) and the initial condition (B.4).

744 Upon writing

$$u(x, \bar{t}) = \bar{T}(1-x, \bar{t}) - \frac{\beta_2 + \bar{G}(\bar{t})}{\beta_3}, \quad (\text{B.13})$$

one has the following system

$$\begin{aligned} \frac{\partial u}{\partial \bar{t}} + \frac{\bar{G}'}{\beta_3} &= \frac{\partial^2 u}{\partial x^2}, \\ \left. \frac{\partial u}{\partial x} \right|_{x=1} + \beta_3 u(1, \bar{t}) &= \left. \frac{\partial u}{\partial x} \right|_{x=0} = 0, \\ u(x, 0) &= -\frac{\beta_2 + \bar{G}(0)}{\beta_3}, \end{aligned} \quad (\text{B.14})$$

whose solution can be written in terms of the following generalised Fourier series:

$$u(x, \bar{t}) = \sum_{n=1}^{\infty} \varphi_n(\bar{t}) \cos(\lambda_n x), \quad (\text{B.15})$$

where λ_n are the positive solutions of

$$\tan(\lambda_n) = \frac{\beta_3}{\lambda_n}. \quad (\text{B.16})$$

745 In what follows we shall use that 1 may be expressed as this same generalised
746 Fourier series like:

$$1 = \sum_{n=1}^{\infty} b_n \cos(\lambda_n x), \quad (\text{B.17})$$

747 where

$$b_n = \frac{2 \sin(\lambda_n)}{\lambda_n + \sin(\lambda_n) \cos(\lambda_n)} = \frac{2\beta_3 \cos(\lambda_n)}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)} \quad (\text{B.18})$$

where we have used (B.16). Therefore, the initial condition is now

$$u(x, 0) = -\frac{\beta_2 + \bar{G}(0)}{\beta_3} \sum_{n=1}^{\infty} b_n \cos(\lambda_n x), \quad (\text{B.19})$$

and so $\varphi_n(\bar{t})$ are the solutions of

$$\begin{aligned} \varphi_n' + \frac{\bar{G}'(\bar{t})}{\beta_3} b_n &= -\lambda_n^2 \varphi, \\ \varphi(0) &= -\frac{\beta_2 + \bar{G}(0)}{\beta_3} b_n. \end{aligned} \quad (\text{B.20})$$

Therefore,

$$\begin{aligned} \varphi_n(\bar{t}) &= \frac{2 \cos(\lambda_n)}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)} \left[\lambda_n^2 \int_0^{\bar{t}} \exp(-\lambda_n^2(\bar{t} - \tau)) \bar{G}(\tau) d\tau \right. \\ &\quad \left. - \bar{G}(\bar{t}) - \beta_2 \exp(-\lambda_n^2 \bar{t}) \right]. \end{aligned} \quad (\text{B.21})$$

Finally, due to the expression of \bar{G} , we can write the integral as

$$\begin{aligned} \int_0^{\bar{t}} \exp(-\lambda_n^2(\bar{t} - \tau)) \bar{G}(\tau) d\tau &= \beta_1 \int_0^{\bar{t}} \exp(-\lambda_n^2(\bar{t} - \tau)) \bar{f}(\tau) d\tau \\ &\quad + (1 - \beta_1) \int_0^{\bar{t}} \exp(-\lambda_n^2(\bar{t} - \tau)) \bar{g}(\tau) d\tau \\ &= \beta_1 \bar{I}_{1,n}(\bar{t}) + (1 - \beta_1) \bar{I}_{2,n}(\bar{t}), \end{aligned} \quad (\text{B.22})$$

where $\bar{I}_{1,n}$, $\bar{I}_{2,n}$ are easily obtained integrating by parts:

$$\begin{aligned} \bar{I}_{1,n}(\bar{t}) &= \frac{\bar{p}_1 \pi}{8\pi^2 + 2\bar{p}_1^2 \lambda_n^4} \left[\bar{p}_1 \lambda_n^2 \sin\left(\frac{2\pi \bar{t}}{\bar{p}_1}\right) - 2\pi \cos\left(\frac{2\pi \bar{t}}{\bar{p}_1}\right) \right. \\ &\quad \left. + 2\pi \exp(-\lambda_n^2 \bar{t}) \right] - \frac{1}{\lambda_n^2} (1 - \exp(-\lambda_n^2 \bar{t})), \end{aligned} \quad (\text{B.23})$$

$$\begin{aligned} \bar{I}_{2,n}(\bar{t}) &= \frac{1}{\lambda_n^2} \left(\frac{T_{\text{mn}}}{T_{\text{av}}} - 1 \right) (1 - \exp(-\lambda_n^2 \bar{t})) + \frac{\Delta T \bar{p}_2^2 \lambda_n^2}{2T_{\text{av}} (4\pi^2 + \bar{p}_2^2 \lambda_n^4)} \\ &\quad \times \left[\exp(-\lambda_n^2 \bar{t}) \cos\left(\frac{2\pi \bar{t}_{\text{min}}}{\bar{p}_2}\right) - \cos\left(\frac{2\pi(\bar{t} - \bar{t}_{\text{min}})}{\bar{p}_2}\right) \right] \\ &\quad - \frac{\Delta T \bar{p}_2 \pi}{T_{\text{av}} (4\pi^2 + \bar{p}_2^2 \lambda_n^4)} \left[\exp(-\lambda_n^2 \bar{t}) \sin\left(\frac{2\pi \bar{t}_{\text{min}}}{\bar{p}_2}\right) \right. \\ &\quad \left. + \sin\left(\frac{2\pi(\bar{t} - \bar{t}_{\text{min}})}{\bar{p}_2}\right) \right]. \end{aligned} \quad (\text{B.24})$$

Then, the solution for the non-dimensional temperature in terms of (\bar{z}, \bar{t}) is given by

$$\begin{aligned} \bar{T}(\bar{z}, \bar{t}) = \sum_{n=1}^{\infty} \frac{2 \cos(\lambda_n) \cos(\lambda_n(1 - \bar{z}))}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)} & \left[\beta_2 (1 - \exp(-\lambda_n^2 \bar{t})) \right. \\ & \left. + \lambda_n^2 (\beta_1 \bar{I}_{1,n}(\bar{t}) + (1 - \beta_1) \bar{I}_{2,n}(\bar{t})) \right], \end{aligned} \quad (\text{B.25})$$

with λ_n as the set of positive solutions of

$$\tan(\lambda_n) = \frac{\beta_3}{\lambda_n}. \quad (\text{B.26})$$

Appendix B.2.1. Non perfect insulation

In this section we consider the case where the thermal insulation of the roof is not perfect and therefore some heat loss takes places at the lower boundary. We will write this as

$$-k_s \frac{\partial T}{\partial z} \Big|_{z=L} = C_1 \epsilon > 0, \quad (\text{B.27})$$

which relates the heat loss in the bottom of the soil to the heat gain occurring at its surface. In the non-dimensional formulation, this boundary condition corresponds to

$$-\frac{\partial \bar{T}}{\partial \bar{z}} \Big|_{\bar{z}=1} = \epsilon. \quad (\text{B.28})$$

In this case we start by writing

$$v(x, \bar{t}) = \bar{T}(1 - x, \bar{t}) - \frac{\beta_2 + \bar{G}(\bar{t})}{\beta_3} + \epsilon \left(\frac{\beta_3 + 1}{\beta_3} - x \right), \quad (\text{B.29})$$

which yields

$$\frac{\partial v}{\partial x} \Big|_{x=0} = -\frac{\partial \bar{T}}{\partial \bar{z}} \Big|_{\bar{z}=1} - \epsilon = 0, \quad (\text{B.30})$$

and

$$\begin{aligned} \frac{\partial v}{\partial x} \Big|_{x=1} &= -\frac{\partial \bar{T}}{\partial \bar{z}} \Big|_{\bar{z}=0} - \epsilon \\ &= \beta_2 + \bar{G}(\bar{t}) - \beta_3 \bar{T}(0, \bar{t}) - \epsilon \\ &= \beta_2 + \bar{G}(\bar{t}) - \beta_3 \left[v(1, \bar{t}) + \frac{\beta_2 + \bar{G}(\bar{t})}{\beta_3} - \frac{\epsilon}{\beta_3} \right] - \epsilon \\ &= -\beta_3 v(1, \bar{t}), \end{aligned} \quad (\text{B.31})$$

and therefore the original system can now be written as

$$\begin{aligned} \frac{\partial v}{\partial \bar{t}} + \frac{\bar{G}'}{\beta_3} &= \frac{\partial^2 v}{\partial x^2}, \\ \frac{\partial v}{\partial x} \Big|_{x=1} + \beta_3 v(1, \bar{t}) &= \frac{\partial v}{\partial x} \Big|_{x=0} = 0, \\ v(x, 0) &= -\frac{\beta_2 + \bar{G}(0)}{\beta_3} + \epsilon \left(\frac{\beta_3 + 1}{\beta_3} - x \right). \end{aligned} \quad (\text{B.32})$$

Similarly as before, we obtain

$$v(x, \bar{t}) = \sum_{n=1}^{\infty} \phi_n(\bar{t}) \cos(\lambda_n x), \quad (\text{B.33})$$

758 where λ_n are again the positive solutions of equation (B.16).

The procedure is now the same as before except for the fact that now

$$v(x, 0) = -\frac{\beta_2 + \bar{G}(0)}{\beta_3} \sum_{n=1}^{\infty} b_n \cos(\lambda_n x) + \epsilon \sum_{n=1}^{\infty} c_n \cos(\lambda_n x), \quad (\text{B.34})$$

759 where b_n is provided in (B.18) and

$$c_n = \frac{2}{\lambda_n^2 + \beta_3 \cos^2(\lambda_n)}. \quad (\text{B.35})$$

Therefore, $\phi_n(\bar{t})$ are the solutions of

$$\begin{aligned} \phi_n' + \frac{\bar{G}'(\bar{t})}{\beta_3} b_n &= -\lambda_n^2 \phi_n, \\ \phi_n(0) &= -\frac{\beta_2 + \bar{G}(0)}{\beta_3} b_n + \epsilon c_n. \end{aligned} \quad (\text{B.36})$$

Therefore,

$$\begin{aligned} \phi_n(\bar{t}) &= \phi_n(0) \exp(-\lambda^2 \bar{t}) - \frac{b_n}{\beta_3} \int_0^{\bar{t}} \exp(-\lambda^2(\bar{t} - \tau)) \bar{G}'(\tau) d\tau \\ &= -\left(\frac{\beta_2}{\beta_3} b_n - \epsilon c_n \right) \exp(-\lambda^2 \bar{t}) - \frac{b_n}{\beta_3} \left[\bar{G}(\bar{t}) \right. \\ &\quad \left. + \lambda_n^2 \int_0^{\bar{t}} \exp(-\lambda^2(\bar{t} - \tau)) \bar{G}(\tau) d\tau \right]. \end{aligned} \quad (\text{B.37})$$

Then, the solution for the non-dimensional temperature in terms of (\bar{z}, \bar{t}) when the roof transfers energy to the house is given by

$$\begin{aligned}
\bar{T}(\bar{z}, \bar{t}) &= \sum_{n=1}^{\infty} \frac{b_n}{\beta_3} \left[\left(\beta_2 - \epsilon \frac{\beta_3 c_n}{b_n} \right) (1 - \exp(-\lambda_n^2 \bar{t})) \right. \\
&\quad \left. + \lambda_n^2 (\beta_1 \bar{I}_{1,n}(\bar{t}) + (1 - \beta_1) \bar{I}_{2,n}(\bar{t})) \right] \cos(\lambda_n(1 - \bar{z})) \\
&= \sum_{n=1}^{\infty} \frac{b_n}{\beta_3} \left[\left(\beta_2 - \frac{\epsilon}{\cos(\lambda_n)} \right) (1 - \exp(-\lambda_n^2 \bar{t})) \right. \\
&\quad \left. + \lambda_n^2 (\beta_1 \bar{I}_{1,n}(\bar{t}) + (1 - \beta_1) \bar{I}_{2,n}(\bar{t})) \right] \cos(\lambda_n(1 - \bar{z})) ,
\end{aligned} \tag{B.38}$$

760 with λ_n is given in (B.26) and $I_{1,n}$, $I_{2,n}$ are provided in (B.23) and (B.24).
761 From equation (B.26) we also find $\cos(\lambda_n) \sim (-1)^n$ for large values of n . In
762 particular, note how the solution (B.38) reduces to (B.25) when we set $\epsilon = 0$.

763 *Appendix B.3. Total energy absorbed*

To compute the expression for the energy absorbed we shall use expressions (23):

$$\begin{aligned}
E(t) &= \int_0^L \rho_s c_s (T(z, t) - T_{\text{mn}}) dz = \frac{L^2 C_1 \rho_s c_s}{k_s} \int_0^1 \bar{T} \left(\bar{z}, \frac{L^2 t}{D_s} \right) d\bar{z} \\
&= \frac{L^2 C_1 \rho_s c_s}{k_s} \sum_{n=1}^{\infty} \left\{ \frac{2 \cos(\lambda_n) \sin(\lambda_n)}{\lambda_n^3 + \beta_3 \lambda_n \cos^2(\lambda_n)} \left[\beta_2 \left(1 - \exp \left(-\lambda_n^2 \frac{L^2 t}{D_s} \right) \right) \right. \right. \\
&\quad \left. \left. + \lambda_n^2 (\beta_1 I_{1,n}(t) + (1 - \beta_1) I_{2,n}(t)) \right] \right\} ,
\end{aligned}$$

764 where $I_{i,n}(t) = \bar{I}_{i,n}(L^2 t / D_s)$. We note that

$$\frac{\cos(\lambda_n) \sin(\lambda_n)}{\lambda_n^3 + \beta_3 \lambda_n \cos^2(\lambda_n)} = \frac{\beta_3}{\lambda_n^4 + \lambda_n^2 \beta_3 (1 + \beta_3)} , \tag{B.39}$$

where we recall that $\beta_3 = C_2 L / k_s$. To compute the energy absorbed at the end of the day we must simply evaluate this expression at t_f . We note that the above expression converges very fast. Indeed, one can even use the expression obtained by just retaining the first term:

$$\begin{aligned}
E(t_f) &\sim \frac{2L^2 C_1 \rho_s c_s \beta_3}{k_s (\lambda_1^4 + \lambda_1^2 \beta_3 (1 + \beta_3))} \left[\beta_2 \left(1 - \exp \left(-\lambda_1^2 \frac{L^2 t_f}{D_s} \right) \right) \right. \\
&\quad \left. + \lambda_1^2 \left(\beta_1 I_{1,1} \left(\frac{L^2 t_f}{D_s} \right) + (1 - \beta_1) I_{2,1} \left(\frac{L^2 t_f}{D_s} \right) \right) \right] ,
\end{aligned} \tag{B.40}$$

765 where $\lambda_1 \in (0, \frac{\pi}{2})$ is the unique solution of (B.26) for $n = 1$.

766 *Appendix B.3.1. Non perfect insulation*

In the case where heat is transferred to the building, we combine (23) and (B.38) and obtain

$$E(t) = \frac{L^2 C_1 \rho_s c_s}{k_s} \sum_{n=1}^{\infty} \frac{2\beta_3}{\lambda_n^4 + \lambda_n^2 \beta_3 (1 + \beta_3)} \left[\left(\beta_2 - \frac{\epsilon}{\cos(\lambda_n)} \right) \left(1 - \exp \left(-\lambda_n^2 \frac{L^2 t}{D_s} \right) \right) + \lambda_n^2 \left(\beta_1 I_{1,n} \left(\frac{L^2 t}{D_s} \right) + (1 - \beta_1) I_{2,n} \left(\frac{L^2 t}{D_s} \right) \right) \right].$$

Therefore, the corresponding approximate expression for the total energy stored at the end of the day reads:

$$E(t_f) \sim \frac{L^2 C_1 \rho_s c_s}{k_s} \frac{2\beta_3}{\lambda_1^4 + \lambda_1^2 \beta_3 (1 + \beta_3)} \left[\left(\beta_2 - \frac{\epsilon}{\cos(\lambda_1)} \right) \left(1 - \exp \left(-\lambda_1^2 \frac{L^2 t_f}{D_s} \right) \right) + \lambda_1^2 \left(\beta_1 I_{1,1} \left(\frac{L^2 t_f}{D_s} \right) + (1 - \beta_1) I_{2,1} \left(\frac{L^2 t_f}{D_s} \right) \right) \right].$$

767 *Appendix B.4. Approximate expression for λ_1*

768 We now focus on (B.26) and we shall derive an approximate explicit
769 expression for the first eigenvalue, λ_1 . Since the first eigenvalue has a value
770 lower than $\pi/2$, we define $z = 2/\pi\lambda_1$ and so z is expected to be small.
771 Therefore, Taylor expanding the tangent one reaches

$$\frac{\pi}{2} z \left(\frac{\pi}{2} z + \frac{\pi^3}{24} z^3 + \frac{\pi^5}{240} z^5 + \mathcal{O}(z^7) \right) = \beta_3, \quad (\text{B.41})$$

772 whose only positive real solution is

$$\lambda_1 \approx \left(\frac{5^{1/3}}{6} y - \frac{5}{6} - \frac{13 \cdot 5^{2/3}}{6y} \right)^{1/2}, \quad (\text{B.42})$$

773 where

$$y = \left(110 + 162\beta_3 + 9\sqrt{285 + 440\beta_3 + 324\beta_3^2} \right)^{1/3}. \quad (\text{B.43})$$

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