$\frac{J / A+A / 568 / A 126}{(\text { Brescia+, 2014) }} \quad$ SDSS-DR9 photometric redshifts

A catalogue of photometric redshifts for the SDSS-DR9 galaxies.

Brescia M., Cavuoti S., Longo G., De Stefano V. <Astron. Astrophys. 568, A126 (2014)> $=2014 \mathrm{~A} \& \mathrm{~A} . \mathrm{F}$ 58A.126B

ADC_Keywords: Galaxy catalogs ; Galaxies, photometry ; Photometry, SDSS ;
Redshifts

Keywords: techniques: photometric - galaxies: distances and redshifts -

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    galaxies: photometry - methods: data analysis
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- catalogs

Abstract:
Accurate photometric redshifts for large samples of galaxies are among
the main products of modern multiband digital surveys. Over the last
decade, the Sloan Digital Sky Survey (SDSS) has become a sort of
benchmark against which to test the various methods. We present an
application of a new method to the estimation of photometric redshifts
for the galaxies in the SDSS Data Release 9 (SDSSDR9). Photometric
redshifts for more than 143 million galaxies were produced. The MLPQNA
(Multi Layer Perceptron with Quasi Newton
Algorithm) model provided
within the framework of the DAMEWARE (DAta Mining and Exploration Web

Application REsource) is an interpolative method derived from machine
learning models. The obtained redshifts have an overall uncertainty of
sigma=0.023 with a very small average bias of about $3 \times 10^{\wedge}-5$, and a
fraction of catastrophic outliers of about 5\%. This result is slightly
better than what was already available in the literature, particularly
in terms of the smaller fraction of catastrophic outliers.

Description:
We present an application of a machine learning method to the
estimation of photometric redshifts for the galaxies in the SDSS Data

Release 9 (SDSS-DR9). Photometric redshifts for more than 143 million
galaxies were produced. The MLPQNA (Multi Layer Perceptron with Quasi

Newton Algorithm) model provided within the framework of the DAMEWARE
(DAta Mining and Exploration Web Application REsource) is an
interpolative method derived from machine learning models. The
obtained redshifts have an overall uncertainty of $\sigma=0.023$ with a
very small average bias of about $3 \times 10^{-5}$ and a fraction of
catastrophic outliers of about 5\%. After removal of the catastrophic
outliers, the uncertainty is about $\sigma=0.017$. The catalogue files
report in their name the range of DEC degrees related to the included
objects.
File Summary:
$\qquad$
$\qquad$
FileName Lrecl Records Explanations

| ReadMe | 80 |  | This file |
| :---: | :---: | :---: | :---: |
| r00_01.dat | 186 | 3174048 | Photo-z for objects with |
| DEC in [+00, | +01 [ | deg |  |
| r01_02.dat | 186 | 2622557 | Photo-z for objects with |
| DEC in [+01, | +02 [ | deg |  |
| r02_03.dat | 186 | 2425949 | Photo-z for objects with |




```
r65_70.dat 186 1677346 Photo-z for objects with
DEC in [+65, +70[ deg
r70_85.dat 186 1337222 Photo-z for objects with
DEC in [+70, +85[ deg
r-01_00.dat 186 3182508 Photo-z for objects with
DEC in [-01, +00[ deg
r-02_01.dat 186 2471559 Photo-z for objects with
DEC ín [-02, -01[ deg
r-04-02.dat 186 3808610 Photo-z for objects with
DEC in [-04, -02[ deg
r-06_-04.dat 186 2637920 Photo-z for objects with
DEC in [-06, -04[ deg
r-08_06.dat 186 2661702 Photo-z for objects with
DEC in [-08, -06[ deg
r-10_08.dat 186 2006992 Photo-z for objects with
DEC in [-10, -08[ deg
r-15_-10.dat 186 2175039 Photo-z for objects with
DEC in [-15, -10[ deg
r-20_-15.dat 186 1693609 Photo-z for objects with
DEC in [-20, -15[ deg
r-30_-20.dat 186 702996 Photo-z for objects with
DEC in [-30, -20[ deg
rspecial.dat 186 18006 *Photo-z for special SDSS
objects
fits/*
5 9 ~ T a b l e s ~ a s ~ f i t s ~ f i l e s
```

Note on rspecial.dat: photo-z for $\operatorname{SDSS}$ special objects with a mismatch between
photometric and spectroscopic class assignment.

See also:
http://www.sdss.org : SDSS Home Page
Byte-by-byte Description of file: r*.dat


Bytes Format Units Label Explanations
$\qquad$
1- 19 I19 --- objID Unique SDSS
identifier

| $\begin{aligned} & 21-29 \\ & (\mathrm{~J} 2000) \end{aligned}$ | F9. 5 | deg | RAdeg | Right Ascension |
| :---: | :---: | :---: | :---: | :---: |
| 31-39 | F9.5 | deg | DEdeg | Declination (J2000) |
| 41-46 | F6. 3 | mag | umag | [0/40] u-band PSF |
| magnitude |  |  |  |  |
| magnitude |  |  |  |  |
| $55-60$ | F6. 3 | mag | rmag | [0/40] r-band PSF |
| $62-67$ | F6. 3 | mag | imag | [0/40] i-band PSF |
| $69-74$ <br> magnitude | F6.3 | mag | zmag | [0/40] z-band PSF |
| $76-81$ | F6.3 | mag | e_umag | ? $=99.999$ u-band PSF |
| $\begin{gathered} \text { magnitude } \\ 83-88 \end{gathered}$ | $\begin{aligned} & \text { error } \\ & \text { F6. } 3 \end{aligned}$ | mag | e gmag | ? $=99.999$ g-band PSF |
| magnitude | error |  |  |  |
| 90-95 | F6. 3 | mag | e_rmag | ? $=99.999$ r-band PSF |
| magnitude | error |  |  |  |
| 97-102 | F6.3 | mag | e_imag | ? $=99.999$ i-band PSF |
| magnitude | error |  |  |  |
| 104-109 | F6.3 | mag | e_zmag | ? $=99.999$ z-band PSF |
| magnitude | error |  |  |  |
| 111-116 | F6. 3 | mag | extu | ? $=99.999$ Extinction |
| in u-band |  |  |  |  |
| 118-123 | F6. 3 | mag | extg | ? $=99.999$ Extinction |
| in $9-b$ and |  |  |  |  |
| 125-130 | F6. 3 | mag | extr | ? $=99.999$ Extinction |
| in r-band |  |  |  |  |
| 132-137 | F6. 3 | mag | exti | ? $=99.999$ Extinction |
| in i-band |  |  |  |  |
| 139-144 | F6. 3 | mag | extz | ? $=99.999$ Extinction |
| in z -band |  |  |  |  |
| 146-152 | F7. 3 | mag | u-g | u-g color index from |
| PSF SDSS magnitudes |  |  |  |  |
| 154-160 | F7. 3 | mag | $g-r$ | g-r color index from |
| PSF SDSS magnitudes |  |  |  |  |
| 162-168 | F7. 3 | mag | r-i | r-i color index from |
| PSF SDSS magnitudes |  |  |  |  |
| 170-176 | F7. 3 | mag | i-z | i-z color index from |
| PSF SDSS magnitudes |  |  |  |  |
| 178-184 | F7. 5 | --- | zphot | Estimated photometric |
| redshift |  |  |  |  |
| 186 | I1 |  | q_zphot | [0/3] Quality Flag of |
| the zphot: | : $0=n$ n |  |  |  |

$$
1=\mathrm{high},
$$

2=medium, 3=1ow

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