# Virial coefficients of trapped and untrapped three-component fermions with three-body forces in arbitrary spatial dimensions 

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#### Abstract

Using a coarse temporal lattice approximation, we calculate the first few terms of the virial expansion of a three-species fermion system with a three-body contact interaction in $d$ spatial dimensions, both in homogeneous space and in a harmonic trapping potential of frequency $\omega$. Using the three-body problem to renormalize, we report analytic results for the change in the fourth- and fifth-order virial coefficients $\Delta b_{4}$ and $\Delta b_{5}$ as functions of $\Delta b_{3}$. Additionally, we argue that in the $\omega \rightarrow 0$ limit the relationship $b_{n}^{\mathrm{T}}=n^{-d / 2} b_{n}$ holds between the trapped (T) and the homogeneous coefficients for arbitrary temperature and coupling strength (not merely in scale-invariant regimes). Finally, we point out an exact, universal (coupling- and frequency-independent) relationship between $\Delta b_{3}^{\mathrm{T}}$ in one dimension with three-body forces and $\Delta b_{2}^{\mathrm{T}}$ in two dimensions with two-body forces.


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## I. INTRODUCTION

Motivated by the recent interest in one-dimensional (1D) Fermi and Bose gases in the fine-tuned situation where only three-body interactions are present [1-9], we explore here the thermodynamics of fermions with a contact three-body interaction in the region of low fugacity (which corresponds to a dilute regime and therefore high temperatures in units of the energy scale set by the density). We focus on the fermionic case but explore the problem in arbitrary dimension $d$. To that end, we implement a semiclassical lattice approximation (SCLA) to calculate the virial coefficients $b_{n}$ and carry out their evaluation up to $n=5$ at leading order (LO) in that approximation.

The LO-SCLA was introduced in Ref. [10] as a way to estimate virial coefficients in two-component Fermi gases. The approximation seems crude in its definition but performs surprisingly well when the lowest nontrivial order in the virial expansion is used as a renormalized coupling constant ( $b_{2}$ for two-body forces, for example, and $b_{3}$ in this work). Not surprisingly, the approximation was seen to work better at weak coupling, which makes sense, as the radius of convergence of the virial expansion was found to be quickly reduced as a result of the interaction. In Ref. [11] the next-to-leading-order SCLA was explored up to $b_{7}$, displaying the convergence properties up to the unitary point (in three dimensions), and in Ref. [12] the LO-SCLA was used for systems in a harmonic trap, showing that the approximation can capture the dependence on the trap frequency $\omega$. In both cases, the analytic dependence of virial coefficients on the dimension was obtained, as is the case here. This is to be contrasted with conventional methods to calculate virial coefficients, which can be very precise but are limited to specific situations

[^0](coupling strength, dimension, etc.) and are typically unable to provide analytic insight, as they are entirely numerical.

Our analytic formulas for the virial coefficients, although approximate, support and shed light on the relationship $b_{n}^{\mathrm{T}} \rightarrow$ $n^{-d / 2} b_{n}$ in the $\omega \rightarrow 0$ limit, where the superindex T indicates the harmonically trapped situation. This connection is well known to be valid in the noninteracting limit and in the socalled unitary limit of spin- $1 / 2$ fermions in three dimensions, both of which feature temperature-independent coefficients $b_{n}$. As we argue, that relationship is actually valid for all temperatures and coupling constants and holds for three-body interactions just as well as for two-body interactions. Finally, we point out an exact, coupling- and frequency-independent relationship between $\Delta b_{3}^{\mathrm{T}}$ in one dimension with three-body forces and $\Delta b_{2}^{\mathrm{T}}$ in two dimensions with two-body forces.

## II. HAMILTONIAN AND VIRIAL EXPANSION

We focus on a nonrelativistic Fermi system with a threebody contact interaction, such that the Hamiltonian for three flavors $1,2,3$ is $\hat{H}=\hat{T}+\hat{V}$, where

$$
\begin{equation*}
\hat{T}=\int d^{d} x \hat{\psi}_{s}^{\dagger}(\mathbf{x})\left(-\frac{\hbar^{2} \nabla^{2}}{2 m}\right) \hat{\psi}_{s}(\mathbf{x}) \tag{1}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{V}=-g_{d} \int d^{d} x \hat{n}_{1}(\mathbf{x}) \hat{n}_{2}(\mathbf{x}) \hat{n}_{3}(\mathbf{x}) \tag{2}
\end{equation*}
$$

where the field operators $\hat{\psi}_{s}$ and $\hat{\psi}_{s}^{\dagger}$ are fermionic fields for particles of types 1,2,3 (summed over $s$ above), and $\hat{n}_{s}(\mathbf{x})$ are the coordinate-space densities. In the remainder of this work, we take $\hbar=k_{\mathrm{B}}=m=1$. Besides the above, we also consider the case in which an external trapping potential term is added to the Hamiltonian, of the form

$$
\begin{equation*}
\hat{V}_{\mathrm{ext}}=\frac{1}{2} m \omega^{2} \int d^{d} x \mathbf{x}^{2}\left[\hat{n}_{1}(\mathbf{x})+\hat{n}_{2}(\mathbf{x})+\hat{n}_{3}(\mathbf{x})\right] . \tag{3}
\end{equation*}
$$

One way to characterize the thermodynamics of the above system is through the virial expansion [13], which is an expansion around the dilute limit $z \rightarrow 0$, where $z=e^{\beta \mu}$ is the fugacity, i.e., it is a low-fugacity expansion, where $\beta$ is the inverse temperature and $\mu$ the chemical potential coupled to the total particle number operator $\hat{N}$. The coefficients accompanying the powers of $z$ in the expansion of the grandcanonical potential $\Omega$ are the virial coefficients; specifically,

$$
\begin{equation*}
-\beta \Omega=\ln \mathcal{Z}=Q_{1} \sum_{n=1}^{\infty} b_{n} z^{n} \tag{4}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{Z}=\operatorname{Tr}\left[e^{-\beta(\hat{H}-\mu \hat{N})}\right]=\sum_{N=0}^{\infty} z^{N} Q_{N} \tag{5}
\end{equation*}
$$

is the grand-canonical partition function, $Q_{1}$ is the one-body partition function, $b_{1}=1$, and the higher-order coefficients require solving the corresponding few-body problems,

$$
\begin{gather*}
Q_{1} b_{2}=Q_{2}-\frac{Q_{1}^{2}}{2!}  \tag{6}\\
Q_{1} b_{3}=Q_{3}-b_{2} Q_{1}^{2}-\frac{Q_{1}^{3}}{3!}  \tag{7}\\
Q_{1} b_{4}=Q_{4}-\left(b_{3}+\frac{b_{2}^{2}}{2}\right) Q_{1}^{2}-b_{2} \frac{Q_{1}^{3}}{2!}-\frac{Q_{1}^{4}}{4!},  \tag{8}\\
Q_{1} b_{5}=Q_{5}-\left(b_{4}+b_{2} b_{3}\right) Q_{1}^{2}-\left(b_{2}^{2}+b_{3}\right) \frac{Q_{1}^{3}}{2} \\
-b_{2} \frac{Q_{1}^{4}}{3!}-\frac{Q_{1}^{5}}{5!} \tag{9}
\end{gather*}
$$

and so forth.
Since $Q_{1} \propto V$, the above expressions display precisely how the volume dependence cancels out in each $b_{n}$. In particular, the highest power of $Q_{1}$ will always involve single-particle (i.e., noninteracting) physics and will therefore cancel in the change due to interactions $\Delta b_{n}$, such that

$$
\begin{gather*}
Q_{1} \Delta b_{2}=\Delta Q_{2}  \tag{10}\\
Q_{1} \Delta b_{3}=\Delta Q_{3}-\Delta b_{2} Q_{1}^{2}  \tag{11}\\
Q_{1} \Delta b_{4}=\Delta Q_{4}-\Delta\left(b_{3}+\frac{b_{2}^{2}}{2}\right) Q_{1}^{2}-\frac{\Delta b_{2}}{2} Q_{1}^{3}  \tag{12}\\
Q_{1} \Delta b_{5}=\Delta Q_{5}-\Delta\left(b_{4}+b_{2} b_{3}\right) Q_{1}^{2} \\
-\frac{1}{2} \Delta\left(b_{2}^{2}+b_{3}\right) Q_{1}^{3}-\frac{\Delta b_{2}}{3!} Q_{1}^{4} \tag{13}
\end{gather*}
$$

and so on. Note that, when only three-body interactions are present, as in the case we consider here, there is no change in the two-body spectrum, i.e., $\Delta b_{2}=0$. Therefore, the above expressions simplify to

$$
\begin{gather*}
Q_{1} \Delta b_{3}=\Delta Q_{3}  \tag{14}\\
Q_{1} \Delta b_{4}=\Delta Q_{4}-\Delta b_{3} Q_{1}^{2}  \tag{15}\\
Q_{1} \Delta b_{5}=\Delta Q_{5}-\left(\Delta b_{4}+b_{2} \Delta b_{3}\right) Q_{1}^{2}-\frac{\Delta b_{3}}{2} Q_{1}^{3} \tag{16}
\end{gather*}
$$

In terms of the partition functions $Q_{M N L}$ of $M$ particles of type $1, N$ of type 2 , and $L$ of type 3 , we have

$$
\begin{gather*}
\Delta Q_{3}=\Delta Q_{111}  \tag{17}\\
\Delta Q_{4}=3 \Delta Q_{211}  \tag{18}\\
\Delta Q_{5}=3 \Delta Q_{311}+3 \Delta Q_{221} \tag{19}
\end{gather*}
$$

From the above equations we see that there is only a small number of nontrivial contributions to each virial coefficient. The main task is calculating each of these terms and for that purpose we use a coarse lattice (or semiclassical) approximation, as explained next.

## III. THE SEMICLASSICAL APPROXIMATION AT LEADING ORDER

To carry out our calculations of virial coefficients we introduce a Trotter-Suzuki factorization of the Boltzmann weight. At the lowest possible order, the Trotter-Suzuki factorization amounts to keeping only the leading term in the formula

$$
\begin{equation*}
e^{-\beta(\hat{T}+\hat{V})}=e^{-\beta \hat{T}} e^{-\beta \hat{V}} \times e^{-\frac{\beta^{2}}{2}[\hat{T}, \hat{V}]} \times \ldots, \tag{20}
\end{equation*}
$$

where higher orders involve exponentials of nested commutators of $\hat{T}$ with $\hat{V}$. Taking the leading order in this expansion is equivalent to setting $[\hat{T}, \hat{V}]=0$, which is why we refer to it as a semiclassical approximation. As Refs. [10-12] have shown, this seemingly crude approximation provides surprisingly good answers, especially at weak coupling, and is therefore useful toward examining the virial expansion in an analytic fashion. Below, we give two explicit examples of the application of our approximation to the calculation of virial coefficients.

## A. A simple example: $\boldsymbol{\Delta} \boldsymbol{b}_{\mathbf{3}}$

As the simplest example, we consider $Q_{111}$,

$$
\begin{align*}
Q_{111} & =\sum_{\mathbf{p}_{j}}\langle\mathbf{P}| e^{-\beta \hat{T}} e^{-\beta \hat{V}}|\mathbf{P}\rangle  \tag{21}\\
& =\sum_{\mathbf{p}_{j}} e^{-\beta\left(p_{1}^{2}+p_{2}^{2}+p_{3}^{2}\right) / 2 m}\langle\mathbf{P}| e^{-\beta \hat{V}}|\mathbf{P}\rangle, \tag{22}
\end{align*}
$$

where we have used a collective momentum index $\mathbf{P}=$ $\left(\mathbf{p}_{1}, \mathbf{p}_{2}, \mathbf{p}_{3}\right)$. Inserting a coordinate-space completeness relation to evaluate the potential energy factor, we obtain

$$
\begin{align*}
e^{-\beta \hat{V}}|\mathbf{X}\rangle & =\prod_{\mathbf{z}}\left(1+C \hat{n}_{1}(\mathbf{z}) \hat{n}_{2}(\mathbf{z}) \hat{n}_{3}(\mathbf{z})\right)|\mathbf{X}\rangle \\
& =|\mathbf{X}\rangle+C \sum_{\mathbf{z}} \delta\left(\mathbf{x}_{1}-\mathbf{z}\right) \delta\left(\mathbf{x}_{2}-\mathbf{z}\right) \delta\left(\mathbf{x}_{3}-\mathbf{z}\right)|\mathbf{X}\rangle \\
& =\left[1+C \delta\left(\mathbf{x}_{1}-\mathbf{x}_{3}\right) \delta\left(\mathbf{x}_{2}-\mathbf{x}_{3}\right)\right]|\mathbf{X}\rangle, \tag{23}
\end{align*}
$$

where $C=\left(e^{\beta g_{d}}-1\right) \ell^{2 d}, \ell$ is an ultraviolet regulator in the form of a spatial lattice spacing, and we have used the fermionic relation $\hat{n}_{s}^{2}=\hat{n}_{s}$. We have also introduced a collective index $\mathbf{X}=\left(\mathbf{x}_{1}, \mathbf{x}_{2}, \mathbf{x}_{3}\right)$. The $C$-independent term yields the noninteracting result, such that we may write

$$
\begin{align*}
\Delta Q_{111}= & C \sum_{\mathbf{p}_{j}, \mathbf{x}_{k}} e^{-\beta\left(p_{1}^{2}+p_{2}^{2}+p_{3}^{2}\right) / 2 m} \\
& \times \delta\left(\mathbf{x}_{1}-\mathbf{x}_{3}\right) \delta\left(\mathbf{x}_{2}-\mathbf{x}_{3}\right)|\langle\mathbf{X} \mid \mathbf{P}\rangle|^{2}, \tag{24}
\end{align*}
$$

which simplifies substantially when using a plane-wave basis since $|\langle\mathbf{X} \mid \mathbf{P}\rangle|^{2}=1 / V^{3}$, where $V$ is the $d$-dimensional volume of the system. We then find

$$
\begin{equation*}
\Delta Q_{111}=C \frac{Q_{100}^{3}}{V^{2}} \tag{25}
\end{equation*}
$$

where

$$
\begin{equation*}
Q_{100}=\sum_{\mathbf{p}_{1}} e^{-\beta p_{1}^{2} / 2 m} \tag{26}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
\Delta b_{3}=C \frac{Q_{100}^{3}}{V^{2} Q_{1}}=C \frac{Q_{1}^{2}}{27 V^{2}}=\frac{1}{3} \frac{C}{\lambda_{T}^{2 d}} \tag{27}
\end{equation*}
$$

where $Q_{1}=3 Q_{100}=3 V / \lambda_{T}^{d}, \lambda_{T}=\sqrt{2 \pi \beta}$ is the thermal wavelength, and $V$ is the system's spatial volume. This relationship between the bare coupling constant $C$ and the physical quantity $\Delta b_{3}$ provides a way to renormalize the problem. In other words, $\Delta b_{3}$ will play the role of the renormalized dimensionless coupling constant.

The general form of the change $\Delta Q_{M N L}$ in the partition function for $M$ type 1 particles, $N$ type 2 particles, and $L$ type 3 particles, with a contact interaction, is given by

$$
\begin{equation*}
\Delta Q_{M N L}=\sum_{\overline{\mathbf{P}}, \overline{\mathbf{X}}} e^{-\beta \overline{\mathbf{P}}^{2} / 2 m}|\langle\overline{\mathbf{X}} \mid \overline{\mathbf{P}}\rangle|^{2}\left(C f_{a}(\overline{\mathbf{X}})+C^{2} f_{b}(\overline{\mathbf{X}})+\cdots\right) \tag{28}
\end{equation*}
$$

where $\overline{\mathbf{P}}$ and $\overline{\mathbf{X}}$ represent all momenta and positions of the $M+N+L$ particles, and the functions $f_{a}, f_{b}, \ldots$, which encode the matrix element of $e^{-\beta \hat{V}}$, depend on the specific $M N L$ case being considered. The wave function $\langle\overline{\mathbf{X}} \mid \overline{\mathbf{P}}\rangle$ is a product of three Slater determinants which, using a planewave single-particle basis, leads to Gaussian integrals over the momenta $\overline{\mathbf{P}}$.

## B. Another example: $\boldsymbol{\Delta} \boldsymbol{b}_{\mathbf{4}}$ in a harmonic trap

In this section we consider the case in which the system is held in a harmonic trapping potential of frequency $\omega$. As the expressions for the virial coefficients in terms of the canonical partition functions carry over to this case, we simply add the superindex ' T ' to denote quantities in the trapped system. To calculate $\Delta b_{4}^{\mathrm{T}}$ we need $\Delta b_{3}^{\mathrm{T}}$ and $Q_{1}^{\mathrm{T}}$. The latter is of course trivial, as there is no interaction in that case (see Ref. [12]),

$$
\begin{align*}
Q_{1}^{\mathrm{T}} & =3 \sum_{\mathbf{n}} e^{-\beta E_{\mathbf{n}}}=3 e^{-\beta \omega d / 2}\left(\frac{1}{1-e^{-\beta \omega}}\right)^{d}  \tag{29}\\
& =3\left(\frac{1}{2 \sinh (\beta \omega / 2)}\right)^{d} \tag{30}
\end{align*}
$$

where $E_{\mathbf{n}}$ is the single-particle energy level of the harmonic oscillator (separated in $d$-dimensional Cartesian coordinates such that $\mathbf{n}$ represents a $d$-dimensional vector of harmonic oscillator quantum numbers).

To obtain $\Delta b_{3}^{\mathrm{T}}$, we proceed as in the previous example to obtain the analog of Eq. (25) for the trapped case:

$$
\begin{align*}
\Delta Q_{111}^{\mathrm{T}}= & C \sum_{\mathbf{n}_{j}, \mathbf{x}_{k}} e^{-\beta\left(E_{\mathbf{n}_{1}}+E_{\mathbf{n}_{2}}+E_{\mathbf{n}_{3}}\right)} \\
& \times \delta\left(\mathbf{x}_{1}-\mathbf{x}_{3}\right) \delta\left(\mathbf{x}_{2}-\mathbf{x}_{3}\right)\left|\left\langle\mathbf{x}_{1} \mathbf{x}_{2} \mathbf{x}_{3} \mid \mathbf{n}_{1} \mathbf{n}_{2} \mathbf{n}_{3}\right\rangle\right|^{2} \tag{31}
\end{align*}
$$

The sums over $\mathbf{x}_{3}$ and $\mathbf{x}_{2}$ can be carried out right away, and moreover,

$$
\begin{equation*}
\left|\left\langle\mathbf{x}_{1} \mathbf{x}_{2} \mathbf{x}_{3} \mid \mathbf{n}_{1} \mathbf{n}_{2} \mathbf{n}_{3}\right\rangle\right|^{2}=\left|\phi_{\mathbf{n}_{1}}\left(\mathbf{x}_{1}\right)\right|^{2}\left|\phi_{\mathbf{n}_{2}}\left(\mathbf{x}_{2}\right)\right|^{2}\left|\phi_{\mathbf{n}_{3}}\left(\mathbf{x}_{3}\right)\right|^{2} \tag{32}
\end{equation*}
$$

where $\phi_{\mathbf{n}}(\mathbf{x})$ is the single-particle harmonic oscillator wave function in $d$-dimensional Cartesian coordinates. Using the above, we obtain

$$
\begin{equation*}
\Delta Q_{111}^{\mathrm{T}}=C \sum_{\mathbf{x}} \rho^{3}(\mathbf{x} ; \beta \omega) \tag{33}
\end{equation*}
$$

where

$$
\begin{equation*}
\rho(\mathbf{x} ; \beta \omega)=\sum_{\mathbf{n}} e^{-\beta E_{\mathbf{n}}}\left|\phi_{\mathbf{n}}(\mathbf{x})\right|^{2} . \tag{34}
\end{equation*}
$$

Note that $\sum_{\mathbf{x}} \rho(\mathbf{x} ; \beta \omega)=Q_{1}^{\mathrm{T}} / 3$.
Using the Mehler kernel (see Ref. [12]) evaluated at equal spatial arguments, we find that

$$
\begin{equation*}
\rho(\mathbf{x} ; \beta \omega)=\omega^{\frac{d}{2}} \frac{e^{-\omega \tanh (\beta \omega / 2) \mathbf{x}^{2}}}{(2 \pi \sinh (\beta \omega))^{\frac{d}{2}}}, \tag{35}
\end{equation*}
$$

where we note that $\tanh (\beta \omega / 2)>0$ for all $\beta \omega>0$. Carrying out the resulting Gaussian integrals and simplifying,

$$
\begin{equation*}
\Delta b_{3}^{\mathrm{T}}=\frac{\Delta Q_{111}^{T}}{Q_{1}^{\mathrm{T}}}=\frac{1}{3^{\frac{d}{2}+1}}\left(\frac{\beta \omega}{\sinh (\beta \omega)}\right)^{d} \frac{C}{\lambda_{T}^{2 d}} \tag{36}
\end{equation*}
$$

where $\lambda_{T}=\sqrt{2 \pi \beta}$.
Note that, as $\beta \omega \rightarrow 0$, we obtain

$$
\begin{equation*}
\Delta b_{3}^{\mathrm{T}}=\frac{1}{3^{\frac{d}{2}+1}} \frac{C}{\lambda_{T}^{2 d}}=\frac{1}{3^{\frac{d}{2}}} \Delta b_{3}, \tag{37}
\end{equation*}
$$

where in the latter equality we have used Eq. (28).
For $\Delta b_{4}^{\mathrm{T}}$, we need $\Delta Q_{211}^{\mathrm{T}}$, which is easily seen to be given by

$$
\begin{align*}
\Delta Q_{211}^{\mathrm{T}} & =C \sum_{\mathbf{x}, \mathbf{x}^{\prime}} \rho^{2}(\mathbf{x} ; \beta \omega)\left[\rho(\mathbf{x} ; \beta \omega) \rho\left(\mathbf{x}^{\prime} ; \beta \omega\right)-\rho^{2}\left(\mathbf{x}, \mathbf{x}^{\prime} ; \beta \omega\right)\right] \\
& =\Delta Q_{111}^{\mathrm{T}} Q_{1}^{\mathrm{T}} / 3-C \sum_{\mathbf{x}, \mathbf{x}^{\prime}} \rho^{2}(\mathbf{x} ; \beta \omega) \rho^{2}\left(\mathbf{x}, \mathbf{x}^{\prime} ; \beta \omega\right), \tag{38}
\end{align*}
$$

where

$$
\begin{equation*}
\rho\left(\mathbf{x}, \mathbf{x}^{\prime} ; \beta \omega\right)=\sum_{\mathbf{n}} e^{-\beta E_{\mathbf{n}}} \phi_{\mathbf{n}}(\mathbf{x}) \phi_{\mathbf{n}}\left(\mathbf{x}^{\prime}\right) \tag{39}
\end{equation*}
$$

which, using the Mehler kernel, becomes

$$
\begin{equation*}
\rho\left(\mathbf{x}, \mathbf{x}^{\prime} ; \beta \omega\right)=\frac{\omega^{\frac{d}{2}} e^{-\omega \operatorname{coth}(\beta \omega)\left(\mathbf{x}^{2}+\mathbf{x}^{\prime 2}\right) / 2+\omega \operatorname{csch}(\beta \omega) \mathbf{x} \cdot \mathbf{x}^{\prime}}}{(2 \pi \sinh (\beta \omega))^{\frac{d}{2}}} \tag{40}
\end{equation*}
$$

Thus, in the continuum limit,

$$
\begin{align*}
\Delta b_{4}^{\mathrm{T}} & =3 \frac{\Delta Q_{211}^{T}}{Q_{1}^{\mathrm{T}}}-\Delta b_{3}^{\mathrm{T}} Q_{1}^{\mathrm{T}} \\
& =-\frac{3 C}{Q_{1}^{\mathrm{T}}} \sum_{\mathbf{x}, \mathbf{x}^{\prime}} \rho^{2}(\mathbf{x} ; \beta \omega) \rho^{2}\left(\mathbf{x}, \mathbf{x}^{\prime} ; \beta \omega\right) \\
& =-\frac{C}{\lambda_{T}^{2 d}} \frac{1}{2^{\frac{d}{2}}}\left[\frac{\beta \omega}{\sinh (\beta \omega)} \frac{1}{(1+3 \cosh (\beta \omega))^{\frac{1}{2}}}\right]^{d}  \tag{41}\\
& =-\frac{3^{\frac{d}{2}+1}}{2^{\frac{d}{2}}} \frac{1}{(1+3 \cosh (\beta \omega))^{\frac{d}{2}}} \Delta b_{3}^{\mathrm{T}} . \tag{42}
\end{align*}
$$

Note that, in the $\beta \omega \rightarrow 0$ limit, our approximation yields

$$
\begin{equation*}
\Delta b_{4}^{\mathrm{T}}=-\frac{3^{\frac{d}{2}}}{2^{d}} \frac{3}{2^{\frac{d}{2}}} \Delta b_{3}^{\mathrm{T}}=-\frac{1}{2^{d}} \frac{3}{2^{\frac{d}{2}}} \Delta b_{3}, \tag{43}
\end{equation*}
$$

which we use below.

## IV. RESULTS IN HOMOGENEOUS SPACE

## A. Virial coefficients

Using the steps outlined above, we have calculated $\Delta b_{4}$ and $\Delta b_{5}$ and obtained

$$
\begin{align*}
\Delta b_{4} & =-C \frac{Q_{1} Q_{1}(2 \beta)}{9 V^{2}}=-3 \frac{Q_{1}(2 \beta)}{Q_{1}} \Delta b_{3},  \tag{44}\\
\Delta b_{5} & =C\left(\frac{\left(Q_{1}(2 \beta)\right)^{2}}{9 V^{2}}+\frac{Q_{1} Q_{1}(3 \beta)}{9 V^{2}}\right) \\
& =\left(\frac{3\left(Q_{1}(2 \beta)\right)^{2}}{Q_{1}^{2}}+\frac{3 Q_{1}(3 \beta)}{Q_{1}}\right) \Delta b_{3} \tag{45}
\end{align*}
$$

for the fermionic three-species system with a three-body contact interaction in $d$ spatial dimensions. In the latter equation, the first term on the right-hand side represents the contribution of $Q_{221}$, and the second term that of $Q_{311}$.

In the continuum limit, it is easy to perform the resulting Gaussian integrals that determine $Q_{1}$ and obtain

$$
\begin{gather*}
\Delta b_{4}=-\frac{3}{2^{\frac{d}{2}}} \Delta b_{3},  \tag{46}\\
\Delta b_{5}=3\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) \Delta b_{3} . \tag{47}
\end{gather*}
$$

Using these results, one may calculate the pressure, density, compressibility, and even Tan's contact (with knowledge of $\Delta b_{3}$ as a function of the interaction strength, e.g., $\beta \epsilon_{B}$ in one or two dimensions, where $\epsilon_{B}$ is the trimer binding energy). To provide a description of the thermodynamics that is as universal as possible across spatial dimensions, we use $\Delta b_{3}$ as the measure of the interaction strength and display our results in terms of that parameter. Furthermore, one may also define a (dimensionless) contact density as

$$
\begin{equation*}
\mathcal{C}=\frac{\lambda_{T}^{d}}{V} \frac{\partial \ln \mathcal{Z}}{\partial \Delta b_{3}} \tag{48}
\end{equation*}
$$

which differs from the conventional definition by a chainrule factor $\partial \Delta b_{3} / \partial \lambda$ (which in turn can be determined by solving the three-body scattering problem), where $\lambda$ is the $d$-dimensional coupling constant. To make the expression dimensionless, we have used the thermal wavelength $\lambda_{T}=$ $\sqrt{2 \pi \beta}$.

## B. Thermodynamics and contact across dimensions

The interaction-induced change in the pressure $\Delta P$ can be written in dimensionless form in arbitrary dimension as

$$
\begin{equation*}
\beta V \Delta P=Q_{1} \sum_{k=1}^{\infty} \Delta b_{k} z^{k} \tag{49}
\end{equation*}
$$



FIG. 1. Density, in units of $\lambda_{T}^{d}=(2 \pi \beta)^{d / 2}$, as a function of $\ln z=\beta \mu$, at $\Delta b_{3}=0.25$, for $d=1$ (top, blue line), $d=2$ (middle, green line), and $d=3$ (bottom, red line).

Similarly, the interaction-induced change in the density can be written as

$$
\begin{equation*}
\lambda_{T}^{d} \Delta n=3 \sum_{k=1}^{\infty} k \Delta b_{k} z^{k} \tag{50}
\end{equation*}
$$

and, using our definition of the contact in Eq. (49),

$$
\begin{equation*}
\Delta \mathcal{C}=3 \sum_{k=1}^{\infty} \frac{\partial \Delta b_{k}}{\partial \Delta b_{3}} z^{k} \tag{51}
\end{equation*}
$$

Implementing our LO-SCLA results, we obtain

$$
\begin{align*}
\beta \lambda_{T}^{d} \Delta P & \simeq 3 \Delta b_{3} z^{3}\left[1-\frac{3}{2^{\frac{d}{2}}} z+3\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) z^{2}\right]  \tag{52}\\
\lambda_{T}^{d} \Delta n & \simeq 9 \Delta b_{3} z^{3}\left[1-\frac{4}{2^{\frac{d}{2}}} z+5\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) z^{2}\right]  \tag{53}\\
\Delta \mathcal{C} & \simeq 3 z^{3}\left[1-\frac{3}{2^{\frac{d}{2}}} z+3\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) z^{2}\right] \tag{54}
\end{align*}
$$

As an example, in Fig. 1 we display the density as a function of the logarithm of the fugacity $\ln z=\beta \mu$ for $\Delta b_{3}=0.25$ and for $d=1,2,3$.

The behavior of $\Delta n$ as a function of $\beta \mu$ in Fig. 1 is as expected for a system with attractive interactions, namely, the interaction-induced change in the density is positive and enhanced by increasing $\beta \mu$ (or, equivalently, washed out at low densities, i.e., for large and negative $\beta \mu$ ). Also as expected (and as observed in Refs. [10] and [11] for two-body interactions), interaction effects are more pronounced in lower dimensions at fixed $\Delta b_{3}$.

## V. RESULTS IN A HARMONIC TRAP

## A. Fourth- and fifth-order virial coefficients

We have generalized our example of $\Delta b_{4}^{\mathrm{T}}$, discussed in a previous section, to $\Delta b_{5}^{\mathrm{T}}$. For future reference, we show both results:

$$
\begin{equation*}
\Delta b_{4}^{\mathrm{T}}=-\frac{3^{\frac{d}{2}+1}}{2^{\frac{d}{2}}} \frac{1}{(1+3 \cosh (\beta \omega))^{\frac{d}{2}}} \Delta b_{3}^{\mathrm{T}}, \tag{55}
\end{equation*}
$$



FIG. 2. $\Delta b_{4}^{T}$ (lower three, blue lines) and $\Delta b_{5}^{T}$ (upper three, red lines), in units of $\Delta b_{3}^{T}$, as a function of $\beta \omega$ in the LO-SCLA. Results are shown in $d=1$ (dotted lines), $d=2$ (dashed lines), and $d=3$ (solid lines).

$$
\begin{align*}
\Delta b_{5}^{\mathrm{T}}= & 3^{\frac{d}{2}+1}\left(\left[\frac{1}{12 \cosh ^{2}(\beta \omega)+4 \cosh (\beta \omega)-1}\right]^{\frac{d}{2}}\right. \\
& \left.+\left[\frac{1}{12 \cosh ^{2}(\beta \omega)+8 \cosh (\beta \omega)}\right]^{\frac{d}{2}}\right) \Delta b_{3}^{\mathrm{T}} \tag{56}
\end{align*}
$$

In Fig. 2 we show these results in $d=1,2,3$ as a function of $\beta \omega$. In contrast to the behavior of $\Delta b_{4}^{T}$ for the case of two-body interactions, explored in Refs. [12] and [14], here both $\Delta b_{4}^{\mathrm{T}}$ and $\Delta b_{5}^{\mathrm{T}}$ display monotonic behavior. Furthermore, at this order in the SCLA, both $\Delta b_{4}^{\mathrm{T}}$ and $\Delta b_{5}^{\mathrm{T}}$ are proportional to $\Delta b_{3}^{\mathrm{T}}$, such that the results in Fig. 2 are universal predictions in the sense of being coupling independent.

## B. A universal relation in the $\beta \omega \rightarrow 0$ limit

Note that, in the $\beta \omega \rightarrow 0$ limit, where the homogeneous system is recovered,

$$
\begin{equation*}
\Delta b_{5}^{\mathrm{T}} \rightarrow 3^{\frac{d}{2}+1} \frac{1}{5^{\frac{d}{2}}}\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) \Delta b_{3}^{\mathrm{T}}=\frac{3}{5^{\frac{d}{2}}}\left(\frac{1}{2^{d}}+\frac{1}{3^{\frac{d}{2}}}\right) \Delta b_{3} . \tag{57}
\end{equation*}
$$

Using Eqs. (44), (47), and (58), we find that trapped and untrapped virial coefficients are related, in the $\beta \omega \rightarrow 0$ limit, as follows:

$$
\begin{align*}
& \Delta b_{3}^{\mathrm{T}}=3^{-\frac{d}{2}} \Delta b_{3},  \tag{58}\\
& \Delta b_{4}^{\mathrm{T}}=4^{-\frac{d}{2}} \Delta b_{4},  \tag{59}\\
& \Delta b_{5}^{\mathrm{T}}=5^{-\frac{d}{2}} \Delta b_{5} \tag{60}
\end{align*}
$$

Although we have only explored $\Delta b_{n}^{\mathrm{T}}$ for $n=3,4$, and 5 here (the cases $n=1$ and 2 are trivially satisfied as well), the fact that the above relationship holds points us to conjecture that the relation

$$
\begin{equation*}
\left.b_{n}^{\mathrm{T}}\right|_{\beta \omega \rightarrow 0}=n^{-\frac{d}{2}} b_{n} \tag{61}
\end{equation*}
$$

is universally valid for all $n$, couplings, and temperatures (it is well known to be satisfied by noninteracting gases).

Other authors (see, e.g., [13,15,16]) have noted (and proven using the local density approximation) that this relationship is satisfied in the unitary limit (where the $b_{n}$ are temperature independent), and the same connection was found for $n=3,4$ in systems with two-body forces in Ref. [12] for arbitrary couplings (within the LO-SCLA). In principle, there is no special reason why $b_{n}^{\mathrm{T}}$ should not approach $b_{n}$ when the trapping potential is removed. That there is a $d$ - and $n$-dependent factor connecting these two quantities in the noninteracting case is merely a geometrical artifact of the choice of basis in which the calculations are performed (namely, the harmonic oscillator basis in the trapped case and plane waves in the homogeneous case), which has no impact on physical quantities. Based entirely on dimensional analysis, however, the natural guess is that $b_{n}^{\mathrm{T}}$ may approach $b_{n}$ times a dimensionless function of temperature and other dynamical scales. (That would actually change the partition function in a nontrivial way, in particular, concerning Tan's contact, but let us put that aside for the moment.) Such a dimensionless function could only result from the interplay between the trapping potential $\hat{V}_{\text {ext }}$ and the interaction $\hat{V}$, possibly leading to subtleties in the $\omega \rightarrow 0$ limit (similar to those arising from degenerate perturbation theory). However, the fact that $\left[\hat{V}_{\text {ext }}, \hat{V}\right]=0$ suggests that there should be no such subtlety and therefore no residual dependence on interaction-related scales in the relationship between $b_{n}^{\mathrm{T}}$ and $b_{n}$ as $\beta \omega \rightarrow 0$. In that limit, the dimensionless quantities $b_{n}^{\mathrm{T}}$ and $b_{n}$ should be related by a coupling- and temperature-independent function; their connection should be entirely geometrical and fully determined by the noninteracting case, for which $b_{n}^{\mathrm{T}}=n^{-\frac{d}{2}} b_{n}$ when $\beta \omega \rightarrow 0$. We therefore conclude that the conjecture is true for all $n$, coupling strengths, and temperatures.

## C. An exact relation across systems and dimensions

Finally, we point out a coupling-independent relationship between the 1D case with a three-body interaction (i.e., the 1D case of the system studied in this work) and the 2D case with only two-body interactions (denoted below by the superindex "2b2D"). As pointed out in Ref. [17], there exists an exact relationship between the three-body problem of the former situation and the two-body problem of the latter. That relationship yields a simple proportionality rule between the corresponding virial coefficients, given by

$$
\begin{equation*}
\Delta b_{3}=\frac{Q_{111}^{\mathrm{c} \cdot \mathrm{~m} .}}{Q_{11}^{\text {c.m. } 2 \mathrm{~b} 2 \mathrm{D}}} \frac{Q_{1}^{2 \mathrm{~b} 2 \mathrm{D}}}{Q_{1}} \Delta b_{2}^{2 \mathrm{~b} 2 \mathrm{D}} \tag{62}
\end{equation*}
$$

where the superscript "c.m." indicates the partition function associated with the center-of-mass motion, which is not affected by the interactions and completely factorizes (both in the spatially homogeneous and in the harmonically trapped case). In the spatially homogeneous case, the proportionality factor between $\Delta b_{3}$ and $\Delta b_{2}^{2 \mathrm{~b} 2 \mathrm{D}}$ is $1 / \sqrt{3}$, as shown in Ref. [17]. On the other hand, in the harmonically trapped case, the relationship becomes

$$
\begin{equation*}
\Delta b_{3}^{\mathrm{T}}=\frac{2}{3} \Delta b_{2}^{\mathrm{T}, 2 \mathrm{~b} 2 \mathrm{D}} \tag{63}
\end{equation*}
$$

We stress that while this relationship is restricted to the 1 D values of $\Delta b_{3}^{\mathrm{T}}$ and $\Delta b_{2}^{\mathrm{T}, 2 \mathrm{~b} 2 \mathrm{D}}$, it is valid for all couplings and all values of $\beta \omega$ and in that sense it is universal.

For completeness and future reference, we provide here details on the origin of this correspondence for the trapped case, which first appeared in Ref. [5], based on the 2D solution in Ref. [18]. The Schrödinger equation for this system takes the form

$$
\begin{equation*}
\left[-\frac{\nabla_{\mathbf{r}}^{2}}{2 m}+g \delta(x-y) \delta(y-z)+\frac{1}{2} m \omega^{2} \mathbf{r}^{2}\right] \psi(\mathbf{r})=E \psi(\mathbf{r}), \tag{64}
\end{equation*}
$$

where $x, y$, and $z$ again indicate the different-flavor particles, $\mathbf{r}=(x, y, z)$, and

$$
\begin{equation*}
\nabla_{\mathbf{r}}^{2}=\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}} \tag{65}
\end{equation*}
$$

Factoring out the center-of-mass motion by defining $Q=$ $\frac{1}{\sqrt{3}}(x+y+z), \quad q_{1}=\frac{1}{\sqrt{2}}(y-x), \quad q_{2}=\frac{1}{\sqrt{6}}(x+y-2 z)$, and $\psi(x, y, z)=\Phi(Q) \phi(\mathbf{q})$, with $\mathbf{q}=\left(q_{1}, q_{2}\right)$, we obtain

$$
\begin{equation*}
\left[-\frac{1}{2 m} \frac{\partial^{2}}{\partial Q^{2}}+\frac{1}{2} m \omega^{2} Q^{2}\right] \Phi(Q)=E_{\mathrm{c} . \mathrm{m} .} \Phi(Q), \tag{66}
\end{equation*}
$$

for the c.m. motion, and

$$
\begin{equation*}
\left[-\frac{\nabla_{\mathbf{q}}^{2}}{2 m}+\tilde{g} \delta(\mathbf{q})+\frac{1}{2} m \omega^{2} \mathbf{q}^{2}\right] \phi(\mathbf{q})=E_{r} \phi(\mathbf{q}) \tag{67}
\end{equation*}
$$

where $\tilde{g}=g / \sqrt{3}$ is the effective coupling and $E_{r}$ is the energy of relative motion, which is identical to that of a single particle in a 2 D harmonic oscillator potential with a $\delta$ potential at the origin. This establishes the exact relationship between our three-body 1D problem and its two-body 2D counterpart with two-body interactions.

As in the spatially homogeneous case, the eigenvalues $\epsilon_{\omega}=E_{r} / \omega$ of the harmonically trapped system are determined implicitly, in this case as solutions to

$$
\begin{equation*}
\frac{1}{\tilde{g}}=\frac{m}{\pi} \sum_{n=0}^{\Lambda_{\omega}} \frac{1}{\epsilon_{\omega}-(2 n+1)} \rightarrow \frac{m}{2 \pi}\left[\psi_{0}\left(\frac{1-\epsilon_{\omega}}{2}\right)-\ln \Lambda_{\omega}\right] \tag{68}
\end{equation*}
$$

where $\psi_{0}(z)$ is the digamma function, where $\Lambda_{\omega}$ is a UV cutoff. Unlike in the untrapped problem, with its unique bound state, the trapped problem admits an infinite set of discrete
excited states (all with positive energy). The problem is renormalized by relating the bare coupling to the $\epsilon_{\omega}$ occurring in the lowest-energy branch.

## VI. SUMMARY AND CONCLUSIONS

In this work we have calculated the high-temperature thermodynamics of three-flavored Fermi gases with a contact three-body interaction in $d$ spatial dimensions, as determined by the virial expansion. We carried out calculations in homogeneous space as well as in a harmonic trapping potential of frequency $\omega$. To that end, we implemented a coarse temporal lattice approximation at leading order (the LO-SCLA) and calculated the change in the virial coefficients $\Delta b_{n}$ due to interaction effects. In this context, we established a relation between the first two nontrivial virial coefficients, namely, $\Delta b_{4}$ and $\Delta b_{5}$, as functions of $\Delta b_{3}$. In addition, we argued that in the $\beta \omega \rightarrow 0$ limit, the relationship $\Delta b_{n}^{\mathrm{T}}=n^{-d / 2} \Delta b_{n}$ holds between the trapped and the homogeneous coefficients for arbitrary $n$, coupling strengths, and temperatures; furthermore, it is valid for systems with two- and three-body interactions. We showed that our calculations reproduce that relationship for $n=3,4,5$. Finally, we showed a relationship between the harmonically trapped case in one dimension with three-body interactions and its analog in two dimensions with two-body interactions, namely, $\Delta b_{3}^{\mathrm{T}}=\frac{2}{3} \Delta b_{2}^{\mathrm{T}, 2 \mathrm{~b} 2 \mathrm{D}}$. In closing, a comment is in order regarding the nature of purely contact (i.e., nonderivative) three-body interactions for $d>1$. The dimension of the coupling constant for such an $n$-body interaction is $L^{d(n-1)-2}$ (where $L$ has units of length), which is always positive for $n=3$ and $d>1$, indicating that the three-body interaction is an irrelevant operator in $d>1$. Therefore, renormalizing to a nontrivial interaction effect will likely require the systematic inclusion of a two-body interaction. The effect of not including the latter would appear when calculating at higher orders in the temporal lattice approximation, where it would manifest as a progressive shrinkage of the range of values of $\Delta b_{3}$ for which renormalization is possible with just a three-body coupling.

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