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# Electronic absorptions of $C_5^+$ detected in the visible through action spectroscopy in a cryogenic trap

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## ARTICLE HISTORY

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## ABSTRACT

The  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  electronic spectrum of  $C_5^+$  in the gas-phase with origin band at 513 nm is reported following experiments in a cryogenic ion trapping instrument. Buffer gas cooled  $C_5^+$  ions, generated by laser vaporisation of graphite, were investigated using two action spectroscopy approaches. Laser induced dissociation of weakly bound  $C_5^+ - He_n$  complexes synthesised in the trap reveal a linear dependence of the absorption energies on  $n$  allowing prediction of those of the bare ion  $C_5^+$ . These results are confirmed in two colour experiments on  $C_5^+$ , by monitoring fragmentation into the  $C_3^+ + C_2$  product channel. The data are supplemented with high level electronic structure calculations which support the assignment of  $D_{\infty h}$  symmetry to the ground electronic state. These laboratory results provide the requisite data needed for spectroscopic detection of this structure in terrestrial and extraterrestrial environments.

## KEYWORDS

cryogenic ion traps; action spectroscopy; carbon chains; interstellar molecules; helium tagging

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## 1. Introduction

Gas phase infrared spectra of the carbon chain  $C_5$  were reported in 1989 by several groups [1–4]. These rotationally resolved data indicate a linear cumulene structure with four identical bond lengths in the ground electronic state,  $X^1\Sigma_g^+$ . The  $\nu_3$  band around  $2164\text{ cm}^{-1}$  was used by Bernath *et al.* to identify  $C_5$  in the carbon star IRC+10216 [1]. The discovery of  $C_5$  in circumstellar regions, as discussed by Heath *et al.*, highlights the importance of carbon chains in the chemistry of astrophysical environments [2]. The cyanopolyacetylenes,  $NC_nH$ , have been known constituents of dense clouds in the interstellar medium since the 1970s (see Ref. 5 and references therein). Their presence led to the influential proposal by Douglas [6], that the photophysical properties of bare carbon chain molecules makes them appealing candidates for the diffuse interstellar bands (DIBs) in the optical region. A recent review discussing the measurement of the electronic transitions of various carbon chains in the laboratory and comparison with observational data is Ref. 7.

Electronic absorptions of  $C_5$  in the gas phase were detected by Maier and colleagues using cavity ring down spectroscopy [8], following earlier observation in a neon matrix [9]. The measured spectrum was assigned to  $^1\Pi_u \leftarrow X^1\Sigma_g^+$  with origin band at 510.94 nm. Additional electronic bands to higher energy were reported in Ref. 10 employing a resonant two-colour two-photon ionization scheme for detection. Rotational lines in the system near 511 nm were not resolved and the broadening attributed to intramolecular processes. The origin band was observed to have a full width at half maximum (FWHM) of around  $1\text{ cm}^{-1}$ , corresponding to an excited state lifetime of 5 ps. The assignment of this electronic transition was made in reference to theoretical calculations [11]. More recent *ab initio* calculations [12] find a 0.5 eV discrepancy between the calculated transition energy and the experimental measurement, and suggest instead that it is a formally forbidden transition, to either  $^1\Sigma_u^-$  or  $^1\Delta_u$ . The implications are discussed in Ref. 10, in relation to the experimentally measured band contour. The detection of gas phase electronic transitions of  $C_5$  enabled searches for its signatures in absorption in interstellar environments where  $C_3$  had been previously identified [13]. Observations through diffuse clouds toward  $\zeta$  Ophiuchi did not lead

to detection, and an upper limit to its column density of  $10^{-11} \text{ cm}^{-2}$  was determined along this line of sight [14, 15], and others [16], at least an order of magnitude below that of  $\text{C}_3$ .

High-resolution photoelectron spectra of  $\text{C}_5^-$ ,  $\text{C}_7^-$  and  $\text{C}_9^-$  at low temperature in the gas phase were recently reported by Neumark and coworkers [17, 18]. In contrast to the neutral and anion, however, there exists no gas phase experimental data on the electronic transitions of the singly charged cation  $\text{C}_5^+$ . Note that electronic spectra of larger positive, neutral, and negatively charged  $\text{C}_n$  fragments generated by dissociation of  $\text{C}_{60}^+$  were recorded by matrix isolation spectroscopy [19]. Theoretical work on the  $\text{C}_5^+$  open shell system relevant to the present work has been reported by several authors [20–23], leading to the question of whether the ground state is  $D_{\infty h}$  or  $C_{\infty v}$ . The most recent reports appear to clarify this, showing that the ground state has the highly symmetric structure [22, 23]. The calculations by Belau *et al.* also indicate the linear isomer is much preferred over cyclic  $\text{C}_{2v}$  structures [23]. The electronic configuration of  $\text{C}_5^+$ ,  $\dots 2\sigma_u^2, 1\pi_u^4, 3\sigma_g^2, 1\pi_g^4, 3\sigma_u^1$ , leads to a  $X^2\Sigma_u^+$  ground electronic state in  $D_{\infty h}$  symmetry. The results of large-scale *ab initio* coupled cluster and multi-reference configuration interaction (MRD-CI) calculations reported in Ref. 22 predict dipole allowed transitions from  $X^2\Sigma_u^+$  to low lying excited states at 0.05 eV ( ${}^2\Sigma_g^+ \leftarrow X^2\Sigma_u^+$ ,  $f = 1.3970$ ), 0.3 eV ( ${}^2\Pi_g \leftarrow X^2\Sigma_u^+$ ,  $f = 0.1686$ ), 2.62 eV ( ${}^2\Pi_g \leftarrow X^2\Sigma_u^+$ ,  $f = 0.0095$ ), and 3.36 eV ( ${}^2\Sigma_g^+ \leftarrow X^2\Sigma_u^+$ ,  $f = 0.0011$ ).

In this contribution, spectroscopic observations of electronic absorptions of  $\text{C}_5^+$  in the range 410 – 513 nm are reported. These measurements were carried out in the gas phase at low temperature through action spectroscopy in a cryogenic ion trap. Two approaches were adopted, the first involved tagging  $\text{C}_5^+$  ions with helium atoms and monitoring dissociation of the weak bond between  $\text{C}_5^+$  and He that occurred following electronic excitation of  $\text{C}_5^+ - \text{He}_n$ . The second spectroscopic method probed  $\text{C}_5^+$  ions more directly in a two colour experiment by monitoring fragmentation into  $\text{C}_3^+ + \text{C}_2$ .

This paper is organised as follows. Experimental and computational details are described in Section 2. Experimental results are presented and discussed in Section 3. Computational results are presented and discussed in relation to the proposed assignment in Section 4, and conclusions are given in Section 5.

## 2. Methods

### 2.1. Experimental

Experiments were carried out with the apparatus described in detail in Ref. 24. Ions were produced in a laser vaporization source containing a translating and rotating graphite rod. Radiation from a pulsed Nd:YAG laser at 355 nm (20 mJ/pulse) was focused with a  $f = 50$  cm lens onto the rod. The produced species were expanded in a helium pulse through a channel in the source block that runs perpendicular to the laser propagation direction. The gas was admitted by pulsing a piezo valve from  $-150$  V to  $+100$  V for  $200 \mu\text{s}$ , and the process repeated at a rate of 10 Hz. A variety of pure carbon cations with  $m/z$  separated by 12 u/e were synthesised, and a typical mass spectrum recorded under these conditions is shown in Figure 1(a).

Mass selected ions with  $m/z = 60$  were turned through  $90^\circ$  and loaded into the s-4PT trap [25] operating at a nominal temperature  $T_{\text{nom}} = 3.5$  K and with helium number densities of  $10^{15} \text{ cm}^{-3}$ .  $\text{C}_5^+$  ions were accumulated for 5 source laser pulses and cooled *via* inelastic collisions. Under these conditions, a significant fraction of  $\text{C}_5^+$  ions could be converted into  $\text{C}_5^+ - \text{He}_n$  complexes (Figure 1(b)). The helium buffer gas was pumped out for several hundred milliseconds before the contents were extracted and analysed using the second quadrupole mass spectrometer. The trapping cycle was repeated at a rate of 1 Hz

Photofragmentation spectra of  $\text{C}_5^+ - \text{He}_n$  were obtained by exposing the trapped ion cloud to tuneable radiation from either a pulsed OPO or a dye laser before extraction. The latter was operated with Coumarin 503 dye, supplying radiation in the visible (480–540 nm). The number of complexes were monitored on alternating storage cycles with ( $N_i$ ) and without ( $N_0$ ) exposure to radiation in order to account for fluctuations in the number of ions trapped. Two colour experiments on  $\text{C}_5^+$  were carried out at elevated trap temperatures ( $T_{\text{nom}} = 8$  K) to avoid  $\text{C}_5^+ - \text{He}_n$  formation. Following electronic excitation using the dye laser, the  $\text{C}_5^+$  ions were fragmented into  $\text{C}_3^+ + \text{C}_2$

by absorption of radiation from a pulsed OPO. Further details on the spectroscopic detection are provided and discussed in Section 3.3.

## 2.2. Computational

Quantum chemical calculations were carried out with the aim of aiding the interpretation of the experimental data. The geometry of  $C_5^+$  was first obtained using the DFT based  $\omega$ B97XD method and the 6-311+G(3df,3pd) basis set using Gaussian09 [26].  $\omega$ B97XD is a range-separated hybrid DFT functional successfully applied in the calculation of excitation energies of various organic molecules [27]. Due to the highly multi-configurational character of the ground state wave function of  $C_5^+$  the complete active space self-consistent field with second order perturbation theory correction (XMS-CASPT2/CASSCF) method and the cc-pVDZ basis set [28] were used to refine the geometry [29, 30]. This method accounts for both dynamic and static electronic correlation and offers systematically more accurate results.

The XMS-CASPT2/CASSCF optimisation was carried out using the Bagel program [31, 32], starting from the structure obtained from the prior optimization with the  $\omega$ B97XD/6-311+G(3df,3pd) method. An active space consisting of 9 electrons correlated in 8 orbitals was selected for the optimisation. The molecular orbitals included in the active space are presented in the Supplemental Material (Figure S1). This active space proved to be very stable for all types of calculations that we performed, and enabled successful geometry optimizations and comparison of the energies of the symmetric and non-symmetric forms of  $C_5^+$ . The active space was also selected with consideration of the rules proposed by Veryazov *et al.* [33], which suggest inclusion of molecular orbitals having occupations between 0.02 and 1.98.

The geometry optimizations performed with the  $\omega$ B97XD method and without any symmetry restrictions returned a less symmetric  $C_\infty$  geometry as opposed to the  $D_{\infty h}$  structure returned by XMS-CASPT2 calculations. However, owing to the lack of an analytical Hessian for the XMS-CASPT2 method, the ground state geometry of  $C_5^+$  at the  $\omega$ B97XD level of theory was reoptimized with imposed  $D_{\infty h}$  symmetry restrictions. This allowed calculation of the harmonic vibrational frequencies of  $C_5^+$  in the ground electronic state, which are presented in the Supplemental Material (Figure S2 and

Table S1).

Vertical excitation energies and single point energies of the DFT and the XMS-CASPT2/CASSCF structure were calculated at the NEVPT2/SA-CASSCF level. It is also worth noting that analytical energy gradients have not been implemented for the NEVPT2 method, so far, and that is why the ground-state geometry optimizations were performed at the equivalent XMS-CASPT2 level. The NEVPT2 approach is similar in nature to the XMS-CASPT2 method and allowed use of the larger cc-pVTZ basis set in the excited-state calculations, which was necessary to more accurately reproduce the excitation energy of the key optically bright excited state. The active space used for vertical excitation energies included 9 electrons correlated in 8 orbitals and the CASSCF calculation involved averaging over 9 electronic states (abbreviated as SA-CASSCF). The orbitals were selected as described for the optimisation procedure based on molecular orbital occupations. Due to the non-negligible effect of spin-orbit coupling in the  $^2\Pi$  ground state of the  $C_5^-$  anion [17], the Breit-Pauli spin-orbit coupling operator was included in the calculations, together with the scalar relativistic 2<sup>nd</sup> order Douglas-Kroll-Hess Hamiltonian. For these calculations, the relativistically contracted cc-pVTZ-DK basis set was used.

Geometry optimisations were attempted for the  $C_5^+ - He$  complex, however, because of its very flat potential energy surface and limited numerical accuracy of the optimization procedure, substantial convergence problems were encountered. No clear minima were located at the XMS-CASPT2/CASSCF/cc-pVDZ level despite attempts using numerous initial geometries. This is a direct consequence of the very low interaction energies between the He atom and the  $C_5^+$  ion. Therefore, to evaluate the effect of helium tagging in qualitative manner, rigid potential energy surface scans were carried out at the NEVPT2/SA-CASSCF/aug-cc-pVTZ level of theory by changing the distance between the He atom the minimum-energy geometry of the  $C_5^+$  molecule. Three different configurations of the tagged species were selected, for which the  $C_5^+ \dots He$  distance either along the molecular axis of  $C_5^+$  (one scan) or perpendicular to it (two different scans) was changed. The different scans and their associated coordinates are indicated in Figure 2 and S3 of the Supplemental Material. All the NEVPT2/SA-CASSCF calculations were carried out using the ORCA 4.2.1 program package [34].

### 3. Results and Discussion

#### 3.1. Photodissociation of $C_5^+ - He$

Weakly bound  $C_5^+ - He$  complexes synthesised in the ion trap were spectroscopically probed by one photon dissociation. The attenuation of  $C_5^+ - He$  ions in the range  $\sim 410 - 520$  nm ( $19230 - 24390$   $cm^{-1}$ ), reveals rich structure, as shown in Figure 3, where an overview spectrum recorded using a  $5$   $cm^{-1}$  bandwidth OPO is presented. The separation of peaks suggests assignment can be made to vibrational activity in the excited electronic state by comparison with the known vibrational frequencies of  $C_5$ , as listed in Table 8 of Ref. 35. The lowest energy absorption near  $19470$   $cm^{-1}$  is not the most intense, and the overall pattern in the region  $19400 - 20000$   $cm^{-1}$  is reminiscent of an electronic transition involving a geometry change. The  ${}^2\Pi_g \leftarrow X {}^2\Sigma_u^+$  band system is expected in this region [22], and assignment of the absorption observed at around  $513$ nm ( $\sim 2.42$  eV) to the origin agrees with the theoretical prediction, as it lies within the uncertainty of about  $\sim 0.2$  eV quoted for calculations at the same level on  $C_5$  [12].

Experimental data recorded using a dye laser ( $\Delta\bar{\nu} \sim 0.07$   $cm^{-1}$ ) across the low energy part of the spectrum is presented in Figure 4(a). The strongest absorption lies  $112$   $cm^{-1}$  to higher energy, and may thus be due to the excitation of the low frequency cis bending mode  $\nu_7$  in the excited electronic state (see Figure S2 for a visualisation of the calculated  $C_5^+$  vibrations). The pattern of absorptions to higher energy indicates a progression in this mode up to  $v' = 4$ . The absorption  $212$   $cm^{-1}$  to the blue of the origin is consistent with the excitation of one quanta of  $\nu_5$ , while the width of the broader feature near  $19825$   $cm^{-1}$  suggests it may comprise two unresolved absorptions. Based on the energy separation from the origin, these may be due to the excitation of 3 quanta of  $\nu_7$  and the fundamental of the  $\nu_5 + \nu_7$  combination band. The combs in Figure 4(a) indicate these tentative assignments, while Table 1 additionally provides the wavenumbers and widths of the absorptions, as determined through Lorentzian fits to the experimental data. The other strong absorptions in the



range 21300 – 21500  $\text{cm}^{-1}$ , resembling the pattern in the region 19400 – 20000  $\text{cm}^{-1}$ , may be combination bands involving the above and the totally symmetric  $\nu_1$  mode.

The  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  electronic transition of  $\text{C}_5^+$  is dipole allowed and the symmetry of the transition dipole moment must transform as  $\Pi_u$  in  $D_{\infty h}$ . Thus vibronic transitions should only be observable if  $\Gamma'_{vib} \otimes \Gamma''_{vib}$  contains  $\Sigma_g^+$ .  $\nu_7$  has  $\pi_u$  symmetry and therefore excitations from the vibrationless  $X\,{}^2\Sigma_u^+$  ground state to the excited  ${}^2\Pi_g$  electronic state with odd quanta in  $\nu_7$  are forbidden. The symmetry of these odd quanta vibronic states is  $\pi_u \otimes \Pi_g \longrightarrow \Sigma_u^+ \oplus \Sigma_u^- \oplus \Delta_u$ . The  $\Sigma_u^+$  component of the electric dipole operator could allow access to excited  ${}^2\Sigma_u^+$  vibronic state(s). In this case, the transition may “borrow” intensity from the nearby  ${}^2\Sigma_g^+ \leftarrow X\,{}^2\Sigma_u^+$  transition (Herzberg-Teller effect). This has also been attributed to the appearance of vibronic bands in the electronic spectrum of  $\text{l-C}_5\text{H}^+$  [36].

The observed electronic spectrum shows no rotational structure. A possible reason is due to the short lifetime of the excited electronic state. Simulations of the  ${}^1\Pi_u \leftarrow X\,{}^1\Sigma_g^+$  electronic transition of  $\text{C}_5$  at 511 nm are reported in Refs. 8 & 10. For this neutral system the rotational structure was also unresolved with each line broadened by 1  $\text{cm}^{-1}$ , corresponding to an excited state lifetime of 5 ps. The observed absorption bands, corresponding to the  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  electronic transition of  $\text{C}_5^+$ , possess Lorentzian profiles with FWHM in the range 15 – 25  $\text{cm}^{-1}$ . This observation indicates the spectrum is rather insensitive to changes to the rotational constant between the involved electronic states and suggests excited state lifetimes of 0.2 – 0.3 ps. Note that for  $\text{C}_5^+$  ions stored in 4 K helium buffer gas, translational temperatures ( $T_{\text{trans}}$ ) as high as 150 K would still lead to rotational temperature below 15 K. Simulations using the rotational constants of  $\text{C}_5$  show only a small increase ( $\sim 1 - 2\text{ cm}^{-1}$ ) in the FWHM upon increasing the rotational temperature from 10 to 100 K in such a lifetime broadened transition. Moreover, the expected Doppler width of rotational lines at  $T_{\text{trans}} = 150\text{ K}$  would be just  $10^{-2}\text{ cm}^{-1}$ , making a negligible contribution to the experimental width.

### 3.2. Photodissociation of $C_5^+ - He_n$

Under conditions of low temperature and high buffer gas number density,  $C_5^+$  ions with multiple helium atoms attached are readily formed in the trap. This enabled photofragmentation spectra of  $C_5^+ - He$ ,  $C_5^+ - He_2$ , and  $C_5^+ - He_3$  to be obtained by monitoring the attenuation of ions with  $m/z = 64$ , 68, and 72. The data presented in Figure 5, recorded in the range  $19300 - 20000 \text{ cm}^{-1}$ , show a clear blue-shift in the wavelengths at band maxima, and a broadening, with increasing  $n$ . Concerning the widths, the effect is most evident from inspection of the peaks near  $19700 \text{ cm}^{-1}$ , which are resolved for  $n = 1$  but much less so for  $n = 3$ . A possible reason for this is due to an isomeric effect, similar to the situation observed for the much larger structure  $C_{60}^+$ .

The lowest energy absorption band, assigned to the band origin, shifts from  $19471 \text{ cm}^{-1}$  for  $n = 1$  to  $19481 \text{ cm}^{-1}$  for  $n = 2$ , and  $19490 \text{ cm}^{-1}$  for  $n = 3$ . This  $\sim 10 \text{ cm}^{-1}$  displacement of the absorption per helium atom is linear, and consideration of other bands in the spectra shown in Figure 5 leads to similar results. Note that the magnitude of the shift is significantly larger than for the electronic transitions of  $C_{60}^+$  in the near infrared region, where the helium perturbs the spectrum by a fraction of a wavenumber [37–39].

The linear relationship between the absorption wavenumbers and  $n$  suggests the data can be used to estimate the values at which the electronic transition of  $C_5^+$  ( $n = 0$ ) occurs. In the Supplemental Material (Figure S4), a fit to the wavenumbers of the origin bands in the  $n = 1 - 3$  spectra presented in Figure 5 is shown. The slope of the fit indicates a shift per helium atom of  $\sim 10 \text{ cm}^{-1}$ . Assuming the linear behaviour holds between  $n = 0$  and  $n = 1$ , the intercept can be used to predict the corresponding absorption band of  $C_5^+$  at  $19463 \text{ cm}^{-1}$ . This prediction is quite accurate, as described in the following Section. The extrapolated  $n = 0$  wavenumbers are listed in Table 2.

### 3.3. Fragmentation of $C_5^+$

Further experiments were carried out to obtain the spectrum of bare  $C_5^+$ , free from

perturbation by helium atom(s). A two-colour approach proved successful, with the absorption bands detected by fragmentation of  $C_5^+$  into  $C_3^+ + C_2$ . In these experiments, a dye laser was scanned across the region  $19300 - 20000 \text{ cm}^{-1}$ , inducing the  ${}^2\Pi_g \leftarrow X {}^2\Sigma_u^+$  transition, while a second, fixed frequency laser operating at 398 nm, was used to fragment the electronically excited molecules. The scheme is shown in Figure 6. Electronic absorption by cold  $C_5^+$  ions was monitored through (1) the attenuation of ions with  $m/z = 60$  and (2) the appearance of  $C_3^+$  ions at  $m/z = 36$ . Note that in these 2-colour fragmentation experiments the temperature of the walls of the ion trap,  $T_{\text{nom}}$ , was raised to 8 K to prevent complications due to the formation of helium complexes. A typical mass-spectrum obtained following irradiation on resonance is presented in the Supplemental Material (Figure S5), showing the appearance of  $C_3^+$  with  $m/z = 36$ .

The number of  $C_3^+$  ions produced as a function of probe pulse energy at 398 nm is shown in Figure 7(a). A small, probe only, signal ( $N_B$ ) was subtracted to reveal the dependence of the two-colour signal. As shown, the data is non-linear and is reproduced with a second order function, suggesting a 2-photon process in the probe step in these measurements. The absorption of one 510 nm and two 398 nm photons increases the internal energy of (internally cold)  $C_5^+$  ions by  $\sim 9 \text{ eV}$ , well above the predicted fragmentation threshold (see Figure 6). The energy of two 398 nm photons lies near threshold and may account for the weak (few  $C_3^+$  ions / pulse) probe only signal.

Spectroscopic data obtained by monitoring the attenuation of ions with  $m/z = 60$  and the production of ions with  $m/z = 36$  are shown in Figure 4(b). Both detection schemes give similar results, however, as the latter method is (almost) background free the signal to noise ratio (S/N) is higher. The origin band maxima determined from Lorentzian fits are at  $19462.2 \pm 1.2 \text{ cm}^{-1}$  and  $19460.6 \pm 0.9 \text{ cm}^{-1}$  for  $m/z = 60$  and  $m/z = 36$ , respectively. The spectroscopic characteristics are listed in Table 2 in comparison with the extrapolated  $n = 0$  values inferred from the  $C_5^+ - \text{He}_n$  data. The band maxima predicted for  $n = 0$  are found to be in remarkable agreement with the results obtained by fragmentation of  $C_5^+$  itself.

The dependence of the number of  $C_3^+$  ions on the delay between the dye laser and

OPO is shown in Figure 7(b). The data are fit with an exponential decay, yielding a time constant of around  $60 \mu\text{s}$ . Remarkably, given the sub-picosecond lifetime of the excited electronic state, two-colour fragmentation can still be observed at a delay of  $100 \mu\text{s}$ . It therefore appears likely that following electronic excitation, intramolecular dynamics leads to population becoming trapped in a long lived state such that the probe, fragmentation, step, is still accessible after  $100 \mu\text{s}$ . It is interesting to compare this result with the electronic transition of  $\text{C}_5$ , [10] where it was suggested that either a nearby triplet state acts as a “sink”, or that the system becomes trapped in vibrationally excited levels in the ground electronic state. Whatever the explanation in the case of the present work on  $\text{C}_5^+$ , the photoexcited population remains available for fragmentation for even longer timescales.

#### 4. Theoretical interpretation of the results

The results of geometry optimisations are presented in the Supplemental Material (Figure S6). The DFT calculations agree well with previous calculations at the B3LYP/cc-pVDZ level of theory [21]. The  $\omega\text{B97XD}/6\text{-}311\text{+G}(3\text{df},3\text{pd})$  optimisation resulted in a structure belonging to the  $C_{\infty v}$  point group, with the four C–C bonds arranged in an alternating long-short manner. Previous optimisation at the CCSD(T)/cc-pVTZ level of theory [22] show that this structure is not the lowest energy arrangement, and resulted in a more symmetric  $D_{\infty h}$  geometry. This has two short terminal bonds and two longer inner C–C bonds. To further test the validity of the more symmetric geometry predicted by the CCSD(T) approach, the multi-configurational XMS-CASPT2/CASSCF method was used to optimise the ground state geometry.

The XMS-CASPT2 geometry optimization, initiated from that obtained at the  $\omega\text{B97XD}/6\text{-}311\text{+G}(3\text{df},3\text{pd})$  level, resulted in a nearly fully symmetric  $D_{\infty h}$  geometry structure, as shown in Figure S6 in comparison with the bond lengths reported in Ref 22. It is worth noting that no symmetry restrictions were involved in the geometry optimizations, and thus, some numerical inaccuracy of the optimization procedure is expected. The XMS-CASPT2 geometry is characterized by shorter terminal and longer inner C-C bond lengths, which is consistent with the trends observed by Schnell *et*

*al.*, and indicates that DFT calculations artificially stabilize the less symmetric  $C_{\infty v}$  structure.

Vertical excitation energies calculated using the NEVPT2/SA-CASSCF/cc-pVTZ approach are presented in Table 3. The molecular orbital characters of specific excited electronic states are consistent with the previous calculations performed at the MRDCI/cc-pVTZ level of theory [22]. Additionally, a splitting of the  $\pi$  orbitals was observed, which can be seen in Figure 8. This leads to two bright states close in energy with a separation of 0.18 eV. These  ${}^2\Pi \leftarrow X {}^2\Sigma$  transitions are expected at 2.35 eV and 2.53 eV, and while they lie close in energy to the two sets of absorption bands shown in Figure 3, they are believed to be caused by the slight deviation from the  $D_{\infty h}$  geometry. A calculation of the vertical excitations based on the fully symmetric structure obtained by Schnell *et al.* showed these states to be degenerate.

To provide a deeper theoretical understanding of the observed absorption spectrum, vertical excitation energies including spin-orbit coupling were calculated (see Section 2.2 for details). Unlike the relatively large effect in the  $C_5^-$  anion [17], spin-orbit coupling has a negligible effect on the vertical excitation energies of all the excited states calculated for  $C_5^+$ . Further calculations were carried out to investigate the Renner-Teller effect on the doubly degenerate  ${}^2\Pi$  state of linear  $C_5^+$ . In particular, splitting of the  ${}^2\Pi$  state is observed along the bending coordinate of the two terminal C-C-C angles, consistent with nuclear motion in the  $\nu_7$  mode that breaks the linearity of  $C_5^+$ . The  ${}^2\Pi$  state splits into  ${}^2A_2$  and  ${}^2B_2$  components ( $C_{2v}$  symmetry), and a systematic increase in the energy difference ( $\Delta E$ ) between these two along this reaction coordinate is observed, reaching  $\Delta E$  of 0.15 eV at  $30^\circ$ . Therefore, the appearance of the two groups of peaks near 19500 and 21500  $\text{cm}^{-1}$  (see Figure 3) may be due to this effect. The results of these calculations can be viewed in the Supplemental Material (Figure S7).

To evaluate the effect of helium tagging on the vertical excitation energy of  $C_5^+$ , rigid potential energy (PE) scans were calculated by varying the distance between the He atom and the  $C_5^+$  molecules, starting from three different initial arrangements of the  $C_5^+ \dots \text{He}$  complex (see Figure 2 and S3 of Supplemental Material). For these scans, distances in the range of 2.5 to 3.8 Å between the He atom and the nearest C atom

to it were considered. These scans demonstrate a very flat potential energy surface resulting from a weak interaction between the two subunits. As shown in Figure S3, the location of the minimum also varies between the different positions. The PE scan with the position of the He atom nearest to the central C atom (direction perpendicular to the  $C_5^+$  molecular axis), exhibits a shallow minimum at approximately 3.1 Å, which is also the shortest minimum-energy distance of the three different scans considered here. This also suggests that the strongest helium-ion interaction is close to the center of the ion.

The vertical excitation energies in the three directions are presented as a function of distance in Figure 2 in comparison to the isolated  $C_5^+$  ion. The strongest effect of He tagging can be observed for the linear complex, for which a clear bathochromic shift is seen, becoming more pronounced for short  $C_5^+ \dots \text{He}$  distances. In contrast, for the two non-linear complexes, a very mild effect of He tagging on the excitation energy of the bright state is observed. In particular, when the He atom is located near the center of one of the terminal C–C bonds, the vertical excitation energy of this state remains nearly unaffected for all the considered distances. When the He atom is placed in near proximity to the central C atom, a very mild hypsochromic shift is seen, which is slightly larger in magnitude at short  $C_5^+ \dots \text{He}$  distances. The latter is consistent with the experimental observation of a mild blue-shift of  $10 \text{ cm}^{-1}$  associated with He tagging. It also indicates that the complexes formed in the experiment are likely to be non-linear, with the He atoms residing somewhere near the center of the molecule. It is worth emphasising that although the calculated energy differences are below the accuracy of the NEVPT2 method, these rigid scans demonstrate qualitative trends and allow evaluation of the direction of the energy shift and the magnitude of the effect, which depend on the location of the He atom. Furthermore, the NEVPT2 results clearly demonstrate that each He atom should have a very subtle effect on the vertical excitation energy of this transition.

## 5. Conclusions

By combining laser vaporisation synthesis with spectroscopic characterisation in a cryogenic trap, the  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  electronic spectrum of the carbon chain  $C_5^+$  has been observed. The experimental data are consistent with theoretical results which indicate a linear, symmetric,  $D_{\infty h}$  structure in the ground electronic state. The pattern of vibronic bands in the observed spectrum is indicative of a geometry change in the excited electronic state. The computational results on the  $C_5^+ - \text{He}$  complex provide an indication of the structure produced in the laboratory. Although the experimental data do not allow  $C_5^+ - \text{He}_n$  structural information to be obtained directly, calculations show that a mild hypsochromic shift of the  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  transition is predicted only for complexes in which the He atom is located near the central C atom of  $C_5^+$ .

The spectra of  $C_n^+$  possessing strong electronic transitions in the visible are sought after in relation to the long standing enigma of the diffuse interstellar bands. These results add to the very limited experimental data set available on the gas phase electronic transitions of bare carbon cations at temperatures relevant to the interstellar medium [25, 40–42]. Moreover, the experimental approaches adopted here pave the way toward obtaining the spectra of larger  $C_n^+$  structures, including chains, rings, and fullerenes, that may be photostable in interstellar regions exposed to a harsh stellar radiation field where the presence of abundant  $C_{60}^+$  has recently been confirmed.

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## References

- [1] P.F. Bernath, K.H. Hinkle and J.J. Keady, *Science* **244**, 562–564 (1989).
- [2] J.R. Heath, A.L. Cooksy, M.H.W. Gruebele, C.A. Schmuttenmaer and R.J. Saykally, *Science* **244**, 564–566 (1989).
- [3] N. Moazzen-Ahmadi, A.R.W. McKellar and T. Amano, *Chem. Phys. Lett.* **157**, 1–4 (1989).
- [4] N. Moazzen-Ahmadi, A.R.W. McKellar and T. Amano, *J. Chem. Phys.* **91**, 2140–2147 (1989).
- [5] H.W. Kroto, *Int. Rev. Phys. Chem.* **1**, 309 (1981).
- [6] A.E. Douglas, *Nature* **269**, 130 (1977).
- [7] E.K. Campbell and J.P. Maier, *J. Chem. Phys.* **146**, 160901 (2018).
- [8] T. Motylewski, O. Vaizert, T.F. Giesen, H. Linnartz and J.P. Maier, *J. Chem. Phys.* **111**, 6161 (1999).
- [9] D. Forney, P. Freivogel, M. Grutter and J.P. Maier, *J. Chem. Phys.* **104**, 4954–4960 (1996).
- [10] A.E. Boguslavskiy and J.P. Maier, *J. Chem. Phys.* **125**, 094308 (2006).
- [11] M. Kolbuszewski, *J. Chem. Phys.* **102**, 3679–3684 (1995).
- [12] M. Hanrath and S.D. Peyerimhoff, *Chem. Phys. Lett.* **337**, 368–374 (2001).
- [13] J.P. Maier, N.M. Lakin, G.A.H. Walker and D.A. Bohlender, *Astrophys. J.* **553**, 267–273 (2001).
- [14] J.P. Maier, G.A.H. Walker and D.A. Bohlender, *Astrophys. J.* **566**, 332–335 (2002).
- [15] J.P. Maier, G.A.H. Walker and D.A. Bohlender, *Astrophys. J.* **602**, 286–290 (2004).
- [16] G.A. Galazutdinov, F.A. Musaev and J. Krelowski, *Mon. Not. R. Astron. Soc.* **325**, 1332–1334 (2001).
- [17] M.L. Weichman, J.B. Kim and D.M. Neumark, *J. Chem. Phys.* **139**, 144314 (2013).
- [18] M.C. Babin, J.A. DeVine, M.L. Weichmana and D.M. Neumark, *J. Chem. Phys.* **149**, 174306 (2018).
- [19] D.V. Strelnikov, M. Link, J. Weippert and M.M. Kappes, *J. Phys. Chem. A.* **123**, 5325 (2019).
- [20] K. Raghavachari, *Y. Phys. D. Atoms Molecules and Cluster* **12**, 61 (1989).
- [21] M.G. Giuffreda, M.S. Deleuze and J.P. Francois, *J. Phys. Chem. A.* **103**, 5137 (1999).
- [22] M. Schnell, M. Mühlhäuser, G.F. Froudakis and S.D. Peyerimhoff, *Chem. Phys. Lett.* **340**, 559–564 (2001).



- [23] L. Belau, S.E. Wheeler, B.W. Ticknor, M. Ahmed, S.R. Leone, W.D. Allen, H.F. Schaefer and M.A. Duncan, *J. Am. Chem. Soc.* **129**, 10229 (2007).
- [24] E.K. Campbell and P.W. Dunk, *Rev. Sci. Instr.* **90**, 103101 (2019).
- [25] E.K. Campbell, M. Holz, D. Gerlich, J.P. Maier, G.A.H. Walker and D. Bohlender, *Astrophys. J.* **822**, 17 (2016).
- [26] M.J. Frisch, G.W. Trucks, H.B. Schlegel, G.E. Scuseria, M.A. Robb, J.R. Cheeseman, G. Scalmani, V. Barone, G.A. Petersson, H. Nakatsuji, X. Li, M. Caricato, A. Marenich, J. Bloino, B.G. Janesko, R. Gomperts, B. Mennucci, H.P. Hratchian, J.V. Ortiz, A.F. Izmaylov, J.L. Sonnenberg, D. Williams-Young, F. Ding, F. Lipparini, F. Egidi, J. Goings, B. Peng, A. Petrone, T. Henderson, D. Ranasinghe, V.G. Zakrzewski, J. Gao, N. Rega, G. Zheng, W. Liang, M. Hada, M. Ehara, K. Toyota, R. Fukuda, J. Hasegawa, M. Ishida, T. Nakajima, Y. Honda, O. Kitao, H. Nakai, T. Vreven, K. Throssell, J.A.M. Jr., J.E. Peralta, F. Ogliaro, M. Bearpark, J.J. Heyd, E. Brothers, K.N. Kudin, V.N. Staroverov, T. Keith, R. Kobayashi, J. Normand, K. Raghavachari, A. Rendell, J.C. Burant, S.S. Iyengar, J. Tomasi, M. Cossi, J.M. Millam, M. Klene, C. Adamo, R. Cammi, J.W. Ochterski, R.L. Martin, K. Morokuma, O. Farkas, J.B. Foresman and D.J. Fox, *Gaussian 09 Revision D.01 2013*, Gaussian Inc. Wallingford CT.
- [27] J.D. Chai and M. Head-Gordon, *Physical Chemistry Chemical Physics* **10** (44), 6615–6620 (2008).
- [28] T.H. Dunning, *The Journal of Chemical Physics* **90** (2), 1007–1023 (1989).
- [29] K. Andersson, P.A. Malmqvist, B.O. Roos, A.J. Sadlej and K. Wolinski, *J. Phys. Chem.* **94** (14), 5483–5488 (1990).
- [30] T. Shiozaki, W. Győrffy, P. Celani and H.J. Werner, *J. Chem. Phys.* **135** (8), 081106 (2011).
- [31] T. Shiozaki, *Wiley Interdiscip. Rev. Comput. Mol. Sci.* **8** (1), e1331 (2018).
- [32] BAGEL, Brilliantly Advanced General Electronic-structure Library. <http://www.nubakery.org> under the GNU General Public License .
- [33] V. Veryazov, P.Å. Malmqvist and B.O. Roos, *Int. J. Quantum Chem.* **111** (13), 3329–3338 (2011).
- [34] F. Wennmohs, D. Aravena, M. Atanasov, U. Becker, G. Bistoni, D. Bykov, V.G. Chilkuri, D. Datta, A. Kumar Dutta, D. Ganyushin, M. Garcia, Y. Guo, A. Hansen, B. Helmich-Paris, L. Huntington, R. Izsák, C. Kollmar, S. Kossmann, M. Krupička, L. Lang, D. Lenk, D.G. Liakos, D. Manganas, D.A. Pantazis, T. Petrenko, P. Pinski, C. Reimann, M. Rete-

gan, C. Riplinger, T. Risthaus, M. Roemelt, M. Saitow, B. Sandhöfer, I. Schapiro, K. Sivalingam, B. De Souza, G. Stoychev, W. Van Den Heuvel, B. Wezislá, M. Kállay, S. Grimme, E. Valeev, G. Chan, J. Pittner, M. Brehm, L. Goerigk, V. Vilhjálmurásgéirsson and W. Schneider, ORCA-An ab initio, DFT and semiempirical SCF-MO package-Version 4.2.1 Design and Scientific Directorship: Frank Neese And contributions from our collaborators, .

- [35] A.V. Orden and R.J. Saykally, *Chem. Rev.* **98**, 2313–2357 (1998).
- [36] J. Fulara, A. Nagy, A. Chakraborty and J.P. Maier, *J. Chem. Phys.* **144**, 244309 (2016).
- [37] E.K. Campbell, M. Holz, and J.P. Maier, *Astrophys. J. Lett.* **826**, L4 (2016).
- [38] M. Kuhn, M. Renzler, J. Postler, S. Ralser, S. Spieler, M. Simpson, H. Linnartz, A.G.G.M. Tielens, J. Cami, A. Mauracher, Y. Wang, F. Martín, M.K. Beyer, R. Wester, A. Lindinger and P. Scheier, *Nature Com.* **7**, 13550 (2016).
- [39] S. Spieler, M. Kuhn, J. Postler, M. Simpson, R. Wester, P. Scheier, W. Ubachs, X. Bacalla, J. Bouwman and H. Linnartz, *Astrophys. J.* **846**, 168 (2017).
- [40] E.K. Campbell, M. Holz, D. Gerlich and J.P. Maier, *Nature* **523**, 322 (2015).
- [41] E.K. Campbell, M. Holz and J.P. Maier, *Astrophys. J.* **835**, 221 (2017).
- [42] E.K. Campbell, *Mol. Phys.* **118**, e1797918 (2020).

**Table 1.** Observed absorption bands in the  ${}^2\Pi_g \leftarrow X\,{}^2\Sigma_u^+$  electronic spectrum of  $\text{C}_5^+ - \text{He}$ .

$\tilde{\nu} / \text{cm}^{-1}$	FWHM / $\text{cm}^{-1}$	$\Delta\tilde{\nu} / \text{cm}^{-1}$	Assignment
$19471.3 \pm 0.9$	$20.6 \pm 2.7$	0	$0_0^0$
$19587.1 \pm 0.4$	$33.5 \pm 1.4$	115.8	$\nu_7 / 7_0^1$
$19683.1 \pm 0.7$	$34.3 \pm 2.5$	211.8	$\nu_5 / 5_0^1$
$19718.5 \pm 1.0$	$23.2 \pm 3.4$	247.2	$2\nu_7 / 7_0^2$
$19812.9 \pm 5.2$	$36.2 \pm 9.7$	341.6	$\nu_5 + \nu_7 / 5_0^1 \ 7_0^1$
$19833.9 \pm 5.3$	$36.9 \pm 9.5$	362.6	$3\nu_7 / 7_0^3$
$19952.1 \pm 0.9$	$33.5 \pm 2.6$	480.8	$4\nu_7 / 7_0^4$

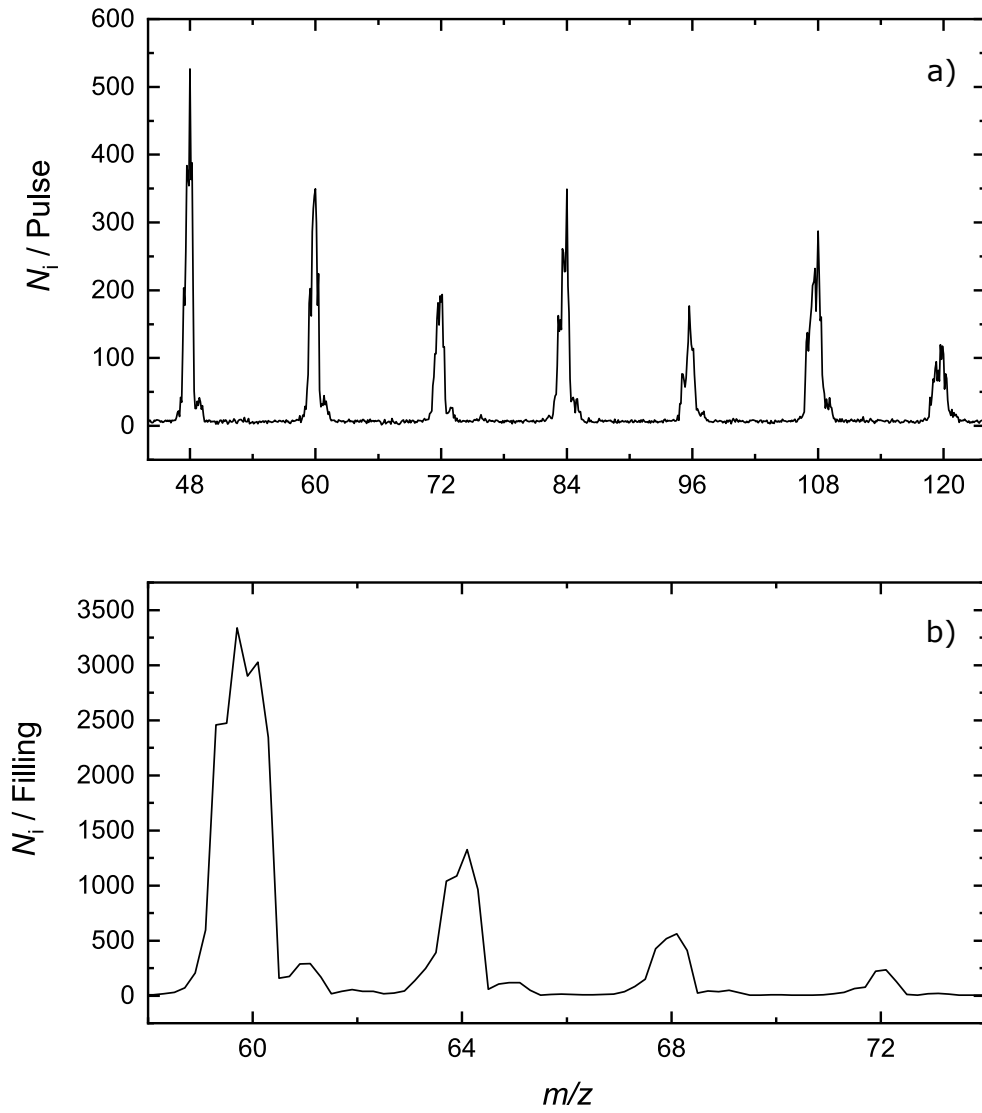
**Table 2.** Comparison of  $C_5^+$  absorption characteristics inferred from different spectroscopic methods.

Extrapolated from $C_5^+ - He_n$		$C_3^+$ Mass Channel		$C_5^+$ Attenuation	
$\tilde{\nu} / \text{cm}^{-1}$	FWHM / $\text{cm}^{-1}$	$\tilde{\nu} / \text{cm}^{-1}$	FWHM / $\text{cm}^{-1}$	$\tilde{\nu} / \text{cm}^{-1}$	FWHM / $\text{cm}^{-1}$
19462.8 $\pm$ 3.0	20.6 $\pm$ 2.7	19460.6 $\pm$ 0.9	14.9 $\pm$ 1.2	19462.2 $\pm$ 1.2	13.4 $\pm$ 3.7
19578.6 $\pm$ 2.5	33.5 $\pm$ 1.4	19575.7 $\pm$ 0.4	24.6 $\pm$ 1.2	19575.2 $\pm$ 0.8	24.7 $\pm$ 2.5
19674.6 $\pm$ 2.8	34.3 $\pm$ 2.5	19675.5 $\pm$ 0.5	22.6 $\pm$ 1.7	19675.1 $\pm$ 0.9	21.7 $\pm$ 3.0
19710.0 $\pm$ 3.1	23.2 $\pm$ 3.4	19706.2 $\pm$ 0.9	16.8 $\pm$ 2.3	19708.0 $\pm$ 1.1	14.0 $\pm$ 4.3
19804.4 $\pm$ 7.3	36.2 $\pm$ 9.7	19801.6 $\pm$ 1.8	19.4 $\pm$ 6.2	19803.9 $\pm$ 5.0	30.6 $\pm$ 12.2
19825.4 $\pm$ 7.4	36.9 $\pm$ 9.5	19821.6 $\pm$ 2.7	29.5 $\pm$ 6.0	19826.5 $\pm$ 2.7	19.9 $\pm$ 9.2
19943.6 $\pm$ 3.0	33.5 $\pm$ 2.6	19940.5 $\pm$ 0.7	22.9 $\pm$ 2.1	19940.6 $\pm$ 1.9	25.3 $\pm$ 5.8

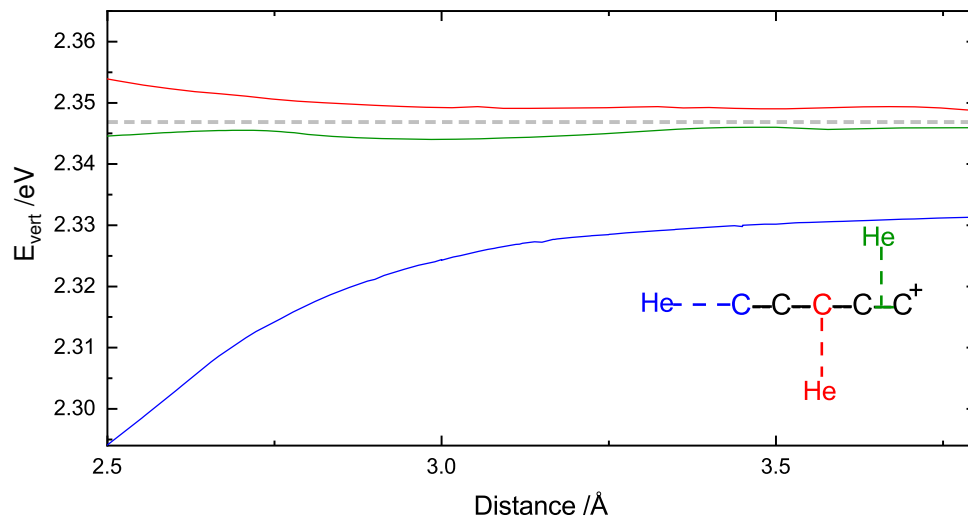
**Table 3.** Vertical excitation energies of isolated  $C_5^+$  calculated using NEVPT2/SA-CASSCF/cc-pVTZ, based on the minimum-energy geometries optimized with the XMS-CASPT2/CASSCF/cc-pVDZ method.

State <sup>1</sup>	Configuration	$E_{\text{exc}}/\text{eV}$	$f_{\text{osc}}$	$\tilde{\nu}/\text{cm}^{-1}$
$X^2\Sigma_u^+$	$3\sigma^{*1}$	-	-	-
$^2\Sigma_g^+$	$3\sigma^1 3\sigma^{*2}$	0.03	$8.01 \times 10^{-4}$	254.7
$^2\Pi_g$	$3\sigma^2 3\sigma^{*2} 1\pi_x^{*1}$	0.67	$3.84 \times 10^{-4}$	5369.2
$^2\Pi_g$	$3\sigma^2 3\sigma^{*2} 1\pi_y^{*1}$	0.78	$7.27 \times 10^{-4}$	6312.0
$^2\Pi_u$	$3\sigma^2 3\sigma^{*0} 2\pi_x^1$	1.38	$2.94 \times 10^{-4}$	11162.8
$^2\Pi_u$	$3\sigma^2 3\sigma^{*0} 2\pi_y^1$	1.66	$5.33 \times 10^{-4}$	13362.0
$^2\Pi_g$	$3\sigma^1 3\sigma^{*1} 2\pi_x^1$	2.35	$5.90 \times 10^{-3}$	18932.7
$^2\Pi_g$	$3\sigma^1 3\sigma^{*1} 2\pi_y^1$	2.53	$9.94 \times 10^{-3}$	20375.5
$^2\Pi_u$	$1\pi_x^2 1\pi_y^1 3\sigma^2 3\sigma^{*2}$	3.10	$5.14 \times 10^{-5}$	25011.3

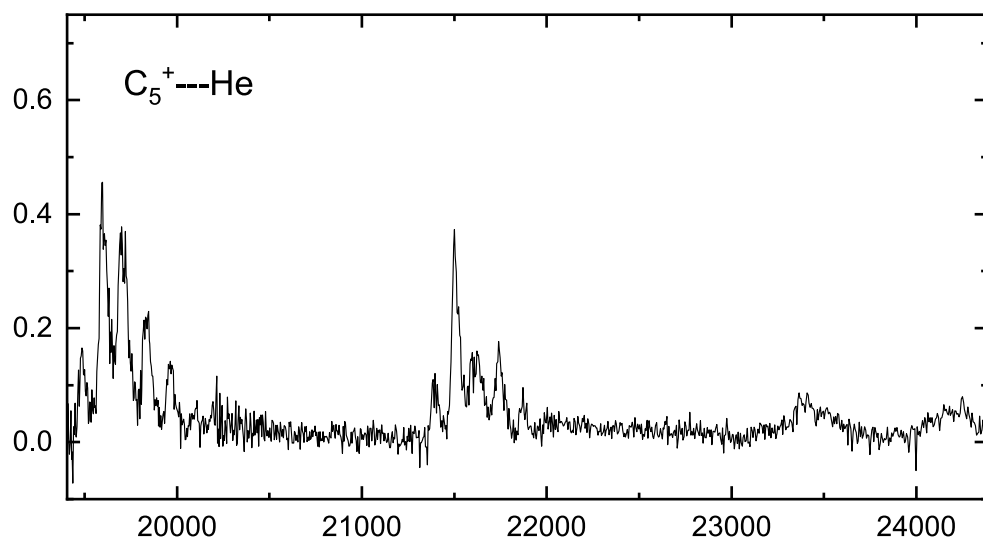
<sup>1</sup> Nomenclature is for  $D_{\infty h}$  symmetry (see the text for a discussion)



**Figure 1.** (a) Mass spectrum demonstrating formation of  $C_n^+$  ( $n = 4 - 10$ ) in the laser vaporisation source. (b) Mass spectrum of trap contents following storage of  $C_5^+$  ions in cold ( $T_{\text{nom}} = 3.7 \text{ K}$ ) and dense ( $10^{15} \text{ cm}^{-3}$ ) helium buffer gas. The data indicate the formation of  $C_5^+ - \text{He}_n$  with  $n = 1 - 3$ .

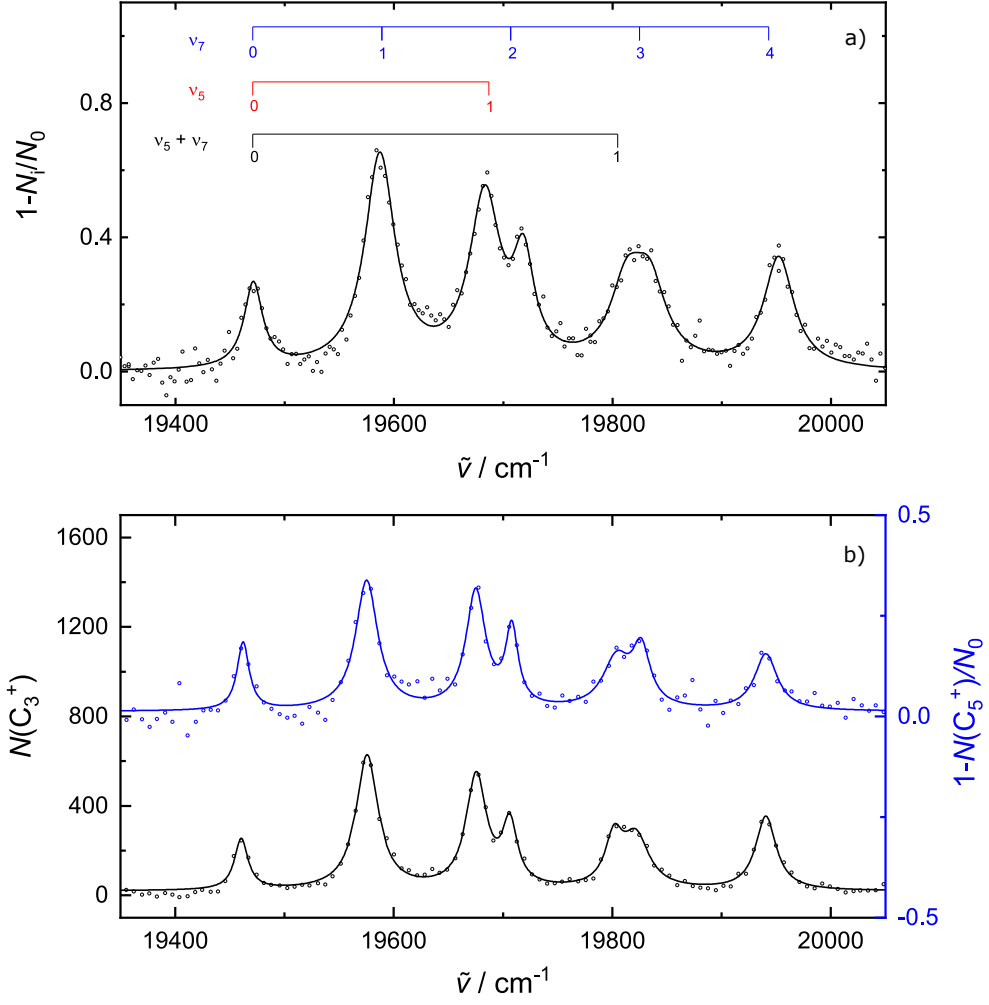


**Figure 2.** Vertical excitation energies, calculated at the NEVPT2/SA-9-CASSCF(8,9)/aug-cc-pVTZ level of theory, for three different isomers of the  $\text{C}_5^+ - \text{He}$  complex, as a function of the distances between the two subunits presented in the inset and described in the text. The grey horizontal dashed line indicates the value for  $\text{C}_5^+$ .

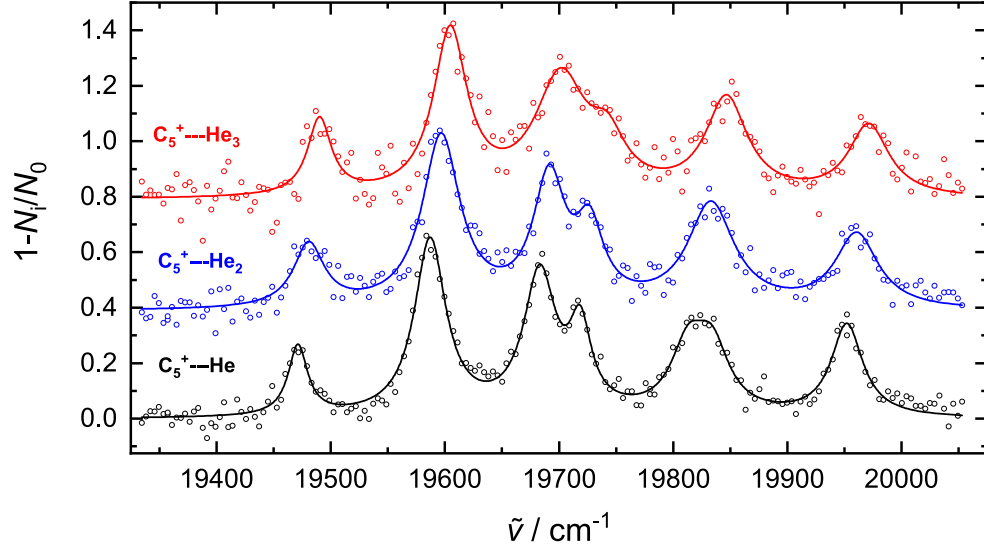


**Figure 3.** Overview spectrum showing  ${}^2\Pi_g \leftarrow X^2\Sigma_u^+$  electronic transition of  $C_5^+ - He$  observed by one-colour fragmentation with an OPO.

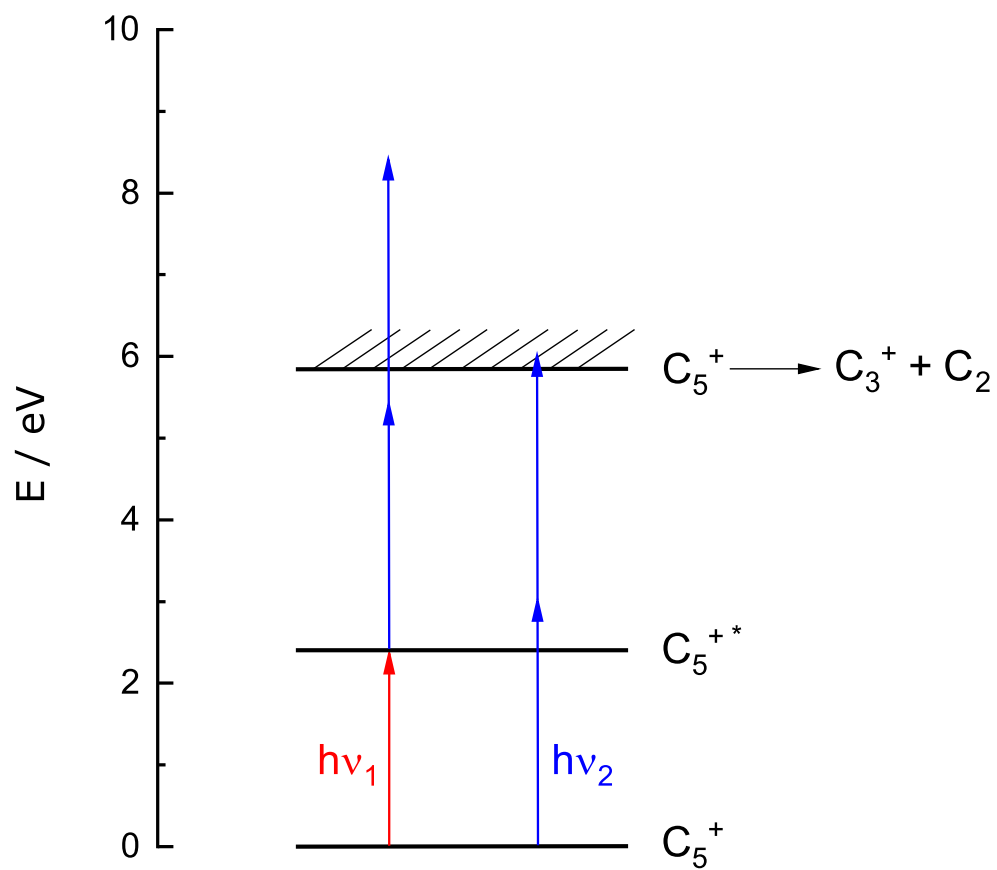




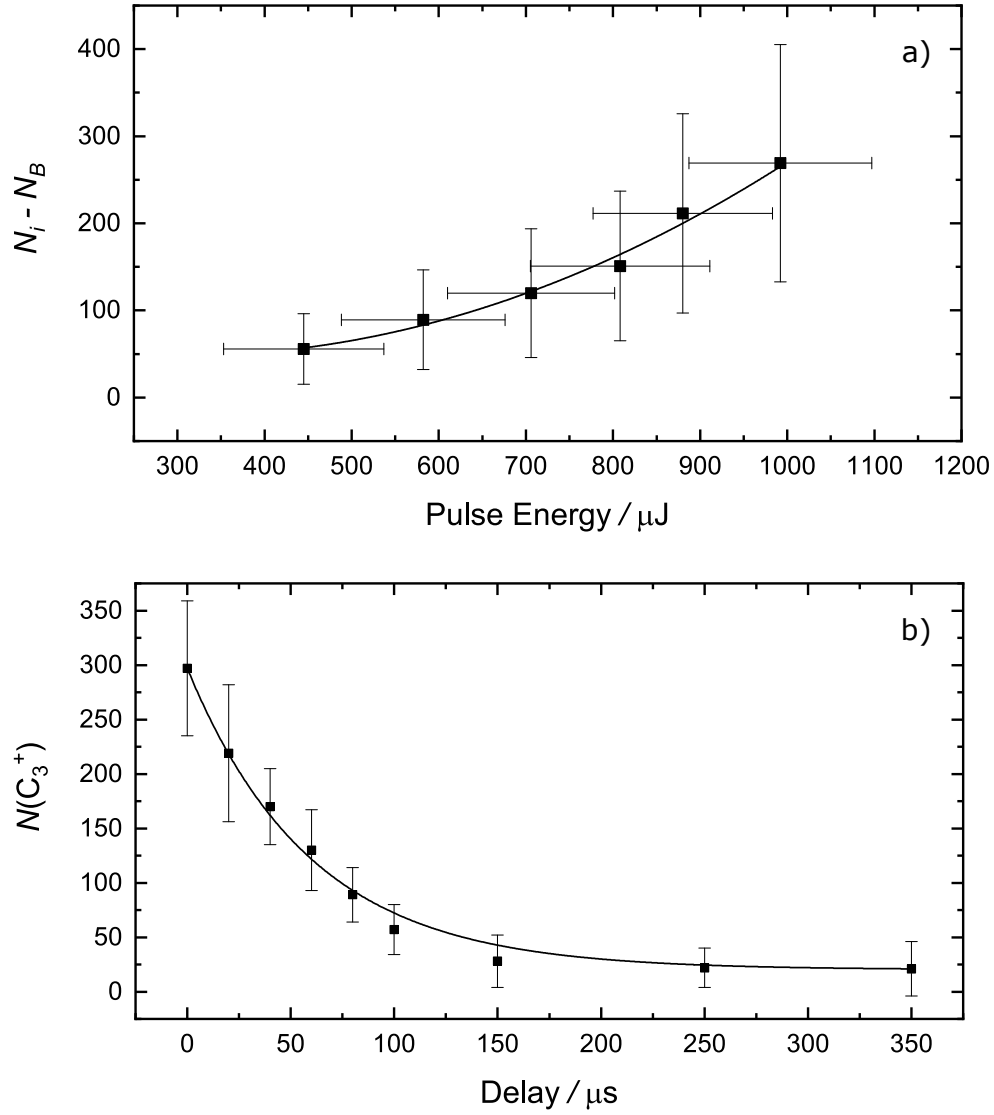
**Figure 4.** (a) Lowest energy portion of  ${}^2\Pi_g \leftarrow X^2\Sigma_u^+$  transition of  $C_5^+ - \text{He}$  recorded using a dye laser. The combs were created with  $\nu_7 = 118 \text{ cm}^{-1}$  and  $\nu_5 = 216 \text{ cm}^{-1}$ . The broader width of the peak near  $19800 \text{ cm}^{-1}$  may indicate more than one unresolved absorption. (b)  ${}^2\Pi_g \leftarrow X^2\Sigma_u^+$  electronic transition of  $C_5^+$  observed by monitoring the production of  $C_3^+$  ions (black) or depletion of  $C_5^+$  ions (blue). Experimental data (circles) have been fit with Lorentzian profiles, the cumulative results are the solid lines.



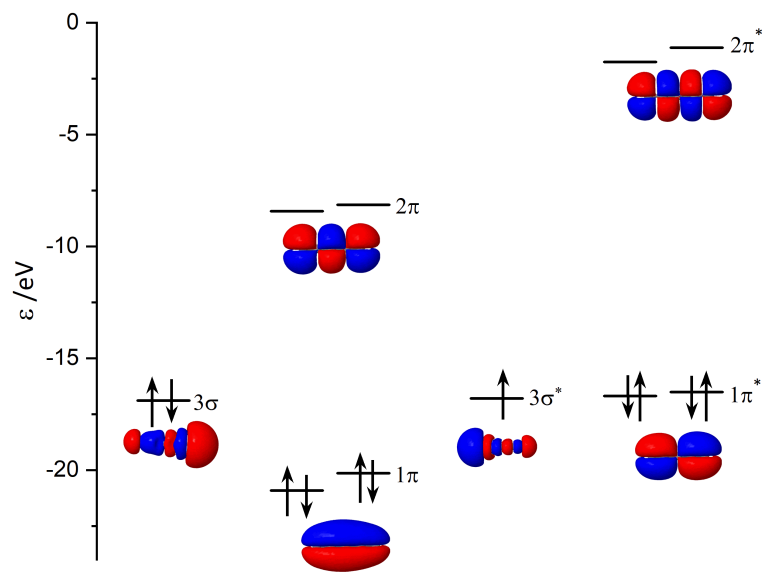
**Figure 5.**  ${}^2\Pi_g \leftarrow X {}^2\Sigma_u^+$  electronic transition of  $C_5^+ - \text{He}_n$  ( $n = 1-3$ ) observed by one-colour fragmentation. The experimental data (circles) have been fit with Lorentzian profiles, the cumulative results are the solid lines.



**Figure 6.**  $C_5^+$  fragmentation scheme. The red arrow (2.4 eV) indicates one photon electronic excitation (pump, dye laser) and the blue arrow (3.1 eV) represents the probe (OPO).



**Figure 7.** (a) Number of  $\text{C}_3^+$  ions produced *via* two colour fragmentation of  $\text{C}_5^+$  as a function of the OPO (probe) pulse energy. Experimental data (squares) have been fit with a second order polynomial (solid line). (b) Number of  $\text{C}_3^+$  ions as a function of the delay between the pump (dye laser) and probe (OPO). An exponential fit (solid line) to typical experimental data (squares) indicates a time constant of around  $60 \mu\text{s}$ .



**Figure 8.** Molecular orbitals obtained at the CASSCF level of theory. See the text for a discussion of the lack of degeneracy of equivalent  $\pi$  and  $\pi^*$  molecular orbitals.