

# Raman and Photoreflectance Study of Cu(In,Ga)S<sub>2</sub> Films and Solar-Cells

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*Abstract:* - The structural and optical properties of  $CuIn_{1-x}Ga_xS_2$  (CIGS), CdS/CuIn\_{1-x}Ga\_xS\_2, and ZnO/CdS/CuIn\_{1-x}Ga\_xS\_2 polycrystalline films, with applications in photovoltaics, were studied by Raman and Photoreflectance (PR) spectroscopy for two different compositions, [Ga]/([In]+[Ga]) =0.04 and 0.12, of the CuIn\_{1-x}Ga\_xS\_2 absorber. The energy band gap of the absorber film was determined by fitting the PR-spectra with a third derivative functional form. Moreover, the thickness of the film was calculated from the interference fringes observed in the PR-spectra below band gap energy. Raman scattering was excited by the 514.5nm line of Ar<sup>+</sup>-laser and the 647.1nm line of Kr<sup>+</sup>-laser. The Raman spectra of the absorber films consist of phonon-modes assigned to CuInS<sub>2</sub>, CuGaS<sub>2</sub> and CuS. The results of the present study are discussed together with the results of SEM and XRD studies of the films and the results of electrical measurements performed on solar cells based on the CIGS absorbers.

*Key-Words:* - CuInS<sub>2</sub>, CuIn<sub>17-x</sub>Ga<sub>x</sub>S<sub>2</sub>, thin film solar cells, Photoreflectance, Raman.

## **1** Introduction

CuInS<sub>2</sub> (CIS) is the most-promising wide-gap chalcopyrite absorber for thin-film solar cells. Its direct band gap  $E_g$  of 1.5 eV is suitable for singlejunction cells. A wider band gap is desirable for the top cell in tandem configurations and can be achieved by incorporating gallium [1]. Cu(In,Ga)S<sub>2</sub> (CIGS)-based solar cells exhibit various benefits such as improved adhesion and open circuit voltage as well as higher flexibility in the choice of process parameters, as reported by several authors [2-5].

It has been shown that the kinetics of film formation in the two-step process lead to an intermediate bi-layer structure of  $CuInS_2$  on  $CuGaS_2$ and that the [Ga]/[In] ratio only partly homogenises by interdiffusion during later stages of sulphurisation [6]. Hence, the gallium concentration is high close to the back contact and decreases towards the film surface. This leads to an inherent grading of the band gap. The two-step process can be modified [7] to give a more uniform distribution of Ga in the active layer. However, the maximum achievable band gap in the active region of the cell is limited to approximately 1.6 eV which corresponds to a x=([Ga]/([In]+[Ga])) ratio of approximately 0.10.

The scope of this paper is to study the properties of  $CuIn_{1-x}Ga_xS_2$ -based films with two different x=[Ga]/([In]+[Ga]) ratios (0.04 and 0.12) and to evaluate the performance of solar cells based on these films. The absorber with x=0.12 is already at the upper limit of the quaternary that can lead to good quality CIGS solar cells with maximum widened energy gap, while the one with x=0.04 is expected to give results very close to the ternary CIS chalcopyrite.

#### 2 Experimental

The investigated samples consist of a soda lime glass coated with sputtered Mo back contact as a substrate, a CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> (CIGS) absorber layer, a CdS buffer layer, a ZnO window and an Al/Ni grid as front contact [7,8]. The  $CuIn_{1-x}Ga_xS_2$  absorber layers are deposited using two different two-stage sequences: for the lower band gap absorber (x=0.04), evaporation takes place with constant rate at constant substrate temperature, while for the absorber with x=0.12, evaporation of Cu (below stoichoiometry), In, Ga, and S at lower substrate temperature (typically 250 °C) is followed by evaporation of Cu and S at higher (typically 500 °C) substrate temperature. The buffer layer is deposited by chemical bath, the window layer is formed by sputtering and the front contact is evaporated [8]. All films were slightly Cu-rich after preparation and the secondary CuS phase formed on the absorber surface was removed by KCN-etching. Further details on the preparation and structural characterization of the films and solar cells are described elsewhere [7].

 $CuInS_2$  polycrystalline films were also prepared using a procedure similar to the one applied in the preparation of the higher band gap quaternary chalcopyrite. The only difference was the absence of Cu in the first stage of the process.

The PR-spectra of the CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> (CIGS) and CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> films were excited at 20K by the 488 nm (or 457.9 nm) line of an argon (Ar<sup>+</sup>) laser as modulation source. The PR-spectra of the ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> films were excited at 20K by the 351 nm line of a Krypton (Kr<sup>+</sup>) laser. A Xenon (or a Halogen) lamp was used as a light source. The reflected beam was spectrally analyzed by a single-grating monochromator (SPEX 1704) and detected by a Si-diode. The detected light was amplified by lock-in techniques. For the low temperature measurements, the samples were cooled down to 20K in a closed-cycle He-cryostat.

The Raman spectra of the CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> (CIGS), CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> polycrystalline films were measured at room temperature using the 514.5nm line of an Ar<sup>+</sup>-laser and the 647.1nm line of a Kr<sup>+</sup>-laser as excitation source. The scattered light was spectrally analyzed by a double grating monochromator and detected by a GaAs-PMT detector.

### **3** Results and discussion

In Fig. 1, the PR-spectra of the  $CuIn_{1-x}Ga_xS_2$  (CIGS) samples (x=0, 0.04, and 0.12) at 20 K are shown. The room temperature PR-spectra could not be evaluated due to their low intensity to noise ratio and they will not be further discussed.

The structure of the fundamental absorption edge of chalcopyrites is well understood [9]. These crystals are semiconductors with a direct fundamental band gap at the Brillouin-zone centre  $\Gamma$ . The threefold valence-band maximum is composed of three nondegenerate states, giving rise to three transitions called Ea, Eb, Ec from smaller to larger energy (Fig. 2). Experimental measurements of these transition energies and their selection rules allow us to determine the symmetry of the three valence states and to calculate the crystal-field ( $\Delta_{cf}$ ) and the spin-orbit ( $\Delta_{so}$ ) parameters, using the quasicubic model [10]. Starting from the threefolddegenerate  $\Gamma_{15}$  state in zinc blende, the introduction of a tetragonal crystal field gives rise to one  $\Gamma_4$  plus a twofold-degenerate  $\Gamma_5$  level. Spin-orbit interaction further splits the  $\Gamma_5$  levels into  $\Gamma_7$  and  $\Gamma_6$  and the selection rules are relaxed.



Fig.1. PR-spectra of  $CuIn_xGa_{1-x}S_2$  (x=0, 0.04, and 0.12).

In  $CuInS_2$  (CIS), with a small distortion from the cubic symmetry, the A and B peaks are so close that they are never resolved at room temperature ( $\Delta_{cf} > 0$ ) [11]. However, Tell et al. [12], by comparing the electro- (ER) and photoreflectance (PR) spectra with reflectance spectra, concluded that the dominating mechanism for CuInS<sub>2</sub> at 2 K is the change in the dielectric functions of two excitonic transitions. Low-temperature studies of exciton reflectivity and photoreflectance have failed to locate a third valence band in any of the studied sulphides, CIS included [13]. The PR spectra of CuInS<sub>2</sub> single crystals below 200 K consist of two well-resolved peaks related to the fundamental transition  $E_0$  and the transition involving spin-orbit splitting  $E_0 + \Delta_0$ [14]. In general, in the low temperature PR spectra, only E<sub>a</sub>- and E<sub>b</sub>- transitions are expected to appear [13]. The two bands observed in the PR spectra of the absorbers in Fig.1 are, thereafter, attributed to the  $E_a$ - and  $E_b$ - excitonic transitions.



Fig.2 Schematic evolution of the energy states at the band gap of chalcopyrite starting from a zinc blende-like gap without spin-orbit coupling.

In the photoreflectance spectra, the modulation of the optical constants is produced by the periodic cancellation of the built-in surface field resulting from the photoexcitation by the laser beam. Basically, the PR spectrum can be described by a third derivative line shape [15] similar to the case of electroreflectance.

The fitting of the PR spectra using a Third Derivative Functional Form (TDFF) gives the positions of the fundamental band gap energies of the ternary and two quaternary polycrystalline absorbers. For the ternary CuInS<sub>2</sub>, the energies of the E<sub>a</sub> and the E<sub>b</sub> bands are: E<sub>a</sub> =(1.533±0.001) eV and E<sub>b</sub> =(1.552±0.001) eV. For the quaternary absorbers CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> with x=0.04 and 0.12, we obtain E<sub>a</sub>=(1.532±0.001) eV, E<sub>b</sub>=(1.580±0.002) eV and E<sub>a</sub>=(1.651±0.002) eV, E<sub>b</sub>=(1.672±0.002) eV,

respectively. The fitted values (Fig. 3) are in good agreement with the linear dependence of the energy band gap on the Ga-fraction reported in experimental PR studies of quaternary chalcopyrites [16].



Fig.3. Compositional dependence of the  $E_a$  and  $E_b$  transition energies.

The PR-spectra of the  $CuIn_{1-x}Ga_xS_2$  layer with the deposited CdS buffer layer on top of the chalcopyrite absorber at 20 K are shown in Fig. 4. Apart from the absorber bands, a band assigned to CdS is observed on the high energy side of these spectra. The band gap energy of the E<sub>a</sub>-transition in CdS was found to be  $E_a=(2.501\pm0.002)$  and (2.487±0.007) eV for x=0.04 and 0.12, respectively, which is in good agreement with the values known from the literature [17]. These spectra exhibit a number of interference fringes at energies below the band gap energy. Similar interference fringes are also present in the reflectance spectra of the  $CuIn_{1-x}Ga_xS_2$  absorbers, as demonstrated in Fig. 5. The fringes are originating from the interference of light beams reflected on the front and back face of the absorber layer. The interferences influence the PR absorber line shapes: the E<sub>a</sub> gap is shadowed, while the  $E_b$  gap is up-shifted to  $E_b=(1.575\pm0.004)$ eV and (1.681±0.008) eV for x=0.04 and 0.12, respectively.

The interference fringes observed in the reflectance spectra (Fig. 5) can be used to calculate the absorber thickness using a relationship derived in ref. [18]:

$$t = \frac{N \lambda_1 \lambda_2}{2(\lambda_1 - \lambda_2)(n^2 - \sin^2 \theta)^{1/2}}$$
(1)

where t is the layer thickness, n is the refractive index of the layer,  $\theta$  is the angle of incidence,  $\lambda_1$ ,  $\lambda_2$  are the start and end wavelength ( $\lambda_1 > \lambda_2$ ), respectively, and N is the number of fringes between  $\lambda_1$  and  $\lambda_2$ .



Fig.4. PR-spectra of CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> (x=0.04 and 0.12).

For the calculation of the absorber thickness, the refractive index of the ternary CuInS<sub>2</sub> is used [19], since the refractive index of the quaternary  $CuIn_{1-x}Ga_xS_2$  is not known. In ref. [20], two different values of n corresponding to the start- and endwavelength of the scanned region were chosen. In the present work, the wavelength dispersion of n has been taken into account by choosing a value in the middle of the scanned spectral region: n=2.87 at 1.5 eV. By adding Ga into CuInS<sub>2</sub>, the refractive index of the quaternary compound is lowered. The refractive index of pure CuGaS<sub>2</sub> is n=2.59 at 1.5 eV [19]. However, addition of 4%-12% Ga does not significantly changes the refractive index of the ternary. In fact, the refractive indices of the quaternary films studied are almost the same as the refractive index of CuInS<sub>2</sub> due to the small gallium content of the films ( $\leq 12\%$ ). The calculated thickness for the non-etched absorbers with x=0.04 and 0.12 are 3.8 and 3.9 µm, respectively. According to the scanning electron micrographs (SEM) in Fig. 6, the films appear to be compact with grain sizes in the order of one µm. Prior to

etching, the film is covered homogenously with CuS, which is important in view of the growth assistance provided by this phase [7].



Fig. 5. Reflectance spectra of  $CuIn_xGa_{1-x}S_2$  (x= 0.04 and 0.12).

The SEM data of the etched absorbers give a film thickness of about 3  $\mu$ m, for both Ga fractions. Using equation (1) for t=3  $\mu$ m, we obtain a refractive index of 3.66 and 3.71 for the absorber films with x=0.04 and 0.12, respectively. This discrepancy between SEM and PR results on the thickness determination is tentatively attributed to In<sub>2</sub>S<sub>3</sub> and GaS secondary phases within the film [7]. CuS segregates on the surface, as stated in the experimentals, and can be removed by KCN-etching. PR-spectroscopy detects only a mean value of the refractive index of the layer (the so called *effective* refractive index).



Fig.6. Scanning electron micrograph of a  $CuIn_{1-x}Ga_xS_2$  film with x=0.04 (cross-section).

The PR-spectrum of the  $ZnO/CdS/CuIn_{1-x}Ga_xS_2$  film with x=0.12 at 20 K is shown in Fig. 7. In the

low energy region, the  $E_a$ -band of the absorber layer is observed, while the narrow band at higher energies is due to the polycrystalline ZnO layer deposited on top of the absorber and the buffer layer. The band gap energy of ZnO is calculated to  $(3.367\pm0.001)$  eV, in agreement with values reported in the literature [21, 22].



Fig.7. Photoreflectance spectrum of  $ZnO/CdS/CuIn_xGa_{1-x}S_2$  with x= 0.12.

In Fig. 8, the Raman spectra of the  $Culn_{1-x}Ga_xS_2$ , CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> samples with x=0.04, excited by two different laser beams, the 514.5nm line of an Ar<sup>+</sup>-laser and the 647.1nm line of a Kr<sup>+</sup>-laser, are shown. The Raman spectra of the CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> samples with x=0.12 are presented in Fig. 9.

The Raman spectra of the absorbers exhibit peaks that can be attributed to the phonon modes of the ternary CuInS<sub>2</sub> and CuGaS<sub>2</sub> chalcopyrites and to the modes of the CuS secondary phase that is formed during the preparation process. The Ramanband frequencies were determined by fitting the Raman spectra with Lorentzians. In Figs. 8(a) and 9(a), the two weak peaks at 262 and 292 cm<sup>-1</sup> are assigned to the E(TO)-mode of CuGaS<sub>2</sub> and to the  $A_1$ -mode of CuInS<sub>2</sub>, respectively. An intense peak at  $300 \text{ cm}^{-1}$ , in the spectra of the CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> layers excited by the 514.5 nm (2.4 eV) laser line, is attributed to the first order  $A_1$  (LO) mode of the CdS buffer layer [23]. The phonon mode at 600 cm<sup>-1</sup>, has twice the frequency of the mode at 300 cm<sup>-1</sup> and is, therefore, assigned to the second order Raman scattering (RS) of CdS. The high intensity of these peaks is due to the resonant Raman scattering of CdS; its energy band

gap of 2.5 eV [17] is close to the excitation energy of 2.4 eV. The mode assignment of the Raman spectra of CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> layers with x=0.04, given in Figs. 8(b) and 9(b), is as follows: the peak at  $63 \text{ cm}^{-1}$ is assigned to the E(TO)-mode of  $CuInS_2$ , the peak at 95 cm<sup>-1</sup> to the B<sub>2</sub>(TO or LO)-mode of CuGaS<sub>2</sub>, the peak at 234 cm<sup>-1</sup> to the  $B_2(TO)$ -mode of CuInS<sub>2</sub>, the peak at 262 cm<sup>-1</sup> to the E(TO)-mode of CuGaS<sub>2</sub>, the peak at 292 cm<sup>-1</sup> to the A<sub>1</sub>-mode of CuInS<sub>2</sub>, the peak at 355 cm<sup>-1</sup> to the E(LO)-mode of CuGaS<sub>2</sub>, the peak at 472 cm<sup>-1</sup> to the A<sub>1</sub>-mode of CuS [24] and the peak at 524 cm<sup>-1</sup> to the second order of the E(TO)-mode of CuGaS<sub>2</sub> [25]. No significant shift of the Raman band frequencies could be observed, as expected by substitution of In by Ga in CuInS<sub>2</sub> and formation of  $CuIn_{1-x}Ga_xS_2$ . This is partially due to the small Ga content of the absorber layers ( $x \le 0.12$ ) and partially due to the formation of a gradient composition layer (for x=0.04) with CuInS<sub>2</sub> on the top,  $CuGaS_2$  at the bottom, and  $CuIn_{1-x}Ga_xS_2$  in between, as has been shown by depth profiling [26]. In fact, previously published results of XRD studies on the same samples [7] have confirmed the presence of a quaternary chalcopyrite phase.



Fig.8. Raman spectra of the CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, and ZnO/CdS/ CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> layers with x=0.12, excited by two different laser beams.



Fig.9. Raman spectra of the  $CuIn_{1-x}Ga_xS_2$ ,  $CdS/CuIn_{1-x}Ga_xS_2$ , and ZnO/CdS/  $CuIn_{1-x}Ga_xS_2$  samples with x=0.04, excited by two different laser beams.

The electrical measurements performed on solar cells based on the CIGS absorbers show that open circuit voltages increase up to 800 mV with increasing absorber band gap. On the other hand, comparison of fill factors and quantum efficiencies of the two different CIGS-solar cells under study reveals [7], [27] that, while the band gap is raised (CIS/x=0, 1.5eV/11.4%, CIGS/x=0.04, 1.53eV/12.3%, CIGS/x=0.12, 1.65eV/10.1%, CGS/x=1, 1.68eV/8.3%), the efficiency drops back. However, the CIGS cells show a better performance than CGS based cells with a comparable band gap, though the latter have been under development for many years.

#### 4 Conclusions

CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> (CIGS), CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub>, and ZnO/CdS/CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> polycrystalline films, with applications in photovoltaics, were studied by Raman and Photoreflectance (PR) spectroscopy for two different compositions (Ga/(In+Ga) =0.04 and 0.12) of the CuIn<sub>1-x</sub>Ga<sub>x</sub>S<sub>2</sub> absorber. The band gap energies of the absorber layers as well as those of

the buffer and the ZnO layer were determined by fitting the PR-spectra. Moreover, the absorber thickness was obtained from the interference fringes of the PR-spectra below band gap energy. The phonon modes observed in the Raman spectra were attributed to the ternary compounds  $CuInS_2$  and  $CuGaS_2$ .

#### Acknowledgements

This project is co-funded by the European Social Fund (75%) and National Resources (25%).

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