

Measurement of the muon charge asymmetry in inclusive $pp \rightarrow W + X$ production at $\sqrt{s} = 7$ TeV and an improved determination of light parton distribution functions

S. Chatrchyan *et al.**

(CMS Collaboration)

(Received 21 December 2013; published 13 August 2014)

Measurements of the muon charge asymmetry in inclusive $pp \rightarrow W + X$ production at $\sqrt{s} = 7$ TeV are presented. The data sample corresponds to an integrated luminosity of 4.7 fb^{-1} recorded with the CMS detector at the LHC. With a sample of more than 20 million $W \rightarrow \mu\nu$ events, the statistical precision is greatly improved in comparison to previous measurements. These new results provide additional constraints on the parton distribution functions of the proton in the range of the Bjorken scaling variable x from 10^{-3} to 10^{-1} . These measurements and the recent CMS measurement of associated $W +$ charm production are used together with the cross sections for inclusive deep inelastic $e^\pm p$ scattering at HERA in a next-to-leading-order QCD analysis. The determination of the valence quark distributions is improved, and the strange-quark distribution is probed directly through the leading-order process $g + s \rightarrow W + c$ in proton-proton collisions at the LHC.

DOI: [10.1103/PhysRevD.90.032004](https://doi.org/10.1103/PhysRevD.90.032004)

PACS numbers: 13.38.-b, 13.85.Qk, 12.38.-t

I. INTRODUCTION

In the standard model (SM), the dominant processes for inclusive W -boson production in pp collisions are annihilation processes: $u\bar{d} \rightarrow W^+$ and $d\bar{u} \rightarrow W^-$ involving a valence quark from one proton and a sea antiquark from the other. Since there are two valence u quarks and one valence d quark in the proton, W^+ bosons are produced more often than W^- bosons. The Compact Muon Solenoid (CMS) experiment at the Large Hadron Collider (LHC) has investigated this production asymmetry in inclusive W -boson production and measured the inclusive ratio of total cross sections for W^+ and W^- boson production at $\sqrt{s} = 7$ TeV to be $1.421 \pm 0.006(\text{stat}) \pm 0.032(\text{syst})$ [1]. This result is in agreement with SM predictions based on various parton distribution functions (PDFs) such as the MSTW2008 and CT10 PDF sets [2,3]. Measurements of the production asymmetry between W^+ and W^- bosons as a function of boson rapidity can provide additional constraints on the d/u ratio and on the sea antiquark densities in the proton. For pp collisions at $\sqrt{s} = 7$ TeV, these measurements explore the PDFs for the proton for Bjorken x from 10^{-3} to 10^{-1} [4]. However, it is difficult to measure the boson rapidity production asymmetry because of the energy carried away by neutrinos in leptonic W -boson decays. A quantity more directly accessible experimentally is the lepton charge asymmetry, defined as

$$\mathcal{A}(\eta) = \frac{\frac{d\sigma}{d\eta}(W^+ \rightarrow \ell^+\nu) - \frac{d\sigma}{d\eta}(W^- \rightarrow \ell^-\bar{\nu})}{\frac{d\sigma}{d\eta}(W^+ \rightarrow \ell^+\nu) + \frac{d\sigma}{d\eta}(W^- \rightarrow \ell^-\bar{\nu})}, \quad (1)$$

where $d\sigma/d\eta$ is the differential cross section for W -boson production and subsequent leptonic decay and $\eta = -\ln[\tan(\theta/2)]$ is the charged lepton pseudorapidity in the laboratory frame, with θ being the polar angle measured with respect to the beam axis.

High precision measurements of the W -boson lepton charge asymmetry can improve the determination of the PDFs. Both the W -boson lepton charge asymmetry and the W -boson production charge asymmetry were studied in $p\bar{p}$ collisions by the CDF and D0 experiments at the Tevatron collider [5–7]. The ATLAS, CMS, and LHCb experiments also reported measurements of the lepton charge asymmetry using data collected at the LHC in 2010 [8–11]. An earlier measurement of the W -boson electron charge asymmetry is based on 2011 CMS data corresponding to an integrated luminosity of 0.84 fb^{-1} [12].

The impact of CMS measurements of the lepton charge asymmetry on the global PDF fits has been studied by several groups [13–17], who concluded that improvements in the PDF uncertainties for several quark flavors could be achieved with more precise data. In this paper, we report a measurement of the muon charge asymmetry using a data sample corresponding to an integrated luminosity of 4.7 fb^{-1} collected with the CMS detector at the LHC in 2011. The number of $W \rightarrow \mu\nu$ events (more than 20 million) in this data sample is 2 orders of magnitude larger than for our previous measurement [10].

This precise measurement of the muon charge asymmetry and the recent CMS measurement of associated $W +$ charm production [18] are combined with the cross sections

* Full author list given at the end of the article.

Published by the American Physical Society under the terms of the [Creative Commons Attribution 3.0 License](https://creativecommons.org/licenses/by/3.0/). Further distribution of this work must maintain attribution to the author(s) and the published articles title, journal citation, and DOI.

for inclusive deep inelastic $e^\pm p$ scattering at HERA [19] in a quantum chromodynamics (QCD) analysis at next-to-leading order (NLO). The impact of these measurements of W -boson production at CMS on the determination of light-quark distributions in the proton is studied and the strange-quark density is determined.

This paper is organized as follows. A brief description of the CMS detector is given in Sec. II. The selection of $W \rightarrow \mu\nu$ candidates is described in Sec. III. The corrections for residual charge-specific bias in the measurement of the muon transverse momentum (p_T) and in the muon trigger, reconstruction, and selection efficiencies are discussed in Sec. IV. The extraction of the $W \rightarrow \mu\nu$ signal is described in detail in Sec. V. Systematic uncertainties and the full correlation matrix are given in Sec. VI. The final measurements are presented in Sec. VII, and the QCD analysis is discussed in detail in Sec. VIII. The summary and conclusion follow in Sec. IX.

II. THE CMS EXPERIMENT

The central feature of the CMS apparatus is a superconducting solenoid 6 m in diameter and 13 m long, which provides an axial field of 3.8 T. Within the field volume are a silicon pixel and strip tracker, a crystal electromagnetic calorimeter (ECAL), and a brass/scintillator hadron calorimeter. Muons are measured in gas-ionization detectors embedded in the steel flux return yoke. The ECAL consists of nearly 76000 lead tungstate crystals that provide coverage in pseudorapidity $|\eta| < 1.479$ in the barrel region and $1.479 < |\eta| < 3.0$ in the two end cap regions. A preshower detector consisting of two planes of silicon sensors interleaved with a total of three radiation lengths of lead is located in front of the ECAL end caps. Muons are selected in the pseudorapidity range $|\eta| < 2.4$, with detection planes constructed of drift tubes (DT), cathode strip chambers (CSC), and resistive plate chambers, and matched to the tracks measured in the silicon tracker resulting in an η -dependent p_T resolution of about 1%–5% for muon p_T up to 1 TeV. The inner tracker, consisting of 1440 silicon pixel and 15148 silicon strip detector modules, measures charged particles within the pseudorapidity range $|\eta| < 2.5$. It provides an impact parameter resolution of $\sim 15 \mu\text{m}$ and a p_T resolution of about 1.5% for 100 GeV particles.

The CMS experiment uses a right-handed coordinate system, with the origin at the nominal interaction point, the x axis pointing toward the center of the LHC, the y axis pointing up (perpendicular to the LHC plane), and the z axis along the counterclockwise-beam direction. The polar angle, θ , is measured from the positive z axis and the azimuthal angle, ϕ , is measured in the x - y plane.

A detailed description of the CMS experiment can be found in Ref. [20].

III. EVENT RECONSTRUCTION

The signature of a $W \rightarrow \mu\nu$ event is a high- p_T muon accompanied by missing transverse momentum \vec{E}_T due to the escaping neutrino. The CMS experiment has utilized a particle-flow algorithm in event reconstruction, and the \vec{E}_T used by this analysis is determined as the negative vector sum of the transverse momenta of all particles reconstructed by this algorithm [21]. The $W \rightarrow \mu\nu$ candidates were collected with a set of isolated single-muon triggers with different p_T thresholds, which is the major difference with respect to the previous CMS measurement where nonisolated single-muon triggers were used [10]. The isolated muon trigger requires that in the neighboring region of the muon trigger candidate both the transverse energy deposits in calorimeters and the scalar sum of the p_T of the reconstructed tracks are small, and it reduces the trigger rate while maintaining a relatively low muon p_T threshold. We use all the data-taking periods during which the isolated muon triggers were not prescaled (i.e. they were exposed to the full integrated luminosity).

Other physics processes can produce high- p_T muons and mimic $W \rightarrow \mu\nu$ signal candidates. We consider the SM background contributions from multijet production (QCD background), Drell-Yan ($Z/\gamma^* \rightarrow \ell^+\ell^-$) production, $W \rightarrow \tau\nu$ production [electroweak (EW) background], and top-quark pair ($t\bar{t}$) production. In addition, cosmic-ray muons can penetrate through the center of the CMS detector and also mimic $W \rightarrow \mu\nu$ candidates.

Monte Carlo (MC) simulations are used to help evaluate the background contributions in the data sample and to study systematic uncertainties. Primarily, we use NLO MC simulations based on the POWHEG event generator [22] where the NLO CT10 PDF model [3] is used. The generated events are interfaced with the PYTHIA (v.6.422) event generator [23] for simulating the electromagnetic finite-state radiation (FSR) and the parton showering. The τ lepton decay in the $W \rightarrow \tau\nu$ process is simulated by the TAUOLA MC package [24]. We simulate the QCD background with the PYTHIA event generator where the CTEQ6L PDF model [25] is used. The CMS detector simulation is interfaced with GEANT4 [26]. All generated events are first passed through the detector simulation and then reconstructed in the same way as the collision data. Pileup is the presence of multiple interactions recorded in the same event. For the data used in this analysis, there are an average of about 7 reconstructed primary interaction vertices for each beam crossing. The MC simulation is generated with a different pileup distribution than we observe in the data. Therefore, the MC simulation is weighted such that the mean number of interactions per crossing matches that in data, using the inelastic pp cross section measured by the CMS experiment [27].

The selection criteria for muon reconstruction and identification are described in detail in a previous report [28]. Therefore, only a brief summary is given here. A muon candidate is reconstructed using two different algorithms:

one starts with a track measured by the silicon tracker and then requires a minimum number of matching hits in the muon detectors, and the other starts by finding a track in the muon system and then matching it to a track measured by the silicon tracker. Muons used in this measurement are required to be reconstructed by both algorithms. A global track fit, including both the silicon tracker hits and muon chamber hits, is performed to improve the quality of the reconstructed muon candidate. The track p_T measured by the silicon tracker is used as the muon p_T and the muon charge is identified from the signed curvature. Cosmic-ray contamination is reduced by requiring that the distance of the closest approach to the leading primary vertex is small: $|d_{xy}| < 0.2$ cm. The remaining cosmic-ray background yield is estimated to be about 10^{-5} of the expected $W \rightarrow \mu\nu$ signal and is therefore neglected [10]. The track-based muon isolation, $\text{Iso}_{\text{track}}$, is defined to be the scalar sum of the p_T of additional tracks in a cone with a radius of 0.3 around the muon candidate [$R = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} < 0.3$, with $\Delta\phi$ and $\Delta\eta$ being the differences between the muon candidate and the track in the η - ϕ plane]. Muons are required to have $\text{Iso}_{\text{track}}/p_T < 0.1$. Only muons within $|\eta| < 2.4$ are included in the data sample.

In each event, muons passing the above selection criteria are ordered according to p_T , and the leading muon is selected as the $W \rightarrow \mu\nu$ candidate. The leading muon is required to be the particle that triggered the event. In addition, the muon is required to have $p_T > 25$ GeV, which is safely above the trigger turn-on thresholds. Events that have a second muon with $p_T > 15$ GeV are rejected to reduce the background from Drell-Yan dimuon events (“Drell-Yan veto”). The rejected events, predominantly $Z/\gamma^* \rightarrow \mu^+\mu^-$ events, are used as a Drell-Yan control sample to study the modeling of the \vec{E}_T and also to provide constraints on the modeling of the p_T spectrum of W and Z bosons. In addition, this sample is used to estimate the level of background from Drell-Yan events where the second muon is not identified. The muon is corrected for a bias in the measurement of the momentum (discussed below) prior to the application of the p_T selection.

The $W \rightarrow \mu\nu$ candidates that pass the above selection criteria are divided into 11 bins in absolute value of muon pseudorapidity $|\eta|$. The bin width is 0.2, except that the last three $|\eta|$ bins are [1.6–1.85], [1.85–2.1], and [2.1–2.4], respectively. The muon charge asymmetry is measured in each of the $|\eta|$ bins, along with the determination of the correlation matrix of the systematic uncertainties between different $|\eta|$ bins.

IV. MUON MOMENTUM CORRECTION AND EFFICIENCY STUDIES

The measured momentum of the muon depends critically on the correct alignment of the tracker system and the details of the magnetic field. Even after the alignment of the

tracker detector a residual misalignment remains, which is not perfectly reproduced in the MC simulation. This misalignment leads to a charge dependent bias in the reconstruction of muon momenta, which is removed by using a muon momentum correction. The detailed description of the method for the extraction of the correction factors using $Z/\gamma^* \rightarrow \mu^+\mu^-$ events is given in Ref. [29]. Here we provide only a short summary of the method. First, corrections to muon momentum in bins of η and ϕ are extracted separately for positively and negatively charged muons using the average of the $1/p_T$ spectra of muons in $Z/\gamma^* \rightarrow \mu^+\mu^-$ events. The mean values of the $1/p_T$ spectra at the MC generator level, varied by the reconstruction resolution, are used as a “reference.” The mean values of the reconstructed $1/p_T$ spectra in data or simulation are tuned to match the reference. Second, the correction factors derived in the previous step are tuned further by comparing the dimuon invariant mass in each bin of muon charge Q , η , and ϕ to the ones at the MC generator level varied by the reconstruction resolution. The same procedure is performed for both data and reconstructed MC events, and correction factors are determined separately. The correction factors are extracted using the same η binning defined above in order to avoid correlations between different η bins.

The data set used to derive the corrections was collected with a double-muon trigger with asymmetric p_T thresholds of 17 and 8 GeV. Both muons are required to have $p_T > 25$ GeV, which exceeds significantly the trigger p_T thresholds. The simulation has been corrected for the muon efficiency difference between data and MC simulation as discussed below. We illustrate the relative size of the derived corrections using a 40 GeV muon as an example. For muons within $|\eta| < 0.2$, the corrections derived using the $1/p_T$ spectra are less than 1.5% and 0.4% for data and MC simulation, respectively. A ϕ modulation of these corrections is observed. The maximum corrections are larger in high- η region, and for muons with $|\eta| > 2.1$ these corrections can be as large as 3.5% and 1.4% for data and MC simulation, respectively. The additional corrections derived using the dimuon invariant mass are smaller. For muons within the complete detector acceptance, the additional corrections are less than 0.5% and 0.2% for data and simulation, respectively. These additional corrections show no evidence of η - ϕ dependence and fluctuate around zero within the statistical uncertainties of the final corrections. The statistical uncertainties of the corrections for various η - ϕ bins are uncorrelated. By comparing the correction factors for positively and negatively charged muons in each bin, we can determine relative corrections from misalignment and from mismodeling of the magnetic field in the tracker system. The mismodeling corrections for muons with $|\eta| > 2.1$, where maximum deviations from zero are evident, are less than 0.3% and 0.4% for data and MC simulation, respectively. In contrast, in the same detector region, the corrections due to misalignment are about 4.4%

and 1.7% for data and MC simulation, respectively. Hence, the bias comes predominantly from misalignment.

Figure 1 shows the average dimuon invariant mass (mass profile) as a function of muon Q and η before and after the correction, which includes both the contributions from tracker misalignment and mismodeling of the magnetic field. The dimuon mass profiles after the correction are compared to the reference mass profile for data and MC simulation. They agree well with the reference, so the muon momentum bias is largely removed. The reference mass profile is expected to be a function of η because of the p_T requirements for the two daughter muons in $Z/\gamma^* \rightarrow \mu^+\mu^-$ decays. Values of the dimuon mass profile as a function of muon η are averaged over ϕ , while the muon scale corrections correct for muon momentum bias in both η and ϕ .

The overall efficiency in the selection of muon candidates includes contributions from reconstruction,

identification (including isolation), and trigger efficiencies. The muon reconstruction efficiency includes contributions from the reconstruction efficiency in the tracker system (“tracking”) and in the muon system. The muon “off-line” efficiency is the product of reconstruction and identification efficiencies. The contribution of each component to the overall efficiency (tracking, muon stand-alone reconstruction, identification, and trigger) is measured directly from the $Z/\gamma^* \rightarrow \mu^+\mu^-$ events using the tag-and-probe method [1,28]. In this method one of the daughter muons is used to tag the $Z/\gamma^* \rightarrow \mu^+\mu^-$ event and the other muon candidate is used as a probe to study the muon efficiencies as a function of Q , η , and p_T . For every event a positively charged muon can be selected as the tag and a negatively charged probe candidate is used to study the efficiencies for negatively charged muons. The same procedure is repeated by selecting a negatively charged

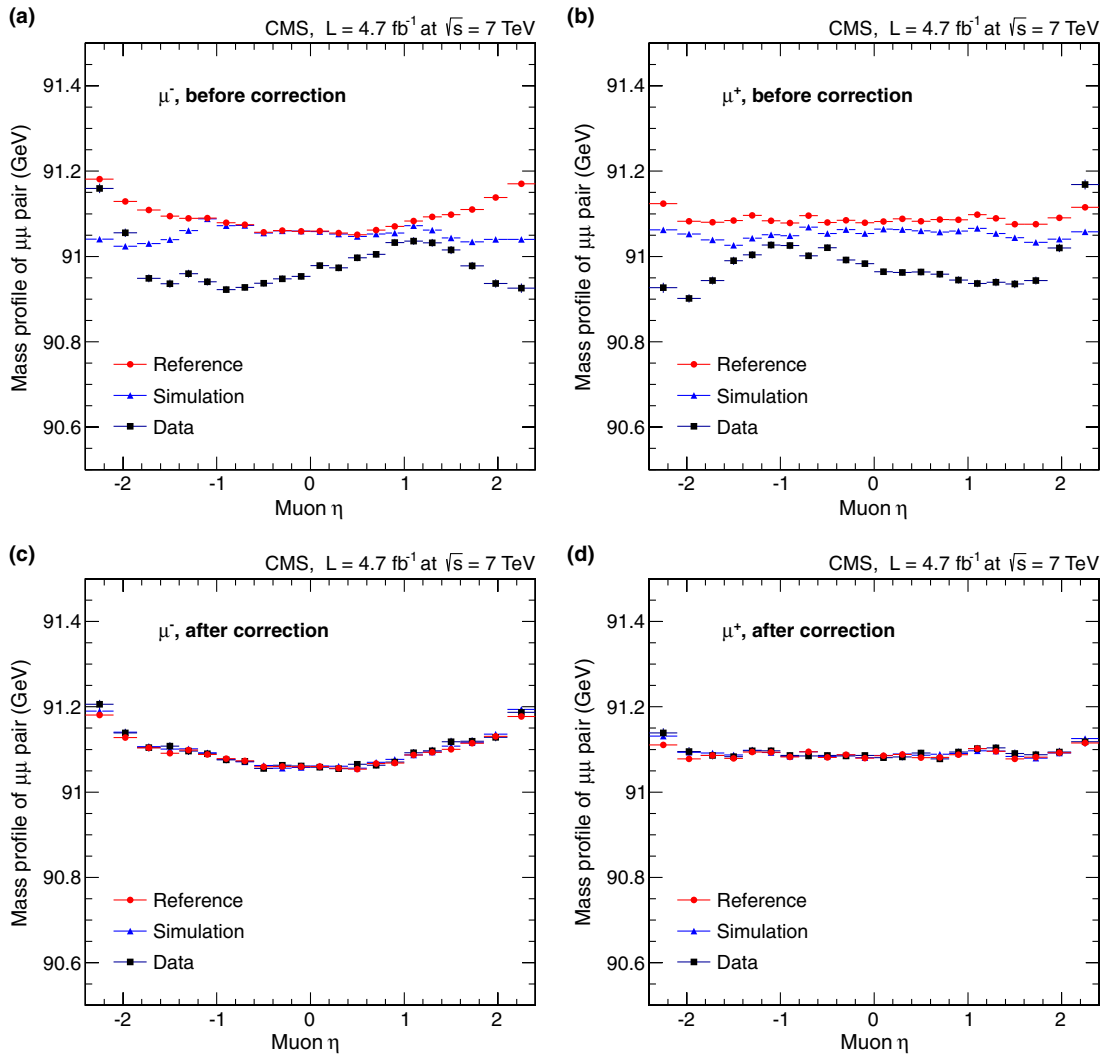


FIG. 1 (color online). The dimuon mass profile as a function of muon η for μ^- (a), (c) and μ^+ (b), (d), where (a) and (b) are before the correction and (c) and (d) are after the correction. The generated muon p_T varied by reconstruction resolution in data is used to obtain the dimuon invariant mass of the reference.

muon as the tag to study efficiencies for positively charged muons. Each individual efficiency is determined in 22 bins of muon η , as defined above, and 7 bins of p_T (15–20, 20–25, 25–30, 30–35, 35–40, 40–45, and > 45 GeV) for both μ^+ and μ^- . The same procedure is applied to both data and MC simulation and scale factors are determined to match the simulation efficiencies to the data.

The measured average tracking efficiency in each η bin varies from 99.6% to 99.9% with a slight inefficiency in the transition regions from the barrel to the end cap segments and at the edge of the tracker system. The ratio of tracking efficiencies for μ^+ and μ^- is consistent with unity within statistical uncertainty. In the transition regions from the DT to the CSC, there is evidence that the muon off-line efficiency has a slight asymmetry between μ^+ and μ^- .

The ratio of efficiencies for positively and negatively charged muons differs from unity by up to $1.0 \pm 0.3\%$. The trigger efficiency ratio is also found to differ from unity in some η regions. The maximum deviation is at $\eta > 2.1$ where the efficiency for μ^+ is about $2.0 \pm 0.5\%$ higher than that for μ^- . Figure 2 shows the η distribution for the leading μ^+ and μ^- in the $Z/\gamma^* \rightarrow \mu^+\mu^-$ sample. The dimuon invariant mass is within $60 < m_{\mu\mu} < 120$ GeV. Here, the MC simulation is corrected for muon momentum bias, efficiency, and modeling of the Z-boson transverse momentum (\vec{q}_T) before normalizing to the measured data. The modeling of Z-boson \vec{q}_T spectrum is discussed in detail in Sec. VID 4. The η dependence effect in data and MC simulation are in good agreement.

V. EXTRACTION OF THE ASYMMETRY

The asymmetry is calculated in bins of $|\eta|$ from the yields of W^+ and W^- . In this section, we explain how the yields are obtained from the E_T distributions, and we discuss corrections to the E_T needed in the accurate estimation of the yields. Finally, we explain how backgrounds are taken into account.

The raw charged asymmetry (A^{raw}) is defined in terms of the numbers N^{W^+} and N^{W^-} of W^+ and W^- signal events,

$$A^{\text{raw}} = \frac{N^{W^+} - N^{W^-}}{N^{W^+} + N^{W^-}}. \quad (2)$$

The yields N^{W^+} and N^{W^-} are obtained from simultaneous binned maximum-likelihood fits of the E_T distributions; the signal yields and the normalization of the QCD background are free parameters. The likelihood is constructed following the Barlow-Beeston method [30] to take into account the limited size of the MC signal event sample. The shapes of the E_T distributions for the $W \rightarrow \mu\nu$ signal and the background contributions are taken from MC simulations after correcting for mismodeling of the detector response and the \vec{q}_T distribution of the W bosons, as discussed further in Sec. VID 4 below. The pileup of each MC sample is matched to the data using an “accept-reject” technique based on the observed and simulated pileup distributions. This technique avoids a large spread of weights that would come from simply reweighting the MC events; the E_T templates are constructed using the accepted MC events.

A total of 12.9 million $W^+ \rightarrow \mu^+\nu$ and 9.1 million $W^- \rightarrow \mu^-\bar{\nu}$ candidate events are selected. The expected backgrounds from QCD, EW, and $t\bar{t}$ events are about 8%, 8%, and 0.5%, respectively. The single top-quark and diboson ($WW/WZ/ZZ$) production is less than 0.1% and is neglected. The variation of the background composition as a function of $|\eta|$ is taken into account.

The estimate for the Drell-Yan background is based on the observed yields in a Drell-Yan control sample. The $W \rightarrow \tau\nu$ background scales with the $W \rightarrow \mu\nu$ signal using a factor determined from a MC simulation. The $t\bar{t}$

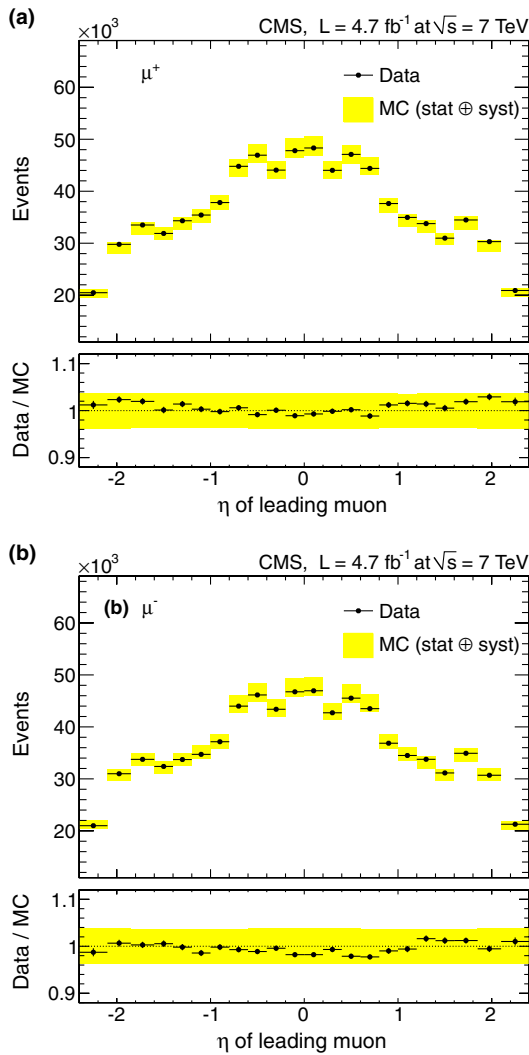


FIG. 2 (color online). The η distribution of the leading μ^+ (a) and μ^- (b) in the $Z/\gamma^* \rightarrow \mu^+\mu^-$ sample. The dimuon invariant mass is within $60 < m_{\mu\mu} < 120$ GeV. The MC simulation is normalized to the data. The light shaded band is the total uncertainty in predicting the $Z/\gamma^* \rightarrow \mu^+\mu^-$ event yields using MC simulation, as described in Sec. VI.

background is normalized to the NLO cross section obtained from MCFM [31–33]. Efficiency correction factors are applied to the simulation before determining the background normalization.

The level of the QCD background is determined by the fit. A constraint on the relative amount of QCD background in the W^+ and W^- samples is obtained from a QCD-enriched control sample collected using a muon trigger with no isolation requirement. This constraint induces a correlation of N^{W^+} and N^{W^-} , and the resulting covariance is taken into account when evaluating the statistical uncertainty on \mathcal{A}^{raw} .

In the following sections, we discuss the corrections to the E_T and then report the results of the fit to the E_T distributions.

A. Corrections of the missing energy measurement

The analysis depends critically on the control of the E_T distributions. Several corrections are needed to bring the simulation into agreement with the observed distributions. The E_T depends on both the measured muon kinematics and the kinematics of the hadrons recoiling against the W boson. The corrections for the calibration of the muon momentum, discussed in Sec. IV above, are applied by adding the \vec{p}_T correction to \vec{E}_T vectorially. The kinematic corrections for the so-called ‘‘hadronic recoil,’’ which are based on the control sample of $Z/\gamma^* \rightarrow \mu^+\mu^-$ events, are explained in detail here.

By definition, the hadronic recoil, \vec{u} , is the vector sum of transverse momenta of all reconstructed particles except for the muon(s). For $Z/\gamma^* \rightarrow \mu^+\mu^-$ events,

$$\vec{u} = -\vec{E}_T - \vec{q}_T, \quad (3)$$

where \vec{q}_T is the transverse momentum of the dimuon system and $\vec{E}_T \approx 0$. The components of \vec{u} parallel and perpendicular to \vec{q}_T are u_{\parallel} and u_{\perp} , respectively. The mean of u_{\perp} , $\langle u_{\perp} \rangle$, is approximately zero, while the mean of u_{\parallel} , $\langle u_{\parallel} \rangle$, is close to the mean of the boson q_T . Differences in the distributions from data and MC are ascribed to detector effects, the simulation of jets, pileup and the underlying event, all of which should be nearly the same for $Z/\gamma^* \rightarrow \mu^+\mu^-$ and $W \rightarrow \mu\nu$ events. The distributions of u_{\parallel} and u_{\perp} in $Z/\gamma^* \rightarrow \mu^+\mu^-$ events are used to derive corrections for the simulation that improve the modeling of E_T for $W \rightarrow \mu\nu$ signal events as well as for backgrounds; this technique was employed previously by the Tevatron experiments and by CMS [34–36]. We correct both the scale and resolution of E_T .

A comparison of the \vec{E}_T distributions for $Z/\gamma^* \rightarrow \mu^+\mu^-$ events in data and MC shows that the agreement is not perfect. Both show a small ϕ modulation, but the phase and amplitude of the modulation are not the same. This modulation follows from the fact that collisions, including hard interactions that produce W events as well as pileup

events, do not occur exactly at the origin of the coordinate system. This modulation can be characterized by a cosine function, $C \cos(\phi - \phi_0)$. The dependence of the amplitude C and phase term ϕ_0 on the number of primary vertices is extracted from the $Z/\gamma^* \rightarrow \mu^+\mu^-$ event sample by fitting a ϕ -dependent profile of $u_{\parallel} - \langle u_{\parallel} \rangle$. The amplitude C is observed to depend linearly on the number of primary vertices, while the phase ϕ_0 is almost independent of pileup. The ϕ modulation of \vec{E}_T can be removed by adding a vector in the transverse plane, $\Delta\vec{E}_T = C \cos \phi_0 \hat{x} + C \sin \phi_0 \hat{y}$ to \vec{E}_T for each event.

The dependence of $\langle u_{\parallel} \rangle$ with Z -boson q_T should be approximately linear, and this behavior is indeed observed in both data and MC. This dependence is further studied according to the direction of the leading jet, namely, in four bins of jet $|\eta|$: [0.0–1.2], [1.2–2.4], [2.4–3.0], and [3.0–5.0]. The jets are formed by clustering particle-flow candidates using the anti- k_T jet clustering algorithm [37] with a distance parameter of 0.5, and the muons are not included in the reconstruction of jets. The $\langle u_{\parallel} \rangle$ behavior with q_T for MC and data agrees very well when the leading jet is in the central region of the detector. When the leading jet is in the forward direction, a modest difference is observed, amounting to less than 10% in the highest $|\eta|$ bin.

The distributions of $u_{\parallel} - \langle u_{\parallel} \rangle$ and u_{\perp} are fit to Gaussian functions whose widths are parametrized as a functions of q_T . They depend strongly on the pileup, so they are also fit as functions of the number of vertices in the event. The weak dependence of $\langle u_{\parallel} \rangle$ on the leading jet $|\eta|$ is neglected. The widths of the $u_{\parallel} - \langle u_{\parallel} \rangle$ and u_{\perp} distributions are slightly larger in data than in MC. For example, when there are seven reconstructed vertices in the event (which corresponds to the mean number for this data set), the widths are 4%–10% larger.

A test of the hadronic recoil corrections is carried out with $Z/\gamma^* \rightarrow \mu^+\mu^-$ events. The hadronic recoil \vec{u} is calculated in each MC event, and the parallel component u_{\parallel} is rescaled by the ratio of $\langle u_{\parallel} \rangle$ in data and in MC. Furthermore, the smearing of u_{\parallel} and u_{\perp} is adjusted to match the resolutions measured with the data. The \vec{E}_T is recalculated according to Eq. (3). Figure 3 shows E_T and the ϕ of the \vec{E}_T after applying the hadronic recoil corrections. The data and MC simulation are in excellent agreement, demonstrating that this empirical correction to E_T works very well for $Z/\gamma^* \rightarrow \mu^+\mu^-$ events.

To apply the hadronic recoil correction determined in $Z/\gamma^* \rightarrow \mu^+\mu^-$ events to other MC simulations, such as $W \rightarrow \mu\nu$ events, requires defining a variable equivalent to the boson \vec{q}_T in $Z/\gamma^* \rightarrow \mu^+\mu^-$ events. In $W \rightarrow \mu\nu$ events, the hadronic recoil is defined to be

$$\vec{u} = -\vec{E}_T - \vec{p}_T, \quad (4)$$

where \vec{p}_T is the muon transverse momentum. The hadronic recoil is decomposed into u_{\parallel} and u_{\perp} components relative to

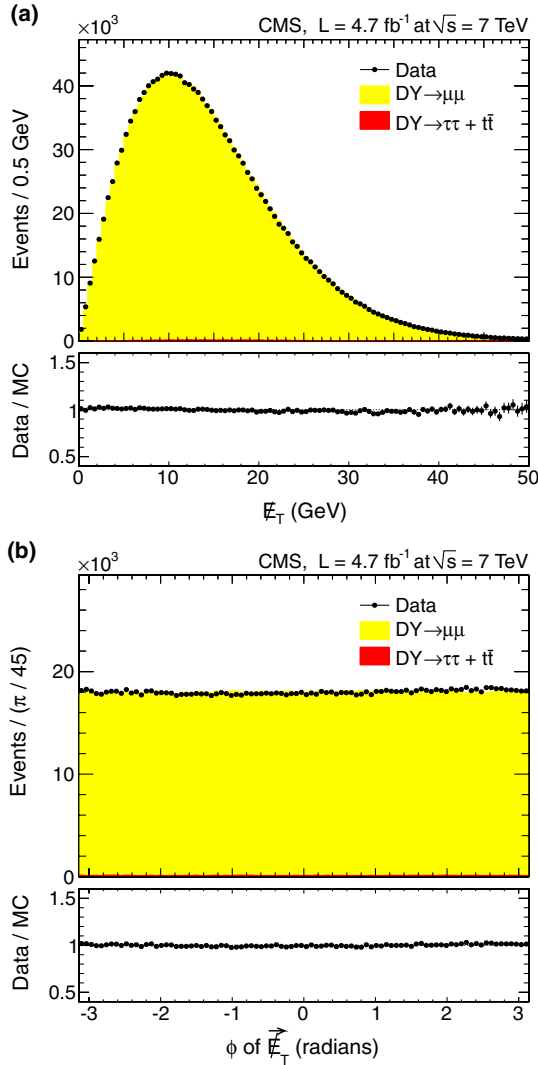


FIG. 3 (color online). Data to simulation comparison for E_T (a) and ϕ distribution of \vec{E}_T (b) in the Drell-Yan control sample. Here, the hadronic recoil derived from the data was used to correct the MC simulation. The $Z/\gamma^* \rightarrow \tau^+\tau^- + t\bar{t}$ contribution (dark shaded region) in data is normalized to the integrated luminosity of the data sample using a MC simulation, and the normalization of the $Z/\gamma^* \rightarrow \mu^+\mu^-$ MC simulation (light shaded region) is taken as the difference between the data and the estimated $Z/\gamma^* \rightarrow \tau^+\tau^- + t\bar{t}$ contribution. In this data sample, the $Z/\gamma^* \rightarrow \tau^+\tau^- + t\bar{t}$ contribution is negligible.

\vec{q}_T . The hadronic recoil correction is applied in the manner above, and \vec{E}_T is recalculated. For the $W \rightarrow \mu\nu$ signal events, \vec{q}_T is the vector sum of the transverse momentum of the reconstructed muon and the generated neutrino. For $W \rightarrow \tau\nu$ events, the generated W -boson \vec{q}_T is used. For selected Drell-Yan background events, one muon is not reconstructed or not identified, so \vec{q}_T is calculated using the \vec{p}_T of the lost muon at the generator level. For the QCD background events, \vec{q}_T is identified with the \vec{p}_T of the reconstructed muon.

Figure 4 shows the E_T distribution for the QCD control sample. We have selected only those events that pass a nonisolated muon trigger but that fail the isolated muon trigger. We also impose an anti-isolation selection cut: $\text{Iso}_{\text{track}}/p_T > 0.1$. With the application of the hadronic recoil corrections, the data and simulation are in very good agreement.

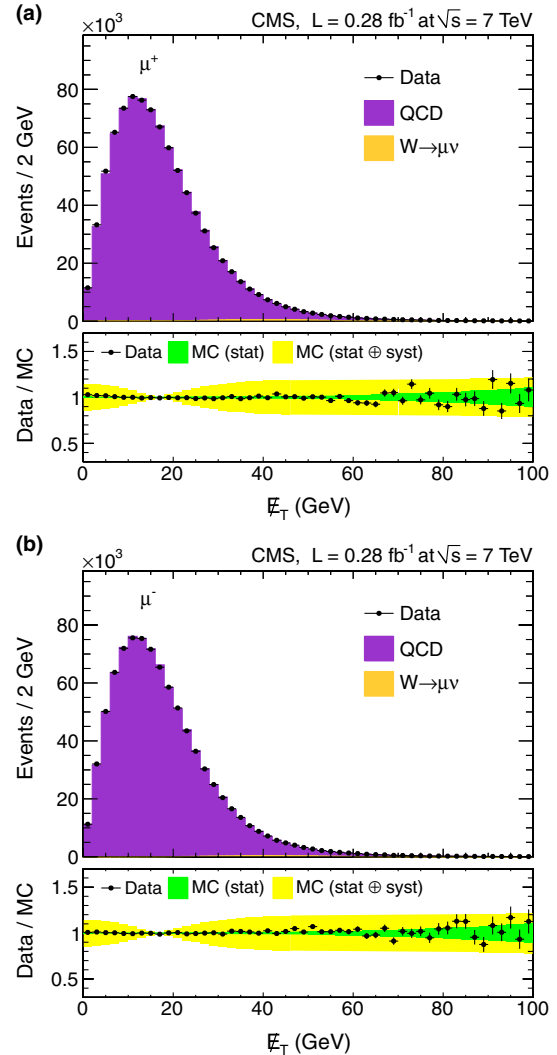


FIG. 4 (color online). The E_T distribution for μ^+ (a) and μ^- (b) in the data sample dominated by the QCD background. The hadronic recoil derived from data has been used to correct the MC simulation. The $W \rightarrow \mu\nu$ contribution (light shaded region) is normalized to the integrated luminosity of the data sample using a MC simulation, and the normalization of the QCD simulation (dark shaded region) is taken as the difference between the data and the estimated $W \rightarrow \mu\nu$ contribution. The $W \rightarrow \mu\nu$ contribution in this data sample is negligible. The dark shaded band in each ratio plot shows the statistical uncertainty in the QCD MC E_T shape, and the light shaded band shows the total uncertainty, including the systematic uncertainties due to QCD E_T modeling as discussed in Sec. VI.

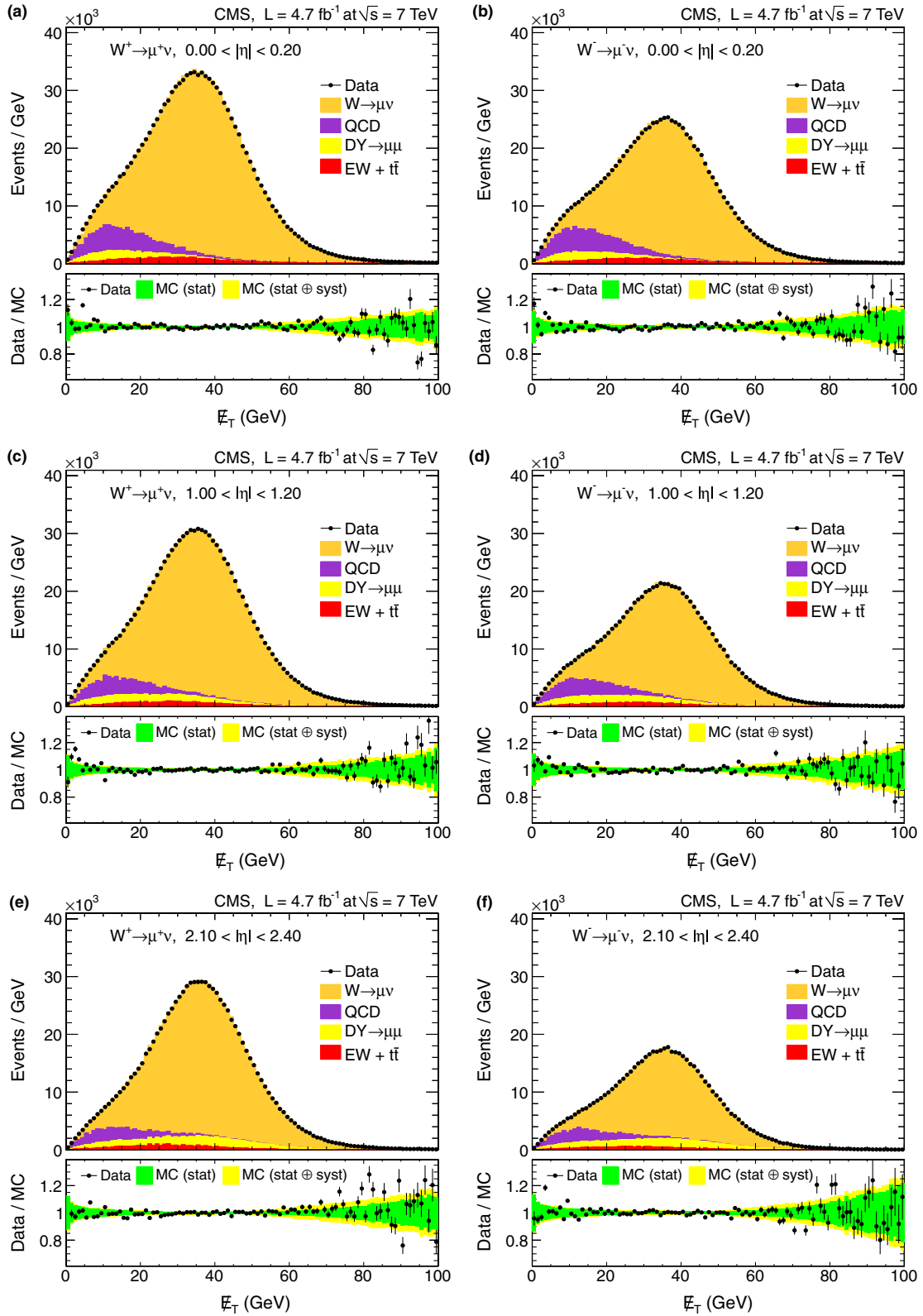


FIG. 5 (color online). Muon $p_T > 25$ GeV data sample. Examples of the extraction of the $W \rightarrow \mu\nu$ signal from fits to E_T distributions of $W \rightarrow \mu\nu$ candidates in data: $0.0 \leq |\eta| < 0.2$ (a), (b), $1.0 \leq |\eta| < 1.2$ (c), (d), and $2.1 \leq |\eta| < 2.4$ (e), (f). The fits to $W^+ \rightarrow \mu^+\nu$ and $W^- \rightarrow \mu^-\bar{\nu}$ candidates are in panels (a), (c), (e) and (b), (d), (f), respectively. The ratios between the data points and the final fits are shown at the bottom of each panel. The dark shaded band in each ratio plot shows the statistical uncertainty in the shape of the MC E_T distribution, and the light shaded band shows the total uncertainty, including all systematic uncertainties as discussed in Sec. VI.

B. Extraction of the asymmetry from fits to the missing transverse energy

The $W \rightarrow \mu\nu$ signal yields are obtained by fitting the E_T distributions with all corrections applied. The events are selected with the default muon p_T threshold of 25 GeV. The fits for W^+ and W^- are shown in Fig. 5 for three ranges of $|\eta|$, namely, $0.0 \leq |\eta| < 0.2$, $1.0 \leq |\eta| < 1.2$, and $2.1 \leq |\eta| < 2.4$. The ratio of the data to the fit result is shown below each distribution, demonstrating good agreement of the fits with the data.

Table I summarizes the fitted yields N^{W^+} and N^{W^-} , the correlation coefficient, the χ^2 value for each fit, and the raw asymmetry \mathcal{A}^{raw} . The χ^2 values indicate that the fits are good. The uncertainty in \mathcal{A}^{raw} takes the covariance of N^{W^+} and N^{W^-} into account. Corrections to \mathcal{A}^{raw} for potential bias are discussed in the next section.

As an important cross-check, we repeat the analysis with a higher muon p_T threshold of 35 GeV. The background compositions change significantly; the QCD background is reduced to about 1%. Furthermore, the predicted asymmetry differs from that predicted for the default analysis with the 25 GeV threshold. The results are summarized in Table I; they can be compared directly to the earlier measurement done with electrons [12].

VI. SYSTEMATIC UNCERTAINTIES AND CORRECTIONS

The systematic uncertainties arise from many sources, including the measurement of the muon kinematics (efficiency, scale, and resolution), the modeling of the E_T distributions, backgrounds, the boson \vec{q}_T distribution, and final-state radiation. In general, the total systematic uncertainty is 2–2.5 times larger than the statistical uncertainty (see Table II) and the main contributions come from the muon efficiency and from the QCD background. In the sections below, we discuss each source of systematic uncertainty, starting with muon-related quantities, followed by the E_T measurement, backgrounds, and boson-related modeling issues.

We evaluate many of these uncertainties using a MC method, in which 400 sets of pseudodata are fitted to obtain the distribution of \mathcal{A}^{raw} values. This method allows us to propagate the uncertainties of the corrections to the measurement in a rigorous manner.

Several sources of potential bias are considered. To evaluate the bias, we defined a “true” muon charge asymmetry, $\mathcal{A}^{\text{true}}$, calculated by taking the muon four-vectors and charge directly from the MC generator.

TABLE I. Summary of the fitted N^{W^+} , N^{W^-} , the correlation between the uncertainties in N^{W^+} and N^{W^-} ($\rho_{(N^{W^+}, N^{W^-})}$), the χ^2 of the fit, and the extracted \mathcal{A}^{raw} for each $|\eta|$ bin. The number of degrees of freedom (n_{dof}) in each fit is 197. Here, $\rho_{(N^{W^+}, N^{W^-})}$ and \mathcal{A}^{raw} are expressed as percentages.

$ \eta $ bin	N^{W^+} (10^3)	N^{W^-} (10^3)	$\rho_{(N^{W^+}, N^{W^-})}$ (%)	χ^2 ($n_{\text{dof}} = 197$)	\mathcal{A}^{raw} (%)
$p_T > 25$ GeV					
0.00–0.20	1033.0 ± 1.4	764.9 ± 1.2	14.5	255	14.912 ± 0.096
0.20–0.40	970.2 ± 1.3	713.9 ± 1.2	14.9	190	15.216 ± 0.098
0.40–0.60	1060.3 ± 1.4	771.5 ± 1.2	14.7	220	15.766 ± 0.094
0.60–0.80	1055.1 ± 1.4	752.4 ± 1.2	14.6	213	16.745 ± 0.093
0.80–1.00	935.8 ± 1.3	652.1 ± 1.1	14.5	245	17.866 ± 0.098
1.00–1.20	931.0 ± 1.3	625.4 ± 1.1	13.9	231	19.636 ± 0.099
1.20–1.40	949.0 ± 1.3	621.6 ± 1.1	14.2	209	20.848 ± 0.099
1.40–1.60	957.1 ± 1.3	607.3 ± 1.1	13.7	202	22.365 ± 0.099
1.60–1.85	1131.8 ± 1.4	687.6 ± 1.2	14.7	225	24.417 ± 0.093
1.85–2.10	1113.4 ± 1.4	656.8 ± 1.1	12.9	237	25.797 ± 0.094
2.10–2.40	843.6 ± 1.2	481.3 ± 1.0	11.8	244	27.341 ± 0.106
$p_T > 35$ GeV					
0.00–0.20	574.3 ± 1.0	459.7 ± 0.9	18.9	203	11.083 ± 0.116
0.20–0.40	538.9 ± 0.9	428.9 ± 0.9	17.4	202	11.371 ± 0.119
0.40–0.60	588.3 ± 1.0	462.8 ± 0.9	18.5	187	11.935 ± 0.114
0.60–0.80	582.9 ± 1.0	453.7 ± 0.9	18.7	205	12.472 ± 0.114
0.80–1.00	513.7 ± 0.9	392.3 ± 0.8	18.7	218	13.406 ± 0.124
1.00–1.20	509.1 ± 0.9	379.2 ± 0.8	15.7	226	14.620 ± 0.121
1.20–1.40	520.2 ± 0.9	376.9 ± 0.8	16.2	191	15.970 ± 0.123
1.40–1.60	522.7 ± 0.9	370.2 ± 0.8	14.7	195	17.074 ± 0.123
1.60–1.85	614.6 ± 1.0	418.8 ± 0.9	17.5	239	18.945 ± 0.118
1.85–2.10	604.7 ± 1.0	395.8 ± 0.9	15.0	192	20.885 ± 0.123
2.10–2.40	464.3 ± 0.9	288.5 ± 0.8	14.7	234	23.357 ± 0.141

TABLE II. Systematic uncertainties in \mathcal{A} for each $|\eta|$ bin. The statistical uncertainty in each $|\eta|$ bin is also shown for comparison. A detailed description of each systematic uncertainty is given in the text. The values are expressed as percentages, the same as for the asymmetries.

$ \eta $ bin	0.0–0.2	0.2–0.4	0.4–0.6	0.6–0.8	0.8–1.0	1.0–1.2	1.2–1.4	1.4–1.6	1.6–1.85	1.85–2.1	2.1–2.4
$p_T > 25$ GeV											
Stat. unc.	0.096	0.098	0.094	0.093	0.098	0.099	0.099	0.099	0.093	0.094	0.106
Efficiency	0.111	0.133	0.121	0.122	0.170	0.175	0.170	0.168	0.165	0.175	0.268
QCD +/-	0.120	0.113	0.110	0.105	0.102	0.103	0.097	0.104	0.108	0.094	0.183
QCD shape	0.070	0.065	0.065	0.067	0.068	0.069	0.078	0.082	0.092	0.083	0.087
Muon scale	0.045	0.050	0.050	0.049	0.051	0.054	0.054	0.058	0.054	0.054	0.055
FSR	0.074	0.077	0.104	0.109	0.089	0.113	0.107	0.091	0.118	0.087	0.077
PDF	0.028	0.026	0.023	0.025	0.018	0.020	0.027	0.031	0.042	0.050	0.069
Drell-Yan bkg.	0.002	0.001	0.002	0.003	0.000	0.007	0.001	0.013	0.019	0.038	0.046
$E_T\phi$ modul.	0.011	0.009	0.033	0.012	0.029	0.034	0.044	0.045	0.055	0.049	0.038
Recoil	0.003	0.003	0.003	0.003	0.003	0.003	0.003	0.003	0.003	0.004	0.003
Pileup	0.017	0.013	0.011	0.005	0.014	0.025	0.022	0.031	0.019	0.028	0.000
Luminosity	0.002	0.003	0.004	0.004	0.006	0.009	0.012	0.017	0.024	0.033	0.040
$t\bar{t}$ bkg.	0.012	0.013	0.012	0.012	0.011	0.011	0.010	0.009	0.008	0.007	0.005
$W \rightarrow \tau\nu$ bkg.	0.026	0.026	0.026	0.026	0.026	0.025	0.025	0.025	0.025	0.025	0.024
Wq_T	0.003	0.004	0.004	0.005	0.008	0.011	0.008	0.009	0.006	0.003	0.000
Total syst. unc.	0.203	0.212	0.217	0.216	0.238	0.255	0.251	0.250	0.266	0.256	0.364
Total unc.	0.225	0.233	0.236	0.235	0.258	0.274	0.270	0.269	0.282	0.273	0.379
$p_T > 35$ GeV											
Stat. unc.	0.116	0.119	0.114	0.114	0.124	0.121	0.123	0.123	0.118	0.123	0.141
Efficiency	0.120	0.138	0.116	0.107	0.159	0.164	0.171	0.176	0.186	0.194	0.325
QCD +/-	0.151	0.138	0.135	0.128	0.133	0.118	0.116	0.122	0.137	0.120	0.168
QCD shape	0.030	0.025	0.017	0.023	0.024	0.022	0.018	0.017	0.031	0.031	0.037
Muon scale	0.122	0.135	0.134	0.141	0.146	0.154	0.162	0.170	0.161	0.172	0.189
FSR	0.028	0.050	0.057	0.078	0.022	0.041	0.076	0.055	0.090	0.109	0.105
PDF	0.008	0.008	0.007	0.011	0.012	0.010	0.017	0.022	0.031	0.040	0.058
Drell-Yan bkg.	0.010	0.009	0.009	0.003	0.006	0.010	0.008	0.009	0.009	0.020	0.040
$E_T\phi$ modul.	0.002	0.009	0.010	0.003	0.008	0.028	0.037	0.035	0.022	0.022	0.001
Recoil	0.005	0.006	0.005	0.004	0.005	0.004	0.005	0.004	0.004	0.006	0.008
Pileup	0.015	0.003	0.005	0.018	0.019	0.002	0.007	0.003	0.013	0.014	0.032
Luminosity	0.001	0.002	0.000	0.000	0.000	0.001	0.004	0.010	0.016	0.025	0.039
$t\bar{t}$ bkg.	0.011	0.013	0.012	0.011	0.011	0.010	0.010	0.009	0.007	0.006	0.005
$W \rightarrow \tau\nu$ bkg.	0.013	0.012	0.013	0.012	0.012	0.012	0.011	0.012	0.011	0.011	0.011
Wq_T	0.004	0.002	0.004	0.004	0.007	0.005	0.006	0.009	0.009	0.001	0.014
Total syst. unc.	0.234	0.245	0.232	0.234	0.258	0.261	0.278	0.283	0.301	0.313	0.436
Total unc.	0.261	0.272	0.259	0.260	0.286	0.288	0.304	0.308	0.323	0.336	0.458

A. Muon kinematics

One source of potential bias for \mathcal{A}^{raw} is the charge of the muon. The rate of charge mismeasurement, w , is very small but not zero. The measured asymmetry will differ from the true asymmetry by a factor $(1 - 2w)$ assuming that the rate of mismeasurement is the same for positive and negative muons. The muon charge misidentification rate has been studied in detail and shown to have a negligible effect on the measured asymmetry [10].

The muon p_T resolution can induce a spread of the measured asymmetry from $\mathcal{A}^{\text{true}}$, which varies from 1.5% to 5.0% [28] as a function of $|\eta|$. The resolution of $|\eta|$ is

several orders of magnitude smaller than the bin 'widths' used in this measurement; consequently, event migration around p_T - η thresholds has a negligible effect on the measured asymmetry.

The muon momentum correction affects both the yields and the shapes of the E_T distributions. To estimate the systematic uncertainty from this source, the muon $1/p_T$ correction parameters in each η - ϕ bin and the muon scale global correction parameters are varied 400 times within their uncertainties. Each time the event yields can be slightly different in both data and MC simulation, and the extraction of the asymmetry is done for each of the 400 cases. The root

mean square (RMS) of the measured \mathcal{A}^{raw} variations in each muon $|\eta|$ bin is taken as the systematic uncertainty, and the bin-to-bin correlations are assumed to be zero.

The systematic uncertainties resulting from the muon momentum corrections are typically less than 40% of those from the uncertainties in the muon efficiencies (discussed below) for the $p_T > 25$ GeV sample. However, the two uncertainties are comparable for the $p_T > 35$ GeV sample for two reasons: first, the charge-dependent bias from the alignment increases with p_T ; second, the Jacobian peak of the $W \rightarrow \mu\nu$ events is close to 35 GeV.

B. Muon efficiency ratio

A difference in the muon efficiencies for positively and negatively charged muons will cause the ratio of the selection efficiencies for W^+ and W^- to differ from unity. This would bias the measured charge asymmetry, and we correct the \mathcal{A}^{raw} for this bias.

As discussed previously, the muon off-line and trigger efficiencies are measured in 7 bins in p_T and 22 bins in $|\eta|$ for both μ^+ and μ^- . The off-line efficiency ratio between μ^+ and μ^- is very close to unity in most of the detector regions. However, there is evidence that this ratio deviates from unity in the transition regions between the DT and CSC detectors.

We correct for this bias using efficiencies for μ^+ and μ^- extracted from the $Z/\gamma^* \rightarrow \mu^+\mu^-$ data and MC samples. For each $|\eta|$ bin, an average W selection efficiency $\epsilon(W^\pm)$ is obtained from the expression

$$\epsilon(W^\pm) = \frac{\sum(k\epsilon_{\text{data}}^\pm(p_T, \eta)/\epsilon_{\text{MC}}^\pm(p_T, \eta))}{\sum(k/\epsilon_{\text{MC}}^\pm(p_T, \eta))}, \quad (5)$$

where $\epsilon_{\text{data}}^\pm(p_T, \eta)$, $\epsilon_{\text{MC}}^\pm(p_T, \eta)$ are total muon efficiencies, k are additional event-by-event weights introduced by W -boson q_T weighting described below, and the sum is over the selected $W \rightarrow \mu\nu$ events. The efficiency ratio ($r^{W^+/W^-} = \epsilon^+/\epsilon^-$) is used to correct the \mathcal{A}^{raw} for the efficiency bias using

$$\mathcal{A}^{\text{true}} = \mathcal{A}^{\text{raw}} - \frac{1 - (\mathcal{A}^{\text{raw}})^2}{2} (r^{W^+/W^-} - 1), \quad (6)$$

which is an expansion to leading order in $(r^{W^+/W^-} - 1)$. In addition, all MC samples are corrected for any data/MC efficiency difference.

To estimate the systematic uncertainty due to the muon efficiencies, the muon efficiency values in data and MC simulation are modified according to their errors in each p_T - η bin independently and 400 pseudoeficiency tables are generated. In each pseudoexperiment the efficiency values are used to correct the MC simulation and \mathcal{A}^{raw} . The raw asymmetry is further corrected for the efficiency ratio r^{W^+/W^-} described above. The RMS of the resulting asymmetries in each $|\eta|$ bin is taken to be the systematic

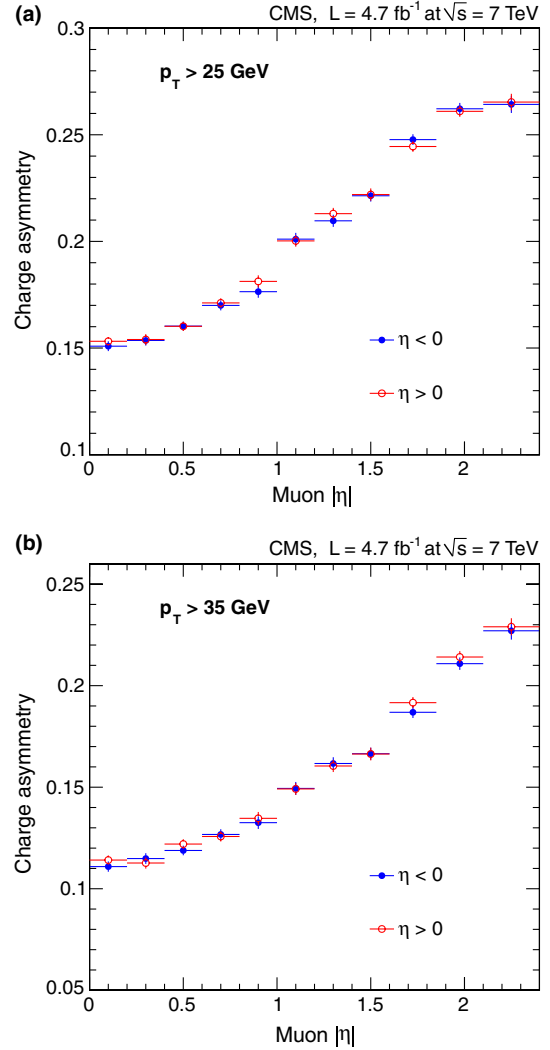


FIG. 6 (color online). Comparison of the final muon charge asymmetry (\mathcal{A}) extracted for the positive pseudorapidity ($\eta > 0$) and negative pseudorapidity ($\eta < 0$) regions with muon $p_T > 25$ GeV (a) and muon $p_T > 35$ GeV (b) samples. The uncertainties include only the statistical uncertainty from the signal extraction and uncertainty in the determination of the efficiencies for positive and negative muons.

uncertainty originating from the determination of the ratio of the muon efficiencies. In this study, the variations for different $|\eta|$ bins are completely independent from each other, so the systematic uncertainties due to the efficiency ratio have zero correlation between different $|\eta|$ bins.

As a cross-check, Fig. 6 shows a comparison of the measured muon charge asymmetry between positive and negative η regions, taken separately and then overlaid. They are in very good agreement with each other, for both muon p_T thresholds.

C. Backgrounds

The QCD background is estimated in part from the data. Nonetheless, a non-negligible systematic uncertainty

remains. We also discuss the uncertainty from the Drell-Yan background, and from the $t\bar{t}$ and $W \rightarrow \tau\nu$ backgrounds. The luminosity uncertainty enters in the estimation of these backgrounds, as discussed below.

1. QCD background

The total QCD background normalization is a parameter in the signal fit. The ratio of the QCD backgrounds in the W^+ and W^- samples is fixed to the ratio observed in the QCD control region for each muon η bin. The ratios are about 1.02 for the first ten η bins and approximately 1.05 for the last η bin, similar for both muon $p_T > 25$ and 35 GeV. There are two sources of the systematic uncertainties in the QCD background. The first is related to the ratio of the backgrounds in the W^+ and W^- samples (“QCD +/−”), and the second is related to the modeling of the shape of the E_T distribution in QCD events (“QCD shape”).

To evaluate the systematic uncertainties QCD +/−, the QCD ratio is varied by $\pm 5\%$ and $\pm 15\%$ for muon p_T thresholds of 25 and 35 GeV, respectively. The resulting shifts in the \mathcal{A}^{raw} are taken as the uncertainties. For the last $|\eta|$ bin, the variations are 10% (25 GeV) and 20% (35 GeV). These variations of the QCD ratio span the maximum range indicated by the QCD MC simulation. As an additional cross-check, we fix the QCD shape to be the same for μ^+ and μ^- and allow the two QCD normalizations to float in the extraction of the signal. We find that the fitted values for the ratio of the QCD backgrounds for W^+ and W^- are within the uncertainties quoted above. The bin-to-bin correlation of these uncertainties in the asymmetries is assumed to be zero.

The second source of systematic uncertainties is a difference in the shape of the QCD background for W^+ and W^- . The QCD E_T shape is taken from the MC simulation, and the recoil correction is applied as discussed in Sec. VA. Two types of variations in the shape of the QCD E_T distribution are considered. First, the shape of the QCD E_T distribution without the hadronic recoil correction is used in the extraction of the signal. This is done in a correlated way for the W^+ and W^- samples. Second, the shape of the E_T distribution for the QCD background is varied separately for the W^+ and W^- samples (within the statistical uncertainties), and the resulting shapes are used in the signal extraction. These two contributions to the uncertainties from the QCD shape are then added in quadrature. The bin-to-bin correlation of the systematic uncertainties due to each shape variation is assumed to be 100%.

2. Drell-Yan background

The $Z/\gamma^* \rightarrow \mu^+\mu^-$ events in the Drell-Yan control region are used to check the Drell-Yan normalization. This is done in bins of dimuon invariant mass: 15–30, 30–40, 40–60, 60–120, 120–150, and > 150 GeV. The $Z/\gamma^* \rightarrow \mu^+\mu^-$ MC simulation in each bin is compared to the data yields after

correcting the simulation for the data/simulation difference in pileup, Z -boson q_T , E_T modeling, and efficiencies. After correcting for the detector bias and physics mismodeling, the MC simulation describes the data well, as shown in Fig. 2 for the dimuon invariant mass between 60 and 120 GeV. The data yield in this mass bin is about 3% higher than the predictions from the next-to-next-to-leading-order (NNLO) cross section as calculated with FEWZ 3.1 [38].

The ratios of data to MC simulation of the $Z/\gamma^* \rightarrow \mu^+\mu^-$ event yields as a function of the dimuon mass are used to rescale the MC prediction of the Drell-Yan background. We take the shift in the \mathcal{A}^{raw} with and without this rescaling as the systematic uncertainty. This and the PDF uncertainties in the $Z/\gamma^* \rightarrow \mu^+\mu^-$ yields are considered as systematic uncertainties due to “Drell-Yan background normalization.” This uncertainty is almost negligible at central $|\eta|$ bins and increases in the forward $|\eta|$ bins. The Drell-Yan background is larger in the forward region because of the lower efficiency of the “Drell-Yan veto” due to less detector coverage. The systematic uncertainties in the Drell-Yan background are assumed to have 100% correlation from bin to bin.

3. The $t\bar{t}$ and $W \rightarrow \tau\nu$ backgrounds

The $t\bar{t}$ and $Z/\gamma^* \rightarrow \tau^+\tau^-$ backgrounds are normalized to the integrated luminosity of the data sample after correcting for the muon efficiency difference between data and MC simulation. The uncertainty of the integrated luminosity is 2.2% [39]. The normalization of all the MC backgrounds is varied by $\pm 2.2\%$, and the resulting maximum shift in \mathcal{A}^{raw} is taken as the systematic uncertainty in the determination of the luminosity. The bin-to-bin correlations are 100%.

The $t\bar{t}$ background estimate also depends on the theoretical prediction [31–33], to which we assign an additional 15%. The bin-to-bin correlation is 100%.

The $W \rightarrow \tau\nu$ background is normalized to the $W \rightarrow \mu\nu$ yields in data with a ratio obtained from a MC simulation. This ratio is largely determined by the branching fraction of τ decaying to μ . A 2% uncertainty is assigned to the $W \rightarrow \tau\nu$ to $W \rightarrow \mu\nu$ ratio [40]. The correlation of this uncertainty is 100% bin to bin.

D. Modeling uncertainties

The remaining systematic uncertainties pertain to the modeling of the detector and the signal process $W \rightarrow \mu\nu$. We discuss first the issues concerning the E_T distribution, then FSR, and finally the q_T distribution.

1. Modeling of missing transverse momentum

To evaluate the systematic uncertainty due to the ϕ modulation of E_T , the correction for the ϕ modulation is removed and the shift in the \mathcal{A}^{raw} is taken as the systematic uncertainty.

The hadronic recoil correction changes the shape of the E_T distribution of all MC samples. To calculate the uncertainties resulting from this source, the average recoil and resolution parameters are varied within their uncertainties, taking into account the correlations between them. This is done 400 times, the RMS of the resulting \mathcal{A}^{raw} variations is taken as systematic uncertainty, and bin-to-bin correlations are calculated.

Pileup can affect the E_T shapes. To estimate the effect of mismodeling the pileup in the simulation, the minimum-bias cross section is varied by $\pm 5\%$ and the pileup distributions expected in data are regenerated. The MC simulation is then weighted to match to data, and the resulting shift in \mathcal{A}^{raw} is treated as a systematic uncertainty due to the pileup. Pileup affects the E_T shapes for all muon η bins in the same direction with a correlation of 100%.

2. Final-state radiation

The emission of FSR photons in W decays reduces the muon p_T and can cause a difference in acceptance between W^+ and W^- . We studied the impact of the FSR on the muon charge asymmetry using the POWHEG $W \rightarrow \mu\nu$ MC sample. In this sample, FSR is implemented using a similar approach to parton showering and is approximate at the leading order (LO). We compare the muon charge

asymmetry before and after FSR, and the difference is found to be within 0.07%–0.12% and 0.03%–0.11% for muon p_T selections of 25 and 35 GeV, respectively. The raw asymmetry values are not corrected for FSR. Instead, the full shift in the muon charge asymmetry predicted by the POWHEG MC is taken as an additional systematic uncertainty, and the bin-to-bin correlation is assumed to be 100%.

3. PDF uncertainty

The evaluation of PDF uncertainties follows the PDF4LHC recommendation [41]. The NLO MSTW2008 [2], CT10 [3], and NNPDF2.1 [42] PDF sets are used. All simulated events are weighted to a given PDF set, and the overall normalization is allowed to vary. In this way both the uncertainties in the total cross sections, as well as in the shape of the E_T distribution are considered. To estimate the systematic uncertainty resulting from the uncertainties in the CT10 and MSTW2008 PDFs, asymmetric master equations are used [2,3]. For the CT10, the 90% confidence level (C.L.) uncertainty is rescaled to 68% C.L. by dividing by a factor of 1.64485. For the NNPDF2.3 PDF set, the RMS of the \mathcal{A}^{raw} distributions is taken. The half-width of the maximum deviation from combining all three PDF uncertainty bands is taken as the PDF uncertainty. The

TABLE III. Correlation matrix of systematic uncertainties between different $|\eta|$ bins. All systematic uncertainties are treated as additive. The values are expressed as percentages.

$ \eta $ bin	0.0–0.2	0.2–0.4	0.4–0.6	0.6–0.8	0.8–1.0	1.0–1.2	1.2–1.4	1.4–1.6	1.6–1.85	1.85–2.1	2.1–2.4
$p_T > 25$ GeV											
0.00–0.20	100.0	28.1	32.4	32.9	27.1	29.0	29.5	28.0	30.5	26.1	16.7
0.20–0.40		100.0	30.7	31.4	25.6	27.5	27.9	26.3	28.9	24.5	15.8
0.40–0.60			100.0	37.4	30.9	33.8	34.5	32.1	36.1	30.3	19.3
0.60–0.80				100.0	31.1	34.0	34.4	32.0	36.3	30.4	20.0
0.80–1.00					100.0	28.5	29.5	28.0	31.2	26.9	17.3
1.00–1.20						100.0	32.6	31.1	34.8	30.2	19.3
1.20–1.40							100.0	32.8	36.9	32.2	20.8
1.40–1.60								100.0	36.0	32.7	21.3
1.60–1.85									100.0	37.1	24.9
1.85–2.10										100.0	24.4
2.10–2.40											100.0
$p_T > 35$ GeV											
0.00–0.20	100.0	4.6	4.8	6.4	3.4	3.6	4.7	3.4	5.4	5.8	4.3
0.20–0.40		100.0	6.4	8.5	3.3	4.3	6.3	4.4	7.2	8.0	5.8
0.40–0.60			100.0	9.8	3.8	5.6	8.4	6.2	8.9	9.9	6.6
0.60–0.80				100.0	5.1	6.9	10.7	7.8	11.9	13.5	9.7
0.80–1.00					100.0	3.2	4.2	3.3	4.7	5.0	3.6
1.00–1.20						100.0	7.0	5.4	7.0	7.5	4.7
1.20–1.40							100.0	8.1	10.8	12.0	7.8
1.40–1.60								100.0	8.8	9.9	6.7
1.60–1.85									100.0	14.2	10.3
1.85–2.10										100.0	12.6
2.10–2.40											100.0

CT10 error set is used to estimate the bin-to-bin correlations. The PDF uncertainties are about 10% of the total experimental uncertainty.

4. W -boson q_T modeling

To improve the agreement between data and simulation, the W -boson q_T spectrum is weighted using weight factors determined by the ratios of the distribution of boson q_T for $Z/\gamma^* \rightarrow \mu^+\mu^-$ events in data and MC simulation. We assume that the corrections are the same for W and Z events. This assumption is tested using two different sets of MC simulations: one from the POWHEG event generator and the other from MADGRAPH [43]. Here, the MADGRAPH simulation is treated as the “data,” and the ratio of Z -boson q_T of the MADGRAPH and POWHEG simulations is compared to the same ratio in simulated W -boson events. This double ratio is parametrized using an empirical function to smooth the statistical fluctuations, and additional weights are obtained using the fitted function. We weight the POWHEG simulation to be close to the MADGRAPH simulation and measure the asymmetry again. The deviation of \mathcal{A}^{raw} is taken as the systematic uncertainty due to mis-modeling of W -boson q_T . The default boson q_T weighting is based on the POWHEG simulation.

E. Total systematic uncertainty

Table II summarizes the systematic uncertainties in all $|\eta|$ bins. For comparison, the statistical uncertainty in each $|\eta|$ bin is also shown. The dominant systematic uncertainties come from muon efficiencies, QCD background, and the muon momentum correction. The correlation matrix of systematic uncertainty among $|\eta|$ bins is reported in Table III. The correlations among $|\eta|$ bins are small and do not exceed 37% and 14% for muon p_T thresholds of 25 and 35 GeV, respectively. Much of the correlation is due to the systematic uncertainties in FSR and QCD background. The total covariance matrix, including both statistical and systematic uncertainties, is provided as Supplemental Material [44].

VII. RESULTS AND DISCUSSION

The measured asymmetries \mathcal{A} , after all the corrections, are shown in Fig. 7 as a function of muon $|\eta|$ and summarized in Table IV. In Fig. 7 both statistical and systematic uncertainties are included in the error bars. These asymmetries are compared to predictions based on several PDF sets. The theoretical predictions are obtained using the FEWZ 3.1 [38] NLO MC calculation interfaced with the CT10 [3], NNPDF2.3 [45], HERAPDF1.5 [46], MSTW2008 [2], and MSTW2008CPdeut [15] PDF sets. No EW corrections are included in these calculations. The numerical values of the theoretical predictions are shown in Table IV. We cross-check the theoretical predictions using the DYNLO 1.0 [47,48] MC tool, and the agreement

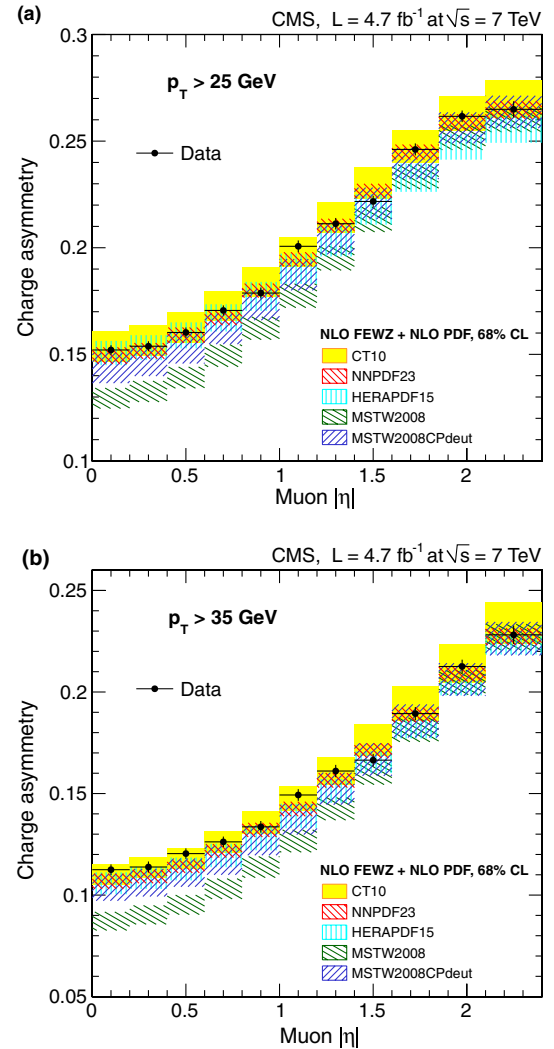


FIG. 7 (color online). Comparison of the measured muon charge asymmetries to the NLO predictions calculated using the FEWZ 3.1 [38] MC tool interfaced with the NLO CT10 [3], NNPDF2.3 [45], HERAPDF1.5 [46], MSTW2008 [2], and MSTW2008CPdeut [15] PDF sets. No EW corrections have been considered in these predictions. Results for muon $p_T > 25$ and > 35 GeV are shown in panels (a) and (b), respectively. The vertical error bars on data points include both statistical and systematic uncertainties. The data points are shown at the center of each $|\eta|$ bin. The theoretical predictions are calculated using the FEWZ 3.1 [38] MC tool. The PDF uncertainty for each PDF set is shown by the shaded (or hatched) band and corresponds to 68% C.L.

between the FEWZ 3.1 and DYNLO 1.0 is within 1%. The predictions using the CT10 and HERAPDF1.5 PDF sets are in good agreement with the data. The predictions using the NNPDF2.3 PDF set (which include the previous CMS electron charge asymmetry result and other LHC experimental measurements [45]) are also in good agreement with the data. The predictions using the MSTW2008 PDF set are not in agreement with the data, as seen in our previous analyses [10,12]. The more recent

TABLE IV. Summary of the final results for muon charge asymmetry \mathcal{A} . The first uncertainty is statistical and the second is systematic. The theoretical predictions are obtained using the FEWZ 3.1 [38] MC tool interfaced with the NLO CT10 [3], NNPDF2.3 [45], HERAPDF1.5 [46], and MSTW2008CPdeut [15] PDF sets. The PDF uncertainty is at 68% C.L. For each $|\eta|$ bin, the theoretical prediction is calculated using the averaged differential cross sections for positively and negatively charged leptons. The numerical precision of the theoretical predictions is less than 10% of the statistical uncertainties of the measurements. The values are expressed as percentages.

$ \eta $	$\mathcal{A} [\pm(\text{stat}) \pm (\text{syst})]$	CT10	NNPDF2.3	HERAPDF1.5	MSTW2008-CP-deut
$p_T > 25 \text{ GeV}$					
0.00–0.20	$15.21 \pm 0.10 \pm 0.20$	$15.35^{+0.74}_{-0.68}$	14.94 ± 0.39	$15.33^{+0.30}_{-0.84}$	$14.34^{+0.75}_{-0.69}$
0.20–0.40	$15.38 \pm 0.10 \pm 0.21$	$15.63^{+0.73}_{-0.69}$	15.16 ± 0.37	$15.58^{+0.32}_{-0.85}$	$14.67^{+0.75}_{-0.69}$
0.40–0.60	$16.03 \pm 0.09 \pm 0.22$	$16.27^{+0.71}_{-0.70}$	15.90 ± 0.36	$16.16^{+0.34}_{-0.88}$	$15.27^{+0.75}_{-0.70}$
0.60–0.80	$17.06 \pm 0.09 \pm 0.22$	$17.27^{+0.68}_{-0.71}$	16.71 ± 0.34	$16.98^{+0.37}_{-0.91}$	$16.19^{+0.74}_{-0.71}$
0.80–1.00	$17.88 \pm 0.10 \pm 0.24$	$18.45^{+0.66}_{-0.74}$	17.99 ± 0.33	$17.98^{+0.42}_{-0.94}$	$17.33^{+0.74}_{-0.73}$
1.00–1.20	$20.07 \pm 0.10 \pm 0.26$	$19.85^{+0.64}_{-0.76}$	19.46 ± 0.33	$19.25^{+0.48}_{-0.95}$	$18.74^{+0.73}_{-0.74}$
1.20–1.40	$21.13 \pm 0.10 \pm 0.25$	$21.50^{+0.63}_{-0.80}$	21.03 ± 0.33	$20.51^{+0.54}_{-0.92}$	$20.45^{+0.72}_{-0.76}$
1.40–1.60	$22.17 \pm 0.10 \pm 0.25$	$23.13^{+0.64}_{-0.84}$	22.66 ± 0.34	$21.92^{+0.59}_{-0.84}$	$22.12^{+0.70}_{-0.78}$
1.60–1.85	$24.61 \pm 0.09 \pm 0.27$	$24.87^{+0.65}_{-0.89}$	24.49 ± 0.35	$23.32^{+0.63}_{-0.70}$	$24.01^{+0.68}_{-0.79}$
1.85–2.10	$26.16 \pm 0.09 \pm 0.26$	$26.42^{+0.67}_{-0.95}$	25.88 ± 0.38	$24.70^{+0.65}_{-0.57}$	$25.70^{+0.65}_{-0.81}$
2.10–2.40	$26.49 \pm 0.11 \pm 0.36$	$27.13^{+0.74}_{-1.03}$	26.46 ± 0.42	$25.40^{+0.81}_{-0.48}$	$26.48^{+0.65}_{-0.87}$
$p_T > 35 \text{ GeV}$					
0.00–0.20	$11.25 \pm 0.12 \pm 0.23$	$11.00^{+0.52}_{-0.48}$	10.68 ± 0.37	$10.80^{+0.32}_{-0.76}$	$10.39^{+0.67}_{-0.67}$
0.20–0.40	$11.38 \pm 0.12 \pm 0.24$	$11.36^{+0.52}_{-0.49}$	10.91 ± 0.33	$11.07^{+0.33}_{-0.77}$	$10.61^{+0.68}_{-0.68}$
0.40–0.60	$12.04 \pm 0.11 \pm 0.23$	$11.80^{+0.52}_{-0.50}$	11.40 ± 0.31	$11.51^{+0.34}_{-0.79}$	$11.10^{+0.70}_{-0.69}$
0.60–0.80	$12.62 \pm 0.11 \pm 0.23$	$12.59^{+0.53}_{-0.53}$	12.18 ± 0.33	$12.17^{+0.36}_{-0.80}$	$11.71^{+0.72}_{-0.71}$
0.80–1.00	$13.36 \pm 0.12 \pm 0.26$	$13.60^{+0.55}_{-0.58}$	13.21 ± 0.35	$13.02^{+0.37}_{-0.82}$	$12.70^{+0.74}_{-0.74}$
1.00–1.20	$14.93 \pm 0.12 \pm 0.26$	$14.79^{+0.59}_{-0.64}$	14.24 ± 0.36	$14.10^{+0.40}_{-0.81}$	$13.75^{+0.77}_{-0.77}$
1.20–1.40	$16.11 \pm 0.12 \pm 0.28$	$16.14^{+0.64}_{-0.73}$	15.65 ± 0.36	$15.31^{+0.41}_{-0.77}$	$15.24^{+0.79}_{-0.79}$
1.40–1.60	$16.64 \pm 0.12 \pm 0.28$	$17.72^{+0.70}_{-0.83}$	17.11 ± 0.36	$16.68^{+0.40}_{-0.68}$	$16.69^{+0.79}_{-0.82}$
1.60–1.85	$18.94 \pm 0.12 \pm 0.30$	$19.53^{+0.77}_{-0.94}$	18.87 ± 0.36	$18.22^{+0.40}_{-0.51}$	$18.62^{+0.77}_{-0.86}$
1.85–2.10	$21.26 \pm 0.12 \pm 0.31$	$21.52^{+0.82}_{-1.06}$	20.89 ± 0.38	$20.15^{+0.41}_{-0.32}$	$20.71^{+0.71}_{-0.90}$
2.10–2.40	$22.81 \pm 0.14 \pm 0.44$	$23.53^{+0.86}_{-1.17}$	22.73 ± 0.42	$22.17^{+0.71}_{-0.33}$	$22.79^{+0.66}_{-0.99}$

MSTW2008CPdeut PDF set is a variant of the MSTW2008 PDF set with a more flexible input parametrization and deuteron corrections [15]. This modification has significantly improved the agreement with the CMS data even though they have not included LHC data, as shown in Fig. 7.

Since the per-bin total experimental uncertainties are significantly smaller than the uncertainty in the current PDF parametrizations, this measurement can be used to constrain PDFs in the next generation of PDF sets.

Figure 8 shows a comparison of the measured muon charge asymmetries to the NNLO predictions. The NNLO HERAPDF1.5 PDF is used. The calculations are performed using the FEWZ 3.1 and DYNLO 1.0 MC tools. Both MC

simulations give consistent results with agreement at the 1% level. With a p_T threshold of 25 GeV, the NLO and NNLO predictions are very similar. The NNLO predictions are slightly higher in high- $|\eta|$ regions. In the same high- $|\eta|$ region at a p_T threshold of 35 GeV, the NNLO predictions are significantly lower than the NLO prediction. However, they agree well within the quoted PDF uncertainty in the HERAPDF1.5 PDFs.

Figure 9 shows a comparison of this result to the previous CMS electron charge asymmetry measurement extracted from part of the 2011 CMS data [12]. The electron charge asymmetry has been measured with a slightly different η binning because of the different sub-detector geometry in the calorimeter and the muon system.

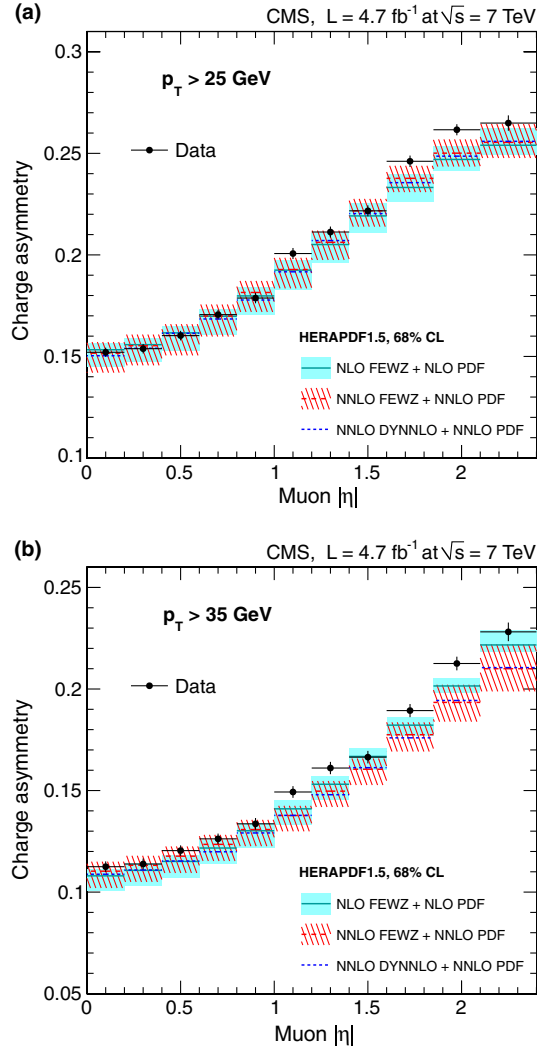


FIG. 8 (color online). Comparison of the measured muon charge asymmetries to the NNLO predictions for muon $p_T > 25$ (a) and muon $p_T > 35$ GeV (b). The NNLO HERAPDF1.5 [46] PDF has been used in the NNLO calculations. The calculations are performed using both the FEWZ 3.1 [38] and DYNLO 1.0 [47,48] MC tools. The NLO prediction based on FEWZ 3.1 is also shown here. The HERAPDF1.5 PDF uncertainties are shown by the shaded (NLO) and hatched (NNLO) bands.

We have calculated the bin-by-bin differences between these two measurements using the first seven η bins, where identical bin definitions are used, and the differences are fitted with a constant. The fitted constant is larger than zero by about 1.7 sigma, and the muon channel exhibits slightly higher asymmetry in these seven η bins than the electron one. The electron charge asymmetry uses a statistically independent data sample. A combination of both results can be used to improve the global PDF fits. The correlation between the electron charge asymmetry and this result is expected to be small. The completely correlated systematic sources of uncertainty include the luminosity measurement, $t\bar{t}$ background, $W \rightarrow \tau\nu$ background, and PDF uncertainty.

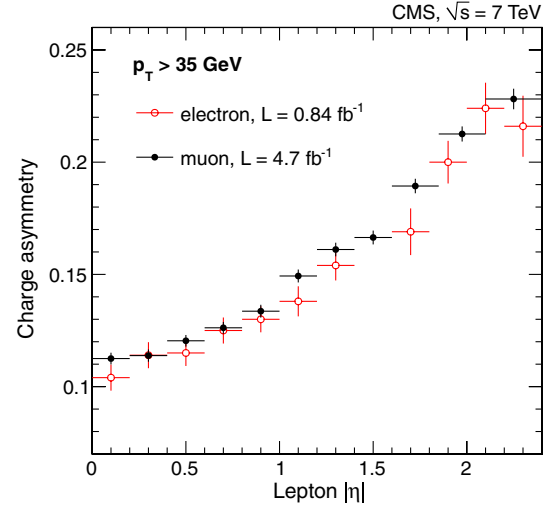


FIG. 9 (color online). Comparison of this measurement to the previous CMS electron charge asymmetry result [12]. Results are shown for lepton $p_T > 35$ GeV.

The theoretical predictions for the lepton charge asymmetry are given for the kinematic region specified by the lepton p_T threshold. The p_T distribution of the W boson affects the acceptance, and hence, the predicted charge asymmetry. However, the effect on W^+ and W^- is largely correlated. Therefore, the impact on the lepton charge asymmetry measurement mostly cancels. Figure 10 shows the comparison of these results to the NLO CT10 PDF predictions based on the FEWZ 3.1 and RESBOS [49–52]. RESBOS does a resummation in boson q_T at NLO (and approximate NNLO) plus approximate next-to-next-to-leading logarithm, which yields a more realistic description of boson q_T than a fixed-order calculation such as the FEWZ 3.1. The difference between the FEWZ 3.1 and RESBOS predictions is negligible and our measurement, however precise, is not sensitive to the difference.

VIII. THE QCD ANALYSIS OF HERA AND CMS RESULTS OF W -BOSON PRODUCTION

The main objective of the QCD analysis presented in this section is to exploit the constraining power and the interplay of the muon charge asymmetry measurements, presented in this paper, and the recent measurements of W + charm production at CMS [18] to determine the PDFs of the proton. These two data sets, together with the combined HERA inclusive cross section measurements [19], are used in an NLO perturbative QCD (pQCD) analysis.

Renormalization group equations, formulated in terms of DGLAP evolution equations [53–58], predict the dependence of the PDFs on the energy scale Q of the process in pQCD. The dependence on the partonic fraction x of the proton momentum cannot be derived from first principles and must be constrained by experimental measurements.

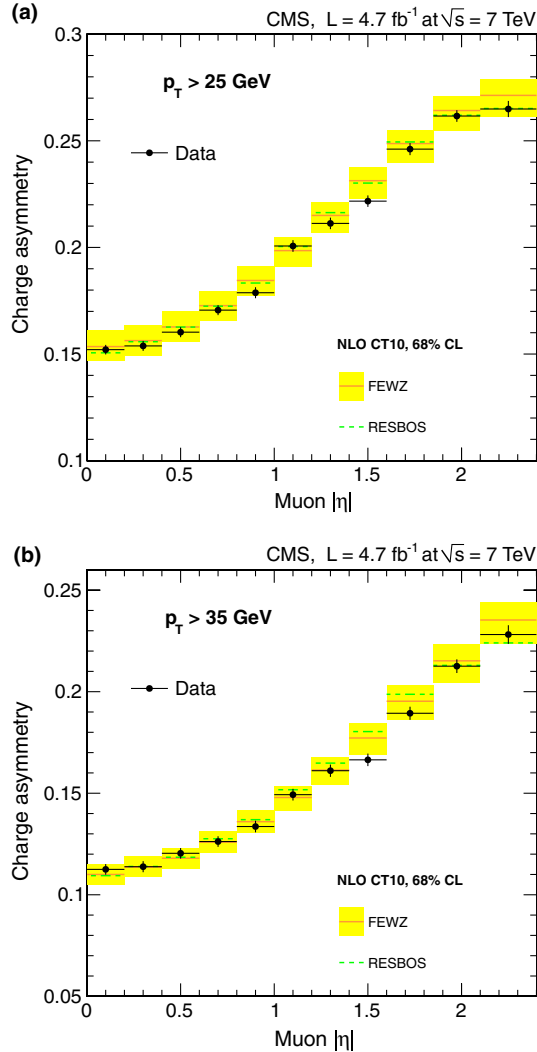


FIG. 10 (color online). Comparison of the measured muon charge asymmetry to theoretical predictions based on the FEWZ 3.1 [38] and RESBOS [49–52] tools. The NLO CT10 PDF is used in both predictions. Results are shown for muon $p_T > 25$ (a) and muon $p_T > 35$ GeV (b). The CT10 PDF uncertainty is shown by the shaded bands.

Deep inelastic lepton-proton scattering (DIS) experiments cover a broad range of the (x, Q^2) kinematic plane. The region of small and intermediate x is probed primarily by the precise data of HERA, which impose the tightest constraints on the existing PDFs. However, some details of flavor composition, in particular the light-sea-quark content and the strange-quark distribution of the proton, are still poorly known. Measurements of the W - and Z -boson production cross sections in proton-(anti)proton collisions are sensitive to the light-quark distributions, and the constraining power of the W -boson measurements is applied in this analysis.

The muon charge asymmetry measurements probe the valence-quark distribution in the kinematic range $10^{-3} \leq x \leq 10^{-1}$ and have indirect sensitivity to the strange-quark

distribution. The measurements of the total and differential cross sections of $W + \text{charm}$ production have the potential to access the strange-quark distribution directly through the LO process $g + s \rightarrow W + c$. This reaction was proposed as a way to determine the strange-quark and antiquark distributions [59–61].

Before the LHC era, constraints on the strange-quark distribution were obtained from semi-inclusive charged-current scattering at the NuTeV [62,63] and CCFR [64] experiments. Dimuon production in neutrino-nucleus reactions is sensitive to strangeness at LO in QCD in reactions such as $W^+ + s \rightarrow c$. These measurements probe the (anti) strange-quark density at $x \approx 10^{-1}$ and Q^2 of approximately 10 GeV^2 , but their interpretation is complicated by nuclear corrections and uncertainties in the charm-quark fragmentation function. The NOMAD Collaboration reported a recent determination of the strange-quark suppression factor

$$\kappa_s(Q^2) = \frac{\int_0^1 x[\bar{s}(x, Q^2) + s(x, Q^2)]dx}{\int_0^1 x[\bar{u}(x, Q^2) + \bar{d}(x, Q^2)]dx}, \quad (7)$$

where the value $\kappa_s(Q^2 = 20 \text{ GeV}^2) = 0.591 \pm 0.019$ is determined at NNLO by using dimuon production [65]. The measurements of semi-inclusive hadron production on a deuteron target at HERMES [66] have been recently reevaluated [67] to obtain the x dependence of the strange-quark distribution at LO at an average $\langle Q^2 \rangle = 2.5 \text{ GeV}^2$. In that analysis the strange-quark distribution is found to vanish above $x = 0.1$, but this result depends strongly on the assumptions of the kaon fragmentation function.

In a recent analysis by the ATLAS Collaboration [68], the inclusive cross section measurements of W - and Z -boson production were used in conjunction with DIS inclusive data from HERA. The result supports the hypothesis of a symmetric composition of the light-quark sea in the kinematic region probed, i.e., $\bar{s} = \bar{d}$.

The LHC measurements of associated production of W bosons and charm quarks probe the strange-quark distribution in the kinematic region of $x \approx 0.012$ at the scale $Q^2 = m_W^2$. The cross sections for this process were recently measured by the CMS Collaboration [18] at a center-of-mass energy $\sqrt{s} = 7 \text{ TeV}$ with a total integrated luminosity of 5 fb^{-1} . The results of the QCD analysis presented here use the absolute differential cross sections of $W + \text{charm}$ production, measured in bins of the pseudorapidity of the lepton from the W decay, for transverse momenta larger than 35 GeV .

A. Details of the QCD analysis

The NLO QCD analysis is based on the inclusive DIS data [19] from HERA, measurements of the muon charge asymmetry in W production for $p_T > 25 \text{ GeV}$, and measurements of associated $W + \text{charm}$ production [18]. The treatment of experimental uncertainties for the HERA data

follows the prescription of HERAPDF1.0 [19]. The correlations of the experimental uncertainties for the muon charge asymmetry and $W + \text{charm}$ data are taken into account.

The theory predictions for the muon charge asymmetry and $W + \text{charm}$ production are calculated at NLO by using the MCFM program [31,32], which is interfaced to APPLGRID [69].

The open source QCD fit framework for PDF determination HERAFITTER [19,70,71] is used and the partons are evolved by using the QCDNUM program [72]. The TR' [2,73] general mass variable flavor number scheme is used for the treatment of heavy-quark contributions with the following conditions: (i) heavy-quark masses are chosen as $m_c = 1.4 \text{ GeV}$ and $m_b = 4.75 \text{ GeV}$, (ii) renormalization and factorization scales are set to $\mu_r = \mu_f = Q$, and (iii) the strong coupling constant is set to $\alpha_s(m_Z) = 0.1176$.

The Q^2 range of HERA data is restricted to $Q^2 \geq Q_{\min}^2 = 3.5 \text{ GeV}^2$ to assure the applicability of pQCD over the kinematic range of the fit. The procedure for the determination of the PDFs follows the approach used in the HERAPDF1.0 QCD fit [19].

The following independent combinations of parton distributions are chosen in the fit procedure at the initial scale of the QCD evolution $Q_0^2 = 1.9 \text{ GeV}^2$: $xu_v(x)$, $xd_v(x)$, $xg(x)$ and $x\bar{U}(x)$, $x\bar{D}(x)$ where $x\bar{U}(x) = x\bar{u}(x)$, $x\bar{D}(x) = x\bar{d}(x) + x\bar{s}(x)$. At Q_0 , the parton distributions are represented by

$$xu_v(x) = A_{u_v} x^{B_{u_v}} (1-x)^{C_{u_v}} (1 + E_{u_v} x^2), \quad (8)$$

$$xd_v(x) = A_{d_v} x^{B_{d_v}} (1-x)^{C_{d_v}}, \quad (9)$$

$$x\bar{U}(x) = A_{\bar{U}} x^{B_{\bar{U}}} (1-x)^{C_{\bar{U}}}, \quad (10)$$

$$x\bar{D}(x) = A_{\bar{D}} x^{B_{\bar{D}}} (1-x)^{C_{\bar{D}}}, \quad (11)$$

$$xg(x) = A_g x^{B_g} (1-x)^{C_g} + A'_g x^{B'_g} (1-x)^{C'_g}. \quad (12)$$

The normalization parameters A_{u_v} , A_{d_v} , A_g are determined by the QCD sum rules, the B parameter is responsible for small- x behavior of the PDFs, and the parameter C describes the shape of the distribution as $x \rightarrow 1$. A flexible form for the gluon distribution is adopted here, where the choice of $C'_g = 25$ is motivated by the approach of the MSTW group [2,73].

Two types of analyses are made. The first is denoted as “fixed- s fit” and is performed by fitting 13 parameters in Eqs. (8)–(12) to analyze the impact of the muon charge asymmetry measurements on the valence-quark distributions. Additional constraints $B_{\bar{U}} = B_{\bar{D}}$ and $A_{\bar{U}} = A_{\bar{D}}(1 - f_s)$ are imposed with f_s being the strangeness fraction, $f_s = \bar{s}/(\bar{d} + \bar{s})$, which is fixed to $f_s = 0.31 \pm 0.08$ as in Ref. [2].

The second analysis is denoted as “free- s fit,” in which the interplay between the muon charge asymmetry measurements and $W + \text{charm}$ production data is analyzed. The strange-quark distribution is determined by fitting 15 parameters in Eqs. (8)–(12). Here, instead of Eq. (11) \bar{d} and \bar{s} are fitted separately by using the functional forms

$$x\bar{d}(x) = A_{\bar{d}} x^{B_{\bar{d}}} (1-x)^{C_{\bar{d}}}, \quad (13)$$

$$x\bar{s}(x) = A_{\bar{s}} x^{B_{\bar{s}}} (1-x)^{C_{\bar{s}}}. \quad (14)$$

Additional constraints $A_{\bar{u}} = A_{\bar{d}}$ and $B_{\bar{u}} = B_{\bar{d}}$ are applied to ensure the same normalization for \bar{u} and \bar{d} densities at $x \rightarrow 0$. The strange-antiquark parameter $B_{\bar{s}}$ is set equal to $B_{\bar{d}}$, while $A_{\bar{s}}$ and $C_{\bar{s}}$ are treated as free parameters of the fit, assuming $x\bar{s} = x\bar{s}$. This parametrization cannot be applied to HERA DIS data alone, because those data do not have sufficient sensitivity to the strange-quark distribution.

B. The PDF uncertainties

The PDF uncertainties are estimated according to the general approach of HERAPDF1.0 [19] in which experimental, model, and parametrization uncertainties are taken into account. A tolerance criterion of $\Delta\chi^2 = 1$ is adopted for defining the experimental uncertainties that originate from the measurements included in the analysis. Model uncertainties arise from the variations in the values assumed for the heavy-quark masses m_b , m_c with $4.3 \leq m_b \leq 5 \text{ GeV}$, $1.35 \leq m_c \leq 1.65 \text{ GeV}$, and the value of Q_{\min}^2 imposed on the HERA data, which is varied in the interval $2.5 \leq Q_{\min}^2 \leq 5.0 \text{ GeV}^2$. The parametrization uncertainty is estimated similarly to the HERAPDF1.0 procedure: for all parton densities, additional parameters are added one by one in the functional form of the parametrizations such that Eqs. (8)–(11) are generalized to $Ax^B(1-x)^C(1+Dx)$ or $Ax^B(1-x)^C(1+Dx+Ex^2)$. In the free- s fit, in addition, the parameters $B_{\bar{s}}$ and $B_{\bar{d}}$ are decoupled. Furthermore, the starting scale is varied within $1.5 \leq Q_0^2 \leq 2.5 \text{ GeV}^2$. The parametrization uncertainty is constructed as an envelope built from the maximal differences between the PDFs resulting from all the parametrization variations and the central fit at each x value. The total PDF uncertainty is obtained by adding experimental, model, and parametrization uncertainties in quadrature. In the following, the quoted uncertainties correspond to 68% C.L.

C. Results of the QCD analysis

The muon charge asymmetry measurements, together with HERA DIS cross section data, improve the precision of the valence quarks over the entire x range in the fixed- s fit. This is illustrated in Fig. 11, where the u and d

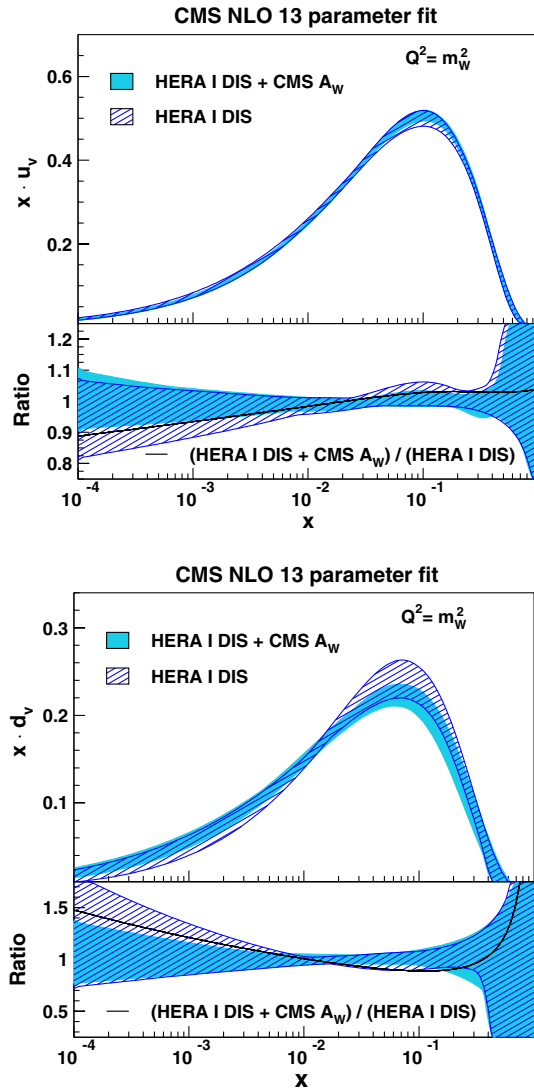


FIG. 11 (color online). Distributions of u valence (top) and d valence (bottom) quarks as functions of x at the scale $Q^2 = m_W^2$. The results of the 13-parameter fixed- s fit to the HERA data and muon asymmetry measurements (light shaded band), and to HERA only (dark hatched band) are compared. The total PDF uncertainties are shown. In the bottom panels the distributions are normalized to one for a direct comparison of the uncertainties. The change of the PDFs with respect to the HERA-only fit is represented by a solid line.

valence-quark distributions are shown at the scale relevant for the W -boson production, $Q^2 = m_W^2$. The results at $Q^2 = 1.9 \text{ GeV}^2$ can be found in supplemental material. A change in the shapes of the light-quark distributions within the total uncertainties is observed. The details of the effect on the experimental PDF uncertainty of u valence, d valence, and d/u distributions are also given in supplemental material.

In the next step of the analysis, the CMS W + charm measurements are used together with the HERA DIS

data and the CMS muon charge asymmetry. Since both CMS W -boson production measurements are sensitive to the strange-quark distribution, a free- s fit can be performed. The advantage of including these two CMS data sets in the 15-parameter fit occurs because the d -quark distribution is significantly constrained by the muon charge asymmetry data, while the strange-quark distribution is directly probed by the associated W + charm production measurements. In the free- s fit, the strange-quark distribution $s(x, Q^2)$ and the strange-quark fraction $R_s(x, Q^2) = (s + \bar{s})/(\bar{u} + \bar{d})$ are determined. The global and partial χ^2 values for each data set are listed in Table V, where the χ^2 values illustrate a general agreement among all the data sets.

In Fig. 12, the resulting NLO parton distributions are presented at $Q_0^2 = 1.9 \text{ GeV}^2$ and $Q^2 = m_W^2$. The strange quark distribution $s(x, Q^2)$ and the ratio $R_s(x, Q^2)$ are illustrated in Fig. 13 at the same values of Q as in Fig. 12. The total uncertainty in Fig. 12 is dominated by the parametrization uncertainty in which most of the expansion in the envelope is caused by the decoupling parameter choice $B_{\bar{s}} \neq B_{\bar{d}}$. The strange-quark fraction rises with energy and reaches a value comparable to that of u and d antiquarks at intermediate to low x . Also, a suppression of R_s at large x is observed, which scales differently with the energy. This result is consistent with the prediction provided by the ATLAS Collaboration [68], where inclusive W - and Z -boson production measurements were used to determine $r_s = 0.5(s + \bar{s})/\bar{d}$. In Ref. [68], the NLO value of $r_s = 1.03$ with the experimental uncertainty $\pm 0.19_{\text{exp}}$ is quoted at $x = 0.023$ and $Q^2 = 1.9 \text{ GeV}^2$. In the framework used, the two definitions of the strange-quark fraction are very similar at the starting scale Q_0^2 and the values R_s and r_s can be directly compared.

TABLE V. Global χ^2/n_{dof} and partial χ^2 per number of data points n_{dp} for the data sets used in the 15-parameter QCD analysis.

Data sets	Global χ^2/n_{dof}	Partial χ^2/n_{dp}
DIS, $\frac{d\sigma_{W+c}}{d\eta^+}$, $\mathcal{A}(\eta_\mu)$	598/593	
NC cross section HERA I H1 + ZEUS e^-p		107/145
NC cross section HERA I H1 + ZEUS e^+p		417/379
CC cross section HERA I H1 + ZEUS e^-p		20/34
CC cross section HERA I H1 + ZEUS e^+p		36/34
CMS W^\pm muon charge asymmetry $\mathcal{A}(\eta_\mu)$		14/11
CMS W + c cross section $\frac{d\sigma_{W+c}}{d\eta^+}$		5/5

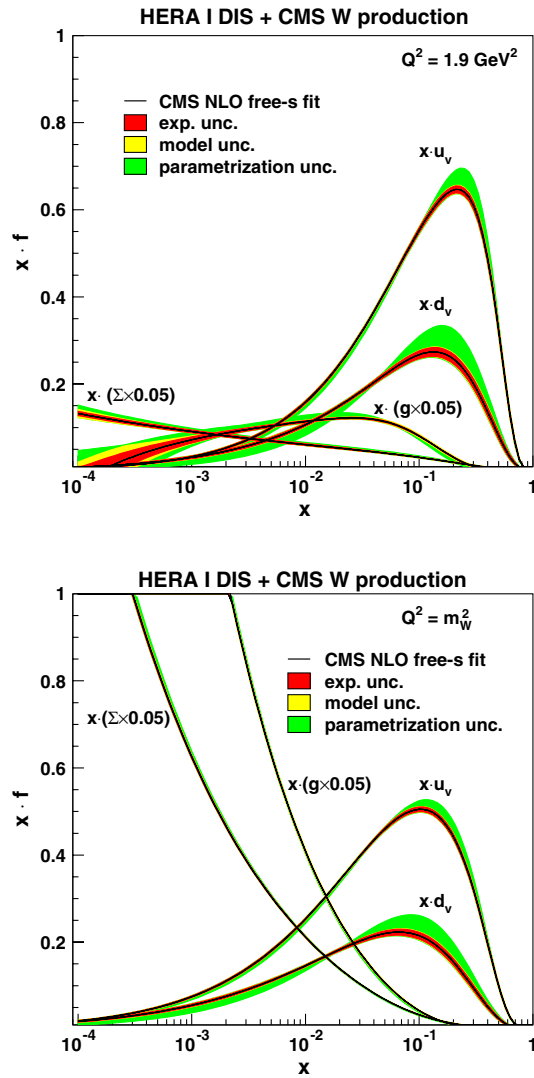


FIG. 12 (color online). Parton distribution functions, shown as functions of x , obtained by using HERA DIS data and CMS measurements of W -boson production in the free- s NLO QCD analysis. Gluon, valence, and sea distributions are presented at the starting scale $Q_0^2 = 1.9 \text{ GeV}^2$ of the PDF evolution (top) and the mass squared of the W boson (bottom). The sea distribution is defined as $\Sigma = 2 \cdot (\bar{u} + \bar{d} + \bar{s})$. The full band represents the total uncertainty. The individual contributions from the experimental, model, and parametrization uncertainties are represented by the bands of different shades. The gluon and sea distributions are scaled down by a factor of 20.

In the free- s fit, the strangeness suppression factor is determined at $Q^2 = 20 \text{ GeV}^2$ to be $\kappa_s = 0.52^{+0.12}_{-0.10}(\text{exp})^{+0.05}_{-0.06}(\text{model})^{+0.13}_{-0.10}(\text{parametrization})$, which is in agreement with the value [65] obtained by the NOMAD experiment at NNLO.

The impact of the measurement of differential cross sections of W + charm production on the strange-quark distribution and strangeness fraction R_s is also examined by using the Bayesian reweighting [13,14] technique. The

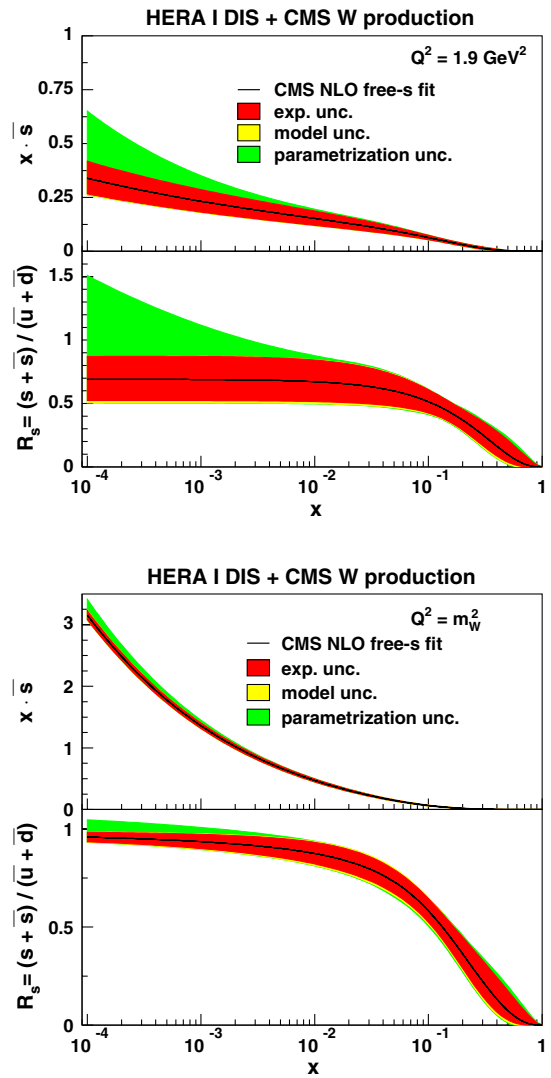


FIG. 13 (color online). Antistrange-quark distribution $\bar{s}(x, Q)$ and the ratio $R_s(x, Q)$, obtained in the QCD analysis of HERA and CMS data, shown as functions of x at the scale $Q^2 = 1.9 \text{ GeV}^2$ (top) and $Q^2 = m_W^2$ (bottom). The full band represents the total uncertainty. The individual contributions from the experimental, model, and parametrization uncertainties are represented by the bands of different shades.

results qualitatively support the main conclusions of the current NLO QCD analysis. Details can be found in supplemental material.

IX. SUMMARY

The $W \rightarrow \mu\nu$ lepton charge asymmetry is measured in pp collisions at $\sqrt{s} = 7 \text{ TeV}$ using a data sample corresponding to an integrated luminosity of 4.7 fb^{-1} collected with the CMS detector at the LHC (a sample of more than 20 million $W \rightarrow \mu\nu$ events). The asymmetry is measured in 11 bins in absolute muon pseudorapidity, $|\eta|$, for two different muon p_T thresholds, 25 and 35 GeV. Compared to the previous CMS measurement, this measurement

significantly reduces both the statistical and the systematic uncertainties. The total uncertainty per bin is 0.2%–0.4%. The data are in good agreement with the theoretical predictions using CT10, NNPDF2.3, and HERAPDF1.5 PDF sets. The data are in poor agreement with the prediction based on the MSTW2008 PDF set, although the agreement is significantly improved when using the MSTW2008CPdeut PDF set. The experimental uncertainties are smaller than the current PDF uncertainties in the present QCD calculations. Therefore, this measurement can be used to significantly improve the determination of PDFs in future fits.

This precise measurement of the $W \rightarrow \mu\nu$ lepton charge asymmetry and the recent CMS measurement of associated $W + \text{charm}$ production are used together with the cross sections for inclusive deep inelastic scattering at HERA in an NLO QCD analysis of the proton structure. The muon charge asymmetry in W -boson production imposes strong constraints on the valence-quark distributions, while the $W + \text{charm}$ process is directly sensitive to the strange-quark distribution.

ACKNOWLEDGMENTS

We congratulate our colleagues in the CERN accelerator departments for the excellent performance of the LHC and thank the technical and administrative staffs at CERN and at other CMS institutes for their contributions to the success of the CMS effort. In addition, we gratefully acknowledge the computing centers and personnel of the Worldwide LHC Computing Grid for delivering so effectively the computing infrastructure essential to our analyses. Finally, we acknowledge the enduring support for the construction and operation of the LHC and the CMS detector provided by the following funding agencies: the Austrian Federal Ministry of Science and Research and the Austrian Science Fund; the Belgian Fonds de la Recherche Scientifique, and Fonds voor Wetenschappelijk Onderzoek; the Brazilian Funding Agencies (CNPq, CAPES, FAPERJ, and FAPESP); the Bulgarian Ministry of Education and Science; CERN; the Chinese Academy of Sciences, Ministry of Science and Technology, and National Natural Science Foundation of China; the Colombian Funding Agency (COLCIENCIAS); the Croatian Ministry of Science, Education and Sport, and the Croatian Science Foundation; the Research Promotion Foundation, Cyprus; the Ministry of Education and Research, Recurrent financing contract SF0690030s09 and European Regional Development Fund, Estonia; the Academy of Finland, Finnish Ministry of Education and Culture, and Helsinki Institute of Physics; the Institut National de Physique Nucléaire et de Physique des Particules/CNRS, and Commissariat à l'Énergie Atomique et aux Énergies Alternatives/CEA, France; the

Bundesministerium für Bildung und Forschung, Deutsche Forschungsgemeinschaft, and Helmholtz-Gemeinschaft Deutscher Forschungszentren, Germany; the General Secretariat for Research and Technology, Greece; the National Scientific Research Foundation, and National Innovation Office, Hungary; the Department of Atomic Energy and the Department of Science and Technology, India; the Institute for Studies in Theoretical Physics and Mathematics, Iran; the Science Foundation, Ireland; the Istituto Nazionale di Fisica Nucleare, Italy; the Korean Ministry of Education, Science and Technology and the World Class University program of NRF, Republic of Korea; the Lithuanian Academy of Sciences; the Mexican Funding Agencies (CINVESTAV, CONACYT, SEP, and UASLP-FAI); the Ministry of Business, Innovation and Employment, New Zealand; the Pakistan Atomic Energy Commission; the Ministry of Science and Higher Education and the National Science Centre, Poland; the Fundação para a Ciência e a Tecnologia, Portugal; JINR, Dubna; the Ministry of Education and Science of the Russian Federation, the Federal Agency of Atomic Energy of the Russian Federation, Russian Academy of Sciences, and the Russian Foundation for Basic Research; the Ministry of Education, Science and Technological Development of Serbia; the Secretaría de Estado de Investigación, Desarrollo e Innovación and Programa Consolider-Ingenio 2010, Spain; the Swiss Funding Agencies (ETH Board, ETH Zurich, PSI, SNF, UniZH, Canton Zurich, and SER); the National Science Council, Taipei; the Thailand Center of Excellence in Physics, the Institute for the Promotion of Teaching Science and Technology of Thailand, Special Task Force for Activating Research and the National Science and Technology Development Agency of Thailand; the Scientific and Technical Research Council of Turkey, and Turkish Atomic Energy Authority; the Science and Technology Facilities Council, UK; the U.S. Department of Energy, and the U.S. National Science Foundation. Individuals have received support from the Marie-Curie programme and the European Research Council and EPLANET (European Union); the Leventis Foundation; the A. P. Sloan Foundation; the Alexander von Humboldt Foundation; the Belgian Federal Science Policy Office; the Fonds pour la Formation à la Recherche dans l'Industrie et dans l'Agriculture (FRIA-Belgium); the Agentschap voor Innovatie door Wetenschap en Technologie (IWT-Belgium); the Ministry of Education, Youth and Sports (MEYS) of Czech Republic; the Council of Science and Industrial Research, India; the Compagnia di San Paolo (Torino); the HOMING PLUS programme of Foundation for Polish Science, cofinanced by EU, Regional Development Fund; and the Thalís and Aristeia programmes cofinanced by EU-ESF and the Greek NSRF.

- [1] CMS Collaboration, *J. High Energy Phys.* **10** (2011) 132.
- [2] A. D. Martin, W. J. Stirling, R. S. Thorne, and G. Watt, *Eur. Phys. J. C* **63**, 189 (2009).
- [3] H.-L. Lai, M. Guzzi, J. Huston, Z. Li, P. M. Nadolsky, J. Pumplin, and C.-P. Yuan, *Phys. Rev. D* **82**, 074024 (2010).
- [4] J. D. Bjorken and E. A. Paschos, *Phys. Rev.* **185**, 1975 (1969).
- [5] T. Aaltonen *et al.* (CDF Collaboration), *Phys. Rev. Lett.* **102**, 181801 (2009).
- [6] V. M. Abazov *et al.* (D0 Collaboration), *Phys. Rev. Lett.* **101**, 211801 (2008).
- [7] V. M. Abazov *et al.* (D0 Collaboration), *Phys. Rev. D* **88**, 091102 (2013).
- [8] G. Aad *et al.* (ATLAS Collaboration), *Phys. Lett. B* **701**, 31 (2011).
- [9] ATLAS Collaboration, *Phys. Rev. D* **85**, 072004 (2012).
- [10] S. Chatrchyan *et al.* (CMS Collaboration), *J. High Energy Phys.* **04** (2011) 050.
- [11] R. Aaij *et al.* (LHCb Collaboration), *J. High Energy Phys.* **06** (2012) 058.
- [12] CMS Collaboration, *Phys. Rev. Lett.* **109**, 111806 (2012).
- [13] R. D. Ball, V. Bertone, F. Cerutti, L. D. Debbio, S. Forte, A. Guffanti, J. I. Latorre, J. Rojo, and M. Ubiali (NNPDF Collaboration), *Nucl. Phys.* **B849**, 112 (2011); **B855**, 927 (E) (2012).
- [14] R. D. Ball, V. Bertone, F. Cerutti, L. D. Debbio, S. Forte, A. Guffanti, N. P. Hartland, J. I. Latorre, J. Rojo, and M. Ubiali, *Nucl. Phys.* **B855**, 608 (2012).
- [15] A. D. Martin, A. J. T. M. Mathijssen, W. J. Stirling, R. S. Thorne, B. J. A. Watt, and G. Watt, *Eur. Phys. J. C* **73**, 2318 (2013).
- [16] S. Alekhin, J. Blumlein, and S. Moch, *Nucl. Phys. B, Proc. Suppl.* **222–224**, 41 (2012).
- [17] S. Alekhin, J. Bluemlein, and S. Moch, *Phys. Rev. D* **89**, 054028 (2014).
- [18] CMS Collaboration, *J. High Energy Phys.* **02** (2014) 013.
- [19] F. D. Aaron *et al.* (H1 and ZEUS Collaborations), *J. High Energy Phys.* **01** (2010) 109.
- [20] CMS Collaboration, *JINST* **3**, S08004 (2008).
- [21] CMS Collaboration, CMS Physics Analysis Summary Report No. CMS-PAS-PFT-10-003, 2010, <http://cdsweb.cern.ch/record/1279347>.
- [22] S. Frixione, P. Nason, and C. Oleari, *J. High Energy Phys.* **11** (2007) 070.
- [23] T. Sjöstrand, S. Mrenna, and P. Skands, *J. High Energy Phys.* **05** (2006) 026.
- [24] Z. Was, *Nucl. Phys. B, Proc. Suppl.* **98**, 96 (2001).
- [25] J. Pumplin, D. R. Stump, J. Huston, H.-L. Lai, P. Nadolsky, and W.-K. Tung, *J. High Energy Phys.* **07** (2002) 012.
- [26] S. Agostinelli *et al.* (GEANT4), *Nucl. Instrum. Methods Phys. Res., Sect. A* **506**, 250 (2003).
- [27] CMS Collaboration, *Phys. Lett. B* **722**, 5 (2013).
- [28] CMS Collaboration, *JINST* **7**, P10002 (2012).
- [29] A. Bodek, A. van Dyne, J. Y. Han, W. Sakumoto, and A. Strelnikov, *Eur. Phys. J. C* **72**, 2194 (2012).
- [30] R. Barlow and C. Beeston, *Comput. Phys. Commun.* **77**, 219 (1993).
- [31] J. M. Campbell and R. K. Ellis, *Phys. Rev. D* **60**, 113006 (1999).
- [32] J. M. Campbell and R. K. Ellis, *Nucl. Phys. B, Proc. Suppl.* **205–206**, 10 (2010).
- [33] J. M. Campbell and R. K. Ellis, [arXiv:1204.1513](https://arxiv.org/abs/1204.1513).
- [34] T. Aaltonen *et al.* (CDF Collaboration), *Phys. Rev. D* **77**, 112001 (2008).
- [35] V. M. Abazov *et al.* (D0 Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **609**, 250 (2009).
- [36] CMS Collaboration, CMS Physics Analysis Summary Report No. CMS-PAS-JME-10-005, 2010, <http://cdsweb.cern.ch/record/1294501>.
- [37] M. Cacciari, G. P. Salam, and G. Soyez, *J. High Energy Phys.* **04** (2008) 063.
- [38] Y. Li and F. Petriello, *Phys. Rev. D* **86**, 094034 (2012).
- [39] CMS Collaboration, CMS Physics Analysis Summary Report No. CMS-PAS-SMP-12-008, 2012, <http://cdsweb.cern.ch/record/1434360>.
- [40] J. Beringer *et al.* (Particle Data Group), *Phys. Rev. D* **86**, 010001 (2012).
- [41] M. Botje, J. Butterworth, A. Cooper-Sarkar, d. A. Roeck, J. Feltesse, S. Forte, A. Glazov, J. Huston, R. McNulty, T. Sjostrand, and R. Thorne, [arXiv:1101.0538](https://arxiv.org/abs/1101.0538).
- [42] R. D. Ball, V. Bertone, F. Cerutti, L. D. Debbio, S. Forte, A. Guffanti, J. I. Latorre, J. Rojo, and M. Ubiali (NNPDF Collaboration), *Nucl. Phys.* **B849**, 296 (2011).
- [43] J. Alwall, M. Herquet, F. Maltoni, O. Mattelaer, and T. Stelzer, *J. High Energy Phys.* **06** (2011) 128.
- [44] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevD.90.032004> for the total covariance matrix, including both statistical and systematic uncertainties, for the measurement of the W muon charge asymmetry and additional information on the QCD analysis.
- [45] R. D. Ball, V. Bertone, S. Carrazza, C. S. Deans, L. D. Debbio, S. Forte, A. Guffanti, N. P. Hartland, J. I. Latorre, J. Rojo, and M. Ubiali, *Nucl. Phys.* **B867**, 244 (2013).
- [46] V. Radescu *et al.* (H1 and ZEUS Collaborations), in *Proc. Sci., ICHEP2010*(2010) 169. Grids available at http://www.desy.de/h1zeus/combined_results/index.php?do=proton_structure.
- [47] S. Catani, L. Cieri, G. Ferrera, D. de Florian, and M. Grazzini, *Phys. Rev. Lett.* **103**, 082001 (2009).
- [48] S. Catani and M. Grazzini, *Phys. Rev. Lett.* **98**, 222002 (2007).
- [49] G. A. Ladinsky and C. P. Yuan, *Phys. Rev. D* **50**, R4239 (1994).
- [50] C. Balazs and C. P. Yuan, *Phys. Rev. D* **56**, 5558 (1997).
- [51] F. Landry, R. Brock, P. M. Nadolsky, and C. P. Yuan, *Phys. Rev. D* **67**, 073016 (2003).
- [52] M. Guzzi, P. M. Nadolsky, and B. Wang, *Phys. Rev. D* **90**, 014030 (2014).
- [53] V. N. Gribov and L. N. Lipatov, *Sov. J. Nucl. Phys.* **15**, 438 (1972).
- [54] G. Altarelli and G. Parisi, *Nucl. Phys.* **B126**, 298 (1977).
- [55] G. Curci, W. Furmanski, and R. Petronzio, *Nucl. Phys.* **B175**, 27 (1980).
- [56] W. Furmanski and R. Petronzio, *Phys. Lett.* **97B**, 437 (1980).
- [57] S. Moch, J. A. M. Vermaseren, and A. Vogt, *Nucl. Phys.* **B688**, 101 (2004).

- [58] A. Vogt, S. Moch, and J. A. M. Vermaseren, *Nucl. Phys.* **B691**, 129 (2004).
- [59] E. L. Berger, F. Halzen, C. S. Kim, and S. Willenbrock, *Phys. Rev. D* **40**, 83 (1989).
- [60] U. Baur, F. Halzen, S. Keller, M. L. Mangano, and K. Riesselmann, *Phys. Lett. B* **318**, 544 (1993).
- [61] H.-L. Lai, P. Nadolsky, J. Pumplin, D. Stump, W.-K. Tung, and C.-P. Yuan, *J. High Energy Phys.* **04** (2007) 089.
- [62] M. Goncharov *et al.* (NuTeV Collaboration), *Phys. Rev. D* **64**, 112006 (2001).
- [63] D. Mason *et al.* (NuTeV Collaboration), *Phys. Rev. Lett.* **99**, 192001 (2007).
- [64] A. O. Bazarko *et al.* (CCFR Collaboration), *Z. Phys. C* **65**, 189 (1995).
- [65] O. Samoylov *et al.* (NOMAD Collaboration), *Nucl. Phys.* **B876**, 339 (2013).
- [66] A. Airapetian *et al.* (HERMES Collaboration), *Phys. Rev. D* **87**, 074029 (2013).
- [67] H. E. Jackson, *Proc. Sci.*, DIS2013 (2013) 056.
- [68] ATLAS Collaboration, *Phys. Rev. Lett.* **109**, 012001 (2012).
- [69] T. Carli, D. Clements, A. Cooper-Sarkar, C. Gwenlan, G. P. Salam, F. Siegert, P. Starovoitov, and M. Sutton, *Eur. Phys. J. C* **66**, 503 (2010).
- [70] F. D. Aaron *et al.* (H1 Collaboration), *Eur. Phys. J. C* **64**, 561 (2009).
- [71] HERAFitter web site, <http://www.herafitter.org>.
- [72] M. Botje, *Comput. Phys. Commun.* **182**, 490 (2011).
- [73] R. S. Thorne, *Phys. Rev. D* **73**, 054019 (2006).

S. Chatrchyan,¹ V. Khachatryan,¹ A. M. Sirunyan,¹ A. Tumasyan,¹ W. Adam,² T. Bergauer,² M. Dragicevic,² J. Erö,² C. Fabjan,^{2,b} M. Friedl,² R. Frühwirth,^{2,b} V. M. Ghete,² C. Hartl,² N. Hörmann,² J. Hrubec,² M. Jeitler,^{2,b} W. Kiesenhofer,² V. Knünz,² M. Krammer,^{2,b} I. Krätschmer,² D. Liko,² I. Mikulec,² D. Rabady,^{2,c} B. Rahbaran,² H. Rohringer,² R. Schöfbeck,² J. Strauss,² A. Taurok,² W. Treberer-Treberspurg,² W. Waltenberger,² C.-E. Wulz,^{2,b} V. Mossolov,³ N. Shumeiko,³ J. Suarez Gonzalez,³ S. Alderweireldt,⁴ M. Bansal,⁴ S. Bansal,⁴ T. Cornelis,⁴ E. A. De Wolf,⁴ X. Janssen,⁴ A. Knutsson,⁴ S. Luyckx,⁴ L. Mucibello,⁴ S. Ochesanu,⁴ B. Roland,⁴ R. Rougny,⁴ H. Van Haevermaet,⁴ P. Van Mechelen,⁴ N. Van Remortel,⁴ A. Van Spilbeeck,⁴ F. Blekman,⁵ S. Blyweert,⁵ J. D'Hondt,⁵ N. Heracleous,⁵ A. Kalogeropoulos,⁵ J. Keaveney,⁵ T. J. Kim,⁵ S. Lowette,⁵ M. Maes,⁵ A. Olbrechts,⁵ D. Strom,⁵ S. Tavernier,⁵ W. Van Doninck,⁵ P. Van Mulders,⁵ G. P. Van Onsem,⁵ I. Vilella,⁶ C. Caillol,⁶ B. Clerbaux,⁶ G. De Lentdecker,⁶ L. Favart,⁶ A. P. R. Gay,⁶ A. Léonard,⁶ P. E. Marage,⁶ A. Mohammadi,⁶ L. Perniè,⁶ T. Reis,⁶ T. Seva,⁶ L. Thomas,⁶ C. Vander Velde,⁶ P. Vanlaer,⁶ J. Wang,⁶ V. Adler,⁷ K. Beernaert,⁷ L. Benucci,⁷ A. Cimmino,⁷ S. Costantini,⁷ S. Dildick,⁷ G. Garcia,⁷ B. Klein,⁷ J. Lellouch,⁷ J. Mccartin,⁷ A. A. Ocampo Rios,⁷ D. Ryckbosch,⁷ S. Salva Diblen,⁷ M. Sigamani,⁷ N. Strobbe,⁷ F. Thyssen,⁷ M. Tytgat,⁷ S. Walsh,⁷ E. Yazgan,⁷ N. Zaganidis,⁷ S. Basegmez,⁸ C. Beluffi,^{8,d} G. Bruno,⁸ R. Castello,⁸ A. Caudron,⁸ L. Ceard,⁸ G. G. Da Silveira,⁸ C. Delaere,⁸ T. du Pree,⁸ D. Favart,⁸ L. Forthomme,⁸ A. Giammanco,^{8,e} J. Hollar,⁸ P. Jez,⁸ M. Komm,⁸ V. Lemaitre,⁸ J. Liao,⁸ O. Militaru,⁸ C. Nuttens,⁸ D. Pagano,⁸ A. Pin,⁸ K. Piotrkowski,⁸ A. Popov,^{8,f} L. Quertenmont,⁸ M. Selvaggi,⁸ M. Vidal Marono,⁸ J. M. Vizan Garcia,⁸ N. Bely,⁹ T. Caebergs,⁹ E. Daubie,⁹ G. H. Hammad,⁹ G. A. Alves,¹⁰ M. Correa Martins Junior,¹⁰ T. Martins,¹⁰ M. E. Pol,¹⁰ M. H. G. Souza,¹⁰ W. L. Aldá Júnior,¹¹ W. Carvalho,¹¹ J. Chinellato,^{11,g} A. Custódio,¹¹ E. M. Da Costa,¹¹ D. De Jesus Damiao,¹¹ C. De Oliveira Martins,¹¹ S. Fonseca De Souza,¹¹ H. Malbouisson,¹¹ M. Malek,¹¹ D. Matos Figueiredo,¹¹ L. Mundim,¹¹ H. Nogima,¹¹ W. L. Prado Da Silva,¹¹ J. Santaolalla,¹¹ A. Santoro,¹¹ A. Sznajder,¹¹ E. J. Tonelli Manganote,^{11,g} A. Vilela Pereira,¹¹ C. A. Bernardes,^{12b} F. A. Dias,^{12a,h} T. R. Fernandez Perez Tomei,^{12a} E. M. Gregores,^{12b} P. G. Mercadante,^{12b} S. F. Novaes,^{12a} S. S. Padula,^{12a} V. Genchev,^{13,c} P. Iaydjiev,^{13,c} A. Marinov,¹³ S. Piperov,¹³ M. Rodozov,¹³ G. Sultanov,¹³ M. Vutova,¹³ A. Dimitrov,¹⁴ I. Glushkov,¹⁴ R. Hadjiiska,¹⁴ V. Kozuharov,¹⁴ L. Litov,¹⁴ B. Pavlov,¹⁴ P. Petkov,¹⁴ J. G. Bian,¹⁵ G. M. Chen,¹⁵ H. S. Chen,¹⁵ M. Chen,¹⁵ R. Du,¹⁵ C. H. Jiang,¹⁵ D. Liang,¹⁵ S. Liang,¹⁵ X. Meng,¹⁵ R. Plestina,^{15,i} J. Tao,¹⁵ X. Wang,¹⁵ Z. Wang,¹⁵ C. Asawatangtrakuldee,¹⁶ Y. Ban,¹⁶ Y. Guo,¹⁶ Q. Li,¹⁶ W. Li,¹⁶ S. Liu,¹⁶ Y. Mao,¹⁶ S. J. Qian,¹⁶ D. Wang,¹⁶ L. Zhang,¹⁶ W. Zou,¹⁶ C. Avila,¹⁷ C. A. Carrillo Montoya,¹⁷ L. F. Chaparro Sierra,¹⁷ C. Florez,¹⁷ J. P. Gomez,¹⁷ B. Gomez Moreno,¹⁷ J. C. Sanabria,¹⁷ N. Godinovic,¹⁸ D. Lelas,¹⁸ D. Polic,¹⁸ I. Puljak,¹⁸ Z. Antunovic,¹⁹ M. Kovac,¹⁹ V. Brigljevic,²⁰ K. Kadija,²⁰ J. Luetic,²⁰ D. Mekterovic,²⁰ S. Morovic,²⁰ L. Tikvica,²⁰ A. Attikis,²¹ G. Mavromanolakis,²¹ J. Mousa,²¹ C. Nicolaou,²¹ F. Ptochos,²¹ P. A. Razis,²¹ M. Finger,²² M. Finger Jr.,²² A. A. Abdelalim,^{23,j} Y. Assran,^{23,k} S. Elgammal,^{23,j} A. Ellithi Kamel,^{23,l} M. A. Mahmoud,^{23,m} A. Radi,^{23,n,o} M. Kadastik,²⁴ M. Müntel,²⁴ M. Murumaa,²⁴ M. Raidal,²⁴ L. Rebane,²⁴ A. Tiko,²⁴ P. Eerola,²⁵ G. Fedi,²⁵ M. Voutilainen,²⁵ J. Härkönen,²⁶ V. Karimäki,²⁶ R. Kinnunen,²⁶ M. J. Kortelainen,²⁶ T. Lampén,²⁶ K. Lassila-Perini,²⁶ S. Lehti,²⁶ T. Lindén,²⁶ P. Luukka,²⁶ T. Mäenpää,²⁶ T. Peltola,²⁶ E. Tuominen,²⁶ J. Tuominiemi,²⁶ E. Tuovinen,²⁶ L. Wendland,²⁶ T. Tuuva,²⁷ M. Besancon,²⁸ F. Couderc,²⁸ M. Dejardin,²⁸ D. Denegri,²⁸ B. Fabbro,²⁸ J. L. Faure,²⁸ F. Ferri,²⁸ S. Ganjour,²⁸

A. Givernaud,²⁸ P. Gras,²⁸ G. Hamel de Monchenault,²⁸ P. Jarry,²⁸ E. Locci,²⁸ J. Malcles,²⁸ A. Nayak,²⁸ J. Rander,²⁸ A. Rosowsky,²⁸ M. Titov,²⁸ S. Baffioni,²⁹ F. Beaudette,²⁹ P. Busson,²⁹ C. Charlot,²⁹ N. Daci,²⁹ T. Dahms,²⁹ M. Dalchenko,²⁹ L. Dobrzynski,²⁹ A. Florent,²⁹ R. Granier de Cassagnac,²⁹ P. Miné,²⁹ C. Mironov,²⁹ I. N. Naranjo,²⁹ M. Nguyen,²⁹ C. Ochando,²⁹ P. Paganini,²⁹ D. Sabes,²⁹ R. Salerno,²⁹ J. Sauvan,²⁹ Y. Sirois,²⁹ C. Veelken,²⁹ Y. Yilmaz,²⁹ A. Zabi,²⁹ J.-L. Agram,^{30,p} J. Andrea,³⁰ D. Bloch,³⁰ J.-M. Brom,³⁰ E. C. Chabert,³⁰ C. Collard,³⁰ E. Conte,^{30,p} F. Drouhin,^{30,p} J.-C. Fontaine,^{30,p} D. Gelé,³⁰ U. Goerlach,³⁰ C. Goetzmann,³⁰ P. Juillot,³⁰ A.-C. Le Bihan,³⁰ P. Van Hove,³⁰ S. Gadrat,³¹ S. Beauceron,³² N. Beaupere,³² G. Boudoul,³² S. Brochet,³² J. Chasserat,³² R. Chierici,³² D. Contardo,^{32,c} P. Depasse,³² H. El Mamouni,³² J. Fan,³² J. Fay,³² S. Gascon,³² M. Gouzevitch,³² B. Ille,³² T. Kurca,³² M. Lethuillier,³² L. Mirabito,³² S. Perries,³² J. D. Ruiz Alvarez,³² L. Sgandurra,³² V. Sordini,³² M. Vander Donckt,³² P. Verdier,³² S. Viret,³² H. Xiao,³² Z. Tsamalaidze,^{33,q} C. Autermann,³⁴ S. Beranek,³⁴ M. Bontenackels,³⁴ B. Calpas,³⁴ M. Edelhoff,³⁴ L. Feld,³⁴ O. Hindrichs,³⁴ K. Klein,³⁴ A. Ostapchuk,³⁴ A. Perieanu,³⁴ F. Raupach,³⁴ J. Sammet,³⁴ S. Schael,³⁴ D. Sprenger,³⁴ H. Weber,³⁴ B. Wittmer,³⁴ V. Zhukov,^{34,f} M. Ata,³⁵ J. Caudron,³⁵ E. Dietz-Laursonn,³⁵ D. Duchardt,³⁵ M. Erdmann,³⁵ R. Fischer,³⁵ A. Güth,³⁵ T. Hebbeker,³⁵ C. Heidemann,³⁵ K. Hoepfner,³⁵ D. Klingebiel,³⁵ S. Knutzen,³⁵ P. Kreuzer,³⁵ M. Merschmeyer,³⁵ A. Meyer,³⁵ M. Olschewski,³⁵ K. Padeken,³⁵ P. Papacz,³⁵ H. Reithler,³⁵ S. A. Schmitz,³⁵ L. Sonnenschein,³⁵ D. Teyssier,³⁵ S. Thüer,³⁵ M. Weber,³⁵ V. Cherepanov,³⁶ Y. Erdogan,³⁶ G. Flügge,³⁶ H. Geenen,³⁶ M. Geisler,³⁶ W. Haj Ahmad,³⁶ F. Hoehle,³⁶ B. Kargoll,³⁶ T. Kress,³⁶ Y. Kuessel,³⁶ J. Lingemann,^{36,c} A. Nowack,³⁶ I. M. Nugent,³⁶ L. Perchalla,³⁶ O. Pooth,³⁶ A. Stahl,³⁶ I. Asin,³⁷ N. Bartosik,³⁷ J. Behr,³⁷ W. Behrenhoff,³⁷ U. Behrens,³⁷ A. J. Bell,³⁷ M. Bergholz,^{37,r} A. Bethani,³⁷ K. Borras,³⁷ A. Burgmeier,³⁷ A. Cakir,³⁷ L. Calligaris,³⁷ A. Campbell,³⁷ S. Choudhury,³⁷ F. Costanza,³⁷ C. Diez Pardos,³⁷ S. Dooling,³⁷ T. Dorland,³⁷ G. Eckerlin,³⁷ D. Eckstein,³⁷ T. Eichhorn,³⁷ G. Flucke,³⁷ A. Geiser,³⁷ A. Grebenyuk,³⁷ P. Gunnellini,³⁷ M. Guzzi,³⁷ S. Habib,³⁷ J. Hauk,³⁷ G. Hellwig,³⁷ M. Hempel,³⁷ D. Horton,³⁷ H. Jung,³⁷ M. Kasemann,³⁷ P. Katsas,³⁷ J. Kieseler,³⁷ C. Kleinwort,³⁷ M. Krämer,³⁷ D. Krücker,³⁷ W. Lange,³⁷ J. Leonard,³⁷ K. Lipka,³⁷ W. Lohmann,^{37,r} B. Lutz,³⁷ R. Mankel,³⁷ I. Marfin,³⁷ I.-A. Melzer-Pellmann,³⁷ A. B. Meyer,³⁷ J. Mnich,³⁷ A. Mussgiller,³⁷ S. Naumann-Emme,³⁷ O. Novgorodova,³⁷ F. Nowak,³⁷ H. Perrey,³⁷ A. Petrukhin,³⁷ D. Pitzl,³⁷ R. Placakyte,³⁷ A. Raspereza,³⁷ P. M. Ribeiro Cipriano,³⁷ C. Riedl,³⁷ E. Ron,³⁷ M. O. Sahin,³⁷ J. Salfeld-Nebgen,³⁷ P. Saxena,³⁷ R. Schmidt,^{37,r} T. Schoerner-Sadenius,³⁷ M. Schröder,³⁷ M. Stein,³⁷ A. D. R. Vargas Trevino,³⁷ R. Walsh,³⁷ C. Wissing,³⁷ M. Aldaya Martin,³⁸ V. Blobel,³⁸ H. Enderle,³⁸ J. Erfle,³⁸ E. Garutti,³⁸ K. Goebel,³⁸ M. Görner,³⁸ M. Gosselink,³⁸ J. Haller,³⁸ R. S. Höing,³⁸ H. Kirschenmann,³⁸ R. Klanner,³⁸ R. Kogler,³⁸ J. Lange,³⁸ T. Lapsien,³⁸ T. Lenz,³⁸ I. Marchesini,³⁸ J. Ott,³⁸ T. Peiffer,³⁸ N. Pietsch,³⁸ D. Rathjens,³⁸ C. Sander,³⁸ H. Schettler,³⁸ P. Schleper,³⁸ E. Schlieckau,³⁸ A. Schmidt,³⁸ M. Seidel,³⁸ J. Sibille,^{38,s} V. Sola,³⁸ H. Stadie,³⁸ G. Steinbrück,³⁸ D. Troendle,³⁸ E. Usai,³⁸ L. Vanelderden,³⁸ C. Barth,³⁹ C. Baus,³⁹ J. Berger,³⁹ C. Böser,³⁹ E. Butz,³⁹ T. Chwalek,³⁹ W. De Boer,³⁹ A. Descroix,³⁹ A. Dierlamm,³⁹ M. Feindt,³⁹ M. Guthoff,^{39,c} F. Hartmann,^{39,c} T. Hauth,^{39,c} H. Held,³⁹ K. H. Hoffmann,³⁹ U. Husemann,³⁹ I. Katkov,^{39,f} A. Kornmayer,^{39,c} E. Kuznetsova,³⁹ P. Lobelle Pardo,³⁹ D. Martschei,³⁹ M. U. Mozer,³⁹ T. Müller,³⁹ M. Niegel,³⁹ A. Nürnberg,³⁹ O. Oberst,³⁹ G. Quast,³⁹ K. Rabbertz,³⁹ F. Ratnikov,³⁹ S. Röcker,³⁹ F.-P. Schilling,³⁹ G. Schott,³⁹ H. J. Simonis,³⁹ F. M. Stober,³⁹ R. Ulrich,³⁹ J. Wagner-Kuhr,³⁹ S. Wayand,³⁹ T. Weiler,³⁹ R. Wolf,³⁹ M. Zeise,³⁹ G. Anagnostou,⁴⁰ G. Daskalakis,⁴⁰ T. Geralis,⁴⁰ S. Kesisoglou,⁴⁰ A. Kyriakis,⁴⁰ D. Loukas,⁴⁰ A. Markou,⁴⁰ C. Markou,⁴⁰ E. Ntomari,⁴⁰ A. Psallidas,⁴⁰ I. Topsis-giotis,⁴⁰ L. Gouskos,⁴¹ A. Panagiotou,⁴¹ N. Saoulidou,⁴¹ E. Stiliaris,⁴¹ X. Aslanoglou,⁴² I. Evangelou,⁴² G. Flouris,⁴² C. Foudas,⁴² J. Jones,⁴² P. Kokkas,⁴² N. Manthos,⁴² I. Papadopoulos,⁴² E. Paradas,⁴² G. Bencze,⁴³ C. Hajdu,⁴³ P. Hidas,⁴³ D. Horvath,^{43,t} F. Sikler,⁴³ V. Veszpremi,⁴³ G. Vesztergombi,^{43,u} A. J. Zsigmond,⁴³ N. Beni,⁴⁴ S. Czellar,⁴⁴ J. Molnar,⁴⁴ J. Palinkas,⁴⁴ Z. Szillasi,⁴⁴ J. Karancsi,⁴⁵ P. Raics,⁴⁵ Z. L. Trocsanyi,⁴⁵ B. Ujvari,⁴⁵ S. K. Swain,⁴⁶ S. B. Beri,⁴⁷ V. Bhatnagar,⁴⁷ N. Dhingra,⁴⁷ R. Gupta,⁴⁷ M. Kaur,⁴⁷ M. Z. Mehta,⁴⁷ M. Mittal,⁴⁷ N. Nishu,⁴⁷ A. Sharma,⁴⁷ J. B. Singh,⁴⁷ A. Kumar,⁴⁸ A. Kumar,⁴⁸ S. Ahuja,⁴⁸ A. Bhardwaj,⁴⁸ B. C. Choudhary,⁴⁸ A. Kumar,⁴⁸ S. Malhotra,⁴⁸ M. Naimuddin,⁴⁸ K. Ranjan,⁴⁸ V. Sharma,⁴⁸ R. K. Shivpuri,⁴⁸ S. Banerjee,⁴⁹ S. Bhattacharya,⁴⁹ K. Chatterjee,⁴⁹ S. Dutta,⁴⁹ B. Gomber,⁴⁹ S. Jain,⁴⁹ S. Jain,⁴⁹ R. Khurana,⁴⁹ A. Modak,⁴⁹ S. Mukherjee,⁴⁹ D. Roy,⁴⁹ S. Sarkar,⁴⁹ M. Sharan,⁴⁹ A. P. Singh,⁴⁹ A. Abdulsalam,⁵⁰ D. Dutta,⁵⁰ S. Kailas,⁵⁰ V. Kumar,⁵⁰ A. K. Mohanty,^{50,c} L. M. Pant,⁵⁰ P. Shukla,⁵⁰ A. Topkar,⁵⁰ T. Aziz,⁵¹ R. M. Chatterjee,⁵¹ S. Ganguly,⁵¹ S. Ghosh,⁵¹ M. Guchait,^{51,v} A. Gurtu,^{51,w} G. Kole,⁵¹ S. Kumar,⁵¹ M. Maity,^{51,x} G. Majumder,⁵¹ K. Mazumdar,⁵¹ G. B. Mohanty,⁵¹ B. Parida,⁵¹ K. Sudhakar,⁵¹ N. Wickramage,^{51,y} S. Banerjee,⁵² S. Dugad,⁵² H. Arfaei,⁵³ H. Bakhshiansohi,⁵³ H. Behnamian,⁵³ S. M. Etesami,^{53,z} A. Fahim,^{53,aa} A. Jafari,⁵³ M. Khakzad,⁵³ M. Mohammadi Najafabadi,⁵³ M. Naseri,⁵³ S. Paktinat Mehdiabadi,⁵³ B. Safarzadeh,^{53,bb} M. Zeinali,⁵³ M. Grunewald,⁵⁴ M. Abbrescia,^{55a,55b} L. Barbone,^{55a,55b} C. Calabria,^{55a,55b} S. S. Chhibra,^{55a,55b} A. Colaleo,^{55a} D. Creanza,^{55a,55c}

N. De Filippis,^{55a,55c} M. De Palma,^{55a,55b} L. Fiore,^{55a} G. Iaselli,^{55a,55c} G. Maggi,^{55a,55c} M. Maggi,^{55a} B. Marangelli,^{55a,55b}
 S. My,^{55a,55c} S. Nuzzo,^{55a,55b} N. Pacifico,^{55a} A. Pompili,^{55a,55b} G. Pugliese,^{55a,55c} R. Radogna,^{55a,55b} G. Selvaggi,^{55a,55b}
 L. Silvestris,^{55a} G. Singh,^{55a,55b} R. Venditti,^{55a,55b} P. Verwilligen,^{55a} G. Zito,^{55a} G. Abbiendi,^{56a} A. C. Benvenuti,^{56a}
 D. Bonacorsi,^{56a,56b} S. Braibant-Giacomelli,^{56a,56b} L. Brigliadori,^{56a,56b} R. Campanini,^{56a,56b} P. Capiluppi,^{56a,56b}
 A. Castro,^{56a,56b} F. R. Cavallo,^{56a} G. Codispoti,^{56a,56b} M. Cuffiani,^{56a,56b} G. M. Dallavalle,^{56a} F. Fabbri,^{56a} A. Fanfani,^{56a,56b}
 D. Fasanella,^{56a,56b} P. Giacomelli,^{56a} C. Grandi,^{56a} L. Guiducci,^{56a,56b} S. Marcellini,^{56a} G. Masetti,^{56a} M. Meneghelli,^{56a,56b}
 A. Montanari,^{56a} F. L. Navarria,^{56a,56b} F. Odorici,^{56a} A. Perrotta,^{56a} F. Primavera,^{56a,56b} A. M. Rossi,^{56a,56b} T. Rovelli,^{56a,56b}
 G. P. Siroli,^{56a,56b} N. Tosi,^{56a,56b} R. Travaglini,^{56a,56b} S. Albergo,^{57a,57b} G. Cappello,^{57a} M. Chiorboli,^{57a,57b} S. Costa,^{57a,57b}
 F. Giordano,^{57a,57c} R. Potenza,^{57a,57b} A. Tricomi,^{57a,57b} C. Tuve,^{57a,57b} G. Barbagli,^{58a} V. Ciulli,^{58a,58b} C. Civinini,^{58a}
 R. D'Alessandro,^{58a,58b} E. Focardi,^{58a,58b} E. Gallo,^{58a} S. Gonzi,^{58a,58b} V. Gori,^{58a,58b} P. Lenzi,^{58a,58b} M. Meschini,^{58a}
 S. Paoletti,^{58a} G. Sguazzoni,^{58a} A. Tropiano,^{58a,58b} L. Benussi,⁵⁹ S. Bianco,⁵⁹ F. Fabbri,⁵⁹ D. Piccolo,⁵⁹ P. Fabbriatore,^{60a}
 R. Ferretti,^{60a,60b} F. Ferro,^{60a} M. Lo Vetere,^{60a,60b} R. Musenich,^{60a} E. Robutti,^{60a} S. Tosi,^{60a,60b} A. Benaglia,^{61a}
 M. E. Dinardo,^{61a,61b} S. Fiorendi,^{61a,61b,c} S. Gennai,^{61a} R. Gerosa,^{61a} A. Ghezzi,^{61a,61b} P. Govoni,^{61a,61b} M. T. Lucchini,^{61a,61b,c}
 S. Malvezzi,^{61a} R. A. Manzoni,^{61a,61b,c} A. Martelli,^{61a,61b,c} B. Marzocchi,^{61a} D. Menasce,^{61a} L. Moroni,^{61a} M. Paganoni,^{61a,61b}
 D. Pedrini,^{61a} S. Ragazzi,^{61a,61b} N. Redaelli,^{61a} T. Tabarelli de Fatis,^{61a,61b} S. Buontempo,^{62a} N. Cavallo,^{62a,62c}
 S. Di Guida,^{62a,62d} F. Fabozzi,^{62a,62c} A. O. M. Iorio,^{62a,62b} L. Lista,^{62a} S. Meola,^{62a,62d,c} M. Merola,^{62a} P. Paolucci,^{62a,c}
 P. Azzi,^{63a} N. Bacchetta,^{63a} D. Bisello,^{63a,63b} A. Branca,^{63a,63b} R. Carlin,^{63a,63b} P. Checchia,^{63a} T. Dorigo,^{63a} U. Dosselli,^{63a}
 M. Galanti,^{63a,63b,c} F. Gasparini,^{63a,63b} U. Gasparini,^{63a,63b} P. Giubilato,^{63a,63b} A. Gozzelino,^{63a} K. Kanishchev,^{63a,63c}
 S. Lacaprara,^{63a} I. Lazzizzera,^{63a,63c} M. Margoni,^{63a,63b} A. T. Meneguzzo,^{63a,63b} J. Pazzini,^{63a,63b} N. Pozzobon,^{63a,63b}
 P. Ronchese,^{63a,63b} E. Torassa,^{63a} M. Tosi,^{63a,63b} A. Triossi,^{63a} S. Vanini,^{63a,63b} S. Ventura,^{63a} P. Zotto,^{63a,63b}
 A. Zucchetta,^{63a,63b} G. Zumerle,^{63a,63b} M. Gabusi,^{64a,64b} S. P. Ratti,^{64a,64b} C. Riccardi,^{64a,64b} P. Vitulo,^{64a,64b} M. Biasini,^{65a,65b}
 G. M. Bilei,^{65a} L. Fanò,^{65a,65b} P. Lariccia,^{65a,65b} G. Mantovani,^{65a,65b} M. Menichelli,^{65a} F. Romeo,^{65a,65b} A. Saha,^{65a}
 A. Santocchia,^{65a,65b} A. Spiezia,^{65a,65b} K. Androsov,^{66a,cc} P. Azzurri,^{66a} G. Bagliesi,^{66a} J. Bernardini,^{66a} T. Boccali,^{66a}
 G. Broccolo,^{66a,66c} R. Castaldi,^{66a} M. A. Ciocci,^{66a,cc} R. Dell'Orso,^{66a} F. Fiori,^{66a,66c} L. Foà,^{66a,66c} A. Giassi,^{66a}
 M. T. Grippo,^{66a,cc} A. Kraan,^{66a} F. Ligabue,^{66a,66c} T. Lomtadze,^{66a} L. Martini,^{66a,66b} A. Messineo,^{66a,66b} C. S. Moon,^{66a,dd}
 F. Palla,^{66a} A. Rizzi,^{66a,66b} A. Savoy-Navarro,^{66a,ee} A. T. Serban,^{66a} P. Spagnolo,^{66a} P. Squillacioti,^{66a,cc} R. Tenchini,^{66a}
 G. Tonelli,^{66a,66b} A. Venturi,^{66a} P. G. Verdini,^{66a} C. Vernieri,^{66a,66c} L. Barone,^{67a,67b} F. Cavallari,^{67a} D. Del Re,^{67a,67b}
 M. Diemoz,^{67a} M. Grassi,^{67a,67b} C. Jorda,^{67a} E. Longo,^{67a,67b} F. Margaroli,^{67a,67b} P. Meridiani,^{67a} F. Micheli,^{67a,67b}
 S. Nourbakhsh,^{67a,67b} G. Organtini,^{67a,67b} R. Paramatti,^{67a} S. Rahatlou,^{67a,67b} C. Rovelli,^{67a} L. Soffi,^{67a,67b} P. Traczyk,^{67a,67b}
 N. Amapane,^{68a,68b} R. Arcidiacono,^{68a,68c} S. Argiro,^{68a,68b} M. Arneodo,^{68a,68c} R. Bellan,^{68a,68b} C. Biino,^{68a} N. Cartiglia,^{68a}
 S. Casasso,^{68a,68b} M. Costa,^{68a,68b} A. Degano,^{68a,68b} N. Demaria,^{68a} C. Mariotti,^{68a} S. Maselli,^{68a} E. Migliore,^{68a,68b}
 V. Monaco,^{68a,68b} M. Musich,^{68a} M. M. Obertino,^{68a,68c} G. Ortona,^{68a,68b} L. Pacher,^{68a,68b} N. Pastrone,^{68a} M. Pelliccioni,^{68a,c}
 A. Potenza,^{68a,68b} A. Romero,^{68a,68b} M. Ruspa,^{68a,68c} R. Sacchi,^{68a,68b} A. Solano,^{68a,68b} A. Staiano,^{68a} U. Tamponi,^{68a}
 S. Belforte,^{69a} V. Candelise,^{69a,69b} M. Casarsa,^{69a} F. Cossutti,^{69a} G. Della Ricca,^{69a,69b} B. Gobbo,^{69a} C. La Licata,^{69a,69b}
 M. Marone,^{69a,69b} D. Montanino,^{69a,69b} A. Penzo,^{69a} A. Schizzi,^{69a,69b} T. Umer,^{69a,69b} A. Zanetti,^{69a} S. Chang,⁷⁰ T. Y. Kim,⁷⁰
 S. K. Nam,⁷⁰ D. H. Kim,⁷¹ G. N. Kim,⁷¹ J. E. Kim,⁷¹ M. S. Kim,⁷¹ D. J. Kong,⁷¹ S. Lee,⁷¹ Y. D. Oh,⁷¹ H. Park,⁷¹ D. C. Son,⁷¹
 J. Y. Kim,⁷² Z. J. Kim,⁷² S. Song,⁷² S. Choi,⁷³ D. Gyun,⁷³ B. Hong,⁷³ M. Jo,⁷³ H. Kim,⁷³ Y. Kim,⁷³ K. S. Lee,⁷³ S. K. Park,⁷³
 Y. Roh,⁷³ M. Choi,⁷⁴ J. H. Kim,⁷⁴ C. Park,⁷⁴ I. C. Park,⁷⁴ S. Park,⁷⁴ G. Ryu,⁷⁴ Y. Choi,⁷⁵ Y. K. Choi,⁷⁵ J. Goh,⁷⁵ E. Kwon,⁷⁵
 B. Lee,⁷⁵ J. Lee,⁷⁵ H. Seo,⁷⁵ I. Yu,⁷⁵ A. Juodagalvis,⁷⁶ J. R. Komaragiri,⁷⁷ H. Castilla-Valdez,⁷⁸ E. De La Cruz-Burelo,⁷⁸
 I. Heredia-de La Cruz,^{78,ff} R. Lopez-Fernandez,⁷⁸ J. Martínez-Ortega,⁷⁸ A. Sanchez-Hernandez,⁷⁸
 L. M. Villasenor-Cendejas,⁷⁸ S. Carrillo Moreno,⁷⁹ F. Vazquez Valencia,⁷⁹ H. A. Salazar Ibarguen,⁸⁰ E. Casimiro Linares,⁸¹
 A. Morelos Pineda,⁸¹ D. Krofcheck,⁸² P. H. Butler,⁸³ R. Doesburg,⁸³ S. Reucroft,⁸³ M. Ahmad,⁸⁴ M. I. Asghar,⁸⁴ J. Butt,⁸⁴
 H. R. Hoorani,⁸⁴ S. Khalid,⁸⁴ W. A. Khan,⁸⁴ T. Khurshid,⁸⁴ S. Qazi,⁸⁴ M. A. Shah,⁸⁴ M. Shoaib,⁸⁴ H. Bialkowska,⁸⁵
 M. Bluj,^{85,gg} B. Boimska,⁸⁵ T. Frueboes,⁸⁵ M. Górski,⁸⁵ M. Kazana,⁸⁵ K. Nawrocki,⁸⁵ K. Romanowska-Rybinska,⁸⁵
 M. Szleper,⁸⁵ G. Wrochna,⁸⁵ P. Zalewski,⁸⁵ G. Brona,⁸⁶ K. Bunkowski,⁸⁶ M. Cwiok,⁸⁶ W. Dominik,⁸⁶ K. Doroba,⁸⁶
 A. Kalinowski,⁸⁶ M. Konecki,⁸⁶ J. Krolikowski,⁸⁶ M. Misiura,⁸⁶ W. Wolszczak,⁸⁶ P. Bargassa,⁸⁷
 C. Beirão Da Cruz E Silva,⁸⁷ P. Faccioli,⁸⁷ P. G. Ferreira Parracho,⁸⁷ M. Gallinaro,⁸⁷ F. Nguyen,⁸⁷ J. Rodrigues Antunes,⁸⁷
 J. Seixas,^{87,c} J. Varela,⁸⁷ P. Vischia,⁸⁷ P. Bunin,⁸⁸ I. Golutvin,⁸⁸ I. Gorbunov,⁸⁸ A. Kamenev,⁸⁸ V. Karjavin,⁸⁸
 V. Konoplyanikov,⁸⁸ G. Kozlov,⁸⁸ A. Lanev,⁸⁸ A. Malakhov,⁸⁸ V. Matveev,^{88,hh} P. Moiseenz,⁸⁸ V. Palichik,⁸⁸ V. Perelygin,⁸⁸

S. Shmatov,⁸⁸ S. Shulha,⁸⁸ N. Skatchkov,⁸⁸ V. Smirnov,⁸⁸ A. Zarubin,⁸⁸ V. Golovtsov,⁸⁹ Y. Ivanov,⁸⁹ V. Kim,⁸⁹ P. Levchenko,⁸⁹ V. Murzin,⁸⁹ V. Oreshkin,⁸⁹ I. Smirnov,⁸⁹ V. Sulimov,⁸⁹ L. Uvarov,⁸⁹ S. Vavilov,⁸⁹ A. Vorobyev,⁸⁹ A. Vorobyev,⁸⁹ Y. Andreev,⁹⁰ A. Dermenev,⁹⁰ S. Gninenko,⁹⁰ N. Golubev,⁹⁰ M. Kirsanov,⁹⁰ N. Krasnikov,⁹⁰ A. Pashenkov,⁹⁰ D. Tlisov,⁹⁰ A. Toropin,⁹⁰ V. Epshteyn,⁹¹ V. Gavrilov,⁹¹ N. Lychkovskaya,⁹¹ V. Popov,⁹¹ G. Safronov,⁹¹ S. Semenov,⁹¹ A. Spiridonov,⁹¹ V. Stolin,⁹¹ E. Vlasov,⁹¹ A. Zhokin,⁹¹ V. Andreev,⁹² M. Azarkin,⁹² I. Dremin,⁹² M. Kirakosyan,⁹² A. Leonidov,⁹² G. Mesyats,⁹² S. V. Rusakov,⁹² A. Vinogradov,⁹² A. Belyaev,⁹³ E. Boos,⁹³ V. Bunichev,⁹³ M. Dubinin,^{93,h} L. Dudko,⁹³ A. Ershov,⁹³ A. Gribushin,⁹³ V. Klyukhin,⁹³ O. Kodolova,⁹³ I. Lokhtin,⁹³ S. Obraztsov,⁹³ V. Savrin,⁹³ A. Snigirev,⁹³ I. Azhgirey,⁹⁴ I. Bayshev,⁹⁴ S. Bitioukov,⁹⁴ V. Kachanov,⁹⁴ A. Kalinin,⁹⁴ D. Konstantinov,⁹⁴ V. Krychkin,⁹⁴ V. Petrov,⁹⁴ R. Ryutin,⁹⁴ A. Sobol,⁹⁴ L. Tourtchanovitch,⁹⁴ S. Troshin,⁹⁴ N. Tyurin,⁹⁴ A. Uzunian,⁹⁴ A. Volkov,⁹⁴ P. Adzic,^{95,ii} M. Djordjevic,⁹⁵ M. Ekmedzic,⁹⁵ J. Milosevic,⁹⁵ M. Aguilar-Benitez,⁹⁶ J. Alcaraz Maestre,⁹⁶ C. Battilana,⁹⁶ E. Calvo,⁹⁶ M. Cerrada,⁹⁶ M. Chamizo Llatas,^{96,c} N. Colino,⁹⁶ B. De La Cruz,⁹⁶ A. Delgado Peris,⁹⁶ D. Domínguez Vázquez,⁹⁶ C. Fernandez Bedoya,⁹⁶ J. P. Fernández Ramos,⁹⁶ A. Ferrando,⁹⁶ J. Flix,⁹⁶ M. C. Fouz,⁹⁶ P. Garcia-Abia,⁹⁶ O. Gonzalez Lopez,⁹⁶ S. Goy Lopez,⁹⁶ J. M. Hernandez,⁹⁶ M. I. Josa,⁹⁶ G. Merino,⁹⁶ E. Navarro De Martino,⁹⁶ J. Puerta Pelayo,⁹⁶ A. Quintario Olmeda,⁹⁶ I. Redondo,⁹⁶ L. Romero,⁹⁶ M. S. Soares,⁹⁶ C. Willmott,⁹⁶ C. Albajar,⁹⁷ J. F. de Trocóniz,⁹⁷ M. Missiroli,⁹⁷ H. Brun,⁹⁸ J. Cuevas,⁹⁸ J. Fernandez Menendez,⁹⁸ S. Folgueras,⁹⁸ I. Gonzalez Caballero,⁹⁸ L. Lloret Iglesias,⁹⁸ J. A. Brochero Cifuentes,⁹⁹ I. J. Cabrillo,⁹⁹ A. Calderon,⁹⁹ J. Duarte Campderros,⁹⁹ M. Fernandez,⁹⁹ G. Gomez,⁹⁹ J. Gonzalez Sanchez,⁹⁹ A. Graziano,⁹⁹ A. Lopez Virto,⁹⁹ J. Marco,⁹⁹ R. Marco,⁹⁹ C. Martinez Rivero,⁹⁹ F. Matorras,⁹⁹ F. J. Munoz Sanchez,⁹⁹ J. Piedra Gomez,⁹⁹ T. Rodrigo,⁹⁹ A. Y. Rodríguez-Marrero,⁹⁹ A. Ruiz-Jimeno,⁹⁹ L. Scodellaro,⁹⁹ I. Vila,⁹⁹ R. Vilar Cortabitarte,⁹⁹ D. Abbaneo,¹⁰⁰ E. Auffray,¹⁰⁰ G. Auzinger,¹⁰⁰ M. Bachtis,¹⁰⁰ P. Baillon,¹⁰⁰ A. H. Ball,¹⁰⁰ D. Barney,¹⁰⁰ J. Bendavid,¹⁰⁰ L. Benhabib,¹⁰⁰ J. F. Benitez,¹⁰⁰ C. Bernet,^{100,i} G. Bianchi,¹⁰⁰ P. Bloch,¹⁰⁰ A. Bocci,¹⁰⁰ A. Bonato,¹⁰⁰ O. Bondu,¹⁰⁰ C. Botta,¹⁰⁰ H. Breuker,¹⁰⁰ T. Camporesi,¹⁰⁰ G. Cerminara,¹⁰⁰ T. Christiansen,¹⁰⁰ J. A. Coarasa Perez,¹⁰⁰ S. Colafranceschi,^{100,jj} M. D'Alfonso,¹⁰⁰ D. d'Enterria,¹⁰⁰ A. Dabrowski,¹⁰⁰ A. David,¹⁰⁰ F. De Guio,¹⁰⁰ A. De Roeck,¹⁰⁰ S. De Visscher,¹⁰⁰ M. Dobson,¹⁰⁰ N. Dupont-Sagorin,¹⁰⁰ A. Elliott-Peisert,¹⁰⁰ J. Eugster,¹⁰⁰ G. Franzoni,¹⁰⁰ W. Funk,¹⁰⁰ M. Giffels,¹⁰⁰ D. Gigi,¹⁰⁰ K. Gill,¹⁰⁰ M. Girone,¹⁰⁰ M. Giunta,¹⁰⁰ F. Glege,¹⁰⁰ R. Gomez-Reino Garrido,¹⁰⁰ S. Gowdy,¹⁰⁰ R. Guida,¹⁰⁰ J. Hammer,¹⁰⁰ M. Hansen,¹⁰⁰ P. Harris,¹⁰⁰ V. Innocente,¹⁰⁰ P. Janot,¹⁰⁰ E. Karavakis,¹⁰⁰ K. Kousouris,¹⁰⁰ K. Krajczar,¹⁰⁰ P. Lecoq,¹⁰⁰ C. Lourenço,¹⁰⁰ N. Magini,¹⁰⁰ L. Malgeri,¹⁰⁰ M. Mannelli,¹⁰⁰ L. Masetti,¹⁰⁰ F. Meijers,¹⁰⁰ S. Mersi,¹⁰⁰ E. Meschi,¹⁰⁰ F. Moortgat,¹⁰⁰ M. Mulders,¹⁰⁰ P. Musella,¹⁰⁰ L. Orsini,¹⁰⁰ E. Palencia Cortezon,¹⁰⁰ E. Perez,¹⁰⁰ L. Perrozzi,¹⁰⁰ A. Petrilli,¹⁰⁰ G. Petrucciani,¹⁰⁰ A. Pfeiffer,¹⁰⁰ M. Pierini,¹⁰⁰ M. Pimiä,¹⁰⁰ D. Piparo,¹⁰⁰ M. Plagge,¹⁰⁰ A. Racz,¹⁰⁰ W. Reece,¹⁰⁰ J. Rojo,¹⁰⁰ G. Rolandi,^{100,kk} M. Rovere,¹⁰⁰ H. Sakulin,¹⁰⁰ F. Santanastasio,¹⁰⁰ C. Schäfer,¹⁰⁰ C. Schwick,¹⁰⁰ S. Sekmen,¹⁰⁰ A. Sharma,¹⁰⁰ P. Siegrist,¹⁰⁰ P. Silva,¹⁰⁰ M. Simon,¹⁰⁰ P. Sphicas,^{100,ll} J. Steggemann,¹⁰⁰ B. Stieger,¹⁰⁰ M. Stoye,¹⁰⁰ A. Tsiros,¹⁰⁰ G. I. Veres,^{100,u} J. R. Vlimant,¹⁰⁰ H. K. Wöhri,¹⁰⁰ W. D. Zeuner,¹⁰⁰ W. Bertl,¹⁰¹ K. Deiters,¹⁰¹ W. Erdmann,¹⁰¹ R. Horisberger,¹⁰¹ Q. Ingram,¹⁰¹ H. C. Kaestli,¹⁰¹ S. König,¹⁰¹ D. Kotlinski,¹⁰¹ U. Langenegger,¹⁰¹ D. Renker,¹⁰¹ T. Rohe,¹⁰¹ F. Bachmair,¹⁰² L. Bäni,¹⁰² L. Bianchini,¹⁰² P. Bortignon,¹⁰² M. A. Buchmann,¹⁰² B. Casal,¹⁰² N. Chanon,¹⁰² A. Deisher,¹⁰² G. Dissertori,¹⁰² M. Dittmar,¹⁰² M. Donegà,¹⁰² M. Dünser,¹⁰² P. Eller,¹⁰² C. Grab,¹⁰² D. Hits,¹⁰² W. Lusterhmann,¹⁰² B. Mangano,¹⁰² A. C. Marini,¹⁰² P. Martinez Ruiz del Arbol,¹⁰² D. Meister,¹⁰² N. Mohr,¹⁰² C. Nägeli,^{102,mm} P. Nef,¹⁰² F. Nessi-Tedaldi,¹⁰² F. Pandolfi,¹⁰² L. Pape,¹⁰² F. Pauss,¹⁰² M. Peruzzi,¹⁰² M. Quittnat,¹⁰² F. J. Ronga,¹⁰² M. Rossini,¹⁰² A. Starodumov,^{102,nn} M. Takahashi,¹⁰² L. Tauscher,^{102,a} K. Theofilatos,¹⁰² D. Treille,¹⁰² R. Wallny,¹⁰² H. A. Weber,¹⁰² C. Amsler,^{103,oo} M. F. Canelli,¹⁰³ V. Chiochia,¹⁰³ A. De Cosa,¹⁰³ C. Favaro,¹⁰³ A. Hinzmann,¹⁰³ T. Hreus,¹⁰³ M. Ivova Rikova,¹⁰³ B. Kilminster,¹⁰³ B. Millan Mejias,¹⁰³ J. Ngadiuba,¹⁰³ P. Robmann,¹⁰³ H. Snoek,¹⁰³ S. Taroni,¹⁰³ M. Verzetti,¹⁰³ Y. Yang,¹⁰³ M. Cardaci,¹⁰⁴ K. H. Chen,¹⁰⁴ C. Ferro,¹⁰⁴ C. M. Kuo,¹⁰⁴ S. W. Li,¹⁰⁴ W. Lin,¹⁰⁴ Y. J. Lu,¹⁰⁴ R. Volpe,¹⁰⁴ S. S. Yu,¹⁰⁴ P. Bartalini,¹⁰⁵ P. Chang,¹⁰⁵ Y. H. Chang,¹⁰⁵ Y. W. Chang,¹⁰⁵ Y. Chao,¹⁰⁵ K. F. Chen,¹⁰⁵ P. H. Chen,¹⁰⁵ C. Dietz,¹⁰⁵ U. Grundler,¹⁰⁵ W.-S. Hou,¹⁰⁵ Y. Hsiung,¹⁰⁵ K. Y. Kao,¹⁰⁵ Y. J. Lei,¹⁰⁵ Y. F. Liu,¹⁰⁵ R.-S. Lu,¹⁰⁵ D. Majumder,¹⁰⁵ E. Petrakou,¹⁰⁵ X. Shi,¹⁰⁵ J. G. Shiu,¹⁰⁵ Y. M. Tzeng,¹⁰⁵ M. Wang,¹⁰⁵ R. Wilken,¹⁰⁵ B. Asavapibhop,¹⁰⁶ N. Suwonjandee,¹⁰⁶ A. Adiguzel,¹⁰⁷ M. N. Bakirci,^{107,pp} S. Cerci,^{107,qq} C. Dozen,¹⁰⁷ I. Dumanoglu,¹⁰⁷ E. Eskut,¹⁰⁷ S. Girgis,¹⁰⁷ G. Gokbulut,¹⁰⁷ E. Gurpinar,¹⁰⁷ I. Hos,¹⁰⁷ E. E. Kangal,¹⁰⁷ A. Kayis Topaksu,¹⁰⁷ G. Onengut,^{107,rr} K. Ozdemir,¹⁰⁷ S. Ozturk,^{107,pp} A. Polatoz,¹⁰⁷ K. Sogut,^{107,ss} D. Sunar Cerci,^{107,qq} B. Tali,^{107,qq} H. Topakli,^{107,pp} M. Vergili,¹⁰⁷ I. V. Akin,¹⁰⁸ T. Aliev,¹⁰⁸ B. Bilin,¹⁰⁸ S. Bilmis,¹⁰⁸ M. Deniz,¹⁰⁸ H. Gamsizkan,¹⁰⁸ A. M. Guler,¹⁰⁸ G. Karapinar,^{108,tt} K. Ocalan,¹⁰⁸ A. Ozpineci,¹⁰⁸ M. Serin,¹⁰⁸ R. Sever,¹⁰⁸ U. E. Surat,¹⁰⁸ M. Yalvac,¹⁰⁸ M. Zeyrek,¹⁰⁸ E. Gülmez,¹⁰⁹

B. Isildak,^{109,uu} M. Kaya,^{109,vv} O. Kaya,^{109,vv} S. Ozkorucuklu,^{109,ww} H. Bahtiyar,^{110,xx} E. Barlas,¹¹⁰ K. Cankocak,¹¹⁰
 Y. O. Günaydin,^{110,yy} F. I. Vardarli,¹¹⁰ M. Yücel,¹¹⁰ L. Levchuk,¹¹¹ P. Sorokin,¹¹¹ J. J. Brooke,¹¹² E. Clement,¹¹²
 D. Cussans,¹¹² H. Flacher,¹¹² R. Frazier,¹¹² J. Goldstein,¹¹² M. Grimes,¹¹² G. P. Heath,¹¹² H. F. Heath,¹¹² J. Jacob,¹¹²
 L. Kreczko,¹¹² C. Lucas,¹¹² Z. Meng,¹¹² D. M. Newbold,^{112,zz} S. Paramesvaran,¹¹² A. Poll,¹¹² S. Senkin,¹¹² V. J. Smith,¹¹²
 T. Williams,¹¹² K. W. Bell,¹¹³ A. Belyaev,^{113,aaa} C. Brew,¹¹³ R. M. Brown,¹¹³ D. J. A. Cockerill,¹¹³ J. A. Coughlan,¹¹³
 K. Harder,¹¹³ S. Harper,¹¹³ J. Ilic,¹¹³ E. Olaiya,¹¹³ D. Petyt,¹¹³ C. H. Shepherd-Themistocleous,¹¹³ A. Thea,¹¹³
 I. R. Tomalin,¹¹³ W. J. Womersley,¹¹³ S. D. Worm,¹¹³ M. Baber,¹¹⁴ R. Bainbridge,¹¹⁴ O. Buchmuller,¹¹⁴ D. Burton,¹¹⁴
 D. Colling,¹¹⁴ N. Cripps,¹¹⁴ M. Cutajar,¹¹⁴ P. Dauncey,¹¹⁴ G. Davies,¹¹⁴ M. Della Negra,¹¹⁴ W. Ferguson,¹¹⁴ J. Fulcher,¹¹⁴
 D. Futyan,¹¹⁴ A. Gilbert,¹¹⁴ A. Guneratne Bryer,¹¹⁴ G. Hall,¹¹⁴ Z. Hatherell,¹¹⁴ J. Hays,¹¹⁴ G. Iles,¹¹⁴ M. Jarvis,¹¹⁴
 G. Karapostoli,¹¹⁴ M. Kenzie,¹¹⁴ R. Lane,¹¹⁴ R. Lucas,^{114,zz} L. Lyons,¹¹⁴ A.-M. Magnan,¹¹⁴ J. Marrouche,¹¹⁴ B. Mathias,¹¹⁴
 R. Nandi,¹¹⁴ J. Nash,¹¹⁴ A. Nikitenko,^{114,nn} J. Pela,¹¹⁴ M. Pesaresi,¹¹⁴ K. Petridis,¹¹⁴ M. Pioppi,^{114,bbb} D. M. Raymond,¹¹⁴
 S. Rogerson,¹¹⁴ A. Rose,¹¹⁴ C. Seez,¹¹⁴ P. Sharp,^{114,a} A. Sparrow,¹¹⁴ A. Tapper,¹¹⁴ M. Vazquez Acosta,¹¹⁴ T. Virdee,¹¹⁴
 S. Wakefield,¹¹⁴ N. Wardle,¹¹⁴ J. E. Cole,¹¹⁵ P. R. Hobson,¹¹⁵ A. Khan,¹¹⁵ P. Kyberd,¹¹⁵ D. Leggat,¹¹⁵ D. Leslie,¹¹⁵
 W. Martin,¹¹⁵ I. D. Reid,¹¹⁵ P. Symonds,¹¹⁵ L. Teodorescu,¹¹⁵ M. Turner,¹¹⁵ J. Dittmann,¹¹⁶ K. Hatakeyama,¹¹⁶ A. Ksmi,¹¹⁶
 H. Liu,¹¹⁶ T. Scarborough,¹¹⁶ O. Charaf,¹¹⁷ S. I. Cooper,¹¹⁷ C. Henderson,¹¹⁷ P. Rumerio,¹¹⁷ A. Avetisyan,¹¹⁸ T. Bose,¹¹⁸
 C. Fantasia,¹¹⁸ A. Heister,¹¹⁸ P. Lawson,¹¹⁸ D. Lazic,¹¹⁸ J. Rohlf,¹¹⁸ D. Sperka,¹¹⁸ J. St. John,¹¹⁸ L. Sulak,¹¹⁸ J. Alimena,¹¹⁹
 S. Bhattacharya,¹¹⁹ G. Christopher,¹¹⁹ D. Cutts,¹¹⁹ Z. Demiragli,¹¹⁹ A. Ferapontov,¹¹⁹ A. Garabedian,¹¹⁹ U. Heintz,¹¹⁹
 S. Jabeen,¹¹⁹ G. Kukartsev,¹¹⁹ E. Laird,¹¹⁹ G. Landsberg,¹¹⁹ M. Luk,¹¹⁹ M. Narain,¹¹⁹ M. Segala,¹¹⁹ T. Sinthuprasith,¹¹⁹
 T. Speer,¹¹⁹ J. Swanson,¹¹⁹ R. Breedon,¹²⁰ G. Breto,¹²⁰ M. Calderon De La Barca Sanchez,¹²⁰ S. Chauhan,¹²⁰ M. Chertok,¹²⁰
 J. Conway,¹²⁰ R. Conway,¹²⁰ P. T. Cox,¹²⁰ R. Erbacher,¹²⁰ M. Gardner,¹²⁰ W. Ko,¹²⁰ A. Kopecky,¹²⁰ R. Lander,¹²⁰
 T. Miceli,¹²⁰ D. Pellett,¹²⁰ J. Pilot,¹²⁰ F. Ricci-Tam,¹²⁰ B. Rutherford,¹²⁰ M. Searle,¹²⁰ S. Shalhout,¹²⁰ J. Smith,¹²⁰
 M. Squires,¹²⁰ M. Tripathi,¹²⁰ S. Wilbur,¹²⁰ R. Yohay,¹²⁰ V. Andreev,¹²¹ D. Cline,¹²¹ R. Cousins,¹²¹ S. Erhan,¹²¹
 P. Everaerts,¹²¹ C. Farrell,¹²¹ M. Felcini,¹²¹ J. Hauser,¹²¹ M. Ignatenko,¹²¹ C. Jarvis,¹²¹ G. Rakness,¹²¹ P. Schlein,^{121,a}
 E. Takasugi,¹²¹ V. Valuev,¹²¹ M. Weber,¹²¹ J. Babb,¹²² R. Clare,¹²² J. Ellison,¹²² J. W. Gary,¹²² G. Hanson,¹²² J. Heilman,¹²²
 P. Jandir,¹²² F. Lacroix,¹²² H. Liu,¹²² O. R. Long,¹²² A. Luthra,¹²² M. Malberti,¹²² H. Nguyen,¹²² A. Shrinivas,¹²² J. Sturdy,¹²²
 S. Sumowidagdo,¹²² S. Wimpenny,¹²² W. Andrews,¹²³ J. G. Branson,¹²³ G. B. Cerati,¹²³ S. Cittolin,¹²³ R. T. D'Agnolo,¹²³
 D. Evans,¹²³ A. Holzner,¹²³ R. Kelley,¹²³ D. Kovalskyi,¹²³ M. Lebourgeois,¹²³ J. Letts,¹²³ I. Macneill,¹²³ S. Padhi,¹²³
 C. Palmer,¹²³ M. Pieri,¹²³ M. Sani,¹²³ V. Sharma,¹²³ S. Simon,¹²³ E. Sudano,¹²³ M. Tadel,¹²³ Y. Tu,¹²³ A. Vartak,¹²³
 S. Wasserbaech,^{123,ccc} F. Würthwein,¹²³ A. Yagil,¹²³ J. Yoo,¹²³ D. Barge,¹²⁴ C. Campagnari,¹²⁴ T. Danielson,¹²⁴
 K. Flowers,¹²⁴ P. Geffert,¹²⁴ C. George,¹²⁴ F. Golf,¹²⁴ J. Incandela,¹²⁴ C. Justus,¹²⁴ R. Magaña Villalba,¹²⁴ N. Mccoll,¹²⁴
 V. Pavlunin,¹²⁴ J. Richman,¹²⁴ R. Rossin,¹²⁴ D. Stuart,¹²⁴ W. To,¹²⁴ C. West,¹²⁴ A. Apresyan,¹²⁵ A. Bornheim,¹²⁵ J. Bunn,¹²⁵
 Y. Chen,¹²⁵ E. Di Marco,¹²⁵ J. Duarte,¹²⁵ D. Kcira,¹²⁵ A. Mott,¹²⁵ H. B. Newman,¹²⁵ C. Pena,¹²⁵ C. Rogan,¹²⁵
 M. Spiropulu,¹²⁵ V. Timciuc,¹²⁵ R. Wilkinson,¹²⁵ S. Xie,¹²⁵ R. Y. Zhu,¹²⁵ V. Azzolini,¹²⁶ A. Calamba,¹²⁶ R. Carroll,¹²⁶
 T. Ferguson,¹²⁶ Y. Iiyama,¹²⁶ D. W. Jang,¹²⁶ M. Paulini,¹²⁶ J. Russ,¹²⁶ H. Vogel,¹²⁶ I. Vorobiev,¹²⁶ J. P. Cumalat,¹²⁷
 B. R. Drell,¹²⁷ W. T. Ford,¹²⁷ A. Gaz,¹²⁷ E. Luigi Lopez,¹²⁷ U. Nauenberg,¹²⁷ J. G. Smith,¹²⁷ K. Stenson,¹²⁷ K. A. Ulmer,¹²⁷
 S. R. Wagner,¹²⁷ J. Alexander,¹²⁸ A. Chatterjee,¹²⁸ N. Eggert,¹²⁸ L. K. Gibbons,¹²⁸ W. Hopkins,¹²⁸ A. Khukhunaishvili,¹²⁸
 B. Kreis,¹²⁸ N. Mirman,¹²⁸ G. Nicolas Kaufman,¹²⁸ J. R. Patterson,¹²⁸ A. Ryd,¹²⁸ E. Salvati,¹²⁸ W. Sun,¹²⁸ W. D. Teo,¹²⁸
 J. Thom,¹²⁸ J. Thompson,¹²⁸ J. Tucker,¹²⁸ Y. Weng,¹²⁸ L. Winstrom,¹²⁸ P. Wittich,¹²⁸ D. Winn,¹²⁹ S. Abdullin,¹³⁰
 M. Albrow,¹³⁰ J. Anderson,¹³⁰ G. Apollinari,¹³⁰ L. A. T. Bauerdick,¹³⁰ A. Beretvas,¹³⁰ J. Berryhill,¹³⁰ P. C. Bhat,¹³⁰
 K. Burkett,¹³⁰ J. N. Butler,¹³⁰ V. Chetluru,¹³⁰ H. W. K. Cheung,¹³⁰ F. Chlebana,¹³⁰ S. Cihangir,¹³⁰ V. D. Elvira,¹³⁰ I. Fisk,¹³⁰
 J. Freeman,¹³⁰ Y. Gao,¹³⁰ E. Gottschalk,¹³⁰ L. Gray,¹³⁰ D. Green,¹³⁰ S. Grünendahl,¹³⁰ O. Gutsche,¹³⁰ D. Hare,¹³⁰
 R. M. Harris,¹³⁰ J. Hirschauer,¹³⁰ B. Hooberman,¹³⁰ S. Jindariani,¹³⁰ M. Johnson,¹³⁰ U. Joshi,¹³⁰ K. Kaadze,¹³⁰ B. Klima,¹³⁰
 S. Kwan,¹³⁰ J. Linacre,¹³⁰ D. Lincoln,¹³⁰ R. Lipton,¹³⁰ J. Lykken,¹³⁰ K. Maeshima,¹³⁰ J. M. Marraffino,¹³⁰
 V. I. Martinez Outschoorn,¹³⁰ S. Maruyama,¹³⁰ D. Mason,¹³⁰ P. McBride,¹³⁰ K. Mishra,¹³⁰ S. Mrenna,¹³⁰ Y. Musienko,^{130,hh}
 S. Nahn,¹³⁰ C. Newman-Holmes,¹³⁰ V. O'Dell,¹³⁰ O. Prokofyev,¹³⁰ N. Ratnikova,¹³⁰ E. Sexton-Kennedy,¹³⁰ S. Sharma,¹³⁰
 W. J. Spalding,¹³⁰ L. Spiegel,¹³⁰ L. Taylor,¹³⁰ S. Tkaczyk,¹³⁰ N. V. Tran,¹³⁰ L. Uplegger,¹³⁰ E. W. Vaandering,¹³⁰ R. Vidal,¹³⁰
 A. Whitbeck,¹³⁰ J. Whitmore,¹³⁰ W. Wu,¹³⁰ F. Yang,¹³⁰ J. C. Yun,¹³⁰ D. Acosta,¹³¹ P. Avery,¹³¹ D. Bourilkov,¹³¹ T. Cheng,¹³¹
 S. Das,¹³¹ M. De Gruttola,¹³¹ G. P. Di Giovanni,¹³¹ D. Dobur,¹³¹ R. D. Field,¹³¹ M. Fisher,¹³¹ Y. Fu,¹³¹ I. K. Furic,¹³¹
 J. Hugon,¹³¹ B. Kim,¹³¹ J. Konigsberg,¹³¹ A. Korytov,¹³¹ A. Kropivnitskaya,¹³¹ T. Kypreos,¹³¹ J. F. Low,¹³¹ K. Matchev,¹³¹

P. Milenovic,^{131,ddd} G. Mitselmakher,¹³¹ L. Muniz,¹³¹ A. Rinkevicius,¹³¹ L. Shchutska,¹³¹ N. Skhirtladze,¹³¹ M. Snowball,¹³¹ J. Yelton,¹³¹ M. Zakaria,¹³¹ V. Gaultney,¹³² S. Hewamanage,¹³² S. Linn,¹³² P. Markowitz,¹³² G. Martinez,¹³² J. L. Rodriguez,¹³² T. Adams,¹³³ A. Askew,¹³³ J. Bochenek,¹³³ J. Chen,¹³³ B. Diamond,¹³³ J. Haas,¹³³ S. Hagopian,¹³³ V. Hagopian,¹³³ K. F. Johnson,¹³³ H. Prosper,¹³³ V. Veeraraghavan,¹³³ M. Weinberg,¹³³ M. M. Baarmand,¹³⁴ B. Dorney,¹³⁴ M. Hohlmann,¹³⁴ H. Kalakhety,¹³⁴ F. Yumiceva,¹³⁴ M. R. Adams,¹³⁵ L. Apanasevich,¹³⁵ V. E. Bazterra,¹³⁵ R. R. Betts,¹³⁵ I. Bucinskaite,¹³⁵ R. Cavanaugh,¹³⁵ O. Evdokimov,¹³⁵ L. Gauthier,¹³⁵ C. E. Gerber,¹³⁵ D. J. Hofman,¹³⁵ S. Khalatyan,¹³⁵ P. Kurt,¹³⁵ D. H. Moon,¹³⁵ C. O'Brien,¹³⁵ C. Silkworth,¹³⁵ P. Turner,¹³⁵ N. Varelas,¹³⁵ U. Akgun,¹³⁶ E. A. Albayrak,^{136,xx} B. Bilki,^{136,eee} W. Clarida,¹³⁶ K. Dilsiz,¹³⁶ F. Duru,¹³⁶ M. Haytmyradov,¹³⁶ J.-P. Merlo,¹³⁶ H. Mermerkaya,^{136,fff} A. Mestvirishvili,¹³⁶ A. Moeller,¹³⁶ J. Nachtman,¹³⁶ H. Ogul,¹³⁶ Y. Onel,¹³⁶ F. Ozok,^{136,xx} S. Sen,¹³⁶ P. Tan,¹³⁶ E. Tiras,¹³⁶ J. Wetzel,¹³⁶ T. Yetkin,^{136,ggg} K. Yi,¹³⁶ B. A. Barnett,¹³⁷ B. Blumenfeld,¹³⁷ S. Bolognesi,¹³⁷ D. Fehling,¹³⁷ A. V. Gritsan,¹³⁷ P. Maksimovic,¹³⁷ C. Martin,¹³⁷ M. Swartz,¹³⁷ P. Baringer,¹³⁸ A. Bean,¹³⁸ G. Benelli,¹³⁸ R. P. Kenny III,¹³⁸ M. Murray,¹³⁸ D. Noonan,¹³⁸ S. Sanders,¹³⁸ J. Sekaric,¹³⁸ R. Stringer,¹³⁸ Q. Wang,¹³⁸ J. S. Wood,¹³⁸ A. F. Barfuss,¹³⁹ I. Chakaberia,¹³⁹ A. Ivanov,¹³⁹ S. Khalil,¹³⁹ M. Makouski,¹³⁹ Y. Maravin,¹³⁹ L. K. Saini,¹³⁹ S. Shrestha,¹³⁹ I. Svintradze,¹³⁹ J. Gronberg,¹⁴⁰ D. Lange,¹⁴⁰ F. Rebassoo,¹⁴⁰ D. Wright,¹⁴⁰ A. Baden,¹⁴¹ B. Calvert,¹⁴¹ S. C. Eno,¹⁴¹ J. A. Gomez,¹⁴¹ N. J. Hadley,¹⁴¹ R. G. Kellogg,¹⁴¹ T. Kolberg,¹⁴¹ Y. Lu,¹⁴¹ M. Marionneau,¹⁴¹ A. C. Mignerey,¹⁴¹ K. Pedro,¹⁴¹ A. Skuja,¹⁴¹ J. Temple,¹⁴¹ M. B. Tonjes,¹⁴¹ S. C. Tonwar,¹⁴¹ A. Apyan,¹⁴² R. Barbieri,¹⁴² G. Bauer,¹⁴² W. Busza,¹⁴² I. A. Cali,¹⁴² M. Chan,¹⁴² L. Di Matteo,¹⁴² V. Dutta,¹⁴² G. Gomez Ceballos,¹⁴² M. Goncharov,¹⁴² D. Gulhan,¹⁴² M. Klute,¹⁴² Y. S. Lai,¹⁴² Y.-J. Lee,¹⁴² A. Levin,¹⁴² P. D. Luckey,¹⁴² T. Ma,¹⁴² C. Paus,¹⁴² D. Ralph,¹⁴² C. Roland,¹⁴² G. Roland,¹⁴² G. S. F. Stephans,¹⁴² F. Stöckli,¹⁴² K. Sumorok,¹⁴² D. Velicanu,¹⁴² J. Veverka,¹⁴² B. Wyslouch,¹⁴² M. Yang,¹⁴² A. S. Yoon,¹⁴² M. Zanetti,¹⁴² V. Zhukova,¹⁴² B. Dahmes,¹⁴³ A. De Benedetti,¹⁴³ A. Gude,¹⁴³ S. C. Kao,¹⁴³ K. Klapoetke,¹⁴³ Y. Kubota,¹⁴³ J. Mans,¹⁴³ N. Pastika,¹⁴³ R. Rusack,¹⁴³ A. Singovsky,¹⁴³ N. Tambe,¹⁴³ J. Turkewitz,¹⁴³ J. G. Acosta,¹⁴⁴ L. M. Cremaldi,¹⁴⁴ R. Kroeger,¹⁴⁴ S. Oliveros,¹⁴⁴ L. Perera,¹⁴⁴ R. Rahmat,¹⁴⁴ D. A. Sanders,¹⁴⁴ D. Summers,¹⁴⁴ E. Avdeeva,¹⁴⁵ K. Bloom,¹⁴⁵ S. Bose,¹⁴⁵ D. R. Claes,¹⁴⁵ A. Dominguez,¹⁴⁵ R. Gonzalez Suarez,¹⁴⁵ J. Keller,¹⁴⁵ D. Knowlton,¹⁴⁵ I. Kravchenko,¹⁴⁵ J. Lazo-Flores,¹⁴⁵ S. Malik,¹⁴⁵ F. Meier,¹⁴⁵ G. R. Snow,¹⁴⁵ J. Dolen,¹⁴⁶ A. Godshalk,¹⁴⁶ I. Iashvili,¹⁴⁶ S. Jain,¹⁴⁶ A. Kharchilava,¹⁴⁶ A. Kumar,¹⁴⁶ S. Rappoccio,¹⁴⁶ G. Alverson,¹⁴⁷ E. Barberis,¹⁴⁷ D. Baumgartel,¹⁴⁷ M. Chasco,¹⁴⁷ J. Haley,¹⁴⁷ A. Massironi,¹⁴⁷ D. Nash,¹⁴⁷ T. Orimoto,¹⁴⁷ D. Trocino,¹⁴⁷ D. Wood,¹⁴⁷ J. Zhang,¹⁴⁷ A. Anastassov,¹⁴⁸ K. A. Hahn,¹⁴⁸ A. Kubik,¹⁴⁸ L. Lusito,¹⁴⁸ N. Mucia,¹⁴⁸ N. Odell,¹⁴⁸ B. Pollack,¹⁴⁸ A. Pozdnyakov,¹⁴⁸ M. Schmitt,¹⁴⁸ S. Stoynev,¹⁴⁸ K. Sung,¹⁴⁸ M. Velasco,¹⁴⁸ S. Won,¹⁴⁸ D. Berry,¹⁴⁹ A. Brinkerhoff,¹⁴⁹ K. M. Chan,¹⁴⁹ A. Drozdetskiy,¹⁴⁹ M. Hildreth,¹⁴⁹ C. Jessop,¹⁴⁹ D. J. Karmgard,¹⁴⁹ N. Kellams,¹⁴⁹ J. Kolb,¹⁴⁹ K. Lannon,¹⁴⁹ W. Luo,¹⁴⁹ S. Lynch,¹⁴⁹ N. Marinelli,¹⁴⁹ D. M. Morse,¹⁴⁹ T. Pearson,¹⁴⁹ M. Planer,¹⁴⁹ R. Ruchti,¹⁴⁹ J. Slaunwhite,¹⁴⁹ N. Valls,¹⁴⁹ M. Wayne,¹⁴⁹ M. Wolf,¹⁴⁹ A. Woodard,¹⁴⁹ L. Antonelli,¹⁵⁰ B. Bylsma,¹⁵⁰ L. S. Durkin,¹⁵⁰ S. Flowers,¹⁵⁰ C. Hill,¹⁵⁰ R. Hughes,¹⁵⁰ K. Kotov,¹⁵⁰ T. Y. Ling,¹⁵⁰ D. Puigh,¹⁵⁰ M. Rodenburg,¹⁵⁰ G. Smith,¹⁵⁰ C. Vuosalo,¹⁵⁰ B. L. Winer,¹⁵⁰ H. Wolfe,¹⁵⁰ H. W. Wulsin,¹⁵⁰ E. Berry,¹⁵¹ P. Elmer,¹⁵¹ V. Halyo,¹⁵¹ P. Hebda,¹⁵¹ J. Hegeman,¹⁵¹ A. Hunt,¹⁵¹ P. Jindal,¹⁵¹ S. A. Koay,¹⁵¹ P. Lujan,¹⁵¹ D. Marlow,¹⁵¹ T. Medvedeva,¹⁵¹ M. Mooney,¹⁵¹ J. Olsen,¹⁵¹ P. Piroué,¹⁵¹ X. Quan,¹⁵¹ A. Raval,¹⁵¹ H. Saka,¹⁵¹ D. Stickland,¹⁵¹ C. Tully,¹⁵¹ J. S. Werner,¹⁵¹ S. C. Zenz,¹⁵¹ A. Zuranski,¹⁵¹ E. Brownson,¹⁵² A. Lopez,¹⁵² H. Mendez,¹⁵² J. E. Ramirez Vargas,¹⁵² E. Alagoz,¹⁵³ D. Benedetti,¹⁵³ G. Bolla,¹⁵³ D. Bortoletto,¹⁵³ M. De Mattia,¹⁵³ A. Everett,¹⁵³ Z. Hu,¹⁵³ M. K. Jha,¹⁵³ M. Jones,¹⁵³ K. Jung,¹⁵³ M. Kress,¹⁵³ N. Leonardo,¹⁵³ D. Lopes Pegna,¹⁵³ V. Maroussov,¹⁵³ P. Merkel,¹⁵³ D. H. Miller,¹⁵³ N. Neumeister,¹⁵³ B. C. Radburn-Smith,¹⁵³ I. Shipsey,¹⁵³ D. Silvers,¹⁵³ A. Svyatkovskiy,¹⁵³ F. Wang,¹⁵³ W. Xie,¹⁵³ L. Xu,¹⁵³ H. D. Yoo,¹⁵³ J. Zablocki,¹⁵³ Y. Zheng,¹⁵³ N. Parashar,¹⁵⁴ A. Adair,¹⁵⁵ B. Akgun,¹⁵⁵ K. M. Ecklund,¹⁵⁵ F. J. M. Geurts,¹⁵⁵ W. Li,¹⁵⁵ B. Michlin,¹⁵⁵ B. P. Padley,¹⁵⁵ R. Redjimi,¹⁵⁵ J. Roberts,¹⁵⁵ J. Zabel,¹⁵⁵ B. Betchart,¹⁵⁶ A. Bodek,¹⁵⁶ R. Covarelli,¹⁵⁶ P. de Barbaro,¹⁵⁶ R. Demina,¹⁵⁶ Y. Eshaq,¹⁵⁶ T. Ferbel,¹⁵⁶ A. Garcia-Bellido,¹⁵⁶ P. Goldenzweig,¹⁵⁶ J. Han,¹⁵⁶ A. Harel,¹⁵⁶ D. C. Miner,¹⁵⁶ G. Petrillo,¹⁵⁶ D. Vishnevskiy,¹⁵⁶ M. Zielinski,¹⁵⁶ A. Bhatti,¹⁵⁷ R. Ciesielski,¹⁵⁷ L. Demortier,¹⁵⁷ K. Goulianos,¹⁵⁷ G. Lungu,¹⁵⁷ S. Malik,¹⁵⁷ C. Mesropian,¹⁵⁷ S. Arora,¹⁵⁸ A. Barker,¹⁵⁸ J. P. Chou,¹⁵⁸ C. Contreras-Campana,¹⁵⁸ E. Contreras-Campana,¹⁵⁸ D. Duggan,¹⁵⁸ D. Ferencek,¹⁵⁸ Y. Gershtein,¹⁵⁸ R. Gray,¹⁵⁸ E. Halkiadakis,¹⁵⁸ D. Hidas,¹⁵⁸ A. Lath,¹⁵⁸ S. Panwalkar,¹⁵⁸ M. Park,¹⁵⁸ R. Patel,¹⁵⁸ V. Rekovic,¹⁵⁸ J. Robles,¹⁵⁸ S. Salur,¹⁵⁸ S. Schnetzer,¹⁵⁸ C. Seitz,¹⁵⁸ S. Somalwar,¹⁵⁸ R. Stone,¹⁵⁸ S. Thomas,¹⁵⁸ P. Thomassen,¹⁵⁸ M. Walker,¹⁵⁸ K. Rose,¹⁵⁹ S. Spanier,¹⁵⁹ Z. C. Yang,¹⁵⁹ A. York,¹⁵⁹ O. Bouhali,^{160,hhh} R. Eusebi,¹⁶⁰ W. Flanagan,¹⁶⁰ J. Gilmore,¹⁶⁰ T. Kamon,^{160,iii} V. Khotilovich,¹⁶⁰ V. Krutelyov,¹⁶⁰ R. Montalvo,¹⁶⁰ I. Osipenkov,¹⁶⁰ Y. Pakhotin,¹⁶⁰ A. Perloff,¹⁶⁰ J. Roe,¹⁶⁰ A. Safonov,¹⁶⁰ T. Sakuma,¹⁶⁰ I. Suarez,¹⁶⁰ A. Tatarinov,¹⁶⁰ D. Toback,¹⁶⁰ N. Akchurin,¹⁶¹ C. Cowden,¹⁶¹ J. Damgov,¹⁶¹

C. Dragoiu,¹⁶¹ P. R. Duderø,¹⁶¹ J. Faulkner,¹⁶¹ K. Kovitanggoon,¹⁶¹ S. Kunori,¹⁶¹ S. W. Lee,¹⁶¹ T. Libeiro,¹⁶¹ I. Volobouev,¹⁶¹ E. Appelt,¹⁶² A. G. Delannoy,¹⁶² S. Greene,¹⁶² A. Gurrola,¹⁶² W. Johns,¹⁶² C. Maguire,¹⁶² Y. Mao,¹⁶² A. Melo,¹⁶² M. Sharma,¹⁶² P. Sheldon,¹⁶² B. Snook,¹⁶² S. Tuo,¹⁶² J. Velkovska,¹⁶² M. W. Arenton,¹⁶³ S. Boutle,¹⁶³ B. Cox,¹⁶³ B. Francis,¹⁶³ J. Goodell,¹⁶³ R. Hirosky,¹⁶³ A. Ledovskoy,¹⁶³ C. Lin,¹⁶³ C. Neu,¹⁶³ J. Wood,¹⁶³ S. Gollapinni,¹⁶⁴ R. Harr,¹⁶⁴ P. E. Karchin,¹⁶⁴ C. Kottachchi Kankanamge Don,¹⁶⁴ P. Lamichhane,¹⁶⁴ D. A. Belknap,¹⁶⁵ L. Borrello,¹⁶⁵ D. Carlsmith,¹⁶⁵ M. Cepeda,¹⁶⁵ S. Dasu,¹⁶⁵ S. Duric,¹⁶⁵ E. Friis,¹⁶⁵ M. Grothe,¹⁶⁵ R. Hall-Wilton,¹⁶⁵ M. Herndon,¹⁶⁵ A. Hervé,¹⁶⁵ P. Klabbers,¹⁶⁵ J. Klukas,¹⁶⁵ A. Lanaro,¹⁶⁵ A. Levine,¹⁶⁵ R. Loveless,¹⁶⁵ A. Mohapatra,¹⁶⁵ I. Ojalvo,¹⁶⁵ T. Perry,¹⁶⁵ G. A. Pierro,¹⁶⁵ G. Polese,¹⁶⁵ I. Ross,¹⁶⁵ A. Sakharov,¹⁶⁵ T. Sarangi,¹⁶⁵ A. Savin¹⁶⁵ and W. H. Smith¹⁶⁵

(CMS Collaboration)

¹*Yerevan Physics Institute, Yerevan, Armenia*

²*Institut für Hochenergiephysik der OeAW, Wien, Austria*

³*National Centre for Particle and High Energy Physics, Minsk, Belarus*

⁴*Universiteit Antwerpen, Antwerpen, Belgium*

⁵*Vrije Universiteit Brussel, Brussel, Belgium*

⁶*Université Libre de Bruxelles, Bruxelles, Belgium*

⁷*Ghent University, Ghent, Belgium*

⁸*Université Catholique de Louvain, Louvain-la-Neuve, Belgium*

⁹*Université de Mons, Mons, Belgium*

¹⁰*Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil*

¹¹*Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil*

^{12a}*Universidade Estadual Paulista, São Paulo, Brazil*

^{12b}*Universidade Federal do ABC, São Paulo, Brazil*

¹³*Institute for Nuclear Research and Nuclear Energy, Sofia, Bulgaria*

¹⁴*University of Sofia, Sofia, Bulgaria*

¹⁵*Institute of High Energy Physics, Beijing, China*

¹⁶*State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing, China*

¹⁷*Universidad de Los Andes, Bogota, Colombia*

¹⁸*Technical University of Split, Split, Croatia*

¹⁹*University of Split, Split, Croatia*

²⁰*Institute Rudjer Boskovic, Zagreb, Croatia*

²¹*University of Cyprus, Nicosia, Cyprus*

²²*Charles University, Prague, Czech Republic*

²³*Academy of Scientific Research and Technology of the Arab Republic of Egypt, Egyptian Network of High Energy Physics, Cairo, Egypt*

²⁴*National Institute of Chemical Physics and Biophysics, Tallinn, Estonia*

²⁵*Department of Physics, University of Helsinki, Helsinki, Finland*

²⁶*Helsinki Institute of Physics, Helsinki, Finland*

²⁷*Lappeenranta University of Technology, Lappeenranta, Finland*

²⁸*DSM/IRFU, CEA/Saclay, Gif-sur-Yvette, France*

²⁹*Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France*

³⁰*Institut Pluridisciplinaire Hubert Curien, Université de Strasbourg, Université de Haute Alsace Mulhouse, CNRS/IN2P3, Strasbourg, France*

³¹*Centre de Calcul de l'Institut National de Physique Nucléaire et de Physique des Particules, CNRS/IN2P3, Villeurbanne, France*

³²*Université de Lyon, Université Claude Bernard Lyon 1, CNRS-IN2P3, Institut de Physique Nucléaire de Lyon, Villeurbanne, France*

³³*Institute of High Energy Physics and Informatization, Tbilisi State University, Tbilisi, Georgia*

³⁴*RWTH Aachen University, I. Physikalisches Institut, Aachen, Germany*

³⁵*RWTH Aachen University, III. Physikalisches Institut A, Aachen, Germany*

³⁶*RWTH Aachen University, III. Physikalisches Institut B, Aachen, Germany*

³⁷*Deutsches Elektronen-Synchrotron, Hamburg, Germany*

³⁸*University of Hamburg, Hamburg, Germany*

³⁹*Institut für Experimentelle Kernphysik, Karlsruhe, Germany*

⁴⁰*Institute of Nuclear and Particle Physics (INPP), NCSR Demokritos, Aghia Paraskevi, Greece*

⁴¹*University of Athens, Athens, Greece*

⁴²*University of Ioánnina, Ioánnina, Greece*

- ⁴³Wigner Research Centre for Physics, Budapest, Hungary
⁴⁴Institute of Nuclear Research ATOMKI, Debrecen, Hungary
⁴⁵University of Debrecen, Debrecen, Hungary
⁴⁶National Institute of Science Education and Research, Bhubaneswar, India
⁴⁷Panjab University, Chandigarh, India
⁴⁸University of Delhi, Delhi, India
⁴⁹Saha Institute of Nuclear Physics, Kolkata, India
⁵⁰Bhabha Atomic Research Centre, Mumbai, India
⁵¹Tata Institute of Fundamental Research - EHEP, Mumbai, India
⁵²Tata Institute of Fundamental Research - HECR, Mumbai, India
⁵³Institute for Research in Fundamental Sciences (IPM), Tehran, Iran
⁵⁴University College Dublin, Dublin, Ireland
^{55a}INFN Sezione di Bari, Bari, Italy
^{55b}Università di Bari, Bari, Italy
^{55c}Politecnico di Bari, Bari, Italy
^{56a}INFN Sezione di Bologna, Bologna, Italy
^{56b}Università di Bologna, Bologna, Italy
^{57a}INFN Sezione di Catania, Catania, Italy
^{57b}Università di Catania, Catania, Italy
^{57c}CSFNSM, Catania, Italy
^{58a}INFN Sezione di Firenze, Firenze, Italy
^{58b}Università di Firenze, Firenze, Italy
⁵⁹INFN Laboratori Nazionali di Frascati, Frascati, Italy
^{60a}INFN Sezione di Genova, Genova, Italy
^{60b}Università di Genova, Genova, Italy
^{61a}INFN Sezione di Milano-Bicocca, Milano, Italy
^{61b}Università di Milano-Bicocca, Milano, Italy
^{62a}INFN Sezione di Napoli, Napoli, Italy
^{62b}Università di Napoli 'Federico II', Napoli, Italy
^{62c}Università della Basilicata (Potenza), Napoli, Italy
^{62d}Università G. Marconi (Roma), Napoli, Italy
^{63a}INFN Sezione di Padova, Padova, Italy
^{63b}Università di Padova, Padova, Italy
^{63c}Università di Trento (Trento), Padova, Italy
^{64a}INFN Sezione di Pavia, Pavia, Italy
^{64b}Università di Pavia, Pavia, Italy
^{65a}INFN Sezione di Perugia, Perugia, Italy
^{65b}Università di Perugia, Perugia, Italy
^{66a}INFN Sezione di Pisa, Pisa, Italy
^{66b}Università di Pisa, Pisa, Italy
^{66c}Scuola Normale Superiore di Pisa, Pisa, Italy
^{67a}INFN Sezione di Roma, Roma, Italy
^{67b}Università di Roma, Roma, Italy
^{68a}INFN Sezione di Torino, Torino, Italy
^{68b}Università di Torino, Torino, Italy
^{68c}Università del Piemonte Orientale (Novara), Torino, Italy
^{69a}INFN Sezione di Trieste, Trieste, Italy
^{69b}Università di Trieste, Trieste, Italy
⁷⁰Kangwon National University, Chunchon, Korea
⁷¹Kyungpook National University, Daegu, Korea
⁷²Chonnam National University, Institute for Universe and Elementary Particles, Kwangju, Korea
⁷³Korea University, Seoul, Korea
⁷⁴University of Seoul, Seoul, Korea
⁷⁵Sungkyunkwan University, Suwon, Korea
⁷⁶Vilnius University, Vilnius, Lithuania
⁷⁷University of Malaya Jabatan Fizik, Kuala Lumpur, Malaysia
⁷⁸Centro de Investigacion y de Estudios Avanzados del IPN, Mexico City, Mexico
⁷⁹Universidad Iberoamericana, Mexico City, Mexico
⁸⁰Benemerita Universidad Autonoma de Puebla, Puebla, Mexico
⁸¹Universidad Autónoma de San Luis Potosí, San Luis Potosí, Mexico

- ⁸²University of Auckland, Auckland, New Zealand
⁸³University of Canterbury, Christchurch, New Zealand
⁸⁴National Centre for Physics, Quaid-I-Azam University, Islamabad, Pakistan
⁸⁵National Centre for Nuclear Research, Swierk, Poland
⁸⁶Institute of Experimental Physics, Faculty of Physics, University of Warsaw, Warsaw, Poland
⁸⁷Laboratório de Instrumentação e Física Experimental de Partículas, Lisboa, Portugal
⁸⁸Joint Institute for Nuclear Research, Dubna, Russia
⁸⁹Petersburg Nuclear Physics Institute, Gatchina (St. Petersburg), Russia
⁹⁰Institute for Nuclear Research, Moscow, Russia
⁹¹Institute for Theoretical and Experimental Physics, Moscow, Russia
⁹²P.N. Lebedev Physical Institute, Moscow, Russia
⁹³Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow, Russia
⁹⁴State Research Center of Russian Federation, Institute for High Energy Physics, Protvino, Russia
⁹⁵University of Belgrade, Faculty of Physics and Vinca Institute of Nuclear Sciences, Belgrade, Serbia
⁹⁶Centro de Investigaciones Energéticas Medioambientales y Tecnológicas (CIEMAT), Madrid, Spain
⁹⁷Universidad Autónoma de Madrid, Madrid, Spain
⁹⁸Universidad de Oviedo, Oviedo, Spain
⁹⁹Instituto de Física de Cantabria (IFCA), CSIC-Universidad de Cantabria, Santander, Spain
¹⁰⁰CERN, European Organization for Nuclear Research, Geneva, Switzerland
¹⁰¹Paul Scherrer Institut, Villigen, Switzerland
¹⁰²Institute for Particle Physics, ETH Zurich, Zurich, Switzerland
¹⁰³Universität Zürich, Zurich, Switzerland
¹⁰⁴National Central University, Chung-Li, Taiwan
¹⁰⁵National Taiwan University (NTU), Taipei, Taiwan
¹⁰⁶Chulalongkorn University, Bangkok, Thailand
¹⁰⁷Cukurova University, Adana, Turkey
¹⁰⁸Middle East Technical University, Physics Department, Ankara, Turkey
¹⁰⁹Bogazici University, Istanbul, Turkey
¹¹⁰Istanbul Technical University, Istanbul, Turkey
¹¹¹National Scientific Center, Kharkov Institute of Physics and Technology, Kharkov, Ukraine
¹¹²University of Bristol, Bristol, United Kingdom
¹¹³Rutherford Appleton Laboratory, Didcot, United Kingdom
¹¹⁴Imperial College, London, United Kingdom
¹¹⁵Brunel University, Uxbridge, United Kingdom
¹¹⁶Baylor University, Waco, USA
¹¹⁷The University of Alabama, Tuscaloosa, USA
¹¹⁸Boston University, Boston, USA
¹¹⁹Brown University, Providence, USA
¹²⁰University of California, Davis, Davis, USA
¹²¹University of California, Los Angeles, USA
¹²²University of California, Riverside, Riverside, USA
¹²³University of California, San Diego, La Jolla, USA
¹²⁴University of California, Santa Barbara, Santa Barbara, USA
¹²⁵California Institute of Technology, Pasadena, USA
¹²⁶Carnegie Mellon University, Pittsburgh, USA
¹²⁷University of Colorado at Boulder, Boulder, USA
¹²⁸Cornell University, Ithaca, USA
¹²⁹Fairfield University, Fairfield, USA
¹³⁰Fermi National Accelerator Laboratory, Batavia, USA
¹³¹University of Florida, Gainesville, USA
¹³²Florida International University, Miami, USA
¹³³Florida State University, Tallahassee, USA
¹³⁴Florida Institute of Technology, Melbourne, USA
¹³⁵University of Illinois at Chicago (UIC), Chicago, USA
¹³⁶The University of Iowa, Iowa City, USA
¹³⁷Johns Hopkins University, Baltimore, USA
¹³⁸The University of Kansas, Lawrence, USA
¹³⁹Kansas State University, Manhattan, USA
¹⁴⁰Lawrence Livermore National Laboratory, Livermore, USA
¹⁴¹University of Maryland, College Park, USA

- ¹⁴²*Massachusetts Institute of Technology, Cambridge, USA*
¹⁴³*University of Minnesota, Minneapolis, USA*
¹⁴⁴*University of Mississippi, Oxford, USA*
¹⁴⁵*University of Nebraska-Lincoln, Lincoln, USA*
¹⁴⁶*State University of New York at Buffalo, Buffalo, USA*
¹⁴⁷*Northeastern University, Boston, USA*
¹⁴⁸*Northwestern University, Evanston, USA*
¹⁴⁹*University of Notre Dame, Notre Dame, USA*
¹⁵⁰*The Ohio State University, Columbus, USA*
¹⁵¹*Princeton University, Princeton, USA*
¹⁵²*University of Puerto Rico, Mayaguez, USA*
¹⁵³*Purdue University, West Lafayette, USA*
¹⁵⁴*Purdue University Calumet, Hammond, USA*
¹⁵⁵*Rice University, Houston, USA*
¹⁵⁶*University of Rochester, Rochester, USA*
¹⁵⁷*The Rockefeller University, New York, USA*
¹⁵⁸*Rutgers, The State University of New Jersey, Piscataway, USA*
¹⁵⁹*University of Tennessee, Knoxville, USA*
¹⁶⁰*Texas A&M University, College Station, USA*
¹⁶¹*Texas Tech University, Lubbock, USA*
¹⁶²*Vanderbilt University, Nashville, USA*
¹⁶³*University of Virginia, Charlottesville, USA*
¹⁶⁴*Wayne State University, Detroit, USA*
¹⁶⁵*University of Wisconsin, Madison, USA*

^aDeceased.

^bAlso at Vienna University of Technology, Vienna, Austria.

^cAlso at CERN, European Organization for Nuclear Research, Geneva, Switzerland.

^dAlso at Institut Pluridisciplinaire Hubert Curien, Université de Strasbourg, Université de Haute Alsace Mulhouse, CNRS/IN2P3, Strasbourg, France.

^eAlso at National Institute of Chemical Physics and Biophysics, Tallinn, Estonia.

^fAlso at Skobeltsyn Institute of Nuclear Physics, Lomonosov Moscow State University, Moscow, Russia.

^gAlso at Universidade Estadual de Campinas, Campinas, Brazil.

^hAlso at California Institute of Technology, Pasadena, USA.

ⁱAlso at Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France.

^jAlso at Zewail City of Science and Technology, Zewail, Egypt.

^kAlso at Suez Canal University, Suez, Egypt.

^lAlso at Cairo University, Cairo, Egypt.

^mAlso at Fayoum University, El-Fayoum, Egypt.

ⁿAlso at British University in Egypt, Cairo, Egypt.

^oAlso at Ain Shams University, Cairo, Egypt.

^pAlso at Université de Haute Alsace, Mulhouse, France.

^qAlso at Joint Institute for Nuclear Research, Dubna, Russia.

^rAlso at Brandenburg University of Technology, Cottbus, Germany.

^sAlso at The University of Kansas, Lawrence, USA.

^tAlso at Institute of Nuclear Research ATOMKI, Debrecen, Hungary.

^uAlso at Eötvös Loránd University, Budapest, Hungary.

^vAlso at Tata Institute of Fundamental Research - HECR, Mumbai, India.

^wAlso at King Abdulaziz University, Jeddah, Saudi Arabia.

^xAlso at University of Visva-Bharati, Santiniketan, India.

^yAlso at University of Ruhuna, Matara, Sri Lanka.

^zAlso at Isfahan University of Technology, Isfahan, Iran.

^{aa}Also at Sharif University of Technology, Tehran, Iran.

^{bb}Also at Plasma Physics Research Center, Science and Research Branch, Islamic Azad University, Tehran, Iran.

^{cc}Also at Università degli Studi di Siena, Siena, Italy.

^{dd}Also at Centre National de la Recherche Scientifique (CNRS) - IN2P3, Paris, France.

^{ee}Also at Purdue University, West Lafayette, USA.

^{ff}Also at Universidad Michoacana de San Nicolas de Hidalgo, Morelia, Mexico.

^{gg}Also at National Centre for Nuclear Research, Swierk, Poland.

^{hh}Also at Institute for Nuclear Research, Moscow, Russia.

- ⁱⁱ Also at Faculty of Physics, University of Belgrade, Belgrade, Serbia.
- ^{jj} Also at Facoltà Ingegneria, Università di Roma, Roma, Italy.
- ^{kk} Also at Scuola Normale e Sezione dell'INFN, Pisa, Italy.
- ^{ll} Also at University of Athens, Athens, Greece.
- ^{mm} Also at Paul Scherrer Institut, Villigen, Switzerland.
- ⁿⁿ Also at Institute for Theoretical and Experimental Physics, Moscow, Russia.
- ^{oo} Also at Albert Einstein Center for Fundamental Physics, Bern, Switzerland.
- ^{pp} Also at Gaziosmanpasa University, Tokat, Turkey.
- ^{qq} Also at Adiyaman University, Adiyaman, Turkey.
- ^{rr} Also at Cag University, Mersin, Turkey.
- ^{ss} Also at Mersin University, Mersin, Turkey.
- ^{tt} Also at Izmir Institute of Technology, Izmir, Turkey.
- ^{uu} Also at Ozyegin University, Istanbul, Turkey.
- ^{vv} Also at Kafkas University, Kars, Turkey.
- ^{ww} Also at Istanbul University, Faculty of Science, Istanbul, Turkey.
- ^{xx} Also at Mimar Sinan University, Istanbul, Istanbul, Turkey.
- ^{yy} Also at Kahramanmaraş Sütcü Imam University, Kahramanmaraş, Turkey.
- ^{zz} Also at Rutherford Appleton Laboratory, Didcot, United Kingdom.
- ^{aaa} Also at School of Physics and Astronomy, University of Southampton, Southampton, United Kingdom.
- ^{bbb} Also at INFN Sezione di Perugia, Università di Perugia, Perugia, Italy.
- ^{ccc} Also at Utah Valley University, Orem, USA.
- ^{ddd} Also at University of Belgrade, Faculty of Physics and Vinca Institute of Nuclear Sciences, Belgrade, Serbia.
- ^{eee} Also at Argonne National Laboratory, Argonne, USA.
- ^{fff} Also at Erzincan University, Erzincan, Turkey.
- ^{ggg} Also at Yildiz Technical University, Istanbul, Turkey.
- ^{hhh} Also at Texas A&M University at Qatar, Doha, Qatar.
- ⁱⁱⁱ Also at Kyungpook National University, Daegu, Korea.