

The Experiment for Space Radiation Analysis: Probing the Earth's Radiation Belts using a CubeSat Platform

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ESRA PDR

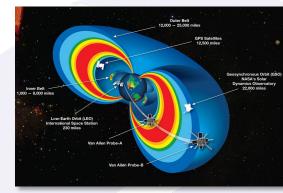
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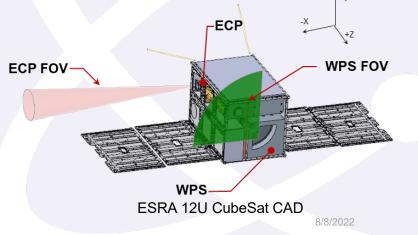


Mission Overview

- The Experiment for Space Radiation Analysis (ESRA) is a pathfinder for rapid Demonstration and Validation (DemVal) missions that LANL will fly on a 3 year cadence
- Primary goal is to provide TRL maturation (to TRL 6-8) and risk reduction for next generation systems
- ESRA will be one of the first CubeSats to operate in GTO
- Utilization of commercial spacecraft and ground station expertise allows for reduced risk and cost
 - Partnering with NanoAvionics for 12U bus
 - KSAT ground stations & services
- Enables LANL technical staff to focus on payload development
- Launch and integration provided by DoD Space Test Program (STP)



Earth's radiation belts illustrating the Van Allen Probe Mission in geostationary transfer orbit (GTO). Image courtesy of NASA





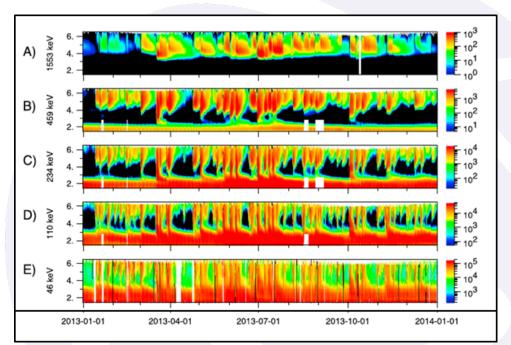
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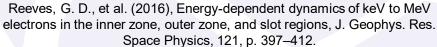
Mission Goals

- Legacy space weather instruments provide a wealth of data but are resource intensive
 - Size, Weight, and Power Cost (SWaP-C)
- Measurements of the local space environment are critical for anomaly resolution and addressing science questions
- A new generation of plasma and energetic particle sensors are being developed to provide the necessary data with lower SWaP-C
 - Wide-field-of-view Plasma Spectrometer (WPS)
 - Energetic Charged Particle (ECP) Telescope
- ESRA will achieve all mission requirements after six months
 - Bus shielding designed to last for 1 year based on the natural radiation environment
 - Payload anticipated to operate well beyond 1 year if the spacecraft can support
- After successfully demonstrating technology and providing risk reduction the payloads will contribute to scientific collection



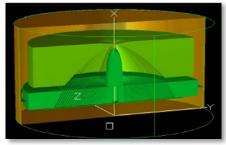
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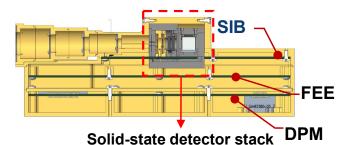


Sensor Payloads

Objective: On-orbit validation of the Wide-field-of-view Plasma Spectrometer (WPS) and Energetic Charged Particle (ECP) sensors' ability to measure proton and electron fluxes associated with the four major types of space environmental hazards: 1) total event dose; 2) single event effects; 3) deep dielectric charging; and 4) surface charging



Cross-section of full-size WPS CAD model



Cross-section of ECP telescope model

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Description:

- The WPS will measure ions 0.1 35 keV/q
 - The WPS design is based on pin-hole camera concept, allowing for significantly increased field-of-view (FOV) compared to standard electrostatic analyzers
- The ECP telescopes will measure protons (p+): ≥100 keV 500 MeV and electrons (e-): ≥100 keV – 20 MeV
 - CZT detectors allow for increased stopping power (shorter telescope with same energy range) when compared to standard silicon detectors
- Sensor performance can be compared against the previous AFRL Demonstration and Science Experiments (DSX), NASA Van Allen Probes, and GTOSat data sets
 - WPS/ECP data will also be compared against the AE9/AP9 model

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Critical Sub-System Payloads

Objective: Study and evaluate new technologies; Payload Processor, Distributed Processor Module (DPM), Flight Software (FSW) architecture, and High Voltage Power Supply (HVPS).



Payload Flight Processor evaluation board during bench top testing

Description:

- Provide technology maturation that can be evaluated for future LANL and NASA, missions
- Payload Processor (PP)
 - Employs a compact 3U Eurocard and open SpaceVPX architecture to reduce form factor while increasing performance
 - Will utilize quad-core LEON4 processor to provide lower cost, lower power, and high performance capability
- Distributed Processor Module (DPM) for "smart sensors"
 - Utilize low-power FPGAs, microcontrollers, and flash memory to provide C&DH at the sensor head location
 - Enables non-conventional layouts for sensor/instruments relative to central flight processors
- High Voltage Power System (HVPS)
 - Develop modular HVPS
 - VPX compliant form factor, standardize low voltage design, & plug in HV elements
- Flight Software (FSW)
 - Validate a generic and re-usable framework for future missions

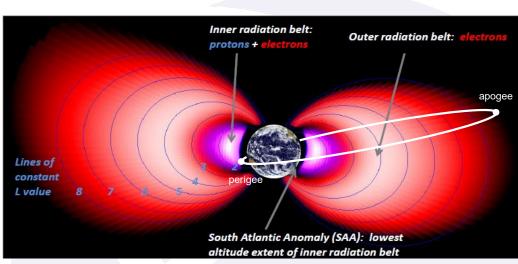


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Mission CONOPS

- Requesting Geosynchronous Transfer Orbit (GTO) through STP
 - Perigee: ~200 km
 - Apogee: 36,000 km
 - Period: ~10 hrs
 - Inclination: 40 +/- 20 deg
- The low perigee is typical for SpaceX and ULA launch vehicles (185-200 km)
- Simultaneous WPS & ECP operation
- High-rate downlink (3.4 Mbps) at perigee (~2×/day)
- Nominal orientation during operations solar panels facing sun
 - WPS faces away from the Sun
 - ECP faces perpendicular to magnetic field lines

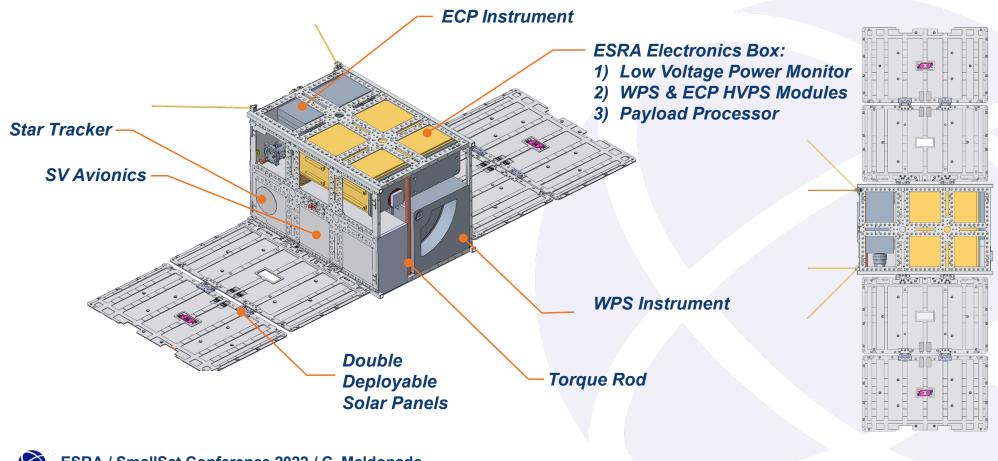


AE9/AP9/SPM: Radiation Belt and Space Plasma Specification Models

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ESRA Space Vehicle Overview



Comparison to DSX and Van Allen Probes

Energy	1 eV 10 eV	100 eV 1 keV 10 keV 100 keV 1 MeV 10 MeV 100 MeV 1 GeV
SWx:DSX Space Weather Experiment (ESPA Class ~600kg)	Key Proton Energies Electron Energies Ion Energies	LEESA: Low Energy ElectroStatic Analyzer w Energy Imaging Particle Spectrometer HIPS: High Energy Imaging Particle Spectrometer HEPS: High Energy Proton Spectrometer EASE: Compact Environment Anomaly SEnsor
Van Allen Probes (2 satellites, ~750 kg each)		HOPE: Helium Oxygen Proton Electron S: Magnetic Electron Ion Spectrometer
		REPT: Relativistic Electron Proton Telescope
ESRA (12U CubeSat~25kg)		WPS: Wide-field-of-view Plasma Spectrometer

Low SWaP ESRA Sensors cover similar Energy Range



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