

## Rotational remanent magnetisation as a magnetic mineral diagnostic tool at low rotation rates

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### Summary

Prior work on rotational remanent magnetisation (RRM) and rotational anhysteretic remanent magnetisation ( $ARM_{ROT}$ ) has demonstrated promise for magnetic mineral identification in earth materials. One challenge has been to calibrate the measurements to magnetic mineral types and microstructural controls, since previous studies have used differing spin rates, alternating field (AF) intensities and decay times, which hinders a comparison of datasets. Using a RAPID magnetometer we show that the range of usable practical rotation rates is 0.25 to 3 Hz [rps] which allows a wide range of RRM and  $ARM_{ROT}$  characteristics to be utilised (at 100 mT AF field, 100  $\mu$ T bias field). Sets of magnetic mineral extracts from sediments, and well characterised rock samples that contain the key magnetic minerals magnetite, pyrrhotite and greigite are used for a calibration of the RRM-  $ARM_{ROT}$  behaviour. Detrital pyrrhotite and pyrrhotite-bearing phyllites have largely small positive effective field ( $B_g$ ) values, with differences in  $B_g$  and  $ARM_{ROT}$  ratios at 0.5 and 2.5 Hz [rps] allowing grain-size discrimination. The positive  $B_g$  values, and changes in RRM and  $ARM_{ROT}$  with rotation rates allow distinction of pyrrhotite from magnetite and diagenetic greigite. Diagenetic greigite has  $B_g$  values of -83 to -109  $\mu$ T (at 0.5 Hz [rps]) and unusual RRM variation at low rotation rates caused by anisotropy affects. In contrast to previous work, based on crushed and sized natural magnetite at high spin rates,  $B_g$  for single domain magnetite from intact bacterial magnetofossils from Upper Cretaceous Chalk has some of the lowest  $B_g$  (0 -1  $\mu$ T) and displays a steep decline in  $ARM_{ROT}$  with increasing rotation rates. A simple tool for particle size characterisation of magnetite may be the ratio of  $ARM_{ROT}$  at spin rates 2.5 and 0.5 Hz [rps]. Stability of RRM is better studied using RRM acquisition with increasing AF field intensity, since static demagnetisation imparts a nuisance gyroremanence along the field axis. Mineral microstructure, dislocations and particle interactions are likely additional effects on RRM behaviour that need more investigation.

**Keywords:** Magnetic properties, Environmental magnetism, Marine magnetism and palaeomagnetism, Rock and mineral magnetism, Biogenic magnetic minerals, Magnetic mineralogy and petrology

### 1. Introduction

Rotational remanent magnetisation (RRM) is related to the mechanisms of gyroremanent magnetisation (GRM) acquisition when samples are rotated in a decreasing intensity alternating magnetic field. (Wilson & Lomax 1972; Stephenson 1980; Potter & Stephenson 1986; Madsen 2003). The RRM produced is either parallel or antiparallel to the rotation vector, and for magnetite may be preferentially acquired by single domain (SD) particles (Potter & Stephenson 1986). The fact that RRM is preferentially acquired in SD particles, or those with high magnetic stability makes it particularly attractive for investigating palaeomagnetic signal carriers, jointly with anhysteretic remanent magnetisation (ARM) properties.

It has been shown that for crushed magnetite, the intensity of RRM acquisition may also be related to particle size (Potter & Stephenson 1986), and for diagenetic greigite the magnitude of RRM is considerably enhanced compared to magnetite (Snowball 1997b; Stephenson & Snowball 2001). This difference provides a simple, diagnostic tool for identification of diagenetic greigite, but not necessarily bacterial magnetofossil-greigite (Snowball 1997a, 1997b; Reinholdsson et al. 2013; Chen 2014). RRM has not been detected in haematite (Wilson & Lomax 1972; Potter & Stephenson 1986) but appears to be acquired by pyrrhotite (Thomson 1990). Despite the apparent utility of RRM for mineral magnetic identification, it has been little used, which may be due to four reasons: 1) the lack of widespread 'off the shelf' equipment for RRM application, 2) the often-weak RRM generated in many earth materials, 3) in many prior studies high spin rates generating the RRM were not conducive to routine study of wet or poorly consolidated sediment sample, and 4) limited understanding of the various kinds of intrinsic and external controls on RRM and GRM production (Madsen 2003, 2004), especially at low rotation rates.

This work firstly addresses some of the practical issues with generating, measuring, and interpreting RRM at low rotation rates. We use measurements from the 2G Enterprises RAPID magnetometer (Kirschvink et al. 2008, hereafter called the RAPID), which has the ability for controllable sample rotation at low rotation rates, allowing routine automated RRM measurement on conventional samples. More than 20 RAPID magnetometers are currently installed worldwide, and therefore this could allow more widespread application of RRM behaviours for mineral magnetic identification. The RRM in many sediments is also weak, and the high sensitivity of SQUID magnetometers allow RRM to be measured in most samples. Prior commercial instrumentation (Stephenson & Molyneux, 1987) tackled the weak RRM intensity issue by rotation at large frequencies (i.e.  $\sim 100$  Hz), which increases the RRM intensity, but is not conducive to studying wet, fragile or imperfectly-shaped samples.

Secondly, we explore the utility of routine mineral magnetic identification, by using RRM behaviours on natural samples of known magnetic composition. These measurements allow us to suggest possible ways to effectively identify magnetic mineralogy in samples of unknown composition.

## 2. Quantification of rotational remanent magnetisation

RRM magnetisations are acquired with an alternating magnetic field (AF) is applied normal to the rotation axis of a specimen (Fig. 1), since RRM's are a special type of the GRM acquisition process (Madsen 2004). To standardize the quantification of RRM, Stephenson (1980) and Potter and Stephenson (1986) suggested the following:

- 1) An RRM magnetisation that is produced parallel to the rotation vector is signed as positive, and when opposite to the rotation vector, is signed as a negative RRM. We follow this convention (Fig. 1). Some studies have followed the opposite convention (Wilson & Lomax 1972; Stephenson 1976; Noel 1988; Edwards 1982b) and when discussing these studies, we use a conversion to the convention of Stephenson (1980).
- 2) The magnitude of the RRM is related to the abundance of magnetic phases, and usefully Potter & Stephenson (1986) introduced the  $B_g$  parameter which normalises the RRM by the ARM acquired under the same AF field and rotation conditions as the RRM. This kind of ARM is here symbolised as  $ARM_{ROT}$  (Fig. 1). The DC bias field used to generate the  $ARM_{ROT}$  is parallel to the rotation axis, but the AF field generating the ARM is aligned  $90^\circ$  to the rotation axis and DC field axis (Fig. 1). More widely used conventional methods of generation of ARM use co-axial AF and DC field axes to produce ARM (here referred as  $ARM_z$ ). For a few samples we also investigated the ARM acquired at zero rotation rate as a proxy for  $ARM_{ROT}$  (See SI section 3.1) as used in some prior studies (Snowball, 1997b).

Rotation rates are here expressed in revolutions per second (SI unit Hz) widely indicated as 'rps' in much prior work, and here symbolised as  $\omega$  in Hz [rps] units.

B<sub>g</sub>, the effective gyromagnetic field in micro-Tesla (Potter and Stephenson, 1986) is defined as:

$$B_g = B_a \times \text{RRM} / \text{ARM}_{\text{ROT}} \quad \text{Eq.1}$$

where B<sub>a</sub> = DC field used in generating the ARM (usually around 50- 150  $\mu\text{T}$ )

### 2.1 Key attributes of RRM

1) The magnitude of RRM is partly dependent on the rotation rate. When the rotation rate exceeds the frequency (in Hz) of the AF field (i.e., the *AF threshold*) most magnetic materials have a greatly enhanced RRM intensity. This issue has led in a strong focus on RRM studies at large rotation rates (Table 1). We refer to low spin rates, as below the AF-field threshold and high spin rates above this—hence the separation of existing RRM datasets into Tables 2 and 3.

2) Magnetite, maghemite and greigite have positive RRM above the AF threshold, and largely negative (but not entirely negative), and much reduced RRM below this threshold. Conversion of B<sub>g</sub> values between high spin to low spin values is open to much uncertainty— Snowball (1997b) used a conversion factor of -0.08 for transforming a magnetite B<sub>g</sub> from 108 to 5 Hz [rps]. The synthetic Mapico black magnetite (Potter & Stephenson, 1986; Edwards & Desta, 1989) has been used in both low and high spin studies suggesting an alternative conversion factor of ca. -0.027 between 108 Hz [rps] and 0.07 Hz [rps]. For diagenetic greigite a conversion factor for B<sub>g</sub> of  $-0.14 \pm 0.06$  for rotation rates between 95 Hz [rps] to 2.5 Hz [rps] is possible with values in Tables 2, 3.

3) Multidomain (MD) magnetite and haematite have no or little appreciable RRM (Wilson & Lomax 1972; Potter & Stephenson 1986; Snowball 1997b). This is exemplified by basalt sample D264-2a (Tables 1, 3) studied by Edwards (1982a, b) and Edwards & Desta (1989). This behaviour of MD magnetite is a useful property of RRM, since it implies that SD grains can be studied in mixed SD-MD earth materials that would acquire appreciable MD-related magnetisations using other magnetic investigation methods.

4) B<sub>g</sub> for maghemite ( $\gamma\text{Fe}_2\text{O}_3$ ) may be in part concentration dependent, with larger B<sub>g</sub> values in samples with larger concentrations of  $\gamma\text{Fe}_2\text{O}_3$  (Madsen 2003). Synthetic  $\gamma\text{Fe}_2\text{O}_3$  has some of the largest |B<sub>g</sub>| values for Fe-oxides at both low and high spin rates, but natural  $\gamma\text{Fe}_2\text{O}_3$  has so far not been studied (Tables 2, 3).

5) CrO<sub>2</sub> recording tape particles have negative RRM above the AF threshold, and a largely positive RRM below the threshold (Madsen 2003), both in the opposite sense to that of magnetite (Table 2). Although Snowball (1997b) measured negative RRM at 5 Hz [rps] (Table 3), and Madsen (2003) negative RRM at 25 Hz [rps].

6) Heated pyrrhotite (either monoclinic or hexagonal) appears to display positive RRM both below and above the AF threshold (Tables 2,3), with rapid changes of RRM intensity close to the AF threshold. Unheated pyrrhotite studied by Thomson (1990) displayed negative RRM at 3 Hz [rps].

7) At AF frequencies below the AF threshold, the frequency of the field used may have little impact on the changes of RRM with rotation rate, or at least at rotation rates up to 0.1 Hz [rps] investigated by Edwards (1980a). However, the impact of differences in AF frequency at rotation rates further towards the AF threshold has not been investigated experimentally. However, some theory suggests the ratio of the rotation rate and AF frequency should in part control the strength of RRM acquisition (Stephenson, 1980, 1985)

8) At low rotation rates the time of decay of the AF field ( $T_D$  in seconds; Fig. 1) has a large impact on the changes of RRM with  $\omega$  (Edwards 1982b), such that changes in RRM properties are best compared when expressed with respect to  $\omega T_D$  (in revolutions) rather than using  $\omega$ .

9) RRM acquisition is not an instantaneous process, some ca.  $90^\circ$  (or larger) of sample rotation at peak AF field is required for RRM to be fully acquired (Edwards, 1982b, Roperch & Taylor 1986). In much demagnetisation equipment this is impacted by the AF peak field hold time ( $T_H$ ), which ideally should be long enough for this rotation 'initialisation' to be achieved.

10) An RRM induced by a particular AF field may require a larger static demagnetising AF field to fully demagnetise the induced RRM (Edwards, 1980b). This effect is likely related to the angular dependency of the switching field (Madsen, 2004). Like demagnetising ARM or IRM (Stephenson, 1983) the intensity-decay curve produced during demagnetisation may differ between static or tumbling demagnetisation methods.

Whilst some is understood about the general RRM behaviour of common magnetic material in earth materials, the range of rotation rates, AF and DC fields used in previous studies (Tables 2 and 3) and the hold and decay rates of the AF inducing field, hinders a general and detailed comparison of the mineral magnetic behaviours.

This work focusses on five aspects of RRM behaviour at rotation rates  $\leq 3$  Hz [rps]: 1) the variation of RRM,  $ARM_{ROT}$  and  $B_g$  with rotation rate, 2) the magnitude and sign of  $B_g$  and RRM, and 3) the stability of the RRM as measured by the median destructive field (MDF)— since MDF of the RRM has been shown to be a useful additional parameter in characterising particle size and mineralogy (Potter & Stephenson, 1986; Snowball, 1997b).; 4) the impact of differences in peak AF hold times ( $T_H$ ); 5) exploration of how RRM datasets could be used for magnetic mineral identification at  $\omega T_D \leq 4.5$  revolutions.

### 3. Mineralogy of test samples

The sample sets used here comprise two types, magnetic concentrates (extracts) and natural rock samples of known magnetic mineralogy. These sample sets comprise:

1) Detrital monoclinic pyrrhotite from river sediment, collected in the upper reaches of the mountainous Zhoushui River (at 23.786836 N, 121.01097 E, Taiwan), which is sourced from the pyrrhotite-bearing metamorphic terranes in the Central Range (Horng and Huh, 2012). Pyrrhotite grains in sediments were extracted using a rare earth magnet and were purified magnetically several times. The purified sample was then subdivided into five micro-sized fractions (<5, 5-10, 10-20, 20-38 and 38-63  $\mu\text{m}$ , samples PY1 to PY 5 respectively; Table 4; Fig. 2a) with sieves and membrane filters. XRD analyses determined their purities (Fig. S1). The larger detrital particle sizes show some minor evidence of oxidation to lepidocrocite and goethite (see supplementary information (SI) Fig. S1).

2) Magnetic extracts using the method of Hounslow & Maher (1996, 1999). The magnetic extracts broadly represent two groups, those from sediments with intact or near-intact magnetic oxide assemblage (e.g., like Franke et al. 2007), and a set with *residual* discrete Fe-oxide components (and Fe-oxides in silicates) after early sulphidic diagenesis or deep-burial diagenesis (e.g., Hounslow et al. 1995; Hounslow, 1996; Maher & Hallam, 2005; Table 4). The extracts are from: a) the late Triassic Lunde Formation (LU codes) from the North Sea, containing a residual suite of silt-sized magnetic minerals after extensive diagenetic dissolution, with in some cases minor discrete magnetic oxides, but abundant Fe-oxide inclusions in associated silicates (Hounslow et al. 1995). B) Quaternary and



Pliocene sediments from the Owen Ridge (OR sample codes) and Madingley Rise (codes MR; on the Mascarene Plateau) both in the Indian Ocean. The Indian Ocean samples have detrital sources exclusively of well mixed, far travelled aeolian origin (Table 4). Some of these extracts also represent residual magnetic assemblages (Table 4). Extract samples with codes LU, OR and MR1 to MR3 are generally poor in fine-grained magnetic particles and represent the magnetic detrital fraction mostly  $>2 \mu\text{m}$  in size (i.e., the  $E_{MP}$  extract type of Hounslow & Maher, 1999). Samples MR4 to MR6 are  $E_{MPT}$  type extracts of Hounslow & Maher (1999) and contain largely detrital grains  $<2 \mu\text{m}$  in size and in some cases contain accessory magnetite magnetofossils (i.e. MR4 and MR5). The  $E_{MPT}$  extracts may contain the bulk of the discrete-grain SD-sized particles. This is not to say that the  $E_{MP}$  extracts contain no SD-like magnetic grains, since these may be within the larger detrital grains. The  $E_{MP}$  extracts contain variable amounts of Fe-oxides as inclusions in silicates, reflected in the variable content of silicates in the extracts (see SI for details). Therefore, these extracts are representative of the key magnetic components often found in sediments of various ages, and those residual magnetic minerals elevated in relative abundance during diagenetic-related Fe-oxide dissolution (Roberts, 2015).

3) Monoclinic pyrrhotite bearing phyllites (samples PY6 to PY9) from Taiwan (Fig. 2B) with relatively coarse grain size. Hysteresis data indicate  $B_{cr}/B_c$  of 1.10 to 1.12 and  $M_{rs}/M_s$  of 0.66 to 0.72 for these samples (SI Table S1). These are typical values for metamorphic pyrrhotite and are not necessarily indicative of SD properties (Horng 2018).

4) Greigite samples from Plio-Pleistocene greigite-containing mudstones from the Lower Gutingkeng Formation, SW Taiwan (Horng et al. 1998; Jiang et al. 2001; sample codes GR1 to GR4; Table 4). Hysteresis data indicate  $B_{cr}/B_c$  of 1.26 to 1.39 and  $M_{rs}/M_s$  of 0.54 to 0.63 for these samples (SI Table S1). Chen (2014) also measured the RRM in two samples from this location at high-spin rates (Table 2). The diagenetic greigite in these samples is very fine-grained and widely dispersed through the phyllosilicate matrix (Fig. 2C).

5) Whole rock samples of the Upper Cretaceous Chalk (CC code samples), from southern England. These are from the stratigraphic level of sample CW9b originally figured by Montgomery et al. (1998) which contains abundant chains of prismatic magnetofossils (Fig. 2D) in extracts. These are also figured in Hounslow & Maher (1999) and Kopp & Kirschvink (2008) (Table 4). The magnetofossil chains in the extracts seem to reflect in-situ preservation, since other extracts from the Chalk from elsewhere (unpublished data of Hounslow) do not show such high abundance of intact chains (Fig. 2D). These samples also contain an associated assemblage of Fe-oxides as inclusions in detrital quartz and feldspars but have very little detrital Fe-oxides outside silicate hosts (see SI for details).

6) Relatively unaltered fine-grained dolerites (SVD code) with a content of typical titanomagnetite, which has undergone high temperature alteration to ilmenite-magnetite intergrowths (Table 4), with margins altered to maghemite. Clots of magnetite and acicular rods of Fe-oxide occur in the groundmass (Gayer et al. 1966; Krumsiek et al. 1968; Halvorsen 1973). Some of the prior work using low rotation rate RRM was performed on similar basaltic rocks (Tables 1, 3).

7) Samples of siltstones and sandstone from the De Geerdalen Formation from several locations in Svalbard (Table 4), which have a rather similar detrital magnetite magnetic mineralogy throughout (Hounslow et al. 2007; 2022). The succession has been subjected to complex diagenetic changes (Mørk 2013; Haile et al. 2018), which will have removed much of the original discrete Fe-oxides.

Synthetic samples have been much used in previous RRM studies (Edwards 1980a; 1980b, Potter and Stephenson 1986; Snowball 1997b; Madsen 2003), but the intention here was to focus on naturally

derived materials, since synthetic samples can have unusual morphologies, microstructure and purity not seen in natural minerals.

To prepare samples of the magnetic extracts and detrital pyrrhotite to measure they were dispersed in a 5 ml droplet of PVA glue on an acetate sheet, which was allowed to dry at room temperature overnight. When dry this was peeled off, folded into a ball, and placed in a plastic pot (with cling-film padding) as the sample to measure. There was some inevitable clumping of magnetic particles during the drying process, an inevitable consequence of using granular magnetic materials, which will have resulted in some magnetic interaction (clumping observed in most) and anisotropy affects (Cisowski 1981). The pyrrhotite bearing phyllites were gently crushed and mounted in a plastic pot. Other samples were core-plugs or cubes with no encapsulation.

#### 4. Experimental procedures and issues

Measurements used the Lancaster University RAPID magnetometer. This RAPID is housed in a large Helmholtz cage for field cancellation, which together with the Mu-metal shields cancels the earth's magnetic field to around 0.03-0.08  $\mu\text{T}$  at the demagnetiser coil position. AF frequencies are 327 Hz on the transverse-axis coil (Fig. 1), and 360 Hz on the Z-axis coil. The Lancaster RAPID has a standard duration of peak AF field of 30 AF cycles giving hold times ( $T_H$ ) of 92 ms and 83 ms for transverse and Z-axis coils respectively.  $T_H$  can be changed but is kept low to limit coil overheating. The ramp-up and ramp-down times are fixed at ca. 0.64 s and 1.53 s (i.e.,  $T_D$ ) respectively, irrespective of maximum field used (See SI section 4 for further details).

The RAPID vacuum system that holds samples onto the silica glass rod for insertion into the measurement space, is unable to routinely hold-onto rock samples at  $\omega > 3$  Hz [rps], unless special precautions are taken to fix the sample onto the quartz-glass rod. This essentially limits the routine practical RRM measurements of usual-sized rock samples to rotation rates of  $\leq 3$  Hz [rps]. The software set 'rotation rate' is half that actually achieved (i.e., set values of 1, 2, 6 correspond to 0.5, 1 and 3 Hz [rps] respectively). This was calibrated at rotation rates 0.1, 0.15, 0.2 Hz [rps], which yielded a linear relationship with time, and extrapolated to rates up to 3 Hz [rps].

All weak magnetisation measurements were performed with a 'measurement blank' appropriate for each sample set (i.e., a blank subjected to all the same steps). This bypassed the 'standard holder-correction procedure' on the RAPID, which does not account for magnetisation and demagnetisation of holders and sample pots during measurements. The GM4Edit software does the blank-type corrections to the data (Hounslow 2019).

All samples were set into an initial standard state by AF demagnetisation, sequentially along magnetometer X, Y axes at 110 mT, followed by demagnetisation only along the Z-axis at 150 mT. A larger field along Z-axis was applied to reduce possible GRM's along the Z-axis from the X and Y AF applications, which can persist into higher coercivity fields than the applied field (Edwards 1982b, Madsen 2003). All RRM and  $\text{ARM}_{\text{ROT}}$  measurements were performed with a 100 mT peak AF field. Here the sign of  $\omega$  indicates the sense of spin coded into the RAPID software. A  $-\omega$  rotation rate has the rotation vector directed **down** (to +Z), and  $+\omega$  rate the rotation vector directed **up** (to -Z; Fig. 1). The sign of  $\omega$  therefore also indicates the down or up-sense of the rotation vector. Hence, positive RRM will be a larger magnetisation along the Z-axis with a  $-\omega$  rotation rate (vice versa for a negative RRM; Fig. 1). It is implicitly assumed that at 1 Hz [rps] one rotation corresponds to 327 cycles of the transverse AF field.

RRM generation follows the procedure of Wilson & Lomax (1972) in which the sample was sequentially spun in opposite senses (first about the Z-axis with  $-\omega$ ). The RRM is  $\frac{1}{2}$  the difference

between the magnetisation measurements at  $+\omega$  and  $-\omega$  (Fig. 1). The sample spins throughout the AF ramp-up, hold and ramp-down interval. Other AF fields use the same hold and ramp-down times, but the AF ramp-down-rate (i.e. mT/sec) is adjusted to fit these times (inset in Fig. 1). For weak RRM's it is often advantageous to perform repeats, since the noise inherent in determining the RRM can be larger than the magnetisation measurement standard deviation (SD) along the Z-axis (See SI for noise details). This problem is amplified if there is residual undemagnetised remanence (e.g., from  $>150$  mT coercivity remanence) along the Z-axis, since the SD of magnetisation measurements along any axis is strongly related to the moment along that axis (See SI Fig. S10). In a practical sense this dictates: 1) the lower limit of useful RRM measurement for those samples with a significant intrinsic remanence which cannot be AF demagnetised. That is, the presence of significant haematite or goethite- magnetisations limits the applicability of using RRM (as does very low abundances of magnetic minerals). 2) The RRM and  $ARM_{ROT}$  measurements should be performed prior to any other large field magnetisations.

The  $ARM_{ROT}$  is determined with a DC bias field of  $100 \mu\text{T}$  along the Z-axis (AF of  $100$  mT along the Y-axis; Fig. 1), and by spins in opposite senses (spun about Z-axis with  $-\omega$  first; Fig. 1). The  $ARM_{rot}$  is the average of the two magnetisation measurements. Theoretically, the difference between these two should also yield the RRM (Stephenson & Molyneux 1987, here symbolised as  $RRM_{100}$ ), although in practice the additional large  $ARM_{rot}$  makes the determination of  $RRM_{100}$  noisier than without the DC field (See SI section 3.2). In addition, we used the same field settings and measured the ARM acquired at zero rotation rate ( $ARM_{trans}$ ). This could be used as a proxy for  $ARM_{ROT}$ , if equipment was not available to rotate samples during ARM acquisition (this is examined in SI section 3.1). Rotational ARM can also be produced by rotating samples in orthogonal weak and strong and declining DC fields as an alternative (Stephenson, 1988). A conventional ARM was also determined along the Z-axis by applying a  $100$  mT AF and  $100 \mu\text{T}$  DC coaxial along the Z-axis (called here  $ARM_z$ ) allowing comparison to other standard ARM datasets. Measuring both  $ARM_{ROT}$  and  $ARM_z$  provides additional mineral-diagnostic values. Care was taken in determining the zero-level for all ARM measurements by measuring residual magnetisation after AF demagnetisation at  $150$  mT (along Z, ramp down time of  $1.39$  sec) before and after the ARM determinations (see SI for details).

For the magnetic extracts a blank comprising a dry PVA droplet + pot + quartz-silica rod has a mean RRM of  $\sim -0.10 \pm 0.2 \times 10^{-10} \text{ Am}^2$  (at  $0.5$  Hz [rps]),  $ARM_{rot}$  of  $\sim 1.7 \pm 0.5 \times 10^{-10} \text{ Am}^2$  (at  $0.5$  Hz [rps]), and  $150$  mT demagnetised Z-axis moment of  $\sim 0.4 \pm 0.3 \times 10^{-10} \text{ Am}^2$  (uncertainties are  $1\sigma$ ). The RAPID quartz-silica rod used has a mean Z-moment of  $\sim 0.15 \pm 0.4 \times 10^{-10} \text{ Am}^2$  ( $1\sigma$ , after  $150$  mT demagnetisation). The average uncertainty ( $1\sigma$ ) of a single magnetisation measurement (average of four in total on the Z-axis) for the quartz-silica rod is  $\sim 0.01 \times 10^{-10} \text{ Am}^2$ .

The median destructive field (MDF) of the RRM was determined in two ways: 1) after RRM acquisition from the  $+\omega$  rotation state (followed by static AF demagnetisation along the Z-axis in  $10$  mT increments, up to  $90$  mT, followed by one at  $150$  mT), and 2) for both  $+\omega$  and then  $-\omega$  rotation states (both with static AF demagnetisation on Z-axis, in  $10$  mT increments to  $150$  mT).

In method-1 the zero-'base line' of the RRM decay was judged using the flat-tail of the RRM decay curve and the  $150$  mT step. In method-2, by subtracting the  $+\omega$  and  $-\omega$  RRM demagnetisation curves, and estimating the zero-'base line'tail of the resulting RRM decay (see SI for details). In each case a linear regression trend using a few data points either side of the MDF was fit to the normalised moment to field intensity changes, from which the median point was estimated in Excel. The MDF of the  $ARM_z$  was similarly determined but using the two points straddling the median point. An alternative simpler proxy for the ARM stability is the proportion of ARM remaining after  $40$  mT

demagnetisation (Peters & Thompson, 1989), which is here symbolised at  $d.ARM_{Z_{40mT}}$ . Comparison of MDF and  $d.ARM_{Z_{40mT}}$  between different methods of demagnetisation can be problematic, since MDF's produced by static, single axis methods can be some 1.3-1.5 times those produced by tumbling demagnetisation (Stephenson, 1983). For example, the seminal work of Dunlop (1983) on basalts seems to have used single axis demagnetisation (West & Dunlop, 1971) and that of Potter and Stephenson (1986) used tumbling. The RAPID protocol `rmg-template` files which run these procedures are in the Supplementary information.

## 5. Results

Since RRM and  $ARM_{ROT}$  vary greatly in magnitude the RRM and  $ARM_{ROT}$  are normalised by the absolute value at 0.5 Hz [rps]. This allows an inter-sample comparison of changes with rotation rates, removing the effect of differences in magnetic mineral abundance. The reasons for using this somewhat arbitrary value for normalisation are outlined in section 5.3.

### 5.1 Changes of RRM, $ARM_{rot}$ and $B_g$ with rotation rate for magnetite.

Like Edwards' (1982b) we find the RRM has predominantly negative values throughout the rotation rate range (Figs. 3a, 4a, 5a). This is except for sample SVD2 at  $\omega > 1.3$  Hz [rps] which has a positive RRM (Fig. 4a). The 3 Hz [rps] step for CC2 (Fig. 4a) is probably subject to a small flux jump on the RAPID, hence its erroneously large, normalised value. One of Edwards' (1982b) samples F37B-1a shows a similar behaviour to SVD2, although a sister specimen measured by Wilson & Lomax (1972)-F37B-2 did not. Most of the samples display a minimum in RRM between 0.3-0.4 Hz [rps]. Minimums have typically also been found in other magnetite bearing rocks at low, but differing rotation rates (Potter & Stephenson 1986; Stephenson & Snowball 2001). This general pattern is consistent except for sample MR2 (rather noisier data), which has a rather flatter RRM- $\omega$  curve (Fig. 3a), like specimens E23a-1 and L-S1-2 of Wilson & Lomax (1972).

The normalised  $ARM_{rot}$  have mostly rather flattish curves with a small decline in  $ARM_{rot}$  at  $\omega > 0.3$  Hz [rps] for the magnetic extracts and the dolerite samples (Fig. 3b, 4b). Contrastingly rather steeply declining  $ARM_{rot}$  for larger  $\omega$  occur for the Chalk and De Geerdalen sandstone/siltstone samples (Figs. 4b, 5b). Significantly,  $ARM_{rot}$  changes with  $\omega$  have not been investigated much previously, although changes observed have been inferred as an interaction between the RRM and the  $ARM_{rot}$  (Stephenson & Snowball 2001). The transverse ARM produced without rotation ( $ARM_{TRAN}$ ) is most often slightly lower than that produced at  $\omega = 0.05$  Hz [rps] (Figs. 3b, 4b, 5b).

Normalised  $B_g$  values largely mirror the changes in normalised RRM for many samples especially those with a negative peak in  $B_g$  less than  $-3 \mu T$  (Figs. 3c, 4c, 5c). The two Chalk samples (CC1, CC2) show the flattest curves and lowest  $B_g$  values between 0 and  $-1 \mu T$  throughout the rotation rate range (Fig. 4c). The changes in  $B_g$  with  $\omega$  are rather noisy due to weak RRM's for MR2 and OR3 (Fig. 3c) and are perhaps closer to  $B_g$  behaviours shown by SVD1, DF2, DF5, DF6 which have stronger RRM's (Figs. 4c, 5c).

The lower  $B_g$  values for the magnetite magnetofossils in the Chalk samples compared to the silt-sized magnetic extracts are opposite to the negative relationship with magnetic grain size noted by Potter & Stephenson (1986) for crushed and sized-magnetite. Chen (2014) also noted the ca.10 to 100 times smaller  $B_g$  values for several sets of magnetite magnetosome samples compared to the samples measured by others at high spin rates (Table 2). The measurements here therefore concur with those of Chen (2014), although ability to directly compare  $B_g$  values between low and high spin datasets is hindered. The sample Pajep8 of Snowball (1997b) containing magnetite magnetofossils

and MD magnetite with a  $B_g$  of  $-0.14 \mu\text{T}$  at 5 Hz [rps] (Table 3) is also consistent with these observations.

The  $B_g$  values and changes with  $\omega$  for samples from the De Geerdalen Formation are rather like the magnetic extracts (Fig. 3c, 5c), a feature which might be expected considering initial sulphidic diagenesis typically removes much of the finer grained discrete magnetite content in sediments (Roberts 2015). Compared to the other samples, dolerite SVD1 has a much larger  $B_g$  value ( $-26 \mu\text{T}$  at 0.5 Hz [rps]). Edwards & Desta (1989) have measured similar values in synthetic  $\gamma\text{Fe}_2\text{O}_3$  ( $-30 \mu\text{T}$ ,  $-37 \mu\text{T}$ ; Table 3) and Roperch & Taylor (1986) in a Miocene basalt sample ( $-25 \mu\text{T}$ ) at low  $\omega$ .

## 5.2 Changes of RRM, $\text{ARM}_{\text{ROT}}$ and $B_g$ with rotation rate for pyrrhotite and greigite.

Greigite produces normalised RRM changes with  $\omega$  like that of magnetite, but with a minimum (negative) RRM at 0.5 Hz [rps], and some larger fluctuations at lower  $\omega$  values (Fig. 6a). Pyrrhotite samples do not display RRM behaviour like magnetite and greigite, having positive RRM for part or all the range of  $\omega$  investigated (Figs. 7a, 8a). Generally, all the pyrrhotite samples had RRM which was challenging to measure (i.e., producing noisy data), since they contained a strong residual remanence remaining after 150 mT demagnetisation, so much of the sample data is from averaging of duplicate measurements. The detrital pyrrhotite  $>20 \mu\text{m}$  in size (PY4, PY5) produces a peak in positive RRM at 0.3-0.4 Hz [rps], and then declines towards near zero or negative RRM by 3 Hz [rps] (Fig. 7a). Detrital pyrrhotite  $<20 \mu\text{m}$  (PY1-PY3) produces a trend towards increasingly more positive RRM through 0.5 -2.5 Hz [rps]. The  $< 2 \mu\text{m}$  detrital pyrrhotite has the largest positive normalised RRM at  $\omega >1$  Hz [rps] and has some possible step-like behaviour in RRM at  $\omega <0.4$  Hz [rps], like greigite (Figs. 6a, 7a)- although this may be measurement noise in this quite weak sample. The pyrrhotite-bearing phyllite samples (PY6 to PY9) show RRM behaviour with  $\omega$  rather like the coarser detrital pyrrhotite (Fig. 8a) suggesting these largely contain pyrrhotite  $> 20 \mu\text{m}$  in size. Such large particle sizes are also seen in electron microscopy (Fig. 2B). A positive RRM was detected by Thomson (1990) in samples heated to 228-315°C but generally much smaller negative RRM's in unheated or those heated to below 228°C (at 3 Hz [rps]). Thomson (1990) attributed these temperature-related changes to creation of hexagonal pyrrhotite during the heating. The RRM produced by Thomson (1990) was weakly if at all dependent on rotation rates at  $\omega <10$  Hz [rps].

Changes in  $\text{ARM}_{\text{ROT}}$  with  $\omega$  for the sulphides are rather similar and flattish at  $> 0.4$  Hz [rps] with all showing a progressive increase in normalised  $\text{ARM}_{\text{ROT}}$  at  $\omega > 0.5$  Hz [rps] (Figs. 6B, & 7B, 8B), and like magnetite have  $\text{ARM}_{\text{TRAN}}$  values rather similar or lower than the value at 0.05 Hz [rps].

Changes in  $B_g$  largely reflect the changes in normalised RRM, since any changes in  $\text{ARM}_{\text{ROT}}$  are rather suppressed compared to RRM changes with  $\omega$  (Fig. 7C). The pyrrhotite samples show mostly positive  $B_g$  at  $\omega \geq 0.5$  Hz [rps]. Slotznick et al. (2016) has apparently detected a range of both positive and negative  $B_g$  values in pyrrhotite bearing rocks (at  $\omega \geq 1$  Hz [rps]), although the methodology used is unclear, since their ' $B_{\text{eff}}$  values' seem to exist for both  $-\omega$  and  $+\omega$ .

## 5.3 Impact of changing AF field hold times on RRM and $\text{ARM}_{\text{ROT}}$

Edwards (1982b) demonstrated that for magnetite the AF field ramp down time ( $T_D$ , in seconds) controlled the position of the peak of the negative RRM in magnetite-bearing samples, such that  $\omega T_D$  (in revolutions) was a more fundamental expression of the RRM changes than the rotation rate. For the magnetite-bearing samples of Edwards (1982b) the peak in negative RRM was at  $0.40 \pm 0.02$  rev. The apparent peak in our magnetite-bearing samples is at  $\omega T_D = 0.4$  to 0.75 rev (Figs. 3,4,5), generally somewhat larger than Edwards observed. However, our RRM initialisation is different to that used by Edwards, since our hold time,  $T_H$  (92 ms) is insufficient to produce the 90° of rotation

initialisation used by Edwards. For three samples we measured RRM and  $ARM_{ROT}$  with rotation rate for  $T_H$  at 0.917 s and 9.17 s (300 and 3000 cycles setting on RAPID; Fig. 9). A  $T_H$  of 0.917 s produces the required 90° (or larger) of initialisation rotation at  $\omega \geq 0.27$  Hz [rps], ( $\omega T_D = 0.42$  revolutions or larger using  $T_D = 1.53$  s) and the later exceeds the 90° threshold at all rotation rates used (Fig. 9A,C,E). Both these larger  $T_H$  produce a peak in RRM close to  $\omega T_D = 0.4$  rev for pyrrhotite and greigite (Fig. 9C,E), as does dolerite sample SVD2 (SI Fig. S19D), but the magnetite sample MR1 has its RRM peak at 0.6 rev (0.4 Hz [rps]; Fig. 9A). These data broadly concur with Edwards (1982a) experimental data but do suggest that perhaps there may be additional mineralogical control on the precise position of the RRM peak in  $\omega T_D$  units.

Broadly the RRM using  $T_H = 92$  ms are a little smaller than that acquired over longer hold times at rotation rates 0.5 to 3 Hz [rps], and diverge more at rates  $< 0.5$  Hz [rps] (Fig. 9,A,C, E). For the magnetite (MR1) and pyrrhotite (PY5) samples at  $\omega = 0.3$ -3 Hz [rps], the RRM at 92 ms hold time is some  $96 \pm 1.6\%$  and  $105 \pm 13\%$  ( $1\sigma$  values) of that for the average of RRM at the longer hold times. Greigite (sample GR1) in contrast, has RRM some  $88 \pm 3\%$  of the RRM for  $T_H = 9.17$  s, over the 0.5- 3 Hz [rps] range (Fig. 9E), but substantial deviations at rates  $< 0.5$  Hz [rps] (Fig. 9E). The greigite RRM is more consistent between the 0.917 s and 9.17 s hold times at  $\omega \geq 0.3$  Hz [rps] above the 90° initialisation threshold (Fig. 9E).

For magnetite (MR1) and pyrrhotite (PY5) the  $ARM_{ROT}$  (for  $T_H = 92$  ms), is on average marginally larger than both the longer hold times (Fig. 9B,D), with the average percentage comparison being  $102 \pm 1.4\%$ ,  $104 \pm 1.2\%$  for  $T_H = 0.917$  s, and  $102 \pm 1.4\%$ ,  $102 \pm 1.1\%$  for  $T_H = 9.17$  s respectively (for  $\omega = 0.1$  to 3 Hz [rps]). In contrast, greigite displays lower  $ARM_{ROT}$  at  $T_H = 92$  ms with the average comparison being  $93 \pm 2.0\%$  and  $89 \pm 2.9\%$  for the 0.917 s and 9.17 hold times (at  $\omega \geq 0.5$  Hz [rps]). The greigite (GR1) behaviour at  $\omega < 0.5$  Hz [rps] is quite different to the other samples and indicates a complexity in  $ARM_{ROT}$  acquisition not seen in other test samples. The impact of  $T_H$  differences on Bg are like those percentage comparisons seen in RRM for magnetite (MR1) and pyrrhotite (see SI Fig. S19). For greigite sample GR1, the peak in Bg is also larger at  $\sim 102$   $\mu$ T and displaced to  $\sim 0.2$  to 0.3 Hz [rps], compared to that in Fig. 6c.

The complexity in the RRM and  $ARM_{ROT}$  behaviour at  $\omega < 0.5$  Hz [rps] (also having the AF hold time ( $T_H$ ) shorter than the optimum for RRM initialisation), led us to choose a blanket normalisation for parameters at 0.5 Hz [rps], like is displayed in Figs. 3 to 8. Values at this rotation rate and higher generally show more consistent changes. Had we used a  $T_H$  of  $\sim 0.92$  s a more optimum value to normalise the RRM,  $ARM_{ROT}$  and Bg values would have been that close to the peak in RRM at  $\omega T_D \approx 0.4$  revolutions (i.e., about 0.26 Hz [rps]). Hence, the optimum hold time on the 327 Hz RAPID demagnetiser is a  $T_H$  of 0.962 s (314 cycles) which would detect the maximum RRM at  $\omega T_D \approx 0.4$  rev (0.26 Hz [rps] at  $T_D = 1.53$  s). However, this has consequences for potential excess coil heating, and the increased refinement has only minor impact on magnetite and pyrrhotite parameters. However, this is a bigger issue for samples containing diagenetic greigite (i.e., Fig. 9E, F).

We suspect the RRM behaviour at  $\omega < 0.5$  Hz [rps] of greigite sample GR1 (at  $T_H = 92$  ms) is due to anisotropy and its impact on RRM and  $ARM_{ROT}$  acquisition (Potter, 2004) at values below the rotation initialisation threshold. Although we could not investigate the directional dependence of RRM and  $ARM_{ROT}$  on GR1, we performed a preliminary study on sample SVD2 which shows similar behaviour to GR1 when rotation was used about the other two axes of this cubic sample, to generate the RRM (See SI section 3.4). For SVD2 this impacts the RRM and  $ARM_{ROT}$  and Bg values, demonstrating that anisotropy is an additional source of changes in Bg with  $\omega$ . However the pattern of changes in Bg,

$ARM_{ROT}$  and RRM are similar with respect to  $\omega$ , if above the rotation initialisation threshold (SI Fig. S19).

In a practical sense these issues suggest the reliable range of  $\omega$  for routine use on the RAPID is 0.26 and 3 Hz [rps], if using an AF hold time sufficient to reach rotation initialisation. Samples with significant ferrimagnetic anisotropy will have RRM-data most-impacted if there is inadequate rotation initialisation. However, for many samples with weak anisotropy the AF hold-time may have a limited impact especially at  $\omega \geq 0.5$  Hz [rps].

#### 5.4 Demagnetisation of the RRM

Demagnetising of the RRM using method 1, produces an additional ‘nuisance’ magnetisation created in most samples, since the ‘background’ 150 mT demagnetisation on the Z-axis and the mean of the two RRM magnetisation measurements (in opposite rotation directions) were not similar. This is demonstrated well by the greigite sample GR1, measured using method 2, which shows this effect very strongly (Fig. 10a), in which coincident with the RRM demagnetisation, an additional magnetisation (referred to here as GRMz) is generated during demagnetisation along the Z-axis. This seems to be a GRM since in the X-Y plane the GRM begins to increase around the start of the GRMz increase (Fig. 10a, 10d). Overall, the GRMz generated are mostly negative (up-directed) magnetisations (Figs. 10b,10c,10d), although sometimes are positive like the demagnetisation of the 0.5 Hz [rps] RRM for sample GR1 (Fig. 10a). Co-axial demagnetisation axis and GRM’s have been observed previously in greigite (Hu et al. 1989; 2002) and magnetite (Stephenson 1981; Roperch & Taylor 1986), and these are thought to be more generally produced using static demagnetisation, due to the angular dependence (with respect to grain anisotropy) of the switching field (Madsen 2003; Finn & Coe 2020). Normalising GRMz by  $ARM_{ROT}$  in the same way as for the RRM (equation 1), produces the effective field  $B_{gZ}$  in  $\mu T$  for this remanence (in this case  $B_{gZ} = 100 \mu T \times GRMz / ARM_{ROT}$ ).  $B_{gZ}$  values are similar but rather smaller than  $B_g$  for RRM induction (Fig. 11a). This is not an ARM from the residual DC field inside the demagnetiser shield, since GRMz is both positive and negative and is variable with respect to  $ARM_z$  (Fig. 11b). The DC field inside the shield should produce an ARM of less than  $\sim 0.1\%$  of the  $ARM_z$ , which is much smaller than the observed GRMz (Fig. 11b). It is also possible that this nuisance remanence may be generated by asymmetry in the AF field, generating an ARM, since small  $\sim 0.03\%$  asymmetry could produce an ARM of this order or larger (Hailwood & Molyneux 1974). However, the variable magnitude with the respect to the  $ARM_z$  and  $ARM_{ROT}$  and the change to opposite direction for some samples, suggests this is less likely than a coaxial GRM.

Therefore, for effective determination of RRM stability (median destructive field, MDF etc) of the RRM the most effective way is to demagnetise both the  $+\omega$  and  $-\omega$  produced RRM, up to maximum fields well beyond the AF inducing field used for the RRM. This allows separation of any coaxial magnetisation produce along the demagnetisation axis, such as GRM’s or spurious ARM’s produced along that axis. An added advantage is that RRM’s can be quite weak in many earth materials, and the resulting RRM demagnetisation curve has both duplicate data and is double the magnitude of a single demagnetisation curve, resulting in a more robust dataset to work with.

This improved RRM demagnetisation (i.e., method-2) was tested on a few samples, which showed a different range of MDF of RRM compared to the MDF of  $ARM_z$  (Fig. 12a). This is expected since these magnetisations activate different grain populations, with the RRM probably activating those



with the higher or the same stability, if the coercivities have a limited range to higher values (Stephenson 1976; Edwards 1984; Potter & Stephenson 1986).

There is a large range in MDF of ARMz from relatively soft MD-type magnetite's with over half the remanence removed at <15 mT (cf. Dunlop, 1983), to many with SD-like stability with MDF of 35- 40 mT (Dunlop 1983), to the greigite sample GR1 with the largest MDF in which half the ARMz remains at 64 mT (Fig. 12a). The differing rotation rates produce rather different MDF of RRM for the same sample (Fig. 12a), with no consistent difference- approximately half (6/13) of the test samples having MDF at 0.5 Hz [rps] lower than that at 2.5 Hz [rps], and half (7 out of 13) the opposite behaviour. A subset of the De Geerdalen Fm samples (circled sample in Fig. 12a) have slightly lower MDF of RRM than MDF of ARMz, which may in part relate to the rather larger uncertainty (order of ca.  $\pm 3$  mT) in estimating stability of RRM than that of ARMz (< 1 mT uncertainty). Hence, the magnetite-bearing samples (except the MD-like SVD2) fall into two sets, firstly a lower stability set in which the RRM and ARMz stability are similar (those with MDF of RRM < ca. 42mT; Fig. 12a) with mostly Bg from 0 to -5  $\mu$ T (circled samples in Fig. 12); and secondly a set with MDF of RRM larger than the MDF of ARMz and a wide range of Bg (Fig. 12b). The two magnetite bearing samples containing common magnetofossils –CC1b and MR5 (a <2  $\mu$ m extract), have the largest MDF of RRM for the magnetite-bearing sets (Fig. 12a), although very different Bg values (Fig. 12b). The differences between CC1b-MR5 perhaps relates to the abundant magnetofossil chains in CC1b.

The greigite sample GR1 has an MDF of RRM of 65-69 mT which is at the lowest end of the range measured for greigite by Snowball (1997b). Nevertheless, the large negative Bg in combination with the high MDF of RRM is a very distinctive feature of greigite (Fig. 12b), confirming the observations of Snowball (1997b). The large-sized pyrrhotites, PY5, PY7 (with MDF of ARMz < 15 mT; Fig.12a) have distinctive positive Bg, and large MDF of RRM comparable to those samples with the smallest magnetite particles (e.g., MR5, CC1b; Fig. 12b).

## 6. Discussion

### 6.1 Comparison with other datasets using low rotation rate

Comparison to other datasets that used low  $\omega$  (less than the AF frequency) is more problematic, since AF hold and ramp-down times have been less clearly described. Potter & Stephenson (1986) used hold and decay times of 5 s and 10 s, and for various sized magnetite's, finding peaks in negative RRM at around 8-10 Hz [rps] ( $\omega T_D \approx 80$ -100 revolutions), considerably larger  $\omega T_D$  than seen here and in Edwards data. Peaks in negative RRM for greigite have not been recognised previously, with existing data from Stephenson & Snowball (2001) showing increasingly more negative RRM from low rotation rates until the AF-frequency threshold.

In Edwards (1982b) magnetite-based data, the negative RRM peak is followed by a trend towards more positive RRM values (both negative and positive RRM) which plateau at around  $\omega T_D \approx 20$ -30 revolutions. These approximately concur with largely negative, but some positive RRM values (in basalts) measured by Roperch & Taylor (1986) for  $\omega T_D$  of  $\approx 30$  rev (Table 3). Contrastingly, Potter & Stephenson (1986) recognised a positive peak in RRM at  $\approx 20$  Hz [rps] ( $\omega T_D \approx 200$  rev) for crushed magnetite. Hence, it is probable that the RRM peaks and troughs seen in our dataset and those of Edwards (1982b) and Wilson & Lomax (1972), between  $\omega T_D = 0.4$  - 30 revolutions, may be hidden in the smallest rotation rates produced by the instrument of Stephenson & Molyneux (1987) at  $\omega < 3$  Hz [rps].



## 6.2 Mineral magnetic identification using RRM

Whilst the comparison of different  $B_g$  values at fixed Hz [rps] (or  $\omega T_D$ ) could be useful for simple mineral magnetic discrimination, the changes in  $B_g$  and  $ARM_{ROT}$  with  $\omega$  potentially indicate a more powerful means of mineral magnetic discrimination.

There are two broad approaches that could be used for magnetic mineral discrimination using low rotation rates. The first is to use the normalised changes in RRM,  $ARM_{ROT}$  and  $B_g$  with respect to  $\omega T_D$ , to build up families of curves for mineral types, granulometry and switching-field behaviours. The data in Figures 3 to 9 are a step in this direction. The second is to parameterise some of the key changes at  $\omega T_D \geq 0.4$  rev by looking at ratios of RRM,  $ARM_{ROT}$  and  $B_g$  at say 0.5 and 2.5 Hz [rps] ( $\omega T_D$  units of 0.77, 3.83 revolutions), as expression of the changes with  $\omega$ . We here explore this later approach, since it potentially expresses the between-sample variability better, allowing exploration of larger datasets, and so is more easily used in palaeomagnetic studies focussed on other aims.

Our somewhat limited data on RRM stability (i.e., MDF) also suggests that this may differ between rotation rates (compare test sample 0.5 Hz [rps] and 2.5 Hz [rps] values in Fig. 12). Although MDF of RRM is time consuming to measure, this may perhaps hold some additional information for magnetic mineral discrimination in magnetically stronger samples (e.g., Fig. 12). Although not explored here, the stability of RRM is more usefully utilised in RRM build-up with increasing AF field (see SI Figs. S15-S17, and Potter and Stephenson, 1986, fig.7)- rather than demagnetisation. Median acquisition is free from issues with co-axial GRM's. Alternatively, single step acquisition (say at 40 mT), as a proxy for stability, is also easy to implement and easy to use on samples displaying weak RRM that may need duplicate RRM measurements to improve accuracy.

### 6.2.1 Magnetic sulphides

Pyrrhotite is recognised by positive  $B_g$  at 0.5 or 2.5 Hz [rps] (Figs. 13a, 7c, 8c). The  $B_g$  ratios at the rotation rates of 2.5 Hz [rps] divided by 0.3 Hz [rps] (ratios at differing  $\omega$  hereafter symbolised like  $\omega\{2.5/0.3\}$ ) can also usefully distinguish large pyrrhotite from pyrrhotite  $< 20 \mu\text{m}$  in size (Figs. 13a); an expression of the flattish shapes of the RRM- $\omega$  curves (Figs. 7c, 8c). An alternative parametrisation is to use the  $ARM_{ROT}$  at 2.5 Hz [rps] normalised by  $ARM_z$ , which clearly separates the  $>1.3$  values for pyrrhotite from magnetite and greigite behaviours (Fig. 14a, b). Pyrrhotite is particularly challenging to identify using conventional coercivity behaviour, since it strongly overlaps the coercivity range of magnetite, and the useful normalisation by magnetic susceptibility (Peters & Thompson, 1998) is hampered in many sediments by paramagnetic contributions to susceptibility.

Single domain diagenetic greigite has the most distinctive signature in RRM behaviour with large negative  $B_g$ , high MDF of RRM and  $ARM_z$  (Figs. 12a, 14b, d). These are features also indicated by Snowball (1997b) and Peters & Thompson (1998) for  $ARM_z$ . The unusually large  $B_g$  and RRM stability probably stem from the magnetic nano-scale composites that make up natural diagenetic greigite crystals (Lesniak et al. 2021). The change of  $ARM_{ROT}$  with  $\omega$  may also be a useful feature, like also seen in the  $< 20 \mu\text{m}$  pyrrhotite, both of which increase through  $\omega = 0.5$  to 3 Hz [rps] (Fig. 6b, 7b). The ratio of  $ARM_{ROT}$  at  $\omega\{2.5/0.5\}$  may also be a useful and simple discriminator (Figs. 13b,c,d, 14d).

In contrast, biogenic greigite present in magnetosomes measured by Chen (2014), appears to have low  $B_g$  values (at 95 Hz [rps]; Table 2) not much larger than for magnetite magnetosomes. However, Reinholdsson et al. (2013) implicated positive  $B_g$  values of 0- 1  $\mu\text{T}$  for greigite magnetofossils at 5 Hz [rps]. A means of using RRM to distinguish these two largely non-interacting SD grain types require more work. Perhaps the change of  $ARM_{ROT}$  with  $\omega$  may be a useful signature, since magnetofossils from the Chalk show particularly dramatic declines in  $ARM_{ROT}$  with increasing  $\omega$  (Fig. 4b), something

that is not seen in our greigite or pyrrhotite samples. Ratios of  $ARM_{ROT}$  at  $\omega\{2.5/0.5\}$  clearly discriminate the Chalk magnetite magnetofossils from other magnetite samples (Figs. 13c, d; 14d). It remains to be seen if similar behaviour is also shown by greigite magnetofossils.

### 6.2.2 Magnetite bearing samples

It is interesting that the SD-sized magnetite in the Chalk sample set has consistently the smallest negative  $B_g$  values, opposite to the behaviour indicated by the crushed magnetite sample set used by Potter & Stephenson (1986). The small  $B_g$  values for the magnetofossils in the Chalk-samples are consistent with the work on magnetite magnetosomes by Chen (2014) at high spin rates. An explanation for the low RRM acquisition in intact magnetofossils/magnetosomes may be related to the model of magnetisation in elongate particles (i.e., equivalent to magnetosome strings) proposed by Potter & Stephenson (2006).

It is challenging to compare the  $B_g$  values at high  $\omega$  with those at low  $\omega$ . Two ways to approximately cross calibrate these, are to use the  $B_g$  for synthetic Mapico magnetite and  $\gamma Fe_2O_3$  which have been measured at both low and high spin rates (Tables 2, 3). Using either the Mapico magnetite or mean of Mapico+  $\gamma Fe_2O_3$  gives two possible conversion factors which allows the  $B_g$  to grain size data of Potter & Stephenson (1986) to be mapped into the low spin rate data here (scales in top Fig. 13a). It is feasible, that excluding the Chalk samples, the remaining magnetite-bearing samples show differing  $B_g$  values corresponding to the particles size changes suggested by Potter & Stephenson (1986), since samples MR5, MR4 have the larger contribution from the finest particle sizes (i.e. SD magnetofossils largely not in chains) in the magnetic extracts (Fig. 13a). Perhaps the particle size,  $d$  (in  $\mu m$ ) dependency of  $d \approx 100/B_g$  (Potter & Stephenson 1986) has compressed much of the apparent grain-size variation in our sample set into a range in  $B_g$  of ca. 0-10  $\mu T$  at 0.5 Hz [rps] (Fig. 13a). An alternative possibility is that the larger  $B_g$  values for the crushed samples of Potter & Stephenson (1986) are related to stress, which can impact coercivity, something clearly expressed in hysteresis data, which strongly modifies the coercivity relationship to grain size (Tauxe et al. 2002). Like hysteresis datasets, we speculate that the presence of dislocations allows additional pinning of domain wall motions and so may enhance the irreversible flip-mechanism which is responsible for RRM acquisition. It seems reasonable that during crushing, dislocation density may increase in smaller particles, so increasing  $B_g$ . Alternatively, the likely particle interactions present in both our magnetic extract samples, and those of Potter & Stephenson (1986) may modify this behaviour- the RRM response to magnetic interactions needs further work to evaluate this.

Excluding the Chalk samples, magnetite-bearing samples show broadly similar declines in  $B_g$  at  $\omega$  from 0.3 to 2.5 Hz [rps], which can be expressed as the  $B_g$  ratio at  $\omega\{2.5/0.3\}$  – with samples clustered in the numerical range ca. 0 to 0.7 for this  $B_g$  ratio (Fig. 13a, b). The ratio of the  $ARM_{ROT}$  for  $\omega\{2.5/0.5\}$  shows a better discrimination of these samples (Figs. 13c, d; 14d), which expresses the decline in  $ARM_{ROT}$  with  $\omega$ . The two dolerite samples (SVD1 and SVD2) show very different RRM behaviour, which is rooted in the differing stability of the natural remanence of the Spitsbergen dolerites which varies much, with some showing MDF of NRM < 10 mT, with others up to 40 mT (Halvorsen, 1973). This is reflected in the  $ARM_z$  stability of SVD1 and SVD2 which are 42mT and 16 mT respectively, with SVD2 showing behaviour closer to MD-like titanomagnetite with small  $B_g$  values and low  $ARM_z$  stability and low MDF of RRM at 0.5 Hz [rps] (Figs. 12a; 14c,d). This variability is rooted in the variable oxidation of titanomagnetite in igneous rocks, which impacts  $ARM_z$  stability (Dunlop 1983; Figs. 14b, c). The larger  $ARM_z$  stability in SVD1 is probably from nanoscale subdivision of the titanomagnetite due to exsolution (cf. Harrison et al. 2002) and titanomagnetite oxidation. This is probably responsible for the large  $B_g$  value (-26  $\mu T$ ) for SVD1 at 0.5 Hz [rps] (Fig. 13a). Sample SVD2 shows similar changes in RRM with  $\omega$ , to basalt sample F37B-1a studied by Edwards (1982b)

derived from R. L. Wilson's samples. All these share a transition into positive RRM values at  $\omega T_D > 3$  revolutions (Figs. 4c). Edward's sample F37B-1a had a Bc of 18 mT and Mrs/Ms of 0.13 (Edwards 1982a), which places F37B-1a close to the MD-magnetite field using the squareness versus Bc plot of Tauxe et al. (2002). Perhaps this positive RRM at larger  $\omega T_D$  represents the RRM response at the border towards truly MD magnetite, but with a small contribution from SD-like material. Truly MD-behaviour (with zero Bg) is shown by Edwards basalt sample D264-26 (Table 3). Hence, increases in  $ARM_{ROT}$  for  $\omega\{2.5/0.5\}$  and  $ARM_{ROT}/ARM_z$  seem to express change towards a more MD state for magnetite-samples, from non-interacting SD-state in the Chalk samples (arrows in Figs. 14b, d). Overall, comparing our dataset to those of basic igneous rocks, not surprisingly shows that the extracts and sediment sample sets fall into regions corresponding to the most oxidized and most stable SD-like behaviour seen in basic igneous rocks (top scale in Fig. 14b, arrows in Fig. 14c). Our data do not therefore provide much in the way of characterisation of the RRM behaviour of low to mid oxidation status basic igneous rocks as characterised by Dunlop (1983)— other than sample SVD2.

Those magnetic extracts corresponding to residual Fe-oxide assemblages, broadly fall into similar intervals as other extracts (Fig. 14a-c). This broadly corresponds with observations of silicate hosted inclusions which have similar chemistry and microstructural characteristics to discrete Fe-oxides (Feinburg et al. 2005; 2006). This is also displayed in the Fe-oxide inclusion-only ARM data of Hounslow & Morton (2004) from UK basement complexes which show a similar range in  $d.ARM_{40mT}$  to the samples here (arrows at top in Fig. 14d). The De Geerdalen Formation samples are remarkably well clustered considering the range of locations and lithology types that occur in these samples (Figs. 13, 14). In addition their properties do not overlap those of the magnetic extracts (Figs. 14a, c, d), but do overlap with the range of  $d.ARM_{40mT}$  seen in silicate hosted inclusions. Hence, presumably these contain more SD-like magnetite which is intermediate towards the Chalk samples, with a larger MDF of  $ARM_z$  (Fig. 14c). Although magnetic extracts have not been performed on these samples, it seems highly unlikely they contain magnetofossils, due to the rather extensive silicate and carbonate diagenesis (Mørk, 2013), but magnetite inclusions in silicates likely make significant contributions.

## 7. Conclusions

Using rotation rates of 0.26 to 3 Hz [rps] there is the possibility for routine magnetic mineral characterisation using RRM and rotational anhysteretic magnetisation ( $ARM_{ROT}$ ). Characterisation is best done using either: 1) normalised  $\omega$ —RRM curve characteristics (or  $\omega$ —Bg,  $\omega$ —  $ARM_{ROT}$  curves), This approach allows more effective characterisation of the mineralogical behaviour, rather than relying on single spin rate-derived values of Bg, RRM or  $ARM_{ROT}$ ; or 2) parametrisation of the variations using values at selected rotation rates (or better at fixed  $\omega T_D$  values). The later approach is likely more widely accessible if rotation rates cannot be continuously varied in equipment available. Optimisation of the AF field hold time to achieve full rotational initialisation of RRM will improve consistency, especially so for investigating greigite (and other strongly anisotropic samples), which shows large changes in RRM and  $ARM_{ROT}$  at rotation angles below the initialisation threshold caused by anisotropy.

Static AF demagnetisation of the RRM, for stability tests, is best achieved by demagnetisation of both  $+\omega$  and  $-\omega$  RRM's which allows removal of gyroremanence acquired along the axis of demagnetisation. A simpler approach free from this problem is to characterise stability using RRM acquisition with increasing AF field- this is probably best standardized at  $\omega T_D = 0.4$  revolutions, since

RRM stability varies with rotation rate, and RRM is maximised at  $\omega T_D = 0.4$ . This allows RRM stability to be parameterized even in samples acquiring a weak RRM.

The sized pyrrhotite and pyrrhotite bearing phyllite samples show largely positive RRM, in contrast to negative RRM largely acquired by test samples of magnetite and greigite. This and the smaller variation of RRM and  $ARM_{ROT}$  with rotation rate are a simple effective tool for detecting pyrrhotite. The median destructive field of RRM produced by pyrrhotite (including large grains) is comparable to greigite and larger than most test samples of magnetite. Samples of diagenetic greigite display large negative  $B_g$  values and a stronger variation of RRM and  $ARM_{ROT}$  with spin rate. The large  $B_g$  values carried by diagenetic greigite are a simple and easy way to identify it, but additionally diagnostic are the ratio of  $ARM_{ROT}$  at  $\omega\{2.5/0.5\}$  and MDF of a conventional static ARM. RRM distinction of greigite magnetofossils from magnetite magnetofossils needs more evaluation, but utilising  $\omega$ - $ARM_{ROT}$  changes or  $ARM_{ROT}$  stability may be diagnostic.

The set of natural magnetite-bearing test samples shows that  $B_g$  is not simply related to grain size as thought previously, but non-interacting SD magnetite in magnetofossils have very small  $B_g$  and little variation with  $\omega$ . A small  $B_g$  value concurs with studies by Chen (2014) measured at high spin rates on magnetite magnetosomes. The ratio of  $ARM_{ROT}$  at  $\omega\{2.5/0.5\}$  may be a better tool for particle size characterisation of magnetite. For magnetite, the  $B_g$  parameter may also be strongly impacted by presence of dislocations, or perhaps interaction-related controls— features that need further study. The prior extensive study of RRM on sized and crushed magnetite by Potter & Stephenson (1986), may have been impacted by particle size related dislocation density in the magnetite samples, which may have enhanced the  $B_g$  variation with grain size. The great advantage of using rotational remanent magnetisation (RRM) characteristics for magnetic mineral identification is the RRM properties are carried by the hardest coercivity grains most relevant to understanding stable palaeomagnetic signals.

## Acknowledgements

Paul Montgomery helped collect the Chalk samples, and Gareth Lord some of the Svalbard samples. Isaac Hilburn helped unravelling some complexities of the RAPID software.

**Author contribution statement:** MWH, VK, C-S.H processed the data. MWH analysed the magnetic results. All contributed to writing of the paper.

**Sample Data Availability Statement:** All data are incorporated into the article and its online supplementary material.

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## Figure captions

Fig. 1. Schematic view of the rotational remanent magnetisation (RRM) and rotational anhysteretic magnetisation ( $ARM_{ROT}$ ) measurement process. The rotation rate ( $\omega$ ) is symbolised also with a + or - sign indicating the down, or up-directed rotation vector. Inset shows the timing relationships between sample rotation and stages in the AF field build-up, hold and decay stages and their symbolised timings.

Fig. 2. Example micrographs of some of the test samples. A) Optical reflected light micrograph of sample PY2 (5-10  $\mu\text{m}$  detrital monoclinic pyrrhotite magnetically separated from river sediment). B) Backscatter scanning electron microscope (BSE) image of pyrrhotite bearing phyllite. The brightest phase is pyrite, and the more abundant slightly greyer phase is pyrrhotite, both embedded in a fine-grained phyllosilicate-quartz matrix. C) BSE image of ca.  $<0.5 \mu\text{m}$  in size, brighter greigite embedded in a clastic-phyllosilicate rich matrix. D) Transmission electron microscope image of chains of magnetofossils in a magnetic extract from Chalk sample at level of CC2. See the supplementary information for more details on the test samples.

Fig. 3. Magnetic extract samples LUD, OR2, OR3, MR1, MR2 and MR5. Variation of: a) normalised RRM, b) normalised rotational ARM ( $ARM_{ROT}$ ) and c)  $B_g$  with revolution rate ( $\omega$  in Hz [rps]). In a) and b) the absolute value for that at 0.5 Hz [rps] is used for normalisation. In a) sample OR3, MR2 share the right-hand scale. The legend in B) applies to all sub-panels. In b) is the ARM generated in the

same way as the RRM, but with zero rotation ( $ARM_{TRAN}$ ). Additional scale in c) is  $\omega T_D$  (in revolutions), where  $T_D$  is the decay time of the AF field in seconds.

Fig. 4. Chalk (CC1, CC2) and dolerite (SVD1, SVD2) test samples. In c) only sample SVD1 has a separate scale for  $B_g$ . See Fig. 3 caption for details.

Fig. 5. De Geerdalen Formation test samples DF1 to DF6. See Fig. 3 caption for details.

Fig. 6. Greigite-bearing siltstone test samples GR1 to GR4. See Fig. 3 caption for details.

Fig. 7. Sized detrital pyrrhotite test samples PY1 to PY5. In a) PY4 and PY5 share the right-hand scale. See Fig. 3 caption for details.

Fig. 8. Pyrrhotite bearing phyllite test samples PY6 to PY9. See Fig. 3 caption for details.

Fig. 9. RRM and  $ARM_{ROT}$  of test samples with three AF field hold times ( $T_H$ ) of 0.092 s, 0.917 s and 9.17 s. A, B) are for magnetite sample MR1, C, D) for pyrrhotite sample PY5, and E, F) for diagenetic greigite sample GR1. The dotted vertical red lines in A,C,E) are the rotation rate thresholds at which the two smaller  $T_H$  values have full  $90^\circ$  of rotation at the maximum field of 100 mT. The longest hold time of 9.17s achieves  $90^\circ$  rotation at maximum field even at 0.05 Hz [rps]. Error bars are the  $1\sigma$  uncertainty (smaller than symbols in C and E) from the remanence measurement. Most of the uncertainty in B, D, F relate to the uncertainty in the zero level of the  $ARM_{ROT}$ . Points are connected by an Excel fitted 'smoothed-curve'.

Fig. 10. Example demagnetisation of RRM obtained for both  $+\omega$  and  $-\omega$  conditions (both 0.5 Hz [rps] and 2.5 Hz [rps]). Static demagnetisation is along the Z-axis parallel to the rotation axis (Fig. 1). Left panels in each case show the RRM and GRMz curves in magnetic moment ( $Am^2$ ) corrected for the subtraction and addition operations on the raw moments. Right panels show the X-Y axis GRM moment ( $\sqrt{X^2+Y^2}$ ) for each of the four spin measurements (two for each  $\omega$ ) during demagnetisation. In each case the baselines have not been adjusted to zero. The measurement order is AFz demagnetisation at 150 mT, RRM ( $-\omega=0.5$  Hz [rps]), progressive AF demagnetisation, RRM ( $+\omega=0.5$  Hz [rps]), progressive AF demagnetisation; to be followed by the same four sets at 2.5 Hz [rps]. A) Arrows show the sign of the GRMz for increasing negative, and positive values respectively.

Fig. 11. A)  $B_g$  versus the effective field ( $B_gZ$ ) of the additional nuisance magnetisation (GRMz) generated during static demagnetisation of the RRM (i.e., Fig. 10). Datasets for both  $\omega=0.5$  Hz [rps] and 2.5 Hz [rps] are shown. Note both negative (upwards directed) and positive (downward directed) GRMz. Arrows indicate the off-graph positions of the greigite sample and its respective  $B_g$  values at  $\omega=0.5$  and 2.5 Hz [rps]. B) Ratio of GRMz and ARMz expressed as a percentage with respect to the ARMz moment. Note the often much larger %GRMz/ARMz values for greigite and pyrrhotite. In each case the error bars include the RRM measurement uncertainties plus the baseline uncertainty for GRMz, both as  $\pm 1\sigma$ .

Fig. 12. Median destructive field (MDF) of the RRM versus MDF of ARMz in a) and  $B_g$  in b). This uses RRM demagnetisations using method -2 for both 0.5 Hz [rps] and 2.5 Hz [rps]. In a) note the soft ARMz (<20 mT) of the 'large'-pyrrhotite and dolerite sample SVD2, the hard ARMz (~64 mT) of greigite sample GR1, and the rather consistent intermediate ARMz stability (35-41 mT) of other samples. Both the 0.5 and 2.5 Hz [rps] measurements of greigite sample GR1 plots off the x-axis in b) at a  $B_g$  of -53  $\mu T$  and -96  $\mu T$  (Fig. 6c).

Fig. 13. Biplots of  $B_g$  and  $ARM_{ROT}$  ratios illustrating possible visualisation of mineral magnetic discrimination. In a) at top are shown two possible scales of  $B_g$  conversion from the high spin rate datasets (sized and crushed magnetite) of Potter and Stephenson (1986)– scale in microns. The lower micron scale is using the Mapico magnetite (Tables 2,3) with  $B_g$  value of -8 mT (at  $\omega_{TD} = 1.05$  revolutions), which gives a high to low spin  $B_g$  conversion factor of -0.022. The upper micron scale is using  $\gamma Fe_2O_3$  (GF01, TDK type D; Tables 2,3), which gives a conversion factor of -0.089. The magnetite-bearing samples marked as ‘residual’ contain a residual magnetic assemblage after extensive diagenetic dissolution (sample LUD only in this figure). The diagenetic greigite samples plot-off to the left of the graph in a), c) and d).

Fig. 14. Biplots of RRM,  $ARM_z$  and demagnetisation stability, illustrating possible visualisation of mineral magnetic discrimination. In b) and c)  $d.ARM_{z40mT}$  is the proportion of the initial  $ARM_z$  remaining after 40 mT static-axis demagnetisation. Regions on the plots for pyrrhotite are loosely defined with a ‘bag’, and the  $B_g=0$  line indicates truly MD magnetite in a) and c). Those marked as ‘residual’ contain a residual magnetic assemblage after extensive diagenetic dissolution (samples LUA, LUD, LUD2, LUD3, OR4, OR5, OR6, MR6; Table 4). Various additional unlabelled samples are shown from the Chalk and De Geerdalen Fm sample sets (See SI datasets). B) top scales show the  $d.ARM_{z40mT}$  values for the petrological oxidation states I to VI (Watkins & Haggerty 1967) of basalts from the Steen’s Mountain and Icelandic basalts (from Dunlop, 1983). C) Arrows on right show the mean (the tick) and standard deviation ( $1\sigma$ ) range of the MDF of  $ARM_z$  for basic igneous rock classed as single domain (SD) type by Dunlop (1983) with the dotted arrow showing the maximum MDF in this class. Dunlop’s (1983) MD-type mean and  $1\sigma$  range falls below 10 mT (so off the graph). D) Top arrow is the mean (tick) and range of  $d.ARM_{z40mT}$  from Fe-oxides inclusions in silicates (Hounslow & Morton, 2004) for nine UK basement complexes (all discrete Fe-oxides outside silicates removed). The greigite samples plot off the graphs to the left in a) and c).

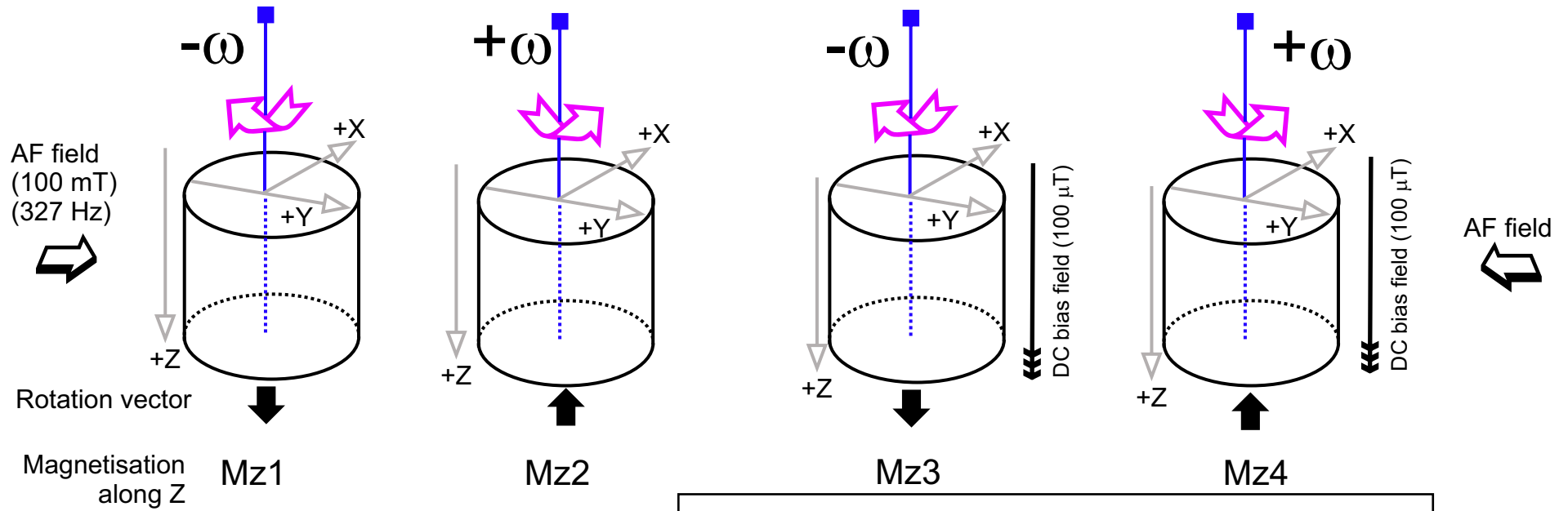
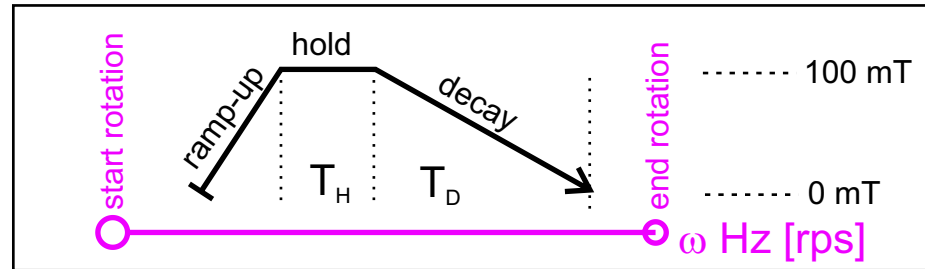


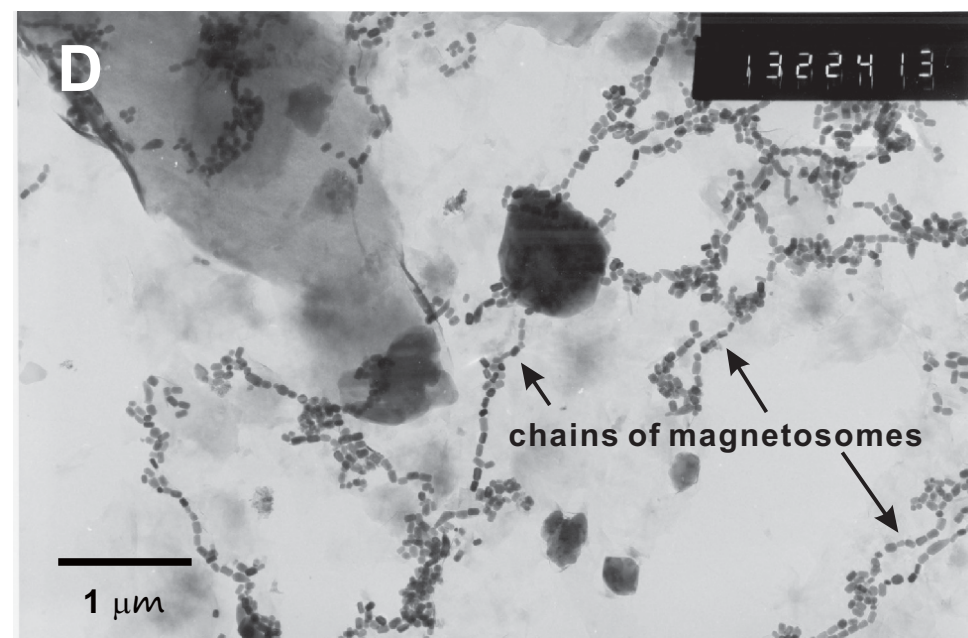
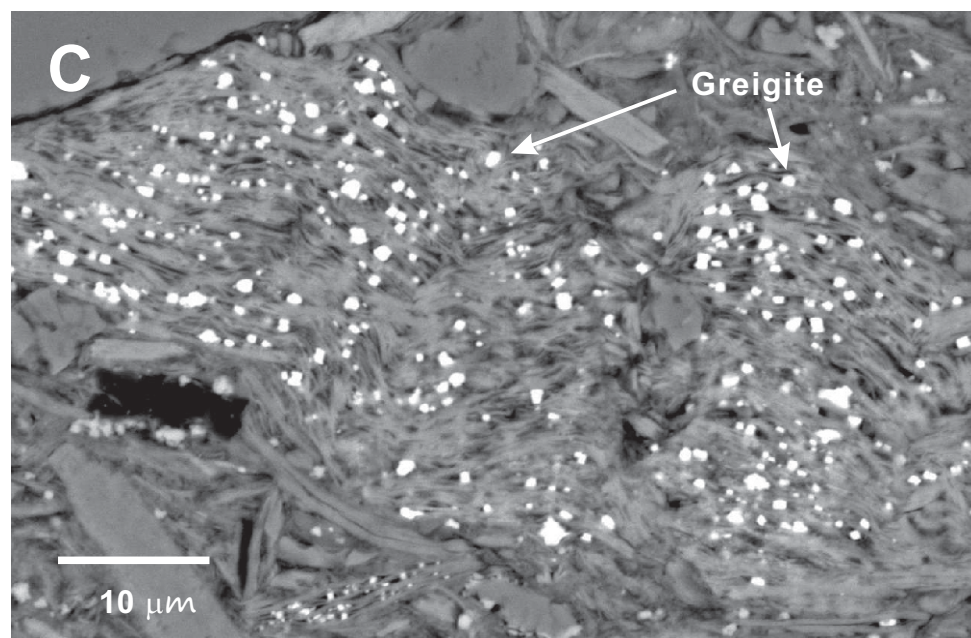
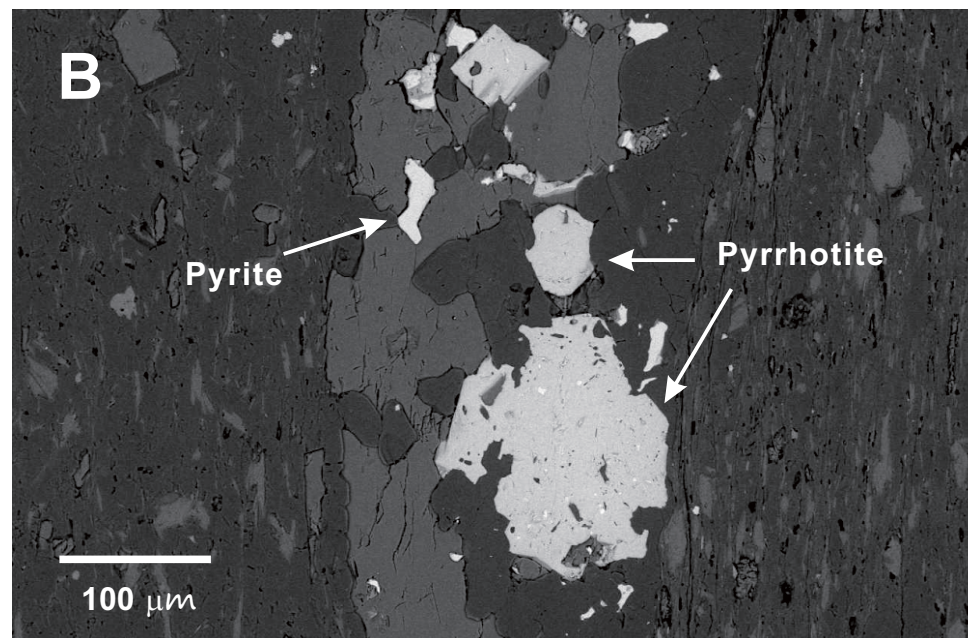
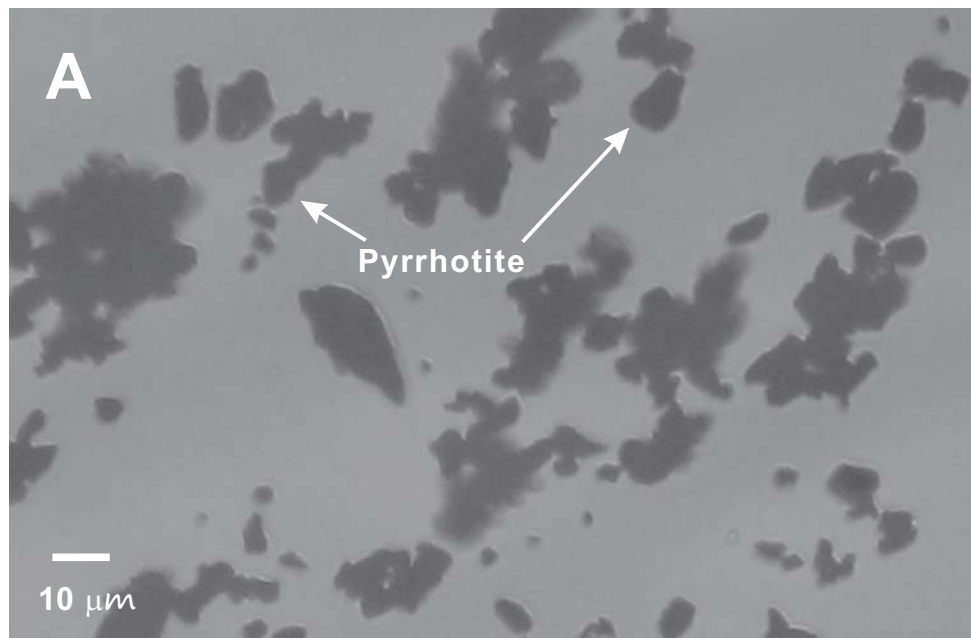
fig. 1

$$RRM = (Mz1 - Mz2) / 2$$

$$ARM_{ROT} = (Mz3 + Mz4) / 2$$

$$RRM_{100} = (Mz3 - Mz4) / 2$$







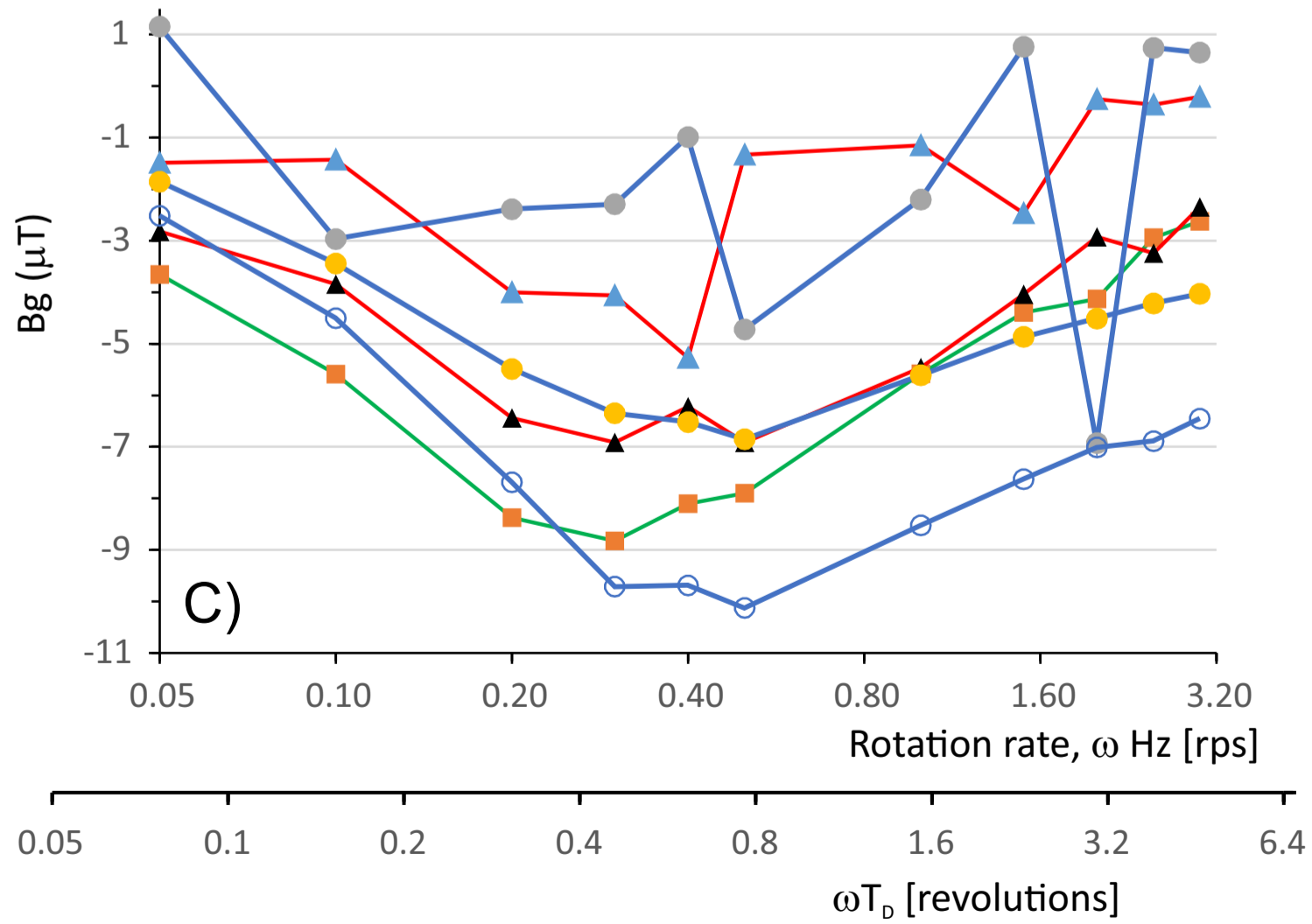
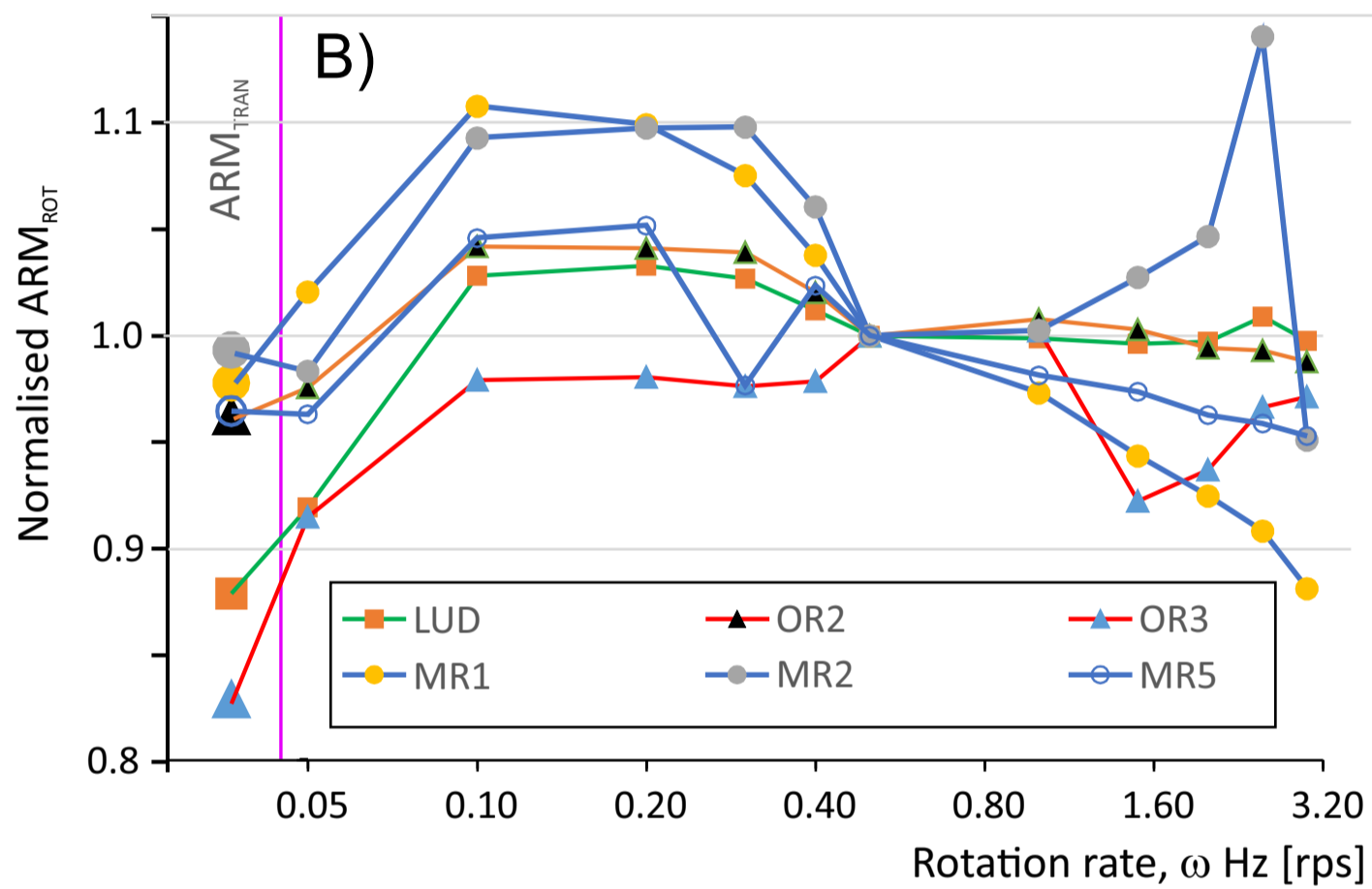
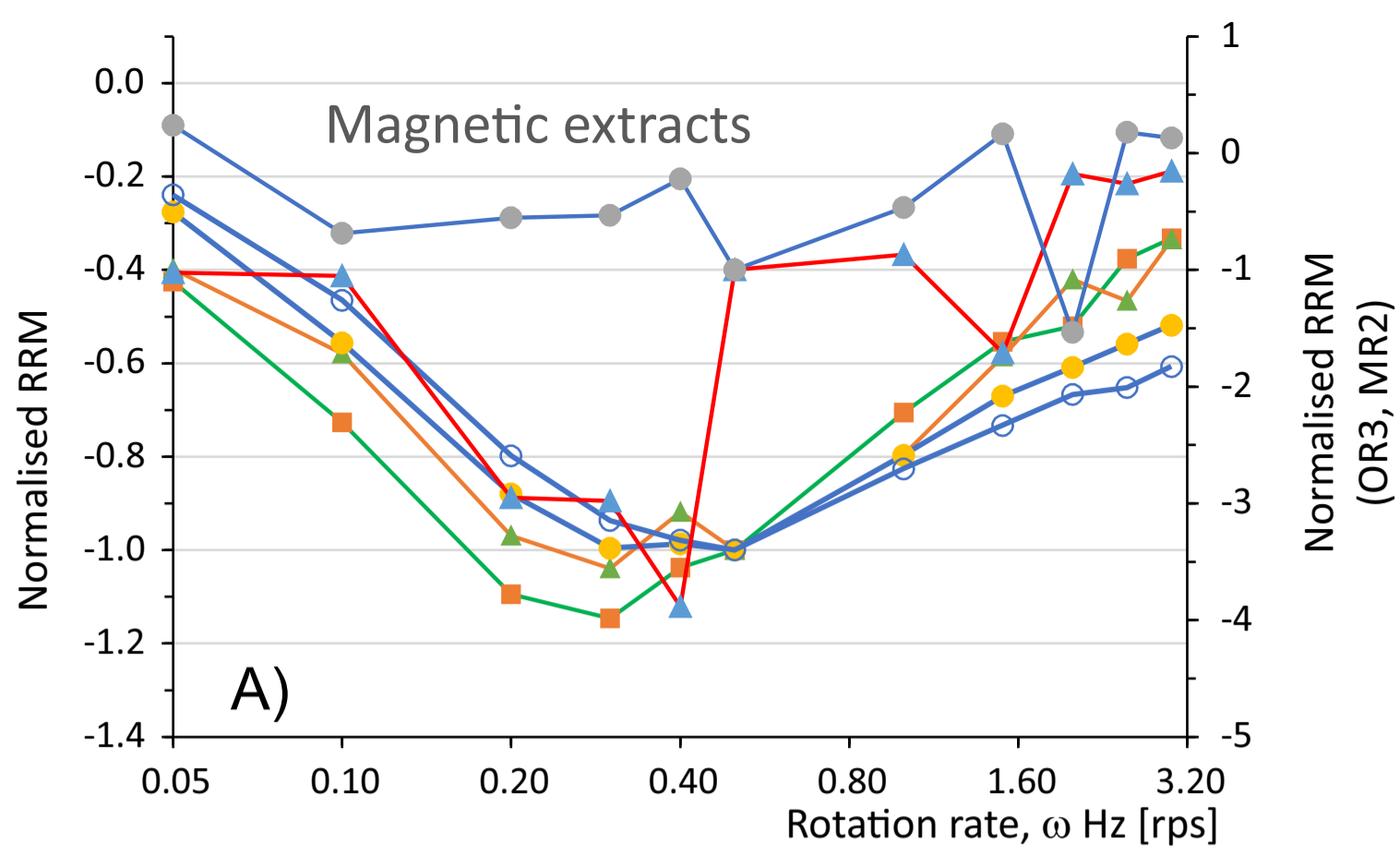


fig. 3

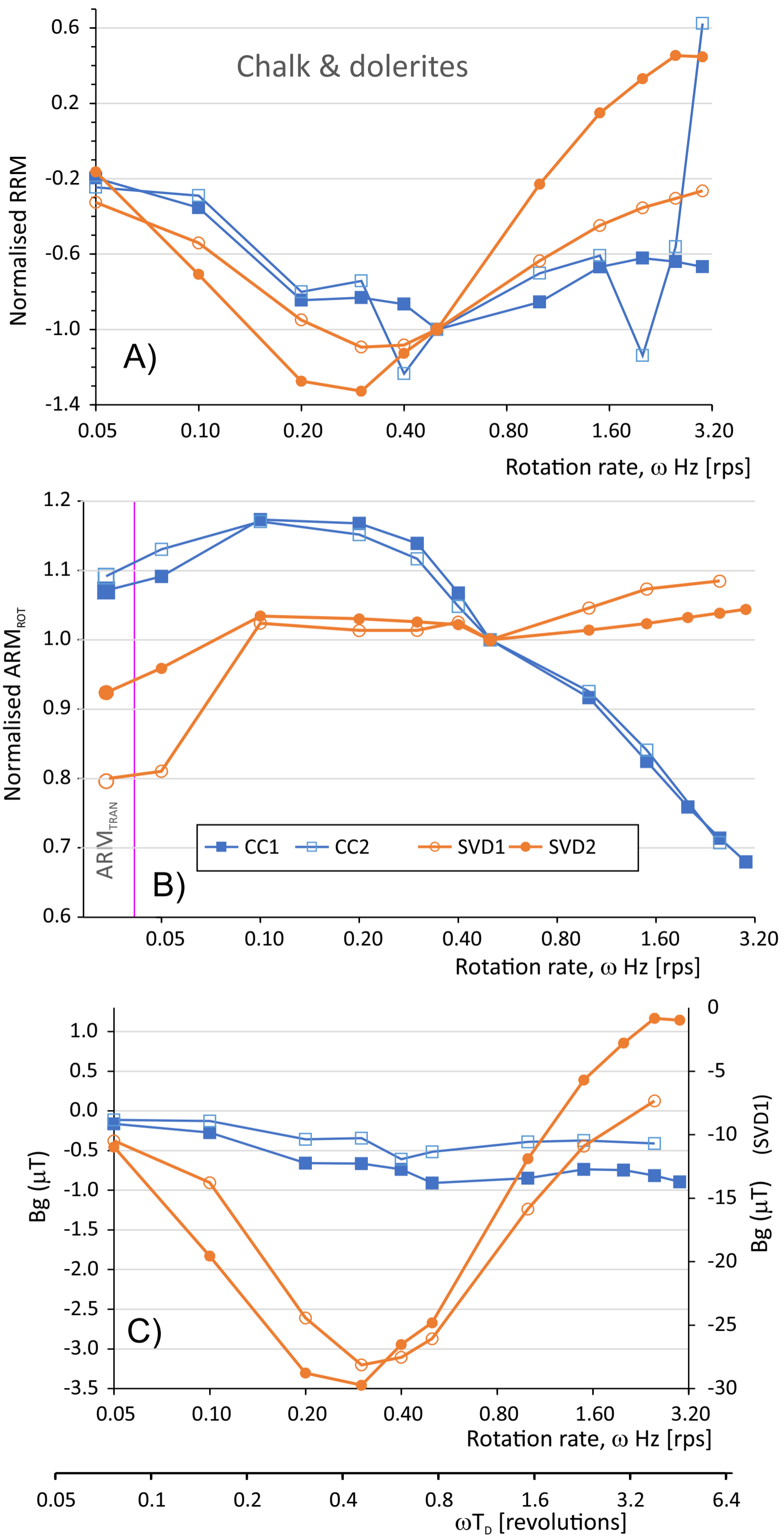


fig. 4

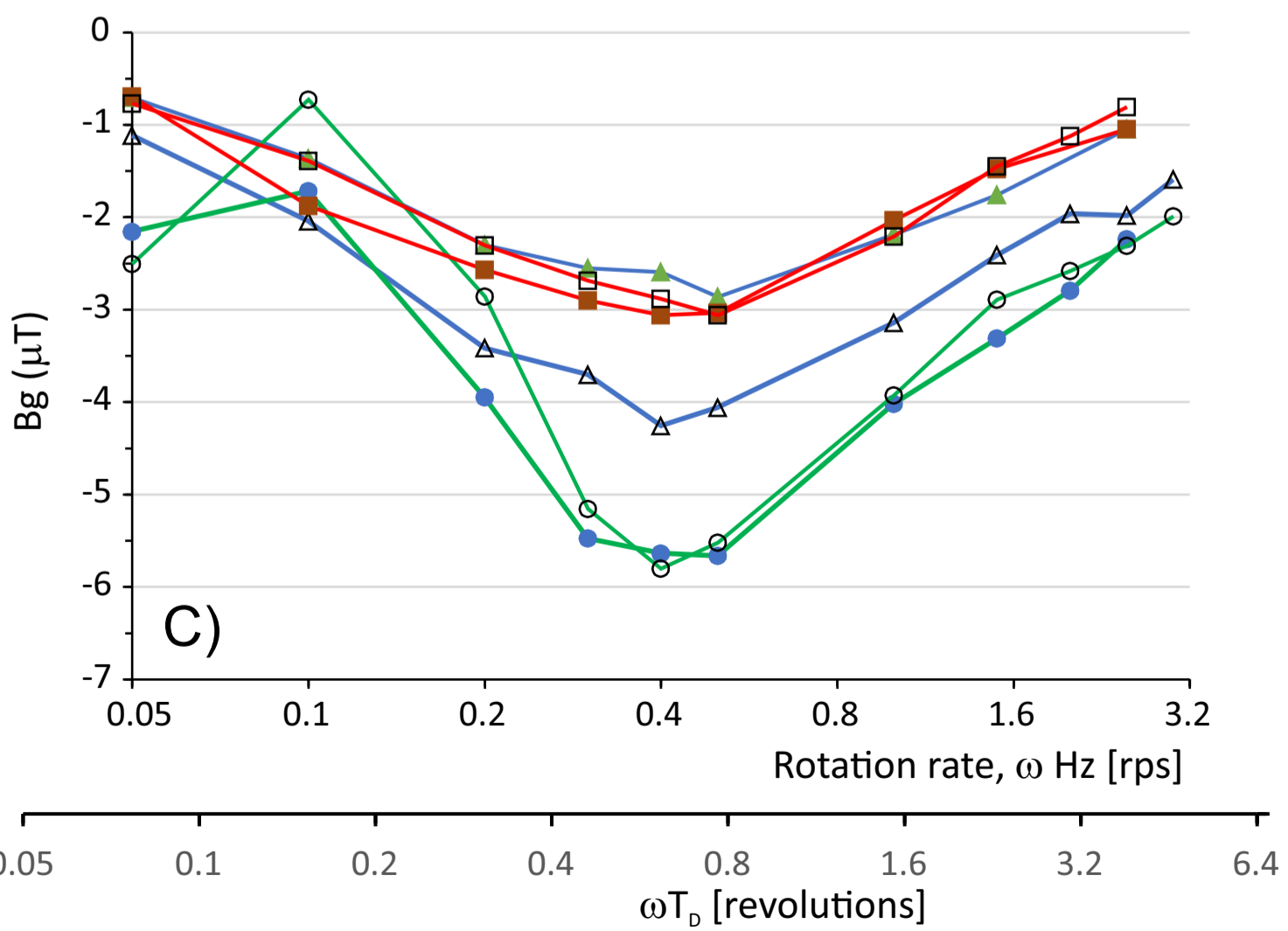
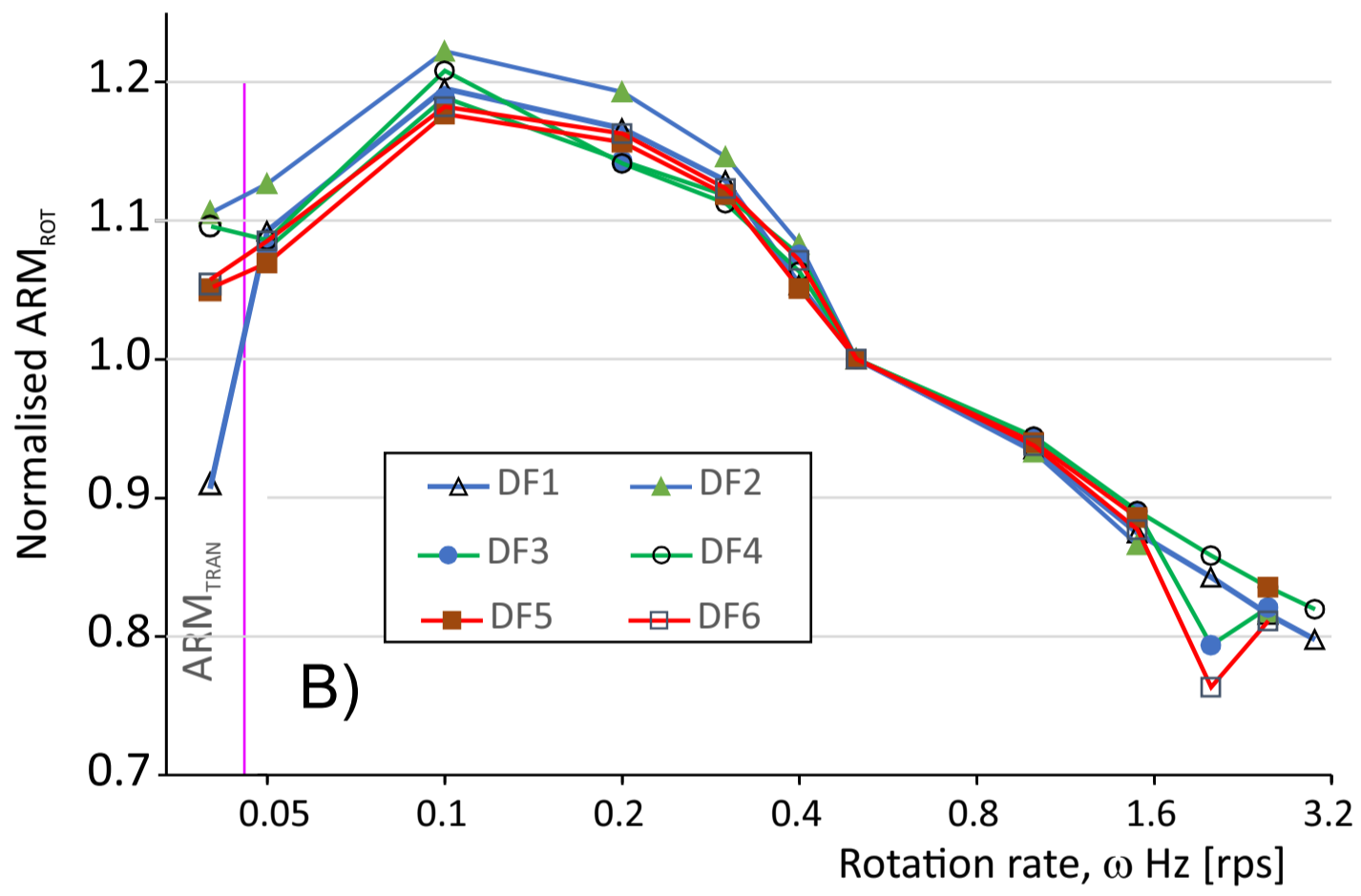
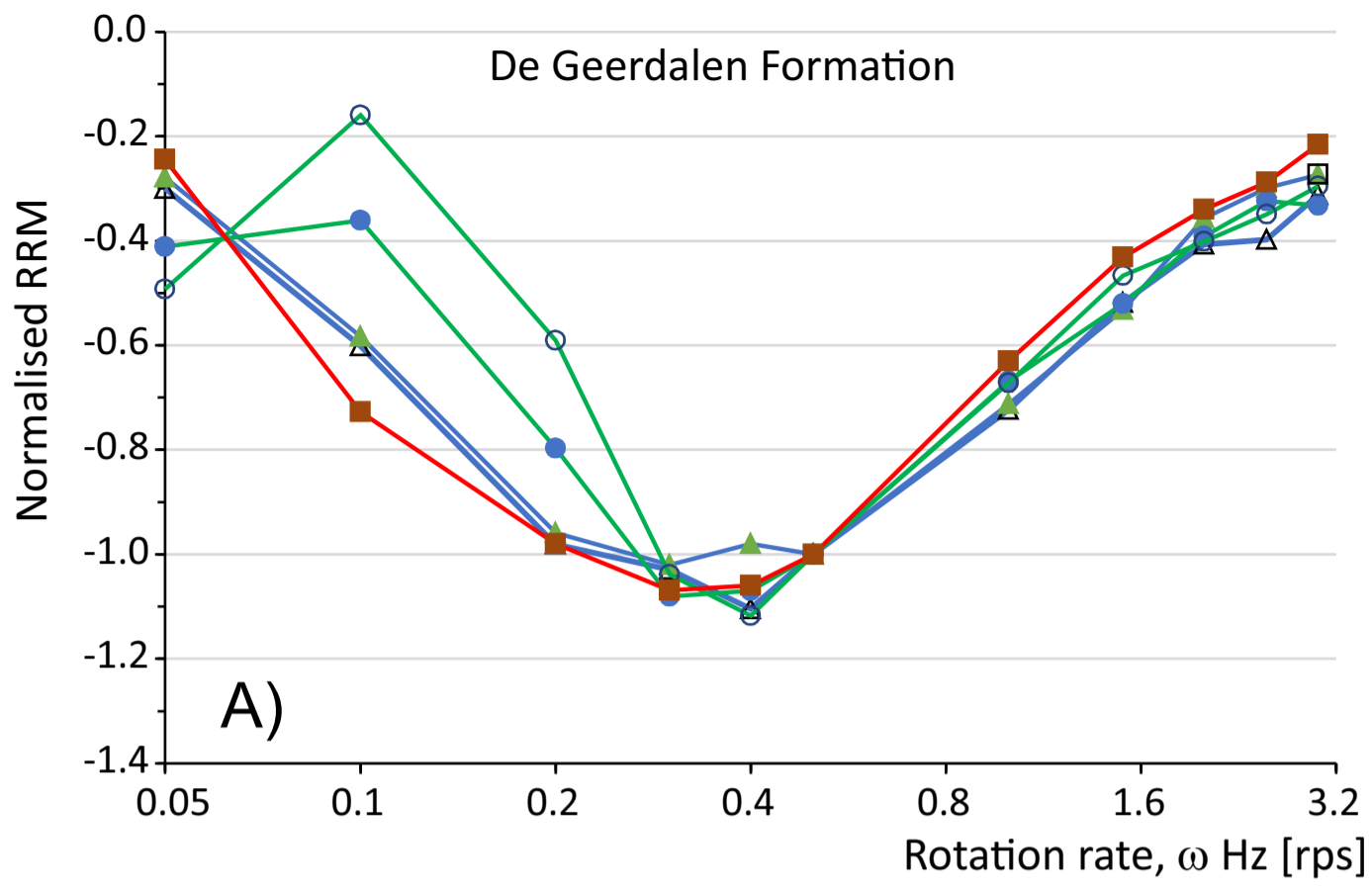
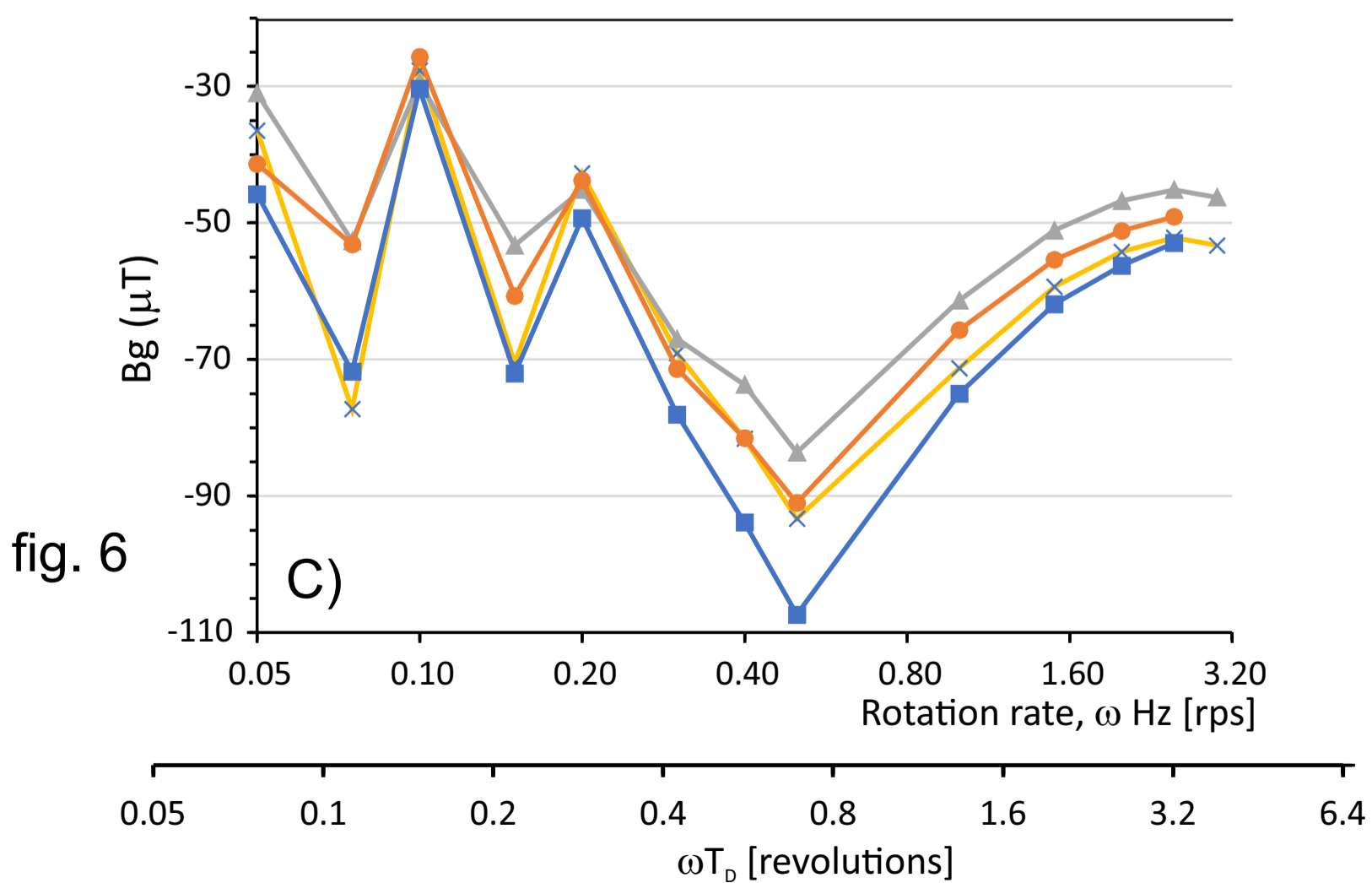
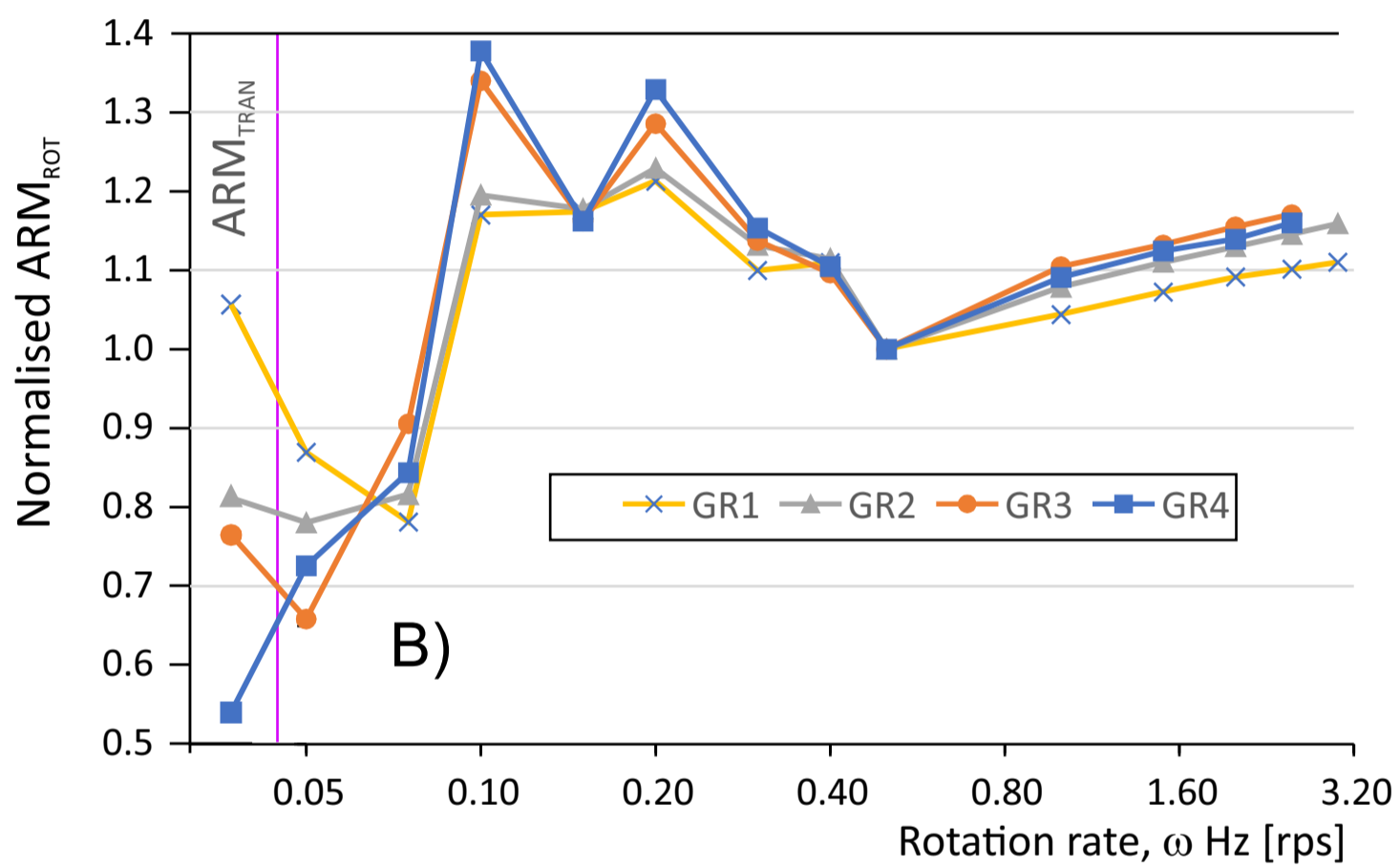
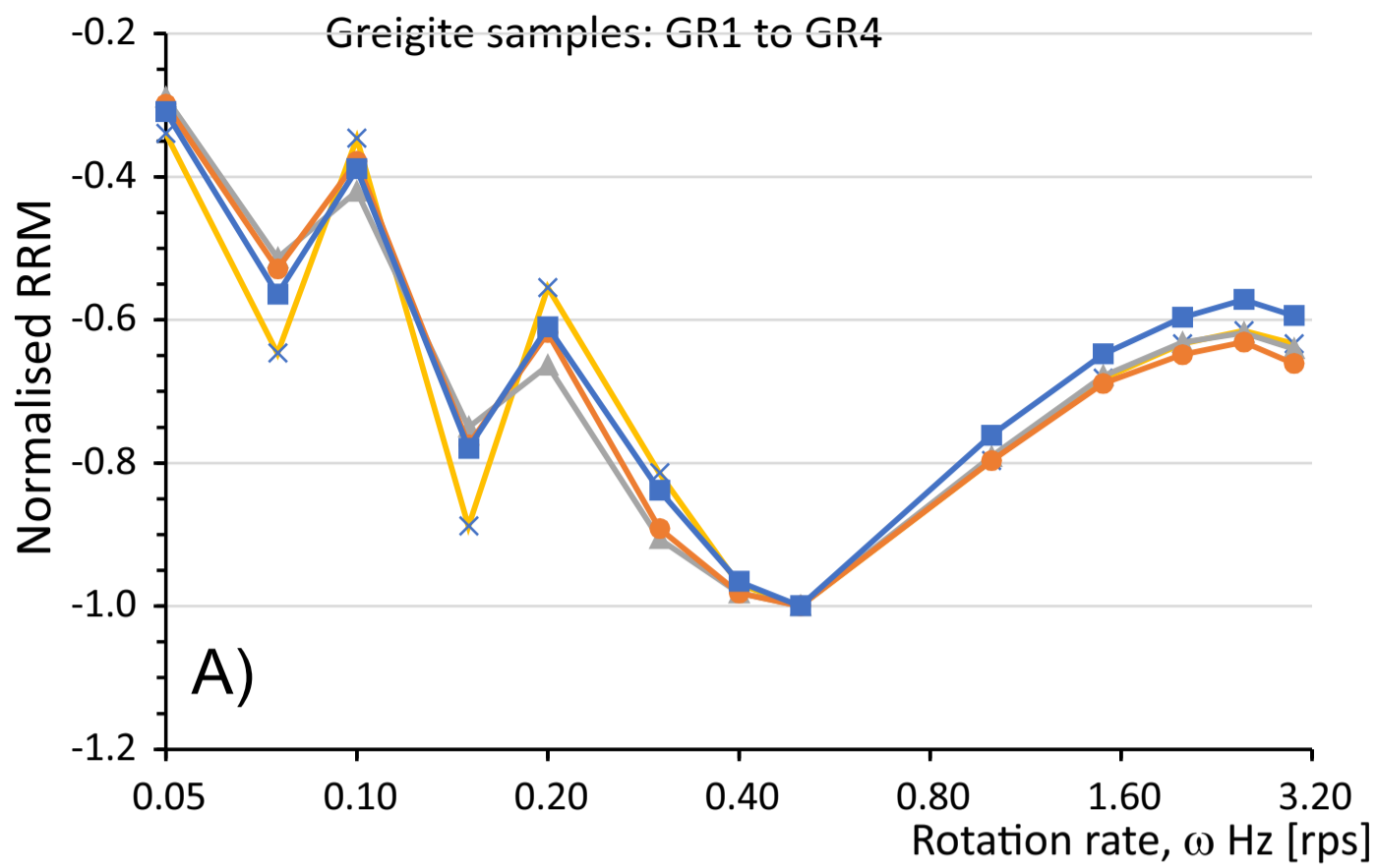


fig. 5





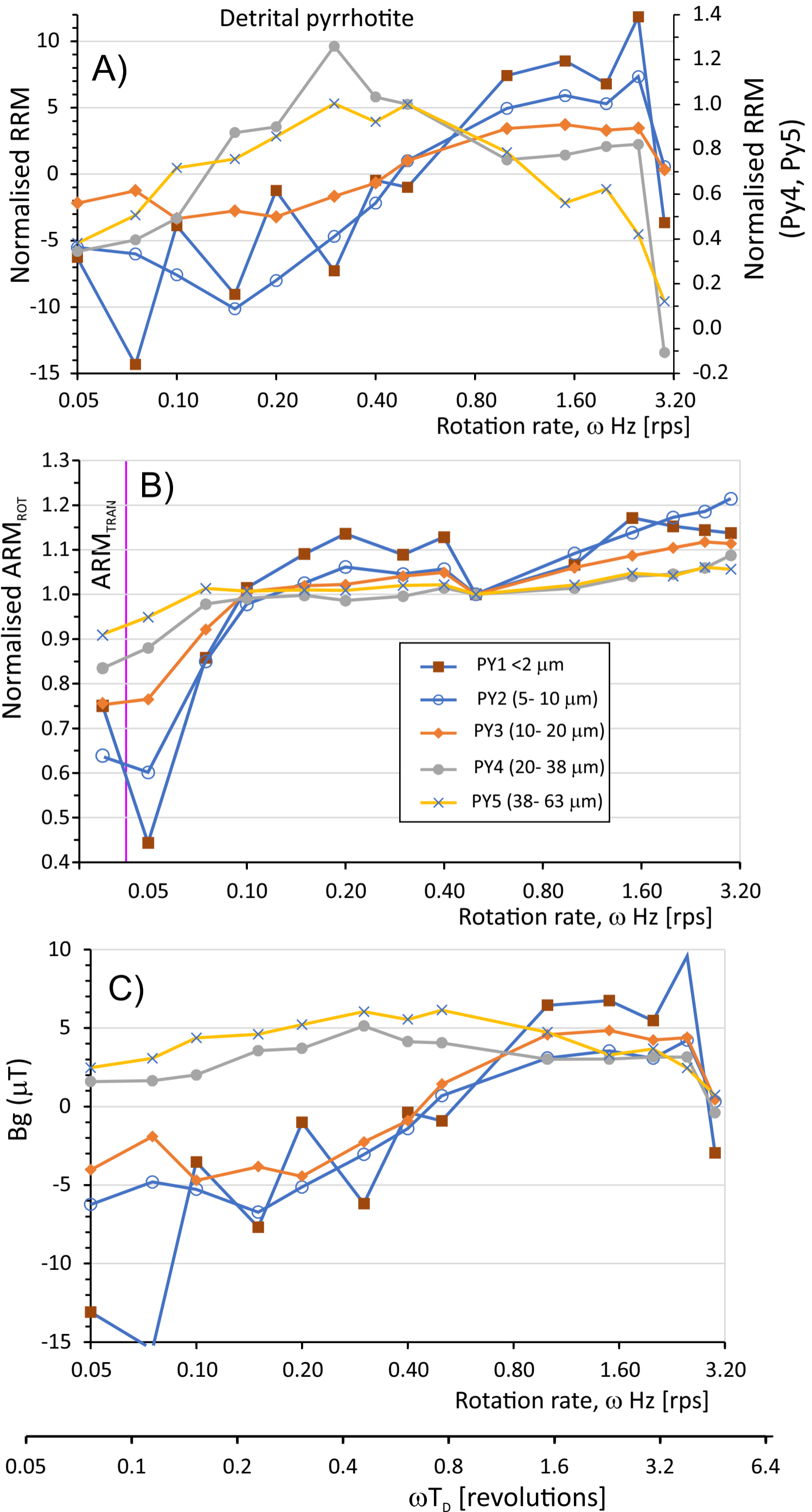


fig. 7

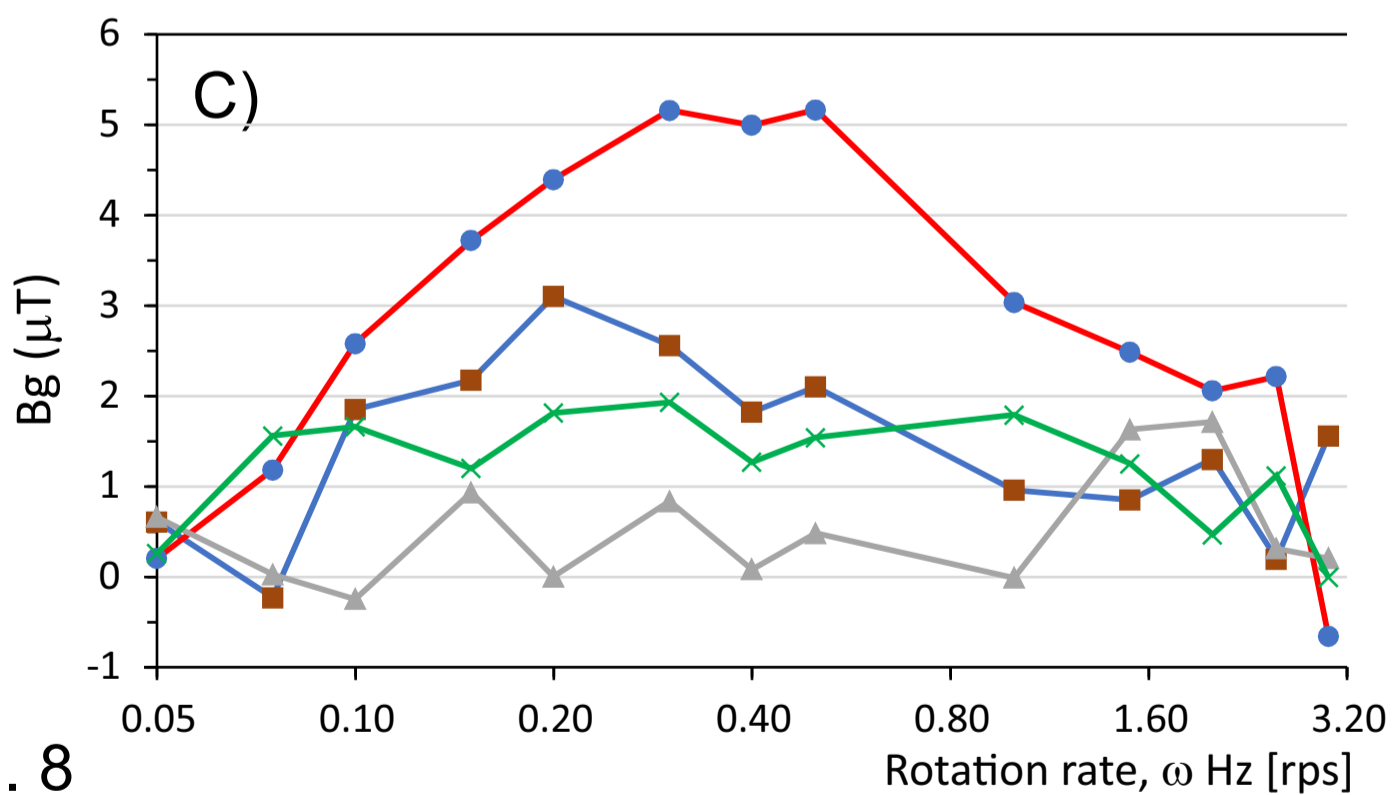
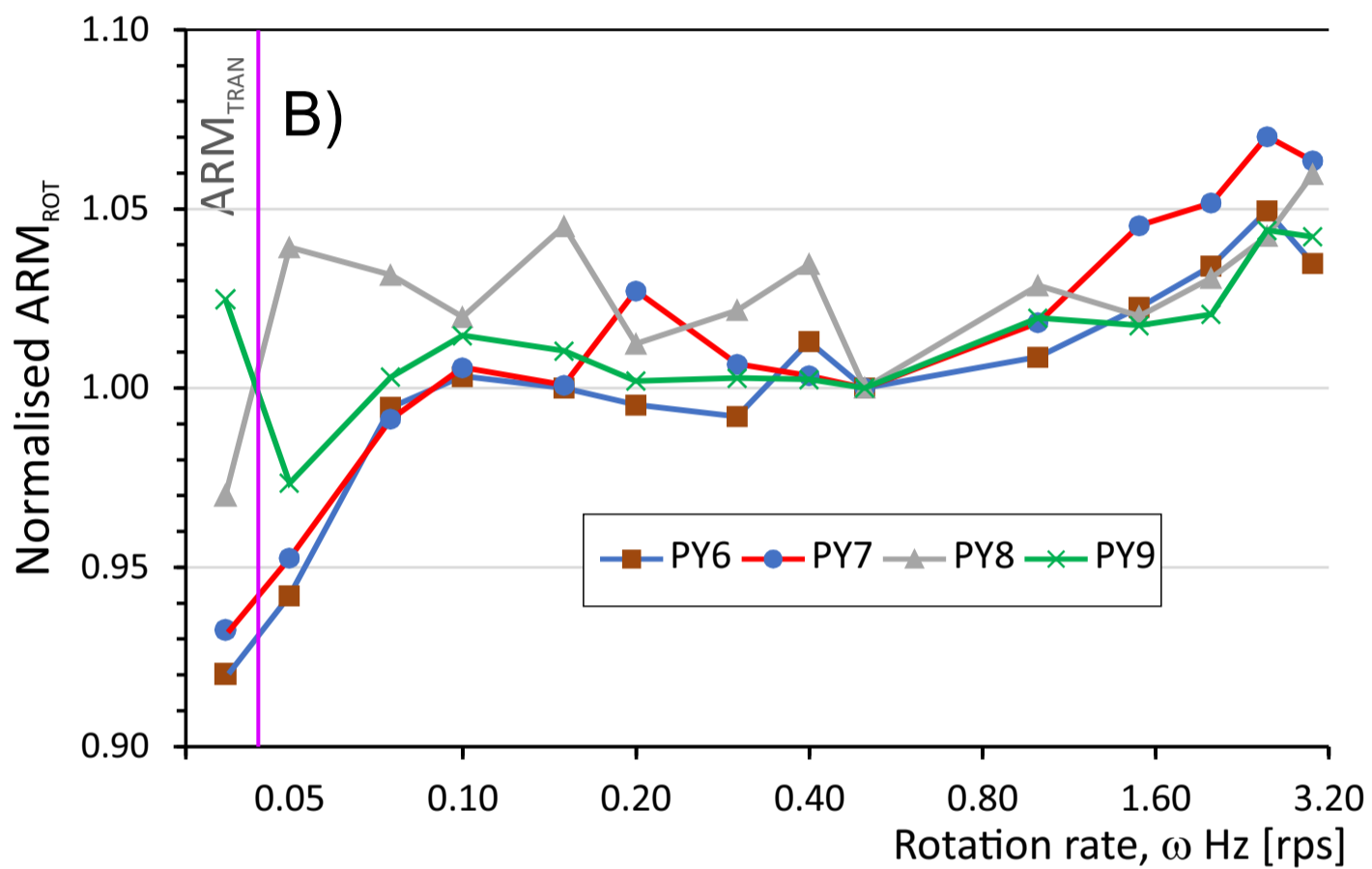
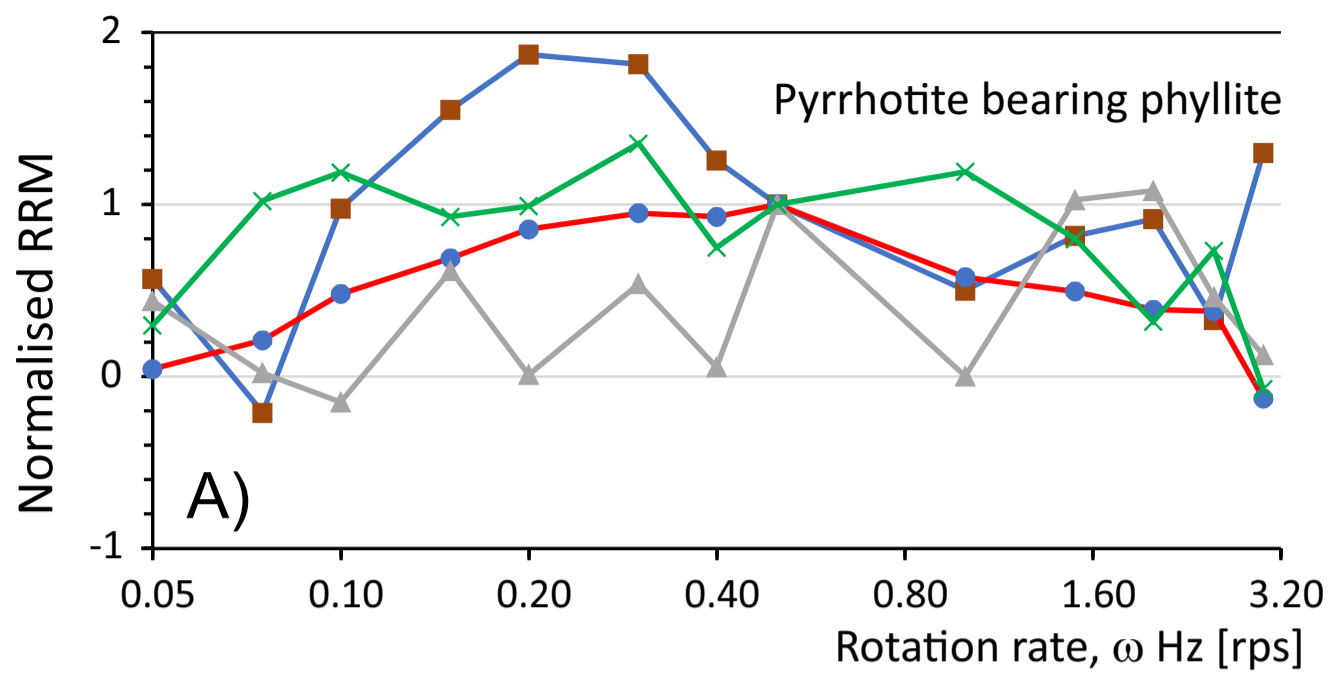
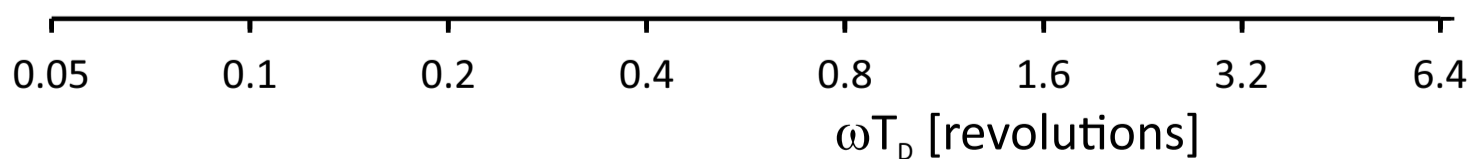
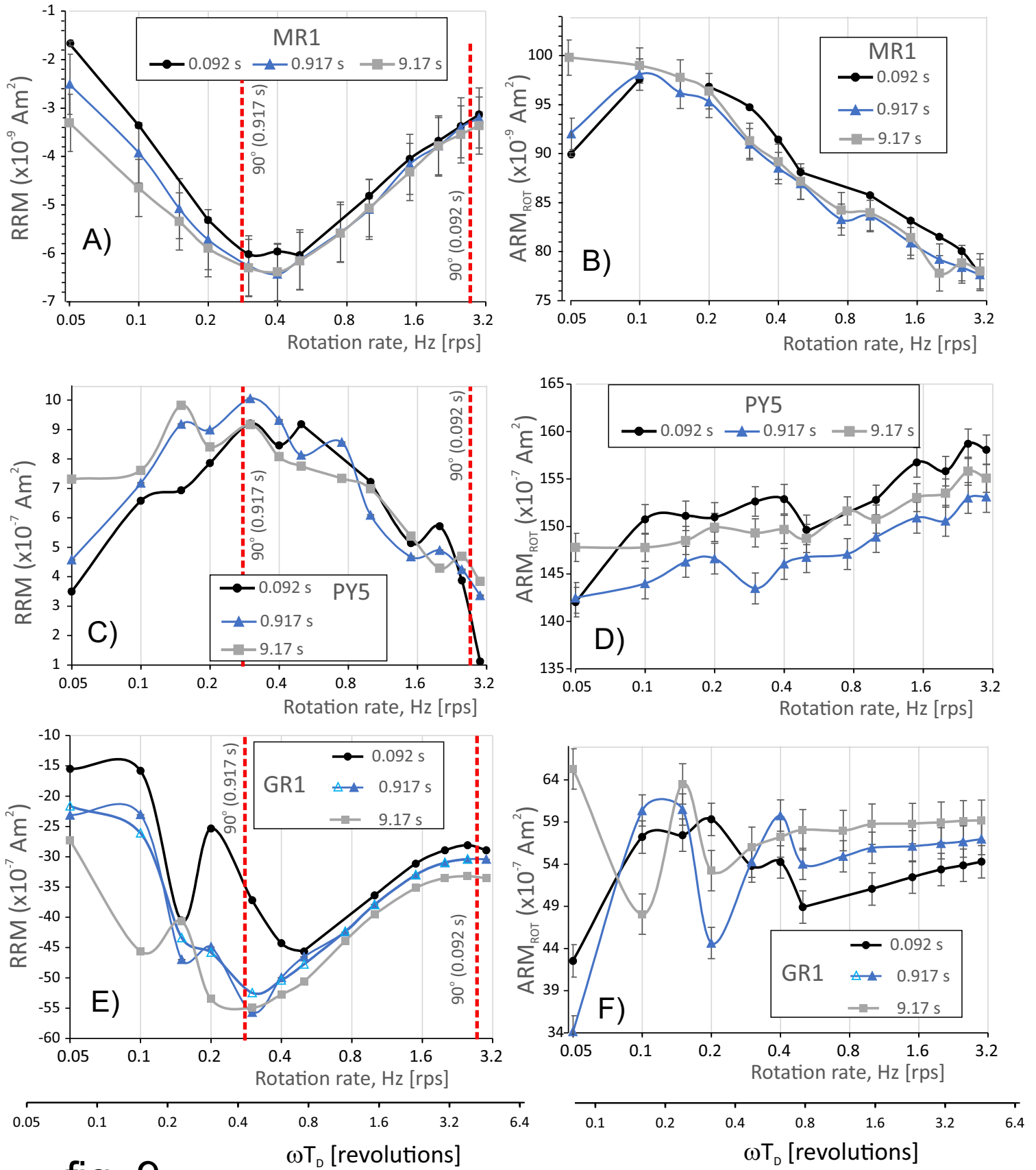


fig. 8





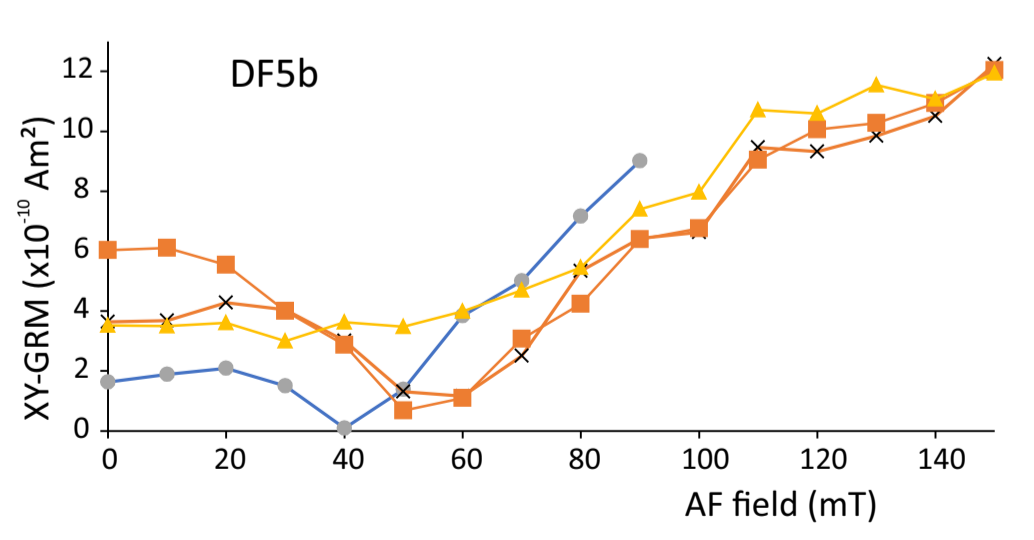
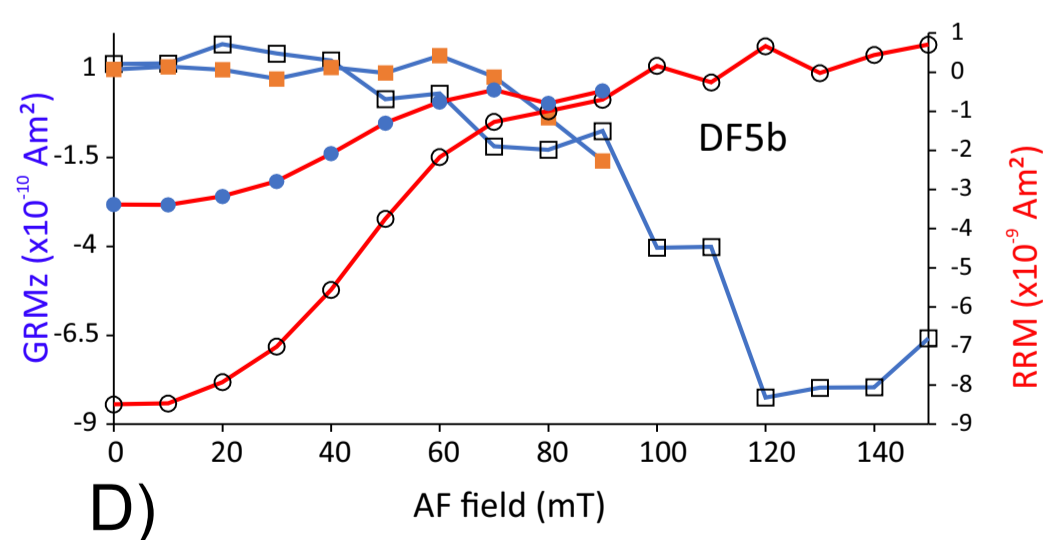
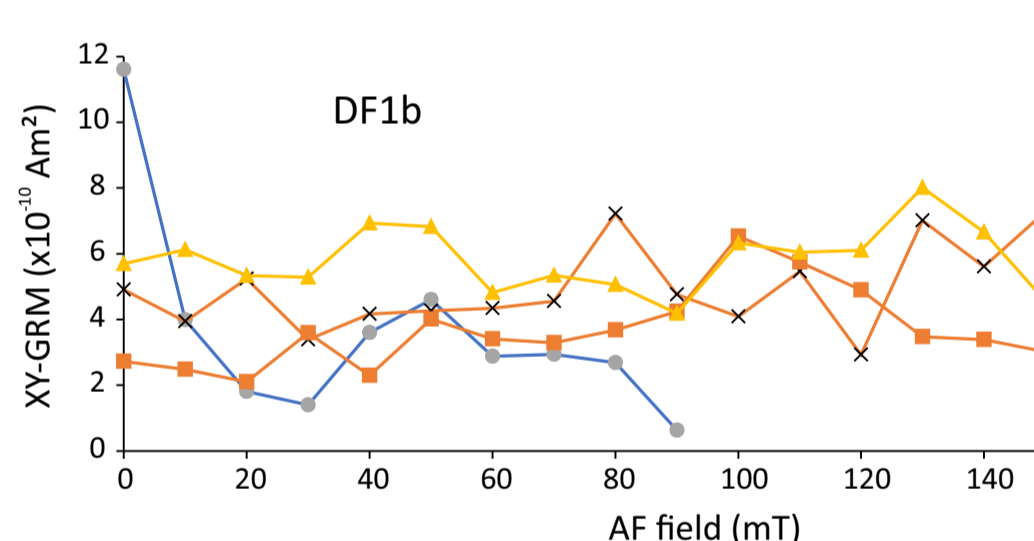
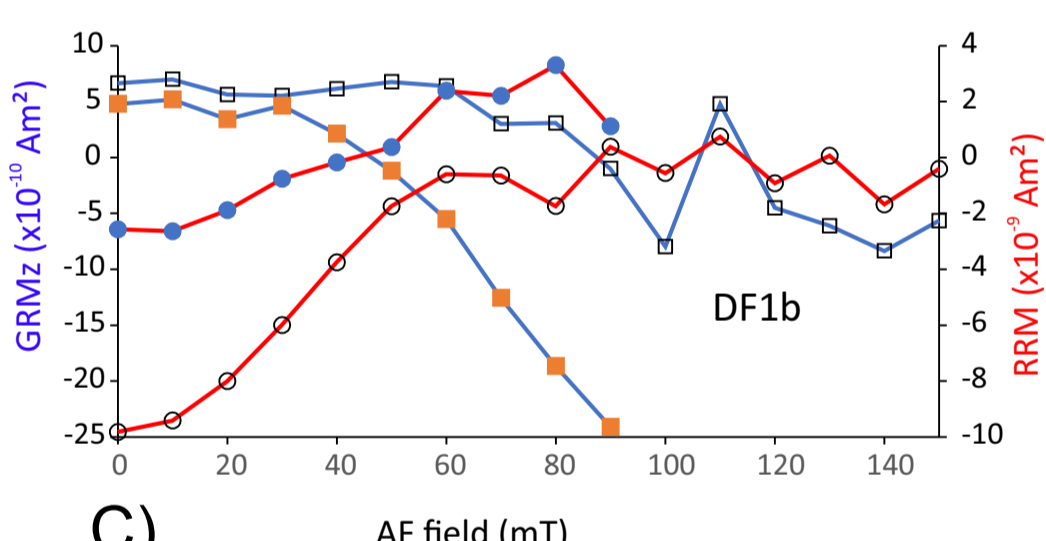
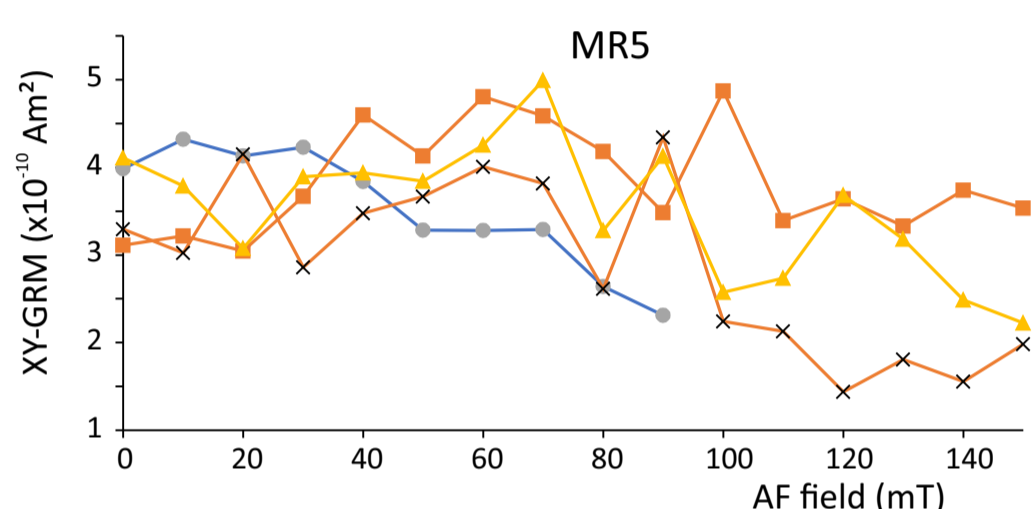
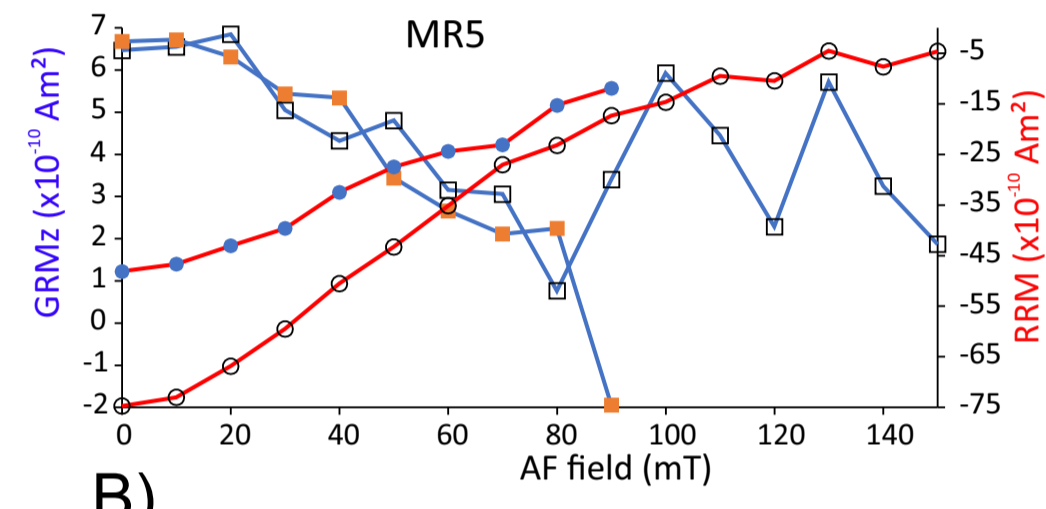
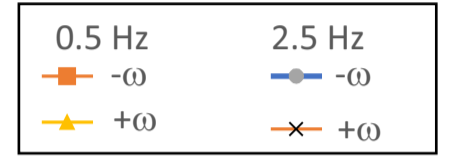
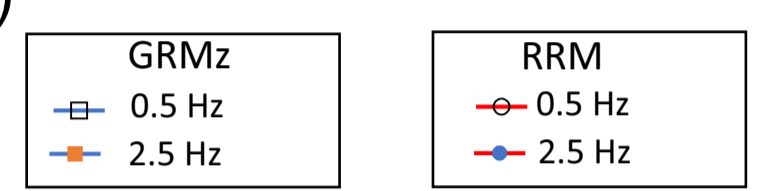
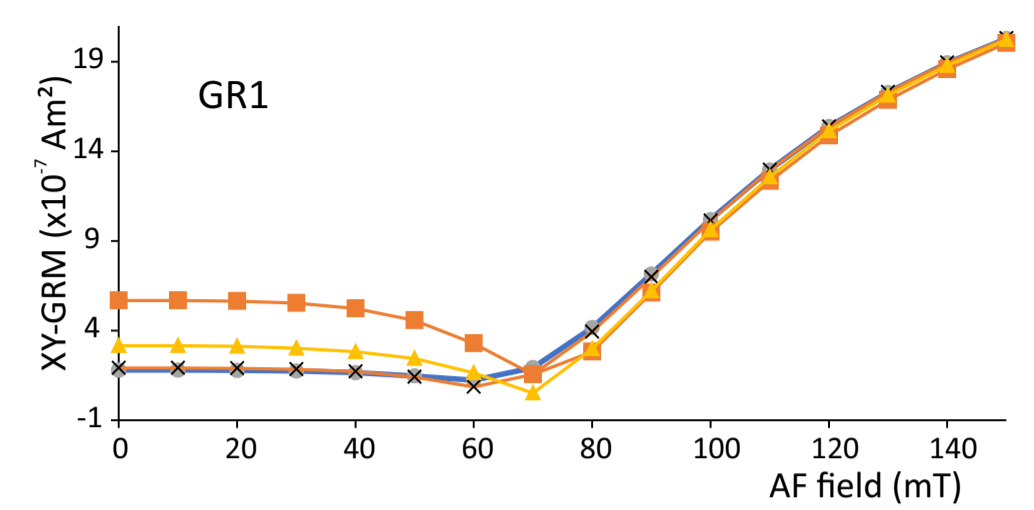
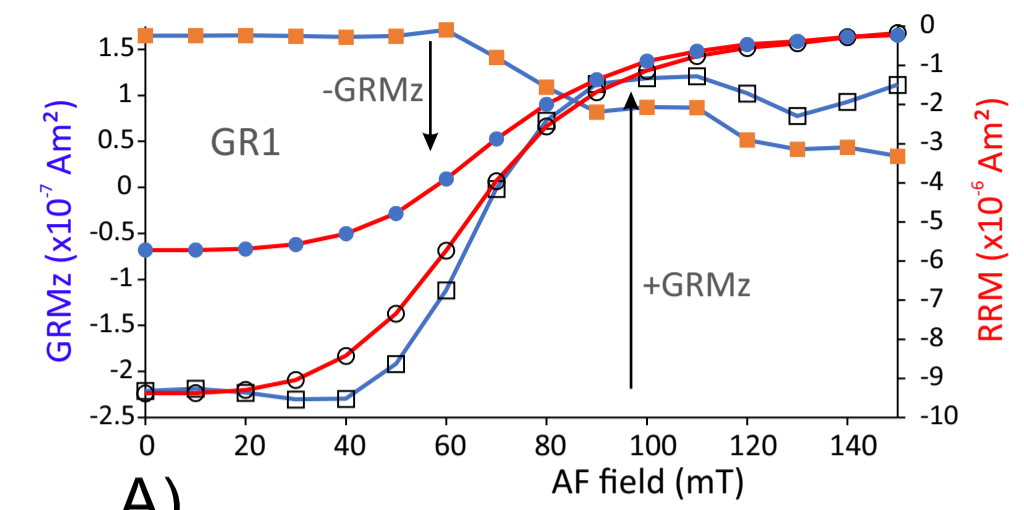


fig. 10

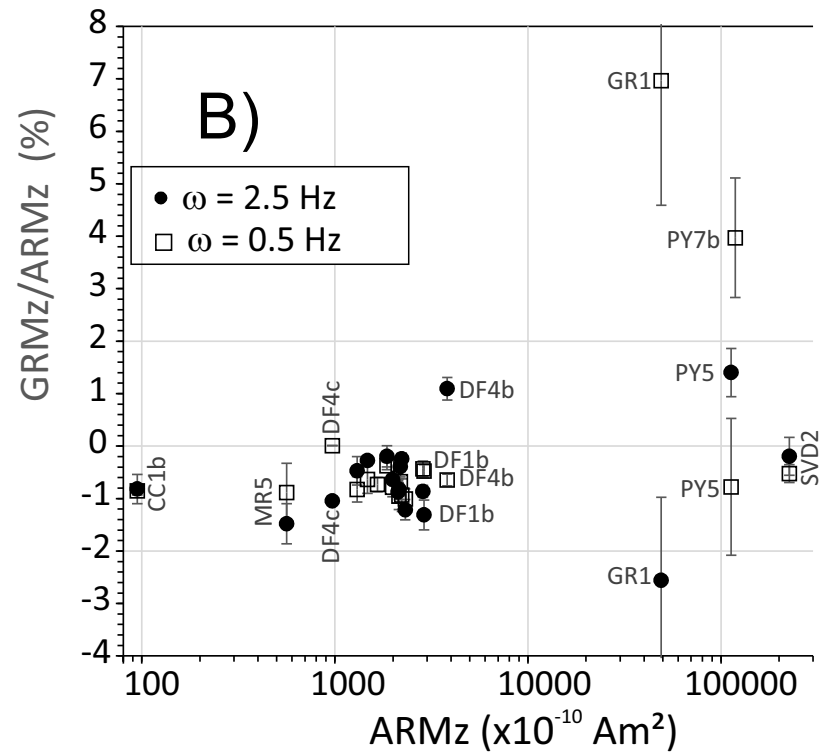
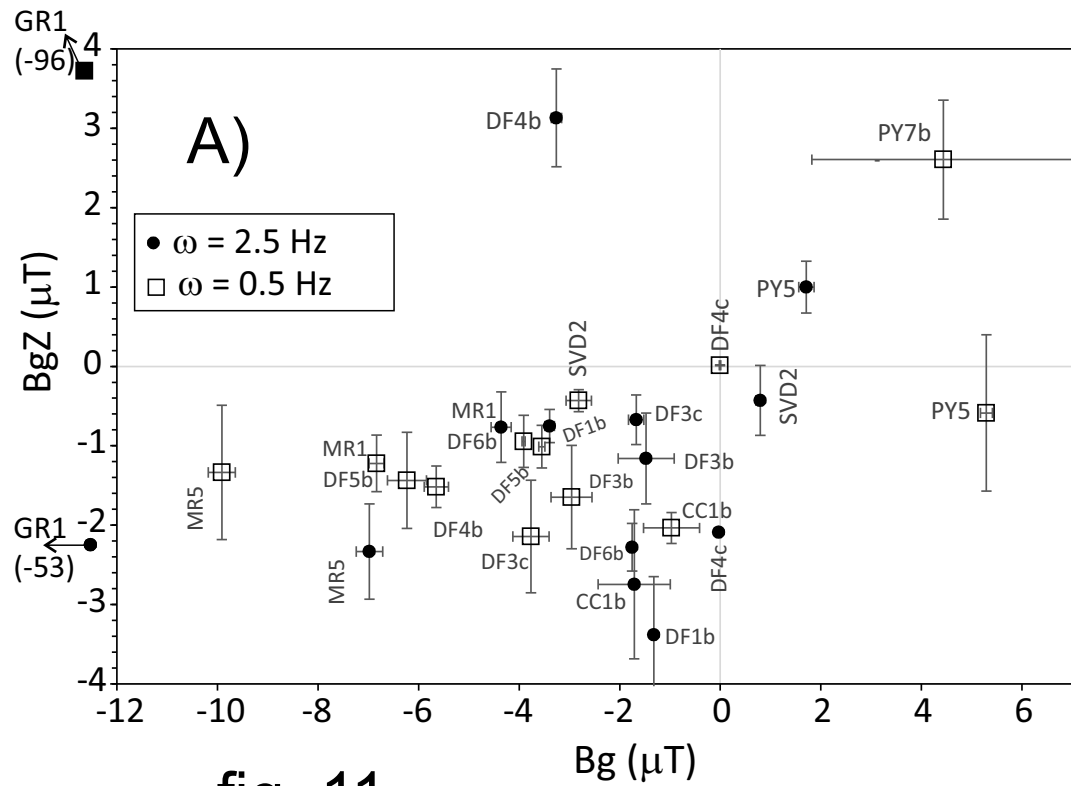
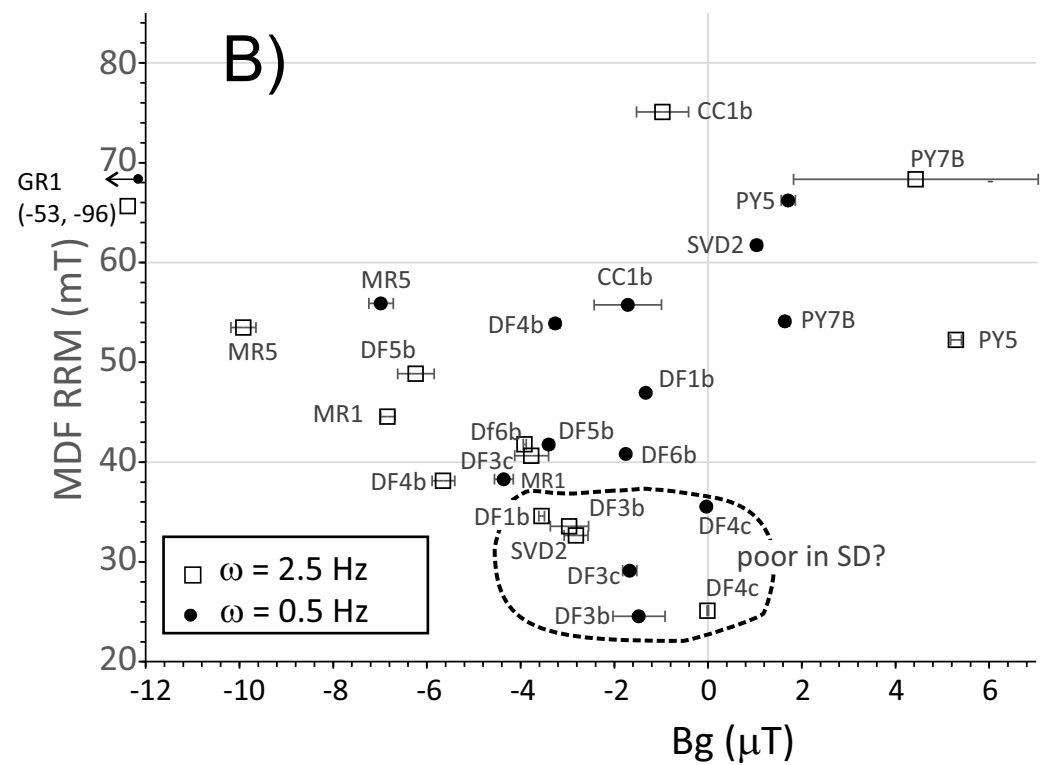
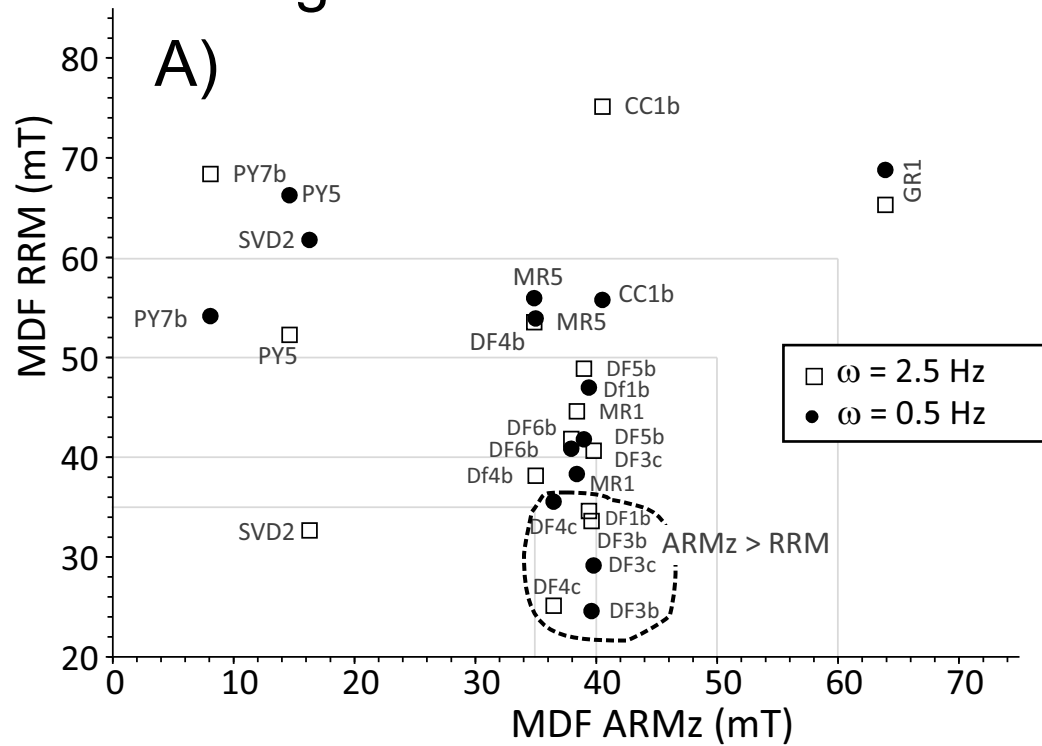


fig. 11

fig. 12



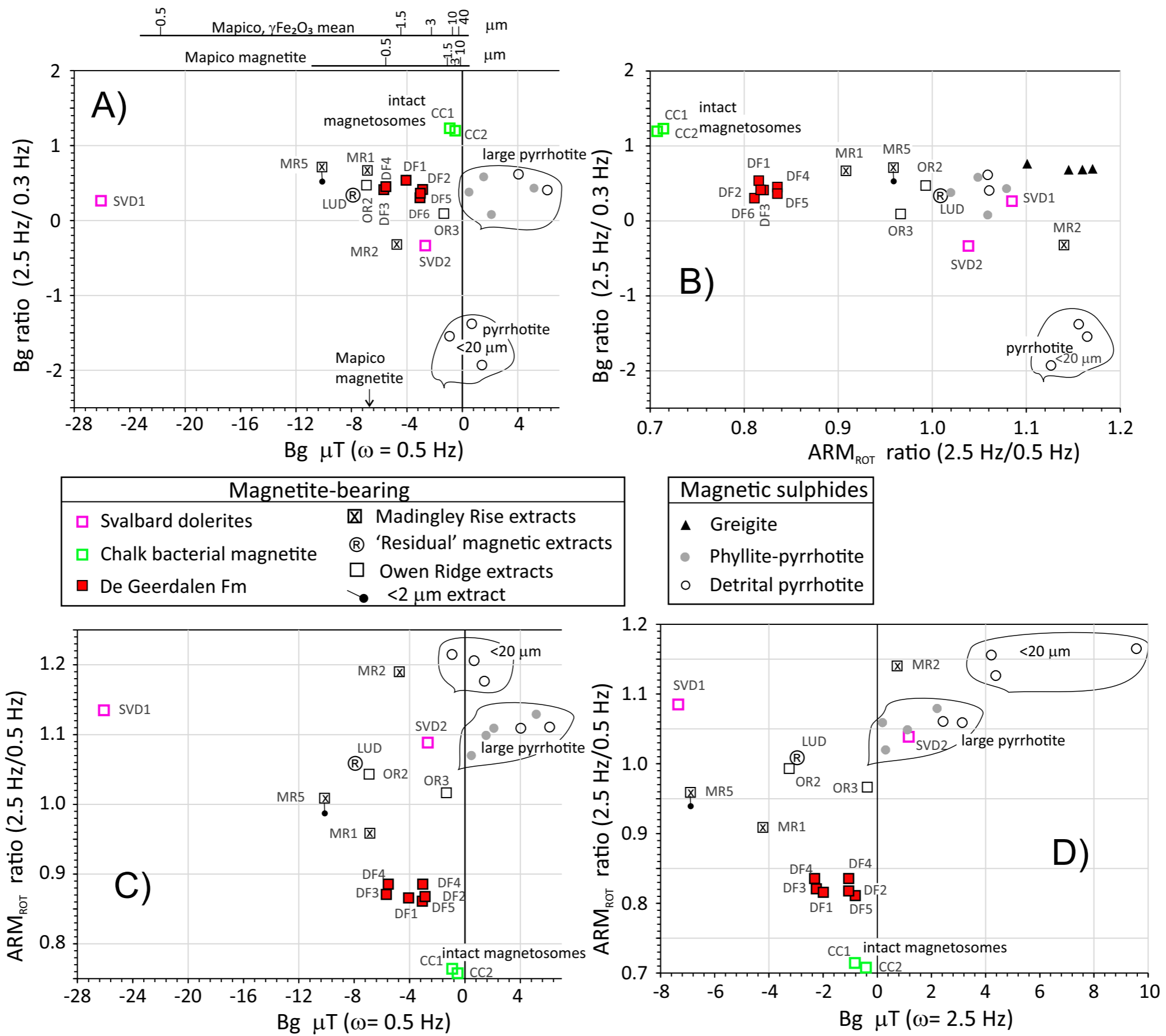


fig. 13



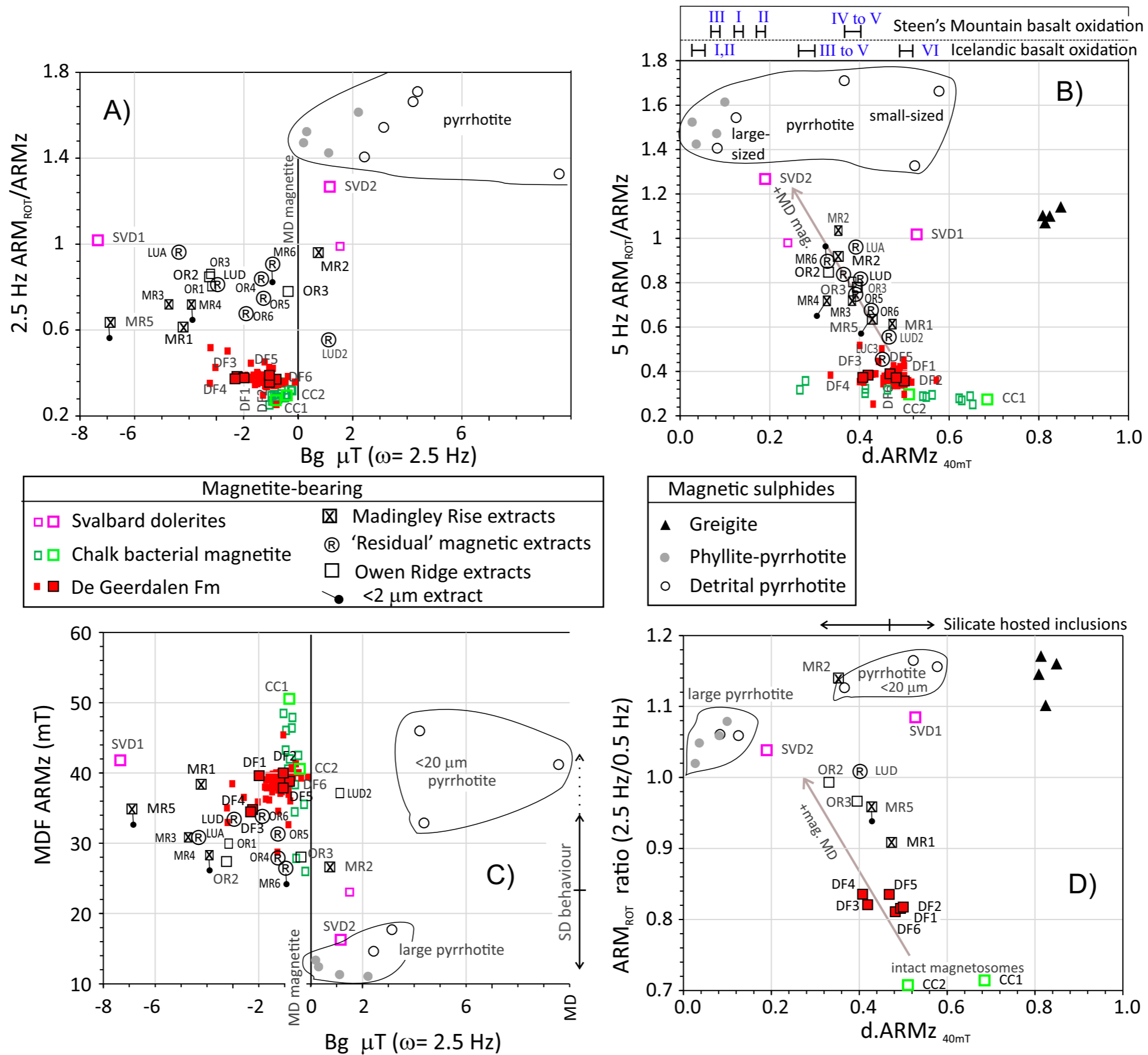


fig. 14

Table 1. Data sources for RRM and  $ARM_{ROT}$  versus rotation rates ( $\omega$  in Hz [rps]).  $T_D$ = Alternating field decay time (i.e. demagnetising field time) in secs. ?= value unspecified. \*Stephenson & Molyneux (1987) device uses decay rate of  $\sim 7 \text{ mT s}^{-1}$  (hold time,  $T_H = 5 \text{ s}$ ), so this is assumed if not given in publication, when this device is used. TM= titanomagnetite.

Mineral/ grain size	AF Hz/mT;	$\omega$ range [ $T_D$ ]	Ref. Fig.	Reference
Crushed magnetite $<0.7 \mu\text{m}$	50/ 40,80	0-180 [ $\sim 10$ ]	fig. 6	Potter & Stephenson 1986
Natural magnetite 13- $0.7 \mu\text{m}$	50/60	0-200[ $\sim 10$ ]	fig. 5	Potter & Stephenson 1986
Mapico magnetite (0.2- $0.8 \mu\text{m}$ ), Cubical	50/60	0-200[ $\sim 10$ ]	fig. 5	Potter & Stephenson 1986
Synthetic magnetite	50/50-60	0.01-6 [?]	fig. 3	Wilson & Lomax 1972
2.2-4.4 magnetite	50/80	0-100[11.4*]	fig. 2	Stephenson & Snowball 2001
Synthetic magnetite (63-125 $\mu\text{m}$ )	63.5,128,603,1210/40-55;	0.003-0.10 [45]	fig. 1	Edwards 1980a
Synthetic magnetite (63-125 $\mu\text{m}$ )	128/55;	0.003-0.40 [6.6,67]	fig. 8	Edwards 1982a,1982b
$\gamma \text{ Fe}_2\text{O}_3$ TDK tape, Type D. elongate	50/60	0-200 [ $\sim 10$ ]	fig. 5	Potter & Stephenson 1986
$\gamma \text{ Fe}_2\text{O}_3$	50/60	0-200 [ $\sim$ minutes?]	fig. 3	Madsen 2003
$\text{CrO}_2$	50/65,80,95	0-200 [ $\sim$ minutes?]	fig. 3	Madsen 2003
Heated pyrrhotite	50/80	0-130 [11.4*]	fig. 3	Thomson 1990
Various pyrrhotite bearing rocks	?/?	-20 to +20 [?]	fig. 4C	Slotznick et al. 2016
BAM8, BMR7 natural greigite	50/80	0-100 [11.4*]	fig. 3, 4	Stephenson & Snowball 2001
Igneous rock sample A66B1	50/51	0-250[ few secs]	figs 4,5	Stephenson 1980
Icelandic basalts C10-2a, S2-1a	50/80	0.01-60 [?]	fig. 5	Stephenson 1976
Igneous samples (TM bearing)	50/50-60	0.01-6 [?]	fig. 3	Wilson & Lomax 1972
Rock sample R1	63.5,128,603,1210/55	0.003-0.16 [45]	fig. 2	Edwards 1980a
Rock R1, Igneous samples F37B-1a, D264-26 (TM bearing)	128/55	0.003-0.4[6.6,34,67,101]	figs. 5,6,7	Edwards 1982b, samples from Wilson & Lomax

Table 2. RRM and Bg values measured at spin rates exceeding the AF frequency threshold. Rotation rate,  $\omega$ , in Hz [rps]. <sup>&</sup> also gives data for AF field down to 20 mT, at approximately linear decrease for all, except for  $\gamma$  Fe<sub>2</sub>O<sub>3</sub> TDK tape. \* in 10<sup>-3</sup> A/m, \*\* in 10<sup>-6</sup> A/m. Negative RRM indicate RRM acquired opposite to rotation vector (Fig. 1).  $\pm$  shows SD or uncertainty on Bg averages. Stephenson & Molyneux (1987) device uses decay rate  $\sim$ 7 mT/s (5 s hold time), so this is assumed if not given in the publication using this machine. T<sub>D</sub>= Alternating field decay time (i.e. demagnetising field) in seconds. ?=decay time unspecified. <sup>c</sup> Chen (2014) inconsistently uses sign on RRM, assume Bg has the correct sign. TM= titanomagnetite.

Mineral/ grain size	AF Hz/mT	$\omega$ [T <sub>D</sub> ], DC Field ( $\mu$ T)	Bg ( $\mu$ T)	RRM, ARM <sub>ROT</sub> ( $\times 10^{-3}$ Am <sup>2</sup> /kg)	Reference
Crushed magnetite <0.7 $\mu$ m	50/ 80	108 [ $\sim$ 10],43	108 <sup>&amp;</sup>	95,	Potter & Stephenson 1986
Crushed magnetite 0.7-2.2 $\mu$ m	50/ 80	108 [ $\sim$ 10], 43	44 <sup>&amp;</sup>	33,	Potter & Stephenson 1986
Mapico magnetite (0.2-0.8 $\mu$ m), Cubical	50/ 80	108 [ $\sim$ 10], 43	301 <sup>&amp;</sup>	177,	Potter & Stephenson 1986
Crushed magnetite 2.2-4.4 $\mu$ m	50/ 80	108 [ $\sim$ 10], 43	28 <sup>&amp;</sup>	15.5,	Potter & Stephenson 1986
Crushed magnetite 2.2-4.4 $\mu$ m	50/80	95 [11.4 <sup>*</sup> ], 70	28	10.9,39	Stephenson & Snowball 2001
Crushed magnetite 4.4-7.6 $\mu$ m	50/ 80	108 [ $\sim$ 10], 43	17.8 <sup>&amp;</sup>	9.2,	Potter & Stephenson 1986
Crushed magnetite 7.6-13.1 $\mu$ m	50/ 80	108 [ $\sim$ 10], 43	8 <sup>&amp;</sup>	2.9,	Potter & Stephenson 1986
Crushed magnetite 13.1-25.5 $\mu$ m	50/ 80	108 [ $\sim$ 10], 43	4.1 <sup>&amp;</sup>	1.01,	Potter & Stephenson 1986
SD3 Crushed magnetite	50/95	90 [minutes], 150	330	6.8,18.3 <sup>*</sup>	Madsen 2004
$\gamma$ Fe <sub>2</sub> O <sub>3</sub> TDK tape, Type D. elongate	50/80	108 [ $\sim$ 10], 43	161 <sup>&amp;</sup>	330,	Potter & Stephenson 1986
$\gamma$ Fe <sub>2</sub> O <sub>3</sub> , BASF FT26 tape	50/95	90 [minutes], 43	85 to 150	167,218	Madsen 2003, 2004
CrO <sub>2</sub> , BASF CK40-13 tape	50/95	63 [minutes], 70	-15	83.5,7098	Madsen 2003,2004
Natural SD magnetite magnetosomes (water column)	50/80	95[11.4], 70	0	0,-	Chen 2014 <sup>c</sup>
MV-1 magnetite magnetosomes	50/80	95[11.4], 70	1, 10.9	0.105, 0.674 <sup>**</sup>	Chen 2014 <sup>c</sup>
SD magnetite and greigite	50/80	95[11.4], 70	2.4, 5.6	0.037,0.461	Chen 2014 <sup>c</sup>
SD greigite magnetosomes	50/80	95[11.4], 70	7.3, 22.5	0.0006, 0.00187	Chen 2014 <sup>c</sup>
Heated (315°C) vein pyrrhotite	50/80	100[11.4-86.4], 70	-	3.3,-	Thomson 1990
<b>Rock and pottery samples below</b>					
DG1 vesicular basalt, Low Ti-TM	50/95	90 [minutes], 150	34	1552,1910 <sup>*</sup>	Madsen 2004
DG17 vesicular basalt, Low Ti-TM	50/95	90 [minutes], 150	26	1014,2384 <sup>*</sup>	Madsen 2004
T97A flood basalt with TM	50/95	65 [minutes], 150	9	327.7,5141 <sup>*</sup>	Madsen 2004
T54A sill, with magnetite	50/95	70 [minutes], 150	38	1368,5824 <sup>*</sup>	Madsen 2004

TC05, TM10, SD TM, Tiva Canyon Tuff	50/80	95[11.4], 70	32	0.804,1.76	Chen 2014 <sup>c</sup>
Supska potsherd	50/80	95[11.4],70	162	-, -	Mahon & Stephenson 1997
Soba potsherds	50/80	95[11.4],70	82 ±16	-, -	Mahon & Stephenson 1997
Baranda potsherds	50/80	95[11.4],70	23	-, -	Mahon & Stephenson 1997
YB7 marine sediment, maghemite	50/95	90 [minutes], 150	89	26.7,41.9*	Madsen 2004
Greigite bearing lake sediments	50/80	95 [11.4],70	482± 221	264,33.1*	Hu et al. 2002
BAM8 natural greigite	50/80	95 [11.4*], 70	1050	998,95	Stephenson & Snowball 2001
BMR7 natural greigite	50/80	95 [11.4*], 70	1070	375,35	Stephenson & Snowball 2001
SD greigite, Gutingkeng Fm (EJC49.1c, EJC50.1c)	50/80	95[11.4], 70	552-790	0.201,0.0210	Chen 2014 <sup>c</sup>

Table 3. RRM and Bg values measured at spin rates below the AF frequency threshold. Rotation rate,  $\omega$ , in Hz [rps]. <sup>AR</sup>= partial rotational ARM (static sample during AF decay). See caption to Table 2 for details.

Mineral/ grain size	AF Hz/mT	$\omega$ [T <sub>D</sub> ], DC Field ( $\mu$ T)	Bg ( $\mu$ T)	RRM, ARM <sub>rot</sub> ( $\times 10^{-3}$ Am <sup>2</sup> /kg)	Reference
<b>Low spin rate data</b>					
$\gamma$ Fe <sub>2</sub> O <sub>3</sub> Sony MC30	50/100	5 [?], 100	-5.5 <sup>Ss</sup>	-, -	Snowball 1997b
$\gamma$ Fe <sub>2</sub> O <sub>3</sub> TDKD1	50/100	5 [?], 100	-8.0 <sup>S</sup>	-, -	Snowball 1997b
$\gamma$ Fe <sub>2</sub> O <sub>3</sub> TDKD2	50/100	5 [?], 100	-7.1 <sup>S</sup>	-, -	Snowball 1997b
MCF0-1, Co- $\gamma$ Fe <sub>2</sub> O <sub>3</sub>	128/46	0.07 [15], 60	-37	-,12 <sup>AR</sup>	Edwards & Desta 1989
GFO-1, $\gamma$ Fe <sub>2</sub> O <sub>3</sub>	128/46	0.07 [15], 60	-30	-,31 <sup>AR</sup>	Edwards & Desta 1989
MBL-1, Mapico Black magnetite	128/46	0.07 [15], 60	-8	-,9.7 <sup>AR</sup>	Edwards & Desta 1989
Magnetite (63-125 $\mu$ m), Mag6	128/18.4	0.4 [17], 60	-3	-,2.13 <sup>AR</sup>	Edwards & Desta 1989
MDM- MD detrital magnetite	50/100	5 [?], 100	0 <sup>S</sup>	-, -	Snowball 1997b
CRO-1 CrO <sub>2</sub> powder	128/46	0.07 [15], 60	-5	-,2.0 <sup>AR</sup>	Edwards & Desta 1989
CrO <sub>2</sub> TDKMF1	50/100	5 [?], 100	-31 <sup>S</sup>	-, -	Snowball 1997b
Rock R1-a	128/36.4	0.4 [17], 45	-15	-,0.21 <sup>AR</sup>	Edwards & Desta 1989
D264-2a (TM bearing)	128/36.4	0.4 [17], 60	0	-,0.12 <sup>AR</sup>	Edwards & Desta 1989

Carboniferous Lavas	800/150	1 [30], 15	-3.1 ±1.9 <sup>S</sup>	-260,1350*	Roperch & Taylor 1986
Miocene Basalt (with -RRM)	800/150	1 [30], 15	-13.6 ±16.7 <sup>S</sup>	-1200,1850*	Roperch & Taylor 1986
Miocene Basalt (with +RRM)	800/150	1 [30], 15	8.9±1.2 <sup>S</sup>	1120,1910*	Roperch & Taylor 1986
Pliocene Basalt	800/150	1 [30], 15	-5.6 ±1.2 <sup>S</sup>	-290,780*	Roperch & Taylor 1986
Serpentinite	800/150	1 [30], 15	-3.8±4.8 <sup>S</sup>	-50,200*	Roperch & Taylor 1986
Pajep8- SD magnetite magnetofossils+detrital	50/100	5 [52?], 100	-0.14 <sup>S</sup>	-, -	Snowball 1997b
Hol81- SD detrital magnetite	50/100	5 [52?], 100	-14 <sup>S</sup>	-, -	Snowball 1997b
Hol252- SD detrital magnetite	50/100	5 [52?], 100	-3.6 <sup>S</sup>	-, -	Snowball 1997b
Embmr1-SD greigite sediment	50/100	5 [52?], 100	-98 <sup>S</sup>	-, -	Snowball 1997b
Embmr5-SD greigite sediment	50/100	5 [52?], 100	-110 <sup>S</sup>	-, -	Snowball 1997b
Embmr6-SD greigite sediment	50/100	5 [52?], 100	-84 <sup>S</sup>	-, -	Snowball 1997b
Embmr8-SD greigite sediment	50/100	5 [52?], 100	-101 <sup>S</sup>	-, -	Snowball 1997b
Bmr9x-SD greigite concentrate	50/100	5 [52?], 100	-108 <sup>S</sup>	-, -	Snowball 1997b
St15-SD greigite sediment	50/100	5 [52?], 100	-119 <sup>S</sup>	-, -	Snowball 1997b
St79x-SD greigite concentrate	50/100	5 [52?], 100	-137 <sup>S</sup>	-, -	Snowball 1997b
Heated (315°C) pyrrhotite	50/80	~5 [11.4-86.4], 70	-	1.9,-	Thomson 1990

Table 4. The test samples used in this work. SD=single domain, MD=multidomain. T<sub>c</sub>= Curie temperature; TM= titanomagnetite, Ti content (wt%). See the SI for further details about these test samples.

Sample codes	Description	Mineralogy	References
PY1	< 5 μm magnetic separate from river sediment, Taiwan	Monoclinic pyrrhotite	Horng & Roberts 2006.
PY2	5-10 μm magnetic separate from river sediment, Taiwan	Monoclinic pyrrhotite	Horng & Roberts 2006.
PY3	10-20 μm magnetic separate from river sediment, Taiwan	Monoclinic pyrrhotite	Horng & Roberts 2006.
PY4	20-38 μm magnetic separate from river sediment, Taiwan	Monoclinic pyrrhotite, quartz, goethite	Horng & Roberts 2006.

PY5	38-63 $\mu\text{m}$ magnetic separate from river sediment, Taiwan	Monoclinic pyrrhotite, quartz, goethite, lepidocrocite, chlorite	Horng & Roberts 2006.
PY6-PY9	Monoclinic pyrrhotite bearing phyllites from Taiwan metamorphic Central Range	Monoclinic pyrrhotite	Horng et al. 2012,
GR1-GR4	Greigite bearing mudstones, Plio-Pleistocene Lower Gutingkeng Formation, SW Taiwan	Fine-grained greigite	Horng et al. 1998; Jiang et al. 2001.
LUA-LUD	<38 $\mu\text{m}$ magnetic extract from Late Triassic Lunde Fm, N. Sea.	Mn and Cr-bearing magnetites, chromite, (magnetite) oxide inclusions in various silicates.	Hounslow et al. 1995; Hounslow 1996
LUD2, LUD3	38-63 $\mu\text{m}$ and 63-250 $\mu\text{m}$ magnetic extract from Late Triassic Lunde Fm, N. Sea	Dominated by Fe-oxide (magnetite) inclusions in silicates, lesser chromite and Mn-magnetite.	Hounslow et al. 1995; Hounslow 1996.
OR1, OR2	<38 $\mu\text{m}$ Magnetic extract from 0.35 m, 1.25 m depth in IODP 722B, Indian Ocean, Owen Ridge	Magnetite as mixed detrital and very minor bacterial magnetite. Fe-oxide inclusions in various silicates (probably minor contribution)	Hounslow & Maher, 1996; 1999
OR3	<38 $\mu\text{m}$ Magnetic extract from 7.8 m, depth in IODP 722B, Indian Ocean, Owen Ridge	Detrital magnetite. Fe-oxide inclusions in various silicates, ilmenite.	Hounslow & Maher, 1996; 1999
OR4, OR5, OR6	<38 $\mu\text{m}$ Magnetic extract from 38.6m, 40.6 m, 60-62 m in IODP 722B, Indian Ocean, Owen Ridge	Residual assemblage of detrital magnetite, ilmenite and major Fe-oxide inclusions in various silicates.	Hounslow & Maher, 1996; 1999
MR1	<38 $\mu\text{m}$ Magnetic extract from 118-120 m in IODP Site 709A, Madingly Rise, Indian Ocean (oxic interval)	Magnetite as mixed detrital and important bacterial magnetite.	Hounslow & Maher, 1996; 1999
MR2, MR3	<38 $\mu\text{m}$ Magnetic extract from 112.9 m and 28-29.7 m in IODP Site 709A, Madingly Rise, Indian Ocean (reduced interval)	Magnetite as mixed detrital and less important bacterial magnetite.	Hounslow & Maher, 1996; 1999
MR4	<~2 $\mu\text{m}$ Magnetic extract from 5.89-8.41 m and IODP Site 709A, sample levels with SIRM/ $\chi$ ARM >390 $\text{Am}^{-1}$ . $E_{\text{mpt}}$ extract.	Major bacterial magnetite with similar content of fine detrital magnetite	Maher & Hounslow 1999, unpubli. data

MR5	<~2 $\mu\text{m}$ magnetic extract from 22.8-23.5m in IODP site 709C, Madingly Rise, Indian Ocean. $E_{\text{mpt}}$ extract.	Minor bacterial magnetite with larger content of fine detrital magnetite	Maher & Hounslow 1999, unpubl. data
MR6	<~2 $\mu\text{m}$ magnetic extract from 45.8-47.3 m in IODP site 709C, Madingly Rise, Indian Ocean. $E_{\text{mpt}}$ extract (sulphide-reduced interval)	Residual assemblage of detrital magnetite, ilmenite, and major Fe-oxide inclusions in various silicates.	Maher & Hounslow 1999; unpubl. data
CC1, CC2	Whole rock Upper Cretaceous Chalk from the level of WC9b at Culver Cliff, Isle of Wight, UK.	Rich in bacterial magnetite in chains, no detrital oxides, but common Fe-oxides as inclusions in silicates.	Montgomery et al. 1998; Kopp & Kirschvink 2008
SVD1, SVD2	Fine-grained dolerites from the Diasbasoddon suite of the High Arctic province, sills near lower contacts at Botneheia, Svalbard	TM (20-31)–ilmenite intergrowths, altered to maghemite at margins. 1–500 $\mu\text{m}$ in size; Low alteration, SD–MD behaviour. Maybe minor pyrrhotite. $T_c$ 520-550°C	Halvorsen, 1973; Vincenz et al. 1984; Nejbort et al. 2011.
DF1 to DF6	Sandstones, siltstones, De Geerdalen Fm, Svalbard. DF1, DF2 from central Spitsbergen, DF3, DF4 from Hopen, DF5, DF6 from E. Svalbard.	Detrital magnetite, $T_c$ ~450-600°C, in part oxidized, possibly wide range of grain sizes.	Hounslow et al. 2007; 2022.





## Supplementary information for: Rotational remanent magnetisation as a magnetic mineral diagnostic tool at low rotation rates

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### The supplementary information consists of:

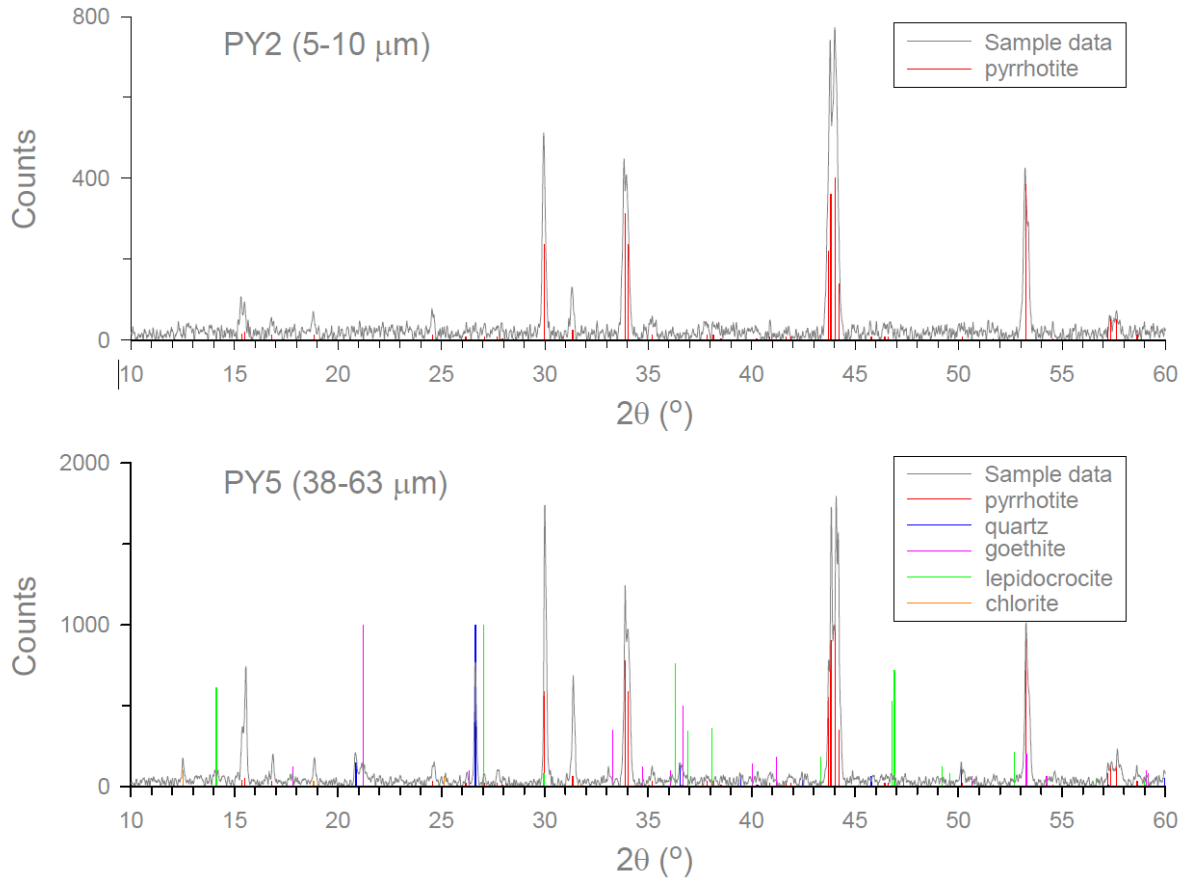
- Section 1- Additional magnetic and mineralogical information about the test samples
- Section 2- Additional information about measurement uncertainty on the RAPID.
- Section 3- Additional characteristics of RRM-related magnetisations
- Section 4- An evaluation of AF hold times and AF frequency on the RAPID
- Associated excel file which contains the data used in this study.

### 1. Further details about the test samples used in this study

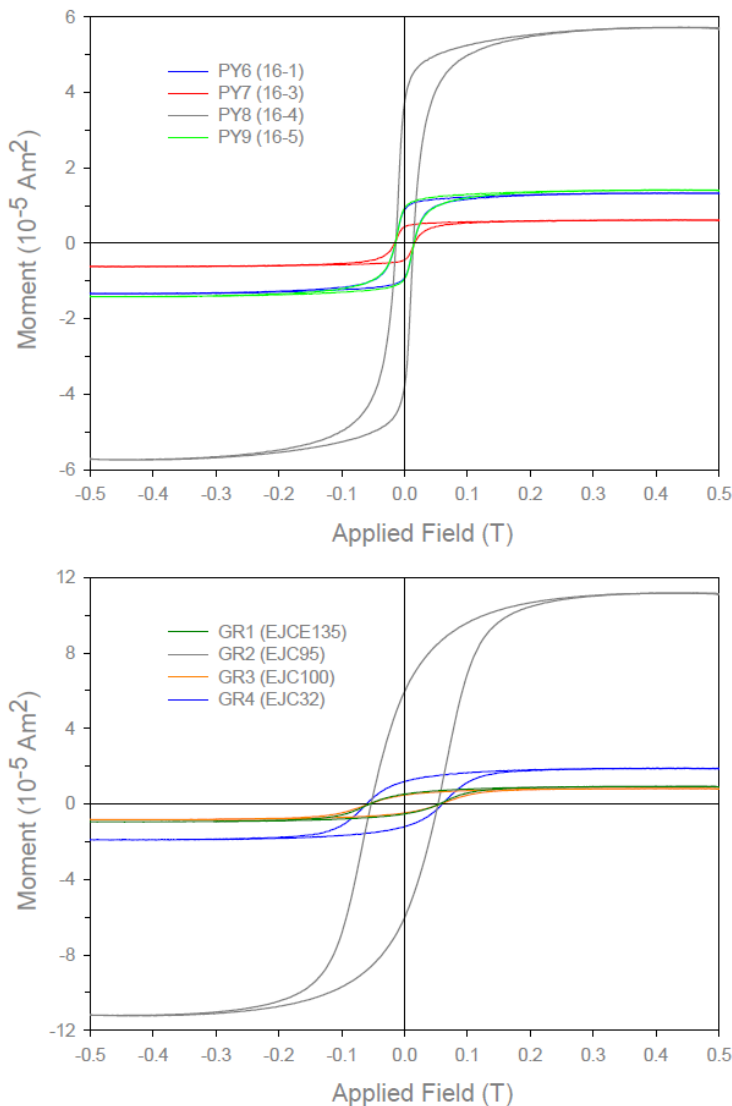
#### 1.1 Rock test samples

The sized pyrrhotite samples are dominated by monoclinic pyrrhotite, with the coarser samples like PY5 having a little quartz and chlorite (probably with associated Fe-oxide inclusions) and small amounts of goethite and lepidocrocite (probably from minor pyrrhotite oxidation, Figs. S1, S2, Table S1).

The Chalk test sample (Campanian in age) come from the CC9b horizon at Culver Cliff (Isle of Wight, UK), sampled across the foreshore within the same horizon (ca.  $\pm 10$  cm error). This bed has locally the highest concentration of bacterial magnetofossils from several Chalk samples tested (unpublished data of Hounslow). A magnetic extract from this bed is figured in Chalk sample data (sample Chalk B) shown in Hounslow & Maher (1996, fig. 7a, 10), Montgomery et al. (1998, plate 1, sample WC9b), Maher et al. (1999, plate 1.10, 1.11), Hounslow & Maher (1999a, fig. 9.9c) and Kopp & Kirschvink (2008, fig 6c-d). The XRD of the  $E_{MP}$  type extract is in Fig. S5a. The  $E_{mpt}$  type magnetic extract from CC9b contains >95% magnetofossils representing the electron dense material in TEM images with an estimated 80% of these in chains. The ARMz extraction efficiency from the sample is around 94% (Hounslow & Maher, 1996). Magnetofossils are dominantly of prismatic type with a modal size of 65 nm and modal aspect ratio of 0.63 (Fig. S3, S4b), which fall into the single domain particle size (Fig. S4a). The Chalk samples contain few detrital Fe-oxide particles (detritals are largely as inclusions in silicates) either in the  $E_{mpt}$  or  $E_{mp}$  extracts, being dominated by a residual Ti-oxide assemblage (Fig. S5a). Using the sample Chalk-A (devoid of magnetofossils, Hounslow & Maher, 1996) for comparison, silicate hosted inclusions + residual discrete detritals may contribute something like 5% of the ARMz in the CC9b-level samples.



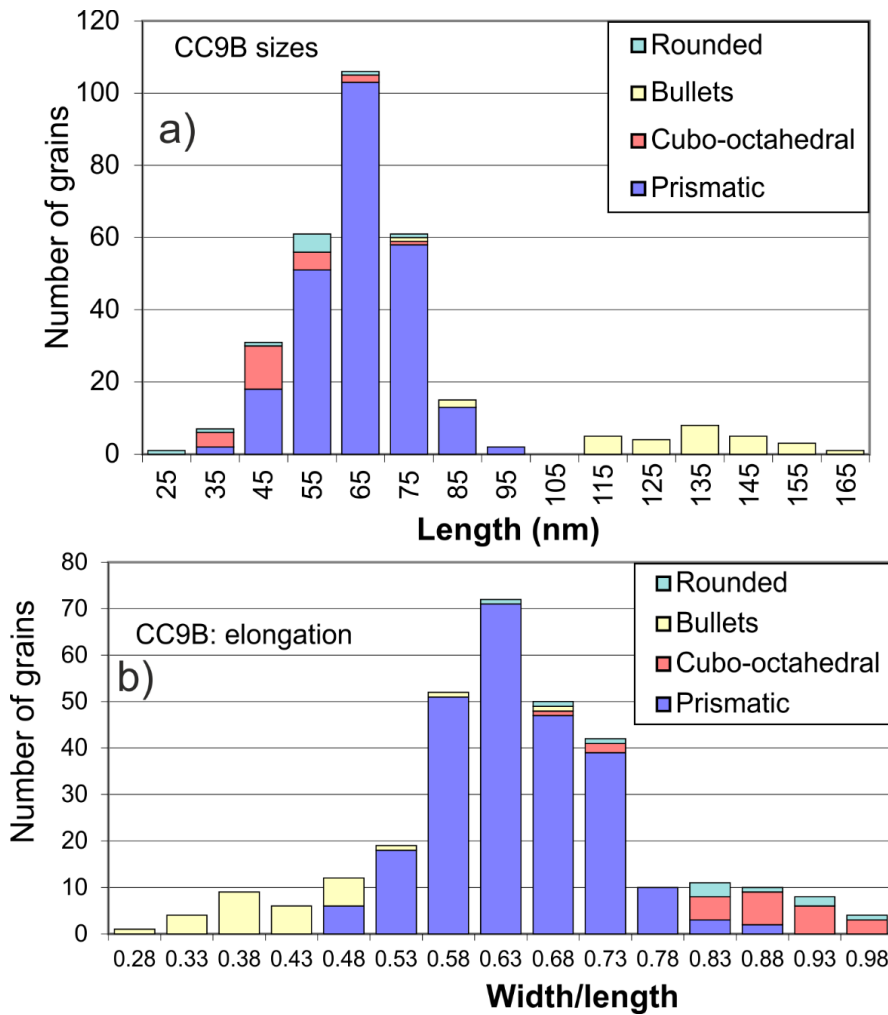
**Fig. S1 (above).** Representative XRD data for the sized natural pyrrhotite samples (5-10  $\mu\text{m}$  and 38-63  $\mu\text{m}$ , respectively) separated from river sediment by using magnetic extraction. Reference XRD minerals were identified using the powder diffraction file (PDF) from the International Centre for Diffraction Data. Minerals with PDF numbers (#) are as follows: monoclinic pyrrhotite (#29-723), quartz (#79-1906), goethite (#29-713), lepidocrocite (#44-1415), and chlorite (#29-701).



**Fig. S2 (left).** Magnetic hysteresis loops after slope correction for a) (top) pyrrhotite-bearing and b) (bottom) greigite-bearing test samples. Hysteresis parameters  $M_s$ ,  $M_{rs}$ ,  $B_c$ , and  $B_{cr}$  are shown in Table S1.

**Table S1. Hysteresis parameters (after paramagnetic slope correction) for the pyrrhotite-bearing (PY6-PY9) and greigite-bearing (GR1-GR4) test samples. The original specimen codes are shown in brackets.**

Sample	Weight (mg)	Hc (mT)	Hcr (mT)	Ms (Am <sup>2</sup> )	Mr (Am <sup>2</sup> )	Hcr/Hc	Mr/Ms
PY6 (16-1)	227.6	15.07	16.69	1.33E-05	8.99E-06	1.11	0.68
PY7 (16-3)	202.0	15.86	17.38	6.11E-06	4.37E-06	1.10	0.72
PY8 (16-4)	227.1	14.00	15.73	5.71E-05	3.76E-05	1.12	0.66
PY9 (16-5)	214.8	15.01	16.69	1.41E-05	9.36E-06	1.11	0.67
GR1 (EJCE135)	801.3	55.76	73.60	9.27E-06	5.25E-06	1.32	0.57
GR2 (EJC95)	752.8	53.31	73.92	1.12E-04	5.99E-05	1.39	0.54
GR3 (EJCE100)	596.2	57.19	79.28	8.19E-06	4.75E-06	1.39	0.58
GR4 (EJC32)	684.4	60.40	76.40	1.88E-05	1.19E-05	1.26	0.63



**Fig. S3. Magnetofossil morphology and sizes in the  $E_{\text{mpt}}$  extract from Chalk sample CC9b. A) The length and type of the magnetofossils. B) The width and length ratio versus magnetofossil morphology. The number on each of the bars represents the mid-point of the class, so for example modal column labelled length 65 nm in a) are those between 60 to 69 nm in size etc.**

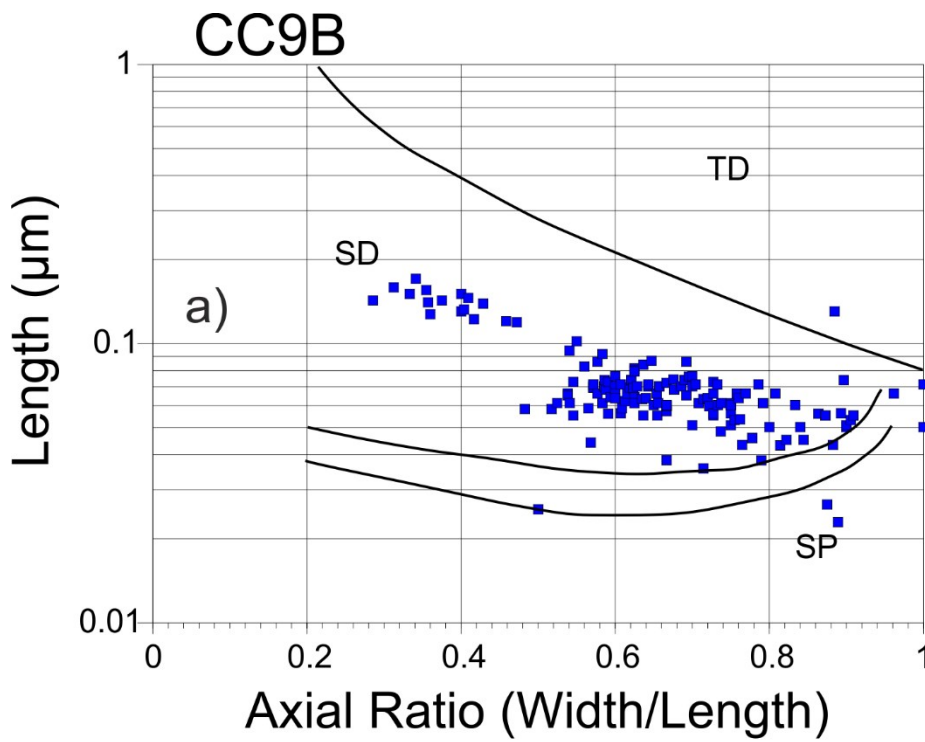
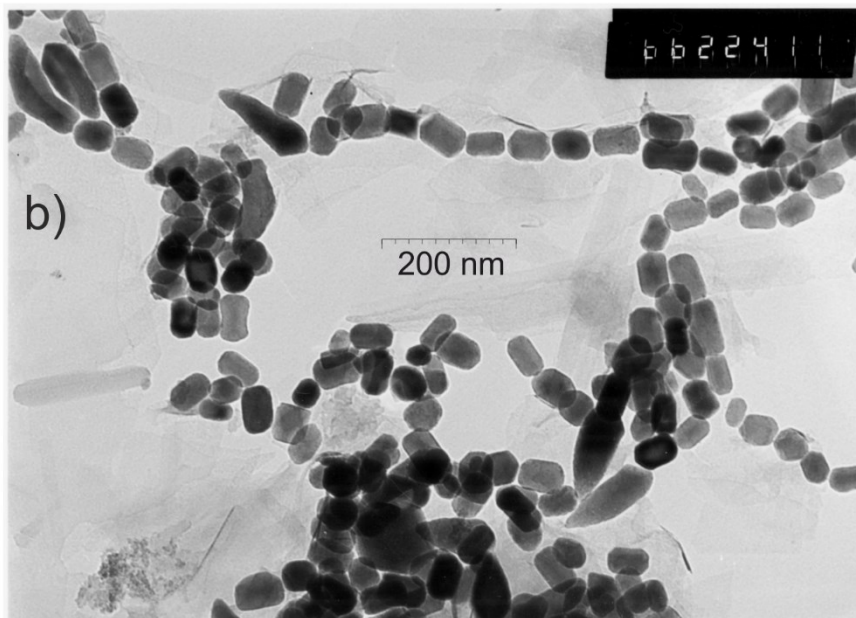


Fig. S4. A) The magnetofossil size and aspect ratio ( $n=312$ ) from Upper Chalk sample CC9b plotted onto the domain state diagram for magnetite. SD=single domain, TD=two domain, SP=superparamagnetic. B) Transmission electron micrograph of magnetofossils of various morphologies in the CC9b  $E_{\text{mpt}}$  magnetic extract.



### 1.2 Magnetic extracts

Further magnetic-related and mineralogical details about the magnetic extracts used here are detailed in Tables S2 and S3. These are taken in part from Hounslow & Maher (1996, table 3), Hounslow et al. (1995, tables 2, 3) and Hounslow & Maher (1999b, table 1, fig 12). The relationship of the simplified sample codes used here and those used in the original publications are shown in Table S4. The variation of magnetic susceptibility at low temperatures (Fig. S6) suggests the extracts may be largely dominated by oxidized forms of  $\text{Fe}_3\text{O}_4$ , due to suppression of the Verwey transition (Özdemir & Dunlop, 2010). However, this behaviour may also be due to minor Ti-substitution, so an alternative interpretation is that the lack of the Verwey transition is related to the lack of substitution-free magnetite and the dominance of Ti-substituted magnetites (Özdemir & Dunlop, 2010). The XRD-derived cell edge spacing suggests the Ti-magnetite oxidation parameter  $z$  ranges from around 0.3 (sample OR2) to around 0.8 (sample OR5) (Table S3)- this assumes the Ti-content is near zero using the data of Readman & O'Reilly (1972). If the Ti content is not minor the cell edge spacing suggests rather more oxidized Ti-magnetites ( $z > 0.7$ ). For the Lunde magnetic extracts some

data collected from thermal demagnetisation of ARMz and IRM indicates that the blocking temperatures of the remanence carrying phase has a broad range of temperatures, with this phase having maximum blocking temperatures at around that of the magnetite Curie temperature (Figs. S6, S7). The broad range of blocking temperatures may be representative of the range of the complex Mn, Ti and Cr substitution in the ferrimagnetic phases in these samples (Hounslow, 1996).

The LUD3 magnetic extract is representative of the magnetic behaviour of the inclusions in the silicates in the Lunde samples, which apart from the Mn-Ti-Cr-Fe oxides in the samples (Fig. S9a-c) are the second main ferrimagnetic source in these samples (Fig. S9d). This inference is based on the observation that the >63  $\mu\text{m}$  fraction has few discrete detrital magnetic oxides. The average SIRM of Fe-oxide inclusions in sediments sourced from various basement terrains (excluding Lewisian gneisses) and sediments in the UK is  $6 \pm 3.3 \times 10^{-5} \text{ Am}^2/\text{Kg}$  (from Hounslow & Morton, 2004)– which agrees broadly with the data from LUD3 (Table S2). An upper range for the magnetic concentration of inclusions in silicates is probably the sand sourced from Lewisian gneisses which have an inclusion-related SIRM of 45-60  $\times 10^{-5} \text{ Am}^2/\text{Kg}$  (Hounslow & Morton, 2004) which is within the SIRM range seen in the reduced intervals in ODP holes 722B (sample OR4, OR5, OR6) and 709 (MR6) (Tables S2 and S4)– although like the Lunde samples these ODP core intervals contain residual discrete Fe-T oxides. The pyrite in some ODP core extracts (Table S3) is associated with the diagenetic pyritization of the Fe-oxides, seen in the Owen Ridge extracts (OR4, OR5 and OR6). However, the impact of sulphate-reduction on the Madingley Rise sediments (extracts MR2, MR3) has not generated pyrite associated with the Fe-Ti oxides (Table S2).

The non-reduced intervals in the ODP cores are represented by extracts OR1, OR2, MR1, MR4, MR5, which have a complex assemblage of largely Fe-Ti oxides as discrete detrital phases (Fig. S9e-g), plus small amounts of bacterial magnetite, largely not in chains (see fig 7b in Hounslow & Maher, 1996; fig. 3 in Maher & Hounslow, 1999), plus Fe-oxide inclusions in many types of silicates, but especially quartz (see fig. 12 in Hounslow & Maher, 1996) and feldspar (Figs. S9f, S5c-d). The silicate-hosted Fe-oxide inclusions largely account for the large contents of quartz and feldspar in the magnetic extracts seen in XRD (Fig. S5) and the many types of minor accessory minerals seen in optical microscopy (Table S3).

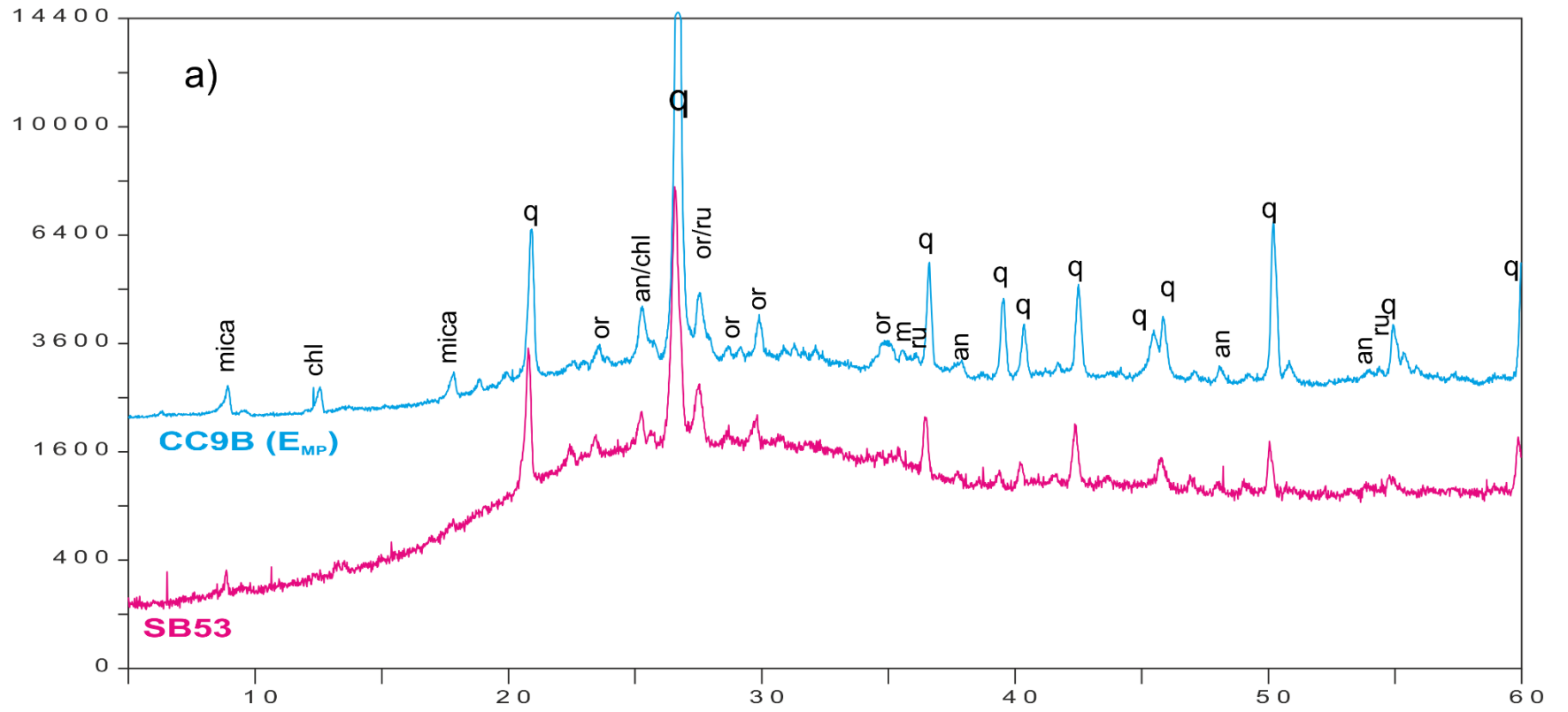
All the magnetic extracts from the Madingley Rise samples contain ‘biogenic’ baryte (see Martinez-Ruiz et al. 2019 for review), largely as 0.5-10  $\mu\text{m}$  ellipsoidal grains (Fig. S8h-l; Fig. S5e-f). It is inferred that these baryte grains contain small amounts of dispersed nanoscale magnetic inclusions. Optical microscopy failed to find any opaque inclusions in these and EDS in TEM and SEM only detected Ba and occasional Sr in these particles; and they are too rounded and electron dense in TEM to see internal structures at grain edges– implying any ferrimagnetic phases are scattered, few in total amount (below EDS detection and not near the grain surface) and too small to see (<0.5  $\mu\text{m}$ ) in optical microscopy. Biogenic baryte grains appear to form progressively in the ocean associated with organic matter films promoted by enhanced bacterial activity (Martinez-Ruiz et al., 2019). We speculate that the apparent net-ferro-magnetic behaviour of some biogenic baryte grains (see Fig S8i) is due to bacterial magnetofossils captured and dispersed through the grains, during the growth phase of the baryte particles associated with their origin in organic matter aggregates and films.

Sample code	$\chi_{LF}$	$\chi_{ARMz}$	SIRM	H'cr	%HIRM	weight%	weight% (x10 <sup>-3</sup> )	<sup>1</sup> Paramagnetic MS	Extraction efficiency (%)			
	x10 <sup>-7</sup> (m <sup>3</sup> /kg)	x10 <sup>-8</sup> (m <sup>3</sup> /kg)	x10 <sup>-5</sup> (Am <sup>2</sup> /kg)	(mT)	(%)	E <sub>mp</sub> /wt-fine	E <sub>mp</sub> /wt-fine		$\chi_{LF}$	ARM	SIRM	0- 100 mT
LUA	2.11	42	392	248	38.2	0.076	1.37	0.57	7	40	12	24
LUD	1.22	29	112.1	57	3	0.064	0.67	~0.90	54	86	92	93
LUD2	0.25	7.9	14.0	55	11.1	0.470 <sup>s</sup>	na	nd	12	~0	31	31
LUD3	0.14	6.9	8.9	62	17.3	0.180 <sup>s</sup>	na	nd	7	5	15	13
OR1*	4.1	286	716	43	6	0.105	19.10	0.36	63	70	82	84
OR2*	4.8	279	715	41	6	0.890	11.50	0.29	66	59	79	85
OR4	1.24	12.5	36.1	53	6.4	0.071	0.97	0.96	17	33	51	53
OR5	1.20	13.1	36.7	53	6.6	0.111	0.56	1.00	10	7	54	56
OR3*	1.23	15.6	59.8	52	5.6	0.192	2.68	0.89	11	15	73	75
OR6	0.96	12.1	28.7	52	6.4	0.503	0.69	0.97	0	38	55	54
MR1	3.7	1203	1026	45	2.7	0.280	13.58	~ 0	44	74	67	66
MR2	1.27	8.8	23.1	48	6.9	0.137	3.77	1.00	0	58	62	67
MR3	1.59	28.7	63.1	50	9.9	0.145	4.51	0.82	31	81	8	83
MR4*	6.0	1196	965	37	2.8	nd	nd	0.06	50	88 <sup>+</sup>	84	72
MR5*	3.5	1813	766	38	1.2	nd	nd	0.27	57	81 <sup>+</sup>	65	~ 100
MR6*	1.58	9.8	34	61	6.1	nd	nd	0.95	41	~0 <sup>+</sup>	62	51

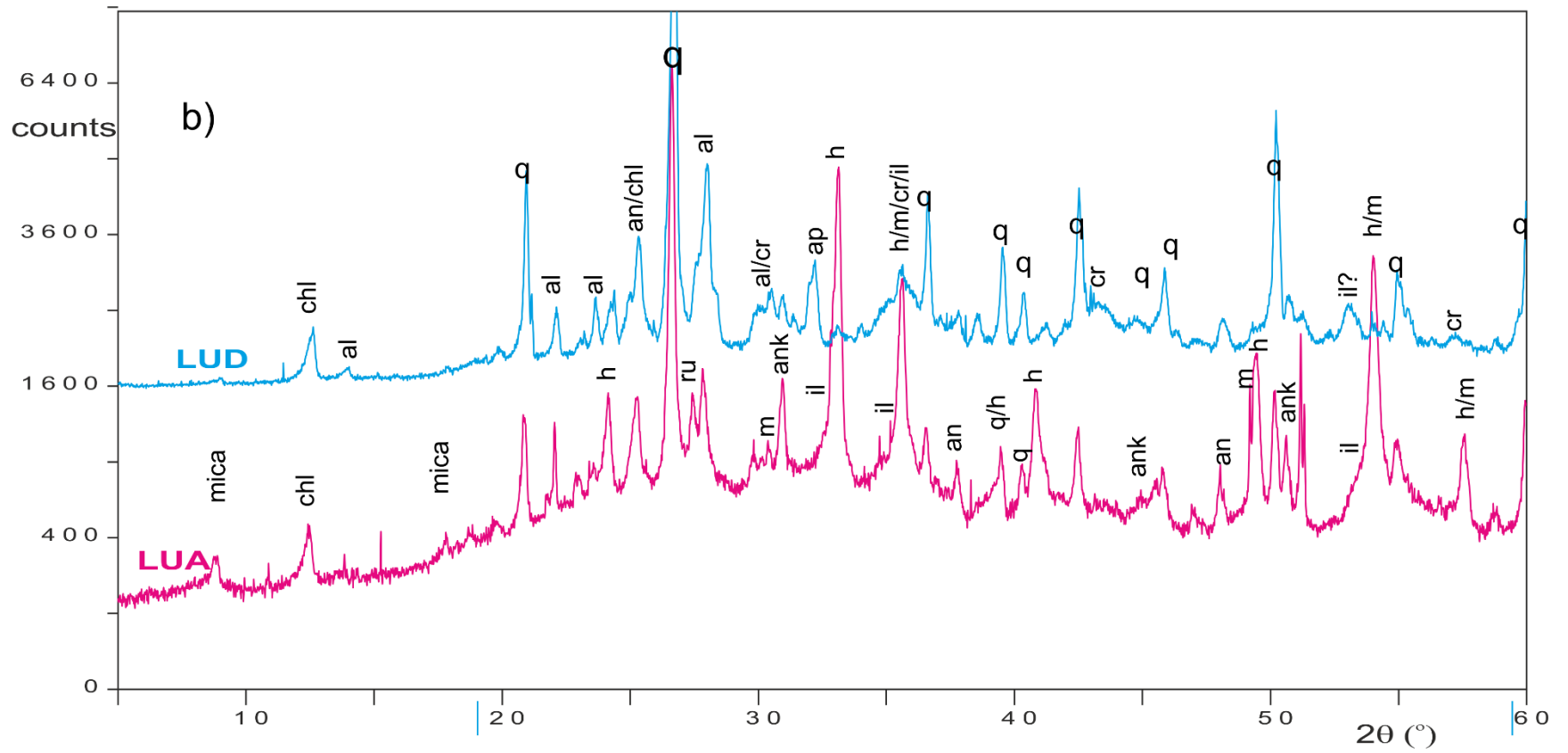
**Table S2. Summary magnetic and magnetic extraction data for the magnetic extracts. The rock magnetic data relate to the material after carbonate dissolution using the method in Hounslow & Maher (1996, 1999a), and before extraction. Most of the OR and MR samples have high carbonate contents >70%. \*= <63  $\mu$ m fraction, all others are sieved at 38  $\mu$ m. + some contamination from carbonate removal chemicals (see Hounslow & Maher (1996, 1999a) has artificially lowered this value. MS=magnetic susceptibility, SIRM= IRM at 1 T. %HIRM is % of IRM in the range 0.3- 1T.  $\chi_{LF}$  =low frequency mass magnetic susceptibility. ARMz applied at 0.08 mT, 90 mT AF field. Wt-fine= weight of fraction using the magnetic extractions (i.e., <38  $\mu$ m or <63  $\mu$ m). H'cr= DC field used to acquire 50% of the SIRM (using forwards fields). E<sub>mp</sub> and E<sub>mp</sub> are the type of extracts described by Hounslow & Maher (1996, 1999a). Nd = no data, na = not applicable. \$= E<sub>me</sub> extract of Hounslow & Maher (1996, 1999a). Paramagnetic MS is the proportion of the total MS (i.e. 1.0= all MS due to paramagnetic behaviour, 0= none or very little) in the original sample (after carbonate removal). <sup>1</sup> estimated uncertainty  $\pm$ 0.03. MR4, MR5 and MR6 are E<sub>mp</sub> type extracts used for the RRM analysis here.**

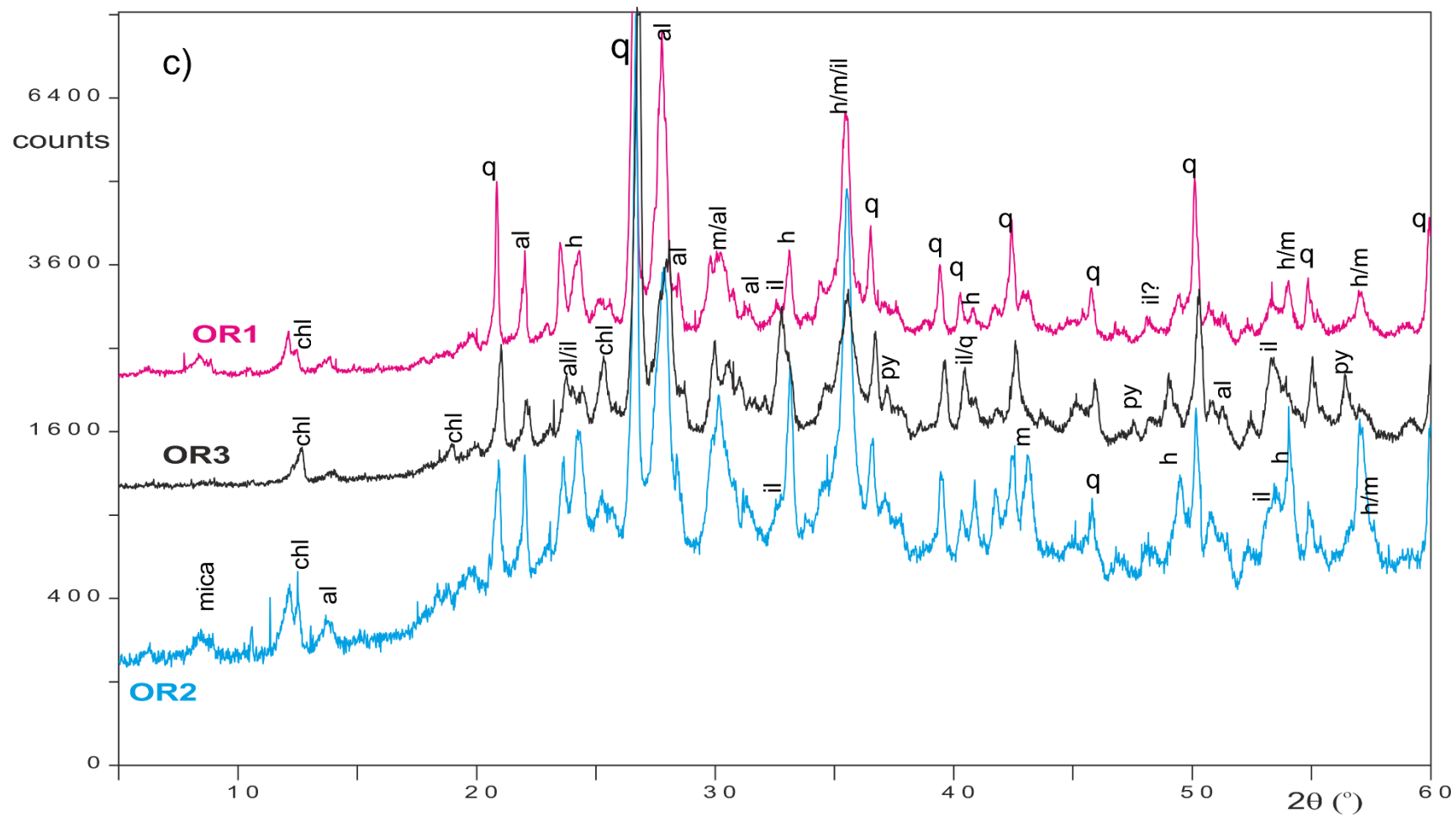
Sample code	Magnetite <sup>2</sup> Cell edge spacing (Å)	Relative percent in extract (from XRD)								Weight % in fraction (x10 <sup>-3</sup> %)		From optical microscopy	
		Quartz	Plag. feldspar	Magnetite + spinels	Haematite	Ilmenite	Pyrite	Rutile	Baryte	Magnetite	hematite	Approx. % opaques <sup>1</sup>	Minor minerals
LUA <sup>§</sup>	nd	28.8	15.1	0	37.7	16.4	0	1.9	0	0.0	318	50	zir, spin, garn
LUD <sup>§</sup>	nd	59.6	32.8	3.3	2	0	0	2.4	0	2.3	1.8	5	rut, spin, zir, ank, mica, apa
LUD2	nd	nd	nd	nd	nd	nd	nd	nd	nd	nd	nd	tr	brown and clear mica, ank
LUD3	nd	nd	nd	nd	nd	nd	nd	nd	nd	nd	nd	tr	brown mica, garn, tour, green mica, rut, zir, ank
OR1*	8.381	33.6	45.5	11.5	6.6	2.7	0	0	0	140	100	22	prx, hor, chl, spin, mic
OR2*	8.379	20.8	45.4	15.8	7.5	7.7	2.8	0	0	165	101	28	prx, hor, tour, zir, rut, mic
OR4 <sup>§</sup>	8.383	26.9	39.4	6	3.1	16.3	6.9	1.4	0	8.0	3.2	nd	mic, chl, hor, rut
OR5 <sup>§</sup>	8.369	28.6	36.3	3.6	0	16.9	8.8	5.8	0	7.1	0.0	nd	mic, chl, hor, rut
OR3*	8.392	29.3	29.1	7	5.9	16.5	12.2	0	0	17.9	18.9	30	rut, mic, hor, chl, apa, spin
OR6 <sup>§</sup>	too low	37.7	41.2	0	0	6.5	12.3	2.4	0	0.0	0.0	nd	mic, chl, hor, rut
MR1 <sup>§</sup>	nd	25.1	25.2	25.8	13.2	0	0	0	10.7	110	88.0	nd	nd
MR2 <sup>§</sup>	nd	34.4	29.5	1.7	0	17.5	0	0	17	3.5	0.0	nd	nd
MR3	nd	31.9	29.4	5.7	4.1	14	0	0	14.9	10.0	9.0	nd	nd
MR4*	nd	19.6	28.7	22.6	2.9	0	0	0	26.2	nd	nd	common	bary, apa, rut, haem
MR5*	nd	32.6	29	5.8	1	0	0	0	31.7	nd	nd	common	rut, apa, bary
MR6*	nd	27.7	36.5	0	0.4	3	0	0	32.4	nd	nd	few	apa, glass, bary

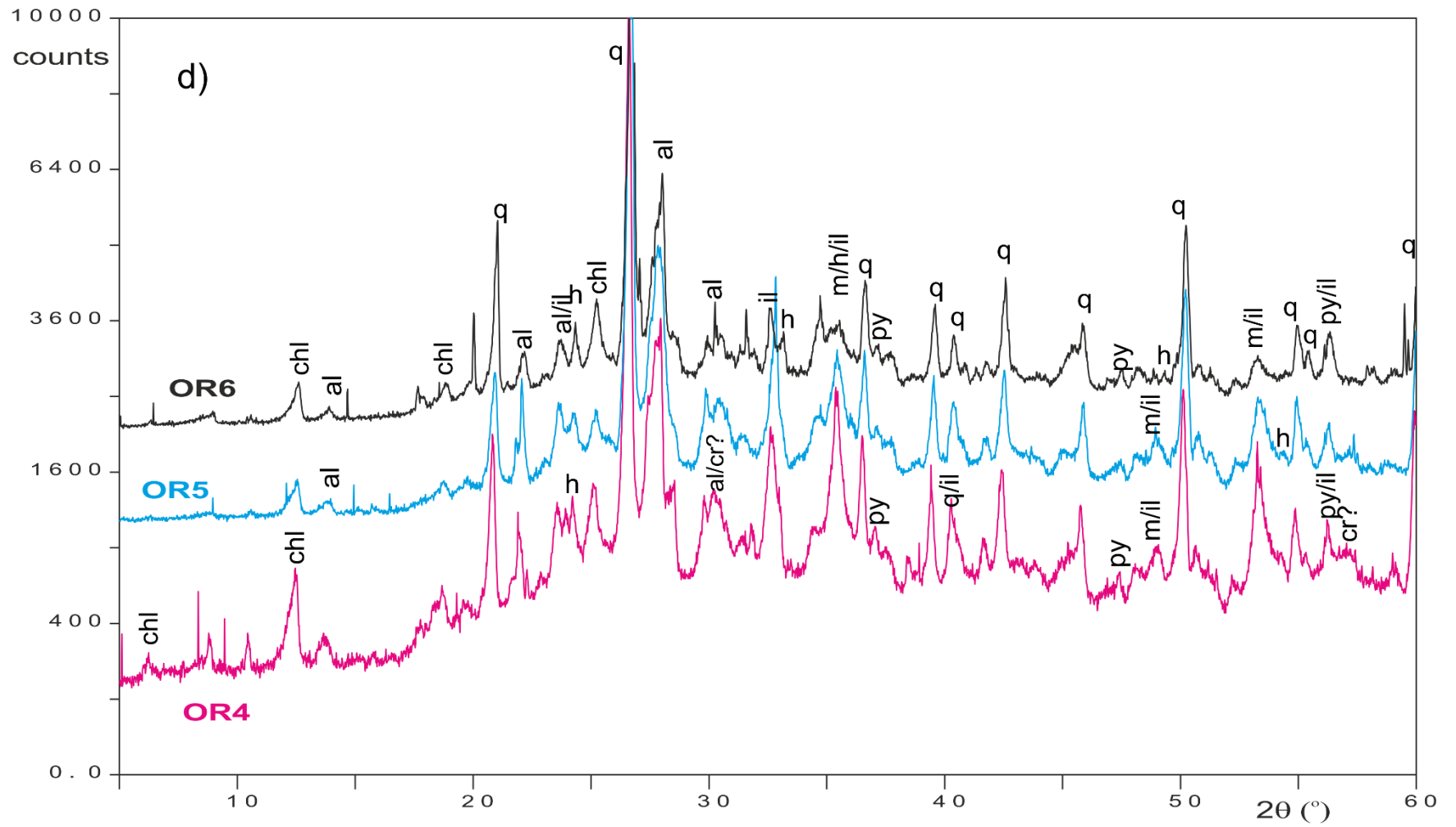
**Table S3. Summary mineralogical data for the magnetic extracts, largely based on semi-quantitative XRD, using the method detailed in Hounslow & Maher (1999a, 1999b). The weight % are those in the appropriate fraction of the sediment (\* <63 or § < 38 µm; LUD2=38-63 µm, LUD3=63-250 µm) <sup>1</sup> excludes leucoxene which may appear opaque in transmitted light but has bright internal reflections in reflected light. <sup>2</sup> using internal quartz reference peaks. nd= no data, tr=trace. Minor minerals: apa= apatite, ank= ankerite, bary=baryte, garn= garnet, rut=rutile, haem=haematite, mic= mica, chl= chlorite/chlorotoid, spin= spinels (most Cr), hor= hornblende, prx= pyroxene, zir=zircon, tour=tourmaline, glass=volcanic glass. See Fig. S5 for some of the source XRD data.**

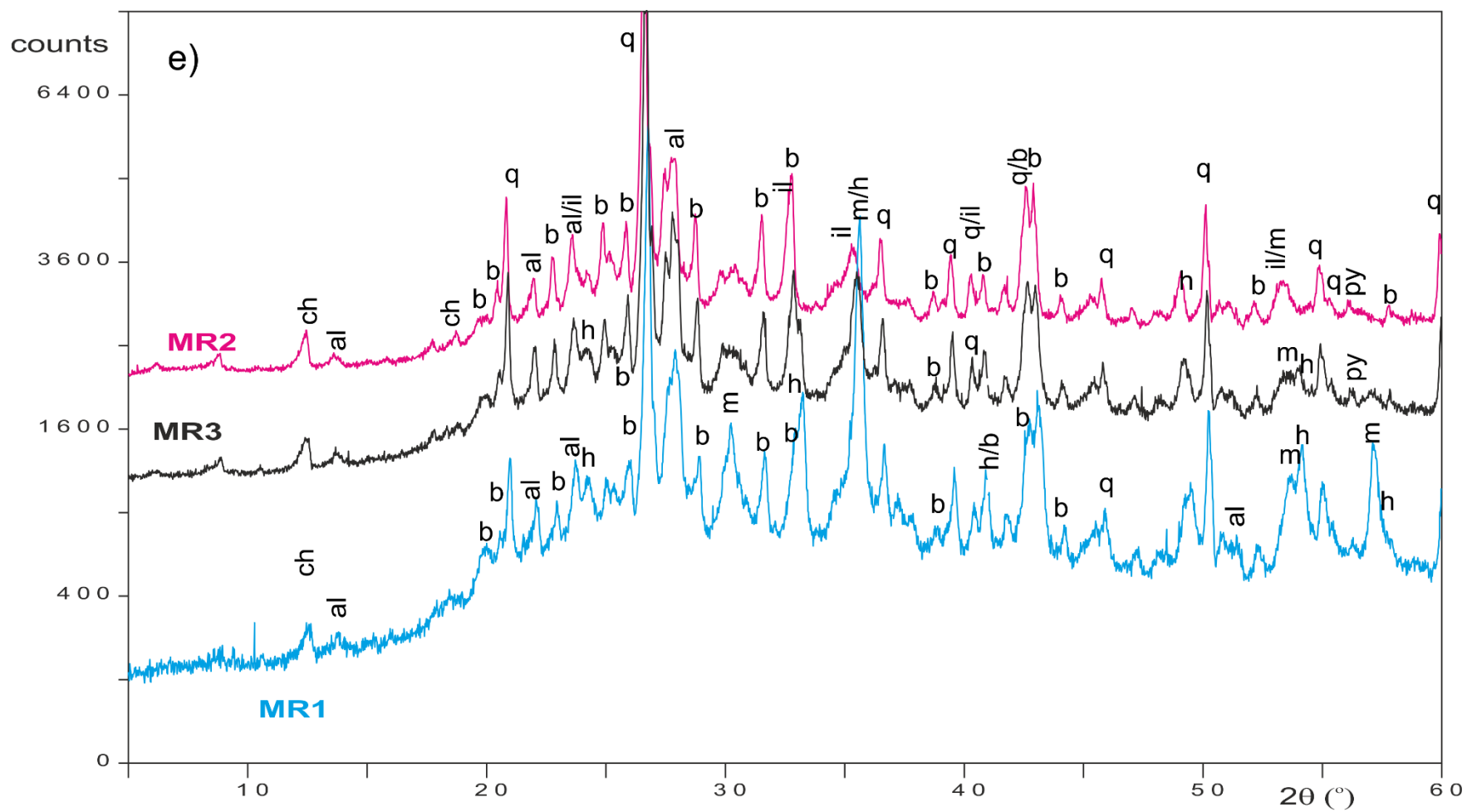












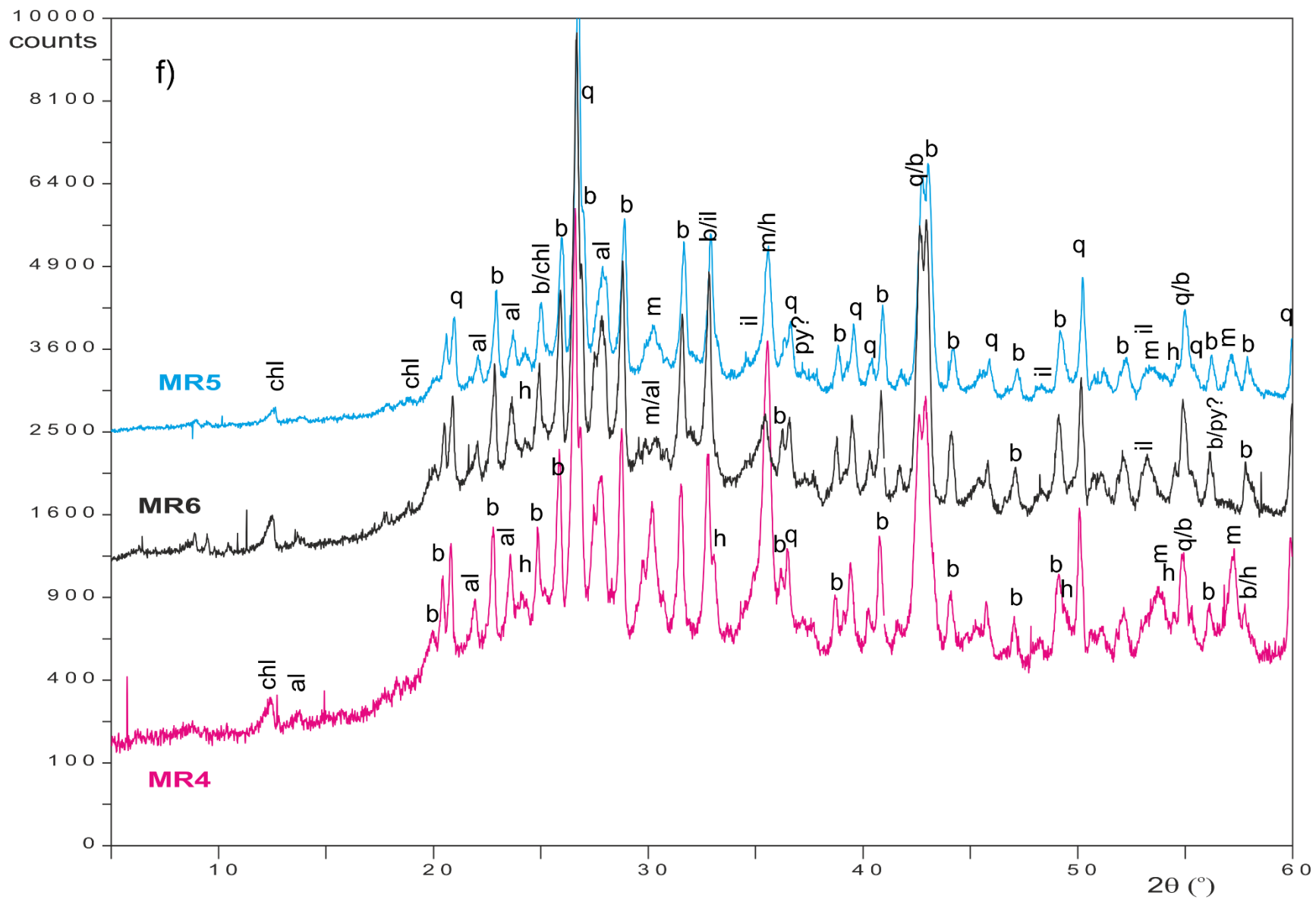
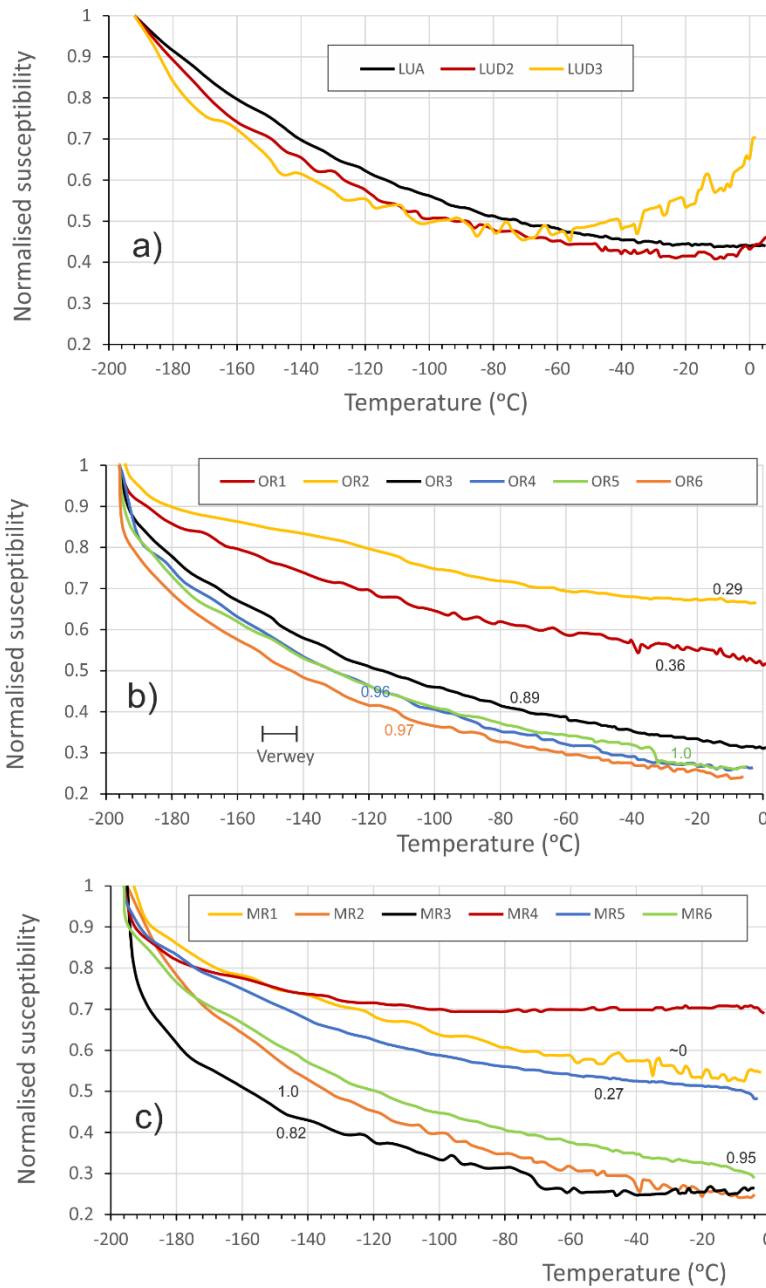


Fig. S5. X-ray diffraction data for the magnetic extracts (b-f), and two example ones from the Chalk (a). In each case these are the extracts from the E<sub>MP</sub> type extracts of Hounslow & Maher (1996), eventhough the MR4,MR5,MR6 test samples use E<sub>MPT</sub> type extracts. The peaks have been labelled with the inferred corresponding minerals: m= magnetite, h=haematite, il=ilmenite, py=pyrite, q=quartz, al=albite (or similar plagioclase feldspar), b=baryte, chl=chlorite, an=anatase, or=orthoclase or similar K-feldspar, ru=rutile, cr=chromite. /=peaks overlap here, ?= alternatively may be from one of the feldspar or phyllosilicate phases. In a) the SB53 sample (no RRM measurements here) is from Scrathell's Bay on the Isle of white, and has similar but lesser abundance of magnetofossils, most of which are in the E<sub>MPT</sub> extract. Extract samples LUD2, LUD3 have no useful XRD data. The large quartz peak has been truncated in a-e.

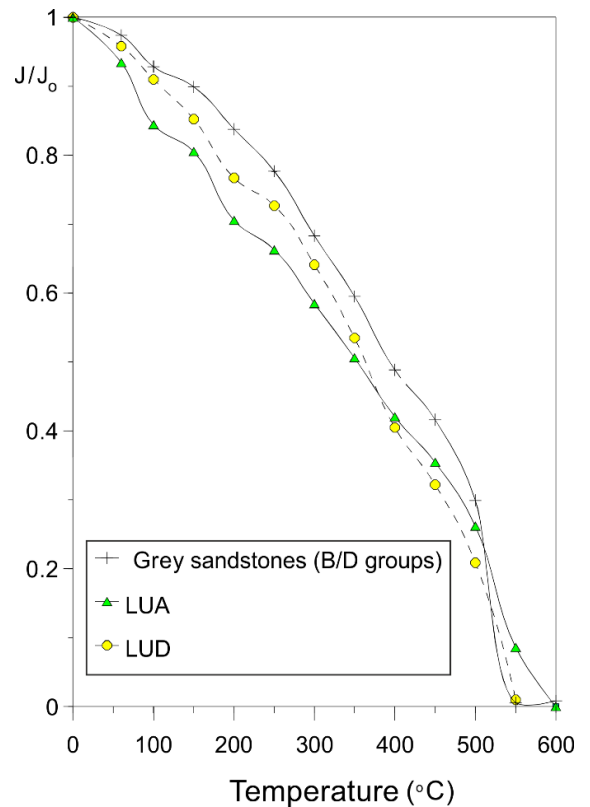
This paper	Native	Other sample codes
+Sample code	Sample code	Code & depths below sea floor
LUA	A36b <38 $\mu\text{m}$	Lunde group A <sup>1</sup>
LUD	A45a <38 $\mu\text{m}$	Lunde group D <sup>1</sup>
LUD2	A45A 38-63 $\mu\text{m}$	Lunde group D <sup>1</sup>
LUD3	A45A 63-250 $\mu\text{m}$	Lunde group D <sup>1</sup>
OR1*	L35 <63 $\mu\text{m}$	722B (0.35 m) <sup>2</sup> zone 1 oxidized
OR2*	L36 <63 $\mu\text{m}$	722B (1.25 m) <sup>2</sup> zone 1 oxidized
OR4	L58 <38 $\mu\text{m}$	722B (38.6 m) <sup>2</sup> , reduced interval, zone 5
OR5	L59 <38 $\mu\text{m}$	722B (40.6 m) <sup>2</sup> , reduced interval, zone 5
OR3*	L42 <63 $\mu\text{m}$	722B (7.8 m) <sup>2</sup> partially reduced interval, zone 3
OR6	L69-70 <38 $\mu\text{m}$	722B (59.5-61.5 m) <sup>2</sup> , reduced interval, zone 5
MR1	L23-24 <38 $\mu\text{m}$	709A (118.4-120.4 m) <sup>2</sup>
MR2	L20 (1) <38 $\mu\text{m}$	709A (112.9 m) <sup>2</sup>
MR3	L10-11 <38 $\mu\text{m}$	709A (28.0-29.7 m) <sup>2</sup>
MR4*	709 A/A	<sup>3</sup> 709A (1H-4, 5.9-8.4 m) high SIRM/ARM subset
MR5*	709C-3H-3	709C (22.8- 23.5 m), partially reduced set
MR6*	709C/D	<sup>3</sup> 709C (6H; 44.3-53.9 m), reduced set

**Table S4.** Shows the relationships of sample codes used here and in prior publications (<sup>1</sup>Hounslow et al. 1995; <sup>1</sup>Hounslow, 1996; <sup>2</sup>Hounslow & Maher, 1996; 1999b) about these extracts. <sup>3</sup> Additional data in Maher & Hounslow (1999, figs. 3, 4). MR4, MR5, MR6 are the  $E_{\text{mpt}}$  extracts, others <63  $\mu\text{m}$  or <38  $\mu\text{m}$  are  $E_{\text{mp}}$  type extracts and LUD2, LUD3 are  $E_{\text{me}}$  type extracts (Hounslow & Maher, 1999a). \* <63  $\mu\text{m}$  sediment fraction used for extraction instead of <38  $\mu\text{m}$ .

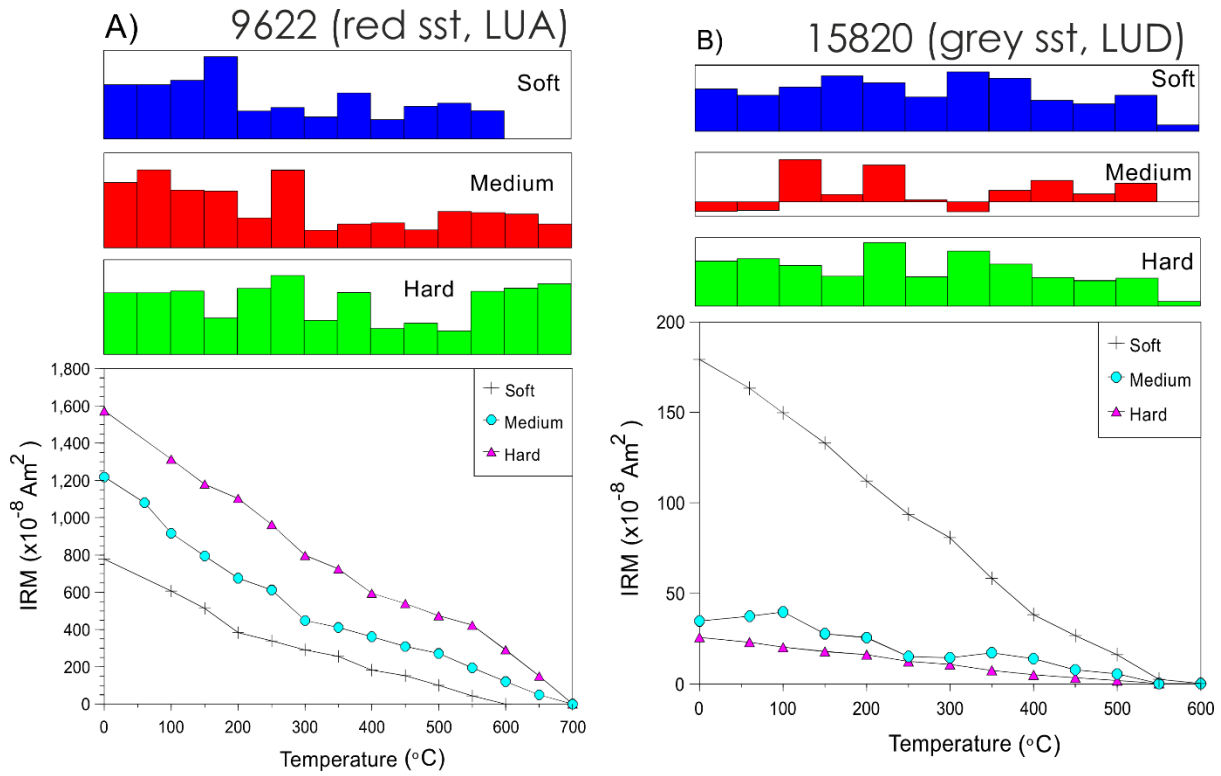


**Fig. S6. Low temperature magnetic susceptibility data for the magnetic extract samples (prior to the magnetic extraction, but after carbonate removal). Note the absence of the Verwey transition, and the variable contribution of paramagnetic minerals to the magnetic susceptibility (proportion at 0°C shown next to curve, using method in Hounslow et al., 1995). The data for LUD2 LUD3 are impacted by some thermal drift in the Bartington meter.**

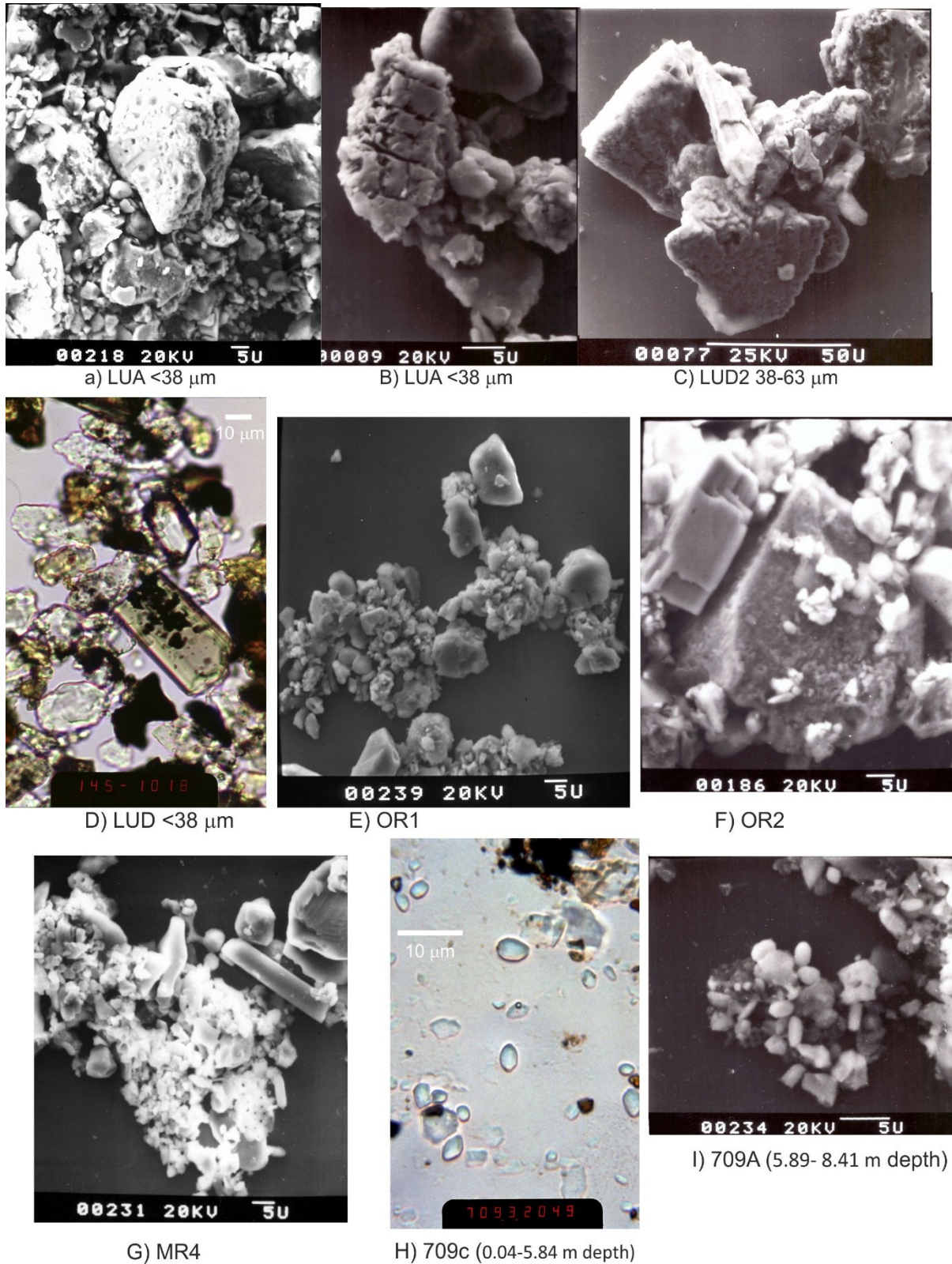
**Fig. S7. Thermal demagnetisation of the normalised ARMz ( $J/J_0$ ) in whole-rock specimens representative of the Lunde extracts. Note the broad range of blocking temperatures in the samples and the Curie temperature  $\sim 550-600^\circ\text{C}$ .**







**Fig. S8. Thermal demagnetisation of a three component IRM in samples representative of the Lunde magnetic extracts (soft= 0-0.1 T, medium= 0.1- 0.3 T; Hard= 0.3-1 T coercivity fractions). In sample 9622 (a red coloured palaeosol), note the broad range of blocking temperatures and the haematite Curie temperatures at ca.700°C for the hard and medium coercivity components. The soft coercivity component is demagnetised at 550-600°C like the ARMz shown in Fig. S7.**

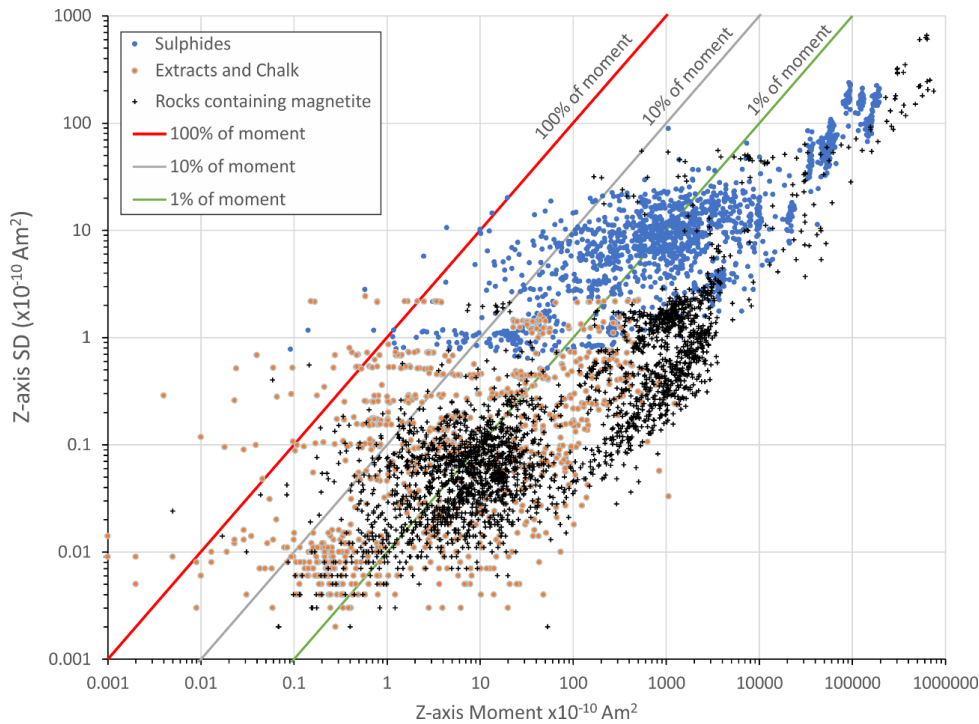


**Fig. S9. Typical features of the magnetic extract test samples. Scanning electron microscope images (a-c, e-g, i). Transmitted light optical micrographs (d, h) under oil immersion. A) Sample LUA (<38  $\mu\text{m}$  fraction) showing a cluster of magnetic particles from  $\sim 1 \mu\text{m}$  to  $\sim 35 \mu\text{m}$  in size. Large pitted grain is Fe-Ti rich and is likely an ilmenite or ilmenco-haematite detrital particle. Of the smaller particles many are angular and not rounded and may in part represent broken and fragmented relicts of original Fe-Ti-Mn-Cr particles (see Hounslow, 1996 for details).**

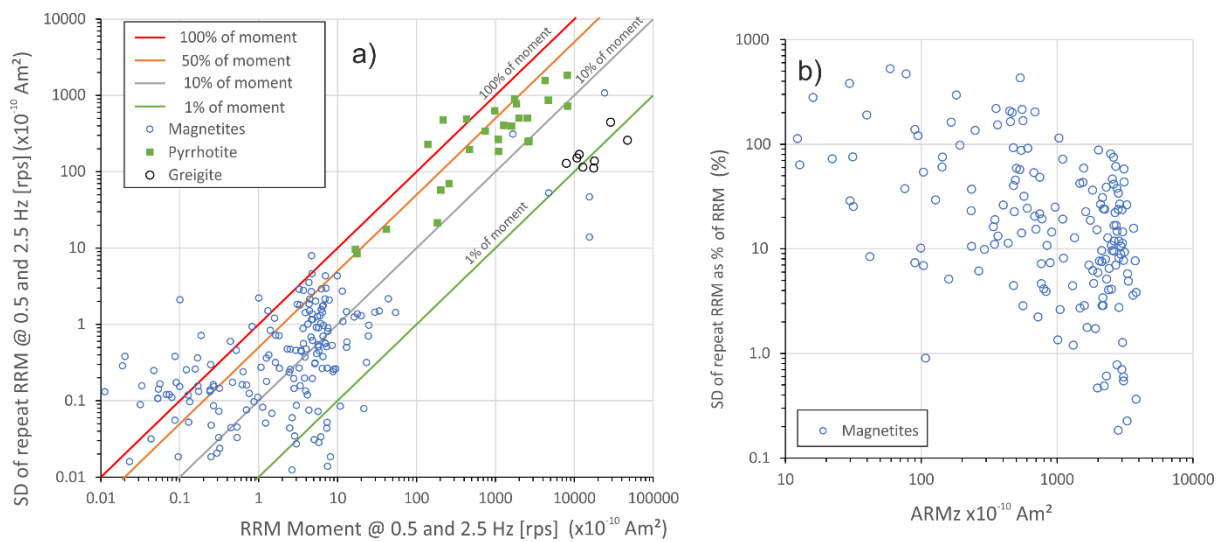
- B) Sample LUA (<38  $\mu\text{m}$  fraction) largest grain is a pitted Fe-Ti grain rich in Ti probably representing an ilmenite or Fe-rich rutile grain in which Fe has been leached from the irregular lamellae like features by diagenesis. Rounded grain at top is a zircon.
- C) LUD2 (38- 63  $\mu\text{m}$  fraction), a magnetic clump of grains, with on the left two large detrital Fe-Mn grains (Mn substituted magnetite), and in the top-right is an Fe-Ti grain (ilmenite or martite). The elongate fragment attached to top of cluster is also a Mn-Fe rich grain, a probable relict of a decomposed primary grain.
- D) LUD (<38  $\mu\text{m}$  fraction) showing a typical view of an extract with a mix of clear quartz and feldspar, some with opaque inclusions, some nearly opaque rutiles (dark ruby colour), scattered detrital opaques, and in the centre a green tourmaline grain with opaques; directly above is a smaller zircon (high relief) particle.
- E) Typical Owen Ridge extracts (OR1) showing magnetic clump of small (0.5- 10  $\mu\text{m}$ ) particles. The larger grains around the edge of the clumps are feldspars (Ca-Mg rich) and the finer-grained clusters are Fe-Ti rich (mix of Ti-magnetite, ilmenite and Ti-hematite).
- F) Sample OR2 showing two large feldspar grains. One on right has a rough surface (and richer in K), and the one top left has cleavage-like surfaces and is angular. Both also give strong Fe-Ti EDS spectra indicating they are heavily included with Fe-Ti oxides and have magnetically (?) attached brighter Fe-rich particles (probably magnetite).
- G) Typical SEM view of magnetic extract from Madingley Rise (MR4), showing a magnetic clump with satellite additional siliceous microfossil fragments. Large grain in top right is Ca-P rich and probably a biogenic phosphate fragment. In TEM the irregular and porous nano-structure of the siliceous fragments sometimes contained Fe-oxide particles (and magnetofossils), which probably accounts for their common presence in the extracts at this site (B. Maher pers comm.).
- H) x1000-Optical micrograph view of typical Madingley Rise extract from the cores (in this case composite of samples 0.04-5.84 m depth) showing a collection of largish typically ellipsoidal biogenic baryte grains (high relief, 'bluish'), and in the top a magnetic clump of Fe-oxides. Some are clumped together (bottom left), other not. The barytes contain no observable opaque grains.
- I) Magnetic clump of largely detrital grains (Madingley Rise extract, 5.89- 8.41 m) with brighter ellipsoidal biogenic baryte attached, apparently magnetically aligned to the cluster, suggesting that for reasons which are unclear, these barytes have a net ferromagnetism. Compare this to H) where many of the barytes are not clumped and therefore have a much weaker net ferromagnetism.

## 2. Further information about noise characteristics of the RRM

Since RRM is often weak in sediment samples (sometimes undetectable in limestones), it is important to evaluate the impact of noise on the measurements, to both consider the uncertainty in the RRM measurements and its impact during demagnetisation. There are two main sources of noise in the RAPID system at Lancaster, the first is the measurement noise which is suitably evaluated during the four SQUID measurements defining the moment (4 on z-SQUID for Z-axis, two on each of the x and y SQUID for X and Y axis components), and secondly noise from demagnetisation.



**Fig. S10.** The standard deviation of the raw measured moments (MSD) along the Z-axis (i.e., that which the RRM and ARMz are acquired along) versus the magnitude of the measured moment (after removal of holder+ rod moment). This clearly shows that overall, the measurements have MSD less than 10% of the net moment ( $\sim 1\%$  on average) when the moment is  $< 10000 \times 10^{-10} \text{ Am}^2$  and around 0.1% at moment  $> 10000 \times 10^{-10} \text{ Am}^2$ . The measurement floor is around  $10^{-12} \text{ Am}^2$  on the Lancaster RAPID on a good day with a well cleaned holder.

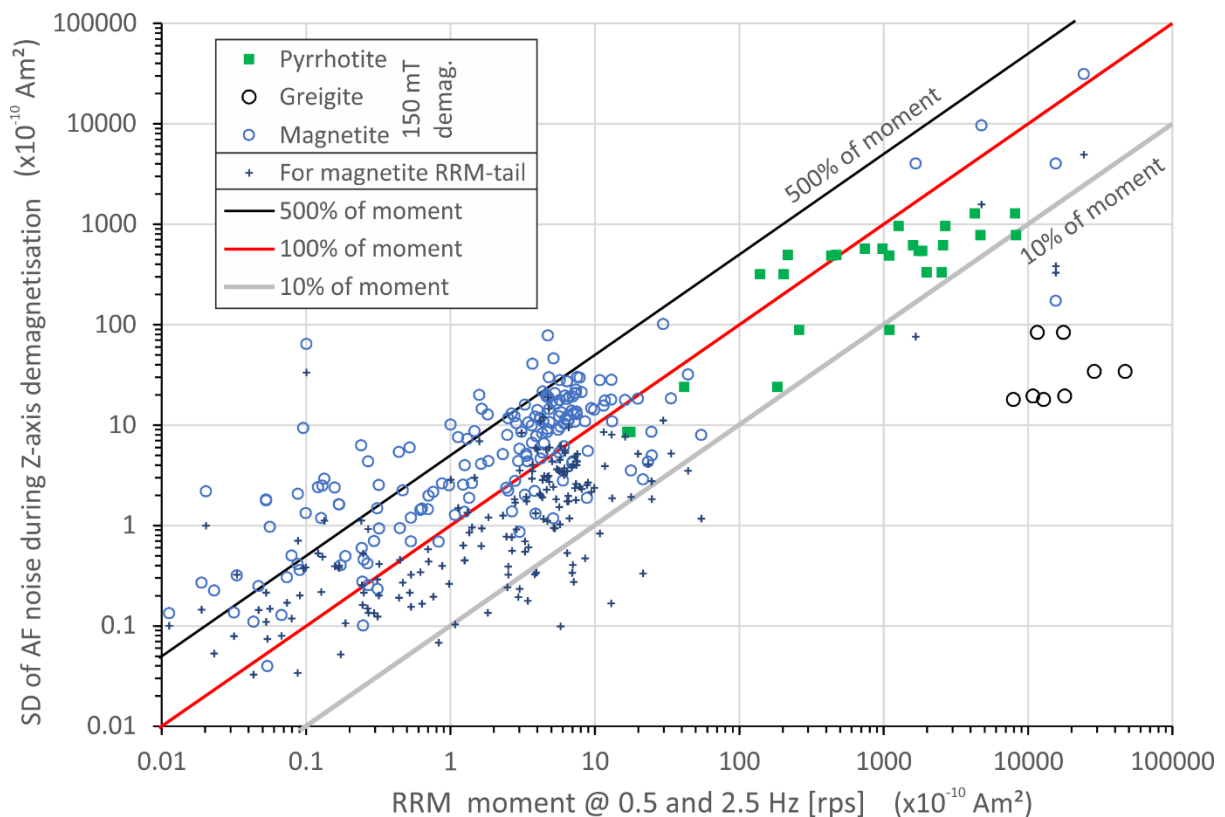


**Fig. S11.** a) The standard deviation (RSD) of the limited number of RRM repeats. RSD shown on both y-axes is the moment in a), and as percentage of RRM in b). The RSD is determined from two repeats for most, but three repeats for some. Compared to the MSD in Fig. S10 the RSD is considerably enhanced at some 1 to 100% of the RRM moment. The RSD of magnetite-bearing samples shown in b) has a rather weak relationship to ARMz moment, which suggests some of the scatter in RSD is inversely related to abundance of magnetic minerals. The four lines shown in a) are the percentages of RSD with respect to the RRM moment.



The measurement standard deviation (MSD) is related to the samples net moment (Fig. S10). The subtraction of the holder and quartz-silica rod moment (with attendant MSD) from the measurement moment will also impact on the final RRM uncertainty, so it might be expected that at moments close to the holder+rod value the uncertainty of the RRM will increase, which is seen in the increased uncertainty in RRM moments at less than ca.  $0.2 \times 10^{-10} \text{ Am}^2$  (Fig. S11A; the quartz-silica rod has a demagnetised NRM of ca.  $0.15 \pm 0.4 \times 10^{-10} \text{ Am}^2$ ). This is probably also the reason for the increased % uncertainty shown in Fig. S11b for ARMz less than ca.  $1000 \times 10^{-10} \text{ Am}^2$ . This suggests that RRM moments below ca.  $1 \times 10^{-10} \text{ Am}^2$  benefit greatly from repeat RRM measurements for enhanced quantification.

However, the scatter in RSD in those samples with RRM moment above  $1 \times 10^{-10} \text{ Am}^2$  is rather larger than might be expected from combining MSD alone, and so there is another, larger source of noise—perhaps generated after the RRM at maximum field, and during sample rotation in the decay stage of the AF field (i.e. RRM-noise). A part of this variability in RSD may be related to the inadequate (only 2 to 3) number of repeats to characterise RSD. Perhaps part may be related to anisotropy in the sample if it moves during RRM acquisitions? Notably, the pyrrhotite samples have a much larger RSD than the greigite and the magnetite bearing rocks with high-moments. The source of the additional noise contributing to RSD is not understood, but seems to be a problem for some other published studies with a pyrrhotite NRM and AF demagnetisation. For those pyrrhotite samples with RRM  $< 100 \times 10^{-10} \text{ Am}^2$  the RSD was not related to either the MDF of ARMz, or to particle size as measured proxied by the ARMz/SIRM ratio. The RRM stability is also similar for both greigite and pyrrhotite (main text Fig. 12a), suggesting this additional RRM-noise is not controlled by RRM stability.



**Fig. S12.** Standard deviation of the sample remanence after being subjected to AF demagnetisation at 150 mT along the Z-axis (ZSD), versus the RRM moment acquired at 100 mT. This is shown for the three types of mineralogy (top part of legend). Also shown as crosses are estimates of the SD for the moment of the tail of the RRM demagnetisation curve (RRM-tail)— this data is largely from the RRM demagnetisation method-1 (see main

text), using data points up to around 100 mT demagnetisation fields. Note the lower RRM-tail SD values than the ZSD indicating high field-strength-dependency of the Z-axis demagnetiser noise. The three lines are the percentages of SD with respect to the RRM moment.

With conventional ARMz demagnetisation it is commonly assumed that AF demagnetisation at or above the AF field used in the ARMz production provides a suitable 'zero' or baseline level for determining the ARMz magnitude (also accounting for any residual remanence). This procedure was adopted here for the ARMz and  $ARM_{ROT}$ , since these remanences are large compared to the uncertainty in determining the 'zero' level. However, the small magnitude of the RRM along with the often-large standard deviation at 150 mT demagnetisation (ZSD; Fig. S12), precluded the use of this simple procedure for RRM demagnetisation (not to mention the additional GRMz generated –see main text). With ZSD typically 1 to 5-times larger than the resulting RRM for magnetite-bearing rocks (blue circles in Fig. S12). A better approximate method is to utilise several points on the tail of the demagnetisation decay curve to estimate the 'zero' level. This was used for assessing the RRM demagnetisation curves and the resulting determination of MDF of RRM. As explained in the main text, method-1 does not isolate the GRMz, whereas method-2 allows GRMz isolation from the RRM decay curve, but either method requires a base-line zero level to be determined to locate the magnitude of RRM or GRMz.

In summary, the uncertainty assessment suggests:

- RRM with magnitude less than about  $0.5 - 1 \times 10^{-10} \text{ Am}^2$  benefit from repeat measurements.
- Uncertainty of the RRM measurements is larger than the sum of the measurement uncertainties, indicating additional sample and/or mineralogy related RRM-noise. This probably makes RRM measurements with respect to rotation rate, a better choice for use in mineral magnetic identification, since RRM measurements at single  $\omega$  may be subject to additional unforeseen uncertainty, which can be identified by additional RRM at adjacent  $\omega$  values. Alternatively repeats at a single rotation rate.
- Conventional zero-level determination seems suitable for determining ARMz and  $ARM_{ROT}$  in the presence of a residual remaining remanence after demagnetisation. However, this procedure is unsuitable to use for RRM demagnetisation, and the tail of the RRM demagnetisation curve should be used at fields beyond the maximum AF field used to generate the RRM decay curve. Method-2 for RRM demagnetisation should be used.
- RRM acquisition with AF field is probably the preferred method for assessing RRM stability, and since RRM can be weak in many sediments it may be best to adhere to a fixed AF field for acquisition (e.g. 40 mT), since it seems more reliable to determine this with repeat measurements, than to determine the median acquisition field from multiple, probably noisier single measurements at various AF fields (see Section 3.3).

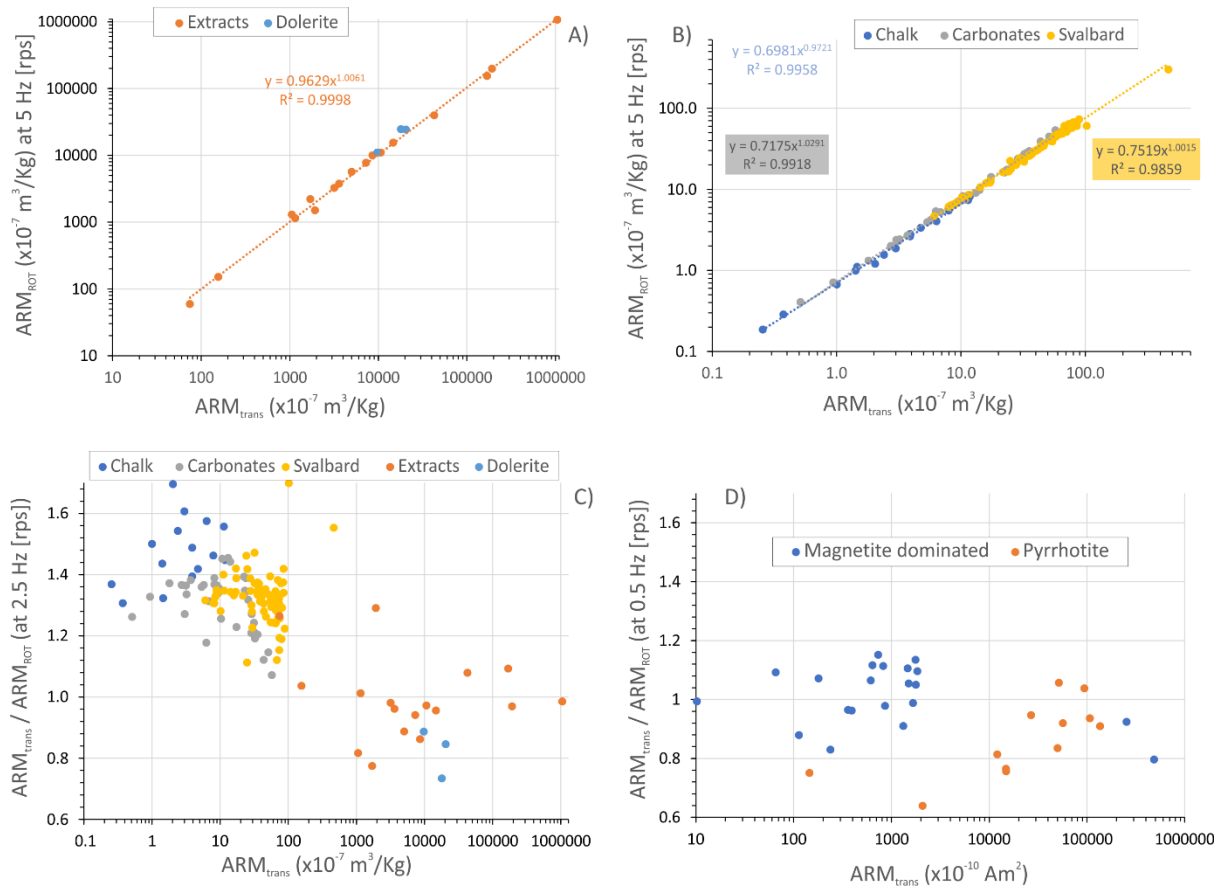
### 3. Additional information about RRM-related magnetic properties

#### 3.1 Estimating the rotational ARM from a static transverse ARM

For kinds of equipment other than the RAPID it may not be possible to determine a rotational ARM, but only a transverse static ARM. Hence, a limited investigation was undertaken on recovering the ARM at a rotation rate of zero ( $ARM_{trans}$ ), using the same configuration as that when acquiring the  $ARM_{ROT}$ . The difference is that  $ARM_{trans}$  is a single measurement obtained by using the same base-line zero level as that for the  $ARM_{ROT}$ . This was done for a large set of the samples (magnetic extracts, dolerites, various rock samples, Fig S13a,b,c) at  $\omega=2.5 \text{ Hz}$  [rps], and for a small set (various stronger magnetite dominated samples and extracts, and pyrrhotite samples;) at  $\omega=0.5 \text{ Hz}$  [rps] (Fig. S13d). Generally, there is a good correspondence between  $ARM_{trans}$  and  $ARM_{ROT}$ , but the

relationships vary somewhat between sample sets, as shown by the equations for the regression fitted lines in Figure S13(a, b). Figures S13(c, d), show that if using  $ARM_{trans}$  to estimate the  $ARM_{ROT}$ , this is subject to an uncertainty of the order of  $\pm 10\%$  in  $ARM_{ROT}$ , and also a dependence on the magnitude of the ARM (Fig. S13c,d). As such the regression equation here for the magnetic extracts (Fig. S13a) may be best for strongly magnetic samples, and one of the rock-based ones for weaker samples (Fig S13b).

**Fig. S13.** The relationship of the transverse static ARM ( $ARM_{trans}$ ) to the  $ARM_{ROT}$  values at rotation rates of  $0.5 \text{ revs}^{-1}$  (panel D) and  $2.5 \text{ Hz}$  [rps] (panels A, B and C;  $T_H = 92 \text{ ms}$ ). The regression equations in A and B are colour coded according to the sample set used (chalk= bacterial magnetite set, carbonates= mixed magnetite and a possible(?) pyrrhotite-like phase (not in main paper) with both negative and positive RRM's, Svalbard=sample set from Svalbard in main paper, extracts= set as in this work). C and D the ratio of  $ARM_{trans} / ARM_{ROT}$  which shows the within-sample set variability, and what appears to be a magnetic mineral abundance control on this ratio (i.e. smaller  $ARM_{trans} / ARM_{ROT}$  for larger ARM). The pyrrhotite dataset is particularly noisy (see later also).



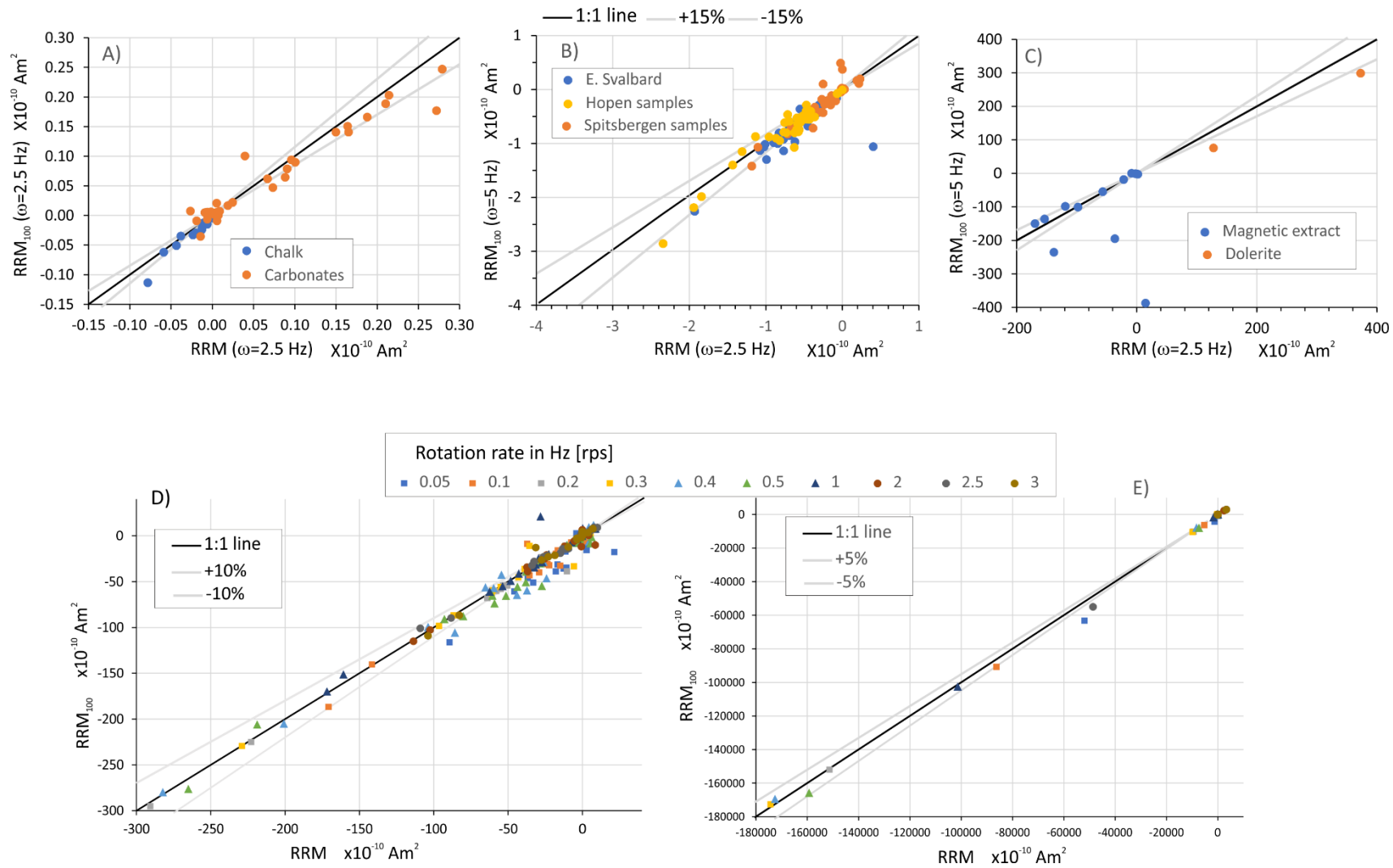
### 3.2 The RRM acquired in a DC bias field of $100 \mu\text{T}$

A limited investigation was undertaken on possibly utilising the RRM which was generated under those conditions of applied DC bias field used for determination of the  $ARM_{ROT}$  value. This is the  $RRM_{100}$  value (i.e. acquired during DC bias field of  $100 \mu\text{T}$ ). This was obtained by subtraction of the rotational ARM values obtained for the  $+\omega$  and  $-\omega$  rotation states.  $RRM_{100}$  are broadly consistent with the RRM obtained with no DC bias field but display larger imposed noise since  $RRM_{100}$  is superimposed on a much larger  $ARM_{ROT}$  (Fig. S13). This also the case for differing peak AF hold times (Fig. S20). In magnetite-dominated samples the  $RRM_{100}$  noise is of the order of 10-15% of the RRM at  $|RRM|$  less than ca.  $300 \times 10^{-10} \text{ Am}^2$  (Fig. S14B,C,d), but larger than this at  $< 0.05 \times 10^{-10} \text{ Am}^2$  (Fig.

S14a). At  $>80,000 \times 10^{-10} \text{ Am}^2$  the noise may be less than ca. 5% (Fig. S14d). Greigite samples as here display comparable scatter in values (Fig. S15c). This is in contrast to pyrrhotite samples which display greatly enhanced noise in both RRM and  $\text{RRM}_{100}$  (Fig. S15a,b).

Overall  $\text{RRM}_{100}$  is of little use for additional evaluation of the RRM process, although the data suggests that  $\text{RRM}_{100}$  may most often underestimate the RRM (see Fig. S14a,d), for a reason that is unclear.





**Fig. S14.** Data for the RRM acquired under zero DC bias and that acquired during a DC bias field of 100 mT (i.e.  $RRM_{100}$ ) at  $\omega=2.5 \text{ Hz}$  [rps] (A,B, C) and with respect to rotation rate in D and E. The datasets have been separated in the graphs due to the large variation in RRM intensity,

to allow the spread of data points to be seen ( $T_H = 92 \text{ ms}$ ).

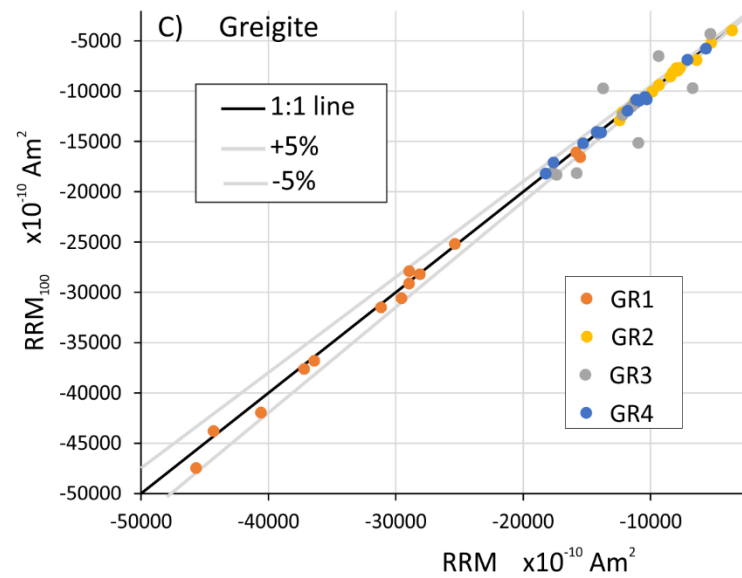
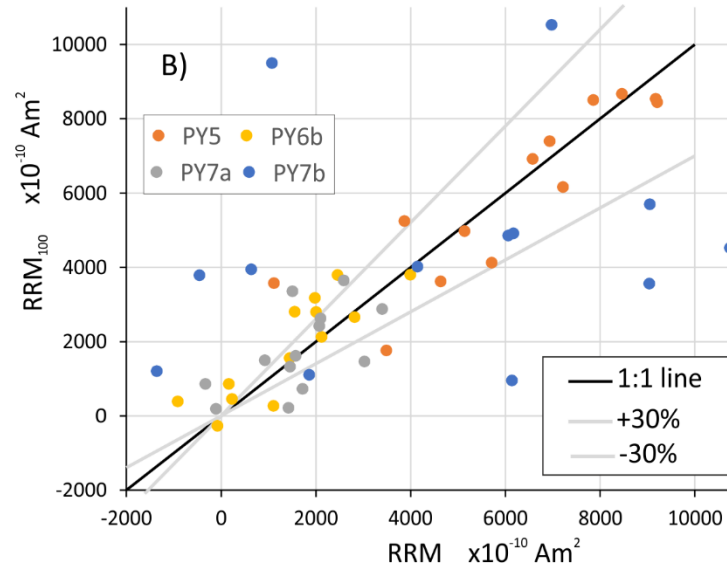
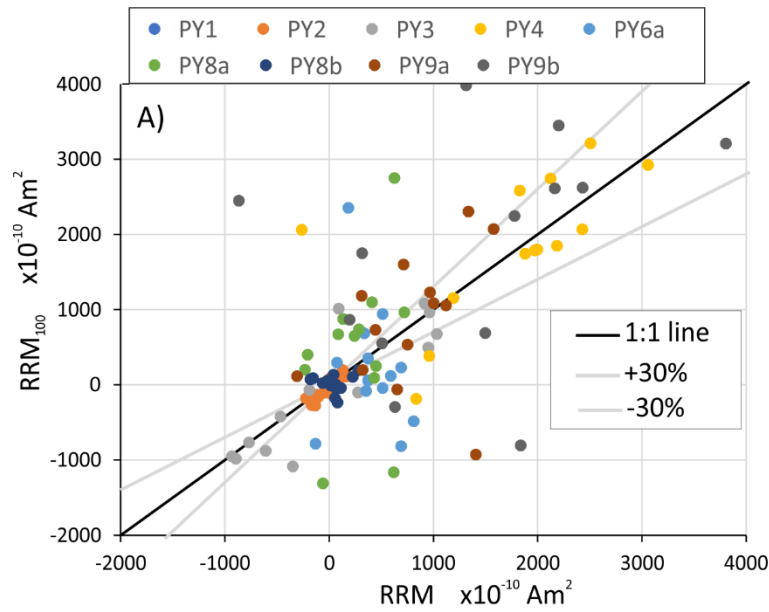


Fig. S15. Data for the RRM of pyrrhotite and greigite acquired under zero DC bias and that acquired with a DC bias field of 100 mT ( $RRM_{100}$ ) at rotation rates  $\omega$  of 0.05 to 3 Hz [rps] (sample codes as in this work,  $T_H = 92$  ms). The dramatically noisier response of pyrrhotite apparent in A and B with  $\pm 30\%$  error bands on RRM, but with  $RRM_{100}$  uncertainty for greigite comparable to that obtained with magnetite at similar RRM intensities in C (see Fig. S14e).

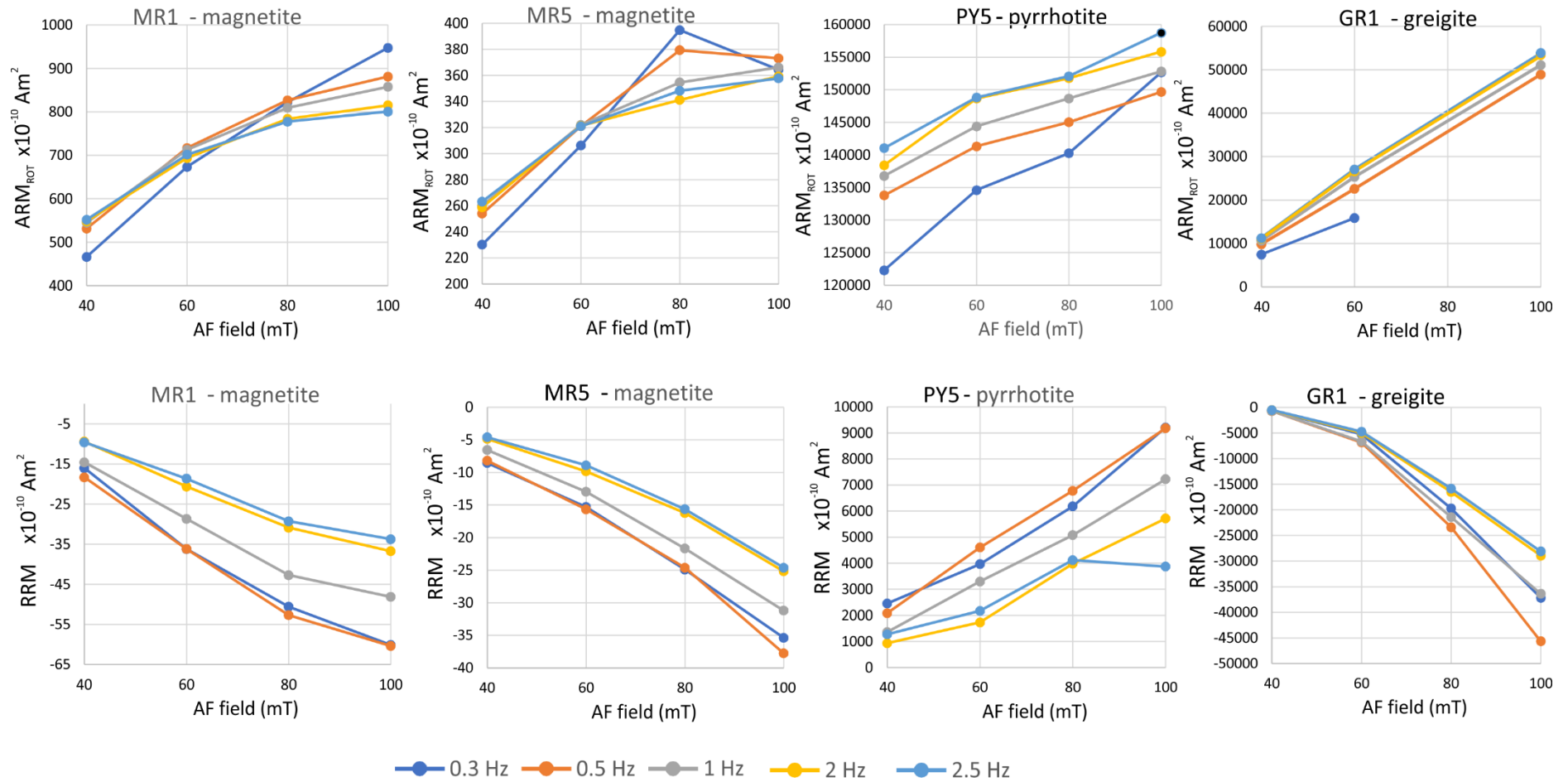
### 3.3 The RRM and ARM<sub>ROT</sub> acquisition process with respect to the AF field intensity.

In comparison to the test magnetite samples used by Potter and Stephenson (1986) our samples display a limited approach to saturation of the ARM<sub>ROT</sub>- theirs were nearly saturated by 50 mT (although their data were produced as a tumbling ARM<sub>ROT</sub>). RRM acquisition is not saturated at 100 mT (Fig. S16), which is usually the case for samples previously studied for RRM at 80-100 mT peak field (Wilson & Lomax, 1972; Potter and Stephenson 1986; Snowball, 1997).

To compare the variation of RRM, ARM<sub>ROT</sub> and Bg with respect to peak AF field and rotation rate, we normalise the values to that at  $\omega=0.5$  Hz [rps] (Fig. S17). These indicate that the RRM shows similar changes with  $\omega$ , with the most similar being samples MR1 and PY5 (over the range of rotation rates). The larger absolute normalised values tend to be consistently the lower AF fields (top row in Fig. S17). The relative changes with peak AF are also similar for the Bg values (bottom row of Fig. S17).

The normalised ARM<sub>ROT</sub> values for the magnetite-bearing samples show the largest variability with respect to rotation rates, with the peak AF at  $\leq 80$  mT having negative slopes with respect to  $\omega$ , and that at 100 mT a small positive slope. This variability in normalised ARM<sub>ROT</sub> must be in part be related to the differing activated coercivity fractions carrying the ARM<sub>ROT</sub>. This demonstrates that standardising the maximum AF field for ARM<sub>ROT</sub> acquisition is an important issue, so that comparison of Bg values obtained with differing peak AF fields will least impact the shapes of the Bg to  $\omega$  relationships (although the relative changes in ARM<sub>ROT</sub> are mostly within  $\pm 7\%$  of the value at 0.5 Hz [rps]).

The fact that the percentage of RRM and ARM<sub>ROT</sub> varies between samples (Fig. S18) are useful additional properties comparable to the stability of the RRM and ARM<sub>ROT</sub> to AF demagnetisation. This suggests that instead of demagnetisation of the RRM as a measure of RRM stability, the proportion of RRM acquired between 40-100 mT (or proportion acquired at 40 mT), or the median acquisition field of the RRM may be better alternatives to RRM demagnetisation, since they will be free of GRMz which complicates the RRM demagnetisation. That part between 0.5-1.5 Hz [rps] appears relatively consistent (Fig. S18) and a rotation rate in this interval may be a useful 'standard' rotation rate to use for both RRM and ARM<sub>ROT</sub> (if using a hold time of 92 ms).



**Fig. S16.** Variation of RRM and ARM<sub>ROT</sub> acquisition with AF field intensity for four test samples. Broadly the acquisition processes are near linear for fields between 40 and 100 mT, with a rather closer approach to saturation of the ARM<sub>ROT</sub> for the two magnetite-bearing samples (T<sub>H</sub>= 92 ms). Data is missing for some steps due to sample-drops during rotation.

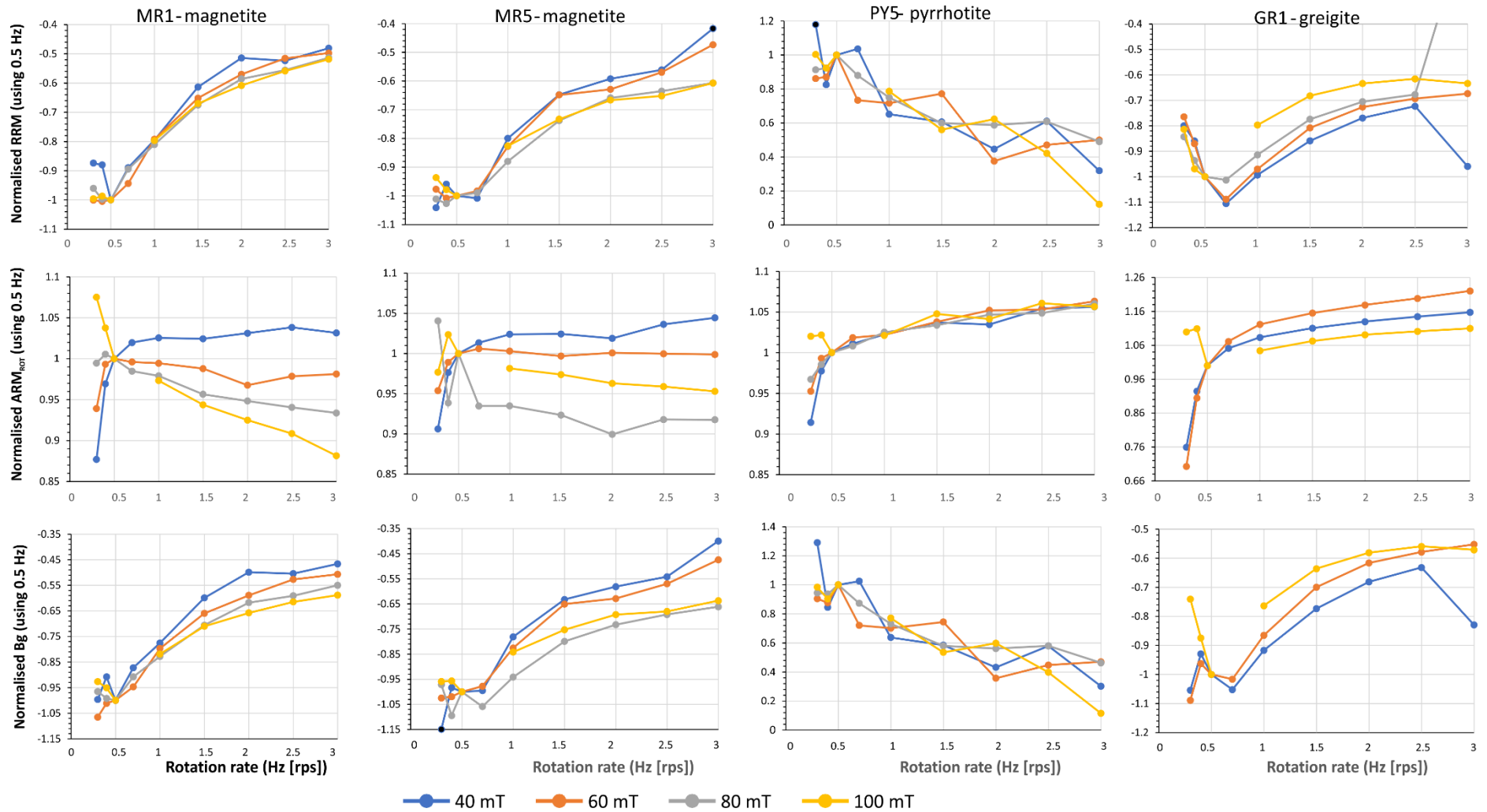
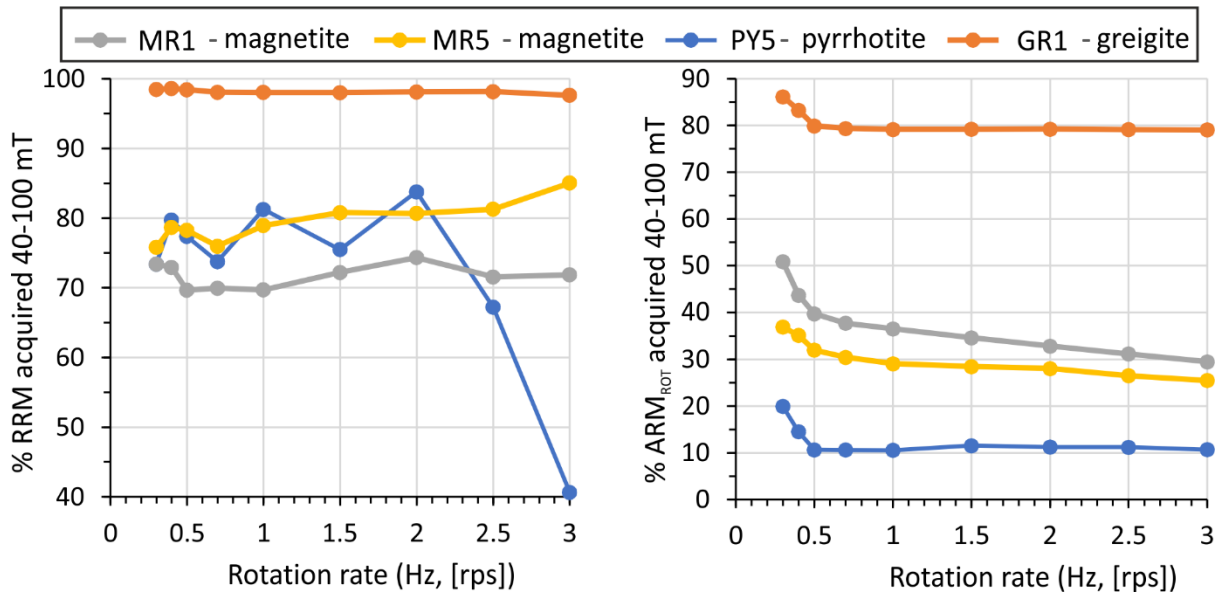


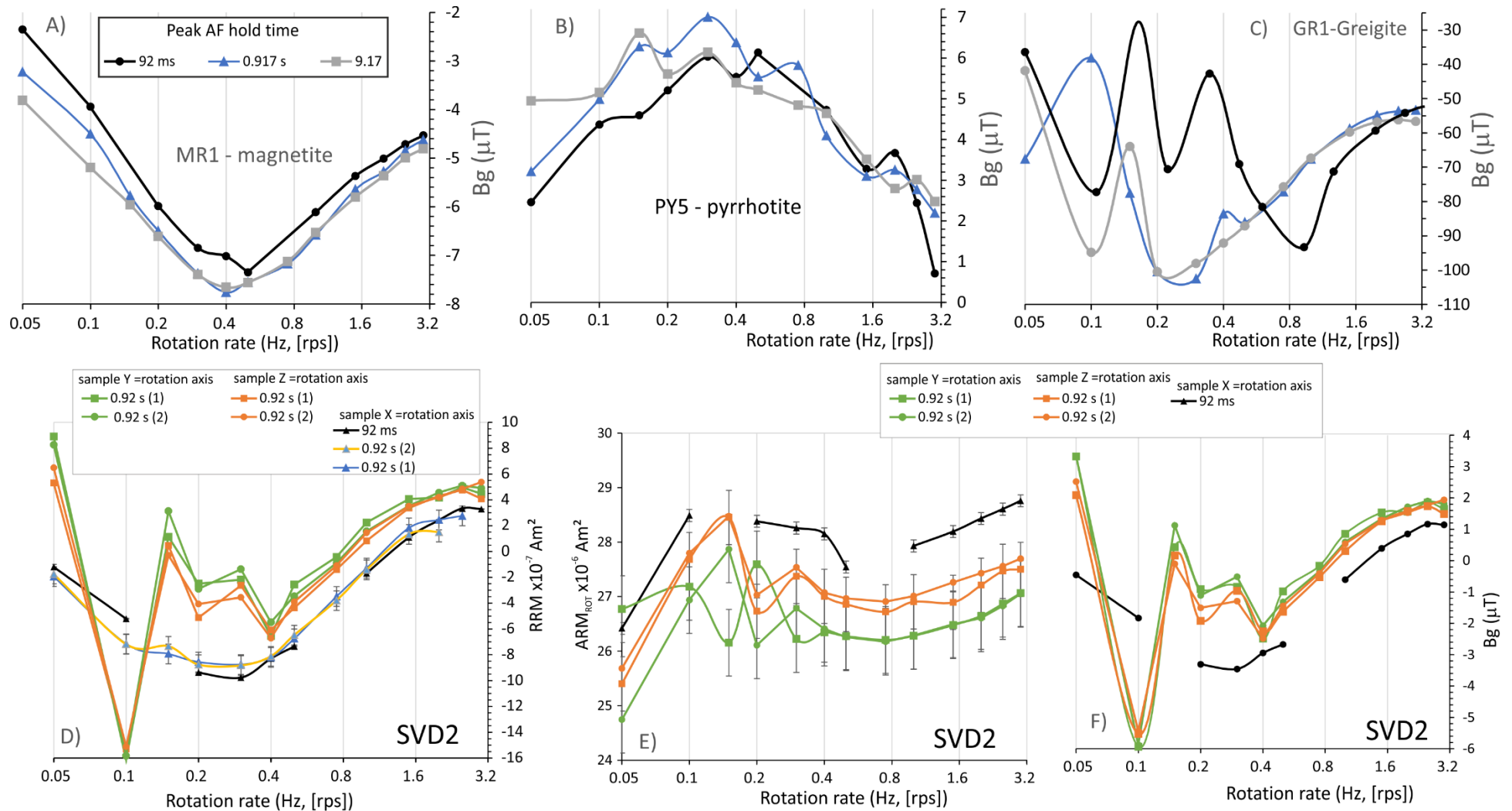
Fig. S17. Variation of normalised RRM,  $ARM_{ROT}$  and  $B_g$  with respect to rotation rates. In each case normalised with respect to the value at 0.5 Hz [rps]. The four columns of graphs apply to each of the samples indicated at the top, and each row of graphs relates to the parameter show on the left ( $T_H = 92$  ms).



**Fig.S18.** The percentage of RRM and  $ARM_{ROT}$  acquired between 40-100mT with respect to the rotation rate. The differing percentages here represent the ability of the RRM and  $ARM_{ROT}$  to be acquired ( $T_H = 92$  ms).

### 3.4 The RRM and $ARM_{ROT}$ acquisition process with changes to the peak AF hold times.

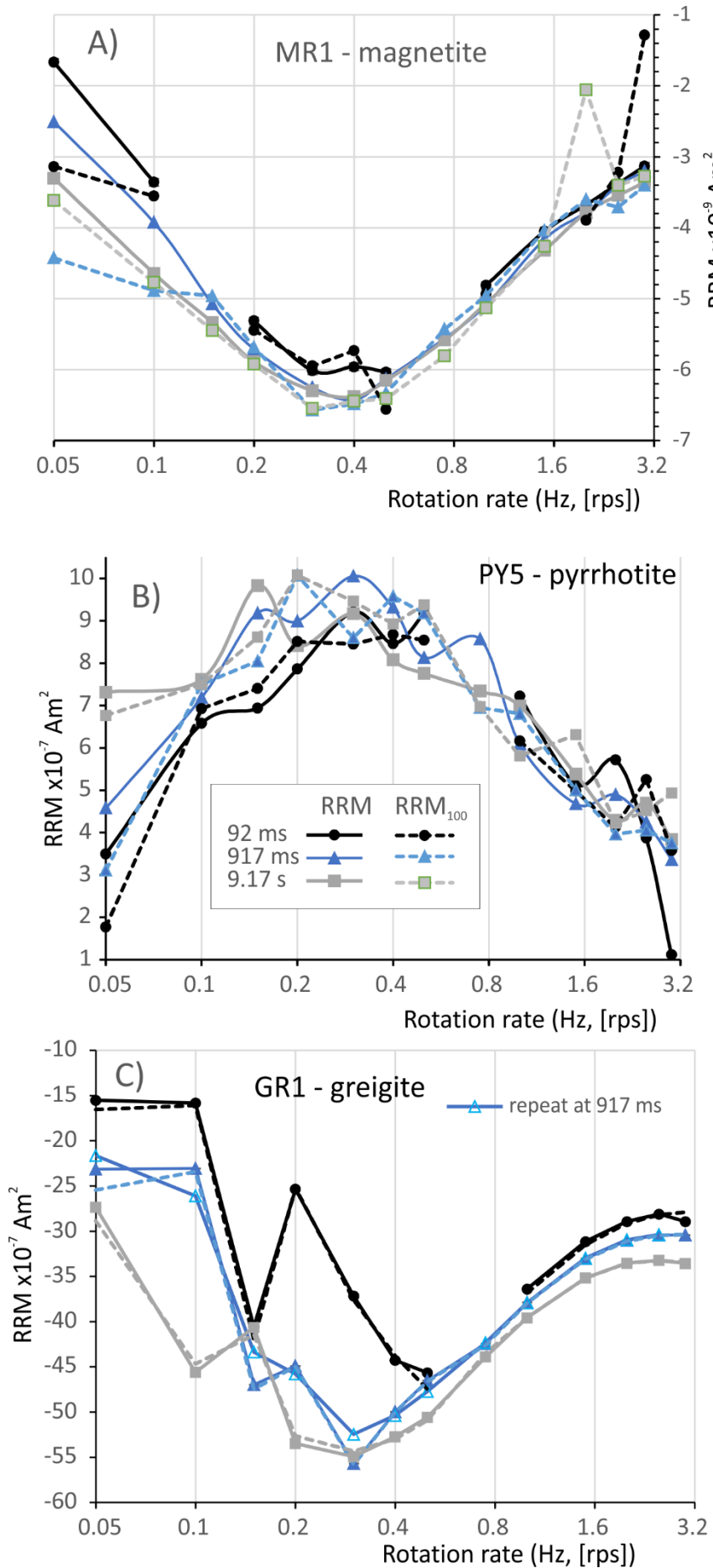
The data in this section displays the impact of changing the hold time of the maximum AF field, to either have incomplete rotation initialisation (92 ms, 30 cycles on RAPID), adequate rotation initialisation (0.917 seconds, 300 cycles on RAPID) at about  $\omega T_D = 0.4$  rev, and full rotation initialisation (9.17 seconds, 3000 cycles) across all rotation rates used.



**Fig. S19.** A,B,C) The variation of  $B_g$  with rotation rate for the three tested values of peak AF hold times. Key to symbols and colours used in A). The 92 ms curve is the normal hold-time setting for most data reported here. A) For MR1 (magnetite) the 92 ms data is approximately 96% of the larger values shown by the data for the longer hold times. B) for PY5 (pyrrhotite) the data is rather noisy and the 92 ms dataset is rather similar to the longer hold times between ~0.4- 3 Hz [rps], but diverges below this. C) The greigite sample (GR1) displays large differences in  $B_g$

for 92 ms, but the maximum is clearly around 0.2-0.3 Hz [rps] for the longer hold times, which for the 0.917 s hold time achieve the full 90° initialisation rotation by around 0.3 Hz [rps]. The large deviations in greigite are likely caused by large anisotropy in this sample which affects Bg at  $\omega$  values below the rotation initialisation threshold. D,E,F) RRM, ARM<sub>ROT</sub> and Bg data for dolerite sample SVD2 at the two smaller hold times versus rotation rate, and with different sample axes parallel to the spin vector (X= sample x-axis parallel to rotation axis etc). The sample is from the margin of near horizontal dolerite sill, and the X-Z plane is the surface approximately parallel to the margin, the Y-axis is near normal to the margin. This sample had been exposed to a 1T IRM between the measurements at 92 ms and 917 ms (responsible for the larger ZSD error bars on the 917 ms measurements, whereas the 92 ms measurements have error bars smaller than the symbols). D) Like the magnetite and basaltic samples studied by Edwards this sample (when spin vector along sample X-axis) has a peak of RRM (using 917 ms hold time) of  $\omega T_D$  between 0.3- 0.45 rev (i.e., 0.2-0.3 Hz [rps] in the scale shown). E) corresponding change in ARM<sub>ROT</sub> which like D) display differences between the axis rotated about, likely due to anisotropy. F) the corresponding changes in Bg. Although we have not independently confirmed the ferrimagnetic anisotropy (Potter, 2004) the curve showing the (negative) maximum in Bg with  $\omega$  is likely that measured such that the rotation axis is near parallel to the maximum anisotropy, possibly represented by an ellipsoid with rather similar minimum and intermediate anisotropy axes, which is why the RRM and Bg along the sample Z and Y axes are similar. A few data points are missing due to sample non-pickup during some tests.





**Fig. S20.** Data for RRM and RRM<sub>100</sub> with respect to rotation rate for the three peak AF hold times tested. Key to symbols and colours used in B). Broadly the RRM<sub>100</sub> values (i.e., the RRM measured with a 100  $\mu$ T DC bias field) track the RRM values acquired with no DC bias field. Like other pyrrhotite data the RRM<sub>100</sub> values are rather noisy in B), and noisy in A) due to weaker intensities. For the greigite sample (GR1) repeats for the 917 ms hold time suggest some additional 'RRM noise', since the measurement noise is smaller than the symbols used (likewise in B). This is perhaps because the sample was mounted differently with respect to the anisotropy axes (in plane normal to rotation axis) on each repeat (or it slipped during rotation).

#### 4. A test of the AF peak hold time and field frequency

Some ambiguity about how to determine the duration of the peak field ( $T_H$ ) during AF demagnetisation, is present in the RAPID software.

1. Indirectly, by reading the parameters "Total Ramp Time (ms)" and "Ramp Up Slope (in ms)" and "Ramp Down Slope (in ms)" (path: Main menu > Diagnostics > AF Demagnetizer > AF Demag Window) when running the AF field in diagnostics mode in the 'AF demag control' window
2. Directly, from the settable parameter  $D$ , "Duration in Sine Wave Periods at the Peak Field" (path: Main menu > View > Settings > AF Demag (2) tab) which controls overall characteristics of the demagnetiser. However, varying the parameter  $D$  does not affect the readings of "Total Ramp time" and "Ramp Up" and "Ramp Down" time as in 1.

A simple test determined that  $T_H$  is governed by  $D$ . Hence, to test the relationship between  $D$ , the AF field frequency and the time of a demagnetisation cycles, we did the following.

- a) Four low AFz (axial) AF fields (between 10 to 30mT) were applied one after another, without any measurements in between.
- b) The total duration of three concatenated AF demagnetization cycles was timed at seven settings of  $D$ . These were 30, 300, 600, 900, 1200, 1500, and 1800 sine-wave cycles.
- c) Each *AF demagnetization cycle* included lowering of the quartz sampling rod to the AFz coil, then the actual AF demagnetization and, finally, the raising of the sampling rod to the starting position.

Figure S21 shows that the lowering and the raising of the sampling rod takes  $14.73 \pm 0.21$  s on average, whilst an increase in  $D$  by 1000 since waves increases  $T_H$  by  $2.687 \pm 0.19$  s on average (uncertainties here are 95% confidence intervals). This corresponds to a frequency of the AFz coil of 372 Hz (95% confidence interval 347-401 Hz) which is close to the 360 Hz frequency indicated in the RAPID software. The conclusion is the hold times and field frequencies seem consistent with the settings in the software.

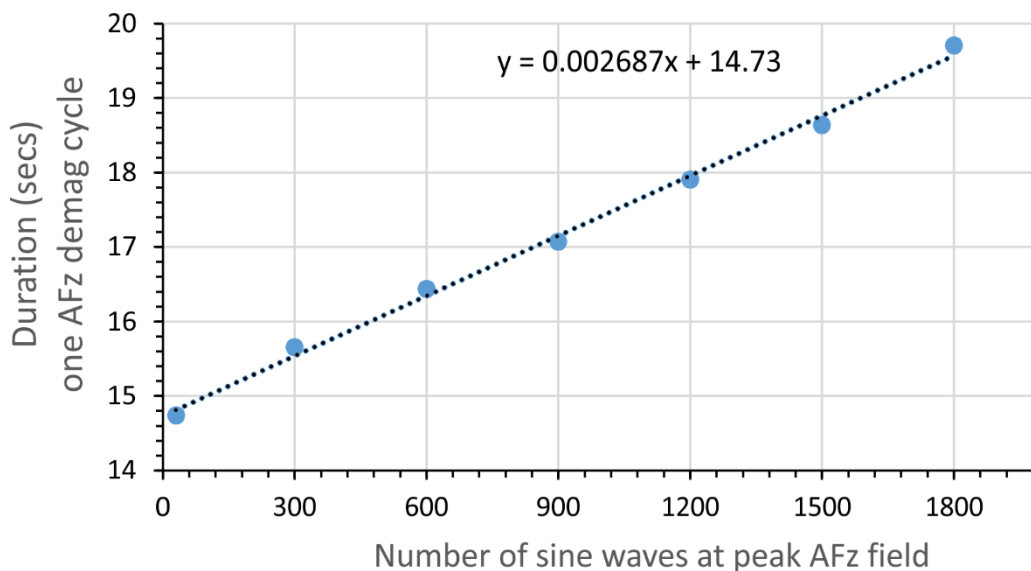


Fig. S21. The data for the test of the  $T_H$  and field frequency. The x-axis is the parameter  $D$  set in the RAPID software. The Y-axis is the average duration of one *AF demagnetization cycle* determined as above.

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