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# Symmetry progression and possible polar metallicity in NiPS3 under pressure

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# ARTICLE OPEN Check for updates Symmetry progression and possible polar metallicity in NiPS<sub>3</sub> under pressure

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van der Waals solids are ideal platforms for the discovery of new states of matter and emergent properties under external stimuli. Under pressure, complex chalcogenides like  $MPS_3$  (M = Mn, Ni, Co, V) host sliding and structural transitions, insulator-to-metal transitions, the possibility of an orbitally-selective Mott state, piezochromism, and superconductivity. In this work, we bring together diamond anvil cell techniques, infrared and Raman scattering spectroscopies, and X-ray diffraction with a detailed symmetry analysis and first-principles calculations to uncover a series of high-pressure phases in NiPS<sub>3</sub>. Remarkably, we find five different states of matter between ambient conditions and 39 GPa—quite different than in the other  $MPS_3$  materials. Even more strikingly, infrared spectroscopy and X-ray diffraction combined with a symmetry analysis reveal both metallicity and loss of the inversion center above ~23 GPa suggesting that NiPS<sub>3</sub> may be a polar metal with a P3m1 space group under these conditions and P1 symmetry under maximum compression. In addition to identifying a candidate polar metal ripe for further inquiry, we suggest that pressure may tune other complex chalcogenides into this elusive state.

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#### INTRODUCTION

Structural phase transitions in complex chalcogenides are attracting considerable attention as drivers of exciting new states of matter with the potential to host emergent properties. As demonstrated in FePS<sub>3</sub> as well as the Mn and V analogs, external stimuli such as pressure (and likely strain) induce a layer-sliding transition and systematic bandgap reduction followed by a volume collapse to the metallic state at room temperature<sup>1-9</sup>. A similar progression might be anticipated to take place in  $NiPS_3^{6,10-13}$ , but even the most cursory inspection shows that this is not the case. Instead, this system hosts an entire series of structural phase transitions and associated symmetry progressions of mysterious origin. Compression thus gives rise to several new states of matter in the  $MPS_3$ 's (M = Mn, Ni, Co, V)—each with properties that can be deterministically controlled. As demonstrated in FePS<sub>3</sub> and CrSiTe<sub>3</sub>, a low-temperature structural phase transition triggers superconductivity as well<sup>2,14</sup>. Whether this is the case in  $NiPS_3$  is not yet apparent, but it is already clear that this system has a rich set of symmetry progressions<sup>12</sup> as well as metallic<sup>11,12,15</sup> and relaxation behavior<sup>15</sup> making it an intriguing platform for deeper study. The recent discoveries of light-induced metallicity<sup>16</sup> and intercalationcontrolled magnetism<sup>17</sup> provide complementary pathways for property regulation, although the structural aspects of these processes are under-explored.

In order to unravel these complexities, we combined highpressure vibrational spectroscopies and X-ray diffraction with an extensive symmetry analysis and first-principles calculations of phase stability, lattice dynamics, and structure to explore the symmetry progression in NiPS<sub>3</sub>. Analysis of the infrared absorption and Raman scattering response reveals that pressure drives a series of structural phase transitions from  $C2/m \rightarrow P\overline{3} \rightarrow P\overline{3}m1 \rightarrow$ a polar phase above ~23 GPa where inversion symmetry is lost. At the same time, NiPS<sub>3</sub> becomes metallic as evidenced by the development of a Drude in the infrared which screens the phonons. X-ray diffraction reveals a P3m1 space group under these conditions, so in addition to becoming metallic, the system breaks inversion symmetry. NiPS<sub>3</sub> remains a polar metal in the two highest pressure states, which X-ray diffraction suggests are P3 and P1, respectively. This complicated symmetry progression and the development of a possible polar metallic phase draws a clear distinction between NiPS<sub>3</sub> and its structural siblings like MnPS<sub>3</sub>, FePS<sub>3</sub>, and V<sub>0.9</sub>PS<sub>3</sub> and signifies unique underlying physics. We discuss the unprecedented number of distinct states of matter in the symmetry progression in terms of stability and stacking sequences as well as the structure-property relationships in the MPS<sub>3</sub> family of materials. The latter is key to understanding prospects for physical and chemical tuning as well as tailoring properties hosted by the ultra-low symmetry phases in these systems.

#### **RESULTS AND DISCUSSION**

## Local lattice distortions and symmetry progression under pressure

Figure 1 summarizes the vibrational response of NiPS<sub>3</sub> under compression at room temperature. Based upon a C2/m structure at ambient conditions [Supplementary Fig. 2, of the Supplementary information], we anticipate 12 infrared-active modes ( $5A_u + 7B_u$ ) and 15 Raman-active modes ( $8A_g + 7B_g$ )<sup>10,18,19</sup>. Overall, the vibrational properties are in good agreement with this symmetry

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**Fig. 1** Vibrational properties of NiPS<sub>3</sub> under compression. a–d Close-up view of the infrared absorption and frequency vs. pressure plots. e-h Close-up view of the Raman scattering response and frequency vs. pressure plots. Spectra for the ambient pressure (AP) phase are shown in gray; those for the high-pressure phases (HP-I, HP-II, and HP-III) are shown in teal, green, and black, respectively. Closed and open circles represent two independent runs. Note that infrared spectra in the HP-III phase do not appear in our measurements because the phonons are screened by the metallic response. The critical pressures ( $P_C$ 's) separate the high-pressure phases and are indicated with dashed vertical gray lines.

analysis and our first-principles lattice dynamics calculations with the exception of (i) an infrared-active  $B_u$  mode near 450 cm<sup>-1</sup> and (ii) a Raman-active lattice mode near 110 cm<sup>-1</sup>, probably due to low oscillator strengths<sup>12,20,21</sup>. NiPS<sub>3</sub> displays several structural phase transitions with increasing pressure. We identify the critical pressures ( $P_c$ 's) by the appearance or disappearance of peaks, splittings or combinations, and frequency shifts as discussed below<sup>22</sup>. Within our investigated range, we find three critical pressures ( $P_c$ 's) and four different phases—the last of which is a polar metal. Clearly, NiPS<sub>3</sub> is much more complex than the Mn and Fe analogs<sup>4,6,9</sup>. The mode assignments, symmetries, and displacement patterns of NiPS<sub>3</sub> are summarized in Supplementary Tables 1 and 2 of the Supplementary information.

The  $C2/m \rightarrow HP-I$  transition at 9 GPa is challenging to unravel because there is spectroscopic evidence for both increasing and decreasing symmetry. This suggests that the initial and final space groups do not have a simple group ↔ subgroup relationship across  $P_{c1}$ . Let's begin by considering the evidence for increasing symmetry. Two different  $2 \rightarrow 1$  peak condensations take place near 260 and 570 cm<sup>-1</sup> [Fig. 1e-h]. We assign these Raman-active modes as in-plane antisymmetric bends and stretches of the  $P_2S_6^{4-}$ dimer units, respectively<sup>20,23</sup>. This means that in addition to the mirror planes, axial glides, inversion centers, two-fold rotations, and the two-fold screw axes that characterize the C2/m space group, the system gains additional symmetry elements across  $P_{c1} = 9 \, \text{GPa.}$ Examination of C2/m group  $\leftrightarrow$  supergroup

relationships in Supplementary Fig. 1 provides several candidates for this intermediate phase including  $P\overline{3}1m$ ,  $P\overline{3}m1$ , and Cmmm. We can distinguish between these different space groups by realizing that a reduction in out-of-plane motion of the phosphorous-dimer atoms (560 cm<sup>-1</sup>) and changes in Ni positions (255 cm<sup>-1</sup>) combined with a decrease in monoclinic angle<sup>4,12</sup> establish new mirror planes, axial glides, two-fold rotations, and two-fold screw axes in the ab-plane. Taken together, these findings support a P31m space group. Vibrational spectroscopy also contains evidence for symmetry breaking across the 9 GPa transition, and the loss of these specific symmetry elements has to be taken into account in order to identify the space group of HP-I. Symmetry reduction is, of course, signaled by the appearance of new peaks. The rise of the out-of-phase intralayer Ni translational mode in the Raman scattering response near 120 cm<sup>-1</sup> across P<sub>c1</sub> is a beautiful example of this effect [Fig. 1e, f]. A second example involves the infrared-active out-of-plane antisymmetric PS<sub>3</sub>-bending modes near 280 cm<sup>-1</sup> [Fig. 1a, b]. The key is to unravel how these modes correspond to the symmetry elements that are lost across  $P_{c1}$ . We do this by considering the relevant mode displacement patterns and how they impact different symmetry elements. These include some of the newly established symmetry elements of  $P\overline{3}1m$  such as axial glide planes, two-fold rotations, and two-fold screw axes. Candidate subgroups include P31m, P3, and P312. Moreover, while many symmetry elements are lost here, the inversion center remains intact and is very robust under compression. Furthermore,



**Fig. 2 Insulator-to-metal transition occurs with the development of a Drude response and screening of the phonons. a** Reflectance of NiPS<sub>3</sub> as a function of frequency showing the development of metallicity under pressure. Dashed lines omit the well-known diamond signature and guide the eye. **b** Close-up view of the 590 cm<sup>-1</sup> P-S stretch which is increasingly screened by the Drude under pressure, consistent with the development of metallicity. The metallicity develops gradually, in line with transport results<sup>12,15</sup> and consistent with a infrared-active second-order transition. This implies that a group-subgroup symmetry analysis is reasonable. **c** Calculated displacement pattern of the P-S stretching mode. Phosphorus and sulfur centers are purple and yellow, respectively. Black arrows indicate both direction and magnitude of the distortion.

we do not see a large increase in the overall number of peaks in the spectra. It is also important to note that the infrared and Raman spectra are never equivalent, thus providing additional evidence for retention of the inversion center. Of the three candidate subgroups of  $P\overline{3}1m$ , only  $P\overline{3}$  retains the inversion center. We therefore assign the 9 GPa structural phase transition as  $C2/m \rightarrow P\overline{3}$ . As shown in Supplementary Fig. 4, X-ray diffraction is consistent with this result.

There is another structural phase transition at  $P_{c2} = 16$  GPa. Increasing symmetry from  $P\overline{3} \rightarrow HP-II$ , group theory provides five space group candidates: P31m, P6/m, P63/m, P3c1, and P3m1. R3 was also considered because it is seen in similar materials including  $\alpha$ -RuCl<sub>3</sub> and CrSTe<sub>3</sub>, although these materials are spectroscopically quite different with an overall fewer number of vibrational modes and likely inconsistent with our data<sup>24,25</sup>. We can remove some of the candidates from this list almost immediately. From a comparison with the intermediate space group from the  $C2/m \rightarrow P\overline{3}$  transition, changes in the spectra indicate that NiPS<sub>3</sub> does not revisit  $P\overline{3}1m$ . Without any evidence for six-fold symmetry, the material does not appear to increase symmetry to either P6/m or  $P6_3/m$ . We are therefore left to select between P3c1 and P3m1. In order to discriminate between these two space groups, we examine experimental evidence in Fig. 1a, b alongside an analysis of the space group-specific symmetry elements in Supplementary Table 3, of the Supplementary information. While the Raman-active vibrational modes change smoothly across  $P_{c2}$ , the infrared response is much more revealing and useful. We focus on the A<sub>u</sub> infrared-active mode at 189 cm<sup>-</sup> The displacement pattern of this mode consists of in-plane Ni translations within the layer and out-of-plane S motions. The disappearance of this mode at  $P_{c2}$  signals increased symmetry in HP-II [Fig. 1a, b]. More specifically, the loss of this motion indicates the development of additional axial glides and mirror planes along the c-axis as well as two-fold rotations and screw axes in the abplane. The addition of these symmetry elements is consistent with  $P\overline{3}m1$ . While two-fold rotations and screw axes are expected to develop in P3c1 as well, diagonal glide planes are also anticipated. There is, however, no evidence for diagonal glide plane formation in HP-II. The smooth behavior of the out-of-plane S motion of the  $B_{\rm u}$  mode at 194 cm<sup>-1</sup> is also inconsistent with significant new interlayer interactions across  $P_{c2} = 16$  GPa. These findings are in

line with our high-pressure X-ray diffraction which refines to  $P\overline{3}m1$  above 16 GPa [Supplementary Fig. 4, of the Supplementary information]. The sequence of pressure-induced transitions thus far is  $C2/m \rightarrow P\overline{3} \rightarrow P\overline{3}m1$ .

NiPS<sub>3</sub> has an additional structural phase transition at  $P_{c3} =$ 23 GPa. As we argue below, the space group progression across this transition (from HP-II to HP-III) is  $P\overline{3}m1 \rightarrow P3m1$ . Unfortunately, the spectra in Fig. 1 provide no mode splittings or frequency shifts across  $P_{c3}$  that can be used to unravel the next step in the symmetry progression. This does not mean, however, that the data are not of any use. The Raman-active modes in Fig. 1 harden continuously across  $P_{c3}$  and remain unaffected by the pressure-driven transition. This indicates that there are no changes in local structure emanating from even-symmetry displacements and that subgroups involving such distortions can be eliminated from consideration. We next turn our attention to the infrared response in order to consider the possible role of odd-symmetry distortions. Unfortunately, the development of a Drude screens the behavior of the infrared-active phonons above 23 GPa. Thus, while the infrared response establishes metallicity (albiet at a slightly lower pressure than that from transport work)<sup>12,15</sup>, it does not provide any information on the behavior of the infrared-active phonons above 23 GPa.

Figure 2 displays the infrared reflectance of NiPS<sub>3</sub> under compression. The overall level rises significantly with increasing pressure, consistent with developing metalicity. The latter also screens the infrared-active P–S stretching mode near 590  $cm^{-1}$ , in line with the discussion above. This metallicity is consistent with the middle-infrared spectra of Ma et al.<sup>12</sup>, and in combination with the loss of the inversion center (established by X-ray scatting, discussed below) suggests that NiPS<sub>3</sub> may be a polar metal at pressures above 23 GPa. Along with metallicity, polar metals require a continuous second-order structural transition, the loss of an inversion center, and the development of a unique polar axis<sup>26</sup>. Screening of long-range Coulombic forces and elimination of polar moment ordering makes this state of matter quite uncommon. Very few polar metals have been experimentally realized. Examples include LiOsO<sub>3</sub>, Cd<sub>2</sub>Re<sub>2</sub>O<sub>7</sub>, the  $T_d$  phases of MoTe<sub>2</sub> and WTe<sub>2</sub>, and antiperovskite Hg<sub>3</sub>Te<sub>2</sub>X<sub>2</sub><sup>27-32</sup>. Although the size of any polarization and whether it is ferroelectric (switchable) or pyroelectric (not switchable) are unexplored, the possibility of



**Fig. 3 High-pressure X-ray scattering results of NiPS<sub>3</sub> up to 39 GPa. a** Integrated X-ray diffraction profiles under pressure. The appearance of X-ray diffraction peaks at the given angle are indicated, alongside dashed lines to guide the eye. Critical pressures, dividing individual pressure phases, are indicated by solid blue horizontal lines. The space groups of individual high-pressure phases are labeled. The asterisk (\*) and plus (+) marks correspond to the diffraction response from NaCl in B1-type structure (below 30 GPa) and B2-type (above 30 GPa)<sup>46</sup>. **b**, **c** The X-ray diffraction images on the detector for HP-I (13 GPa) and HP-III (24 GPa). The peaks from NiPS<sub>3</sub> are enclosed with squares alongside the *hkl* indices for each phase. The gray labels in **c** indicate the diffraction from HP-II. The Bragg diffraction from the diamond anvils is masked with red circles.

ferroelectricity in the metallic state of a chalcogenide like NiPS<sub>3</sub> opens the door to many intriguing opportunities for studying this intriguing state of matter. Furthermore, if the polar metal state can be stabilized under low pressures or strain, there will be applications in nanoelectronics since these effects take place at room temperature.

A subgroup analysis for  $P\overline{3}m1$  suggests several possible pathways by which symmetry in NiPS<sub>3</sub> can be reduced. As before, some can be immediately eliminated. For instance, the spectroscopic results are different from the ambient pressure phase and HP-I, indicating that HP-III does not possess C2/m symmetry or revisit  $P\overline{3}$ . To differentiate between the remaining two space groups, P3m1 and P321, we bring our spectroscopic findings together with X-ray diffraction as discussed below. Importantly, both of these candidate space groups are polar and lack an inversion center.

# Establishing loss of the inversion center above 23 GPa by X-ray scattering

Figure 3a displays the X-ray diffraction profile of NiPS<sub>3</sub> as a function of pressure at room temperature. Similar to the vibrational properties work, we identify the  $P_{C}$ 's by the appearance/disappearance of various resonances as well as changes in peak position and splitting. The critical pressures extracted from the X-ray diffraction studies are in excellent agreement with those obtained from the infrared absorption and Raman scattering measurements. The phases are again labeled from ambient

pressure (AP) through the various high-pressure phases (HP-I to HP-V). Figure 3b, c displays diffraction images recorded at 13 and 24 GPa, in high-pressure phases HP-I and HP-III, respectively. These images indicate that NiPS<sub>3</sub> remains single-crystal-like across the full series of structural transitions up to 39 GPa [Supplementary Fig. 4, of the Supplementary information]. One of the challenges that arise from the experimental configuration and angle of incidence is that we do not collect all of the diffraction peaks. This issue prevents a full refinement of the high-pressure structures. On the other hand, there are usually enough peaks to distinguish between symmetry candidates. We therefore carried out the least-squares fitting of the lattice constants [Table 1] to investigate the symmetries suggested by vibrational spectroscopy. The space groups are labeled accordingly in Fig. 3.

To complement our spectroscopic work, we performed highpressure X-ray diffraction measurements between 0 and 39 GPa and carried out detailed refinements of each phase that were compared with minimum energy calculations. We discuss the analysis of HP-III as an illustration of this process. As a reminder, infrared spectroscopy reveals that this phase is metallic [Fig. 2]. The predicted space group candidates are P3m1 and P321. Both of these space group candidates fit similarly well to the X-ray data based upon the values of their lattice constants [Table 1]. For example, the peaks at  $2\Theta = 4.8^\circ$ ,  $8.3^\circ$ ,  $14.5^\circ$ , and  $14.8^\circ$  correspond to (h, k, l) = (1,0,0), (1,1,0), (3,0,0), and (3,0,1), respectively [Fig. 3c]. We discovered, however, that P321 does not retain the hexagonal structure centered around the  $P_2S_6^{4-}$  dimer, suggesting P3m1 as the most probable space group for HP-III. A schematic structure of this phase is provided in Supplementary Fig. 5 of the Supplementary information. The sequence of transitions under pressure is therefore  $C2/m \rightarrow P\overline{3} \rightarrow P\overline{3}m1 \rightarrow P3m1$ .

We carried out a similar analysis for all of the phases from AP to HP-V. Although perfect fits are not always obtained [Table 1], the extracted space groups from AP to HP-III are in excellent agreement with the spectroscopic results. The data also suggest that the symmetry of NiPS<sub>3</sub> continues to decrease with increasing pressure [Fig. 3 and Supplementary Fig. 4]. With evidence for symmetries lower than P3m1 under additional compression, we anticipate the space groups for HP-IV and HP-V to be P3 and P1 respectively. In principle, P1 should have a larger electric polarization than either P3m1 or P3.

### Evidence for the series of pressure-induced transitions from ab initio calculations

In order to provide insight into the structural transitions and lattice dynamics of NiPS<sub>3</sub>, we carried out a combinatorial search of all possible stacking sequences up to and including three-layer periodicity followed by first-principles density functional theory calculations including on-site Coulomb repulsion, magnetism, and

<b>Table 1.</b> Symmetry and lattice constants for the ambient pressure(AP) and high-pressure (HP) phases of NiPS3.			
P (GPa)	Phase	Symmetry	Lattice constants
0.7	AP	C2/m	a = 5.800(1) Å, b = 10.057(1) Å, c = 6.605(2) Å
			$a = \gamma = 90^{\circ}, \ \beta = 106.99(2)^{\circ}$
13	HP-I	P <u>3</u>	a = b = 5.638(5) Å, $c = 5.88(2)$ Å
			$\alpha = \beta = 90^\circ$ , $\gamma = 120^\circ$
16	HP-II	P <u>3</u> m1	a = b = 9.720(1) Å, $c = 4.94(1)$ Å
			$\alpha = \beta = 90^{\circ}$ , $\gamma = 120^{\circ}$
24	HP-III	P3m1	a = b = 5.582(3) Å, $c = 15.03(2)$ Å
			$\alpha = \beta = 90^{\circ}$ , $\gamma = 120^{\circ}$
27	HP-IV	P3	a = b = 5.488(1) Å, $c = 9.92(4)$ Å
			$\alpha = \beta = 90^\circ$ , $\gamma = 120^\circ$

van der Waals correction. Three potential candidates were identified: (i) the well-known bulk C2/m structure with staggered phosphorus dimers where each layer slips by a/3 along the monoclinic *a*-axis, (ii) the trigonal  $P\overline{3}$  structure with three-layer periodicity along the c-axis, and (iii) the triclinic P1 structure, the lowest symmetry space group with the same *c*-axis periodicity as  $P\overline{3}$ . Note that the suggested intermediate high-pressure space groups  $P\overline{3}m1$  and P3m1 could not be accommodated within our model crystal structures. We also mention in passing that sticking between adjacent NiPS<sub>3</sub> layers, reported in ref.<sup>12</sup> but unobserved in our current calculations, can be reproduced only when the onsite Coulomb repulsion and magnetism are ignored. It is our experience that neglecting the Coulomb repulsion and magnetism in MnPS<sub>3</sub> and NiPS<sub>3</sub> leads to substantial underestimation of the lattice parameters as well<sup>11</sup>. Further computational detail is available in the Methods section.

Figure 4a displays enthalpy versus pressure for the three candidate structures mentioned above. At ambient conditions, the C2/m phase is energetically stable—consistent with the experiment. The other two structures are slightly less favorable with energies that are 3-4 meV higher. We find that the enthalpies of both P1 and  $P\overline{3}$  drop below that of the C2/m phase near 4 GPa. The critical pressure varies significantly depending upon the choice of the U parameter and the type of van der Waals correction. That said, we do not find evidence for a  $P\overline{3}$  ground state up to 20 GPa. The origin of this fast stabilization of the P1 phase is unclear at this time and is under study now. Note that the calculated enthalpy difference between C2/m,  $P\overline{3}$ , and P1 is less than 20 meV per formula unit even up to 20 GPa, smaller than the energy scale of thermal fluctuations at room temperature ( $\approx$ 30 meV). The incorporation of temperature effects in the lattice dynamics calculations may stabilize  $P\overline{3}$  in the intermediate pressure range. By contrast, enthalpy differences between competing phases in  ${\rm MnPS_3}^9$  and  ${\rm FePS_3}^8$  are greater than 50 meV per formula unit. The small enthalpy difference between competing structural phases along with a narrower van der Waals gap<sup>20</sup> may be the origin of the sluggish<sup>15</sup> and complicated set of pressure-induced structural phase transitions in NiPS<sub>3</sub>. Finally we note in passing that the  $P\overline{3}1m$ -type stacking which is reported to occur in the high-pressure regimes of MnPS<sub>3</sub><sup>9</sup>, FePS<sub>3</sub><sup>4,7,8</sup>, and  $V_{0.9}PS_3^5$  is not stabilized in NiPS<sub>3</sub>, and in our first-principles



**Fig. 4 Computed enthalpies and infrared/Raman frequencies versus pressure. a** Relative enthalpies of the  $P\overline{3}$  and P1 phases with respect to C2/m. **b**, **c** Comparison between frequencies of computed and experimental **b** infrared and **c** Raman modes. For clarity,  $A_{u/g}$  (blue) and  $B_{u/g}$  (violet) modes from the C2/m phase and  $E_{u/g}$  (orange) modes from the  $P\overline{3}$  phase are selected for display. The experimental frequencies are overlaid in open circles (cyan). The experimental frequencies were shifted from the theoretical data by 5 and 8 cm<sup>-1</sup> in panels **b** and **c**, respectively, to overlap the theoretical curves. Frequencies from the C2/m ( $P\overline{3}$ ) phase are depicted as solid (dashed) lines below 8 GPa and dashed (solid) lines above 8 GPa.

calculations, it becomes increasingly unfavorable under pressure. At 20 GPa, the relative enthalpy of the  $P\overline{3}1m$  phase with respect to C2/m reaches between 30 and 60 meV per formula unitdepending on the type of magnetism and exchange-correlation functionals employed (not shown). Figure 4b, c displays a comparison of calculated and experimental infrared- and Raman-active mode frequencies as a function of pressure. Overall, the computed frequencies are in excellent agreement with our experimental results. The measured infrared response, for instance, reveals a gradual change in slope of the  $\{A, B\}_{ij}$  modes close to 230 cm<sup>-1</sup> near 12 GPa. Based on our computed phonon frequencies, we can understand this structure as the merging of  $\{A, B\}_{\mu}$  modes (blue and violet lines) into an  $E_{\mu}$  doublet (orange solid line beyond 8 GPa); this merge takes place with a  $10 \text{ cm}^{-1}$ red shift. In addition, the experimental infrared response shows the sudden appearance of a peak around 280 cm<sup>-1</sup> beyond 12 GPa, which matches well with the theoretical orange line and can be identified as another  $E_{u}$  mode. The agreement between computational and experimental findings is less evident in the Raman data due to the overall density of modes [Fig. 4c], but we comment that merging the two separate modes near 280 cm<sup>-</sup> close to 8 GPa, possibly the  $A_{q}$  and  $B_{q}$  modes, into a single  $E_{q}$ doublet can account for our results.

Overall, our theoretical results are reasonably consistent with the experimental observations. The theory correctly predicts C2/mas the ambient pressure phase as well as a structural transition from C2/m to  $P\overline{3}$ . It also identifies P1 as the final state. There are also, unfortunately, some inconsistencies. These include: (i) X-ray and spectroscopic measurements show the P1 phase should appear beyond 30 GPa, whereas theory always favors P1 over  $P\overline{3}$ , (ii) the *c*-axis periodicites of the  $P\overline{3}$  structure are different, with computational structures tending to prefer more complex stacking patterns for  $P\overline{3}$ . and (iii) we do not identify  $P\overline{3}m1$  and P3m1 space groups as low energy candidates in the intermediate pressure range-despite their identification in the experiment. These discrepancies may have their origins in our model structures which do not contain components of in-planesymmetry lowering and the resulting multiplications of in-plane periodicity.

Finally, it is worth mentioning that the behavior of NiPS<sub>3</sub> is vastly different from MnPS<sub>3</sub>, FePS<sub>3</sub>, and V<sub>0.9</sub>PS<sub>3</sub> under pressure where structural and insulator-metal transitions drive the formation of phosphorous chains along the layer-normal direction<sup>4,5,9</sup>. We speculate that (i) charge-transfer involving Ni<sup>2+33,34</sup> may disfavor phosphorous ion participation in the insulator-metal transition thereby preventing the simple  $P\overline{3}1m$ -type stacking observed in the Mn, Fe, and V analogs, and (ii) the much narrower van der Waals gap in NiPS<sub>3</sub><sup>20</sup>, possibly originating from partially filled e<sub>g</sub> orbitals and the consequent stronger inter-plane interactions<sup>11</sup>, may prevent layers from sliding easily to form the more familiar stacking patterns. Taken together, these mechanisms may contribute to more complex stacking patterns of the Ni honeycomb layers as well as local lattice distortions within the layers.

### Possible polar metallicity in NiPS<sub>3</sub>, other complex chalcogenides, and the two-dimensional limit

The unique symmetry progression in NiPS<sub>3</sub> under compression provides an opportunity to explore a variety of different states of matter not present in other members of the *M*PS<sub>3</sub> family of materials. By combining high-pressure vibrational spectroscopies and X-ray diffraction with an extensive symmetry analysis and first-principles calculations, we show that this system displays a series of five different structures from *C*2/*m* at ambient conditions to *P*1 above 35 GPa. Strikingly, NiPS<sub>3</sub> appears to be a polar metal above ~23 GPa due to the development of metallicity (evidenced by a Drude in the infrared that screens the phonons) and loss of the center of inversion (established by a combined symmetry and

structural analysis). Whether this phase is ferroelectric (switchable) or pyroelectric (not switchable) is not yet clear; future work will surely address this question. Although of contemporary interest, polar metals are uncommon and currently not well understood. Physical systems are in short supply due to the need to overcome the innate opposition between metallicity and the presence of polar moments<sup>27,30,32</sup>. The most common way to overcome these competitive effects is with a unique polar axis<sup>26,35</sup>. The latter is often a result of an anisotropic crystal environment, can be induced through the application of external stimuli: pressure, strain, and magnetic field. Such a mechanism is possible in NiPS<sub>3</sub> if the unique polar axis is along with c. In any case, the discovery of a candidate physical manifestation of this elusive state of matter is useful and important, offering a real system on which to carry out experiments and test predictions. Additionally, exploring the opposite effect of pressure by dimensionality-lowering layer exfoliation<sup>20</sup> or thin-film growth may reveal unprecedented states of matter<sup>36</sup>, considering that strong spin-orbital-lattice coupling<sup>33,34</sup> and interlayer interactions<sup>11</sup> in NiPS<sub>3</sub> may result in a complicated interplay in the two-dimensional limit. We anticipate that our findings will inspire further exploration of NiPS<sub>3</sub> and a search for polar metallicity in other complex chalcogenides.

#### METHODS

#### Crystal growth and loading of the diamond anvil cell

 $\mathsf{NiPS}_3$  single crystals were prepared by chemical vapor transport techniques as described previously<sup>37</sup>. In order to fulfill the need to access high pressures alongside the requirement for the necessary throughput for the experiments, various diamond culets and gasket sizes were employed. Typically, single-crystal samples were loaded into a symmetric diamond anvil cell with 300 µm anvils using a stainless steel gasket with a thickness on the order of 50 µm and a 100 µm hole along with a pressure medium and an annealed ruby ball. Type IIas and ultra-low fluorescence diamonds were employed for infrared absorption and Raman scattering, respectively. Signal response for each measurement was maximized for the excitations of interest. Depending on the measurement, a guasihydrostatic pressure environment was obtained using KBr (Raman and middle-infrared) or petroleum jelly (far-infrared) as a pressure-transmitting medium. Fluorescence from the R1 peak of the ruby ball was used for pressure determination<sup>38</sup>. Similarly, we loaded single-crystalline NiPS<sub>3</sub>, NaCl pressure-transmitting medium, and ruby chips for pressure calibration<sup>38</sup> into a diamond anvil cell for X-ray diffraction measurements. In the present X-ray diffraction measurement, the single crystals were loaded with the ab-plane parallel to the diamond face. A pre-compressed NaCl flake was placed underneath the sample. This pressure-transmitting medium assured that the pressure distribution across the sample was as homogeneous as possible.

#### Spectroscopic measurements

High-pressure infrared measurements were carried out in both transmittance and reflectance mode over the 50–8000 cm<sup>-1</sup> frequency range with a resolution of 2 or 4 cm<sup>-1</sup> for far and middle-infrared, respectively. A helium-cooled bolometer was employed in the far-infrared. Raman scattering for NiPS<sub>3</sub> was performed in a back-scattering geometry using a green laser ( $\lambda_{excit} = 532$  nm;  $\approx 1$  mW power) and a 1200 g/mm grating, with 30 second integration times, averaged three times at each pressure. Compression was reversible within our sensitivity. These experiments were performed at the 22-IR-1 beamline at the National Synchrotron Light Source II at Brookhaven National Laboratory. All data were collected at room temperature. The critical pressures were determined from splittings, combinations, and inflection points in the mode frequency vs. pressure plots in conjunction with our X-ray analysis.

#### X-ray diffraction measurements

High-pressure X-ray diffraction measurements were performed at 16-ID-B at the Advanced Photon Source, Argonne National Laboratory. The wavelength of the X-ray beam was 0.4066 Å, and the beam spot size was 10  $\mu$ m in diameter for all measurements. Diffraction data were collected on a flat panel detector array (Dectris Pilatus 1M-F, pixel size:

 $172 \times 172 \ \mu\text{m}$ ) in the forward scattering geometry at room temperature. For two-dimensional X-ray diffraction data reduction, Dioptas was used<sup>39</sup>. Samples were single-crystal-like up to the highest pressure. Measurements were performed with the X-ray beam perpendicular to the *ab*-plane, in this manner X-ray diffraction peaks for the restricted hkl indices were obtained. To maximize the number of reflections from the single crystal, data was recorded by rotating the cell about the vertical axis (by ±10 degrees). This process is illustrated in Supplementary Fig. 5, in the Supplementary information.

#### Symmetry analysis

We point out that a group-subgroup analysis is valid only for a secondorder phase transition. That is why we did more than move up and down the subgroup/supergroup symmetry tree—even though we suspect that the transitions of interest here are second order. For instance, we tested a number of space groups that have been seen in another van der Waals materials such as *Cmmm* and *R*3. And we tested them in terms of calculating their energies relative to other candidates under those conditions, trying to relax the structure to one with a suitable match to the experiment, and comparing the calculated lattice dynamics of that candidate structure to our measured results.

#### **First-principles calculations**

Density functional theory calculations were carried out using projectoraugmented-wave (PAW) method as implemented in Vienna ab initio Simulation Package (vasp)<sup>40,41</sup>. Perdew-Burke-Ernzerhof generalized gradient approximation was chosen as the exchange-correlation functional<sup>42</sup>. On-site Coulomb repulsion and van der Waals interactions were incorporated via the choice of a rotationally-invariant flavor of DFT+U formalism<sup>43</sup> and optB86b functional, respectively<sup>44</sup>. The total energy of each magnetic structure is calculated by assuming Néel-type antiferomagnetic order. A non-shifted  $\Gamma$  centered  $9 \times 9 \times 4$  k-mesh was used to sample Brillouin zone. For all self-consistent calculations, plane-wave energy cutoff and energy convergence criterion were set to 500 and  $10^{-8}$  eV, respectively. Internal atomic coordinates of NiPS3 were relaxed until Hellmann-Feynman force on each atom satisfies convergence criteria,  $10^{-4} \text{ eV/Å}$ . To compute infrared and Raman-active modes of vibration, we employed the phonopy code interfaced with vasp<sup>45</sup>. The mode decomposed phonon frequencies were calculated as a function of pressure to compare with the experimental phonon frequencies.

#### DATA AVAILABILITY

The datasets generated and/or analyzed during the present study are available from the corresponding author on reasonable request.

#### **CODE AVAILABILITY**

The Vienna ab initio Simulation Package (vasp, Ver. 5.4, see https://www.vasp.at) and Wien2k (Ver. 2019, see http://www.wien2k.at) are commercial codes, while DFT +embedded DMFT Functional code (see http://hauleweb.rutgers.edu/tutorials/) is an open-source one which runs on top of the Wien2k package.

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#### **AUTHOR CONTRIBUTIONS**

This project was developed by J.L.M., H.-S.K., and N.C.H. Crystal growth and sample characterization were carried out by A.H. with guidance from D.M. The spectroscopic measurements were performed by N.C.H., A.J.C., K.A.S., Z.L., and J.L.M. The X-ray measurements were carried out by T.M., E.F., J.S., and H.C. First-principles calculations

were performed by S.S. with guidance from H.-S.K. The manuscript was written by N.C.H., T.M., S.S., H.-S.K., and J.L.M. All authors commented on the final document.

#### **COMPETING INTERESTS**

The authors declare no competing interests.

#### ADDITIONAL INFORMATION

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