

Correlated-Gaussian calculations of the ground and low-lying excited states of the boron atomSergiy Bubin¹ and Ludwik Adamowicz^{2,3}¹*Department of Physics and Astronomy, Vanderbilt University, Nashville, Tennessee 37235, USA*²*Department of Chemistry and Biochemistry, University of Arizona, Tucson, Arizona 85721, USA*³*Department of Physics, University of Arizona, Tucson, Arizona 85721, USA*

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Benchmark variational calculations of the four lowest 2P and 2S states of the boron atom (including the ground state) have been performed. The wave functions of the states have been expanded in terms of all-particle explicitly correlated Gaussian basis functions and the finite mass of the nucleus has been explicitly accounted for. Variational upper bounds for the nonrelativistic finite- and infinite-nuclear-mass energies of all considered states have been obtained with the relative convergence of the order of 10^{-7} – 10^{-8} . Expectation values of the powers of the inter-particle distances and Dirac δ functions depending on those distances have also been computed. These calculations provide reference values that can be used to test other high-level quantum chemistry methods.

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I. INTRODUCTION

There are several reasons for which finding accurate solutions of the Schrödinger equation for small atoms and molecules is important. Probably the most compelling one is to keep up with constantly improving experimental measurements. Much of the up-to-date atomic and molecular spectroscopic data have a resolution that exceeds 0.01 – 0.001 cm^{-1} or even better. Unfortunately, theoretical techniques are usually unable to match such high accuracy values. With this, a good amount of potentially interesting physical phenomena remain inaccessible for quantitative (and often even qualitative) analysis and interpretation. For some atomic properties, the theory is orders of magnitude behind the experiment. It should be mentioned that the results obtained in high accuracy calculations of small few-electron systems when combined with high-resolution spectroscopic data make it possible to precisely determine such quantities as the fine-structure constant, electron-to-proton mass ratio, nuclear radii, quadrupole moments, etc.

Another important reason for performing accurate calculations on small atomic and molecular systems is that they can provide valuable reference data for the development and testing of less accurate quantum-chemical methods. A simple example is the total nonrelativistic energy of the system, which is one of the most commonly calculated quantities. This energy is difficult to determine very precisely based on purely experimental data because, even if very accurate measurements of the ionization potentials and/or the dissociation energies are available, the procedure to determine the energy requires knowledge of the binding energies of subsystems and the exact contribution of relativistic and QED effects. These latter quantities, while readily available from the calculations, are not directly obtained in the experiment.

However desirable very precise calculations on small atoms are, the difficulties associated with obtaining a very accurate solution of the Schrödinger equation are quite substantial. The main cause of these difficulties is rooted in the multi-dimensional nature of the wave function. The amount of the computational work necessary for such a task grows very fast with the number of particles in the system. Among the methods that are capable of effectively describing the inter-particle

correlation effects, the class of explicitly correlated methods has proven to be particularly powerful. In this class of methods, the wave function of the system is represented in terms of basis functions that explicitly depend not only on the positions of the particles but also on the inter-particle distances. When this type of wave function is combined with the Rayleigh-Ritz variational scheme, it can describe either the ground state or an excited state of a small atom, a molecule, or any other quantum-mechanical system with an unmatched accuracy. In the case of the most studied three-particle atomic system, the helium atom, the accuracy of the best recent calculations has been truly impressive and exceeded 40 decimal figures in the total energy [1–3]. Such accuracy was possible due to the use of a basis set that, apart from other necessary components, also included logarithmic terms that correctly describe the behavior of the wave function at three-particle coalescence points.

Unfortunately, including basis functions that are capable of exactly representing all fine nuances of the exact wave function is only possible for three-particle systems. Including such functions in the calculations of larger systems leads to extremely complicated Hamiltonian and overlap matrix elements. For this reason, the next smallest atom, Li, has been computed with “only” 12- to 13-digit accuracy [4–6] using wave functions expanded in terms of the Hylleraas-type basis functions. These basis functions do not have proper behavior at the three-particle coalescence points, but the wave function constructed with them can still satisfy the Kato cusp conditions [7]. The calculations with Hylleraas-type basis functions are currently limited to four-particle (or three-electron) atomic systems. The extension of the approach to larger atoms again faces difficulties with the analytic evaluation of the multidimensional integrals. The same is true in the case of yet another type of basis function, i.e., explicitly correlated exponential functions (also known as explicitly correlated Slater-type functions) [8–11].

The achieved precision in the calculations on atoms larger than Li drops even further. The accuracy in the calculations of the Be atom and other four-electron atomic systems had remained significantly lower than the accuracy of the available spectroscopic data until very recently, when a series of calculations matching the experimental error bars were

reported [12–18]. All those calculations were performed using the explicitly correlated Gaussian functions (ECGFs). These functions have the advantage that the algorithms for the Hamiltonian matrix elements with them can be derived in the general form for any number of electrons, thus overcoming the difficulties associated with the use of the Hylleraas-type functions and of the explicitly correlated Slater-type functions.

In this paper, we show that, if sufficient computational effort is invested, the correlated Gaussians are capable of producing results for the ground and some lower-lying excited states of a five-electron atomic system, the boron atom, which almost match the accuracy achieved for the four-electron systems. As will be described in the next section, the key component of the variational energy minimization procedure that leads to such accurate results is the analytical energy gradient determined with respect to the Gaussian nonlinear parameters, which are optimized.

II. METHOD

The standard atomic quantum-mechanical calculations are usually performed with infinite nuclear mass, i.e., with assuming the Born-Oppenheimer (BO) approximation. However, in calculations where very high accuracy is desired, the energy needs to include the effect of the finite mass of nucleus. This can be done by using a Hamiltonian, which represents the internal state of the system, to explicitly include the dependency on the nuclear mass. Such a Hamiltonian, called here the internal Hamiltonian, is obtained from the laboratory-frame nonrelativistic Hamiltonian by rigorously separating out the center-of-mass motion. For an atom with n electrons, this is done by a transformation from a laboratory coordinate frame, which describes the positions and the nucleus and the electrons of the atom, to a new set of coordinates comprising three Cartesian coordinates of the center of mass and n internal Cartesian coordinates describing the positions of the electrons with respect to the nucleus. By separating out the center-of-mass motion, we obtain the following internal Hamiltonian (for details, see [19–22]):

$$\hat{H} = -\frac{1}{2} \left(\sum_{i=1}^n \frac{1}{\mu_i} \nabla_{\mathbf{r}_i}^2 + \sum_{\substack{i,j=1 \\ i \neq j}}^n \frac{1}{m_0} \nabla_{\mathbf{r}_i} \cdot \nabla_{\mathbf{r}_j} \right) + \sum_{i=1}^n \frac{q_0 q_i}{r_i} + \sum_{i>j=1}^n \frac{q_i q_j}{r_{ij}}, \quad (1)$$

where \mathbf{r}_i is the distance between the i th electron and the nucleus, m_0 is the nucleus mass [$m_0(^{10}\text{Be}) = 18\,247.4689m_e$ and $m(^{11}\text{Be}) = 20\,063.7375m_e$, where m_e the electron mass], q_0 is its charge, q_i are electron charges, and $\mu_i = m_0 m_i / (m_0 + m_i)$ are electron reduced masses. The Hamiltonian (1) describes the motion of n (pseudo)electrons, the masses of which are the reduced masses, in the central field of the nuclear charge. This motion is coupled through the Coulombic interactions between the electrons and the nucleus $\sum_{i=1}^n \frac{q_0 q_i}{r_i}$ and through the interactions between the electrons $\sum_{i>j=1}^n \frac{q_i q_j}{r_{ij}}$, where $r_{ij} = |\mathbf{r}_j - \mathbf{r}_i|$, as well as through the mass polarization term $-\frac{1}{2} \sum_{\substack{i,j=1 \\ i \neq j}}^n (1/m_0) \nabla_{\mathbf{r}_i} \cdot \nabla_{\mathbf{r}_j}$.

In the calculations of atomic states with only s electrons (in the leading configuration), we use the following all-electron,

spherical ECGFs:

$$\phi_k(\mathbf{r}_1, \mathbf{r}_2, \dots, \mathbf{r}_n) = \exp[-\mathbf{r}'(A_k \otimes I_3)\mathbf{r}], \quad (2)$$

where \mathbf{r} is a vector formed by the $\mathbf{r}_1, \mathbf{r}_2, \dots, \mathbf{r}_n$ vectors stacked on top of each other, A_k is a $n \times n$ symmetric matrix, I_3 is a 3×3 identity matrix, \otimes is the Kronecker product symbol, and the prime indicates the matrix or vector transpose. As the basis functions used in describing bound states must be square integrable, restrictions must be imposed on the A_k matrices. Each A_k matrix must be positive definite. Rather than imposing restrictions on the A_k matrix elements, which can be quite costly to handle in the computational implementation, we use the following Cholesky factored form of A_k : $A_k = L_k L_k'$, where L_k is a lower triangular matrix. With this representation, A_k is automatically positive definite for any values of the L_k matrix elements ranging from $-\infty$ to ∞ . Thus, the variational energy minimization with respect to the L_k parameters can be carried out without any restrictions imposed on the optimized parameters. It should be noted that the $L_k L_k'$ representation of A_k matrix does not limit the flexibility of basis functions, because any symmetric positive matrix can be represented in a Cholesky factored form.

To describe states with $(n-1)$ s electrons and a single p electron, we used the following ECGFs in the wave-function expansion:

$$\phi_k(\mathbf{r}_1, \mathbf{r}_2, \dots, \mathbf{r}_n) = z_{m_k} \exp[-\mathbf{r}'(A_k \otimes I_3)\mathbf{r}]. \quad (3)$$

Here m_k is an integer that depends on k with values ranging from 1 to n . m_k is also subject to optimization.

Appropriate symmetry projections are used to make the wave function of the atom antisymmetric with respect to the permutation of the electron labels. In this paper, we use the spin-free formalism. In this formalism, the symmetry projections acting on the spatial parts of the wave function, i.e., the basis functions, can be represented using the Young projection operators \hat{Y} , which are linear combinations of permutational operators \hat{P}_ν . As the Hamiltonian is invariant with respect to all permutations of the electrons, in calculating the overlap and Hamiltonian matrix elements, the permutational operators can be applied to the ket (or the bra) only. More specifically, the ket basis functions in those matrix elements are operated on with the operator $\hat{Y}^\dagger \hat{Y}$ (the dagger stands for conjugate), where the \hat{Y} operator is derived using the appropriate Young tableaux for the state under consideration (for details of the formalism see, for example, [23]). For the doublet states of boron, the Young operator can be chosen as $\hat{Y} = (\hat{1} + \hat{P}_{23})(\hat{1} + \hat{P}_{45})(\hat{1} - \hat{P}_{24})(\hat{1} - \hat{P}_{26} - \hat{P}_{46})(\hat{1} - \hat{P}_{35})$, where the nucleus is labeled as particle 1, and the electrons are labeled as particles 2, \dots , 6, $\hat{1}$ is the identity operator, and \hat{P}_{ij} is the permutation operator of the spatial coordinates of the i th and j th electrons.

III. RESULTS AND DISCUSSION

The two lowest 2S and two lowest 2P states of boron are considered in these calculations. The lowest 2P state is the boron ground state. In Fig. 1, we use the experimental data taken from NIST Atomic Spectra Database [24] to show where those considered states are positioned in the boron spectrum.

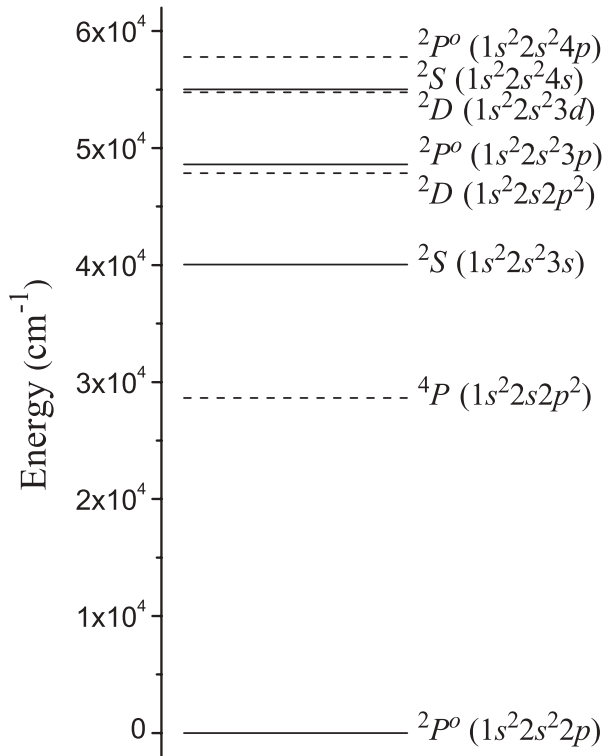


FIG. 1. The energy level diagram of the low-lying states of the boron atom. The four levels computed in this work are marked with the solid line.

The calculations have been carried out for ^{10}B and ^{11}B isotopes as well as for $^{\infty}\text{B}$; the latter one corresponds to the boron atom with an infinitely heavy nucleus. As the majority of quantum-chemical calculations are traditionally performed with an infinite nuclear mass, including the $^{\infty}\text{B}$ results here may provide a useful set of reference values for such calculations. The basis sets for different states have only been generated for the main ^{11}B isotope and reused in the calculations of ^{10}B and $^{\infty}\text{B}$. The reason for this time-saving simplification is due to the fact that, upon a small change of the wave function (the change remains small as long as the mass of the nucleus is significantly larger than that of electrons), it is sufficient to readjust only the linear coefficients of the wave function in order to recover the shift in the total energy. The

simplification has virtually no effect on the accuracy of the calculation.

These calculations have been performed using the standard variational approach. In the variational energy minimization, the matrix elements of the L_k matrices of the basis functions have been optimized. As mentioned, this was only done for the ^{11}B isotope. For 2P states, the m_k powers have been partially optimized. The optimization was carried out independently for each state and the basis set for each state has been grown incrementally from a small set of functions to the size of 5100 functions. Some preliminary results for ground state only and the basis size up to 2000 functions were reported in one of our previous works [25]. The initial selection of L_k parameters (and m_k for P states) was done using a stochastic scheme similar to that described in other works [26–28]. After each function was added to the set, its L_k parameters were optimized using a method that employs an analytical energy gradient determined with respect to the optimized parameters. After a subset of 10 functions were added to the basis, all functions in the set were reoptimized one by one using the gradient-based method. When the basis size reached 5000, we further increased the computational effort by changing the number of cycles performed after the addition of each new 10 basis functions from one to three. Also, starting from approximately the 5080 function basis, we switched from standard double precision (64-bit) to extended precision (80-bit) to improve the quality of the optimization, which is sensitive to the numerical accuracy of the computed eigenvalues and their derivatives.

The results of the calculations are presented in Table I. For the ^{11}B isotope, the energies obtained with 1000, 2000, 3000, 4000, 5000, and 5100 ECGFs are shown. This allows for an analysis of the energy convergence. We estimate that the energies for all considered states are converged to about 10^{-7} – 10^{-8} relative to the respective energy values. For ^{10}B and $^{\infty}\text{B}$, we only show the energies obtained with 5100 ECGFs. They are as well converged as the ^{11}B energies.

Using the total energies of the states from Table I, we can determine the energies corresponding to transitions between states. These, in turn, can be directly compared with the experimental data taken from Ref. [24]. Such a comparison is presented in Table II. First let us examine the convergence of the calculated transition energies. This analysis can be done based on the transition energy values obtained for the ^{11}B isotope for the different basis-set sizes. It shows that

TABLE I. The convergence of the total nonrelativistic energies (in Hartrees) with the basis size for the main isotope of boron atom ^{11}B . Energies obtained for ^{10}B and $^{\infty}\text{B}$ with 5100 basis functions are also shown. The values in parentheses are estimates of the remaining uncertainty due to finite basis size used in the calculations.

Isotope	Basis	$^2P^o (1s^22s^22p)$	$^2S (1s^22s^23s)$	$^2P^o (1s^22s^23p)$	$^2S (1s^22s^24s)$
^{11}B	1000	−24.652 494 17	−24.470 106 32	−24.430 979 54	−24.401 841 39
	2000	−24.652 598 09	−24.470 136 27	−24.431 073 08	−24.401 923 14
	3000	−24.652 615 70	−24.470 140 92	−24.431 088 39	−24.401 936 14
	4000	−24.652 621 14	−24.470 142 33	−24.431 093 26	−24.401 939 96
	5000	−24.652 623 43	−24.470 142 90	−24.431 095 31	−24.401 941 47
	5100 ^a	−24.652 623 87(250)	−24.470 143 16(50)	−24.431 095 74(250)	−24.401 941 83(150)
^{10}B	5100 ^a	−24.652 500 24(250)	−24.470 018 76(50)	−24.430 971 75(250)	−24.401 817 83(150)
$^{\infty}\text{B}$	5100 ^a	−24.653 866 08(250)	−24.471 393 06(50)	−24.432 341 44(250)	−24.403 187 67(150)

^aThis basis set was generated with a more extensive optimization of the nonlinear parameters.

TABLE II. Nonrelativistic transition frequencies (in cm^{-1}) for different isotopes of B compared to the experimental values.

Isotope	Basis	${}^2P^o(2p) \rightarrow {}^2S(3s)$	${}^2P^o(2p) \rightarrow {}^2P^o(3p)$	${}^2P^o(2p) \rightarrow {}^2S(4s)$	${}^2S(3s) \rightarrow {}^2P^o(3p)$	${}^2S(3s) \rightarrow {}^2S(4s)$	${}^2P^o(3p) \rightarrow {}^2S(4s)$
${}^{11}\text{B}$	1000	40 029.504	48 616.840	55 011.926	8587.336	14 982.422	6395.086
	2000	40 045.740	48 619.119	55 016.792	8573.378	14 971.051	6397.673
	3000	40 048.584	48 619.623	55 017.802	8571.039	14 969.218	6398.179
	4000	40 049.469	48 619.749	55 018.160	8570.280	14 968.691	6398.411
	5000	40 049.847	48 619.802	55 018.331	8569.956	14 968.484	6398.529
	5100 ^a	40 049.887(200)	48 619.806(50)	55 018.349(100)	8569.920(200)	14 968.462(100)	6398.542(70)
${}^{10}\text{B}$	5100 ^a	40 050.054(200)	48 619.883(50)	55 018.428(100)	8569.828(200)	14 968.374(100)	6398.545(70)
${}^{\infty}\text{B}$	5100 ^a	40 048.200(200)	48 619.039(50)	55 017.551(100)	8570.839(200)	14 969.351(100)	6398.513(70)
Experiment ^b	5100 ^a	40 039.650	48 611.817	55 010.181	8572.167	14 970.531	6398.364

^aThis basis set was generated with a more extensive optimization of the nonlinear parameters.

^bFor transitions frequencies involving P states, there is significant fine-structure splitting (up to several cm^{-1} in magnitude). In this table, we show the frequencies corresponding to the transitions involving lowest-lying sublevels.

our nonrelativistic transition energies are converged to about $0.05\text{--}0.20 \text{ cm}^{-1}$.

Let us now examine the agreement of the calculated transition energies with the experimental values. As these calculations have included neither relativistic nor QED corrections, one may expect some deviations between the calculated and the experimental values. They indeed appear, but never exceed about 10 cm^{-1} . It is interesting to examine the transition energies in the series of the following consecutive excitations: ${}^2P^o(2p) \rightarrow {}^2S(3s)$, ${}^2S(3s) \rightarrow {}^2P^o(3p)$, and ${}^2P^o(3p) \rightarrow {}^2S(4s)$. The corresponding calculated ${}^{11}\text{B}$ transition energies are $40\,049.887(200)$, $8569.920(200)$, and $6398.542(70) \text{ cm}^{-1}$, respectively. As those excitations correspond to promoting the valence electron between increasingly higher pairs of adjacent S and P states, one would expect to see the calculated transition energies to become closer to the experimental values. This is because the relativistic contribution is mainly due to the state of the core electrons and, as the valence electron becomes more removed from the atom (which happens when it gets excited to increasingly higher states), the state of the core is affected increasingly less. This means that the relativistic contribution to the transition energy should decrease for higher S - P and P - S transitions. With this, the calculated transition energies should become closer to the experimental transitions, and this

indeed is what happens in the series of the three mentioned transitions. The differences between the calculated and the experimental transition energy values are 10.237 , 7.753 , and 0.178 cm^{-1} , respectively. As one sees, the transition energy for the highest transition is very close to the experimental value.

As the goal of this work has been to provide more accurate calculated reference values for boron, aside from the energy we also calculated some other expectation values that provide some additional characterization of the wave functions of the considered states. The expectations values are shown in Table III. They have been calculated using the basis sets of 5100 ECGFs. Let us first comment on the expectation value of the nucleus-electron distance $\langle r_i \rangle$. It characterizes the spatial extent of the wave function. As expected, as the level of excitation increases, $\langle r_i \rangle$ gets bigger. For the highest ${}^2S(1s^22s^24s)$ state, it is almost three times larger than for the ground ${}^2P(1s^22s^22p)$ state. Also, $\langle r_i \rangle$ shows that, when the nuclear mass increases, the electron density contracts. This contraction is smaller for the ground state than for the higher states. However, interestingly, the contraction does not increase monotonically with the level of excitation. It is slightly larger for the ${}^2S(1s^22s^23s)$ state than for the ${}^2P(1s^22s^23p)$ state, even though the latter has higher energy.

TABLE III. Expectation values of powers of the inter-particle distances and contact densities. All values are in a.u. The results obtained with 5100 function basis sets are shown. The estimated uncertainties due to finite size of the basis are given in parentheses.

State	Isotope	$\langle 1/r_i \rangle$	$\langle 1/r_{ij} \rangle$	$\langle r_i \rangle$	$\langle r_{ij} \rangle$	$\langle r_i^2 \rangle$	$\langle r_{ij}^2 \rangle$	$\langle \delta(\mathbf{r}_i) \rangle$	$\langle \delta(\mathbf{r}_{ij}) \rangle$
${}^2P^o(1s^22s^22p)$	${}^{10}\text{B}$	2.278 870 27(5)	0.766 675 6(4)	1.348 148(2)	2.245 196(5)	3.101 86(2)	6.707 26(5)	14.3659(30)	0.353 99(5)
	${}^{11}\text{B}$	2.278 881 37(5)	0.766 678 6(4)	1.348 143(2)	2.245 188(5)	3.101 84(2)	6.707 21(5)	14.3661(30)	0.353 99(5)
	${}^{\infty}\text{B}$	2.278 992 96(5)	0.766 709 2(4)	1.348 088(2)	2.245 104(5)	3.101 60(2)	56.70672(5)	14.3683(30)	0.354 04(5)
${}^2S(1s^22s^23s)$	${}^{10}\text{B}$	2.225 981 84(20)	0.670 950 9(3)	2.103 834(4)	3.660 657(8)	10.673 68(3)	21.771 87(6)	14.5047(30)	0.358 29(9)
	${}^{11}\text{B}$	2.225 993 01(20)	0.670 953 9(3)	2.103 824(4)	3.660 639(8)	10.673 57(3)	21.771 64(6)	14.5049(30)	0.358 29(9)
	${}^{\infty}\text{B}$	2.226 105 21(20)	0.670 984 4(3)	2.103 715(4)	3.660 453(8)	10.672 44(3)	21.769 39(6)	14.5071(30)	0.358 34(9)
${}^2P^o(1s^22s^23p)$	${}^{10}\text{B}$	2.217 176 77(30)	0.656 747 6(4)	2.483 814(15)	4.404 871(30)	17.067 35(10)	34.544 84(20)	14.4788(10)	0.357 55(5)
	${}^{11}\text{B}$	2.217 187 83(30)	0.656 750 5(4)	2.483 806(15)	4.404 856(30)	17.067 24(10)	34.544 63(20)	14.4790(10)	0.357 55(5)
	${}^{\infty}\text{B}$	2.217 298 94(30)	0.656 779 1(4)	2.483 719(15)	4.404 712(30)	17.066 21(10)	34.542 57(20)	14.4812(10)	0.357 60(5)
${}^2S(1s^22s^24s)$	${}^{10}\text{B}$	2.206 234 20(80)	0.635 222 1(15)	3.598 680(50)	6.611 520(100)	44.840 99(100)	89.997 69(300)	14.4906(30)	0.357 90(7)
	${}^{11}\text{B}$	2.206 245 18(80)	0.635 224 7(15)	3.598 666(50)	6.611 495(100)	44.840 61(100)	89.996 94(300)	14.4908(30)	0.357 90(7)
	${}^{\infty}\text{B}$	2.206 355 57(80)	0.635 251 6(15)	3.598 524(50)	6.611 241(100)	44.836 85(100)	89.989 43(300)	14.4930(30)	0.357 95(7)

Another property worth noting is the electron density at the nucleus [i.e., the $\langle\delta(\mathbf{r}_i)\rangle$ expectation value]. With the increasing level of excitation this density increases, which is understandable because, as the valence electron moves further away from the nucleus, the core electrons contract leading to their higher density at the nucleus. Also, for all the states, the higher nuclear mass leads to a slightly higher $\langle\delta(\mathbf{r}_i)\rangle$.

IV. SUMMARY

In this work, we obtained the four lowest 2S and 2P states of the two stable isotopes of the boron atom ^{10}B

and ^{11}B . High accuracy has been achieved by employing large basis sets of explicitly correlated Gaussian functions and optimizing their nonlinear parameters with an approach that utilizes the analytical gradient of the energy determined with respect to those parameters. The results demonstrate that five-electron atomic systems can now be calculated with a comparable accuracy as for atoms with four electrons. The lack of accounting for relativistic and QED effects results in some discrepancies between the calculated and the experimental transition energies. Thus, the next task in our work will be to develop algorithms for computing the relativistic corrections.

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