

# THE UNIVERSITY of EDINBURGH

## Edinburgh Research Explorer

## Formation and Stability of Dense Methane-Hydrogen Compounds

#### Citation for published version:

Ranieri, U, Conway, LJ, Donnelly, M-E, Hu, H, Wang, M, Dalladay-Simpson, P, Pena-Alvarez, M, Gregoryanz, E, Hermann, A & Howie, RT 2022, 'Formation and Stability of Dense Methane-Hydrogen Compounds', *Physical Review Letters*, vol. 128, no. 21, 215702, pp. 1-6. https://doi.org/10.1103/PhysRevLett.128.215702

#### **Digital Object Identifier (DOI):**

10.1103/PhysRevLett.128.215702

#### Link:

Link to publication record in Edinburgh Research Explorer

**Document Version:** Peer reviewed version

**Published In: Physical Review Letters** 

#### **General rights**

Copyright for the publications made accessible via the Edinburgh Research Explorer is retained by the author(s) and / or other copyright owners and it is a condition of accessing these publications that users recognise and abide by the legal requirements associated with these rights.

Take down policy The University of Edinburgh has made every reasonable effort to ensure that Edinburgh Research Explorer content complies with UK legislation. If you believe that the public display of this file breaches copyright please contact openaccess@ed.ac.uk providing details, and we will remove access to the work immediately and investigate your claim.



## **Important Notice to Authors**

No further publication processing will occur until we receive your response to this proof.

Attached is a PDF proof of your forthcoming article in *Physical Review Letters*. The article accession code is LL17763.

Your paper will be in the following section of the journal: LETTERS - Condensed Matter: Structure, etc.

Please note that as part of the production process, APS converts all articles, regardless of their original source, into standardized XML that in turn is used to create the PDF and online versions of the article as well as to populate third-party systems such as Portico, Crossref, and Web of Science. We share our authors' high expectations for the fidelity of the conversion into XML and for the accuracy and appearance of the final, formatted PDF. This process works exceptionally well for the vast majority of articles; however, please check carefully all key elements of your PDF proof, particularly any equations or tables.

Figures submitted electronically as separate files containing color appear in color in the online journal.

However, all figures will appear as grayscale images in the print journal unless the color figure charges have been paid in advance, in accordance with our policy for color in print (https://journals.aps.org/authors/color-figures-print).

#### Specific Questions and Comments to Address for This Paper

The numbered items below correspond to numbers in the margin of the proof pages pinpointing the source of the question and/or comment. The numbers will be removed from the margins prior to publication.

- Please review the funding information section of the proof's cover letter and respond as appropriate. We must receive confirmation that the funding agencies have been properly identified before the article can publish.
- 2 This query was generated by an automatic reference checking system. References [18,25,56,59,62,63] could not be located in the databases used by the system. While the references may be correct, we ask that you check them so we can provide as many links to the referenced articles as possible.
- **3** NOTE: External links, which appear as blue text in the reference section, are created for any reference where a Digital Object Identifier (DOI) can be found. Please confirm that the links created in this PDF proof, which can be checked by clicking on the blue text, direct the reader to the correct references online. If there is an error, correct the information in the reference or supply the correct DOI for the reference. If no correction can be made or the correct DOI cannot be supplied, the link will be removed.
- 4 Please note that the references have been reordered so that they are cited in the text in numerical order.
- 5 Please check and confirm updated for Ref. [44].
- Please note that Ref. [45] have been set up in the style used by APS for Supplemental Material references. The URL link will be activated at the time of publication.

#### **Titles in References**

The editors now encourage insertion of article titles in references to journal articles and e-prints. This format is optional, but if chosen, authors should provide titles for *all* eligible references. If article titles remain missing from eligible references, the production team will remove the existing titles at final proof stage.

### ORCIDs

Please follow any ORCID links ((b)) after the authors' names and verify that they point to the appropriate record for each author.

### **Crossref Funder Registry ID:**

Information about an article's funding sources is now submitted to Crossref to help you comply with current or future funding agency mandates. Crossref's Funder Registry (https://www.crossref.org/services/funder-registry/) is the definitive registry of funding agencies. Please ensure that your acknowledgments include all sources of funding for your article following any requirements of your funding sources. Where possible, please include grant and award ids. Please carefully check the following funder information we have already extracted from your article and ensure its accuracy and completeness:

- European Synchrotron Radiation Facility, FundRef ID http://dx.doi.org/10.13039/501100001671 (Republic of France/ FR)
- H2020 European Research Council, FundRef ID http://dx.doi.org/10.13039/100010663 (European Union/EU)
- Horizon 2020 Framework Programme, FundRef ID http://dx.doi.org/10.13039/100010661 (European Union/EU)
- UK Research and Innovation, FundRef ID http://dx.doi.org/10.13039/100014013 (United Kingdom of Great Britain and Northern Ireland/GB)
- United Kingdom Car-Parrinello consortium
- United Kingdom Materials and Molecular Modelling Hub

#### **Other Items to Check**

- Please note that the original manuscript has been converted to XML prior to the creation of the PDF proof, as described above. Please carefully check all key elements of the paper, particularly the equations and tabular data.
- Title: Please check; be mindful that the title may have been changed during the peer-review process.
- Author list: Please make sure all authors are presented, in the appropriate order, and that all names are spelled correctly.
- Please make sure you have inserted a byline footnote containing the email address for the corresponding author, if desired. Please note that this is not inserted automatically by this journal.
- Affiliations: Please check to be sure the institution names are spelled correctly and attributed to the appropriate author(s).
- Receipt date: Please confirm accuracy.
- Acknowledgments: Please be sure to appropriately acknowledge all funding sources.
- References: Please check to ensure that titles are given as appropriate.
- Hyphenation: Please note hyphens may have been inserted in word pairs that function as adjectives when they occur before a noun, as in "x-ray diffraction," "4-mm-long gas cell," and "*R*-matrix theory." However, hyphens are deleted from word pairs when they are not used as adjectives before nouns, as in "emission by

x rays," "was 4 mm in length," and "the *R* matrix is tested."

Note also that Physical Review follows U.S. English guidelines in that hyphens are not used after prefixes or before suffixes: superresolution, quasiequilibrium, nanoprecipitates, resonancelike, clockwise.

- Please check that your figures are accurate and sized properly. Make sure all labeling is sufficiently legible. Figure quality in this proof is representative of the quality to be used in the online journal. To achieve manageable file size for online delivery, some compression and downsampling of figures may have occurred. Fine details may have become somewhat fuzzy, especially in color figures. The print journal uses files of higher resolution and therefore details may be sharper in print. Figures to be published in color online will appear in color on these proofs if viewed on a color monitor or printed on a color printer.
- Overall, please proofread the entire *formatted* article very carefully. The redlined PDF should be used as a guide to see changes that were made during copyediting. However, note that some changes to math and/or layout may not be indicated.

### Ways to Respond

• Web: If you accessed this proof online, follow the instructions on the web page to submit corrections.

- Email: Send corrections to aps-robot@luminad.com. Include the accession code LL17763 in the subject line.
- Fax: Return this proof with corrections to +1.855.808.3897.

#### If You Need to Call Us

You may leave a voicemail message at +1.855.808.3897. Please reference the accession code and the first author of your article in your voicemail message. We will respond to you via email.

**Editors' Suggestion** 

1

2

3 4

5 6

7

8

9

10

11 12

13

14 15

16

17

#### Formation and Stability of Dense Methane-Hydrogen Compounds

Umbertoluca Ranieri<sup>®</sup>,<sup>1,2</sup> Lewis Conway<sup>®</sup>,<sup>3</sup> Mary-Ellen Donnelly,<sup>1</sup> Huixin Hu,<sup>1</sup> Mengnan Wang,<sup>1</sup> Philip Dalladay-Simpson,<sup>1</sup> Miriam Peña-Alvarez<sup>®</sup>,<sup>3</sup> Eugene Gregoryanz,<sup>1,3,4</sup> Andreas Hermann<sup>®</sup>,<sup>3</sup> and Ross T. Howie<sup>1,3,\*</sup>

<sup>1</sup>Center for High Pressure Science and Technology Advanced Research, 1690 Cailun Road, Shanghai, 201203, China

<sup>2</sup>Dipartimento di Fisica, Università di Roma La Sapienza, Piazzale Aldo Moro 5, 00185 Rome, Italy

<sup>3</sup>Centre for Science at Extreme Conditions and The School of Physics and Astronomy, The University of Edinburgh,

<sup>4</sup>Key Laboratory of Materials Physics, Institute of Solid State Physics, CAS, Hefei, China

(Received 25 November 2021; revised 2 February 2022; accepted 20 April 2022)

Through a series of x-ray diffraction, optical spectroscopy diamond anvil cell experiments, combined with density functional theory calculations, we explore the dense  $CH_4$ - $H_2$  system. We find that pressures as low as 4.8 GPa can stabilize  $CH_4(H_2)_2$  and  $(CH_4)_2H_2$ , with the latter exhibiting extreme hardening of the intramolecular vibrational mode of H<sub>2</sub> units within the structure. On further compression, a unique structural composition,  $(CH_4)_3(H_2)_{25}$ , emerges. This novel structure holds a vast amount of molecular hydrogen and represents the first compound to surpass 50 wt % H<sub>2</sub>. These compounds, stabilized by nuclear quantum effects, persist over a broad pressure regime, exceeding 160 GPa.

18

DOI:

Hydrogen and methane are, besides water, the most 19 prevalent small molecules in the outer Solar System. Their 20 interplay under extreme conditions is of key interest to 21 22 understanding the evolution and interior dynamics of Neptune and Uranus, as well as of Earth and exoplanets 23 [1-4]. The simplest molecule, H<sub>2</sub>, has been shown to 24 exhibit rich physical phenomena, including quantum 25 rotational phases and pressure-induced steps toward an 26 27 atomic metallic solid, while methane  $(CH_4)$ , the simplest 28 hydrocarbon, has been proposed to polymerize during compression to form long-chain hydrocarbons and at 29 further extremes decompose into diamond plus  $H_2$  [5–19]. 30

Hydrogen reacts with a number of materials at high 31 32 pressures and temperatures but also has a propensity to form van der Waals compounds that can be stable far beyond 33 100 GPa [20–32]. Methane and hydrogen were first reported 34 35 to crystallize into inclusion (host-guest) compounds over 36 25 years ago at pressures between 5 and 7 GPa. A range of 37 compositions were claimed:  $(CH_4)_2H_2$ ,  $CH_4H_2$ ,  $CH_4(H_2)_2$ , and  $CH_4(H_2)_4$  [21]. However, the experimental evidence for 38 these compounds was limited, with only the shift of the 39 40 hydrogen intramolecular vibrational mode  $\nu_1$ -H<sub>2</sub> (vibron) relative to that of pure hydrogen being reported together with 41 42 the lattice parameters of potential structures. Assuming that these structural compositions were correct, a later exper-43 imental study explored the recoverability of  $CH_4(H_2)_4$  at 44 low temperature to investigate its potential as a hydrogen 45 46 storage media [33,34].

In contrast to experimental work, the CH<sub>4</sub>-H<sub>2</sub> system has 47 recently had a surge of theoretical investigations, predicting a 48 49 variety of extremely hydrogen-rich compositions to emerge 50 such as triclinic  $(CH_4)_2(H_2)_3$ , trigonal  $(CH_4)_2(H_2)_7$ , and hexagonal  $CH_4(H_2)_9$ , as well as several  $CH_4H_2$  structures [18,35-37]. Furthermore, it was suggested that CH<sub>4</sub>-H<sub>2</sub> structures could be the most stable form of carbon and hydrogen at pressures up to 200 GPa [18].

51

52

53

54

55

56

57

58

59

60

61

62

63

64

65

66

67

68

69

70

Surprisingly, given these predictions and the advent of technological advances over the past 25 years, no further experimental exploration has been conducted despite undetermined CH<sub>4</sub>-H<sub>2</sub> compound signatures emerging as a by-product in many prolific studies of hydrocarbons at planetary conditions [3,19,38–41]. Moreover, the doping of carbon (or methane) could even enhance the properties of materials at extreme pressures, with the unprecedented claim of room temperature superconductivity in the carbonaceous sulfur hydride system [42–44]. The lack of knowledge regarding CH<sub>4</sub>-H<sub>2</sub> compound formation inhibits our understanding of these more complex ternary systems and how doping could induce novel properties. As such, it is of fundamental interest to know which methane-hydrogen compounds are actually formed and test their pressure stability limits.

In this Letter, we have conducted a series of high-71 pressure synchrotron powder x-ray diffraction and Raman 72 spectroscopy experiments in diamond anvil cells com-73 bined with density functional theory (DFT) calculations 74 and structural searching to explore the formation and 75 properties of CH<sub>4</sub>-H<sub>2</sub> compounds from mixtures of 76 methane and hydrogen. In H<sub>2</sub>-rich mixtures, we observe 77 the formation of hexagonal  $CH_4(H_2)_2$  above 5 GPa, 78 before partially transforming into monoclinic  $(CH_4)_3$ 79  $(H_2)_{25}$ .  $(CH_4)_3(H_2)_{25}$  represents a unique composition 80 and contains an unprecedented 51.1 wt % H<sub>2</sub>, which is 81 the highest hydrogen content of any currently known 82

Peter Guthrie Tait Road, Edinburgh, United Kingdom



F1:1 FIG. 1. Pressure-composition phase diagram of the CH<sub>4</sub>-H<sub>2</sub> F1:2 binary system. Filled square symbols indicate the formation F1:3 pressures of each compound:  $(CH_4)_2H_2$  (orange),  $CH_4(H_2)_2$ F1:4 (purple), and  $(CH_4)_3(H_2)_{25}$  (green). The green open squares F1:5 represent the pressures at which  $(CH_4)_3(H_2)_{25}$  is first observed. F1:6 The black squares, black triangle, and blue circle represent F1:7 the formation pressure of  $CH_4$ -I,  $CH_4$ -A, and  $H_2$ -I, respectively. F1:8 The gray line represents the liquidus curve and is adapted from F1:9 Ref. [21]. Below 7 GPa, the error in pressure is  $\pm 0.2$  GPa and F1:10 smaller than the size of the symbol. Initial gas mixtures have a F1:11 tolerance of 1%.

stoichiometric compound. In CH<sub>4</sub>-rich mixtures, tetragonal (CH<sub>4</sub>)<sub>2</sub>H<sub>2</sub> forms, which undergoes extreme hardening
of the H<sub>2</sub> intramolecular vibrational mode with pressure.
All three compounds exhibit remarkable stability, exceeding pressures of 160 GPa.

At pressures below 1.5 GPa, all mixture concentrations 88  $(30, 50, 70, 80, and 90 \text{ mol }\% \text{ H}_2)$  are homogeneous well-89 90 mixed fluids and display the Raman signatures of both 91 constituent species (see Fig. 1 for the pressure-composition 92 phase diagram, Supplemental Material [45] for experimental methods, and Figs. S1-S9 [45] for Raman spectra of all 93 mixtures). On compression of hydrogen-rich CH<sub>4</sub>-H<sub>2</sub> mix-94 tures (50–90 mol %  $H_2$ ) above 5 GPa, the fluid mixture 95 crystallizes into a solid which we identify as  $CH_4(H_2)_2$ . 96 97 X-ray diffraction measurements reveal that this compound adopts a hexagonal MgZn<sub>2</sub> Laves phase structure (space 98 group  $P6_3/mmc$ ), where the CH<sub>4</sub> molecules occupy the Mg 99 100 sites and the H<sub>2</sub> molecules occupy the Zn sites (see Fig. 2), the lattice parameters of which are a = 4.981 Å and 101 c = 8.125 Å at 16.6 GPa. This structure and composition 102 was suggested previously but attributed to a compound 103 104 forming in a narrow pressure regime (between approxi-105 mately 6 and 7 GPa) and between 35 and 65 mol % H<sub>2</sub> mixture concentrations [21]. Here, we find that  $CH_4(H_2)_2$  is 106 the dominant phase across all H<sub>2</sub>-rich mixtures and persis-107 108 tent over a large pressure regime. Raman spectroscopy reveals two distinct H<sub>2</sub> vibrons,  $\nu_{1H}$  and  $\nu_{2H}$ , the latter of which is approximately 50 times lower in intensity. The two most intense CH<sub>4</sub> stretching modes,  $\nu_{1M}$  and  $\nu_{2M}$ , are shifted to higher frequency compared to pure CH<sub>4</sub>, while the wagging mode is shifted to lower frequency (see Figs. 3 and 4).

109

110

111

112

113

114

On further compression, we observe another compound 115 emerge evidenced by visual changes in the sample mor-116 phology and the appearance of three distinct H<sub>2</sub> vibrational 117 modes (the lowest in frequency being the most intense by 2 118 orders of magnitude) and CH<sub>4</sub> stretching bands that can be 119 isolated from  $CH_4(H_2)_2$  (see Figs. 3 and 4 and Supplemental 120 Material [45]). The most intense  $H_2$  vibron is lower in 121 frequency than that of  $CH_4(H_2)_2$ , suggestive that the 122 compound has a higher H<sub>2</sub> content. X-ray diffraction 123 measurements indicate the compound is weakly scattering; 124 however, this could be in part due to only partial trans-125 formation of the sample. The patterns could be indexed to a 126 hexagonal cell, with a = 7.804 Å and c = 11.199 Å at 127 17.3 GPa (see Fig. 2). Systematic absences, unit cell volume, 128 and c/a ratio suggest this compound has a close resem-129 blance with the previously reported R-3m structure of 130  $Xe(H_2)_8$  [23,26] and an initial composition estimate of 131  $CH_4(H_2)_8$ . This seemed plausible given  $CH_4$  and Xe have 132 similar van der Waals diameters (3.78 and 4.32 Å, 133 respectively). 134

We subsequently performed DFT calculations to help 135 identify this phase. We constructed a series of structures of 136 variable composition, starting by populating the hexago-137 nal cell with randomly oriented CH<sub>4</sub> molecules on the Xe 138 sites of  $Xe(H_2)_8$ , before adding  $H_2$  molecules on hex-139 140 agonal close packed (hcp) arrangements commensurate with the cell, using up to  $3 \times 3 \times 4$  repeats of the hcp 141 lattice and a random offset against the CH<sub>4</sub> sublattice. H<sub>2</sub> 142 molecules too close to CH<sub>4</sub> were removed before fully 143 optimizing the remaining atoms and all lattice parameters. 144 Several thousand of these structures were generated with 145 stoichiometries ranging from  $CH_4(H_2)_7$  to  $CH_4(H_2)_{11}$ . 146 Of these,  $CH_4(H_2)_{8,33}$  [or  $(CH_4)_3(H_2)_{25}$ ] emerged as the 147 most energetically competitive and was found to be 148 dynamically stable at and above 20 GPa (see Figs. S19 149 and S20 [45]). It is also the only computationally obtained 150 structure which reproduces the positions and intensities of 151 the experimental diffraction peaks across the studied 152 pressure range (see Fig. S14 [45]), which is meaningful 153 given that no symmetry restrictions were applied in the 154 structures' construction and subsequent optimization. While 155 we find both  $CH_4(H_2)_2$  and  $(CH_4)_2H_2$  stable or close to 156 stability within the CH<sub>4</sub>-H<sub>2</sub> phase diagram across a wide 157 pressure range, once zero point energies and vibrational 158 entropies are accounted for,  $(CH_4)_3(H_2)_{25}$  remains meta-159 stable throughout (see Supplemental Material [45]). 160

Considering methane and hydrogen molecules to be 161 spherical, the structure of  $(CH_4)_3(H_2)_{25}$  is monoclinic with 162 space group C2/m (coordinates are given in Supplemental 163



FIG. 2. (a) Representative x-ray diffraction patterns of the three compounds plotted as a function of exchanged wave vector and their F2:1 F2:2 Le Bail refinements. Refinements include  $(CH_4)_3(H_2)_{25}$ -R-3m (a = 7.804 Å and c = 11.199 Å),  $(CH_4)_2H_2$ -I4/mcm (a = 7.195 Å and c = 5.909 Å), CH<sub>4</sub>-A-R3 (a = 12.306 Å and c = 15.520 Å), CH<sub>4</sub>(H<sub>2</sub>)<sub>2</sub>-P6<sub>3</sub>/mmc (a = 4.981 Å and c = 8.125 Å), and F2:3 Au-Fm-3m (a = 3.976 Å). Excluded regions in the pattern of  $(CH_4)_3(H_2)_{25}$  correspond to rhenium (gasket material) and rhenium F2:4 hydride (\*) and excess H<sub>2</sub> (#). (b) Volumes per formula unit as a function of pressure. Symbols represent experimental data (the error F2:5 bars are smaller than the symbol size), and black full lines represent their best second-order Birch-Murnaghan fits: CH<sub>4</sub>(H<sub>2</sub>)<sub>8.33</sub>, F2:6  $V_0 = 246 \pm 17$  Å<sup>3</sup> and  $K = 3.0 \pm 0.6$  GPa; (CH<sub>4</sub>)<sub>2</sub>H<sub>2</sub>,  $V_0 = 101 \pm 2$  Å<sup>3</sup> and  $K = 10.2 \pm 0.8$  GPa; CH<sub>4</sub>(H<sub>2</sub>)<sub>2</sub>,  $V_0 = 90 \pm 3$  Å<sup>3</sup> and F2:7  $K = 5.0 \pm 0.4$  GPa. Dashed lines represent volumes derived from our DFT calculations, and gray dotted lines represent the volumes of F2:8 ideal mixtures of  $CH_4$ - $A/CH_4$ -B and  $H_2$ -I using the previously determined equations of state [14,17,69]. The experimental volumes F2:9 F2:10 given in Ref. [21] are represented by gray circles. (c) Structural models of the three compounds, where  $CH_4$  and  $H_2$  are represented by brown and white spheres, respectively, and lines indicate CH<sub>4</sub>-H<sub>2</sub> nearest neighbors. F2:11



F3:1 FIG. 3. Representative vibrational Raman spectra of (a) 70%  $H_2$ , (b) 50%  $H_2$ , and (c) 30%  $H_2$  mixtures. Colors indicate the modes F3:2 assigned to  $CH_4(H_2)_2$  (purple),  $(CH_4)_3(H_2)_{25}$  (green),  $(CH_4)_2H_2$  (orange), and excess  $H_2$  (black).



F4:1 FIG. 4. Raman shift as a function of pressure of the H<sub>2</sub> F4:2 vibrational modes and C-H stretching modes of CH<sub>4</sub>(H<sub>2</sub>)<sub>2</sub> and F4:3  $(CH_4)_3(H_2)_{25}$  (left panel, purple and green symbols, respectively), and  $(CH_4)_2H_2$  (right panel, orange symbols). Frequencies F4:4 F4:5 are collated from 30% (squares), 50% (circles), 70% (triangles), F4:6 80% (diamonds), and 90% (hexagons) H<sub>2</sub> mixtures, with the F4:7 most intense modes represented by larger symbols. Raman F4:8 frequencies of individual mixtures are given in Supplemental F4:9 Material [45]. The dot-dashed lines represent the calculated F4:10 frequencies of the H<sub>2</sub> modes and C-H stretching bands of each F4:11 compound, with the thickest line representing the most intense F4:12 modes. Black and gray solid lines represent pure  $H_2$  [6] and  $CH_4$ F4:13 [15], respectively. The black dotted line represents the frequency F4:14 of  $H_2$  in a Ne matrix [70]. The error in pressure ranges from  $\pm 0.2$ below 50 GPa,  $\pm 1$  below 150 GPa, and  $\pm 5$  above. The error bars F4:15 F4:16 in frequency are smaller than the symbol size. Photomicrographs F4:17 show examples of morphology of synthesised samples within the F4:18 DAC sample chamber.

Material [45]). Le Bail refinements of the experimental 164 165 diffraction patterns were performed using the trigonal space group R-3m, which is obtained considering the carbon atoms 166 only, and the hexagonal setting, which resulted in good fits 167 to the data. The obtained volume per methane molecule is 168 close to  $V_{CH_4}$  + 8.33 $V_{H_2}$  (see Fig. 2), the volume of an ideal 169 mixture of CH<sub>4</sub> and H<sub>2</sub> with the same composition. A 170 comparison between the experimental and computational 171 vibrational Raman modes of  $CH_4(H_2)_2 - PG_3/mmc$  and 172 173  $(CH_4)_3(H_2)_{25}$ -C2/m is given in Fig. 4, where we see qualitatively good agreement between the number of modes, 174 relative intensity, and frequency for both compounds. 175

176 In methane-rich concentrations of 30 mol % H<sub>2</sub>, we 177 observe a body-centered tetragonal Al<sub>2</sub>Cu-type structure (space group I4/mcm) to form above 4.6 GPa (with a =178 7.195 Å and c = 5.909 Å at 5.4 GPa), which is consistent 179 180 with the previously reported composition of  $(CH_4)_2H_2$  [21] (see Fig. 1). Raman spectroscopy measurements show that 181 182 the CH<sub>4</sub> stretching bands of  $(CH_4)_2H_2$  are close to the 183 frequencies of pure solid CH<sub>4</sub>, while two wagging modes are observed. The number of modes and frequencies are in good 184 agreement with our calculated values (see Fig. 4). The  $\nu_{1H}$ 185 mode, corresponding to H<sub>2</sub> units within the structure, 186 exhibits the most extreme hardening of any stoichiometric 187 molecular compound [25,28]. At a pressure of 207 GPa, 188  $(CH_4)_2H_2-\nu_{1H}$  has a frequency of 4798 cm<sup>-1</sup>, over 189 900 cm<sup>-1</sup> higher than H<sub>2</sub>- $\nu_1$  (see Fig. 4). Remarkably, the 190 frequencies we observe reach similar values to an impurity 191 H<sub>2</sub> molecule isolated in a noble gas matrix, albeit at higher 192 pressure [70]. Up to pressures of 207 GPa, we do not observe 193 a turnover of  $\nu_{1H}$  (and 500 GPa in our calculations), unlike in 194 pure  $H_2$ , which exhibits a maximum frequency at approx-195 imately 38 GPa [5]. 196

These results highlight the enormous impact that the 197 local environment around a H<sub>2</sub> molecule can have on its 198 vibron frequency. In Supplemental Material [45], we 199 present a simple molecular model demonstrating that 200 the vibron frequency of a H<sub>2</sub> surrounded by two CH<sub>4</sub> 201 (respectively, H<sub>2</sub>) will continuously increase (respectively, 202 decrease) if the surrounding molecules are pushed closer. 203 A molecular orbital analysis reveals that, in H<sub>2</sub>-dominated 204 environments, intermolecular interactions lead to occu-205 pancy of antibonding H-H  $\sigma^*$  states, which does not 206 happen in  $CH_4$ -dominated environments. In  $(CH_4)_2H_2$ , 207 every H<sub>2</sub> has two nearest H<sub>2</sub> neighbors and eight nearest 208 CH<sub>4</sub> neighbors; as a result, the vibron exhibits extreme 209 hardening [see Fig. 2(c) and Supplemental Material [45] 210 for graphical representations of the unit cells and inter-211 molecular distances]. This local environment is reversed 212 in  $(CH_4)_3(H_2)_{25}$  with an average of 9.8 H<sub>2</sub> neighbors but 213 only 2.24 CH<sub>4</sub> neighbors per H<sub>2</sub> molecule; still, it remains 214 very different from H<sub>2</sub>-I, having a coordination number of 215 12 and considerably shorter H<sub>2</sub>-H<sub>2</sub> distances [28]. For 216  $(CH_4)_3(H_2)_{25}$ , the  $\nu_{1H}$  mode has a maximum at about 217 70 GPa, and the vibron frequencies tend toward the values 218 found for pure H<sub>2</sub>. In CH<sub>4</sub>(H<sub>2</sub>)<sub>2</sub>, where each H<sub>2</sub> has six 219 nearest H<sub>2</sub> neighbors and six nearest CH<sub>4</sub> neighbors, the 220 vibron frequencies are interpolated between  $(CH_4)_3(H_2)_{25}$ 221 222 and  $(CH_4)_2H_2$ .

We find that the pressure-composition phase diagram is 223 simpler than previously reported [21], with only three 224 stable CH<sub>4</sub>-H<sub>2</sub> compositions. In 50 mol % H<sub>2</sub> mixtures, 225 we observe the coexistence of both  $(CH_4)_2H_2$  and 226  $CH_4(H_2)_2$ , together with some excess  $H_2$ , indicative that 227 equilibrium has not been reached. Furthermore, we observe 228 trace amounts of  $(CH_4)_3(H_2)_{25}$ , which can be attributed to 229 macroscopic inhomogeneity within the sample chamber. In 230 70 mol % mixtures and 80 mol %  $H_2$  mixtures, we first 231 observe the Raman signatures of  $(CH_4)_3(H_2)_{25}$  at 14.5 GPa 232 and in 13.6 GPa (shown as open symbols in Fig. 1), 233 respectively, while in 90 mol % mixtures, we observe 234  $(CH_4)_3(H_2)_{25}$  at 10.9 GPa. It is possible that, in 70 mol 235 % mixtures and 80 mol % H<sub>2</sub> mixtures, formation occurs at 236 lower pressure but is below the detection limit in the Raman 237 spectra. We find that even in 90 mol% mixtures the 238 239 transformation from  $CH_4(H_2)_2$  and  $H_2$  to  $(CH_4)_3(H_2)_{25}$  is never complete, suggestive that the transformation is 240 kinetically sluggish and due to hindered diffusion of H<sub>2</sub> 241 242 in the solid state. Although our liquidus curve is in good agreement with that reported previously, we do not observe 243 244 either  $CH_4H_2$  and  $CH_4(H_2)_4$  and attribute this to a misinterpretation of  $CH_4(H_2)_2$  [21]. The experimental volume 245 given for hexagonal wurtzite CH4H2 is identical to our 246 experimental volume for  $CH_4(H_2)_2$  at the same given 247 248 pressure. Furthermore, after extensive structure searches, with randomly orientated molecules in the reported wurt-249 250 zite center-of-mass positions, we do not find any version of the hexagonal wurtzite structure to be dynamically stable. 251 The evidence for  $CH_4(H_2)_4$  was predominantly based on 252 253 Raman spectroscopy, and we find to be a misinterpretation 254 of the coexistence between the CH<sub>4</sub>-H<sub>2</sub> mixed fluid and solid  $CH_4(H_2)_2$  (see Fig. S8 [45]). 255

 $(CH_4)_3(H_2)_{25}$  possesses 51.1 wt % molecular hydrogen 256 (63.4 wt % if the hydrogen of methane is included), which is 257 the highest hydrogen content of any currently known 258 stoichiometric compound. Notwithstanding the good agree-259 ment between the experimental and calculated data, 260  $(CH_4)_3(H_2)_{25}$  represents an unusual and unique composi-261 262 tion. All three compounds exhibit remarkable stability, with 263 the Raman signatures of  $(CH_4)_2H_2$  and  $CH_4(H_2)_2$  detectable to at least 215 GPa and  $(CH_4)_3(H_2)_{25}$  up to 165 GPa. 264 We estimate the melting temperatures  $T_m$  of  $(CH_4)_2H_2$  and 265  $CH_4(H_2)_2$  using Lindemann's equation based on calculated 266 267 Debye temperatures and found  $T_m$  significantly higher than molecular hydrogen at all pressures (see Supplemental 268 269 Material [45]). This would imply that hydrogen mixed with methane would potentially not possess the melting line 270 turnover that is observed in pure hydrogen [7]. Taken 271 272 together in a planetary context, this could influence critical 273 properties of planetary matter such as thermal conductivities and viscosities. In a materials science context, high Debye 274 temperatures coupled with extremely large vibron frequen-275 cies  $\nu_H$  are promising ingredients for high- $T_c$  superconduc-276 tivity—provided a reasonable density of states  $N(E_F)$  at the 277 Fermi energy. The molecular compounds described here are 278 279 insulators (all remain transparent in the visible up to the 280 highest pressures reached), but the presence of an electron or hole dopant would metallize the system, with partially 281 charged entities  $H_2^{\pm\delta}$  present. This motif, at least in calcu-282 lations, can lead to  $T_c$  close to 200 K [71]. 283

Parts of this research were carried out at P02.2 at DESY, a 284 1 member of the Helmholtz Association (HGF), and we thank 285 H.-P. Liermann and K. Glazyrin for assistance. We also 286 acknowledge the European Synchrotron Radiation Facility 287 for provision of synchrotron radiation facilities at the ID15B 288 beam line and assistance from M. Hanfland and D. 289 Comboni. R. T. H. acknowledges that the project has 290 received funding from the European Research Council 291 292 (ERC) under the European Union's Horizon 2020 research and innovation program (Grant Agreement No. 948895 293

"MetElOne"). M.P.-A. acknowledges the support of the 294 European Research Council (ERC) Grant "Hecate" reference 295 No. 695527 and the UKRI Future leaders fellowship Mrc-296 Mr/T043733/1. Computational resources provided by the 297 United Kingdom's National Supercomputer Service through 298 the United Kingdom Car-Parrinello consortium (EP/ 299 P022561/1) and Project ID No. d56 "Planetary Interiors" 300 and by the United Kingdom Materials and Molecular 301 Modelling Hub (EP/P020194) are gratefully acknowledged. 302 The authors thank J. Binns for useful discussions. 303

<sup>\*</sup>ross.howie@hpstar.ac.cn

W. B. Hubbard, W. J. Nellis, A. C. Mitchell, N. C. Holmes, 2
 S. S. Limaye, and P. C. McCandless, Science 253, 648 (1991).

306 38₹

308

309

310

311

312

313

314

315

316

317

318

319

320

321

322

323

324

325

326

327

328

329

330

331

332

333

334

335

336

337

338

339

340

341

342

343

344

345

346

347

348

349

- [2] R. Chau, S. Hamel, and W. J. Nellis, Nat. Commun. 2, 203 4 (2011).
- [3] S. S. Lobanov, P.-N. Chen, X.-J. Chen, C.-S. Zha, K. D. Litasov, H.-K. Mao, and A. F. Goncharov, Nat. Commun. 4, 2446 (2013).
- [4] R. Helled and J. J. Fortney, Phil. Trans. R. Soc. A 378, 20190474 (2020).
- [5] H.-k. Mao and R. J. Hemley, Rev. Mod. Phys. 66, 671 (1994).
- [6] R. T. Howie, C. L. Guillaume, T. Scheler, A. F. Goncharov, and E. Gregoryanz, Phys. Rev. Lett. 108, 125501 (2012).
- [7] R. T. Howie, P. Dalladay-Simpson, and E. Gregoryanz, Nat. Mater. 14, 495 (2015).
- [8] P. Dalladay-Simpson, R. T. Howie, and E. Gregoryanz, Nature (London) **529**, 63 (2016).
- [9] E. Gregoryanz, C. Ji, P. Dalladay-Simpson, B. Li, R. T. Howie, and H.-K. Mao, Matter Radiat. Extremes 5, 038101 (2020).
- [10] P. Loubeyre, F. Occelli, and P. Dumas, Nature (London)577, 631 (2020).
- [11] R. M. Hazen, H.-K. Mao, L. W. Finger, and P. M. Bell, Appl. Phys. Lett. **37**, 288 (1980).
- [12] H. Hirai, K. Konagai, T. Kawamura, Y. Yamamoto, and T. Yagi, Phys. Earth Planet. Inter. **174**, 242 (2009).
- [13] L. Sun, A. L. Ruoff, C.-S. Zha, and G. Stupian, J. Phys. Chem. Solids 67, 2603 (2006).
- [14] L. Sun, W. Yi, L. Wang, J. Shu, S. Sinogeikin, Y. Meng, G. Shen, L. Bai, Y. Li, J. Liu, H.-K. Mao, and W. L. Mao, Chem. Phys. Lett. **473**, 72 (2009).
- [15] J. E. Proctor, H. E. Maynard Casely, M. A. Hakeem, and D. Cantiah, J. Raman Spectrosc. 48, 1777 (2017).
- [16] H. E. Maynard-Casely, C. L. Bull, M. Guthrie, I. Loa, M. I. McMahon, E. Gregoryanz, R. J. Nelmes, and J. S. Loveday, J. Chem. Phys. **133**, 064504 (2010).
- [17] H. E. Maynard-Casely, L. F. Lundegaard, I. Loa, M. I. McMahon, E. Gregoryanz, R. J. Nelmes, and J. S. Loveday, J. Chem. Phys. **141**, 234313 (2014).
- [18] L. J. Conway and A. Hermann, Geosciences bulletin Series A 9, 227 (2019).
- M. Peña-Alvarez, A. V. Brovarone, M.-E. Donnelly, M.
   Wang, P. Dalladay-Simpson, R. Howie, and E. Gregoryanz, Nat. Commun. 12, 6387 (2021).
   352

- 353 [20] P. Loubeyre, R. Letoullec, and J.-P. Pinceaux, Phys. Rev. Lett. 72, 1360 (1994). 354
- [21] M. S. Somayazulu, L. W. Finger, R. J. Hemley, and H. K. 355 Mao, Science 271, 1400 (1996). 356
- [22] T. A. Strobel, P. Ganesh, M. Somayazulu, P. R. C. Kent, and 357 R. J. Hemley, Phys. Rev. Lett. 107, 255503 (2011). 358
- 359 [23] M. Somayazulu, P. Dera, A. F. Goncharov, S. A. Gramsch, 360 P. Liermann, W. Yang, Z. Liu, H.-K. Mao, and R. J. Hemley, 361 Nat. Chem. 2, 50 (2010).
- 362 [24] D. K. Spaulding, G. Weck, P. Loubeyre, F. Datchi, P. 363 Dumas, and M. Hanfland, Nat. Commun. 5, 5739 (2014).
- 364 [25] A. K. Kleppe, M. Amboage, and A. P. Jephcoat, Sci. Rep. 4, 365 <del>1</del> (2014).
- 366 [26] M. Somayazulu, P. Dera, J. Smith, and R. J. Hemley, J. Chem. 367 Phys. 142, 104503 (2015).
- [27] E. J. Pace, J. Binns, M. Peña Alvarez, P. Dalladay-Simpson, 368 E. Gregoryanz, and R.T. Howie, J. Chem. Phys. 147, 369 370 184303 (2017).
- 371 [28] C. Ji, A. F. Goncharov, V. Shukla, N. K. Jena, D. Popov, B. 372 Li, J. Wang, Y. Meng, V.B. Prakapenka, J.S. Smith, R. 373 Ahuja, W. Yang, and H.-k. Mao, Proc. Natl. Acad. Sci. 374 U.S.A. 114, 3596 (2017).
- [29] D. Laniel, V. Svitlyk, G. Weck, and P. Loubeyre, Phys. 375 Chem. Chem. Phys. 20, 4050 (2018). 376
- 377 [30] R. Turnbull, M.-E. Donnelly, M. Wang, M. Peña-Alvarez, 378 C. Ji, P. Dalladay-Simpson, H.-k. Mao, E. Gregoryanz, and 379 R. T. Howie, Phys. Rev. Lett. 121, 195702 (2018).
- 380 [31] J. Binns, P. Dalladay-Simpson, M. Wang, G. J. Ackland, E. 381 Gregoryanz, and R. T. Howie, Phys. Rev. B 97, 024111 382 (2018).
- 383 [32] M. Ceppatelli, D. Scelta, M. Serrano-Ruiz, K. Dziubek, G. 384 Garbarino, J. Jacobs, M. Mezouar, R. Bini, and M. Peruz-385 zini, Nat. Commun. 11, 6125 (2020).
- [33] W. L. Mao and H. K. Mao, Proc. Natl. Acad. Sci. U.S.A. 386 101, 708 (2004). 387
- [34] W. L. Mao, V. V. Struzhkin, H. K. Mao, and R. J. Hemley, 388 389 Chem. Phys. Lett. 402, 66 (2005).
- 390 [35] Y. Liu, D. Duan, F. Tian, H. X, D. Li, Z. Zhao, X. Sha, B. Chu, 391 H. Zhang, B. Liu, and T. Cui, RSC Adv. 4, 37569 (2014).
- 392 [36] G. Saleh and A. R. Oganov, Sci. Rep. 6, 32486 (2016).
- 393 [37] A. S. Naumova, S. V. Lepeshkin, and A. R. Oganov, J. Phys. 394 Chem. C 123, 20497 (2019).
- 395 [38] A. Zerr, G. Serghiou, R. Boehler, and M. Ross, High Press. 396 Res. 26, 23 (2006).
- [39] H. Kadobayashi, S. Ohnishi, H. Ohfuji, Y. Yamamoto, M. 397 Muraoka, S. Yoshida, N. Hirao, S. Kawaguchi-Imada, and 398 399 H. Hirai, Sci. Rep. 11, 8165 (2021).
- 400 [40] A. Kolesnikov, V. G. Kutcherov, and A. F. Goncharov, Nat. 401 Geosci. 2, 566 (2009).
- 402 [41] D. A. Kudryavtsev, T. -. Fedotenko, E. G. Koemets, S. E. Khandarkhaeva, V.G. Kutcherov, and L.S. Dubrovinsky, 403 404 Sci. Rep. 10, 1483 (2020).
- 405 [42] E. Snider, N. Dasenbrock-Gammon, R. McBride, M. Debessai, H. Vindana, K. Vencatasamy, K. V. Lawler, A. 406 407 Salamat, and R. P. Dias, Nature (London) 586, 373 (2020).
- 408 [43] W. Cui, T. Bi, J. Shi, Y. Li, H. Liu, E. Zurek, and R. J. 409 Hemley, Phys. Rev. B 101, 134504 (2020).
- 410 [44] A. F. Goncharov, E. Bykova, M. Bykov, X. Zhang, Y. Wang,
- 411 S. Chariton, and V.B. Prakapenka, J. Appl. Phys. 131, 412 **5** 025902 (2022).

- [45] See Supplemental Material at http://link.aps.org/ 413 supplemental/10.1103/PhysRevLett.000.000000 for fur-414 ther methodological details and additional experimental 415 and computational figures, which includes Refs. 416 417 [46-68]. 418
- [46] E. Aprà et al., J. Chem. Phys. 152, 184102 (2020).
- [47] S. B. Ramsey, M. Pena-Alvarez, and G. J. Ackland, Phys. Rev. B 101, 214306 (2020).

419

420

421

422

423

424

425

426

427

428

429

430

431

432

433

434

435

436

437

438

439

440

441

442

443

444

445

446

447

448

449

450

451

452

453

454

455

456

457

458

459

460

461

462

463

464

465

466

467

468

469

470

471

472

- [48] V. Labet, R. Hoffmann, and N. W. Ashcroft, J. Chem. Phys. 136, 074502 (2012).
- [49] V. Labet, R. Hoffmann, and N. W. Ashcroft, J. Chem. Phys. 136, 074503 (2012).
- [50] H.-K. Mao, J. Xu, and P. M. Bell, J. Geophys. Res. 91, 4673 (1986).
- [51] Y. Akahama and H. Kawamura, J. Appl. Phys. 100, 043516 (2006).
- [52] Y. Grin, F. R. Wagner, M. Armbrüster, M. Kohout, A. Leithe-Jasper, U. Schwarz, U. Wedig, and H.G. von Schnering, J. Solid State Chem. 179, 1707 (2006).
- [53] S. Schaack, U. Ranieri, P. Depondt, R. Gaal, W. F. Kuhs, P. Gillet, F. Finocchi, and L. E. Bove, Proc. Natl. Acad. Sci. U.S.A. 116, 16204 (2019).
- [54] H. T. Stokes, D. M. Hatch, and B. J. Campbell, FINDSYM, ISOTROPY Software Suite (iso.byu.edu).
- [55] H. T. Stokes and D. M. Hatch, J. Appl. Crystallogr. 38, 237 (2005).
- [56] T. Ohba, Y. Kitano, and Y. Komura, Acta Cryst. C 40, 1 (1984).
- [57] K. Ishizaki, P. Bolsaitis, and I. Spain, Solid State Commun. 15, 1591 (1974).
- [58] J. J. Gilvarry, Phys. Rev. 102, 308 (1956).
- [59] F.A. Lindemann, Phys. Z. 11, 609 (1910).
- [60] K. Refson, P. R. Tulip, and S. J. Clark, Phys. Rev. B 73, 155114 (2006).
- [61] A. Tkatchenko and M. Scheffler, Phys. Rev. Lett. 102, 073005 (2009).
- [62] G. Gao, A. R. Oganov, Y. Ma, H. Wang, P. Li, Y. Li, T. Iitaka, and G. Zou, J. Chem. Phys. 133, 4 (2010).
- [63] S. J. Clark, M. D. Segall, C. J. Pickard, P. J. Hasnip, M. I. J. Probert, K. Refson, and M. C. Payne, Z. Krist. - Cryst. Mater, 220, 567 (2005).
- [64] K. Momma and F. Izumi, J. Appl. Crystallogr. 44, 1272 (2011).
- [65] J. Gonzalez-Platas, M. Alvaro, F. Nestola, and R. Angel, J. Appl. Crystallogr. 49, 1377 (2016).
- [66] J. Rodríguez-Carvajal, Physica (Amsterdam) 192B, 55 (1993).
- [67] Y. Fei, A. Ricolleau, M. Frank, K. Mibe, G. Shen, and V. Prakapenka, Proc. Natl. Acad. Sci. U.S.A. 104, 9182 (2007).
- [68] C. Prescher and V. B. Prakapenka, High Press. Res. 35, 223 (2015).
- [69] P. Loubeyre, R. LeToullec, D. Hausermann, M. Hanfland, R. J. Hemley, H. K. Mao, and L. W. Finger, Nature (London) 383, 702 (1996).
- [70] P. Loubeyre, R. Letoullec, and J. P. Pinceaux, Phys. Rev. Lett. 67, 3271 (1991).
- [71] B. Chen, L. J. Conway, W. Sun, X. Kuang, C. Lu, and A. Hermann, Phys. Rev. B 103, 035131 (2021).