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**Filtered Data** 

The 29<sup>th</sup> Annual International Astronomical Data Analysis Software & Systems (ADASS) Conference

# Visualizing High-Dimensional Chemical Abundance Space in GALAH DR2

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Aim: Visualize high-dimensional data to find interesting patterns and underlying structures

**High-Dimensional Data** 

Filter high-dimensional data



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Landmark Multidimensional Scaling (LMDS [1]):

Clusters are **not well separated**.

Method is **fast**.



Proposed method (LGC+LMDS):

Clusters are well separated in the 2D projection.

Method is **fast**.

*t*-Stochastic Neighbor Embedding (t-SNE [2]): Clusters are **well separated.** Method is **slow**.

# **GALAH DR2**

Dataset: 10K observations are randomly chosen from the second data release of GALactic Archaeology with HERMES survey (GALAH DR2) [4] cross-matched with Gaia DR2 [5-6]. 10-D data set that consists of the following 10 stellar abundances are used: [Fe/H], [Mg/Fe], [Al/Fe], [Si/Fe], [Ca/Fe], [Ti/Fe], [Cu/Fe], [Zn/Fe], [Y/Fe], and [Ba/Fe]
 Results: LGC+LMDS shows cleaner separation of substructures in the 2D abundance-space than the original LMDS and t-SNE



## Key idea

Filter the high-dimensional data so that potential clusters are well separated even after dimensionality reduction

Summary

### Method

- Estimate density using Epanechnikov kernel [7-8]
- Shift points upstream in kernel density gradient, resulting in cluster contraction [9]
- III. Perform LMDS [1]

### **Advantages**

Clusters are well separated after the projection by preprocessing the data with local-based gradient clustering

### References

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- **Predictable** outcome with one parameter
- More computationally scalable than t-SNE, in terms of wall-clock time



### **Future Work**

A more sophisticated analysis of the different substructures gained from the LGC+LMDS results using GALAH DR2



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