The Open University

Open Research Online

The Open University's repository of research publications and other research outputs

The Stac Fada "impact ejecta" layer: not what it seems

Conference or Workshop Item

How to cite:

Osinski, G. R.; Preston, L.; Ferrière, L.; Prave, T.; Parnell, J.; Singleton, A. and Pickersgill, A. E. (2011). The Stac Fada "impact ejecta" layer: not what it seems. In: 74th Annual Meeting of the Meteoritical Society, 8-12 Aug 2011, London, UK.

For guidance on citations see \underline{FAQs} .

 \odot 2011 Not known

Version: Version of Record

Link(s) to article on publisher's website: http://www.lpi.usra.edu/meetings/metsoc2011/pdf/5451.pdf

Copyright and Moral Rights for the articles on this site are retained by the individual authors and/or other copyright owners. For more information on Open Research Online's data <u>policy</u> on reuse of materials please consult the policies page.

oro.open.ac.uk

THE STAC FADA "IMPACT EJECTA" LAYER: NOT WHAT IT SEEMS.

G. R. Osinski¹, L. Preston¹, L. Ferrière^{1,2}, T. Prave³, J. Parnell⁴, A. Singleton¹, and A. E. Pickersgill¹. ¹Dept. Earth Sciences, University of Western Ontario, London, ON, Canada. E-mail: gosinski@uwo.ca. ²Natural History Museum, Burgring 7, A-1010 Vienna, Austria. ³Dept. Earth Sciences, University of St. Andrews, St. Andrews, UK. ⁴Dept. Geology & Petroleum Geology, University of Aberdeen, Aberdeen, UK.

Introduction: The Stac Fada Member (SFM) forms part of the Stoer Group of the Torridonian of NW Scotland. The SFM is unique in the Torridonian, being characterized by the presence of greenish altered glass clasts. Its origin has been debated for decades with several hypotheses being proposed but all invoking some connection with volcanic activity in the region (e.g., [1]). More recently, Amor et al. [2] suggested that the SFM represents "a chord section through the continuous ejecta blanket surrounding an impact crater". Here, we confirm the presence of shocked material within the SFM and then discuss its origin.

Methods: Fieldwork was carried out in August 2008 and June 2009. We performed optical microscopy on 26 polished thin sections using an optical microscope and a four-axis universal stage. Quantitative analyses and investigation of textures were carried out on a JEOL JXA-8900 L electron microprobe.

Confirmation of shocked material within the SFM: The investigated breccia samples display a large heterogeneity of shock effects in quartz grains; including grains with PFs and/or PDFs (mostly 1-2 sets; up to 4 sets per grain). It is notable that some samples contain no shocked quartz grains. The crystallographic orientations of 90 PF and PDF sets in 59 quartz grains were measured by U-stage. A large proportion of the measured PDFs show orientations parallel to $\omega\{10\overline{1}3\}$ and $\sim 12\%$ of the measured PDFs are parallel to the $\pi\{10\overline{1}2\}$ orientation. Amor et al. [2] identified possible PDFs in 25 quartz grains from 9 thin sections. Our observations confirm the presence of PDFs in quartz in the SFM; although several discrepancies are notable.

Impact origin? We have confirmed that the SFM contains shocked material. However, based on a review of all terrestrial impact ejecta deposits and considerations of impact ejecta emplacement [3] we find several observations and properties of the SFM that are not consistent with it being a "continuous ejecta blanket" as proposed by [2]. Most importantly, it contains an order of magnitude less shocked material than other proximal impact melt-bearing ejecta layers (e.g., "suevite" from the Ries impact structure Germany [4]); and it is very well sorted compared to other impact ejecta deposits. In this latter respect, it shares many similarities to the Onaping Formation of the Sudbury impact structure, Canada, which has been recently reinterpreted as hot impact melt that interacted with seawater [5]. As such, we propose that the SFM is not of primary impact origin but is more akin to volcaniclastic rocks, which represent volcanic materials that have been transported and reworked by wind or water. This has implications for the recognition of impact ejecta deposits and their emplacement in H₂O-rich environments.

References: [1] Young G.M. 2002. *Trans. Earth Sci.* 93:1-16. [2] Amor K. et al. 2008. *Geology* 36:303-306. [3] Osinski G.R. et al. 2011. *EPSL* in press. [4] Engelhardt W.v. 1997. *MAPS* 32:545-554. [5]. Grieve R.A.F. et al. 2010. *MAPS* 45:759-782.

Additional Information: This work was funded by the Canadian Space Agency CARN Program and NSERC.