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# A FULL-VECTOR PALEOMAGNETIC SECULAR VARIATION RECORD FROM 55,000 TO 33,000 CAL. YEARS BP FROM RÍO VALDÉZ GLACIOLACUSTRINE OUTCROP (TIERRA DEL FUEGO, ARGENTINA)

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#### Abstract

High-resolution paleomagnetic investigations were performed on Río Valdez outcrop at the Fagnano Lake, central Isla Grande de Tierra del Fuego, southernmost South America. Our aims were to develop a full-vector paleomagnetic secular variation (PSV) record and establish paleomagnetic chronostratigraphy for the outcrop based on radiocarbon dating and the correlation of the PSV record with other dated PSV records in the region. We detected two distinct anomalous directional intervals at 35,400 and 33,800 cal. years BP and at 41,500 and 39,000 cal. years BP, which could be associated with Mono Lake and Laschamp excursions.

Rock magnetic investigations revealed slight changes in concentration, mineralogy, and magnetic grain size along the sedimentary sequence, but the reconstruction of past direction and relative paleointensity variations in the geomagnetic field remained feasible. The record provides new insights into the behavior of the geomagnetic field at high latitudes in the Southern Hemisphere, about which very little is currently known.

**Keywords**: paleomagnetic secular variation, excursion, Late-Pleistocene, sediments, Patagonia

#### 1. Introduction

The main geomagnetic field and its secular variation can 'ell us a great deal about the Earth's deep interior, and main-field and secular-variation models (particularly spherical harmonic models) are widely used in studies on the dynamics of the outer layers of the Earth's core. The data on which such models are baser' must be of the highest quality and distributed as homogeneously as possible worldwide (Barraclough, 1991). According to Lisé-Pronovost (2013), global stacks are truly derived from a majority of records located in the Northern Hemisphere (e.g., GLOPIS-75 La, et al., 2004; SINT-200, Guyodo and Valet, 1996) and geomagnetic field models lark c dibration data from the Southern Hemisphere to better understand the core geodynamics (e.g., Korte et al., 2005; Roberts, 2008; Korte and Constable, 2011; Pavón-Carrasco et al., 2014). Besides, high-resolution records of the paleomagnetic field made available a series of appropriate markers for global-scale correlation of paleomagnetic and paleoclimatic archives. In particular, the interval 30,000-50,000 cal. years BP comprising significant evidence of the variability of the climate during the Marine Isotope Stage (MIS) 3 justifies more detailed studies (Blanchet et al., 2006). Over the last two decades, we have contributed to building a more comprehensive record of Pleistocene-Holocene geomagnetic variability in southwestern Argentina and filling the gaps in data distribution for the Southern Hemisphere (Gogorza et al., 2012; Lisé-Pronovost et al.,

2013; Gogorza et al., 2018; Palermo et al., 2019; to cite some of the latest papers) by studying lake sediments, evaluating rock magnetic parameters, and yielding increasingly precise chronostratigraphy for these records. This study presents paleomagnetic data from a Late-Pleistocene glaciolacustrine record located at the Fagnano Lake (54° 35′S; 67° 20′W), central Isla Grande de Tierra del Fuego, southernmost South America. In particular, we searched for distinct features in the secular variation of the geomagnetic field and, possibly field excursions, which can be useful as chronostratigraphic markers. This paleosecular variation record would provide a correlation tool for this region and geomagnetic time-series data for a previously unrepresented part of the globe.

#### 2. Study area

The Fagnano Lake is located in central Isla Gr .n.  $^{\circ}$   $^{\circ}$ 

surrounding glaciodeltaic deposits by Bujalesky et al. (1997) and till deposits by Coronato et al. (2009).

The defined lithofacies in the studied profile (Sanci et al., 2020) made it possible to delineate the evolution of the paleolacustrine environment during the deposition of the studied record (Fig. 2). The base of the outcrop is represented by Unit A, composed of silty-clayey sediments distributed in two lithofacies formed by deposition from suspension, one of them during periods with no current activity in the water body and the other one under alternating deposition conditions of low energy tractive currents. Unit B complises a succession of fine sand silty-clay layers with thin and very thin bedding; two hun tacies were recognized, and they were generated when incomplete small sand ripples were formed in a mud-silt substratum and when wavy stratification was formed due to slightly alternating conditions. Unit C, at the top of the lake sequence, is comported of clay and silt sediments. The observed lithofacies represent successively massive sediments deposited from suspended load and slight bed load movements that form different sedimentary structures such as ripple beds or lenticular bedding. All the sediments and their characteristics are compatible with a glacial lacustrine record. Also, the succession of sedimentary facies represents an exceptionally continuous outcropping record of lake sediments. The total absence of fossils indicates lake water temperature relate 4 to direct contact with ice mass (Sanci et al, 2020); consequently, any kind of bioturbation can be expected. The characteristics of the diamititic sediments that cover the outcrop indicate that they were laid down by ice, thus forming a till. These diamititic sediments were not sampled for paleomagnetic studies.

#### 3. Methods

3.1 Field work

Two parallel profiles, a few meters apart from each other, were sampled at the Río Valdez outcrop during the 2012 Austral summer: Río Valdez 1 (RV1) and Río Valdez 2 (RV2), about 6.5 m and 4.5 m thick, respectively. Standard paleomagnetic sampling boxes (external size 2.2x2.2x2.2 cm and internal volume of 8 cm<sup>3</sup>) were pushed into the sediments at about 2 cm intervals at the site (n=224 for RV1 and n=144 for RV2) and then transported in boxes to the laboratory (see Subsection 3.2 below) for magnetic measurements.

#### 3.2 Paleomagnetic and rock magnetic measurements

Paleomagnetic and rock magnetic measurements were performed at the laboratory of Paleomagnetism and Paleoenvironmental Magnetism in 'L.ndil (CIFICEN, Buenos Aires, Argentina). Magnetic susceptibility was measured at low ( $\kappa_{low}$ ) and high frequency ( $\kappa_{high}$ ) with the MS2 Bartington susceptibility instrum en. quipped with the MS2B dual-frequency sensor at 0.47 kHz and 4.7 kHz. The frequer cy dependence of low-field susceptibility was estimated by the frequency-dependent tartor ( $\kappa_{fd}$ ), in terms of the low frequency and highfrequency measurements (Dearing, + al., 1999). The natural remanent magnetization (NRM, D, and I) was measured using a 'R6A Dual Speed Spinner Magnetometer. The orientation of the samples was preserved which respect to the reference line pre-marked in the field. The NRMs of all samples we. demagnetized and measured sequentially in alternating magnetic fields (AF) of 10, 15, 20, 25, 30, and 40 mT, which reduced most initial NRMs to less than 20% of their initial values. The selected samples were also demagnetized and measured at 60, 80 mT, and 100 mT. The results of stepwise AF demagnetization were plotted and analyzed using Zijderveld diagrams (Zijderveld, 1967). The characteristic remanent magnetization (ChRM), i.e., paleomagnetic inclination, declination, intensity, and maximum angular deviation (MAD), were calculated by principal component analysis using the least square method (Kirschvink, 1980). After each demagnetization cycle, anhysteretic remanent

magnetization (ARM) was applied to a 0.05 mT bias field, with an in-line single-axis direct current (DC) coil, combined with an axial peak AF of 100 mT. Finally, Isothermal Remanent Magnetization (IRM) for all samples was acquired at room temperature in increasing steps up to 1.2 T reaching saturation (SIRM), and in increasing steps back until the magnetic remanence was canceled using an IM-10-30 Pulse Magnetizer (ASC Scientific). Subsequently, AF demagnetizations of ARM and SIRM were measured using the same steps as for NRM demagnetization. Different magnetic properties and inter-parametric ratios were used to characterize the sediments in terms of type, concentration, and grain size of magnetic minerals. Magnetic susceptibility was used as a first-order indicator of the concentration of magnetic (sensu lato) minerals. Sratio was determined by a viding the back field measurement IRM-300mT with SIRM (Thompson and Oldfield, 1986). Combined magnetic parameters were calculated (SIRM/k and ARM/SIRM). In addition, the median destructive field of the NRM  $(MDF_{NRM})$  was determined as a first estimate of relative magnetic grain-size variations. Susceptibility vs. temperature measurements (k-T curve) under air were carried out using a Bartington susceptibility meter equipped with a furnace. The pilot samples were heated up to about 700 °C at a heating rate of 20 °C/min and then cooled at the same rate. Hysteresis loops were measured in fields up to conaximum of 1 T using a Molspin VSM (Vibration Sample Magnetometer) at the Institute of Basic, Applied, and Environmental Geosciences of Buenos Aires (IGEBA; Buenos Aires, Argentina) to determine the domain state of the magnetic minerals (Day et al., 1977; Dunlop, 2002).

#### 3.3 Age-depth model

The age model used for this study was based on the radiocarbon age-depth model constructed by Sanci et al. (2020) and refined by tuning the PSV records of Río Valdez outcrop to the PSV records of Potrok Aike (Lisé-Pronovost et al., 2013), a high-resolution paleomagnetic record constructed from the long sedimentary archive of Laguna Potrok Aike (PASADO-ICDP) covering the last 51.2 kyr. cal. BP.

#### 4. Results

#### 4.1 Rock magnetic results

Rock magnetic parameters were investigated to characterize the magnetic properties of the sediments and evaluate the stability of paleomagnetic records. The concentration magnetic parameters (K, NRM, ARM, and SIRM) vs depth, along with the sedimentary units, are plotted in Fig. 3.a. The sediments are generally weakly magnetized with  $\kappa$ , NRM, ARM, and SIRM intensities in the range of  $9 - 23 \times 10^{-5}$  SI ( $8 - 2^{\circ} \times 10^{-5}$  SI), 0.2 - 14 mA/m (0.2 - 60mA/m, except for three samples with NRM of 206, 2 J7, and 514 mA/m), 3 - 10 mA/m (0.1 -30 mA/m), and 101 – 243 mA/m (57 – 1400 r A/m) for RV1 (RV2), respectively. An up-core decrease (increase) and high-frequency va i ations in κ, NRM, and ARM (SIRM) are observed in Unit A. In particular,  $\kappa$ , NRM, and  $A\kappa$  display conspicuous oscillations superimposed on the general trend. In Unit B, all concentration magnetic parameters show oscillations although these changes are less noticeable than in Unit A, except for NRM in RV1. Finally, in Unit C, all parameters show small v. riations and follow an increasing trend at the top of the unit. Further above, significantly decreased values for all concentration-dependent magnetic parameters are observed in RV2 associated with unstratified deposits at the top of the outcrop. Interestingly, in general, all magnetic parameters show more pronounced fluctuations in RV2. A detailed analysis of the magnetic parameters obtained allows us to determine the variety of magnetic minerals in the studied material. The following analysis is presented for the whole sequence in general, focusing on some levels with distinctive characteristics that differ from the means. The  $\kappa_{fd}$  (not shown), a parameter sensitive to magnetic grains in superparamagnetic (SP) grain size range (Maher, 1988; Dearing, 1999), is highly fluctuating

and varies between 0.003 and 0.05, suggesting no presence of SP grains.  $S_{ratios}$  are typically higher than 0.9, indicating that the mineral magnetic composition is dominated by a soft coercivity component (e.g. (titano)magnetite). Anyway, lower values of  $S_{ratio}$  are observed in Units A and B, which suggests a slightly higher presence of a high coercivity mineral in addition to the (titano)magnetite (Fig. 3.a and b).

As detailed by Nilsson et al. (2018), ARM/SIRM is a parameter sensitive to the remanent carrying stable single domain (SSD)/PSD grains (Maher, 1988) and varies inversely with magnetic grain size (Maher and Taylor, 2007); higher values indicate finer magnetic grain size (SSD, PSD) and lower values correspond to coarse grains (MD). ARM/SIRM ranges from 0.01 to 0.05 (0.005 to 0.03) in RV1 (RV2) (Fig. 3 a and b). ARM/SIRM shows a decreasing trend, suggesting a decrease in the abundance of SSD/PSD-sized magnetic grains. Besides, the SIRM/ $\kappa$  ratio is a parameter sensitive to magnetic grains in the PSD range. Such inferences cannot be drawn when other form agnetic minerals such as greigite are present. Indeed, Roberts (1995) showed that greigite is characterized by large SIRM/ $\kappa$  values. SIRM/ $\kappa$  varies from 1 to 5 kA/m, and the top part of RV2. SIRM/ $\kappa$  shows a meth more even pattern, which points to a more evenly distributed abundance of are PSD-sized magnetic grains (Fig. 3.b). In particular, the presence of ferrimagnetic iron sulfites (greigite) at Unit C of RV2 (300-302 cm, Fig. 3.b) is supported by high values of SIRM/ $\kappa$ , an increase in the remanence coercivity (up to 40 mT), and a slight drop in the S<sub>ratio</sub>.

The SIRM of the pilot samples was reached below ca. 300 mT, indicating a magnetic assemblage dominated by low coercivity minerals (Fig. 4.a). Progressive removal of SIRM by back-field demagnetization indicates that coercivities (B<sub>CR</sub>) vary approximately between 31 and 42 mT (27 and 40 mT) in RV1 (RV2), suggesting that a low coercive magnetic mineral, such as magnetite of pseudo-single domain size (PSD), might be predominant within the

sediments (e.g., Dunlop and Özdemir, 1997). Higher values are observed between 772 and 803 cm in RV1 and towards the upper part (<10 cm) of RV2, in both cases reaching ca. 55 mT (Fig. 3.a and b). These values, not characteristic for pure magnetite, could be caused by the presence of partially oxidized (titano)magnetite (Roberts and Turner, 1993) and/or antiferromagnetic minerals in low concentrations, or by the decrease in the size of the grain. The increase of ARM/IRM in these parts of the profiles supports the latter hypothesis. The Day plot of hysteresis loop parameters (Fig. 5) for the pilot samples exhibits  $M_{RS}/M_{S}$ ratios ranging between 0.1 and 0.4 and  $H_{CR}/H_C$  ratios varying between 3.1 and 3.9, lying roughly on a (titano)magnetite grain size mixing line within up PSD field (Dunlop, 2002). There appears to be a slight displacement toward higher v<sub>a</sub><sup>1</sup>ues of  $H_{CR}/H_C$  in a group of samples. These results show that, as is usually the care n. sediment mixtures, the average grain size corresponds to the PSD range. It should be noted, however, that such observation could also derive from a mixture of two populations of single domain and multidomain grains (Blanchet et al., 2006). The continuous thermomagnetic experiments showed that the pilot samples present two magnetic phases. Ir the shown sample, the first magnetic mineral fraction has a Curie temperatur, of around 567±19° C and 452±47°C, compatible with (titano)magnetite with variable tranium content (Prévot et al., 1983). The second magnetic phase has a Curie temp. aure of about 675°C, suggesting the presence of a hard magnetic carrier like hematite in low proportion (Dunlop and Özdemir, 1997) (Fig. 4.b).

#### 4.2 Paleomagnetic directions

The AF demagnetization results of representative samples are illustrated and analyzed using Zijderveld diagrams (Zijderveld, 1967), also known as vector end plot diagrams. ChRM information is extracted from the orthogonal projection of the Zijderveld diagram using least square analysis (Kirschvink, 1980). The Zijderveld plot for selected samples of RV1 and RV2

(Fig. 6) shows that the viscous remanence magnetization -if present at all- is eliminated by applying demagnetization fields of 5 mT. Most of the samples have a single paleomagnetic direction that is demagnetized between 10-40 mT, which demagnetizes toward the origin. The ChRM MAD values are generally <5°, indicating a well-defined component magnetization. Beyond 941 cm (RV1) and for the top 10 cm (RV2), we could not retrieve reliable PSV records due to the scattered behavior of NRM intensity. Inclination and declination data below 968 cm are too noisy and many samples did not exhibit stable demagnetization paths; MAD also shows large scatter and an increase to 10–20°. Therefore, the ChRM data below 941 cm are not interpreted in the present study. Besides, samples in both profiles that show scatter behavior during the AF den agnetization were also discarded. The MDF is defined as the value of the peak AF required to reduce the remanence intensity to half its initial value and mostly depends on the cumposition and/or grain size of the minerals carrying these remanences. MDF<sub>NRM</sub> lie. be<sup>+</sup> ween 21 and 35 mT and 13 and 38 mT in RV1 (mean = 21 mT) and RV2 (mean = 19 h.T), respectively (Fig. 3.b) and indicates a fairly stable remanent magnetization. Stacking of individual curves is commonly employed on multiple paleomagnetic records to obtain more reliable estimates of the temporal variation in the directions of the geomagnetic field (Gogorza et al., 2012 and references therein). The mean inclination is -57.5° for Rv1 and -60.9° for RV2; inclinations range between -32.8° and  $-71.8^{\circ}$  and between  $-38.2^{\circ}$  and  $-79.8^{\circ}$ , respectively. Declination varies between  $-109^{\circ}$  and  $35.6^{\circ}$  in RV1 and  $-66^{\circ}$  and  $62.7^{\circ}$  in RV2 (Fig. 7). The average inclination (-58.6°) is lower than the inclination  $(-70.4^{\circ})$  of the field produced by an axial dipole at the site latitude and results from a series of intervals with lower inclinations (I-1 to I-5, Fig. 7). The inclination lows I-1, I-2, and I-3 depart around 30° from the GAD (vertical dashed line in Fig. 7) whereas I-4 and I-5 depart up to approximately 20° and 18°, respectively. I-1 is the interval displaying the lowest inclination values, and it is found from 35,500 to 33,700 cal. years BP with a

minimum inclination recorded at 34,600 cal. years BP ( $-38.6^\circ$ ). The other interval with inclinations deviating around 30° from the GAD is I-2, and it represents the interval from 37,700 to 35,700 cal. years BP. Besides, I-4 is the longest feature ranging from 41,600 to 38,500 cal. years BP, showing two narrow peaks and a small drop in between. Finally, the other two less prominent features, I-4 and I-5, are recorded at 43,950 ( $-50.2^\circ$ ) and 45,700 ( $-52.6^\circ$ ) cal. years BP, respectively. Rapid fluctuations in declination observed between 46,500 and 45,000 cal. years BP (D-5) and 41,000 and 39,000 cal. years BP (D-3) are probably associated with low inclinations, I-5 and I-3, respectively. Minor upward swings are observed in declination in the intervals 44,300–43,500 and 50,700–33,700 cal. years BP (D-4 and D-1), which could be associated with the anomalous 1. Clinations I-4 and I-1, respectively. We found no anomaly in declination in the vicinity of 37,700–35,700 cal. years BP inclination anomaly (I-2).

#### 4.3 Relative Paleointensity

The use of normalized NRM intensity to evaluate relative paleointensity (RPI) variations is conventionally based on the reliability criteria established by Banerjee et al. (1981) and King et al. (1982, 1983) and furthe. If fined by Tauxe (1993). The concentration of magnetic minerals in the sediment between 55,000 and 33,000 cal. years BP can be considered relatively stable, suggesting that the trends in the RPI within this period could reflect the relative intensity variations in the paleomagnetic field. In order to meet the criteria for uniformity in terms of concentration (Tauxe, 1993), those samples whose concentrations are beyond a factor of 10 were rejected. Paleointensity variability was estimated by normalizing the NRM to  $\kappa$ , ARM and SIRM. The NRMs after 10 mT AF demagnetization were normalized by dividing the NRM by either ARM or SIRM demagnetized at the same AF level and then renormalized by dividing the values by their mean value. NRM<sub>10mT</sub>/ $\kappa$ ,

NRM<sub>10mT</sub>/ARM<sub>10mT</sub>, NRM<sub>10mT</sub>/SIRM<sub>10mT</sub> are displayed in Fig. 8. Due to the markedly decreasing trend in the relative abundance of small magnetic grains observed in both profiles (Fig. 3.b), the use of the ARM as a normalizer may not be suitable. On the contrary, both  $\kappa$  and IRM are not affected in the same way as the ARM (Fig. 3.a) and are therefore likely more appropriate for these sediments. It is observed in Fig. 8 that NRM<sub>10mT</sub>/ $\kappa$ , and NRM<sub>10mT</sub>/SIRM<sub>10mT</sub> show similar results both in terms of fluctuation and amplitude, whereas NRM<sub>10mT</sub>/ARM<sub>10mT</sub> displays a more pronounced long-term trend and lower amplitude fluctuations, in particular from 40,500 to 55,000 cal. years BP

We carried out spectral analyses and a coherence test of norma'ized remanences and normalization parameters to confirm that the normalized records are free from environmental influence. Spectral analysis was performed using the MCCLEAN algorithm (Heslop and Dekkers, 2002). Coherence tests were carried out ...cording to the methods described by Tauxe and Wu (1990) by means of MA1. 43 6.1 software. The best normalization parameter is the one that displays no coherence while its normalizer parameter (Tauxe, 1993; Tauxe and Wu, 1990). At the 95% confidence involve the NRM<sub>10mT</sub>/SIRM<sub>10mT</sub> normalized intensity is not coherent with its normalizer except at periods corresponding to the interval 450-850 years. However, neither NRM<sub>10  $\mu_1$ </sub>/SIXM<sub>10mT</sub> nor SIRM<sub>10mT</sub> shows significant power in these periods. Therefore, they 1'2 not represent the influence of climate on lithology (Fig. 9). In agreement with the sediment cores from Laguna Potrok Aike (Gogorza et al., 2012), we consider that SIRM<sub>10mT</sub> should be the best normalizer for the sediments. We also show a closer resemblance between NRM and SIRM demagnetization curves compared to ARM demagnetization curves (Supplementary Material). This behavior was observed in most of the analyzed pilot samples taken along the whole core. We conclude that SIRM<sub>10mT</sub> is the most appropriate normalizer to construct the RPI proxies in these sediments. RPI mostly fluctuates between 0.4 and 1.2 for the period 55,000–41,900 cal. years BP (Fig.

8). We note a significant drop in paleointensity in the vicinity of the observed inclination anomaly I-4 at 41,500–40,400 cal. and a slight fall in the interval 46,200–45,200 years BP, which would coincide with the anomalous inclination I-5. From 40,400 cal. years BP to the present, we observe a clear increase and conspicuous oscillations. In particular, a remarkable minimum is displayed around 35,500–34,500 cal. years BP. However, we do not observe any anomaly in paleointensity in the vicinity of I-2 inclination anomaly.

#### 5. Discussion

#### 5.1. Some highlights about magnetic parameters

The mixed magnetic mineralogy determined in all sedime. tary units is described in Subsection 4.1. However, given the variation in the r agretic parameters obtained throughout the studied sedimentary sequence, some environ. enal considerations can be suggested. It is interesting to note that, in general, all magnetic parameters show more pronounced fluctuations in RV2. Also, the top of Unit C in this profile, which represents the youngest record of the sequence, shows both a correase in S<sub>ratio</sub> and an increase in B<sub>CR</sub> (Fig. 3.a and 3.b). These magnetic signals in Vicate a relatively more oxidizing environment, which could be explained by the position of the sequence.

Magnetic parameters in Unit A show high-frequency variations in  $\kappa$ , NRM, and ARM (SIRM), displaying conspicuous oscillations superimposed on the general trend. Also, a trend to higher  $\kappa$  and B<sub>CR</sub> (and lower S<sub>ratio</sub>) would indicate mixed magnetic particles slightly different from the rest of the sequence (Fig. 3.a). These magnetic variations are associated with lithofacies formed under alternating deposition from suspension during periods with no current activity in the water body and deposition with low-energy tractive currents. Very fine gravel intercalated between the thin facies could correspond to dropstones from ice-drifted

processes (Sanci et al., 2020). Based on the age model, the top of Unit A would be assigned to circa 55,000 cal. years BP. Therefore, the geological and magnetic evidence indicates that this unit could have developed under a different climatic scenario from that of the rest of the sequence, possibly related to the end of Marine Isotope Stage (MIS) 4 or more likely the beginning of MIS 3 (Lisiecki and Raymo, 2005). Unit B also shows a variation in magnetic parameters (Fig. 3.a and 3.b), although these changes are less conspicuous than those of the other units. A certain cyclicality can be perceived, which is associated with the rhythmites detected in the stratigraphic description. Another interesting level is represented in Unit C, namely a level with greigite detected in RV2 profile (Fig. ?.o, '1nis finding is supported mainly by high SIRM/ $\kappa$  values associated with an increase in the remanence coercivity (up to 40 mT); this level with a diagenetic mineral is representative of particular environmental conditions (Fig. 3.b). Considering an authigenic crigan, greigite is an intermediate iron sulfide on the polysulfide pathway to pyrite (Rocart's et al. 2011, among others). Greigite has become a common component in both lacustrine anoxic sedimentary environments (Blanchet et al. 2009, among others). It is formed balo v the oxic-to-anoxic transition zone (OATZ). Below the OATZ, degradation of organic matter must involve a reaction with sulfates. To form greigite in a lacustrine system, the lake water must be stratified and the sediment/water interface must be under . ducing conditions (Evans and Heller, 2003). In the studied deposit, lake stratification may have been caused by an increase in the lake level and/or low wind speeds, which prevent water mixture and/or unusual low winter temperatures. Taking into account its depth into the sequence and the age model, this level could represent prevailing 35,000 cal. years BP conditions. The aforementioned rise in the water lake level is not clearly related to any conspicuous global climate change but could be the consequence of a local process or represent some of the climatic variations occurring during MIS 3 (Lisiecki and Raymo, 2005).

#### 5.2 Mono Lake and Laschamp Excursions

Apart from reversals of the geomagnetic field, wide departures from its usual value have been observed in the geomagnetic field direction at a single locality. Such departures, when the field does not appear to change polarity but returns to its previous state, have been named geomagnetic excursions. Geomagnetic excursions during Brunhes chron, such as Mono Lake (Teanby et al., 2002; Nowaczyk et al., 2013; Lund et al., 2017); Laschamp (Teanby et al., 2002; Thouveny et al., 2004; Nowaczyk et al., 2013; Lund et al., 2017; Usapkar et al., 2018); Denmark (Abrahamsen and Knudsen, 1979); Norwegian-Creerland (Nowaczyk et al., 2013; Usapkar et al., 2018); Blake (Zhu et al., 1994; Thouvenv et al., 2004), and Icelandic Basin excursions (Thouveny et al., 2004), have been identified vorldwide. The youngest recorded excursions are of particular interest because they offer the potential for identifying them in several places on the Earth's surface and nv stigating their global morphology (Merrill and Mc Elhinny, 1983; Korte et al., 2019). In sedimentary sequences, it is possible to obtain a good continuous record of a transition if the sedimentation has been rapid. In marine sediments, the magnetic record is locked in over a 10 to 30 cm range so that the magnetic signal recorded by them is finared. However, sediment records have proved to be one of the most valuable sources c. information on the transitional behavior of the geomagnetic field (Merrill and Mc Elhinny, 1983).

The PSV recorded in the sediments of Río Valdez outcrop contains three main intervals of anomalous directions: 35,400–33,800 cal. years BP, 41,500–39,000 cal. years BP, and 46,600–45,300 cal. years BP (Fig. 7). We consider that the youngest of these anomalous PSV intervals (I-1) could be associated with the Mono Lake excursion, the large directional swing (I-3) at intermediate ages could be related to Laschamp geomagnetic excursion, and the oldest but less conspicuous one (I-5) could correspond to the directional swing recorded at 46,000

cal. years BP at Laguna Potrok Aike (Lisé-Pronovost et al., 2013). According to Lisé-Pronovost et al. (2013), the last one could represent an important geomagnetic feature in the Southern Hemisphere. The center ages of these intervals of anomalous directions are 34,600, 40,250, and 45,650 cal. years BP, which implies a good agreement with the recent dating of Mono Lake excursion (34,100±300 cal. years BP, Lund et al., 2017) and Laschamp geomagnetic excursion (41,100±350 cal. years BP, Lascu et al., 2016; Lund et al., 2017). Further support for the interpretation of I-1 being the Mono Lake excursion is the presence of a coeval declination swing (D-1) and a relative paleointensity intrinum (Fig. 8), as expected during a geomagnetic excursion. Similar departures are observed in inclination values. The inclination swing (D-2) and a minimum in RPI value our interpretation. Finally, a coeval declination swing (D-5) is associated v. th 2.5, although the behavior in paleointensities is quite ambiguous.

#### 5.3 A comparison of global-scale pulpunagnetic directions

To compare the global-scale pair omagnetic directions, the following available directional records are displayed: Poteck rake (51°58′S 70°22′W) (Lisé-Pronovost et al., 2013); MD94-103 (45°35′S 86°31′E) (Azaud et al., 2002); Crevice Cave (37° 40′N 89°50′W) (Lascu et al., 2016); Wega (Macri et al. 2005), and Hawaiian lava flows (19°29′N 154°54′W) (Teanby et al., 2002). In particular, a very good centennial-scale agreement is observed with the inclination record of Laguna Potrok Aike, as expected due to its location at a distance of ~400 km. However, declination does not show a very similar pattern. Differences in the amplitude of variations could, in turn, be related to the different sedimentation rates and lock in depths of the different records. The inclination records are generally consistent and most peak values of Río Valdez can be correlated among the records (I-1 to I-5). In the case of declination,

following the numbers of inclination, the main patterns are identified as D-1, D-3, D-4, and D-5. The inclination low event I-5 in Rio Valdez (Fig. 10.a) is associated with the sharp declination swing D-5 (Fig. 10.b). Similar direction swings are observed in Potrok Aike (Lisé-Pronovost et al., 2013), Crevice Cave (Lascu et al., 2016), MD94-103 (Mazaud et al., 2002), and Hawaiian lava flows (Teanby et al., 2002). The comparison reveals a good correlation between the inclination lows I-1, I-2, I-3, I-4, I-5 (Fig 10.a), (with, however, different amplitudes for I-2 and I-5) and the declination swings D-2 and D-5 (Fig. 10.b). Concerning the directional feature (I-5 and D-5) at 46,000 cal. years BP, which according to Lisé-Pronovost et al. (2013) seems to be mainly observed in the potthern Hemisphere, a similar pattern is displayed in Crevice Cave, located about 38°N.

5.3 A comparison of global-scale paleomagnetic intensities

In Fig. 10.c, the relative paleointensity reform a from Río Valdéz is compared for the period 55,000–33,000 cal. years BP with the lagustrine records of Laguna Potrok Aike (Lisé-Pronovost et al., 2013), the marine records from the Indian and Atlantic sectors of the Southern Ocean (MD94-103; Nazaud et al., 2002, and SAPIS; Stoner et al., 2002), from the western Equatorial Pacific (ND97-2134; Blanchet et al., 2006), and the WEGA stack near Antarctica (Macrì et al., 2005). This comparison also includes a speleothem geomagnetic record from North America (Crevice Cave; Lascu et al., 2016) and the VADMs derived from Black Sea sediments (Nowaczyk et al., 2013) and lacustrine sediments of Lake Baikal (Peck et al., 1996). Three periods can be clearly identified. A decreasing trend from 55,000 to ~45,500 cal. years BP is followed by a small increase from 45,500 to 41,000 cal. years BP. From 41,000 to 33,000 cal. years BP, the intensity shows an oscillating behavior superimposed on a value that doubled the intensity of the previous intervals. Although the records are generally consistent, it should be noted that the described decreasing trend, from

55,000 to ~45,500 cal. years BP, is not as noticeable in the records from Argentina as observed in the remaining records. In particular, an intensity low (P-3) in Río Valdéz (Fig. 10.c) is associated with the inclination low event I-3 (Fig. 10.b) and the sharp declination swing D-3 (Fig. 10.b). Similar direction swing and intensity low are observed in the records: MD94-103 (Mazaud et al., 2002), MD97-2134 (Blanchet et al., 2006), and Crevice Cave (Lascu et al., 2016). Besides, the low P-1 in Río Valdéz (Fig. 10.c) associated with the low inclination event I-1 (Fig. 10.b) and the less conspicuous declination swing D-1 (Fig. 10.b) is displayed in Potrok Aike (Lisé-Pronovost et al., 2013) and MD24-193 (Mazaud et al., 2002). The intensity lows P-1 and P-3 are also observed in the RP1 of SAPIS (Stoner et al., 2002) and ODP 1089 (Stoner et al., 2003). Some temporal shifts are observed and could be linked to the uncertainties of the respective chronologies. Add<sup>2</sup>.ion.<sup>3</sup>lly, the minimum in intensity at about 46,000 cal. years BP, which is related to the irclination low event I-5 (Fig. 10.b) and the declination swing D-5, is absent at R<sub>1</sub> V aldez (Fig. 10.c). However, it is displayed at Potrok Aike (Lisé-Pronovost et al., 2012) SAPIS (Stoner et al., 2002), and ODP 1089 (Stoner et al., 2003). Based on the comparative analysis, we cannot discard a rock-magnetic influence on the RPI recorded in the sediments from Río Valdéz for the period 55,000 to 47,000 cal. years B, although all the string out requirements for RPI studies were fulfilled.

#### 6. Conclusions

We performed a rock magnetic and paleomagnetic study of Late-Pleistocene sediments of Río Valdez outcrop in Southern South America. On the basis of our analysis, we constructed a full-vector PSV record for the interval 55,000 to 33,000 calibrated years BP. A global-scale paleomagnetic intensity and direction comparison of the Río Valdéz record supports the validity of the geomagnetic signal. We confirm the existence of two intervals of anomalous paleomagnetic directions, at 34,600  $\pm$  800 cal. years BP and at 40,250  $\pm$  1250 cal. years BP,

which could be associated with evidence of Mono Lake and Laschamp Excursion in the records of Río Valdéz. It is important to point out that this is the first record with these characteristics at this latitude.

The following are the supplementary data related to this article.

Supplementary Table 1. Modelled age for RV1 and RV2 profiles of Río

Valdéz outcrop, including 14C date and calibrate 1 at es (RV1-22; RV1-17;

RV2-5; RV2-11; RV2-15). Modified from Sanc. et al. (2020).

Supplementary material 1

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#### **AUTHORSHIP STATEMENT**

**P. Palermo:** Conception and design of study, acquisition of data, analysis and/or interpretation of data, drafting the manuscript, revising the manuscript critically for important intellectual content, approval of the version of the manuscript to be published.

**C.S.G. Gogorza:** Conception and design of study, acquisition of data, analysis and/or interpretation of data, drafting the manuscript, revising the manuscript critically for important intellectual content, approval of the version of the manuscript to be published.

María J. Orgeira: Conception and design of study, acquisition of data, analysis and/or interpretation of data, revising the manuscript critically for important intellectual content, approval of the version of the manuscript to be published.

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Romina Sanci: Support for the stratigraphic characterization of Río Valdez outcrop, approval of the version of the manuscript to be published.

Andrea Coronato: Sedimentary exposure discovered and conception of its paleoenvironmental significance, acquisition of data, revising the manuscript critically for important intellectual content, approval of the version of the manuscript to be published.

conflicts of interests The authors report no conflicts of interests

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Figure 1. Location maps: (a) geographical location of Río Valdez outcrop, (b) locations of the sampling sites.

Figure 2. Stratigraphic integrated scheme from the sedimentary record of RV1 and RV2.

Figure 3. (a) Down-core variations in concentration-magnetic parameters: magnetic susceptibility ( $\kappa$ ), NRM, ARM, SIRM, and MDF<sub>NRM</sub>from RV1 and RV2 along with the sedimentary units. (b) Down-core variations in rock-magnetic parameters: SIRM/ $\kappa$ , ARM/SIRM, BCR, and SratioOf RV1 and RV2, along with  $\kappa$  and the sedimentary units.

Figure 4. Rock magnetic properties of the Río Valdez sediments: (a) Isothermal Remanent Magnetization (IRM) acquisition curves, for representative samples from RV1 and RV2. (b) Thermomagnetic curves (κvs. T) of a pilot sample.

Figure 5. Day plot (Day et al., 1977) for the sediment samples of Río Valdez outcrop. The mixing reference lines for single and multi-domain (SD and MD) are from Dunlop (2002).

Figure 6. Normalized intensity decay curves and orthogonal vector plots for four pilot samples from RV1 and RV2, produced by stepwise AF demagnetization at 5-10 mT steps, up to a maximum field of 100 mT. Closed (open) circles denote the projection on the horizontal (vertical) plane.

Figure 7. (a) Inclination and (b) acc ination plots for RV1 (open squares), RV2 (closed squares), and stacked profile (orange line).(c) Latitudes of the related virtual geomagnetic point (VGP).

Figure 8. Normalized estimate of relative paleointensity (RPI):

NRM<sub>10mT</sub>/ $\kappa$ , NRM<sub>10mT</sub>/S!KM<sub>10mT</sub> and NRM<sub>10mT</sub>/ARM<sub>10mT</sub> versus Age. Figure 9. Spectral analysis of three normalization parameters (ARM<sub>10mT</sub>, SIRM<sub>10mT</sub> and three normalized remanences (NRM<sub>10mT</sub>/ARM<sub>10mT</sub>, NRM<sub>10mT</sub>/SIRM<sub>10mT</sub> NRM<sub>10mT</sub>/ $\kappa$ ). Coherence tests of relative paleointensity curves with respect to the normalization parameters, showing that the normalized intensity curves are generally not coherent with the normalization parameters.

Figure 10. (a) Comparison of global-scale of paleomagnetic (a) inclination, (b) declination and (c) relative paleointensity records. Directional profiles include Río Valdez (this study); Potrok Aike (Lisé-Pronovost et al., 2013); Crevice Cave (Lascu et al., 2016); Wega (Macri et al. 2005); MD94-103 (Mazaud et al., 2002), Hawaiian lava flows (Teanby et al., 2002); MD97-2134 (Blanchet et al., 2006), and Lake Baikal (52°N 106°E) (Peck et al., 1996). (c) Comparison of normalized intensity comprises the relative paleointensity from Río Valdez (this study), the marine records from the

Indian and Atlantic sectors of the Southern Ocean (MD94-103; Mazaud et al., 2002, and SAPIS; Stoner et al., 2002), from the western Equatorial Pacific (MD97-2134; Blanchet et al., 2006), the WEGA stack near Antarctica (Macrì et al., 2005), a speleothem geomagnetic record from North America (Crevice Cave; Lascu et al., 2016),the VADMs derived from Black Sea sediments (Nowaczyk et al., 2013), and lacustrine sediments from Lake Baikal (Peck et al., 1996). Notable features of the paleomagnetic inclination (I-1 to I-5), declination (D-1 to D-5), and intensity (P-1 to P-2) are indicated. The Laschamp and Mono Lake geomagnetic excursions are underlined in green and the event at ca 46,000cal. years BP in purple (see text for details).

#### **Graphical Abstract**

Highlights

- We report new paleomagnetic results for the 'agnono Lake, central Isla Grande de Tierra del Fuego, southernmost South Ainerina.
- The results confirm the existence of t vo intervals of anomalous paleomagnetic directions, which could be associated with evidence of Mono Lake and Laschamp Excursions.
- This is the first record with the c characteristics reported at this latitude.



Río Valdéz Outcrop (Argentina)





**Graphics Abstract** 



Figure 1









Figure 3B



Figure 4

Theoretical Day plot curves for magnetite





![](_page_37_Figure_0.jpeg)

Figure 7

![](_page_38_Figure_0.jpeg)

Figure 8

![](_page_39_Figure_0.jpeg)

Figure 9

![](_page_40_Figure_0.jpeg)

![](_page_40_Figure_1.jpeg)

Río Valdez (54°S)

120

0

-120 l

150

D-1

Mono Lake

(a)

![](_page_40_Figure_3.jpeg)

(b)

49000

D-5

D-5

D-5

100 T

0

J<sub>-100</sub>

30

0

-30

-60

60

C

-60

53000

Potrok Aike (52°S)

MD97-2134 (10°S)

Baikal (53°N)

D-4

C-3

D-3

(C)