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1 Transcurrent displacement of the Cadomian magmatic arc

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9

10 Abstract

11 The Ossa-Morena Zone in Iberia, and equivalent terranes in northern France and
12 Central Europe, are thought to be paleogeographically linked to the West African
13 Craton at Ediacaran to early Paleozoic times. Evidence is mainly based on
14 metasedimentary rock detrital zircon age spectra, characterized by two dominant
15 populations of Paleoproterozoic and Cryogenian-Ediacaran ages, with a systematic lack
16 of late Stenian – early Tonian zircon grains. We report here U-Pb-Hf results on detrital
17 zircon grains from six Ordovician-Devonian metasedimentary rocks from the Ossa-
18 Morena Zone. In addition to the Cryogenian-Ediacaran and Paleoproterozoic
19 populations already recognized in previous studies on uppermost Ediacaran and lower
20 Cambrian rocks, the samples ubiquitously show a late Stenian – early Tonian
21 population centered at ≈ 1 Ga and representing ≈ 20 % of the concordant dates. The
22 ϵHf *versus* age plot of the studied samples mainly depicts two vertical arrays
23 corresponding to the Cryogenian-Ediacaran and Stenian-Tonian detrital zircon
24 populations, which spread from ϵHf values of ≈ 10 down to ≈ -20 . The detrital zircon
25 age distribution and Hf isotope signature of the studied rocks point to the Sahara
26 metacraton as the most plausible sediment source. The eastward translation of the
27 continental ribbon represented by the Ossa-Morena Zone and equivalent domains of
28 the Variscides from an original position close to the West African Craton to an
29 Ordovician-Devonian location close to the Sahara metacraton probably occurred at
30 latest Ediacaran – earliest Cambrian times in a dextral strike-slip tectonic setting that
31 post-dated Pan-African collision and Cadomian subduction.

32

33 **Key-words:** Detrital zircon grains, U-Pb geochronology, Hf isotope signature, Ossa-
34 Morena Zone, Sahara metacraton, West African Craton.

35

36 1. Introduction

37 Paleogeographic and paleotectonic reconstruction of pre-Mesozoic orogens is
38 subjected to a number of uncertainties regarding the number of intervening
39 plates/microplates, as well as the extent and age of subducted oceanic domains in-
40 between the amalgamated continental pieces. A first step to addressing this classic
41 issue is producing good quality geological maps at the scale of the orogen, where the
42 location and lateral continuity of ophiolitic and/or high-pressure belts must be
43 carefully established. Obviously, different transects of the same orogen often offer
44 very different perspectives on the number of putative oceanic and continental
45 domains implied, so correlation at the scale of the whole orogen becomes
46 controversial. The combination of field geology with other geological, paleontological,
47 geophysical and geochronological data can sometimes shed some light on the orogenic
48 picture, though consensus is rarely reached. Among these new data, the systematic U-
49 Pb dating of detrital zircon populations has become a powerful tool to elucidate the
50 derivation of the different continental terranes welded during collisional and pre-
51 collisional orogenic evolution (e.g., Miller et al., 2006; Wang et al., 2007; Gehrels, 2012
52 and 2014).

53 The Variscides represent a paradigmatic example of an orogen where uncertainty
54 remains regarding the number and extent of oceanic domains subducted at Late
55 Paleozoic time, before assembly of the supercontinent Pangea. Interpretations of the
56 paleogeographic/paleotectonic evolution vary from those considering only a single and
57 wide oceanic realm (Rheic Ocean) between two main continental plates (Laurussia and
58 Gondwana; e.g., Robardet, 2002; Kroner and Romer, 2013; Romer and Kroner, 2019;
59 Stephan et al., 2019a and b), to the existence of a number of different oceanic and
60 continental domains between the two main continental blocks (e.g., Matte, 2001;
61 Stampfli et al., 2013; Franke, 2000; Franke et al., 2017 and 2019). The variety of
62 models results from overemphasis on one kind of data over others. Proposals relying
63 only on faunal differences (e.g., Robardet, 2002; Robardet and Gutiérrez-Marco, 2004),
64 geochemical data (e.g., Romer and Kroner, 2019), or detrital zircon data (Stephan et
65 al., 2019a and b) favour a single and wide oceanic domain. In contrast, those models
66 based on a multi-method approach support that at least two oceanic domains of
67 varying importance were involved in Variscan orogenesis (e.g., Franke, 2000; Matte,
68 2001; Simancas et al., 2002 and 2009).

69 Independently of the number of oceanic domains involved, all available
70 paleogeographic reconstructions consider that the Variscides resulted from the
71 amalgamation of a number of Gondwana-derived terranes, whose primary position at
72 the periphery of the supercontinent is well established in some cases, but unclear in
73 others. Studies on detrital zircon populations from Ediacaran to Paleozoic rocks have
74 provided key paleogeographic evidence (e.g., Linnemann et al., 2008; Pereira et al.,
75 2011 and 2012a, Pérez-Cáceres et al., 2017; Collet et al., 2020). Avalonian terranes are
76 characterized by detrital zircon spectra with a distinctive multi-peak distribution
77 between 1 and 1.8 Ga, as well as a zircon-forming event at Late Ordovician-Silurian
78 time (Hamilton and Murphy, 2004; Braid et al., 2012; Pérez-Cáceres et al., 2017;
79 Pereira et al., 2017; Herbosch et al., 2020; Sorger et al., 2020). These terranes were

80 attached to the western segment of Northern Gondwana at Ediacaran time, but during
81 Rheic Ocean in the Early Paleozoic, they drifted from Gondwana and collided with
82 Laurentia during the Silurian (e.g., Nance et al., 2002). The remaining Gondwana-
83 derived continental pieces involved in the Variscan orogenesis would have never
84 drifted far from Gondwana, having evolved either as ribbon continental pieces
85 separated by minor oceanic realms from the mainland, or as aborted rifts and passive
86 margins attached to it. These peri-Gondwanan terranes have detrital zircon spectra
87 dominated by a prominent Ediacaran peak attributed to subduction-related
88 Cadomian/Avalonian arc magmatism developed all along the northern Gondwana
89 margin (e.g., Gutiérrez-Alonso et al., 2003; Linnemann et al., 2014). Minor peaks at 2
90 Ga are always present, while a 1 Ga population is only present in some zones (Central
91 Iberian Zone, West Asturian-Leonese Zone, and Cantabrian Zone in Iberia; e.g.,
92 Gutiérrez-Alonso et al., 2015) and has been attributed to eastern Arabian-Nubian
93 shield or Sahara metacraton zircon sources (Bea et al., 2010; Díez-Fernández et al.,
94 2010; Meinhold et al., 2013). The absence of a 1 Ga detrital zircon population of in the
95 Ossa-Morena Zone (OMZ) and equivalent terranes in northern Brittany and Central
96 Europe (Saxo-Thuringian Zone) has served to locate the OMZ close to the West African
97 Craton (WAC) in Ediacaran to early Paleozoic times, west of the other Variscan zones
98 (Fernández-Suárez et al., 2002; Linnemann et al., 2008; Pereira et al., 2011 and 2012a).
99 Nevertheless, there is no consensus on this paleogeographic attribution. Cambeses et
100 al. (2017) used the Sm-Nd isotope signature of Ediacaran to Cambro-Ordovician rocks
101 to propose that the OMZ was located close to the Tuareg Shield in Cambrian-
102 Ordovician times, in a more easterly position than previous studies have claimed. In
103 contrast, López-Guijarro et al. (2008), Fuenlabrada et al. (2020) and Rojo-Pérez et al.
104 (2021) also utilised Sm-Nd data to propose a western OMZ paleoposition, close to the
105 WAC.

106 This paper aims to elucidate the paleogeographic and paleotectonic meaning of the
107 OMZ, based on the previous geological knowledge, new U-Pb ages and the new Hf
108 isotope signatures of detrital zircon grains from its Ordovician-Devonian sequence. In
109 particular, our study seeks evidence of a 1Ga zircon population in Paleozoic OMZ
110 samples in an effort to evaluate the Paleozoic location of this piece of the Iberian
111 Massif, and its counterparts in other regions of the Variscides.

112

113 **2. Geological setting**

114 The Variscan Orogen extends from Central and Western Europe to Northwestern
115 Africa, with the northern foreland (Laurussia) cropping out in England, Belgium and
116 Germany. The southern one (Gondwana) is mostly obscured due to Alpine reworking,
117 except in Morocco where it is exposed in the Anti-Atlas ranges (Fig. 1A). The Iberian
118 Massif (Central-Western Spain and Portugal) constitutes the largest outcrop of the
119 Variscan orogen, having been traditionally divided into a number of zones with
120 different stratigraphic, tectonometamorphic, and magmatic features (e.g., Julivert et
121 al., 1974; Simancas, 2019 and references therein). This classic division (Fig. 1B)

122 comprises the Cantabrian (CZ), Western Asturian-Leonese (WALZ), Central Iberian
123 (CIZ), Galicia - Tras-Os-Montes (GTMZ), Ossa-Morena (OMZ) and South Portuguese
124 (SPZ) Zones. The correlation of these zones with other European Variscan massifs
125 through the Cantabrian arc usually considers the CIZ, WALZ and CZ to represent the
126 Gondwanan margin, while the SPZ represents the Avalonian foreland with
127 counterparts in Southern England, Belgium and Germany (Reno-Hercynian Zone)
128 (e.g., Matte, 1991; Franke, 2000; Simancas et al., 2005; Murphy et al., 2016). The OMZ
129 extends through the Cantabrian arc in northernmost Brittany, Belgium, Germany and
130 Czech Republic, where it is referred to as the Saxo-Thuringian Zone.

131 The OMZ is characterized by particular stratigraphic and magmatic features, which
132 have served, despite controversies, to elucidate its paleogeographic and paleotectonic
133 evolution. In the stratigraphic record, both Ediacaran and Lower Paleozoic sequences
134 in the OMZ are distinctive with respect to other Variscan Zones in the Iberian Massif.
135 The OMZ is also characterized by an abundance of upper Ediacaran and lower-middle
136 Cambrian to Ordovician volcanic and plutonic rocks (e.g., Simancas et al., 2004;
137 Sarrionandia et al., 2020).

138 On a broad scale, the OMZ can be considered as a continental segment deformed
139 between two Variscan suture-type contacts, namely the contacts with the CIZ (Burg et
140 al., 1981; Azor et al., 1994 and 2019; Simancas et al., 2001; Gómez-Pugnaire et al.,
141 2003; López Sánchez-Vizcaíno et al., 2003) and SPZ (Bard, 1977; Crespo-Blanc and
142 Orozco, 1991; Fonseca and Ribeiro, 1993; Quesada et al., 1994; Castro et al., 1996;
143 Azor et al., 2008 and 2019; Pérez-Cáceres et al., 2015 and 2020). These two tectonic
144 boundaries attest to oceanic and/or continental subduction and subsequent collision
145 at Devonian-Carboniferous times, which deformed the whole OMZ and SPZ interiors
146 (e.g., Simancas et al., 2013). The OMZ/SPZ boundary is unanimously considered as the
147 Rheic Ocean suture, though the contact is marked by a Carboniferous MORB-featured
148 amphibolitic unit that postdates ocean consumption (Azor et al., 2008) and makes the
149 suture appear as cryptic (Pérez-Cáceres et al., 2015). On the contrary, the OMZ/CIZ
150 boundary, namely Badajoz-Córdoba Shear Zone (Figs. 1B and 2), has been the subject
151 of an extensive debate regarding its tectonic significance. Some authors have
152 considered this boundary as a Variscan intracontinental shear zone and, hence, have
153 claimed paleogeographic continuity of the CIZ and OMZ at Early Paleozoic times along
154 the Gondwana margin (Abalos et al., 1991; Quesada, 1991; Robardet, 2002; Ribeiro et
155 al., 2007). Others have interpreted this contact as representing a second-order
156 Variscan suture, where the ophiolite-bearing allochthonous units of NW Iberia would
157 be rooted (Matte, 1991; Simancas et al., 2002). The available petrological,
158 geochronological and geochemical data have shown that the entire metamorphic
159 evolution, including an HP event, is Variscan in age (Ordóñez-Casado, 1998; Pereira et
160 al., 2010a and 2012b; Abati et al., 2018). Furthermore, some of the MORB-featured
161 mafic rocks have Cambrian to Early Ordovician ages (Ordóñez-Casado, 1998; Gómez-
162 Pugnaire et al., 2003), thus confirming that the CIZ/OMZ boundary records the closure
163 of a minor oceanic realm between Gondwana and a ribbon continental segment (Azor
164 et al., 2019).

166 **2.1. Stratigraphic framework**

167 The OMZ shows an almost continuous stratigraphic record from late Ediacaran to late
168 Carboniferous times (Figs. 2 and 3), with several unconformities of diverse importance
169 and meaning (e.g., Quesada et al., 1990; Azor et al., 2004; Robardet and Gutiérrez-
170 Marco, 2004). The whole sequence is dominated by siliciclastic rocks, with some
171 carbonate-dominated periods (Early Cambrian and Mississippian); volcanic-plutonic
172 rocks with variable geochemical signatures are common at latest Ediacaran, Cambrian,
173 early Ordovician, and Mississippian times (e.g., Sánchez Carretero et al., 1990; Ordóñez
174 Casado, 1998; Galindo and Casquet, 2004). From a tectonic point of view, the pre-
175 orogenic sequence is mainly affected by km-scale SW-vergent recumbent folds, with
176 an associated axial-plane foliation; a second folding event gave way to NW-SE striking
177 upright folds with an associated crenulation cleavage (e.g., Expósito, 2000; Azor et al.,
178 2019). This second event is the only ductile deformation observed in the syn-orogenic
179 sequence. Left-lateral strike-slip faults occurred during a late Variscan deformational
180 event (e.g., Pérez-Cáceres et al., 2016).

181 The oldest outcropping rocks belong to the so-called Serie Negra Group (Black Series),
182 which consists of slates, schists and greywackes, with amphibolite and black quartzite
183 intercalations (e.g., Quesada et al., 1990) (Fig. 3). Based on the age of the youngest
184 detrital zircon population in samples from the Serie Negra Group (Schäfer et al., 1993;
185 Fernández-Suárez et al., 2002; Linnemann et al., 2008; Pereira et al., 2011 and 2012a),
186 a late Ediacaran maximum depositional age was established. This age is in agreement
187 with the stratigraphic position of the Serie Negra Group (underlying paleontologically-
188 dated Cambrian rocks), as well as with some radiometric ages obtained on the
189 amphibolites (Sánchez-Lorda et al., 2016) and cross-cut relationships with Early
190 Cambrian granitic rocks (Pereira et al., 2011; Sánchez-García et al., 2014). A bimodal
191 volcano-sedimentary sequence with abundant associated plutonic rocks (Malcocinado
192 Formation) unconformably overlies the Serie Negra Group. Numerous radiometric ages
193 on the volcanic and plutonic rocks associated with this formation yielded latest
194 Ediacaran to earliest Cambrian ages (Oschner, 1993; Ordóñez-Casado, 1998; Salman,
195 2004; Sarrionandia et al., 2020). The Malcocinado Formation is traditionally
196 interpreted as representing a supra-subduction zone / Cadomian volcanic arc
197 developed all along the northern Gondwanan continental margin (Sánchez Carretero
198 et al., 1990; Pin et al., 2002; Simancas et al., 2004).

199 The Cambrian sequence unconformably overlies the Serie Negra Group and the
200 Malcocinado Formation (Fig. 3), attesting to the onset of a new Wilson cycle that gave
201 way to the opening of the Rheic ocean. The lowermost Cambrian formation
202 (Torreárboles Formation) is made up of arkosic sandstones with conglomerate and
203 slate intercalations (Liñán, 1978). Onto the Torreárboles formation, a lower Cambrian
204 succession of limestones with slate intercalations crops out (Pedroche formation;
205 Liñán, 1974). A lower-middle Cambrian slate-sandstone dominated sequence overlies
206 the Pedroche Formation. Finally, a thick sequence of middle-upper Cambrian slate and

207 fine-grained sandstones with abundant felsic/mafic volcanic intercalations (Sánchez-
208 García, 2001) witnesses the rifting stage that preceded Rheic ocean opening (Sánchez-
209 García et al., 2003, 2010 and 2019; Chichorro et al., 2008; Palacios et al., 2021).

210 The entire Ordovician-Devonian succession is interpreted as a passive-margin
211 sequence deposited on the Gondwana margin coevally with Rheic –and other minor–
212 ocean expansion (e.g., Robardet and Gutiérrez-Marco, 1990). In the localities of Cerrón
213 del Hornillo and El Valle, the Ordovician (upper Floian-Hirnantian) consists of a ≈500 m-
214 thick sequence dominated by slates, siltstones and sandstones, with a 15-20 m-thick
215 limestone intercalation of Katian age (e.g., Robardet, 1976; Robardet et al., 1998;
216 Robardet and Gutiérrez-Marco, 2004) (Fig. 3); in other outcrops, the Ordovician
217 sequence is entirely siliciclastic and reaches thicknesses of up to 1000 m. The Silurian
218 comprises a 100-150 m-thick condensed sequence dominated by graptolite-bearing
219 black slates, with some black chert, sandstone and limestone intercalations. Among
220 these intercalations, the most important is the so-called Scyphocrinites Limestone
221 (lower Pridoli), a 10-15 m-thick alternating black limestone and slate series that divides
222 the Silurian sequence into a 120 m-thick lower part (lower graptolitic slates) and a 20
223 m-thick upper part (upper graptolitic slates). However, in some of the transects the
224 Silurian does not intercalate any carbonatic level. The Devonian succession starts with
225 black slates of Lochkovian age in stratigraphic continuity with Silurian slates. The
226 remaining Lower Devonian pre-orogenic sequence is dominated by green to brown
227 shales with siltstone intercalations and has been dated as Praguian-Emsian according
228 to the abundant fossil content (e.g., Robardet et al., 1998; Robardet and Gutiérrez-
229 Marco, 2004).

230 The syn-orogenic deposits of the OMZ start with marine greywackes and
231 conglomerates concentrated in the so-called Terena Flysch syncline. Actually, the
232 Terena Flysch comprises two different successions (Expósito, 2000): the Lower
233 Devonian Terena sediments (Piçarra 1996; Piçarra et al. 1998) are the first record of
234 syn-orogenic deposits; the upper Terena sediments, which unconformably lie over the
235 lower Terena or directly overlie the pre-orogenic succession, are of Famennian to
236 Visean age (Boogaard and Vazquez, 1981; Giese et al., 1994). In other outcrops there
237 are only lower Carboniferous (Tournaisian-Visean) conglomerates, shales, greywackes
238 and limestones, with volcanic intercalations (e.g., Azor et al., 2004).

239

240 **2.2. Previous detrital zircon geochronological data**

241 A number of studies have reported detrital zircon ages of late Ediacaran (Serie Negra
242 and Malcocinado formations) and Cambrian rocks of the OMZ (Fernández-Suárez et al.,
243 2002; Linnemann et al., 2008; Pereira et al., 2011 and 2012a). Despite the limited
244 number of analyzed grains, these studies showed a dominant peak of Ediacaran-
245 Cryogenian age (550-750 Ma), attributed to the Pan-African orogeny and/or the
246 Cadomian arc magmatism, and minor peaks centered at 2 Ga, which are interpreted as
247 material derived from the West African Craton (WAC). Furthermore, the absence of a

248 late Stenian – early Tonian (≈ 1 Ga) detrital zircon population has been used to
249 reinforce the western paleoposition of the OMZ, close to the WAC along the northern
250 Gondwanan margin.

251 The detrital zircon ages of the Carboniferous syn-orogenic deposits have also been
252 investigated (Pereira et al., 2012c and 2020; Dinis et al., 2018). The results show the
253 presence of Early Carboniferous, Late Devonian, Ordovician, Cambrian, Neoproterozoic
254 and Paleoproterozoic detrital zircon populations at variable percentages, which are
255 compatible with the amalgamation and recycling of variably sourced grains during the
256 Variscan collision.

257

258 **3. Samples and methods**

259 This study focuses on the OMZ Ordovician-Devonian sequence at two localities,
260 namely the Cerrón del Hornillo and El Valle synclines (**Error! Reference source not**
261 **found.** 2 and 3), where both sedimentary facies and faunistically-based stratigraphic
262 ages are very well known (Robardet, 1976; Robardet et al., 1998; Robardet and
263 Gutiérrez-Marco, 2004). The whole Ordovician-Devonian succession was sampled,
264 though the Silurian rocks are monotonous black shales and ampelites with very minor
265 fine-grained silt intercalations that did not yield detrital zircon grains. In contrast,
266 Ordovician and Devonian rocks are mostly sandy levels with abundant detrital zircon
267 grains. We analyzed 4 samples from the Cerrón del Hornillo syncline (CH2, CH3, and
268 CH4 of Middle Ordovician age, and CH6 of Early Devonian age) and 2 samples from the
269 El Valle syncline (EV1 and EV2 of Late Ordovician age; Figs. 2 and 3; Table 1).

270 Under the microscope, all of the studied samples are non-foliated quartz-rich
271 sandstones. Quartz represents up to 90 % of the rock content, with opaque minerals
272 being also present in all of the samples, often at interstitial positions together with
273 phyllosilicate minerals. Quartz grains generally have angular shapes and variable
274 diameters from 100 to 400 microns, with mean values of ≈ 200 -250 microns. Feldspar,
275 zircon, rutile and tourmaline have been recognized as accessory minerals.

276 For each sample site, 4-5 kg of rock were collected and processed at the laboratories of
277 the University of Granada (Spain). A statistically significant amount of detrital zircon
278 was separated from each sample using mechanical smashing in a jaw-crusher, sorting
279 by sieving, manual panning, and, finally, handpicking. Cathodoluminescence (CL)
280 images (Fig. 4) were taken on a Carl Zeiss SIGMA HD VP Field Emission SEM at the
281 School of GeoSciences, University of Edinburgh (Scotland, United Kingdom), and a
282 Mira3 FESEM instrument at the John de Laeter Centre (JdLC), Curtin University (Perth,
283 Australia) in order to recognize zoning and alteration textures within the zircon grains.
284 U-Th-Pb analyses were conducted using a Laser Ablation Inductively Coupled Plasma
285 Mass Spectrometry (LA-ICPMS) at the GeoHistory Facility, JdLC. In order to have a
286 representative number of analyses, some additional U-Th-Pb analyses were performed
287 on small zircon grains of sample EV2 using a Secondary Ion Mass Spectrometry (SIMS)

288 at the NERC Ion Microprobe Facility of the University of Edinburgh (UK) (Table 1). A
289 detailed description of the analytical methods can be found in Appendix A.

290 Raw data were analyzed using Isoplot (Ludwig, 2003 and 2009) and SQUID II (for
291 SHRIMP analyses). All the dates with a discordance higher than 10% were discarded.
292 $^{206}\text{Pb}/^{238}\text{U}$ dates were used for zircon grains younger than 1.5 Ga and $^{207}\text{Pb}/^{206}\text{Pb}$ dates
293 for >1.5 Ga zircon grains because of the significant increase in error of $^{206}\text{Pb}/^{238}\text{U}$ ratios
294 for > 1.5 Ga zircon grains. The mean square weighted deviation (MSWD) of the
295 youngest detrital zircon population was calculated using IsoplotR online (Vermeesch,
296 2018). Finally, Kernel Density Estimators (KDE) and histograms were calculated using
297 the software DensityPlotter 8.4 (Vermeesch, 2012) and applying a KDE bandwidth and
298 histogram bin of 20 Ma. Errors are expressed at 1σ level.

299 Hf isotope analysis was undertaken in the GeoHistory Facility, JdLC. The analyses were
300 performed on previously dated zircon grains using a Resonetics resolution M-50A
301 excimer laser, coupled to a Nu Plasma II multicollector inductively coupled plasma
302 mass spectrometer. The analyzed isotopes were: ^{180}Hf , ^{179}Hf , ^{178}Hf , ^{177}Hf , ^{176}Hf , ^{175}Lu ,
303 ^{174}Hf , ^{173}Yb , ^{172}Yb , and ^{171}Yb . The ^{176}Lu decay constant of Scherer et al. (2001) and the
304 Chondritic Uniform Reservoir (CHUR) values of Blichert-Toft and Albarède (1997) were
305 applied in the calculation of ϵHf values. Errors are expressed at 2σ level. A full
306 description of this method can be found in Appendix A.

307

308 **4. Results**

309

310 **4.1. U-Pb geochronology of detrital zircon grains**

311 Sample CH2 (Middle Ordovician; Fig. 3): We carried out 150 analyses on 147 zircon
312 grains, having obtained 138 concordant dates (Fig. 5A; Appendix B). Most ages define a
313 dominant Cryogenian-Ediacaran population peaked at ≈ 630 Ma (c. 554-703 Ma; 51
314 analyses; 37 % of the data); two second-order peaks within this population occur at \approx
315 590 and 680 Ma. The youngest detrital zircon population in this sample, based on 7
316 Ediacaran dates, has a mean $^{206}\text{Pb}/^{238}\text{U}$ age of 592.6 ± 1.7 Ma (MSWD = 1.24). A late
317 Stenian – early Tonian population peaked at ≈ 1 Ga accounts for 27 dates and
318 represents 19.6 % of the data, while second-order maxima appear at ≈ 780 and 885
319 Ma. Furthermore, third-order peaks appear at ≈ 1070 , 1880 (c. 1955-1808 Ma; 8
320 analyses; 5.8 % of the data) and 2025 Ma (c. 2170-1986 Ma; 10 analyses; 7.2 % of the
321 data); the two Paleoproterozoic peaks can be interpreted as featuring an Orosirian
322 population that would include 18 dates (13 % of the data). Finally, a few scattered data
323 yielded Neo-Archean dates (2.5-2.8 Ga).

324 Sample CH3 (Middle Ordovician; Fig. 3): 150 analyses were performed on 145 zircon
325 grains, yielding 137 concordant dates (Fig. 5B; Appendix B). The majority of data can be
326 grouped into a Cryogenian-Ediacaran population with two maxima at ≈ 600 and 635
327 Ma (c. 721-541 Ma; 64 analyses; 46.7 % of the data; Table 2). The $^{206}\text{Pb}/^{238}\text{U}$ mean age

328 of the youngest detrital zircon population in this sample is 592.8 ± 1 Ma (MSWD =
329 1.07), being defined by 12 Ediacaran dates. A late Stenian – early Tonian population is
330 also present, peaked at 992 Ma (c. 1085-881 Ma; 24 analyses; 17.5 % of the data; Fig.
331 5B). A minor Paleoproterozoic (Orosirian) population centered at c. 2 Ga is also present
332 (16 dates; 11.6 % of the data), with peaks at 1874 (c. 1911-1834 Ma), 1970 (c. 1986-
333 1957 Ma) and 2063 Ma (c. 2079-2045 Ma). Finally, a very minor Neo-Archean-Siderian
334 population can be defined (11 dates; 8 % of the data), with peaks at ≈ 2455 (c. 2497-
335 2428 Ma) and 2650 Ma (c. 2693-2601 Ma).

336 Sample CH4 (Middle Ordovician; Fig. 3): We carried out 150 analyses on 150 zircon
337 grains, yielding 139 concordant dates (Appendix B). A prominent Cryogenian-Ediacaran
338 population peaked at ≈ 595 and 645 Ma dominates this sample (c. 734-565 Ma; mean
339 age 628.8 ± 0.6 Ma; 53 dates; 38.1 % of the data; Fig. 5C and Table 2); a second-order
340 peak centered at ≈ 710 Ma can also be included within this population. The youngest
341 detrital zircon population of this sample is late Ediacaran in age (585.6 ± 1.7 Ma; 6
342 dates; MSWD = 1.03). The late Stenian – early Tonian population defines a narrow
343 band centered at ≈ 1 Ga (c. 1071-913 Ma; mean age 988.0 ± 1.3 Ma; 30 dates; 21.6 % of
344 the data). A third population of Paleoproterozoic (Rhyacian-Orosirian) age is also
345 observed in this sample (21 dates; 15.2 % of the data), with maxima at ≈ 1865 (c. 1892-
346 1848 Ma), 1970 (c. 1981-1965 Ma) and 2080 Ma (2154-2021 Ma). The older dates
347 define a very minor Neo-Archean population centered at ≈ 2630 Ma (c. 2678-2563 Ma;
348 8 dates; 5.8 % of the data). There are also a few scattered analyses at ≈ 1225 , 1425 and
349 2340 Ma, and 4 grains with early Cambrian dates (c. 545-520 Ma; Table 2).

350 Sample CH6 (Lower Devonian; Fig. 3) yielded 137 concordant dates obtained from 150
351 analyses on 135 zircon grains (Appendix B). The main population of this sample is
352 Cryogenian-Ediacaran in age and peaks at ≈ 640 Ma, (c. 715-546 Ma; mean age
353 639.1 ± 0.7 Ma; 54 analyses; 39.4%; Fig. 5D and Table 2); a third-order late Tonian peak
354 centered at ≈ 775 Ma (c. 813-755 Ma; 7 dates; 5.1 % of the data) appears very close to
355 the Cryogenian-Ediacaran population. The youngest detrital population of this sample,
356 based on 7 analyses, is Ediacaran in age (590.2 ± 1.7 Ma; MSWD = 1.26). A second
357 noticeable population is of late Stenian – early Tonian age, with a main peak at ≈ 1 Ga
358 (c. 1091-886 Ma; 37 dates; 27 % of the data). A Paleoproterozoic (Rhyacian-Orosirian)
359 population is also present in this sample, with four separate peaks at ≈ 1890 , 1970,
360 2055 and 2135 Ma (26 dates; 11.6 % of the data). Finally, a Neo-Archean population is
361 defined by a single peak centered at ≈ 2615 Ma (5 dates; 3.6 % of the data).

362 Sample EV1 (Upper Ordovician; Fig. 3): We obtained 133 concordant dates from 148
363 analyses on 131 zircon grains (Appendix B). Most dates form a Cryogenian-Ediacaran
364 population (c. 770-578 Ma; mean age 648.1 ± 0.5 Ma; 67 dates, 50.4% of the data; Fig.
365 5E and Table 2) peaking at ≈ 650 Ma, with two second-order maxima at ≈ 600 and 690
366 Ma. A late Stenian – early Tonian population (c. 1141-803 Ma; 30 dates; 23.6 % of the
367 data) peaks at ≈ 1070 Ma, with second-order maxima at ≈ 985 and 825 Ma.
368 Paleoproterozoic (Rhyacian-Orosirian) and Neo-Archean minor populations are also
369 present with ages ranging from c. 2138 to 1963 Ma (9 dates; 6.8 % of the data) and

370 from 2719 to 2506 Ma (13 dates; 9.8 % of the data), respectively. Finally, a few dates
371 define an early Cambrian peak centered at ≈ 500 Ma (5 dates with a mean age of
372 498.7 ± 1.5 Ma; 3.8 % of the data). Paradoxically, the youngest detrital zircon
373 population of this sample is Ediacaran (591.4 ± 1.8 Ma; 4 dates; MSWD = 1.17).

374 Sample EV2 (Upper Ordovician; Fig. 3): We obtained 114 concordant dates out of 122
375 analyses (80 by LA-ICPMS and 42 by SIMS; Table 1) of 122 zircon grains (Appendix B).
376 The dominant population is Cryogenian-Ediacaran in age (c.714-546 Ma; mean age
377 632.1 ± 0.5 Ma; 46 dates; 40.4% of the data), with a main peak at ≈ 630 Ma and minor
378 peaks at ≈ 580 and 670 Ma (Fig. 5F; a third-order latestmost Tonian peak centered at \approx
379 735 Ma (10 dates; 8.8 % of the data) appears very close to the Cryogenian-Ediacaran
380 population and separated from an older Stenian-Tonian population. This latter
381 population is centered at ≈ 1 Ga (c. 1081-979 Ma; mean age 1023.2 ± 1.2 Ma; 22 dates;
382 19.3% of the data) and shows a second-order peak at ≈ 910 Ma (c. 958-862 Ma; 10
383 dates; 8.8% of the data). A third and less relevant population is of Paleoproterozoic
384 (Rhyacian-Orosirian) age, characterized by two maxima at ≈ 1900 and 2000 Ma (19
385 dates; 16.7 % of the data). Finally, a few scattered dates group at ≈ 2650 Ma and can be
386 considered to represent a very scarce Neo-Archean population. Two dates of ≈ 500 Ma could
387 be taken as evidence of the presence of Cambrian detrital zircon grains. The youngest
388 detrital population of this sample, based on 4 dates, is late Ediacaran (582.4 ± 1.9 Ma;
389 MSWD = 0.79).

390 To sum up, all of the studied samples show a quite similar detrital zircon record
391 characterized by (i) a dominant Ediacaran-Cryogenian population with maxima at c.
392 595-650 Ma, (ii) a noteworthy late Stenian – early Tonian population with an almost
393 invariable peak at ≈ 1 Ga, (iii) a subordinate Paleoproterozoic population usually
394 depicting several peaks between c. 1.8 and 2.1 Ga, and (iv) very scarce Neo-Archean-
395 Siderian scattered dates that range from c. 2.4 to 2.7 Ga.

396

397 **4.2. Hf isotope signature**

398 In all of the studied samples, Lu–Hf isotopes were analyzed on the same zones of the
399 zircon grains where concordant U–Pb zircon ages were obtained.

400 Sample CH2 (Middle Ordovician; Fig. 3): One hundred and thirty-eight Lu-Hf isotope
401 measurements were carried out in this sample, yielding $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.280865–
402 0.28271 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial ratios of 0.280689–0.282677 (Appendix C). Seventy-six
403 per cent of the CHUR-normalized (ϵHf) values are negative, ranging from -0.1 to -27.3,
404 while supra-CHUR values vary from 0 to 10.9. Cryogenian-Ediacaran zircon grains show
405 a vertical arrangement that extends from the Depleted Mantle (DM) line at ϵHf 10.9
406 down to ϵHf -27.3 (Fig. 6A; Table 2); negative ϵHf values clearly dominate over positive
407 values, with most ϵHf clustered between 3.4 and -20. The late Stenian – early Tonian
408 population depicts ϵHf values between 10.5 and -22.2 (Fig. 6A; Table 2); the zircon
409 grains peaked at ≈ 1 Ga define an almost continuous and vertical trend with a range of
410 ϵHf from 6.3 to -18.8; as for the two minor Tonian peaks, the younger (≈ 780 Ma)

411 shows dominant positive ϵHf values between 7 and 3, while the older peak (≈ 885 Ma)
412 plots following a vertical arrangement with rather dispersedly distributed values from
413 10.9 to -18.8. The Orosirian zircon population dominantly shows negative ϵHf values in
414 the range -0.2 to 18.5. Finally, the scarce and scattered Neo-Archean zircon grains have
415 ϵHf values that vary between 4.3 and -10.0.

416 Sample CH3 (Middle Ordovician; Fig. 3): One hundred and eight Lu-Hf isotope
417 measurements were carried out in this sample, yielding $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.280902–
418 0.282667 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial ratios of 0.280859–0.282645 (Appendix C). Seventy-two
419 per cent of the CHUR-normalized (ϵHf) values are negative, ranging from -0.1 to -27.7,
420 while supra-CHUR values vary from 0 to 11.7. The Cryogenian-Ediacaran zircon grains
421 are organized in a vertical array in the ϵHf *versus* age plot that covers an ϵHf interval of
422 9.7 to -27.7, with most data concentrated between ϵHf 5 and -15 (Fig. 6B). The late
423 Stenian – early Tonian population also displays a vertical arrangement that extends
424 from ϵHf 11.7 down to -23.8; the younger and older subpopulations (centered at ≈ 875
425 and 1070 Ma, respectively) dominantly show supra-CHUR ϵHf values, while ≈ 1 Ga
426 zircon grains define a sub-CHUR vertical array ($\epsilon\text{Hf} = 0$ to -16.9). The Paleoproterozoic
427 (Orosirian) and Neo-Archean-Siderian zircon grains mostly have sub-CHUR ϵHf values,
428 ranging from 0 to -20.7 (Fig. 6B).

429 Sample CH4 (Middle Ordovician; Fig. 3): One hundred and thirty-nine Lu-Hf isotope
430 analyses were carried out in this sample, yielding $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.280927–
431 0.28262 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial ratios of 0.280826–0.282595 (Appendix C). Seventy-one
432 per cent of the CHUR-normalized (ϵHf) values are negative, ranging from -0.1 to -24.2,
433 while supra-CHUR values vary from 0 to 11.7. The Cryogenian-Ediacaran detrital zircon
434 population depicts a vertical array in the ϵHf *versus* age diagram (Fig. 6C); more
435 precisely, peak ages (≈ 630 Ma) cluster between ϵHf 5.7 and -21.2, while older ages (\approx
436 715 Ma) mostly show supra-CHUR values (3.2-8.7). The late Stenian – early Tonian
437 population also defines a vertical trend in the ϵHf *versus* age diagram, with most of the
438 values clustering between ϵHf 4.4 and -13 (Fig. 6C). Paleoproterozoic zircon grains are
439 distributed between ϵHf 4.1 and -13.7, with older ages rather clustered at supra-CHUR
440 values and younger ages progressively plotted at sub-CHUR values. Finally, the scarce
441 Neo-Archean zircon grains show a subvertical spread, covering an ϵHf interval from 2.3
442 to -10.2.

443 Sample CH6 (Lower Devonian; Fig. 3): One hundred and thirty-six Lu-Hf isotope
444 analyses were carried out in this sample, yielding $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.280984–
445 0.282692 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial ratios of 0.280917–0.282658 (Appendix C). Seventy-nine
446 per cent of the CHUR-normalized (ϵHf) values are negative, ranging from -0.2 to -24.5,
447 while positive values vary from 0 to 10.3. The dominant Cryogenian-Ediacaran
448 population defines a heterogeneously distributed vertical array, extending from ϵHf
449 10.3 down to -24.5, with most values clustered between ϵHf 2.3 and -12.4 (Fig. 6D); a
450 minor late Tonian population shows a similar vertical trend (ϵHf from 5.1 to -13). The
451 late Stenian – early Tonian zircon grains also depict a remarkable vertical array with

452 ϵ Hf values extending from 7 to -18.6 (Fig. 6D). The Paleoproterozoic zircon grains are
453 distributed from ϵ Hf 4.1 down to -14.4, with older ages (peaked at \approx 2135 Ma) defining
454 a vertical trend (ϵ Hf from 4.1 to -3.6) and younger ones mostly clustered at sub-CHUR
455 ϵ Hf values (0 to -14.4). Finally, the scarce Neo-Archean zircon grains arrange vertically
456 in the sub-CHUR field (0.1 to -7; Fig. 6D).

457 Sample EV1 (Upper Ordovician; Fig. 3): One hundred and thirty Lu-Hf isotope analyses
458 yielded $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.280618–0.282637 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial ratios of
459 0.280525–0.282629 (Appendix C). Seventy-eight per cent of the CHUR-normalized
460 (ϵ Hf) values are negative, ranging from -0.1 to -32.1, while positive values vary from
461 0.3 to 9.3. The Cryogenian-Ediacaran population is organized in a vertical array that
462 extends from ϵ Hf values of 9.3 down to -24.3, with most data concentrated in the sub-
463 CHUR field between -3.3 and -12.5 (Fig. 6E); late Tonian zircon grains also distribute
464 vertically, but supra-CHUR ϵ Hf values dominate over sub-CHUR ones. The scarce
465 Cambrian zircon grains mostly show ϵ Hf values close to CHUR (0.7 to -3.2), except for
466 one of the measurements (-32.1). The late Stenian – early Tonian population defines
467 an apparent vertical array with ϵ Hf values varying from 4.8 to -23.7 (Fig. 6E). The
468 subordinate Paleoproterozoic and Neo-Archean populations dominantly show negative
469 ϵ Hf values, though a few supra-CHUR values are also present.

470 Sample EV2 (Upper Ordovician; Fig. 3): The small size of a good number of the dated
471 zircon grains did not allow Hf isotope analysis. Thus, only 59 Lu-Hf isotope analyses
472 were conducted, yielding $^{176}\text{Hf}/^{177}\text{Hf}$ ratios of 0.281068–0.282675 and $^{176}\text{Hf}/^{177}\text{Hf}$ initial
473 ratios of 0.281013–0.282666 (Appendix C). Sixty-four per cent of the CHUR-normalized
474 (ϵ Hf) values are negative (-0.4 to -32.7) and the remaining positive (0.1 to 10.4). The
475 analysed Cryogenian-Ediacaran and late Stenian – early Tonian zircon grains depict
476 vertical trends in the ϵ Hf *versus* age diagram, extending from 10.4 to -32.7 and 9.1 to -
477 16.0 ϵ Hf values, respectively (Fig.6F). In-between the two dominant detrital zircon
478 populations, late Tonian detrital zircon grains cluster in the supra- (ϵ Hf 3.5 to 7.3) and
479 sub-CHUR (ϵ Hf -3.7 to -6.2) fields. Finally, the scarce Paleoproterozoic and Neo-
480 Archean zircon grains appear scattered between ϵ Hf values of 2.8 and -9.4.

481 To sum up, all of the samples studied are characterized in the ϵ Hf vs age diagram by
482 two well-defined vertical arrays corresponding to the Cryogenian-Ediacaran and late
483 Stenian – early Tonian detrital zircon populations, which extend from supra-CHUR
484 values of \approx 10 down to sub-CHUR values of \approx -20 (Fig. 6). Paleoproterozoic and Neo-
485 Archean detrital zircon populations are also mostly arranged in vertical trends, with
486 supra-CHUR ϵ Hf values clustered at the \approx 0 to 5 interval and sub-CHUR values ranging
487 from 0 to -10.

488

489 5. Discussion

490

491 5.1. Stability of detrital zircon sources at Ordovician-Devonian time

492 All of the samples studied show very similar detrital zircon age spectra, with
493 comparable percentages of both dominant and subordinate populations (Fig. 5). The
494 Hf isotope signature is also uniform (Fig. 6), independent of the stratigraphic age of the
495 sample. The simplest way of interpreting these results is by invoking stable detrital
496 zircon sources at Ordovician-Devonian time, which, in turn, can be related to the
497 tectonic setting: a passive-margin sequence is the obvious tectonic setting for the OMZ
498 Ordovician-Devonian rocks.

499 The youngest detrital zircon populations can serve to constrain maximum depositional
500 ages (MDA) of the studied sedimentary rocks. True depositional ages (TDA), usually
501 based on paleontological dating of the sedimentary rocks and/or direct
502 geochronological dating of volcanic intercalations, can be slightly delayed –or even be
503 geologically coeval– with respect to MDA in some cases, but they can also be
504 separated by hundreds of millions of years in others (e.g., Sharman and Malkoswki,
505 2020). In our case, the scarcity of Cambrian - Lower Ordovician detrital zircon grains in
506 the OMZ Ordovician-Devonian means that there is a long time lapse between MDA and
507 TDA, in agreement with the inferred passive-margin setting of the OMZ throughout the
508 Ordovician - Devonian period. The most plausible explanation for the scarcity of
509 Cambrian - Lower Ordovician detrital zircon grains is that the present-day exposed
510 igneous rocks of Cambrian – Early Ordovician age did not reach the surface until the
511 collisional and/or post-collisional stages of the Variscan orogeny, i.e., they would not
512 have been exposed at surface until Carboniferous time.

513

514 **5.2. Potential sources of detrital zircon grains in the Ordovician-Devonian** 515 **sedimentary sequence of the Ossa-Morena Zone**

516 The most important –and unexpected- finding of our study is the recognition of an
517 important late Stenian – early Tonian detrital zircon population peaked at ≈ 1 Ga, with
518 percentages of ≈ 20 % in all of the studied samples. The presence of Tonian detrital
519 zircon grains in the OMZ was previously reported by Pereira et al. (2014), who noted a
520 progressive increase from the middle-late Cambrian (Ossa and Fatuquedo Formations)
521 to the Ordovician-Silurian (Colorada Formation). The appearance of a noticeable
522 Stenian – early Tonian detrital zircon population in Ordovician-Devonian OMZ rocks
523 has important paleogeographic and tectonic consequences, which will be discussed
524 separately in sub-section 5.3.

525

526 *5.2.1. Neo-Archean and Paleoproterozoic detrital zircon populations*

527 These populations are always present in the studied samples and depict three-order
528 peaks in the OMZ Ordovician-Devonian rocks (Fig. 5). These populations are
529 traditionally thought to have been derived from the WAC, where igneous and
530 metamorphic rocks of 1.8-2.1 and 2.5-2.75 Ga crop out extensively (Abati et al., 2010
531 and 2012 and references therein; Bea et al., 2020). However, other cratons and

532 metacratons cropping out in northern Africa (Arabian-Nubian shield, Sahara
533 metacraton, Tuareg shield) also contain rocks of these two age intervals (see for
534 instance Pereira et al., 2008, Drost et al., 2011, and Cambeses et al., 2017, for
535 compilations), and could also be invoked as primary sources of both Paleoproterozoic
536 and Neo-Archean detrital zircon grains. Therefore, the pre-Mesoproterozoic detrital
537 zircon populations in the Ordovician-Devonian OMZ rocks cannot be used in isolation
538 to unequivocally constrain their primary source.

539

540 *5.2.2. Cryogenian-Ediacaran detrital zircon population*

541 This population is dominant in all of the studied samples, representing \approx 40-50 % of the
542 concordant analyses (Figs. 5 and 7). It is also prevalent in upper Ediacaran- lower
543 Cambrian samples of the OMZ (Fernández-Suárez et al., 2002; Linnemann et al., 2008;
544 Pereira et al., 2011 and 2012a), and some Carboniferous samples (Dinis et al., 2018;
545 Pereira et al., 2012c and 2020). Beyond the OMZ, Cryogenian-Ediacaran detrital zircon
546 grains are the most conspicuous population in the CIZ, WALZ, CZ and GT MZ (and
547 counterparts through the Cantabrian arc (e.g., Martínez Catalán et al., 2020), as well as
548 in the Moroccan mesetas (Accotto et al., 2019 and 2021; Fig. 7). The source of this
549 population is unanimously attributed to a prominent Cadomian/Avalonian magmatic
550 arc that would have extended all along the northern Gondwanan continental margin at
551 Cryogenian-Ediacaran time (e.g., Murphy et al., 2018, and references therein). This arc
552 was apparently long-lasting (\approx 550 to 700 Ma) and laterally continuous along
553 thousands of km, developed over different segments of older continental crust from
554 the WAC to the Arabian Nubian shield (e.g., Fig. 3 of Murphy et al., 2018). The ϵ_{Hf}
555 isotope signature of the Cryogenian-Ediacaran detrital zircon grains can serve to
556 discriminate between zircon grains sourced from different segments of the
557 Cadomian/Avalonian magmatic arc. In this regard, the OMZ Cryogenian-Ediacaran ϵ_{Hf}
558 values compare quite well with those from the WAC and the Sahara metacraton (Fig.
559 8; Henderson et al., 2016 and references therein). In contrast, the Cryogenian-
560 Ediacaran ϵ_{Hf} signature from the Arabian-Nubian shield does not show the strong
561 vertical distribution on an age *versus* ϵ_{Hf} plot that was observed in the OMZ, being
562 dominated by supra-CHUR values (Henderson et al., 2016). Therefore, the ϵ_{Hf} data
563 suggest that the OMZ detrital zircon of Ediacaran-Cryogenian age might have been
564 derived from either the WAC or the Sahara metacraton segments of the
565 Cadomian/Avalonian magmatic arc. A more careful analysis of the ϵ_{Hf} data shows that
566 the older ages of the Cryogenian-Ediacaran population mainly plot in the supra-CHUR
567 field in the case of the OMZ and the Sahara metacraton, while the West African Craton
568 values do not (Fig. 8). The evidence is meager, but it might favor a dominant derivation
569 of Cryogenian-Ediacaran detrital zircon grains in the OMZ Ordovician-Devonian rocks
570 from a source located in the Sahara metacraton segment of the long-lived Cadomian/
571 Avalonian magmatic arc.

572

573 *5.2.3. Cambrian – Early Ordovician detrital zircon grains*

574 The very scarce Cambrian – Early Ordovician detrital zircon grains found in the studied
575 samples were probably derived from local igneous sources (e.g., Táliga-Barcarrota and
576 Salvatierra de los Barros plutons; Fig. 2; see Simancas et al., 2004, for a compilation),
577 formed during the rifting episode that preceded the passive margin stage. The ϵHf
578 isotopic signature of one of the Cambrian zircon grains is compatible with a juvenile
579 mantle-derived origin, while the remainder show sub-CHUR ϵHf values indicative of
580 recycling of older crustal rocks (Fig. 6). Therefore, the ϵHf signature supports the
581 provenance of these zircons from igneous rocks formed mostly by the partial melting
582 of the crust underlying the OMZ at that time.

583

584 **5.3. Late Stenian – early Tonian detrital zircon population: new insights into the**
585 **paleogeography and tectonics of the OMZ during late Neoproterozoic – early**
586 **Cambrian times**

587 The late Ediacaran – early Cambrian rocks of the OMZ include a dominant Cryogenian-
588 Ediacaran detrital zircon population and a subordinate Paleoproterozoic population,
589 with a systematic late Stenian – early Tonian (≈ 1 Ga) gap (Fernández-Suárez et al.,
590 2002; Linnemann et al., 2008; Pereira et al., 2011 and 2012a). However, the new U-Pb-
591 Hf data reported here show the presence of a significant ($\approx 20\%$) late Stenian – early
592 Tonian (≈ 1 Ga) detrital zircon population in Ordovician-Devonian OMZ rocks. To
593 explore the paleogeographic and tectonic consequences of this finding, the context
594 provided by additional geological data is essential.

595 The discussion that follows addresses two main issues. The first is establishing the
596 primary source of the late Stenian – early Tonian detrital zircon grains, bearing in mind
597 that despite the potential existence of intermediate sediment repositories (ISR)
598 (Pereira and Gama, 2021), a direct connection with a primary source seems justified by
599 the high percentage of this population in the studied Ordovician-Devonian OMZ
600 samples. In this respect, the middle Cambrian siliciclastic rocks from the Anti-Atlas
601 contain a meager 1 Ga detrital zircon population, which cannot be invoked as an ISR
602 for the Stenian – early Tonian detrital zircon grains found in the Ordovician-Devonian
603 OMZ rocks (Fig. 7), since sediment recycling could maintain the detrital zircon content
604 in the primary/intermediate source (Pereira and Gama, 2021) but not selectively
605 increase the content of one of the populations. The second issue to be addressed is
606 why late Stenian – early Tonian detrital zircon grains are not found in the late
607 Ediacaran – early Cambrian rocks of the OMZ, while they constitute a significant
608 population in the Ordovician-Devonian rocks.

609

610 *5.3.1. Primary late Stenian – early Tonian detrital zircon sources*

611 Two potential late Stenian – early Tonian primary sources exist, separated by the
612 Paleoproterozoic WAC: i) to the west the Grenville terranes (Henderson et al., 2016);
613 and ii) to the east the Sahara metacraton / Arabian-Nubian shield (Bea et al., 2010,
614 Cambeses et al., 2017). The fact that Mesoproterozoic detrital zircon ages in

615 Grenvillian terranes show a multi peak distribution between 1 and 1.6 Ga (e.g.,
616 Henderson et al., 2016; Pérez-Cáceres et al., 2017) can serve to cast doubt on a
617 Grenvillian provenance of late Stenian – early Tonian zircon grains in the Ordovician-
618 Devonian OMZ rocks, which invariably show a single peak population centered at 1 Ga
619 (Figs. 5 and 7). On this ground, the northern Africa cratonic areas seem to be more
620 plausible sources. Furthermore, the ϵHf isotopic signature of the late Stenian – early
621 Tonian detrital zircon population in the OMZ Ordovician-Devonian rocks shows a
622 remarkably similar distribution than those of the Sahara metacraton and the Arabian-
623 Nubian shield (Iizuka et al., 2013; Henderson et al., 2016; Fig. 8).

624

625 *5.3.2. Right-lateral translation of the Ossa-Morena Zone at late Ediacaran – early* 626 *Cambrian times*

627 In addition to the arguments given above, the geological context provides support for
628 the origin of late Stenian – early Tonian OMZ zircon grains being the Sahara
629 metacraton. The most plausible key mechanisms to explain the 1 Ga detrital zircon
630 provenance are putative late Ediacaran – early Cambrian right-lateral displacements of
631 the OMZ, modifying its position with respect to the CIZ and the Precambrian cratons
632 along the northern Gondwanan margin (Fig. 9). However, a first consideration to be
633 made is whether the 1 Ga detrital zircon grains were supplied by a relatively close
634 source or by a distant one. In this respect, the OMZ might have been located close to
635 the WAC all along the Ediacaran-Devonian timespan, with the 1 Ga detrital zircon
636 grains coming from a distant Sahara metacraton source. This possibility is undermined
637 by the systematic presence and relative abundance of Stenian – early Tonian zircon
638 grains with respect to Paleoproterozoic ones in the Ordovician-Devonian OMZ rocks
639 (Fig. 5). Distant but intermittent NE African detrital zircon sources have been invoked
640 to explain the presence of a 1 Ga population in Ordovician-Devonian rocks from the
641 Eastern Moroccan Meseta (Accotto et al., 2019). Nevertheless, in the Eastern
642 Moroccan Meseta samples the 1 Ga detrital zircon population is neither systematic nor
643 so abundant as in the similar age OMZ samples (Fig. 7). On this ground, we can
644 plausibly discard a long-lasting and distant NE African provenance for the 1 Ga detrital
645 zircon population in the Ordovician-Devonian OMZ rocks. Hence, in the following
646 paragraphs we will assume a relatively close Sahara metacraton source for the Stenian
647 – early Tonian detrital zircon population in the studied samples.

648 The proposed tectonic evolution is based on a critical review of three main issues: i)
649 the boundary between the OMZ and the CIZ (Figs. 1 and 2); ii) the correlation between
650 the OMZ and the STZ in Central Europe (Figs. 1A and 9A); and iii) the kinematics of the
651 Cadomian late Ediacaran subduction (Fig. 9B and C).

652 i) According to the Variscan evolution, the boundary between the OMZ and the CIZ is
653 the Badajoz-Córdoba shear zone (BCSZ; Fig. 2), with convergent-transcurrent left-
654 lateral kinematics and Devonian-to-earliest Carboniferous activity (Burg et al., 1981;
655 Azor et al., 1994 and 2019; Simancas et al., 2001). The BCSZ is a first-order tectonic
656 structure developed over an early Paleozoic rifting band, which eventually gave way to
657 MORB-type mafic rocks (e.g., Gómez-Pugnaire et al., 2003). These mafic rocks
658 underwent eclogite facies metamorphism during early Variscan tectonic evolution

659 (López Sánchez-Vizcaíno et al., 2003). Despite this dominant view, some authors have
660 located the OMZ/CIZ boundary north of the BCSZ, along the late Carboniferous
661 Pedroches Batholith (Fig. 2). Actually, the Ediacaran formations on both sides of this
662 batholith show significant differences (San José et al., 2004): southwards, the Serie
663 Negra formation contains Paleoproterozoic detrital zircon grains but not late Stenian –
664 early Tonian grains (Fernández-Suárez et al., 2002; Pereira et al., 2011 and 2012a;
665 Cambeses et al., 2017), while northwards the Lower Alcludian formation yielded a late
666 Stenian – early Tonian detrital zircon population (Talavera et al., 2012; Fernández-
667 Suárez et al., 2014). Therefore, the existence of a hidden tectonic line underneath the
668 Pedroches Batholith has been proposed (e.g., Pérez-Cáceres et al., 2017; Fig. 9A). This
669 concealed tectonic boundary would have been active at latest Ediacaran - earliest
670 Cambrian time, since the Cambrian formations are similar on both sides. Consequently,
671 this tectonic line should not be used to differentiate two Variscan zones of the Iberian
672 massif, i.e. the Variscan zonation should be based on the Paleozoic evolution.
673 Nevertheless, the hidden tectonic line seems to have played the main role in the
674 lateral translation of the OMZ to a location where late Stenian – early Tonian detrital
675 zircon grains were available (see below).

676 ii) The along-strike orogenic correlation between the OMZ in Iberia and the STZ in
677 Central Europe is implicit in most of Variscan reconstructions (e.g., Matte, 2001;
678 Simancas et al., 2005; Franke, 2006; Martínez Catalán et al., 2020). Thus, the outcrops
679 of the OMZ and the STZ would form part of a ribbon-shaped continental terrane (Fig.
680 9A), which, based on the nature of its basement, would be linked to the
681 Paleoproterozoic WAC. Direct evidence of the Paleoproterozoic age of the OMZ/STZ
682 basement has been found in the granulites of the Galicia Bank (Guerrot et al., 1989;
683 Gardien et al., 2000) and in Proterozoic outcrops of the Cherbourg-Trégor region in
684 NW France (Calvez and Vidal, 1978; Samson and D’Lemos, 1998; Inglis et al., 2004).
685 Furthermore, the presence of Paleoproterozoic detrital zircon grains but not late
686 Stenian – early Tonian grains characterizes the Ediacaran rocks of both OMZ and STZ
687 (Linnemann et al., 2014; Pérez-Cáceres et al., 2017). Finally, Sm-Nd isotope
688 geochemistry on the Ediacaran rocks of the OMZ and STZ supports an affinity between
689 these rocks and the Paleoproterozoic WAC (Fuenlabrada et al., 2020; Rojo-Pérez et al.,
690 2021).

691 iii) The broad-scale plate tectonic displacement related to the Cadomian orogeny is
692 very difficult to constrain, since almost no direct data on the subduction kinematics are
693 available. Nevertheless, important -though indirect- pieces of evidence can be inferred
694 from the strip geometry of the OMZ/STZ terrane, which connects westwards with the
695 Paleoproterozoic WAC (Fig. 9A), depicting a tail-shaped terrane. We suggest that the
696 OMZ/STZ terrane was primarily a continental ribbon formed by dextral strike-slip
697 displacement of a piece of the northern border of the Paleoproterozoic WAC.
698 Regarding direct Cadomian kinematic data, Linnemann et al. (2008) proposed a
699 tectonic frame characterized by oblique left-lateral Ediacaran subduction, which would
700 have turned into right-lateral displacements during latest Ediacaran - early Cambrian

701 times (Fig. 9B). Furthermore, considering the Pan-African suture cropping out in the
702 Anti-Atlas (El Hadi et al., 2010), a broad-scale kinematic scenario can be proposed by
703 combining the convergence vector of the Pan-African collision with the oblique left-
704 lateral Cadomian subduction on the northern border of the WAC (Fig. 9B; Linnemann
705 et al., 2008). Thus, the blockade of the Pan-African convergence at around 580 Ma (El
706 Hadi et al., 2010) would result in dextral kinematics, which, in turn, would have
707 displaced eastwards a strip from the northern border of WAC, namely the OMZ/STZ
708 terrane with Paleoproterozoic basement and Ediacaran magmatic arc rocks (Fig. 9B;
709 Pérez-Cáceres et al., 2017). Indeed, present-day oblique subduction in many areas
710 worldwide has usually given way to strain partitioning with decoupling of segments of
711 the overriding magmatic arc along transcurrent faults subparallel to the trench (e.g.,
712 Sumatra, Fitch, 1972; Kurile arc, Kimura 1986; Burma, Maung, 1987). Alternatively,
713 drawing on the recent evolution of western North America (e.g., Atwater, 1970), other
714 authors have proposed that the subduction of a ridge triggered the shift from
715 convergence to lateral displacement at the end of the Cadomian orogenesis (e.g.,
716 Linnemann et al., 2008; Sánchez-García et al., 2019). Independently of the reason
717 behind dextral displacements, a subsequent kinematic change would have occurred at
718 early-middle Cambrian time, when a rifting environment existed at the northern
719 Gondwanan margin, lasting until the Early Ordovician and marking the onset of the
720 Variscan cycle (e.g., Sánchez-García et al., 2019; Simancas, 2019).

721 In summary, the above-presented data lead us to propose that the OMZ-STZ domain is
722 part of an originally continuous and thin continental terrane with Paleoproterozoic
723 basement and located close to the WAC. This terrane would have been displaced
724 eastwards from the northern part of the WAC to a position close to the Sahara
725 metacraton. In SW Iberia, the transcurrent fault or shear zone permitting dextral
726 displacement would be currently concealed underneath the late Carboniferous
727 Pedroches Batholith (Figs. 2 and 9A), which bounds two contrasting types of Ediacaran
728 rocks. The transcurrent tectonics must have occurred at latest Ediacaran - earliest
729 Cambrian time.

730 As a corollary, we can now answer the two main issues raised at the beginning of this
731 section regarding the presence of late Stenian – early Tonian c detrital zircons in
732 Ordovician-Devonian OMZ rocks. The primary source of the ≈ 1 Ga detrital zircon
733 population was probably the Sahara metacraton. Before early Cambrian time, the
734 OMZ-STZ domain was located at the border of the WAC and received detrital zircon
735 grains from the Paleoproterozoic Eburnian crust and the Neoproterozoic magmatic arc,
736 developed at the northern Gondwanan margin. Once the dextral displacement that
737 placed the OMZ-STZ terrane close to the Sahara metacraton occurred, Mesoproterozoic
738 detrital zircon grains coming from this source started to appear in the Paleozoic
739 sedimentary rocks of the OMZ. The neighboring CIZ would have been located at
740 Ordovician-Devonian time close to the Sahara metacraton too, but westwards with
741 respect to the OMZ (Fig. 9B), in order to account for the important left-lateral

742 displacement related to the whole Variscan collisional evolution of SW Iberia (e.g.,
743 Pérez-Cáceres et al., 2016).

744 A latest Ediacaran - earliest Cambrian age of the OMZ-STZ dextral displacement fits
745 with late Stenian – early Tonian detrital zircon grains occurring in the Cambrian rocks
746 of this terrane. Nevertheless, this detrital zircon population was not detected in the
747 two lower Cambrian samples analyzed until now (Linnemann et al., 2008; Pereira et al.,
748 2012a). This is an open issue that needs more data to be properly explored, but
749 considering putative changes in the sedimentary paleogeography due to the Cambro-
750 Ordovician rifting seems to be an appropriate approach.

751

752 **6. Concluding remarks**

753 i) Ordovician-Devonian sedimentary rocks from the OMZ show a similar detrital zircon
754 content characterized by a dominant Cryogenian-Ediacaran population with a
755 maximum at 595-650 Ma, a noteworthy late Stenian – early Tonian population with an
756 almost invariable peak at ≈ 1 Ga, and a subordinate Paleoproterozoic population
757 usually depicting several peaks between 1.8 and 2.1 Ga.

758 ii) The Hf isotope signature of the studied samples features two well-defined vertical
759 arrays with age corresponding to the Cryogenian-Ediacaran and Stenian-Tonian detrital
760 zircon populations, which extend from supra-CHUR ϵ_{Hf} values of ≈ 10 down to sub-
761 CHUR values of ≈ -20 . Neo-Archean and Paleoproterozoic detrital zircon grains have a
762 vertical spread, with supra-CHUR ϵ_{Hf} values clustered at the ≈ 0 to 5 interval and sub-
763 CHUR values ranging from 0 to -10.

764 iii) The similarity of both detrital zircon spectra and Hf isotope signature in all of the
765 studied samples fits with the passive margin setting that other geological data
766 (particularly stratigraphy) also suggest for the OMZ at Ordovician–Devonian times.

767 iv) The OMZ and equivalent zones in Central Europe were located close to the WAC
768 along the northern Gondwanan margin at Ediacaran times, according to the presence
769 of ≈ 2 Ga orthogneiss remnants and the detrital zircon record of Ediacaran rocks,
770 characterized by a dominant Cryogenian-Ediacaran population, a subordinate
771 Paleoproterozoic one, and a systematic lack of late Stenian – early Tonian zircon
772 grains.

773 v) The unexpected presence of a ≈ 1 Ga detrital zircon population in Ordovician-
774 Devonian OMZ sedimentary rocks requires relocating this continental piece at that
775 time in a position close to the Sahara metacraton, which represents the most probable
776 source of late Stenian – early Tonian detrital zircon grains.

777 vi) The translation of the OMZ and equivalent zones eastwards from an original
778 position close to the WAC to an Ordovician-Devonian location close to the Sahara
779 metacraton probably occurred at latest Ediacaran – earliest Cambrian time in a dextral

780 strike-slip tectonic scenario that postdated Pan-African collision and Cadomian
781 subduction.

782

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798

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1350

1351 **Figure captions**

1352 **Figure 1: A)** Plate reconstruction at latest Paleozoic time to show the distribution of
1353 Variscan sutures, the deformation fronts, and the location of the Iberian Massif and

1354 other European Variscan domains. **B)** Main Variscan zones in the Iberian Massif, with
1355 location of the map shown in Figure 2 (red polygon).

1356 **Figure 2:** Schematic lithological map of the Ossa-Morena Zone and the southernmost
1357 Central Iberian Zone. The two blue stars locate the areas sampled in this study (EV: El
1358 Valle; CH: Cerrón del Hornillo). Two of the numerous Cambrian-Ordovician intrusive
1359 bodies have been marked (T-B: Táliga-Barcarrota plutons; S: Salvatierra de los Barros
1360 pluton). The boundary between the Ossa-Morena and the Central Iberian zones
1361 (Badajoz-Córdoba Shear Zone) has also been labelled (BCSZ).

1362 **Figure 3:** General stratigraphic column of the Ossa-Morena Zone (left), and detailed
1363 Ordovician-Devonian sequence adapted from Robardet et al. (1998) (right) with
1364 location of the studied samples (CH1, CH2, CH4, CH6, EV1 and EV2).

1365 **Figure 4:** Cathodoluminescence images of detrital zircon grains from the studied
1366 samples with the U-Pb analytical spots marked by circles and the obtained ages and ϵ_{Hf}
1367 values. High-U sectors of the crystals correspond to darker zones in the images.

1368 **Figure 5: A-F)** U-Pb detrital zircon age distribution in the studied Ordovician-Devonian
1369 rocks of the Ossa-Morena Zone. The results are presented as Kernel Density Estimates
1370 (KDE, black lines) and histograms (grey bars) on frequency *versus* age plots, as well as
1371 on pie-charts. Colors in the frequency *versus* age plots and pie-charts correspond to
1372 the different periods as specified in the legend.

1373 **Figure 6: A-F)** ϵ_{Hf} values *versus* U-Pb ages for the studied samples. CHUR: Chondritic
1374 Uniform Reservoir (Bouvier et al., 2008); NC: new crust (Dhuime et al., 2011). Colors in
1375 the plots correspond to the different periods as specified in Figure 5.

1376 **Figure 7:** Comparison among KDE of detrital zircon spectra for different ages and from
1377 different areas of the Variscides and surrounding regions. The following references
1378 have been used as primary sources of U-Pb detrital zircon ages: Abati et al. (2010);
1379 Accotto et al. (2019, 2021 and in press); Avigad et al. (2012); Díez Fernández et al.
1380 (2010); Fernández-Suárez et al. (2002 and 2014); Gutiérrez-Alonso et al. (2003);
1381 Linnemann et al. (2008 and 2011); Meinhold et al. (2011); Ordóñez-Casado (1998);
1382 Pereira et al. (2010a and b, 2011, 2012a and d); Schäfer (1990); Shaw et al. (2014);
1383 Talavera et al. (2012).

1384 **Figure 8:** Comparison of ϵ_{Hf} results in the studied samples (**A**) with those of the West
1385 African Craton (**B**), Sahara metacraton (**C**) and Arabian-Nubian shield (**D**), redrawn
1386 from Henderson et al. (2016). The following references have been used as primary
1387 sources of ϵ_{Hf} versus age data: Abati et al. (2012); Ali et al. (2013); Avigad et al. (2012);
1388 Be'eri-Shlevin et al. (2014); Gärtner et al. (2014); Iizuka et al. (2013); Linnemann et al.
1389 (2014); Meinhold et al. (2014); Morag et al. (2011 and 2012); Robinson et al. (2014).
1390 CHUR: Chondritic Uniform Reservoir (Bouvier et al., 2008); NC: new crust (Dhuime et
1391 al., 2011). Colors in the plots correspond to the different periods as specified in Figure
1392 5.

1393 **Figure 9: A)** Broad-scale reconstruction of the Variscan belt at latest Carboniferous
1394 time, emphasizing the connection of the Ossa-Morena / Saxo-Thuringian ribbon-
1395 shaped terrane with the West African Craton. The evidence supporting this connection
1396 is discussed in the text. **B)** Tentative kinematic scenario for the Cadomian orogenesis
1397 depicting the change from variably oblique subduction throughout the northern
1398 Gondwanan margin at Ediacaran time to a single dextral subduction at the vanishing
1399 stages of the Pan-African convergence (latest Ediacaran – earliest Cambrian). Once
1400 Pan-African subduction was blocked, strain partitioning would have given way to
1401 transcurrent tectonics, translating eastwards a strip of the Cadomian/Avalonian
1402 magmatic arc. The translated ribbon-shaped arc fragment represents the crustal
1403 basement of the Ossa-Morena and Saxo-Thuringian Variscan zones. The dextral
1404 kinematics would have relocated the Ossa-Morena/Saxo-thuringian zones close to the
1405 Sahara metacraton, which, in turn, would have sourced the detrital zircon grains
1406 analyzed in the Ordovician-Devonian rocks. $V_{NA/CS}$: assumed motion of the Cadomian
1407 subducted plate with respect to the North Africa complex basement (Tuareg shield,
1408 Sahara metacraton and Arabian-Nubian shield); $V_{WA/NA}$: assumed motion of the North
1409 Africa complex basement with respect to the West African Craton; $V_{WA/CS}$: assumed
1410 motion of the Cadomian subducted plate with respect to the West African Craton. See
1411 text for further explanations and alternative proposals.

1412

1413 **Table Captions**

1414 **Table 1:** Location of the samples and number and type of analyses carried out; (*)
1415 Total number of U-Pb analyses performed and concordant values (in bold).

1416 **Table 2:** Summary of U-Pb-Hf results.

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