



Hartree–Fock many-body perturbation theory for nuclear ground-states



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ABSTRACT

We investigate the order-by-order convergence behavior of many-body perturbation theory (MBPT) as a simple and efficient tool to approximate the ground-state energy of closed-shell nuclei. To address the convergence properties directly, we explore perturbative corrections up to 30th order and highlight the role of the partitioning for convergence. The use of a simple Hartree–Fock solution for the unperturbed basis leads to a convergent MBPT series for soft interactions, in contrast to the divergent MBPT series obtained with a harmonic oscillator basis. For larger model spaces and heavier nuclei, where a direct high-order MBPT calculation is not feasible, we perform third-order calculations and compare to advanced *ab initio* coupled-cluster results for the same interactions and model spaces. We demonstrate that third-order MBPT provides ground-state energies for nuclei up into the tin isotopic chain in excellent agreement with the best available coupled-cluster calculations at a fraction of the computational cost.

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1. Introduction

The solution of the Schrödinger equation for atomic nuclei using realistic nuclear interactions is at the heart of *ab initio* nuclear structure theory. In practice this problem is addressed by constructing approximate methods for a truncated, i.e., finite-dimensional Hilbert space. However, for the calculation of ground-state energies of heavy nuclei significant algorithmic and computational efforts are needed. There is a plethora of different *ab initio* methods, e.g., coupled cluster (CC) theory [1–6], in-medium similarity renormalization group (IM-SRG) [7–11], or self-consistent Green's function methods [12–14]. However, it is desirable to have an alternative, light-weight framework available. A conceptually simple method to solve for the eigenenergies of a physical system is many-body perturbation theory (MBPT) [15–17]. A perturbative treatment is the standard approach for many problems from different fields of theoretical physics. The advantage of MBPT compared to other *ab initio* approaches is its simplicity, which also allows for straightforward generalizations to excited states and open-shell nuclei [18] without the need of sophisticated equation-of-motion techniques. The reasons, why MBPT usually is not considered as

an *ab initio* technique, are convergence issues of the underlying perturbation series. Several studies of high-order MBPT based on Slater determinants constructed from harmonic oscillator (HO) single-particle states (HO-MBPT) have shown that the perturbation series is divergent in almost every case [18,19]. In such cases one heavily relies on the use of resummation techniques, e.g., Padé approximants, that enable a robust extraction of observables although the perturbative expansion diverges [19–21].

In this Letter, we formulate MBPT based on Hartree–Fock (HF) single-particle states (HF-MBPT), and, for the first time, investigate the convergence behavior of the perturbation series up to 30th order. We compare the ground-state energies of ⁴He and ^{16,24}O to results from exact diagonalizations in the configuration interaction (CI) approach using the same model space [22–24]. Based on the rapidly converging perturbation series resulting from the use of HF basis states, we study ground-state energies of selected closed-shell medium-mass and heavy nuclei at third-order MBPT, and compare to recent CC calculations [6].

2. The nuclear Hamiltonian

For all following investigations we start from the chiral nucleon–nucleon (NN) interaction at next-to-next-to-next-to leading order (N³LO) by Entem and Machleidt [25] combined with the three-nucleon (3N) interaction at next-to-next-to leading order N²LO

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in its local form [26] with three-body cutoff $\Lambda_{3N} = 400 \text{ MeV}/c$. Additionally, we use the similarity renormalization group (SRG) to soften the Hamiltonian through a continuous unitary transformation controlled by a flow parameter α [27–31]. In principle this transformation induces beyond-3N operators, which we have to neglect. To avoid the complication of dealing with explicit 3N interactions, we make use of the normal-ordered two-body approximation (NO2B) of the 3N interaction that was found to be very accurate for medium-mass nuclei, see Refs. [32,33]. For the matrix-element preparation we adopt the procedure introduced in Ref. [6], in particular, we use large SRG model spaces and exploit the iterative scheme where necessary. Thus, the matrix elements and the treatment of the chiral NN+3N interaction are identical to Ref. [6] and we can compare directly to the CC results presented there.

3. Many-body perturbation theory

The essence of Rayleigh–Schrödinger perturbation theory is the definition of an additive splitting, referred to as partitioning, of a given Hamiltonian H into an unperturbed part H_0 and a perturbation W . By introducing an auxiliary parameter λ we obtain a one-parameter family of operators,

$$H_\lambda = H_0 + \lambda W, \quad (1)$$

where the perturbation is defined by $W = H - H_0$. As ansatz for the solution of the eigenvalue problem of H we take a power series expansion of the energy and eigenstate in terms of an auxiliary parameter λ , where the expansion coefficients are given by the energy corrections and state corrections, respectively. We choose H_0 to be the HF Hamiltonian arising from an initial NN+3N interaction. We have shown in Refs. [19,18] that high-order MBPT corrections are accessible by means of a recursive scheme, allowing for detailed investigations of the convergence characteristics of the perturbation series. In general we cannot expect that a perturbation series is convergent [34–36], but one can exploit resummation-theory techniques to extract information on the observables of interest. There are different schemes and transformations that can be used to extract, e.g., the ground-state energy from a divergent expansion [37–39]. Padé approximants have proven to be particularly useful in the treatment of high-order HO-MBPT [19,18]. Additionally, they are well-known to mathematicians especially in the field of convergence acceleration [21,36,37]. However, the calculation of energy corrections up to sufficiently high orders is only feasible for light nuclei due to increasing computational requirements. When proceeding to the medium-mass region, one has to choose a different strategy. Depending on the rate of convergence, one might expect low-order partial sums of the perturbation series to be reasonable approximations to the exact ground-state energy. Having only low-order information available, resummation methods are less effective, because one is limited to a small number of approximants that yield valid approximations only if the transformed sequence converges sufficiently fast [18]. However, an alternative is to exploit the freedom in the partitioning, i.e., the choice of the unperturbed basis, to improve the convergence of the perturbation series.

We specifically explore a partitioning defined by a prior HF calculation, which optimizes the single-particle basis [17]. Note that the HF ground-state energy corresponds to the first-order partial sum,

$$E_{\text{HF}} = E^{(0)} + E^{(1)}. \quad (2)$$

Therefore, the first contribution to the correlation energy appears in second-order HF-MBPT. The second- and third-order contribu-

tions to the ground-state energy for a two-body operator W are given by [40]

$$\begin{aligned} E^{(2)} &= \frac{1}{4} \sum_{ab}^{>\epsilon_F} \sum_{ij}^{<\epsilon_F} \frac{\langle ab|W|ij\rangle \langle ij|W|ab\rangle}{(\epsilon_a + \epsilon_b - \epsilon_i - \epsilon_j)}, \\ E^{(3)} &= \frac{1}{8} \sum_{abcd}^{>\epsilon_F} \sum_{ij}^{<\epsilon_F} \frac{\langle ab|W|ij\rangle \langle ij|W|cd\rangle \langle cd|W|ab\rangle}{(\epsilon_a + \epsilon_b - \epsilon_c - \epsilon_d)(\epsilon_a + \epsilon_b - \epsilon_i - \epsilon_j)} \\ &\quad + \frac{1}{8} \sum_{ab}^{>\epsilon_F} \sum_{ijkl}^{<\epsilon_F} \frac{\langle ab|W|ij\rangle \langle ij|W|kl\rangle \langle kl|W|ab\rangle}{(\epsilon_a + \epsilon_b - \epsilon_i - \epsilon_j)(\epsilon_a + \epsilon_b - \epsilon_k - \epsilon_l)} \\ &\quad + \sum_{abc}^{>\epsilon_F} \sum_{ijk}^{<\epsilon_F} \frac{\langle ab|W|ij\rangle \langle cj|W|kb\rangle \langle ik|W|ac\rangle}{(\epsilon_a + \epsilon_b - \epsilon_i - \epsilon_j)(\epsilon_a + \epsilon_c - \epsilon_i - \epsilon_k)}. \end{aligned} \quad (3)$$

In the third-order energy correction the first, second, and third term are called particle–particle (pp), hole–hole (hh), and particle–hole (ph) correction, respectively. The ϵ_i correspond to the HF single-particle energies and all matrix elements are taken to be antisymmetrized. Summation indices a, b, c, \dots correspond to particle indices, i.e., states above the Fermi level ϵ_F , whereas i, j, k, \dots correspond to hole indices up to the Fermi level. The zero and one-body parts of the normal-ordered Hamiltonian only enter in the first-order energy correction. Brillouin’s theorem states that there is no mixing of the HF state with singly-excited determinants [17] and by orthogonality the zero-body part is only present in the expectation value of the perturbation. In principle, the derivation of energy corrections beyond third order is straightforward. However, considering a diagrammatic approach in terms of Hugenholtz diagrams, the number of contributing diagrams at a given perturbation order p increases rapidly [41] such that it becomes challenging to go beyond third-order in practice. Additionally, terms from higher-order corrections involve expressions that are notoriously hard to compute, because their efficient implementation, e.g., by means of BLAS-enabled matrix operations, is not obvious. The computational power needed to perform third-order MBPT calculations up into the medium-mass region can be provided by a single compute node within 1–3% of the computing time needed for state-of-the-art CC calculations.

4. Convergence characteristics of Hartree–Fock many-body perturbation theory

We start by comparing perturbation series from HO and HF-MBPT, and we focus on their convergence characteristics and sensitivity to the SRG flow parameter. In Fig. 1 we present a direct comparison of the order-by-order behavior for the two partitionings up to 30th order for ^{16}O . For these high-order calculations we use an N_{max} -truncation of the many-body model space, similar to the no-core shell model (NCSM) [23]. The left-hand column of Fig. 1 shows the high-order partial sums and the right-hand column the individual energy corrections for each order. Panel (a) shows the partial sums from HO-MBPT for a sequence of model spaces with fixed SRG flow parameter $\alpha = 0.08 \text{ fm}^4$. The partial sums are divergent for every model space. The divergence is also apparent from panel (c) which reveals exponentially increasing energy corrections. In contrast, panel (b) shows the partial sums arising from HF-MBPT that are convergent for all model spaces. Furthermore, the converged values agree with direct CI results. As seen in panel (d), the energy corrections are exponentially suppressed for higher orders, giving rise to a robust convergence.

In Fig. 2 we show the high-order partial sums and energy corrections in HF-MBPT for different SRG flow parameters. Panels (a), (b) and (c) show the convergent perturbation series for

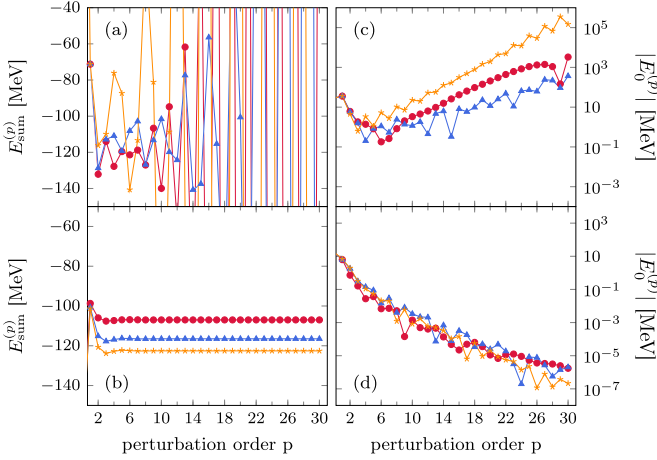


Fig. 1. Partial sums for the ground-state energy of ^{16}O in the HO basis (a) and the HF basis (b) for the NN+3N-full interaction with $\alpha = 0.08 \text{ fm}^4$ and model-space truncation parameters $N_{\text{max}} = 2$ (●), 4 (▲), and 6 (★). The corresponding energy corrections for each order are displayed in panels (c) and (d), respectively. All calculations are performed at frequency $\hbar\Omega = 24 \text{ MeV}$.

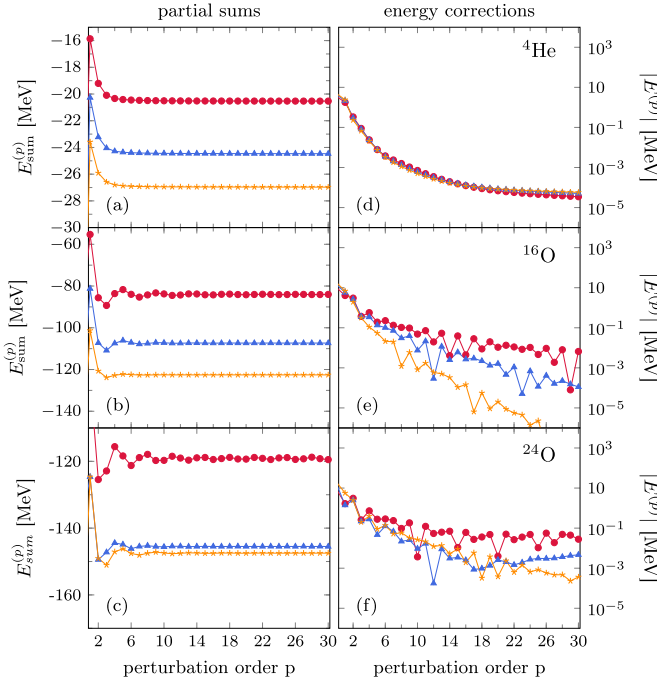


Fig. 2. Partial sums for varying flow parameters in HF-MBPT for ^4He (a), ^{16}O (b), and ^{24}O (c). The corresponding energy corrections are shown in panels (d), (e) and (f), respectively. The model space is truncated at $N_{\text{max}} = 6$ for ^4He and ^{16}O , and at $N_{\text{max}} = 4$ for ^{24}O . The flow parameters for the different data sets are $\alpha = 0.02 \text{ fm}^4$ (●), 0.04 fm^4 (▲), and 0.08 fm^4 (★). All calculations use a NN+3N-full interaction and frequency $\hbar\Omega = 24 \text{ MeV}$.

^4He , ^{16}O and ^{24}O , respectively. The calculations are performed for fixed $N_{\text{max}} = 6$ for ^4He , ^{16}O and $N_{\text{max}} = 4$ for ^{24}O , and the flow-parameter dependence of the absolute energies results from the varying degree of convergence with respect to the many-body model space.

A more interesting flow-parameter dependence can be observed for the individual energy corrections in panels (d), (e), and (f). There is a clear dependence of the convergence rate on the flow parameter for the oxygen isotopes. For ^{16}O the series converges exponentially in all three cases. The larger the flow parameter, i.e. the softer the Hamiltonian, the more rapid the convergence—as

Table 1

Ground-state energies for ^4He , ^{16}O and ^{24}O in units of [MeV] obtained in HF-MBPT for different orders up to $p = 30$ and in CI calculations with NN+3N-full interactions for different flow parameters α . The model spaces are truncated by $N_{\text{max}} = 6$ for ^4He and ^{16}O and $N_{\text{max}} = 4$ for ^{24}O . The HO frequency is $\hbar\Omega = 24 \text{ MeV}$.

		α [fm ⁴]		
		0.02	0.04	0.08
^4He	$E_{\text{sum}}^{(2)}$	-19.204	-20.269	-23.588
	$E_{\text{sum}}^{(3)}$	-20.334	-23.224	-26.589
	$E_{\text{sum}}^{(10)}$	-20.507	-24.444	-26.947
	$E_{\text{sum}}^{(20)}$	-20.526	-24.462	-26.964
	$E_{\text{sum}}^{(30)}$	-20.537	-24.469	-26.971
	CI	-20.539	-24.483	-26.994
	^{16}O	$E_{\text{sum}}^{(2)}$	-85.620	-107.241
$E_{\text{sum}}^{(3)}$		-89.315	-110.861	-123.863
$E_{\text{sum}}^{(10)}$		-83.780	-107.199	-122.561
$E_{\text{sum}}^{(20)}$		-84.180	-107.341	-122.577
$E_{\text{sum}}^{(30)}$		-84.018	-107.331	-122.577
CI		-84.043	-107.330	-122.577
^{24}O		$E_{\text{sum}}^{(2)}$	-125.460	-124.459
	$E_{\text{sum}}^{(3)}$	-122.880	-126.670	-151.059
	$E_{\text{sum}}^{(10)}$	-119.705	-121.233	-147.446
	$E_{\text{sum}}^{(20)}$	-119.335	-121.314	-147.508
	$E_{\text{sum}}^{(30)}$	-119.483	-120.948	-147.489
	CI	-119.131	-120.947	-147.488

might be naively expected. For ^{24}O the behavior is slightly more complicated. For the softest interaction with $\alpha = 0.08 \text{ fm}^4$ there is still a clear exponential convergence. However, for the harder interactions, i.e., $\alpha = 0.02$ and 0.04 fm^4 , we observe no systematic decrease of the high-order perturbative contributions anymore, they remain approximately constant and cause a small-amplitude oscillatory behavior of the partial sums. However, even in these cases we can easily extract a robust estimate for the asymptotic value. In the case of ^4He the suppression is independent of α and we observe the same rapid convergence for all interactions.

The numerical values of the partial sums for selected orders of HF-MBPT for the three nuclei and the different flow parameters are summarized in Table 1 together with the results of direct CI calculations for the same Hamiltonians and model spaces. The higher-order partial sums are in good agreement with the CI results—in most cases the deviation of the ground-state energy is much smaller than 0.1%.

Based on our detailed analysis of high-order HF-MBPT and due to the exponential suppression of the energy corrections, we can take low-order partial sums as a reasonable approximation to the converged results. This motivates the investigation of third-order HF-MBPT for medium-mass and heavy closed-shell nuclei in the following.

5. Explicit summation for heavy nuclei

For heavier nuclei and larger model spaces we cannot compute the high-order perturbation series explicitly and, thus, we cannot investigate the convergence characteristics explicitly. We can, however, evaluate the perturbative contributions up to third order very efficiently. To demonstrate the validity of a low-order perturbative approximation, we need to compare our results to established *ab initio* techniques, in our case, CC calculations with sophisticated triples corrections.

We consider a sequence of closed-shell nuclei ranging from ^4He to ^{132}Sn and perform calculations in second and third-order HF-MBPT in a large model space truncated with respect to the single-particle principal quantum number $e_{\text{max}} = 12$. We restrict ourselves to SRG-evolved Hamiltonians with flow parameter $\alpha =$

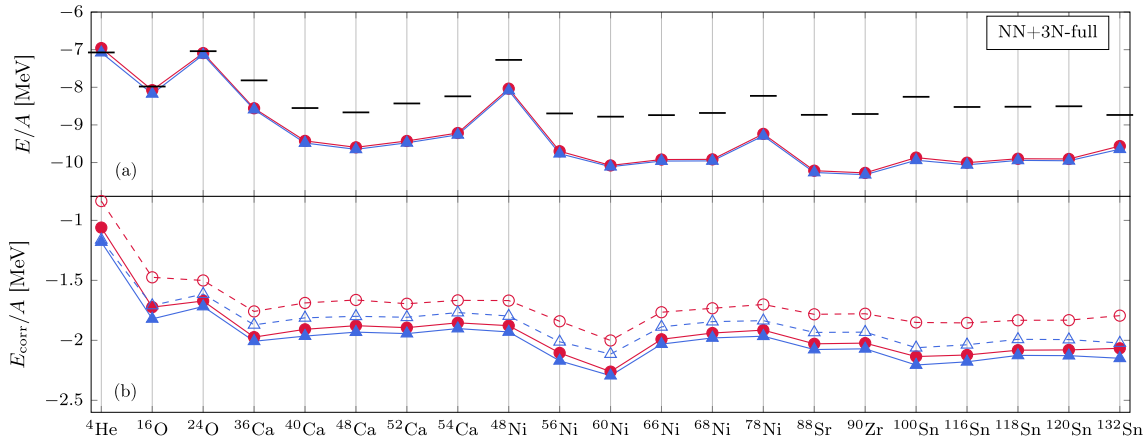


Fig. 3. Panel (a) shows the ground-state energies per nucleon from third-order HF-MBPT (●) in comparison to CR-CC(2, 3) (▲) results for selected closed-shell nuclei. Panel (b) shows the correlation energies per nucleon, $E_0^{(2)}$ (○) as well as $E_0^{(2)} + E_0^{(3)}$ (●) for HF-MBPT. Additionally, the correlation energy per nucleon for CCSD (▲) and CR-CC(2,3) (▲) are shown. All calculations are performed with the NN+3N-full interaction with $\alpha = 0.08 \text{ fm}^4$, $\hbar\Omega = 24 \text{ MeV}$ in an $e_{\text{max}} = 12$ model space. Experimental values are indicated by black bars.

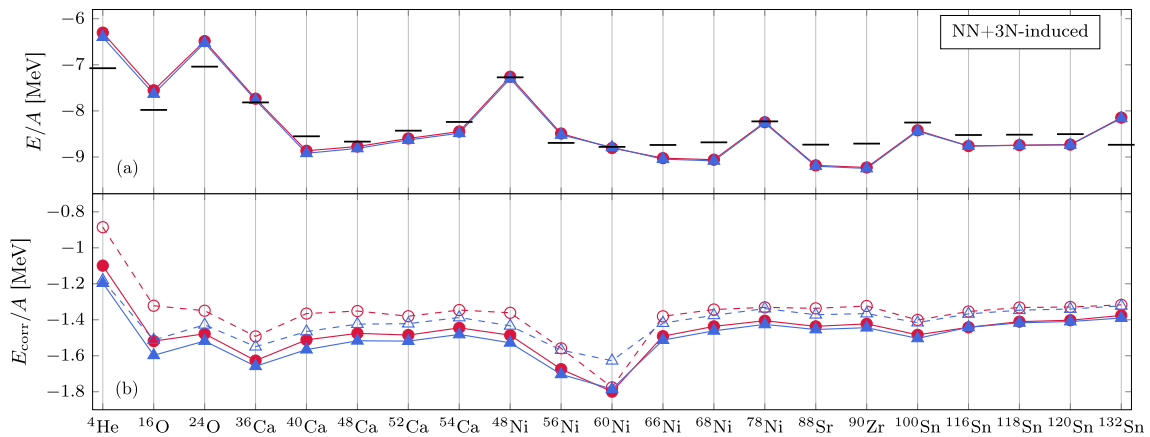


Fig. 4. Ground-state and correlation energies for the NN+3N-induced interaction. All other parameters as in Fig. 3.

0.08 fm^4 , which were used extensively in previous calculations and showed favorable order-by-order convergence in our high-order studies. We cannot perform CI calculations for these large spaces, however, the coupled-cluster framework has proven to provide accurate results for ground-state energies of closed-shell nuclei [1–4]. We compare the HF-MBPT results to recent CC calculations at the CCSD and the CR-CC(2, 3) level [5,6,42]. Starting from a HF reference state, this approach provides a complete inclusion of singly and doubly excited clusters on top of the reference state and, in the case of CR-CC(2, 3) an approximate non-iterative inclusion of triply excited clusters [43–46].

In Figs. 3 and 4 the ground-state energies per nucleon (a) as well as the correlation energies $E_{\text{corr}} = E - E_{\text{HF}}$ per nucleon (b) from HF-MBPT and CR-CC(2, 3) are depicted for an initial chiral NN+3N and an initial chiral NN interaction. The SRG-induced three-nucleon contributions are taken into account in both cases, leading to the NN+3N-full and NN+3N-induced interactions, respectively.

These figures show a remarkable result: The binding energies in third-order HF-MBPT and CR-CC(2,3) are in excellent agreement with each other. The relative differences are much smaller than 1% in most cases. The same observation holds for the correlation energy, i.e., the correction to the HF energy. The third-order energy corrections contribute approximately 0.2 MeV to the overall binding energy per nucleon and are, therefore, one order of magnitude smaller than the second-order correction but not negligible.

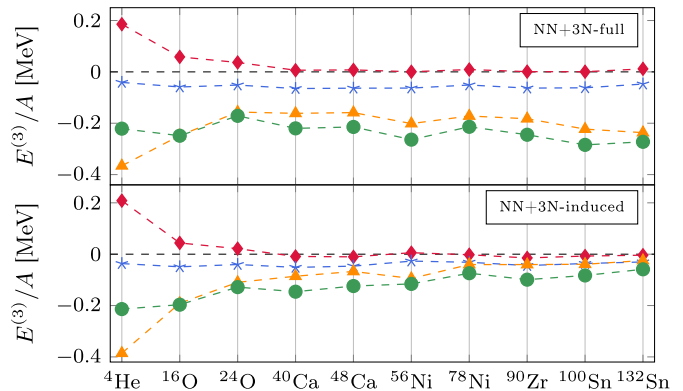


Fig. 5. Individual contributions of the diagrams appearing at third-order perturbation theory. Shown are the contributions per nucleon from the pp diagram (◆), the hh diagram (★), and the ph diagram (▲). The overall contribution of the third-order correction is depicted in (●). The first panel corresponds to the NN+3N-full interaction and the second panel to a NN+3N-induced interaction with $\alpha = 0.08 \text{ fm}^4$, $\hbar\Omega = 24 \text{ MeV}$, and $e_{\text{max}} = 12$.

The third-order energy contribution (3) consists of three terms corresponding to three Hugenholtz diagrams. Fig. 5 disentangles their individual contributions to the overall third-order energy correction. The contribution of the pp, hh, and ph terms vary mildly over the entire mass range. For the tin isotopes, the total third-

order energy correction contributes 3% to the overall binding energy and is not negligible. In particular we observe that the main contribution to the third-order energy correction arises from the ph diagram. In the case of a NN+3N-induced interaction all three terms are suppressed with increasing mass number, whereas for the NN+3N-full interaction the ph contribution remains sizeable. These systematic dependencies of the individual third-order contributions on the input Hamiltonian show that a partial inclusion of selected third-order terms may lead to wrong estimates.

6. Conclusions

We have discussed Rayleigh–Schrödinger MBPT as an efficient approach to compute ground-state energies for closed-shell nuclei throughout the medium-mass region. The use of a HF basis has enabled us to overcome convergence problems that generally arise in HO-MBPT. Investigating ^{16}O in different model spaces showed convergent partial sums when using HF-MBPT coinciding with the results from explicit CI calculations. Additionally, we found systematic dependencies of the convergence rate on the SRG parameter, i.e. the softness of the interaction, in the case of ^{16}O and ^{24}O . Thus, in HF-MBPT we can improve the convergence behavior of the perturbation series by further evolving the Hamiltonian, whereas the divergence of the HO-MBPT series will not be cured.

We can identify a hierarchy of elements influencing the convergence properties of the perturbation series. Defining a partitioning, or equivalently, defining a starting point for the perturbative treatment is the most important part. The radically different behavior of the perturbation series in HF-MBPT and HO-MBPT shows that the order-by-order convergence of the partial sums is very sensitive to the partitioning. When using HF-MBPT we can improve the convergence by using softer interactions corresponding to larger SRG flow parameters. Even for HF basis sets harder interactions can spoil convergence. The ‘softness’ of the interaction has been characterized in terms of Weinberg eigenvalues, which are connected to the spectrum of two-body Green’s functions [47–49]. Similar expressions also appear in the equations for the first-order state correction. Though the general connection seems obvious, one should be careful with conclusions about the convergence of MBPT for a finite nucleus based on the softness of the interaction. Our work has shown that the partitioning is key for convergence. Our observation that the convergence of HF-MBPT deteriorates for harder interactions could simply be explained by the fact that the unperturbed HF solution becomes a much worse approximation for the ground state in these cases.

The superior convergence properties of HF-MBPT is a motivation to use low-order approximations to investigate nuclei in the medium-mass region. We have validated these low-order approximations by comparing to the most sophisticated CC calculations and found excellent agreement of third-order HF-MBPT and CR-CC(2,3) at the level of better than 1%. The consistency of high-order partial summations with exact CI diagonalizations as well as the agreement of low-order summations with CC results may qualify HF-MBPT as an *ab initio* approach. However, the strong dependence of the convergence on the partitioning should be a reason for caution. The HF partitioning seems to be robust for sufficiently soft interactions, but there is no formal guarantee for convergence.

The great advantage of low-order HF-MBPT is its simplicity: Computationally, the third-order calculations are much cheaper than CC or IM-SRG calculations. They are, therefore, ideal for survey calculations over a large range of medium-mass nuclei, e.g., to explore the ground-state systematics with new interactions. Formally, the underlying equations and algorithms are trivial compared to CC or IM-SRG. As a result, extensions to the description of excited states and open-shell nuclei are straight-forward. We have

demonstrated this already for light nuclei using high-order degenerate HO-MBPT [18]. Alternative multi-configurational formulations for open-shell nuclei are under investigation.

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