

21 **Abstract**

22 Much of the interior of the Tibetan Plateau is characterized by internal drainage, low
23 relief topography and high altitude. How and when this landscape formed is
24 controversial. In this study, we use new zircon U-Pb data and low temperature
25 thermochronological data (apatite and zircon (U-Th/He), apatite fission track (AFT))
26 from the Late Cretaceous to Cenozoic Nima basin sedimentary rocks and Xiabie
27 granite in the adjacent Muggar Thrust hanging wall (part of the regional
28 Shiquanhe-Gaize-Amdo thrust system), to determine the palaeodrainage and timing of
29 exhumation in the region. Individual AHe and ZHe cooling ages range from 9 to 60
30 Ma and 58-118 Ma, and the AFT ages range from 30-90 Ma. The thermal history
31 derived from the Northern Nima Basin sediments and Xiabie granite require a period
32 of exhumation between 70 and 40 Ma in the thrust fault hanging wall, and 40 to 30
33 Ma in the Nima Basin. Across the region this event was followed by low rates of
34 exhumation and the deposition of locally-sourced sediment, lacustrine and evaporitic
35 deposits that are indicative of an internal drainage system. We suggest that the
36 exhumation event is associated with development of thrust-elevated relief that may
37 have disrupted the drainage network favouring the development of an endorheic
38 system. This system, sediment accumulation and/or post-30 Ma tectonic quiescence
39 led to the generation of low relief topography.

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42 **Plain Language Summary**

43 The Tibetan Plateau is the highest altitude low topographic relief region on Earth.
44 There is no consensus on how the flat topography formed. Low-temperature
45 thermochronology is widely used to establish the time of exhumation of rocks to the
46 Earth's surface. In this study, we found that the rocks in the Nima area of the central
47 Tibetan Plateau experienced moderate exhumation from Late Cretaceous to early
48 Oligocene (70-30 Ma), after which time the exhumation rate slowed. Such low rates
49 imply slowing of local tectonic activity. The Sedimentology and Stratigraphy of the
50 basin indicate that an internal drainage system developed in the Nima area since 30
51 Ma. Given the arid climate conditions, we therefore propose that the fault activity
52 (70-30 Ma) represented by rapid exhumation resulted in a change to topography that
53 triggered the change of water systems. The low relief topography subsequently
54 developed in the arid and tectonically quiescent environments in this area.

55 **Keywords:** Low temperature thermochronology; exhumation; internal drainage; low
56 relief topography; Tibet

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62 **1 Introduction**

63 Much of the Tibetan Plateau is characterized by low relief topography (LRT) and high
64 elevation ([Fielding et al., 1994](#)). In the externally drained regions of the plateau
65 margins, such as the southeast Tibetan Plateau, the low relief surface has been
66 attributed to several mechanisms ([Fox et al., 2020](#)), ranging from deep Earth
67 processes such as lower crustal flow ([McKenzie and Jackson, 2002](#); [Clack et al.,](#)
68 [2006](#)), to drainage area reorganization resulting from horizontal tectonic shortening
69 ([Yang et al., 2015](#)) or propagating uplift ([Yuan et al., 2021](#)). However, there are
70 relatively few studies on the formation mechanisms of high-elevation, low-relief
71 surfaces in the internally drained regions of the modern Tibetan Plateau interior
72 ([Fielding et al., 1994](#); [Liu-Zeng et al., 2008](#); [Hetzel et al., 2011](#); [Han et al., 2019](#)).
73 [Liu-Zeng et al. \(2008\)](#) proposed development of a low-relief surface through
74 establishment of an endorheic river system that promoted bath-tub infilling of the
75 basins from adjacent topography in the central Tibetan Plateau, based on satellite
76 observation of modern plateau landforms. In contrast [Hetzel et al. \(2011\)](#) suggested
77 that peneplanation was driven by an external drainage system removing material to
78 the ocean at low elevation by 50 Ma, that was followed by rapid uplift. [Fielding et al.](#)
79 [\(1994\)](#) suggested that the LRT in the Tibetan interior formed in response to the
80 cessation of tectonic activity and onset of climate aridity in the late Cenozoic, and that
81 the slow erosion has managed to remove the pre-existing relief. In a further
82 complexity, the development of low relief has been asynchronous across the Tibetan
83 Plateau ([Liu-Zeng et al., 2008](#); [Rohrmann et al., 2012](#)), and has been shown to

84 develop soon after cessation of significant exhumation (e.g. [Law and Allen, 2020](#)).

85 Clearly determining the timing of formation of internal drainage, and interaction
86 between the evolution of the drainage system and the low relief topography, are
87 central to understanding the formation and evolution of the plateau landscape.

88 Low-temperature thermochronology (LTT) methods, such as zircon and apatite
89 (U-Th)/He and apatite fission track, record rock cooling between 200 and 60°C
90 ([Reiners et al., 2005](#)). They are commonly used to quantify the long-term exhumation
91 history of mountain belts and determine the climate- and tectonic-driven processes
92 responsible ([Reiners and Brandon, 2006](#)). Changes in exhumation may be driven by
93 erosion associated with tectonics or climate. However, as erosion requires potential
94 energy to remove material downhill, higher rates of exhumation require relief,
95 particularly in an internally-drained region where exhumation cannot be effected in
96 high altitude low relief settings by rivers incising to sea level. We use LTT to
97 reconstruct the denudation history of the study area, broadly associating periods of
98 higher exhumation with the presence of topographic relief. We combine our data with
99 sedimentological and palaeoclimate data ([DeCelles et al., 2007a, 2007b](#)) to identify
100 the timings and modes of formation of the present landscape.

101 The Nima and adjacent Lunpola basins are the largest Cenozoic non-marine basins in
102 Central Tibet, whilst the modern day Siling Co, located between the two basins, is the
103 largest saline lake in Tibet (Figure 1b). The Nima Basin has accumulated extensive
104 non-marine deposits from the Late Cretaceous to the Cenozoic, and several studies
105 have investigated the structural geology ([Kapp et al., 2007](#)), sedimentary

106 environments (DeCelles et al., 2007a), basin sedimentary provenance (DeCelles et al.,
107 2007a; Kong et al., 2019; Han et al., 2019), paleo-latitude (Meng et al., 2017), and
108 paleo-climate and paleo-elevation (DeCelles et al., 2007b; Wang and Wu, 2015; Deng
109 and Jia et al., 2018) of the region. Here we build on this earlier research with the first
110 LTT study applied to both basin sedimentary rocks and adjacent thrust belt to
111 elucidate the interaction between drainage system, low relief topography and tectonic
112 uplift. We combine zircon U-Pb ages, apatite and zircon (U-Th)/He (AHe and ZHe),
113 and fission track (AFT) thermochronology of granitic and sandstone clasts from the
114 Late Cretaceous to Oligocene Muggar unit in the North Nima Basin, and Xiabie
115 granite in the adjacent Muggar Thrust hanging wall, to elucidate the history of
116 exhumation and long-term evolution of central Tibet landscape.

117 **2 Geological background**

118 The central Tibetan Plateau is mainly comprised of the Lhasa and Qiangtang terranes
119 (Figure 1a). The Bangong–Nujiang and Indus-Yalu sutures in central and southern
120 Tibet, respectively, represent the Late Jurassic to Early Cretaceous Lhasa-Qiangtang
121 and early Cenozoic India-Lhasa collisions (Girardeau et al., 1984; DeCelles et al.,
122 2014; Ding et al., 2016; Zhu et al., 2016; Hu et al., 2016; Li et al., 2019). Marine
123 sedimentation ceased at ~118 Ma (Kapp et al., 2007) in the Bangong-Nujiang suture
124 region because of considerable Cretaceous tectonic shortening (Murphy et al., 1997;
125 Kapp et al., 2003; Volkmer et al., 2007). The N-dipping Shiquanhe-Gaize-Amdo
126 thrust system (SGA) follows the trace of the Bangong-Nujiang suture zone for ~1200
127 km from Shiquanhe to Amdo (Yin and Harrison, 2000) (Figure 1a). Although the SGA

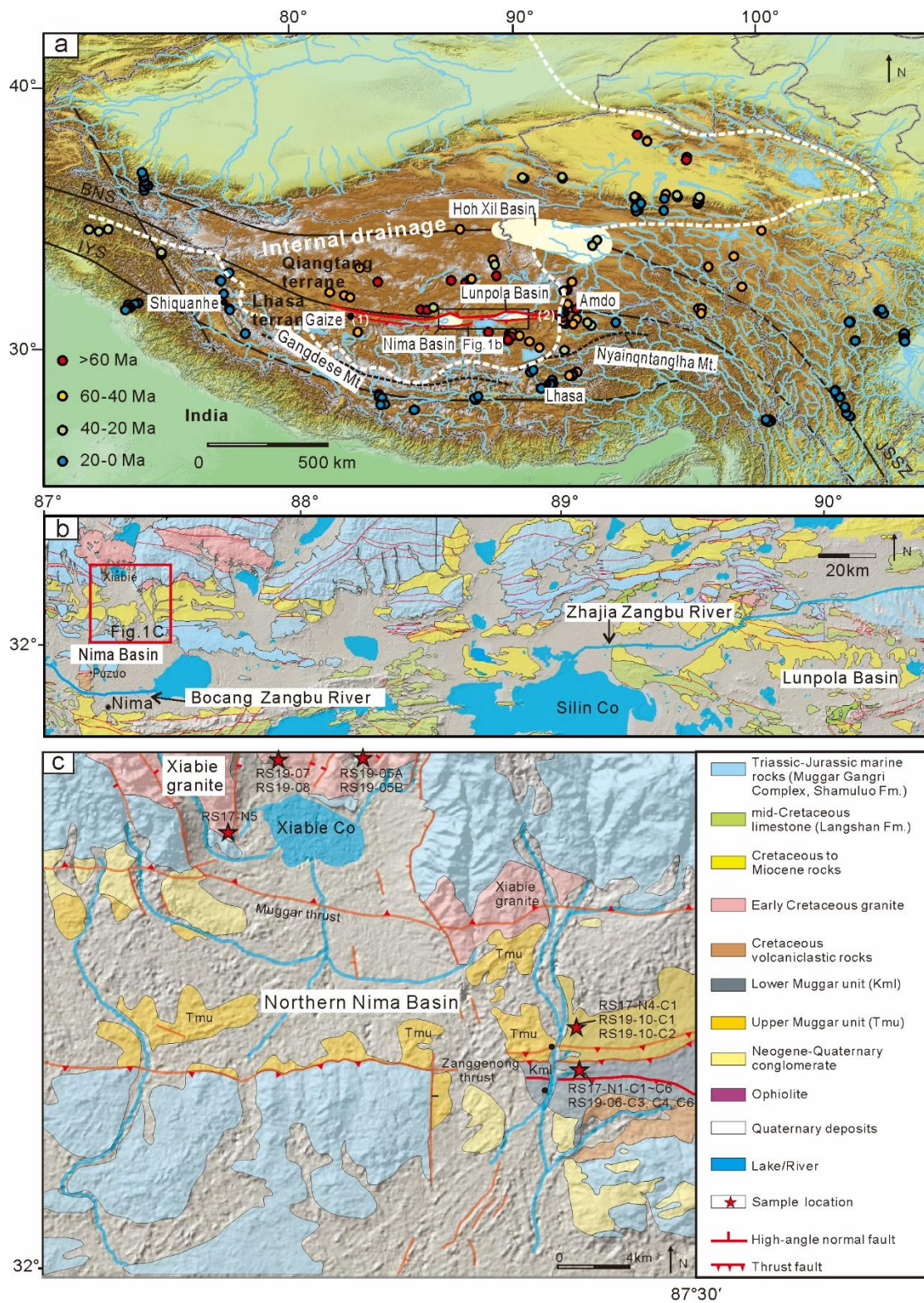
128 thrust initially slipped in the Early Cretaceous, likely driven by Lhasa-Qiangtang
129 collision, it was re-activated during Cenozoic times during the India-Lhasa collision
130 (e.g. [Kapp et al., 2005, 2007](#); [Dewey et al., 1988](#)). An Early Cretaceous granite belt
131 (100-130 Ma), extending from the west in Gaize through northern Nima to northwest
132 Amdo, along the Bangong-Nujiang suture (Figure 1b; [Zhu et al., 2016](#)) was most
133 likely related to slab rollback ([Yang et al., 2019](#)).

134 The Nima Basin lies within the Bangong-Nujiang suture zone (Figure 1a). The Early
135 Cretaceous Xiabie granite lies to the north of the Nima Basin ([Kapp et al., 2007](#)) in
136 the hanging wall of the Muggar Thrust, which forms the northern boundary to the
137 Basin. The Muggar Thrust is part of the regional system of N-dipping thrust faults of
138 the SGA thrust system ([Kapp et al., 2007](#)). The Muggar Thrust separates the Xiabie
139 granite and Triassic-Jurassic deep marine clastics and ophiolitic rocks (Muggar
140 Gangri Complex and Shamuluo Fm.) in the hanging wall, from Late Cretaceous to
141 Cenozoic basinal sedimentary rocks in the footwall. The southern boundary of the
142 Nima Basin is marked by the regionally extensive, southward-dipping Gaize-Siling
143 Co Thrust (Figure 1a).

144 The Nima Basin comprises Cretaceous-late Miocene sedimentary rocks, which are
145 almost exclusively non-marine, and include volcanic flows, tuffs, alluvial fan deposits,
146 shallow braided stream, eolian dunes and ephemeral lacustrine deposits ([DeCelles et
147 al., 2007a](#)). The Lower Muggar unit (Kml; [Kapp et al., 2007](#)) is only present in the
148 Northern Nima Basin (Figure 2). It consists of ~400 m of Upper Cretaceous to
149 Paleocene fluvial and eolian sandstones and conglomerates ([DeCelles et al., 2007a](#));

150 the deposits were dated using palynomorphs. The Upper Muggar unit (Tmu) (Figures
151 1c and 2; [Kapp et al., 2007](#)) is largely composed of lacustrine and fluvial red siltstone,
152 marl, evaporites and conglomerates that were deposited in the late Eocene-Oligocene,
153 dated on the basis of palynomorph data ([DeCelles et al., 2007a](#)). Lower and Upper
154 Muggar strata are disrupted by three closely spaced (< 1 km) S-dipping reverse faults,
155 collectively referred to as the Zanggenong Thrusts, likely active at ~40-20 Ma
156 ([DeCelles et al., 2007a](#); Figure 1c).

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159 Figure 1. (a) Topographic map of the Tibetan Plateau showing the tectonic framework,

160 modern drainage, and (coloured circles) compilation of timings of cessation of rapid

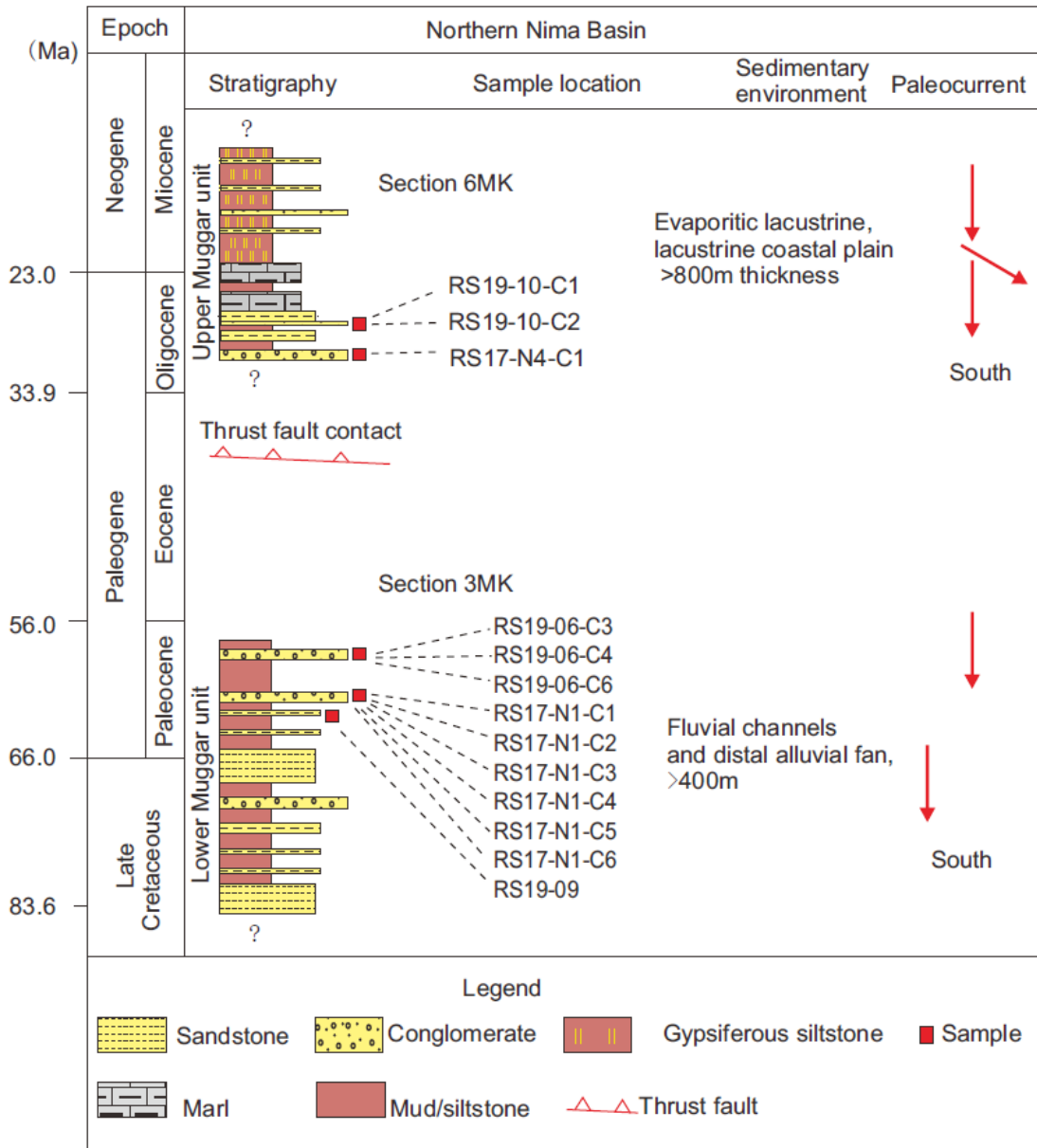
161 exhumation as determined from published thermal models derived from low

162 temperature thermochronology data (references provided in Table S1); (1):
163 Gaize-Siling Co Thrust; (2): Shiquanhe-Gaize-Amdo thrust system; IYS: Indus-Yalu
164 suture zone; BNS: Bangong-Nujiang suture zone; JSSZ: Jinshajiang suture zone. (b)
165 Geological map of the Nima-Lunpola area modified from [ITGS \(2002a, 2002b, 2003a,](#)
166 [2003b, 2006a, 2006b\)](#). (c) Geological map of the northern Nima area ([DeCelles et al.,](#)
167 [2007a](#)) showing sample locations.

168 **3 Sampling and Methods**

169 Eighteen samples were collected from the Nima area as detailed below (Figure 1c;
170 Table 1). Five samples (RS17-N5, RS19-05A, RS19-05B, RS19-08, RS19-07) were
171 collected from the Xiabie granite. One sandstone sample (RS19-09) was collected
172 from the Lower Muggar unit. The twelve conglomerate clasts included 6 granitoid
173 clasts (samples RS19-06-C3, -C4, -C6, RS17-N1-C4, -C5, -C6) and 3 sandstone clasts
174 (RS17-N1-C1, -C2, -C3) from the Lower Muggar unit, and three granitoid clasts
175 (RS19-10-C1, -C2, RS17-N4-C1) from the Upper Muggar unit (Table 1 and Figure 2).
176 Apatite fission track and/or (U-Th)/He analysis were carried out on all samples except
177 sandstone RS19-09, which contained no apatite. The granite (RS17-N5) was used for
178 zircon U-Pb analysis and four granitoid clasts (RS17-N4-C1; RS17-N1-C4, -C5, -C6)
179 were used for zircon (U-Th)/He and U-Pb analysis. In addition, we carried out U-Pb
180 analysis on detrital zircon from the Late Cretaceous sandstone (RS19-09) to
181 complement the existing data from [Han et al. \(2019\)](#).

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183

184 Figure 2. Stratigraphic logs, facies summary and paleocurrent data from the Northern
 185 Nima Basin from DeCelles et al. (2007a), showing sample locations for this study.
 186 Some samples were collected from the same stratigraphic level, such as granitic clasts
 187 (RS17-N1-C6) and sandstone clasts (RS17-N1-C1).

188

189 Table 1 Location, analytical methods employed and lithological details of samples from the Northern Nima Basin and Xiabie granite

Number	Sample	Longitude (°E)	Latitude (°N)	Sample position	Lithology	AFT	AHe	Zircon U-Pb	ZHe
1	RS17-N1-C4	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
2	RS17-N1-C5	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
3	RS17-N1-C6	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
4	RS19-06-C3	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
5	RS19-06-C6	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
6	RS19-06-C4	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
7	RS17-N1-C1	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
8	RS17-N1-C2	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
9	RS17-N1-C3	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
10	RS19-09	87.4171	32.1017	Lower Muggar Unit	sandstone			√	
11	RS17-N4-C1	87.4225	32.1139	Upper Muggar Unit	granitoid conglomerate clast	√	√	√	√
12	RS19-10-C1	87.4210	32.1119	Upper Muggar Unit	granitoid conglomerate clast	√	√		
13	RS19-10-C2	87.4210	32.1119	Upper Muggar Unit	granitoid conglomerate clast		√		
14	RS17-N5	87.2241	32.2186	Xiabie granite	granite	√	√	√	
15	RS19-05B	87.2973	32.2615	Xiabie granite	granite	√	√		
16	RS19-08	87.2694	32.2599	Xiabie granite	granite	√			
17	RS19-07	87.2945	32.2607	Xiabie granite	granite	√			
18	RS19-05A	87.3042	32.2447	Xiabie granite	granite	√			

Note: Tick √ represents the analysis technique carried out for every sample

190 **3.1 Zircon U-Pb**

191 Zircon and apatite crystals were extracted by standard crushing and separation
192 techniques. U-Pb dating of zircons was conducted using a LA-ICP-MS at the
193 Laboratory of Earth Surface Process and Environment in Nanjing University ([Zhang
194 et al., 2018](#)). Cathodoluminescence (CL) images were used to record the internal
195 structure of zircon grains and select target domains for in-situ U-Pb dating. U-Pb
196 dating of zircons was conducted using an ICP-MS (Agilent 7700x), which was
197 attached to a New Wave 193 nm laser ablation system, following the methods
198 described by [Jackson et al. \(2004\)](#). The laser beam diameter was 25 μm , with a 10-Hz
199 repetition rate and energy of 2–3 J/cm^2 . Zircon 91500 was used as the external
200 standard for isotopic fractionation correction, and GJ-1 zircon was used to monitor
201 instrumental reproducibility and stability. The average age of the former (1062 ± 17
202 Ma, $n=60$), and latter (603 ± 8 Ma, $n=20$) overlap the accepted ages ([Wiedenbeck et
203 al., 1995](#); [Liu et al., 2010](#)). The software GLITTER (version 4.4) was used to
204 calculate raw data ([Griffin, 2008](#)). Common Pb correction was implemented
205 following the method of [Andersen \(2002\)](#). Isoplot 4 software was used for plotting
206 concordance curves and calculating weighted mean age. Zircons with discordance
207 of $>10\%$ were excluded in data presentation and interpretation. The complete dataset
208 is reported in Tables S2 and S3.

209 **3.2 Apatite and zircon (U-Th)/He**

210 (U-Th)/He analysis was conducted at the Scottish Universities Environmental
211 Research Centre following procedures of [Foeken et al. \(2006\)](#). In order to avoid the

212 influence of physical properties of apatite and zircon grains such as fragmentation, on
213 age (Brown et al., 2013), we used inclusion-free euhedral crystals with one or two
214 terminations, and a diameter of at least 50 μm . At least five single grains were
215 analyzed for each sample. Grains were packed in Pt tubes and degassed in ultra-high
216 vacuum using an 808 nm diode laser. ^4He abundances determined with a Hiden
217 HAL3F quadrupole mass spectrometer (Foeken et al., 2006). Apatite grains were
218 degassed for 60 seconds at 500-600°C and zircon grains were degassed for 20 minutes
219 at 1100-1300°C (Foeken et al., 2006). The degassed apatite packets were spiked with
220 ^{235}U , and ^{230}Th and natural U and Th was measured by ICP-MS (Agilent 7500ce).
221 Where U and Th contents were within 3 times the background value, analytical
222 uncertainties are large and the data are excluded. Degassed zircon grains were
223 extracted from packets prior to dissolution and U-Th analysis were performed
224 following procedures in Dobson et al. (2008). Durango apatite (n=20) and Fish
225 Canyon Tuff zircon (n=15) standards were analysed; the average age of the former is
226 31.1 ± 1.7 Ma, the latter is 28.3 ± 1.1 Ma, indistinguishable from the ages reported in
227 the literature (Farley et al., 2000; Dobson et al., 2008). AHe and ZHe ages are
228 reported in Tables 2 and 3.

229 3.3 Apatite fission track determinations

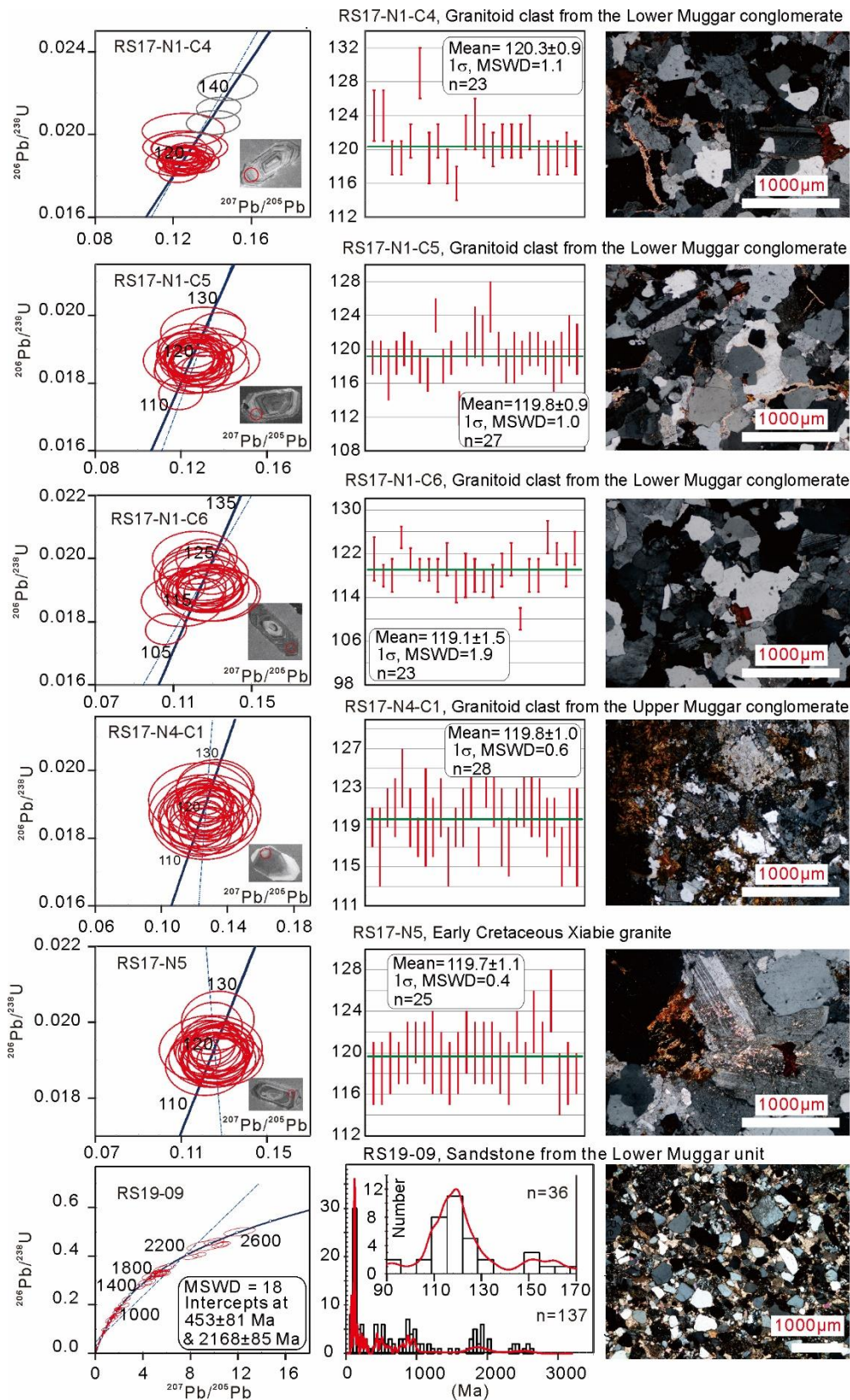
230 Apatite fission track analysis was performed at the GLOW laboratory, at the
231 University of Glasgow, using either the external detector method (Wildman et al.,
232 2016) or the LA-ICPMS, for U determination (Chew et al., 2016). The U
233 concentration derived from the LA-ICPMS method is consistently lower than that

234 determined from the EDM (Table 4); the discrepancy may be due to internal variation
235 in the concentration of U in apatite grains; given that both methods reproduce the ages
236 of the standards, the AFT ages of the unknown samples are considered accurate in
237 both cases. Apatite grains were mounted in epoxy resin and then polished to reveal a
238 horizontal, internal surface for each crystal; the polished mounts were etched in a
239 5.5M HNO₃ acid solution, for 20s at 20°C (Carlson et al., 1999). Tracks were counted
240 using an Axioplan Zeiss microscope (x1250), in both transmitted and reflected light;
241 for each crystal, three Dpar (the diameter of the track etched pits parallel to the c-axis
242 of crystal) measurements were also taken. Either the external detector method
243 (Wildman et al., 2016) or LA-ICPMS were used for U determination. The accuracy of
244 the isotopic determinations derived from the LA-ICP-MS was controlled using the
245 glass standard NIST SRM612 and Durango apatite. The weighted age of Durango
246 apatite is 32.1 ± 1.2 Ma (n=42). The zeta calibration, for the LA-ICPMS method,
247 was performed using Durango fragments (Chew et al., 2016). For the external
248 detector method (EDM), a mica external detector is used; the zeta determination used
249 Durango apatites (31.4 ± 0.5 Ma; Persano, 2003), Mt Dromedary (98.7 ± 0.6 Ma;
250 Persano, 2003), Fish Canyon Tuff (27.9 ± 0.5 Ma; Persano, 2003) and the Shap
251 Granite (75.0 ± 3.1 Ma; Luszczak et al., 2017) standards. Radial plots, created using
252 the free software Isoplot R (Vermeesch, 2018), were used to visualize and analyse the
253 single grain age dispersion (Figure S5).

254 **4 Results and interpretations**

255 **4.1 Zircon U-Pb data**

256 We obtained 270 U-Pb zircon dates from six samples. Three granitoid clasts
257 (RS17-N1-C4, RS17-N1-C5, RS17-N1-C6) from the Lower Muggar unit yield
258 weighted mean U-Pb zircon ages of 120.3 ± 0.9 Ma, 119.8 ± 0.9 Ma, and 119.1 ± 1.5
259 Ma (Figure 3). Granitoid clast RS17-N4-C1 from the Upper Muggar unit yields a
260 U-Pb zircon age of 119.8 ± 1.0 Ma. Xiabie granite sample RS17-N5 from the Muggar
261 Thrust hanging wall yields a weighted mean zircon U-Pb age of 119.7 ± 1.1 Ma. The
262 137 zircon U-Pb ages from a sandstone sample (RS19-09) from the Lower Muggar
263 unit exhibit a distinct cluster at 100-130 Ma, with ages extending to >2500 Ma
264 (Figure 3). The weighted mean age (67.0 ± 2.0 Ma) of the three youngest grains of
265 sample RS19-09 constrains the maximum depositional ages of the Lower Muggar unit
266 ([Dickinson and Gehrels, 2009](#)).



267

268 Figure 3. Zircon U-Pb dating results and thin section photos for granitoid clasts
 269 (RS17-N1-C4, N1-C5, N1-C6, N4-C1), and the sandstone (RS19-09) and granite
 270 sample (RS17-N5) from the Northern Nima Basin and Xiabie granite.

271 4.2 Apatite/zircon U-Th/He data

272 As the recoil correction (F_t) factor may over or under-correct AHe or ZHe depending
273 on rates of diffusion and parent element distribution (e.g., [Wildman et al., 2016](#)),
274 uncorrected AHe and ZHe ages are reported below, and the effect of alpha ejection
275 incorporated during the thermal history inversion modeling. Although some samples
276 yield dispersed single crystal ages, there is no analytical reason to discard them. For a
277 more comprehensive report of all (U-Th)/He ages including the ‘outliers’, Table 2 and
278 3 presents all the single-grain (U-Th)/He ages; the mean is calculated using only those
279 ages that are within 20% of the standard deviation ([Flowers and Kelley, 2011](#)).

280 Single-grain apatite (U-Th)/He analyses were performed on fourteen samples (Table
281 2). Five apatite grains from RS19-05B (Xiabie granite) yield AHe ages ranging from
282 30 to 50 Ma. Sample RS17-N5 from the Xiabie granite has a mean AHe age of $41 \pm$
283 6 (1σ) Ma. Apatite grains from two granitoid clasts (RS17-N4-C1, RS19-10-C1) from
284 the Upper Muggar unit yield mean AHe ages of 36 ± 4 Ma and 28 ± 5 Ma,
285 respectively. Sample RS19-10-C2 from the Upper Muggar unit yields individual AHe
286 ages ranging from 13 to 43 Ma. Four samples RS17-N1-C4, RS17-N1-C5,
287 RS19-06-C3, RS19-06-C6 from the Lower Muggar unit yield mean AHe ages of $38 \pm$
288 3 Ma, 31 ± 6 Ma, 31 ± 6 Ma, 25 ± 3 Ma, respectively. Two samples (RS17-N1-C6
289 and RS19-06-C4) from the Lower Muggar unit yield individual AHe ages ranging
290 from 20 to 47 Ma, and 35 to 62 Ma, respectively. Fifteen apatite grains from three
291 sandstone clasts (RS17-N1-C1, -C2, -C3) from the Lower Muggar unit yield
292 individual AHe ages ranging from 9 to 21 Ma, 12 to 29 Ma and 8 to 19 Ma,

293 respectively.

294 Seven of the fourteen samples yield single-grain AHe ages with > 20% standard
295 deviation of the mean. Four of fourteen samples show a weak positive correlation
296 between uncorrected AHe ages and eU (RS17-N1-C1, RS17-N1-C5, RS19-06-C6,
297 RS17-N4-C1). Two samples (RS17-N1-C2, RS19-10-C2) show a weak negative
298 correlation between age and eU. Similarly, except for RS17-N1-C2, RS17-N1-C3,
299 RS17-N1-C6 and RS17-N4-C1 (Figures S2 and S3), there is no evident correlation
300 between AHe age and grain size. The absence of clear correlations suggests that the
301 age dispersion reflects some combination of eU variation (Gautheron., 2009; Flower
302 et al., 2009; Guenther et al., 2013), differences in grain radii (Reiners et al., 2001),
303 grain fragmentation (Brown et al., 2013), undiscovered inclusions (Vermeesch et al.,
304 2007), and/or implantation of ⁴He (Farely et al., 1996). The (U-Th)/He age of apatites
305 from sandstone clasts of the Lower Muggar unit are generally younger than those of
306 granitoid conglomerate clasts. This may be caused by the fact that most of the detrital
307 apatites in the sandstone clasts are broken and with a heavily pitted surface that may
308 suggest U precipitation from circulating fluids (Brown et al., 2013).

309 A few Ft-corrected AHe ages from the granitoid clasts from the Lower Muggar unit
310 are slightly older than same-sample AFT ages. This has been previously recorded in
311 samples that experienced a period of burial followed by erosional unroofing (Flowers
312 et al., 2009; Ault et al., 2009). Radiation damage increases the closure temperature of
313 ⁴He in apatite (e.g., Flowers et al., 2009; Gautheron et al., 2009) and its effect on
314 samples characterized by reheating in the PRZ and PAZ are known to be amplified.

315 Whereas the dispersion of AHe ages from rocks that experienced no or very little
316 burial (e.g., the Xiabie granite and the Upper Muggar unit) can be accounted for in the
317 thermal modelling (Figure S1), in the case of the Lower Muggar sedimentary rocks
318 that were buried at temperatures high enough to almost completely reset the AFT ages,
319 some of the AHe ages are not reproduced by the inversion models (Figure S1). These
320 observations suggest that a major driver of the over-dispersion is our inability to
321 capture the diffusion process when radiation damage and partial annealing occur
322 contemporaneously, at different rates, depending on their original track lengths and
323 apatite chemical composition (Flowers et al., 2009; Gautheron et al., 2009).

324 Eleven zircon grains from two granitoid clasts (RS17-N1-C4 and -C6) from the
325 Lower Muggar unit yield individual He ages ranging from 58 to 107 Ma and 74 to
326 105 Ma (Table 3). Seven zircon grains from sample RS17-N1-C5 from the Lower
327 Muggar unit yield a mean ZHe ages of 84 ± 9 Ma. Another seven zircon grains from
328 sample RS17-N4-C1 from the Upper Muggar unit yield a mean age of 67 ± 6 Ma. A
329 negative uncorrected ZHe ages-eU correlation is observed for three samples
330 (RS17-N1-C4, -C5, -C6) from the Lower Muggar unit (Figure S4), suggesting that
331 radiation damage facilitates He diffusion and, at least partially, drives age dispersion
332 (Guenther et al., 2013; 2014). Sample RS17-N1-C4 displays a positive ZHe age-eU
333 correlation for concentrations of U up to 800-1000 ppm; for higher concentrations, the
334 correlation is negative (Figure S4). This suggests that, at low U concentrations,
335 radiation damage is not sufficient to affect He diffusion and the U concentration
336 increases radiation damage to form a network of linked voids that promotes diffusion

337 ([Guenthner et al., 2013](#)). A few Ft corrected ZHe ages are slightly older than the time
338 of intrusion, probably an effect of still poorly understood interaction between alpha
339 ejection ([Farley et al., 1996, 2002](#)), radiation damage effects ([Guenthner et al., 2013](#)),
340 U-Th-zonation ([Dobson et al., 2008](#); [Danisik et al., 2017](#)) and He diffusion ([Reiners et](#)
341 [al., 2006](#); [Guenthner et al., 2013](#)).

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Table 2. Results of Apatite (U-Th)/He Analysis

Grain name	⁴ He (cc)	²³⁸ U (ng)	²³² Th (ng)	eU ppm	²¹ T	^b L (um)	^c W (um)	Ft	Th/U	Raw age (Ma)	error (Ma)	Cor. Age (Ma)	error (Ma)	Ave. Raw age and SD (Ma)	Ave. Cor. Age And SD (Ma)
RS17-N1-C1-02	2.3E-10	0.019	0.308	28.5	0	120	103	0.7	15.8	20.7	0.8	28.6	1.2		
RS17-N1-C1-03	2.1E-10	0.021	0.229	33.6	0	110	90	0.7	10.7	23.4	0.3	34.1	0.5		
RS17-N1-C1-05	2.8E-10	0.051	0.202	48.2	0	100	90	0.7	4.0	23.7	0.3	34.9	0.4		
RS17-N1-C1-06	2.3E-10	0.037	0.293	26.7	0	130	110	0.7	8.0	17.8	0.2	24.0	0.3		
RS17-N1-C1-07	8.0E-11	0.019	0.244	41.8	0	100	85	0.7	13.1	8.7	0.1	13.0	0.2	/	/
RS17-N1-C2-02	6.0E-11	0.008	0.040	14.3	0	90	73	0.6	5.2	29.1	1.4	47.6	2.2		
RS17-N1-C2-03	8.7E-11	0.015	0.050	16.4	0	100	80	0.7	3.4	27.2	1.0	41.9	1.5		
RS17-N1-C2-04	9.2E-11	0.012	0.087	26.0	0	100	70	0.6	7.4	23.4	0.6	38.1	0.9		
RS17-N1-C2-05	5.2E-11	0.003	0.075	25.4	0	90	60	0.6	23.5	20.8	0.6	37.4	1.1		
RS17-N1-C2-06	9.9E-11	0.035	0.140	94.4	0	80	60	0.5	3.9	11.9	0.4	22.0	0.8	/	/
RS17-N1-C3-01	7.4E-11	0.017	0.131	37.7	0	90	75	0.6	7.6	12.6	0.2	20.3	0.4		
RS17-N1-C3-02	2.7E-11	0.012	0.026	19.4	0	90	65	0.6	2.1	12.0	0.7	20.7	1.2		
RS17-N1-C3-03	4.6E-11	0.018	0.126	23.2	0	100	90	0.7	7.1	8.0	0.1	11.8	0.2		
RS17-N1-C3-04	7.1E-11	0.018	0.081	20.3	0	100	85	0.7	4.5	15.7	0.4	23.7	0.6		
RS17-N1-C3-05	1.7E-10	0.028	0.181	28.2	0	110	95	0.7	6.5	19.4	0.3	27.7	0.4	/	/
RS17-N1-C4-1	1.8E-10	0.012	0.121	48.1	1	110	55	0.5	10.1	35.9	0.5	65.8	0.9		
RS17-N1-C4-2	3.1E-10	0.032	0.170	44.2	1	140	68	0.6	5.3	35.4	0.5	55.8	0.8		
RS17-N1-C4-3	6.6E-10	0.021	0.455	101.0	1	140	60	0.6	21.7	42.2	0.6	70.8	1.0	37.8 ± 3.1	64.2 ± 6.2
RS17-N1-C5-1A	3.7E-10	0.042	0.272	34.4	1	160	88	0.7	6.5	28.7	0.3	40.5	0.5		
RS17-N1-C5-2A	1.5E-10	0.024	0.127	44.7	1	160	55	0.6	5.2	23.1	0.4	40.3	0.7		
RS17-NI-C5-1	3.8E-10	0.046	0.236	24.9	2	200	90	0.7	5.1	30.4	0.4	41.7	0.5		

RS17-NI-C5-3	7.5E-10	0.096	0.354	31.1	1	200	107	0.8	3.7	34.2	0.4	45.1	0.5		
RS17-NI-C5-4	8.5E-10	0.068	0.429	76.0	1	180	70	0.7	6.3	41.5	0.5	63.0	0.7		
RS17-NI-C5-5	6.1E-10	0.106	0.313	14.5	1	300	128	0.8	3.0	28.1	0.3	35.1	0.4	31.0 ± 5.7	44.3 ± 8.8
RS17-N1-C6-2	5.6E-10	0.043	0.375	54.8	2	170	75	0.7	8.6	35.1	0.5	52.0	0.7		
RS17-N1-C6-3	6.3E-10	0.043	0.307	60.3	2	155	70	0.6	7.2	45.1	0.6	69.4	0.9		
RS17-N1-C6-4	3.2E-10	0.066	0.279	59.9	2	155	75	0.7	4.3	20.1	0.2	30.0	0.3		
RS17-N1-C6-2a	1.1E-09	0.108	0.488	157.3	1	137	64	0.6	4.5	41.8	1.0	67.9	1.6		
RS17-N1-C6-3a	3.3E-10	0.033	0.237	49.1	1	140	72	0.6	7.1	30.8	1.5	47.5	2.3		
RS17-N1-C6-4a	3.8E-10	0.028	0.163	44.7	2	140	65	0.6	5.8	46.5	0.7	74.9	1.1		
RS17-N1-C6-6a	5.2E-10	0.098	0.423	72.6	1	150	85	0.7	4.3	21.5	0.2	30.8	0.3	/	/
RS19-06-C3-3	2.8E-10	0.011	0.264	39.0	2	160	68	0.60	23.64	31.62	0.5	52.8	0.8		
RS19-06-C3-5	2.9E-10	0.019	0.242	32.5	2	157	77	0.63	12.78	30.96	0.5	48.8	0.7		
RS19-06-C3-1c	9.6E-11	0.011	0.121	18.3	2	130	81	0.63	10.90	19.94	0.5	31.7	0.8		
RS19-06-C3-2c	1.1E-10	0.014	0.051	11.3	2	157	77	0.65	3.58	35.98	1.6	55.6	2.5		
RS19-06-C3-3c	1.7E-10	0.022	0.075	10.3	2	190	89	0.70	3.46	36.10	1.1	51.7	1.6	30.9 ± 5.9	48.1 ± 8.5
RS19-06-C4-1	6.1E-10	0.058	0.359	21.2	2	210	113	0.74	6.14	34.89	0.3	46.4	0.4		
RS19-06-C4-2	4.8E-10	0.024	0.196	22.4	2	180	83	0.67	8.11	55.50	0.6	82.2	0.8		
RS19-06-C4-3	7.9E-10	0.054	0.209	11.7	2	236	122	0.77	3.87	62.40	0.6	80.5	0.8		
RS19-06-C4-4	9.2E-10	0.099	0.454	42.4	1	232	91	0.73	4.58	36.60	0.3	49.6	0.4	/	/
RS19-06-C6-1	8.4E-11	0.013	0.122	20.2	2	130	80	0.63	9.18	16.10	0.2	26.0	0.3		
RS19-06-C6-2	1.1E-10	0.014	0.076	12.8	2	152	80	0.65	5.54	28.44	0.3	44.5	0.5		
RS19-06-C6-3	2.5E-10	0.026	0.207	29.3	0	163	79	0.65	7.82	27.33	0.3	41.8	0.4		
RS19-06-C6-4	1.3E-10	0.015	0.147	24.4	1	159	72	0.65	9.61	21.28	0.2	32.8	0.4		
RS19-06-C6-5	2.0E-10	0.014	0.166	26.6	1	130	78	0.67	11.49	29.93	0.3	44.3	0.5		
RS19-06-C6-1C	4.4E-10	0.050	0.406	26.7	2	183	109	0.72	8.17	24.78	0.2	34.2	0.3	24.6 ± 4.7	37.3 ± 6.8
RS19-10-C1-1	3.3E-10	0.021	0.341	60.3	2	130	72	0.60	15.96	26.69	0.3	44.9	0.6		

RS19-10-C1-2	2.8E-10	0.011	0.296	56.8	2	123	68	0.57	27.77	28.13	0.4	49.3	0.7		
RS19-10-C1-3	1.8E-10	0.020	0.101	22.8	2	160	69	0.62	4.98	32.71	0.8	52.8	1.3		
RS19-10-C1-4	3.5E-10	0.029	0.258	40.2	2	150	77	0.63	9.03	32.45	0.4	51.4	0.7		
RS19-10-C1-5	1.1E-10	0.016	0.122	26.3	1	118	76	0.66	7.49	19.53	0.4	29.6	0.7	27.9 ± 4.8	45.6 ± 8.4
RS19-10-C2-1	2.0E-10	0.037	0.355	71.1	2	147	68	0.60	9.62	13.38	0.2	22.4	0.3		
RS19-10-C2-2	4.5E-10	0.029	0.389	37.2	2	184	84	0.67	13.34	30.37	0.4	45.7	0.5		
RS19-10-C2-3	4.4E-10	0.038	0.222	29.7	1	183	81	0.69	5.87	40.36	0.6	58.2	0.8		
RS19-10-C2-4	2.6E-10	0.021	0.127	18.9	2	162	81	0.66	6.12	42.93	0.9	65.3	1.4		
RS19-10-C2-5	2.4E-10	0.020	0.175	26.6	1	141	80	0.68	8.96	32.22	0.6	47.4	0.8	/	/
RS17-N4-C1-1A	3.0E-10	0.031	0.142	26.8	1	170	75	0.7	4.6	38.3	0.6	56.8	0.9		
RS17-N4-C1-1	2.8E-10	0.033	0.154	16.4	2	200	92	0.7	4.6	33.2	0.4	45.3	0.5		
RS17-N4-C1-2	8.3E-10	0.074	0.380	65.4	1	250	63	0.6	5.2	41.9	0.4	74.8	0.8		
RS17-N4-C1-3	8.4E-10	0.070	0.613	55.1	1	200	88	0.7	8.7	32.1	0.4	44.1	0.5		
RS17-N4-C1-4	9.7E-10	0.128	0.510	42.9	1	180	113	0.8	4.0	31.9	0.4	41.4	0.5	35.7 ± 7.4	51.6 ± 3.5
RS19-05B-1	4.2E-10	0.054	0.219	45.3	2	146	80	0.65	4.06	32.97	0.5	50.9	0.7		
RS19-05B-2A	1.7E-10	0.021	0.075	18.6	2	170	70	0.63	3.53	36.20	1.1	57.5	1.8		
RS19-05B-3	5.0E-10	0.073	0.265	73.4	1	149	70	0.62	3.64	30.10	0.4	48.5	0.6		
RS19-05B-4	3.1E-10	0.046	0.165	58.3	2	110	72	0.60	3.60	29.58	0.5	49.6	0.8		
RS19-05B-5	3.6E-10	0.036	0.099	26.3	1	180	71	0.64	2.75	49.80	1.2	77.8	1.8	/	/
RS17-N5-1	1.8E-10	0.021	0.087	22.4	2	175	65	0.6	4.1	35.0	0.5	55.1	0.7		
RS17-N5-2	3.0E-10	0.028	0.098	40.7	1	140	60	0.6	3.5	47.1	0.7	79.2	1.1	41.1 ± 6.1	67.1 ± 12.0

Note. a. T—termination; b. L—length; c. W—width; For samples with standard deviation greater than 20%, the mean age is not calculated (Flowers and Kelley, 2011).

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Table 3. Results of Zircon (U-Th)/He Analysis

Grain name	⁴ He (cc)	²³⁸ U (ppm)	²³² Th (ppm)	eU (ppm)	[°] T	^a L (um)	^b W (um)	F(T)	Th/U	Raw age (Ma)	Error (Ma)	Corrected age (Ma)	Error (Ma)	Ave. Raw age and SD (Ma)	Ave. Cor. Age and SD (Ma)
RS17-N1-C4-Z1	2.49E-08	588.88	361.23	673.77	2	200	75	0.74	0.61	106.70	5.95	113.11	8.01		
RS17-N1-C4-Z2	6.92E-09	373.14	424.29	472.85	2	180	65	0.71	1.14	62.66	4.45	88.73	6.31		
RS17-N1-C4-Z3	1.64E-08	1107.38	516.58	1228.78	2	190	53	0.67	0.47	82.73	4.61	97.41	6.90		
RS17-N1-C4-Z4	2.31E-09	217.60	318.86	292.54	2	150	55	0.65	1.47	56.55	4.06	86.41	6.22		
RS17-N1-C4-Z5	1.04E-08	925.97	533.22	1051.27	2	170	50	0.64	0.58	75.55	5.35	117.72	8.35		
RS17-N1-C4-Z6	1.05E-08	980.32	461.13	1088.68	2	130	53	0.64	0.47	87.39	4.88	107.25	7.61		
RS17-N1-C4-Z7	9.86E-09	1249.02	651.77	1402.18	2	130	55	0.65	0.52	58.21	4.12	89.13	6.31	/	/
RS17-N1-C5-Z1	8.15E-09	870.23	767.37	1050.57	2	140	50	0.63	0.88	72.04	5.11	114.90	8.15		
RS17-N1-C5-Z2	2.68E-08	1866.43	1223.56	2153.97	2	150	60	0.68	0.66	74.85	5.29	109.98	7.77		
RS17-N1-C5-Z3	2.38E-08	1596.10	723.78	1766.19	2	190	55	0.67	0.45	76.21	5.38	113.89	8.04		
RS17-N1-C5-Z4	1.04E-08	560.04	398.59	653.71	2	180	55	0.67	0.71	94.72	6.71	142.33	10.08		
RS17-N1-C5-Z5	1.61E-08	847.21	542.20	974.63	2	180	65	0.71	0.64	70.79	5.01	100.01	7.07		
RS17-N1-C5-Z6	2.17E-08	1346.13	755.57	1523.68	2	160	60	0.68	0.56	80.52	5.69	118.61	8.38		
RS17-N1-C5-Z7	8.08E-09	418.03	516.81	539.48	2	150	53	0.64	1.24	117.41	8.36	182.36	12.98	78 ± 9	117 ± 14
RS17-N1-C6-Z1	1.57E-08	1168.61	414.97	1266.13	2	170	53	0.66	0.36	86.17	6.09	131.54	9.29		
RS17-N1-C6-Z2	1.24E-08	1105.77	732.77	1277.97	2	170	50	0.64	0.66	74.17	5.25	115.93	8.20		
RS17-N1-C6-Z3	1.60E-08	1532.07	830.99	1727.35	2	150	50	0.64	0.54	80.08	5.66	125.99	8.90		
RS17-N1-C6-Z4	5.24E-09	440.67	409.93	537.00	2	120	50	0.62	0.93	105.48	7.54	169.78	12.13	86 ± 14	136 ± 24
RS17-N4-C1-Z1	3.83E-09	198.38	185.40	241.95	2	210	59	0.69	0.93	70.34	5.02	102.02	7.28		
RS17-N4-C1-Z2	6.94E-09	198.17	215.01	248.70	2	220	78	0.75	1.08	68.65	4.88	92.01	6.54		
RS17-N4-C1-Z3	1.30E-08	354.01	331.52	431.91	2	280	70	0.73	0.94	71.16	5.04	97.09	6.88		
RS17-N4-C1-Z4	1.16E-08	228.66	211.63	278.39	2	250	95	0.79	0.93	59.94	4.25	75.62	5.36		

RS17-N4-C1-Z5	7.10E-09	293.59	245.53	351.29	2	200	75	0.74	0.84	58.47	4.15	78.81	5.59		
RS17-N4-C1-Z6	1.99E-08	474.62	469.85	585.04	2	260	75	0.75	0.99	75.77	5.37	101.11	7.16		
RS17-N4-C1-Z7	4.35E-09	152.32	162.25	190.45	2	240	70	0.73	1.07	63.24	4.50	86.26	6.14	67 ± 6	90 ± 11

348 Note. a. L—length; b. W—width; c. T—termination. For samples with standard deviation greater than 20%, the mean age is not calculated (Flowers and Kelley, 2011).

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351 4.3 Apatite fission track data

352 Sixteen samples were dated using single-grain apatite fission track (AFT) analysis
353 (Table 4). Five Xiabie granite samples (RS19-05A, -05B, -07, -08, and RS17-N5)
354 were analysed; four yielded AFT ages between 43 ± 5 Ma and 57 ± 4 Ma, with one
355 significantly older (89 ± 10 Ma; Table 4). Two granitoid clasts from the Upper
356 Muggar Unit yielded AFT ages of 51 ± 5 Ma and 67 ± 5 Ma. Four of the six granitoid
357 clasts (RS17-N1-C6, RS19-06-C3, -C4, -C6) from the conglomerate beds (Lower
358 Muggar unit) yielded AFT ages ranging from 31 ± 3 Ma to 34 ± 3 Ma, younger than
359 their stratigraphic age. Two other granite clasts (RS17-N1-C4, -N1-C5) from the
360 Lower Muggar unit yielded older apatite fission track ages, 62 ± 10 Ma to 77 ± 15 Ma,
361 which overlap the conglomerate depositional age. The AFT ages of 3 sandstone clasts
362 (RS17-N1-C1, -C2, -C3) from the Lower Muggar unit range from 50 ± 7 Ma to $83 \pm$
363 10 Ma. In summary, all the AFT central ages are either younger than or overlap the
364 depositional age of the Lower Muggar unit.

365 The low U concentration of many apatite crystals and their small size hampered
366 identification of horizontal confined tracks ([Donelick et al., 2005](#)). Consequently, few
367 tracks were measured, and no samples had a statistically significant track length
368 distribution; they were nonetheless used in thermal modelling. Four of the granitoid
369 samples (RS19-05B; RS17-N5, RS17-N1-C4, RS17-N1-C5) failed the chi-square test
370 ($P(\chi^2) < 0.05$), and three of them (RS17-N1-C5, RS17-N1-C4, RS17-N5) have high
371 dispersion ($>37\%$). These three samples were measured by the LA-ICP-MS method.
372 [Ketcham et al. \(2018\)](#) noted that LA-ICP-MS derived AFT ages are commonly

373 characterized by a dispersion of the single grain ages that is larger than those obtained
374 by the EDM method. We suggest that this is due to the local variability of U
375 concentration sampled by LA-ICP-MS spot analysis that may not be representative of
376 the bulk apatite value ([Cogne et al., 2021](#)). This is supported by the observation that
377 the three samples (RS17-N1-C5, RS17-N1-C4, RS17-N5) with the lowest U
378 concentration (2.1-4.8 ppm; Table 4) have the oldest AFT ages (cf. [McDannell et al.,
379 2019](#); [Femie et al., 2018](#)). Irrespective of the method used to calculate the AFT ages,
380 high dispersion is likely related to prolonged residence within the PAZ, rather than to
381 the preservation of the original detrital age population, as the AFT ages are similar or
382 younger than the stratigraphic age.

Table 4. Results of Apatite Fission Track Analysis

Sample	No. of Grains	^a ρ _s (10 ⁵ /cm ²)	Total ^b N _s	^a ρ _i (10 ⁵ /cm ²)	Total ^b N _i	^a ρ _d (10 ⁵ /cm ²)	^b N _d	^c P(χ) ²	^d D _{par} (μm)	²³⁸ U ppm	Central AFT Age (Ma)	±1σ	Pooled AFT Age (Ma)	±1σ	Dis. (%)	MTL c Axis Correction (μm)	SD (μm)	N ^e
RS19-05-A	16	5.1	250	32.0	1562	17.7	14503	0.50	1.5	25.3	42.5	5.3	42.5	2.9	0	13.5	1.4	8
RS19-05-B	20	7.4	368	31.1	1556	16.1	14503	0.02	1.5	28.4	57.1	4.4	57.1	3.3	20	13.96	1.2	27
RS19-08	17	3.1	263	17.8	1488	18.8	14503	0.60	1.6	14.9	50.0	3.3	49.0	3.3	0	14.33	1.1	9
RS19-07	12	5.4	209	25.5	990	18.1	14503	0.17	1.6	23.6	56.9	7.7	57.2	4.4	11	13.41	1.3	15
RS17-N5*	22	5.2	179	-	-	-	-	0.00	1.1	2.1	89.0	10	96.8	7.3	37	12.48	1.3	5
RS19-10-C1	20	5.7	196	23.5	810	18.4	14503	0.43	1.5	17.4	67.0	5.3	67.0	5.3	0	13.44	1.2	4
RS17-N4-C1*	14	4.8	174	-	-	-	-	0.09	1.4	3.8	51.3	4.8	51.3	3.9	19	12.29	0.5	2
RS19-06-C3	12	2.1	93	15.8	697	16.5	14503	0.23	1.6	15.8	33.0	4.1	33.1	3.7	18	-	-	0
RS19-06-C4	20	2.0	189	14.2	1317	14.5	14503	0.10	1.5	13.5	31.2	2.9	31.3	2.4	19	13.24	0.6	3
RS19-06-C6	20	2.0	170	15.6	1313	17.3	14503	0.10	1.6	12.9	33.4	3.4	33.6	2.7	24	14.45	1.0	3
RS17-N1-C6*	20	2.0	124	-	-	-	-	0.69	1.3	2.2	33.6	3.0	33.6	3.0	0	-	-	0
RS17-N1-C4*	18	3.4	129	-	-	-	-	0.00	1.2	2.3	62.0	10.0	71.0	6.3	56	12.70	0.2	3
RS17-N1-C5*	9	4.5	57	-	-	-	-	0.03	1.2	4.8	77.0	15.0	77.0	10	39	-	-	0
RS17-N1-C1	26	5.4	313	19.8	1143	12.5	8790	0.04	1.7	23.5	53.1	4.3	51.7	10.8	22	13.30	0.6	6
RS17-N1-C2	11	4.6	69	17.5	262	12.4	8790	0.89	1.8	20.6	49.7	6.7	49.3	10.3	0	12.26	1.0	5
RS17-N1-C3	20	13.6	481	30.8	1091	12.2	8790	0.00	2.0	36.1	83.0	10.0	81.0	16.9	47	12.78	1.5	15

*Concentration of U measured by LA-ICP-MS

- 385 a. $\rho_{(i,s,d)}$ are track density of induced, spontaneous, dosimeter tracks.
- 386 b. N_i , N_s , and N_d are the number of induced, spontaneous, and dosimeter tracks counted.
- 387 c. $P(\chi^2)$ value of the chi-square age homogeneity test.
- 388 d. D_{par} measurements are etch pit diameters used as a proxy for the influence of chemical composition on track annealing [[Donelick et al., 2005](#)]. Three to five D_{par} measurements were used for each grain.
- 389 e. The number of confined tracks
- 390

391 5 Thermal histories

392 5.1 Thermal Modelling method

393 Although the number of measured track lengths was often low, the combination of the
394 apatite fission ages and single-grain ZHe and AHe ages made it possible to invert
395 thermal histories. Single-sample AFT and AHe/ZHe ages were combined in the
396 software QTQt (5.4.2) and used to derive thermal histories that best fit the
397 thermochronological data (Gallagher, 2012). We used the radiation damage models of
398 Flowers et al. (2009), Gautheron et al. (2009) and Guenther et al. (2013) to describe
399 ⁴He diffusion of apatite and zircon. For the apatite fission track data, the multi-kinetic
400 annealing model of Ketcham et al. (2007) was applied, using Dpar values as the
401 kinetic parameter. Few tracks were measured, thus the models were run without track
402 length data. Uncorrected ZHe and AHe ages were used. Constraint boxes of $10 \pm 10^\circ$
403 C (assumed surface temperature) at 70 ± 10 Ma, and $10 \pm 10^\circ\text{C}$ at 30 ± 10 Ma were
404 used for modelling of the Lower Muggar unit and Upper Muggar unit respectively,
405 based on their stratigraphic age. Using the analytical uncertainty on individual ages
406 likely underrepresents the true uncertainty (Flowers et al., 2015), so a 10%
407 uncertainty was given by QTQt software for each single grain (AHe and ZHe) age
408 when running the thermal model (Gallagher, 2012).

409 The thermal modelling strategy and model parameters are reported in the supporting
410 information (Text S1, Table S4). The data were scrutinized following the
411 recommendations of Flowers and Kelley (2011); of the thirteen samples that yielded

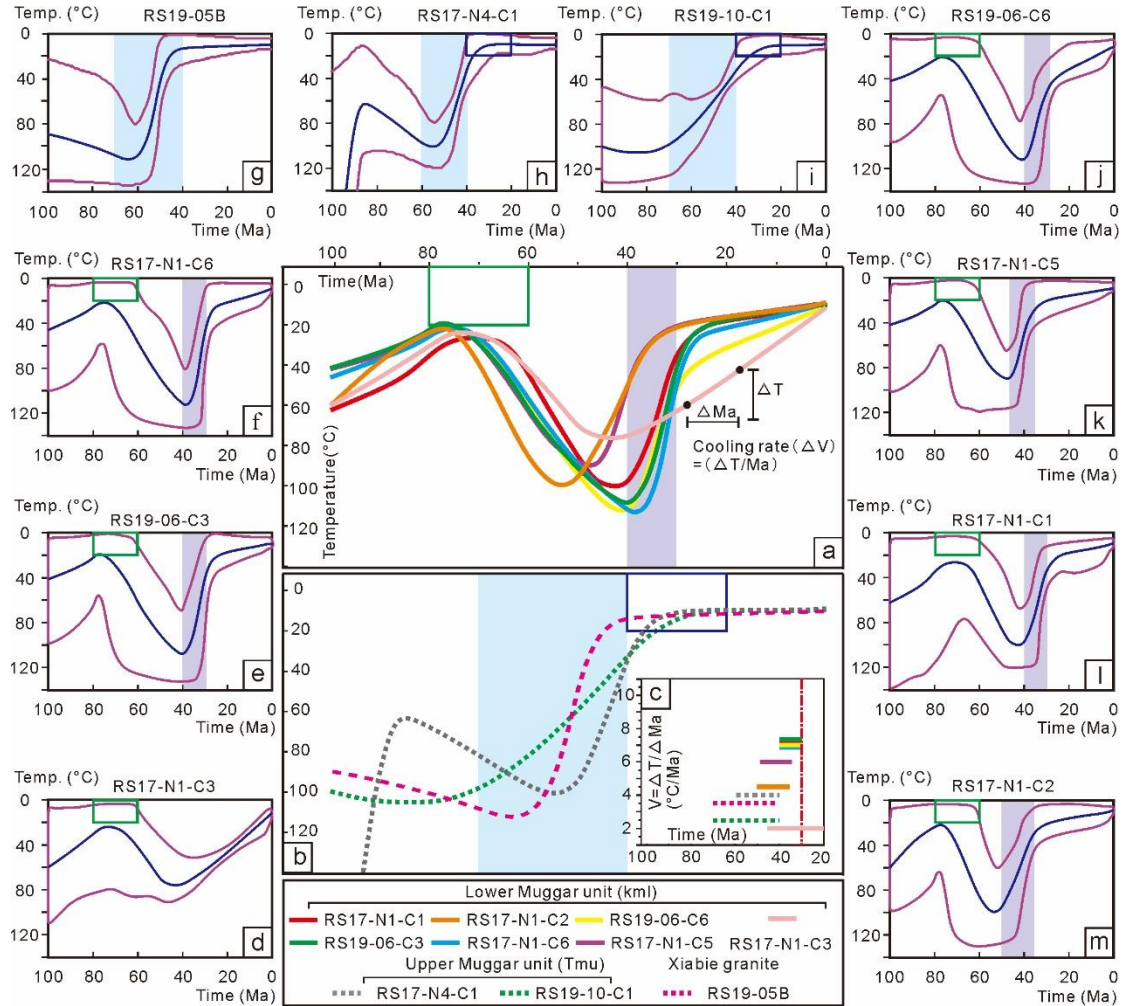
412 both AFT and AHe ages, only 10 were modelled. One of these (RS17-N4-C1)
413 includes ZHe data. Three samples were discarded due to the overdispersion of either
414 the single grain AFT ages (granitic clast sample RS17-N1-C4 and granite RS17-N5),
415 or the AHe determinations (granitoid clast RS19-06-C4).

416 **5.2 Thermal modelling results**

417 Sample RS19-05B from the Xiabie granite produced a thermal history characterized
418 by cooling between 70-40 Ma, from 100 to 20°C at a rate of ~3°C/Ma, followed by a
419 transition to slow cooling (0.25°C/Myr), with one order of magnitude difference
420 between the two values. The two granite clasts (RS17-N4-C1 and RS19-10-C1) from
421 the Oligocene Upper Muggar unit show a similar history, recording cooling at 70-40
422 Ma, with sample RS17-N4-C1 residing for a long time in the apatite partial annealing
423 zone prior to cooling (Figure 4).

424 The seven conglomerate clasts from the Late Cretaceous Lower Muggar unit show
425 similar thermal histories (Figure 4). These seven clasts are characterized by moderate
426 heating at ~70-40 Ma, but exhumation occurs at different rates. Four granite clasts
427 (RS17-N1-C5, -C6, and RS19-06-C3, -C6) show rapid cooling through the apatite
428 partial retention zone (80-40°C), beginning at 40 Ma and reaching near-surface
429 temperatures by 30 Ma. The two sandstone clasts RS17-N1-C1 and RS17-N1-C2
430 show pulses of rapid cooling at 40-30 Ma and 50-30 Ma respectively, from 100 to
431 30°C (Figure 4). RS17-N1-C3 records monotonic cooling starting at ~40 Ma. Except
432 for sample RS17-N1-C3, the rapid cooling rate of the other six samples from the

433 Lower Muggar unit ranges from $\sim 6^\circ\text{C}$ to $\sim 8^\circ\text{C}/\text{Myr}$, decreasing to approximately
 434 $0.25^\circ\text{C} \sim 1^\circ\text{C}/\text{Myr}$ after ~ 30 Ma.



435

436 Figure 4. Thermal history models for Late Cretaceous to Oligocene conglomerate
 437 clasts of the North Nima Basin and for the Xiabie granite. (a) The expected model for
 438 4 granitoid and 3 sandstone clasts from the Lower Muggar unit, highlighting 40-30
 439 Ma cooling episode. The green box represents the sample at surface temperature from
 440 80-60 Ma. (b) Expected model for one granite sample from the Xiabie granite and 2
 441 granitoid clasts from the Upper Muggar unit, highlighting 70-40 Ma cooling. The blue
 442 box represents sample at surface temperature from 40-20 Ma. (c) Cooling rate of each

443 sample over specified time period. (d-m) Expected thermal history model with 95%
444 confidence interval for each sample.

445 **6 Discussion**

446 **6.1 Age and provenance of the Late Cretaceous-Cenozoic sediments in the** 447 **Nima-Lunpola Basin**

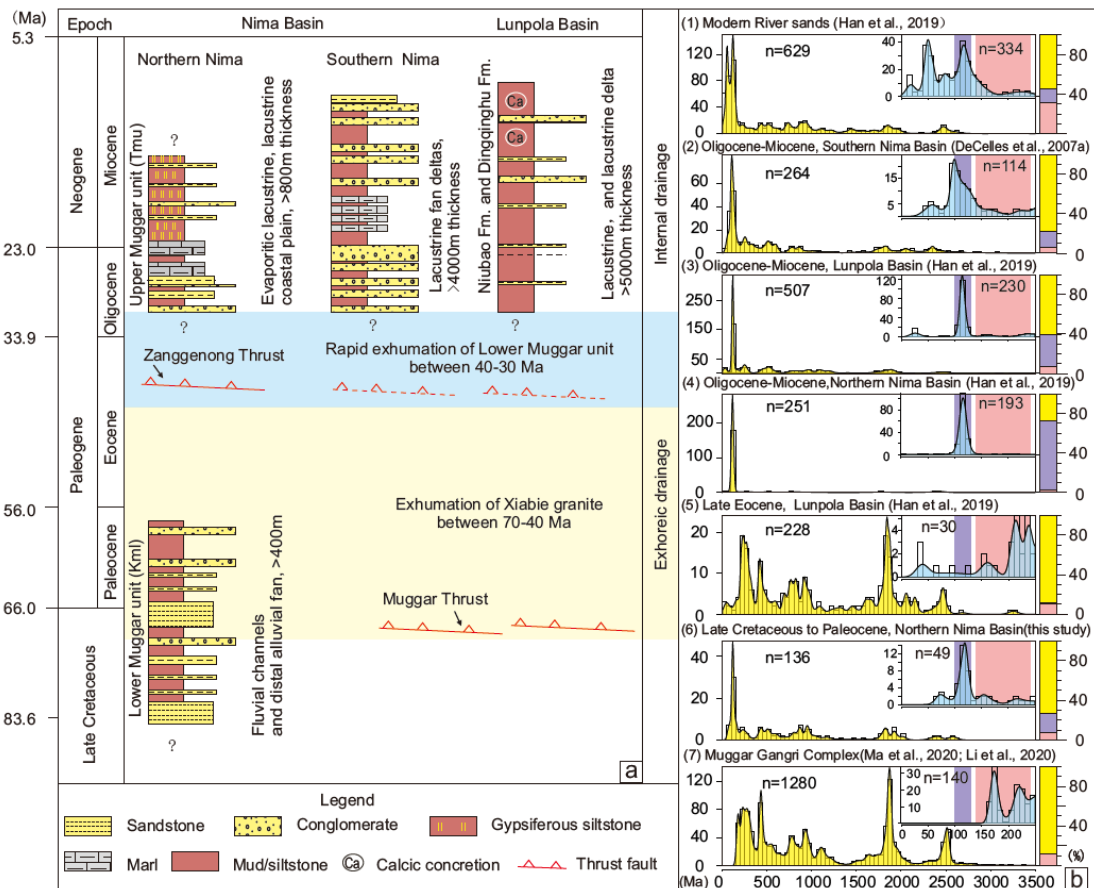
448 The weighted mean age of the three youngest detrital zircon grains from the Lower
449 Muggar unit (67 ± 2 Ma; section 4.1) is consistent with palynomorph results
450 (DeCelles et al., 2007a), indicating it was deposited during the Late Cretaceous
451 (Senonian) to Early Paleocene. Likewise, the weighted mean of the two youngest
452 U-Pb zircon ages of the Upper Muggar unit (35 ± 2 Ma; Han et al., 2019) is
453 consistent with palynomorph results indicating it was deposited in the late
454 Eocene-Oligocene (DeCelles et al., 2007a).

455 In the Nima-Lunpola basin, the provenance of zircons from all sandstone samples (i.e.
456 the Late Cretaceous-Paleocene Lower Muggar sandstone (RS19-09); and published
457 Eocene and Oligo-Miocene sandstones (DeCelles et al., 2007a; Han et al., 2019),
458 Figure 5b (2-6)) can be explained by a mixture of input from the Muggar Gangri
459 Complex and the Xiabie granite, both in the hanging wall of the Muggar Thrust; the
460 former has wide age distribution from 150-2,500 Ma (Figure 5b (7)), the latter has a
461 100-130 Ma peak (Figure 3). The zircon U-Pb ages from the granitic clasts of the
462 Lower Muggar (RS17-N1-C4, -C5, -C6) and Upper Muggar unit (RS17-N4-C1),
463 overlap the emplacement age of the Xiabie granite (Figure 3; 120 Ma); southward

464 paleocurrents and gravel facies suggest that the latter was the proximal source for
465 these clasts (DeCelles et al., 2007a).

466 Han et al. (2019) noted that detrital zircon U-Pb age spectra from Eocene to Miocene
467 sedimentary rocks of the Lunpola and Nima basins and from the two largest internally
468 draining modern rivers of Tibet, the Zhajia Zangbo and Bocang Zangbo (Figure 1b)
469 lack zircon populations (140-240 Ma) typical of the more distally-located Qiangtang
470 Central Uplift (240-200 Ma), Amdo microcontinent (190-160 Ma) and SW Qiangtang
471 terrain (160-140 Ma) (Figure 5b). They therefore concluded that a short transport
472 internal drainage, similar to modern day was established by late Eocene times (~35
473 Ma). However, a close examination of the zircon U-Pb age spectra show that a
474 component with ages of 140-240 Ma accounts for ~10% of the total zircon population
475 in their Cretaceous to Oligo-Miocene samples (Figure 5b). Whether such a percentage
476 might be a reasonable proportion delivered by a major long distance-transport river is
477 difficult to assess without knowing the geology of the hypothetical catchment area,
478 and such an age population (140-240 Ma) also has been found in the proximal
479 Muggar Gangri complex in the hanging wall of the Muggar Thrust (Figure 5b).
480 Regardless, the relatively strong peak at 100-130 Ma, most-likely derived from the
481 proximal Xiabie granite, suggests dominant local derivation.

482



483

484 Figure 5. (a) Stratigraphic logs and facies summary from DeCelles et al. (2007a), Han
 485 et al. (2019), Fang et al. (2020), showing sampling locations. (b) Kernel density
 486 estimation plots with histograms for detrital zircons U-Pb data, including published
 487 data where referenced, from modern river sands (plot 1), Cretaceous-Cenozoic
 488 sedimentary rocks from the Nima and Lunpola basins (plots 2-6) and the Mugar
 489 Thrust hanging wall (plot 7). Purple and pink colors highlight the 100-130 Ma and
 490 140-240 Ma populations, characteristic of the Xiabie granite and more distal regions
 491 (see text) respectively. The rectangular boxes to the right show the percentages of
 492 populations 100-130 Ma (purple), 140-240 Ma (pink), 0-3500 Ma (yellow).

493 **6.2 The timing and mechanism of exhumation of the Xiabie granite and the**
 494 **Northern Nima Basin**

495 The thermal history of the Early Cretaceous Xiabie granite records a cooling event at
496 70-40 Ma, with a cooling rate of 3-4°C/Ma (Figure 4c). For an assumed geothermal
497 gradient of 25-30°C/km, this corresponds to a denudation rate of 0.10-0.16 mm/yr.
498 The thermal histories of most of the granitoid and sandstone clasts from the Late
499 Cretaceous Lower Muggar unit reveal a broadly similar pattern: heating to $100 \pm$
500 20°C during 70-40 Ma, at a rate of $\sim 3^\circ\text{C}/\text{Myr}$, followed by cooling to $20 \pm 10^\circ\text{C}$ at
501 40-30 Ma, at a cooling rate between 6-8°C/Myr, then a transition to at least one order
502 of magnitude slower cooling ($0.25^\circ\text{C}-1^\circ\text{C}/\text{Myr}$) at ~ 30 Ma.

503 Cooling of a pluton can result from denudation of overlying rocks or heat released
504 during equilibration with the country rock (McInnes and Evans, 2005). Cooling of the
505 Xiabie granite at 70-40 Ma has not been previously reported, although Zhao et al.
506 (2020) assumed the existence of this event and used it as a constraint in their thermal
507 modelling based on a regional thermal history (Rohrmann et al., 2012). Provenance
508 studies indicate that the Xiabie granite provided detritus to the Late Cretaceous to
509 Oligocene Lower and Upper Muggar units (DeCelles et al., 2007a), corroborating our
510 hypothesis that the Xiabie granite was exhumed at 70-40 Ma. The 89 Ma AFT age
511 from the Xiabie granite sample RS17-N5 may record early exhumation of the pluton
512 or represent its cooling to the temperature of the country rock (Kapp et al., 2007;
513 Zhao et al., 2020).

514 The old AFT ages of granite clasts (RS19-10-C1, RS17-N4-C1) from the Upper
515 Muggar unit (67 ± 5 Ma and 51 ± 5 Ma) pre-date the Oligocene stratigraphic age
516 (DeCelles et al., 2007a), suggesting only partial resetting of the fission tracks

517 occurred during burial. The clasts likely retain information on the 70-40 Ma cooling
518 event recorded by the Xiabie granite.

519 The average AFT age of all granitoid clasts (RS19-06-C3, -C4, -C6, RS17-N1-C4,
520 -C5, -C6) from the Lower Muggar unit are younger than, or overlap, the depositional
521 age, indicating that apatites from these clasts are near-completely annealed. The
522 sandstone clasts (RS17-N1-C1, -N1-C3) failed the chi-square test; RS17-N1-C3
523 yielded two populations (160 ± 37 Ma, $n=6$; 58 ± 13 Ma, $n=14$, respectively),
524 indicating that at least some grains were not completely annealed, whereas the
525 youngest population (58 ± 13 Ma) is, within uncertainty, indistinguishable from the
526 central age of other samples (Figure S5; 53 Ma (RS17-N1-C1) and 50 Ma
527 (RS17-N1-C2)). That the sandstone clasts retained pre-burial signals, whereas the
528 granitoid samples did not, is consistent with a higher average D-par value of the
529 former (Donelick et al., 2005). Both the granite and sandstone clasts show a similar
530 post-depositional thermal history, characterized by moderate heating at ~70-40 Ma,
531 followed by exhumation at 40-30 Ma (Figure 4). This heating event could be due to
532 burial and/or to a magmatic event that drastically increased the geothermal gradient of
533 the basin. However, there are no 70-40 Ma magmatic events reported in the Nima
534 Basin and surrounding areas, and the population of 62~75 Ma zircon U-Pb ages in the
535 sandstone sample (RS19-09) of the Lower Muggar unit is limited (4%; Figure 3). For
536 these reasons we interpret the 70-40 Ma heating of the Lower Muggar unit as
537 predominantly the result of burial, even if a slight increase in the geothermal gradient,
538 due to a blanketing effect of the cover, cannot be discounted (e.g. Luszczak et al.,

539 [2017](#)). Cooling can be translated into denudation if the geothermal gradient is known,
540 a value of 30°C/km can be assumed for sedimentary basins. The present geothermal
541 gradient in the nearby Lunpola Basin is around 50°C/km ([Liu et al., 2020](#)). Using
542 these two limiting values, Eocene exhumation rates would be 0.27 mm/yr to 0.12
543 mm/yr.

544 Given that the climate influence is weak during these times (see section 6.3), we
545 interpret the initiation of the exhumation of the Xiabie granite (70-40 Ma) and the
546 Lower Muggar unit (40-30 Ma) as related to erosion associated with activity of thrust
547 faults in the region. The Xiabie granite is in the hanging wall of the Muggar Thrust;
548 thus we assign the 70-40 Ma exhumation event to tectonic activity of the Muggar
549 Thrust during the latest Cretaceous to early Cenozoic. This is in agreement with the
550 interpretation of [Kapp et al. \(2007\)](#), [Rohrmann et al. \(2012\)](#) and [Zhao et al. \(2020\)](#) for
551 major tectonic activity in the Nima area between 100-40 Ma. The preceding burial
552 stage of the Lower Muggar unit corresponds to the exhumation stage of the Xiabie
553 granite and activity of the Muggar Thrust; we suggest that the latter drove the Lower
554 Muggar unit to subside at 70-40 Ma, allowing the Xiabie granite to be a detrital
555 source for the Lower Muggar unit, which was experiencing burial to 2-4 km. The
556 rapid exhumation at 40-30 Ma of the Lower Muggar unit may be related to movement
557 on the S-dipping Zanggenong Thrusts (Figure 1c), which [Kapp et al. \(2007\)](#)
558 recognized as having deformed the Lower Muggar unit. This interpretation, together
559 with the early Cenozoic activity along the Muggar Thrust, implies that the Nima area
560 was continuously tectonically deformed from ~70 to ~30 Ma. This is in contrast to the

561 existence of a tectonic quiescence period between 50 and 30 Ma proposed by [Kapp et](#)
562 [al. \(2007\)](#), and explained as due to preexisting thick crust in southern Tibet was
563 sufficient to inhibit upper-crustal shortening in this area over the period of 50-30 Ma.

564 The exhumation rate of the Nima area has been very slow since 40-30 Ma. At ~40 Ma,
565 the exhumation rate of the hanging wall decreased by ten-fold (~0.01 mm/yr; Figure
566 4), while basin inversion in the footwall (0.12-0.27 mm/yr) started, possibly due to
567 movement on the Zanggenong Thrust. No Eocene strata are preserved between the
568 Lower and Upper Muggar units, likely removed by basin exhumation at that time
569 (Figure 5a). By 30 Ma, this exhumation event ceased, as marked by the LTT-derived
570 thermal history and the restart of sedimentation with the deposition of the Upper
571 Muggar unit (Figure 5a). In the last 30 Ma or so, rates of denudation in the Xiabie
572 granite have been low (~0.01 mm/yr) (Figure 4g), and southward-directed
573 palaeocurrents ([DeCelles et al., 2007a](#)) suggest that the pluton may have continued to
574 be a source for sediments of the Upper Muggar unit. The lack of annealing in the
575 apatites from the Upper Muggar unit and the LTT-derived thermal history suggest
576 passive exhumation of the hanging wall of the Muggar Thrust throughout the
577 Oligocene (Figure 4).

578 **6.3 Development of internal drainage in the Nima Basin**

579 Several factors influence erosion rates, such as slope, precipitation, glacial cover and
580 tectonic activity ([Harrison, 2000](#); [Molnar, 2001](#); [Lal et al., 2004](#)). We interpret that the
581 relatively high rates of denudation of the Xiabie granite at 70-40 Ma and the Northern

582 Nima Basin at 40-30 Ma was predominantly driven by tectonic activity possibly along
583 the Muggar and Zanggenong thrusts, which led to an increase in relief and river
584 incision. We discount any effect of glacial cover since such an environment has not
585 been recorded in the early Cenozoic in central Tibet. We also consider that any
586 strengthening of precipitation is not the most likely cause for the increase in
587 denudation rates; whilst we acknowledge that there is considerable debate
588 surrounding the climate of the region in the Paleogene (e.g. [DeCelles et al., 2007a](#),
589 [2007b](#); [Su et al., 2019](#); [Fang et al., 2020](#)), we adopt the sedimentological arguments of
590 [DeCelles et al. \(2007a\)](#) since they apply directly to Nima Basin, rather than Lunpola
591 Basin more than 200 km east. [DeCelles et al. \(2007a\)](#) describe sedimentology
592 indicative of an arid climate from Late Cretaceous onwards in the Nima Basin as
593 evidenced by the presence of paleosols, calcareous nodules and large eolian dune
594 fields in the Late Cretaceous Lower Muggar unit; restricted evaporitic lakes in the
595 Late Eocene to Oligocene Upper Muggar unit; and the reconstructed paleogeography
596 and deposystems since Late Cretaceous which are similar to those of the modern arid
597 plateau interior ([DeCelles et al., 2007a](#)).

598 We propose that the evidence for tectonic uplift and consequent erosion in the Muggar
599 thrust hanging wall (Xiabie granite), and subsidence and inversion of the Northern
600 Nima Basin in the footwall, suggest that over the interval of 70-30 Ma, landscape
601 evolution was driven by tectonic events and, in particular, by compressive
602 deformation along the Muggar and Zanggenong thrusts. These deformation events
603 may have led to the establishment of internal drainage in the Nima area. Several lines

604 of stratigraphic and sedimentological evidence support the contention that the Nima
605 area developed internal drainage by 40-30 Ma:

606 The early Eocene sedimentary record of the Nima area is discontinuous, a
607 consequence of sediment removal during uplift. Section 3MK (Figure 2) records the
608 loss of at least 2 km of section (required to have reset the AHe system and almost
609 completely anneal the tracks in the apatites of the Lower Muggar Unit) suggesting the
610 persistence of an exoreic system in the early Eocene, capable of transporting away the
611 material previously deposited in the Nima Basin. A Late Cretaceous relatively
612 energetic transport environment in the area of the Nima Basin is also evidenced by the
613 sedimentology of the Lower Muggar unit, which is characterized by gravel-dominated
614 fluvial channel and distal alluvial fan deposits (DeCelles et al., 2007a). Following the
615 exhumation of the sedimentary basin, the environmental conditions drastically
616 changed (Figure 5a). The Oligocene Upper Muggar unit is characterized by deposition
617 of massive red gypsiferous siltstone, previously interpreted as shallow, restricted
618 evaporitic lakes, and fluvial systems flowing over a lacustrine coastal plain (DeCelles
619 et al., 2007a). The presence of such massive evaporites and gypsiferous siltstone
620 generally indicate the persistence of an internal drainage system and arid conditions
621 (Langbein, 1961; Vandervoort et al., 1995).

622 We suggest that the deformation events we document, drove the transformation of the
623 Nima Basin from an exorheic to an endorheic basin by the late Eocene to early
624 Oligocene.

625 **6.4 Implications for the development of Low Relief Topography**

626 Numerous studies have shown that the formation of internal drainage can play an
627 important role in the development of low relief landforms in plateau interiors, such as
628 in Tibet and in the Andes (e.g. [Metivier et al. 1998](#); [Garcia-Castellanos, 2007](#);
629 [Liu-Zeng et al., 2008](#); [Strecker et al., 2009](#)). The evolution of an intra-mountain
630 internal drainage traps the eroded material from the orogenic belt in a depression,
631 forming a low-lying plateau landform through a mechanism called “bathtub filling”
632 ([Meyer et al. 1998](#)). This process may have occurred in the Nima area, when after ~30
633 Ma, slow exhumation and the development of an internal drainage system could have
634 promoted the formation of a low relief landscape that is similar to the present day ([Lal](#)
635 [et al., 2004](#)).

636 However, we note that other regions have similar deformation and exhumations
637 histories to the Nima Basin and are externally drained. For example, the Fenghuoshan
638 Group of the Hoh Xil Basin was deposited at 75-51 Ma prior to basin inversion at
639 50-27 Ma ([Staisch et al., 2014](#); [2016](#)). This may suggest that instead of internal
640 drainage development being closely linked to low relief topography, it is onset of
641 tectonic quiescence, in concert with the development of an arid climate, that is the
642 more major influence on the development of low relief topography ([Fielding et al.,](#)
643 [1994](#)). In the Nima Basin, the thermal histories clearly indicate that around 40 Ma, the
644 exhumation rate decreased in the hanging wall and by 30 Ma regional erosion rates in
645 the basin were low (0.01-0.04 mm/yr). Such low rates imply a low river erosion
646 capacity and, given that the climate remained dry over this interval ([DeCelles et al.,](#)

647 [2007b](#)), it suggests that the cause of the decreasing exhumation rate is slowing of
648 local tectonic activity.

649 The extent to which these events can be applied more widely across the region is
650 debatable. The topographic surface expression of the 40-30 Ma thrust activity along
651 the Zanggenong Thrust, which we propose may have resulted in the change from
652 external to internal drainage in the Nima Basin, may be local. Such a proposition is
653 supported firstly by the work of [Kapp et al. \(2007\)](#) who suggested the magnitude of
654 slip on the Zanggenong Thrust is low compared with Cretaceous thrusts. Secondly,
655 magnetostratigraphic studies indicate that the Lunpola Basin experienced relatively
656 slow subsidence (100 m/Ma) from 40 Ma to 25 Ma, suggesting that deformation was
657 limited ([Fang et al., 2020](#), but note the debate on the depositional age dating of this
658 basin in [Su et al., 2019](#)). Thirdly, [Rohrmann et al. \(2012\)](#) suggested that the limited
659 number of cooling ages less than 40 Ma in central Tibet be attributed to localized
660 thrust reactivation.

661 Unlike the Zanggenong Thrust which is a localized feature, the Muggar Thrust is part
662 of the major 1200 km long Shiquanhe-Gaize-Amdo thrust system ([Yin and Harrison,](#)
663 [2000](#)). The cessation of activity in the Nima area at 40 Ma is in agreement with the
664 regional synthesis of [Rohrmann et al. \(2012\)](#) and also with our own regional
665 compilation of such LTT data (Figure 1a) which indicates that the central part of the
666 plateau experienced little exhumation and tectonic quiescence since this time. In
667 summary, we suggest that the formation mechanism of low relief topography of
668 central Tibet may be closely related to the extensive tectonic quiescence, the

669 persistence of an arid climate and/or the establishment of internal drainage since
670 ~40-30 Ma. Such a suggestion is in agreement with the observation of [Law and Allen](#)
671 [\(2020\)](#) who determined that of low relief topography occurred soon after cessation of
672 exhumation in northern and central Tibet.

673 **7 Conclusions**

674 The Muggar Thrust, part of the 1200 kms long Shiquanhe-Gaize-Amdo thrust system
675 along the Bangong-Nujiang Suture Zone, experienced an exhumational event between
676 70-40 Ma in the Nima area. Reactivation of local thrusts in the adjacent Nima Basin
677 to the south occurred between 40-30 Ma, after which time exhumation in the region
678 slowed. Consistent with the localized provenance and restricted facies in the Nima
679 Basin, we propose that the thrust-generated topographic expression of relief triggered
680 initiation of internal drainage by ~30 Ma, with subsequent basinal aggradation and
681 development of low relief topography.

682

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696 **Data Availability Statement**

697 Supporting texts, figures, tables and datasets (Text S1, Figure S1-S6 and Table S1-S6)
698 are provided in Supplementary materials 1 and 2, which can be found at the figshare
699 (<https://doi.org/10.6084/m9.figshare.14995134>).

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Figure 1.

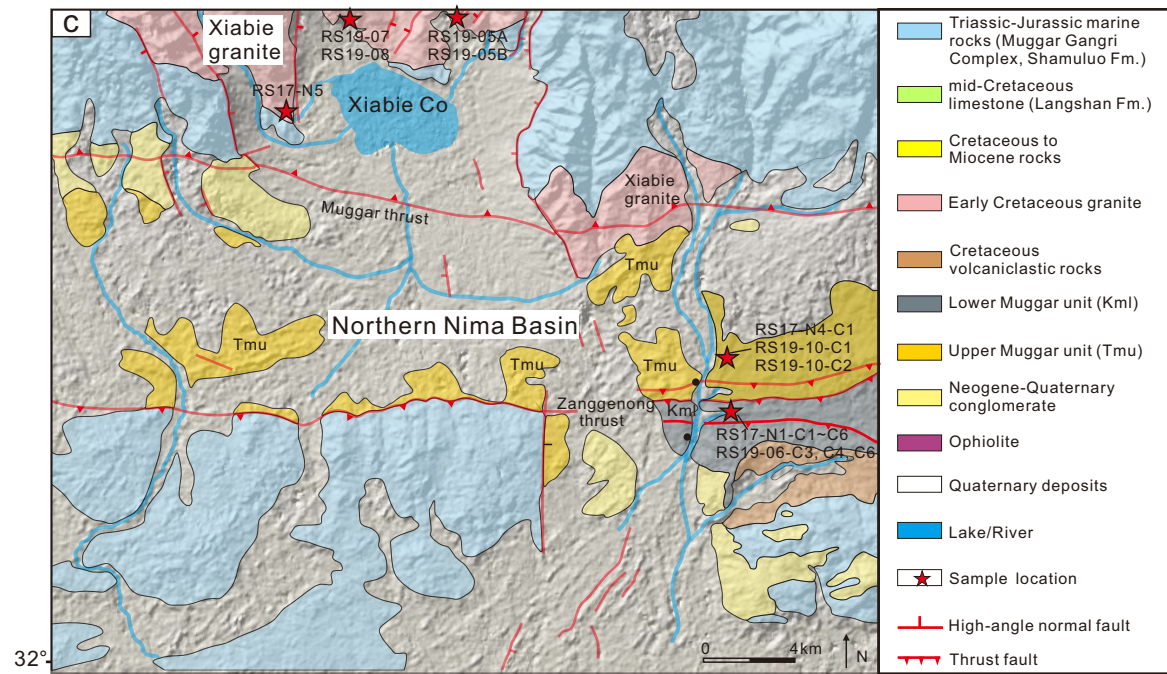
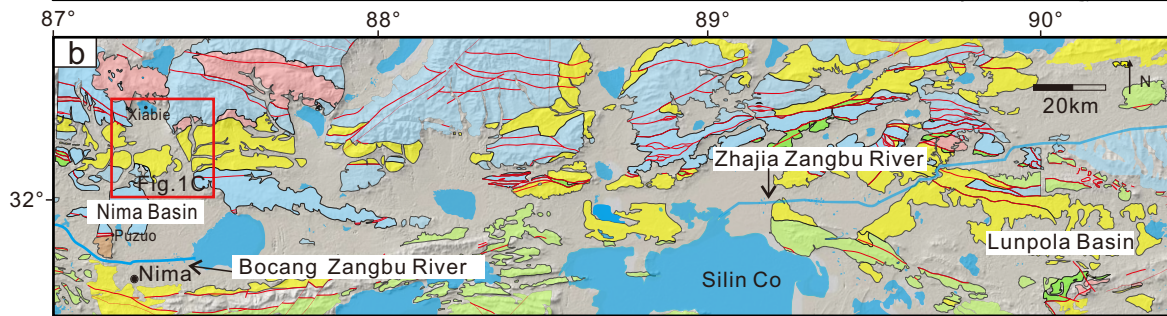
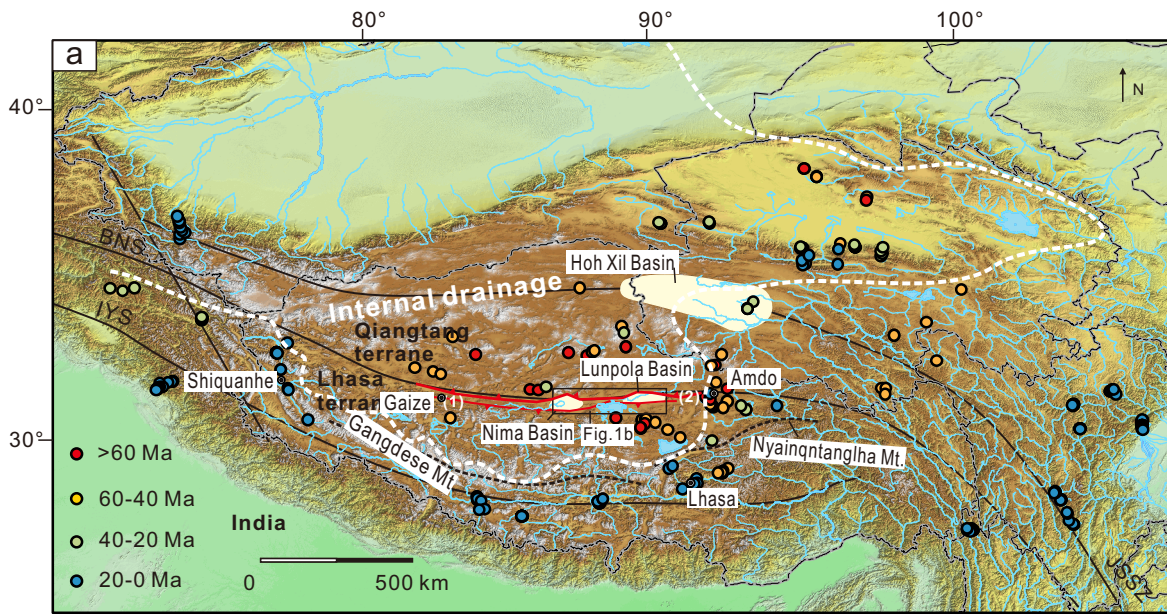


Figure 2.

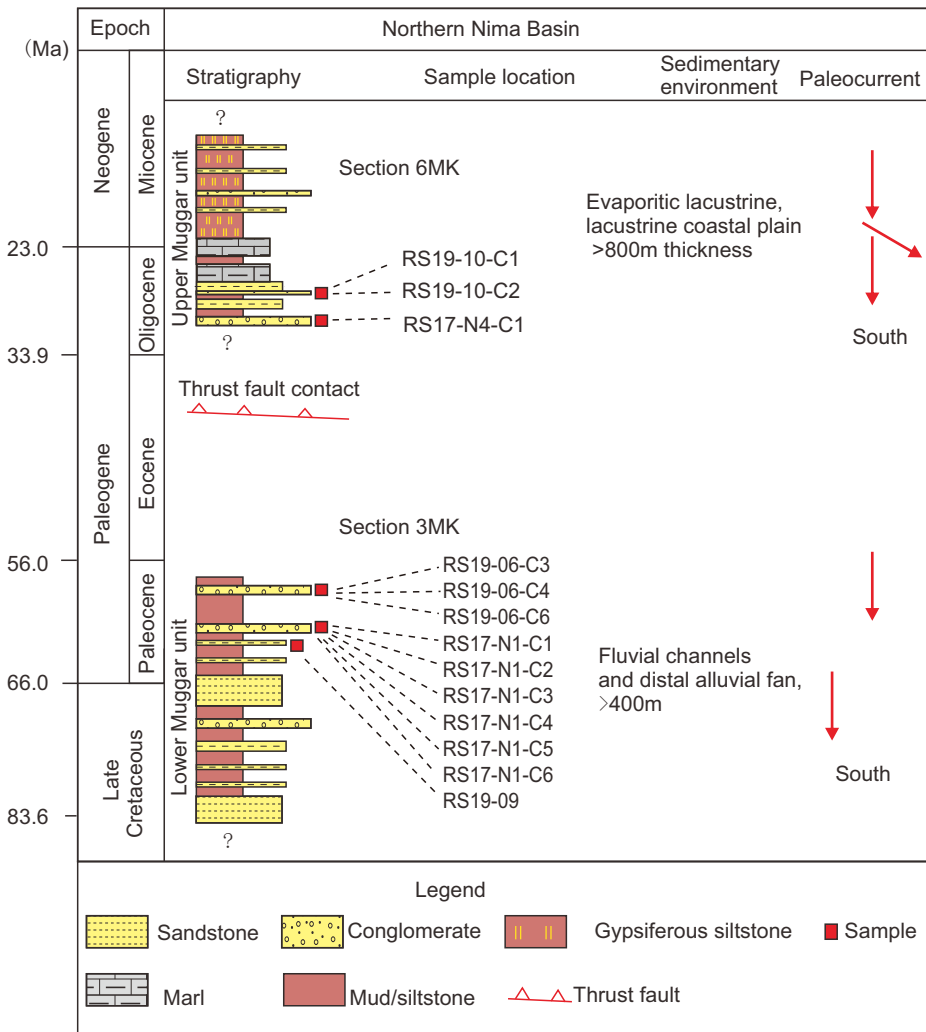


Figure 3.

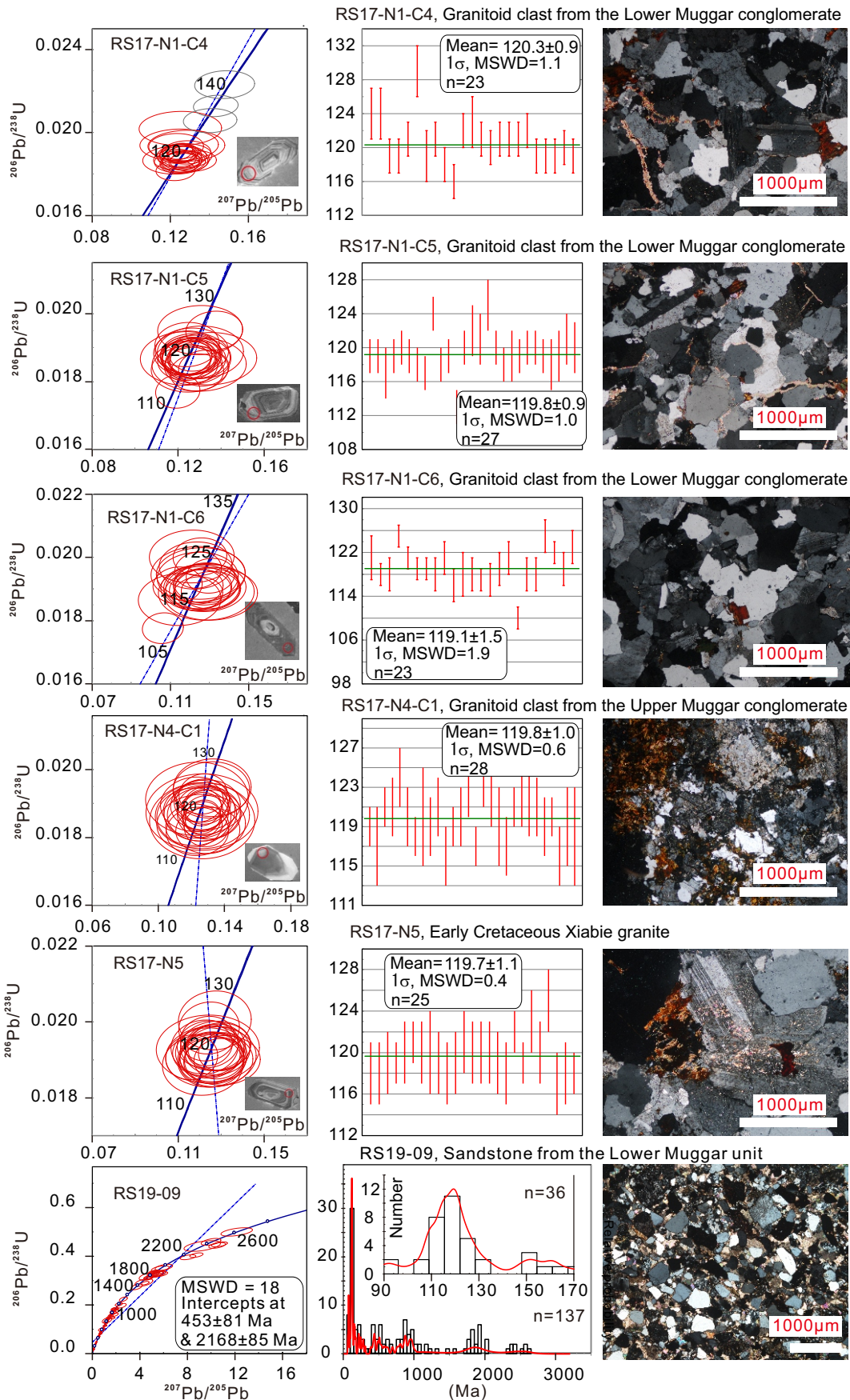


Figure 4.

Figure 5.

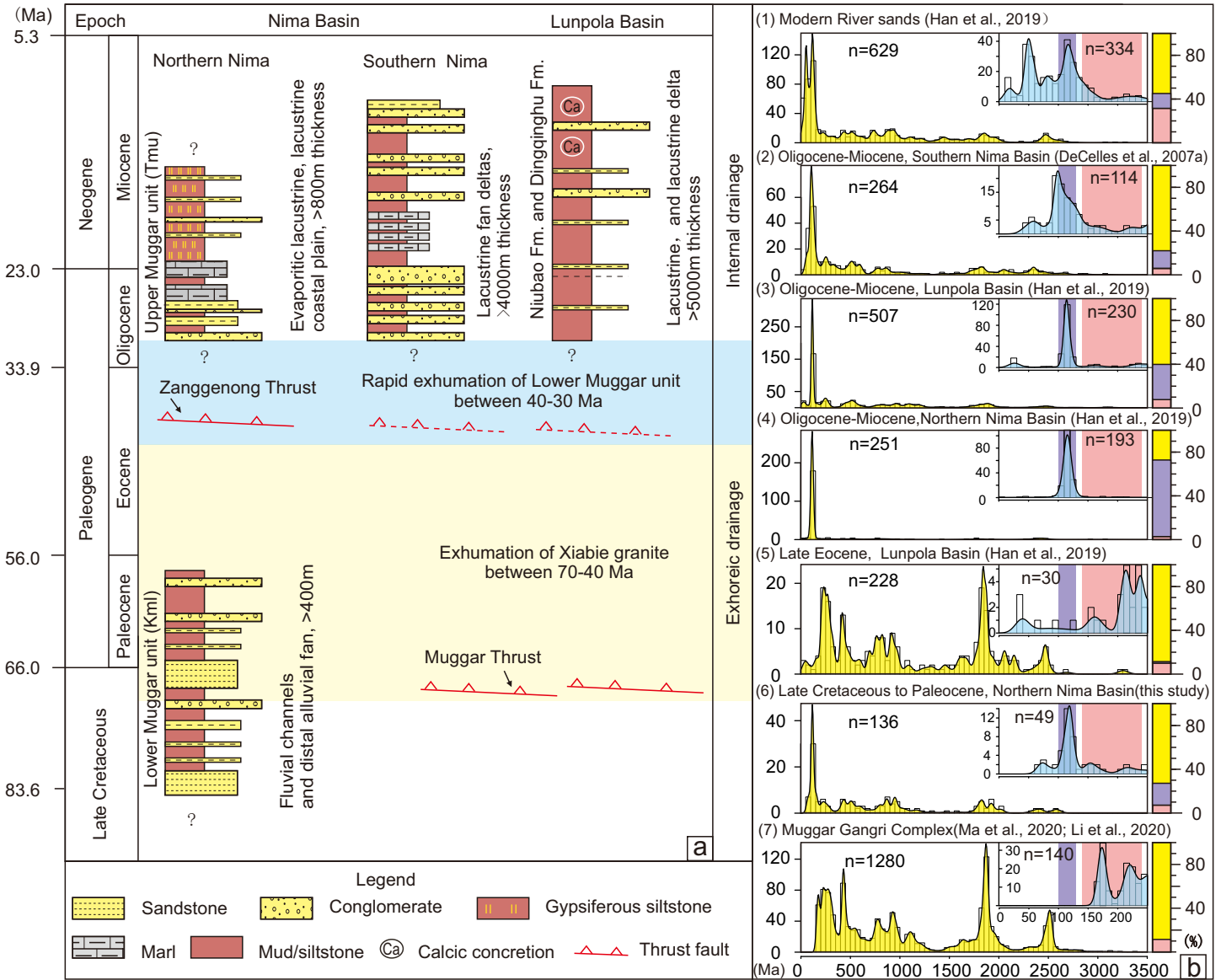


Table 1 Location, analytical methods employed and lithological details of samples from the Northern Nima Basin and Xiabie granite

Number	Sample	Longitude (°E)	Latitude (°N)	Sample position	Lithology	AFT	AHe	Zircon U-Pb	ZHe
1	RS17-N1-C4	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
2	RS17-N1-C5	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
3	RS17-N1-C6	87.4168	32.1014	Lower Muggar Unit	granitoid conglomerate clast	√	√	√	√
4	RS19-06-C3	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
5	RS19-06-C6	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
6	RS19-06-C4	87.4171	32.1017	Lower Muggar Unit	granitoid conglomerate clast	√	√		
7	RS17-N1-C1	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
8	RS17-N1-C2	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
9	RS17-N1-C3	87.4168	32.1014	Lower Muggar Unit	buff sandstone conglomerate clast	√	√		
10	RS19-09	87.4171	32.1017	Lower Muggar Unit	sandstone			√	
11	RS17-N4-C1	87.4225	32.1139	Upper Muggar Unit	granitoid conglomerate clast	√	√	√	√
12	RS19-10-C1	87.4210	32.1119	Upper Muggar Unit	granitoid conglomerate clast	√	√		
13	RS19-10-C2	87.4210	32.1119	Upper Muggar Unit	granitoid conglomerate clast		√		
14	RS17-N5	87.2241	32.2186	Xiabie granite	granite	√	√	√	
15	RS19-05B	87.2973	32.2615	Xiabie granite	granite	√	√		
16	RS19-08	87.2694	32.2599	Xiabie granite	granite	√			
17	RS19-07	87.2945	32.2607	Xiabie granite	granite	√			
18	RS19-05A	87.3042	32.2447	Xiabie granite	granite	√			

Note: Tick √ represents the analysis technique carried out for every sample

Table 2. Results of Apatite (U-Th)/He Analysis

Grain name	⁴ He (cc)	²³⁸ U (ng)	²³² Th (ng)	eU ppm	²¹ T	^b L (um)	^c W (um)	Ft	Th/U	Raw age (Ma)	error (Ma)	Cor. Age (Ma)	error (Ma)	Ave. Raw age and SD (Ma)	Ave. Cor. Age And SD (Ma)
RS17-N1-C1-02	2.3E-10	0.019	0.308	28.5	0	120	103	0.7	15.8	20.7	0.8	28.6	1.2		
RS17-N1-C1-03	2.1E-10	0.021	0.229	33.6	0	110	90	0.7	10.7	23.4	0.3	34.1	0.5		
RS17-N1-C1-05	2.8E-10	0.051	0.202	48.2	0	100	90	0.7	4.0	23.7	0.3	34.9	0.4		
RS17-N1-C1-06	2.3E-10	0.037	0.293	26.7	0	130	110	0.7	8.0	17.8	0.2	24.0	0.3		
RS17-N1-C1-07	8.0E-11	0.019	0.244	41.8	0	100	85	0.7	13.1	8.7	0.1	13.0	0.2	/	/
RS17-N1-C2-02	6.0E-11	0.008	0.040	14.3	0	90	73	0.6	5.2	29.1	1.4	47.6	2.2		
RS17-N1-C2-03	8.7E-11	0.015	0.050	16.4	0	100	80	0.7	3.4	27.2	1.0	41.9	1.5		
RS17-N1-C2-04	9.2E-11	0.012	0.087	26.0	0	100	70	0.6	7.4	23.4	0.6	38.1	0.9		
RS17-N1-C2-05	5.2E-11	0.003	0.075	25.4	0	90	60	0.6	23.5	20.8	0.6	37.4	1.1		
RS17-N1-C2-06	9.9E-11	0.035	0.140	94.4	0	80	60	0.5	3.9	11.9	0.4	22.0	0.8	/	/
RS17-N1-C3-01	7.4E-11	0.017	0.131	37.7	0	90	75	0.6	7.6	12.6	0.2	20.3	0.4		
RS17-N1-C3-02	2.7E-11	0.012	0.026	19.4	0	90	65	0.6	2.1	12.0	0.7	20.7	1.2		
RS17-N1-C3-03	4.6E-11	0.018	0.126	23.2	0	100	90	0.7	7.1	8.0	0.1	11.8	0.2		
RS17-N1-C3-04	7.1E-11	0.018	0.081	20.3	0	100	85	0.7	4.5	15.7	0.4	23.7	0.6		
RS17-N1-C3-05	1.7E-10	0.028	0.181	28.2	0	110	95	0.7	6.5	19.4	0.3	27.7	0.4	/	/
RS17-N1-C4-1	1.8E-10	0.012	0.121	48.1	1	110	55	0.5	10.1	35.9	0.5	65.8	0.9		
RS17-N1-C4-2	3.1E-10	0.032	0.170	44.2	1	140	68	0.6	5.3	35.4	0.5	55.8	0.8		
RS17-N1-C4-3	6.6E-10	0.021	0.455	101.0	1	140	60	0.6	21.7	42.2	0.6	70.8	1.0	37.8 ± 3.1	64.2 ± 6.2
RS17-N1-C5-1A	3.7E-10	0.042	0.272	34.4	1	160	88	0.7	6.5	28.7	0.3	40.5	0.5		
RS17-N1-C5-2A	1.5E-10	0.024	0.127	44.7	1	160	55	0.6	5.2	23.1	0.4	40.3	0.7		
RS17-NI-C5-1	3.8E-10	0.046	0.236	24.9	2	200	90	0.7	5.1	30.4	0.4	41.7	0.5		

RS17-NI-C5-3	7.5E-10	0.096	0.354	31.1	1	200	107	0.8	3.7	34.2	0.4	45.1	0.5		
RS17-NI-C5-4	8.5E-10	0.068	0.429	76.0	1	180	70	0.7	6.3	41.5	0.5	63.0	0.7		
RS17-NI-C5-5	6.1E-10	0.106	0.313	14.5	1	300	128	0.8	3.0	28.1	0.3	35.1	0.4	31.0 ± 5.7	44.3 ± 8.8
RS17-N1-C6-2	5.6E-10	0.043	0.375	54.8	2	170	75	0.7	8.6	35.1	0.5	52.0	0.7		
RS17-N1-C6-3	6.3E-10	0.043	0.307	60.3	2	155	70	0.6	7.2	45.1	0.6	69.4	0.9		
RS17-N1-C6-4	3.2E-10	0.066	0.279	59.9	2	155	75	0.7	4.3	20.1	0.2	30.0	0.3		
RS17-N1-C6-2a	1.1E-09	0.108	0.488	157.3	1	137	64	0.6	4.5	41.8	1.0	67.9	1.6		
RS17-N1-C6-3a	3.3E-10	0.033	0.237	49.1	1	140	72	0.6	7.1	30.8	1.5	47.5	2.3		
RS17-N1-C6-4a	3.8E-10	0.028	0.163	44.7	2	140	65	0.6	5.8	46.5	0.7	74.9	1.1		
RS17-N1-C6-6a	5.2E-10	0.098	0.423	72.6	1	150	85	0.7	4.3	21.5	0.2	30.8	0.3	/	/
RS19-06-C3-3	2.8E-10	0.011	0.264	39.0	2	160	68	0.60	23.64	31.62	0.5	52.8	0.8		
RS19-06-C3-5	2.9E-10	0.019	0.242	32.5	2	157	77	0.63	12.78	30.96	0.5	48.8	0.7		
RS19-06-C3-1c	9.6E-11	0.011	0.121	18.3	2	130	81	0.63	10.90	19.94	0.5	31.7	0.8		
RS19-06-C3-2c	1.1E-10	0.014	0.051	11.3	2	157	77	0.65	3.58	35.98	1.6	55.6	2.5		
RS19-06-C3-3c	1.7E-10	0.022	0.075	10.3	2	190	89	0.70	3.46	36.10	1.1	51.7	1.6	30.9 ± 5.9	48.1 ± 8.5
RS19-06-C4-1	6.1E-10	0.058	0.359	21.2	2	210	113	0.74	6.14	34.89	0.3	46.4	0.4		
RS19-06-C4-2	4.8E-10	0.024	0.196	22.4	2	180	83	0.67	8.11	55.50	0.6	82.2	0.8		
RS19-06-C4-3	7.9E-10	0.054	0.209	11.7	2	236	122	0.77	3.87	62.40	0.6	80.5	0.8		
RS19-06-C4-4	9.2E-10	0.099	0.454	42.4	1	232	91	0.73	4.58	36.60	0.3	49.6	0.4	/	/
RS19-06-C6-1	8.4E-11	0.013	0.122	20.2	2	130	80	0.63	9.18	16.10	0.2	26.0	0.3		
RS19-06-C6-2	1.1E-10	0.014	0.076	12.8	2	152	80	0.65	5.54	28.44	0.3	44.5	0.5		
RS19-06-C6-3	2.5E-10	0.026	0.207	29.3	0	163	79	0.65	7.82	27.33	0.3	41.8	0.4		
RS19-06-C6-4	1.3E-10	0.015	0.147	24.4	1	159	72	0.65	9.61	21.28	0.2	32.8	0.4		
RS19-06-C6-5	2.0E-10	0.014	0.166	26.6	1	130	78	0.67	11.49	29.93	0.3	44.3	0.5		
RS19-06-C6-1C	4.4E-10	0.050	0.406	26.7	2	183	109	0.72	8.17	24.78	0.2	34.2	0.3	24.6 ± 4.7	37.3 ± 6.8
RS19-10-C1-1	3.3E-10	0.021	0.341	60.3	2	130	72	0.60	15.96	26.69	0.3	44.9	0.6		

RS19-10-C1-2	2.8E-10	0.011	0.296	56.8	2	123	68	0.57	27.77	28.13	0.4	49.3	0.7		
RS19-10-C1-3	1.8E-10	0.020	0.101	22.8	2	160	69	0.62	4.98	32.71	0.8	52.8	1.3		
RS19-10-C1-4	3.5E-10	0.029	0.258	40.2	2	150	77	0.63	9.03	32.45	0.4	51.4	0.7		
RS19-10-C1-5	1.1E-10	0.016	0.122	26.3	1	118	76	0.66	7.49	19.53	0.4	29.6	0.7	27.9 ± 4.8	45.6 ± 8.4
RS19-10-C2-1	2.0E-10	0.037	0.355	71.1	2	147	68	0.60	9.62	13.38	0.2	22.4	0.3		
RS19-10-C2-2	4.5E-10	0.029	0.389	37.2	2	184	84	0.67	13.34	30.37	0.4	45.7	0.5		
RS19-10-C2-3	4.4E-10	0.038	0.222	29.7	1	183	81	0.69	5.87	40.36	0.6	58.2	0.8		
RS19-10-C2-4	2.6E-10	0.021	0.127	18.9	2	162	81	0.66	6.12	42.93	0.9	65.3	1.4		
RS19-10-C2-5	2.4E-10	0.020	0.175	26.6	1	141	80	0.68	8.96	32.22	0.6	47.4	0.8	/	/
RS17-N4-C1-1A	3.0E-10	0.031	0.142	26.8	1	170	75	0.7	4.6	38.3	0.6	56.8	0.9		
RS17-N4-C1-1	2.8E-10	0.033	0.154	16.4	2	200	92	0.7	4.6	33.2	0.4	45.3	0.5		
RS17-N4-C1-2	8.3E-10	0.074	0.380	65.4	1	250	63	0.6	5.2	41.9	0.4	74.8	0.8		
RS17-N4-C1-3	8.4E-10	0.070	0.613	55.1	1	200	88	0.7	8.7	32.1	0.4	44.1	0.5		
RS17-N4-C1-4	9.7E-10	0.128	0.510	42.9	1	180	113	0.8	4.0	31.9	0.4	41.4	0.5	35.7 ± 7.4	51.6 ± 3.5
RS19-05B-1	4.2E-10	0.054	0.219	45.3	2	146	80	0.65	4.06	32.97	0.5	50.9	0.7		
RS19-05B-2A	1.7E-10	0.021	0.075	18.6	2	170	70	0.63	3.53	36.20	1.1	57.5	1.8		
RS19-05B-3	5.0E-10	0.073	0.265	73.4	1	149	70	0.62	3.64	30.10	0.4	48.5	0.6		
RS19-05B-4	3.1E-10	0.046	0.165	58.3	2	110	72	0.60	3.60	29.58	0.5	49.6	0.8		
RS19-05B-5	3.6E-10	0.036	0.099	26.3	1	180	71	0.64	2.75	49.80	1.2	77.8	1.8	/	/
RS17-N5-1	1.8E-10	0.021	0.087	22.4	2	175	65	0.6	4.1	35.0	0.5	55.1	0.7		
RS17-N5-2	3.0E-10	0.028	0.098	40.7	1	140	60	0.6	3.5	47.1	0.7	79.2	1.1	41.1 ± 6.1	67.1 ± 12.0

Note. a. T—termination; b. L—length; c. W—width; For samples with standard deviation greater than 20%, the mean age is not calculated (Flowers and Kelley, 2011).

Table 3. Results of Zircon (U-Th)/He Analysis

Grain name	⁴ He (cc)	²³⁸ U (ppm)	²³² Th (ppm)	eU (ppm)	[°] T	^a L (um)	^b W (um)	F(T)	Th/U	Raw age (Ma)	Error (Ma)	Corrected age (Ma)	Error (Ma)	Ave. Raw age and SD (Ma)	Ave. Cor. Age and SD (Ma)
RS17-N1-C4-Z1	2.49E-08	588.88	361.23	673.77	2	200	75	0.74	0.61	106.70	5.95	113.11	8.01		
RS17-N1-C4-Z2	6.92E-09	373.14	424.29	472.85	2	180	65	0.71	1.14	62.66	4.45	88.73	6.31		
RS17-N1-C4-Z3	1.64E-08	1107.38	516.58	1228.78	2	190	53	0.67	0.47	82.73	4.61	97.41	6.90		
RS17-N1-C4-Z4	2.31E-09	217.60	318.86	292.54	2	150	55	0.65	1.47	56.55	4.06	86.41	6.22		
RS17-N1-C4-Z5	1.04E-08	925.97	533.22	1051.27	2	170	50	0.64	0.58	75.55	5.35	117.72	8.35		
RS17-N1-C4-Z6	1.05E-08	980.32	461.13	1088.68	2	130	53	0.64	0.47	87.39	4.88	107.25	7.61		
RS17-N1-C4-Z7	9.86E-09	1249.02	651.77	1402.18	2	130	55	0.65	0.52	58.21	4.12	89.13	6.31	/	/
RS17-N1-C5-Z1	8.15E-09	870.23	767.37	1050.57	2	140	50	0.63	0.88	72.04	5.11	114.90	8.15		
RS17-N1-C5-Z2	2.68E-08	1866.43	1223.56	2153.97	2	150	60	0.68	0.66	74.85	5.29	109.98	7.77		
RS17-N1-C5-Z3	2.38E-08	1596.10	723.78	1766.19	2	190	55	0.67	0.45	76.21	5.38	113.89	8.04		
RS17-N1-C5-Z4	1.04E-08	560.04	398.59	653.71	2	180	55	0.67	0.71	94.72	6.71	142.33	10.08		
RS17-N1-C5-Z5	1.61E-08	847.21	542.20	974.63	2	180	65	0.71	0.64	70.79	5.01	100.01	7.07		
RS17-N1-C5-Z6	2.17E-08	1346.13	755.57	1523.68	2	160	60	0.68	0.56	80.52	5.69	118.61	8.38		
RS17-N1-C5-Z7	8.08E-09	418.03	516.81	539.48	2	150	53	0.64	1.24	117.41	8.36	182.36	12.98	78 ± 9	117 ± 14
RS17-N1-C6-Z1	1.57E-08	1168.61	414.97	1266.13	2	170	53	0.66	0.36	86.17	6.09	131.54	9.29		
RS17-N1-C6-Z2	1.24E-08	1105.77	732.77	1277.97	2	170	50	0.64	0.66	74.17	5.25	115.93	8.20		
RS17-N1-C6-Z3	1.60E-08	1532.07	830.99	1727.35	2	150	50	0.64	0.54	80.08	5.66	125.99	8.90		
RS17-N1-C6-Z4	5.24E-09	440.67	409.93	537.00	2	120	50	0.62	0.93	105.48	7.54	169.78	12.13	86 ± 14	136 ± 24
RS17-N4-C1-Z1	3.83E-09	198.38	185.40	241.95	2	210	59	0.69	0.93	70.34	5.02	102.02	7.28		
RS17-N4-C1-Z2	6.94E-09	198.17	215.01	248.70	2	220	78	0.75	1.08	68.65	4.88	92.01	6.54		
RS17-N4-C1-Z3	1.30E-08	354.01	331.52	431.91	2	280	70	0.73	0.94	71.16	5.04	97.09	6.88		
RS17-N4-C1-Z4	1.16E-08	228.66	211.63	278.39	2	250	95	0.79	0.93	59.94	4.25	75.62	5.36		

RS17-N4-C1-Z5	7.10E-09	293.59	245.53	351.29	2	200	75	0.74	0.84	58.47	4.15	78.81	5.59		
RS17-N4-C1-Z6	1.99E-08	474.62	469.85	585.04	2	260	75	0.75	0.99	75.77	5.37	101.11	7.16		
RS17-N4-C1-Z7	4.35E-09	152.32	162.25	190.45	2	240	70	0.73	1.07	63.24	4.50	86.26	6.14	67 ± 6	90 ± 11

Note. a. L—length; b. W—width; c. T—termination. For samples with standard deviation greater than 20%, the mean age is not calculated (Flowers and Kelley, 2011).

Table 4. Results of Apatite Fission Track Analysis

Sample	No. of Grains	^a ρ _s (10 ⁵ /cm ²)	Total ^b N _s	^a ρ _i (10 ⁵ /cm ²)	Total ^b N _i	^a ρ _d (10 ⁵ /cm ²)	^b N _d	^c P(χ) ²	^d D _{par} (μm)	²³⁸ U ppm	Central AFT Age (Ma)	±1σ	Pooled AFT Age (Ma)	±1σ	Dis. (%)	MTL c Axis Correction (μm)	SD (μm)	N ^e
RS19-05-A	16	5.1	250	32.0	1562	17.7	14503	0.50	1.5	25.3	42.5	5.3	42.5	2.9	0	13.5	1.4	8
RS19-05-B	20	7.4	368	31.1	1556	16.1	14503	0.02	1.5	28.4	57.1	4.4	57.1	3.3	20	13.96	1.2	27
RS19-08	17	3.1	263	17.8	1488	18.8	14503	0.60	1.6	14.9	50.0	3.3	49.0	3.3	0	14.33	1.1	9
RS19-07	12	5.4	209	25.5	990	18.1	14503	0.17	1.6	23.6	56.9	7.7	57.2	4.4	11	13.41	1.3	15
RS17-N5*	22	5.2	179	-	-	-	-	0.00	1.1	2.1	89.0	10	96.8	7.3	37	12.48	1.3	5
RS19-10-C1	20	5.7	196	23.5	810	18.4	14503	0.43	1.5	17.4	67.0	5.3	67.0	5.3	0	13.44	1.2	4
RS17-N4-C1*	14	4.8	174	-	-	-	-	0.09	1.4	3.8	51.3	4.8	51.3	3.9	19	12.29	0.5	2
RS19-06-C3	12	2.1	93	15.8	697	16.5	14503	0.23	1.6	15.8	33.0	4.1	33.1	3.7	18	-	-	0
RS19-06-C4	20	2.0	189	14.2	1317	14.5	14503	0.10	1.5	13.5	31.2	2.9	31.3	2.4	19	13.24	0.6	3
RS19-06-C6	20	2.0	170	15.6	1313	17.3	14503	0.10	1.6	12.9	33.4	3.4	33.6	2.7	24	14.45	1.0	3
RS17-N1-C6*	20	2.0	124	-	-	-	-	0.69	1.3	2.2	33.6	3.0	33.6	3.0	0	-	-	0
RS17-N1-C4*	18	3.4	129	-	-	-	-	0.00	1.2	2.3	62.0	10.0	71.0	6.3	56	12.70	0.2	3
RS17-N1-C5*	9	4.5	57	-	-	-	-	0.03	1.2	4.8	77.0	15.0	77.0	10	39	-	-	0
RS17-N1-C1	26	5.4	313	19.8	1143	12.5	8790	0.04	1.7	23.5	53.1	4.3	51.7	10.8	22	13.30	0.6	6
RS17-N1-C2	11	4.6	69	17.5	262	12.4	8790	0.89	1.8	20.6	49.7	6.7	49.3	10.3	0	12.26	1.0	5
RS17-N1-C3	20	13.6	481	30.8	1091	12.2	8790	0.00	2.0	36.1	83.0	10.0	81.0	16.9	47	12.78	1.5	15

*Concentration of U measured by LA-ICP-MS

- a. $\rho_{(i,s,d)}$ are track density of induced, spontaneous, dosimeter tracks.
- b. N_i , N_s , and N_d are the number of induced, spontaneous, and dosimeter tracks counted.
- c. $P(\chi^2)$ value of the chi-square age homogeneity test.
- d. D_{par} measurements are etch pit diameters used as a proxy for the influence of chemical composition on track annealing [[Donelick et al., 2005](#)]. Three to five D_{par} measurements were used for each grain.
- e. The number of confined tracks