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# Results for a turbulent system with unbounded viscosities: weak formulations, existence of solutions, boundedness, smoothness 

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#### Abstract

We consider a circulation system arising in turbulence modelling in fluid dynamics with unbounded eddy viscosities. Various notions of weak solution are considered and compared. We establish existence and regularity results. In particular we study the boundedness of weak solutions. We also establish an existence result for a classical solution.


Key words: degenerate elliptic system, weak solution, regularity, oceanography. 1991 MSC: 35J70, 35D05, 35D10, 86A05

## 1 Introduction

Let $\Omega$ be an open bounded set in $\mathbb{R}^{3}$, with a Lipschitz boundary. We consider the following turbulent circulation model :

$$
\text { (P) }\left\{\begin{array}{l}
-\operatorname{div}(\nu(k) \nabla u)=f \quad \text { in } \Omega \\
-\operatorname{div}(a(k) \nabla k)=\nu(k)|\nabla u|^{2} \quad \text { in } \Omega \\
u=0 \quad \text { on } \partial \Omega \\
k=0 \quad \text { on } \partial \Omega
\end{array}\right.
$$

Here $f, a$ and $\nu$ are given, and the functions $u, k: \Omega \rightarrow \mathbb{R}$ are the unknowns.

[^0]We study Problem $(P)$ under the following main assumption:

$$
\left(H_{0}\right)\left\{\begin{array}{l}
f \in L^{r}(\Omega), \text { with } r>\frac{3}{2} \\
a, \nu: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+} \text {are continuous } \\
\exists \delta>0: a(s), \nu(s) \geq \delta \quad \forall s \in \mathbb{R}^{+}
\end{array}\right.
$$

Problem (P) is a simplified scalar version of the RANS model arising in oceanography (see $[10,11,1]$ ): the function $u$ is an idealisation of the mean velocity of the fluid and $k$ is the turbulent kinetic energy. The mathematical analysis of $(\mathrm{P})$ is a step towards better understanding the RANS model. Various studies were made in this direction. Some existence results were established in [10,7].
In this paper we focus on the case where the viscosity functions $a$ and $\nu$ are not a priori bounded. In fact (see [11,7]), in the relevant physical situation, we have

$$
\left(H_{p}\right)\left\{\begin{array}{l}
a(s)=a_{1}+a_{2} \sqrt{s} \\
\nu(s)=\nu_{1}+\nu_{2} \sqrt{s}
\end{array}\right.
$$

We will establish an existence result for a weak solution for ( P ) under less restrictive assumptions than in [7]. An important feature is that our assumptions are satisfied under $\left(H_{p}\right)$, contrarily to the assumptions made in [7].
Moreover we give additional regularity results for the weak solution we obtain. In particular, under $\left(H_{0}\right)$ and the following additional assumption: $a$ is proportional to $\nu, \partial \Omega$ is of class $\mathcal{C}^{2, \alpha}, f \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$ and $\nu \in \mathcal{C}^{1, \alpha}\left(\mathbb{R}^{+}\right)$, we prove the existence of a classical solution for (P).

We also compare our results with the results presented in [10].

Another feature of our work is to considere various notions of weak solution for Problem (P): $W$-solution, $H$-solution, distributional solution, renormalized solution, 'energy solution', classical solution. We give some relations between these notions.

### 1.1 Notions of weak solution for ( $P$ )

We can reformulate equation (P). 2 by using the Kirchoff transform. Let

$$
A(s):=\int_{0}^{s} a(t) d t
$$

Instead of (P).2, we can consider

$$
(P) \cdot 2^{\prime} \quad-\Delta K=\nu \circ A^{-1}(K)|\nabla u|^{2} \quad \text { on } \Omega,
$$

where $K=A(k)$.
In fact, from every distributional solution $K \in W^{1}(\Omega)$ of (P).2' we obtain a distributional solution $k$ of $(\mathrm{P}) .2$ by setting $k=A^{-1}(K)$. This property is related to the facts that $A$ is invertible, $A^{-1}(0)=0$ and $\left|A^{-1}(s)\right| \leq C . s$ (this can be seen by using the assumptions made on $\nu$ in $\left.\left(H_{0}\right)\right)$.

The situation is more complicated for equation (P).1, where the a priori unbounded coefficient $\nu(k)$ appears in the principal part of the operator and cannot be removed. Hence we have to restrict $u$ to satisfy the energy condition

$$
\begin{equation*}
\int_{\Omega} \nu(k)|\nabla u|^{2}<\infty \tag{1}
\end{equation*}
$$

Nevertheless we will see later on that various non equivalent notions of weak solution can be considered for (P).1.
We will introduce the notions of W -solution and H -solution. It is also possible to consider the notion of renormalized solution (see [10] chap.5). In [7] the authors defined another notion that they call energy solution.
We will give some relations between these notions in the Appendix I.

Remark now that under the restriction (1), the right hand side in (P). 2 (or in (P). $2^{\prime}$ ) is only a priori in $L^{1}(\Omega)$. Hence (see [2]) it is natural to seek $k$ in the space $\cap_{p<3 / 2} W_{0}^{1, p}(\Omega)$.
We want to find a function $u$ vanishing on $\partial \Omega$ that satisfies the energy condition (1). This leads to considering the following spaces:

$$
\begin{aligned}
W_{k} & =\left\{v \in H_{0}^{1}(\Omega):[v]_{k}<\infty\right\} \\
H_{k} & =\text { closure of } \mathcal{C}_{c}^{\infty}(\Omega) \text { with respect to }[\cdot]_{k}
\end{aligned}
$$

where we used the notation

$$
[v]_{k}=\left(\int_{\Omega} \nu(k)|\nabla v|^{2}\right)^{1 / 2}
$$

For any measurable function $k$, the map $[.]_{k}$ defines a norm on $W_{k}$. In the general situation $H_{k}$ and $W_{k}$ are not equal. Moreover $W_{k}$ is not necessarily complete and a function in $H_{k}$ does not always have a uniquely defined gradient (see [14]). If we assume that $\nu(k) \in L^{1}(\Omega)$ then $W_{k}$ is complete and in fact $H_{k} \subset W_{k}$ are Hilbert spaces (see $[6,14,13]$ ) when they are equipped with the scalar product

$$
(v, w)=\int_{\Omega} \nu(k) \nabla v \nabla w
$$

Consequently, we will consider the following two distinct notions of solution for
(P).1:
$u$ is called a $H_{k}$-solution of (P). 1 if $u \in H_{k}$ and $\int_{\Omega} \nu(k) \nabla u \nabla v=\int_{\Omega} f v \quad \forall v \in H_{k}$
$u$ is called a $W_{k}$-solution of (P). 1 if $u \in W_{k}$ and

$$
\int_{\Omega} \nu(k) \nabla u \nabla v=\int_{\Omega} f v \quad \forall v \in W_{k}
$$

Finally, we define the following notions of weak solution for $(\mathrm{P})$ :
$(u, k)$ is called a $\mathbf{H}$-solution of $(\mathrm{P})$ if
$k \in \cap_{p<3 / 2} W_{0}^{1, p}(\Omega), u \in H_{k}$,
$k$ is a distributional solution of $(\mathrm{P}) .2$ and $u$ is a $H_{k}$-solution of ( P ). 1
$(u, k)$ is called a $\mathbf{W}$-solution of ( P ) if
$k \in \cap_{p<3 / 2} W_{0}^{1, p}(\Omega), u \in W_{k}$
$k$ is a distributional solution of $(\mathrm{P}) .2$ and $u$ is a $W_{k}$-solution of (P). 1

### 1.2 Main results

Let $\left(H_{1}\right)$ and $\left(H_{2}\right)$ denote the following conditions:

$$
\begin{array}{lll}
\left(H_{1}\right) \quad \exists \gamma>0: a(s) \geq \gamma \nu(s) & \forall s \in \mathbb{R}^{+} \\
\left(H_{2}\right) \quad \exists \gamma>0: a(s)=\gamma \nu(s) & \forall s \in \mathbb{R}^{+}
\end{array}
$$

We will establish:
Theorem 1 Assume that $\left(H_{0}\right)$ and $\left(H_{1}\right)$ hold. Then there exists at least one $W$ solution $(u, k)$ for $(P)$ such that

$$
\begin{equation*}
u \in L^{\infty}(\Omega) \text { and } \int_{\Omega} a(k)|\nabla k|^{2}<\infty . \tag{2}
\end{equation*}
$$

Corollary 1 Assume that in addition to $\left(H_{0}\right)$ and $\left(H_{1}\right)$ we have

$$
\begin{equation*}
\exists \nu_{0}>0: \quad \nu(s) \leq \nu_{0}\left(1+s^{6}\right), \quad \forall s \in \mathbb{R}^{+} \tag{3}
\end{equation*}
$$

Then the $W$-solution $(u, k)$ given in Theorem 1 is a distributional solution of ( $P$ ).
Theorem 2 Assume that $\left(H_{0}\right)$ and $\left(H_{2}\right)$ hold. Then the $W$-solution $(u, k)$ given in Theorem 1 satisfies

$$
\begin{equation*}
u, k \in \mathcal{C}^{0, \alpha}(\bar{\Omega}), \quad \text { for some } \alpha \in(0,1) \tag{4}
\end{equation*}
$$

Moreover $(u, k)$ is also a H-solution of $(P)$ (and in fact a classical weak solution). If in addition to $\left(H_{0}\right)$ and $\left(H_{2}\right)$ we assume that $\partial \Omega$ is of class $\mathcal{C}^{2, \alpha}, f \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$ and $\nu \in \mathcal{C}^{1, \alpha}\left(\mathbb{R}^{+}\right)$then

$$
u, k \in \mathcal{C}^{2, \beta}(\bar{\Omega}), \quad \text { for some } \beta \in(0,1)
$$

and $(u, k)$ is a classical solution of $(P)$.

### 1.3 Discussion of the results

In Theorem 1 we give an existence result of a W -solution. We next give some regularity results: firstly the property (2) and secondly (in Theorem 2) the property (4). Finally, in Theorem 2 we give an existence result for a classical solution for ( P ).

The main previous studies of Problem (P) are presented in [10] chap. 5 and in [7].

In [10] chap.5, the authors prove the existence of a renormalized solution for (P) under the assumptions $\left(H_{0}\right)$ and $\left(H_{2}\right)$. It seems that their proof also works under $\left(H_{0}\right)$ and $\left(H_{1}\right)$. Nevertheless the notion of renormalized solution is very weak. A renormalized solution $(u, k)$ for $(\mathrm{P})$ is a distributional solution if $\nu(k) \in L^{\infty}(\Omega)$, whereas a H- or a W-solution is a distributional solution if $\nu(k) \in L^{1}(\Omega)$ (see the Appendix I).

In [7] the authors introduced a notion of solution that they call 'energy solution' (see the Appendix I). In fact an 'energy solution' is a W-solution which satisfies an additional property ensuring that $H_{k}=W_{k}$ (the additionnal property imposed is sufficient but not necessary to have this equality). Under this point of view an 'energy solution' is slighty stronger than a W-solution. However, their existence result is obtained by assuming complicated conditions on the coefficients $a$ and $\nu$ which are not exactly satisfied in the physically relevant situation $\left(H_{p}\right)$, but only in the following approximate situation:

$$
\left(H_{p}^{\prime}\right)\left\{\begin{array}{l}
\text { for some } \epsilon>0 \text { we have: } \\
a(s)=a_{1}+a_{2} \sqrt{s+\epsilon} \\
\nu(s)=\nu_{1}+\nu_{2} \sqrt{s+\epsilon}
\end{array}\right.
$$

On the contrary, our assumptions in Theorem 1 and Corollary 1 are very simple, and they are satisfied in $\left(H_{p}\right)$.
Note also that we establish the regulartity property (2) which are not established in [7] (or in [10]).
In the Appendix I we also give a new existence result for an 'energy solution'.

In Theorem 2 we assume that $\left(H_{0}\right)$ and $\left(H_{2}\right)$ hold. These assumptions are fulfilled in the physical situation $\left(H_{p}\right)$ if $a_{2} \nu_{1}=a_{1} \nu_{2}$. We then prove that $u$ and $k$ are Hölder
continuous. In particular we give here a positive answer to a central question put in [7]: $k$ is bounded. Note that in this situation we clearly have $W_{k}=H_{k}$.

We next establish the existence of a classical solution for Problem (P) by assuming some differentiability properties for $a$ and $\nu$. These properties are fulfilled in the situation $\left(H_{p}^{\prime}\right)$ if $a_{2} \nu_{1}=a_{1} \nu_{2}$.
It seems that this result is completely new: the existence of a classical solution for $(\mathrm{P})$ was not studied in any previous work.

### 1.4 Organization of the paper

In the sequel $n$ will always denote an arbitrary integer greater or equal to one, and $C$ (possibly with subscript) will denote a positive real that does not depend on $n$, but that can differ from one part to another.
We always consider the space $H_{0}^{1}(\Omega)$ equipped with the gradient norm.
The condition $\left(H_{0}\right)$ is always assumed.

- In section 2 we introduce an approximate sequence $\left(u_{n}, k_{n}\right)$ of solutions obtained by truncating the coefficients $a$ and $\nu$.
We immediatly obtain the basic estimates :

$$
\begin{array}{r}
\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} \leq C \\
\forall p<\frac{3}{2}: \int_{\Omega}\left|a_{n}\left(k_{n}\right) \nabla k_{n}\right|^{p} \leq C
\end{array}
$$

The point is that we establish the following fundamental estimates:

$$
\begin{align*}
\left\|u_{n}\right\|_{L^{\infty}(\Omega)} & \leq C \\
\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2} & \leq C \tag{*}
\end{align*}
$$

The first estimate above is proved by developping further a technique due to Stampacchia.
The second is obtained under the assumption $\left(H_{1}\right)$. The proof is based on the following idea: if $(u, k)$ is a solution of $(\mathrm{P})$, we formally have ${ }^{1}$

$$
\begin{equation*}
\nu(k)|\nabla u|^{2}=\underbrace{-\operatorname{div}(\nu(k) \nabla u) \cdot u}_{=f u}+\operatorname{div}(\nu(k) u \nabla u) . \tag{5}
\end{equation*}
$$

In other words one can hope that the second member in the second equation in (P) is more regular than it seems.

[^1]In fact, we prove that a similar relation to (5) holds for the approximate sequence. By using next that $\left(u_{n}\right)$ is uniformly bounded in $L^{\infty}(\Omega)$, we obtain $(*)$ which is the key estimate to prove Theorem 1.

- In section 3 we extract from $\left(u_{n}, k_{n}\right)$ a subsequence converging to some element denoted by $(u, k)$. Under the assumptions $\left(H_{0}\right)$ and $\left(H_{1}\right)$, we directly obtain that

$$
u \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega), \quad k \in H_{0}^{1}(\Omega)
$$

We prove that moreover we have:

$$
\int_{\Omega} \nu(k)|\nabla u|^{2}<\infty, \quad \int_{\Omega} a(k)|\nabla k|^{2}<\infty .
$$

- In section 4 we pass to the limit in the approximating Problems. In a first step we prove that $u$ is a $W_{k}$-solution of (P).1. To do this, we use the test functions $v=h_{q}\left(k_{n}\right) \varphi$ (where $\varphi \in W_{k} \cap L^{\infty}(\Omega)$ and $\left(h_{q}\right)$ is a sequence of functions that cut off the large values), and we pass to the limits $n \rightarrow \infty, q \rightarrow \infty$.
We next prove that the energies of the approximating sequence converge to the energy $\int_{\Omega} \nu(k)|\nabla u|^{2}$.
Finally we can pass to the limit in the second equation in order to prove that $k$ is a distributional solution of (P).2. We then obtain Theorem 1 and Corollary 1 follows.
- In section 5 we assume that $\left(H_{0}\right)$ and $\left(H_{2}\right)$ hold. In a first step we obtain the estimate

$$
\left\|k_{n}\right\|_{L^{\infty}(\Omega)} \leq C
$$

Hence $k \in L^{\infty}$ and by using the De Giorgi-Nash Theorem we prove the Hölder continuity of $u$ and $k$.
Next, by assuming additional regularity on $\nu, \partial \Omega$ and $f$ we can apply the Schauder's estimates and we prove Theorem 2.

- In the Appendix I we study some relations between the notions of $W$-solution, $H$-solution, distributional solution, renormalized solution and 'energy solution' for Problem (P). We continue the discussion begun in Subsection 1.3 and we also establish a new existence result for an 'energy solution' for Problem (P).
In the Appendix II we recall some basic properties of Hölder continuous functions.


## 2 Approximating sequence and estimates

We assume that $\left(H_{0}\right)$ holds and we set

$$
\begin{align*}
& \nu_{n}(s)=T_{n}(\nu(s))  \tag{6}\\
& a_{n}(s)=T_{n}(a(s)) \tag{7}
\end{align*}
$$

where $T_{n}$ is the truncated function defined by $T_{n}(t)=\min (n, t)$.
We consider the Problem of finding $\left(u_{n}, k_{n}\right) \in\left(H_{0}^{1}(\Omega)\right)^{2}$ such that

$$
\left(P_{n}\right)\left\{\begin{array}{l}
\int_{\Omega} \nu_{n}\left(k_{n}\right) \nabla u_{n} \nabla v=\int_{\Omega} f v \quad \forall v \in H_{0}^{1}(\Omega) \\
\int_{\Omega} a_{n}\left(k_{n}\right) \nabla k_{n} \nabla \varphi=\int_{\Omega} T_{n}\left(\nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right) \varphi \quad \forall \varphi \in H_{0}^{1}(\Omega)
\end{array}\right.
$$

For any $n \geq 1$, Problem $\left(P_{n}\right)$ is well posed because $a_{n}, \nu_{n} \in L^{\infty}(\mathbb{R})$ and $a_{n}^{-1}, \nu_{n}^{-1} \in$ $L^{\infty}(\mathbb{R})$ by construction.
It is proved in [7] that a solution $\left(u_{n}, k_{n}\right)$ exists for any $n \geq 1$. Moreover, the following basic properties were established:

$$
\begin{align*}
k_{n} & \geq 0  \tag{8}\\
\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} & \leq C_{1}  \tag{9}\\
\forall p<\frac{3}{2}: \int_{\Omega}\left|a_{n}\left(k_{n}\right) \nabla k_{n}\right|^{p} & \leq C_{2} \tag{10}
\end{align*}
$$

We now establish
Lemma 3 The sequence $u_{n}$ is uniformly bounded in the $L^{\infty}(\Omega)$-norm, that is,

$$
\begin{equation*}
\left\|u_{n}\right\|_{L^{\infty}(\Omega)} \leq C_{3} \tag{11}
\end{equation*}
$$

Before proving this lemma we point out that the assumption $f \in L^{r}(\Omega)$, with $r>\frac{3}{2}$ made in $\left(H_{0}\right)$ implies that

$$
\begin{equation*}
f \in W^{-1, \rho}(\Omega), \text { with } \rho=\frac{3 r}{3-r}>3 \tag{12}
\end{equation*}
$$

This last property is easy to prove by using the Sobolev injection Theorem.

## Proof

We will obtain the estimate (11) by using the technique presented on p. 108 in [12]. In order to prove that $C_{3}$ is independent of $n$ we have to detail the technique of Stampacchia.
Let

$$
b_{n}(u, v):=\int_{\Omega} \nu_{n}\left(k_{n}\right) \nabla u \nabla v
$$

Recall that $f$ satisfies (12) and then by using a classical result (see [3]) there exists $g \in\left(L^{\rho}(\Omega)\right)^{3}$ such that $-\operatorname{div}(g)=f$ and $\|g\|_{\left(L^{\rho}(\Omega)\right)^{3}} \leq C\|f\|_{L^{r}(\Omega)}$.
Hence the sequence $u_{n}$ satisfies

$$
\begin{equation*}
b_{n}\left(u_{n}, v\right)=\int_{\Omega} g \nabla v \quad \forall v \in H_{0}^{1}(\Omega) \tag{13}
\end{equation*}
$$

For $s \geq 0$, we define the measurable set $A_{n}(s) \subset \Omega$ by setting

$$
A_{n}(s)=\left\{x \in \Omega:\left|u_{n}(x)\right| \geq s\right\}
$$

We also introduce

$$
\begin{equation*}
\varphi:=\max \left(\left|u_{n}\right|-s, 0\right) \operatorname{sgn}\left(u_{n}\right) \tag{14}
\end{equation*}
$$

It is proved in [12] that $\varphi \in H_{0}^{1}(\Omega)$ and

$$
\begin{aligned}
& \nabla \varphi=\nabla u_{n} \quad \text { in } A_{n}(s) \\
& \nabla \varphi=0 \quad \text { in } \Omega \backslash A_{n}(s)
\end{aligned}
$$

By testing (13) with $v=\varphi$, we obtain

$$
\begin{equation*}
b_{n}(\varphi, \varphi)=b_{n}\left(u_{n}, \varphi\right)=\int_{A_{n}(s)} g \nabla \varphi \tag{15}
\end{equation*}
$$

Remark now that assumption $\nu(s) \geq \delta>0$ in $\left(H_{0}\right)$ implies that $\nu_{n}\left(k_{n}\right) \geq \min (\delta, 1)$. Consequently the bilinear form $b_{n}$ is uniformly coercive on $H_{0}^{1}(\Omega)$. By using this property together with the Hölder inequality, we obtain from (15):

$$
\|\varphi\|_{H_{0}^{1}(\Omega)}^{2} \leq \tilde{C}\left(\int_{A_{n}(s)}|g|^{2}\right)^{1 / 2}\|\varphi\|_{H_{0}^{1}(\Omega)}
$$

Hence by using the Cauchy inequality together with the Hölder inequality we obtain

$$
\begin{equation*}
\|\varphi\|_{H_{0}^{1}(\Omega)^{2}}^{2} \leq \tilde{C}_{1}\|g\|_{L^{\rho}(\Omega)}^{2}\left|A_{n}(s)\right|^{\frac{\rho-2}{\rho}} \tag{16}
\end{equation*}
$$

On the other hand, the Poincaré-Sobolev inequality gives

$$
\begin{equation*}
\left(\int_{A_{n}(s)}|\varphi|^{6}\right)^{1 / 3} \leq \tilde{C}_{2}\|\varphi\|_{H_{0}^{1}(\Omega)}^{2} \tag{17}
\end{equation*}
$$

Let now $t>s$. It is clear that $A_{n}(t) \subset A_{n}(s)$ and consequently

$$
\begin{equation*}
\left(\int_{A_{n}(s)}|\varphi|^{6}\right)^{1 / 3} \geq\left(\int_{A_{n}(t)}|\varphi|^{6}\right)^{1 / 3} \geq\left(\int_{A_{n}(t)}|t-s|^{6}\right)^{1 / 3} \geq|t-s|^{2}\left|A_{n}(t)\right|^{1 / 3} \tag{18}
\end{equation*}
$$

We set

$$
\psi_{n}(s):=\left|A_{n}(s)\right|, \quad \forall s \geq 0
$$

For fixed $n, \psi_{n}$ is a decreasing function, and from the estimates (16)-(18), we obtain

$$
\psi_{n}(t) \leq \tilde{C}_{3}\left|\psi_{n}(s)\right|^{\beta}(t-s)^{-6} \quad \forall t>s \geq 0
$$

where we have used the notation $\beta:=\frac{3(\rho-2)}{\rho}>1$ and where $\tilde{C}_{3}=\tilde{C}_{3}\left(\tilde{C}_{1}, \tilde{C}_{2},\|f\|_{L^{r}}\right)$. Both quantity $\beta$ and $\tilde{C}_{3}$ do not depend on $n$. Hence by using Lemma 4.1 in [12] it follows:

$$
\psi_{n}(\theta)=0
$$

where $\theta=2^{\beta /(\beta-1)}\left(\tilde{C}_{3}|\Omega|^{\beta-1}\right)^{1 / 6}<\infty$ does not depend on $n$.
This property tells precisely that (11) holds true with $C_{3}=\theta$.
Notice that the bilinear form

$$
(u, v) \rightarrow \int_{\Omega} a_{n}\left(k_{n}\right) \nabla u \nabla v
$$

is also uniformly coercive on $H_{0}^{1}(\Omega)$. Moreover, the sequence

$$
h_{n}:=T_{n}\left(\nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right)
$$

is imbedded in $L^{\infty}(\Omega)$. We can then apply again the technique of Stampacchia detailed in the proof of lemma 3, and obtain:

$$
\begin{equation*}
\text { for } n \geq 1: \quad k_{n} \in L^{\infty}(\Omega) \tag{19}
\end{equation*}
$$

Nevertheless the control we have on $\left\{h_{n}\right\}$ is obtained from (9), which gives a uniform bound in the $L^{1}$-norm for the sequence. This is not enough to obtain a uniform estimate for $\left\{k_{n}\right\}$ in the $L^{\infty}$-norm.
However we can establish:
Lemma 4 Assume that $\left(H_{0}\right)$ and $\left(H_{1}\right)$ hold. Then we have

$$
\begin{align*}
a_{n}(s) & \geq \gamma_{1} \nu_{n}(s), \quad \gamma_{1}=\min (1, \gamma)  \tag{20}\\
\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2} & \leq C_{5} \tag{21}
\end{align*}
$$

## Proof

The estimate (20) is easy to obtain. Its verification is left to the reader.
Let $\left(u_{n}, k_{n}\right)$ be the chosen approximating sequence. We have from (11) and (19) that

$$
\forall n \geq 1: \quad u_{n}, k_{n} \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega)
$$

It follows (see [3]) that $v:=u_{n} \cdot k_{n} \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega)$ is admissible for $\left(P_{n}\right) .1$ and we get ${ }^{2}$

$$
\begin{equation*}
\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} k_{n}=\int_{\Omega} f u_{n} k_{n}-\int_{\Omega} \nu_{n}\left(k_{n}\right) u_{n} \nabla u_{n} \nabla k_{n} \tag{22}
\end{equation*}
$$

By testing $\left(P_{n}\right) .2$ with $\varphi=k_{n}$, we obtain:

$$
\begin{equation*}
\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}=\int_{\Omega} T_{n}\left(\nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right) k_{n} \leq \int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} k_{n}, \tag{23}
\end{equation*}
$$

by using the properties $T_{n}(s) \leq s$ and (8).
Hence, by combining (22) with (23) we have:

$$
\begin{equation*}
I:=\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2} \leq \underbrace{\int_{\Omega}\left|f u_{n} k_{n}\right|}_{:=I I}+\underbrace{\int_{\Omega}\left|\nu_{n}\left(k_{n}\right) u_{n} \nabla u_{n} \nabla k_{n}\right|}_{:=I I I} \tag{24}
\end{equation*}
$$

We can estimate the term II as follows:
$\overline{2 \text { more generally: } \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}=\underbrace{-\operatorname{div}\left(\nu_{n}\left(k_{n}\right) \nabla u_{n}\right) \cdot u_{n}}_{=f u_{n}}+\operatorname{div}\left(\nu_{n}\left(k_{n}\right) u_{n} \nabla u_{n}\right) \text { in } \mathcal{D}^{\prime}(\Omega), ~(\Omega)}$

$$
\begin{aligned}
& I I \quad \leq \quad C_{3} \int_{\Omega}\left|f k_{n}\right| \stackrel{\text { Hölder Ineq. }}{\leq} C_{3}\|f\|_{L^{3 / 2}}\left\|k_{n}\right\|_{L^{3}} \\
& \stackrel{\text { Poincaré-Sobolev Ineq. }}{\leq} \tilde{C}_{1}\|f\|_{L^{3 / 2}}\left(\int_{\Omega}\left|\nabla k_{n}\right|^{2}\right)^{1 / 2} \leq \frac{\tilde{C}_{1}}{\delta}\|f\|_{L^{3 / 2}}\left(\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}\right)^{1 / 2} \\
& \underset{\leq}{\text { Young Ineq. }} \quad \frac{\tilde{C}_{1}}{\delta}\left(\frac{1}{\epsilon}\|f\|_{L^{3 / 2}}^{2}+\epsilon \int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}\right) \quad \text { for any } \epsilon>0 \text { given } \\
& \leq \quad \frac{1}{3} \int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}+\tilde{C}_{2}\|f\|_{L^{3 / 2}}^{2}
\end{aligned}
$$

where $\delta>0$ is the constant given in $\left(H_{0}\right)$. The last inequality was obtained by choosing $\epsilon=\delta /\left(3 \tilde{C}_{1}\right)$, using the estimate (11) and by setting $\tilde{C}_{2}=3 \tilde{C}_{1}^{2} / \delta^{2}$.

We next estimate the term III:

$$
\begin{aligned}
I I I & =\int_{\Omega}\left|u_{n} \sqrt{\nu_{n}\left(k_{n}\right)} \nabla u_{n} \sqrt{\nu_{n}\left(k_{n}\right)} \nabla k_{n}\right| \\
& \leq \tilde{C}_{3} \int_{\Omega}\left|\sqrt{\nu_{n}\left(k_{n}\right)} \nabla u_{n} \sqrt{a_{n}\left(k_{n}\right)} \nabla k_{n}\right|, \quad \tilde{C}_{3}=C_{3} \gamma_{1}^{-1 / 2} \\
& \leq \frac{1}{3} \int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}+\tilde{C}_{4} \int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}, \quad \tilde{C}_{4}=\tilde{C}_{4}\left(\tilde{C}_{3}\right)
\end{aligned}
$$

where $C_{3}, \gamma_{1}$ are the constants that appear in (11) and (20). The last inequality follows from the Young inequality.
Recall now the inequality (24) and use the estimates established for the terms II and III. We obtain:

$$
\begin{equation*}
\frac{1}{3} \int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2} \leq \tilde{C}_{2}\|f\|_{L^{3 / 2}(\Omega)}^{2}+\tilde{C}_{4} \int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} . \tag{25}
\end{equation*}
$$

By using (25) together with (9) we finally obtain (21).

## 3 Basic convergence results for $\left(u_{n}, k_{n}\right)$

The estimates established in the previous section allow us to extract a converging subsequence from $\left(u_{n}, k_{n}\right)$. We have

## Lemma 5

1. Assume that $\left(H_{0}\right)$ holds. Then we can extract a subsequence (still denoted by $\left.\left(u_{n}, k_{n}\right)\right)$ such that

$$
\begin{array}{rl}
a_{n}\left(k_{n}\right) \nabla k_{n} & \rightharpoonup a(k) \nabla k \text { in } L^{p}(\Omega), p<\frac{3}{2} \\
k_{n} & \rightarrow k \text { a.e in } \Omega \\
u_{n} & \rightharpoonup u \text { in } H_{0}^{1}(\Omega) \\
u_{n} & * \tag{29}
\end{array} u \text { in } L^{\infty}(\Omega)
$$

2. If in addition the condition $\left(H_{1}\right)$ is fulfilled then we may assume that

$$
\begin{equation*}
k_{n} \rightharpoonup k \text { in } H_{0}^{1}(\Omega) \tag{30}
\end{equation*}
$$

## Proof

1. The properties (26) and (27) are obtained from (10). The property (28) is obtained by using the estimate (9) together with the assumption $\nu(s) \geq \delta>0$ in $\left(H_{0}\right)$. We establish (29) from the estimate (11).
2. By using Lemma 4 together with the assumption $a(s) \geq \delta>0$ in $\left(H_{0}\right)$ we obtain (30). Notice that the $k$ appearing in (26), (27) and (30) is necessarily the same in the three situations.

We are able to prove additional regularity results for the element $(u, k)$ introduced in Lemma 5. For technical reasons we introduce the sequence $\left\{h_{q}\right\}_{q \in \mathbb{N}}$ of real functions defined in [10] p. 185. It satisfies:

$$
\begin{align*}
\left|h_{q}(s)\right| & \leq 1 \quad \forall(q, s) \in \mathbb{N} \times \mathbb{R}  \tag{31}\\
h_{q}(s) & =0 \quad \text { when }|s|>2 q  \tag{32}\\
\left|h_{q}^{\prime}(s)\right| & \leq \frac{1}{q} \quad \forall q \in \mathbb{N}, \text { and a.e } s \in \mathbb{R}  \tag{33}\\
h_{q} & \rightarrow 1 \quad \text { uniformly on the compacts } \tag{34}
\end{align*}
$$

## Lemma 6

1. Assume that $\left(H_{0}\right)$ holds. Then the element $(u, k)$ given in Lemma 5 satisfies

$$
\begin{equation*}
\int_{\Omega} \nu(k)|\nabla u|^{2}<\infty \tag{35}
\end{equation*}
$$

2. Assume that in addition $\left(H_{1}\right)$ holds. Then

$$
\begin{equation*}
\int_{\Omega} a(k)|\nabla k|^{2}<\infty \tag{36}
\end{equation*}
$$

## Proof

1. We take over the arguments presented in [10] p. 192.

For $q \geq 1$, we set

$$
\eta_{n, q}:=\left(h_{q}\left(k_{n}\right) \nu_{n}\left(k_{n}\right)\right)^{1 / 2} \nabla u_{n}
$$

Let now $q$ be fixed. The sequence $\left\{\left(h_{q}\left(k_{n}\right) \nu_{n}\left(k_{n}\right)\right)^{1 / 2}\right\}_{n \geq 1}$ is uniformly bounded in $L^{\infty}(\Omega)$. Consequently, $\left\{\eta_{n, q}\right\}_{n \geq 1}$ is bounded in $\left(L^{2}(\Omega)\right)^{3}$ and we can extract a subsequence weakly convergent to some $\eta_{q} \in\left(L^{2}(\Omega)\right)^{3}$.
On the other hand, we have

$$
\begin{aligned}
\left(h_{q}\left(k_{n}\right) \nu_{n}\left(k_{n}\right)\right)^{1 / 2} & \rightarrow\left(h_{q}(k) \nu_{n}(k)\right)^{1 / 2} \quad \text { a.e in } \Omega \\
\nabla u_{n} & \rightharpoonup \nabla u \quad \text { in } L^{2}(\Omega)
\end{aligned}
$$

and thus $\eta_{q}=\left(h_{q}(k) \nu(k)\right)^{1 / 2} \nabla u$.
We now use a classical property of the weak convergence in $L^{2}(\Omega)$ :

$$
\left\|\eta_{q}\right\|_{L^{2}(\Omega)} \leq \liminf _{n \rightarrow \infty}\left\|\eta_{n, q}\right\|_{L^{2}(\Omega)} \leq \liminf _{n \rightarrow \infty}\left(\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right)^{1 / 2} \leq C_{1}^{1 / 2},
$$

where $C_{1}$ is a constant independent of $q$ given in (9).
By using properties (34) and (31) we can see that

$$
\begin{aligned}
& \eta_{q}^{2} \rightarrow \infty \\
& \eta_{q}^{2} \leq \nu(k)|\nabla u|^{2}
\end{aligned}
$$

Hence by the Fatou Lemma we finally obtain:

$$
\int_{\Omega} \nu(k)|\nabla u|^{2} \leq \liminf _{q \rightarrow \infty}\left\|\eta_{q}\right\|_{L^{2}}^{2} \leq C_{1}
$$

2. If the additional assumption $\left(H_{1}\right)$ holds, then we have the estimate (21) and the previous reasoning allows us to obtain (36)

## 4 The proof of theorem 1

In the previous section we have proved that under $\left(H_{0}\right)$ we can extract a converging subsequence of $\left(u_{n}, k_{n}\right)$. If moreover $\left(H_{1}\right)$ holds then the limit $(u, k)$ obtained satisfies:

$$
\begin{align*}
& u \in W_{k} \cap L^{\infty}(\Omega)  \tag{37}\\
& \left.k \in H_{0}^{1}(\Omega) \quad \text { and in fact } k \in W_{k}\right) \tag{38}
\end{align*}
$$

4.1 Passing to the limit in $\left(P_{n}\right) .1$

We recall that the space $W_{k}$ was defined by

$$
W_{k}=\left\{v \in H_{0}^{1}(\Omega):[v]_{k}<\infty\right\}
$$

We now establish:

Lemma 7 Assume that $\left(H_{0}\right)$ and ( $H_{1}$ ) hold. Then the element $(u, k)$ given in Lemma 5 satisfies (37), (38) and:

$$
\begin{equation*}
\int_{\Omega} \nu(k) \nabla u \nabla v=\int_{\Omega} f v \quad \forall v \in W_{k} \tag{39}
\end{equation*}
$$

## Proof

Let $n \geq 1, q \in \mathbb{N}$ and $\varphi \in W_{k} \cap L^{\infty}(\Omega)$. We consider the function $v:=h_{q}\left(k_{n}\right) \varphi$. By recalling the properties (31)-(34) of $h_{q}$, we can verify that $h_{q}\left(k_{n}\right) \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega)$. Consequently $v \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega)$. By testing $\left(P_{n}\right) .1$ with $v$, we obtain:

$$
\begin{equation*}
I:=\int_{\Omega} \nu_{n}\left(k_{n}\right) h_{q}\left(k_{n}\right) \nabla u_{n} \nabla \varphi+\underbrace{\int_{\Omega} h_{q}^{\prime}\left(k_{n}\right) \nu_{n}\left(k_{n}\right) \nabla u_{n} \nabla k_{n} \varphi}_{:=I I}=\underbrace{\int_{\Omega} f h_{q} k_{n} \varphi}_{:=I I I} \tag{40}
\end{equation*}
$$

In a first step we fix $q$ and we study the behaviour of terms I, II and III when $n$ tends to infinity.
By using the property (32) we see that

$$
\left|\nu_{n}\left(k_{n}\right) h_{q}\left(k_{n}\right)\right| \leq \max _{s \in[0,2 q]} \nu(s):=C_{q},
$$

and by using (32) together with (27) we obtain

$$
\nu_{n}\left(k_{n}\right) h_{q}\left(k_{n}\right) \underset{n \rightarrow \infty}{\rightarrow} \nu(k) h_{q}(k) \quad \text { a.e in } \Omega .
$$

Consequently

$$
\nu_{n}\left(k_{n}\right) h_{q}\left(k_{n}\right) \nabla \varphi \underset{n \rightarrow \infty}{\rightarrow} \nu(k) h_{q}(k) \nabla \varphi \quad \text { in }\left(L^{2}(\Omega)\right)^{2},
$$

and by also employing (28) we get:

$$
\begin{equation*}
I \underset{n \rightarrow \infty}{\rightarrow} \int_{\Omega} \nu(k) h_{q}(k) \nabla u \nabla \varphi \tag{41}
\end{equation*}
$$

We now estimate II. From (33) we obtain:

$$
\begin{align*}
I I & \leq \frac{1}{q} \int_{\left\{q \leq k_{n} \leq 2 q\right\}}\left|\nu_{n}\left(k_{n}\right) \nabla u_{n} \nabla k_{n} \varphi\right| \\
& \leq\|\varphi\|_{L^{\infty}} \frac{C}{q}\left(\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right)^{1 / 2}\left(\int_{\Omega} a_{n}\left(k_{n}\right)\left|\nabla k_{n}\right|^{2}\right)^{1 / 2} \leq \frac{C}{q}, \tag{42}
\end{align*}
$$

where the second inequality is obtained by using (20).
For the last term we get

$$
\begin{equation*}
I I I \underset{n \rightarrow \infty}{\rightarrow} \int_{\Omega} f h_{q}(k) \varphi \tag{43}
\end{equation*}
$$

By using the estimates (41)-(43) together with (40) we obtain that for any fixed $\varphi \in W_{k} \cap L^{\infty}(\Omega)$ the following holds

$$
\begin{equation*}
\underbrace{\int_{\Omega} \nu(k) h_{q}(k) \nabla u \nabla \varphi}_{:=J_{1}}=\underbrace{\int_{\Omega} f h_{q}(k) \varphi}_{:=J_{2}}+\mathcal{O}\left(\frac{1}{q}\right) \tag{44}
\end{equation*}
$$

We next remark that the integrand in $J_{1}$ converges for a.e. $x \in \Omega$ to $\nu(k) \nabla u \nabla \varphi$ when $q$ tends to infinity. Moreover by using (31) together with the fact that $\varphi \in W_{k}$ we can see that the integrand in $J_{1}$ is dominated by $|\nu(k) \nabla u \nabla \varphi| \in L^{1}(\Omega)$. Consequently, by the Dominated Convergence Theorem we get

$$
J_{1} \underset{q \rightarrow \infty}{\rightarrow} \int_{\Omega} \nu(k) \nabla u \nabla \varphi
$$

Similarly we can see that

$$
J_{2} \underset{q \rightarrow \infty}{\rightarrow} \int_{\Omega} f \varphi
$$

At this stage we have proved that

$$
\begin{equation*}
\int_{\Omega} \nu(k) \nabla u \nabla \varphi=\int_{\Omega} f \varphi \quad \forall \varphi \in W_{k} \cap L^{\infty}(\Omega) \tag{45}
\end{equation*}
$$

and it remains to show that the condition $\varphi \in L^{\infty}(\Omega)$ is not necessary.
Let $\varphi \in W_{k}$ and $i \in \mathbb{N}$. We consider $\varphi_{i} \in W_{k} \cap L^{\infty}(\Omega)$ given by $\varphi_{i}=T_{i}(\varphi)$. By using some basic properties of $T_{i}$ (see [7]), we see that $\left|\varphi_{i}\right| \leq|\varphi|,\left|\nabla \varphi_{i}\right| \leq|\nabla \varphi|, \varphi_{i} \rightarrow \varphi$ a.e, and $\nabla \varphi_{i} \rightarrow \nabla \varphi$ a.e in $\Omega$. Consequently, if we take $\varphi_{i}$ as test function in (45), we can pass to the limit $i \rightarrow \infty$ and we obtain (39).

In Lemma 7 we have showed that $u$ is a $W_{k}$-solution of ( P$) .1$. In order to prove Theorem 1 we have to prove that $k$ is a distributional solution of ( P ).2. We need first to establish:

Lemma 8 Assume that $\left(H_{0}\right)$ and $\left(H_{1}\right)$ hold. Then, in addition to the results presented in Lemma 5, we may assume:

$$
\begin{equation*}
\nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} \underset{n \rightarrow \infty}{\rightarrow} \nu(k)|\nabla u|^{2} \quad \text { in } L^{1}(\Omega) \tag{46}
\end{equation*}
$$

## Proof

We test $\left(P_{n}\right) .1$ with the function $u_{n}$. By using (28) we obtain:

$$
\begin{equation*}
\int_{\Omega} \nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2} \underset{n \rightarrow \infty}{\rightarrow} \int_{\Omega} f u=\int_{\Omega} \nu(k)|\nabla u|^{2} \tag{47}
\end{equation*}
$$

where the latter equality is obtained by testing (39) with $u$.
We set $\eta_{n}:=\sqrt{\nu_{n}\left(k_{n}\right)} \nabla u_{n}$ and $\eta:=\sqrt{\nu(k)} \nabla u$. The relation (47) tells us that

$$
\begin{equation*}
\left\|\eta_{n}\right\|_{L^{2}(\Omega)}^{\longrightarrow}\|\eta\|_{L^{2}(\Omega)} \tag{48}
\end{equation*}
$$

We can next take over the arguments presented in [10] Lemma 5.3.4 in order to obtain:

$$
\begin{equation*}
\eta_{n} \underset{n \rightarrow \infty}{\rightharpoonup} \eta \quad \text { in }\left(L^{2}(\Omega)\right)^{2} \tag{49}
\end{equation*}
$$

Finally properties (49) and (48) imply that the convergence is strong in (49), and (46) follows.

### 4.2 The Proofs of Theorem 1 and Corollary 1

Assume that $\left(H_{0}\right)$ and $\left(H_{1}\right)$ hold. In Lemma 5 we have extracted a subsequence $\left(u_{n}, k_{n}\right)$ which converges in a certain sense to an element $(u, k)$. This element has the properties (37)-(38). Next we have established (39).
Let now $\varphi \in \mathcal{C}_{c}^{\infty}(\Omega)$. By using (26) we get:

$$
\begin{equation*}
\int_{\Omega} a_{n}\left(k_{n}\right) \nabla k_{n} \nabla \varphi \underset{n \rightarrow \infty}{\rightarrow} \int_{\Omega} a(k) \nabla k \nabla \varphi \tag{50}
\end{equation*}
$$

We next remark that the property (46) ensures that

$$
\begin{equation*}
\int_{\Omega} T_{n}\left(\nu_{n}\left(k_{n}\right)\left|\nabla u_{n}\right|^{2}\right) \varphi \underset{n \rightarrow \infty}{\rightarrow} \int_{\Omega} \nu(k)|\nabla u|^{2} \varphi \tag{51}
\end{equation*}
$$

Recall that the sequence $\left(u_{n}, k_{n}\right)$ satisfies $\left(P_{n}\right) .2$. Then relation (50) together with (51) allows to take the limit in $\left(P_{n}\right) \cdot 2$. We get:

$$
\begin{equation*}
\int_{\Omega} a(k) \nabla k \nabla \varphi=\int_{\Omega} \nu(k)|\nabla u|^{2} \varphi \quad \forall \varphi \in \mathcal{C}_{c}^{\infty}(\Omega) \tag{52}
\end{equation*}
$$

Thus (P). 2 is fulfilled in the distributional sense.
At this point we have obtained (37), (38), (39) and (52). The proof of Theorem 1 is complete.

Assume now that the condition (3) in Corollary 1 is fulfilled. By using (38) together with the Sobolev Injection Theorem we get $k \in L^{6}(\Omega)$ and thus $\nu(k) \in L^{1}(\Omega)$. Then we can conclude the proof of Corollary 1 by using Proposition 9 in the Appendix I: $(u, k)$ is a distributional solution of $(\mathrm{P})$.

## 5 The proof of Theorem 2

We assume in this section that $\left(H_{0}\right)$ and $\left(H_{2}\right)$ hold.
In this situation all the results presented in section 2 and section 3 are valid. For technical reasons we slightly modify the definition of $a_{n}$ by setting

$$
\begin{equation*}
a_{n}(s):=\gamma \nu_{n}(s) \tag{53}
\end{equation*}
$$

where $\gamma>0$ is the constant appearing in $\left(H_{2}\right)$ and $\nu_{n}$ is defined as before. We will now consider Problems $\left(P_{n}\right)$ modified by the new definition (53) of $a_{n}$. Nevertheless the modification is very sligth, and all the results presented in the previous section can be recovered easly. The verifications are left to the reader.

We now prove that we have the new estimate:

$$
\begin{equation*}
\left\|k_{n}\right\|_{L^{\infty}(\Omega)} \leq C_{6} \tag{54}
\end{equation*}
$$

In order to prove this result we set

$$
\begin{equation*}
\chi_{n}:=k_{n}+\frac{\gamma}{2} u_{n}^{2}, \tag{55}
\end{equation*}
$$

and we remark that $\left(P_{n}\right) .2$ leads to

$$
\int_{\Omega} a_{n}\left(k_{n}\right) \nabla \chi_{n} \nabla \varphi=\int_{\Omega} f u_{n} \varphi \quad \forall \varphi \in H_{0}^{1}(\Omega)
$$

Recall that $a_{n}\left(k_{n}\right) \geq \gamma \min (1, \delta)>0, a_{n}\left(k_{n}\right) \in L^{\infty}(\Omega)$ and note that the sequence $f u_{n}$ is uniformly bounded in $L^{r}(\Omega)$ with $r>3 / 2$. These properties are sufficient (see the proof of Lemma 3) to get the estimate

$$
\begin{equation*}
\left\|\chi_{n}\right\|_{L^{\infty}(\Omega)} \leq C \tag{56}
\end{equation*}
$$

where $C$ does not depend on $n$.
The estimate (54) is finally obtained by using Lemma 3 together with (56).
Consequently, in addition to the properties in Lemma 5 we may assume that

$$
\begin{equation*}
k_{n} \stackrel{*}{\rightharpoonup} k \quad \text { in } L^{\infty}(\Omega) . \tag{57}
\end{equation*}
$$

We will now prove that

$$
\begin{equation*}
u, k \in \mathcal{C}^{0, \alpha}(\bar{\Omega}) \quad \text { for some } \alpha \in(0,1) \tag{58}
\end{equation*}
$$

Let $\lambda:=\nu(k)$. We have $\lambda, \lambda^{-1} \in L^{\infty}(\Omega)$ and

$$
\begin{equation*}
\int_{\Omega} \lambda \nabla u \nabla \phi=\int_{\Omega} f \phi \quad \forall \phi \in H_{0}^{1}(\Omega) \tag{59}
\end{equation*}
$$

Recall also that $f$ have the property (12). Hence we can apply the De Giorgi-Nash Theorem (see for instance [5] Prop. 6 p. 683 or [8] Th. 8.22 and Th. 8.29). We obtain that $u \in \mathcal{C}^{0, \alpha_{1}}(\bar{\Omega})$ for some $\alpha_{1} \in(0,1)$. We next set $\chi:=k+(\gamma / 2) u^{2}$. Then $\chi \in H_{0}^{1}(\Omega)$ and we have

$$
\begin{equation*}
\int_{\Omega} \frac{\lambda}{\gamma} \nabla \chi \nabla \phi=\int_{\Omega} f u \phi \quad \forall \phi \in H_{0}^{1}(\Omega) \tag{60}
\end{equation*}
$$

By using the fact that $u \in L^{\infty}(\Omega)$ in (60), we can again apply the De Giorgi-Nash Theorem to get $\chi \in \mathcal{C}^{0, \alpha_{2}}(\bar{\Omega})$ for some $\alpha_{2} \in(0,1)$. Hence also $k$ is Hölder continuous, and (58) follows.

Let $\alpha \in(0,1)$ be a generic parameter that can differ from one part to another. We assume now that $\partial \Omega$ is of class $\mathcal{C}^{2, \alpha}, f \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$ and $\nu \in \mathcal{C}^{1, \alpha}\left(\mathbb{R}^{+}\right)$.
We will prove the second part of Theorem 2 by iterating the Schauder estimates.
We have $\lambda=\nu(k) \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$ (see the Appendix II) and then, by applying the Schauder estimate (see [4] Theorem 2.7 p. 154) on (59) we get $u \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$. Similarily, from equation (60) we obtain $\chi \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$ and thus $k \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$.
Hence (see Appendix II) $\lambda \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$. By iterating again the Schauder estimates (see now Theorem 2.8 p. 154 in [4]) we obtain that $u$ and $k$ are in $\mathcal{C}^{2, \alpha}(\bar{\Omega})$.
Finally we see that $(u, k)$ is a classical solution of $(\mathrm{P})$. Theorem 2 is proven.

## Appendix I: Some relations between the notions of weak solution

We give here some relations between the various notions of weak solution: Wsolution, $H$-solution, distributional solution, renormalized solution, 'energy solution'.

Comparison with renormalized solution

We have:

## Proposition 9

1. Any $W$ - or $H$-solution $(u, k)$ of Problem ( $P$ ) that satisfies in addition $k \in H_{0}^{1}(\Omega)$, is also a renormalized solution.
2. If $\nu(k) \in L^{1}(\Omega)$ then any $W$ - or $H$-solution of Problem $(P)$ is a also a distributional solution of $(P)$.

## Proof

1. Let $(u, k)$ be a $W$-solution of (P). Then the conditions (5.2.1)-(5.2.5) in [10] chap. 5 are satisfied. We have to prove that (5.2.6) holds.
Let $h \in \mathcal{C}_{c}^{\infty}(\mathbb{R})$ and $\phi \in \mathcal{C}_{c}^{\infty}(\Omega)$ be arbitrarily chosen. We set $v:=h(k) \phi$. Then $v \in L^{\infty}(\Omega)$ and $\nabla v=h(k) \nabla \phi+h^{\prime}(k) \nabla k \phi$. Let $M<\infty$ be such that the support of $h$ being included in $[-M, M]$. We have

$$
\begin{gathered}
\int_{\Omega} \nu(k) h^{2}(k)|\nabla \phi|^{2} \leq \max _{[0, M]} \nu\|h\|_{L^{\infty}}^{2} \int_{\Omega}|\nabla \phi|^{2}<\infty \\
\int_{\Omega} \nu(k)\left(h^{\prime}(k)\right)^{2}|\nabla k|^{2}|\phi|^{2} \leq \max _{[0, M]} \nu\left\|h^{\prime}\right\|_{L^{\infty}}^{2}\|\phi\|_{L^{\infty}}^{2} \int_{\Omega}|\nabla k|^{2}<\infty
\end{gathered}
$$

Hence $v \in W_{k}$. By testing (39) with $v$ we obtain the relation (5.2.6). a in [10].
We remark that $v$ is also admissible in (52). This allows us to obtain the condition (5.2.6).b in [10]. Consequently $(u, k)$ is a renormalized solution of ( P ).

If we consider a $H$-solution $(u, k)$ of $(\mathrm{P})$ we can take over the previous argument because the function $v$ is now in $H_{k}$.
2. If $\nu(k) \in L^{1}(\Omega)$ then we have $\mathcal{C}_{c}^{\infty}(\Omega) \hookrightarrow H_{k} \hookrightarrow W_{k}$. In consequence a $W_{k^{-}}$or a $H_{k}$-solution of $(\mathrm{P}) .1$ is also a distributional solution of this equation. Hence $(h, k)$ is a distributional solution of $(\mathrm{P})$.

## Remarks

1. The first point in Proposition 9 tells that the notions of $H$ - or $W$-solution are stronger that the notion of renormalized solution. This fact is coherent with the second point established in Proposition 9: a $H$ - or $W$-solution is a distributional solution if $\nu(k) \in L^{1}(\Omega)$ whereas a renormalized solution is only a priori a distributional solution if $\nu(k) \in L^{\infty}(\Omega)$ (see [10] p.185).
2. if we have $k \in H_{0}^{1}(\Omega)$ and if $\nu$ satisfies the growth condition (3) then $\nu(k) \in L^{1}(\Omega)$.

Comparison with 'energy solution'

We have seen that when $\nu(k) \in L^{1}(\Omega)$ then any $W$ - (or $H$-) solution is a distributional solution. Moreover the notion of $W$-solution coincides with the notion of $H$-solution iff $W_{k}=H_{k}$ (see [14]).

Some sufficient conditions to have this last equality were established in [14] and in [7], but necessary and sufficient conditions are not known.

Let us consider the following condition:

$$
(R)\left\{\begin{array}{l}
\sqrt{\nu(k)} \in H^{1}(\Omega) \\
T_{n}(k) \in H_{0}^{1}(\Omega), \quad \forall n \in \mathbb{N}
\end{array}\right.
$$

It was schown in [7] that the first condition in $(R)$ together with the property $\nu^{-1} \in L^{\infty}(\mathbb{R})$ (which is assumed in $\left(H_{0}\right)$ ) implies that $W_{k}=H_{k}$.

In [7] the authors introduced the notion of 'energy solution'. They impose $\left(H_{0}\right)$ as the basic assumption. Then an 'energy solution' $(u, k)$ for $(\mathrm{P})$ is in fact a $W$-solution which satisfies $(R)$. This implies that $W_{k}=H_{k}$. The energy solution is also a $H$ solution, and moreover a distributional solution (because the first assumption in ( $R$ ) implies that $\left.\nu(k) \in L^{1}(\Omega)\right)$.

We see then that the notion of 'energy solution' (in the sense of [7]) has the advantage of unifying various notions by putting us in the situation where $\sqrt{\nu(k)} \in H^{1}(\Omega)$. The disadvantage is that we have to impose more complicated conditions on the coefficients $a$ and $\nu$, in order to obtain a solution. In particular in [7] Theorem 2.1, the authors prove the existence of an 'energy solution' under the assumptions ( $H_{0}$ ) and $\left(H_{3}\right)$ (see below).

$$
\left(H_{3}\right)\left\{\begin{array}{l}
\nu \in \mathcal{C}^{1}\left(\mathbb{R}^{+}\right) \\
\exists C>0 \text { and } \gamma>1 / 2 \text { such that: } \\
\left|\nu^{\prime}(s)\right| \leq C \quad \forall s \in[0,1] \\
\frac{\left|\nu^{\prime}(s)\right|}{\sqrt{a(s) \nu(s)}} \leq C . s^{-\gamma} \quad \forall s \geq 1
\end{array}\right.
$$

This condition is not verified in the physical situation $\left(H_{p}\right)$, but only in the approximate situation $\left(H_{p}^{\prime}\right)$.

In Theorem 1 we obtain a $W$-solution under much simpler conditions which are satisfied by $\left(H_{p}\right)$. This solution is a distributional solution under an additionnal simple
assumption (see Corollary 1) which is again satisfied in $\left(H_{p}\right)$.

Note also that in the first part of Theorem 2 we prove that under the assumptions $\left(H_{0}\right)$ and $\left(H_{2}\right)$ (which are satisfied in $\left(H_{p}\right)$ if $a 1 \nu_{2}=a_{2} \nu 1$ ), the functions $u$ and $k$ are Hölder continuous. In particular $\nu(k) \in L^{\infty}$ which implies that $W_{k}=H_{k}$, and the notions of $H$-solution, $W$-solution, distributional solution and renormalized solution coincide in this case.

In order to conclude this Appendix we give a last existence result. Let $\left(H_{4}\right)$ be the following condition:

$$
\left(H_{4}\right)\left\{\begin{array}{l}
\nu \in \mathcal{C}^{1}\left(\mathbb{R}^{+}\right) \\
\exists C>0 \text { s.t. } \frac{\left|\nu^{\prime}(s)\right|}{\nu(s)} \leq C \quad \forall s \in \mathbb{R} .
\end{array}\right.
$$

We have:
Proposition 10 Assume that $\left(H_{0}\right),\left(H_{1}\right)$ and $\left(H_{4}\right)$ hold. Then the $W$-solution given in Theorem 1 is an 'energy solution' (in the sense of [7]).

## Proof

We have assumed that $\left(H_{0}\right),\left(H_{1}\right)$ hold and consequently all the results presented in the sections 2, 3 and 4 can be recovered.

Let $(u, k)$ be the $W$-solution given by Theorem 1. By using (2) we see that the second condition in (R) is satisfied. Nevertheless we cannot directly conclude that $\sqrt{\nu(k)} \in H^{1}(\Omega)$, but we can obtain a new estimate for the approximating sequence $\left(u_{n}, k_{n}\right)$. More precisely, we have:

$$
\begin{equation*}
\left\|\sqrt{\nu_{n}\left(k_{n}\right)}\right\|_{H^{1}(\Omega)} \leq C \tag{61}
\end{equation*}
$$

In fact, by using the property that $k_{n} \in H_{0}^{1}(\Omega) \cap L^{\infty}(\Omega)$ together with $\nu \in \mathcal{C}^{1}\left(\mathbb{R}^{+}\right)$ we obtain $\nu\left(k_{n}\right) \in H^{1}(\Omega) \cap L^{\infty}(\Omega)$, with $\nabla \nu\left(k_{n}\right)=\nu^{\prime}\left(k_{n}\right) \nabla k_{n}$. Recall now that $\nu_{n}\left(k_{n}\right)=T_{n}\left(\nu\left(k_{n}\right)\right)$. Hence we have

$$
\nabla \nu_{n}\left(k_{n}\right)=1_{\left\{\nu_{n}\left(k_{n}\right)<n\right\}} \nu^{\prime}\left(k_{n}\right) \nabla k_{n} .
$$

It follows that:

$$
\begin{aligned}
\nabla \sqrt{\nu_{n}\left(k_{n}\right)} & =1_{\left\{\nu\left(k_{n}\right)<n\right\}} \frac{\nu^{\prime}\left(k_{n}\right) \nabla k_{n}}{2 \sqrt{\nu_{n}\left(k_{n}\right)}}=1_{\left\{\nu_{n}\left(k_{n}\right)<n\right\}} \frac{\nu^{\prime}\left(k_{n}\right)}{2 \sqrt{\nu_{n}\left(k_{n}\right) a_{n}\left(k_{n}\right)}} \sqrt{a_{n}\left(k_{n}\right)} \nabla k_{n} \\
\quad \text { by }(20) & \leq 1_{\left\{\nu_{n}\left(k_{n}\right)<n\right\}} \frac{\nu^{\prime}\left(k_{n}\right)}{\nu_{n}\left(k_{n}\right)} \sqrt{a_{n}\left(k_{n}\right)} \nabla k_{n}=C \frac{\nu^{\prime}\left(k_{n}\right)}{\nu\left(k_{n}\right)} \sqrt{a_{n}\left(k_{n}\right)} \nabla k_{n} .
\end{aligned}
$$

Hence, by using (21) we obtain

$$
\left\|\nabla \sqrt{\nu_{n}\left(k_{n}\right)}\right\|_{L^{2}(\Omega)} \leq C
$$

Moreover $\sqrt{\nu_{n}\left(k_{n}\right)}=\sqrt{\nu(0)}$ on $\partial \Omega$ and thus we obtain (63) by using a Poincaré inequality.

## Remark

The hypotheses made in Proposition 10 are verified under assumption $\left(H_{p}^{\prime}\right)$. In the hypotheses, we require only very weak growth condition at infinity for $\nu$. For instance (contrarily to the result presented in [7]) the Proposition 10 works if we have:

$$
\nu(s)=\nu_{1}+\nu_{2} e^{\beta_{1} s}, \quad a(s)=a_{1}+a_{2} e^{\beta_{2} s}, \quad \beta_{1} \leq \beta_{2}
$$

## Appendix II: Hölder continuity and composition

Let $\Lambda \subset \mathbb{R}^{d}$ and $\alpha \in(0,1)$. We recall that the space $\mathcal{C}^{0, \alpha}(\Lambda)$ of Hölder continuous (with exponent $\alpha$ ) functions on $\Lambda$ is defined by:

$$
\mathcal{C}^{0, \alpha}(\Lambda)=\left\{f: \Lambda \rightarrow \mathbb{R} \quad \text { s.t. } \forall x_{0} \in \Lambda: \sup _{x \in \Lambda} \frac{\left|f(x)-f\left(x_{0}\right)\right|}{\left|x-x_{0}\right|^{\alpha}}<\infty\right\}
$$

More generally, for any integer $k$, the space $\mathcal{C}^{k, \alpha}(\Lambda)$ is the space of those $f \in \mathcal{C}^{k}(\Lambda)$ whose kth derivative is in $\mathcal{C}^{0, \alpha}(\Lambda)$.

A first elementary result tells that the product of two Hölder continuous functions is an Hölder continuous function. More precisely we have (see relation (4.7) in [8]):
Lemma 11 Assume that $f_{1}, f_{2} \in \mathcal{C}^{0, \alpha}(\Lambda)$. Then $f_{1} \cdot f_{2} \in \mathcal{C}^{0, \alpha}(\Lambda)$
In Section 5 we used a function defined as a composition of two Hölder continuous functions. We needed the following result:

Lemma 12 Let $\bar{\Omega}$ be a compact in $\mathbb{R}^{d}$ and $\alpha \in(0,1)$. We consider the following three conditions:
(A) $\quad \lambda \in \mathcal{C}^{1}(\mathbb{R})$ and $k \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$
(B) $\quad \lambda \in \mathcal{C}^{0, \alpha}(\mathbb{R})$ and $k \in \mathcal{C}^{1}(\bar{\Omega})$
(C) $\lambda \in \mathcal{C}^{1, \alpha}(\mathbb{R})$ and $k \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$

We have:

1. Assume that $(A)$ or $(B)$ is satisfyed. Then $\lambda(k) \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$.
2. Assume that $(C)$ is satisfyed. Then $\lambda(k) \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$.

## Proof

1. In this situation we clearly have $\lambda(k) \in \mathcal{C}^{0}(\bar{\Omega})$ and

$$
\begin{equation*}
M_{1}:=\sup _{x \in \bar{\Omega}}|k(x)|<\infty \tag{62}
\end{equation*}
$$

Let

$$
I\left(x, x_{0}\right):=\frac{\left|\lambda(k(x))-\lambda\left(k\left(x_{0}\right)\right)\right|}{\left|x-x_{0}\right|^{\alpha}} .
$$

We want to prove that

$$
\begin{equation*}
\sup _{x, x_{0} \in \bar{\Omega}} I\left(x, x_{0}\right)<\infty . \tag{63}
\end{equation*}
$$

- Assume that (A) holds. Then in addition of (62) we have:

$$
M_{2}:=\sup _{t, t_{0} \in\left[-M_{1}, M_{1}\right]} \frac{\left|\lambda(t)-\lambda\left(t_{0}\right)\right|}{\left|t-t_{0}\right|}<\infty \quad \text { and } \quad M_{3}:=\sup _{x, x_{0} \in \bar{\Omega}} \frac{\left|k(x)-k\left(x_{0}\right)\right|}{\left|x-x_{0}\right|^{\alpha}}<\infty .
$$

Consequently:

$$
I\left(x, x_{0}\right) \leq M_{2} \frac{\left|k(x)-k\left(x_{0}\right)\right|}{\left|x-x_{0}\right|^{\alpha}} \leq M_{2} \cdot M_{3}
$$

Hence (63) is satisfyed.

- Assume now that (B) holds. Then in addition of (62) we have:

$$
M_{4}:=\sup _{x, x_{0} \in \bar{\Omega}} \frac{\left|k(x)-k\left(x_{0}\right)\right|}{\left|x-x_{0}\right|}<\infty \quad \text { and } \quad M_{5}:=\sup _{t, t_{0} \in\left[-M_{1}, M_{1}\right]} \frac{\left|\lambda(t)-\lambda\left(t_{0}\right)\right|}{\left|t-t_{0}\right|^{\alpha}}<\infty
$$

In this situation we can estimate $I\left(x, x_{0}\right)$ as follows:

$$
I\left(x, x_{0}\right) \leq \frac{\left|\lambda(k(x))-\lambda\left(k\left(x_{0}\right)\right)\right|}{\left|k(x)-k\left(x_{0}\right)\right|^{\alpha}} \cdot \frac{\left|k(x)-k\left(x_{0}\right)\right|^{\alpha}}{\left|x-x_{0}\right|^{\alpha}} \leq M_{5} M_{4}^{\alpha} .
$$

Hence (63) is again satisfyed.
2. Assume that (C) holds and let $\mu:=\lambda(k)$. Clearly $\mu \in \mathcal{C}^{1}(\bar{\Omega})$ and $\nabla \mu=\lambda^{\prime}(k) \nabla k$. We remark that $\lambda^{\prime} \in \mathcal{C}^{0, \alpha}(\mathbb{R})$ and $k \in \mathcal{C}^{1, \alpha}(\bar{\Omega})$. We can then apply the first point of this lemma to obtain: $\lambda^{\prime}(k) \in \mathcal{C}^{0, \alpha}(\bar{\Omega})$. Moreover $\nabla k \in\left(\mathcal{C}^{0, \alpha}(\bar{\Omega})\right)^{d}$. Hence the product $\lambda^{\prime}(k) \nabla k$ is Hölder continuous.

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[^1]:    1 We thank Michel Chipot for this remark

