

### Project of 1 DoF Attitude Control System of 1U Cubesat Based On Reaction Wheel

### Thesis

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#### Abstract

The following project is undertaken within **PLATHON** (*PLATform of Optical communications in Nanosatellites*) project in **DISEN** and **TIEG** Research Group lead by Dr. David González and Dr. Javier Gago in the ESEIAAT (Escola Superior d'Enginyeries Industrial, Audiovisual i Aeronàutica de Terrassa) School of the Technical School of Catalonia - BarcelonaTech.

PLATHON's main mission is to simulate a Networks System of CubeSats that collect information and communicate with other different orbital group satellites using IoT sensors to retrieve the information towards a ground station minimizing delay and maximizing global coverage. Final year degree and master students from several departments are the main contributors to the project and most of the system components are designed and built by students.

The project's overview encompasses an analytical estimation of the worst-case disturbance torques calculations for the mission. Moreover, a detailed control algorithm is developed for fine pointing and coarse pointing modes in the orbit as well as a 3D real-time monitoring interface. Finally, the design of the Reaction Wheel (RW) is also made following the technical requirements set by the project specifications.

The final stage of the project focuses on the testing and simulation of the implemented control algorithms. To perform the simulation, the CubeSat is introduced in an Air bearing in the centre of a magnetic simulator, which will generate a magnetic field similar to the conditions that the satellite will be subjected in a Low Earth Orbit (LEO). The Onboard Computer (OBC) of the nanosatellite communicates via Bluetooth with the ground computer waiting for commands. The ground station has full access to the satellite's attitude and full control over the distinct modes of the satellite (i.e. Detumbling mode, Pointing mode, Normal mode, among others). Once the magnetic field and other sources of disturbances are set, the reaction wheels and magnetorquers will activate to control the CubeSat's attitude. Analogously, this attitude data measured by the Inertial Measurement Unit (IMU) of the Attitude Determination and Control Subsystem (ADCS) is sent back to the ground station's computer and further displayed with a computer-generated 3D model in real-time. The CubeSat is intended to be powered with both solar panels or a LiPo battery. Running a software-in-the-loop and a hardware-in-the-loop test has shown that the system requires some modifications to achieve more precise results.

#### Resumen

El siguiente proyecto se realiza dentro de **PLATHON** (*Plataforma de comunicaciones ópticas en nanosatélites*) del proyecto **DISEN** y del Grupo de Investigación **TIEG** liderado por el Dr. David González y el Dr. Javier Gago en la ESEIAAT (Escola Superior d'Enginyeries Industrial, Audiovisual i Aeronàutica de Terrassa) de la Escuela Politécnica de Cataluña - BarcelonaTech.

La misión principal de PLATHON es simular un Sistema de redes de CubeSats que recoja información y se comunique con otros satélites de diferentes grupos orbitales utilizando sensores IoT para recuperar la información hacia una estación terrestre minimizando el retraso y maximizando la cobertura global. Estudiantes de último curso de grado y máster de varios departamentos son los principales colaboradores del proyecto y la mayoría de los componentes del sistema están diseñados y construidos por los estudiantes.

La siguiente tesis comprende una estimación analítica de los cálculos de los pares de perturbación en el peor de los casos para la misión. Además, se desarrolla un algoritmo de control detallado para los modos de apuntamiento fino y grueso en la órbita, así como una interfaz de monitorización 3D en tiempo real. Por último, el diseño de la Rueda de Reacción (RW) también se realiza siguiendo los requisitos técnicos establecidos por las especificaciones del proyecto.

La etapa final del proyecto se centra en la realización de pruebas y la simulaciones de los algoritmos de control implementados. Para realizar la simulación, se introduce el CubeSat en un cojinete de aire en el centro de un simulador magnético, que generará un campo magnético similar a las condiciones a las que estará sometido el satélite en una órbita baja terrestre (LEO). El ordenador de a bordo (OBC) del nanosatélite se comunica vía Bluetooth con el ordenador de tierra a la espera de órdenes (en este caso, el ordenador central del laboratorio). La estación de tierra tiene acceso total a la actitud del satélite y control total sobre los distintos modos del satélite (es decir, modo de Detumbling, modo de apuntamiento, modo normal, entre otros). Una vez fijado el campo magnético y otras fuentes de perturbación, las ruedas de reacción y los magnetorquers se activarán para controlar la actitud del CubeSat. Análogamente, estos datos de actitud medidos por la Unidad de Medición Inercial (IMU) del Subsistema de Determinación y Control de Actitud (ADCS) se envían de vuelta al ordenador de la estación de tierra y se visualizan además con un modelo 3D generado por ordenador en tiempo real. El CubeSat está pensado para ser alimentado tanto con paneles solares como con una batería LiPo. La realización de una prueba de software-in-the-loop ha demostrado que el sistema requiere algunas modificaciones para lograr resultados más precisos.

#### 摘要

以下项目是在 **DISEN**和**TIEG**研究组领导的**PLATHON** (*PLATform of Optical communications in Nanosatellites*) 项目中进行的。项目中进行的,该项目由David González博士和Javier Gago博士领导的 加泰罗尼亚技术学校的ESEIAAT (Escola Superior d'Enginyeries Industrial, Audiovisual i Aeronàutica de Terrassa) 学院的研究小组负责。

PLATHON的主要任务是模拟一个立方体卫星的网络系统,该系统使用IoT传感器收集信息并与其他不同的轨道组卫星进行通信,将信息检索到地面站,最大限度地减少延迟并最大化全球覆盖。来自几个系的毕业班学生和硕士生是这个项目的主要贡献者,大部分的系统部件都是由学生设计和建造的。

该项目的概述包括对任务中最坏情况下的干扰扭矩计算的分析估计。此外,还为轨道上的精细指向和粗略指向模式开发了一个详细的控制算法,以及一个三维实时监控界面。最后,反应轮(RW)的设计也是按照项目规范所规定的技术要求进行的。

项目的最后阶段主要是对已实现的控制算法进行测试和仿真。为了进行模拟,CubeSat被引入到一个磁 性模拟器中心的空气轴承中,这将产生一个类似于卫星在低地球轨道(LEO)上所受条件的磁场。纳米 卫星的机载计算机(OBC)通过蓝牙与地面计算机进行通信,等待指令。地面站可以完全了解卫星的姿 态,并完全控制卫星的不同模式(即脱离模式、指向模式、正常模式等)。一旦磁场和其他干扰源被设 定,反应轮和磁轮就会启动,以控制CubeSat的姿态。类似地,由姿态确定和控制子系统(IMU)的惯 性测量单元(ADCS)测量的这些姿态数据被送回地面站的计算机,并进一步用计算机生成的三维模型 实时显示。CubeSat打算用太阳能电池板或LiPo电池供电。运行软件在环和硬件在环的测试表明,该系 统需要进行一些修改,以获得更精确的结果。

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# **Declaration on Honour**

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# Nomenclature

#### Acronyms

- 1U 1-Unit CubeSat
- 2U 2-Unit CubeSat
- 3U 3-Unit CubeSat
- AD0 Address pin
- ADCS Attitude Determination and Control Subsystem
- ADC Analog to Digital Converter
- ADC Attitude Determine and Control
- AOCS Attitude and Orbit Control Systems
- C&DH Command and Data Handling
- CAC CubeSat Acceptance Checklist
- CCD Charged Coupled Devices
- CCW Counterclockwise
- CDH Command & Data Handling
- CFS Core Flight System
- COM Center of Mass
- COM Centre of Mass
- COTS Off-The-Shelf Components
- CPU Central Processing Unit
- CW Clockwise
- DAC Digital to Analog Converter
- DISEN Distributed Sensor Network
- DMP Digital Motion Processor

#### NOMENCLATURE

DOF	Degree of Freedom
ECEF	Earth-Centered Earth Fixed (ECEF)
ECI	Earth-Centered Inertial
FCC	Federal Communication Comission
FHSS	Frequency-Hopping Spread-Spectrum
GNC	Guidance, Navigation & Control
GND	Ground
HEO	High Earth Orbit
HiTL	Hardware-in-the-Loop
I/O	Input / Output
$I^2C$	Inter-Integrated Circuit
IMU	Inertial Measurement Unit
IoT	Internet of Things
ISM	Industrial, Scientific and Medical
ISR	Interrupt Service Routine
LEO	Low Earth Orbit
LV	Launch Vehicle
MCU	MicroController Unit
MCU	Microcontroller Unit
MEO	Medium Earth Orbit
MISO	Master In Slave Out
MOSE	ET Metal–Oxide Silicon Transistor
MOSI	Master Out Slave In
N/A	Not Applicable
NASA	National Aeronautics and Space Administration
NCS	Chip Select
NOAA	National Oceanic and Atmospheric Administration
OBC	Onboard Computer

- OSK OpenSatKit
- $\mathbf{P}-\mathbf{POD}$ Poly Picosatellite Orbital Deployer

- PCB Printed Circuit Board
- PC Personal Computer
- PLATHON PLATform of Optical communications in Nanosatellites
- PPOD Poly-Picosatellite Orbital Deployer
- PSoC Programmable System on Chip
- RBF Remove Before Flight
- RCS Reaction Control System
- RF Radio Frequency
- RWA Reaction Wheel Assembly (RW + motor)
- RW Reaction Wheel
- RXD Receive data
- SCL Serial Clock
- SDA Serial Data
- SiTL Software-in-the-Loop
- SPI Serial Peripheral Interface
- TIEG Terrassa Industrial Electronics Group
- TTC Tracking, Telemetry & Command
- TXD Transmit data
- UART Universal Asynchronous Receiver-Transmitter
- UART Universal Asynchronous Receiver/Transmitter
- UHF Ultra High Frequency
- Vin Input voltage

#### **Physics Constants**

$\mu_{\mathrm{Earth}}$	Earth's gravitational constant	$3.9860 \cdot 10^{1}4 \text{ m}^{3}/\text{s}^{2}$
с	Speed of light in a vacuum inertial system	299792458  m/s
G	Gravitational Constant	$6.67408 \cdot 10^{-3} \text{ Nm}^3/\text{kgs}^2$
$M_{\rm E}$	Earth's mass	$5.9742 \cdot 10^{24} \text{ kg}$
$R_{\rm Earth}$	Earth's mean radius	$6371.000 \cdot 10^3 \text{ m}$

#### Variables and Parameters

- Angle between the incident ray and  $\hat{n}$  $\alpha$ Differential of mass  $\mathrm{d}m$  $\mathrm{d}t$ Change in time  $\dot{\Omega}$ Angular acceleration  $\dot{r}$ Velocity of the spacecraft Normal vector representing the position of the north geomagnetic pole  $\hat{m}_{\rm north}$ Unit vector normal to the incident surface  $dA_{inc}$  $\hat{n}$  $\hat{s}$ Unit vector perpendicular to  $\hat{n}$  co-planar to the incident ray Gravitational parameter μ Ω Angular velocity Angular velocity ω Angular velocity of the motor  $\omega_{\rm motor}$ Reaction Wheels angular velocity  $\omega_{\mathrm{RW}}$  $\phi$ Pitch angle Pitch angular velocity  $\phi$ Yaw angle ψ Yaw angular velocity ψ Density ρ Normal Exchange coefficient  $\sigma_n$ Momentum Exchange coefficient  $\sigma_t$  $au_D$ Drag torque  $\tau_{\rm ext,dist}\,$ External disturbances torque Satellite's torque  $\tau_{\rm sat}$  $\theta$ Roll angle  $\theta$ Roll angular velocity Maximum deviation angle of the Z axis and therefore is maximum at  $\theta_{dev} = 45^{\circ}$  in body frame  $\theta_{\rm dev}$ from its local vertical Gravity gradient torque  $\vec{\tau}_{\rm gg}$
- $\vec{\tau}_{\rm m}$  Magnetic torque
- $\vec{a}$  Acceleration
- $\vec{B}_{\text{Earth}}$  Earth's magnetic field

$\vec{F}$	Force
$\vec{p}$	Impulse
$\vec{r}$	Position of the satellite's COM with respect to Earth's COM
$\vec{r}^B$	Position of the satellite in ECEF reference frame
$A_{\rm inc}$	Incident satellite surface area
$A_{\rm inc}$	Incident surface area
$C_d$	Drag coefficient
$c_{rd}$	Coefficient of specular diffusion
$c_{rs}$	Coefficient of diffuse reflection
D	Drag
e	Eccentricity
$F_{\rm SRP}$	Solar Radiation force
h	Altitude above the surface of Earth
h	Height
$h_{\rm disk}$	Disk height
$H_{\rm G}$	Angular momentum computed from its center of mass
$h_{\mathrm{ring}}$	Ring height
$H_{\rm RW}$	Angular momentum generated by the Reaction Wheel
$I_{\rm disk}$	Reaction Wheel's disk inertia
$I_{\rm ring}$	Reaction Wheel's ring inertia
$I_{\rm RW}$	Reaction Wheel's Inertia
$I_{\rm RW}$	Reaction Wheels inertia
$I_{\rm sat,max}$	Satellite's maximum moment of inertia matrix
$I_{\rm sat,min}$	Satellite's minimum moment of inertia matrix
$I_{\rm sat}$	Satellite's moment of inertia matrix
$I_{\rm SRP}$	Energy per unit of time through a cross sectional unit area
$K_d$	Derivative gain constant
$K_i$	Integral gain constant
$K_p$	Proportional gain constant
m	Mass

#### NOMENCLATURE

$M_{\rm Earth}$	Satellite's local magnetic induction
$m_{\rm rw,dis}$	$_{\rm k}$ Reaction Wheel's disk mass
$m_{\rm rw,rin}$	$_{\rm g}$ Reaction Wheel's ring mass
$m_{\rm rw,tot}$	al Reaction Wheel's total mass (disk and
$m_{ m RW,tc}$	tal Reaction Wheel's total mass
$n_{\mathrm{d}A}$	Normal vector of surface $dA$
0	Origin of axis variable
q	Quaternion
R	Rotation matrix
$R_{\rm Earth}^3$	$H_0$ Strength of geomagnetic moment
$r_{\mathrm{d}A}$	Location of the centroid
$r_{\rm cg}$	Drag torque
$r_{\rm disk}$	Disk radius
$r_{\rm hollow}$	Reaction Wheel hollow radius
$r_{\rm ring}$	Ring radius
$T_{\rm RW}$	Reaction Wheels distribution matrix
u(t)	PID control variable
V	Velocity
$v^T$	Spacecraft tangential velocity
x	x variable
y	y variable
z	z variable

ring)

### Chapter 1

### Introduction

The present Bachelor Thesis covers the study of the attitude determination and control system (ADCS) of a future Cubesat, which is the main goal of TIEG, the Electronics Aerospace Research Group of the Polytechnical University of Catalonia.

Cubesats are tiny satellites made up of one or more  $10 \times 10 \times 10$  cm cubes. Although reaching Earth orbit may seem like an unattainable achievement, numerous universities and related organizations and institutions have already designed, constructed and launched a CubeSat into orbit.

The CubeSat that is being developed consists of a 1-Unit Cubesat printed using additive manufacturing method with 2 electromagnetic coils (magnetorquers) embedded as actuators that will interact with Earth's geomagnetic field reaction wheel. The symbiotic relation between the two actuators will be responsible for controlling the satellite's attitude and orientation in space. This Cubesat also carries a set of communication modules to interact with the ground station. The main matter is the study of rigid solid mechanics, which will provide the mathematical equations that control the satellite's motion and the interaction with Earth's magnetic field and other disturbances. The CubeSat is intended to be powered using either solar panels or a LiPo battery.

Ultimately, the control algorithm and several simulations are performed for the different pointing modes. The thesis embraces classical Newtonian mechanics, programming several control algorithms, sizing and calibrating sensors and actuators, PCB and CAD designs and assembling all the work done within the rest of the department members. The result of the work is a fully operative automatic controlled CubeSat with the ability to send and receive telemetry data.

### 1.1 Aim of the project

The following thesis aims to study and design both the hardware and the control software of the Attitude Determination and Control Subsystem of a 1DOF 1U Cubesat using reaction wheels. This project is part of **PLATHON** (*PLATform of Optical communications in Nanosatellites*) Research Group the goal of which is to establish and simulate communications in a constellation of satellites.

A 1 degree of freedom (1DOF) control of a Cubesat means there is only one system variable that is unbound (free). Degrees of freedom are used instead of variables to clarify the freedom of a system instead of a specific number of variables.

For instance, suppose a 2D grid plane. To get the exact location of any particle at any location of the plane there always must be 2 coordinates. For instance, using Cartesian coordinates the position of the particle is specified by (x, y), or analogously, using polar coordinates the position of the particle is also determined using two variables  $(r, \beta)$ , where r is the radius and  $\beta$  is the angle. Thus, there are  $x, y, r, \beta$  multiple variables that can define the position but only using two of them the system's position is fully defined.

The final goal is to send a telecommand with the desired rotation to the CubeSat and the CubeSat's response shall turn into the desired angle and the correct direction. To do so, two main pointing modes were developed, a coarse pointing mode which will trigger the motor driver to make an impulse and then, when the CubeSat is approaching the desired set-point, fine pointing mode comes in with a PD controller (see Section 7).

### 1.2 Justification

#### 1.2.1 PLATHON Project

The successful development of missions with nanosatellite clusters requires tools that integrate orbital propagation, command and telemetry transmission between the nanosatellite and the ground station and the operational status of the satellites. For missions where nanosatellites are communication nodes, it is also necessary to integrate the communication network model.

The applications that constitute the mission object of the satellite can be validated by simulating the behaviour of all the hardware that surrounds it and emulating the flight conditions (position, inclination, velocity, etc.). This form of software validation is called Software-in-the-Loop (SiTL).

A more advanced step to ensure mission success is to include testing of the vehicle hardware and its correct use by the applications developing the mission. This is covered when the actual hardware can be incorporated into the application development and validation environment. This is known as HiTL (Hardware-in-the-Loop).

The **PLATHON** (*PLATform of Optical communications in Nanosatellites*) Research Group system aims to meet these needs by developing a HiTL (Hardware-in-the-Loop) platform integrated with a network and orbital simulator that allows the verification of the communication network of a nanosatellite cluster considering the operational constraints of the nodes that form it. Specifically, the aim is to address the current needs for the design of low orbit missions that could establish a communications network between satellites in the cluster itself and links with ground stations and with other satellites in other orbits.

All development will be based on the OpenSatKit software developed in open-source by NASA for the simulation of satellite missions. The software will be adapted to the inclusion of SiTL for simulations in the first instance, and subsequently to emulations with HiTL. For this purpose, several real CubeSat

prototypes (equipped with power, communications and ADCS subsystems) will be placed on air-bearing platforms. In this way, the operation of the ADCS system is emulated in conditions as close as possible to an actual flight. These satellite models will communicate with each other via a real UHF link and an optical communication link.

CubeSats appeared as a teaching tool in universities during the '90s but immediately shifted from university labs to commercial applications and business. At present, they have become a mature standard in the aerospace industry. Recently the first CubeSat based constellations intended to offer global communication coverage have appeared. However, we consider that there are some lacks in two areas related to such constellations based missions:

- Simulation tools. Current communication networks simulators used by the scientific community (NS2, SNS3 and Opensand) do not have the features required to design communications among nodes in placed LEO, between LEO and GEO and all above mentioned and ground stations as well. Neither the power availability of nodes to establish proper communications are considered.
- Reliability: CubeSat missions still have a large fail rate (55% in CubeSat from Universities and 33% from Companies). Most of these failures (40%) are attributed to a malfunction of the Electric Power Systems (EPS).



Figure 1.1 PLATHON real architecture. Source: PLATHON [1].

PLATHON project proposes to develop an integrated simulation platform intended as a test bench for validation of engineering models before flight and network communication architectures. This platform considers the following aspects:

- 1. Integration of an orbital propagator and the restrictions imposed by the visibility among network nodes placed in different orbits (LEO, GEO) and the power availability on each node to establish a reliable optical communication.
- 2. Design and manufacturing of an emulator (HiL) to test engineering models of CubeSat. This emulator is a CubeSat with EPS, ADCS and COMS, which will be connected to the integrated simulator above mentioned.



Figure 1.2 PLATHON diagram. Source: PLATHON [1].

This new tests facility will be a helpful assessment tool in the design of more reliable constellation missions.



Figure 1.3 PLATHON hardware diagram. Source: PLATHON [1].

Hitherto, the common communication with satellites was using a Ground Station with a specific radiofrequency band to communicate with the satellite. Nevertheless, there are some limitations as the satellite is only visible to the Ground Station in a certain period (4-6 min) in general). The major problem that arises was that a huge amount of information needed to be transmitted in an acutely reduced amount of time. Thus, PLATHON project is to use communication between satellites to broaden the effective communication area. At the same time, the operative state of the satellite can be increased as part of the constellation of satellites will always be in the sunlight region.



Figure 1.4 PLATHON Open Sat Kit diagram. Source: PLATHON [1].

#### Step by Step Operation of the PLATHON system

We can synthesise a PLATHON emulation with the following steps:

- 1. Creating the orbits of the satellites in 42 from TLE data or manually.
- 2. Designing the mission and translating it into communication links to be carried out at each instant or point of the orbit to bring the data from the original point (e.g. the IoT in our example) to the end (the corresponding GS).
- 3. Apply NS3 to establish the optimal routes at each instant or point of the orbit.
- 4. Design the Command and Telemetry messages to be sent between GS and satellite each time they have visibility. Only one GS is selected as a tracking station.
- 5. There is one satellite simulated as SiTL in Computer 1, and N hardware satellites as HiTL.
- 6. The communication of 42 and COSMOS with the CFS of the SiTL satellite in Computer 1 is via sockets or software communication ports, as all the software is on the same machine.

- 7. If 42 manages the satellite attitude control, the attitude and position data is passed from 42 to CFS through the TCP socket to the I42 application.
- 8. If it is the CFS that performs the attitude control, this data is provided by the F42 application.
- 9. The NS3 provides it, synchronised with the 42, with the information on the attitude setpoint to be followed at any given moment. Given that we have previously planned the mission, we know at each instant how we have to rotate the antenna or the optical link to focus on another satellite that will be the next node of the communication following the optimal route. This information is sent through the UDP socket to the NS3 App of the CFS.
- 10. The NS3, once it has estimated that it has already made the turn, tells it which message to send via the antenna or the optical link. It is assumed that the same previous operation has been performed on the receiving satellite, and therefore, they are oriented, and when sending the message, it must be received correctly on the receiving satellite.
- 11. Each time the tracking station is passed, COSMOS activates the Command Ingest application, to send command messages, and Telemetry Output, to receive the status of the satellite's sensors. This is managed through the scheduler table.
- 12. The communication of the 42 and COSMOS with the CFS of the HiTL satellite of each satellite is by means of sockets or communication ports through the IP Network with Wi-Fi access points.
- 13. The 42 sends the position data (GPS) via Wi-Fi to the satellite's Wi-Fi through the IP network. And the Wi-Fi injects it through the I2C bus to the GPS application that transforms it into the CFS format for the satellite's CFS bus software.
- 14. The NS3 provides it, synchronised with the 42, with the information on the attitude setpoint to be followed at each instant and the message (MSG) to be sent, via Wi-Fi and IP network. The satellite's Wi-Fi will inject the attitude setpoint data through the I2C bus to the ADCS application, and the message (MSG) through the SPI or UART bus to the NS3 application. These applications transform the data format and inject it into the software bus.
- 15. The ADCS board detects and changes the attitude of the satellite according to the data received, orienting it to the satellite, IoT or GS that corresponds to it at that moment according to the design of the mission.
- 16. The corresponding data is sent by the COMS board, either in RF or by optical link and is received by the receiving satellite, which will also have carried out the attitude setpoint operations to be oriented to the transmitter. Or they are received by the GS. Or they are transmitted from the IoT, whose hardware is also an SDR.
- 17. Communication with COSMOS to send the information to or from the tracking station is also done through CFS applications, the Command Ingest or Telemetry Output of the satellite that sends the data by UART or SPI bus to the Wi-Fi or COMS boards.
- 18. The Command Ingest or Telemetry Output data is sent to Wi-Fi when there is no SDR hardware connected to Computer 1. In this case, the data is sent via IP Network to the Wi-Fi board of Computer 1.
- 19. If the SDR is connected via USB to Computer 1, the Command Ingest or Telemetry Output data is sent via the COMS RF antenna to the SDR RF antenna.

- 20. The Command Ingest or Telemetry Output data arriving either by Wi-Fi or by SDR are injected into COSMOS through the corresponding Targets.
- 21. Finally, the position data of the 42 (GPS) and satellite attitude (ADCS) are sent via IP network to Computer 2 to be displayed on the Render 42 in real-time.

### 1.3 Scope of the project

The scope of the project encompasses the design and integration of the full assembly inside the 1U Cubesat. All hardware used must be Open Source.

The reaction wheel model must consider all different constraints as the satellite will be subjected to many perturbations and the goal is to obtain the most realistic approach to its attitude control while performing its mission in orbit.

### 1.4 Project Requirements

- The reaction wheel must not exceed the size of a 1U CubeSat  $10 \times 10 \times 10$  cm.
- Every electronic component must fit into the ADCS PCB which carries the microcontroller, motor drivers, a motor and two motor drivers, a Bluetooth module and two magnetorquers.
- The microcontroller must have enough pins to work with several peripherals at the same time and numerous communication interfaces.
- The reaction wheels design must meet the weight constraints set by 1U CubeSat California University requirements [2].
- The code must be a Commercial-Off-The-Shelf (COTS) cod, meaning it shall be used without any customization.
# Chapter 2

# State of the art

## 2.1 What is a Cubesat?

Cubesats are small-sized satellites based on a standardized unit of mass and volume. The most basic Cubesat unit has  $10 \times 10 \times 10$  cm, conforming to specific interfaces for allowing a standardized containerized launch and had a maximum mass of 1 kg (the mass was later on increased to 1.33 kg) [3].

The purpose of the Cubesats is to provide a standard for the design of nanosatellites to reduce cost and development time, increase accessibility to space, and sustain frequent launches. Additionally, thanks to their size and weight, basic Cubesats units could be combined for supplying larger objectives for major missions while adhering to the same constraints and requirements. Therefore, a 3-Unit Cubesat consists of 3 different standard 1-Unit Cubesat stacked together [4].



(a) Basic 1U Cubesat structure. Source: Cubesatworld [5].



(b) CELESTA Cubesat. Source: CERN [6].

Figure 2.1 Cubesat examples. Source: Cubesatworld [5] and [6].

Cal Poly and Stanford University developed the Poly-Picosatellite Orbital Deployer (P-POD) launch

#### CHAPTER 2. STATE OF THE ART

dispenser standard to provide a low-cost solution to develop and safely launch CubeSats. Since Cubesats are vastly light-weighted satellites, the overall mission costs are reduced exponentially as the amount of fuel needed is significantly minor. Besides, Cubesats have a handful of useful applications such as remote sensing or communications, but as engineers are beginning to get used to the technology, CubeSats are beginning to venture farther afield.

The CubeSat is planned to be modular, with each module or unit measuring  $100 \times 100 \times 100$  mm, weighing approximately 1 kg for a 1U CubeSat, and having a functional capacity of 1 L. Its modularity enables relatively flexible scaling, allowing it to accommodate the required components for a specific application in several connected units with little effort or additional cost [2].

Meteorological research, atmospheric analysis, high-energy particle research, disaster monitoring, spacecraft damage research, and spacecraft attitude control systems are among the current experiments conducted in low-earth orbit. In point of fact, because these smaller satellites have the same essential components as their bigger counterparts (communication systems, processing units, and solar panels), the only limitations imposed on t are related to experiments that have lower power requirements.

At the same time, in 1998, at a joint meeting of students between the United States and Japan, Dr, Jordi Puig proposed the idea of launching a spaceship the size of soda water (Coca-Cola) into space. This original concept would subsequently evolve into other nanosatellite projects and influence the CubeSat concept. At the time, the plan was to take a Coke can, add some electronics in it, mount it in a high-powered amateur rocket, launch it, and eject it on a parachute. Professor Shinici Nakazuka and his students at the University of Tokyo began to develop CanSat [7].



Figure 2.2 Cubesat sizes. Source: Alen Space [8].



Figure 2.3 Cubesat components. Source: Alen Space [8].

Nanosatellites are typically launched into low circular or elliptical orbits (altitudes ranging from 400 to 650 km) and travels at around 8 km/s. At this altitude and height, it takes them around 90 minutes to orbit the Earth, with 14 to 16 orbits completed every day. These are great configurations for nanosatellites. They are better shielded from solar and cosmic radiation by circling closer to the Earth [8].

Satellites move in circular or elliptical orbits around the Earth because of the balance of gravitational and escape pull at launch. Because air friction is very small, they can stay in orbit for a long operating time. When a nanosatellite's operational life comes to an end, it re-enters the atmosphere and disintegrates.

#### 2.1.1 Satellite Constellations

Usually, nanosatellites work in constellations in which they give support, redundancy, and granularity to the services they deliver. Each satellite in a constellation is replaced every 4-5 years, ensuring that the operator always has an optimized, low-risk service that gets continual technical advances. As a result, nanosatellite constellations are systems in which the ideas of obsolescence and useful life are no longer relevant [8].



Figure 2.4 Concept of a satellite constellation. Source: Space News [9].

#### 2.1.2 CubeSat Applications

The key driving factors behind the worldwide adoption and growth of CubeSats can be approximately abbreviated as follows [7]:

- 1. Affordability: The intention is to create an instructional tool for a research group and university students. With these platforms, students can develop, design, launch, operate satellites to further research. As simplicity and profitability is key, cost-effectiveness factors are achieved by creating a simplified design, using recommended affordable COTS (Off-The-Shelf Components) and accepted specifications and requirements that streamlined various stages of the development cycle such as deployment, structural design, and some verification requirements.
- 2. Standardization: As the technology becomes more popular design configurations arrived over time. Several engineering solutions and industry standards, such as the PPOD mechanism of common development of electronic devices, were table alternatives and industry standards under the limitations of the design specification restrictions. COTS products for Cubesats enable mass manufacture, and mass production components of whole parts and subsystems, as well as lowering development costs for developers.
- 3. **High-Return value**: Over time, CubeSats become high-return investments for space applications. This is what allows the CubeSat to perform such a wide range of functions in such a wide range of mission circumstances, thanks to a diverse set of developers and customers.
- 4. **Technology**: New technologies are implemented day after day in a range of missions such as advancements in material research, manufacture, energy, software algorithms, mission architectures (both segments of space and terrestrial segments) and communications.

## 2.2 Satellite Subsystems

Alike any other satellite, CubeSats incorporates numerous subsystems such as Structures, Attitude Determination and Control Subsystem, Power Subsystem, among others. For instance, the On-Board Computer (OBC) belongs to the communication subsystem. NASA provides a multi-facet platform Core Flight System (CFS) that can serve as a command and control platform using 42 Simulator.

The principal subsystems are listed below:

Subsystem	Principal Function	Other denominations
Propulsion	Provides thrust to adjust orbit and attitude, and to manage angular momentum	Reaction Control System (RCS)
Attitude Determination & Control System (ADCS)	Provides determination and control of attitude and orbit position, plus pointing of spacecraft and appendages	Guidance, Navigation & Control (GNC) Attitude and Orbit Control Systems (AOCS)
Communication	Communicates with ground and other spacecraft Spacecraft tracking	Tracking, Telemetry & Command (TTC)
Command & Data Handling (CDH)	Processes and distributes commands Processes, stores and formats data	Spacecraft Computer System Spacecraft Processor
Thermal	Maintains equipment and         Thermal       structure within         environmental allowed temperature ranges	
Structures and Mechanisms	Provides support structure, booster adapter, and moving parts	Structure Subsystem

Table	<b>2.1</b>	Principal	subsystems.	Source:	[10]	l
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Regarding dry-mass allocation [10],

- Payload = 28% of total satellite mass.
- Power = 23% of total satellite mass.
- Structure = 18% of total satellite mass.
- Propulsion = 12% of total satellite mass.

#### 2.2.1 Attitude Determination and Control Subsystem

Attitude determination is the process of combining available sensor inputs with spacecraft dynamics information to offer an accurate and unique answer for the attitude state as a function of time, either

on-board for immediate use or after the fact (i.e. post-processing). Most attitude algorithms that were previously conducted as post-processing may now be implemented as on-board computations, due to the powerful microprocessors available for spaceflight. As a result, while there are still compelling engineering reasons for certain procedures to be carried out only by ground-based attitude systems [11].

To sum up, the ADCS initials yields for [12]:

#### • Attitude

- The orientation of a spacecraft in space.
- The angular orientation of the spacecraft body frame with respect to an external frame.
- Describes the rotational motion of the body of the spacecraft about its center of mass.

#### • Determination

- The process of computing the orientation of the spacecraft relative to either an inertial reference or some object of interest, involving:
  - \* Several types of sensors on each spacecraft
  - \* Sophisticated data processing procedures
- Control
  - Attitude stabilization: the process of maintaining an existing orientation.
  - Attitude maneuver control: the process of controlling the reorientation of the spacecraft from one attitude to another.

#### 2.2.2 Attitude determination

Reviewing Wertz's documentation [12], there are different instruments that aids in the process of attitude determination.

#### 2.2.2.1 Sun Sensors

Sun sensors are visible-light or infrared detectors that measure one or two angles between their mounting base and incident sunlight. These devices have a thin entry slit at the top surface which is sensitive to light. When the Sun's light rays enter the thin cabin on the base of the chamber, it illuminates a sequence of photo-cell detectors. Eventually, the digital output can be mapped onto a unique entry angle, allowing the plane in which the Sun lies to be determined. Sun sensor systems are designed to give two fundamental output signal types, analogue and digital, even though hardware designs vary greatly. The output of an analogue system is a continuous function of the angle of incidence, whereas the output of a digital system is discrete [13]. Sun sensors can be quite accurate (< 0.01 deg), but it is not always possible to take advantage of that feature [11] since most low earth orbits include eclipse periods.



Figure 2.5 Analog and digital sun sensors. Source: NASA [13].

#### 2.2.2.2 Magnetometers

Magnetometers are simple, reliable, lightweight sensors that detect the amount and direction of the Earth's magnetic field. The magnetometer output assists in determining the spacecraft's attitude relative to the local magnetic field. A flux-gate magnetometer measures the strength of the magnetic field in one direction. Thereby, adding a 3-axis magnetometer, which consists of three orthogonal magnetometers will permit measurements of the three components of the magnetic field in the spacecraft body frame. Subsequently, the attitude is determined by comparing the measured field to an inertial magnetic model held in the onboard processor [14] [11].

#### 2.2.2.3 Earth Horizon Sensors

Earth Horizon sensors are infrared devices that measure the difference in temperature between deep space and the Earth's atmosphere (approximately 40 km above the surface in the detected range). Pippers (simple narrow field-of-view fixed-head) are used to detect the Earth phase. These devices are static sensors that can provide only roll and pitch information. Their operating principle consists in detecting Infra-Red (IR) radiation in a given range, for instance 14 to 16  $\mu$ m (the emission band of the CO<sub>2</sub> CO<sub>2</sub> molecule) [15] [11].

A mirror (or prism) focuses a narrow pencil of light onto a sensing element. Afterwards, the mirror

spins as a result of spinning the spacecraft. As the mirror or prism rotates its field of view sweeps out a cone. Then, electronics in the sensor detect when the infrared signal from the Earth is first received or finally lost during each sweep of the scan cone. Finally, the time between the arrival and loss of signal determines the Earth width.



Figure 2.6 Earth horizon sensor. Source: NASA [15].

#### 2.2.2.4 Star sensors/trackers

Star sensors are the most accurate reference sensors for measuring attitude. The light from a star strikes a light-sensitive surface. The point of impact on the surface is determined using Charged Coupled Devices (CCDs), similar to the optical element in a video camera. Then, using a star catalogue, internal processing calculates a three-axis attitude. Many systems can calculate an extremely precise attitude within seconds of turning on [11].



Figure 2.7 Star sensor. Source: ESA [16].

#### 2.2.2.5 Gyroscopes

Gyroscopes are any instrument that uses a rapidly spinning mass to sense and responds to changes in the inertial orientation of its spin axis. These inertial sensors measure the speed or angle of rotation from an initial reference but lack knowledge of an external, absolute reference [11].

#### 2.2.2.6 Inertial Measurement Unit

Inertial Measurement Units are devices that measure a spacecraft's accelerations within a stable frame of reference provided by key ingredients: gyroscopes and accelerometers.

The gyroscope's role is to detect instantaneous motions that might disrupt the spacecraft's stability and respond with compensating drive signals to spin the gimbals properly. The accelerometers' duty is to monitor acceleration forces operating on the spacecraft – and so changes in spacecraft direction and position – within a gyro-stabilized frame of reference. Because gyroscopes and accelerometers are mechanical instruments, the gyros may drift a tad over lengthy missions [17].



Figure 2.8 Apollo Inertial Measurement Unit. Source: We Hack the Moon [17].

#### 2.2.3 Attitude Control

Once the attitude is determined by the aforementioned instruments and databases, the next step is to properly adjust the orientation of the spacecraft to stabilize or point towards the desired direction. Here is where attitude control comes into play. The following set of instruments helps the Cubesat to perform orbit manoeuvres and each of them has different torques as the goal may be a fine or coarse pointing.

#### 2.2.3.1 Magnetorquers

Magnetorquers are often used as actuation actuators in spacecraft. These torques are magnetic coils (electromagnets) that produce magnetic dipole moments of a specified magnitude. The magnetic field generated by the spacecraft interacts with the local field from a Planet, producing an external torque on the vehicle. Then electromagnets may be used to provide a controllable external torque. Their strength can be controlled by employing current. When three orthogonal magnetorquers are attached to a spacecraft, they may generate a magnetic dipole in any direction and magnitude up to the magnetorquers' strength. Magnetic torquers can adjust for residual magnetic fields or attitude drift caused by modest disturbance torques on a spacecraft. Besides, they can also be used to desaturate momentum-exchange systems. However, they require much more run time than other actuators since their torque is quite small and requires longer run-times [11].

#### 2.2.3.2 Reaction Wheels

Reaction wheels are essentially devices that provide attitude control and stability on a spacecraft. They allow the spacecraft to rotate quickly towards the desired location, in other words, they are torque motors with high inertia rotors. By adding or removing energy from the flywheel, torque is applied to a single axis of the spacecraft, causing it to react by rotating. These wheels can spin at different speeds using electric motors. If the wheel turns clockwise direction, according to Newton's Third Law, the CubeSat will rotate in the opposite direction (counterclockwise). By maintaining flywheel rotation, known as momentum, a single axis of the spacecraft is stabilized. Several reaction/momentum wheels can be used to provide full three-axis attitude control and stability [11].

One major problem appears when slowing down the wheels, reaction wheels cannot spin indefinitely and if they slow down, the spacecraft will turn in the opposite direction.

Let's define the differences between reaction wheels and momentum wheels to avoid future confusion with the terminology. Both reaction and momentum wheels are flywheels that spin and creates torque by changing their momentum. However, there is a major dissimilarity between the two actuators:

- Reaction wheels are spun to create the torque and hence, force the spacecraft to rotate.
- Momentum wheels, on the other hand, are constantly rotating at very high speed to provide a nearly constant angular momentum which creates a stabilization of the spacecraft, thus, generating an angular momentum by creating a resistance in the change of the spacecraft's attitude.

#### 2.2.3.3 Thrusters

Because of its dual utility in altering orbital parameters, thrusters (i.e. rocket engines) are arguably the most commonly flown attitude actuator. Almost every spacecraft that has to conduct orbital manoeuvres will utilize thrusters to do so, and in many cases, part of the thrusters utilized will be for attitude control. Thrusters exert force on the spaceship by ejecting a high-velocity material known as propellant from its exit nozzles [11].

Thrusters produce torque in proportion to their moment torque, which is the distance by which the direction of the force is offset from the vehicle's centre of mass. So, while a thruster may generate a lot of force, the torque it can produce is limited by the actual size of the vehicle and how the thrusters are placed on it. Thrusters have the advantage of being able to provide large instantaneous manoeuvres at any time in orbit. Notwithstanding, the other issue with thrusters is that the amount of stored gas is not unlimited.

# Chapter 3

# Mathematical description of the physical system

The following section intends to summarize the mathematical model expressions of the basic astrodynamics equations and the development of it.

## 3.1 Orbit Classification

The preceding section's characterization of orbital elements allows for the establishment of an orbital taxonomy. As a matter of fact, orbits are classified based on their altitude (relative to the semi-major axis a), eccentricity e, and inclination i. Low Earth orbits (LEO) are those with a maximum distance from the Earth's centre of mass of less than 2000 km; medium Earth orbits (MEO) are those between 2000 and 20,000 km; and high Earth orbits (HEO) are those with a maximum distance more than 20,000 km [7].

Closed and open orbits are differentiated by eccentricity (see Table 3.1): circular (e = 0) and elliptical (e > 1) orbits are closed trajectories, whereas parabolic (e = 1) and hyperbolic (e > 1) orbits are open paths that can escape the gravitational attraction of the celestial body. Finally, orbits can be classified as equatorial (i = 0 degrees) or polar (i = 90 degrees) based on their inclination.

Table 3.1	$\operatorname{Conic}$	$\operatorname{sections}$	of an	orbit.	Source:	Own.

Eccentricity	Conic section	Open/closed trajectory
e = 0	Circle	Closed
0 < e < 1	Ellipse	Closed
e = 1	Parabola	Open
e > 1	Hyperbola	Closed

Some of the most important orbits are produced by combining some of the properties mentioned be-

forehand. For instance, geosynchronous orbits have a duration equal to one sidereal day. Geostationary orbits are zero-inclination geosynchronous circular orbits. Taking the point of view of an observer on Earth, a geostationary satellite seems fixed in the sky, which is why these orbits are utilized for communications missions, since they are always in the same position and communication can be performed without problems. CubeSats are unlikely to be launched into geostationary orbits, but the majority of them have been launched into a different type of synchronous orbit known as the Sun-synchronous orbit, which is nearly a polar orbit.

Notice how in a Sun-synchronous orbit, the orbital plane rotates to maintain a consistent orientation concerning the Sun throughout the year ( $\dot{\Omega} = 360$  degrees per year. The spaceship will travel over the Earth at the same local solar time at all times of the year (points 1, 2, and 3). This orbit is extremely useful for remote sensing missions since Earth's surface are hugely susceptible to variations in lighting conditions.



Figure 3.1 Orientation of the orbital plane in a Sun-Synchronous orbit. Source: Cubesat Handbook [7].

For low orbits, orbits with inclinations around 90 degrees are the most accessible for CubeSats. As a result, a satellite in a low Sun-synchronous orbit traverses the whole Earth's surface multiple times per day, providing several possibilities for interaction with ground stations. This is critical, especially if the spacecraft can only be operated by a single station, as is the situation with many academic missions. The Earth is always in the same local solar time zone and, as mentioned before, this is critical for distant sensing missions since the Earth's surface is incomparable lighting conditions at each passing.

## 3.2 Attitude description

First of all, before beginning to work on the actual equations a reference frame must be defined. This is crucial since the actual parameters depend on the reference frames. In both reference systems and attitude representation methods and rotation matrices are to be well defined to present accurately the satellite's dynamic equations.

#### 3.2.1 Reference frames

Flight mechanics use a series of specific reference systems with the aim of projecting the positions, rotations linear and angular velocities, linear and angular accelerations, forces and torques in them. These systems can be represented in a generic way as F(O, x, y, z) with its origin in (O) and 3 perpendicular axis (x, y, z) [18] [19] [20].

#### 3.2.1.1 Earth-Centered Inertial (ECI) Reference Frame

The ECI  $F(O_{eci}, x_{eci}, y_{eci}, z_{eci})$  is a fixed in the spatial reference frame, this means that there is not any kind of acceleration so Newton's second is applicable <sup>1</sup>.

- The origin of this reference frame is located at the centre of the Earth (the geo-centre of Earth)
- The z-axis points towards the North Pole, specifically, towards Epoch J2000 which is near the Pole Star.
- The x-axis points towards the Vernal Equinox at J2000 (where equatorial and ecliptic planes intersect).
- The y-axis is orthogonal to both of the prior axis following the right-hand rule.

Note the Equatorial plane and Ecliptic plane are inclined at  $23^{\circ} 27'$  to each other.



# Figure 3.2 ECI reference frame with $x_{ECEF}$ pointing towards the mean Equinox. Source: Own.

<sup>&</sup>lt;sup>1</sup>Newton's laws apply only in "inertial" non-rotating systems that are not rotating or non-accelerating concerning the distant galaxies or accelerating. All inertial reference frames can be seen as being in a state of constant velocity or rectilinear (straight-line) motion concerning one another. Consequently, an accelerometer moving with any inertial frame of reference should always detect zero acceleration.

#### 3.2.1.2 Earth-Centered Earth Fixed (ECEF) Reference Frame

The ECEF reference frame ( $F(O_{\text{ecef}}, x_{\text{ecef}}, z_{\text{ecef}})$ ) is linked to Earth itself in a sense that, in a generic instant, the inertial system rotates with the same angular velocity as the rotation of the Earth around its axis with an angular velocity of  $\omega = 7.27 \cdot 10^{-5}$  rad/s [21].

- Just as the ECI, the ECEF's origin is also located at the center of the Earth.
- The z-axis points towards the North Pole, specifically, towards Epoch J2000 which is near the Pole Star.
- The *x*-axis is contained in the intersection between the Greenwich Meridian and the Equator plane.
- The y-axis is orthogonal to both of the prior axis following the right hand rule.



**Figure 3.3** ECEF reference frame with the *x*-axis pointing to the fixed point on the Prime Meridian. Source: Own.

#### 3.2.1.3 Orbit Reference Frame

This reference frame is also known as the satellite's reference frame. This coordinate system is very useful for satellites that have to perform terrestrial imaging manoeuvres.

- The origin of coordinates corresponds to the Centre of Mass (COM) of the satellite itself.
- The z-axis points towards the centre of Earth.
- The x-axis follows the satellite's velocity vector in the same direction.
- The y-axis is orthogonal to both of the prior axis following the right-hand rule.

At any given time, the Satellite Reference Frame and Satellite Body Frame will be at the same location, i.e., the origin of the Satellite Body Frame will always coincide with that of the Satellite Reference Frame,

however, the Satellite Reference Frame may not be aligned with Satellite Body Frame. The ADCS of a satellite works to align the Satellite Body Frame with the Satellite Reference Frame. To simplify our test cases and to understand ADCS easily, we will be considering a satellite whose reference frame is the Orbit Reference Frame.



Figure 3.4 Orbit Reference Frame. Source: Own.

#### 3.2.1.4 Body Reference Frame

Unlike the Orbit reference frame, the Body reference frame is fixed to the satellite. The orientation in space is defined with respect to the Orbit frame using *Euler angles* thus representing the satellite's attitude.

- The origin of coordinates correspond to the Centre of Mass (COM) of the satellite itself.
- The x y z axis points along the principal directions of the spacecraft, normal to the satellites surfaces forming an orthogonal Cartesian coordinate system.

Though, depending on the mission, these axis can be set according to each mission's specifications. Each mission shall bear in mind to adapt the dynamics equations in consonant of the selected reference frame. Eventually, by doing this will only affect the representation of the motion but the entire dynamics of the system remain the same no matter which reference frame it is used.

This reference frame will be used for acquiring data for the CubeSat with the incorporated IMU. Ans the orientation angles will be referenced to this particular frame.



Figure 3.5 Body Reference Frame. Source: Own.

#### 3.2.2 Euler Angles

So far, in the latter section was explained how the orientation of the Body frame with respect to the satellite's Orbit frame describes its attitude. This is expressed in Euler angles. According to *Euler's rotation theorem*, any rotation may be described using three angles. If the rotations are written in terms of rotation matrices D, C, and B, then a general rotation A can be written as

$$\mathbf{A} = R_B^O = R_{x,\psi} R_{x',\theta} R_{z',\phi} = \mathrm{BCD}$$
(3.2.1)

These three rotations matrices provides three different angles called *Euler angles*. The convention used to define the Euler angles is illustrated below. Let A be

$$A = R_B^O = \begin{bmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ a_{31} & a_{32} & a_{33} \end{bmatrix}$$
(3.2.2)

The rotation given by Euler angles  $(\phi, \theta, \psi)$  is as follows, where

- 1. The first rotation is by an angle  $\psi$  about the z-axis using D.
- 2. The second rotation is by an angle  $\theta \in [0, \pi]$  about the former x-axis (now x') using C.
- 3. The third rotation is by an angle  $\phi$  about the former z-axis (now z') using B.

Here, the notation  $(\psi, \theta, \phi)$  is used. In the x -convention, the component rotations are then given by the following rotations:

$$R_{z',\phi} = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos\phi & -\sin\phi \\ 0 & \sin\phi & \cos\phi \end{bmatrix} \quad R_{x',\theta} = \begin{bmatrix} \cos\theta & 0 & \sin\theta \\ 0 & 1 & 0 \\ -\sin\theta & 0 & \cos\theta \end{bmatrix} \quad R_{x,\psi} = \begin{bmatrix} \cos\psi & -\sin\psi & 0 \\ \sin\psi & \cos\psi & 0 \\ 0 & 0 & 1 \end{bmatrix}$$
(3.2.3)

Finally, by developing the expressions (3.2.3) the rotation matrix is described as a rotation across all the axis. The final matrix can be expressed as



Figure 3.6 Euler angles rotations. Source: Stack-exchange [22].

Eventually, the angular velocities  $\omega$  can be determined by the Euler angles derivatives [20].

Since the position is uniquely defined by Euler's angles, angular velocity is expressible in terms of these angles and their derivatives. The angular velocity vector  $\Omega_B^O = [\Omega_1 \Omega_2 \Omega_3]^T$ , which components are the rates is determined by the Euler angles and the derivative of  $\phi$ ,  $\theta$ ,  $\psi$ . The angle rates are respectively are known as precession rate  $\dot{\phi}$ , nutation rate  $\dot{\theta}$ , and spin  $\dot{\psi}$  [20].

To do so, the strategy is is to find the angular velocity components along the body axes  $x_1, x_2, x_3$  of  $\dot{\phi} \dot{\theta}$  $\dot{\psi}$  in turn.  $\dot{\theta}$  is along the line ON, and hence, in the  $x_1, x_2$  plane. Next,  $\dot{\phi}$  is about the Z axis.

Thus, the Euler angle angular velocities (components along the body's principal axes) are:

$$\vec{\theta} = (\dot{\theta}\cos\psi, -\dot{\theta}\sin\psi, 0)$$
  

$$\vec{\phi} = (\dot{\phi}\sin\theta\sin\psi, \dot{\phi}\sin\theta\cos\psi, \dot{\phi}\cos\theta)$$
(3.2.5)  

$$\vec{\psi} = (0, 0, \dot{\psi})$$

and the angular velocity components velocity components along those in-body axes  $x_1, x_2, x_3$  are:

$$\Omega_B^O = \begin{cases} \Omega_1 = \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \\ \Omega_2 = \dot{\phi} \sin \theta \cos \psi - \dot{\theta} \sin \psi \\ \Omega_3 = \dot{\phi} \cos \theta + \dot{\psi} \end{cases}$$
(3.2.6)

The above expressions (3.2.6) can be solved in terms of precession, nutation and spin rates:

$$\dot{\phi} = \frac{1}{\sin\theta} \left( \Omega_1 \sin\psi + \Omega_2 \cos\psi \right)$$
  
$$\dot{\theta} = \Omega_1 \cos\psi + \Omega_2 \sin\psi$$
  
$$\dot{\psi} = -\frac{1}{\tan\theta} \left( \Omega_1 \sin\psi + \Omega_2 \cos\psi \right) + \Omega_3$$
  
(3.2.7)

It is crucial to avoid nutation angle through 90  $^{\circ}$  so as to avoid the singularity that exists in the precession angle [23].

Euler angles are astoundingly useful to get an idea of a satellite 's attitude. Notwithstanding, a particular scenario can cause Euler angles to provide erroneous attitude due to singularities for some angles, e.g.  $\cos \pi/2$  when calculating the resulting rotation matrix. As attitude exists in  $SO^3$ , no three-parameter representation, such as Euler angles or the Gibbs Vector is singularity free. For this reason *Euler parameters*, also called *quaternions* are used in the numerical computations in this project. However, for ease of interpretation, Euler angles are often used to represent the results.

#### 3.2.2.1 Euler Gimbal Lock

Euler gimbal lock occurs when the orientation of the sensor cannot be uniquely described using Euler Angles. The exact orientation at which the gimbal lock occurs is determined by the rotational sequence used. When the pitch angle  $\alpha$  is 90 degrees, the sequencing of actions on the IMU sensor results in a gimbal lock. This results in a loss of a degree of freedom since two axes are coupled and there is no third axis to turn on.

To avoid Euler Gimbal Lock singularities, the correct way to approach this problem is using *quaternions* (see Section 3.2.3) by taking the Euler parameters (quaternions) and performing a transformation, the rotation matrix expressed in terms of quaternions is obtained as

$$R_B^O = \begin{bmatrix} q_0^2 + q_1^2 - q_2^2 - q_3^2 & 2(q_1q_2 + q_0q_3) & 2(q_1q_3 - q_0q_2) \\ 2(q_1q_2 - q_0q_3) & q_0^2 - q_1^2 + q_2^2 - q_3^2 & 2(q_2q_3 + q_0q_1) \\ 2(q_1q_3 + q_0q_2) & 2(q_2q_3 - q_0q_1) & q_0^2 - q_1^2 - q_2^2 + q_3^2 \end{bmatrix}$$
(3.2.8)



Figure 3.7 Euler Angles. Source: Own.

#### 3.2.3 Quaternions

A quaternion q (a.k.a Euler parameters) is a singularity free attitude representation hyper complex number. To avoid confusion, the term quaternion will be used from now on. Quaternions are used in computation to overcome the singularity problem in the attitude representation and is defined as a complex number composed of a scalar real part  $q_0$  and an imaginary three component vector  $q_{1:3}$ , with components spanning  $\mathbb{R}^3$  defined by [24] [25]:

$$q = \begin{bmatrix} q_0 \\ q_{1:3} \end{bmatrix} = \begin{cases} q_0 = \cos \phi/2 \\ q_1 = e_1 \sin \phi/2 \\ q_2 = e_2 \sin \phi/2 \\ q_3 = e_3 \sin \phi/2 \end{cases}$$
(3.2.9)

they represent a rotation about a unit vector through and angle  $\phi$ , where

$$e_1^2 + e_2^2 + e_3^2 = 1 (3.2.10)$$

and q satisfies the holonomic constraint  $q^T q = 1$  which also means

$$q_0^2 + q_1^2 + q_2^2 + q_3^2 = 1 (3.2.11)$$

Note that this constraint geometrically describes a four-dimensional unit sphere.

Quaternions can also be written  $(q_0, q_1, q_2, q_3)$ . It is an element of  $\mathbb{R}^4$  and mathematically not defined in

ordinary linear algebra. The hyper imaginary numbers i, j and k must satisfy the following conditions

$$i^{2} = j^{2} = k^{2} = ijk = -1$$

$$ij = k = -ji$$

$$jk = i = -kj$$

$$ki = j = -ik$$

$$(3.2.12)$$

Subsequently, the transformation from Euler angles to quaternions can be calculated as

$$q = \begin{bmatrix} q_0 \\ q_1 \\ q_2 \\ q_3 \end{bmatrix} = \begin{bmatrix} (\cos(\phi/2)\cos(\theta/2)\cos(\psi/2) + \sin(\phi/2)\sin(\theta/2)\sin(\psi/2) \\ (\sin(\phi/2)\cos(\theta/2)\cos(\psi/2) - \cos(\phi/2)\sin(\theta/2)\sin(\psi/2) \\ (\cos(\phi/2)\sin(\theta/2)\cos(\psi/2) + \sin(\phi/2)\cos(\theta/2)\sin(\psi/2) \\ (\cos(\phi/2)\cos(\theta/2)\sin(\psi/2) - \sin(\phi/2)\sin(\theta/2)\cos(\psi/2) \end{bmatrix}$$
(3.2.13)

The inverse relation, to transform from quaternions to Euler angles is expressed as  $^2$ 

$$\begin{bmatrix} \phi \\ \theta \\ \psi \end{bmatrix} = \begin{bmatrix} \operatorname{atan} 2\left(2\left(q_0q_1 + q_2q_3\right), 1 - 2\left(q_1^2 + q_2^2\right)\right) \\ \operatorname{arcsin}\left(2\left(q_0q_2 - q_1q_3\right)\right) \\ \operatorname{atan} 2\left(2\left(q_0q_3 + q_1q_2\right), 1 - 2\left(q_2^2 + q_3^2\right)\right) \end{bmatrix}$$
(3.2.14)

#### 3.2.3.1 Time derivation of a quaternion

The calculation of the time derivative of a quaternion q in terms of the quaternion itself and the corresponding angular rate is defined by Kuipers [24] as

$$\dot{q} = \frac{1}{2}\vec{\omega}' \otimes q = \frac{1}{2}\Omega'(\vec{\omega})q \tag{3.2.15}$$

$$\dot{q} = \frac{1}{2}q^* \otimes \vec{\omega}' = \frac{1}{2}\Xi(q)\vec{\omega}', \qquad (3.2.16)$$

where

$$\Omega'(\vec{\omega}) = \begin{bmatrix} 0 & \omega_z & -\omega_y & \omega_x \\ -\omega_z & 0 & \omega_x & \omega_y \\ \omega_y & -\omega_x & 0 & \omega_z \\ -\omega_x & -\omega_y & -\omega_z & 0 \end{bmatrix}, \\ \Xi(q) = \begin{bmatrix} q_4 & -q_3 & q_2 & q_1 \\ q_3 & q_4 & -q_1 & q_2 \\ -q_2 & q_1 & q_4 & q_3 \\ -q_1 & -q_2 & -q_3 & q_4 \end{bmatrix} \text{ and } \vec{\omega}' = \begin{bmatrix} \vec{\omega} \\ 0 \end{bmatrix} = \begin{bmatrix} \omega_x \\ \omega_y \\ \omega_z \\ 0 \end{bmatrix}$$
(3.2.17)

Based on prior results, the equation for calculating angular rates in terms of a rotation quaternion and its derivative is as follows:

$$\vec{\omega}' = 2\dot{q} \otimes q^* \tag{3.2.18}$$

Similarly, Kuipers c<sup>[24]</sup> proposes the second derivative of a quaternion with respect to time.

$$\ddot{\mathbf{q}} = \frac{1}{2} \left( \dot{\vec{\omega}}' \otimes q + \vec{\omega}' \otimes \dot{q} \right) \tag{3.2.19}$$

<sup>&</sup>lt;sup>2</sup>Four-quadrant inverse tangent which is also known as arc tangent of two numbers.  $\tan 2(y, x)$  returns the arctangent of the two numbers x and y. It is similar to calculating the arctangent of y/x, except that the signs of both arguments are used to determine the quadrant of the result. The result is an angle expressed in radians rad. Notice how arctan and arcsin functions implemented in computer languages only produces values from  $-\pi/2$  to  $\pi/2$ , for all the possible rotation between those values one does not obtain all possible orientations. This is why it is needed to replace arctan to atan 2 [26].

## 3.3 Disturbance analysis

Any spacecraft orbiting Earth is subject to some perturbations and disturbance forces and torques. These perturbations come from various sources and their magnitude and direction depend on different physical properties of the satellite such as its weight, shape, material, etc. Whenever a force is not acting through the COM of the satellite it will result in a net torque.

Four major sources of disturbance torques shall be taken into account when designing the ADCS hardware and software:

- Aerodynamic drag (atmospheric density)
- Geomagnetic field disturbances
- Gravity gradient
- Solar radiation pressure

For LEO orbits, the ones that have a larger impact are the aerodynamic drag and the geomagnetic field disturbances in terms of torque disturbances [12]. Nonetheless, the following sections will present an overview of all these four sources of disturbances as well as analytical estimation models of them as a function of the position and orbital velocity. In all the sources, the worst-case scenario will be taken into account.

#### 3.3.1 Aerodynamic drag

The aerodynamic drag is one of the major contributors to the disturbance torque. This is due to the number of air particles present at the altitude at which the CubeSat will be operating.

NASA's Space Vehicle Design Criteria SP-8058 [27] is used to formulate the aerodynamic drag. To compute the torque, the aerodynamic pressure acting on the surface of the CubeSat is to be integrated over the exposed forward-facing ("wetted") area of the spacecraft.

To assay how the atmosphere interacts with the spacecraft, the interaction between a body and an atmosphere through which it is moving is to be analyzed in the subsequent section. At low orbit altitudes, the interaction of the vehicle and the atmosphere can be characterized by the *free-molecular flow regime* of gas dynamics, i.e., when the molecular mean free path is much greater than a characteristic spacecraft dimension.

The incident flow is considered to be undisturbed by the presence of the spacecraft. The net aerodynamic torque can be calculated by summing up all the contributions of each spacecraft element surface. The gross value of the aerodynamic force may be obtained using the following expression:

$$D = \frac{1}{2}\rho V^2 C_d A_{\rm inc} \tag{3.3.1}$$

 $\rho~[\rm kg/m^3]$ : is the atmospheric density at given altitude.

V [m/s]: is the spacecraft velocity also depending on the orbit altitude.

 $C_d$  [adim]: is the drag coefficient. A reference value between 2-2.5 is taken for CubeSat.

 $A_{\rm inc}$  [m<sup>2</sup>]: is the incident satellite surface area.

The atmospheric drag is predominant for Earth orbits up to 600 km, which are the most common for CubeSat missions. In general, for spacecraft in Earth orbit, the radiation force on a given surface becomes more significant than the aerodynamic force at orbital altitudes above 1000 km. As seen previously, nanosatellites missions are launched in low circular or elliptical orbits LEO (altitudes of between 400 and 650 km) and travel at around 8 km/s. [8] [7]. Between 600 and 1000 km, torques resulting from solar radiation and aerodynamic forces are likely to be of the same order of magnitude. Nevertheless, aerodynamic torques have a direct impact on the attitude control of a spacecraft orbital lifetime.

The major drag induced by the friction with the upper atmosphere's particles will be produced in the lowest layer of the orbit's altitude. Hence, the main hypothesis used in this part is to consider the nanosatellite's orbit with a constant altitude of h = 400 km above the surface of Earth. Furthermore, the ISA model cannot be applied at altitudes above 100 km, however, Bowman suggests a more accurate empirical thermospheric density model of the atmosphere for altitudes ranging between 175 to 1000 km [28].

According to [27], using a simplified particle/surface interaction models, based on free-molecular flow theory. The aerodynamic force on an elemental area dA is:

$$dD = \rho \left[ \left( 2 - \sigma_{\rm n} - \sigma_t \right) \left( v^{\rm T} n_{\rm dA} \right) n_{\rm dA} + \sigma_{\rm t} v \right] v^{\rm T} n_{\rm dA} dA$$
(3.3.2)

where

- $\sigma_n$  [adim]: is the normal exchange coefficient.
- $\sigma_t$  [adim]: is the tangential momentum exchange coefficient.
- $v^{\rm T}$  [m/s<sup>2</sup>]: is the spacecraft tangential velocity.
- $\rho$  [kg/m<sup>3</sup>]: is the atmospheric density.
- $n_{\mathrm{d}A}$ : is the normal vector of a surface dA (outwards the CubeSat).

As regards the momentum exchange coefficients, a good empirical values is around  $\sigma_n = \sigma_t = 0.8$ . The linear relationship between the dD and the density  $\rho$  in the equation above calls for an accurate atmospheric model density.

The elemental torque  $d\tau_D$  generated from a force dD by an elemental area dA is given by equation (3.3.3), where  $r_d A$  is the location of the centroid of the elemental area and  $r_{cg}$  is the location of the centre of mass:

$$d\tau_D = (r_{dA} - r_{cg}) \times dD \tag{3.3.3}$$

Total torque can be computed by integrating the above expression over the whole wetted surface that is affected by the aerodynamic drag using

$$D = \int_{\mathcal{A}_{wetted}} \mathrm{d}D \tag{3.3.4}$$

Particularly, PLATHON's 1U CubeSat dimensions are equal on all sides, as it is a square CubeSat, thus, the six faces of the satellite can be considered as an elemental area with the net force acting through the centroid of each surface. Hence, the net aerodynamic torque is the result of summing all elemental forces on all surfaces where  $v^T \cdot n_{dA} > 0$ . However, if the Cubesat is not square, the wetted area is not the same for all orientations. Combining this explanation to equations (3.3.2) and (3.3.4),

$$D_j = \rho \left[ (2 - \sigma_n - \sigma_t) \left( v^T n_j \right) n_j + \sigma_t v \right] v^T \max(0, v^T n_j) n_j A_j \quad j = \pm x, \pm y, \pm z$$
(3.3.5)

$$\tau_{D,j} = \sum_{j=\pm x,\pm y,\pm z} \left( r_{A_j} - r_{\rm cg} \right) \times D_j \tag{3.3.6}$$

In the above expression,  $n_i$  indicates the normal vector of the surface in j direction,  $A_i$  is the surface facing  $n_j$  direction and  $r_{A,j}$  represents the centroid of surface j.

Now, the maximum wetted area of a 1U CubeSat is  $0.01 \text{ m}^2$ . The density of the spacecraft around 400 km is  $\rho$  = and the satellite's orbital velocity ranges from 7.6738 km/s and 7.5629 km/s at an orbital altitude of 600 km (see expression (3.3.7))

$$v = \sqrt{\frac{GM_{\rm E}}{R_{\rm E} + h}} = \sqrt{\frac{6.67408 \cdot 10^{-11} \cdot 5.9742 \cdot 10^{24}}{6371 \cdot 10^3 + 400 \cdot 10^3}} = 7673.7770 \text{ m/s}$$
(3.3.7)

Jacchia-Bowman Atmosphere Model [28] provides density values from the last decade (see Figures 3.8), which shows how difficult is to model Earth's atmospheric density. There is are no tabular values of density. Instead, the barometric equation and diffusion equation are integrated numerically using the Newton-Coates method to produce the density profile up to the input position. (check Appendix K for the MATLAB code):



(b) Temperature from 1999 to 2015. Source: Own K.2.

Figure 3.8 Density and temperature fluctuation from 1999 to 2015 using JB2008. Source: Own K.2.

Besides, the density fluctuates according to solar activity (solar cycles) and JB2008 takes into account those events. Higher values of density means that there are more particles in the atmosphere colliding with the satellite. Thus, according to Figure 3.8, a value of

$$\rho_{h_{400}} \approx 4.58268 \cdot 10^{-12} \text{ kg/m}^3 \tag{3.3.8}$$

will be used for the calculations, which is the maximum density at an altitude of h = 400 km. The maximum wetted area for a 1U CubeSat is  $A_{\text{max}} = 0.01$  m<sup>2</sup> with the largest offset on the centroid is approximately 0.03 m.

With these assumptions, one can estimate the aerodynamic drag using expression (3.3.5) and (3.3.6):

$$D_{\max} = 4.58268 \cdot 10^{-12} \left[ (2 - 0.8 \cdot 0.8) \cdot 7673.7770 + 0.8 \cdot 7673.7770 \right] \cdot 7673.7770 \cdot 0.01$$

 $= 1.0978 \cdot 10^{-7} \text{ N} \quad (3.3.9)$ 

$$\tau_D = (0.05 - 0) \cdot 1.0978 \cdot 10^{-7} = 3.9939 \cdot 10^{-7} \text{ Nm}$$
(3.3.10)



Figure 3.9 Solar activity over the past 50 years. Source: SILSO [29].

Since the value of atmospheric density is affected by altitude and solar activity, the following plot 3.10 simulates atmospheric torque as a function of height h and  $C_D$  drag coefficient in terms of solar activity. The value of the atmospheric density is presented in this model for various levels of solar activity (low, moderate, and high).

In this case, equations (3.3.1) (3.3.4) may be used to obtain the following expression:

$$\tau_{D} = \frac{1}{2} C_{D} \rho A v^{2} \left( r_{dAF} - r_{cg} \right) \frac{3}{\sqrt{2}}$$
(3.3.11)

Since the value of  $C_D$  has yet to be determined, Wertz and Larson [30] offers a table with typical drag coefficients for various missions flown in past years, where it is seen that  $C_D$  typically fluctuates between 1 and 4 (see Appendix K for the code and JB2008 algorithm).





Figure 3.10 Evolution of the atmospheric torque in terms of h,  $C_D$  and the solar activity using JB2008. Source: Own.

Closer examination of the simulations in Figure 3.10 shows clearly how the aerodynamic torque increases as drag coefficient grows. Besides when solar activity F10.7 index increases it also influences the induced torque.

#### 3.3.2 Gravity gradient torque

Any non-symmetrical object with finite dimensions in orbit is subject to a force known as gravitational torque because of the variance in the Earth's gravitational pull over the object. The inverse square gravitational force field causes this gravity-gradient torque; in a homogeneous gravitational field, there would be no gravitational torque [12].

Equation (3.3.12) may be used to calculate the torque caused by a gravity gradient. The question to be asked here is how the gravitational force affects the satellite itself since the lower part of the satellite is closer to Earth while the upper side of the satellite is further from Earth.

Newton's law shows how gravitational force decreases with the square root of the distance  $1/r^2$ . Although at first, this height difference may be inappreciable, a constant torque acting on a body in space where friction is considerably smaller can end up rotating the whole CubeSat upside down if not taken into consideration.

NASA [31] and Wertz [12] approach shows that gravity gradient torque can be defined as

$$\vec{\tau}_{\rm gg} = \left(\frac{3Gm_{\oplus}}{\|\vec{r}\|^3}\right) \left[\hat{r} \times (I_{\rm sat} \cdot \hat{r})\right] = \left(\frac{3\mu_{\rm Earth}}{\|r\|_2^5}\right) \left[\vec{r} \times (I_{\rm sat}\vec{r})\right]$$
(3.3.12)

where

 $\mu_{\text{Earth}} [\text{m}^3/\text{s}^2] = 3.9860 \cdot 10^{14} \text{ m}^3/\text{s}^2$ : is Earth's gravitational constant.

 $\vec{r}$  [m]: is the position of the satellite's COM with respect to Earth's COM.

 $I_{\rm sat}$  [kgm<sup>2</sup>]: is the satellite's moment of inertia matrix.

All torques that tend to upset a spacecraft's attitude must be addressed in the design of spacecraft attitude control systems. Gravity gradient is one of these torques and it derives from variations in the gravitational force over the distributed mass of the spacecraft.

Considering the worst case scenario, expression (3.3.12) can be simplified as:

$$\tau_{\rm gg,max} = \frac{3\mu_{\rm Earth}}{2R_{\rm Orbit}^3} \left( I_{\rm sat,max} - I_{\rm sat,min} \right) \cdot \sin(2\theta_{\rm dev}) \tag{3.3.13}$$

in which  $I_{\text{sat,max}}$  and  $I_{\text{sat,max}}$  are the maximum and minimum moment of the inertia;  $\theta_{\text{dev}}$  is the maximum deviation of the Z axis and therefore is maximum at  $\theta_{\text{dev}} = 45^{\circ}$  in body frame from its local vertical (refer to section 3.2.1) and  $R_{\text{Orbit}}$  represents the orbital altitude, in which  $R_{\text{Orbit}} = R_{\text{Earth}} + h$ .

The inertia of the satellite can be deduced using expression (3.4.5), since it is not a massive cube the following subtraction must be made (considering solely the frame with the total components' mass):

$$I_{\rm sat,min} = m_{\rm sat} \frac{d_{\rm out}^2 - d_{\rm int}^2}{6} = 1.3 \cdot \frac{0.01^2 - 0.008^2}{6} = 6 \cdot 10^{-6} \text{ kg} \cdot \text{m}^2$$
(3.3.14)

For the maximum inertia, let's suppose the satellite is has some mass shifted outside the center of mass (where  $m_{\text{shift}}$  can be some electric component such as the battery):

$$I_{\text{sat,max}} = m_{\text{sat}} \frac{d_{\text{out}}^2 - d_{\text{int}}^2}{6} + m_{\text{shift}} \cdot d_{\text{shift}}^2 = 1.3 \cdot \frac{0.01^2 - 0.008^2}{6} + 0.3 \cdot 0.02^2 = 1.278 \cdot 10^{-4} \text{ kg} \cdot \text{m}^2 \quad (3.3.15)$$

where  $m_{\text{sat}}$  is the satellite's mass and d represents the side length of the CubeSat. Since it is a 1U CubeSat, the Inertia is the same for all principal axis. Considering the worst scenario case, the CubeSat's maximum weight shall be around 1.3 kg

Accounting for the following parameters:

$$I_{\text{sat,max}} = 2.957 \text{ kgm}^2$$
  
 $I_{\text{sat,min}} = 2.899 \text{ kgm}^2$   
 $R_{\text{Earth}} = 6371.000 \text{ km}$   
 $h = 400 \text{ km}$   
 $R_{\text{Orbit}} = 6771.0000 \text{ km}$   
 $\theta_{\text{dev}} = 45^{\circ}$   
(3.3.16)

which results in a maximum gravitational torque of

$$\tau_{\rm gg,max} = \frac{3 \cdot 3.9860 \cdot 10^{14}}{2 \cdot (6771 \cdot 10^3)^3} \left( 1.274 \cdot 10^{-4} - 6 \cdot 10^{-6} \right) \cdot \sin(2 \cdot 45^\circ) = \boxed{2.3594 \cdot 10^{-10} \,\,\mathrm{Nm}} \tag{3.3.17}$$

Since the CubeSat is symmetric, it's gravity gradient contribution shall shall not affect too much.

#### 3.3.3 Geomagnetic Field

The magnetic torque results from the interaction between the magnetic properties of the spacecraft and the ambient magnetic field (the Earth's magnetic field, in the case that applies to this project), is the subject of section [32].

Sources of magnetic disturbance torques to be considered should include but should not be limited to:

- Spacecraft magnetic moments: permanent magnetism in the spacecraft.
- Eddy currents: spacecraft generated current loop.
- Hysteresis: Magnetism induced by external fields or currents induced by external fields, i.e., when magnetic induction lags behind the magnetizing force.

For missions conducted within the Earth's magnetosphere (which is approximately 10 times Earth's radius from the geocenter), the magnetic field can be described by:

- 1. A **steady-state** nominal analytical model that defines magnitude and direction as a function of location with corrections distortions.
- 2. **Distortions** in the field caused by influences from outside the magnetosphere (effects of solar plasma) not accounted for by the nominal model.
- 3. Time dependent disturbances that alter and are superimposed on the model.

For missions that go beyond the Earth's magnetosphere, a value of 40 nT in an arbitrary direction is an appropriate upper-bound characterisation of the interplanetary magnetic field at about 1 AU for missions performed beyond the Earth's magnetosphere. With increasing distance from the Sun, the influence of Earth's magnetic field weakens.

The mathematical model of Earth's geomagnetic field is given by Wertz [12]:

$$B_{\text{Earth}} = \frac{R_{\text{Earth}}^3 H_0}{\vec{r}^3} \left[ 3 \left( \hat{m}_{\text{north}} \cdot \hat{r} \right) \hat{r} - \hat{m}_{\text{north}} \right]$$
(3.3.18)

where

 $R_{\rm Earth}^3 H_0 ~[{\rm Tm}^3]$  represents the strength of the geomagnetic moment.

 $\vec{r}^3$  [m]: is the position of the satellite in ECEF reference frame.

 $\hat{m}_{\text{north}}$ : is the normal vector representing the position of the north geomagnetic pole (which is updated by IAGA (International Association of Geomagnetism and Aeronomy) every 5 years [33]).

 $R_{\rm Earth}$  [m]: denotes the mean Radius of Earth.

$$H_0 = \sqrt{\left(g_1^0\right)^2 + \left(g_2^1\right)^2 + \left(h_1^1\right)^2} \tag{3.3.19}$$

where  $g_n^m$  and  $h_n^m$  are called Gauss coefficients which are functions of time and are conventionally given in units of nano-Tesla [nT]. and the most recent values found in [33] are:

$$g_1^0 = -29442 \cdot 10^{-9} \text{ T}$$
  

$$g_1^1 = -1501 \cdot 10^{-9} \text{ T}$$
  

$$h_1^1 = 4797.1 \cdot 10^{-9} \text{ T}$$
  
(3.3.20)

Subsequently, the average geomagnetic moment is calculated using the preceding values in equation

$$R_{\text{Earth}}^3 H_0 = 7.7238 \cdot 10^{15} \text{ T} \cdot \text{m}^3 \tag{3.3.21}$$

The interaction between the Earth's magnetic field and any other magnetic field created inside the satellite produces the magnetic disturbance torque. Eddy currents, hysteresis, and the magnetic dipole of the satellite's structure are the three fundamental processes that lead to the formation of magnetic torques. However, the torque generated by Eddy currents is insignificant in magnitude when compared to the torque generated by the CubeSat's internal structure [32].

The instantaneous magnetic disturbance torque  $\vec{\tau}_{\rm m}$  is the vector cross-product of the spacecraft effective dipole moment  $\vec{M}$  and the local magnetic induction or flux density  $\vec{B}$ :

$$\vec{\tau}_{\rm m} = \dot{M}_{\rm sat} \times \dot{B}_{\rm Earth}$$
 (3.3.22)

where

 $M_{\text{Earth}}$  [Am<sup>2</sup>]: satellite's local magnetic induction.  $\vec{B}_{\text{Earth}}$  [Wb/m<sup>2</sup>] or [T]: Earth's magnetic field.

The maximum magnetic torque occurs in a polar orbit (when the inclination is 90  $^{\circ}$ ). Therefore, expression (3.3.18) can be simplified as,

$$B_{\text{Earth,max}} = \frac{2 \cdot R_{\text{Earth}}^3 H_0}{\vec{R}_{\text{Orbit}}^3}$$
(3.3.23)

For values of

$$R_{\text{Earth}}^{3} H_{0} = 7.7238 \cdot 10^{15} \text{ T} \cdot \text{m}^{3}$$

$$R_{\text{Earth}} = 6371.0000 \text{ km}$$

$$h = 400 \text{ km}$$
(3.3.24)

$$R_{\rm Orbit} = 6771.0000 \text{ km}$$

and (3.3.22) can be reduced to

$$\tau_{\rm m,max} = M_{\rm sat} \cdot B_{\rm Earth} \tag{3.3.25}$$

This way, replacing the above values into equation (3.3.23) As for the moment there is no data available for PLATHON's CubeSat, the magnetic moment shall be as:

$$\vec{M}_{\text{sat}} = \left[\frac{1}{\sqrt{3}}\frac{1}{\sqrt{3}}\frac{1}{\sqrt{3}}\right]^T \text{ Am}^2 \implies M_{\text{sat}} = 1 \text{ Am}^2$$
(3.3.26)

Thus, the maximum torque due to the geomagnetic field is

$$\tau_{\rm m,max} = M_{\rm sat} \cdot \frac{2 \cdot R_{\rm Earth}^3 H_0}{\vec{R}_{\rm Orbit}^3} = 1 \cdot \frac{2 \cdot 7.7238 \cdot 10^{15}}{(6771 \cdot 10^3)^3} = \boxed{4.9763 \cdot 10^{-5} \,\,\rm Nm} \tag{3.3.27}$$

#### 3.3.4 Solar radiation pressure

Radiation incident on a spacecraft's surface generates forces that may cause a torque about the spacecraft's mass centre. Surface characteristics are dominant among the factors causing both the radiation and aerodynamic torques that act on a spacecraft [34].

The direct solar illumination is the primary source of radiation force. Earth-reflected sunlight and infrared emission from the Earth and its atmosphere are additional sources for spacecraft in Earth orbit. Asymmetrical emission of electromagnetic energy (typically heat or radio signals) from onboard the spacecraft should also be considered as a radiation source.

Major factors in the determination of radiation torques are

- The intensity, spectrum, and direction of the incident or emitted radiation.
- The shape of the surface and the location of the Sun face with respect to the mass centre of the spacecraft.
- The optical properties of the surface upon which the radiation is incident or from which it is emitted.

Since solar irradiation varies as the inverse square of the distance between the spacecraft and the Sun, for a spacecraft in near-Earth orbit, this contribution to the radiation force is essentially independent of altitude. Because most other disturbance torques tend to diminish with increasing altitude, radiation torque is most likely to be a significant factor in the design of spacecraft with a large surface area that operate at orbital altitudes above about 1000 km.

NASA's Space Vehicle Design Criteria on Radiation torques provides a good approximation of the force from solar radiation pressure acting on an illuminated elemental area dA. The assumption is made that all the incident radiation is partly absorbed, partly reflected diffusely, partly reflected specularly and negligible scattering occurs.

$$dF_{\rm SRP} = \frac{I_{\rm SRP}}{c} \left\{ -\left[ (1+c_{rs})\cos\alpha + \frac{2}{3}c_{rd} \right] \hat{n} + (1-c_{rs})\sin\alpha \,\hat{s} \right\} \cos\alpha \, dA_{\rm inc}$$
(3.3.28)

where

 $I_{\text{SRP}} [\text{W/m}^2] = 1396 \text{ W/m}^2$ : is the energy per unit of time through a cross sectional unit area [34]. c [m/s] = 299792458 m/s: is the speed of light in vacuum.

 $c_{rd}$ : is the coefficient of specular diffusion.

 $c_{rs}$ : is the coefficient of diffuse reflection.

 $\hat{n}$ : is the unit vector normal to the incident surface  $dA_{inc}$ .

 $\hat{s}$ : is the unit vector perpendicular to  $\hat{n}$  co-planar to the incident ray.

 $\alpha$  [rad]: is the angle between the incident ray and  $\hat{n}$ .

 $A_{\rm inc}$  [m<sup>2</sup>]: is the incident satellite surface area.

with

$$0 \le c_a + c_{rd} + c_{rs} \le 1^3 \tag{3.3.29}$$

According to [35], for a solar panel the coefficient of specular diffusion and diffuse reflection are  $c_{rd} = 0.042$ and  $c_{rs} = 0.168$ , and it is assumed for the whole spacecraft.

Analogously, the elemental torque  $dT_{SRP}$  generated from a force  $dF_{SRP}$  on elemental area dA is given by

$$d\tau_{SRP} = (r_{dA} - r_{cg}) \times dF_{SRP}$$
(3.3.30)

The total torque acting on the spacecraft can be computed by integrating Equation (3.3.30) over the entire spacecraft.

$$\tau_{\rm SRP} = \int_{A_{\rm inc}} d\tau_{\rm SRP} \tag{3.3.31}$$

It is important to mention that the disturbance force generated by the albedo effect (the radiation reflected by Earth) is not taken into account since its contribution is way more reduced.

Thus, the worst case scenario can be computed as

$$F_{\rm SRP,j} = \frac{I_{\rm SRP}}{c} \left\{ -\left[ (1+c_{rs})\cos\alpha + \frac{2}{3}c_{rd} \right] \hat{n}_j + (1-c_{rs})\sin\alpha \,\hat{s} \right\} \cos\alpha \,\mathrm{d}A_{\rm inc,j} \quad j = \pm x, \pm y, \pm z$$

$$(3.3.32)$$

$$\tau_{\rm SRP,j} = \sum_{j=\pm x,\pm y,\pm z} \left( r_{A_j} - r_{\rm cg} \right) \times F_{\rm SPR,j}$$
(3.3.33)

The incident area of a 1U CubeSat is  $0.01 \text{ m}^2$ . The six faces can be accounted as one elemental area and hence, the integral in can be replaced by the summation over the centroid of each face. With the following assumptions,

$$I_{sr} = 1396 \text{ W/m}^2$$

$$c = 299792458 \text{ m/s}$$

$$c_{rd} = 0.042$$

$$c_{rs} = 0.168$$

$$\alpha = 0^{\circ}$$

$$A_{inc} = 0.01 \text{ m}^2$$
(3.3.34)

the solar radiation torque yields to a maximum radiation torque of

$$F_{\rm SRP,max} = \frac{1396}{299792458} \left\{ -\left[ (1+0.168)\cos 0 + \frac{2}{3}0.168 \right] \hat{n}_j + (1-0.168)\sin 0 \ \hat{s} \right\} \cos 0.01 = 4.3958 \cdot 10^{-8}$$

$$(3.3.35)$$

$$T_{\rm SRP,max} = (0.05-0) \cdot = \boxed{2.1979 \cdot 10^{-9} \ \rm Nm^2}$$

$$(3.3.36)$$

 $<sup>{}^{3}</sup>c_{a}$  is the absorption coefficient



Figure 3.11 Absorption, Reflection and Diffusion over a surface. Source Own.

### 3.3.5 Total Disturbance

The total disturbance is the sum of all disturbances in their worst-case scenario. This comprises the worst-case scenario for all disturbance torques. As claimed by Dr. Brown [36], the satellite would be able to produce twice as much torque as the sum of all the external torques analyzed hitherto.

Table 3.2 summarizes the calculations obtained in previous sections. These results shall be taken into account when choosing the right motor and the design of the reaction wheel.

	Disturbance Torque	$\mathbf{Magnitude} \; [\mathrm{Nm}]$
$ au_D$	Aerodynamic	$3.9939 \cdot 10^{-7}$
$ au_{ m gg}$	Gravity Gradient	$2.3594 \cdot 10^{-10}$
$ au_{ m m}$	Geomagnetic Field	$4.9763 \cdot 10^{-5}$
$ au_{\mathrm{SRP}}$	Solar Radiation Pressure	$2.1979 \cdot 10^{-9}$
τ	TOTAL	$5.0165 \cdot 10^{-5}$

Table 3.2 Summary of external disturbance torques and their contribution. Source: Own.

This contribution of external torque shall be added to the control algorithm. However, the algorithm implemented in section 7 is not considered this contribution since the CubeSat will be tested in an Air bearing. Notwithstanding, for a future mission, these disturbance torques are highly important and

should not be disregarded.

# 3.4 Rotational Dynamics

#### 3.4.1 Moment of Inertia

In a rigid body rotating with an angular velocity  $\omega$ , the angular momentum  $H_{\rm G}$  computed from its center of mass [20] yields:

$$H_{\rm G} = \int \rho \times \dot{r} \, \mathrm{d}m = \int \rho \times (\omega \times \rho) \, \mathrm{d}m \tag{3.4.1}$$

where

 $\rho{:}$  is the density of the spacecraft.

 $\mathrm{d}m:$  is a differential of mass.

 $\dot{r}$ : is the velocity of the spacecraft.

The angular momentum  $H_{\rm G}$  can also be expressed with respect to the moment of inertia as

$$\vec{H} = [I]\vec{\omega} \rightarrow \begin{pmatrix} H_x \\ H_y \\ H_z \end{pmatrix} = \begin{bmatrix} I_{xx} & I_{xy} & I_{xz} \\ I_{yx} & I_{yy} & I_{yz} \\ I_{zx} & I_{zy} & I_{zz} \end{bmatrix} \begin{pmatrix} \omega_x \\ \omega_y \\ \omega_z \end{pmatrix}$$
(3.4.2)

Eventually, the components of the Inertia tensor about the center of mass are:

$$I_x = \int (y^2 + z^2) dm \quad I_{xy} = -\int xy dm \qquad I_{xz} = -\int xz dm$$
  

$$I_{yx} = -\int yx dm \qquad I_y = \int (x^2 + z^2) dm \qquad I_{yz} = -\int yz dm$$
  

$$I_{zx} = -\int zx dm \qquad I_{zy} = -\int zy dm \qquad I_z = \int (x^2 + y^2) dm$$
  
(3.4.3)

If ij plane is a plane of symmetry of the body, then  $I_{ij} = I_{ji} = 0$ . The inertia of a mass expresses:

- How the mass of a rigid body is distributed
- How the body responds to applied torques

$$\frac{1}{12}m\begin{bmatrix} (a^2+l^2) & 0 & 0\\ 0 & (b^2+l^2) & 0\\ 0 & 0 & (a^2+b^2) \end{bmatrix}$$
(3.4.4)

However, in the case of a uniform density cube, the above expression can be simplified as:



Figure 3.12 Moments of inertia for ta rectangular parallelepiped. Source [20].

with d side and mass  $m_{sat}$ , the inertial matrix for a cube results as:

$$I = m_{\rm sat} \frac{d^2}{6} \begin{bmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix}$$
(3.4.5)

Notice that the diagonal terms are much higher than non-diagonal ones, then the homogeneous uniform density assumption would not be correct.

#### 3.4.2 Dynamics for Satellite Model

The subsequent section derives the dynamics and kinematics of the satellite. The description of the dynamics will employ Newton-Euler formulation where the angular momentum changes as a function of applied torque.

$$H = r \times p \tag{3.4.6}$$

where H is the angular momentum, r is the position and  $\vec{p} = m\vec{v}$  is the impulse.

Thereby, developing the derivative of the angular momentum one can see that

$$\frac{\mathrm{d}}{\mathrm{d}t}H = \left(\frac{\mathrm{d}}{\mathrm{d}t}r \times p\right) + \left(r \times \frac{\mathrm{d}}{\mathrm{d}t}p\right) = (v \times mv) + (r \times ma) = r \times F = \tau \tag{3.4.7}$$

where  $v \times v = 0$  and it is proved that

$$\frac{\mathrm{d}}{\mathrm{d}t}H = r \times F = \tau \tag{3.4.8}$$

Recall that the angular momentum H can be expressed with respect to the moment of inertia tensor and the angular velocity vector  $\omega$  as

$$\vec{H} = [I]\vec{\omega} \tag{3.4.9}$$

Now, if the CubeSat is assumed to be a rigid solid that only has reaction wheels as control actuators, the total angular momentum of the satellite in body reference  $H_{\text{sat}}^b$  yields,

$$H_{\rm sat}^b = I\omega_{\rm sat}^b + H_{\rm RW} \tag{3.4.10}$$

where  $H_{\rm RW}$  is the angular momentum generated by the reaction wheels, and  $\omega_{\rm sat}^b$  is the vehicle's angular velocity in body reference frame.

The angular momentum generated by the reaction wheel is defined as

$$H_{\rm RW} = T_{\rm RW} I_{\rm RW} \omega_{\rm RW} \tag{3.4.11}$$

where  $T_{\rm RW}$  is the reaction wheels distribution matrix,  $I_{\rm RW}$  expresses the inertia tensor of the reaction wheel and  $\omega_{\rm RW}$  is the angular velocity of the reaction wheel. By using Euler moment equations [20] [37].

$$\frac{\mathrm{d}}{\mathrm{d}t}H^b = -\omega^b(t) \times H^b(t) + \tau_{\mathrm{ext,dist}}(t)$$
(3.4.12)

where  $\tau_{\text{ext,dist}}$  encompasses all the external disturbance torques calculated in the prior section.

Eventually, joining expressions (3.4.2) (3.4.11) and (3.4.12),

$$\dot{H}^{b} = \left[\frac{dH}{dt}\right]^{b} = I\frac{\mathrm{d}\omega^{b}}{\mathrm{d}t} + H_{\mathrm{RW}}$$
(3.4.13)

$$I\frac{\mathrm{d}\omega^{b}}{\mathrm{d}t} = -\omega^{b}(t) \times \left(I\omega^{b} + H_{w}\right) + \tau_{\mathrm{ext,dist}}(t) - H_{\mathrm{RW}}$$
(3.4.14)

Since momentum exchange of the reaction wheel and the satellite is the same but in opposite direction  $\tau_{\rm sat}$ , the above expression yields

$$\dot{H}_{\rm RW} = -\tau_{\rm sat} \tag{3.4.15}$$

Finally, the derivative of the angular velocity can be derived

$$\frac{\mathrm{d}\omega^b}{\mathrm{d}t} = I^{-1} \left[ -\omega^b(t) \times \left( I\omega^b + H_{\mathrm{RW}} \right) + \tau_{\mathrm{ext,dist}}(t) + \tau_{\mathrm{sat}} \right]$$
(3.4.16)

Furthermore, the calculation of the time derivative of a quaternion q in terms of the quaternion itself and the corresponding angular rate is defined by Kuipers [24] as

$$\dot{q} = \frac{1}{2}\vec{\omega}' \otimes q = \frac{1}{2}\Omega'(\vec{\omega})q \tag{3.4.17}$$

$$\dot{q} = \frac{1}{2}q^* \otimes \vec{\omega}' = \frac{1}{2}\Xi(q)\vec{\omega}',$$
 (3.4.18)

where

$$\Omega'(\vec{\omega}) = \begin{bmatrix} 0 & \omega_z & -\omega_y & \omega_x \\ -\omega_z & 0 & \omega_x & \omega_y \\ \omega_y & -\omega_x & 0 & \omega_z \\ -\omega_x & -\omega_y & -\omega_z & 0 \end{bmatrix}, \\ \Xi(q) = \begin{bmatrix} q_4 & -q_3 & q_2 & q_1 \\ q_3 & q_4 & -q_1 & q_2 \\ -q_2 & q_1 & q_4 & q_3 \\ -q_1 & -q_2 & -q_3 & q_4 \end{bmatrix} \text{ and } \vec{\omega}' = \begin{bmatrix} \vec{\omega} \\ 0 \end{bmatrix} = \begin{bmatrix} \omega_x \\ \omega_y \\ \omega_z \\ 0 \end{bmatrix}$$
(3.4.19)

Based on prior results, the equation for calculating angular rates in terms of a rotation quaternion and its derivative is as follows:

$$\vec{\omega}' = 2\dot{q} \otimes q^* \tag{3.4.20}$$

Similarly, Kuipers proposes the second derivative of a quaternion with respect to time [24].

$$\ddot{\mathbf{q}} = \frac{1}{2} \left( \dot{\vec{\omega}}' \otimes q + \vec{\omega}' \otimes \dot{q} \right) \tag{3.4.21}$$

# Chapter 4

# Reaction wheel design

The core of the thesis resides in the design of a reaction wheel that is able to be fit into the 1U Cubesat. A major source of limitation is the size and mass specifications. A secondary aim of the design was to broaden the current knowledge of inertia wheels integration in CubeSats and reduce its interference with other components, as well as preserving a minimal weight and energy consumption.

Recapitulating, reaction wheels operate by accelerating an inertia wheel in one direction and thereby forcing the satellite to rotate in the other direction by conservation of angular momentum. Three reaction wheels can be mounted in an orthogonal system providing three-axis control of a spacecraft. However, according to PLATHON research Group's needs, the first approximation was to build a 1 DOF CubeSat, thus using only 1 reaction wheel.

## 4.1 Theoretical basis

Oland's [38] approach seem to be well-founded. The equations given are an accurate and comprehensive methodology for designing a Reaction wheel. In sizing wheels, two performance quantities must always be considered: angular momentum capacity, and torque authority.

The intended reaction wheel was designed to minimize the inertia when it begins to spin. In a few words, the design comprises a ring along the periphery of the wheel where most of the mass is placed, which reduces the total weight while maintaining a high level of momentum. Hence, as the size and mass are the most limiting factors, the reaction wheel can be decomposed into

$$m_{\rm RW,total} = m_{\rm RW,disk} + m_{\rm RW,ring} \tag{4.1.1}$$

The mass of the reaction wheel can be found by using the expressions for mass of a cylinder and a ring as

$$m_{\rm RW,total} = \rho \pi r_{\rm RW,disk}^2 h_{\rm RW,disk} + \rho \pi \left( r_{\rm RW,ring}^2 - r_{RW,disk}^2 h_{\rm RW,ring} \right)$$
(4.1.2)

where  $\rho$  is the density of the material and r and h are the radius and height, respectively.

Since the radius and height are constrained by the dimensions of the CubeSat, the only parameter that
seems to be left with freedom of choice is the material. However, just as the other parameters, the material selection is also constrained by the maximum weight of the CubeSat.

The proposed design values for the reaction wheels are illustrated in the Table below (check Appendix H for the drawing):

Part	Parameter	$\mathbf{Value}\;[\mathrm{mm}]$
Disk radius	$r_{\rm disk}$	60
Ring radius	$r_{ m ring}$	80
Disk height	$h_{ m disk}$	2.5
Ring height	$h_{ m Ring}$	7.5

Table 4.1 Reaction Wheel design parameters

These values were chosen according to the restrictions made by PLATHON's project requirements. To start with, the ring's maximum diameter shall not surpass 80 mm since the size of the PCB is a  $9.10 \times 9.30$  cm plate. In order to maximize inertia, the ring shall concentrate all the inertia in the reaction wheel.



Figure 4.1 Reaction wheel CAD design. Source: Own.

Some particularities of this design are holes in the centre of the disk. These four hollow spaces were left for operational purposes. When the ring is attached to the motor, and the motor is fixed in the PCB, if the motor happens to be damaged during the testing phase. Then, it is very simple to replace the motor for a new one unscrewing the fixation parts from the PCB (see Figures 6.6 and 6.8).

The weights of the other components can be added up to be 1 kg approximately (considering the worstcase scenario). First, the hollow part of the reaction wheel shall be calculated. This can be approximated as the difference between an outer circle and the inner circle.

$$r_{\text{hollow}} = r_{\text{outer,hollow}} - r_{\text{inner,hollow}} = 20 - 7.5 = 12.5 \text{ mm}$$

$$(4.1.3)$$

There are many materials that are suitable for the reaction wheel. The table below shows the main properties and differences between each (see Table 4.2).

	<b>ABS</b> [39]	<b>PLA</b> [39]	Aluminium [40]	<b>Iron</b> [41]
<b>Density</b> $[g/cm^3]$	1.04	1.24	2.70	7.87
Ultimate Strength [GPa]	0.04	0.065	68.3	212
Young's Modulus [GPa]	1.1 - 2.9	3.5	70.3	204
$\begin{array}{c} \mathbf{Maximum Service} \\ \mathbf{S} \end{array}$	98	52	127	267
Temperature				
$\mathbf{Melting \ point} \ [^\circ]$	105	88-128	450	1532
Coefficient of [K <sup>-1</sup> ]	$73.8 \cdot 10^{-6}$	$41 \cdot 10^{-6}$	$16 \cdot 10^{-6}$	$12 \cdot 10^{-6}$
Thermal Expansion				
Thermal [W/mK]	0.15	0.13	205	70
Conductivity [117]	0.20	0.20		
Other characteristics	Impact and heat resistant	High durability	Light and good conductor	High density and good conductor

**Table 4.2** Different materials properties. Source: Various [39] [40] [41] [42] [43] [44] [45].

The material selected must have high density values so the flywheel's momentum can beat the whole satellites' momentum. The only material available in the laboratory was either to use PLA or ABS with additive manufacturing. PLA was chosen as the material for the reaction wheel since it has higher density  $\rho$  of 1.24 g/cm<sup>3</sup>. In the future, for a 3DOF control, the CubeSat won't have enough space to fit the designed reaction wheels. Hence, other materials such as aluminium or iron should be considered to reduce volume.

The mass of the hollow is then,

$$m_{\text{hollow}} = \rho \pi r_{\text{hollow}}^2 \cdot 0.0025 = 1240 \cdot \pi \cdot 0.0125 = 1.522 \cdot 10^{-3} \text{ kg}$$
(4.1.4)

Thus, the mass of the disk then can be calculated as:

$$m_{\rm RW,disk} = \rho \pi r_{\rm RW,disk}^2 h_{\rm RW,disk} - m_{\rm hollow} = 1240 \cdot \pi \cdot 0.06 \cdot 0.0025 - 1.522 \cdot 10^{-3} = 0.0335 \text{ kg} \quad (4.1.5)$$
  
Now, the ring's mass yields,

$$m_{\rm RW,ring} = \rho \pi \left( r_{\rm RW,ring}^2 - r_{RW,disk}^2 h_{\rm RW,ring} \right) = 1240 \cdot \pi \left( 0.08^2 - 0.06^2 \right) \cdot 0.0075 = \boxed{0.0818 \text{ kg}} \quad (4.1.6)$$

Finally, recovering expression (4.1.1), maximum weight of a single reaction wheel is found,

$$m_{\rm RW,total} = m_{\rm RW,disk} + m_{\rm RW,ring} = 0.1153 \text{ kg}$$

$$(4.1.7)$$

The inertial momentum of the disk is given by the following equation:

$$I_{\rm RW,disk} = \frac{m_{\rm RW,disk} \ r_{\rm disk}^2}{2} = \frac{0.0335 \cdot 0.06^2}{2} = \boxed{6.03 \cdot 10^{-5} \ \rm{kg} \cdot \rm{m}^2}$$
(4.1.8)

The inertial momentum of the ring is given by the following equation:

$$I_{\rm RW,ring} = \frac{m_{\rm RW,ring} \ r_{\rm ring}^2}{2} = \frac{0.0818 \cdot (0.08^2 + 0.06^2)}{2} = \boxed{4.09 \cdot 10^{-4} \ \rm{kg} \cdot \rm{m}^2}$$
(4.1.9)

Finally, with the known design parameters, the inertial momentum of the reaction wheel can be deduced:

$$I_{RW} = I_{disk} + I_{ring} = 4.693 \cdot 10^{-4} \text{ kg} \cdot \text{m}^2$$
(4.1.10)

Since the motor used is the RF-300EA-1D390 which achieves an angular velocity of 3250 rpm, the maximum angular momentum this reaction wheel can achieve is deduced by the following expression (see Chapter 3.4.2):

$$H_{\rm RW} = \omega_{\rm motor} \cdot I_{\rm RW} = 3520 \cdot \frac{2\pi}{60} \cdot 4.693 \cdot 10^{-4} = \boxed{0.173 \text{ N} \cdot \text{m}}$$
(4.1.11)

Then, the maximum torque generated by the reaction wheel is found as (3.4.15) (without considering external disturbance torques):

$$H_{\rm RW} = \omega_{\rm motor} \cdot I_{\rm RW} \tag{4.1.12}$$

which means that the torque (i.e. the time derivative of the angular momentum) is

$$\tau_{RW} = \frac{\mathrm{d}I}{\mathrm{d}t}\omega + I_{\mathrm{RW}}\frac{\mathrm{d}\omega_{\mathrm{RW}}}{\mathrm{d}t} \tag{4.1.13}$$

Because the moment of inertia of a cylinder is  $\frac{1}{2}mr^2$ , it follows that

$$\frac{\mathrm{d}H_{RW}}{\mathrm{d}t} = m_{\mathrm{RW}} \cdot r_{\mathrm{ring}} \cdot v_{\mathrm{RW}} + I_{\mathrm{RW}} \frac{\mathrm{d}\omega_{\mathrm{RW}}}{\mathrm{d}t}$$
(4.1.14)

which reduces to

$$\tau_{\rm RW} = m_{\rm RW} \cdot r_{\rm ring} \cdot v_{\rm RW} + I_{\rm RW} \cdot \alpha = 0.1153 \cdot 0.08 \cdot 29.49 + 0 = 0.2720 \text{ N} \cdot \text{m}$$
(4.1.15)

considering the Cubesat is in stationary phase (no angular acceleration  $\alpha = 0$ ) and  $v_{\rm RW} = \omega_{\rm RW} r_{\rm ring} = 29.49 \text{ m/s}$ 

The total angular rotation  $\theta_{rot}$  produced by a reaction wheel within a given time

$$t_{\rm rotation} = \frac{2 \cdot I_{\rm max} \theta_{\rm rotation}}{h_{\rm max}} \tag{4.1.16}$$

This is,

Table 4.3 Angular rotation produced the reaction wheel. Source: Own.

Rotation angle $\theta_{\text{rotation}}$ [°]	Time of rotation $t_{\rm rotation}$ [s]
90	0.31
180	0.62
270	0.93
360	1.24

Later in Section 8, we'll see the results differ from these times since the motor has a transient phase before reaching the maximum rotation speed of 3520 rpm.

#### CHAPTER 4. REACTION WHEEL DESIGN

Another remark shall be noticed in the design of the Reaction Wheel design. Take a look of Figure 4.2. Notice there is a lateral hole. This lateral hole was added for a highly important issue. Whenever a piece is manufactured using additive manufacturing techniques, the final result not always is perfect and probably there might be some imperfections, i.e., leftover support material or leftover material inside drill holes. These imperfections are undesired since they will make the reaction wheel's axis to be shifted from the principal z-axis of the CubeSat's principal axis, hence, ending up with a motion lead by precession phenomena.



Figure 4.2 Reaction Wheel lateral hole. Source: Own.

The final printed pieces are shown in Figure 4.3,



Figure 4.3 3D printed pieces (Reaction wheel and supports). Source: Own.

# Chapter 5

# Attitude Determination and Control Subsystem

This chapter will revise all the different components that the ADCS board will have as well as showing the assembly process of each component.

# 5.1 Electrical Components

### 5.1.1 Microcontroller

Before diving into the different hardware available, the first thing one must choose is a **MicroController Unit**. A <u>MicroController Unit</u> (or MCU) is a compact integrated circuit designed to govern a specific operation in an embedded system. A typical microcontroller includes a processor, memory and input/output (I/O) peripherals on a single chip [46].

In simple terms, a microcontroller is similar to a computer. The difference is that a computer can perform several tasks at the same time, whereas a microcontroller, on the other hand, is typically dedicated to a single task. However, taking a deeper look at the basic components that make both the PC and a microcontroller it can be seen they share many similarities.

The fundamental components of a microcontroller are [47]:

• Central Processing Unit (or CPU): The CPU is the "brain" of the computer or microcontroller and is in charge of controlling everything that the computer or the microcontroller does. In other terms, the CPU is a microprocessor with a series of associated circuitry that controls the computer's software programs. The CPU extracts each program instruction from memory and executes that instruction. After completing the next instruction, the CPU moves on to the next and in most cases, it can operate on more than one instruction at the same time (depending on the number of cores and threads). This extraction and execution are repeated until all instructions are carried out.

- System Clock Speed: An essential component of any computer is its clock, this system controls the speed of the different operations the CPU handles. A microcontroller also has a system clock based on an oscillator. This is one of the essential features to consider when building projects. The oscillator's speed will determine how fast the microcontroller will run which can impact the resolution of the response or power consumption. Clock speed is measured in Hz.
- Memory: A microcontroller also has a memory, just like a PC has a hard drive (ROM) or RAM. The memory allows the microcontroller to store information so it can be used at a later time. Besides, just like a PC, a microcontroller's memory allows it to store programs to perform specific tasks. Memory capacity is measured in bytes (i.e. MB).
- **Peripherals**: These items are used to have bidirectional communication with the microcontroller. A mouse, keyboard or any sensor module are used to input information. For instance, a monitor ca the output of the system. Peripherals on a microcontroller are usually associated with a particular pin on a package.

The most common peripheral interfaces for microprocessors are USB, High-Speed Internet or UART. However, a microcontroller uses interfaces such as  $I^2C$ , SPI or UART.

Apart from the above, there are other supporting elements a microcontroller usually includes [46] [47] such as:

- Analog to Digital Converter (ADC): An ADC is a circuit that converts analogue signals to digital signals. This device allows the processor at the centre of the microcontroller to interface with external analogue devices, such as sensors. In other words, an ADC performs a conversion from analogue (continuous, infinitely variable) signals to digital (discrete-time, discrete-amplitude) signals. It does this by mapping the continuous set of values to a smaller (countable) set of values by rounding or interpolating. Thus, this conversion will always involve some noise.
- Digital to Analog Converter (DAC): A DAC performs the inverse function of an ADC and allows the processor at the centre of the microcontroller to communicate its outgoing signals to external analogue components.
- **System bus**: The system bus is the connective wire that links all components of the microcontroller together where data is sent and received.
- Serial port: The serial port is one example of an I/O port that allows the microcontroller to connect to external components. Data travels through the serial port.

There is a common misconception between microprocessors and microcontrollers. Briefly, microcontrollers can be said to function on their own, with a direct connection to sensors and actuators and incorporating RAM and ROM memories as well as the CPU in a single place. Whereas microprocessors are designed to maximise the compute power on the chip with internal bus connections to support hardware such as RAM and serial ports (rather than direct I/O) [46].

Parameters	Microprocessors	Microcontrollers	
C+	Hard Disk	Flash memory	
Storage	(128  GB to up to  2  TB)	(32  kB up to  2  MB)	
Internal	Memory and I/O devices	CPU, memory and IO devices	
Structure	connected externally	connected internally	
Clock	High clock speed	Lower clock speed	
Speed	(1  GHz to  4  GHz)	(1  MHz up to  300  MHz)	
Memory	519 MP to 29 CP	2 hP up to 256 hP	
(RAM)	512 MD to 52 GD	2 KB up to 250 KB	
Power	Uigh	Low	
Consumption	mgn	Low	
Cost	High	Low	
Bit	It is available in 32-Bit	It is available in 8-bit,	
Size	and 64-bit.	16-bit, and 36-bit.	
Peripheral	USB, High Speed Internet,	I <sup>2</sup> C SDI UAPT et a	
Interfaces	UART, etc.	1 C, SF I, UARI, etc.	
Applications	Gaming, web browsing,	Specific tasks	
Applications	document writing, etc.	(i.e. camera, washing machine, etc.)	

 Table 5.1 Comparison between Microprocessors and Microcontrollers.

 Source:
 Components101 [48].

There are different microcontrollers and microprocessors in the market. The most famous ones are the Raspberry Pi microprocessor, Arduino microcontroller and STM32 microcontroller. These devices are widely used for different purposes.



Figure 5.1 Microcontrollers and Microprocessors. Source: Arduino [49], Raspberry Pi [50], STM32 [51].

When selecting the microcontroller for the project, several aspects were taken into account. First and foremost, the microprocessor must fit into a PCB and leave room for other components. Even though the Raspberry Pi is by far the most powerful option, the size and weight of both the Arduino Nano and

Bluepill are most suitable for the CubeSat. Therefore, the comparison was restricted to these two boards.

The Arduino Nano is one of the most popular microcontrollers on the market. It offers a large amount of flexibility and purposes. Since this is from Arduino's family, it uses the Arduino IDE, one of the comprehensive open-source interfaces for microcontroller programming. The built-in microprocessor is the ATmega328, which is highly powerful for the initial phase of the project (see Appendix B for more detailed information).

Conversely, on the other hand, the Bluepill microcontroller shares a similar dimension to the Arduino Nano board. This microcontroller, though, has a slightly bigger dimension than the Arduino Nano board but offers faster clock speed numbers, RAM and flash memory. Since ST-Microelectronics is the manufacturer of this board it uses its own interface. Nevertheless, this board can be programmed to work with the Arduino IDE using an ST-Link device and a custom library (see Appendix B).

While the Arduino is powered via 5 V, thus, it will need a logic level converter to convert the 3.3 V to 5 V and vice-versa to work with the motor driver as this latter one functions with 3.3 V.

Besides, it is important to highlight the usage of the pins. Since the board will be connected to multiple devices at the same time it must ensure to have a sufficient amount of pins to manage the communication with the On-Board Computer and also the actuators. Later, in section 5.1.3 it is shown how the drivers use  $I^2C$  protocols and also the communication with the OBC. To avoid disturbances and perturbations, it is highly recommended to use different channels for communication and actuators. As seen in Figure B.2, there are multiple  $I^2C$  pins with different channels. In the case of the Inertial Measurement Unit, while the MPU9250 can make use of the SPI protocol and  $I^2C$ , the BNO055 is limited to work only with  $I^2C$ .

Finally, balancing the advantages and disadvantages of both boards, the STM32 Bluepill (see Figure 5.1 (c)). microcontroller suits best for the needs of the project. It has more communication pins as well as faster clock speed, namely, faster computing speeds. This microcontroller has been programmed using Arduino IDE with ST-Link device (check Appendix B to see how it has been programmed.)

	Arduino Nano	Bluepill
Manufacturer	Arduino	ST-Microelectronics
Microcontroller	ATmega328	STM32F103C8T6
Clock Speed	$16 \mathrm{~MHz}$	32  MHz(72MHzmax)
Flash Memory	32  kB	64  kB
SRAM	2  kB	$20 \mathrm{kB}$
Operating Voltage	5V	3.3 V
Analog IN Pins	8	10
Digital I/O Pins	$22\ (6\ have\ pwm$	29 (10  have pwm)
I2C buses	1	2 + 1 (I2C2)
SPI buses	1	3
Size	$18\times45~\mathrm{mm}$	$23 \times 53 \mathrm{mm}$
Weight	$7\mathrm{g}$	$20 \mathrm{~g}$

Table 5.2 Comparison between Arduino and Bluepill. Source: Arduino and Bluepill [49] [51]

# 5.1.2 Inertial Measurement Unit

An Inertial Measurement Unit is a device in charge of obtaining the attitude data for the ADCS. As shown in section 2.2.2, several devices are used to retrieve the attitude data. Moreover, these devices are selected depending on the needs in each situation and missions usually holds more than one attitude determination instrument for precision purposes.

The IMU is perhaps the core of the ADCS subsystem. It contains sensors such as accelerometers, magnetometers, gyroscopes and even thermometers. To increase the reliable output movement, IMUs collects data from numerous distinct sensor:

- Accelerometer: A device that monitors velocity and acceleration.
- Gyroscope: determines rotation and rotational rate.
- Magnetometer: identifies cardinal directions (directional heading).

However, despite the corrections and all the sensors, IMU's have a majors disadvantage that has to be taken into account. The main downside of an IMU is that it is prone to error accumulation over time, which is also known as often known as "drift" (check tests in Section 8). This happens because the IMU is continually monitoring changes relative to itself (rather than triangulating against an absolute or known outside device), it continuously rounds off tiny fractions in its computations, which add up over time. If left unaddressed, these minor inaccuracies might compound up to large mistakes. To address this problem, one way of solving it is to use other attitude instruments as a complement to constantly correct the tiny errors that may appear.

Two different IMUs were studied for PLATHON's project, these two IMUs are: the BNO055 and the GY-9250 module which incorporates the sensor MPU9250. To analyze them, magnetorquer's team studied MPU9250 while the reaction wheels team performed an analysis with BNO05.



(a) GY-9250 module with MPU-9250 sensor. Source: Components101 [52].



(b) BNO055 module sensor. Source: Adafruit [53].

Figure 5.2 Inertial Measurement Units. Source: [52] [53].

The first IMU, BNO055 from Adafruit was selected for the capabilities it provided. This sensor has 9-DOF and offers a fusion technique that combines accelerometer, magnetometer, and gyroscope data

to provide consistent three-axis orientation output. By combining a Micro Electrical-Mechanical System (MEMS) accelerometer, magnetometer, and gyroscope on a single built-in die with a high-speed ARM Cortex-M0 processor to process all sensor data, the abstraction of the information in quaternions, Euler angles or vectors are just extremely fast and accurate. One small downside of the BNO055 relies upon the fact it only operates with I<sup>2</sup>C communication protocols, which later in the project will the reason to shift to an alternative IMU that has SPI interface.

As the project moved on, the team found that there was a crucial problem with respect to the drivers and microcontroller. To summarize, the problem was due to the coupling of channels on the  $I^2C$  address of the IMU and the drivers. Several tests were done to find out where the problem resided. In the end, the team opted for another IMU with SPI interface.

The second tested IMU was the GY-9250 module. Similar to the prior one, the MPU9250 sensor has 9-DOF and both SPI and I<sup>2</sup>C protocols. Just as the BNO055, this IMU has a built-in Digital Motion Processor (DMP) which makes the IMU able to make calculations and corrections and the measurements instantaneously (check Appendix C for details).

# 5.1.3 Motor Driver

The Cubesats need to generate rotational inertial to orientate in space. As seen before, one way of doing such control is using reaction wheels. Nevertheless, reactions wheels alone are just a mass with a certain amount of inertia and thus, it does not have the capacity of generating a moment of inertia. To do so, it is needed a motor driver.

A motor driver is a device that is used to control the power and therefore, the velocity of a motor. The principle behind a motor driver can be simplified as a device that turns a low current signal into a high signal for driving the motor. The main reason why is that motors cannot be connected directly to a microcontroller pin is that they do not provide sufficient power from those pins. For instance, an Arduino nano standard 3.3 V or 5, 5 V pin can output only a maximum of 50 mA of current [54] which is insufficient for most of the motors that are used in different applications, especially, on startup and stall currents. Hence, the main goal is to drive that larger load from the small signal of the microcontroller and amplify that to be able to control a much larger load. The motor driver acts like a gateway where it accepts this small input and can control larger loads using a special type of transistors called MOSFET (*Metal–Oxide Silicon Transistor*). The MOSFET transistor acts as a switch, when a current is applied onto the base connector, the transistor lets the current flow through it. In fact, MOSFETs are usually integrated within the motor driver.

It's important to remark that when referring to motor drivers, the motor itself must be taken into account. A brushed DC motor does not require any control or timing whereas servo motors require a specific type of signal to operate correctly [55].

The correct way of controlling a motor is to connect the motor driver to the motor first and then supply power to the motors using an external battery through the motor driver. Besides, the most common form of speed control systems is the **H-bridge**.

As a brushed motor always rotates in one direction, it is mandatory to provide the motor with the ability

to rotate both in a clockwise direction and counterclockwise direction. Precisely, an H-bridge system configuration permits the motor to rotate in either direction. The motor driver selected for this project uses this configuration (see Figure 5.3).



Figure 5.3 H-bridge configuration schematic. Source: Core Electronics [55]

The electrical schematic (Figure 5.3) shows an H-bridge with two pairs of MOSFETs to control the flow of the electrical current through the motors. Consequently, switching on and off a specific combination of MOSFETs will permit the rotor to rotate in the desired direction. Notice how two pairs of diodes are added as well, these diodes restricts the flow in a certain direction and are commonly added for safety issues to protect against transient current and voltages produced by the motor as it turns. Depending on the application, the H-bridge can have another set of variants with more components. In the Figure above, the additional resistors were left out to make it easier to understand.



(a) H-bridge configuration schematic in clockwise direction. Source: Core Electronics [55]

(b) H-bridge configuration schematic for counterclockwise direction. Source: Core Electronics [55]

Figure 5.4 H-bridge clockwise and counter clockwise rotation direction. Source: Core Electronics [55]

Let's look at Figures 5.4a and 5.4b, the motor is placed in the middle and the flows of the current through the motor takes one way or other depending on the MOSFETs. In the first state, the motor's positive potential is located on his right side and the lower potential is located on the left side. On the other hand, the second scenario depicts the same flow but has interchanged the positive and negative connectors.

There is another configuration that is useful if the user desires the motor to stop rotating. Despite turning off the motor from the power, there is a remaining momentum of the rotor that prevents it to stop immediately. so it will gradually come to a complete stop. Thus, to perform a complete stop and set the rotor in a static position it is needed to brake the rotor. By applying an equal voltage to both sides of the rotor it will make sure the rotor stops its motion instantly.





One important use-case that should be avoided at all times is shown in Figure 5.5b which will generate a direct path from voltage supply to ground. Not only this will shortcircuit out the circuit but also damage other peripherals connected to it. The same reasoning applies to the other side.

The driver that was chosen for the project was the SparkFun Qwiic Rob-15451 motor driver [56] which has an H-bridge configuration and suits the requirements of the project with its dimensions, weights and the communication protocol it has which can be either  $I^2C$  or SPI.



Figure 5.6 SparkFun Qwiic Motor driver ROB 15451. Source: SparkFun [57].

The Sparkfun Qwiic motor driver will use  $I^2C$  communication protocol which translates to a fast, simple, high speed and cost-effective protocol for communication [58]. Furthermore, taking into account the PCB that must be less than  $10 \times 10$  cm, its size  $(2.5 \times 2.5 \text{ cm})$  and weight fits perfectly in the PCB with all the other components. It allows using two  $I^2C$  connections via two ports at the sides simultaneously, meaning that various drivers can be connected one to another in the same  $I^2C$  buses. Each driver can connect 2 devices, which translates that for a full 3-axis control of a CubeSat, it would be needed 4 drivers (3 magnetorquers and 3 reaction wheels). The driver's address can be changed by changing a pin (see Appendix G). The output voltage ranges from 3 to 11 V, and powered at 3.3 V. The size of these drivers is relatively small, measuring  $1 \times 1$  in  $(25.4 \times 24.5 \text{ mm})$  which gives more versatility, reduces weight and space.

As regards the connection of the motor to the driver, the motor will be connected directly to the two outputs it has. On the other hand, the control signal will be connected to a Qwiic connector located on the side, so no soldering is required to connect it to the rest of the system. The Qwiic connector part of the Sparkfun ecosystem and gathers the needed pins for both  $I^2C$  and SPI protocols. As shown later, this reduces the amount of required PCB space.

Among other properties it has, it is notable that this motor driver supports 1.2 A of steady-state drive per channel with 1.5 A peak. Since the driver is a 3.3 V logic device, is compulsory to use a logic level converter to interface to 5 V. Also, the I<sup>2</sup>C address of the Qwiic Motor Driver is default to 0x5D, however, the address can effortlessly be changed with its jumper to another address (please refer to Appendix G for more information about the documentation of the motor and the connections).

#### 5.1.4 Motor

The motor is the fundamental piece for the project. The motor has to be able to generate a sufficient amount of torque to rotate the entire CubeSat and also offset the external disturbance torques.

The motor selection must agree with the type of driver it has to function with. There are two types of motors in the market: Brushed DC motors (BDC) and Brushless DC motor (BLDC):

- Brushed DC motor: The armature of a brushed DC motor is a structure of coiled wire coils that acts as a two-pole electromagnet. The commutator, a mechanical rotary switch, reverses the directionality of the current twice every cycle. As a result, the electromagnet's poles pull and push against the permanent magnets along the exterior of the motor, allowing current to flow through the armature. The commutator then switches the polarity of the electromagnet in the armature as its poles cross the poles of the permanent magnets.
- Brushless DC motor: This type of motors, by contrast, has a permanent magnet as its exterior rotor. Furthermore, it employs three stages of driving coils as well as a specific sensor that measures rotor position. The sensor delivers reference signals to the controller as it tracks the rotor position. In turn, the controller systematically activates the coils one phase after the other.

Overall, both types of motors have their advantages and disadvantages. The advantages of the brushed motors are that they are more economic, simple and uses inexpensive controllers whereas the major downside is that they are usually less efficient than brushless motors as constant switching action with the commutators creates electromagnetic noise and they wear out due to the perpetual physical contact with the shaft. However, in accordance with the requirements set for this project, a BDC motor is sufficient. In the next generation of the project, a BLDC will be used.

Considerations and priority on the motor research:

- **Cost**. The primary goal of the CubeSat is to be as economic as possible, so the cost motor shall be reduced.
- Efficiency / Low consumption: Low consumption is essential in space missions. When in standby mode, the motor's consumption must be as reduced as possible. This translates into better performance and saves energy in case of occlusion.
- **Dimensions**. The sizing of the motor is a key factor. Although the vast types of CubeSats, the motor has to be as flat as possible and compact. The motor must leave enough space for other components in the PCB.
- **Great torque**. The torque this motor must supply shall be enough to rotate the entire CubeSat when loaded with all the components.
- Lifespan. For space missions, the lifespan of a CubeSat is rather small. However, the motor needs to be fully operative during the time of the mission and even more than the missions lifespan. Thereby, it shall not require any maintenance.
- Great range of velocities. This one has to be fulfilled otherwise it would limit CubeSat's rotation. The faster it spins, the less time it spends on manoeuvres, but it will more energy consumption in a shorter period of time. It shall have notorious responses and the intention is to reduce the control time. A greater range of velocities leaves more margin of manoeuvrability.

Two motors were analysed for the project, the first one is the Mabuchi RF-500TB-14415 and the second motor is the Mabuchi RF-300EA-1D390. The RF-500TB-14415 was studied during the initial state of the project.

The Mabuchi RF-500TB-14415 [59] was bigger in size with a diameter of  $\emptyset$  32 × 19.5 mm and a mass of 45 g. Its maximum efficiency is around 55% depending on the torque, which is expected for brushed motors. In terms of speed, it is able to rotate at incredibly high revolutions per minute, achieving a maximum of 3100 rpm without loads. Although this motor had a great amount of torque, its size was far larger than the limits of the PCB, which restricted the amount of maximum diameter available for the motor.



Figure 5.7 Mabuchi RF-500TB-14415 (left) vs Mabuchi RF-300EA-1D390 (right). Source: Own.

Next motor analysed was the Mabuchi RF-300EA-1D390 [60]. This second motor solves the size problem of the prior motor and possesses high speeds up to 4400 rpm with lower nominal voltage consumption of 3.9 V rather than 5 V. However, due to the smaller size restriction, the amount of torque it can generate also is affected and reduced to 0.47 mN  $\cdot$  m. The size of this motor is  $\emptyset$  24 × 12.8 mm and only 22 g of mass, half the mass of the prior one. As regards the performance, the efficiency remains practically the same at around 52%. Eventually, as the bench tests demonstrated great results, this motor was selected to work with the driver.

	RF-500TB-14415	RF-300EA-1D390
Voltage Range	$1.5 - 9.0 { m V}$	2.8 - 7.0 V
Nominal Voltage	5 V	$3.9\mathrm{V}$
No Load Speed	$3100 \mathrm{rpm}$	$4400~\mathrm{rpm}$
No Load Current	26  mA	$21 \mathrm{mA}$
Maximum Efficiency	55.0%	52.5%
Max Efficiency Speed	2540rpm	3100rpm
Max Efficiency Current	120  mA	84  mA
Max Efficiency Torque	$1.23 \text{ mN} \cdot \text{m}$	$0.47~\mathrm{mN}\cdot\mathrm{m}$
Stall Torque	$6.86 \text{ mN} \cdot \text{m}$	$2.35~\mathrm{mN}\cdot\mathrm{m}$
Stall Current	54  mA	34  mA
Size	$\varnothing$ 32 × 19.5mm	$\varnothing$ 24.4 $\times$ 12.8 mm
Weight	$45 \mathrm{g}$	$22 \mathrm{g}$

Table 5.3 Characteristics of the two motors analyzed for the project. Source: Mabuchi [59] [60].

Hence, the Mabuchi RF-300EA-1D390 was selected for driving up the reaction wheel. In section 5.2.3 is explained how to connect the motor to the microcontrollers.

## 5.1.5 Logic Level Converter

Most devices are designed to be powered at 5 V. Albeit, the newest components use the newest standards (5 V). The main advantage of it is that as the voltage is higher, noise susceptibility will be significantly lower since higher noise levels are needed to disturb the 5 V. The major downside is power consumption. See Appendix D).

The bi-directional logic level converter is a small device that securely levels down 5V signals to 3.3 V while simultaneously stepping up 3.3 V to 5 V which allows you to adapt the SPI, I<sup>2</sup>C, UART or any digital signal. The board needs to be powered by the two voltages to be converted and only works with digital signals (high and low levels). This model does NOT work with analogue signals.



Figure 5.8 Logic level converter. Source: Naylamp Mechatronics SAC [61].



Figure 5.9 Logic level converter schematic. Source: Naylamp Mechatronics SAC [61].

Despite Arduino Nano operates with 5 V supply, the Bluepill still operates with a 3.3 V supply. When connecting the STM32 Bluepill via USB, the USB outputs 5 V voltage. One could ask how is it possible for the STM32 Bluepill to output a 5 V supply, this is due to the fact that it has an internal logic level converter that controls the entire microcontroller.

# 5.1.6 Bluetooth Module

The next step is to establish wireless communication. Thus, this will be extremely useful for sending data without the need of a physical wire connection to the OBC as the final step is to test it on an air bearing and send it to outer space.

The fundamental advantage of satellite communication is that it gives communication services to any location on the planet. Besides, satellites are not affected by distance. Nonetheless, the following issues are related to satellites.

There are several methods for accomplishing wireless communication for testing purposes in the ground station, such as WiFi (Wireless Fidelity), IR (Infrared), microwave or mobile communications such as 4G or 5G. However, the design of the communication using high band antennas is out of the scope of this project. Furthermore, there are many problems related to satellite communication, including high propagation delays, poor bandwidths compared to terrestrial media, and noise due to the effect of rain and atmospheric disturbances.

Instead, to get wireless communication working, the team has decided to take advantage of Bluetooth communication, since wireless communication was a requirement of the project. The Bluetooth module HC-06 from DSD Tech will be used for telemetry and telecommunication.

The HC-06 module allows any Bluetooth device, such as a computer or a smartphone, to connect wirelessly to the Arduino Nano or the STM32 Bluepill. It has a 3.3 V supply voltage, though it is usually equipped with a regulator that allows for a supply voltage of 3.6 to 6 V.

The Bluetooth module HC-06 has 4 pins, 2 for power and 2 to establish a connection. Project PLATHON's PCB is designed to work with either the HC-06 or an HC-05. The only difference between both modules is that the HC-05 can be used as a slave module (equivalent to HC-06) and as a master module which means that it can pair itself to another device on its own. Due to this fact, the HC-05 have 2 more pins, one that enters to configuration mode and the other one to connection state (check Appendix E for a detailed data-sheet and configuration).

The communication protocol of the HC-06 is UART (Universal Asynchronous Receiver/Transmitter) which uses an asynchronous serial transmission at a variable speed. This means that unlike  $I^2C$  for SPI, there is no clock signal to synchronize the transmitting device's output bits to the receiving end.

In Bluetooth communications, there are two types of roles: host and slave. This module is a slave, which means it waits for the host to issue the two execute command. The host, on the other hand, selects the slave with whom it wishes to communicate. Finally, the host/slave, which can play both roles depending on its configuration in the code, combines these two roles.

The module is factory setup as a "Slave" device. However, using an HC-05 which the same as the HC-06 but with the added ability to act as a master as well. Since the laboratory have both of them, the PCB was designed to support both devices as they share the same pins.

# 5.2 ADCS board assembly

Once reviewed all the electrical components, the following section intends to show how to connect every component to the microcontroller and prepare it for testing both the hardware and control software before assembling all the pieces into a custom PCB as well as the code for its correct setup and configuration.

As explained in prior sections, the microcontroller used for this part is the Bluepill, however, several tests were performed using the Arduino board. A lot of component and code testing was done during the initial stage of the project. There are two ADCS teams (Reaction wheels team and Magnetorquer team) within the ADCS Subsystem. Both teams analysed different components and once the component testing phase was settled, the global team decided which component to use. Figure 5.10 presents the diagram of all the components the ADCS board must have and the communication protocols.



Figure 5.10 ADCS Board diagram. Source: Own.

# 5.2.1 The Inertial Measurement Unit

#### **BNO055 IMU** assembly

At the initial stage, the IMU tested was the BNO055 from Adafruit. This IMU was first considered since the characteristics of the Arduino board and the rest of the components have not been dealt with in-depth yet. Figure J.1 and L.1 shows the Arduino board and the BNO055 IMU.



Figure 5.11 Schematic of the BNO055 and the Arduino Nano connections. Source: Own.

As for the connections, the BNO055 uses  $I^2C$  communication protocol and A4 and A5 pins of the Arduino board supports this communication. The synchronization and data transmission are in separate pins, where SDA is the line for the master and slave to send and receive data and SCL refers to the serial clock port that carries the clock signal.

Arduino Nano	BNO055 IMU
5 V	Vin
GND	GND
A4	SDA
A5	SCL

Table 5.4 Arduino Nano and BNO055 IMU connections. Source: Own.

 $I^2C$ , like SPI, is synchronous, which means that the output of bits is synced to the sampling of bits by a clock signal shared by the master and slave. The master is always in charge of the clock signal.

Firstly, the IMU must be calibrated and configured (see Appendix L). Once the IMU is configured and calibrated, the next step is to get all the raw data from all the sensors, namely, accelerations in all three axes (accelerometer), magnetic field measurements in all three axes (magnetometer) and angular velocity (gyroscope). Note that these measurements are in the satellites body reference frame (see 3.2.1).

After this, one can proceed to calculate the Euler angles from the raw data of the sensors. Nevertheless, the IMU itself has also a built-in Digital Motion Processor (DMP) which can provide the Euler angles and even quaternions directly without any computation. The complete code for this test is detailed later in Section 8.

#### MPU9250 IMU assembly

The connections for the MPU9250 follows the same procedure, however, this time, an SPI communication protocol has been chosen since it is able to transfer data at faster rates. However, it is also possible to use the I<sup>2</sup>C protocol. In this case, the MPU is connected via the 3.3 V. The data transfer and synchronization are done by SPI serial data ports where MOSI and MISO are the pins for sending and receiving data, SS enables the transmission data transmission circuit, SCK serves as the serial clock and finally, FSYNC is the pin for digital input frame synchronization. Below is shown the schematic and connections of the Arduino board and the MPU9250.



Figure 5.12 Schematic of the MPU9250 and the Arduino Nano connections. Source: Own.

The connections are summarized in the table below 5.5:

Arduino Nano	BNO055 IMU
3.3 V	VCC
GND	GND
D10 (SS)	NCS
D11 (MOSI)	SDA
D12 (MISO)	AD0
D13 (SCK)	SCL
FSYNC (SS)	GND

Table 5.5 Arduino Nano and MPU9250 connections. Source: Own.

Analogously, the microcontroller used at the end of the project was the STM32 Bluepill. The connections remain the same but the pins for the data transmission protocol are different.





Eventually, the pin connections is outlined in the table below 5.6:

STM32 Bluepill	BNO055 IMU
3.3 V	VCC
GND	GND
A4 (SS)	NCS
A5 $(SCK)$	SCL
A6 (MISO)	ADO
A7 $(MOSI)$	SDA
FSYNC	GND

Table 5.6 Arduino Nano and MPU9250 connections. Source: Own.

# 5.2.2 Bluetooth module assembly

Bluetooth communication technology is designed as a short-range connectivity solution for personal, portable and handheld electronic devices such as smartphones, laptops and smartwatches [62]. This technology supports both synchronous and asynchronous data flow transmission over links with speeds of 1 Mb/s. It operates in the Industrial, Scientific and Medical (ISM) radio band spectrum of 2.4 GHz utilizing low transmit power radios and Frequency-Hopping Spread-Spectrum technique (FHSS). The FHSS is a signal modulation technique that involves that a signal generated with a particular bandwidth is deliberately spread within the frequency band thus reducing interference and jamming and most importantly, preventing detection or interception. This communication protocol functioning is illustrated in Figure 5.14,



Figure 5.14 Frequency-Hopping Spread-Spectrum diagram. Source: Sharda University [63].

The HC06 Bluetooth module uses UART communication protocol, thus, the setup of this module is done by sending AT commands (Attention Commands). AT commands mode is used to modify the Bluetooth module's default settings. To alter the BT device name, device role such as master or slave, the password of the device BT module must be established in AT Command mode, and the default settings can be changed. This configuration code can be found on E.1.



Figure 5.15 Bluetooth setup schematic with STM32 Bluepill. Source: Own.

It is important to mention that the HC06 is connected to a 5 V output pin from the Bluepill. Depending on the manufacturer, they mention using a 3.3 V pin to supply power to the module. Hence, as regards the Bluepill microcontroller is as simple as switching to a 3.3 V output pin. Nevertheless, when using an Arduino Nano microcontroller, one has to implement a voltage divider since there is no 3.3 V output pin. Even so, the one used for the project is capable of working at 5 V.

Figure J.4, schematic 5.15 and the Table 5.7 illustrate how to connect the pins with the Bluepill microcontroller:

STM32 Blue Pill	HC06 Bluetooth module
3.3 V	VCC
GND	GND
A9	RXD
A10	TXD

 Table 5.7 STM32 Bluepill and HC06 Bluetooth module connections. Source: Own.

To connect with the STM32, UART1 interface has been used where pins A9 (TXD) and A10 (RXD) is connected to the Bluetooth module's RXD and TXD pins, respectively. Once everything is set up, the next step is to send the data from the Bluepill to any device that supports Bluetooth.

### 5.2.3 Motor and Motor Driver assembly

To assembly both the motor and motor driver, the diagram and schematic below present the connections to the STM32 Bluepill (see Figures J.5 and 5.16).

The motor driver, as seen previously, has the role of powering the motor since the output pins of the microcontroller do not have enough range of voltage to use the motor at its full operational range potential. The testing control code can be found in G.8. For both IMU configurations (SPI and  $I^2C$ ) the code is very similar as the only parameters changed are from the IMU readings.

The driver is connected to the Bluepill microcontroller via  $I^2C$  port. However, Sparkfun has developed its own pin for this type of connection, the Qwiic cable provides the needed functionalities of  $I^2C$  but reduces the space needed for a PCB assembly (check Appendix G for more details).

Diagram J.5 corresponds to 5.16 in which the IMU is connected via SPI protocol.



Figure 5.16 Schematic of the motor driver and motor and the STM32 Bluepill connections with SPI. Source: Own.

The connections are (check Appendix G for Qwiic cable pinout):

**Table 5.8** STM32 Bluepill and Qwiic motor driver. Source: Own.

STM32 Blue Pill	Qwiic motor driver
3.3 V	VCC
GND	GND
3.3	3.3 V
B6	SDA
B7	SCL

Other connections left of this driver is the motor, which can be connected to either A1 and A2 pins or B1 and B2 pins of the driver. This means that more than one motor can be used at the same time as well as two possible configurations can be used depending on the distribution of the PCB. In order to change the configurations is as simple as setting either a 0 or a 1 in the setDrive() function. The first parameter of this function represents the peripheral, the second one is also a Boolean which represents the direction of the rotation and the last parameter is an integer that ranges from [0, 255] which is the amount of power given to the motor. Finally, an external battery connected to VIN and GND will provide sufficient energy to move the respective motor. In contrast to the Arduino Nano microcontroller, where a DC-DC converter is necessary, for the Bluepill there is no DC-DC converter since the output voltage pin 3.3 V is the same as the necessary voltage of the driver.

To use the driver, a set of libraries must be installed which are provided by Sparkfun itself. These libraries are SCMD.h and SCMD\_config.h. Next, the driver object is created. Following this step, the  $I^2C$  address is set for the communication to finally enable the driver.

Analogously, if the user wants to connect the IMU using  $I^2C$ , the following schematic shall be considered 5.17:



Figure 5.17 Schematic of the motor driver and motor and the STM32 Bluepill connections with  $I^2C$ . Source: Own.

# 5.2.4 Full assembly

The following diagram shows the final assembly of the components into a breadboard. This information will be used to design the custom PCB that holds all the following components:

- $1 \times$  STM32 Bluepill controller
- $1 \times$  HC06 Bluetooth module
- $1 \times$  Mabuchi motor RF-300EA-1D390
- 1× Sparkfun Qwiic motor driver
- $1 \times$  External battery



Figure 5.18 Final assembly schematic. Source: Own.



 ${\bf Figure \ 5.19} \ {\rm Final \ assembly \ physical \ assembly. \ Source: \ Own.}$ 

Next chapter 6 shows the design of the PCB. This PCB will not only take into account the components for the reaction wheel but also the pair of magnetorquers designed by the magnetorquer team.

# Chapter 6

# PCB Design

In the following chapter, it is shown the design of the Printed Circuit Board (PCB) and the circuits that will allow the communication.

A printed circuit board, or PCB, is a type of circuit board that is used to physically support and electrically link electronic components by connecting conductive paths, tracks, or signal traces from copper sheets bonded onto a non-conductive substrate. The PCB is specifically designed by PLATHON's project electronics requirements. For a more in-depth overview of the schematics see Appendix I.

# 6.1 Schematic

Prior to designing the schematic, several CubeSat prototypes were already available within PLATHON's project (see Figures 8.3). The particularity of the 8.3 Cubesat prototype resides in its ability to access the PCB with ease, namely, this CubeSat can be classified as a modular 1U Cubesat. For instance, whenever the user is willing to change the PCB or make some rearrangements the process is seamless.

This Cubesat complies with the structural restrictions listed in A. Its dimensions is  $9.1 \times 9.3$  cm and the components chosen for the project has already been thought to fit inside the PCB.



Figure 6.1 PCB layout top view (left) and bottom view (right). Source: Own.

Notice how the PCB is not fully squared and a slight material is missing in its sides. The reason behind this design was to enable the power cables to connect from the solar panel cells to the main board and supply power to the system.

The PCB is designed following the corresponding technical specifications. Each pin is connected via copper paths with a separation width of 0.3 mm and a path width of 0.6 mm. This model was chosen because it is the most feasible and practical ways to avoid potential errors, i.e. when paths are too close there might be some interactions between pins that can cause short circuits.

Additionally, as regards the PCB, it consists of 4 layers in which all connections are settled on side of the board. This design facilitates the welding process of the different components such as the motor and the IMU.



(a) PCB Gerber file. Source: Own.

(b) PCB Excellon file. Source: Own.

Figure 6.2 PCB Gerber and Excellon files. Source: Own.

*Gerber* files are ASCII vector files that provide information about each physical board layer of the PCB design. This include, for instance, copper paths, solder mask and silkscreen graphics and they are all represented by a drawing code and specified by a sequence of vector coordinates. *Excellon* drill format is a file format used for drilling and routing equipment and it is an industry standard. Both files are necessary for manufacturing the PCB. These files were sent to *JLC PCB*, a company specialized in manufacturing all kind of PCBs for different needs and applications.



(a) Physical PCB (front). Source: Own.



(b) Physical PCB (back). Source: Own.

Figure 6.3 Physical PCB. Source: Own.

The PCB is made out of an FR-4 sandwich layer between two thin layers of copper lamination. FR-4 is by far the most used material when it comes to manufacturing PCBs. The letter "FR' denotes that the substance is Flame Retardant, and the number "4" indicates that it is woven glass-reinforced epoxy resin. Additional prepreg layers are sandwiched between the central core and the top and bottom copper layers in this board. FR-4 boards offer good strength and water resistance characteristics that allow them to be used as an insulator in many electrical applications. It insulates neighbour planes and provides bending strength to the structure.

Before engaging in welding the components, several tests were performed to ensure the paths are correctly connected and there are no short circuits.

Two design errors were found during this phase. The first one is that all the system is connected to a 5 V input from an external battery and despite there is a 9.6 V input pin (see Figure 6.1) that pin is connected as well in the 5 V paths and not to the VIN input for the drivers. However, after performing some tests with 5 V. The results shows it is enough to rotate the entire satellite even when using 20% of the supplied voltage (i.e. 20% of 5 V (see 8)).

The second design error resides in the VIN and GND pins of the drivers. These pins are switched. However, a practical solution proposed was to connect the pins with a short cable that connects the pins instead of placing the drivers on the other side of the board as there might be some friction with the reaction wheel.

# 6.2 Final Design

Eventually, KiCad software enables to render the pieces in a 3D view and the final assembly of the components shall look like as Figure 6.4:



Figure 6.4 Final PCB assembly render. Source: Own.

In Figure 6.4 there is the Reaction wheel on the backside of the board and it carries two magnetorquers at the side of the board. The vertical module in front of the reaction wheel is the Bluetooth module. Also, the two red pieces in the lower right side of the figure are the two set of drivers that controls both the magnetorquers and the reaction wheels.

Finally, Figures 6.5 and 6.6 shows the top and bottom sides of the assembled PCB. Figures 6.7 6.8 shows the physical assembly of the board with all the components.



Figure 6.5 Final PCB assembly render (top view). Source: Own.



Figure 6.6 Final PCB assembly render (bottom view). Source: Own.



Figure 6.7 PCB layout top view (left) and bottom view (right). Source: Own.



Figure 6.8 PCB layout top view (left) and bottom view (right). Source: Own.
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Figure 6.9 Reaction wheel mounted on PCB. Source: Own.

## Chapter 7

# **ADCS** software

The following chapter intends to explain the control algorithm developed for the ADCS subsystem of the CubeSat. This control is developed for a 1DOF Cubesat which means it only controls the heading of the CubeSat. However, the code is implemented so it can be easily extrapolated to a 3DOF control in the future. This part of the project is done in collaboration with [64] and [65] which forms the ADCS reaction wheels team. To download the entire control algorithm visit Github.

## 7.1 Control Algorithm

Firstly, as seen in Section 3.2.1, the reference frame used for the CubeSat is the Orbit reference frame. However, more local references shall be noted. Figure 7.1 illustrates the set of reference systems used for the development of the control software.



Figure 7.1 CubeSat reference frames. Source: Own.

Four axis systems are shown in Figure 7.1,

- $X_a, Y_a, Z_a$  are the **absolute local system of reference** which corresponds to orbit reference frame but with z-axis pointing outwards Earth center.
- $X_a, Y_a, Z_a$  are the **CubeSat's system of reference** which are aligned with the IMU's reference frames due to symmetry.
- $X_i, Y_i, Z_i$  are the **IMU's system of reference** which are set by the own Inertial Measurement Unit.
- $X_m, Y_m, Z_m$  are the **motor's reference frame**. This reference frame is coupled with the motor's orientation.

The reference systems are extremely important to bear in mind since depending on the orientation or the setup of the board will change the code and some signs will be needed to be changed. Again, the developed code is only valid for a 1U CubeSat in the same orientation as Figure 8.5, 8.6.

Once the assembly is done, the next step is to calibrate the IMU. This calibration process is located within the main code and calibrations are needed for both the accelerometers and gyroscopes (see Code in O).

## 7.1.1 ADCS Structure

The control software of the ADCS is shown in flowchart 7.2. The system is organized in 4 different modes of operation. Magnetorquer's team is in charge of detumbling mode, nadir pointing whereas Reaction wheels team is in charge of positioning modes since the reaction wheels torque is stronger than magnetorquers.

Take a look at the flowchart in Figure 7.2, each mode is coloured in orange (except emergency mode which is in green).

#### 7.1.2 Control algorithm

Then after all configuration is done, the general control algorithm flowchart is shown in Figure 7.2:



Figure 7.2 ADCS control flowchart. Source: Own.

The control is briefly explained here-under (see Appendix O for the code and O for the Table with all

variables and O for the Table with all functions and their explanation).

First, the program sets up different definitions and parameters. All the variables used in the code are global variables as the program does not make use of functions, instead, Timers were employed for every mode. The reason why the code uses only Timers is that microcontrollers such as Arduino Nano and Bluepill cannot do multitasking. Thus, using Timers can emulate multitasking and perform several actions at the same time (see Section 7.1.3 for a detailed explanation on Timers).

After Timers functions are defined, the microcontroller enters to mode\_OBC\_Input\_Wait, which is the default mode for receiving a mode from the On-Board Computer (OBC). This mode constantly checks the Seriall port from UART to see if the user has sent any command and then execute the corresponding mode.

Once the microcontroller receives the data from the OBC, depending on the input from the user, it enters to either Positioning mode mode\_Positioning\_RW(), IMU reading mode mode\_IMU\_reading or mode\_motor\_on\_off().

- mode\_OBC\_Input\_Wait(): Default mode for OBC reading.
- mode\_Positioning\_RW(): This mode allows the user to turn the CubeSat into the desired position using reaction wheels. The inputs are the direction of turn (whether clockwise or counterclockwise) and the desired angle to turn. Then inside this mode. Depending on the tolerance, the microcontroller will enter to either Fine or Coarse pointing mode. Once set-point is reached and the CubeSat remains still. The process goes back to Waiting mode.
- mode\_IMU\_reading(): This mode displays the IMU's values to let the user check the orientation values (roll, pitch and yaw).
- mode\_motor\_on\_off(): THis mode triggers an impulse to the motor and after a time it turns off.

#### 7.1.3 Timers

First, the program sets up different definitions and parameters. All the variables used in the code are global variables as the program does not make use of functions. Instead, Timers were employed for every mode. The reason why Timers are used is that they can provide virtual "multitasking" to the microcontroller.

Neither the Arduino Nano nor the Bluepill microcontrollers have the ability to real multitasking as they only have 1 core and no operative system. When we talk about timers, in reality, we are taking advantage of the fact that they offer the possibility to temporize sub-tasks without blocking the main task. Thus, microprocessors usually have a set of timers which is a mechanism that permits the execution of an interruption, known as Interrupt Service Routine (ISR). When the program executes this callback function, it goes off the normal flux of the program and executes the associated ISR ignoring all other functions.

There are hardware interruptions that are triggered by a physical change in an event on a physical pin and software interruptions that are in sync with the microcontrollers internal clock. Interruption functions (ISR) must take as short amount of time as possible as the main program is stopped. Nevertheless, in the code implemented, it is only featured to use interruptions and the main void() function is left blank. Eventually, building a program like this takes full advantage of the microprocessor and emulates a "multitasking".

To configure the timers, two concepts must be defined. Dividing the STM32 Bluepill's internal clock's frequency to the Prescale Factor (PF) determines the frequency of the Timer.

$$dt = \frac{1}{\frac{\text{STM32}_{\text{C}}\text{LOCK}}{\text{PF}}} \cdot \text{Ov}$$
(7.1.1)

Multiplying the frequency of the Timer by the Overflow (Ov) one gets the period. In the project, the internal clock is set to 72 MHz, so a PF of 7200 and an Overflow of Ov of 1000 are used to obtain a period of 100 ms. See Figure 7.3 example:



Figure 7.3 Timer and Prescale Factor diagram. Source: Adrià Pérez [64].

The first wave represents the original STM32 clock without any PF nor Ov. The second wave introduces a prescale factor and the third wave introduces the Ov to see the effect on the Timer and the ISR calls.

On important notice is that a Timer can have multiple channels. Timer1 was used for the program and channels CH3 and CH4 were used for OBC readings and IMU readings, respectively.

#### 7.1.4 Coarse Pointing

Coarse pointing mode is the manoeuvre that allows the CubeSat to turn into the desired angle in a fast way. The principle behind it is to give a big impulse to rotate the CubeSat very quickly which produces an acceleration phase. Then to stop the motor when it is approaching the desired angle, the same impulse is injected in the opposite direction of rotation provoking a deceleration.



Figure 7.4 Coarse pointing mode manoeuvre. Source: Own.

Hence, the coarse pointing mode is divided into three phases:

- 1. Acceleration phase: An impulse is generated to arrive at the angle set-point.
- 2. Waiting phase: In this phase, the acceleration is zero and the motor rotates at a constant velocity. If the current yaw angle is higher than half of the desired angle it goes to the deceleration phase.
- 3. **Deceleration phase**: This phase is analogous to the acceleration phase but with an acceleration in the opposite direction.

There are many possible cases in the acceleration phase whether the turn is clockwise (CW) or counterclockwise (CCW). Let's illustrate an example in Figure 7.5. The  $\Delta \psi$  is the final yaw angle minus the initial yaw angle  $\psi_{\text{final}} - \psi_{\text{initial}}$ .

- If the CubeSat goes from  $30^{\circ}$  to  $330^{\circ}$  in the CubeSat CW direction, the difference in angle is  $330 30 = -300^{\circ}$ .
- If the CubeSat goes from 30° to 330° in the CubeSat CCW direction, the difference in angle is
  -(330-30) = -300°, or 60°.
- If the CubeSat goes from  $330^{\circ}$  to  $30^{\circ}$  in the CubeSat CW direction, the difference in angle is  $30 330 = -300^{\circ}$ , or  $60^{\circ}$ .
- If the CubeSat goes from  $330^{\circ}$  to  $30^{\circ}$  in the CubeSat CCW direction, the difference in angle is  $-(30 330) = 300^{\circ}$ .



Figure 7.5 Turn diagram. Source: Own.

If the CubeSat orientation passes through the zero angle, the Yaw angle would decrease its value drastically from 360 to 0 (see Figure 7.6).



Figure 7.6 Yaw angle diagrams (Sketch). Source: Own.

Figure 7.7 shows the angular speed of the system associated with the three phases clearly visible which are acceleration, waiting and deceleration phase.



Figure 7.7 Angular speed diagram (Sketch). Source: Own.

As regards the angular acceleration, Figure 7.8 shows the diagram impulse and theoretical and real acceleration values. In Section 8 we'll see if the results coincide with the theoretical diagram.



Figure 7.8 Angular acceleration diagram (Sketch). Source: Own.

At the end of the Coarse manoeuvre, if the CubeSat is not correctly oriented the next manoeuvre that it will enter is the Fine pointing mode. If the CubeSat happens to get to the desired angle within a tolerance set in the Huge impulses will rarely meet the results and consequently, thus, fine control will try to minimize the error with a PD control.



Figure 7.9 Coarse pointing mode flowchart. Source: Own.

#### 7.1.5 Fine Pointing

Fine pointing mode activates when the desired angle and the actual CubeSats angle are  $< 5^{\circ}$ . Fine pointing mode's goal is to increase the accuracy of the orientation that an impulse might not achieve.



Figure 7.10 Rotation cubesat fine full. Source: Own.

### 7.1.6 PID controller

A PID algorithm consists of the three components before, a proportional constant, integral and a derivative constant which are tuned to get the optimal response.

The basic principle of this algorithm is to calculate the proportional, integral and derivative responses to sum all of them to the output response [66]. A *close loop system* control is a system that is constantly reading sensors to provide constant feedback and calculates the desired output continuously at a fixed loop rate. The *process variable* is the parameter that the system must control, in this case, the process variable is the yaw angle  $\psi$ . And the user inputs the desired set-point.

The formula of a PID controller is defined as follows:

$$u(t) = \underbrace{K_p e(t)}_{\text{Proportional}} + \underbrace{K_i \int e(t) dt}_{\text{Integral}} + \underbrace{K_d \frac{de}{dt}}_{\text{Derivative}}$$
(7.1.2)

where

- u(t): PID control variable.
- e(t): Error value.
- $K_p$ : Proportional gain.
- $K_i$ : Integral gain.
- de: Change in error value.
- dt: Change in time.

**Proportional contribution** depends only on the difference between the actual measurement and the setpoint. The proportional gain is multiplied by the error term so, for instance, if the error has a value of 10 and the proportional constant has a value of 5 it means the system will produce a response of 50. The higher the response the faster the systems responds. However, this value cannot be too high since the system can enter an oscillatory phase with no control.

**Integral contribution** fixes the steady-state error left in the system since it sums the error over time. Steady-state error can be defined as the final difference between the set-point and the actual measurement. In the case of the CubeSat, the integral contribution is not applied since this force usually tends to destabilize the spacecraft instead of correcting it.

**Derivative contribution** considers the rate of approximation of the process variable and the set-point. If the process variable is increasing rapidly, the derivative component will decrease the rate of change of the process variable. A small step size dt provides more accurate and fast responses. For bigger step sizes, the process variable will certainly go faster to the set-point and hence increase the overall speed of the system. An important issue to note is that for very slow control rates or a noisy signal, the derivative response will make the control unstable.

The process of choosing the best constants to match the desired performance is called *controller tuning*. Ziegler and Nichols proposed criteria for tuning PID controllers as one of the most used ways to tune the variables. The basic idea relies on trial and error technique. First, both integral  $K_i$  and derivative  $K_d$  are set to 0. Subsequently, the proportional gain is tuned so it oscillates a bit to a point where the response overshoots but not for too much. Then, the critical constant  $K_{p,cr}$  and period  $T_{p,cr}$  are noted to calculate the other two controller's gains using Table 7.1.

**Table 7.1** Ziegler–Nichols Tuning Rule Based on Critical Gain  $K_{p,cr}$  and Critical Period. Source: Ogata [66].

Control	P	Ti	Td
Р	$0.5K_{p,cr}$	-	-
$\mathbf{PI}$	$0.45K_{p,cr}$	$T_{p,cr}/1.2$	-
PID	$0.60K_{p,cr}$	$0.5T_{p,cr}$	$T_{p,cr}/8$

The response of the controller can be shown in Figure 7.11:



Figure 7.11 System's possible response outcomes [yaw angle] (Sketch). Source: Own.

The CubeSat enters the Fine mode when the angle difference between the set-point and the process variable is  $< 5^{\circ}$ . The final position of the coarse phase shall be set to the initial fine position. Two considerations are made for the final position:

- 1. The angles of the local reference frame are limited from  $[0^{\circ}, 360^{\circ}]$ . This is done to avoid huge values and has a visual representation of the number of laps.
- 2. The second aspect to take into consideration is whether the CubeSat passes through the 0° angle. In this case, if not accounted for, the CubeSat's PD controller can set a huge impulse when the error is actually small. For instance, if the set-point of the CubeSat is small:
  - If the current Yaw angle is 1° which is inside the tolerance set for the Fine control. PID is triggered instead of the coarse pointing mode and the PD response will be small, hence, stabilizing the CubeSat.
  - If the current Yaw angle is 359°, the difference between the angles is very large so the PD will trigger a huge impulse to control it. However, in reality, 359° and 0° are below < 1° of difference so the PD shall set a small impulse. To solve this problem, the angles are shifted 180°. Consequently, the setpoint now becomes 180° and the actual yaw angle is 179° which will produce a difference of 1°.</li>



Figure 7.12 Fine Pointing diagram. Source: Own.



Figure 7.13 Fine pointing mode flowchart. Source: Own.

## Chapter 8

# Performance and Tests

This last chapter's objective is to perform tests emulating an orbit environment. The CubeSat is placed on an air bearing that will emulate outer space conditions. However, limitations on the compressor's capacity restricted the duration time of the simulations.

## 8.1 3D visualization

STM32 or Arduino output data from the Serial port can be pass into a Python script for real-time 3D visualization. Also, real-time orientation angles can be extracted to plot the evolution over time of those variables. However, limitations on Bluetooth communication restricted the number of bits of information sent.

The program can extract the data using the following code (this code works with the IMU\_data\_acquisition\_orientation\_all\_a Arduino file from Serial port).

Projecting the yaw and pitch angles onto vPython's reference frame,



Figure 8.1 vPython reference frame. Source: Own.

The x, y and z position of any point can be expressed as:

$$x = r\cos\phi\cos\theta \tag{8.1.1}$$

$$y = r\sin\phi\cos\theta \tag{8.1.2}$$

$$z = r\sin\theta \tag{8.1.3}$$

To get the Euler angles from quaternions, Euler-Rodriguez formulas were used.



Figure 8.2 Real-time Python visualization. Source: Own.

This python code can be modified to use Euler angles directly with the complementary filter implemented in the code M. However, in the future, it is highly recommended to use quaternions as they provide more accurate results.

## 8.2 Setup

Figures 8.3 shows the Air Bearing used in the project. This testbed includes a pivoting platform that allows the free movement of a nanosatellite and the design of a CubeSat 1U fully equipped to control its spatial orientation [67].



(a) Air bearing structure. Source: Own.

(b) Air bearing with CubeSat mount. Source: Own.



A compressor is located below the air bearing, the compressor blows air into the air bearing structure so when the CubeSat placed on the Air bearing a near-zero friction environment system is created. This system will allow reproducing the CubeSat condition in space. Take a look at Figures 8.4 where the fully modular CubeSat is shown with its solar panels deployed and closed. The final assembly of the CubeSat is represented in Figures 8.5 and 8.6.



(a) Cubesat solar panels. Source: Own.

(b) CubeSat structure. Source: Own.

Figure 8.4 Modular CubeSat. Source: Own.



Figure 8.5 PCB inside the CubeSat. Source: Own.



Figure 8.6 PCB inside the CubeSat (opened). Source: Own.

The final mounted CubeSat on in the air bearing is shown in 8.7. This will permit 3DOF control.



(a) Cubesat mounted on Air Bearing. Source:

Own.



(b) Full CubeSat structure mounted on Air Bearing. Source: Own.

Figure  $8.7\ {\rm Cubesat}$  on Air Bearing. Source: Own.

However, the project focuses on a 1DOF control. Consequently, the testing phase was done using 8.8.



Figure 8.8 Reaction Wheel control test with CubeSat configuration. Source: Own.

## 8.3 Testing phase

First, before engaging in the testing phase, the user must calibrate the magnetometer and acceleration sensors. In order to calibrate the magnetometer, one must follow O guidelines inside the code.

Magnetorquers are calibrated by drawing an "eight" in the air. The calibration usually takes 1-2 minutes and if succeeded the values of calibration are returned to the user so the user must annotate the newest values to the code.



Figure 8.9 GY-9250 IMU magnetometer calibration. Source: Own.

It is important to mention that sending very large messages at a high transfer speed rate may cause the Bluetooth module to freeze, thus, some tests were done twice so in the first one we obtain the orientation angles and in the second one, we obtain accelerometer and gyroscopes measurements as well.

### 8.3.1 IMU Reading mode test

The first test performed was a IMU reading mode mode\_IMU\_reading(). This test aims to examine the Air bearing's stability and how is the data from the different sensors.

Figure 8.10 shows how the CubeSat begins to rotate after 12 seconds when the air bearing was engaged. The conclusions are that if the user does not inject air to the air bearing smoothly, there might be an initial perturbation due to the asymmetry of weight distribution that causes the CubeSat to rotate. Thus, the air injection must be performed gradually and smoothly.



Figure 8.10 Attitude orientation. Source: Own.



(a) Pitch angle. Source: Own.

(b) Roll angle. Source: Own.

Figure 8.11 Pitch and roll angles. Source: Own.

Notice how pitch and roll angles are oscillating back and forth between a set of angles. These results thus need to be interpreted with attention. This phenomenon is due to the vibrations of the CubeSat. As the motor is spinning, the vibrations are not damped by any means so the whole PCB vibrates and, indeed, the IMU attached to it. However, these angles are nearly 0, showing that the CubeSat is certainly

horizontal.

### 8.3.2 Motor ON OFF mode test

The following test aims to turn on the motor until it reaches some velocity and then stops the motor. Notice Figure 8.12, at time t = 0 s the user sets the motor to rotate at it's maximum speed and it is clearly seen how the yaw  $\psi$  angle goes from the initial angle  $\psi_0 = 140^\circ$  approximately to  $\psi = 360^\circ$  completing a lap. When the emergency stops kicks in at t = 12s approximately, the CubeSat's angular rotation slows down, which can be interpreted as the slope of the yaw angle. Contrary to our expectations, in which the team thought the CubeSat to maintain a constant angular velocity once the emergency stop is activated, Results show the CubeSat not only slows down but rotates in the opposite direction.



Figure 8.12 Attitude orientation. Source: Own.



Figure 8.13 Pitch and roll angle. Source: Own.

The physical interpretation of the occurrence can be understood in the following way. After the motor is turned, the CubeSat eventually will arrive at an equilibrium phase where the entire CubeSat and the reaction wheels inertia is countervailed. However, when the reaction wheel slows down its velocity rotating it causes a disruption in the equilibrium of forces and after a while, the CubeSat's rotation due to leftover inertia is bigger than the reaction wheel's leftover inertia. Thereby, the CubeSat rotates in the opposite direction.

#### 8.3.3 Motor ON OFF mode test

Beneath is presented the same motor ON and OFF mode as the prior test. However, this time, other set pf data are recovered instead of the three orientation angles, accelerometer measurements on x and yaxis are shown as well as the angular velocity in z direction.

When the motor is on, one can clearly notice how the accelerometers are induced a lot of noise.



Figure 8.14 Acceleration in x and y direction. Source: Own.



Figure 8.15 Yaw angle. Source: Own.



Figure 8.16 Angular velocity in z direction. Source: Own.

In Figure 8.16 we can observe clearly when the motors are stopped where the angular velocity drops down. Taking a look at the yaw angle, one notice how the CubeSat turns makes 4 complete laps before stopping at  $t \approx 22$  s to rotate in the opposite direction.

### 8.3.4 Coarse pointing control mode test

The next test encompasses the ADCS control. For this test, the intention is to perform a full coarse and fine control. The initial angle  $\psi_0 = 150$  approximately and the desired angle is to rotate 100 degrees in the clockwise direction to achieve 250 degrees approximately.



Figure 8.17 Acceleration in x and y direction. Source: Own.



 $Figure \ 8.18 \ {\rm Yaw \ angle. \ Source: \ Own.}$ 



Figure 8.19 Angular velocity in z direction. Source: Own.

An appreciable change in yaw and angular velocity happens after the deceleration ramp is started (DEC Ramp start). The angular velocity rapidly decreases in  $t \approx 2.5$  s as it surpassed the set point. One important remark is that the impulse gives to the motor is just an impulse. A simple impulse has practical constraints as it sets the motor to its maximum speed. However, ramps will provide better precision. Then when fine pointing mode is activated at t = 12 - 15 s where PD control kicks in.

Nevertheless, the proportional, derivative constants are yet still to be corrected since the CubeSat oscillates back and forth. One method to reduce the oscillation is by increasing the derivative contribution so it reduces the speed if it is approaching the setpoint too fast.

#### 8.3.5 Fine pointing control mode test

This time, only the Fine control was tested on the CubeSat. To do so, the initial turn angle input was less  $< 5^{\circ}$ . This way, as the angle is less than the fine tolerance angle, it will perform a Fine control just from the beginning. The initial angle is  $\psi \approx 350$  and the setpoint is  $346^{\circ}$  approximately.



Figure 8.20 Acceleration in x and y direction. Source: Own.



Figure  $8.21\ {\rm Yaw}\ {\rm angle.}\ {\rm Source:}\ {\rm Own.}$ 



Figure 8.22 Angular velocity in z direction. Source: Own.

Figure 8.22 shows how the CubeSat tries to stay still within the setpoint zone. As mentioned earlier, PD constants need to be tuned accordingly since the proportional gain is yet very strong and overshoots the setpoint.

### 8.3.6 Fine pointing control mode test (PCB with magnetorquer)

Finally, the last test intends to add the weight of the magnetorquers to the CubeSat configuration. As the magnetorquers were placed in one corner of the PCB, it destabilizes the CubeSat. This design error is not complex to solve. The solution is to place some counterweight on the other side of the CubeSat.



Figure 8.23 Bluetooth setup schematic. Source: Own.



Figure 8.24 Bluetooth setup schematic. Source: Own.



Figure 8.25 Bluetooth setup schematic. Source: Own.

Eventually, the CubeSat's PD controller seems to work as expected. Again, the PD controller still needs some tuning since more weight is added to ensure the CubeSat avoid overshooting and oscillatory motion.

## Chapter 9

# Planning of the project

This chapter illustrates the whole planning for the project. A Work Breakdown Structure is developed as well as a Gantt Chart to reproduce the schedule and deadlines for the project.

## 9.1 Work Breakdown Structure

The deliverable-based Work Breakdown Structure demonstrates the relationship between the project deliverables (i.e., results, products) and the scope (i.e., work to be executed).

## 9.1.1 Tasks identification

Table 9.1 shows the tasks performed during the project:

TASK CODE	TASK IDENTIFICATION
1.	Cubesats
1.1.	State of the art
1.2.	Applications
1.3.	ADCS Instruments
2.	Attitude Determination and Control Subsystem
2.1.	Orbital Mechanics
2.2.	Inertial Measurement Unit
2.3.	Euler Angles and Quaternions
2.4.	Disturbance torques
2.5.	Control Algorithm
3.	Communication and Telemetry
3.1.	Tracking
3.2.	Point-to-point Communication
3.3.	Data acquisition and processing
3.4.	Real-time Data plotting
4.	Reaction wheels
4.1.	Physics analysis
4.2.	Motor Selection
4.3.	Motor Driver
4.4.	CAD Design
5.	Assembly
5.1.	PCB Design
5.2.	Final Assembly
6.	Optimization
6.1.	Testing
6.2.	Data Verification
6.3.	Sensibility and precision
7	Report writing

 ${\bf Table \ 9.1 \ Work \ Breakdown \ Structure \ task \ identification. \ Source: \ Own.}$ 

The amount of hours spend in each task can is shown in Table 9.2.

TASK CODE	Duration
1.1.	5
1.2.	5
1.3.	5
2.1.	5
2.2.	10
2.3.	5
2.4.	10
2.5.	20
3.1.	5
3.2.	5
3.3.	5
3.4.	5
4.1.	10
4.2.	5
4.3.	5
4.4.	5
5.1.	5
5.2.	10
6.1.	15
6.2.	10
6.3.	10
7.	222
TOTAL	382

 ${\bf Table \ 9.2} \ {\rm Tasks \ and \ working \ hours. \ Source: \ Own.}$ 

## 9.1.2 Work Breakdown Structure



Figure 9.1 Work Breakdown Structure. Source: Own.

## 9.2 Gantt Diagram

The following Gantt diagram is used to plan the project management and schedule. This tool permits to show the activities (i.e., tasks and events) displayed against time. On the left of the chart is a list of the activities and along the top is a suitable and appropriate time scale. Each activity is presented with its time allocation (see next page).

This allows to get an overview of the number of activities there are, when they begin and how long each activity is scheduled to last and also keep in mind when the project is intended to end.
			1/21 27 31	7	2/	21	21	28	7	3/21 14	21	28	4	11	4/21	2	25	2	9	21	23 30		6 1	6/21	20	27	7/21	11
Thesis - ADCS Cubesat	start	end	_																									
Project Charter	28/01/21	19/03/21									1																	
Documentation and information Res	28/01	11/03					 																					
Aim of the Project	08/02	23/02																										
Scope of the Project	08/02	11/03					1																					
Project Requirements	08/02	11/03																										
State of the art	01/02	16/03																										
Work Breakdown Structure	12/03	16/03 16/03								_																		
Project Charter Submission	19/03	19/03								•	>																	
Development Structural requirements	08/02/21	18/06/21 27/02																										
Attitude Control physics and mathe	24/02	17/04						-																				
Disturbance torques	27/02	17/04							_																			
Inertial Measurement Unit and code	10/03	14/05														I												
Real time data plot and code	10/03	16/04																										
Reaction Wheel Motor & Driver	11/03	24/04																										
Reaction Wheel Design	21/03	07/05																										
Telemetry and Telecommands	03/04	14/05															I											
Coarse and Fine pointing mode	03/04	21/05													1				1									
Control Algorithm	03/04	14/06																										
Error, stability, detumbling and man	16/05	18/06																				1						
Optimization Thesis Report Writing	16/05	18/06																							_			
Introduction	01/02	28/02																										
Documentation and information Res	01/02	19/06					1		1					_														
Mathematical description of the phys	01/02	06/04					1		1				-	1														
Attitude Determination and Control	10/03	30/04																										
Hardware and Software Development	23/03	18/06																										
Budget Environmental aspects	01/06	18/06																										
Report Revision	14/06	20/06																										
Report Writing	02/02	20/06																										
1st Follow-up	16/04	16/04													<b></b>													
3rd Follow-up	28/05	28/05																· ·			•							
Final Subsmission	22/06	22/06																							•			
Presentation	12/07/21	12/07/21																										
Project Exposition	12/07	12/07																										<b>♦</b>

### Chapter 10

## **Environmental Impact**

This chapter intends to analyse the environmental impact of the scope of the project. Despite the major advances in the aerospace area, there are still secondary effects that need to be taken into account.

Sustainability is defined as the capacity of preserving an activity within a long period of time without harming the environment. However, the most accepted definition of **sustainability** is provided by the *World Commission on Environment and Development Definition*.

The major dilemma of putting satellites into orbit resides in the space debris it produces. At the same time, one of the major concern among scientists around the world is not only the impact it might have on other exploration missions but also the light pollution it has in the celestial vault.

Any project that implies a satellite shall analyse the effect among other outer space activities. for instance, the accumulation of space debris can lead to a chain reaction that can cause severe damage to other missions. Besides, the carbon footprint of the project is nearly zero as it has solar panels and most of the operational time it will not rely on any other source.

Since this project is undertaken entirely in the laboratory there are no more environmental issues than the  $CO_2$  emitted during the analysis and development of it. See Table K

$$m_{\rm CO_2} = \text{Impact factor} \left[ \frac{\text{kg CO}_2}{\text{kW} \cdot \text{h}} \right] \times \text{Energy} [\text{kW} \cdot \text{h}]$$
(10.0.1)

Activity	$\mathbf{Time}\;[h]$	$\mathbf{Power}\ \mathrm{kW}$	Impact factor $[\mathrm{kg}~\mathbf{CO_2}/\mathrm{kg}\cdot\mathrm{h}]$	$\mathrm{kg}~\mathbf{CO_2}$
Laptop	382	0.30	0.2	22.92
3D manufacturing	32	0.35	0.2	2.24

Table 10.1 Power consumption and kg  $CO_2$  emission

Thus, an approximate value of 25.16 kg of  $\rm CO_2$  was emitted during the development of this project.

## Chapter 11

# Budget

This section intends to analyze the budget of the personnel time invested in working in this investigation performed and the overall budget to conduct the present project.

### 11.1 Materials budget

To undertake the project, several instruments were used and many components were utilized. Table 11.1 shows the electrical components bought to assembly the ADCS Subsystem.

The price of the components are subject to vary depending on the shopping site. The corresponding values are taken from amazon.es.

Electrical components					
Concept	Quantity	Unitary price $(\mathbf{\epsilon})$	Total price $(\mathbf{\epsilon})$		
Breadboard	2	4.99 €	9.98 €		
Cables	1	6.99 €	6.99 €		
Arduino Nano	1	21.78 €	21.78 €		
STM32 Blue Pill	1	13.49 €	13.49 €		
USB ST-Link V2	1	10.27 €	10.27 €		
BNO055 IMU	1	34.95 €	34.95 €		
GY-9250 IMU	1	8.23 €	8.23 €		
SparkFun Qwiic Motor Driver	1	12.52 €	12.52 €		
Qwiic cable	1	1.50 €	1.50 €		
Motor Mabuchi RF-300EA-1D390	1	6.91 €	6.91 €		
Logic Level converter	1	9.88 €	9.88 €		
Bluetooth module HC-06	1	10.99 €	10.99 €		
Air storage tank Kunshan Abama	1	200 €	200 €		
TOTAL			347.49 €		

Table 11.1Materials budget.Source: Own.

As regards the additive manufactured pieces the associated costs depend on the manufacturing techniques. The air bearing system was manufactured using *Hewlett-Packard* Multi-Jet Fusion 3D printing technologies with PA12CB as the material. However, reaction wheels and supports were made using PLA.

3D printed pieces					
Concept	Quantity	Unitary price $(\mathbf{\xi})$	Total price $(\mathbf{E})$		
Reaction Wheel	2	1.00€	2.00 €		
Reaction wheel support ring top	2	0.10 €	0.20 €		
Reaction wheel support ring bottom	2	0.08 €	0.16 €		
Magnetorquer support	8	0.02 €	0.16 €		
Air Bearing	1	800 €	800 €		
TOTAL			802.52 €		

Table 11.2 3D printed pieces budget. Source: Own.

The PCB was entirely designed by PLATHON's team, however, the manufacturing was sent to JLCPCB as the university could only make 1 layered PCBs.

PCB						
Concept	Quantity	Unitary price $(\mathbf{\xi})$	Total price $(\mathbf{\xi})$			
PCB	5	1.312 €	6.56 €			
Shipping cost (JLCPCB)	1	30.40 €	30.40 €			
TOTAL			36.96 €			

 Table 11.3 PCB manufacturing budget. Source: Own.

Regarding the software tools used to develop the control algorithm, the CAD design and the simulation analysis are shown below 11.4. However, being a student of the Technical University of Catalonia -BarcelonaTech these software tools were no charged since they were educational licensed, nevertheless, for external investigators or companies these budgets must be accounted.

Solidworks was used to design the reaction wheel model and the support pieces. MATLAB was used to develop the atmospheric model and the disturbance torques calculations. Arduino IDE was used to develop the control algorithm for the CubeSat. VS Code was used to perform python 3D simulation. Fritzing tool was used to develop the diagrams and finally, KiCad was used to develop the PCB.

Software licenses						
Concept	Quantity	Unitary price $(\mathbf{\xi})$	Total price $(\mathbf{\xi})$			
Solidworks	1	1295 €/year	1295 €/year			
MATLAB	1	800 €/year	800 €/year			
Arduino IDE	1	0 €	0 €			
VS Code	1	0 €	0 €			
Fritzing	1	8 €	8 €			
KiCad	1	30.40 €	30.40 €			
TOTAL			<b>2133</b> .40 €			

Table 11.4 Software licences budget. Source: Own.

According to the Idescat (Statistics Institute of Catalonia), the mean hourly salary for an professional investigator, scientist or engineer [68] is around  $25.42 \notin /h$ . Thus, the time invested in this research project was about 382 hours (see Table 9.2).

Personnel salary budget						
Concept	Hours (h)	Hourly salary $(\mathbf{E}/\mathbf{h})$	Total (€)			
Aerospace Engineer	382	$25.42~{\rm €/h}$	9710.44 €			
TOTAL			9710.44 €			

Table 11.5 Personnel salary budget. Source: Own.

Finally, summing up all the budgets of this project, the total budget can be calculated:

Total budget				
Concept	Budget (€)			
Electrical components	347.49 €			
3D printed pieces	802.52 €			
PCB manufacturing	36.96 €			
Software licenses	2133.40 €			
Personnel salary	9710.44 €			
TOTAL	13030.81 €			

Table 11.6 Total budget. Source: Own.

## Chapter 12

# Conclusions

To sum up, the present thesis describes the attitude determination and control subsystem for PLATHON's project. According to the requirements set in the initial phase of the project, the reaction wheels were designed.

The first phase of the project encompassed a review of astrodynamics in which it reviews some important concepts such as reference frames and how to describe the attitude with information about Euler angles and Quaternions. After that, the most important disturbance torques were studied. The spacecraft in space is subjected to different disturbance torques which if not considered can cause undesired torques or inaccurate measurements. The main contribution was due to

The second phase of the project provides an in-depth analysis of the Reaction wheels design that must comply with the requirements of the project and develop the code for the control. All the components studied were assembled into the PCB with no issues. Finally, as regards the control algorithm developed has been divided into two principal modes, fine and coarse pointing mode. As for the coarse pointing mode, an impulse is set to achieve the desired set-point. When the CubeSat orientation is near the desired set-point, fine control steps into using a PD controller.

The main objective of the thesis is accomplished, which was to design a fully operational 1DOF attitude control system for a CubeSat. All software is open-source and the different modes of operation work perfectly with some tuning. The code is accessible for further development and can be extrapolated to a 3DOF system.

### 12.1 Future work and recommendations

The following section gives some suggestions based on the study during this thesis and recommendations for future hardware and software improvements.

• Although GY-9250 IMU module provides accurate measurements, the code provided does not use the internal DMP for extracting the orientation angles. These angles are calculated from raw instruments so the error is slightly higher and it loads more code for the microprocessor to calculate. This restriction is faced since there are no libraries that benefit the IMU's DMP using SPI communication protocol but there is a library that uses  $I^2C$  protocol which enables DMP. Nevertheless, it is always better to use quaternions for attitude determination and then transforming them into Euler angles.

- Calibration method for the IMU is a slow process and despite the measurements have small errors, the drift on the gyroscope for long-term missions cannot be avoided. Alternative instruments can be used to readjust the gyroscope's error or a self-calibration mode in the control algorithm can be added.
- Additional attitude determination sensors, such as a gyroscope or an Earth horizon sensor, are required to improve angular estimation accuracy during an eclipse.
- Reaction wheel design performs as expected and has a lot of inertia. However, the implemented wheel is for a 1DOF system. If a 3DOF system is added, the reaction wheel shall be smaller and perhaps new motors with higher torques should be considered, for instance, brushless motors.
- If the attachment of the reaction wheel and motor is not perfectly aligned, the Cubesat can rotate off-axis causing precession. This issue can be solved by designing a piece that ensures the alignment of the axis in both objects.
- Results show that vibrations induce the accelerometers measurements to oscillate, thus, a damping system may be useful to retrieve smoother and more accurate data.
- As the code now employs only impulses for activating coarse pointing mode, a ramp input will reduce the overshooting and obtain more precise results.
- Extra functionalities such as a graphical interface should be developed to make a better
- Detumbling and nadir pointing modes were developed by magnetorquer team. A further step is to merge both team's software and develop control with both reaction wheels and magnetorquers which will drastically increase the accuracy of the project.

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# Appendices

## Appendix A

## Requirements

### A.1 Technical requirements

In this section, we aim to provide strict and measurable statements that will guide and limit the process, to minimise weak spots and vague actions that could lead to misunderstandings:

One of the most important limitations comes from the maximum weight permitted for a 1U Cubesat. The Cubesat standard specification [2] dictates the maximum dimensions and weights for a 1U Cubesat, which cannot exceed a  $10 \times 10$  cm side cube with a maximum weight of 1,3 kg. Thereby, the current Cubesat is categorized as a nanosatellite since its weight is higher than 1 kg and less than 10 kg.

According to Dr. Griffin and Dr. French [37], the ADCS subsystem weight is around 10 - 15% of the total weight of the satellite. This means the upper boundary limitation is around 200 g out of the whole satellite's weight.

An additional requirement shall be to reduce the response time of the reaction wheel once a duty is set. The final aim is to pursue the optimal response time and performance while preserving energy and accuracy.

The Cubesat must meet The Belgian Law on the Activities of Launching, Flight Operation or Guidance of Space Objects 2005 regulation Belgian Air and Space Policy [69].

The CubeSat must also comply with technical constraints as specified by FCC regulations Federal Communications Commission [70].

The experimental-licensed Cubesat spacecraft that use frequency bands allocated to the Amateur-Satellite Service must coordinate with the International Amateur Radio Union (IARU) [71].

If the Cubesat includes an imager or a camera, for non-Government owned Cubesat with an imager must contact the National Oceanic and Atmospheric Administration (NOAA) for a remote sensing license if needed [72].

#### APPENDIX A. REQUIREMENTS

The Cubesat design must always be designed following the United States Government (USG) Orbital Debris Mitigation Standard Practices (ODMSP) to address the increase in orbital debris in the near-Earth space environment. The goal is to limit the generation of new, long-lived debris by the control of debris released during normal operations, minimizing debris generated by accidental explosions, the selection of safe flight profile and operational configuration to minimize accidental collisions, and post-mission disposal of space structure from NASA [73].

Small satellites should also fall under the scope of national space legislation United Nations Office for Outer Space Affairs (UNOOSA) [74].

Finally, all of the aforementioned standards have a lot in common because their purpose is the same: to propose an organized system for getting things done correctly. If the project is space-related and takes place in Europe, the ECSS standard should be followed. Nonetheless, if it is located in the United States, it must adhere to the NASA standard. To consider which standard to use depends on many factors: where is the mission located, what is its purpose and whether it is part of a larger project.

The following are some formal document-based and model-based systems engineering standards:

- European Cooperation for Space Standardization (ECSS): ECSS-E-ST-10C [75]
- The National Aeronautics and Space Administration (NASA): Systems Engineering Handbook [76].
- International Council on Systems Engineering (INCOSE): Guide to the Systems Engineering Body of Knowledge (SEBoK) [77].
- ISO/IEC/IEEE 15288 : Systems and software engineering [78].

No exact requirements for attitude control or response time are set. Nonetheless, the aim will always be to increase the performance and response time ensuring fast and reliable maneuvers while preserving the Cubesat Standard constraints [3] [79] [80] [81].

### A.2 Structure Subsystem

Cubesat specifications can be found in California's University's specification [2]. The CubeSat Design Specification document presents a deep overview of all general, mechanical, electrical and operational requirements.

Below, general and mechanical requirements are listed and explained:

#### A.2.1 General requirements of the Cubesat

1. CubeSats which incorporate any deviation from the CDS will submit a DAR and adhere to the waiver process.

This means that any changes to the standardized specifications shall be notified, that is, the mission shall fill out a form and send it to the launch provider.

2. All parts shall remain attached to the CubeSats during launch, ejection and operation. No additional space debris will be created.

To ensure this requirement, the structure of the Cubesat must ensure to be able to support the forces it is subjected.

- 3. No pyrotechnics shall be permitted.
- 4. Any propulsion systems shall be designed, integrated, and tested in accordance with AFSPCMAN 91-710 Volume 3 [82].
- 5. Propulsion systems shall have at least 3 inhibits to activation.

This translates that 3 different mechanisms or buttons must be engaged to inhibit activation.

6. Total stored chemical energy will not exceed 100 Watt-Hours.

6.1. Note: Higher capacities may be permitted, but could potentially limit launch opportunities.

- 7. CubeSat hazardous materials shall conform to AFSPCMAN 91-710, Volume 3 [82].
- 8. CubeSat materials shall satisfy the following low out-gassing criterion to prevent contamination of other spacecraft during integration, testing, and launch. A list of NASA approved low out-gassing materials can be found at: http://outgassing.nasa.gov.

8.1. Note: Higher capacities may be permitted, but could potentially limit launch opportunities.

Some materials, most common plastics, are designed to operate a certain level of atmospheric pressure. When exposed to outer space conditions, i.e. vacuum and zero pressure levels, the material can release a cloud of mist or condensate which can end up damaging other electronic components including solar panels. For instance, outgassing can contribute to the degraded performance of charge-coupled-device (CCD) sensors of some missions. NASA maintains a list of materials with low-outgassing properties suitable for use in spacecraft [83].

- 9. The latest revision of the CubeSat Design Specification will be the official version to which all CubeSat developers will adhere to. The latest revision is available at http://www.cubesat.org.
- 10. The CubeSat shall be designed to accommodate ascent venting per ventable volume/area <2000 inches

#### A.2.2 Mechanical requirements of the Cubesat

CubeSats are cube-shaped nanosatellites with dimensions and features outlined in the CubeSat Specification Drawing (see Figure A.1) for reference.



Figure A.1 P-POD Cubesat coordinate system. Source: California Polytechnic State University [2].

General features of all CubeSats include:

- 1. The CubeSat shall use the coordinate system as defined in Appendix B of [2] (see Figure A.2) for the appropriate size. The CubeSat coordinate system will match the P-POD coordinate system while integrated into the P-POD. The origin of the CubeSat coordinate system is located at the geometric centre of the CubeSat.
  - 1.1. The CubeSat configuration and physical dimensions shall be per the appropriate section of Appendix B [2].
  - 1.2. The extra volume available for 3U CubeSats is shown in Figure 6 of Appendix B [2].
- 2. The Z face of the CubeSat will be inserted first into the P-POD.
- 3. No components on the green and yellow shaded sides shall exceed 6.5 mm normal to the surface.
  - 3.1. When completing a CubeSat Acceptance Checklist (CAC), protrusions will be measured from the plane of the rails.
- 4. Deployables shall be constrained by the CubeSat, not the P-POD.
- 5. Rails shall have a minimum width of 8.5 mm.
- 6. Rails will have a surface roughness less than 1.6  $\mu \mathrm{m}.$
- 7. The edges of the rails will be rounded to a radius of at least 1 mm.
- 8. The ends of the rails on the +/-Z face shall have a minimum surface area of 6.5 mm × 6.5 mm contact area for neighbouring CubeSat rails (as per Figure 6).

- 9. At least 75% of the rail will be in contact with the P-POD rails. 25% of the rails may be recessed and no part of the rails will exceed the specification.
- 10. The maximum mass of a 1U CubeSat shall be 1.33 kg.

10.1. Note: Larger masses may be evaluated on a mission to mission basis.

11. The maximum mass of a 1.5U CubeSat shall be 2.00 kg.

11.1. Note: Larger masses may be evaluated on a mission to mission basis.

12. The maximum mass of a 2U CubeSat shall be 2.66 kg.

12.1. Note: Larger masses may be evaluated on a mission to mission basis.

- 13. The maximum mass of a 3U CubeSat shall be 4.00 kg.
  - 13.1. Note: Larger masses may be evaluated on a mission to mission basis.
- 14. The CubeSat center of gravity shall be located within 2 cm from its geometric center in the X and Y direction.
  - 14.1. The 1U CubeSat center of gravity shall be located within 2 cm from its geometric center in the Z direction.
  - 14.2. The 1.5U CubeSat center of gravity shall be located within 3 cm from its geometric center in the Z direction.
  - 14.3. The 2U CubeSat center of gravity shall be located within 4.5 cm from its geometric center in the Z direction.
  - 14.4. 3U and 3U CubeSats' center of gravity shall be located within 7 cm from its geometric center in the Z direction.
- 15. Aluminum 7075, 6061, 5005, and/or 5052 will be used for both the main CubeSat structure and the rails.
  - 15.1. If other materials are used the developer will submit a DAR and adhere to the waiver process.

These materials were selected considering their weight and their ability to support structural loads.

- 16. The CubeSat rails and standoff, which contact the P-POD rails and adjacent CubeSat standoffs, shall be hard anodized aluminium to prevent any cold welding within the P-POD.
- 17. The 1U, 1.5U and 2U CubeSats shall use separation springs to ensure adequate separation (see Table A.1 and Figure A.3).
  - 17.1. The 1U, 1.5U and 2U CubeSats shall use separation springs to ensure adequate separation.
  - 17.2. The compressed separation springs shall be at or below the level of the standoff.
  - 17.3. The 1U, 1.5U and 2U CubeSats separation spring will be centred on the end of the standoff on the CubeSat's -Z face as per Figure 7 [2].
  - 17.4. Separation springs are not required for 3U CubeSats.



Figure A.2 1U CubeSat Design Specification Drawing. Source: California Polytechnic State University [2].

Characteristics	Value
Plunger Material	Stainless Steel
End Force Initial/Final	$0.14~\mathrm{lbs}$ / $0.9~\mathrm{lbs}$
Throw Length	0.16 in minimum above the standoff surface
Thread Pitch	8-36 UNF-2B

 Table A.1 CubeSat Separation Spring Characteristics. Source: California

 Polytechnic State University [2].



Figure A.3 Deployment switches and separation spring locations. Source: California Polytechnic State University [2].

#### A.2.3 Electrical Requirements

Electronic systems will be designed with the following safety features.

- The CubeSat power system shall be at a power-off state to prevent CubeSat from activating any powered functions while integrated with the P-POD from the time of delivery to the Launch Vehicle (LV) through on-orbit deployment. CubeSat powered function includes a variety of subsystems such as Command and Data Handling (C&DH), RF Communication, Attitude Determine and Control (ADC), deployable mechanism actuation. CubeSat power systems include all battery assemblies, solar cells, and coin cell batteries.
- 2. The CubeSat shall have, at a minimum, one deployment switch on a rail standoff, per Figure A.3.
- 3. In the actuated state, the CubeSat deployment switch shall electrically disconnect the power system from the powered functions; this includes real-time clocks (RTC).
- 4. The deployment switch shall be in the actuated state at all times while integrated in the P-POD.
  - 4.1. In the actuated state, the CubeSat deployment switch will be at or below the level of the standoff.
- 5. If the CubeSat deployment switch toggles from the actuated state and back, the transmission and deployable timers shall reset to t = 0.

- 6. The Remove Before Flight (RBF) pin and all CubeSat umbilical connectors shall be within the designated Access Port locations, green shaded areas shown in Appendix B [2].
  - 6.1. Note: All diagnostics and battery charging within the P-POD will be done while the deployment switch is depressed.
- 7. The CubeSat shall include an RBF pin.
  - 7.1. The RBF pin shall cut all power to the satellite once it is inserted into the satellite.
  - 7.2. The RBF pin shall be removed from the CubeSat after integration into the P-POD.
  - 7.3. The RBF pin shall protrude no more than 6.5 mm from the rails when it is fully inserted into the satellite.
- 8. CubeSats shall incorporate battery circuit protection for charging/discharging to avoid unbalanced cell conditions.
- 9. The CubeSat shall be designed to meet at least one of the following requirements to prohibit inadvertent radio frequency (RF) transmission. The use of three independent inhibits is highly recommended and can reduce required documentation and analysis. An inhibit is a physical device between a power source and a hazard. A timer is not considered an independent inhibit.
  - 9.1. The CubeSat will have one RF inhibit and RF power output of no greater than 1.5 W at the transmitting antenna's RF input.
  - 9.2. The CubeSat will have two independent RF inhibits

### A.3 Operational Requirements

CubeSats will meet certain requirements pertaining to integration and operation to meet legal obligations and ensure the safety of other CubeSats [2].

- 1. Operators will obtain and provide documentation of proper licenses for use of radio frequencies.
  - 1.1. For amateur frequency use, this requires proof of frequency coordination by the International Amateur Radio Union (IARU). Applications can be found at www.iaru.org.
- 2. CubeSats will comply with their country's radio license agreements and restrictions.
- 3. CubeSats mission design and hardware shall be in accordance with NPR 8715.6 to limit orbital debris.
  - 3.1. Any CubeSat component shall re-enter with energy less than 15 J.
  - 3.2. Developers will obtain and provide documentation of approval of an orbital debris mitigation plan from the Federal Communication Commission (FCC) (or NOAA if imager is present).
    - 3.2.1. Note: To view FCC amateur radio regulations, go to http://www.arrl.org/part-97-amateur-radio
    - 3.2.2. Note: Analysis can be conducted to satisfy the above with NASA DAS, available at http://orbitaldebris.jsc.nasa.gov/mitigate/das.html

- 4. All deployable such as booms, antennas, and solar panels shall wait to deploy a minimum of 30 minutes after the CubeSat's deployment switch(es) are activated from P-POD ejection.
- 5. No CubeSats shall generate or transmit any signal from the time of integration into the P-POD through 45 minutes after on-orbit deployment from the P-POD. However, the CubeSat can be powered on the following deployment from the P-POD.
- 6. Private entities (non-U.S. Government) under the jurisdiction or control of the United States who propose to operate a remote sensing space system (satellite) may need to have a license as required by U.S. law. For more information visit http://www.nesdis.noaa.gov/CRSRA/licenseH ome.html. Click on the Application Process link under the Applying for a License drop-down section to begin the process.
- 7. Cal Poly will conduct a minimum of one fit check-in in which developer hardware will be inspected and integrated into the P-POD or TestPOD. A final fit check will be conducted before launch. The CubeSat Acceptance Checklist (CAC) will be used to verify compliance with the specification (Found in the appendix of this document or online at http://cubesat.org/index.php/do cuments/developers).

### A.4 Testing requirements

Testing will be performed to meet all launch provider requirements as well as any additional testing requirements deemed necessary to ensure the safety of the CubeSats, P-POD, and the primary mission. The General Environmental Verification Standard (GEVS, GSFC-STD-7000) and MIL-STD-1540 can be used to derive testing requirements[2].

The general testings are listed below:

#### 1. Random Vibration:

Random vibration testing shall be performed as defined by the launch provider

#### 2. Thermal Vacuum Bakeout:

Thermal vacuum bakeout shall be performed to ensure proper outgassing of components. The test specification will be outlined by the launch provider.

#### 3. Shock testing:

Shock testing shall be performed as defined by the launch provider.

#### 4. Visual inspection:

Visual inspection of the CubeSat and measurement of critical areas will be performed per the appropriate CAC (Appendix C) [2].

#### 5. Cubesat testing Philosophy:

The CubeSat shall be subjected to either a qualification or protoflight testing as defined in the CubeSat Testing Flow Diagram, shown in Figure A.4. The test levels and durations will be supplied by the launch provider or P-POD integrator.

#### 5.1. Qualification:

Qualification testing is performed on an engineering unit hardware that is identical to the flight model CubeSat. Qualification levels will be determined by the launch vehicle provider or P-POD integrator. Both MIL-STD-1540 and LSP-REQ-317.01 are used as guides in determining testing levels. The flight model will then be tested to Acceptance levels in a TestPOD then integrated into the flight P-POD for a final acceptance/workmanship random vibration test. Additional testing may be required if modifications or changes are made to the CubeSats after qualification testing.

#### 5.2. Protoflight:

Protoflight testing is performed on the flight model CubeSat. Protoflight levels will be determined by the launch vehicle provider or P-POD integrator. Both MIL-STD-1540 and LSPREQ-317.01 are used as guides in determining testing levels. The flight model will be tested to Protoflight levels in a TestPOD then integrated into the flight P-POD for a final acceptance/workmanship random vibration test. The flight CubeSat SHALL NOT be disassembled or modified after protoflight testing. Disassembly of hardware after protoflight testing will require the developer to submit a DAR and adhere to the waiver process prior to disassembly. Additional testing will be required if modifications or changes are made to the CubeSats after protoflight testing.

#### 5.3. Acceptance:

After delivery and integration of the CubeSat into the P-POD, additional testing will be performed with the integrated system. This test ensures proper integration of the CubeSat into the P-POD. Additionally, any unknown, harmful interactions between CubeSats may be discovered during acceptance testing. The P-POD Integrator will coordinate and perform acceptance testing. Acceptance levels will be determined by the launch vehicle provider or P-POD integrator. Both MIL-STD-1540 and LSP-REQ-317.01 are used as guides in determining testing levels. The PPOD SHALL NOT be deintegrated at this point. If a CubeSat failure is discovered, a decision to disintegrate the P-POD will be made by the developers, in that P-POD, and the P-POD Integrator based on safety concerns. The developer is responsible for any additional testing required due to corrective modifications to deintegrated P-PODs and CubeSats.



Figure A.4 CubeSat General Testing Flow Diagram. Source: California Polytechnic State University [2].

## Appendix B

## Microcontroller

The following section provides the pin-outs of the microcontrollers discussed in the project.

### B.1 Arduino Nano

The Arduino board is built in such a manner that newcomers may easily get started with microcontrollers. This board, in particular, is breadboard friendly, making it very simple to manage the connections. Let us begin by energizing the Board.

To power the Arduino Nano, there are three ways:

- USB Jack: Connect the mini USB jack to a phone charger or computer through a cable and it will draw the power required for the board to function
- Vin Pin: The Vin pin can be supplied with an unregulated 6-12V to power the board. The on-board voltage regulator regulates it to +5V
- +5V Pin: If you have a regulated +5V supply then you can directly provide this with the +5V pin of the Arduino.

For the input and outputs pins:

Arduino Nano board has 14 digital pins and 8 analog pins in total. The digital pins can be used to interconnect sensors (as input pins) or to control loads (as output pins). To regulate their functionality, basic functions such as pinMode() and digitalWrite() can be used. For digital pins, the operational voltage is 0 V and 5 V. Analog voltages ranging from 0V to 5 V may be measured using any of the 8 Analog pins and a simple function called analogRead()

Also some pins have special purposes:

• Serial Pins 0 (Rx) and 1 (Tx): Rx and Tx pins are used to receive and transmit TTL serial data. They are connected with the corresponding ATmega328P USB to TTL serial chip.

- External Interrupt Pins 2 and 3: These pins can be configured to trigger an interrupt on a low value, a rising or falling edge, or a change in value.
- **PWM Pins 3, 5, 6, 9 and 11**: These pins provide an 8-bit PWM output by using analogWrite() function.
- SPI Pins 10 (SS), 11 (MOSI), 12 (MISO) and 13 (SCK): These pins are used for SPI communication.
- In-built LED Pin 13: This pin is connected with an built-in LED, when pin 13 is HIGH LED is on and when pin 13 is LOW, its off.
- I2C A4 (SDA) and A5 (SCA): Used for I2C communication using Wire library.
- AREF: Used to provide reference voltage for analog inputs with analogReference() function.
- Reset Pin: Making this pin LOW, resets the microcontroller.

These special functions and their respective pins are illustrated in the Arduino Nano pin diagram shown in Figure B.1.



Figure B.1 Arduino Nano Pinout. Source: Robu.in Arduino [84]











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### B.2 STM32 Bluepill

STM- Microelectronics provides a diverse variety of general-purpose microcontrollers for both 8-bit MCU (STM8) and 32-bit ARM Cortex-M-based microcontrollers (STM32). STM32 Blue Pill is based on STM32F103C8T6, which contains a Cortex-M3 ARM CPU that operates at 72 MHz, 20 kB of RAM, and 64kB of flash memory. The microcontroller has one USB port (with no additional chip support), three UART ports, 16-bit PWM pins, and other features. It is a 3.3V microcontroller with several 5V tolerant pins [51].

To power the Bluepill, one can power it with several ways listed below:

- Using the built-in USB micro connector.
- Supplying 5 V to the 5 V pin as external supply.
- Supplying 3.3 V directly to the 3.3 V pin.

For the input and output pins: The Blue Pill features 37 GPIO pins distributed across four ports: A and B (16 pins), C (3 pins), and D (no pins) (2 pins). Each pin has a 6 mA current sink/source capability. On each of the pins, pull-up and pull-down resistors can be activated.

Most pins have extra functionality as well:

- Serial ports receive and transmit data via the UART protocol
- $I^2C$  ports two-wire communication via the  $I^2C$  protocol
- SPI serial communication
- PWM
- Pin 13 has a built-in LED

These special functions and their respective pins are illustrated in the Blue Pill pin diagram shown below:



Figure B.2 Arduino Nano Pinout. Source: Robu.in Arduino [84]



#### B.2.1 ST-Link

To upload and download code from the PC to the STM32 Bluepill an external device is needed called STLink.



Figure B.3 STLink v2 pinout diagram. Source: Freeelectron [85].

Table B.1 ST-Link and STM32 pin connections. Source: Own.

ST-Link	STM32 Bluepill
GND	GND
SWCL	$\operatorname{SCL}$
SWDIO	SWDIO
3,3 V	$3.3\mathrm{V}$

#### B.2.2 Arduino IDE setup configuration for STM32

Having seen the advantages of the STM32 microcontroller explained in the precedent section. The next step encompasses the configuration of the Arduino IDE for developing on the STM32 microcontroller. The STM32 microcontroller is compatible with the Arduino IDE. However, some configurations need to be met in order to satisfy the compatibility.

The next paragraphs will summarize the STM32 configuration to meet the Arduino IDE requirements. The whole process can be found in [86].

- First, the Arduino IDE must be opened.
- Go to File then Preferences.
- In "Additional Boards Manager URL" field add: http://dan.drown.org/stm32duino/package\_STM32duino\_index.json
- Next click on Tools, Board, Board Manager and search for "STM32".
- Install "STM32F1xx/GD32F1xx" by stm32duino.
- Then select the corresponding board in Tools and make sure upload method is "ST-LINK".

## Appendix C

## **Inertial Measurement Unit**

The following chapter shows the two IMU analyzed and tested within the project. It has everything to check

### C.1 MPU9025 module

This module is based on the MPU9250 sensor and includes several sensors to monitor 9-axis (9-DoF) motion. It is composed of a 3-axis gyroscope, a 3-axis accelerometer, and a 3-axis magnetometer. Furthermore, it has a DMP (Digital Motion Processor) that can run complex 9-axis motion capture algorithms [52].



Figure C.1 MPU9250 Pin diagram. Source: Components101 [52].

It communicates with the microcontrollers through an  $I^2C$  interface and includes a widely used library for simple usage. This sensor has 9 degrees of freedom and has a 3.3V voltage regulator as well as pull-up resistors for direct  $I^2C$  application.

The accelerometer sensibility can be adjusted to work in different ranges from  $\pm 2g$ ,  $\pm 4g$ ,  $\pm 8g$ ,  $\pm 16g$ . I the case of the magnetometer it has a sensibility of  $\pm 4800 \ \mu\text{T}$  and  $\pm 250^{\circ}/\text{s}$ ,  $\pm 500^{\circ}/\text{s}$ ,  $\pm 1000^{\circ}/\text{s}$ ,  $\pm 2000^{\circ}/\text{s}$  gyroscope.

The technical specifications are summarized below:

- Operating voltage: 3 V, 3.3 V or 5 V
- Accelerometer Range:  $\pm 2g$ ,  $\pm 4g$ ,  $\pm 8g$ ,  $\pm 16g$
- Gyroscope Range:  $\pm 250^{\circ}/s$ ,  $\pm 500^{\circ}/s$ ,  $\pm 1000^{\circ}/s$ ,  $\pm 2000^{\circ}/s$
- Magnetometer range:  $\pm 4800 \ \mu T$
- Interface: I<sup>2</sup>C and SPI
- AD converter: 16 Bits (digital output)
- Degrees of Freedom (DoF): 9
- Sensor: MPU9250
- Onboard voltage regulator
- Size:  $2.0 \text{cm} \ge 1.6 \text{cm} \ge 0.3 \text{cm}$

Table C.1 MPU9250 Pin-outs. Source: Components101 [52].

Pin Number	Pin Name	Description
1	VCC	Power Supply
2	GND	Ground Reference
3	SCL	I <sup>2</sup> C Serial Clock
4	SDA	I <sup>2</sup> C Serial Data
5	EDA	Auxiliary Serial Data
6	ECL	Auxiliary Serial Clock
7	AD0	I <sup>2</sup> C/SPI Address Select
8	INT	Interrupt
9	NCS	SPI Chip Select
10	FSYNC	Frame Synchronization



Figure C.2 MPU9250 Internal Circuit. Source: Components101 [52].
#### C.2 BNO055 module

The BNO055 is a System in Package (SiP) that combines in a single package a triaxial 14-bit accelerometer, a triaxial 16-bit gyroscope with a range of 2000 degrees per second, a triaxial geomagnetic sensor, and a 32-bit cortex M0+ microprocessor running Bosch Sensortec sensor fusion software.

The chip-sets are housed in a single 28-pin LGA chassis measuring 3.8mm x 5.2mm x 1.1mm. The BNO055 has digital bidirectional I2C and UART interfaces for optimal system integration. The BNO055 data output is versatile, it can provide both quaternions or Euler angles as well as providing directly the use of the RAW output as well (see items below).

The BNO055 can output the following sensor data [87]:

- Absolute Orientation (Euler Vector, 100 Hz)
- Three axis orientation data based on a  $360^{\circ}$  sphere
- Absolute Orientation (Quaternion, 100 Hz)
- Four point quaternion output for more accurate data manipulation
- Angular Velocity Vector (100 Hz)
- Three axis of 'rotation speed' in rad/s
- Acceleration Vector (100 Hz)
- Three axis of acceleration (gravity + linear motion) in  $m/s^2$
- Magnetic Field Strength Vector (20 m/s<sup>2</sup>)
- Three axis of magnetic field sensing in micro Tesla  $(\mu T)$
- Linear Acceleration Vector (100 Hz)
- Three axis of linear acceleration data (acceleration minus gravity) in m/s<sup>2</sup>
- Gravity Vector (100 Hz)
- Three axis of gravitational acceleration (minus any movement) in m/s<sup>2</sup>
- Temperature (1 Hz)
- Ambient temperature in degrees Celsius

With respect to the pin-out of this sensor [88]:

Pins	Pin Name	Description			
	VIN	3.3-5.0V power supply input			
Power pins	31/0	3.3V output from the on-board linear voltage regulator,			
	310	up to about 50mA as necessary			
	GND	The common/GND pin for power and logic			
		I2C clock pin, connect to your microcontrollers			
	SCI	I2C clock line.			
I2C	501	This pin can be used with 3V or 5V logic,			
		and there's a 10K pullup on this pin			
		I2C data pin, connect to your microcontrollers			
	SDA	I2C data line.			
	SDA	This pin can be used with 3V or 5V logic,			
		and there's a 10K pullup on this pin.			
	BST	Hardware reset pin. Set this pin low the high to perform a rest.			
	1651	This pin is 5V safe.			
Other pins		Set this pin high to change the default I2C			
		address for the BNO055 if you need to			
	ADB	connect two ICs on the same I2C bus.			
	ADIt	The default address is 0x28.			
		If this pin is connected to 3V,			
		the address will be $0x29$			
		The HW interrupt output pin, which can be configured to			
		generate an interrupt signal when certain events			
	INT	occur like movement detected by the accelerometer, etc.			
	1111	(not currently supported in the Adafruit library,			
		but the chip and HW is capable of generating this signal).			
		The voltage level out is 3V			
		These pins can be used to change the mode of the device			
		(it can also do HID-I2C and UART)			
	PS0 and PS1	and also are provided in case Bosch provides a firmware update			
		at some point for the ARM Cortex M0 MCU inside the sensor.			
		They should normally be left unconnected.			

Table C.2 BNO055 Pin-out connections. Source:[87].

### Appendix D

### Logic Level Converter

The logic level converter is a fundamental piece when converting high voltage to low voltage and vice-versa. Processors, run at several voltages, the most popular of which are 5 V, 3.3 V, and, to a lesser extent, 2.8 and 1.8 V. 5 V. However, most current microcontrollers, though (Raspberry, Arduino, etc.), run at 3.3 V.

Nevertheless, the majority of the electrical components (sensors, controllers, and displays) operate at various nominal voltages, namely 5 V and 3.3 V. To link digital devices with various nominal voltages, the voltage levels must be adjusted. Otherwise, the assembly is unlikely to function well, and we may even damage a component.



Figure D.1 Logic Level converter. Source: Luis Llamas [89].

To connect to a logic level shifter is a simple process. There are two sides, one for the device with the highest voltage (identified as HIGH) and one for the device with the lowest voltage (identified as LOW).

First, connect the GND of both devices to the GND of the board on the respective sides of the end adapter. GND has to be common for both devices, so the GND of both sides of the board are linked internally. If not sure, check with a multimeter on both sides of the level adapter.

On the other hand, both devices to the level shifter must be connected to the two reference voltages of both devices to the level shifter. For this, there is the HV pin, for the Vcc voltage of the higher voltage device, and LV for Vcc of the lower voltage device.

Finally, GPIO pins (channels) can be used as many time as needed. The channels are bidirectional and work for both digital and analogue signals, or communication systems (UART, I<sup>2</sup>C, SPI).

### Appendix E

## **Bluetooth Module**

The Bluetooth module selected in the present project is the HC-06 model from DSD Tech. The following section is a guide for setting up the Bluetooth module for both Arduino and STM32 microcontrollers.

As regards the pins on the HC-06 Bluetooth module [90],

- RXD: Serial Data Receive Pin. Used for serial input, 3.3V logic. Typically hooked up to transmission pin (TX) of the Arduino.
- TXD: Serial Data Transmit Pin. Used for serial output. 3.3V logic. Typically hooked up to reception pin (RX) of the Arduino
- GND: Ground. Typically hooked up to GND pin of the Arduino
- VCC: +5V. Power supply. Typically hooked up to 5V pin of the Arduino.

The connections are [90]:

$$\begin{array}{rccc} \mathrm{RXD} & \longrightarrow & \mathrm{TX} \\ \mathrm{TXD} & \longrightarrow & \mathrm{RX} \\ \mathrm{GND} & \longrightarrow & \mathrm{GND} \\ \mathrm{VCC} & \longrightarrow & 5\mathrm{V} \end{array}$$



Figure E.1 HC-05 & HC-06 pin-out. Source: Components101 [90].

The characteristics of the module are [90]

- HC-06 Features and Electrical characteristics
- Bluetooth protocol: Bluetooth V2.0 protocol standard
- Power Level: Class2(+6dBm)
- Band: 2.40GHz—2.48GHz, ISM Band
- Receiver sensitivity: -85 dBm
- USB protocol: USB v1.1/2.0
- Modulation mode: Gauss frequency Shift Keying
- Safety feature: Authentication and encryption
- Operating voltage range:+3.3V to +6V
- Operating temperature range: -20°C to +55°C
- Operating Current: 40mA

On the other hand, the HC-05 shares similar pins [91]

Table E.1 HC-05 pins an	l its purpose. Se	ource: AranaCorp [91]
-------------------------	-------------------	-----------------------

Pin	Name	Function
1 V		The pin state determines whether the module
1	Key	works in AT command mode or normal mode <sup>*</sup>
2	Vcc	$+5\mathrm{V}$ Positive supply needs to be given to this pin for powering the module
3	Gnd	Connect to ground
4	TYD	Serial data is transmitted by module through
4	IAD	this pin (at 9600bps by default), 3.3V logic
5	BXD	Serial data is received by module through
0	IIAD	this pin (at 9600bps by default),3.3V logic
6	State	The pin is connected to the LED on the
0	State	board to represent the state of the module

\*High=AT commands receiving mode(Commands response mode), Low or NC= Bluetooth module normally working

When setting up the Bluetooth, the process is detailed in [90],

- 1. The user must set the default baud rate of UART serial communication to 9600 in the IDE. The value is the module's default setting and can be changed in the application.
- 2. Because the module is a slave, a master is required to create a successful wireless link. Another [arduino + module (with master feature)] setup is required for this, alternatively, the user may use a mobile phone as a master and search for HC-06 slave. The HC-06 module comes with a default password of '1234,' which may be altered.

- 3. The user may also use libraries to facilitate communication. Just need to download them and add them to the header. After it is set, then it is possible to command the Arduino to transmit or receive data. This data is sent to the master over wireless Bluetooth by the module. If the module gets data from the master, it will send it to the Arduino through UART serial connection.
- 4. To connect with the module, use the AT command on a serial terminal, which may be obtained here.

This section aims to develop a code that uses HC06 Bluetooth module for telemetry and telecommand.

#### E.1 Initial Bluetooth module code

Make sure to configure the Bluetooth using a Arduino device.

```
1 /* HC06 bluetooth module setup
2
   */
3
4 // Libraries
5 #include <SoftwareSerial.h> // This library enables bluetooth communcication
   // HC06 parameters
7
                                // Connect Arduino Digital Pin 2 to HCO6 Pin Tx
s const int Rx = 2;
9 const int Tx = 3;
                                // Connect Arduino Digital Pin 3 to HCO6 Pin Rx
10 SoftwareSerial hc06(Rx, Tx); // Set Receive and Transmission pins
11
12 // Configuration parameters
  char BPS = '4';
                                      // 1=1200 , 2=2400, 3=4800, 4=9600, 5=19200, ...
13
       6=38400, 7=57600, 8=115200
14 char NAME_ID[20] = "HC06_Plathon"; // 30 characters maximum
15 char PASS[10] = "1234";
                                      // 4 bit number
16
17 // Setup
   void setup()
18
   {
19
     //Initialize Serial Monitor
20
     Serial.begin(9600);
^{21}
22
   //Initialize Bluetooth Serial Port
^{23}
    hc06.begin(9600);
^{24}
25
     // AT commands for configuration
26
     hc06.print("AT"); // Initiate sequence
27
     delay(1000); // Wait 1 second
^{28}
^{29}
     hc06.print("AT+NAME");
30
     hc06.print(NAME_ID); // Configure bluetooth device name
31
    delay(1000); // Wait 1 second
32
33
    hc06.print("AT+BAUD");
34
    hc06.print(BPS); // Transmission baud rate velocity
35
     delay(1000); // Wait 1 second
36
```

```
37
     hc06.print("AT+PIN");
38
    hc06.print(PASS); // Set Password
39
    delay(1000); // Wait 1 second
40
41
    // Print setup message
42
     Serial.println("Module set up");
^{43}
     Serial.print("PIN:");
^{44}
     Serial.println(PASS);
45
    Serial.print("Name: ");
46
    Serial.println(NAME_ID);
47
48
    // Wait 1000 ms to ensure the sensor starts correctly
49
    delay(1000);
50
51 }
52
53
   // Send and receive data
54 void loop()
55 {
     //Write data from HCO6 to Serial Monitor
56
   if (hc06.available())
57
58
   {
       Serial.write(hc06.read());
59
60
     }
61
    //Write from Serial Monitor to HCO6
62
     if (Serial.available())
63
64
    {
65
     hc06.write(Serial.read());
    }
66
67
  }
68
```

Code E.1 HC06 Bluetooth Module configuration. Source: Own.

#### E.2 Send and Receive data from module (Arduino Nano) code

```
_{\rm 1} /* This code reads all sensors data from the BNO055 IMU Sensor
   * and uses bluetooth communication
2
3 */
4
  // Libraries
5
6 #include <Wire.h>
                                // This library allows to communicate with I2C / TWI \ldots
       devices.
  #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
7
       Adafruit Unified Sensor Driver
8 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
  #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
9
       includes matrix.h, quaternions.h and vector.h)
10 #include <math.h> // Math functions
11 #include <SoftwareSerial.h> // This library enables bluetooth communcication
12
13 // Global parameters and objects
```

```
14 #define BN0055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
      (sample every 100 ms)
15
16 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
17
                               // Connect Arduino Digital Pin 2 to HC06 Pin Tx
18 const int Rx = 2;
19 const int Tx = 3;
                               // Connect Arduino Digital Pin 3 to HC06 Pin Rx
20 SoftwareSerial hc06(Rx, Tx); // Set Receive and Transmission pins
21
22 /* Function Prototypes*/
23 // Calibration function
24 void displayCalStatusSerial(void);
25 void displayCalStatusBluetooth(void);
26 void printSerial(void);
27 void prinBluetooth(void);
^{28}
29 void setup()
30 {
   //Initialize Serial Monitor
31
32 Serial.begin(9600);
33
34 // Start the IMU sensor
   IMU.begin();
35
36
    //Initialize Bluetooth Serial Port
37
    hc06.begin(9600);
38
39
   // Wait 1000 ms to ensure the sensor starts correctly
40
   delay(1000);
41
42
   // Change the time clock on the chip to the time clock on board of the IMU
43
   IMU.setExtCrystalUse(true);
44
45 }
46
47 void loop()
48 {
49
   // Print calibration data through serial
50
51 // displayCalStatusSerial();
52 // Print calibration data
    displayCalStatusBluetooth();
53
54
   // Print data over Bluetooth port
55
     printBluetooth();
56
     // Print data over Serial Port
57
    // printSerial();
58
59
60
   //Write data from HCO6 to Serial Monitor
61
   if (hc06.available())
   {
62
     Serial.write(hc06.read());
63
   }
64
65
    //Write from Serial Monitor to HC06
66
67
    if (Serial.available())
68
   {
69
     hc06.write(Serial.read());
```

70

}

```
71
     // Wait the specified delay before requesting next data
72
    delay(BN0055_SAMPLERATE_DELAY_MS);
73
74 }
75
76 // Calibration function
   void displayCalStatusSerial(void)
77
78 {
     /* Get the four calibration values (0..3)
79
80
         Any sensor data reporting 0 should be ignored,
          3 means 'fully calibrated" */
81
     uint8_t system, gyros, accel, mg = 0;
82
     IMU.getCalibration(&system, &gyros, &accel, &mg);
83
84
     // The data should be ignored until the system calibration is > \; 0
85
86
     // Display the individual values
     Serial.print(system, DEC);
87
     Serial.print(",");
88
     Serial.print(gyros, DEC);
89
    Serial.print(",");
90
   Serial.print(accel, DEC);
91
92
    Serial.print(",");
     Serial.print(mg, DEC);
93
94 }
95
   // Calibration function for bluetooth
96
97 void displayCalStatusBluetooth(void)
98 {
     /* Get the four calibration values (0..3)
99
         Any sensor data reporting 0 should be ignored,
100
          3 means 'fully calibrated" */
101
    uint8_t system, gyros, accel, mg = 0;
102
     IMU.getCalibration(&system, &gyros, &accel, &mg);
103
104
105
     // The data should be ignored until the system calibration is > 0
     // Display the individual values
106
     hc06.print(system, DEC);
107
    hc06.print(",");
108
    hc06.print(gyros, DEC);
109
    hc06.print(",");
110
111
   hc06.print(accel, DEC);
112
    hc06.print(",");
113
    hc06.print(mg, DEC);
114 }
115
116 // Print serial
117 void printSerial(void)
118 {
119
     // Work with the imu sensor from Adafruit library
    // Go to BN0055's 'imu' and bring back a vector of 4 components into 'quat' ...
120
        (quaternion) for the specific object IMU
121
     imu::Quaternion quat = IMU.getQuat();
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'acc' ...
122
        (accelerometer) for the specific object IMU
123
     imu::Vector<3> acc = IMU.getVector(Adafruit_BN0055::VECTOR_ACCELEROMETER);
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'gyro' ...
124
        (gyroscope) for the specific object IMU
```

```
125
      imu::Vector<3> gyro = IMU.getVector(Adafruit_BN0055::VECTOR_GYROSCOPE);
      // Go to BN0055's 'imu' and bring back a vector of 3 components into 'magn' \ldots
126
       (magnetometer ) for the specific object IMU
      imu::Vector<3> magn = IMU.getVector(Adafruit_BN0055::VECTOR_MAGNETOMETER);
127
128
      // Print Ouaternion data
129
130
      Serial.print(quat.w()); // Real part
      Serial.print(",");
131
      Serial.print(quat.x()); // 'i'
132
     Serial.print(",");
133
134
     Serial.print(quat.y()); // 'j'
     Serial.print(",");
135
136
     Serial.print(quat.z()); // 'k'
     Serial.print(",");
137
138
      // Print Acceleration data
139
140
     Serial.print(acc.x());
141
      Serial.print(",");
     Serial.print(acc.y());
142
     Serial.print(",");
143
     Serial.print(acc.z());
144
     Serial.print(",");
145
146
147
     Serial.print(gyro.x());
148
     Serial.print(",");
     Serial.print(gyro.x());
149
150
     Serial.print(",");
     Serial.print(gyro.x());
151
     Serial.print(",");
152
153
     Serial.print(magn.x());
154
155
   Serial.print(",");
    Serial.print(magn.y());
156
     Serial.print(",");
157
158
      Serial.println(magn.z());
159 }
160
161 // Print Bluetooth
162 void printBluetooth(void)
163 {
     // Work with the imu sensor from Adafruit library
164
    // Go to BN0055's 'imu' and bring back a vector of 4 components into 'quat' ...
165
        (quaternion) for the specific object IMU
     imu::Quaternion quat = IMU.getQuat();
166
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'acc' ...
167
        (accelerometer) for the specific object IMU
      imu::Vector<3> acc = IMU.getVector(Adafruit_BN0055::VECTOR_ACCELEROMETER);
168
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'gyro' ...
169
        (gyroscope) for the specific object IMU
170
     imu::Vector<3> gyro = IMU.getVector(Adafruit_BN0055::VecTOR_GYROSCOPE);
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'magn' \dots
171
        (magnetometer ) for the specific object IMU
172
     imu::Vector<3> magn = IMU.getVector(Adafruit_BN0055::VECTOR_MAGNETOMETER);
173
174
     // Print Quaternion data
175
     hc06.print(quat.w()); // Real part
     hc06.print(",");
176
177
     hc06.print(quat.x()); // 'i'
```

```
178
     hc06.print(",");
179
     hc06.print(quat.y()); // 'j'
     hc06.print(",");
180
     hc06.print(quat.z()); // 'k'
181
182
     hc06.print(",");
183
     // Print Acceleration data
184
     hc06.print(acc.x());
185
     hc06.print(",");
186
     hc06.print(acc.y());
187
188
     hc06.print(",");
     hc06.print(acc.z());
189
190
     hc06.print(",");
191
     hc06.print(gyro.x());
192
193
     hc06.print(",");
194
     hc06.print(gyro.x());
     hc06.print(",");
195
     hc06.print(gyro.x());
196
     hc06.print(",");
197
198
    hc06.print(magn.x());
199
    hc06.print(",");
200
    hc06.print(magn.y());
201
202
     hc06.print(",");
203
    hc06.println(magn.z());
204 }
```

 ${\bf Code \ E.2} \ {\bf Send \ data \ through \ Bluetooth \ module \ with \ Arduino \ Nano. \ Source: \ Own.$ 

#### E.3 Send and Receive data from module (STM32 Bluepill) code

```
1 /*
   * UART Bluetooth Communication with STM32
2
    */
3
4
   // Libraries
5
6 #include <Arduino.h>
7 #include <time.h>
9 // Parameters
10 char inputData_serial1 = 0; // Initialize input data from Serial 1
11 char inputData_serial = 0; // Initialize input data from Serial
12
13 // Main setup
14 void setup()
15 {
    Serial1.begin(9600);
                                               // Begin UART serial port
16
   Serial.begin(9600);
                                               // Begin Serial port
17
   Serial1.println("Bluetooth initialized"); // Print initilize message
18
19 }
20
21 // Main loop
```

```
22 void loop()
^{23}
   {
     // Check if serial is available
^{24}
     if (Serial1.available() > 0)
25
26
      {
        inputData_serial1 = Serial1.read();
27
        Serial.write(inputData_serial1); // Display message thround Serial port
^{28}
^{29}
      }
      // If Serial
30
     if (Serial.available() > 0)
31
^{32}
      {
        inputData_serial = Serial.read();
33
        Serial1.write(inputData_serial); // Display message thround Serial 1 port
34
35
      }
36
      // Delay 100 ms
37
38
     delay(100);
39
   }
```

 ${\bf Code~E.3}$  Send and receive data through HC06 with STM32. Source: Own.

Let's see an example of Bluetooth and the Arduino. First, it is needed to configure the Bluetooth parameters, this can be done using the code provided in E.1

After that, it is possible to test data sending and receiving via the HC-06 module as shown in E.2.

In the code implementation, a baud rate of 9600 was set since otherwise the Bluetooth module does not show up on the phone.



Figure E.2 Bluetooth setup schematic. Source: Own.



(a) Bluetooth setup schematic. Source: Own.

(b) Bluetooth setup schematic. Source: Own.

Figure E.3 Bluetooth Serial COM. Source: Own.

### Appendix F

### Motor

Here is presented the datasheet of each motor RF-300EA and RF-500TB from Mabuchi Motors [60] [59].

#### F.1 Motor Designations and their Meanings

Let's revise first the nomenclature of the motor. Each motor has a designation that collects the most important aspects of it.

- The types of motors are normally designated by a seven-digit code.
- The following 3 to 6 digits designate the armature winding specifications.
- The meaning of each character or figure is as shown in Figure F.1.
- However, since there are some exceptions, contact our sales office for further information.



Figure F.1 Motor designation. Source: Mabuchi Motors [92].

#### Shape of motor-housing



[Exception] Powerwindow Lift Motor Inner-rotor-type Brushless Motor

Figure F.2 Motor shape. Source: Mabuchi Motors [92].



Figure F.3 Motor brush construction. Source: Mabuchi Motors [92].

Code number for armature diameter



Figure F.4 Motor armature diameter code. Source: Mabuchi Motors [92].

Code number for housing length (mainly)



Figure F.5 Motor housing length code. Source: Mabuchi Motors [92].

#### Number of armature pole



Figure F.6 Motor number of armature pole. Source: Mabuchi Motors [92].

A: 10 poles B: 12 poles

#### Type of magnet (mainly)



Figure F.7 Motor type of magnet. Source: Mabuchi Motors [92].



Diameter of magnet wire

Figure F.8 Motor diameter of magnet wire. Source: Mabuchi Motors [92].

Number of turns of armature winding per slot



Figure F.9 Motor number of turns of armature winding per slot. Source: Mabuchi Motors [92].

Hereafter is presented the datasheets for each motor:

# **RF-300EA**



WEIGHT: 22g (APPROX)



#### OUTPUT : 0.02W~1.8W (APPROX)

Typical Applications Office Automation Equipment : CD/DVD-ROM Drive Audio and Visual Equipment : Car CD Player / DVD Player

	VOLTAGE		NO LOAD		AT MAXIMUM EFFICIENCY					STALL		
MODEL	OPERATING RANGE	NOMINAL	SPEED	CURRENT	SPEED	CURRENT	TOR	QUE	OUTPUT	TOR	QUE	CURRENT
			r/min	A	r/min	A	mN∙m	g∙cm	w	mN∙m	g∙cm	A
RF-300EA-1D390	2.8~7.0	3.9V CONSTANT	4400	0.021	3520	0.084	0.47	4.8	0.17	2.35	24	0.34
RF-300EA-8Z485	1.6~11.0	8V CONSTANT	7100	0.021	5760	0.090	0.62	6.3	0.37	3.28	33	0.39

Also available with terminals (without leadwires).







UNIT: MILLIMETERS

8.0V

Usable machine screw length 1.1 max. from motor mounting surface.

#### **RF-300EA-1D390**

25

**EFFICIENCY** [%]

#### **RF-300EA-8Z485**



MABUCHI MOTOR CO., LTD. Headquarters 430 Matsuhidai, Matsudo City, Chiba, 270-2280 Japan. Tel:81-47-710-1177 Fax:81-47-710-1132 (Sales Dept.)

3.9V

# **RF-500TB**



#### OUTPUT : 0.01W~2.0W (APPROX)

Precious metal-brush motors

MABUCHI MOTOR

WEIGHT : 45g (APPROX)

Typical Applications Audio and Visual Equipment : CD Player / DVD Player / VCR

	VOLTAGE		NO LOAD AT MA		AXIMUM EFFICIENCY			STALL				
MODEL OPERATIN RANGE	OPERATING	TING NOMINAL	SPEED	CURRENT	SPEED	CURRENT	TOR	QUE	OUTPUT	TOR	QUE	CURRENT
	RANGE		r/min	A	r/min	A	mN∙m	g∙cm	W	mN∙m	g∙cm	A
RF-500TB-14415	1.5~9.0	5V CONSTANT	3100	0.026	2540	0.12	1.23	12.6	0.33	6.86	70	0.54
RF-500TB-12560	1.5~12.0	6V CONSTANT	2700	0.020	2180	0.084	1.13	11.6	0.26	5.88	60	0.35



Usable machine screw length 1.5 max. from motor mounting surface.

UNIT: MILLIMETERS



MABUCHI MOTOR CO., LTD. Headquarters 430 Matsuhidai, Matsudo City, Chiba, 270-2280 Japan. Tel:81-47-710-1177 Fax:81-47-710-1132 (Sales Dept.)



### Appendix G

## SparkFun Qwiic Motor Driver

A motor driver is used to convert a low strength signal to high strength current signal to adequately control the power with a sufficient amount of power from an external source. This section aims to describe all the features of the SparkFun Qwiic Rob-15451 motor driver [56].

#### G.1 SparkFun Qwiic Rob-15451

As mentioned earlier, the SparkFun Qwiic Rob-15451 motor driver was chosen to control the motor.







(b) SparkFun Qwiic Motor driver ROB 15451 bottom view. Source: Sparkfun [56].



Table G.1 main shows the main characteristics and features of this motor.

SparkFun's Qwiic motor driver specifications and features
1.5 A peak drive per channel, 1.2 A steady state
Operates from 3 to 11 volts with 12 V absolute max $% \mathcal{A}$
3.3  V default VCC and logic
127 levels of DC drive strength.
Controllable by I <sup>2</sup> C or TTL UART signals
Direction inversion on a per motor basis
Global Drive enable
Exposed small heat sink shape
Several I <sup>2</sup> C addresses, default UART bauds available

 Table G.1 SparkFun's Qwiic motor driver specifications and features.

 Source: SparkFun [56].

#### G.2 Power

There are two separate power circuits on this board. The power for the motors is supplied through the VIN Connectors and it supports voltage ranges from 3.3 V to 11 V with the MAX11V and GND connections. On the other hand, power for the PSoC (Programmable System on Chip), i.e. the STM32, and the logic circuits are provided by the 3.3 V input on the Qwiic connectors. All these two power sources are needed for proper functioning.



Figure G.2 Qwiic motor driver power port. Source: SparkFun [56].

#### G.3 Qwiic connectors

SparkFun's Qwiic Connect System uses 4-pin JST connectors to quickly interface development boards with sensors, LCDs, relays and more. There are two Qwiic connectors on the board such that two motors can be controlled by a single motor driver (see Figure G.3).



Figure G.3 Qwiic motor driver port. Source: SparkFun [2].

Unlike traditional connectors, this type of connector is thinner and lighter. The 28 AWG wire's length is 150 *millim* and its properties are shown in Tables G.2 and G.3.

AWG	Conductor	Conductor	Conductor cross	Ohms	Ohms
gauge	Diameter Inches	$\mathbf{Diameter} \ \mathrm{mm}$	section in $mm^2$	$\mathbf{per} \ 1000 \ \mathrm{ft}$	$\mathbf{per} \ \mathrm{km}$
8	0.0126	0.32004	0.080	64.9	212.872

**Table G.2** 8 AWG gauge cable specifications. Source: Handbook of Electronic Tables and Formulas for American Wire Gauge [93].

Maximum A	Maximum A	Maximum frequency for	Breaking force
for	for	100% skin depth for	Soft Annealed
chassis wiring	power transmission	solid conductor copper	Cu <b>37000</b> PSI
1.4	0.226	170 kHz	4.5 lbs

Table G.3 8 AWG gauge cable specifications. Source: Handbook of Electronic Tables and Formulas for American Wire Gauge [93].

The Qwiic adapter's wires have been colour coded to red, black, blue and yellow (see Table G.4).

Color scheme	Arrangement		
Black	GND		
Red	<b>3.3</b> V		
Blue	SDA		
Yellow	SCL		

Table G.4 Color scheme. Source: SparkFun [94].



Figure G.4 Qwiic cable. Source: SparkFun [95]

As shown in the specification by the AWG standard the maximum current on a Qwiic cable is 226 mA. Despite it is possible to push it up to 1.4 A for chassis wiring in an isolated, unbundled wire in free air conditions as per Handbook of Electronic Tables and Formulas for American Wire Gauge, though. It is not recommended to stress the cable to that level but that cable can easily support a few hundreds of mA.

#### G.4 Motor ports

There are screw pin terminals at the top of the board that allows two motor connections. They are labelled on the backside of the board.



Figure G.5 Qwiic motor driver power port. Source: SparkFun [56].

Table  $\operatorname{G.5}$  summarizes the connection name and communication protocols.

Group	Name	Direction	Description	UART I <sup>2</sup> C
	A1	О	Winding of first addressable location	Motor A winding
Motor Port	A2	Ο	Winding of first addressable location	Motor A winding
	B1	Ο	Winding of second addressable location	Motor B winding
	B2	Ο	Winding of second addressable location	Motor B winding

Table G.5 Qwiic Motor driver motor ports. Source: SparkFun [56]

#### G.5 Jumpers

There are 2 sets of jumpers to configure on this board. There are pull-up enables for  $I^2C$  and 4 config bits that select operational mode.

- **I<sup>2</sup>C Jumpers**: I<sup>2</sup>C pull-up enable. Opening these disables the I<sup>2</sup>C pull-up resistors used for I<sup>2</sup>C communication. If multiple I<sup>2</sup>C devices are being used, these pull-ups should be disabled on all but one device. If UART is being used, the pull-up resistors should be disabled.
- Address Jumpers: Serial and function selection. The configuration bits are 4 bits that form a configuration nybble (half-byte). A closed jumper is a '1' and an open jumper is a '0'. See configuration Table G.6 for more information.

A1 A2 B1 B2	
<b>sparkfun</b>	<b>.</b>
	218
	0

Figure G.6 Qwiic motor driver pull up jumpers. Source: SparkFun [56].

#### G.5.1 Address bits

The configuration is set by encoding a number into the 4 config bits on the bottom of the board. Close a jumper to indicate a 1, or leave it open to indicate a 0.



Figure G.7 Qwiic motor driver jumpers. Source: SparkFun [56].

Use this table as a reference to see what the user port, address, and expansion port will become in each configuration (the default configuration is shown in **bold**):

Pattern	Mode and User port	User Address	Expansion Port
0000	UART at 9600	N/A	Master
0011	$I^2C$	0x58	Master
0100	$I^2C$	0x59	Master
0101	$I^2C$	0x5A	Master
0110	$I^2C$	0x5B	Master
0111	$I^2C$	0x5C	Master
1000	$I^2C$	0x5D	Master
1001	$I^2C$	0x5E	Master
1010	$I^2C$	0x5F	Master
1011	$I^2C$	0x60	Master
1100	$I^2C$	0x61	Master
1101	UART at $57600$	N/A	Master
1110	UART at $115200$	N/A	Master
1111	N/A	N/A	N/A

Table G.6 Motor driver configurations. Source: SparkFun [56].

#### G.6 Thermal Conduction Area

The Qwiic Motor Driver is designed to operate small robot drive motors without a heatsink; we were able to run up to about 1.1 A continuous current without going above 100 °C. Nonetheless, if a heatsink is needed, Theragrip Thermal Tape can be used to attach three Small Heat Sinks across the thermal conduction area on the back of the board.



(a) Qwiic motor driver thermal conduction area. Source: Sparkfun [95].

(b) Qwiic motor driver heatsink. Source: Sparkfun [95].

Figure G.8 Thermal system of the motor driver. Source: SparkFun [95]

#### G.7 Board dimensions

The following blueprint (Figure G.9) depicts the dimensions of the board (please note that all dimensions are in inches).



Figure G.9 Qwiic motor driver dimensions. Source: SparkFun [56].



#### G.8 Simple Impulse test code

```
1 /*
^{2}
   \star This code sets an impulse to the motor at maximum speed.
3
   * SCMD library by SparkFun used ...
4
       (https://github.com/sparkfun/SparkFun_Serial_Controlled_Motor_Driver_Arduino_Library)
   * STM32 package from http://dan.drown.org/stm32duino/package_STM32duino_index.json
5
6
    *
7 */
8
9 #include "SCMD.h"
   #include "SCMD_config.h" // Contains #defines for common SCMD register names and ...
10
       values
11
12 SCMD DriverOne; // Driver object
13
14 void setup()
15 {
     DriverOne.settings.commInterface = I2C_MODE; // Driver mode definition
16
   DriverOne.settings.I2CAddress = 0x5D; // Driver adress definition (0x5D ...
17
      by default)
^{18}
19
     while (DriverOne.begin() != 0xA9)
    { // Wait for idle
20
     Serial.println("ID Mismatch");
^{21}
     delay(200);
22
^{23}
    }
^{24}
    Serial.println("ID Match");
25
   Serial.println("Waiting for enumeration"); // Wait for peripherals (enumeration)
26
    while (DriverOne.ready() == false)
27
28
     ;
     Serial.println("Ready");
29
30
    while (DriverOne.busy())
31
     ; // Enables the driver
32
     DriverOne.enable();
33
34 }
35
36 void loop()
37 {
     DriverOne.setDrive(0, 0, 255); // Upload of values to motor
38
    DriverOne.setDrive(1, 1, 0);
39
40 }
```

Code G.1 Motor driver Simple Impulse test. Source: Own.

## Appendix H

# CAD drawings

H.1 DC Motor



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	E
	D
	С
DO NOT CHANGE THE SCALE REVISION	Final
Polytechical University of C	atalonia
DC Motor	A
endices)	A3
SCALE: 2:1 SHEET 1	

### H.2 Reaction Wheel



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perty.	DO NOT CHANGE THE SCAL	E REVISION: Final		В
	Polytechical	University of Cato	alonia	
Reaction Wheel			el	A
	N.º DRAWING 02 - Rectio	on Wheel	A3	
	SCALE: 1:1 2	SHEET 2		
	Ĺ	I		

### H.3 Top support ring


## H.4 Bottom support ring



# Appendix I

# PCB connections drawings

Below is presented the schematic of the ADCS PCB board used in PLATHON's project.

### I.1 ADCS PCB schematic

Designed by magnetorquer team [96].



## I.2 PCB drawing

Designed by magnetorquer team [97].



2		F											2
				This drawing is our property. It can't be reproduced or communicated without our written agreement.			PLACA ADCS						
				DRAWN BY Miguel Re	urer	DATE 25/05/2021	DRAWII	NG TITLE	PCE	}			
1	Taladr	°0S		CHECKED BY Jordan Mor	rales	DATE 31/05/2021	SIZE A3	DRAWING	NUMBER	1			1
	8 agujeros	R1 mm		DESIGNED E	BY I	DATE	/.0						
	14 agujeros	R1.5 mm		Miguel Reu	irer	31/05/2021	•		Escala 2:1		SHEET 1/	2	
	D						<b>I</b>	1 1			А		

## I.3 Bottom support ring

Designed by magnetorquer team [97].



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D

# Appendix J

# **ADCS** assembly diagrams

Diagram of the BNO055 and the Arduino Nano connections



Figure J.1 Diagram of the BNO055 and the Arduino Nano connections. Source: Own.

## Diagram of the MPU9250 and the Arduino Nano connections



Figure J.2 Diagram of the MPU9250 and the Arduino Nano connections. Source: Own.

Diagram of the MPU9250 and the STM3 Bluepill connections



Figure J.3 Diagram of the MPU9250 and the STM3 Bluepill connections. Source: Own.



## Bluetooth setup diagram with STM32 Bluepill

Figure J.4 Bluetooth setup diagram with STM32 Bluepill. Source: Own.

# Diagram of the motor driver and motor and the STM32 Bluepill connections



Figure J.5 Diagram of the motor driver and motor and the STM32 Bluepill connections. Source: Own.

## Final assembly diagram



 ${\bf Figure ~J.6}~{\rm Final~assembly~diagram.~Source:~Own.}$ 

# Appendix K

# Jacchia-Bowman JB2008 Atmospheric Density Model

The following code uses [28] to compute the values for density  $\rho$  and temperature T for different heights h. See Github Github

Jacchia-Bowman model uses several inputs such as EUV activity and geomagnetic indices, below is presented all the indices:

- MJD: Date and Time in Modified Julian Days and fraction (MJD = JD 2400000.5).
- Right Ascension of Sun [rad]
- Declination of Sun [rad]
- Right Ascension of Position [rad]
- Geocentric Latitude of Position [rad]
- Height of Position [km]
- F10:  $F_{10.7}$  in  $[10^{-22} \text{ W m}^{-2} \text{ Hz}^{-1}]$  (Tabular time 1.0 day earlier)
- F10B:  $F_{10.7}$  solar flux average, 81 -day centered on the input time (Tabular time 1.0 day earlier)
- S10: EUV index (26 34 nm) scaled to F10 (Tabular time 1.0 day earlier)
- S10B: EUV 81 -day ave. centered index (Tabular time 1.0 day earlier)
- XM10: MG2 index scaled to F10 (Tabular time 2.0 days earlier)
- XM10B: MG2 81 -day averaged centered index (Tabular time 2.0 days earlier)
- Y10: Solar X-Ray & Lya index scaled to F10 (Tabular time 5.0 days earlier)
- Y10B: Solar X-Ray & Lya 81-day average centered index (Tabular time 5.0 days earlier)
- $dT_c$ : Temperature change computed from Dst geomagnetic storm index

2 % Jacchia-Bowman 2008 Model Atmosphere

3 % 4 응 This is the CIRA "Integration Form" of a Jacchia Model. There are no tabular values of density. Instead, the barometric 5 ę equation and diffusion equation are integrated numerically using 6 ÷ 9 the Newton-Coates method to produce the density profile up to the 7 ŝ input position. 8 2 9 Ŷ INPUT: 10 2 11 12 % MJD : Date and Time, in modified Julian Days 13 % and Fraction (MJD = JD-2400000.5) SUN(1) : Right Ascension of Sun (radians) 14응 15 ę SUN(2) : Declination of Sun (radians) SAT(1) : Right Ascension of Position (radians) 8 16 SAT(2) : Geocentric Latitude of Position (radians) 8 17SAT(3) : Height of Position (km)  $^{18}$ 8 19 응 F10 : 10.7-cm Solar Flux (1.0E-22\*Watt/(M\*\*2\*Hertz)) 응 (Tabular time 1.0 day earlier) 2021 % F10B : 10.7-cm Solar Flux, ave. 81-day centered on the input time 22 % 23 % (Tabular time 1.0 day earlier) S10 : EUV index (26-34 nm) scaled to F10 24 % 25 % (Tabular time 1.0 day earlier) S10B : EUV 81-day ave. centered index 26 % 27응 (Tabular time 1.0 day earlier) : MG2 index scaled to F10 ÷ XM10  $^{28}$ 응 (Tabular time 2.0 days earlier) 2900 XM10B : MG2 81-day ave. centered index 30 (Tabular time 2.0 days earlier) 31 % 32 % Y10 : Solar X-Ray & Lya index scaled to F10 (Tabular time 5.0 days earlier) 33 ÷ 8 Y10B : Solar X-Ray & Lya 81-day ave. centered index 34 (Tabular time 5.0 days earlier) 2 35 DSTDTC : Temperature change computed from Dst index 36 2 37 8 38 ÷ OUTPUT: 39 2 TEMP(1): Exospheric Temperature above Input Position (deg K) 응 4041 % TEMP(2): Temperature at Input Position (deg K) : Total Mass-Desnity at Input Position (kg/m\*\*3) 42 % RHO 43 % 44 % 45 % JB2008 Model Development: (Ref. 7) 2 4647 8 2 A. Development of the JB2006 model: 48 49 8 50 % 1. Started with the CIRA72 model (Jacchia 71). 51 % 52응 2. Converted to CIRA70 model replacing equations from Jacchia 70 model (Ref. 5) 8 53 54 % 55 2 3. Replaced Tc equation using new solar indices (Ref. 1 and 2) ÷ 56578 4. Replaced semiannual equation with new global model based 8 on F10B (Ref. 1 and 3) 58 592 60 % 5. Added correction for local solar time and latitude errors

#### APPENDIX K. JACCHIA-BOWMAN JB2008 ATMOSPHERIC DENSITY MODEL

```
e
S
               (Ref. 1)
61
62
   ę
               Added smooth transition between altitude bands
63
    Ŷ
            6. Added high altitude ( z > 1500 \text{ km} ) correction
64
   ÷
   8
                (Ref. 1 and 4)
65
    2
66
    ÷
            7. REV A of JB2006 - Oct 2006
67
                      Smoothing of density corrections and scale height
68
    2
    9
                      through different altitude bands in the latitude-
69
    8
                      local time correction subroutine DTSUB
70
   8
                      dTx correction replaced with dTc correction
71
72
   응
73
   응
          B. Modification to develop JB2008 model:
   2
74
            1. Replaced Tc equation in JB2006 using new solar indices
    2
75
76
    2
                (Ref. 7)
77
    2
    ę
            2. Replaced semiannual equation with new global model based
78
               on F10B and S10B (Ref. 6)
    2
79
80
    2
            3. Use dTc value based on Dst geomagnetic storm index
81
   ę
               (This replaces ap use) (Ref. 7)
82
   ÷
   8
83
   2
84
   2
              All equation references below refer to the original
85
    9
              Jacchia 1971 (CIRA 1972) model papers.
86
    ÷
87
    9
88
89
   8
          References:
90
   8
           1. Bowman, Bruce R., etc. : "A New Empirical Thermospheric
91
   응
   9
              Density Model JB2006 Using New Solar Indices",
92
              AIAA/AAS Astrodynamics Specialists Conference, Keystone, CO,
    2
93
^{94}
              21-24 Aug 2006, (Paper AIAA 2006-6166).
    2
95
           2. Bowman, Bruce R., etc. : "Improvements in Modeling
96
              Thermospheric Densities Using New EUV and FUV Solar Indices",
97
    8
              AAS/AIAA Space Flight Mechanics Meeting, Tampa, FL,
98
    8
    Ŷ
              23-26 Jan 2006, (Paper AAS 06-237).
99
100
   ę
           3. Bowman, Bruce R.: "The Semiannual Thermospheric Density
   응
101
              Variation From 1970 to 2002 Between 200-1100 km",
102
    ŝ
              AAS/AIAA Space Flight Mechanics Meeting, Maui, HI,
103
    8
              8-12 Feb 2004, (Paper AAS 04-174).
104
    8
105
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    9
           4. Bowman, Bruce R.; "Atmospheric Density Variations at
106
              1500 km to 4000 km Height Determined from Long Term
107
   8
108
   8
              Orbit Perturbation Analysis", AAS/AIAA Space Flight
109
   ę
              Mechanics Meeting, Santa Barbara, CA, 11-14 Feb 2001,
110
   응
              (Paper AAS 01-132).
111
   2
           5. Jacchia, Luigi G.; "New Static Models of the
112
   e
S
113
    ÷
              Thermosphere and Exosphere with Empirical Temperature
              Profiles", (Smithsonian Astrophysical Observatory
114
   응
115
    8
              Special Report 313), 6 May 1970.
116
   2
           6. Bowman, Bruce R., etc. : "The Thermospheric Semiannual Density
117 %
118 %
              Response to Solar EUV Heating, " JASTP, 2008
```

#### APPENDIX K. JACCHIA-BOWMAN JB2008 ATMOSPHERIC DENSITY MODEL

```
119 %
120 %
          7. Bowman, Bruce R., etc. : "A New Empirical Thermospheric
             Density Model JB2008 Using New Solar and Geomagnetic Indices",
121 %
             AIAA/AAS 2008, COSPAR CIRA 2008 Model
122 %
123 %
124 % Last modified: 2018/01/27 M. Mahooti
125
   126
127 function [TEMP, RHO] = JB2008 (MJD, SUN, SAT, F10, F10B, S10, S10B, XM10, XM10B, Y10, Y10B, DSTDTC)
128
129 % The alpha are the thermal diffusion coefficients in Eq. (6)
130 ALPHA = [0., 0., 0., 0., -0.38];
131
132 % AL10 is log(10.0)
133 AL10 = 2.3025851;
134
135 % The AMW are the molecular weights in order: N2, O2, O, Ar, He & H
AMW = [28.0134,31.9988,15.9994,39.9480,4.0026,1.00797];
137
138 % AVOGAD is Avogadro's number in mks units (molecules/kmol)
139 AVOGAD = 6.02257e26;
140
141 TWOPI = 2*pi;
142 PIOV2 = 1.5707963;
143
144 % The FRAC are the assumed sea-level volume fractions in order:
145 % N2, O2, Ar, and He
146 FRAC = [0.78110,0.20955,9.3400e-3,1.2890e-5];
147
148 % RSTAR is the universal gas-constant in mks units (joules/K/kmol)
149 RSTAR = 8314.32;
150
151 % The R# are values used to establish height step sizes in
152 % the regimes 90km to 105km, 105km to 500km and 500km upward.
153 R1 = 0.010;
154 R2 = 0.025;
155 R3 = 0.075;
156
157 % The WT are weights for the Newton-Cotes Five-Point Quad. formula
158 WT =[0.31111111111111,1.4222222222222,0.533333333333333,...
         1.42222222222222,0.31111111111111];
159
160
161 % The CHT are coefficients for high altitude density correction
162 CHT = [0.22, -0.20e-2, 0.115e-2, -0.211e-5];
163 DEGRAD = pi/180.;
164
165 % Equation (14)
166 \text{ FN} = (F10B/240)^{(1/4)};
167 if (FN>1)
168
       FN = 1.0;
169 end
170 \text{ FSB} = \text{F10B} \times \text{FN} + \text{S10B} \times (1 - \text{FN});
   TSUBC = 392.4 + 3.227*FSB + 0.298*(F10-F10B) ...
171
                  + 2.259*(S10-S10B) + 0.312*(XM10-XM10B) ...
172
173
                  + 0.178*(Y10-Y10B);
174
175 % Equation (15)
176 \text{ ETA} = 0.5 \star \text{abs}(\text{SAT}(2) - \text{SUN}(2));
```

```
177 THETA = 0.5 * abs(SAT(2) + SUN(2));
178
179 % Equation (16)
180 H = SAT(1) - SUN(1);
181 TAU = H - 0.64577182 + 0.10471976 + \sin(H + 0.75049158);
_{182} GLAT = SAT(2);
183 ZHT
          = SAT(3);
184 GLST = H + pi;
185 GLSTHR = (GLST/DEGRAD) * (24/360);
186
187 if (GLSTHR≥24)
       GLSTHR = GLSTHR - 24;
188
189 end
190 if (GLSTHR< 0)
       GLSTHR = GLSTHR + 24;
191
192 end
193
194 % Equation (17)
195 C = \cos(ETA)^2.5;
196 S = sin(THETA)^{2.5};
197
198 DF = S + (C - S) \star abs(cos(0.5 \star TAU))^3;
199 TSUBL = TSUBC \star (1 + 0.31 \star DF);
200
201 % Compute correction to dTc for local solar time and lat correction
202 DTCLST = DTSUB (F10, GLSTHR, GLAT, ZHT);
203
204 % Compute the local exospheric temperature.
205 % Add geomagnetic storm effect from input dTc value
206 TEMP(1) = TSUBL + DSTDTC;
207 TINF = TSUBL + DSTDTC + DTCLST;
208
209 % Equation (9)
210 TSUBX = 444.3807 + 0.02385 * TINF - 392.8292 * exp(-0.0021357 * TINF);
211
212 % Equation (11)
213 GSUBX = 0.054285714 * (TSUBX - 183);
214
215 % The TC array will be an argument in the call to
216 % XLOCAL, which evaluates Equation (10) or Equation (13)
217 TC(1) = TSUBX;
_{218} TC(2) = GSUBX;
219
220 % A AND GSUBX/A OF Equation (13)
221 TC(3) = (TINF - TSUBX)/PIOV2;
_{222} TC(4) = GSUBX/TC(3);
223
224 % Equation (5)
225 Z1 = 90;
226 Z2 = min(SAT(3),105);
_{227} AL = log(Z2/Z1);
228 N = floor(AL/R1) + 1;
229 ZR = exp(AL/N);
230 AMBAR1 = XAMBAR(Z1);
231 TLOC1 = XLOCAL(Z1,TC);
232 ZEND = Z1;
233 SUM2 = 0;
234 AIN = AMBAR1 * XGRAV(Z1)/TLOC1;
```

```
236
   for I = 1:N
       Z = ZEND;
237
       ZEND = ZR * Z;
^{238}
       DZ = 0.25 * (ZEND-Z);
239
        SUM1 = WT(1) * AIN;
240
        for J = 2:5
241
             Z = Z + DZ;
242
             AMBAR2 = XAMBAR(Z);
243
             TLOC2 = XLOCAL(Z, TC);
244
^{245}
            GRAVL = XGRAV(Z);
            AIN = AMBAR2 * GRAVL/TLOC2;
^{246}
^{247}
             SUM1 = SUM1 + WT(J) * AIN;
        end
248
        SUM2 = SUM2 + DZ * SUM1;
249
250 end
251
252 FACT1 = 1000/RSTAR;
253 RHO = 3.46e-6 * AMBAR2 * TLOC1 * exp(-FACT1*SUM2)/AMBAR1/TLOC2;
254
255 % Equation (2)
256 ANM = AVOGAD * RHO;
257 AN = ANM/AMBAR2;
258
259 % Equation (3)
260 FACT2 = ANM/28.960;
261 ALN(1) = log(FRAC(1) * FACT2);
262 ALN(4) = log(FRAC(3) * FACT2);
263 ALN(5) = log(FRAC(4)*FACT2);
264
265 % Equation (4)
266 \text{ ALN}(2) = \log(\text{FACT2} * (1 + \text{FRAC}(2)) - \text{AN});
267 \text{ ALN}(3) = \log(2 \star (AN - FACT2));
268
269
   if (SAT(3) > 105)
270 else
        \text{TEMP}(2) = \text{TLOC2};
271
        % Put in negligible hydrogen for use in DO-LOOP 13
272
273
       ALN(6) = ALN(5) - 25;
        % Equation (24) - J70 Seasonal-Latitudinal Variation
274
       TRASH = (MJD - 36204) / 365.2422;
275
       CAPPHI = mod(TRASH,1);
276
        DLRSL = 0.02 \times (SAT(3) - 90) \times exp(-0.045 \times (SAT(3) - 90))...
277
               * sign_(1,SAT(2)) * sin(TWOPI * CAPPHI+ 1.72)...
278
279
               * sin(SAT(2))^2;
        % Equation (23) - Computes the semiannual variation
280
        DLRSA = 0;
281
282
        if (Z<2000)
283
             YRDAY = TMOUTD (MJD);
284
             % Use new semiannual model
             [FZZ,GTZ,DLRSA] = SEMIAN08 (YRDAY,ZHT,F10B,S10B,XM10B);
285
             if (FZZ<0)
286
287
                 DLRSA = 0;
             end
288
289
        end
290
        \ Sum the \Delta\text{-log-rhos} and apply to the number densities.
291
292
        % In CIRA72 the following equation contains an actual sum,
```

```
% namely DLR = AL10 * (DLRGM + DLRSA + DLRSL)
293
        % However, for Jacchia 70, there is no DLRGM or DLRSA.
294
        DLR = AL10 * (DLRSL + DLRSA);
295
296
        for I = 1:6
297
            ALN(I) = ALN(I) + DLR;
298
299
        end
300
        % Compute mass-density and mean-molecular-weight and
301
        % convert number density logs from natural to common.
302
303
        SUMN = 0;
        SUMNM = 0;
304
305
        for I = 1:6
306
            AN = exp(ALN(I));
307
308
            SUMN = SUMN + AN;
309
            SUMNM = SUMNM + AN * AMW(I);
            AL1ON(I) = ALN(I) / AL10;
310
311
        end
312
313
       RHO = SUMNM/AVOGAD;
314
315
        % Compute the high altitude exospheric density correction factor
        FEX = 1;
316
317
        if ((ZHT>1000)&&(ZHT<1500))
318
            ZETA = (ZHT - 1000) * 0.002;
319
            ZETA2 = ZETA * ZETA;
320
            ZETA3 = ZETA * ZETA2;
321
322
            F15C = CHT(1) + CHT(2)*F10B + CHT(3)*1500 + CHT(4)*F10B*1500;
            F15C_ZETA = (CHT(3) + CHT(4) *F10B) * 500;
323
            FEX2 = 3 * F15C - F15C_ZETA - 3;
324
            FEX3 = F15C_ZETA - 2 * F15C + 2;
325
                   = 1 + FEX2 * ZETA2 + FEX3 * ZETA3;
326
            FEX
327
        end
328
        if (ZHT \geq 1500)
329
            FEX = CHT(1) + CHT(2)*F10B + CHT(3)*ZHT + CHT(4)*F10B*ZHT;
330
331
        end
332
        % Apply the exospheric density correction factor.
333
       RHO = FEX * RHO;
334
335
        return
336 end
337
338 % Equation (6)
339 Z3 = min(SAT(3), 500);
_{340} AL = log(Z3/Z);
341 N = floor(AL/R2) + 1;
342 ZR = \exp(AL/N);
_{343} SUM2 = 0;
344 AIN = GRAVL/TLOC2;
^{345}
   for I = 1:N
346
347
       Z = ZEND;
       ZEND = ZR * Z;
348
      DZ = 0.25 * (ZEND - Z);
349
350
       SUM1 = WT(1) * AIN;
```

```
for J = 2:5
352
           Z = Z + DZ;
           TLOC3 = XLOCAL(Z,TC);
353
            GRAVL = XGRAV(Z);
354
            AIN = GRAVL/TLOC3;
355
            SUM1 = SUM1 + WT(J) * AIN;
356
357
        end
       SUM2 = SUM2 + DZ * SUM1;
358
359 end
360
_{361} Z4 = max(SAT(3), 500);
_{362} AL = log(Z4/Z);
363 R = R2;
364
   if (SAT(3) > 500)
365
    R = R3;
366
367
   end
368
_{369} N = floor(AL/R) + 1;
370 ZR = \exp(AL/N);
371 SUM3 = 0;
372
373 for I=1:N
      Z = ZEND;
374
375
       ZEND = ZR * Z;
376
       DZ = 0.25 * (ZEND - Z);
        SUM1 = WT(1) * AIN;
377
       for J = 2:5
378
379
            Z = Z + DZ;
380
           TLOC4 = XLOCAL(Z,TC);
           GRAVL = XGRAV(Z);
381
           AIN = GRAVL/TLOC4;
382
           SUM1 = SUM1 + WT(J) * AIN;
383
384
       end
385
        SUM3 = SUM3 + DZ * SUM1;
386
   end
387
388 if (SAT(3) > 500)
389
       T500 = TLOC3;
       TEMP(2) = TLOC4;
390
       ALTR = log(TLOC4/TLOC2);
391
      FACT2 = FACT1 * (SUM2 + SUM3);
392
      HSIGN = -1;
393
394 else
      T500 = TLOC4;
395
        TEMP(2) = TLOC3;
396
397
       ALTR = log(TLOC3/TLOC2);
398
      FACT2 = FACT1 * SUM2;
399
       HSIGN = 1;
400 end
401
   for I = 1:5
402
    ALN(I) = ALN(I) - (1 + ALPHA(I)) * ALTR - FACT2 * AMW(I);
403
404
   end
405
406 % Equation (7) - Note that in CIRA72, AL10T5 = log10(T500)
407 AL10T5 = log10 (TINF);
408 ALNH5 = (5.5 * AL10T5 - 39.40) * AL10T5 + 73.13;
```

```
409 ALN(6) = AL10 * (ALNH5 + 6) + HSIGN * (log(TLOC4/TLOC3) + FACT1 * SUM3 * AMW(6));
410
411 % Equation (24) - J70 Seasonal-Latitudinal Variation
412 TRASH = (MJD - 36204) / 365.2422;
413 CAPPHI = mod(TRASH, 1);
414 DLRSL = 0.02 * (SAT(3) - 90) * exp(-0.045 * (SAT(3) - 90))...
          * sign_(1,SAT(2)) * sin(TWOPI * CAPPHI+ 1.72)...
415
          * sin(SAT(2))^2;
416
417
418 % Equation (23) - Computes the semiannual variation
419 DLRSA = 0;
420 if (Z<2000)
421
        YRDAY = TMOUTD (MJD);
        % Use new semiannual model
422
       [FZZ,GTZ,DLRSA] = SEMIAN08 (YRDAY,ZHT,F10B,S10B,XM10B);
423
424
       if (FZZ<0)
425
            DLRSA = 0;
       end
426
427 end
428
429 % Sum the \Delta-log-rhos and apply to the number densities.
430 % In CIRA72 the following equation contains an actual sum,
431 % namely DLR = AL10 * (DLRGM + DLRSA + DLRSL)
432 % However, for Jacchia 70, there is no DLRGM or DLRSA.
433 DLR = AL10 \star (DLRSL + DLRSA);
434
   for I = 1:6
435
       ALN(I) = ALN(I) + DLR;
436
437 end
438
   % Compute mass-density and mean-molecular-weight and
439
440 % convert number density logs from natural to common.
441 SUMN = 0.;
442 SUMNM = 0.;
443
   for I = 1:6
444
445
       AN = exp(ALN(I));
       SUMN = SUMN + AN;
446
447
       SUMNM = SUMNM + AN*AMW(I);
       AL1ON(I) = ALN(I) / AL10;
448
449 end
450
451 RHO = SUMNM/AVOGAD:
452
   % Compute the high altitude exospheric density correction factor
453
454 FEX = 1;
455
456
   if ((ZHT>1000)&&(ZHT<1500))
457
        ZETA = (ZHT - 1000) * 0.002;
458
        ZETA2 = ZETA * ZETA;
       ZETA3 = ZETA * ZETA2;
459
       F15C = CHT(1) + CHT(2)*F10B + CHT(3)*1500 + CHT(4)*F10B*1500;
460
       F15C_ZETA = (CHT(3) + CHT(4) * F10B) * 500;
461
        FEX2 = 3 * F15C - F15C_ZETA - 3;
462
              = F15C_ZETA - 2 * F15C + 2;
463
        FEX3
        FEX
               = 1 + FEX2 * ZETA2 + FEX3 * ZETA3;
464
465 end
466
```

```
467 if (ZHT \geq 1500)
     FEX = CHT(1) + CHT(2)*F10B + CHT(3)*ZHT + CHT(4)*F10B*ZHT;
468
469
  end
470
471 % Apply the exospheric density correction factor.
472 RHO = FEX * RHO;
473
474
  end
475
  $*****
476
477 function XAMBAR2 = XAMBAR(Z)
478 % Evaluates Equation (1)
479
480 C = [28.15204, -8.5586e-2, +1.2840e-4, -1.0056e-5, -1.0210e-5, +1.5044e-6, +9.9826e-8];
481 DZ = Z - 100;
482 AMB = C(7);
483
484 for I = 1:6
     T = 7 - T:
485
     AMB = DZ \star AMB + C(J);
486
487 end
488
489 XAMBAR2 = AMB;
490
491
   end
492
   493
494 function XGRAV2 = XGRAV(Z)
495 % Evaluates Equation (8)
496 XGRAV2 = 9.80665/(1 + Z/6356.766)^2;
497
498 end
500 function XLOCAL2 = XLOCAL(Z,TC)
501
   % Evaluates Equation (10) or Equation (13), depending on Z
502 DZ = Z - 125;
503 if (DZ > 0)
     XLOCAL2 = TC(1) + TC(3) * atan(TC(4)*DZ*(1 + 4.5e-6*DZ^2.5));
504
505
      return
506
  end
      XLOCAL2 = ((-9.8204695e-6 * DZ - 7.3039742e-4) * DZ^{2} + 1) * DZ * TC(2) + TC(1);
507
508 end
509
   function DTC = DTSUB (F10,XLST,XLAT,ZHT)
510
511
   % COMPUTE dTc correction for Jacchia-Bowman model
512
513
   8
514
  8
       Calling Args:
515
   8
516
   S
       F10
              = (I) F10 FLUX
       XLST
              = (I) LOCAL SOLAR TIME (HOURS 0-23.999)
517
   8
              = (I) XLAT = SAT LAT (RAD)
   8
       XLAT
518
               = (I) ZHT = HEIGHT (KM)
519
   8
       ZHT
   Ŷ
       DTC
               = (0)
                     dTc correction
520
521
522 B = [-0.457512297e1, -0.512114909e1, -0.693003609e2,...
      0.203716701e3, 0.703316291e3, -0.194349234e4,...
523
524
       0.110651308e4, -0.174378996e3, 0.188594601e4,...
```

```
-0.709371517e4, 0.922454523e4, -0.384508073e4,...
526
         -0.645841789e1, 0.409703319e2, -0.482006560e3,...
          0.181870931e4, -0.237389204e4, 0.996703815e3,...
527
528
          0.361416936e2];
529
530 C = [-0.155986211e2, -0.512114909e1, -0.693003609e2,...
          0.203716701e3, 0.703316291e3, -0.194349234e4,...
531
          0.110651308e4, -0.220835117e3, 0.143256989e4,...
532
         -0.318481844e4, 0.328981513e4, -0.135332119e4,...
533
          0.199956489e2, -0.127093998e2, 0.212825156e2,...
534
535
         -0.275555432e1, 0.110234982e2, 0.148881951e3,...
         -0.751640284e3, 0.637876542e3, 0.127093998e2,...
536
537
         -0.212825156e2, 0.275555432e1];
538
539 DTC = 0:
540 tx = XLST/24;
541
   ycs = cos(XLAT);
542 F
       = (F10 - 100) / 100;
543
544 % calculates dTc
545 if (ZHT≥120 && ZHT≤200)
       H = (ZHT - 200)/50;
546
        DTC200 = C(17)
                                   + C(18) *tx*ycs
                                                      + C(19)*tx^2*ycs...
547
               + C(20)*tx^3*ycs + C(21)*F*ycs
                                                        + C(22) *tx*F*ycs...
548
549
               + C(23) *tx^2*F*ycs;
        sum = C(1) + B(2) *F + C(3) *tx*F
                                             + C(4) *tx^2*F...
550
            + C(5)*tx^3*F + C(6)*tx^4*F + C(7)*tx^5*F...
551
                              + C(9)*tx^2*ycs + C(10)*tx^3*ycs...
            + C(8) *tx*vcs
552
553
            + C(11) *tx^4*ycs + C(12) *tx^5*ycs + C(13) *ycs...
            + C(14) *F*ycs + C(15) *tx*F*ycs + C(16) *tx^2*F*ycs;
554
        DTC200DZ = sum;
555
       CC = 3*DTC200 - DTC200DZ;
556
       DD = DTC200 - CC;
557
       ZP = (ZHT - 120) / 80;
558
559
        DTC = CC \star ZP \star ZP + DD \star ZP \star ZP \star ZP;
   end
560
561
562 if (ZHT>200 && ZHT≤240)
       H = (ZHT - 200) / 50;
563
        sum = C(1) * H + B(2) * F * H + C(3) * tx * F * H + C(4) * tx^2 * F * H...
564
           + C(5)*tx^3*F*H + C(6)*tx^4*F*H + C(7)*tx^5*F*H...
565
            + C(8) *tx*vcs*H
                              + C(9) *tx^2*vcs*H + C(10) *tx^3*vcs*H...
566
            + C(11)*tx^4*ycs*H + C(12)*tx^5*ycs*H + C(13)*ycs*H...
567
            + C(14)*F*ycs*H
                              + C(15)*tx*F*ycs*H + C(16)*tx^2*F*ycs*H...
568
                                                    + C(19)*tx^2*ycs...
                                 + C(18)*tx*ycs
569
            + C(17)
            + C(20) *tx^3*ycs + C(21) *F*ycs
                                                     + C(22) *tx*F*ycs...
570
            + C(23) *tx^2*F*ycs;
571
572
       DTC = sum;
573 end
574
   if (ZHT>240 && ZHT<300)
575
       H = 40/50:
576
        sum = C(1) * H + B(2) * F * H + C(3) * tx * F * H
                                                  + C(4)*tx^2*F*H...
577
            + C(5) *tx^3*F*H + C(6) *tx^4*F*H
                                                    + C(7) *tx^5*F*H...
578
579
            + C(8) *tx*ycs*H
                                + C(9) *tx^2*ycs*H + C(10) *tx^3*ycs*H...
            + C(11)*tx^4*ycs*H + C(12)*tx^5*ycs*H + C(13)*ycs*H...
580
            + C(14) *F*ycs*H + C(15) *tx*F*ycs*H + C(16) *tx^2*F*ycs*H...
581
582
            + C(17)
                                 + C(18) *tx*ycs
                                                    + C(19) *tx^2*ycs...
```

```
583
           + C(20) *tx^3*ycs + C(21) *F*ycs + C(22) *tx*F*ycs...
584
            + C(23) *tx^2*F*ycs;
       AA = sum;
585
       BB = C(1) + B(2) *F + C(3) *tx*F
586
                                             + C(4) *tx^2*F...
          + C(5) * tx^3 * F + C(6) * tx^4 * F + C(7) * tx^5 * F...
587
           + C(8) *tx*ycs + C(9) *tx^2*ycs + C(10) *tx^3*ycs...
588
           + C(11) *tx^4*ycs + C(12) *tx^5*ycs + C(13) *ycs...
589
                         + C(15)*tx*F*ycs + C(16)*tx^2*F*ycs;
           + C(14) *F*ycs
590
          = 300/100;
591
       Н
       sum = B(1) + B(2) *F + B(3) *tx*F
                                                   + B(4) *tx^2*F...
592
593
           + B(5)*tx^3*F
                             + B(6)*tx^4*F
                                                 + B(7) *tx^5*F...
            + B(8)*tx*ycs
                               + B(9)*tx^2*ycs + B(10)*tx^3*ycs...
594
595
            + B(11)*tx^4*ycs + B(12)*tx^5*ycs + B(13)*H*ycs...
            + B(14) *tx*H*ycs + B(15) *tx^2*H*ycs + B(16) *tx^3*H*ycs...
596
            + B(17)*tx^4*H*ycs + B(18)*tx^5*H*ycs + B(19)*ycs;
597
598
      DTC300 = sum;
599
       sum = B(13) * ycs...
            + B(14) *tx*ycs
                             + B(15)*tx^2*ycs + B(16)*tx^3*ycs...
600
            + B(17) *tx^4*ycs + B(18) *tx^5*ycs;
601
       DTC300D7 = sum:
602
       CC = 3.*DTC300 - DTC300DZ - 3.*AA - 2.*BB;
603
       DD = DTC300 - AA - BB - CC;
604
       ZP = (ZHT - 240) / 60;
605
       DTC = AA + BB * ZP + CC * ZP * ZP + DD * ZP * ZP * ZP;
606
607 end
608
   if (ZHT>300 && ZHT≤600)
609
       H = ZHT/100;
610
       sum = B(1) + B(2) *F + B(3) *tx*F
                                                 + B(4)*tx^2*F...
611
           + B(5)*tx^3*F + B(6)*tx^4*F
                                                 + B(7)*tx^5*F...
612
            + B(8) *tx*ycs
                              + B(9) *tx^2*ycs + B(10) *tx^3*ycs...
613
           + B(11) *tx^4*ycs + B(12) *tx^5*ycs + B(13) *H*ycs...
614
            + B(14) *tx*H*ycs + B(15) *tx^2*H*ycs + B(16) *tx^3*H*ycs...
615
            + B(17)*tx^4*H*ycs + B(18)*tx^5*H*ycs + B(19)*ycs;
616
617
           DTC = sum;
618
   end
619
620 if (ZHT>600 && ZHT≤800)
       ZP = (ZHT - 600) / 100;
621
       HP = 600./100.;
622
       AA = B(1) + B(2) *F + B(3) *tx*F
                                                  + B(4)*tx^2*F...
623
                              + B(6)*tx^4*F + B(7)*tx^5*F...
            + B(5)*tx^3*F
624
           + B(8)*tx*ycs
                              + B(9)*tx^2*ycs + B(10)*tx^3*ycs...
625
            + B(11) *tx^4*ycs + B(12) *tx^5*ycs + B(13) *HP*ycs...
626
                              + B(15)*tx^2*HP*ycs+ B(16)*tx^3*HP*ycs...
627
            + B(14) *tx*HP*vcs
            + B(17)*tx^4*HP*ycs + B(18)*tx^5*HP*ycs + B(19)*ycs;
628
629
       BB = B(13) * vcs...
630
           + B(14)*tx*ycs + B(15)*tx^2*ycs + B(16)*tx^3*ycs...
631
            + B(17)*tx^4*ycs + B(18)*tx^5*ycs;
632
       CC = -(3 * AA + 4 * BB) / 4;
       DD = (AA+BB)/4;
633
       DTC = AA + BB \star ZP + CC \star ZP \star ZP + DD \star ZP \star ZP \star ZP;
634
635 end
636
637
   end
638
639
    function [FZZ,GTZ,DRLOG] = SEMIAN08(DAY,HT,F10B,S10B,XM10B)
640
```

```
641 %
642 % COMPUTE SEMIANNUAL VARIATION (DELTA LOG RHO)
643 % INPUT DAY, HEIGHT, F10BAR
644 % 025. 650. 150.
645 % OUTPUT FUNCTIONS FZ, GT, AND DEL LOG RHO VALUE
646 %
647 % DAY
             (I)
                  DAY OF YEAR
   % HT
             (I)
                   HEIGHT (KM)
648
649 % F10BAR (I) AVE 81-DAY CENTERED F10
650 % FZZ (O) SEMIANNUAL AMPLITUDE
651 % GTZ
             (O) SEMIANNUAL PHASE FUNCTION
652 % DRLOG (O) DELTA LOG RHO
653
654 TWOPI = 2*pi;
655
656 % FZ GLOBAL MODEL VALUES
657 % 1997-2006 FIT:
658 FZM = [0.2689,-0.1176e-1, 0.2782e-1, ...
         -0.2782e-1, 0.3470e-3];
659
660
661 % GT GLOBAL MODEL VALUES
662 % 1997-2006 FIT:
663 GTM = [-0.3633, 0.8506e-1, 0.2401,-0.1897, ...
          -0.2554,-0.1790e-1, 0.5650e-3,-0.6407e-3, ...
664
665
          -0.3418e-2,-0.1252e-2];
666
667 % COMPUTE NEW 81-DAY CENTERED SOLAR INDEX FOR FZ
668 FSMB = F10B - 0.70*S10B - 0.04*XM10B;
669 \text{ HTZ} = \text{HT}/1000;
670
671 FZZ = FZM(1) + FZM(2) *FSMB + FZM(3) *FSMB*HTZ ...
      + FZM(4) *FSMB*HTZ^2 + FZM(5) *FSMB^2*HTZ;
672
673
674 % COMPUTE DAILY 81-DAY CENTERED SOLAR INDEX FOR GT
675 FSMB = F10B - 0.75*S10B - 0.37*XM10B;
676
677 \text{ TAU} = (DAY-1)/365;
678 SIN1P = sin(TWOPI*TAU);
679 COS1P = cos(TWOPI*TAU);
680 SIN2P = sin(2*TWOPI*TAU);
681 \quad COS2P = \cos(2 \star TWOPI \star TAU);
682
683 GTZ = GTM(1) + GTM(2) * SIN1P + GTM(3) * COS1P ...
     + GTM(4)*SIN2P + GTM(5)*COS2P ...
684
        + GTM(6) *FSMB ...
685
       + GTM(7)*FSMB*SIN1P + GTM( 8)*FSMB*COS1P ...
686
       + GTM(9) *FSMB*SIN2P + GTM(10) *FSMB*COS2P;
687
688
689
   if (FZZ<1e-6)
      FZZ = 1e-6;
690
691 end
692
693 DRLOG = FZZ*GTZ;
694
695
   end
696
697
   $********
   function doy = TMOUTD(MJD)
698
```

```
699
700 [year,month,day,hr,min,sec] = invjday(MJD+2400000.5);
701 doy = finddays(year,month,day,hr,min,sec);
702
703 end
```

**Code K.1** Jacchia-Bowman JB2008 Atmospheric Density Model. Source: Matlab and Jacchia-Bowman [98] [28].

```
1 %------%
2 % Temperature and density plots using Jacchia-Bowman 2008 Model Atmosphere
3
4 % Date: 17/04/2021
5 % Author/s: Yi Qiang Ji Zhang
6
  7
8
9 clc;
10 close all;
11 clear all;
12 format long g
^{13}
14 % Set interpreter to latex
15 set(groot, 'defaultAxesTickLabelInterpreter', 'latex');
16 set(groot, 'defaulttextinterpreter', 'latex');
17 set(groot,'defaultLegendInterpreter','latex');
18
 global const PC
19
20
21 SAT_Const
22 constants
23 load DE430Coeff.mat
24 PC = DE430Coeff;
25
26 % read Earth orientation parameters
27 fid = fopen('eop19620101.txt','r');
  8 ...
^{28}
 % | Date MJD x y UT1-UTC LOD dPsi dEpsilon dX ...
^{29}
    dY DAT
                                               ....
                  ....
                         " S S
  % |(Oh UTC)
30
          .....
               s
31 % ...
32 eopdata = fscanf(fid, '%i %d %d %i %f %f %f %f %f %f %f %f %f %i ,[13 inf]);
33 fclose(fid);
^{34}
35 % read space weather data
36 fid = fopen('SOLFSMY.txt','r');
37 % -----
38
  % | YYYY DDD JulianDay F10 F81c S10 S81c M10 M81c Y10 Y81c
39 % -----
40 SOLdata = fscanf(fid, '%d %d %f ,[11 inf]);
41 fclose(fid);
42
^{43}
```

```
44 %% Density at h=400km for different years
^{45}
46 % Dates vector
47 dates = 1999:1:2015;
48
49 % Temperature vector
  temp_dates = zeros(1,length(dates));
50
51
52 % Density vector
53 rho_dates = zeros(1, length(dates));
54
55 % Height vector
56 height = [400 500 600];
57
58\, % Strings (for automatic plotting)
59 str = strings([1, length(height)]);
60
61 % Loop for each height
62 for k=1:1:length(height)
63
      % Loop for each date
      for j=1:1:length(dates)
64
          year = dates(j);
65
          doy = 200;
66
67
          [month, day, hour, minute, sec] = days2mdh(year, doy);
          MJD = Mjday(year,month,day,hour,minute,sec);
68
69
          % READ SOLAR INDICES
70
           % USE 1 DAY LAG FOR F10 AND S10 FOR JB2008
71
72
          JD = floor(MJD-1+2400000.5);
          i = find(JD==SOLdata(3,:),1,'first');
73
          SOL = SOLdata(:,i);
74
          F10 = SOL(4);
75
          F10B = SOL(5);
76
77
          S10 = SOL(6);
           S10B = SOL(7);
^{78}
79
          % USE 2 DAY LAG FOR M10 FOR JB2008
80
          SOL = SOLdata(:,i-1);
81
          XM10 = SOL(8);
82
          XM10B = SOL(9);
83
84
          % USE 5 DAY LAG FOR Y10 FOR JB2008
85
          SOL = SOLdata(:,i-4);
86
          Y10 = SOL(10);
87
           Y10B = SOL(11);
88
89
           % READ GEOMAGNETIC STORM DTC VALUE
90
           fid = fopen('DTCFILE.txt','r');
^{91}
           § _____
92
           % | YYYY DDD DTC1 to DTC24
93
           010
94
           95
       %d %d %d %d %d %d %d',[26 inf]);
           fclose(fid);
96
97
           doy = finddays(year,month,day,hour,minute,sec);
98
           i = find(year==DTCdata(1,:) & floor(doy)==DTCdata(2,:),1,'first');
99
100
           DTC = DTCdata(:,i);
```

```
ii = floor(hour) + 3:
101
102
            DSTDTC = DTC(ii);
103
             % CONVERT POINT OF INTEREST LOCATION (RADIANS AND KM)
104
             % CONVERT LONGITUDE TO RA
105
             [x_pole, y_pole, UT1_UTC, LOD, dpsi, deps, dx_pole, dy_pole, TAI_UTC] = ...
106
        IERS(eopdata,MJD,'1');
             [UT1_TAI,UTC_GPS,UT1_GPS,TT_UTC,GPS_UTC] = timediff(UT1_UTC,TAI_UTC);
107
             [DJMJD0, DATE] = iauCal2jd(year, month, day);
108
            TIME = (60 * (60 * hour + minute) + sec) / 86400;
109
110
            UTC = DATE+TIME;
            TT = UTC+TT_UTC/86400;
111
112
            TUT = TIME+UT1_UTC/86400;
            UT1 = DATE+TUT;
113
            GWRAS = iauGmst06(DJMJD0, UT1, DJMJD0, TT);
114
115
            XLON = 60*const.Rad;
116
            SAT(1) = mod(GWRAS + XLON, 2*pi);
            SAT(2) = -70 \star const.Rad;
117
            SAT(3) = height(k);
118
119
            % SET Sun's right ascension and declination (RADIANS)
120
             % Difference between ephemeris time and universal time
121
122
             % JD = MJD_UTC+2400000.5;
             % [year, month, day, hour, minute, sec] = invjday(JD);
123
124
             % days = finddays(year, month, day, hour, minute, sec);
             % ET_UT = ETminUT(year+days/365.25);
125
             % MJD_ET = MJD_UTC+ET_UT/86400;
126
             % [r_Mercury,r_Venus,r_Earth,r_Mars,r_Jupiter,r_Saturn,r_Uranus, ...
127
128
             % r_Neptune,r_Pluto,r_Moon,r_Sun,r_SunSSB] = JPL_Eph_DE430(MJD_ET);
129
            MJD_TDB = Mjday_TDB(TT);
130
            [r_Mercury,r_Venus,r_Earth,r_Mars,r_Jupiter,r_Saturn,r_Uranus, ...
131
             r_Neptune,r_Pluto,r_Moon,r_Sun,r_SunSSB] = JPL_Eph_DE430(MJD_TDB);
132
133
             ra_Sun = atan2(r_Sun(2), r_Sun(1));
134
             dec_Sun = atan2(r_Sun(3), sqrt(r_Sun(1)^2+r_Sun(2)^2));
                    = ra_Sun;
             SUN(1)
135
             SUN(2) = dec_Sun;
136
137
            % COMPUTE DENSITY KG/M3 RHO
138
             [TEMP, RHO] = JB2008 (MJD, SUN, SAT, F10, F10B, S10, S10B, XM10, XM10B, Y10, Y10B, DSTDTC);
139
140
            % Save temperature
141
             temp_dates(k,j) = TEMP(2);
142
143
             % Save density in density vector
144
             rho_dates(k, j) = RHO;
145
146
147
        end
148
        % Concatenate strings (for automatic plotting)
        str(k) = {strcat('$h$ = ', '\,', num2str(SAT(3)), '\,', '$\mathrm{km}$')};
149
150
   end
151
152
153
154 % Plots
155
156 title1 = strcat('\textbf{Temperature $T$ }');
157 title2 = strcat('\textbf{Density $\rho$ }');
```

```
159 plot_pdf1 = figure;
160 plot(dates,temp_dates);
161 grid on;
162 grid minor;
163 title(title1);
164 legend(str(:));
165 xlabel('Date [year]');
166 ylabel('Temperature [$^\circ K$]');
167
168 plot_pdf2 = figure;
169 plot(dates,rho_dates);
170 grid on;
171 grid minor;
172 title(title2);
173 legend(str(:));
174 xlabel('Date [year]');
175 ylabel('Density [$\mathrm{kq}/\mathrm{m^3}$]');
176
177
178 % Save plots in .pdf and .png
179
180 % % Figure 1
181 % set(plot_pdf1, 'Units', 'Centimeters');
182 % pos = get(plot_pdf1, 'Position');
183 % set(plot_pdf1, 'PaperPositionMode', 'Auto', 'PaperUnits', 'Centimeters', ...
184 % 'PaperSize', [pos(3), pos(4)]);
185 % print(plot_pdf1, 'JB2008_temperature_vs_height.pdf', '-dpdf', '-r0');
186 %
187 % % Save png
188 % print (plot_pdf1,'JB2008_temperature_vs_height.png','-dpng','-r1000');
189 %
190 % % Figure 2
191 % set(plot_pdf2, 'Units', 'Centimeters');
192 % pos = get(plot_pdf2, 'Position');
193 % set(plot_pdf2, 'PaperPositionMode', 'Auto', 'PaperUnits', 'Centimeters', ...
194 응
            'PaperSize', [pos(3), pos(4)]);
195 % print(plot_pdf2, 'JB2008_density_vs_height.pdf', '-dpdf', '-r0');
196 %
197 % % Save png
198 % print(plot_pdf2,'JB2008_density_vs_height.png','-dpng','-r1000');
```

Code K.2 Jacchia-Bowman JB2008 Atmospheric Density Model. Source: Own.

```
1 %------%
2 % Aerodynamic torque as a function of height, CD and Solar Acticity
3
4 % Date: 17/04/2021
5 % Author/s: Yi Qiang Ji Zhang
6
7 %------%
8
9 clc;
10 close all;
11 clear all;
12 format long g
```

```
13
14 % Set interpreter to latex
15 set(groot, 'defaultAxesTickLabelInterpreter', 'latex');
16 set(groot,'defaulttextinterpreter','latex');
17 set(groot,'defaultLegendInterpreter','latex');
18
19 global const PC
^{20}
21 SAT_Const
22 constants
23 load DE430Coeff.mat
24 PC = DE430Coeff;
25
26 % read Earth orientation parameters
27 fid = fopen('eop19620101.txt','r');
28 % ...
29 % | Date MJD x y UT1-UTC LOD dPsi dEpsilon
                                                                          dX ...
         dy dat
                                                                     " ...
                      .....
                             S S
  % |(Oh UTC)
30
            .....
                 s
31 % ...
32 eopdata = fscanf(fid, '%i %d %d %i %f %f %f %f %f %f %f %f %i ',[13 inf]);
33 fclose(fid);
34
35 % read space weather data
36 fid = fopen('SOLFSMY.txt','r');
37 % -----
38 % | YYYY DDD JulianDay F10 F81c S10 S81c M10 M81c Y10 Y81c
39 🖁 -----
40 SOLdata = fscanf(fid, '%d %d %f ,[11 inf]);
41 fclose(fid);
42
^{43}
44 %% Aerodynamic torque as a function of height, CD and Solar Acticity
45
46 % Dates vector
47 dates = [2005];
^{48}
49 % Temperature vector
50 temp_dates = zeros(1, length(dates));
51
52 % Density vector
53 rho_dates = zeros(1, length(dates));
54
55 % Height vector
56 height = 400:50:700; % [km]
57
58 % Strings (for automatic plotting)
59 str = strings([1,length(height)]);
60
61 % Drag coefficient
62 CD = 1:0.1:4;
63
64 % F10_7 Solar activity
65 F10_7 = 150; % Use [65 150 250 300]
66
```

```
67 % Cubesat Wetted Area
68 Area = 0.01; % [m<sup>2</sup>]
69
70 % Cubesat velocity
G = 6.67408e - 11;
72 M_E = 5.9742e24;
73 R_E = 6371e3;
74 velocity = zeros(1, length(height));
75
76~ % Location of the center of the atmospheric force
r_dA = 0.03; % [m]
78
79 % Exchange coeff
so sigma_n = 0.8;
81 sigma_t = 0.8;
82
83 % Density data
s4 rho_data = zeros(length(height), length(CD));
85
86
87 % Loop for each height
ss for k=1:1:length(height)
       % Loop for every CD
89
       for drag_counter=1:1:length(CD)
90
           % Loop for each date
91
           for j=1:1:length(dates)
92
               year = dates(j);
93
               doy = 200;
94
95
               [month, day, hour, minute, sec] = days2mdh(year, doy);
               MJD = Mjday(year, month, day, hour, minute, sec);
96
97
               % READ SOLAR INDICES
98
               % USE 1 DAY LAG FOR F10 AND S10 FOR JB2008
99
               JD = floor(MJD-1+2400000.5);
100
101
               i = find(JD==SOLdata(3,:),1,'first');
102
               SOL = SOLdata(:,i);
               F10 = F10_7;
103
               F10B = 0;
104
105
               S10 = SOL(6);
               S10B = SOL(7);
106
107
               % USE 2 DAY LAG FOR M10 FOR JB2008
108
               SOL = SOLdata(:,i-1);
109
               XM10 = SOL(8);
110
               XM10B = SOL(9);
111
112
               % USE 5 DAY LAG FOR Y10 FOR JB2008
113
114
               SOL = SOLdata(:,i-4);
115
               Y10 = SOL(10);
116
               Y10B = SOL(11);
117
               % READ GEOMAGNETIC STORM DTC VALUE
118
               fid = fopen('DTCFILE.txt','r');
119
120
               § _____
121
               % | YYYY DDD DTC1 to DTC24
               8
122
               123
      %d %d %d %d %d %d %d',[26 inf]);
```

```
fclose(fid);
124
125
                 doy = finddays(year,month,day,hour,minute,sec);
126
127
                 i = find(year==DTCdata(1,:) & floor(doy)==DTCdata(2,:),1,'first');
                 DTC = DTCdata(:,i);
128
                 ii = floor(hour) + 3:
129
                 DSTDTC = DTC(ii);
130
131
                 % CONVERT POINT OF INTEREST LOCATION (RADIANS AND KM)
132
                 % CONVERT LONGITUDE TO RA
133
                 [x_pole,y_pole,UT1_UTC,LOD,dpsi,deps,dx_pole,dy_pole,TAI_UTC] = ...
134
        IERS(eopdata,MJD,'l');
                 [UT1_TAI,UTC_GPS,UT1_GPS,TT_UTC,GPS_UTC] = timediff(UT1_UTC,TAI_UTC);
135
                 [DJMJD0, DATE] = iauCal2jd(year, month, day);
136
                 TIME = (60*(60*hour+minute)+sec)/86400;
137
138
                 UTC = DATE+TIME;
139
                 TT = UTC+TT_UTC/86400;
                 TUT = TIME+UT1_UTC/86400;
140
                 UT1 = DATE+TUT:
141
                 GWRAS = iauGmst06(DJMJD0, UT1, DJMJD0, TT);
142
                 XLON = 60*const.Rad;
143
                 SAT(1) = mod(GWRAS + XLON, 2*pi);
144
                 SAT(2) = -70 \star const.Rad;
145
                 SAT(3) = height(k);
146
147
                 % SET Sun's right ascension and declination (RADIANS)
148
                 % Difference between ephemeris time and universal time
149
                 % JD = MJD_UTC+2400000.5;
150
151
                 % [year, month, day, hour, minute, sec] = invjday(JD);
                 % days = finddays(year, month, day, hour, minute, sec);
152
                 % ET_UT = ETminUT(year+days/365.25);
153
                 % MJD_ET = MJD_UTC+ET_UT/86400;
154
                 % [r_Mercury,r_Venus,r_Earth,r_Mars,r_Jupiter,r_Saturn,r_Uranus, ...
155
156
                 % r_Neptune,r_Pluto,r_Moon,r_Sun,r_SunSSB] = JPL_Eph_DE430(MJD_ET);
157
                 MJD_TDB = Mjday_TDB(TT);
158
                 [r_Mercury,r_Venus,r_Earth,r_Mars,r_Jupiter,r_Saturn,r_Uranus, ...
159
                 r_Neptune,r_Pluto,r_Moon,r_Sun,r_SunSSB] = JPL_Eph_DE430(MJD_TDB);
160
                 ra_Sun = atan2(r_Sun(2), r_Sun(1));
161
                 dec_Sun = atan2(r_Sun(3), sqrt(r_Sun(1)^2+r_Sun(2)^2));
162
                 SUN(1) = ra_Sun;
163
                 SUN(2) = dec Sun;
164
165
                 % COMPUTE DENSITY KG/M3 RHO
166
                 [TEMP, RHO] = JB2008 (MJD, SUN, SAT, F10, F10B, S10, S10B, XM10, XM10B, Y10, Y10B, DSTDTC);
167
168
169
                 % Density for each height and drag
170
                 rho_data(k,drag_counter) = RHO;
171
                 velocity(k) = sqrt((G*M_E)/(R_E+height(k))); % [m/s]
172
            end
        end
173
174
    end
175
    % Estimate a mean orbital velocity
176
177
   mean_velocity = mean(velocity);
178
179 plot_pdf1 = figure;
   [X_CD, Y_height] = meshgrid(CD,height);
180
```

### APPENDIX K. JACCHIA-BOWMAN JB2008 ATMOSPHERIC DENSITY MODEL

```
181 % Calculate aero_torque
182 aero_torque = 0.5.*X_CD.*rho_data*Area*mean_velocity^2*r_dA*3/sqrt(2);
183 % Countourf to plot the data
184 contourf(X_CD,Y_height,aero_torque,'ShowText','on');
185 % shading interp (optional)
186 title('\textbf{Moderate Solar activity: F10.7 = 150 $\mathbf{s.f.u.}$}');
187 xlabel('Drag coefficient $C_D$ [adim]');
188 ylabel('Orbital Altitude $h$ [$\mathrm{km}$]');
189 colorbar_label = colorbar;
190 colorbar_label.Label.Interpreter = 'latex';
191 set(colorbar_label, 'FontSize',11);
192 colorbar_label.Label.String = 'Atmospheric drag torque [$\mathrm{N \cdot m}$]';
193 grid on;
194
195 % Save plots in .pdf and .png
196
197 % % Figure 1
198 % set(plot_pdf1, 'Units', 'Centimeters');
199 % pos = get(plot_pdf1, 'Position');
200 % set(plot_pdf1, 'PaperPositionMode', 'Auto', 'PaperUnits', 'Centimeters', ...
201 % 'PaperSize', [pos(3), pos(4)]);
202 % print(plot_pdf1, 'JB2008_F107_150.pdf', '-dpdf', '-r0');
203 %
204 % % Save png
205 % print(plot_pdf1,'JB2008_F107_150.png','-dpng','-r1000');
```

**Code K.3** Jacchia-Bowman JB2008 Evolution of the atmosphere torque in terms of  $h C_D$  and solar activity. Source: Own.

## Appendix L

# IMU Data Acquisition code

### L.1 BNO055 IMU

The following code collects the data from the BNO055 IMU sensor [53].

First, the IMU must be calibrated:

```
_{\rm 1} /* This code reads calibrates the BNO055 IMU Sensor
   */
^{2}
3
4 // Libraries
5 #include <Wire.h> // This library allows to communicate with I2C / TWI devices.
6 #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
       Adafruit Unified Sensor Driver
7 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
8 #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
       includes matrix.h, quaternions.h and vector.h)
9
   // Global parameters and objects
10
11 #define BN0055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
        (sample every 100 ms)
12
13 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
14
15 // Calibration function
   void displayCalStatus(void)
16
17
   {
     /* Get the four calibration values (0..3)
18
         Any sensor data reporting 0 should be ignored,
19
         3 means 'fully calibrated" */
20
^{21}
   uint8_t system, gyros, accel, mg =0;
^{22}
   IMU.getCalibration(&system, &gyros, &accel, &mg);
23
     // The data should be ignored until the system calibration is > 0
24
```

```
Serial.print("\t");
25
   if (!system)
26
^{27}
   {
     Serial.print("! ");
^{28}
     }
29
30
   // Display the individual values
^{31}
    Serial.print("Sys:");
32
    Serial.print(system, DEC);
33
   Serial.print(" Gyros:");
34
   Serial.print(gyros, DEC);
35
   Serial.print(" Accel:");
36
37
   Serial.print(accel, DEC);
   Serial.print(" Magne:");
38
    Serial.println(mg, DEC);
39
40 }
41
42 void setup() {
   // Set the baud rate speed. This is how fast the data is to be sent through the \ldots
43
      USB connection
44
   Serial.begin(115200);
45
46
   // Print initial message
   Serial.println("BNO055 IMU Sensor Raw Data Initialized");
47
^{48}
    Serial.println("");
49
    // Start the IMU sensor
50
    IMU.begin();
51
52
53
   // Wait 1000 ms to ensure the sensor starts correctly
   delay(1000);
54
55
    // Change the time clock on the chip to the time clock on board of the IMU
56
    IMU.setExtCrystalUse(true);
57
58 }
59
60 void loop() {
61
62
   displayCalStatus();
63
   // Wait the specified delay before requesting next data
64
   delay(BNO055_SAMPLERATE_DELAY_MS);
65
66
67 }
```

Code L.1 Arduino IMU calibration code. Source: Own.



Figure L.1 Configuration of the BNO555 and the Arduino Nano connections. Source: Own.

The following code collects the data from the BNO055 IMU sensor [53].

To begin to use the code, two libraries must be installed:

```
    Adafruit Unified Sensor Driver
    Adafruit Unified BN0055 Driver (AHRS/Orientation)
```

Code L.2 Adafruit Libraries for Arduino IDE. Source: Own

The first library incorporates all Adafruit Sensor Drivers, from accelerometers, gyroscopes, magnetometers, barometric pressure, etc.. Besides, the Adafruit Unified Sensor Library provides a common interface and data type for any supported sensor. It defines some basic information about the sensor (sensor limits, etc.), and returns standard SI units of a specific type and scale for each supported sensor type. [87].

The second library is the driver for the Adafruit BNO055 Breakout, and is based on Adafruit's Unified Sensor Library (Adafruit\_Sensor) [53].

First, the IMU must be calibrated using L calibration function. The calibration process is simple, 3 steps are needed to calibrate the IMU:

- To calibrate the **gyroscope**, just let the IMU rest in a flat surface.
- To calibrate the **accelerometer**, tilt the IMU 45 ° in all 3 axis x, y, z.
- To calibrate the **magnetometer**, swing the IMU and rotate it in all axis.

Once all of the instruments are correctly calibrated (calibration are ranged from 0-3, when the output
is 3, the sensor is correctly calibrated), the system parameter should be calibrated as well.

The code implemented to extract the Acceleration, Gyroscope and Magnetometer's data is shown below [99]:

```
1 /* This code reads [ACCELEROMETER, GYROSCOPE and MAGNETOMETER'S] raw data from the ...
       BNO055 IMU Sensor
2
      Connections
3
      _____
4
\mathbf{5}
      Connect SCL to SCL pin (analog 5 on Arduino UNO)
     Connect SDA to SDA pin (analog 4 on Arduino UNO)
6
     Connect Vin to 3-5V DC (depending on your board's logic level)
7
     Connect GROUND to common ground
8
9
10
   */
11
12 /* Units
13
   - VECTOR_ACCELEROMETER - m/s^2
14
   - VECTOR_MAGNETOMETER - uT
15
16
   - VECTOR_GYROSCOPE - rad/s
   - VECTOR_EULER
                          - degrees
17
    - VECTOR_LINEARACCEL - m/s^2
18
   - VECTOR_GRAVITY
                           - m/s^2
19
   */
20
^{21}
22 // Libraries
23 #include <Wire.h> // This library allows to communicate with I2C / TWI devices.
24 #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
       Adafruit Unified Sensor Driver
25 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
26 #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
       includes matrix.h, quaternions.h and vector.h)
^{27}
   // Global parameters and objects
^{28}
  #define BN0055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
29
       (sample every 100 ms)
30
31 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
32
   // Calibration function
33
   void displayCalStatus(void)
34
35
   {
     /* Get the four calibration values (0..3)
36
         Any sensor data reporting 0 should be ignored,
37
38
         3 means 'fully calibrated" */
39
   uint8_t system, gyros, accel, mg =0;
   IMU.getCalibration(&system, &gyros, &accel, &mg);
40
41
     // The data should be ignored until the system calibration is > 0 \,
42
     Serial.print("\t");
43
     if (!system)
44
45
     {
```

```
Serial.print("! ");
46
47
     }
48
   // Display the individual values
49
   Serial.print("Sys:");
50
   Serial.print(system, DEC);
51
    Serial.print(" Gyros:");
52
     Serial.print(gyros, DEC);
53
     Serial.print(" Accel:");
54
     Serial.print(accel, DEC);
55
   Serial.print(" Magne:");
56
     Serial.print(mg, DEC);
57
58 }
59
60 void setup() {
    // Set the baud rate speed. This is how fast the data is to be sent through the \ldots
61
       USB connection
     Serial.begin(115200);
62
63
   // Print initial message
64
   Serial.println("BNO055 IMU Sensor Raw Data Initialized");
65
    Serial.println("");
66
67
    // Start the IMU sensor
68
     IMU.begin();
69
70
     // Wait 1000 ms to ensure the sensor starts correctly
71
     delay(1000);
72
73
     // Get temperature from the IMU and save it into a 8-bit int variable named 'temp'
74
     int8_t temp = IMU.getTemp();
75
76
     // Print the temperature through the Serial Monitor
77
     Serial.println(temp);
78
79
     // Change the time clock on the chip to the time clock on board of the IMU
80
     IMU.setExtCrystalUse(true);
81
82 }
83
84 void loop() {
85
     // Display Calibration status (This line is optional)
86
     // displayCalStatus();
87
88
     // Work with the imu sensor from Adafruit library
89
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'acc' ...
90
       (accelerometer) for the specific object IMU
     imu::Vector<3> acc = IMU.getVector(Adafruit_BN0055::VECTOR_ACCELEROMETER);
^{91}
^{92}
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'gyro' ...
       (gyroscope) for the specific object IMU
     imu::Vector<3> gyro = IMU.getVector(Adafruit_BN0055::VECTOR_GYROSCOPE);
93
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'mag' \ldots
^{94}
        (magnetometer) for the specific object IMU
     imu::Vector<3> mag = IMU.getVector(Adafruit_BNO055::VECTOR_MAGNETOMETER);
95
96
     // Print Acceleration data
97
    Serial.print(acc.x());
98
     Serial.print(",");
99
```

```
100
      Serial.print(acc.y());
101
      Serial.print(",");
      Serial.print(acc.z());
102
      Serial.print(",");
103
104
      // Print Gyroscope data
105
      Serial.print(gyro.x());
106
      Serial.print(",");
107
      Serial.print(gyro.y());
108
      Serial.print(",");
109
110
      Serial.print(gyro.z());
      Serial.print(",");
111
112
     // Print Magnetometer data
113
     Serial.print(mag.x());
114
115
     Serial.print(",");
116
     Serial.print(mag.y());
      Serial.print(",");
117
     Serial.println(mag.z());
118
119
     // Wait the specified delay before requesting next data
120
     delay(BN0055_SAMPLERATE_DELAY_MS);
121
122
123 }
```

Code L.3 Arduino and BNO055 IMU Raw data acquisition code. Source: Own.

The IMU has an on-board algorithm for continuous real-time generating quaternions.

### L.2 Pitch and roll from accelerometers code

```
/* This code reads [ACCELEROMETER] from the BN0055 IMU Sensor and calculates pitch ...
1
       and roll angles
2
  */
3
   // Libraries
4
  #include <Wire.h> // This library allows to communicate with I2C / TWI devices.
5
6 #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
       Adafruit Unified Sensor Driver
7 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
   #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
8
       includes matrix.h, quaternions.h and vector.h)
  #include <math.h> // Math functions
9
10
   // Global parameters and objects
11
12 #define BNO055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
       (sample every 100 ms)
13
14 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
15
16 const float pi = 3.1415926535; // Number Pi
```

```
17
18 //Variables
19 float theta, phi;
20
21 void setup() {
    // Set the baud rate speed. This is how fast the data is to be sent through the \ldots
22
       USB connection
     Serial.begin(115200);
^{23}
^{24}
     // Print initial message
25
     Serial.println("BNO055 IMU Sensor Raw Data Initialized");
26
     Serial.println("");
27
^{28}
     // Start the IMU sensor
29
    IMU.begin();
30
^{31}
32
     // Wait 1000 ms to ensure the sensor starts correctly
     delay(1000);
33
34
     // Change the time clock on the chip to the time clock on board of the IMU
35
     IMU.setExtCrystalUse(true);
36
37 }
38
   void loop() {
39
40
     // Work with the imu sensor from Adafruit library
^{41}
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'acc' ...
42
        (accelerometer) for the specific object IMU
     imu::Vector<3> acc = IMU.getVector(Adafruit_BN0055::VECTOR_ACCELEROMETER);
43
44
     // Parameters
45
     // Tilt approximation and normalize it to '1q' and convert into degrees
46
     theta = - \operatorname{atan2}(\operatorname{acc.x}()/9.81, \operatorname{acc.z}()/9.81)/(2*\operatorname{pi})*360;
47
     phi = - atan2(acc.y()/9.81, acc.z()/9.81)/(2*pi)*360;
^{48}
49
     // Print Acceleration data
50
     Serial.print(acc.x()/9.81);
51
     Serial.print(",");
52
     Serial.print(acc.y()/9.81);
53
     Serial.print(",");
54
    Serial.print(acc.z()/9.81);
55
    Serial.print(",");
56
     Serial.print(theta);
57
     Serial.print(",");
58
     Serial.println(phi);
59
60
     // Wait the specified delay before requesting next data
61
^{62}
    delay(BNO055_SAMPLERATE_DELAY_MS);
63
64 }
```

Code L.4 Arduino BNO055 IMU pitch and roll code. Source: Own.

### L.3 Complementary filter code

```
_{\rm 1} /* This code reads [ACCELEROMETER] and [GYROSCOPE] data from the BN0055 IMU Sensor
2
   * with a Complementary Filter for the output signal
3 */
4
5 // Libraries
6 #include <Wire.h> // This library allows to communicate with I2C / TWI devices.
7 #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
       Adafruit Unified Sensor Driver
8 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
9 #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
       includes matrix.h, quaternions.h and vector.h)
  #include <math.h> // Math functions
10
11
12 // Global parameters and objects
13 #define BN0055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
       (sample every 100 ms)
14
15 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
16
17 const float pi = 3.1415926535; // Number Pi
18
19 //Variables
20 float theta, phi;
21 float theta_meas, phi_meas;
22 float theta_filt_old = 0, phi_filt_old = 0;
23 float theta_filt_new, phi_filt_new;
^{24}
25 float theta_gyro = 0, phi_gyro = 0;
26 float Dt; // Change in time
  unsigned long millisOld; // Marker of t[n-1]
27
28
29 void setup() {
   // Set the baud rate speed. This is how fast the data is to be sent through the ...
30
      USB connection
    Serial.begin(115200);
^{31}
32
    // Print initial message
33
34
    Serial.println("BN0055 IMU Sensor Raw Data Initialized");
    Serial.println("");
35
36
    // Start the IMU sensor
37
    IMU.begin();
38
39
    // Wait 1000 ms to ensure the sensor starts correctly
40
    delay(1000);
41
42
    // Change the time clock on the chip to the time clock on board of the IMU
43
     IMU.setExtCrystalUse(true);
44
45
    // Start millis counter
46
    millisOld = millis();
47
48 }
```

```
49
50 void loop() {
51
    // Work with the imu sensor from Adafruit library
52
    // Go to BN0055's 'imu' and bring back a vector of 3 components into 'acc' ...
53
        (accelerometer) for the specific object IMU
      imu::Vector<3> acc = IMU.getVector(Adafruit_BN0055::VECTOR_ACCELEROMETER);
54
     // Go to BN0055's 'imu' and bring back a vector of 3 components into 'gyro ...
55
        (gyroscope) for the specific object IMU
     imu::Vector<3> gyro = IMU.getVector(Adafruit_BN0055::VECTOR_GYROSCOPE);
56
57
     // Parameters
58
     // Tilt approximation and normalize it to '1g' and convert into degrees
59
     theta_meas = - atan2(acc.x()/9.81,acc.z()/9.81)/(2*pi)*360;
60
      phi_meas = - atan2(acc.y()/9.81,acc.z()/9.81)/(2*pi)*360;
61
62
63
      theta_filt_new = theta_filt_old*0.9 + theta_meas*0.1;
     phi_filt_new = phi_filt_old*0.9 + phi_meas*0.1;
64
65
     // Change in time (seconds)
66
     Dt = (millis()-millisOld)/1000.;
67
     millisOld = millis();
68
69
70
     // Complementary filter
     theta = (theta + gyro.y()*Dt)*0.95 + theta_meas*0.05;
^{71}
      phi = (phi - gyro.x()*Dt)*0.95 + phi_meas*0.05; // Minus sign is for the ...
72
       particular orientation
73
74
      theta_gyro = theta_gyro + gyro.y()*Dt;
      phi_gyro = phi_gyro - gyro.x()*Dt; // Negative sign due to body axis
75
76
     // Print Acceleration data
77
     Serial.print(acc.x()/9.81);
78
     Serial.print(",");
79
80
      Serial.print(acc.y()/9.81);
      Serial.print(",");
81
      Serial.print(acc.z()/9.81);
82
     Serial.print(",");
83
     // Raw data
84
     Serial.print(theta_meas);
85
     Serial.print(",");
86
     Serial.print(phi_meas);
87
     Serial.print(",");
88
      // Filtered data
89
      Serial.print(theta_filt_new);
90
      Serial.print(",");
91
     Serial.print(phi_filt_new);
92
93
^{94}
     // Gryoscope data
     Serial.print(",");
95
     Serial.print(gyro.y()); // Pitch Rotational velocity
96
      Serial.print(",");
97
      Serial.print(gyro.x()); // Roll Rotational velocity
98
99
100
      Serial.print(",");
     Serial.print(theta_gyro); // Pitch
101
102
     Serial.print(",");
103
      Serial.print(phi_gyro); // Roll
```

```
104
105
      // Complementary filter data
     Serial.print(",");
106
     Serial.print(theta); // Pitch
107
     Serial.print(",");
108
     Serial.println(phi); // Roll
109
110
      // Next iteration
111
     theta_filt_old = theta_filt_new;
112
     phi_filt_old = phi_filt_new;
113
114
     // Wait the specified delay before requesting next data
115
116
    delay(BN0055_SAMPLERATE_DELAY_MS);
117
118 }
```

Code L.5 Arduino BNO055 IMU Complementary Filter. Source: Own.

### L.4 Quaternion

```
1 /* This code reads [QUATERNIONS] data from the BNO055 IMU Sensor
2 */
3
4 // Libraries
5 #include <Wire.h> // This library allows to communicate with I2C / TWI devices.
6 #include <Adafruit_Sensor.h> // Library with drivers that are based on the ...
       Adafruit Unified Sensor Driver
7 #include <Adafruit_BN0055.h> // This is a library for the BN0055 orientation sensor
   #include <utility/imumaths.h> // Inertial Measurement Unit Maths Library (it ...
       includes matrix.h, guaternions.h and vector.h)
   #include <math.h> // Math functions
9
10
11 // Global parameters and objects
  #define BN0055_SAMPLERATE_DELAY_MS (100) // Define how fast the sensor sample rate ...
12
       (sample every 100 ms)
13
14 Adafruit_BN0055 IMU = Adafruit_BN0055(); // Create IMU object and set what the ...
       object is
15
16
  void setup() {
17
    // Set the baud rate speed. This is how fast the data is to be sent through the \ldots
18
      USB connection
    Serial.begin(115200);
19
20
   // Start the IMU sensor
21
^{22}
    IMU.begin();
23
     // Wait 1000 ms to ensure the sensor starts correctly
^{24}
     delay(1000);
^{25}
26
27
     // Change the time clock on the chip to the time clock on board of the IMU \,
28
     IMU.setExtCrystalUse(true);
```

```
29
30
  }
^{31}
  void loop() {
32
33
     // Work with the imu sensor from Adafruit library
34
     // Go to BN0055's 'imu' and bring back a vector of 4 components into 'quat' \ldots
35
        (quaternion) for the specific object IMU
     imu::Quaternion quat=IMU.getQuat();
36
37
     // Print Quaternion data
38
     Serial.print(quat.w()); // Real part
39
     Serial.print(",");
40
     Serial.print(quat.x()); // 'i'
41
     Serial.print(",");
42
^{43}
     Serial.print(quat.y()); // 'j'
^{44}
     Serial.print(",");
     Serial.println(quat.z()); // 'k'
45
46
    // Wait the specified delay before requesting next data
47
    delay(BN0055_SAMPLERATE_DELAY_MS);
48
49
50 }
```

Code L.6 Arduino and BNO055 IMU Quaternion code. Source: Own.

### L.5 MPU9250 IMU

To setup the IMU for  $I^2C$  communication protocol, first, it is needed to setup the address for the communication:

```
1 /*
    * This code configures the MPU9250 IMU for I2C communication with the STM32 Bluepill
2
3
    */
4
   #include <Wire.h>
5
6
   const int MPU = 0x68;
                                              // I2C address of the MPU9250
7
   int16_t AcX, AcY, AcZ, Tmp, GyX, GyY, GyZ; // MPU values
8
9
  void setup()
10
  {
11
                                  // Initiate the library and join the I2C bus
     Wire.begin();
12
    Wire.beginTransmission(MPU); // Begin a transmission with the given address
13
     Wire.write(0x6B);
                                  // Send value in bytes
14
     Wire.write(0);
15
     Wire.endTransmission(true); // End the transmission
16
17 }
18
19 void loop()
20 {
   Wire.beginTransmission(MPU); // Begin a transmission with the given address
21
```

```
Wire.write(0x3B);
22
^{23}
    Wire.endTransmission(false);
     Wire.requestFrom(MPU, 12, true);
^{24}
^{25}
   AcX = Wire.read() << 8 | Wire.read();</pre>
26
    AcY = Wire.read() << 8 | Wire.read();</pre>
27
     AcZ = Wire.read() << 8 | Wire.read();</pre>
28
     GyX = Wire.read() << 8 | Wire.read();</pre>
29
     GyY = Wire.read() << 8 | Wire.read();</pre>
30
     GyZ = Wire.read() << 8 | Wire.read();</pre>
31
^{32}
     Serial.print("IMU values: ");
33
   Serial.print(AcX);
^{34}
    Serial.print(",");
35
36
37
     // Repeat for each value
38
     Serial.print(GyY);
39
     Serial.print(",");
40
     Serial.print(GyZ);
41
42 }
```

Code L.7 STM32 I<sup>2</sup>C configuration of the MPU9250. Source: Own.

```
1 /*
   \star This code configures the MPU9250 IMU for SPI communication with the STM32 ...
2
       Bluepill and prints data.
   *
3
   * MPU9250 library by Brian Chen used (https://github.com/brianc118/MPU9250)
4
   */
5
 6
7
   #include "Wire.h"
s #include <SPI.h>
9 #include <MPU9250.h>
10
11 #define SPI_CLOCK 1000000 // 1MHz clock works.
12 #define CS1 PA4 // STM32 NCS pin
13 #define WAITFORINPUT() { // Function to read serial imputs
14 while (!Serial.available())
15 {
17 while (Serial.available())
18 {
   Serial.read();
19
20 };
21 }
22
23 MPU9250 mpu(SPI_CLOCK, CS1); // MPU object
^{24}
25 void setup()
26 {
    pinMode(CS1, OUTPUT); // Pin definition
27
   SPI.begin();
                         // Initiate the library and join the SPI bus
^{28}
^{29}
   Serial.println("Press any key to continue");
30
```

```
WAITFORINPUT();
31
32
   mpu.init(true); // Initiate the IMU
33
^{34}
   uint8_t wai = mpu.whoami(); // Connection status with mpu
35
    if (wai == 0x71)
36
     {
37
     Serial.println("Successful connection");
38
39
     }
     else
40
41
   {
      Serial.print("Failed connection: ");
42
      Serial.println(wai, HEX);
^{43}
    }
44
45
     uint8_t wai_AK8963 = mpu.AK8963_whoami(); // Connection status with mag
46
47
     if (wai_AK8963 == 0x48)
    {
^{48}
     Serial.println("Successful connection to mag");
49
50
   }
    else
51
52
   {
     Serial.print("Failed connection to mag: ");
53
     Serial.println(wai_AK8963, HEX);
54
55
    }
56
    mpu.calib_acc(); // Calibrations
57
    mpu.calib_mag();
58
59
   Serial.println("Send any char to begin main loop.");
60
   WAITFORINPUT();
61
62 }
63
64 void loop()
65 {
66
    mpu.read_all();
   // Data is stored in variables from the library
67
   Serial.print(mpu.gyro_data[0]);
68
69
   Serial.print('\t');
   Serial.print(mpu.gyro_data[1]);
70
   Serial.print('\t');
71
   Serial.print(mpu.gyro_data[2]);
72
   Serial.print('\t');
73
     Serial.print(mpu.accel_data[0]);
74
     Serial.print('\t');
75
     Serial.print(mpu.accel_data[1]);
76
     Serial.print('\t');
77
   Serial.print(mpu.accel_data[2]);
78
79
   Serial.print('\t');
   Serial.print(mpu.mag_data[0]);
80
   Serial.print('\t');
81
   Serial.print(mpu.mag_data[1]);
82
    Serial.print('\t');
83
     Serial.print(mpu.mag_data[2]);
84
85
     Serial.print('\t');
     Serial.print(mpu.temperature);
86
87 }
```

Code L.8 STM32 SPI configuration of the MPU9250. Source: Own.

```
/*
1
     * SCMD library by SparkFun used ...
^{2}
        (https://github.com/sparkfun/SparkFun_Serial_Controlled_Motor_Driver_Arduino_Libraty)
     * Rest of libraries from Arduino
3
     * STM32 package from http://dan.drown.org/stm32duino/package_STM32duino_index.json
4
5
     * ----- TEST SMT32 1 MOTOR + IMU, SAME I2C PORT (MOTOR IN I2C, IMU IN I2C2)
6
7
     */
8
9
10 #include <Arduino.h>
11 #include <stdint.h>
12 #include "SCMD.h"
13 #include "SCMD_config.h" //Contains #defines for common SCMD register names and values
14 #include<Wire.h>
15
16
   #define LEDPIN PC13 //STM32 LED pin
17
  const int MPU=0x68; //I2C adress of the MPU9250
18
19 int16_t AcX, AcY, AcZ, Tmp, GyX, GyY, GyZ; //MPU values
^{20}
  SCMD DriverOne; //Driver object
21
22
  void setup() {
23
^{24}
    Wire.begin(); //Initiate the library and join the I2C bus
25
     Serial.begin(9600); //Data rate in bytes per second //Serial Monitor initiation
26
    pinMode(LEDPIN, OUTPUT); //LED definition
27
    Serial.println("Start");
28
29
    //IMU
30
   Wire.beginTransmission(MPU); //Begin a transmission with the given adress
^{31}
    Wire.write(0x6B); //Send value in bytes
32
    Wire.write(0);
33
34
    Wire.endTransmission(true); //End the transmission
35
     //MOTOR
36
     DriverOne.settings.commInterface = I2C_MODE; //Driver mode definition
37
     DriverOne.settings.I2CAddress = 0x5D; //Driver adress definition (0x5D by default)
38
39
     while(DriverOne.begin() != 0xA9) { //Wait for idle
40
       Serial.println("ID Mismatch");
41
       delay(200);
42
^{43}
     }
     Serial.println("ID Match");
^{44}
45
     Serial.println("Waiting for enumeration"); //Wait for peripherals (enumeration)
46
     while(DriverOne.ready() == false);
47
48
     Serial.println("Ready");
49
    while(DriverOne.busy()); //Enables the driver
50
     DriverOne.enable();
51
```

```
52 }
53
54 void loop() {
55
    int vel=255;
56
     int steps=20;
57
 58
      for(int i= 0; i<=vel; i++) {</pre>
59
60
        DriverOne.setDrive(0,0,i);
61
        Wire.beginTransmission(MPU); //Begin a transmission with the given adress
^{62}
        Wire.write(0x3B);
63
        Wire.endTransmission(false);
64
        Wire.requestFrom(MPU, 12, true);
65
66
        AcX=Wire.read()<<8|Wire.read();</pre>
67
68
         AcY=Wire.read() <<8 |Wire.read();
        AcZ=Wire.read() << 8 | Wire.read();</pre>
69
        GyX=Wire.read() <<8 |Wire.read();</pre>
70
        GyY=Wire.read() <<8 |Wire.read();</pre>
71
        GyZ=Wire.read()<<8|Wire.read();</pre>
72
73
        Serial.print("IMU values: ");
74
        Serial.print(AcX);
75
        Serial.print(",");
 76
        Serial.print(AcY);
 77
        Serial.print(",");
78
        Serial.print(AcZ);
79
80
        Serial.print(",");
        Serial.print(GyX);
81
        Serial.print(",");
82
        Serial.print(GyY);
83
        Serial.print(",");
84
        Serial.print(GyZ);
 85
 86
        Serial.print(", Speed: ");
        Serial.println(i);
87
        delay(steps);
88
89
      }
90
      delay(5000);
91
      for(int i= vel; i>=0; i--) {
92
93
        DriverOne.setDrive(0,0,i);
94
        Wire.beginTransmission(MPU); //Begin a transmission with the given adress
 95
         Wire.write(0x3B);
96
        Wire.endTransmission(false);
97
        Wire.requestFrom(MPU,12,true);
98
99
100
        AcX=Wire.read()<<8|Wire.read();</pre>
101
        AcY=Wire.read() <<8 | Wire.read();</pre>
        AcZ=Wire.read() <<8 |Wire.read();</pre>
102
        GyX=Wire.read()<<8|Wire.read();</pre>
103
104
        GyY=Wire.read() <<8 |Wire.read();
        GyZ=Wire.read() <<8 | Wire.read();</pre>
105
106
        Serial.print("IMU values: ");
107
        Serial.print(AcX);
108
109
        Serial.print(",");
```

1	10	Serial.print(AcY);
1	11	<pre>Serial.print(",");</pre>
1	12	<pre>Serial.print(AcZ);</pre>
1	13	<pre>Serial.print(",");</pre>
1	14	<pre>Serial.print(GyX);</pre>
1	15	<pre>Serial.print(",");</pre>
1	16	<pre>Serial.print(GyY);</pre>
1	17	<pre>Serial.print(",");</pre>
1	18	<pre>Serial.print(GyZ);</pre>
1	19	<pre>Serial.print(", Speed: ");</pre>
1	20	<pre>Serial.println(i);</pre>
1	21	<pre>delay(steps);</pre>
1	22	}
1	23	delay(5000);
1	<b>24</b>	
1	25	}

Code L.9 Motor driver test with IMU using I<sup>2</sup>C. Source: Own.

```
1 /*
    \star SCMD library by SparkFun used ...
2
       (https://github.com/sparkfun/SparkFun_Serial_Controlled_Motor_Driver_Arduino_Library)
     * MPU9250 library by Brian Chen used (https://github.com/brianc118/MPU9250)
3
     * Rest of libraries from Arduino
 4
     * STM32 package from http://dan.drown.org/stm32duino/package_STM32duino_index.json
5
 6
     * ----- TEST IMU IN SPI + MOTOR IN I2C + millis
 7
 8
     * Definitions stablished made to adapt to STM32 ports instead of Arduino ...
9
       (Library is designed for Arduino)
10
     \star 1MHz of clock because SPI conection with IMU is max 1MHz (by datasheet)
11
     */
12
13 #include <Arduino.h>
14 #include <stdint.h>
15 #include "SCMD.h"
16 #include "SCMD_config.h" //Contains #defines for common SCMD register names and values
17 #include "Wire.h"
18 #include <SPI.h>
19 #include <MPU9250.h>
20
21 #define LEDPIN PC13 //STM32 LED pin
22 #define SPI_CLOCK 1000000 // 1MHz clock works.
23 #define CS1 PA4 //STM32 NCS pin
24 //MOSI on PA7, MISO on PA6 and CLK on PA5 by default on STM32
25
26 #define WAITFORINPUT() {
    while(!Serial.available()){};
27
    while(Serial.available()){
^{28}
      Serial.read();
^{29}
30
    };
31 }
^{32}
33 MPU9250 mpu(SPI_CLOCK, CS1); //MPU object
34 SCMD DriverOne; //Driver object
```

```
35
36 unsigned long prevMillis=0; //Keeps track of previous MIllis() value
37 unsigned long currMillis=0; //Keeps track of current MIllis() value
38
39 int counter=0; //Counter for speed
40 bool laststate=true; //Current/Latest state of motor (false=decceleration, ...
       true=acceleration)
41 bool waiting=false; //Check if the motor is at a const speed
42
43 int maxspeed=255; //Max velocity to reach
44 int delayramp=100; //Delay for increase in speed, in milisec
45 int delaywait=5000; //Delay on waiting
46
47 void setup() {
^{48}
    counter=0; //Reset of counter
49
50
    pinMode(LEDPIN, OUTPUT); //LED definition
    pinMode(CS1, OUTPUT); //Pin definitions
51
    digitalWrite(LEDPIN, HIGH);
52
   Serial.begin(9600); //Serial Monitor initiation
53
   Serial.println("Start");
54
    SPI.begin();
55
56
      Serial.println("Press any key to continue");
57
    WAITFORINPUT();
58
59
    //IMU
60
    mpu.init(true);
61
62
      uint8_t wai = mpu.whoami(); //Connection with mpu
63
    if (wai == 0x71){
64
      Serial.println("Successful connection");
65
    }
66
67
     else{
68
       Serial.print("Failed connection: ");
       Serial.println(wai, HEX);
69
    }
70
71
   uint8_t wai_AK8963 = mpu.AK8963_whoami(); //Connection with mag
72
   if (wai_AK8963 == 0x48){
73
       Serial.println("Successful connection to mag");
74
    }
75
    else{
76
      Serial.print("Failed connection to mag: ");
77
       Serial.println(wai_AK8963, HEX);
78
79
     }
80
    mpu.calib_acc();
^{81}
82
    mpu.calib_mag();
83
     //MOTOR
84
     DriverOne.settings.commInterface = I2C_MODE; //Driver mode definition
85
     DriverOne.settings.I2CAddress = 0x5D; //Driver adress definition (0x5D by default)
86
87
    while(DriverOne.begin() != 0xA9){ //Wait for idle
88
      Serial.println("ID Mismatch");
89
       delay(200);
90
^{91}
     }
```

```
92
      Serial.println("ID Match");
93
     Serial.println("Waiting for enumeration"); //Wait for peripherals (enumeration)
94
     while(DriverOne.ready() == false);
95
     Serial.println("Ready");
96
97
     while(DriverOne.busy()); //Enables the driver
98
      DriverOne.enable();
99
100
        Serial.println("Send any char to begin main loop.");
101
     WAITFORINPUT();
102
103
   }
104
105 void loop() {
106
107 //Rollover is done in the ifs, with currMilils-prevMIllis. if it reaches rollover ...
        the difference remains the sameas it is an unsigned value
     currMillis=millis();
108
109
     if (unsigned int (currMillis-prevMillis)>=delayramp && waiting==false) { //State ...
110
        of accel or decel. Accounts for millis() override
      prevMillis=currMillis;
111
112
       if(counter<=maxspeed && laststate==true){ //Acceleration</pre>
          counter=counter+1;
113
114
          if(counter>maxspeed) { //Max speed reach
115
            waiting=true;
            counter=maxspeed;
116
117
          }
        }else if(counter>=0 && laststate==false) { //Deceleration
118
119
         counter=counter-1;
          if(counter<0){ //0 speed reach</pre>
120
121
           waiting=true;
            counter=0;
122
123
          }
124
        }
125
      }
126
     if(unsigned int (currMillis-prevMillis)>=delaywait && waiting==true){ //State of ...
127
        waiting. Accounts for millis() override
128
       prevMillis=currMillis;
129
       waiting=false;
       if(laststate) { //Start of decel
130
          laststate=false;
131
        }else{
132
          laststate=true; //Start of accel
133
134
        }
135
      }
136
137
      DriverOne.setDrive(0,0,counter);
138
     mpu.read_all();
139
      Serial.print(mpu.gyro_data[0]); Serial.print('\t');
140
141
      Serial.print(mpu.gyro_data[1]);
                                          Serial.print('\t');
      Serial.print(mpu.gyro_data[2]);
                                          Serial.print('\t');
142
143
      Serial.print(mpu.accel_data[0]); Serial.print('\t');
      Serial.print(mpu.accel_data[1]); Serial.print('\t');
144
      Serial.print(mpu.accel_data[2]); Serial.print('\t');
145
146
      Serial.print(mpu.mag_data[0]);
                                          Serial.print('\t');
```

```
147
     Serial.print(mpu.mag_data[1]); Serial.print('\t');
    Serial.print(mpu.mag_data[2]);
                                        Serial.print('\t');
^{148}
149
    Serial.println(mpu.temperature);
150
   Serial.print("Millis: ");
151
    Serial.print(currMillis);
152
    Serial.print(", State: ");
153
154
     if(waiting==false){
       if(laststate==true){
155
          Serial.print("ACCELERATION");
156
       }else{
157
158
         Serial.print("DECELERATION");
159
        }
     }else{
160
      Serial.print("WAITING");
161
162
     }
     Serial.print(", Speed: ");
163
164
      Serial.println(counter);
     delay (10);
165
166 }
```

Code L.10 Motor driver test with IMU using SPI. Source: Own.

# Appendix M

# Complementary filter design

# M.1 Pitch $\theta$ , Roll $\phi$ and Yaw $\psi$ determination from 3 - axis accelerometer and gyroscope

This section aims to approximate roll, pitch and yaw angles with an approximation.

Let  $\theta$  be the pitch angle and  $\phi$  be the roll angle.



Figure M.1 Acceleration vector decomposition. Source: Own.

The following trigonometric relations can be extracted from the figure above:

$$\tan \theta = \frac{a_x}{a_z} \tag{M.1.1}$$

From the latter expression, the angle  $\theta$  can be found as:

$$\theta = \arctan \frac{a_x}{a_z} \tag{M.1.2}$$

The same methodology can be used for the roll movement:

$$\phi = \arctan \frac{a_y}{a_z} \tag{M.1.3}$$

NOTE: This is a first approximation, thus, the major limitation found is that this only works for tilt angles from 0 - 89°. For angles above this range, the data returned is not relevant. The problem is caused by the trigonometric function tan(), when it reaches 90° the function has a vertical asymptote at that angle (see Figure M.2).



Figure M.2 Tangent function. Source: Own.

See L.2 for the code to get the pitch and roll from the IMU sensor.

Notice the code has a negative sign before the atan operation. This is because the IMU sensor's  $\hat{x}$  axis is in the opposite direction of the  $\hat{x}$  direction that the student has set initially.



Own.

of the IMU sensor. Source: Own.

Figure M.3 Pitch and roll of the IMU Sensor. Source: Own.

The orange data is pitch and purple data is roll and both pitch and roll are successfully measured. Nevertheless, one issue that affects the accelerometer is that it is highly susceptible to vibrations and noise. For instance, shaking the sensor will cause the accelerometer misread a pitch, thus, causing an invalid data. The accelerometer is highly sensitive to all kind of vibrations and the system inappropriately interprets vibration as tilt. In other words, even if the IMU does not pitch but suffer vibrations or brusque motion, the output signal will measure a tilt when in fact it is not tilting.

To approach this problem a low pass filter must be added. A low pass filter enables to get rid f the higher frequency movements that the accelerometer senses and takes into account low frequency movements.

The accelerometer measures a pitch  $\theta_{meas}$  angle and a roll angle  $\phi_{meas}$  (measured values) and let  $\theta_{filt}$ and  $\phi_{filt}$  be the filtered values of pitch and roll angles, respectively.

The following low pass filter is added:

$$\theta_{filt,new} = [\theta_{filt,old}] \cdot P_1 + [\theta_{meas}] \cdot P_2 \tag{M.1.4}$$

where

$$P_1 + P_2 = 1 \tag{M.1.5}$$

The above expressions can be interpreted as follows, given a measured angle, the newest pitch angle (most recent angle) is a sum of the prior iteration pitch angle times a parameter  $P_1$  and the actual measured pitch angle times a parameter  $P_2$ . Notice that setting a  $P_1 \approx 1$  value and  $P_2 \approx 0$  shall be understood as the prior iteration angle is most susceptible to remain the actual movement of the IMU (for the test, a  $P_1 = 0.9$  and  $P_2 = 0.1$  was used). In the end, the most recent pitch angle is set to the old pitch angle for the subsequent iteration.

After adding the filter, Figure M.4 show the unfiltered and the filtered data of the pitch angle. The result now provides evidence to the fact that the accelerometer behaves poorly when it is subjected to high frequency movements. It is important to notice that the filtered value is slower than the unfiltered data. They both get to the same value but at a different pace. Thereby, in order to find the optimal parameters, there must be a balance between accurate data and how fast the IMU accelerometer sensor detects.



Figure M.4 Filtered positive pitch of the IMU sensor. Source: Own.

Nonetheless, the major advantage is that the filtered data is no susceptible to vibration motions. Overall, it tends to have more stability and smoother outputs.

#### Measuring rotational velocity with gyroscope

Now, due to the limitations given by the accelerometer. The next step is to use the built-in gyroscope to measure the angular velocity. By means of the uniform angular equation of motion:

$$\theta_{gyro,new} = \theta_{gyro,old} + \omega_y \Delta t \tag{M.1.6}$$

$$\phi_{gyro,new} = \theta_{gyro,old} + \omega_x \Delta t \tag{M.1.7}$$

where the angular position  $\theta_{gyro,new}$  is the prior step position plus a change in angular velocity times the variation of time t. Since pitch is a rotation around the y axis, the angular velocity for pitch is  $\omega_y$  and  $\omega_x$  for roll.



Figure M.5 Pitch and roll from the BNO055 Gyroscope sensor. Source: Own.

Figure M.5 shows the same values of pitch and tilt extracted from the gyroscope sensor. This plot shows first a positive pitch and then a negative roll between  $x \in 850 - 875$ . Additionally, from between  $x \in 930 - 960$ , some vibration and high frequency motion was induced to the sensor. Results suggest that kind of high frequency motion does not affect the gyroscope's sensibility. However, notice how the gyroscope shows a drift in the value after keeping it at rest in a flat surface as the original situation  $(x \in 960 - 1000)$ . Thus, it is not possible to only use a gyroscope to get the tilt and roll data.

On balance the following conclusions are extracted:

• The accelerometer is useful for low frequency motion but is highly sensitive to noise and susceptible to vibration.

- The gyroscope is not sensitive to noise but its values drifts over time.
- Gyroscopes short-term data is accurate and accelerometers long-term data is unerring.

To address the aforementioned problem, the upcoming step resides in finding a solution by getting the best from the two worlds: accelerometer and gyroscope. The subsequent section proposes the use of a complementary filter combining both sensors data to acquire an accurate stable tilt and roll.

#### Design of a complementary filter

At the moment so far, 4 different parameters are extracted from the BNO055 sensor. The pitch and roll angles from the accelerometer sensor ( $\theta_{acc}$  and  $\phi_{acc}$ ) and the pitch and roll angles from the gyroscope sensor  $\theta_{gyro}$  and  $\phi_{gyro}$ .

$$\begin{array}{c} \theta_{acc} \\ \phi_{acc} \end{array} \right\} \quad \text{Trust long-term data} \qquad (M.1.8)$$

In order to design a complementary filter, all four parameters are kept and the main idea is to combine this four outputs which will allow the sensor to behave accurately in both sort and long term time period [100].

The complementary filter using all four parameters is calculated as follows:

$$\theta = \underbrace{\left[\theta + \omega_y \Delta t\right]}_{\theta_{avera}} \cdot P_A + \left[\theta_{acc}\right] \cdot P_B \tag{M.1.10}$$

$$\phi = \underbrace{[\phi + \omega_x \Delta t]}_{\phi_{auro}} \cdot P_A + [\phi_{acc}] \cdot P_B \tag{M.1.11}$$

The aforementioned expression can be interpreted as follows. On the short term, the sensor will use the data from the gyroscope, because it is very precise and not susceptible to external forces. On the long term, the sensor uses the data from the accelerometer, as it does not drift. Parameters  $P_A \approx 1$  and  $P_B \approx 0$  in order to filter out high frequencies data for the accelerometer and filter out low frequencies for the gyroscope. It must be noticed that  $\theta_{acc}$  and  $\phi_{acc}$  are NOT the filtered values but the raw data.

The complementary filter has to be used in a infinite loop. For every iteration, the pitch and roll angles must be updated with the new data of the gyroscope values by means of integration over time. At the same time, the filter checks if the magnitude of the force seen by the accelerometer has a reasonable value that could be the real g-force vector. If the value is too small or too big, it is known for sure that it is a disturbance that shall not be taken into account. Afterwards, pitch and roll angles will be updated with the accelerometer data by taking  $P_A = 95\%$  of the gyroscopes angle and adding  $P_B = 5\%$  of the angle calculated by the accelerometer. This will ensure that the measurement will not drift, but that it will be very accurate on the short term [100].

Code L.3 implements the complementary filter explained beforehand.



Figure M.6 Pitch and roll from the BNO055 Gyroscope sensor with a Complementary Filter. Source: Own.

Figure M.6 clearly shows the responsiveness of responsiveness of the gyroscope data at short-term motion and movements, keeping out the possible noise due to the accelerometer. Besides, stability is reached at long-term as the sensor trusts the accelerometers data rather than the gyroscopes values. Nonetheless, there is a modest overshoot when the IMU is brought back to the original position but after that it immediately compensates out.

#### Adding Yaw $\psi$ using the magnetometer

So far both pitch  $\theta$  and roll  $\phi$  were determined thanks to the Earth gravitational force vector as this vector can be decomposed in two components at the IMU's body axis for any roll or pitch movement. However, the gravitational force remains the same for any yaw motion, thus, by means of the magnetic North Pole location, yaw angle  $\psi$  can be extracted.



(a) Magnetic North vector aligned with the Magnetometer's x axis. Source: Own.



Figure M.7 Magnetometer's reference frame. Source: Own.

Keeping in mind the Earth magnetic North's vector, its components can easily be expressed in term of the magnetometer's reference frame (see Figure M.7). Thus,

$$\tan\psi = \frac{y_m}{x_m} \tag{M.1.12}$$

$$\psi = \arctan \frac{y_m}{x_m} \tag{M.1.13}$$

Nevertheless, the above expressions only works in a 2D motion with the IMU's magnetometer in a flat surface. If the sensor has a positive pitch angle  $\theta > 0$  and then it yaws, the above expression does not take into account three dimensional motion. Likewise, projecting the vector into a three dimensional reference frame this problem can plainly be extrapolated to a 3D reference frame.

Here-under are the resultant data:



Figure M.8 Magnetometer's yaw plot. Source: Own.

Let's have a closer look at the plot (see Figure M.8). When the IMU's x axis is aligned with the Earth's magnetic North Pole vector, yaw angles is 0°. Furthermore, as the IMU rotates the resultant yaw angle decreases up to a point where it is pointing towards the South then, a whole rotation is made. This is the reason why there is a jump from -180° and 180°. In a unit circle, they are the same value. Notwithstanding, if the IMU rolls, the sensor interprets the roll as a change in heading (yaw) as the sensor is no longer aligned with the x - y plane.

## Appendix N

# Python visualization code

This section shows how to extract data from Arduino and pass it into a Python script for further processing. The main requirement is to have python installed as well as serial library:

1 pip install serial

Code N.1 Libraries to be installed. Source: Own

Once done that, the program can extract the data using the following code (this code works with the IMU\_data\_acquisition\_orientation\_all\_axis.ino arduino file).

### N.1 Import Arduino data code

```
# This code imports ARDUINO data into python
 1
 2
 3
   # Libraries
   import time
 4
 \mathbf{5}
   import serial
 6
 \overline{7}
  # Code
   arduinoData = serial.Serial('com5', 115200) # COM and baud rate
 8
 9
  time.sleep(1) # Delay of 1 s to ensure the serial port working
10
11
   # Loop forever
^{12}
   while True:
^{13}
        # Loop until there is data into serial port (hang on until there is data)
14
       while (arduinoData.inWaiting() == 0):
15
16
            pass
17
^{18}
        # Since there is data in serial port, read the data
       data = arduinoData.readline()
19
       data = str(data, 'utf-8') # Format the data
20
```

```
data = data.split(',') # Split the data into a list
21
^{22}
        # print(data)
^{23}
^{24}
       # Extract values from the accelerometer
25
       acc_x = float(data[0])
26
       acc_y = float(data[1])
27
       acc_z = float(data[2])
28
29
       # Extract filtered values (Complementary filter) of pitch and roll
30
^{31}
       pitch = float(data[11])
       roll = float(data[12])
32
33
       # Extract yaw
34
       yaw = float(data[13])
35
36
37
       print(
           f"acc_x={acc_x}, acc_y={acc_y}, acc_z={acc_z}, pitch={pitch}, roll={roll}, ...
38
      yaw={yaw}")
```

Code N.2 Python code to extract Arduino data. Source: Own.

### N.2 vPython 3D representation code

```
1 # The following code rotates a reference frame
2
3 # Libraries
4 from vpython import * # vPython library
5 import numpy as np # Numpy library
  import time # Time library
6
  import math # Math library
7
8
9 # Libraries
10 import serial
11
12 # Code
13 arduinoData = serial.Serial('com3', 115200) # COM and baud rate
14
15 time.sleep(1) # Delay of 1 s to ensure the serial port working
16
17 # View orientation
18 scene.forward = vector (-1, -1, -1)
19
20 scene.width = 600
21 scene.height = 600
22
23
24 # Adding the Arduino
25 breadBoard = box(length=6, width=6, height=6,
                    color=color.white, opacity=0.25)
^{26}
27
28 # Compound object
29 object = compound([breadBoard])
```

```
30
  x_arrow = arrow(length=4, shaftwidth=0.25,
^{31}
                   color=color.red, axis=vector(1, 0, 0))
32
33 y_arrow = arrow(length=4, shaftwidth=0.25,
                   color=color.green, axis=vector(0, 1, 0))
34
  z_arrow = arrow(length=4, shaftwidth=0.25,
35
                    color=color.cyan, axis=vector(0, 0, 1))
36
37
  x_body_arrow = arrow(length=6, shaftwidth=0.25,
38
                         color=color.purple, axis=vector(1, 0, 0))
39
  y_body_arrow = arrow(length=6, shaftwidth=0.25,
^{40}
                         color=color.magenta, axis=vector(0, 1, 0))
41
  z_body_arrow = arrow(length=6, shaftwidth=0.25,
^{42}
                         color=color.white, axis=vector(0, 0, 1))
43
44
^{45}
46
   # Loop forever
  while True:
47
       # Loop until there is data into serial port (hang on until there is data)
48
       while (arduinoData.inWaiting() == 0):
49
50
           pass
51
       # Since there is data in serial port, read the data
52
       data = arduinoData.readline()
53
       data = str(data, 'utf-8') # Format the data
54
       data = data.split(',') # Split the data into a list
55
56
       # Ouaternions
57
58
       q0 = float(data[0])
       q1 = float(data[1])
59
       q2 = float(data[2])
60
       q3 = float(data[3])
61
62
63
       # Convert quaternions to euler angles
64
       roll = -atan2(2*(q0*q1+q2*q3), 1-2*(q1**2+q2**2))
       pitch = asin(2*q0*q2-q3*q1)
65
       yaw = -atan2(2*(q0*q3+q1*q2), 1-2*(q2**2+q3**2))
66
67
       rate(50) # 50fps
68
       k = vector(cos(yaw)*cos(pitch),
69
70
                   sin(pitch), sin(yaw)*cos(pitch))
71
       y = vector(0, 1, 0)
72
       s = cross(k, y) # Side vector
73
       v = cross(s, k) # Up vector
74
75
76
       vrotate = v*cos(roll) + cross(k, v)*sin(roll)
77
78
       x\_body\_arrow.axis = k
79
       z_body_arrow.axis = cross(k, vrotate)
       y_body_arrow.axis = vrotate
80
       object.axis = k
81
       object.up = vrotate # Up direction
82
       x\_body\_arrow.length = 4
83
84
       z_body_arrow.length = 4
       y\_body\_arrow.length = 4
85
```

Code N.3 3D CubeSat visualization. Source: Own.

# Appendix O

# **Reaction Wheel Control code**

Here is presented the full control algorithm, the control algorithm flowchart is presented again in a bigger scale:















To download the code, check GitHub: Github

```
1 /* Reaction Wheel controller
2
3
   * The following code provides a control for PLATHON project's CubeSat.
4
5
   * The following libraries are used:
6
   * STM32 Bluepill microcontroller package from ...
7
      http://dan.drown.org/stm32duino/package_STM32duino_index.json
   * SCMD Driver library by SparkFun used (https:// ...
       github.com/sparkfun/SparkFun_Serial1_Controlled_Motor_Driver_Arduino_Library)
   * MPU9250 IMU library by Rafa Castalla used ...
9
       (https://github.com/rafacastalla/MPU9250-1)
    * read_IMU() and IMU setup content code extracted from Andres Gomez and Miquel ...
10
       Reurer, from the PLATHON group (magnetorquers section)
11
   */
12
13 // ...
      DEFINITIONS
14 // ...
       LIBRARIES
15 #include <Arduino.h> // Arduino library
16 #include <stdint.h> // Integer types library
17 #include <SCMD.h> // Serial Controlled Motor Driver library
18 #include <SCMD_config.h> // Serial Controlled Motor Driver Configuration library
19 #include <SPI.h> // SPI library
20 #include <MPU9250.h> // IMU library
21 #include <Wire.h>
22
23 // ...
       DEFINITIONS
                                     // Integrated LED of the Bluepill
24 #define LEDPIN PC13
25 #define CS1 4
                                     // STM32 NCS Chip select (SPI mode only) pin. ...
      MOSI on PA7, MISO on PA6 and CLK on PA5 by default on STM32
26 #define STM32_CLOCK 72000
                                    // Internal STM32 Clock (72MHz, in kHz so ...
       period is millisec)
                                     // Number PI
27 #define PI 3.14159265
28 #define Rad_to_deg 57.29577951
                                     // Convert radians to degrees
  #define Deg_to_rad 0.01745329
                                      // Convert degrees to radians
29
30 #define Final_pointing_tolerance 1 // Tolerance for final pointing (+-1 degree of ...
       tolerance) (Check if position is within the margin)
31 #define Pointing_mode_tolerance 5 // Tolerance for selecting pointing mode (+-5 ...
       degree of tolerance) (Select whether to use coarse or fine mode)
32 #define Accel_tolerance 0.5 // Tolerance for acceleration measurement ...
       (+-0.5 unit of acceleration of tolerance) (Acceptable error of acceleration)
33 #define Gyro_tolerance 1 // Tolerance for gyro measurement (+-1 unit of ...
       gyroscope Z tolerance) (Accounts to know if it is rotating or not at a ...
       constant speed for the ramp)
34
                        // Driver Object definition
35 SCMD DriverOne;
36 MPU9250 IMU(SPI, CS1); // MPU object definition
```

```
37
38 // ...
      GLOBAL VARIABLES
39 float IMU_accel_data_X, IMU_accel_data_Y, IMU_accel_data_Z; // Values of local IMU ...
      accelerometer (m/s^2)
40 float IMU_gyro_data_X, IMU_gyro_data_Y, IMU_gyro_data_Z;
                                                            // Values of local IMU ...
      gyroscope (degrees/s)
41 float IMU_mag_data_X, IMU_mag_data_Y, IMU_mag_data_Z;
                                                            // Values of local IMU ...
      magnetometer (uT)
42 float Pitch_deg, Roll_deg, Yaw_deg = 0;
                                                             // Pitch, Roll and Yaw ...
      (degrees)
43
44 // Declare variables outside IMU, since variables initiates to 0 without defining ...
      them in function
45 float Accel_total_vector_modulus; // Accelerometer total vector magnitude ...
      (m/s^2)
46 float Accel_pitch_deg, Accel_roll_deg; // Accelerometer pitch and roll angles ...
      (degrees)
47 float Gyro_pitch_deg, Gyro_roll_deg; // Gyroscope pitch and roll angles (degrees)
48 bool Gyro_Accel_sync = false;
                                         // Check if Gyroscope and Accelerometer ...
      are synchronized
49 float Mag_X_calc, Mag_Y_calc;
                                   // Calculated Magnetometer value (uT)
50 float Mag_X_corrected, Mag_Y_corrected; // Corrected value Magnetometer value ...
      (with damping) (uT)
51 float Pitch_rad, Roll_rad;
                                         // Pitch and Roll angles (rad)
52
                      // Value of Reaction Wheel speed [from -255 to 255]
53 int RW speed = 0;
54 int OBC_mode_value = 0; // Value of On Board Computer mode (waiting, positioning, ...
      detumbling, etc.)
55 int OBC_data_value = 0; // Value of On Board Computer data (desired angle turn, etc.)
56 int Stop_state = 0; // If Stop_state != 0, motor stops rotating
57
58 // PID control definitions
59 double PID_error, PID_last_error; // Initialize error and previousError
60 double PID_cumulative_error, PID_rate_error; // Initialize the cumulative Error ...
      (Integral) and the rate of Error (Derivative)
61 float Current_Yaw_deg = 0; // Addition of 180 deg to all values. ...
      The error is a substract, so the difference is the same
62
63 double kp = 1; // Proportional contribution
64 double ki = 0; // Integral contribution
65 double kd = 1; // Derivative contribution
66
67 float Deg_to_reach = 0; // Setpoint angle (degrees)
68 bool Zero_state = false; // Check if PID encompass yaw of 0 degrees (to turn the ...
      value to a workable zone)
69 int PID_output = 0; // Output value of the PID [from 0 to 255]
70
71 // ...
      FUNCTIONS
72 // ...
      GENERAL USE FUNCTIONS
73 // ...
      OBC Functions
```
```
74 void OBC_mode_receive()
75 {
76
    /*
      Function to set the Bluepill to receiving mode. Uses Timer CH3. Serial1 as ...
77
       UART communication
78
      INPUT: None, but gets mode from User Seriall input
79
      OUTPUT: None, but saves OBC_mode_value
80
     */
81
82
    // Check if UART Serial1 is available
83
    read_IMU();
84
85
    if (Serial1.available() > 0)
86
87
    {
       String bufferString = ""; // String for buffer of Serial1
88
89
       // Keep saving input prompt
       while (Serial1.available() > 0)
90
       {
91
        bufferString += (char)Serial1.read(); // Adds chars to the Serial1 buffer
92
       }
93
       // Conversion from String to int
94
       OBC_mode_value = bufferString.toInt();
95
       Serial1.print("Mode Number: ");
96
97
       Serial1.println(OBC_mode_value);
     }
98
99
   }
100
101 void OBC_data_receive()
102 {
103
    /*
      Function to set the Bluepill to receiving data mode. Uses Timer CH3. Seriall ...
104
       as UART communication
105
106
       INPUT: None, but gets data from User Seriall input
107
       OUTPUT: Save OBC_data_value
108
     */
    read_IMU();
109
    // Check if UART Seriall is available
110
111
   if (Serial1.available() > 0)
112
   {
      String bufferString = ""; // String for buffer of Serial1
113
114
      // Keep saving input prompt
115
       while (Serial1.available() > 0)
116
       {
         bufferString += (char) Serial1.read(); // adds chars to the Serial1 buffer
117
118
       }
119
       OBC_data_value = bufferString.toInt(); // Conversion from String to int
120
     }
121 }
122
123 void EmergencyStop()
124 {
125
    /*
126
      Function to STOP the motor from rotating.
127
      INPUT: None, but gets data from User Seriall input
128
       OUTPUT: Save OBC_data_value
129
```

```
*/
130
131
   // Check if UART Serial1 is available
132
   if (Serial1.available() > 0)
133
134
    {
      String bufferString = ""; // String for buffer of Serial1
135
       // Keep saving input prompt
136
       while (Serial1.available() > 0)
137
138
       {
        bufferString += (char)Serial1.read(); // adds chars to the Serial1 buffer
139
140
       }
      Stop_state = bufferString.toInt(); // Conversion from String to int
141
142
      Serial1.print("Emergency Stop activated"); // In this case turn, but can be ...
       what is needed
143
     }
144
145
     // Timer1.attachInterrupt(TIMER_CH4, read_show_IMU);
     // read_show_IMU();
146
     // mode_Select(OBC_mode_value);
147
148
   if (Stop_state != 0)
149
150
   {
151
     // If it stopped
      set_impulse(0, 0);
152
      Serial1.flush();
153
      OBC_data_value = 0;
154
      mode_OBC_Input_Wait();
155
156
    }
157 }
158
159 // ...
       Driver Functions
160 void set_impulse(bool RW_direction, int New_RW_speed)
161 {
162
      Function to set the motor at a fixed direction and speed (B = port 1, A = port 0)
163
164
165
      INPUT:
        RW_direction: Reaction Wheel direction [Clockwise or Counter Clockwise] in A
166
        New_RW_speed: New reaction wheel speed [from 0 to 255]
167
      OUTPUT:
168
        None, but saves speed of the Reaction Wheel
169
170
    */
171
     // Account for change in direction
172
     // If motor's z-axis direction is coincident with CubeSat's z-axis direction
173
174
     /**
175
     * RW_direction == true ---> CounterClockwise, motor direction = 0
176
      * RW_direction == false ---> Clockwise, motor direction = 1
177
      * */
178
179
     // If motor's z-axis direction is CONTRARY to CubeSat's z-axis direction
     /**
180
      * RW_direction == true ---> CounterClockwise, motor direction = 1
181
      * RW direction == false ---> Clockwise, motor direction = 0
182
183
      *
184
       * */
```

```
185
186
      if (RW_direction) // Reaction Wheel Counter Clock Wise
187
     {
       DriverOne.setDrive(1, 0, New_RW_speed); // Change direction depending on motor ...
188
        connection
       RW_speed = New_RW_speed;
189
190
      }
      else // Reaction Wheel Clock Wise
191
192
     {
        DriverOne.setDrive(1, 1, New_RW_speed);
193
194
        RW_speed = -New_RW_speed;
195
      }
196
    }
197
    void generate_ramp(bool RW_direction, int Acc_ramp_time_duration, int RW_ramp_speed_reach, bool Acc_D
198
199
    {
200
      /*
        Function to generate ramps. In this case it is a Single Impulse.
201
        As we dont know the time it lasts, we have to see if speed changes on the ...
202
        cubesat. That is the use of the while.
203
        INPUT:
204
         RW_direction: Reaction Wheel direction [Clockwise or Counter Clockwise]
205
         Acc_ramp_time_duration: Duration of ramp
206
207
          RW_ramp_speed_reach: Final speed reached
          Acc_Dec_state: Check wether in acceleration state ('1') or deceleration ...
208
        state ('0')
        OUTPUT:
209
210
          None
211
      */
212
      Serial1.println("Ramp Start");
213
     bool waiting = true;
214
215
216
      set_impulse(RW_direction, RW_ramp_speed_reach);
217
      Timer1.attachInterrupt(TIMER_CH4, read_show_IMU);
218
     waiting = true;
219
220
     while (waiting)
221
     { // Range of tolerance
222
        if (Acc Dec state)
223
        { // If accelerating
224
          if (abs(IMU_gyro_data_Z) > Gyro_tolerance)
225
          { // If gyro is not 0 or so, means it has constant speed of rotation
226
            if (abs(IMU_accel_data_X) < Accel_tolerance)</pre>
227
            { // Stop of acceleration
228
              waiting = false;
229
230
            }
^{231}
          }
        }
232
233
        else
        { // If decelerating
234
          if (abs(IMU_gyro_data_Z) < Gyro_tolerance)</pre>
235
236
          { // If gyro is not 0 or so, means it has constant speed of rotation
            if (abs(IMU_accel_data_X) < Accel_tolerance)</pre>
237
            { // Stop of acceleration
238
              waiting = false;
239
```

```
240
            }
^{241}
         }
       }
242
       delay(1); // Delay for ensuring getting inside the while loop
^{243}
    }
244
245
     waiting = true;
     Timer1.detachInterrupt(TIMER_CH4);
246
247 }
248
249 // ...
        IMU Functions
250 void read_IMU()
251 {
      /*
252
253
        Function to read IMU_Data (and refreshes the global variables for IMU data).
254
        INPUT:
255
       None
256
       OUTPUT:
257
       None, but saves IMU variables and Roll, Pitch and Yaw
258
259
      */
260
     // Begin reading IMU and pitch, yaw, roll
261
     IMU.readSensor();
262
263
      // Accelerations
264
      IMU_accel_data_X = IMU.getAccelX_mss();
265
266
     IMU_accel_data_Y = IMU.getAccelY_mss();
     IMU_accel_data_Z = IMU.getAccelZ_mss();
267
268
     // Gvroscope
     IMU_gyro_data_X = (IMU.getGyroX_rads() * Rad_to_deg);
269
     IMU_gyro_data_Y = (IMU.getGyroY_rads() * Rad_to_deg);
270
271
     IMU_gyro_data_Z = (IMU.getGyroZ_rads() * Rad_to_deg);
272
      //Magnetometer
      IMU_mag_data_X = IMU.getMagX_uT();
273
274
      IMU_mag_data_Y = IMU.getMagY_uT();
     IMU_mag_data_Z = IMU.getMagZ_uT();
275
276
     // GetTime of IRS, in case value is correct if timer definition changes.
277
     // The dT is the time step of each interrupt, thus, each read of the IMU
278
      // First division of clock by preescaler (frequency of timer), then inversion ...
279
        (period of timer), then multiplication by overflow (period of channel)
280
      . . .
       float dT = ((1 / ((float)STM32_CLOCK / (float)Timer1.getPrescaleFactor())) * Timer1.getOverflow()
        dT in seconds
281
      // Get the angle Pitch and Roll angle(w = w0 + dT*angular_vel)
282
283
      Gyro_pitch_deg += dT * (IMU.getGyroY_rads()) * Rad_to_deg;
284
      Gyro_roll_deg += dT * (IMU.getGyroX_rads()) * Rad_to_deg;
285
      // Gimbal lock compensation
286
      Gyro_pitch_deg = Gyro_pitch_deg + Gyro_roll_deg * sin(dT * (IMU.getGyroZ_rads()));
287
      Gyro_roll_deg = Gyro_roll_deg - Gyro_pitch_deg * sin(dT * (IMU.getGyroZ_rads()));
288
289
     // Modulus of the acceleration vector, calculate the total (3D) vector
290
291
      . . .
        Accel_total_vector_modulus = sqrt((IMU.getAccelX_mss() * IMU.getAccelX_mss()) + (IMU.getAccelY_ms
```

```
292
      . . .
       Accel_pitch_deg = asin((float)IMU.getAccelX_mss() / Accel_total_vector_modulus) * Rad_to_deg; //
        Calculate the pitch angle from accelerometer
293
      . . .
        Accel_roll_deg = asin((float)IMU.getAccelY_mss() / Accel_total_vector_modulus) * Rad_to_deg; //
        Calculate the roll angle from accelerometer
294
      /* Accelerometer calibration */
295
296
      // To calibrate the accelerometer, put the values \hfill 0\, . 0 \, in the variables and \ldots
297
       uncomment the 3 Serial1.prints
      // When compiling, leave the IMU immobile so that the accelerometer calibrates \ldots
298
        properly.
      // Once the values
                            are obtained, they are noted and it is recompiled as it ...
299
        had been before.
300
301
      Accel_pitch_deg -= -11.17;
      Accel_roll_deg -= -0.45;
302
303
          Serial1.print(Accel_pitch_deg,6);
304
          Serial1.print("\t");
305
          Serial1.println(Accel_roll_deg,6);
306
307
     // If gyroscope and accelerometer are synchronized
308
     if (Gyro_Accel_sync)
309
310
      {
        // ----- Gyro & accel have been synchronized
311
        Gyro_pitch_deg = Gyro_pitch_deg * 0.95 + Accel_pitch_deg * 0.05; // Correct ...
312
        the drift of the gyro pitch angle with the accelerometer pitch angle. ...
        Original values were 0.9996 and 0.0004m, respectively.
       Gyro_roll_deg = Gyro_roll_deg * 0.95 + Accel_roll_deg * 0.05; // Correct ...
313
        the drift of the gyro roll angle with the accelerometer roll angle
      }
314
315
      else
316
      { // The 0.98 and 0.02 values are taken from http:// \ldots
        www.pieter-jan.com/node/11else{
        // ----- Synchronize gyro & accel
317
       Gyro_pitch_deg = Accel_pitch_deg; // Set the gyro pitch angle equal to the ...
318
        accelerometer pitch angle
       Gyro_roll_deg = Accel_roll_deg; // Set the gyro roll angle equal to the ...
319
        accelerometer roll angle
       Gyro_Accel_sync = true;
                                         // Set the IMU started flag
320
321
      }
322
      // Final corrected Pitch and Roll angle (degrees)
323
      Pitch_deg = Pitch_deg * 0.9 + Gyro_pitch_deg * 0.1;
324
      Roll_deg = Roll_deg * 0.9 + Gyro_roll_deg * 0.1;
325
326
327
      // Final corrected Pitch and Roll angle (radians)
328
      Pitch_rad = -Roll_deg * Deg_to_rad;
      Roll_rad = Pitch_deg * Deg_to_rad;
329
330
      // Calculated values of Magnetometer along X and Y axis
331
332
        Maq_X_calc = IMU.getMaqX_uT() * cos(Pitch_rad) + IMU.getMaqY_uT() * sin(Roll_rad) * sin(Pitch_rad)
     Mag_Y_calc = IMU.getMagY_uT() * cos(Roll_rad) + IMU.getMagZ_uT() * sin(Roll_rad);
333
334
      // Correct drif (same as Pitch_deg)
335
```

```
336
      Mag_X_corrected = Mag_X_corrected * 0.9 + Mag_X_calc * 0.1;
337
      Mag_Y_corrected = Mag_Y_corrected * 0.9 + Mag_Y_calc * 0.1;
338
     // The magnetic north has been modified, originally, it was from X to Y (towards \dots
339
        which the reading 0 was 90 degrees offset) and without the negative sign of ...
        the beginning (towards which the reading would add the angles ...
        counterclockwise instead of clockwise).
340
      Yaw_deg = -atan2(Mag_Y_corrected, Mag_X_corrected) * Rad_to_deg;
341
342
^{343}
      // If you want to be precise, add the declination of the geographical place \dots
        where you are (https: // www.ign.es/web/gmt-declinacion-magnetica)
344
      // Yaw_deg += Declination;
345
      // The next two lines contain the value of Yaw\_deg between 0 and 360 degrees.
346
347
     if (Yaw_deg < 0)
348
       Yaw_deg += 360;
349
     if (Yaw_deg >= 360)
       Yaw_deg -= 360;
350
351 }
352
353 void show_IMU()
354 {
     /*
355
356
       Function to show IMU_Data on the Serial1 Monitor
357
       INPUT:
358
       None
359
       OUTPUT:
360
       None
361
      */
362
363
     // Print accelerometer values
364
      // Serial1.print("Ax: ");
365
366
      Serial1.print(IMU_accel_data_X, 4);
367
      Serial1.print(';');
      Serial1.print(IMU_accel_data_Y, 4);
368
     Serial1.print(';');
369
370
     // Serial1.print(IMU_accel_data_Z,4);
     // Serial1.println("");
371
372
     // Print Gyro values
373
      // Serial1.print(" / G: ");
374
      // Serial1.print(IMU_gyro_data_X,4); Serial1.print(';');
375
      // Serial1.print(IMU_gyro_data_Y,4); Serial1.print(';');
376
      Serial1.print(IMU_gyro_data_Z, 4);
377
      Serial1.print(';');
378
379
     // Serial1.println("");
380
381
     // Print Mag values
     // Serial1.print("MAGS: ");
382
      // Serial1.print(IMU_mag_data_X,4); Serial1.print(';');
383
          Serial1.print(IMU_mag_data_Y,4);
                                                Serial1.print(';');
384
          Serial1.print(IMU_mag_data_Z,4);
385
          Serial1.println("");
386
387
     // Print orientation angles
388
     // Serial1.print(" / D: ");
389
```

```
// Serial1.print(Pitch_deg, 4);
390
     // Serial1.print(';');
391
    // Serial1.print(Roll_deg, 4);
392
    // Serial1.print(';');
393
394
   Serial1.print(Yaw_deg, 4);
    Serial1.println("");
395
396 }
397
398 void read_show_IMU()
399 {
400
    /*
       Function to both read and show IMU_Data.
401
402
       INPUT:
403
      None
404
       OUTPUT:
405
       None, but saves IMU variables and Roll, Pitch and Yaw
406
407
     */
408
    read_IMU();
409
410
    show_IMU();
411 }
412
413 // ...
       PID Functions
414 void computePID()
415 {
    / *
416
417
      Function to calculate PID
418
      INPUT:
419
      None
420
       OUTPUT:
421
422
       None, but saves PID values
423
     */
424
    // Time step
425
426
    . . .
       float dT = ((1 / ((float)STM32_CLOCK / (float)Timer1.getPrescaleFactor())) * Timer1.getOverflow()
       in seconds
427
    read_show_IMU();
428
429
     Current_Yaw_deg = Yaw_deg;
430
431
     // Check if PID passes through 0 degrees
432
433
     if (Zero_state)
434
     {
435
       // Saves if the pid pass through 0, if it does it moves the zone away from 0
       // New variable to not modify the Yaw_deg value
436
       Current_Yaw_deg = Yaw_deg + 180; // Addition of 180 deg to all values. The ...
437
        error is a substract, so the difference is the same
438
      }
439
     if (Current_Yaw_deg >= 360)
440
      Current_Yaw_deg -= 360; // One of them will increase over 360, it is a correction
441
442
```

```
// Percentage
443
     volatile float Yaw_deg_perc = (Current_Yaw_deg - 360) / (-360) * 100;
444
445
   // Transform angle to percentage (Deg_to_reach == setPoint)
446
447
     volatile float Deg_to_reach_perc = (Deg_to_reach - 360) / (-360) * 100;
448
    // Errors
449
     PID_error = Deg_to_reach_perc - Yaw_deg_perc; // Calculate error ...
450
       (Proportional)
    PID_cumulative_error += PID_error * dT; // Calculate the cumulative ...
451
      error (Integral)
     PID_rate_error = (PID_error - PID_last_error) / dT; // Calculate the rate of ...
452
      error (Derivative)
453
    // PID Control
454
     float PID_P = kp * PID_error;
455
                                            // Proportional
456
     float PID_I = ki * PID_cumulative_error; // Integral
     float PID_D = kd * PID_rate_error; // Derivative
457
458
    /*
459
     // Limit
460
      // Ensure not overflow
461
      if (PID_P > 255)
462
       PID_P = 255;
463
      if (PID_P < -255)
464
465
       PID_P = -255;
       if (PID_I > 255)
466
       PID_I = 255;
467
      if (PID_I < -255)
468
       PID_{I} = -255;
469
      if (PID_D > 255)
470
        PID_D = 255;
471
      if (PID D < -255)
472
        PID_D = -255;
473
474
    */
475
     PID_output = PID_P + PID_I + PID_D; // PID control
476
477
478
   // Prevent overflow
479
   if (PID_output > 255)
      PID_output = 255;
480
    if (PID output < -255)
481
      PID_output = -255;
482
483
    bool RW_direction = true; // true=positive (CounterClockwise), false=negative ...
484
      (Clockwise)
485
486
    if (PID_output < 0)
487
    {
488
      PID_output = -PID_output;
      RW_direction = false;
489
    }
490
491
     set_impulse(RW_direction, PID_output);
492
493
     // Save current error and time for next iteration
494
     PID_last_error = PID_error; // Save current error
495
496 }
```

```
497
498 // ...
       MODES OF OPERATION FUNCTIONS
499 void mode_OBC_Input_Wait()
500 {
501
    /*
502
      0. Default mode, OBC Reading
503
      INPUT:
504
505
      None
      OUTPUT:
506
507
      None, but exits to mode_Select
    */
508
    if (Stop_state != 0)
509
510
    {
511
      Stop_state = 0;
512
      Timer1.detachInterrupt(TIMER_CH3);
513
    }
514
   OBC_mode_value = 0;
515 Timer1.attachInterrupt(TIMER_CH3, OBC_mode_receive);
   while (OBC_mode_value == 0)
516
517
    {
      delay(1); // If not used the while function does not work
518
519
      // it can be added more conditions to evade being blocked until a data is ...
       received.
       // for example, it could function an interrupt with a forced exit and an if ...
520
       after or something
521
    }
522 Timer1.detachInterrupt(TIMER_CH3);
   Serial1.println(OBC_mode_value);
523
524
   mode_Select(OBC_mode_value);
525 }
526
527 void mode_Select(int mode_value)
528 {
    /*
529
      Function to select Mode
530
531
      INPUT:
532
      mode_value ['0': Waiting; '1': Positioning; '2': Reading]
533
      OUTPUT:
534
      None, but exits to selected mode
535
     */
536
537
     switch (mode_value)
538
539
    {
540
     default: // Mode OBC Input Waiting (0): Waiting for OBC
541
     Serial1.println("Reading mode from OBC");
542
      mode_OBC_Input_Wait();
      break;
543
    case 1: // Mode Positioning RW only
544
545
      Serial1.println("Mode Positioning");
       mode_Positioning_RW();
546
547
       break;
    case 2: // Mode IMU reading (TEST mode)
548
      Serial1.println("Reading IMU");
549
550
      mode_IMU_reading();
```

```
551
       break;
    case 3: // Mode Motor ON OFF (TEST mode)
552
      Serial1.println("Mode motor ON OFF");
553
554
      mode_motor_on_off();
555
      break;
    }
556
557 }
558
559 void mode_Positioning_RW()
560 {
561
    /*
      1. Mode Positioning (RW only)
562
563
       INPUT:
564
      None
565
       OUTPUT:
566
567
       None, but exits to suitable positioning type
     */
568
569
    OBC_mode_value = 0;
570
571 Serial1.println("Positioning begin");
572 // Insert code to select if the positioning will be coarse (impulses) or fine (PD)
573
    /* It should be something like:
   * read IMU degree value
574
575
     \star get OBC value from OBC and save in global variable
     * comparison to OBC value (wanted position), add to the void function definition
576
     * if (comparison<acceptable_error) {</pre>
577
    * positioning_fine();
578
579
    * }else{
580
    * positioning_coarse();
    * }
581
    */
582
     read_IMU();
583
    Serial1.print("Axis Z position: ");
584
585
     Serial1.println(Yaw_deg);
      Serial1.println("Insert degree value to turn");
586
     OBC_data_value = 0;
587
     Timer1.attachInterrupt(TIMER_CH3, OBC_data_receive);
588
589
     while (OBC_data_value == 0)
590
     {
      delay(1); // If not used the while function does not work
591
592
      }
     Timer1.detachInterrupt(TIMER_CH3);
593
594
      Serial1.print ("Value to turn: "); // In this case turn, but can be what is needed
595
      Serial1.print(OBC_data_value);
596
597
598
      // Initiate EmergencyStop routine
599
     Timer1.attachInterrupt(TIMER_CH3, EmergencyStop);
600
      // Depend on the tolerance enable Fine or Coarse positioning
601
      if (abs(OBC_data_value) <= Pointing_mode_tolerance)</pre>
602
     {
603
      positioning_Fine();
604
605
      }
     else
606
607
     {
608
      positioning_Coarse();
```

```
609
     }
610 }
611
612 void positioning_Coarse()
613 {
614
     /*
615
      1.1. Mode Positioning Coarse: For now simple
616
       INPUT:
617
       None
618
       OUTPUT:
619
       None, but exits to positioning_Fine or mode_select
620
621
      */
622
     Serial1.println("Mode Positioning Coarse");
623
624
625
     // Stores a new variable to not overwrite the original value, also for motor as ...
       it needs a positive value and direction
     float degree_turn_value = OBC_data_value;
626
     bool RW_direction = true; // true=positive (CCW), false=negative (CW) along ...
627
       CubeSat's z-axis
                                // In case OBC_data_value is positive, value remains ...
628
       the same and rw direction is true, the default value of the variable
                                // In case OBC_data_value is negative, value change to ...
629
        positive and rw direction is false
630
     if (degree_turn_value < 0)</pre>
631
632
     {
633
       degree_turn_value = -degree_turn_value;
       RW_direction = false;
634
635
     }
636
637
       float Initial_IMU_degree_value, Final_IMU_degree_value, Delta_degree_ramp, Degree_$top_wait;
638
     float Prev_Yaw_deg, Check_Yaw_deg;
     int overlap_count = 0;
639
     bool waiting = true;
640
641
     // -----ACCELERATION
642
643
644
     read_show_IMU();
     Initial_IMU_degree_value = Yaw_deg; // Stores initial degree, only in Z
645
     int initial_RW_speed = RW_speed;
646
     generate_ramp(RW_direction, 0, 30, 1);
647
     // O will not be used as we dont define the time the motor lasts to do an \ldots
648
       impulse. 255 is the max value
     // This values will be substituted by how we want the time and speed to reach in \ldots
649
       the ramp.
650
     read_show_IMU();
651
     Final_IMU_degree_value = Yaw_deg; // Stores final degree, only in Z
652
     Delta_degree_ramp = Final_IMU_degree_value - Initial_IMU_degree_value; // degree ...
653
       turnt on acc, only in Z.
654
655
     if (RW_direction == false) // RW Clockwise
656
     {
       Delta_degree_ramp = Final_IMU_degree_value - Initial_IMU_degree_value; // ...
657
        degree turnt on acc, only in Z. Should be positive
```

```
658
     }
659
    if (RW_direction)
660
661
     {
      Delta_degree_ramp = Final_IMU_degree_value - Initial_IMU_degree_value; // ...
662
       degree turnt on acc, only in Z. Should be positive
      }
663
     else
664
665
     {
       Delta_degree_ramp = - (Final_IMU_degree_value - Initial_IMU_degree_value); // ...
666
        degree turnt on acc, only in Z. Should be positive
667
     }
668
     if (Delta_degree_ramp < 0)</pre>
669
670
     {
      Delta_degree_ramp += 360; // In case its negative adds 360
671
       // Negative cases: changes goes by 0 . EX: 330 to 30 when CCW (should be 60 \dots
672
       but calculus is 30-330= -300)
                                               EX: 30 to 330 when CW (should be 60 ...
673
       but calculus is 30-330= -300)
674
     }
     // Serial1.print("ID: ");
675
676
     // Serial1.print(Initial_IMU_degree_value);
     // Serial1.print(" / T: ");
677
     // Serial1.print(Delta_degree_ramp); // Difference of ange
678
     // Serial1.print(" / ED: ");
679
         Serial1.println(Final_IMU_degree_value); // Final, End angle
680
     // ------WAITING
681
     if ((degree_turn_value - 2 * Delta_degree_ramp) > 0) // if it is <0, skip the ...
682
       waiting phase, deceleration must be done immediately after, and still it ...
        would be too much turn.
683
     {
       waiting = true;
684
       if (RW_direction)
685
686
       {
687
         // CASE CCW
688
         /*
689
         RW CCW
690
         CB CW
691
         IMU positive
692
693
         Angulo inicial 300
694
695
          gira 150 en sentido CB CW = degree_turn_value
696
697
         angulo final acc 350 = Final_IMU_degree_value
698
699
700
          Delta_degree_ramp=350-300=50
701
          Degree_stop_wait=350+(150-2*50)=400; 40
702
703
         */
704
705
          . . .
        Degree_stop_wait = Final_IMU_degree_value + (degree_turn_value - 2 * Delta_degree_tamp); // ...
        Get value of degree to start stop
706
707
         Timer1.attachInterrupt(TIMER_CH4, read_show_IMU);
```

```
708
          Serial1.println("Waiting");
709
          while (waiting)
          { // Stays as long as waiting is true.
710
711
712
            // CAUTION WITH READING VALUES; AS IT IS ALWAYS FROM 0 TO 360
            // Degree_stop_wait will always be > Yaw_deg. If Yaw_deg>> (ex: 359 ), ...
713
        Degree_stop_wait can be >360. Thus the value of If Yaw_deg would never reach ...
        Degree_stop_wait
            // Degree_stop_wait cant be decreased, as the way to check if its reached ...
714
        is by a greater. If it is decreased by 360 (so it stays in relative place), ...
        it could be < Yaw_deg and would immediately exit the while without waiting.
            // the way to do is check if there is a heavy change on Yaw_deg (pass on ...
715
        0) to check if adding or substracting a lap, making it go out of the 0 and ...
        360 range.
            // It should not be a high acceleration enough to make a jump of degree of \ldots
716
        180
             in such short time, so it should be fine.
717
            // Caution disconnection from IMU, could give a false jump, that is why a ...
        check on Yaw_deg first.
718
            if (Yaw_deg < 0.0001 && Yaw_deg > -0.0001)
719
            \{ // \text{ if it is almost exactly 0 then most probably there is a ... }
720
        disconnection. Close numbers should not trigger it.
            }
721
            else if (Yaw_deg - Prev_Yaw_deg < -180)</pre>
722
            {
                                  // Checks for a heavy change (greater than half ...
723
        turn). Done by changes in 0, like jumping from 359 to 0.
              overlap_count += 1; // adds a lap
724
725
            }
726
            else if (Yaw_deg - Prev_Yaw_deg > +180)
            {
                                   // Checks for a heavy change (greater than half ...
727
        turn). Done by changes in 0, like jumping from 0 to 359. (Not possible in ...
        theory, as it increases, but deviations could mess it up)
             overlap_count -= 1; // substracts a lap
728
                                  // if neither are triggered, change has been samall \ldots
729
            }
        or no change has been done yet
            Prev_Yaw_deg = Yaw_deg;
730
            Check_Yaw_deg = Yaw_deg + (360 * overlap_count); // Rewriting of value, in ...
731
        new variable so it does not add itself.
           if (Check_Yaw_deg > Degree_stop_wait)
732
                               // As it is CCW, degree increases. When reading is > to ...
733
            {
        stop value, exits the while
              waiting = false; // exit condition
734
735
            }
            delay(1); // To solve errors
736
            // No writing of read_IMU as it is already done by a timer. Just to remind ...
737
        Yaw_deg is constantly reading values.
738
         }
         Timer1.detachInterrupt(TIMER_CH4);
739
740
        }
741
        else
742
        {
          // CASE CW
743
          // if its <0, it must be done immediately after, and still it would be too ...
744
        much turn.
745
          . . .
        Degree_stop_wait = Final_IMU_degree_value - (degree_turn_value - 2 * Delta_degree_tamp); // ...
        Get value of degree to start stop
746
```

```
Timer1.attachInterrupt(TIMER_CH4, read_show_IMU);
747
748
          while (waiting)
          { // Stays as long as waiting is true.
749
            Serial1.println("Waiting");
750
751
            // CAUTION WITH READING VALUES; AS IT IS ALWAYS FROM 0 TO 360
752
            // Degree_stop_wait will always be < Yaw_deg. If Yaw_deg<< (ex: 1 ), ...</pre>
753
        Degree_stop_wait can be <0. Thus the value of If Yaw_deg would never reach ...
        Degree_stop_wait
            // Degree_stop_wait can't be increased, as the way to check if its ...
754
        reached is by a greater. If it is increased by 360 (so it stays in relative ...
        place), it could be > Yaw_deg and would immediately exit the while without ...
        waiting.
            // the way to do is check if there is a heavy change on Yaw_deg (pass on ...
755
        0) to check iff adding or substracting a lap, making it go out of the 0 and \dots
        360 range.
756
            // It should not be a high acceleration enough to make a jump of degree of ...
        180
             in such short time, so it should be fine.
            // Caution disconnection from IMU, could give a false jump, that is why a \ldots
757
        check on Yaw deg first.
758
            Check_Yaw_deg = Yaw_deg; // New variable so it does not change during an ...
759
        iteration
            if (Check_Yaw_deg < 0.0001 && Check_Yaw_deg > -0.0001)
760
            { // if it is almost exactly 0 then most probably there is a \ldots
761
        disconnection. Close numbers should not trigger it.
            }
762
            else if (Check_Yaw_deg - Prev_Yaw_deg > +180)
763
                                   // Checks for a heavy change (greater than half \ldots
764
            {
        turn). Done by changes in 0, like jumping from 359 to 0. (Not possible in ...
        theory, as it decreases, but deviations could mess it up)
              overlap_count += 1; // adds a lap
765
            }
766
            else if (Check_Yaw_deg - Prev_Yaw_deg < -180)</pre>
767
768
            {
                                  // Checks for a heavy change (greater than half ...
        turn). Done by changes in 0, like jumping from 0 to 359.
             overlap_count -= 1; // substracts a lap
769
                                   // if neither are triggered, change has been samall \ldots
770
            }
        or no change has been done yet
771
772
            Prev_Yaw_deg = Check_Yaw_deg;
            Check_Yaw_deg = Check_Yaw_deg + (360 * overlap_count); // Rewriting of ...
773
        value. It does not add itself as it gets the read value in each iteration. In \ldots
        other words, it will not go through the iteration with numbers outside 0 and ...
        360 range
            if (Check_Yaw_deg < Degree_stop_wait)</pre>
774
                                // As it is CCW, degree increases. When reading is > to \ldots
775
            {
        stop value, exits the while
776
              waiting = false; // exit condition
777
            }
            delay(1); // To solve errors
778
            // No writing of read_IMU as it is already done by a timer. Just to remind \ldots
779
        Yaw_deg is constantly reading values.
780
781
          Timer1.detachInterrupt(TIMER_CH4);
        }
782
783
      }
784
```

```
785
                            -----DECELERATION
786
     // Serial1.print("ID: ");
     // Serial1.print (Final_IMU_degree_value);
787
     // Serial1.print(" / ED: ");
788
789
      // Serial1.println(Degree_stop_wait);
790
      // Initial_RW_Speed returns speed to original value instead of 0
791
      generate_ramp(RW_direction, initial_RW_speed, 0, 0);
792
     // This is to assure the impulse is the same, as if there is some momentum ...
793
       accumulation, returning to 0 would be a different impulse
      // O will not be used as we dont define the time the motor lasts to do an \ldots
794
       impulse. 255 is the max value
795
     // This values will be substituted by how we want the time and speed to reach in \ldots
       the ramp.
796
797
     read_show_IMU();
798
      Final_IMU_degree_value = Yaw_deg; // can be overwritten, final degree of manoeuvre
      if (RW_direction)
799
800
     {
      Delta_degree_ramp = Final_IMU_degree_value - Initial_IMU_degree_value; // ...
801
       Should be positive
802
    }
     else
803
804
     {
      Delta_degree_ramp = Initial_IMU_degree_value - Final_IMU_degree_value; // ...
805
       Should be positive
806
      }
     if (Delta_degree_ramp < 0)</pre>
807
808
     {
                                  // can be overwritten, real turn of manoeuvre
      Delta_degree_ramp += 360; // In case its negative adds 360
809
       // Negative cases: changes goes by 0 . EX: 330 to 30 when CCW (should be 60 ...
810
       but calculus is 30-330= -300)
                                               EX: 30 to 330 when CW (should be 60 ...
811
       but calculus is 30-330= -300)
812
      }
813
     // Serial1.print("ID: ");
814
     // Serial1.print(Degree_stop_wait);
815
     // Serial1.print(" / T: ");
816
     // Serial1.print(Delta_degree_ramp);
817
     // Serial1.print(" / ED: ");
818
     // Serial1.println(Final_IMU_degree_value);
819
820
821
     . . .
       if ((Delta_degree_ramp - degree_turn_value) < Final_pointing_tolerance && (Delta_degree_ramp - de
822
      {
       // Detach EmergencyStop
823
       Timer1.detachInterrupt(TIMER_CH3);
824
825
       // Real turn vs wanted turn, checks if it is considered good
826
       mode_OBC_Input_Wait(); // Valid position
827
828
      }
      else
829
830
     {
       // Recalculate the OBC_data_value for PID
831
       OBC_data_value = OBC_data_value - Delta_degree_ramp;
832
      positioning_Fine(); // Correction
833
      }
834
```

```
835 }
836
837 void positioning_Fine()
838 {
839
    /*
      1.2. Mode Positioning Fine: PID
840
841
       INPUT:
842
       None
843
       OUTPUT:
844
845
       None, but exits to mode_select
      */
846
847
     Serial1.println("Mode Positioning Fine");
848
849
850
     bool waiting = true;
851
     read_show_IMU();
852
     Deg_to_reach = OBC_data_value + Yaw_deg; // Get value to reach, contained in 360
853
     if (Deg_to_reach < 0)</pre>
854
855
    {
      Deg_to_reach += 360;
856
857
      Zero_state = true; // Used to store if it passes 0.
       // A PD passing through 0 could give great problems, as it has a very big \ldots
858
        change in value. Further used in computePID()
859
      }
      else if (Deg_to_reach >= 360)
860
861
     {
      Deg_to_reach -= 360; // These two lines contain the value of the Yaw_deg in ...
862
       0-360 degrees.
       Zero_state = true;
863
864
      }
     else
865
866
     {
867
      Zero_state = false;
     }
868
869
    if (Zero_state)
870
871
    {
      Deg_to_reach + 180;
872
      if (Deg_to_reach >= 360)
873
         Deg_to_reach -= 360;
874
875
      }
876
      Timer1.attachInterrupt(TIMER_CH4, computePID);
877
      while (waiting)
878
     { // Range of tolerance
879
880
        . . .
        if (abs(IMU_gyro_data_Z) < Gyro_tolerance && abs(IMU_accel_data_X) < Accel_toleran¢e && abs(Deg_t
881
       { // we consider it is stopped, modify values to be accurate
         waiting = false;
882
       }
883
884
        delay(1); // Change
885
      }
886
      waiting = true;
      Timer1.detachInterrupt(TIMER_CH4);
887
888
889
      // At exit, RW_speed could not be 0
```

```
890
     OBC_data_value = 0;
891
     Serial1.println("End of manoeuvre");
892
893 // Detach EmergencyStop
894
   Timer1.detachInterrupt(TIMER_CH3);
895
896
    mode_Select(OBC_mode_value);
897 }
898
   void mode_IMU_reading()
899
900
   {
    /*
901
902
      2- IMU reading (TEST PURPOSE Function)
903
      INPUT:
904
905
      None
906
       OUTPUT:
       None, but saves IMU variables and Roll, Pitch and Yaw
907
     */
908
909
910 OBC_mode_value = 0;
911 Timer1.attachInterrupt(TIMER_CH4, read_show_IMU);
912 Timer1.attachInterrupt (TIMER_CH3, OBC_data_receive);
913
    while (OBC_data_value == 0)
914
     {
915
     delay(1); // If not used the while function does not work
916
     }
   Timer1.detachInterrupt(TIMER_CH3);
917
   Timer1.detachInterrupt(TIMER_CH4);
918
919
   OBC_data_value = 0;
   mode_Select(OBC_mode_value);
920
921 }
922
923 void mode_motor_on_off()
924 {
925
    /*
      3- Motor ON/OFF (TEST PURPOSE Function)
926
927
      INPUT:
928
      None
929
      OUTPUT:
930
      None
931
   */
932
   OBC_mode_value = 0;
933
     set_impulse(true, 100);
934
     Timer1.attachInterrupt (TIMER_CH3, OBC_data_receive);
935
     while (OBC_data_value == 0)
936
937
    {
938
      delay(1); // If not used the while function does not work
939
    }
   Timer1.detachInterrupt(TIMER_CH3);
940
   set_impulse(true, 0);
941
     OBC_data_value = 0;
942
943
     mode_OBC_Input_Wait();
944 }
945
946 // ...
```

```
VOID SETUP
947 void setup()
948 {
    delay(1000); // wait to let open the serial
949
    // ...
950
       Timers
     Timer1.pause();
951
     Timer1.setPrescaleFactor(7200); // 72MHz Clock / 7200 = 10KHz timer
952
     Timer1.setOverflow(1000); // Overflow occurs at 1000, each 100 ms timer ...
953
      restarts
954
955
     Timer1.setMode(TIMER_CH3, TIMER_OUTPUT_COMPARE); // Configure channel to ...
       OUTPUTCOMPARE: Channel for OBC read values
     Timer1.setMode(TIMER_CH4, TIMER_OUTPUT_COMPARE); // Channel for IMU read values
956
                                                      // Phase value in Overflow range
957
     Timer1.setCompare(TIMER_CH3, 1);
958
     Timer1.setCompare(TIMER_CH4, 1);
959
     Timer1.refresh(); // Refresh timer and start over
960
     Timer1.resume();
961
962
     // ...
963
       Initiations
964
     Serial1.begin(9600);
     SPI.begin();
965
966
     Serial1.println("START");
967
968
     // ...
969
       Setups
970
     pinMode(LEDPIN, OUTPUT); // Integrated LED
971
972
     pinMode(CS1, OUTPUT); // NCS Pin definition
973
     // ...
974
       Driver Setup
     Serial1.println("Configuring Driver...");
975
     DriverOne.settings.commInterface = I2C_MODE; // Driver Comm Mode
976
     DriverOne.settings.I2CAddress = 0x5D; // Driver Address (0x5D by Default)
977
978
     while (DriverOne.begin() != 0xA9)
979
     { // Driver wait for idle
980
       Serial1.println("ID Mismatch");
981
       delay(200);
982
983
     }
984
     Serial1.println("ID Match");
985
     Serial1.println("Waiting for enumeration"); // Driver wait for peripherals
986
     while (DriverOne.ready() == false)
987
988
       ;
     Serial1.println("Ready");
989
990
     while (DriverOne.busy())
991
      ; // Driver enable
992
993
     DriverOne.enable();
```

```
994
      Serial1.println("Driver Ready to Use!");
995
      // ...
996
        IMU Setup
      Serial1.println("Configuring MPU9250...");
997
      int status = IMU.begin();
998
      if (status < 0)
999
000
      {
        Serial1.println("IMU initialization unsuccessful");
001
       Serial1.println("Check IMU wiring or try cycling power");
002
        Serial1.print("Status: ");
1003
       Serial1.println(status);
1004
       while (1)
1005
        {
006
007
        }
008
      }
009
      IMU.setGyroRange(IMU.GYRO_RANGE_500DPS);
010
011
      // ...
1012
        IMU Calibration
      // Use for getting the values on calibration setting below
1013
014
      // To calibrate the magnetometer or the compass, everything is commented except ...
015
       the last 3 lines.
      // It moves in the shape of an eight approximately 2 min. It is recommended to \ldots
1016
       carry out several times until the measurements are fine-tuned.
      // Uncomment everything and enter the values obtained from the MagBias and ...
1017
       ScaleFactor to view the results of the magnetometer.
1018
              IMU.calibrateMag();
019
      11
              Serial1.println("Done");
020
021
.022
              Serial1.print(IMU.getMagBiasX_uT());
              Serial1.print(",");
023
              Serial1.print(IMU.getMagBiasY_uT());
024
     11
              Serial1.print(",");
025
     11
              Serial1.println(IMU.getMagBiasZ_uT());
026
027
             Serial1.print(IMU.getMagScaleFactorX());
028
              Serial1.print(",");
029
              Serial1.print(IMU.getMagScaleFactorY());
.030
              Serial1.print(",");
.031
              Serial1.println(IMU.getMagScaleFactorZ());
032
033
034
     // Reactioni wheels + Magnetorquer
1035
     IMU.setMagCalX(9.42, 0.91); // The first value corresponds to the MagBias, and ...
       the second the ScaleFactor.
     IMU.setMagCalY(41.82, 0.98);
1036
      IMU.setMagCalZ(-6.59, 1.14);
037
038
      // Only reaction wheels
039
     // IMU.setMagCalX(14.06, 0.84); // The first value corresponds to the MagBias, ...
040
       and the second the ScaleFactor.
     // IMU.setMagCalY(29.85, 1.36);
041
      // IMU.setMagCalZ(-32.30, 0.94);
1042
```

```
1043
    Serial1.println("MPU9250 Ready to Use!");
1044
1045
1046
    Serial1.println("Reading mode from OBC");
    mode_OBC_Input_Wait();
1047
1048 }
049
050 // ...
        VOID LOOP
1051 void loop()
1052
   {
1053
   }
```

 $Code \ O.1$  Full control of the Cubesat with coarse and fine pointing mode. Source: Own.

Name	Description	Type	Units	Ambit	
LEDPIN	Integrated LED of the Bluepill	#define	-	Bluepill configuration	
CS1	STM32 NCS Chip select (SPI mode only) pin	#define	-	Bluepill configuration	
STM32_CLOCK	Internal STM32 Clock	#define	kHz	Bluepill configuration	
PI	Number PI	#define	-	Math values	
Rad_to_deg	Convert radians to degrees	#define	-	Math values	
Deg_to_rad	Convert degrees to radians	#define	-	Math values	
Final_pointing_tolerancy	Tolerancy for final pointing	#define	$Deg (\hat{A}\check{z})$	Tolerancies	
Pointing_mode_tolerancy	Tolerance for selecting pointing mode	#define	$Deg (\hat{A}\check{z})$	Tolerancies	
Accel_Tolerance	Tolerance for acceleration measurement	#define	$m/s^2$	Tolerancies	
Gyro_Tolerance	Tolerance for gyro measurement	#define	Deg/s	Tolerancies	
DriverOne	Driver Object definition	SCMD	-	Object initiations	
IMU	MPU Object definition	MPU9250	-	Object initiations	
IMU_accel_data_X	Values of local IMU accelerometer along X	float	$m/s^2$	IMU Values	
	axis				
IMU_accel_data_Y	Values of local IMU accelerometer along Y	float	$m/s^2$	IMU Values	
	axis				
IMU_accel_data_Z	Values of local IMU accelerometer along Z	float	$m/s^2$	IMU Values	
	axis				
IMU_gyro_data_X	Values of local IMU gyroscope along X axis	float	Deg/s	IMU Values	
IMU_gyro_data_Y	Values of local IMU gyroscope along Y axis	float	Deg/s	IMU Values	
IMU_gyro_data_Z	Values of local IMU gyroscope along Z axis	float	Deg/s	IMU Values	
IMU_mag_data_X	Values of local IMU magnetometer along X	float	uT	IMU Values	
	axis				
IMU_mag_data_Y	Values of local IMU magnetometer along Y	float	uT	IMU Values	
	axis				
IMU_mag_data_Z	Values of local IMU magnetometer along Z	float	uT	IMU Values	
	axis				

## ${\bf Table \ O.1} \ {\rm Variables \ explanation \ of \ the \ control \ algorithm. \ Source: \ Own.}$

Pitch_deg	Pitch angle	float	$Deg (\hat{A}\check{z})$	Orientation angles
Roll_deg	Roll angle	float	$Deg (\hat{A}\check{z})$	Orientation angles
Yaw_deg	Yaw angle	float	$Deg (\hat{A}\check{z})$	Orientation angles
Accel_total_vector_modulus	Accelerometer total vector magnitude	float	$m/s^2$	Angles determination
Accel_pitch_deg	Accelerometer Pitch angle	float	$Deg (\hat{A}\check{z})$ Angles determinat:	
Accel_roll_deg	Accelerometer Roll angle	float	$Deg (\hat{A}\check{z})$ Angles determinat	
Gyro_pitch_deg	Gyroscope Pitch angle	float	$Deg (\hat{A}\check{z})$ Angles determinat	
Gyro_roll_deg	Gyroscope Roll angle	float	$Deg (\hat{A}\check{z})$ Angles determina	
Gyro_Accel_sync	Check if Gyroscope and Accelerometer are	bool	-	Angles determination
	synchronized			
Mag_X_calc	Calculated Magnetometer value along X axis	float	uT	Angles determination
Mag_Y_calc	Calculated Magnetometer value along Y axis	float	uT	Angles determination
Mag_X_corrected	Corrected Magnetometer value along X axis	float	uT	Angles determination
Mag_Y_corrected	Corrected Magnetometer value along Y axis	float	uT	Angles determination
Pitch_rad	Pitch angle	float	rad	Angles determination
Roll_rad	Roll angle	float	rad	Angles determination
RW_speed	Value of Reaction Wheel speed	int	-	Motor values
OBC_mode_value	Value of mode sent by the On Board Com-	int	-	OBC readings
	puter			
OBC_data_value	Value of turn sent by the On Board Com-	int	$Deg (\hat{A}\check{z})$	OBC readings
	puter			
Stop_State	Check if an Emergency Stop is performed	int	$Deg (\hat{A}\check{z})$	OBC readings
PID_error	Proportional error from PID	double	-	Fine Positioning (PID)
PID_last_error	Previous proportional error from PID	double	-	Fine Positioning (PID)
PID_cumulative_error	Cumulative (Integral) error from PID	double	-	Fine Positioning (PID)
PID_rate_error	Rate (Derivative) error from PID	double	-	Fine Positioning (PID)
kp	Proportional constant	double	-	Fine Positioning (PID)
ki	Integral constant	double	-	Fine Positioning (PID)

kd	Derivative constant	double	-	Fine Positioning (PID)	
Deg_to_reach	Setpoint angle	float	$Deg (\hat{A}\check{z})$	Fine Positioning (PID)	
Zero_state	Check if PID encompass yaw of 0 degrees	bool	-	Fine Positioning (PID)	
PID_output	Output value of the PID	int	bytes	Fine Positioning (PID)	
RW_direction	Direction of the Reaction Wheel	rection of the Reaction Wheel bool -		(Inside functions) Mo-	
				tor settings	
New_RW_speed	Speed to set on the Reaction Wheel	int	bytes	(Inside functions) Mo-	
				tor settings	
Acc_ramp_time_duration	Duration of ramp	int	s	(Inside functions) Mo-	
				tor settings	
RW_ramp_speed_reach	Final speed reached	int	bytes	(Inside functions) Mo-	
				tor settings	
Acc_Dec_state	Check wether in acceleration or deceleration	bool	-	(Inside functions) Mo-	
	state			tor settings	
dT	Time step of each interrupt, calculated from	float	s	(Inside functions) IMU	
	Timer			reading and Fine Posi-	
				tioning (PID)	

Name	Description	Input	Output	Ambit
OBC_mode_receive()	Get mode from OBC	None, but gets mode from	None, but saves	OBC readings
		User Serial1 input	OBC_mode_value	
OBC_data_receive()	Get data from OBC	None, but gets mode from	None, but saves	OBC readings
		User Serial1 input	OBC_data_value	
EmergencyStop	Stops the motor if the user calls an	None, but gets stop state	None	OBC readings
	emergency stop	from User Serial1 input		
<pre>set_impulse()</pre>	Set the motor at a fixed direction	RW_direction,	None, but saves RW_speed	Motor settings
	and speed	New_RW_speed		
generate_ramp()	Creates a ramp	RW_direction,	-	Motor settings
		$Acc\_ramp\_time\_reach,$		
		$RW\_ramp\_speed\_reach,$		
		Acc_Dec_state		
read_IMU()	Read data from the IMU	None	None, but saves IMU vari-	IMU reading
			ables, pitch, roll and yaw	
show_IMU()	Show data from the IMU on the Se-	None	None	IMU reading
	rial1 monitor			
read_show_IMU()	Read then show data from the IMU	None	None, but saves IMU vari-	IMU reading
			ables, pitch, roll and yaw	
computePID()	Compute the PID and updates the	None	None, but saves PID val-	Fine Positioning
	motor speed and direction		ues	(PID)
<pre>mode_OBC_Input_Wait()</pre>	Default mode, waits for a mode in-	None	None, but exits to	Modes
	put from OBC		mode_Select()	
<pre>mode_Select()</pre>	Decides the mode to execute de-	mode_value	None, but exits to selected	Modes
	pending on the mode read		mode	
<pre>mode_Positioning_RW()</pre>	Waits for a data input from OBC	None	None, but exits to suitable	Modes
	and decides the positioning type		positioning type	
	(coarse or fine)			

 Table 0.2 Functions of the control algorithm. Source: Own.

positioning_Coarse()	Executes a coarse pointing control	None	None, but exits to	Modes
	(ramps and waiting)		positioning_Fine()	
			or mode_Select()	
positioning_Fine()	Executes a fine pointing control	None	None, but exits to	Modes
	(PID)		mode_Select()	
<pre>mode_IMU_reading()</pre>	Test purpose function to read and	None	None, but saves IMU	Modes
	show IMU values constantly		values and exits to	
			mode_Select()	
mode_motor_on_off	Test purpose function to start and	None	None, but exits to	Modes
	stop the motor		mode_Select()	