

BACHELOR'S DEGREE IN AEROSPACE TECHNOLOGIES ENGINEERING

# EVALUATION OF PULSE DETECTION OF THE GEOSTATIONARY LIGHTNING MAPPER (GLM)

BACHELOR'S THESIS

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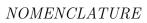
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# Nomenclature

| $\mathbf{CCD}$          | Charge Coupled Device                              |
|-------------------------|--|
| CHU                     | Camera Head Unit                                   |
| FOV                     | Field Of View                                      |
| GLM                     | Geostationary Lightning Mapper                     |
| GOES                    | Geostationary Operational Environmental Satellites |
| GPST                    | Global Positioning System Time                     |
| ISS                     | International Space Station                        |
| LCFA                    | Lightning Cluster Filter Algorithm                 |
| LMA                     | Lightning Mapping Array                            |
| MMIA                    | Modular Multispectral Imaging Array                |
| MXGS                    | Modular X- and Gamma- ray Sensor                   |
| $\operatorname{NetCDF}$ | Network Common Data Form                           |
| NOAA                    | National Oceanic and Atmospheric Administration    |
| OTD                     | Optical Transient Detector                         |
| РНОТ                    | Photometer   |
| TGF                     | Terrestrial Gamma-ray Flashes                      |
| TLE                     | Transient Luminous Events                          |
| TRMM                    | Tropical Rainfall Measuring Mission                |



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Date

Title of the Thesis:

Evaluation of pulse detection of the Geostationary Lightning Mapper (GLM)



#### Abstract

This study evaluates the sensitivity of the Geostationary Lightning Mapper (GLM) versus the Modular Multispectral Imaging Array (MMIA)'s photometer 3, both operating at the oxygen band at 777.4nm of wavelength. To do so, both GLM's and MMIA's data is extracted from pre-processed data files by a Lightning Cluster Filter Algorithm (LCFA) -data classified by time, geolocalization, detection intensity and with an ID number- to be treated by classification of their signals according to Linet detections in Colombia, signal synchronization and peak comparison.

#### Resumen

En este estudio se evalúa la sensibilidad del sistema GLM (Geostationary Lightning Mapper) frente a la de los fotómetros de MMIA (Modular Multispectral Imaging Array), en concreto el fotómetro 3 captando la luz en 777.4 nm de longitud de onda. Para ello, tanto los datos de GLM como de MMIA son extraídos de los ficheros con información preprocesada por LCFA (Lightning Cluster Filter Algorithm) -datos clasificados por tiempo, geolocalización, intensidad de la detección y con un número identificativo- para posteriormente ser tratados mediante la clasificación sus señales en detecciones dadas por la red de detección de rayos Linet en Colombia, sincronización de señales y comparación de picos de señal.

# INTRODUCTION



# 1 Introduction

### 1.1 Object

The objective of this study is to develop a program to evaluate the pulse detection sensitivity of the Geostationary Lightning Mapper (GLM) aboard the GOES-16 satellite against the sensitivity of the 777.4nm photometer of the Modular Multispectral Imaging Array (MMIA) on the Columbus Module aboard the International Space Station (ISS) over lightning detections in Colombia.

To achieve that goal, given data from both instruments is broken into data snippets according to Colombian Lightning Location Network (LINET) and compared snippet-to-snippet, cross-correlating their signals and outputting their relative delay as well as their detected and non-detected peaks.

### 1.2 Justification

As comprehension of atmospheric phenomena increase, more effort is put into the study of their consequences and possible potential or risk. One of those atmospheric phenomena is thunderstorm processes, especially lightning activity.

Understanding of those phenomena allows for improvement in many different areas of knowledge as well as it has many direct practical applications. Lightning activity monitoring given by different systems can allow for, for instance, better comprehension of severe thunderstorm lead times and dynamics, early warning of lightning ground strikes or even provide valuable data for improving numerical weather prediction models, as well as decreasing weather prediction uncertainty and false alarm probabilities. In more direct applications, this knowledge can be applied to improving aero- and nautical routing over oceanic regions where lightning activity information is scarce and work as an assistant to storm radar systems in locations where radar coverage is poor [7].

All that valuable information must be given by a lightning detection system in



which to rely on to assure detection accuracy in geolocation, time and magnitude. In this study the GLM instrument is evaluated against a MMIA's photometer, both space-based, to account just how reliable its detections really are and how many of them are ignored by this system. This would allow for better future analysis and results in lightning detection studies as well as for better tuning in future detection systems.

While GLM is an operational-purpose instrument with built-in GPS Time (GPST), MMIA is a high time-resolution and sensitivity scientific instrument with own time. This leads to a GLM time accuracy of 2ms, while MMIA's time accuracy is to up to 20ms. With this study, high quality lightning detections from MMIA are cross-correlated with time-accurate GLM detections in order to have ASIM data with a time accuracy of 2ms, 10 times better than before, reporting high-detailed, time-accurate detection data in a high sample rate of 100kHz.

### 1.3 Scope

In order to compare signals from GLM and MMIA using cross-correlation, both different datasets must be treated and driven into a similar structure to allow direct analysis. This section schematizes the key elements to proceed from study on the field to results outputting.

#### 1.3.1 Preliminary study

- Review on previous research on lightning detection.
- Review on GLM and MMIA instruments.
- Review on future lightning detection systems.

#### 1.3.2 Lightning detection data study

• Development of a general resolution algorithm.



- Development of tools for extracting GLM and MMIA data from their given file formats to similar data structures.
- Development of tools for trimming GLM and MMIA data into snippets from LINET data.
- Development of tools for GLM and MMIA data conditioning before cross-correlation.
- Development of tools for rejection of trivial snippets.
- Development of a snippet synchronisation system via cross-correlation.
- Development of signal time delay counter.

### 1.3.3 Lightning detection performance

- Development of tools for peak detection in GLM and MMIA cross-correlated signals.
- Development of tools for accounting detected and non-detected peaks among those instruments.
- Development of tools for accounting difference in order of magnitude in common detected peaks.

### 1.4 Requirements

Requirements and restrictions on the development of the architecture of the solution algorithm, the result presentation on this report and the developed code is presented in Table (1.1).

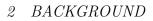




| TABLE $1.1$ : | <b>Project Requirements</b> |
|---------------|-----------------------------|
|---------------|-----------------------------|

| Requirement ID | Description   |
|----------------|---|
| REQ-1          | Comparisons must be made independently of the differ-<br>ent instruments' nature.   |
| REQ-2          | Data output must be equivalent among different instru-<br>ments in order to be able to make a comparison.                           |
| REQ-3          | The developed program and its functions should be<br>based on free software for compatibility reasons.                              |
| REQ-4          | Accuracy of results must be inside acceptable margins<br>as to base a scientific study upon them.                                   |
| REQ-5          | Verification plots must be included in the report as to<br>understand all processes and their accuracy.                             |
| REQ-6          | Result plots must be included in the report as to show<br>the final output of the program.  |
| REQ-7          | The program must check a minimum correlation among different instruments' data to better determine a match.                         |
| REQ-8          | The program must be fully modular in order to make it<br>comprehensive and easy to retouch.   |
| REQ-9          | Statistics about GLM and MMIA peak detections must<br>be outputted from the program.  |
| REQ-10         | The analysis must accept automatization of routines to<br>study $n$ cases without problems, regardless of the size<br>of the input. |
| REQ-11         | A full documentation of the code should be made in the report.  |
| REQ-12         | Code documentation must explain clearly every step of the script.   |

Development



## 2 Background

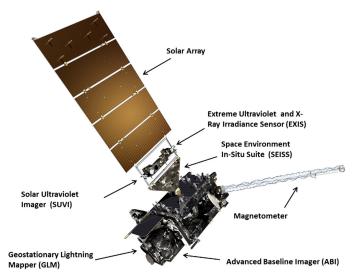
In this section a brief introduction to GLM and MMIA instruments is made, as these are the instruments to be analysed and compared along the study.

#### 2.1 GLM

The Geostationary Lighting Mapper (GLM) (Fig.(2.1)) is a high-speed event detector operating at near-infrared wavelength to study both cloud and cloud-to-ground lightning activity 24 hours a day over the American continent. It is a major aid in detecting potentially dangerous storms or weather elements that may affect aviation [5]. The GLM is designed to operate on The Geostationary Operational Environmental Satellite R-series (GOES-R) which becomes part of the original GOES constellation. This set of satellites deployed by the National Aeronautics and Space Administration (NASA) and the National Oceanic and Atmospheric Administration (NOAA) is responsible for providing weather forecasts and warnings. The aim of adding the GLM is to obtain greater accuracy in both timing and prediction.

Currently there are other methods for lightning detection such as the NASA Lightning Imaging Sensor (LIS) and the Optical Transient Detector (OTD) that operate in low Earth orbit. The GLM is able to obtain very similar data and extend its combined climatology to study long-term effects such as climate change for about the next 20 years [4, 1].

To perform well, the GLM has many more requirements than a simple imager. The transient nature of lightning, daylight sampling when there may be solar reflections or spectral characteristics are some of the challenges it faces. A field-of-view (FOV) lens is used together with a narrow-band interference filter and is focused to a high speed Charge Coupled Device (CCD) focal plane. The data is then sent to and processed on the satellite's Local Area Network (LAN). Even though, the LIS also had similar characteristics and made use of the same techniques, with the GLM they have been considerably improved to obtain a much higher accuracy.



(a) Illustration of the GOES-16 geostationary satellite, carrying GLM

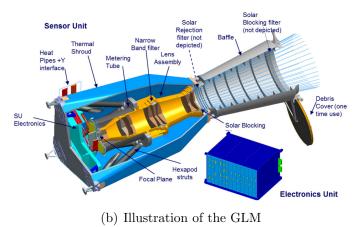


FIGURE 2.1: Illustrations of GLM instrument aboard the GOES-16 satellite, both from [5]

Through the two satellites GOES-E (75 W) and GOES-W (135 W), the focal CCD of the GLM with a resolution of  $1372 \times 1300$  pixels can focus on the storms at any time (see Fig.(2.2)). Being in a geostationary orbit, it has a field of view of practically the whole hemisphere with a resolution at nadir of about 8 km and 14 km for the FOV edge. To achieve such a uniform coverage, the pixel distribution has been densified at the extremes by using smaller pixels that compensate the resolution [3].

The device can detect approximately 86% of the lightning strikes and can even reach 90%. To achieve such good performance it is necessary to use a solar blocking filter at the aperture of the instrument together with a solar rejection filter to limit



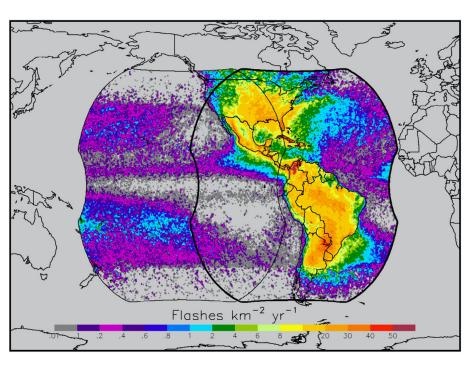


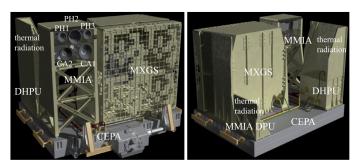
FIGURE 2.2: Combined FOV view from the GOES-R series constellation (75 W, 137 W) superimposed on 10-yr of lightning observations from the NASA Lightning Imaging Sensor on board the Tropical Rainfall Measuring Mission (TRMM/LIS) and Optical Transient Detector (OTD) low earth-orbiting satellites, from [5]

the light outside the band from entering the instrument. In addition, a 1-nm narrowband interference filter is added to ensure that the 777.4 OI oxygen triplet passes to the detector [5].

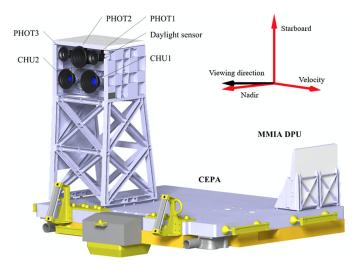
### 2.2 ASIM

The latest discoveries related with thunderstorms caused a huge amusement in the scientific community. These discoveries show that lightning and thunderstorm can produce electrical breakdown above storms. This event is called Transient Luminous Events (TLEs). TLE can be shown in different forms such as "sprites" (electrical discharges in the mesosphere at 50 to 80 km) and the "blue jets" (streamer type discharges propagating upward from clouds) among others. The electrical band in which these phenomena occur mostly are 337nm, near ultraviolet (NUV).

Due to the great interest of TLEs, a scientific experiment was designed to study this phenomenon, the Atmosphere-Space interactions Monitor (ASIM) (Fig.(2.3)). The ASIM mission's major scientific objectives are to examine thunderstorm electri-



(a) Illustration of the ASIM payload, from [6]



(b) Illustration of MMIA instrument, from [2]

FIGURE 2.3: Illustrations of ASIM payload and MMIA ubication

cal activity such as lightning, Transient Luminous Emissions (TLEs), and Terrestrial Gamma-ray Flashes (TGFs) by studying the associated emissions in the UV, nearinfrared, x-, and gamma-ray spectral bands. The ASIM is formed by two main instruments placed on the Colombus module of the European Space Agency on the International Space Station; the The Modular Multispectral Imaging Array (MMIA) and the Modular X- and Gamma- ray Sensor (MXGS).

The MMIA consists of an array of optical sensors [2]. It is in charge of analysing the TLE phenomenon using three co- aligned photometers and two camaras with the highest sensity, dynamical range and temporal resolution. The instruments are so sensitive to light that they are only operated during nighttime. The two cameras called Camera Head Units (CHUs) are composed by three key elements; an optical assembly consisting of a baffle to reduce stray light and optics hosting a narrow band





filter, a focal plane assembly containing an Electron Multiplication Charge Coupled Device (EM-CCD) of high sensitivity, and control and readout electronics capable of reading out up to 12 full frames per second from the sensor. The three co- aligned photometers (PHOTs) are made up of an optical assembly that includes a baffle to limit stray light, lenses that concentrate on the photocathode of a Photo-Multiplier Tube (PMT) in photon counting mode, proximity electronics, and a calibration light emitting diode (LED).

The software is separated into two parts: Boot Software (BSW) and Application Software (ASW), with the BSW running when the computer turns on. Both BSW and ASW are implemented in Ada 2012 using AdaCore GNAT Pro for LEON Bare Board, eliminating the need for an operating system.

MMIA may function in a variety of operational modes focused on main and secondary science goals. The triggered data collection mode enables for the recording of rapid changes in light for the major research objectives of observing transient bright flashes of emissions from thunderstorms. This is used to collect information about lightning strikes, TLEs, meteors, and TGFs. The timed data collection mode enables the performance of programmed periodic observations of set length for secondary research purposes such as auroras.

In conclusion, MMIA is composed of two cameras imaging in the 337 nm and 777.4 nm bands, with a frame rate up to 12 frames per second, and three high-speed photometers in the 180–230 nm, 337 nm and 777.4 nm bands, sampling at rates up to 100 kHz.

# 3 Explanation of the resolution algorithm

In this section the approach to the problem is explained with detail. Section 3.1 presents a visual representation of the main\_TFG.py program architecture, sections 3.2 and 3.3 describe the initial approach to data handling and its information extraction for analysis, accordingly. Finally, section 3.4 explains data synchronisation and peak detection on the signals for later result outputting by the program.

# 3.1 General resolution algorithm

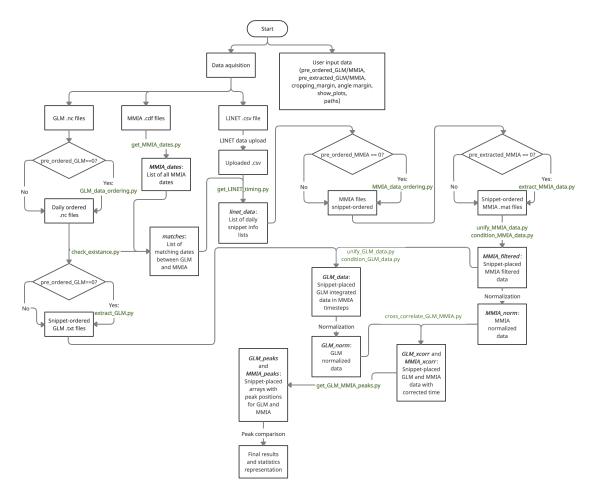


FIGURE 3.1: Visual representation of the global resolution algorithm



#### 3.2 Initial data management

As explained in section 3, this section explains how LINET's, GLM's and MMIA's data is being given as well as how its information is treated into easily manipulable data in order to identify those vector sections with useful information to study.

#### 3.2.1 GLM's data order

GLM data is given by National Oceanic and Atmospheric Administration (NOAA) via a Google Cloud Platform Repository, where thousands of .nc files are stored with GLM information for 200ms each and ordered by year, day-of-year and hour-of-day. Network Common Data From (NetCDF, giving .nc files) is a vastly used multiplatform open-source binary file format to share large amounts of array-oriented data in a way that is self-describing, portable and efficient [11]. In the case of GLM data, those .nc files contain information of events, flashes and groups as well as positioning and configuration of the satellite at every moment. Fig.(3.2) gives an idea of the structure of the information inside a GLM .nc file and its contents, using the specialized program *Panoply*. As it can be seen it hosts many different variables (rows) with a short description each in the form of arrays (1D).

Once the files of the dates to analyse have been downloaded from the repository, an algorithm to analyse each GLM date range was developed as follows. The program main\_TFG.py asks for having them all together inside a single directory. It demands the existence of two more void directories, one for hosting multiple daily directories with files inside, and another to host extracted snippet data. Inside USER INPUT DATA section in main\_TFG.py (see appendix A) the paths to those directories has to be written as a string<sup>1</sup>. Code snippet 3.1 shows an example of this input.

Function GLM\_data\_ordering.py (see appendix B.1) then sweeps along all .nc files inside the directory GLM\_files\_path where they are located creating a new

<sup>&</sup>lt;sup>1</sup>IMPORTANT: Those paths must NOT contain a final '/' (slash) as it will be added when needed.



| ate Plot Combine Plot Open Dataset       |   | Remove All Show |
|--|---|-----------------|
| tasets Catalogs Bookmarks                |   |                 |
| Name                                     | Long Name   | Туре            |
|  | 6_s2 GLM L2 Lightning Detections: Events, Groups, and Flashes           | Local File      |
| 🗢 algorithm_dynamic_input_data_container | container for filenames of dynamic algorithm input data                 | _               |
| algorithm_product_version_container      | container for algorithm package filename and product version            | -               |
| 🗢 event_count                            | number of lightning events in product                                   | -               |
| 🗢 event_energy                           | GLM L2+ Lightning Detection: event radiant energy                       | 1D              |
| 🗢 event_id                               | product-unique lightning event identifier                               | 1D              |
| 🗢 event_lat                              | GLM L2+ Lightning Detection: event latitude coordinate                  | 1D              |
| 🗢 event_lon                              | GLM L2+ Lightning Detection: event longitude coordinate                 | 1D              |
| 🗢 event_parent_group_id                  | product-unique lightning group identifier for one or more events        | 1D              |
| event_time_offset                        | GLM L2+ Lightning Detection: event's time of occurrence                 | 1D              |
| 🗢 flash_area                             | GLM L2+ Lightning Detection: flash area coverage (pixels containing a   | 1D              |
| flash_count                              | number of lightning flashes in product                                  | -               |
| 🗢 flash_energy                           | GLM L2 + Lightning Detection: flash radiant energy                      | 1D              |
| flash_frame_time_offset_of_first_event   | GLM L2 + Lightning Detection: time of occurrence of first constituent e | 1D              |
| flash_frame_time_offset_of_last_event    | GLM L2 + Lightning Detection: time of occurrence of last constituent ev | 1D              |
| 🗢 flash_id                               | product-unique lightning flash identifier                               | 1D              |
| 🗢 flash_lat                              | GLM L2+ Lightning Detection: flash centroid (mean constituent event I   | 1D              |
| 🗢 flash_lon                              | GLM L2+ Lightning Detection: flash centroid (mean constituent event I   | 1D              |
| flash_quality_flag                       | GLM L2+ Lightning Detection: flash data quality flags                   | 1D              |
| flash_time_offset_of_first_event         | GLM L2+ Lightning Detection: time of occurrence of first constituent e  | 1D              |
| flash_time_offset_of_last_event          | GLM L2+ Lightning Detection: time of occurrence of last constituent ev  | 1D              |
| flash_time_threshold                     | lightning flash maximum time difference among lightning events in a fl  | -               |
| goes_lat_lon_projection                  | GOES-R latitude / longitude projection                                  | _               |
| 🗢 group_area                             | GLM L2+ Lightning Detection: group area coverage (pixels containing     | 1D              |
| group_count                              | number of lightning groups in product                                   | _               |
| group_energy                             | GLM L2 + Lightning Detection: group radiant energy                      | 1D              |
| group frame time offset                  | GLM L2+ Lightning Detection: mean time of group's constituent events    | 1D              |
| 🗢 group_id                               | product-unique lightning group identifier                               | 1D              |
| 🗢 group_lat                              | GLM L2+ Lightning Detection: group centroid (mean constituent event     | 1D              |
| group_lon                                | GLM L2+ Lightning Detection: group centroid (mean constituent event     | 1D              |
| group_parent_flash_id                    | product-unique lightning flash identifier for one or more groups        | 1D              |
| group_quality_flag                       | GLM L2 + Lightning Detection: group data quality flags                  | 1D              |
| group time offset                        | GLM L2+ Lightning Detection: mean time of group's constituent events    | 1D              |

FIGURE 3.2: Example of a Panoply capture for GLM data of October 30th, 2019

LISTING 3.1: Example of user input for GLM directories

```
# Path where GLM's .nc files are located
107
    GLM_files_path =
108
        '/Users/jaimemorandominguez/Desktop/Final/GLM_archivos/nc'
     \hookrightarrow
109
    # Path where you want your daily ordered GLM's .nc files to be
110
        located
     \hookrightarrow
    GLM_ordered_dir = '/Users/jaimemorandominguez/Desktop/Final/GLM_arc
111
     → hivos/Dairy_dir'
112
    # Path where you want your daily ordered extracted GLM's .txt
113
     \hookrightarrow files to be located
    GLM_ordered_outputs = '/Users/jaimemorandominguez/Desktop/Final/GLM
114
     → _archivos/GLM_output'
```

directory inside GLM\_ordered\_dir for every different date with existing GLM data, and moving all files with data of that day inside. This new directory is named after the date it contains files of, in the form YearMonthDay.



#### 3.2.2 LINET's data handling

In this study, lightning location data from LINET are used as the ground-truth, and allowed to find the corresponding GLM and MMIA space-based detection as is discussed below. Tab.(3.1) shows the basic structure of the *.csv* file where LINET data is stored in a graphical way for an example of 5 rows.

| $date_{-}trunc$     | $ms\_linet$ | ka   | lat     | lon      | type      | id    | $group\_id$ |
|---------------------|-------------|------|---------|----------|-----------|-------|-------------|
| 2020-02-22 08:37:46 | 46.294      | 10.3 | 3.4295  | -75.1912 | Lightning | 6005  | 246325      |
| 2020-02-22 08:37:46 | 46.294      | 10.3 | 3.4295  | -75.1912 | Lightning | 5990  | 246325      |
| 2020-05-12 23:58:16 | 16.583      | 12.6 | 5.2109  | -71.8171 | Lightning | 5751  | 313357      |
| 2020-05-12 23:58:16 | 16.583      | 12.6 | 5.2109  | -71.8171 | Lightning | 5752  | 313357      |
| 2020-02-23 07:48:37 | 37.816      | 16.9 | 0.4223  | -74.8342 | Lightning | 6957  | 245729      |
| 2020-06-06 02:06:25 | 25.798      | 11.3 | 5.1966  | -72.5463 | Lightning | 25277 | 0           |
| 2020-06-06 02:06:54 | 53.235      | 39.4 | 10.2725 | -79.1761 | Lightning | 25713 | 0           |
|                     |             |      |         |          |           |       |             |

TABLE 3.1: Example of the structure of a given .csv file with LINET data

As it can be seen LINET data comes with important time (columns date\_trunc and ms\_linet) and location (columns lat and lon) information of detections and their corresponding ID for MMIA's .cdf files (column id). Ideally, rows are ordered with increasing date and increasing time, i.e. from January 1st at 00:00:00.000 to December 31st at 23:59:59.999, where every row corresponds to an existing LCFAprocessed MMIA trigger and belongs into a group marked by its group\_ID column. Group ID's can be shown, of course, in one or more rows, depending on time and location of its row trigger.

Once all GLM files have been ordered into different daily directories as explained in the previous section, function get\_MMIA\_dates.py (see appendix B.2) returns a list of dates with existing MMIA data as strings in the form YearMonth-Day. Function check\_existance.py (appendix B.3) returns the list matches with those dates with existing GLM and MMIA data (it already accounts for different number of dates for each instrument, only returning those that match).

With those date matches computed and with LINET's data already uploaded as a matrix, function get\_linet\_timing.py first gets those lines of the LINET



.csv that correspond to a date inside matches. Then separates those lines with an existing value for Group ID from those with a value of 0 (see Tab.(3.1) as an example). For those lines with an existing value a little subset is created storing in a list the starting time and end time (lowest and highest time value of the rows with same Group ID, respectively), minumum and maximum latitude and longitude (following the same philosophy as with time) and a list of all MMIA trigger ID's with that Group ID. In case multiple lines had the same Group ID incorrectly, the program creates new Group ID's for those lines that do not match that group<sup>2</sup>. This list of data defines a *snippet*. Its structure can be seen as:

### $[ \ \textbf{start\_time}, \ \textbf{end\_time}, \ \textbf{min\_lat}, \ \textbf{max\_lat}, \ \textbf{min\_lon}, \ \textbf{max\_lon}, \ [\textbf{ID}_1, \ \textbf{ID}_2, ...] ]$

A snippet is a fraction of a signal corresponding to one single group, which is the minimum set of data to compare between GLM and MMIA. All those snippets are stored into a variable called  $events^3$ , whose structure is widely used along the program. Tab.(3.2) shows a graphical visualization of the structure of this variable. As

| matches[0]   | matches[1]   | matches[2]   | ••• | matches[-1]  |
|--------------|--------------|--------------|-----|--------------|
| snippet 0    | snippet 0    | snippet 0    |     | snippet 0    |
| snippet $1$  | snippet $1$  | snippet $1$  |     | snippet $1$  |
| snippet $2$  | snippet $2$  | snippet $2$  |     | snippet 2    |
|              |              |              |     |              |
| snippet $-1$ | snippet $-1$ | snippet $-1$ |     | snippet $-1$ |

TABLE 3.2: Basic structure of events (linet\_times) variable

seen, events variable is just a list of daily lists of snippets. Every snippet information list is stored inside the day position in matches of its information (columns), and its position inside the day column is the snippet identification number. Of course every different day will have a different number of snippets according to available detection data for that day. This structure of storing snippets is maintained all over the program, as seen in upcoming sections.

 $<sup>^{2}</sup>$ In the case there were highly separated lines with the same Group ID, the program checks if those lines have a maximum separation of 20 lines. If not, new Group ID's are created for those last lines in order not to generate future problems.

<sup>&</sup>lt;sup>3</sup>Note that in events is the name of the variable *inside* the function get\_linet\_timing.py. In main\_TFG.py this variable is called linet\_times.



#### 3.2.3 MMIA's data handling

During the development of this study, all MMIA data has been given due to the need of credentials for their download from the server. This data comes in *.cdf* files and, similarly to GLM case, each file is defined by its detection date and hour, and more importantly, its trigger ID number. As in GLM case, the program asks for a directory with all *.cdf* files inside, a void directory for creating new directories and another void directory to store extracted data files. In section USER INPUT DATA of main\_TFG.py (appendix A) paths to those directories are asked in a similar way as in code snippet 3.1 in section 3.2.1.

Once all snippets have been identified and delimited using LINET's data, MMIA .cdf files are ordered in different directories according to the snippet where their ID number appears (see snippet structure in section 3.2.2 and function MMIA\_data \_ordering.py in appendix B.5). This new snippet directory is named after the day of the detections in the form YearMonthDay and the position of the snippet they contain (0 to n-1), while copied MMIA's .cdf files' names are changed to their ID number. It is important to note that, as MMIA data is ordered following LINET's information just for matches, only those MMIA .cdf files regarding a matching date with GLM will be ordered (and extracted, processed and compared), acting as a first filter for MMIA data.

As this GLM and MMIA ordering process is done just once per dataset and all new directories are already installed, boolean variables called pre\_ordered\_GLM and pre\_ordered\_MMIA let bypass this process once it has been done, accelerating the program and avoiding a new unnecessary ordering.

#### 3.3 GLM and MMIA data extraction and conditioning

This section explains how ordered MMIA's and GLM's data is extracted from given files into easily-working arrays and how those arrays are treated and conditioned to be cross-correlated to one another. It is important to note that both functions for



explicitly extracting data from the GLM's *.nc* and MMIA's *.cdf* files were given, and no major changes were made to them.

#### 3.3.1 MMIA data extraction and conditioning

Having all MMIA's .cdf files ordered by snippet in different directories, the program proceeds to extract their data using a given MatLab script that concatenates all .cdf files inside a directory and saves their important information as a variable. Function extract\_MMIA.py (see appendix B.6) hovers over all MMIA's snippet directories calling a MatLab engine for every snippet, executing the given .m script and returning only a .mat file per directory with just time and 777.4 photometer data vectors. This file is called after the directory of the snippet (i.e. 'YearMonthDay\_index.mat', being index the position of the snippet in that day column) and stored in the outputs directory of MMIA. Some snippets do not output a .mat file due to false triggers in the instrument for photometers usually caused by triggers in MMIA's 'CHU' cameras. When a camera starts a trigger using the first frame for storing its data, a decompensation in photometer's signal and time vectors occur causing an error while saving photometer data, so if the case was presented this step is simply avoided and no *.mat* file is outputted. The program prints the day and snippet index for those snippets with no data while writing the variable MMIA\_raw\_data, which in that position remains as type None.

A new function, unify\_MMIA\_data.py (appendix B.7) reads all those snippet data tables and writes them inside a similar structure as seen in Tab.(3.2) for LINET's data variable linet\_times. Tab.(3.3) shows how this data vectors are stored inside MMIA\_raw\_data. As seen, the variable is a list of daily lists of snippet vectors. For every daily position of the variable (i.e. MMIA\_raw\_data[0], MMIA\_raw\_data[1], ..., MMIA\_raw\_data[-1]) a list of snippets is stored. The order of daily positioning is the same as in linet\_times variable, following matches order, and the order of snippets is also the same. Snippet information is then given by MMIA\_raw\_data[i][j][:,0] represents the time vector of points of the snippet and



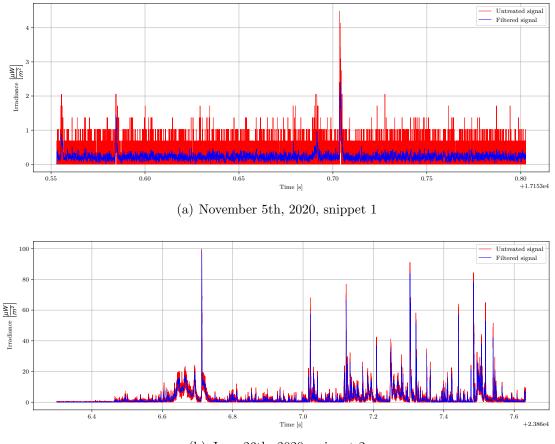
MMIA\_raw\_data[i][j][:,1] represents its signal vector. All this stored data coming

| Snippet day position | matches[0]  | matches[1]  | <br>matches[-1]   |
|----------------------|---|---|---|
| 0                    | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  |
| 1                    | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  | [time[0], signal[0]<br>time[1], signal[1]<br><br>time[-1], signal[-1]]  |
|                      |   |   |   |
| -1                   | $ \begin{array}{l} [\operatorname{time}[0],  \operatorname{signal}[0] \\ \operatorname{time}[1],  \operatorname{signal}[1] \\ \dots \\ \operatorname{time}[-1],  \operatorname{signal}[-1]] \end{array} $ | $ \begin{array}{l} [\operatorname{time}[0],  \operatorname{signal}[0] \\ \operatorname{time}[1],  \operatorname{signal}[1] \\ \dots \\ \operatorname{time}[-1],  \operatorname{signal}[-1]] \end{array} $ | $ \begin{array}{l} [\operatorname{time}[0],  \operatorname{signal}[0] \\ \operatorname{time}[1],  \operatorname{signal}[1] \\ \dots \\ \operatorname{time}[\text{-1}],  \operatorname{signal}[\text{-1}]] \end{array} $ |

TABLE 3.3: Representation of the MMIA\_raw\_data-type variable structure

from the .cdf files need to be treated in order to improve the signal quality and the correlation with GLM signal. Function condition\_MMIA\_data.py (appendix B.8) applies a filter to every snippet signal vector to reduce noise, as well as applies an absolute threshold of 1.75  $\frac{\mu W}{m^2}$  to bypass those snippets whose signal vector does not contain any important detection data, just noise. This noise snippets can be caused by triggers in other photometers or cameras of MMIA, while 777.4nm photometer does not detect any significant peak. All snippets are then input to function fit\_vector\_in\_MMIA\_timesteps.py to assure all time and signal vectors have a sample every exactly 0.00001s, as sometimes MMIA data comes with little time jumps or missing timesteps. Those unexisting samples are filled by a linear regression taking the last and next existing samples as reference points. New time and filtered signal pair of vectors per snippet are stored inside the variable MMIA\_filtered, which has the exactly same structure as the previous MMIA\_raw\_data structure shown in Tab.(3.3).

Fig.(3.3) shows two examples of unaltered signal vectors as extracted from the *.cdf* files and stored in MMIA\_raw\_data as well as the filtered signals. It is notorious how noise in signal is reduced and how important peaks are more easily detectable, easily seen in Fig.(3.3a). Of course, as signal detection values increase the noise



(b) June 26th, 2020, snippet 2

FIGURE 3.3: Examples of unaltered (red) and filtered (blue) MMIA 777.4nm photometer detection signals

contribution to the final vector is reduced, translating into a clearer curve. It is also important to note how peak values are also altered by their own prominence with respect to the noisy signal, as the filter crops part of their absolute magnitude. Comparing Fig.(3.3a) with Fig.(3.3b) it can be seen how peak magnitude is better conserved in those snippets with less noise (i.e. higher detection energies) and more prominent peaks, where the filtered signal fits the unaltered vector much better.

#### 3.3.2 GLM data extraction and conditioning

With all GLM's .nc files daily ordered in separate directories as explained in section 3.2.1, a similar extraction process as followed with MMIA's .cdf files is computed with GLM data. Function extract\_GLM.py (appendix B.9) uses a given Python script that concatenates all .nc files inside a given directory, extracting all usefull



| Second       | Lat.     | Lon.       | ID    | Lat.     | Lon.       | Radiance    |
|--------------|----------|------------|-------|----------|------------|-------------|
| of Day       | Event    | Event      | ID    | Flash    | Flash      | [J]         |
| 23866.493860 | 3.527283 | -71.354897 | 53555 | 3.701137 | -71.503380 | 1.40757e-15 |
| 23866.496149 | 3.527283 | -71.354897 | 53555 | 3.701137 | -71.503380 | 1.78805e-15 |
| 23866.498056 | 3.527283 | -71.354897 | 53555 | 3.701137 | -71.503380 | 1.38854e-15 |
|              |          |            |       |          |            |             |
| 23867.720686 | 3.600410 | -71.137550 | 53578 | 3.530246 | -71.177681 | 1.80707e-15 |

TABLE 3.4: Example of an output *.txt* for June 27th, 2020, snippet 2

data from those files that fit inside some restrictions and transcripting it into a .txt output for every date directory. The structure of this .txt file can be seen in Tab.(3.4) with an example. As every daily directory contains all GLM's .nc files for that day, restrictions in data extraction into the .txt file stand for a better delimitation in both time and space for every snippet. Data for those snippet restrictions is given by MMIA variable MMIA\_filtered as well as by LINET's data variable linet\_times. For doing so, main\_TFG.py program hovers over every snippet inside

MMIA\_filtered[i][j] = 
[ time[0], signal[0]
time[1], signal[1]
...
time[-1], signal[-1] ]



FIGURE 3.4: Fast visualization of GLM's data extraction restriction fonts for a generic snippet, day matches[i] position j

MMIA\_filtered variable, calling the given data extraction function for every MMIA non-None-type snippet position (positions that lacked of a *.mat* file as explained in the previous section). For those positions with data (of type numpy.ndarray), restrictions for GLM data extraction are the minimum and maximum latitudes and longitudes given by LINET's data, and the first and last time points of MMIA time vector for that particular snippet. A plus of time cropping\_margin (default 0.2s)



and angle angle\_margin (default 0.5°) are given to increase probability of matches due to MMIA time uncertainty. Fig.(3.4) summarizes this data acces for a generic snippet. Note again that as linet\_times and MMIA\_filtered variables follow the same structure (day order after matches and same snippet order inside each day list) every day position - snippet position combination [i][j] refers to the same snippet. As in the case of MMIA's .cdf files, every GLM snippet .txt file is named afer its day and snippet index inside that day (i.e. 'YearMonthDay\_index.txt'), as well as the ordering and data extraction processes can be bypassed once done by setting boolean variables pre\_ordered\_GLM and pre\_extracted\_GLM to 1.

Once GLM data has been extracted, function unify\_GLM\_data.py (appendix B.10) reads all snippet's *.txt*'s, sorts lines by ascending time and writes their information inside the respective snippet position in variable GLM\_raw\_data. This variable has exactly the same structure as MMIA\_raw\_data, seen in Tab.(3.3), but with more information per snippet (all the information from the *.txt*).

Having all GLM data uploaded to the program, its conditioning needs to be done. For every snippet the only important data is the time and radiance vectors, as these are the ones to compare with MMIA. Function condition\_GLM\_data.py (appendix B.11) first checks if the resulting *.txt* had information in it or if the information inside is too poor as to generate a vector (low number of different timesteps) by checking the snippet position in GLM\_raw\_data. If the information contained for that snippet is enough as to generate a vector, it integrates every 0.002s. This step allows for a smoother more reliable curve as well as creates a continuous time vector. Fig.(3.5) shows how the original time vector lacks of some timesteps, while the integrated dataset has a perfect time vector with a sample every 0.002s. The integrated snippet is stored inside variable GLM\_int\_data just before calling function fit\_vector\_in\_MMIA\_timesteps.py (see appendix B.12). As one of the functions of the program is to cross-correlate GLM and MMIA signals, and cross-correlation works on samples, this is an important step as fit\_vector\_in\_MMIA\_timesteps.py function expands the GLM snippet into a set of longer time and signal vectors in MMIA timesteps of 0.00001s. To do so, it places the existing GLM timesteps into



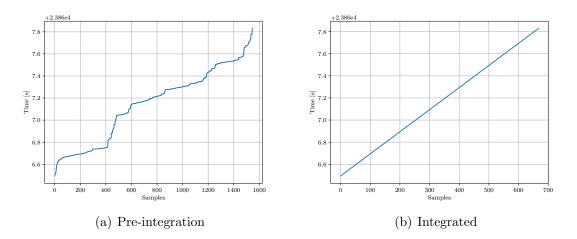


FIGURE 3.5: Pre-integration and integrated snippet time vector VS samples

the longer vectors and fills the new blank timesteps with a linear regression between the last and next timestep with data from the previous shorter vector. Of course, this function just adds resolution to the vector without changing its contents or varying the lineality in time. The resulting time and signal vectors are stored into variable GLM\_data. Both variables GLM\_int\_data and GLM\_int\_data have the exact same structure as shown for MMIA\_raw\_data in Tab.(3.3). Fig.(3.6) shows an example of an integrated and expanded GLM signal.

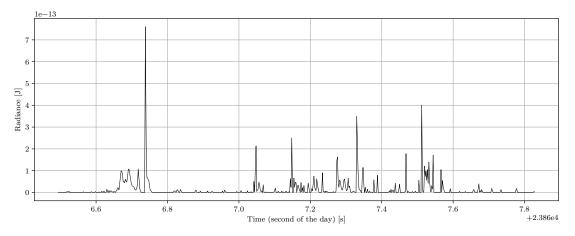


FIGURE 3.6: Example of an integrated GLM signal in MMIA timesteps for June 27th, 2020, snippet 2



#### 3.4 Cross-correlation of signals and peak detections

In this section it is explained how basic cross-correlation works and how it is implemented to correlate GLM and MMIA signals, as well as how those signals are treated to detect their peaks and then extract some statistics values as is discussed below.

#### 3.4.1 Cross-correlation basic functioning

Cross-correlation is a method for computing the similarity between two signals. In order to do so, two arrays of data Vector1 and Vector2 are inputted into the correlation function (in this study correlate() function is used from *Python* package scipy.signal), which slides Vector1 over Vector2 over every overlapping position from Vector1[0]-Vector2[-1] to Vector1[-1]-Vector2[0][10]. For every position a correlation factor is computed based on the similarity of the two signals (higher similarity translates into a higher correlation factor), outputting a new vector of correlation factors and length len(xcorr\_factors) = len(Vector1)+len(Vector2)-1<sup>4</sup>. See Fig.(3.7) for a visual respresentation of the process.

Knowing when those signals resemble the most, one can easily move one of them to make it fit perfectly between them. Taking as an example two simple signals of same length (for example, 10) and only one peak of value 1 in a different position for every signal (Fig.(3.8a) plots those signals):

Vector1 = [0,0,0,0,0,0,0,1,0,0]Vector2 = [0,0,0,1,0,0,0,0,0,0]

One can see how Vector1 is delayed by 4 samples with respect to Vector2, as it has its peak 4 samples after Vector2's peak. This is translated into a maximum in

<sup>&</sup>lt;sup>4</sup>The correlation function allows for 3 different modes for cross-correlation. Mode 'full' is the one used in the study and therefore explained, where the two vectors slide for every overlapping position. Mode 'same' returns a vector of  $len(xcorr_factors) = max(len(Vector1), len(Vector2))$  and centered with respect to the 'full' mode, and mode 'valid' computes cross-correlation just for those overlepping positions where every sample of one vector overlaps a sample of the other, returning a vector with length  $len(xcorr_factors) = max(len(Vector1), len(Vector2)) - min(len(Vector1), len(Vector2)) + 1 [8].$ 



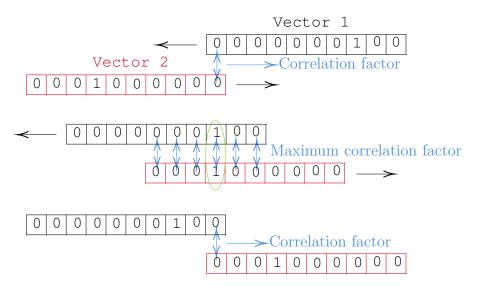


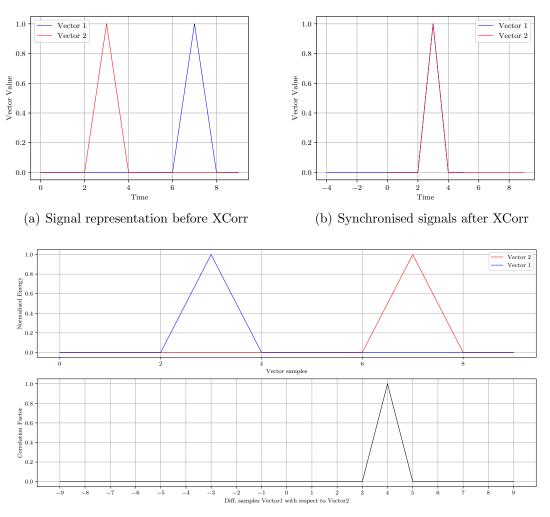
FIGURE 3.7: Fast visualization of cross-correlation dynamics for 'full' mode

xcorr\_factors's position where the two positions with 1 in Vector1 and Vector2 match, as the correlation factor in that position increases. Assigning a value to every xcorr\_factors position where the center position gets a 0, values on the left of the center position get single-spaced negative numbers and values on the right get single-spaced positive positions, this new vector x tells the delay samples of Vector1 with respect to Vector1. Following the previous example, xcorr\_factors and x are presented below. Note how the value inside the x vector in position where xcorr\_factors has its absolute maximum tells the delay. Fig.(3.8c) show both signals overlapped represented over samples, as well as xcorr\_factors over x. Finally, Fig.(3.8b) shows Vector1 and Vector2 after cross-correlation after using the delay value to align them properly.

#### 3.4.2 Cross-correlation of GLM and MMIA vectors

While synchronising GLM and MMIA signals some more steps must be done in order to get an overall good agreement. Although the cross-correlation process follows the same philosophy as in the previous section, the given example considered same-length vectors, an even value of len(Vector1) + len(Vector2) and most





(c) Signal representation by samples and correlation factor per overlapping position

FIGURE 3.8: Representation of example signals Vector1 (blue) and Vector2 (red) before and after being cross-correlated and synchronised

importantly, both vectors followed the same time vector. While analysing GLM and MMIA's signals, any length of those vectors can be input to the cross\_correlate\_ GLM\_MMIA.py function (see appendix B.14) and there is no need for GLM signal vector to be in the same time vector as MMIA's.

Firstly, cross\_correlate\_GLM\_MMIA.py function calls signal\_delay.py function (appendix B.13) to compute the real delay of the GLM signal with respect to MMIA's. The first step to compute this delay is exactly the same as in the example before, but considering even and odd values for len(Vector1) + len(Vector2), which would cause problems. Code snippet 3.2 shows how the x vector is generated



27

with the consideration in mind, deleting the even or odd problem. In this case, data1 corresponds to GLM signal vector, while data2 corresponds to MMIA signal Note that when len(Vector1) + len(Vector2) is an even number the vector.

LISTING 3.2: Generation of the Cross-Correlation x vector regarding even and odd values of len(Vector1) + len(Vector2)

```
len_x = len(data1) + len(data2) - 1
        x = np.empty(len_x)
28
29
        for i in range(len_x):
30
            if (len_x % 2) == 0: # Even number
31
                 x[i] = (i - (len_x/2))
32
            if (len_x % 2) != 0: # Odd number
33
                 x[i] = (i - (len_x/2 - 0.5))
34
```

resulting  $\mathbf{x}$  vector has an odd length and has the exactly same structure as in the example, with a 0 in the center position. If the length sum is odd, the  $\mathbf{x}$  vector has even length and 0 value falls on the first position of the second half of the vector:

$$x_{\text{odd length sum}} = [0,0,0,0,0,1,0,0,0,0]$$
  
$$x_{\text{even length sum}} = [0,0,0,0,0,1,0,0,0,0,0]$$

The next step is to account for different lengths of Vector1 and Vector2. Code snippet 3.3 shows how the delay value given by vector  $\mathbf{x}$  in the position where xcorr\_factors has its maximum value does not match the real delay when the signal vectors have different lengths. Note how a correction of 0.5 samples is made to delay when len(Vector1) + len(Vector2) is odd to account for the position in Once the even/odd length sum and different lengths problems are solved, the x. given delay represents the number of samples GLM signal is shifted with respect to MMIA signal if both vectors started from the same timestep. As this is not the case (every signal vector has its own time vector) another readjustment is needed. Code snippet 3.4 shows this correction, adjusting delay\_samples value by getting the difference in samples between same times in GLM and MMIA time vectors.

This value, real\_delay\_samples, is the delay in samples of MMIA signal with



LISTING 3.3: Variation of delay\_samples value considering different lengths of Vector1 and Vector2

```
# Delay samples accounting actual positioning due to
52
            different lengths:
        \hookrightarrow
        max_factor_pos = np.where(xcorr_factors ==
53
            max(xcorr_factors))[0][0]
54
        if ((len(data1)+len(data2)) \% 2 == 0): # len(x) is Odd
55
            delay_samples = x[max_factor_pos]+(len(data1)-len(data2))/2
56
57
        if ((len(data1)+len(data2)) \% 2 != 0): # len(x) is Even
58
            delay_samples = x[max_factor_pos]+(len(data1)-len(data2))/2
59
             → + 0.5
```

LISTING 3.4: Variation of delay\_samples value considering samples in between same times in GLM and MMIA time vectors

```
# Delay samples accounting actual positioning due to time:
61
        if data1[0,0] > data2[0,0]: # GLM vector starts later
62
             pos_MMIA_start_GLM = np.where(data2[:,0] <=</pre>
63
             \rightarrow data1[0,0])[0][-1]
             real_delay_samples = delay_samples + pos_MMIA_start_GLM
64
        elif data1[0,0] < data2[0,0]: # GLM vector starts earlier</pre>
65
             pos_GLM_start_MMIA = np.where(data1[:,0] <=</pre>
66
             \rightarrow data2[0,0])[0][-1]
             real_delay_samples = delay_samples - pos_GLM_start_MMIA
67
        else:
68
             real_delay_samples = delay_samples
69
```

respect to GLM signal accounting for different lengths of their signal vectors, different even/odd value for their length sum and considering different time vectors for each signal vector. A positive value of this variable means GLM signal is behind MMIA's (MMIA anticipated), while a negative value means GLM signal came before MMIA signal (MMIA delayed). After signal\_delay.py functions returns this value to cross\_correlate\_GLM\_MMIA.py, it just moves the MMIA signal by changing its time vector, adding 'real\_delay\_samples' times MMIA's period, 0.00001s. GLM time is taken as reference because of its better time resolution of 2ms. That

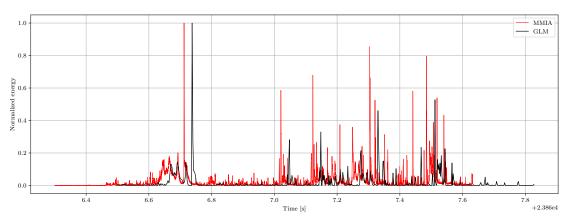


is the reason behind always moving MMIA signal instead of GLM's, regardless of the sign of the delay. Delay value per snippet is stored in variable **delays** as it is an important piece of data to know how time-shifted is GLM data with respect to MMIA's.

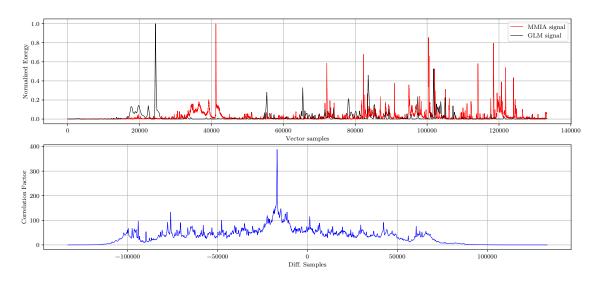
Fig.(3.9) shows an example of uncorrelated signals between GLM and MMIA signals ready to be synchronised. In Fig.(3.9a) both signals can be seen plotted with their respective time vector. Fig.(3.9b) shows those signals plotted by samples and the correlation factor for every position. Note that as those signals are plotted by samples before the delay corrections, time vector is ignored, and the delay seen is not the real one as neither is the difference in samples given by the position of the maximum correlation factor as explained before. Finally, Fig.(3.9c) shows the cross-correlated, synchronised GLM and MMIA signals over their time vectors. For this example, June 27th 2020, snippet 2, the delay is of 2587 samples (positive, MMIA anticipated as seen in the figure) or 0.02587s (very close to MMIA maximum time accuracy of 20ms).

#### 3.4.3 Detection of peaks in GLM and MMIA signals

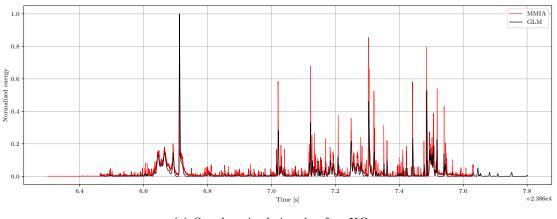
After cross-correlating GLM and MMIA signals both can be directly compared. Function get\_GLM\_MMIA\_peaks.py (appendix B.15) first trims correlated GLM and MMIA signals on every snippet to the same length and finds all prominent peaks on both signals, returning a list of indexes where peaks have been detected for every snippet, storing this data in variables GLM\_peaks and MMIA\_peaks, both following a similar structure as seen in previous snippet-structured variables (a list of snippet data inside a daily list following matches order). Fig.(3.10) shows an example of both GLM and MMIA signals over samples with their detected peaks marked.



(a) Signal representation before XCorr



(b) Signal representation by samples and correlation factor per overlapping position



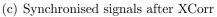


FIGURE 3.9: Representation cross correlation of GLM and MMIA signals for June 27th, 2020, snippet 2



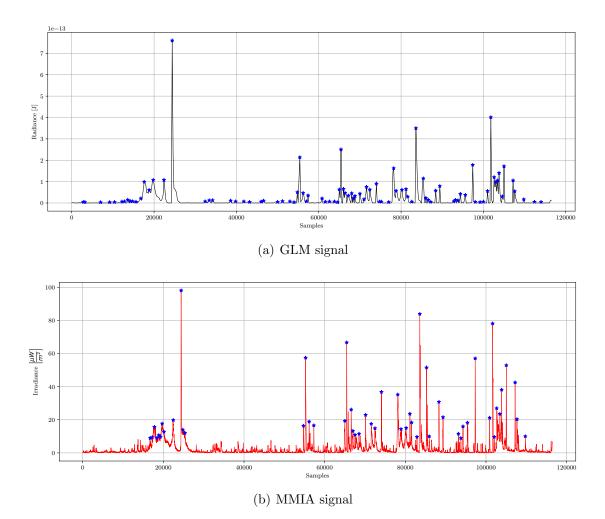


FIGURE 3.10: Example of GLM and MMIA detected peaks for June 27th, 2020, snippet 2

# RESULTS



### 4 Results

In this section final results are presented after all data management and treatment. Section 4.1 explains all results from the analysis, section 4.2 presents final conclusions while section 4.3 shows possible avenues for improvement for better and more consistent outputs, and finally, section 4.4 details the total budget for developing this study.

#### 4.1 Presentation of final results

This study covers the detection of lightning for 114 different dates in 2020. For every day, an average of 5.39 snippets have been found. A total of 1507 MMIA triggers were studied, obtaining 615 different snippets, out of which only 340 were suited for cross-correlation and peak comparison (55.28%). Snippets may be discarded if lack of MMIA or GLM data is found after extraction, or if the extracted data is too poor as to conform a snippet. This lack of data may be given by no GLM or MMIA datafiles for a particuar event, or by no detections in the time and space restrictions for a snippet (LINET captured a lightning while MMIA did not have Colombia in its FOV or simply GLM and/or MMIA didn't detect anything, for example). After comparing those snippets that could be analysed, some important statistics can be extracted. In the following sections delays between GLM and MMIA data are studied for a better comprehension of which instrument delays, how often and how much, as well as a better study of signal peaks is explained.

#### 4.1.1 Delays between GLM and MMIA

While cross-correlating signals the delay of MMIA with respect to GLM signal in samples for every snippet was stored in order to compare the results. The total average delay (a positive delay means that MMIA signal anticipated as shown in the simple example of Fig.(4.1)) is  $\overline{total\_delay}_{samples} = -1423.72$  samples or  $\overline{total\_delay}_s = -0.014$  seconds, meaning the average MMIA displacement among

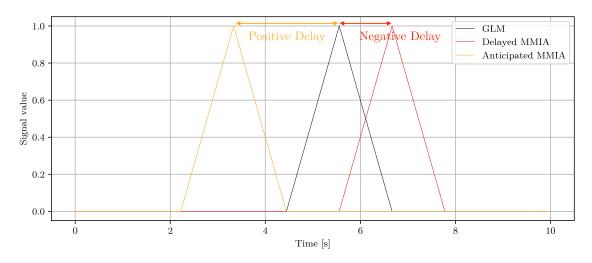
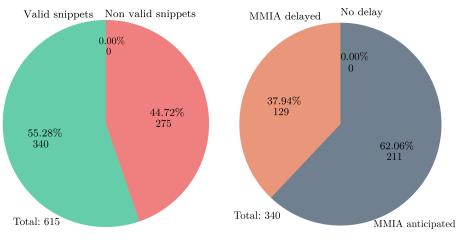


FIGURE 4.1: Simple example of delay sign convention

all 340 studied snippets. Although it may seem that MMIA uses to delay more often (negative value), 62.06% of the snippets show it actually anticipates more (211 of GLM over 129 for MMIA). This means that average MMIA delay time is (in absolute value) higher than anticipation time. Indeed, average MMIA anticipation is  $MMIA\_anticipation_{samples} = 11541.42$  samples (0.115 seconds) while average delay is  $\overline{MMIA\_delay}_{samples} = -22630.26$  samples (-0.226 seconds), with standard deviations being  $\sigma_{GLM} = 14644.11$  and  $\sigma_{MMIA} = 18760.87$  samples (0.146 and 0.188 seconds), respectively. Fig.(4.2) shows this data for easier understanding. It is also interesting to analyse the relationship between the energy of the detections and the probability of being delayed. Fig.(4.3) shows separated scattered data for MMIA-anticipated snippets and MMIA-delayed snippets, where all snippet delays are displayed as a function of the average radiance of the snippet and its standard deviation. As clearly seen, most of the snippets fall into a low-energy, low-delay zone. It can also be seen how as average energy increase, or more peaks are present (higher standard deviation), the delay in the signal decreases. In those cases where MMIA was anticipated a bigger concentration of snippets is found in the low-low zone, but the gradient of delay vs average energy and delay vs average energy is higher than in MMIA's delays (it is easier to have a low delay if the snippet has a high average energy or many peaks on MMIA anticipated-snippets than in MMIA delayed-snippets).



(a) Valid snippets over total snippets (b)

(b) Delayed instrument per snippet

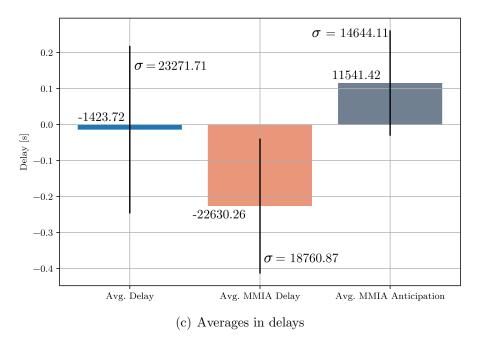
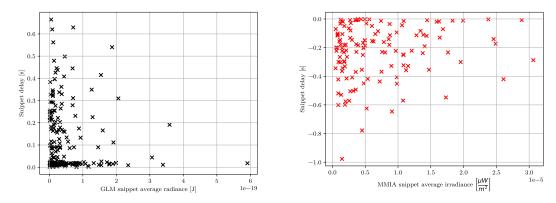


FIGURE 4.2: Basic delay statistics for GLM and MMIA signals

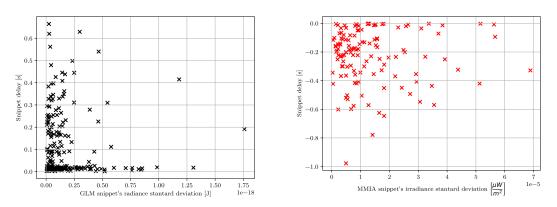
#### 4.1.2 Peak correspondence

Once GLM and MMIA have correlated their signals their peak values and frequency distribution are calculated.

After computing all 114 dates, GLM has, in average, 10.84 peaks per snippet, while MMIA presents an average of 24.29 peaks per snippet (MMIA detects more peaks than GLM of about 44.63%). A function compares both peak vectors in order to determine what peaks on one instrument were detected by the other instrument



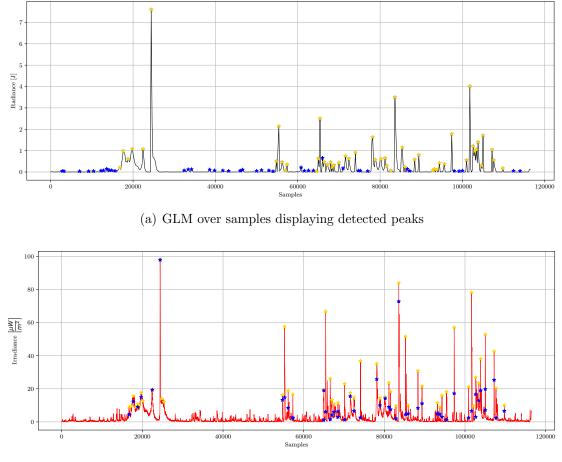
(a) MMIA anticipations VS Average Radiance (b) MMIA delays VS Average Radiance distridistribution bution



(c) MMIA anticipations VS Std. Deviation dis- (d) MMIA delays VS Std. Deviation distribution tribution

FIGURE 4.3: Distributions of MMIA delays separated for positive and negative, based on average energy of the whole signal and its standard deviation

and vice versa. On GLM, from every average of 10.84 peaks per snippet, 7.6 peaks are detected by MMIA. That means that 70.42% of GLM maxima are detected by MMIA. On the other way around, numbers flip. For every MMIA average of 24.29 peaks per snippet, only those 7.6 are detected (of course, average detected peaks on both instruments simultaneously does not vary). Exactly, just 36.97% of MMIA's peaks are detected by GLM. Fig.(4.4) shows the example snippet displaying common peaks. The function that compares peaks starts by hovering over all GLM detections and trying to find the nearest MMIA detection in vector position, giving preference to higher peaks.



1e - 13

(b) MMIA over samples displaying detected peaks

FIGURE 4.4: Example of GLM and MMIA signals displaying their detected peaks in blue, and with common peaks in yellow



#### 4.2 Conclusions

To conclude this report, some considerations have to be taken. One of the main purposes of this project was to develop useful tools for determine the detection sensitivity of the Geostationary Lightning Mapper with respect to the Modular Multispectral Imaging Array. Another important step this study has made is to reduce MMIA's uncertainty from about 20ms to 2ms thanks to LINET's and GLM's data, of much more timing precision.

In the development of those tools several difficulties have been found. First of all, the main architecture of the analysis was designed, showing what functions where needed for the whole analysis. Snippets were defined and trimmed to evaluate their single-piece-of-analysis capability.

As seen in previous sections, GLM instrument lacks of sensitivity when compared to MMIA's photometer of much more resolution. When given data contains signals of low energy, it has been proved how MMIA data is more likely to come more delayed than in those high energy signals with prominent peaks. It also has been shown how GLM's low sensitivity concatenates a lower peak detection number than MMIA.

This study also opens an avenue to consolidate a program to easily automatise the process of MMIA normalisation using GLM data and its comparison, as it is a widely frequented process during thunderstorm and lightning activity studies.



#### 4.3 Future Continuation

Although this study has been performed in a period of 7 months and great attention to detail has been put into every function, there are some key points in the way to continue with the study. As the main weight of the process has been the programming process of the resolution, future continuation on this work should follow the same line.

All the process is maintained by the importance of the snippets, as they are the minimum datasets for comparison. As so, their definition is key to a good crosscorrelation and posterior analysis, as well as their initial treatment. One of the main points to retouch in the program is MMIA filter's parameters, when the initial non-processed data is smoothered (see section 3.3.1). The optimal filter should straighten the signal as much as possible without losing any shape, specially when finding prominent peaks.

Focusing on snippets, their characterization must be perfect in order to get good results in the following steps as mentioned before. Not all snippets are valid for analysis, as many of them are noise signals in MMIA, for example, or extremely short data vectors in GLM. Although the program detects most of those problematic snippets and takes them out of next steps, a good way of continuing the study would be to better characterize those non-valid snippets and even looking for a better way to understand their procedence. If not deleted as soon as possible, final statistics may be highly biased by those non-constructive snippets.

Another important step for the analysis of the signals is done in cross-correlation. So far the presented function cross-correlates and synchronises any two sets of data regardless of their length or time vectors, as explained in section 3.4.2, successfully. In order to mantain a good cross-correlation, good quality snippets should be used, returning to the main previous idea. If not, extreme values of delays may be seen in further statistics that can lead to great confusion regarding GLM timing.

One of the most improvable points of all the program is executed as peak detection. Due to the noisy MMIA signal, which is not constant for all snippets, it is



difficult to detect a real peak from a noise peak. This problem leads to snippets with large amounts of 'ghost' peaks than can alter future statistics reducing drastically GLM detection rate, when those peaks are just fake. So far a threshold is computed as the 90th percentile of all snippet signal vector, assuming most of the samples are located as noise in the lowest energy levels. A retouch on peak detection functions' parameters would probably allow for more accurate peak detections, improving results drastically. It would be even useful to cross-correlate the peak position vectors to compare how many peaks were detected by both instruments even better.

Finally, although some other considerations such as output data quantity, code cleanup and reestructure or optimization for faster code using less computer resources could be done, there is a key point during data ordering processes, explained in section 3.2. All file ordering is done by using the os Python library, which allows for direct command line orders using the function os.system(). This function takes as a input a command line order as a string, and just executes it on a terminal. As all the study has been developed using a Unix-like OS, those commands are witten in a Unix-like command interpreter (namely *zsh*) that are not directly compatible with Microsoft's *cmd* interpreter. This means that the initial steps of file ordering should be run by a Unix-like OS, opening a new future work of translating those commands to be used in Windows NT OS. After the ordering steps (that can be bypassed with variables pre\_ordered\_GLM and pre\_ordered\_MMIA), the only uso for the os library is to list directories entries, which based on Python documentation, just calls the native OS directory iteration system [9].



#### 4.4 Budget

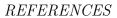
In this section a complete breakdown of the study's budget is given, detailing every major cost with a brief description.

As this study does not require any type of material or support apart from the computer, the major cost is given by the salary of a junior engineer, estimated as  $15 \in /h$ . Tab.(4.1) shows a simple breakdown of the main work packages developed and the time dedicated to each one.

| Work Package          | Dedication [h] | Unitary Cost $[\in/h]$ | Cost $[\in]$ |
|-----------------------|----------------|------------------------|--------------|
| Documentation         | 20             | 15                     | 600          |
| Code design           | 50             | 15                     | 750          |
| Code development      | 200            | 15                     | 3000         |
| Debugging and testing | 100            | 15                     | 1500         |
| Code documentation    | 50             | 15                     | 750          |
| Runtime               | 20             | 0.015                  | 0.45         |
|                       |                | Total:                 | 6600.5€      |

TABLE 4.1: Main work packages and cost associated

Where documentation accounts for initial approach with the area, article consultation and coding package documentation. Design of the code accounts for the structure of the program, how steps are placed in order to achieve the final desired output. Code development and debbuging and testing packages are the longest by far as this study required a great dedication to developing the program. Runtime package is the cost derived from having a computer constantly plugged-in to compute operations while alone (some processes took long running times before some optimization). Finally, code documentation accounts for the development of this report as an introductory manual to the code as well as code presentation.



## References

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- [8] NUMPY, numpy.convolve documentation page.
- [9] PYTHON, Python os library Documentation.
- [10] SCIPY, scipy.signal.correlate Documentation.
- [11] UNIDATA, Network Common Data Form (NetCDF).

# APPENDICES



## A Code of the main\_TFG.py script

LISTING A.1: Full code for the main\_TFG.py script of the program

```
1 ....
2 README
3
4 This script extracts data from GLM .nc files and MMIA .cdf files
     and prepares
5 it for comparison to extract detection sensitivity of GLM.
7 To do so, a time snippet for every day with GLM and MMIA files is
     extracted
s to use only instead of the whole signal vectors using LINET's data
     from a
9 .csv file, where date, time, latitude, longitude, trigger ID and
    group_ID
10 are taken.
11
12 Every file from a given path 'GLM_files_path' (for GLM files) and
13 'MMIA_files_path' (for MMIA files) is classified in a different
     directory
14 according to the snippet (given by LINET's data) if MMIA, or date (
     day) of
15 the event if GLM, and stored in given paths 'GLM_ordered_dir' and
16 'MMIA_Ordered_dir' (for GLM and MMIA, respectively).
17
                               CAUTION!
18
19 This step needs to be run by a UNIX-like OS, as explicit terminal
     commands
20 are given using Python's library "os".
21
22 As the files have been ordered, data extraction is done using a
     Python
23 script for GLM and a MatLab script for MMIA, both given by Jesús Ló
     pez, only
24 for the matching dates between the existing GLM and MMIA files.
     This main
25 script initialises the MatLab engine by its own. This extracted
     data is
26 stored in different snippet .txt files for GLM and .mat files for
     MMIA at
27 given paths 'GLM_ordered_outputs' and 'MMIA_ordered_outputs'.
28
                               IMPORTANT
29
_{\rm 30} Extracting data from GLM's .nc files and MMIA's .cdf files takes a
     lot of time
31 and is a one-time step, specially for MMIA .cdf files where a
    MatLab engine
32 has to be started for every snippet. Because of that, one can set
     special
33 parameters 'pre_extracted_GLM' and 'pre_extracted_MMIA' to '0' to
  make that
```



```
34 extraction once, and then set those parameters to '1' to bypass the
     extracting
_{35} operations as the data has already been stored in ' \ldots
     _ordered_outputs'.
36 Again, this process needs to be run by a UNIX-like OS.
37
38 This data is then uploaded to the Python workspace and treated to
    allow
39 a comparison analysis using cross-correlation.
40
41 For GLM data, an integration of radiance is made every 0.002s (GLM
    sample rate)
42 and time and signal vectors are accommodated to MMIA sample rate of
      0.0001s.
43
44 For MMIA data, afilter is applied to the 777.4nm photometer data
     vector in
45 order to reduce signal noise, and again accommodated to MMIA sample
     rate
46 in case any time jump was there.
47
48 Both types of data are then normalised and cross-correlated to get
     the time
49 shift between them, aligned and compared by counting peaks in their
      signals.
50
51 All functions are imported from 'TFG_library.py' to make this
    script simple
52 and clean.
53
54
55 @ Jaime Francisco Morán Domínguez, 2021
56
57 """
58
59 import TFG_library as TFG
60 import pandas as pd
61 import numpy as np
62 import matplotlib.pyplot as plt
63 import pickle
64
65 # Just for plot presentation in LaTeX Style (slows the program)
66 plt.rc('font', **{'family': 'serif', 'serif': ['latin modern roman'
     ]})
67
68
69
70 111
USER INPUT DATA
                                                  ##
72 ##
74 111
75
76 ### GENERAL ###
77
78 # Boolean variable for pre-ordered files
```



```
79 \text{ pre_ordered_GLM} = 1
80 \text{ pre_ordered_MMIA} = 1
81
82 # Boolean variable for pre-extracted files
83 pre_extracted_GLM = 1
84 pre_extracted_MMIA = 1
85
86 # Boolean variable for generating plots
87 \text{ show_plots} = 0
88
89
90 ### LINET ###
91
92 # CAUTION !! Make sure to write the 'r' before the path string
       # Example: LINET_path = r'path_to_csv_file'
93
94 LINET_path = r'/Users/jaimemorandominguez/Desktop/Final/
      MMIA_match_pos_LINET_2020.csv'
95
96 # Time in seconds to analyze GLM and MMIA before and after LINET's
      time snippet
      # Recommended 0.2
97
98 cropping_margin = 0.15
99
100 # Plus of angle in latitude and longitude to snip GLM data (with
      respect to LINET data)
       # Recommended 0.5
101
102 angle_margin = 0.5
103
104
105 ### GLM ###
106
107 # Path where GLM's .nc files are located
108 GLM_files_path = '/Users/jaimemorandominguez/Desktop/Final/
      GLM_archivos/nc'
109
110 # Path where you want your daily ordered GLM's .nc files to be
      located
111 GLM_ordered_dir = '/Users/jaimemorandominguez/Desktop/Final/
      GLM_archivos/Dairy_dir'
112
113 # Path where you want your daily ordered extracted GLM's .txt files
       to be located
114 GLM_ordered_outputs = '/Users/jaimemorandominguez/Desktop/Final/
      GLM_archivos/GLM_output'
115
116
117 ### MMIA ###
118
119 # Path where MMIA's .cdf files are located
120 MMIA_files_path = '/Users/jaimemorandominguez/Desktop/Final/
      MMIA_archivos/cdf'
121
122 # Path where you want your snippet ordered MMIA's .cdf files to be
      located
123 MMIA_ordered_dir = '/Users/jaimemorandominguez/Desktop/Final/
     MMIA_archivos/MMIA_dairy'
```



124

```
125 # Path where you want your snippet ordered extracted MMIA's .mat
     files to be located
126 MMIA_ordered_outputs = '/Users/jaimemorandominguez/Desktop/Final/
     MMIA_archivos/MMIA_output'
127
128
  . . .
129
131 ##
              END OF USER INPUT DATA
                                              ##
1.1.1
133
134
135
136
  137
138
139 # Uploading Linet's data from .csv
140
141 print(' ')
142 print("Uploading Linet's data...")
143 linet_data = pd.read_csv (LINET_path)
144 linet_data = linet_data.to_numpy()
145 print('Done!')
146 print(' ')
147
148
150
  if pre_ordered_GLM == 0:
151
152
      # Ordering GLM .nc files in dairy directories
153
     TFG.GLM_data_ordering(GLM_files_path, GLM_ordered_dir)
154
155
  else:
156
     print('All GLM .nc files are already daily ordered')
157
      print(' ')
158
159
160
162
163 MMIA_dates= TFG.get_MMIA_dates(MMIA_files_path)
164
165 # Searching for dates that both GLM and MMIA have data from
166 matches = TFG.check_existance(GLM_ordered_dir, MMIA_dates)
167
168
169 ########### EXTRACTING SNIPPET DATA FROM LINET'S DATAFRAME
     ###########
170
171 # Getting times from LINET data
172 [linet_times, indx] = TFG.get_LINET_timing(linet_data, matches)
173 del linet_data
174
175 ###### MMIA'S DATA ORDERING, EXTRACTION, UPLOAD AND CONDITIONING
     ######
```



176

```
177 # Ordering MMIA data files into daily snippet folders
178 if pre_ordered_MMIA == 0:
179
       TFG.MMIA_data_ordering(MMIA_files_path, MMIA_ordered_dir,
180
      linet_times, matches)
  else:
181
       print('All MMIA .cdf files are already daily ordered')
182
       print(' ')
183
184
185 # Extracting MMIA data into dairy .mat files
186
       # CAUTION! Only matching dates with GLM are being extracted
  if pre_extracted_MMIA == 0:
187
188
       TFG.extract_MMIA(MMIA_files_path, MMIA_ordered_dir,
189
      MMIA_ordered_outputs, matches, linet_times)
190
  else:
191
       print('All data from MMIA .cdf files has already been extracted
192
      1)
       print(' ')
193
194
195 # Unifying all data in a structure of lists
196 MMIA_raw_data = TFG.unify_MMIA_data(MMIA_ordered_outputs,
      linet_times, matches, show_plots)
197
198 # Conditioning MMIA data for further analysis
199 MMIA_filtered = TFG.condition_MMIA_data(MMIA_raw_data, matches,
      show_plots)
200
201
  del MMIA_raw_data
202
203
204 ########### GLM'S DATA EXTRACTION, UPLOAD AND CONDITIONING
      ###########
205
206 # Extracting GLM data into dairy .txt files
       # CAUTION! Only snippets with MMIA data are being extracted
207
  if pre_extracted_GLM == 0:
208
209
       TFG.extract_GLM(GLM_ordered_dir, GLM_ordered_outputs,
210
      linet_times, matches, MMIA_filtered, angle_margin,
      cropping_margin)
211
212 else:
       print('All data from GLM .nc files has already been extracted')
213
       print(' ')
214
215
216 del linet_times
217
218 # Unifying all data in a structure of matrices
219 GLM_raw_data = TFG.unify_GLM_data(GLM_ordered_outputs,
      MMIA_filtered, matches, show_plots)
220
221
222 # Conditioning GLM data for further analysis
```



```
223 GLM_data = TFG.condition_GLM_data(GLM_raw_data, matches, show_plots
      )
  del GLM_raw_data
224
225
226
228
229 # Normalizing GLM data to cross-correlate with MMIA data
230 print('Normalizing GLM data...')
231 GLM_norm = [None] * len(GLM_data)
  for i in range(len(GLM_data)):
232
233
       snip = [None] * len(GLM_data[i])
       GLM_norm[i] = snip
234
235
   for i in range(len(GLM_data)):
236
       for j in range(len(GLM_data[i])):
237
           if type(GLM_data[i][j]) == np.ndarray:
238
               snippet = np.zeros((len(GLM_data[i][j]),2))
239
               GLM_norm[i][j] = snippet
240
               GLM_norm[i][j][:,0] = GLM_data[i][j][:,0]
241
               GLM_norm[i][j][:,1] = TFG.normalize(GLM_data[i][j
242
      ][:,1])
243 print('Done!')
244 print(' ')
245
246 # Normalizing MMIA data to cross-correlate with GLM data
247 print ('Normalizing MMIA data...')
248 MMIA_norm = [None] * len(MMIA_filtered)
  for i in range(len(MMIA_filtered)):
249
       snip = [None] * len(MMIA_filtered[i])
250
251
       MMIA_norm[i] = snip
252
  for i in range(len(MMIA_filtered)):
253
       for j in range(len(MMIA_filtered[i])):
254
           if type(MMIA_filtered[i][j]) == np.ndarray:
255
               snippet = np.zeros((len(MMIA_filtered[i][j]),2))
256
               MMIA_norm[i][j] = snippet
257
               MMIA_norm[i][j][:,0] = MMIA_filtered[i][j][:,0]
258
               MMIA_norm[i][j][:,1] = TFG.normalize(MMIA_filtered[i][j
259
      ][:,1])
260 print('Done!')
261 print(' ')
262
263 # Cross-correlating snippets
_{264} show_plots = 1
265 [GLM_xcorr, MMIA_xcorr, delays] = TFG.cross_correlate_GLM_MMIA(
      GLM_data, MMIA_filtered, GLM_norm, MMIA_norm, matches,
      show_plots)
266 del GLM_data
267 del MMIA_filtered
268
269 # Saving cross-correlated data
270 f = open('xcorr_data.pckl', 'wb')
271 pickle.dump([GLM_xcorr,MMIA_xcorr, GLM_norm, MMIA_norm, matches,
      delays], f)
272 f.close()
```



# B Code for important functions in order of appearance in main\_TFG.py

#### B.1 Function GLM\_data\_ordering.py

LISTING B.1: Full code for GLM\_data\_ordering.py function

```
1 def GLM_data_ordering(read_path, dir_path):
      1.1.1
2
      This function gets a directory where unordered .nc files are
3
     located
      and gets them daily ordered in diferent directories inside
4
     dir_path.
      Those new directories are named by year-month-day.
5
6
      Parameters
7
      _____
8
      read_path : string
9
          Path of the directory where the unordered .nc files are
10
     located.
      dir_path : string
11
          Path to the directory where the new daily-ordered
12
     directories with
          ordered separated .nc files will be located.
13
14
      Returns
15
16
      None. Daily-ordered GLM files in new directories.
17
      1.1.1
18
19
      print('Ordering GLM data into separate dates...')
20
      with os.scandir(read_path) as files:
21
          files = [file.name for file in files if file.is_file() and
22
     file.name.endswith('.nc')]
      if len(files)==0:
23
          print('Error: No GLM .nc files found to process!')
24
25
      for i in range(len(files)):
26
          date = datetime.datetime.strptime(files[i][44:48] + "-" +
27
     files[i][48:51], "%Y-%j").strftime("%Y%m%d")
          if i==0:
                           # First file does not have any folder yet
28
               os.system('mkdir ' + dir_path+'/'+date)
29
              os.system('cp ' + read_path+'/'+files[0] + ' ' +
30
     dir_path+'/'+date)
                           # All the other files
          else:
31
               subfolders = [f.path for f in os.scandir(dir_path) if f
32
     .is_dir()]
              if subfolders.count(date)==0: # If there is not a
33
     folder with that date
                   os.system('mkdir '+dir_path+'/'+date)
34
```



| 35 | os.system('cp ' + read_path+'/'+files[i] + ' ' + |
|----|--|
|    | dir_path+'/'+date)                               |
| 36 | else: # If there already exists a folder with    |
|    | that date  |
| 37 | os.system('cp ' + read_path+'/'+files[i] + ' ' + |
|    | dir_path+'/'+date)                               |
| 38 | <pre>print('Done')</pre>                         |
| 39 | <pre>print(' ')</pre>                            |

#### B.2 Function get\_MMIA\_dates.py

LISTING B.2: Full code for get\_MMIA\_dates.py function

```
1 def get_MMIA_dates(read_path):
       1.1
2
      This function just hovers over MMIA .cdf files to extract a
3
     list with
      all existing dates with MMIA data.
4
\mathbf{5}
      Parameters
6
7
      read_path : string
8
          Path to the directory where all MMIA .cdf files are stored.
9
10
      Returns
11
12
      MMIA_dates : list
13
14
          List of strings with all dates with existing MMIA data, in
     the form
          YearMonthDay.
15
      1.1.1
16
17
      print('Getting the list of MMIA dates with existing data..')
18
19
      with os.scandir(read_path) as files:
20
          files = [file.name for file in files if file.is_file() and
21
     file.name.endswith('.cdf')]
      if len(files) == 0:
22
           print('Error: No MMIA .cdf files found to process!')
23
24
      MMIA_dates = []
25
26
      for i in range(len(files)):
27
           date = files[i][50:54]+files[i][55:57]+files[i][58:60]
28
29
          if i == 0:
                                # First file does not have any existing
30
      date
               MMIA_dates.append(date)
31
                                # All the other files
           else:
32
               if MMIA_dates.count(date) == 0: # If there is no
33
     register of that date
                   MMIA_dates.append(date)
34
35
```

36 print('Done')
37 print('')
38
39 return MMIA\_dates

#### B.3 Function check\_existance.py

LISTING B.3: Full code for check\_existance.py function

```
1 def check_existance(GLM_ordered_dir, MMIA_dates):
       1.1.1
2
      This function determines which dates are present among both GLM
3
       and MMIA
      sets of data. It returns a vector with the matching dates as
4
     strings in
      the form YearMonthDay.
5
6
      Parameters
7
8
      GLM_ordered_dir : str
9
           Path to the directory where daily ordered GLM .nc files are
10
      located.
      MMIA_dates : array
11
           Vector containing all the dates in the MMIA's set of data.
12
13
      Returns
14
       _ _ _ _ _ _ _
15
16
      matches : array
          Vector containing just the dates that are present in GLM
17
     and MMIA
          sets of data.
18
       1.1.1
19
20
      print('Checking for matching dates between GLM and MMIA data...
21
      •)
      print(' ')
22
23
      GLM_dates = os.listdir(GLM_ordered_dir)
24
25
      if GLM_dates.count('.DS_Store')!=0:
26
           GLM_dates.remove('.DS_Store')
27
28
      GLM_len = len(GLM_dates)
29
      MMIA_len = len(MMIA_dates)
30
      matches = []
31
32
      if GLM_len <= MMIA_len:</pre>
33
           for i in range(GLM_len):
34
               if GLM_dates[i] in MMIA_dates:
35
                    matches.append(GLM_dates[i])
36
37
      elif MMIA_len < GLM_len:</pre>
38
           for i in range(MMIA_len):
39
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.4 Function get\_linet\_timing.py

if MMIA\_dates[i] in GLM\_dates: 40 matches.append(MMIA\_dates[i])  $^{41}$ 4243if len(matches) == GLM\_len and len(matches) == MMIA\_len: print('Both GLM and MMIA data correspond to same days. All 44 data will be analyzed') elif len(matches) < GLM\_len and len(matches) < MMIA\_len:</pre> 45print('There is/are just %d matching date/s, which will be 46analyzed. Both GLM and MMIA daily ordered folders contain more dates that do not match each other' % len(matches)) elif len(matches) == GLM\_len and len(matches) < MMIA\_len:</pre> 47print('All GLM dates correspond to MMIA dates and will be 48analyzed. The folder where MMIA daily ordered folders are located contais %d more date/s than GLMs' % (MMIA\_len-len( matches))) elif len(matches) == MMIA\_len and len(matches) < GLM\_len:</pre> 49print('All MMIA dates correspond to GLM dates and will be 50analyzed. The folder where GLM daily ordered folders are located contais %d more date/s than MMIAs' % (GLM\_len-len(matches))) print(' ') 5152return matches 53

#### B.4 Function get\_linet\_timing.py

LISTING B.4: Full code for get\_linet\_timing.py function

| 1  | <pre>def get_LINET_timing(linet_data, matches):</pre>          |
|----|--|
| 2  |  |
| 3  | This function gets the data from linet_data and generates a    |
|    | list of lists  |
| 4  | where information of different events is stored only for dates |
|    | with   |
| 5  | existing GLM and MMIA data. Date order is the same as in       |
|    | matches.   |
| 6  |  |
| 7  | Parameters   |
| 8  |  |
| 9  | linet_data : matrix  |
| 10 | Linet's data from the .csv file ordered in a numpy array.      |
| 11 | matches : list   |
| 12 | List of dates with existing GLM and MMIA data files.           |
| 13 |  |
| 14 | Returns  |
| 15 |  |
| 16 | events : list  |
| 17 | List of lists containing different events for every            |
|    | existing GLM   |
| 18 | and MMIA data. For every day layer, the structure is:          |
| 19 | Events structure (for every day layer):                        |
| 20 | 1st column: Event starting time                                |
| 21 | 2nd column: Event ending time                                  |
| 22 | 3rd column: Event min latitude                                 |

В CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.4 Function get\_linet\_timing.py

23

```
4th column: Event max latitude
              5th column: Event min longitude
24
25
              6th column: Event max longitude
              7th column: List of MMIA trigger ID's in this snippet
26
      1.1.1
27
28
      print('Getting time snippets from LINET data...')
29
30
      # Getting only dates with existing GLM and MMIA data
31
      linet_subset = []
32
      # linet_subset structure:
33
          #
              1st column: Event date
34
          #
               2nd column: Event hour
35
          #
              3rd column: Event latitude
36
          #
              4th column: Event longitude
37
          #
              5th column: Event ID
38
              6th column: Group ID
          #
39
40
      for i in range(len(linet_data)):
41
          date = datetime.datetime.strptime(linet_data[i,0][0:10], "%
42
     Y-%m-%d").strftime("%Y%m%d")
          if date in matches:
43
               hour = float(linet_data[i,0][11:13])*3600 + float(
44
     linet_data[i,0][14:16])*60 + linet_data[i,1]
               linet_subset.append([date, hour, linet_data[i,3],
45
     linet_data[i,4], linet_data[i,6], linet_data[i,7]])
      if len(linet_subset) == 0:
46
          print('Your .csv file does not contain any data of the
47
     dates of existing GLM and MMIA data files')
      else:
48
          # Extracting times from linet data, generation of days:
49
          events = []
50
          for i in range(len(matches)):
51
52
               events.append([])
          # events structure (per day layer):
53
               #
                   1st column: Event min hour
54
               #
                   2nd column: Event max hour
55
               #
                   3rd column: Event min latitude
56
                   4th column: Event max latitude
               #
57
                   5th column: Event min longitude
               #
58
                   6th column: Event max longitude
               #
59
                   7th column: Event MMIA ID's
               #
60
61
          # Extracting data from those lines with group ID != 0
62
          linet_subset_gID = []
63
          gIDs = []
64
          no_gID = []
65
          for i in range(len(linet_subset)):
66
               if linet_subset[i][5] != 0: # If group ID is not 0 (no
67
     data)
                   linet_subset_gID.append(linet_subset[i])
68
                   if i == 0:
69
                       gIDs.append(linet_subset[i][5])
70
                   else:
71
                       if gIDs.count(linet_subset[i][5]) == 0:
72
                            gIDs.append(linet_subset[i][5])
73
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.4 Function get\_linet\_timing.py

else: 74 75no\_gID.append(i) 76# Removing all lines with group\_ID from linet\_subset 77 linet\_subset = [linet\_subset[index] for index in no\_gID] 7879 80 # Checking for wrongly repeated Group ID's 81 indx = []82 for i in range(len(gIDs)): 83 indx.append([]) 84 85 for i in range(len(indx)): # Creating lists of indexes with 86 same GroupID for j in range(len(linet\_subset\_gID)): 87 if linet\_subset\_gID[j][5] == gIDs[i]: 88 indx[i].append(j) 89 90 for i in range(len(indx)): # Checking incorrect index jumps 91 if len(indx[i]) > 1: 92 jumps = [] 93 for j in range(1,len(indx[i])): 94if indx[i][j] > indx[i][j-1]+5: 95 jumps.append(j) 96 if len(jumps) != 0: 97 new\_gIDs = [None] \* len(jumps) 98 new\_groups = [None] \* len(jumps) 99 for k in range(len(jumps)): 100 new\_gIDs[k] = int(str(i)+str(k))\*100000 # 101 Assign a non-existent GroupID of 6 digits gIDs.append(new\_gIDs[k]) 102 if k != len(jumps)-1: 103 new\_groups[k] = indx[i][jumps[k]:jumps[ 104k+1]] else: 105 if jumps[k] < len(indx[i])-1:</pre> 106 new\_groups[k] = indx[i][jumps[k 107 ]:-1] new\_groups[k].append(indx[i][-1]) 108 else: 109 new\_groups[k] = [indx[i][-1]] 110 for k in range(len(new\_groups)): 111 for m in range(len(new\_groups[k])): 112 linet\_subset\_gID[new\_groups[k][m]][5] = 113 new\_gIDs[k] 114 day\_per\_gID = [] 115hour\_per\_gID = [] 116 lat\_per\_gID = [] 117 lon\_per\_gID = [] 118 ID\_per\_gID = [] 119 for i in range(len(gIDs)): 120 day\_per\_gID.append([]) 121 hour\_per\_gID.append([]) 122lat\_per\_gID.append([]) 123 lon\_per\_gID.append([]) 124

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.4 Function get\_linet\_timing.py

```
ID_per_gID.append([])
125
126
127
           for i in range(len(linet_subset_gID)):
                pos = gIDs.index(linet_subset_gID[i][5])
128
                day_per_gID[pos] = linet_subset_gID[i][0]
129
                hour_per_gID[pos].append(linet_subset_gID[i][1])
130
                lat_per_gID[pos].append(linet_subset_gID[i][2])
131
                lon_per_gID[pos].append(linet_subset_gID[i][3])
132
                ID_per_gID[pos].append(linet_subset_gID[i][4])
133
134
           for i in range(len(gIDs)):
135
                day_pos_in_matches = matches.index(day_per_gID[i])
136
                list_to_events = [min(hour_per_gID[i]), max(
137
      hour_per_gID[i]), min(lat_per_gID[i]), max(lat_per_gID[i]), min(
      lon_per_gID[i]), max(lon_per_gID[i]), ID_per_gID[i]]
                events[day_pos_in_matches].append(list_to_events)
138
139
140
           # Extracting data from those lines with group ID == 0
141
           event_pos = [None] * len(matches)
142
           # Accounting for positions taken by lines with group_ID !=
143
      0
           for i in range(len(events)):
144
                event_pos[i] = len(events[i])
145
146
           row_0 = [linet_subset[0][1], linet_subset[0][1],
147
      linet_subset[0][2], linet_subset[0][2], linet_subset[0][3],
      linet_subset[0][3], [linet_subset[0][4]]]
148
           events[matches.index(linet_subset[0][0])].append(row_0)
149
150
           for i in range(1,len(linet_subset)):
151
152
153
                current_day = linet_subset[i][0]
                prev_day = linet_subset[i-1][0]
154
                same_day = current_day == prev_day
155
                day_pos = matches.index(current_day)
156
157
                current_hour = linet_subset[i][1]
158
                prev_hour = linet_subset[i-1][1]
159
                same_hour = prev_hour <= current_hour and current_hour</pre>
160
      <= prev_hour + 1
161
                current_lat = linet_subset[i][2]
162
163
                current_lon = linet_subset[i][3]
164
165
                current_ID = linet_subset[i][4]
166
167
                same_event = [same_day, same_hour]
168
169
                if all(same_event) == True:
170
                    # Refresh end time
171
                    events[day_pos][event_pos[day_pos]][1] =
172
      current_hour
                    # Refresh max and min latitude if necessary
173
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.5 Function MMIA\_data\_ordering.py

```
if current_lat < events[day_pos][event_pos[day_pos</pre>
174
      ]][2]:
                        events[day_pos][event_pos[day_pos]][2] =
175
      current_lat
                    elif current_lat > events[day_pos][event_pos[
176
      day_pos]][2]:
                        events[day_pos][event_pos[day_pos]][3] =
177
      current_lat
                    # Refresh max and min longitude if necessary
178
                    if current_lon < events[day_pos][event_pos[day_pos</pre>
179
      ]][4]:
                        events[day_pos][event_pos[day_pos]][4] =
180
      current_lon
                    elif current_lon > events[day_pos][event_pos[
181
      day_pos]][5]:
                        events[day_pos][event_pos[day_pos]][5] =
182
      current_lon
                    # Add trigger ID to snippet list of ID's
183
                    events[day_pos][event_pos[day_pos]][6].append(
184
      current_ID)
185
                else:
186
                    events[day_pos].append([current_hour, current_hour,
187
       current_lat, current_lat, current_lon, current_lon, [current_ID
      ]])
                    if same_day == True:
188
                        event_pos[day_pos] = event_pos[day_pos] + 1
189
190
       print('Done!')
191
       print(' ')
192
       return [events, indx]
193
```

# B.5 Function MMIA\_data\_ordering.py

```
LISTING B.5: Full code for MMIA_data_ordering.py function
```

```
1 def MMIA_data_ordering(read_path, dir_path, linet_times, matches):
      1.1.1
2
      This function gets a directory where unordered .cdf files are
3
     located
      and gets them snippet-ordered in diferent directories inside
4
     dir_path.
      Those new directories are named by YearMonthDay_snippetIndex,
\mathbf{5}
     and files
      are named after their ID number.
6
7
      Parameters
8
9
      read_path : string
10
          Path of the directory where the unordered .cdf files are
11
     located.
      dir_path : string
12
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.5 Function MMIA\_data\_ordering.py

```
Path to the directory where the new daily-ordered
13
     directories with
          ordered separated .cdf files will be located.
14
      linet_times : list_to_events
15
          List of daily lists of snippet info (time, geolocation and
16
     MMIA Id's)
      matches : list
17
          List of dates with existing GLM and MMIA files.
18
19
      Returns
20
21
      Snippet-ordered files in new directories.
22
      1.1.1
23
24
      print('Ordering .cdf files according to ID in snippet...')
25
26
      with os.scandir(read_path) as files:
27
           files = [file.name for file in files if file.is_file() and
28
     file.name.endswith('.cdf')]
29
      to_analize = []
30
      for i in range(len(files)):
31
           file_date = files[i][50:54] + files[i][55:57] + files[i
32
     ][58:60]
           if matches.count(file_date) != 0: # Exists in matches
33
               to_analize.append(i)
34
      files = [files[index] for index in to_analize]
35
36
      if len(files)==0:
37
           print('Error: No matching MMIA .cdf files found to process!
38
      • )
39
      ID_path = read_path+'/cdf_ID'
40
      os.system('mkdir ' + ID_path)
41
42
      for i in range(len(files)):
43
          if len(files[i]) == 92:
44
               ID = files[i][85:88]
45
           elif len(files[i]) == 93:
46
               ID = files[i][85:89]
47
           elif len(files[i]) == 94:
48
               ID = files[i][85:90]
49
           elif len(files[i]) == 108:
50
               ID = files[i][85:88]
51
           elif len(files[i]) == 109:
52
               ID = files[i][85:89]
53
          elif len(files[i]) == 110:
54
               ID = files[i][85:90]
55
56
           os.system('cp ' + read_path+'/'+files[i] + ' ' + ID_path+'/
57
      '+ID+'.cdf')
58
      for i in range(len(linet_times)):
59
           for j in range(len(linet_times[i])):
60
               os.system('mkdir '+dir_path+'/'+matches[i]+'_'+str(j))
61
               IDs = linet_times[i][j][-1]
62
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.6 Function extract\_MMIA.py

| 63 | <pre>for k in range(len(IDs)):</pre>               |  |
|----|--|--|
| 64 | os.system('mv '+ID_path+'/'+str(IDs[k])+'.cdf' + ' |  |
|    | <pre>' + dir_path+'/'+matches[i]+'_'+str(j))</pre> |  |
| 65 | os.system('rm -rf '+ID_path)                       |  |
| 66 | print('Ordering done!')                            |  |
| 67 | <pre>print(' ')</pre>                              |  |

## B.6 Function extract\_MMIA.py

LISTING B.6: Full code for extract\_MMIA.py function

```
1 def extract_MMIA(MMIA_files_path, MMIA_ordered_dir,
     MMIA_ordered_outputs, matches, linet_times):
      1.1.1
2
      This function gets the ordered .cdf files and calls a MatLab
3
     script
      to process them, returning the .mat files to an output folder
4
     whose
      names match the file's date.
\mathbf{5}
6
      Parameters
7
8
      MMIA_files_path : string
9
          Path to the directory where the ordered .cdf files are
10
     located.
      MMIA_ordered_outputs : string
11
          Path to the directory where the snippet-ordered files are
12
     located.
      MMIA_ordered_outputs : string
13
          Path to the directory where the processed files will be
14
     located.
      matches : list
15
          List of common dates with GLM and MMIA data, as strings in
16
     the form
          YearMonthDay.
17
      linet_times : list
18
          List of daily lists of important snippet information.
19
20
      Returns
21
22
      Snippet-ordered and extracted MMIA .cdf files into a folder
23
     with
      .mat extension.
24
      1.1.1
25
26
      print('Extracting data from MMIA .cdf files...')
27
      print(' ')
28
29
      for i in range(len(linet_times)):
30
31
          print('Processing MMIA data, date %d %d / %d...' % (int(
32
     matches[i]), i+1, len(matches)))
33
```

```
for j in range(len(linet_times[i])):
34
35
               current_snippet_path = MMIA_ordered_dir + '/' + matches
36
     [i] + '_' + str(j)
37
               with os.scandir(MMIA_ordered_outputs) as pre_done:
38
                   pre_done = [file.name for file in pre_done if file.
39
     is_file() and file.name.endswith('.mat')]
40
               if pre_done.count(matches[i] + '_' + str(j)+'.mat') ==
41
     0:
42
                   with os.scandir(current_snippet_path) as files:
43
                       files = [file.name for file in files if file.
44
     is_file() and file.name.endswith('.cdf')]
                   size = len(files)
45
46
                   if size != 0:
47
                       print('Starting the MatLab engine and
48
     extracting data from .cdf files for snippet %d, %d / %d...' % (
     int(matches[i]),j, len(linet_times[i])))
                       eng = matlab.engine.start_matlab()
49
                       path = MMIA_ordered_dir + '/' + matches[i] + '_
50
      ' + str(j) + '/'
                       eng.workspace['str'] = path
51
                       eng.MMIA_symplified_v4(nargout=0)
52
                       eng.quit()
53
                       wd = os.getcwd()
54
                       os.system('mv '+wd+'/MMIA_data.mat ' +
55
     MMIA_ordered_outputs+'/'+matches[i]+'_' + str(j) +'.mat')
                   else:
56
                       print('There is no MMIA data for date %d,
57
     snippet %d' % (int(matches[i]), j))
58
               else:
                   print('MMIA data for day %d snippet %d was pre-
59
     extracted' % (int(matches[i]), j))
          print('Date', matches[i], ' done')
60
          print(' ')
61
      print('Your processed .mat files can be accessed at ',
62
     MMIA_ordered_outputs)
      print(' ')
63
```

# B.7 Function unify\_MMIA\_data.py

LISTING B.7: Full code for unify\_MMIA\_data.py function

```
1 def unify_MMIA_data(output_path, linet_times, matches, show_plots):
2 '''
3 This functions gets all the MMIA's extracted data .mat files
4 from the directory output_path and creates and returns list
MMIA_raw_data.
5
6 Parameters
```

\_ \_ \_ \_ \_ \_ \_ \_ \_ \_ \_ \_

7

```
8
      output_path : string
9
          Path to the existing directory where the resulting daily .
     mat files
          are located.
10
      linet_times : list
11
          List of daily lists of important snippet information.
12
      matches : list
13
          List of common dates with GLM and MMIA data, as strings in
14
     the form
          YearMonthDay.
15
      show_plots : bool
16
          Boolean variable for showing plots all through the program.
17
18
      Returns
19
       _____
20
      MMIA_data : list
21
          A list of tables with all the information fount in the .mat
22
      files.
      1.1.1
23
24
      # Creation of the list of lists for MMIA data
25
      MMIA_raw_data = [None] * len(linet_times)
26
      for i in range(len(linet_times)):
27
           snippets = [None] * len(linet_times[i])
28
           MMIA_raw_data[i] = snippets
29
30
      with os.scandir(output_path) as files:
31
           files = [file.name for file in files if file.is_file() and
32
     file.name.endswith('.mat')]
      size = len(files)
33
      if size == 0:
34
          print('No MMIA .mat files found!')
35
36
      print('Uploading MMIA data from .mat files...')
37
      print(' ')
38
39
      for i in range(size):
40
           current_path = output_path + '/' + files[i]
41
           day = files[i][0:8]
42
           day_pos = matches.index(day)
43
           ID = int(files[i][9:-4])
44
45
          mat = sio.loadmat(current_path)
46
          current_data = mat.get('MMIA_all')
47
48
          time_jump = 0
49
          min_jump_time = 5 # [s]
50
           pos = 1
51
           while time_jump == 0:
52
               # If there is a time jump
53
               if current_data[pos,0]-current_data[pos-1,0] >=
54
     min_jump_time:
                   time_jump = 1
55
               # If there is no time jump and position is not last
56
```

```
elif current_data[pos,0]-current_data[pos-1,0] <</pre>
57
     min_jump_time and pos != len(current_data)-1:
58
                   pos = pos+1
               # If there is no time jump and position is last
59
               elif current_data[pos,0]-current_data[pos-1,0] <</pre>
60
     min_jump_time and pos == len(current_data)-1:
                   time_jump = 2
61
62
          if time_jump == 1:
63
               print('Day %s snippet %d has a time jump!' % (day, ID))
64
               print("This means that two or more lines in Linet's .
65
     csv have the same ID while they do not belong to the same group"
     )
               current_data = None
66
67
          MMIA_raw_data[day_pos][ID] = current_data
68
69
          if show_plots == 1 and time_jump !=1:
70
               plt.figure()
71
               plt.plot(MMIA_raw_data[day_pos][ID][:,0],MMIA_raw_data[
72
     day_pos][ID][:,1],linewidth=0.1, color='r')
               plt.title('MMIA 777.4nm Photometer Detections for day %
73
     d snippet %d with no filter applied' % (int(day), ID))
               plt.xlabel('Time [s]')
74
               plt.ylabel('Energy')
75
               plt.grid('on')
76
               plt.show()
77
78
      print('Done!')
79
      print(' ')
80
      return MMIA_raw_data
81
```

### B.8 Function condition\_MMIA\_data.py

```
LISTING B.8: Full code for condition_MMIA_data.py function
```

```
1 def condition_MMIA_data(MMIA_data, matches, show_plots):
       1.1.1
2
      This functions takes 'MMIA_data', a list of MMIA tables of
3
     information
      and applies a filter in 777.4nm photometer information vector
4
     to reduce
      noise.
\mathbf{5}
      It also plots every signal with and without the filter applied
6
      if 'show_plots' is True.
7
8
      Parameters
9
       _ _ _ _ _ _ _ _ _ _ _
10
      MMIA_data : list
11
          List of MMIA tables of information.
12
      matches : list
13
          List of common dates with GLM and MMIA data, as strings in
14
     the form
```

```
YearMonthDay.
15
16
      show_plots : bool
          Boolean variable for plotting. Not ploting makes the
17
     program faster.
18
      Returns
19
20
      MMIA_filtered : list
21
          A list of MMIA tables of information with a filter applied
22
          and only regarding time and 777.4nm photometer information
23
      1.1.1
24
25
      # Creation of the new list of lists of MMIA data with Moving
26
     Average
      MMIA_filtered = [None] * len(MMIA_data)
27
      for i in range(len(MMIA_data)):
28
          snippets = [None] * len(MMIA_data[i])
29
          MMIA_filtered[i] = snippets
30
31
      for i in range(len(MMIA_data)): # For every day of MMIA data
32
          print('Date %d, %d / %d...' % (int(matches[i]), i+1, len(
33
     matches)))
          for j in range(len(MMIA_data[i])):
34
35
               if type(MMIA_data[i][j]) == np.ndarray:
36
37
                   print('Applying a filter to reduce noise to MMIA
38
     signal, date %d snippet %d / %d...' % (int(matches[i]), j, len(
     MMIA_data[i])))
39
                   current_data=np.zeros((len(MMIA_data[i][j]),2))
40
                   current_data[:,0] = MMIA_data[i][j][:,0]
41
                   current_data[:,1] = MMIA_data[i][j][:,1]
42
43
44
                   if show_plots == 1:
45
                       plt.figure(figsize=(9, 3))
46
                       plt.plot(current_data[:,0],current_data[:,1],
47
     linewidth=0.5, color='r')
                   #
48
                   n = 15 # the larger n is, the smoother curve will
49
     be
                   b = [1.0 / n] * n
50
                   a = 1
51
                   current_data[:,1] = lfilter(b,a,current_data[:,1])
52
53
                   if (current_data[:,1] < 1.75).all() == True:</pre>
54
                       MMIA_filtered[i][j] = None
55
                       print('MMIA data for day %d snippet %d was just
56
      noise!' % (int(matches[i]), j))
                       print(' ')
57
                   else:
58
59
                       # Assuring continuity in MMIA_MA timesteps
60
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.9 Function extract\_GLM.py

```
MMIA_filtered[i][j] = TFG.
61
     fit_vector_in_MMIA_timesteps(current_data, int(matches[i]), j,
     1, 1)
62
                       #MMIA_filtered[i][j] = current_data
63
                       if show_plots == 1:
64
                            # MMIA representation with filter and with
65
     rectified time
                            #plt.figure()
66
                            plt.plot(MMIA_filtered[i][j][:,0],
67
     MMIA_filtered[i][j][:,1],linewidth=0.5, color='b')
                            plt.title("Untreated (red) and filtered (
68
     blue) MMIA 777.4nm photometer detections of day %d snippet %d" %
      (int(matches[i]), j))
                            plt.xlabel('Time [s]')
69
                            plt.grid('on')
70
                            plt.ylabel(r"Irradiance $\left[\dfrac{\mu W
71
     }{m^2}\right]$")
72
                            plt.legend(['Untreated signal', 'Filtered
     signal'])
                            plt.show()
73
74
               else:
75
                   print('There is no MMIA data for day %d, snippet %d
76
      ' % (int(matches[i]), j))
                   print(' ')
77
78
      print('Done!')
79
      print(' ')
80
      return MMIA_filtered
81
```

### B.9 Function extract\_GLM.py

LISTING B.9: Full code for extract\_GLM.py function

```
1 def extract_GLM(dir_path, output_path, linet_times, matches,
     MMIA_MA, angle_margin, cropping_margin):
      1.1.1
2
      This function calls every directory with .nc files and extracts
3
      the data
      of all the files in it via GLM_processing function.
4
5
      Parameters
6
       . _ _ _ _ _ _ .
7
      dir_path : string
8
          Path to the directory where the daily-ordered directories
9
     with
          ordered separated .nc files are located.
10
      output_path : string
11
          Path to the existing directory where the resulting daily .
12
     txt files
          will be located.
13
      linet_times : list_to_events
14
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.9 Function extract\_GLM.py

```
List of daily lists of snippet info (time, geolocation and
15
     MMIA Id's)
      matches : list
16
          List of dates with existing GLM and MMIA files
17
      MMIA_MA : list
18
          List of daily lists of MMIA's time and signal (with a
19
     filter
          applied) vectors for every snippet
20
      angle_margin : float
21
          Plus of latitude and longitude angle for extracting GLM
22
     data with
          respect to Linet's data.
23
      cropping_margin : float
24
          Plus of time before and afer MMIA snippet times (or Linet
25
     times) for
           extracting GLM data.
26
27
28
      Returns
29
      .txt files for every snippet with GLM data prepared to be
30
     analyzed.
31
32
      print('Extracting data from GLM .nc files into snippet .txt...'
33
     )
      print(' ')
34
35
      for i in range(len(linet_times)): # Analyzing each directory's
36
      .nc files
          for j in range(len(linet_times[i])):
37
38
               if type(MMIA_MA[i][j]) == np.ndarray:
39
40
                   print('Extracting GLM data for date %d snippet %d
41
     ....' % (int(matches[i]), j))
42
                   min_lat = linet_times[i][j][2] - angle_margin
43
                   max_lat = linet_times[i][j][3] + angle_margin
44
45
                   min_lon = linet_times[i][j][4] - angle_margin
46
                   max_lon = linet_times[i][j][5] + angle_margin
47
48
                   start_time = MMIA_MA[i][j][0,0] - cropping_margin
49
                   end_time = MMIA_MA[i][j][-1,0] + cropping_margin
50
51
                   # If willing to use Linet's timing, uncomment below
52
                   #start_time = linet_times[i][j][0] .
53
     cropping_margin
                   #end_time = linet_times[i][j][1] + cropping_margin
54
55
                   TFG.GLM_processing(dir_path+'/'+matches[i]+'/',
56
     output_path, matches[i]+'_'+str(j), min_lat, max_lat, min_lon,
     max_lon, start_time, end_time)
57
                   print('Date %s snippet %d done' % (matches[i], j))
58
                   print(' ')
59
```



| 60 | else:   |
|----|---|
| 61 | <pre>print('GLM data for date %d snippet %d will not be</pre>     |
|    | extracted due to lack of MMIA data' % (int(matches[i]), j))       |
| 62 | <pre>print(' ')</pre>   |
| 63 |   |
| 64 | <pre>print('Your processed .txt files can be accessed at ',</pre> |
|    | output_path)  |
| 65 | <pre>print(' ')</pre>   |

# B.10 Function unify\_GLM\_data.py

LISTING B.10: Full code for unify\_GLM\_data.py function

```
1 def unify_GLM_data(output_path, MMIA_MA, matches, show_plots):
       1.1.1
2
      This function gets all the GLM's extracted data .txt files from
3
      the
      directory output_path and creates and returns list GLM_data.
4
\mathbf{5}
      Parameters
6
7
      output_path : string
8
          Path to the existing directory where the resulting daily .
9
     txt files
          are located.
10
      MMIA_MA : list
11
          List of daily lists of MMIA's time and signal (with a
12
     moving average
          applied) vectors for every snippet
13
      matches : list
14
          List of dates with existing GLM and MMIA files
15
      show_plots : bool
16
          Boolean value for outputting plots all through the program.
17
18
      Returns
19
       _____
20
      GLM_data : list
21
          A list of daily lists with all the information found in the
22
      .txt files,
          ordered by snippets.
23
      1.1.1
24
25
      print('Uploading GLM data from .txt files...')
26
      # Creation of a new list of daily GLM snippets
27
      GLM_data = [None] * len(MMIA_MA)
28
      for i in range(len(MMIA_MA)):
29
          snippets = [None] * len(MMIA_MA[i])
30
          GLM_data[i] = snippets
31
32
      with os.scandir(output_path) as files:
33
          files = [file.name for file in files if file.is_file() and
34
     file.name.endswith('.txt')]
      size = len(files)
35
```

```
if size == 0:
36
           print('No GLM .txt files found!')
37
38
      column_subset = ['Time', 'Event_lat', 'Event_lon', 'Event_ID',
39
      'Flash_lat', 'Flash_lon', 'Radiance']
40
      for i in range(size):
41
42
           day = files[i][0:8]
43
           day_pos = matches.index(day)
44
           snip = int(files[i][9:-4])
45
46
           current_path = output_path + '/' + files[i]
47
48
           # Uploading the GLM's .txt using Pandas to sort it by time
49
           current_data = pd.read_csv(current_path, names=
50
     column_subset, sep=' \setminus s+')
          # Sorting data by time
51
           current_data = current_data.sort_values(by='Time')
52
           # Translating Pandas Dataframe to Numpy Matrix for easy
53
     data access
           current_data = current_data.to_numpy()
54
           # Appending current day to GLM_data
55
           GLM_data[day_pos][snip] = current_data
56
57
           if show_plots == 1:
58
               # Showing event map
59
               plt.figure()
60
               plt.scatter(current_data[:,2],current_data[:,1])
61
               plt.scatter(current_data[:,5],current_data[:,4],marker=
62
      'x')
               plt.axis('equal')
63
               plt.grid('on')
64
               plt.title('Event grid for day %d snippet %d' % (int(day
65
     ), snip))
               plt.xlabel('Longitude [deg]')
66
               plt.ylabel('Latitude [deg]')
67
               plt.show()
68
      print('Done!')
69
      print(' ')
70
71
      return GLM_data
72
```

### B.11 Function condition\_GLM\_data.py

LISTING B.11: Full code for condition\_GLM\_data.py function

```
5
\mathbf{6}
      Parameters
7
         ----
8
      GLM_total_raw_data : list
          List of daily GLM tables of data.
9
      matches : list
10
          List of dates with existing GLM and MMIA files
11
12
      show_plots : bool
          Boolean variable for plotting. Not ploting makes the
13
     program faster.
14
      Returns
15
        ____
16
17
      GLM_data : list
          List of daily lists of snippets with integrated GLM
18
     radiance in MMIA
          sample rate.
19
      1.1.1
20
21
      # Creating a new set of data
22
      GLM_data = [None] * len(GLM_total_raw_data)
23
      for i in range(len(GLM_total_raw_data)):
24
          snippets = [None] * len(GLM_total_raw_data[i])
25
          GLM_data[i] = snippets
26
27
      # Integrating and extending GLM vectors by date
^{28}
      for i in range(len(GLM_total_raw_data)): # For every date
29
     with GLM data
          print('Integrating and extending GLM data vector for day %d
30
     ...' % int(matches[i]))
          for j in range(len(GLM_total_raw_data[i])):
31
32
               if type(GLM_total_raw_data[i][j]) == np.ndarray and len
33
     (GLM_total_raw_data[i][j]) <= 1:
                   print('GLM detection for day %d snippet %d is void!
34
      ' % (int(matches[i]), j))
                   print(' ')
35
                   GLM_total_raw_data[i][j] = None
36
37
               elif type(GLM_total_raw_data[i][j]) == np.ndarray and
38
     len(GLM_total_raw_data[i][j]) != 0:
                   just_one_timestep = 1
39
                   check_pos = 1
40
                   while just_one_timestep == 1:
41
                       # If last position and same as timestep before
42
     in the .txt
                       if check_pos == (len(GLM_total_raw_data[i][j])
43
     -1) and GLM_total_raw_data[i][j][check_pos-1,0] ==
     GLM_total_raw_data[i][j][check_pos,0]:
                            just_one_timestep = 2
44
                       else:
45
                            # Different timestep in the .txt as in line
46
      before
                            if GLM_total_raw_data[i][j][check_pos-1,0]
47
     != GLM_total_raw_data[i][j][check_pos,0]:
                                just_one_timestep = 0
48
```

```
else: # Same timestep in the .txt as in
49
     line before
50
                                check_pos = check_pos + 1
51
               if type(GLM_total_raw_data[i][j]) == np.ndarray and
52
     just_one_timestep == 2:
                   print('GLM detection for day %d snippet %d contains
53
      only 1 timestep and will not be compared' % (int(matches[i]), j
     ))
                   print(' ')
54
                   GLM_total_raw_data[i][j] = None
55
56
               if type(GLM_total_raw_data[i][j]) == np.ndarray:
57
58
                   # Showing non-inflated time vector
59
                   if show_plots == 1:
60
                       plt.figure()
61
                       plt.plot(GLM_total_raw_data[i][j][:,0])
62
                       plt.xlabel('Samples')
63
                       plt.ylabel('Time [s]')
64
                       plt.title('Original GLM Time VS Samples for
65
     date %d snippet %d' % (int(matches[i]), j))
                       plt.grid('on')
66
                       plt.show()
67
68
                   # Integration
69
70
                   print('Integrating GLM data, date %d snippet %d / %
71
     d' % (int(matches[i]), j, len(GLM_total_raw_data[i])))
72
                   GLM_length = math.ceil((GLM_total_raw_data[i][j
73
     ][-1,0] - GLM_total_raw_data[i][j][0,0]) / 0.002)
74
                   # Current table of data (current day)
75
76
                   GLM_int_data = np.zeros((GLM_length, 2))
77
78
                   pos_0 = 0
79
                   for k in range(GLM_length): # For every sample
80
     accounting zeroes at GLM rate
                       GLM_int_data[k,0] = round(GLM_total_raw_data[i
81
     ][j][0,0] + k*0.002, 3)
                       t_min = GLM_int_data[k,0]
82
                       t_{max} = t_{min} + 0.002
83
                       inside = True
84
                        count = 0
85
86
                       while inside == True:
87
                            raw_pos = pos_0 + count
88
89
                            if GLM_total_raw_data[i][j][raw_pos,0] >=
90
     t_min and GLM_total_raw_data[i][j][raw_pos,0] < t_max:</pre>
                                GLM_int_data[k,1] = GLM_int_data[k,1] +
91
      GLM_total_raw_data[i][j][raw_pos,6]
92
```

```
# Check if the next GLM_total_raw_data
93
      sample will be added
94
                             raw_end = (raw_pos == (len(
95
      GLM_total_raw_data[i][j])-1))
96
                             if raw_end == False:
97
                                  inside = (GLM_total_raw_data[i][j][
98
      raw_pos+1,0]) < t_max
                                  if inside == True:
99
                                      count = count + 1
100
                                  else:
101
                                      count = count + 1
102
103
                                      pos_0 = raw_pos
                             else:
104
                                  inside = False
105
                    print('Done!')
106
107
108
                    GLM_data[i][j] = fit_vector_in_MMIA_timesteps(
      GLM_int_data, int(matches[i]), j, show_plots, 0)
109
                    # Check for too short snippet vectors
110
                    if len(GLM_data[i][j]) <=100:</pre>
111
                         print('Data for day %s snippet %d is too poor,
112
      only %d samples. This snippet will be omitted.' % (matches[i], j
      , len(GLM_data[i][j])))
                         GLM_data[i][j] = None
113
114
                    if show_plots == 1 and type(GLM_data[i][j]) == np.
115
      ndarray:
                         # Plotting lineality in GLM time vector with
116
      GLM sampling rate
                         plt.figure()
117
                         plt.plot(GLM_int_data[:,0])
118
                         plt.title('GLM Time vector of day %d snippet %d
119
       with 0.002s period' % (int(matches[i]), j))
                         plt.xlabel('Samples')
120
                         plt.ylabel('Time [s]')
121
                         plt.grid('on')
122
                         plt.show()
123
124
                         # Integrated GLM radiance vs time graph
125
      representation
                         plt.figure()
126
                        plt.plot(GLM_int_data[:,0],GLM_int_data[:,1],
127
      linewidth=0.5, color='black')
                         plt.grid('on')
128
                         plt.title('GLM signal of day %d snippet %d with
129
       GLM sample rate (0.002s) ' % (int(matches[i]), j))
                         plt.xlabel('Time [s]')
130
                         plt.ylabel('Radiance [J]')
131
                         plt.show()
132
133
       print('Integration of GLM vectors done!')
134
       print(' ')
135
136
```

137 return GLM\_data

#### B.12 Function fit\_vector\_in\_MMIA\_timesteps.py

LISTING B.12: Full code for fit\_vector\_in\_MMIA\_timesteps.py function

```
1 def fit_vector_in_MMIA_timesteps(GLM_int_data, day, snippet,
     show_plots, is_MMIA):
      1.1.1
2
      This function takes a time and signal pair of vectors and
3
     accommodates it
      into MMIA timesteps of 0.00001s. Inexistent values in between
4
     are filled
      by simple linear regression.
5
6
      Parameters
7
8
      GLM_int_data : list
9
          List of daily lists of data snippets. NOT necessarily GLM
10
     data.
      day : int
11
          Date of the form YearMonthDay.
12
      snippet : int
13
          Index of the current snippet to expand inside day "day".
14
      show_plots : bool
15
          Boolean variable for showing plots all through the program.
16
      is_MMIA : bool
17
          Boolean variable for sepparating GLM expansion from MMIA
18
     expansion.
19
      Returns
20
21
      GLM_current_data : list
22
          List of daily lists of snippets like "GLM_int_Data" input,
23
     but with
          accomodation to 0.00001s timesteps done.
24
      1.1.1
25
26
      # Expanding snippet to fit missing MMIA timesteps to cross-
27
     correlate data
      if is_MMIA == 0:
28
          print('Fitting GLM data in MMIA timesteps date %d snippet %
29
     d...' % (day, snippet))
      else:
30
          print('Completing MMIA data in MMIA timesteps date %d
31
     snippet %d...' % (day, snippet))
32
                               # New length of the timestep-wise
     new_length = 1
33
     matrix
      acumulated_voids = 0
                            # Number of non-existing timesteps up
34
     to current line
      void_info = np.zeros((len(GLM_int_data),4)) # Matrix of special
35
     info for each line:
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.12 Function fit\_vector\_in\_MMIA\_timesteps.py

```
# 1st column: .txt row number
36
          # 2nd column: void timesteps after that row until next
37
     existing timestep
          # 3rd column: Accumulated void timesteps before current
38
     line
          # 4th column: Differential energy between existing
39
     timesteps ([i]-[i-1])
40
      # Updating new_length value to make a new table with 1st
41
     dimension being new_length
42
      GLM_int_data[0,0] = round(GLM_int_data[0,0],5)
                                                           # Rounding
43
     to MMIA period
44
      for j in range(1,len(GLM_int_data)):
45
46
          GLM_int_data[j,0] = round(GLM_int_data[j,0],5) # Rounding
47
     to MMIA period
          void_info[j][0] = j  # Filling first void_info column
48
49
          if GLM_int_data[j,0] == GLM_int_data[j-1,0] + 0.00001: #
50
     Exactly one timestep ahead
              new_length = new_length + 1
51
              void_info[j][2] = acumulated_voids
52
53
          elif GLM_int_data[j,0] < GLM_int_data[j-1,0] + 0.00001: #</pre>
54
     Less than a whole timestep (sometimes occur)
              new_length = new_length + 1
55
              void_info[j][2] = acumulated_voids
56
57
                # There are missing timesteps in between current
          else:
58
     and last row
              void_timesteps = round((GLM_int_data[j,0] -
59
     GLM_int_data[j-1,0])/0.00001) - 1
              new_length = new_length + 1 + void_timesteps
60
              void_info[j-1][1] = void_timesteps
61
              acumulated_voids = acumulated_voids + void_timesteps
62
              void_info[j][2] = acumulated_voids
63
              void_info[j-1][3] = GLM_int_data[j,1] - GLM_int_data[j
64
     -1,1]
65
      # Filling the new time-wise matrix
66
67
      GLM_current_data = np.zeros((new_length,2)) # New matrix with
68
     void lines for non-existing timesteps
69
      for j in range(0,len(GLM_int_data)):
70
          new_j = int(j + void_info[j,2])
                                              # Row position in the
71
     new matrix
          GLM_current_data[new_j,:] = GLM_int_data[j,:] # Filling
72
     rows with existing data
73
          if void_info[j,1] != 0: # Lines with non-existing
74
     timesteps afterwards
              counter = 1
                                     # Adds 0.00001s and a linear
75
     energy fraction
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.13 Function signal\_delay.py

```
for k in range(new_j+1, new_j+1+int(void_info[j,1])):
76
                    GLM_current_data[k,0] = GLM_int_data[j,0] + counter
77
       * 0.00001
                    GLM_current_data[k,1] = GLM_int_data[j][1] +
78
      counter * (void_info[j][3]/void_info[j][1])
                   counter = counter + 1
79
80
       if show_plots == 1 and is_MMIA == 0:
81
           # GLM time representation at MMIA sample rate
82
           plt.figure()
83
           plt.plot(GLM_current_data[:,0])
84
           plt.title('GLM Time vector of day %d snippet %d with
85
      0.00001s period' % (day, snippet))
           plt.xlabel('Samples')
86
           plt.ylabel('Time [s]')
87
           plt.grid('on')
88
           plt.show()
89
90
           # Radiance vs time graph representation
91
           plt.figure()
92
           plt.plot(GLM_current_data[:,0],GLM_current_data[:,1],
93
      linewidth=0.5, color='black')
           plt.grid('on')
94
           plt.title('GLM signal of day %d snippet %d with MMIA sample
95
       rate (0.00001s)' % (day, snippet))
           plt.xlabel('Time (second of the day) [s]')
96
           plt.ylabel('Radiance [J]')
97
           plt.show()
98
99
       print('Date %d snippet %d fit' % (day, snippet))
100
       print(' ')
101
102
       return GLM_current_data
103
```

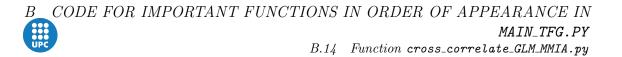
### B.13 Function signal\_delay.py

LISTING B.13: Full code for signal\_delay.py function

```
1 def signal_delay(data1, data2, show_plots, day, snip):
      1.1.1
2
      This function determines the delay in samples of two signals
3
     using a
      cross-correlation method.
\mathbf{4}
5
      Parameters
6
7
      data1 : Array
8
          First signal to be analyzed. In this case, GLM signal.
9
      data2 : Array
10
          Second signal to be analyzed. In this case, MMIA signal.
11
      show_plots : bool
12
           Boolean variable for showing plots all through the program.
13
      day : int
14
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.13 Function signal\_delay.py

```
Day of the snip to be cross-correlated in shape
15
     YearMonthDay
      snip : int
16
          Position of snip in day "day"
17
18
      Returns
19
20
      real_delay_samples : The number of samples that data1 is
21
     shifted with
          respect to data2.
22
      1.1.1
23
24
      xcorr_factors = correlate(data1[:,1], data2[:,1], mode='full',
25
     method = 'auto')
26
      len_x = len(data1) + len(data2) - 1
27
      x = np.empty(len_x)
28
29
      for i in range(len_x):
30
          if (len_x % 2) == 0: # Even number
31
               x[i] = (i - (len_x/2))
32
          if (len_x % 2) != 0: # Odd number
33
               x[i] = (i - (len_x/2 - 0.5))
34
35
      if show_plots == 1:
36
          plt.subplot(2, 1, 1)
37
          plt.plot(data2[:,1],'-r', linewidth = 0.5)
38
          plt.plot(data1[:,1],'-k', linewidth = 0.5)
39
          #plt.title('Non-correlated GLM (black) and MMIA (red)
40
     signals, day %d snippet %d' % (day, snip))
          plt.ylabel('Normalized Energy')
41
          plt.xlabel('Vector samples')
42
          plt.legend(['MMIA signal', 'GLM signal'])
43
44
          plt.grid('on')
45
          plt.subplot(2, 1, 2)
46
          plt.plot(x, xcorr_factors, '-b', linewidth = 0.5)
47
          plt.xlabel('Diff. Samples')
48
          plt.ylabel('Correlation Factor')
49
          plt.grid('on')
50
51
      # Delay samples accounting actual positioning due to different
52
     lengths:
      max_factor_pos = np.where(xcorr_factors == max(xcorr_factors))
53
      [0][0]
54
      if ((len(data1)+len(data2)) % 2 == 0): # len(x) is Odd
55
          delay_samples = x[max_factor_pos]+(len(data1)-len(data2))/2
56
57
      if ((len(data1)+len(data2)) % 2 != 0): # len(x) is Even
58
          delay_samples = x[max_factor_pos]+(len(data1)-len(data2))/2
59
      + 0.5
60
      # Delay samples accounting actual positioning due to time:
61
      if data1[0,0] > data2[0,0]: # GLM vector starts later
62
```



| 63 | <pre>pos_MMIA_start_GLM = np.where(data2[:,0] &lt;= data1[0,0])</pre>   |
|----|---|
|    | [0] [-1]  |
| 64 | real_delay_samples = delay_samples + pos_MMIA_start_GLM                 |
| 65 | <pre>elif data1[0,0] &lt; data2[0,0]: # GLM vector starts earlier</pre> |
| 66 | <pre>pos_GLM_start_MMIA = np.where(data1[:,0] &lt;= data2[0,0])</pre>   |
|    | [0] [-1]  |
| 67 | real_delay_samples = delay_samples - pos_GLM_start_MMIA                 |
| 68 | else:   |
| 69 | real_delay_samples = delay_samples                                      |
| 70 |   |
| 71 | <pre>return (real_delay_samples)</pre>                                  |

#### B.14 Function cross\_correlate\_GLM\_MMIA.py

LISTING B.14: Full code for cross\_correlate\_GLM\_MMIA.py function

```
1 def cross_correlate_GLM_MMIA(GLM_snippets, MMIA_snippets, GLM_norm,
      MMIA_norm, matches, show_plots):
      1.1.1
2
      This function gets snippets from GLM and MMIA and cross-
3
     correlates them
      to syncronize the signals and compare peaks.
4
\mathbf{5}
      Parameters
6
      _____
7
      GLM_snippets : list
8
          List of daily lists of GLM snippets.
9
10
      MMIA_snippets : list
          List of daily lists of MMIA snippets.
11
      GLM_norm : list
12
          List of daily lists of GLM normalized snippets.
13
      MMIA_norm : list
14
          List of daily lists of MMIA normalized snippets.
15
      matches : list
16
          List of existing GLM and MMIA dates.
17
      show_plots : bool
18
          Boolean variable for plotting. Not ploting makes the
19
     program faster.
20
      Returns
21
       _ _ _ _ _ _ _
22
      GLM_xcorr : list
23
          List of daily lists of synchronized GLM data.
24
      MMIA_xcorr : list
25
          List of daily lists of synchronized MMIA data.
26
      delays : list
27
          List of daily lists of delay between GLM and MMIA signal
28
     per snippet.
       1.1.1
29
30
      print('Starting cross-correlation of snippets and
31
     syncronization of signals...')
32
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.14 Function cross\_correlate\_GLM\_MMIA.py

```
# Creation of new lists for daily GLM and MMIA cross-correlated
33
      data
      GLM_xcorr = [None] * len(GLM_snippets)
34
      MMIA_xcorr = [None] * len(MMIA_snippets)
35
      delays = [None] * len(GLM_snippets)
36
      for i in range(len(GLM_snippets)):
37
          snips = [None] * len(GLM_snippets[i])
38
          GLM_xcorr[i] = snips
39
      for i in range(len(MMIA_snippets)):
40
          snips = [None] * len(MMIA_snippets[i])
41
          MMIA_xcorr[i] = snips
42
      for i in range(len(GLM_snippets)):
43
          snips = [None] * len(GLM_snippets[i])
44
          delays[i] = snips
45
46
      for i in range(len(GLM_snippets)):
47
          print('Date %d, %d / %d...' % (int(matches[i]), i+1, len(
48
     matches)))
          for j in range(len(GLM_snippets[i])):
49
               if i == 6 and j == 2:
50
                   show_plots = 1
51
               else:
52
                   show_plots = 0
53
               # If there's no info for this snippet due to lack of .
54
     nc or .cdf files
55
               if type(GLM_snippets[i][j]) == np.ndarray and type(
56
     MMIA_snippets[i][j]) == np.ndarray:
57
                   current_GLM = GLM_norm[i][j]
58
                   current_MMIA = MMIA_norm[i][j]
59
60
                   if show_plots == 1:
61
62
                       # Plotting cross-correlated and syncronized GLM
      and MMIA normalized signals
                       plt.figure()
63
                       plt.plot(current_MMIA[:,0], current_MMIA[:,1],
64
     color = 'r', linewidth = 0.5)
                       plt.plot(current_GLM[:,0], current_GLM[:,1],
65
     color = 'black', linewidth = 1)
                       plt.legend(['MMIA','GLM'])
66
                       plt.title('GLM (black) and MMIA (red)
67
     correlated normalized signals for day %d snippet %d' % (int(
     matches[i]), j))
                       plt.xlabel('Time [s]')
68
                       plt.ylabel('Normalized energy')
69
                       plt.grid('on')
70
                       plt.show()
71
72
                   # Calculation of delay in samples of GLM with
73
     respect to MMIA
                   delay = int(TFG.signal_delay(current_GLM,
74
     current_MMIA, show_plots, int(matches[i]), j))
75
                   delays[i][j] = delay
76
77
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.14 Function cross\_correlate\_GLM\_MMIA.py

```
MMIA_xc = current_MMIA
78
                    GLM_xc = np.zeros((len(current_GLM),2))
79
80
                    for k in range(len(current_GLM)):
81
                        if delay != 0: # There is delay
82
                            # Adjust Normalized vector
83
                            GLM_xc[k,0] = current_GLM[k,0] - delay
84
      *0.00001
                            GLM_xc[k,1] = current_GLM[k,1]
85
                            # Adjust original vector
86
                            GLM_snippets[i][j][k,0] = GLM_snippets[i][j
87
      ][k,0] - delay*0.00001
                        else: # delay==0 so no delay at all
88
                            GLM_xc[k,:] = current_GLM[k,:]
89
90
                    if show_plots == 1:
91
                        # Plotting cross-correlated and syncronized GLM
92
       and MMIA normalized signals
                        plt.figure()
93
                        plt.plot(MMIA_xc[:,0], MMIA_xc[:,1], color = 'r
94
      ', linewidth = 0.5)
                        plt.plot(GLM_xc[:,0], GLM_xc[:,1], color = '
95
      black', linewidth = 1)
                        plt.legend(['MMIA','GLM'])
96
                        plt.title('GLM (black) and MMIA (red)
97
      correlated normalized signals for day %d snippet %d' % (int(
      matches[i]), j))
                        plt.xlabel('Time [s]')
98
                        plt.ylabel('Normalized energy')
99
                        plt.grid('on')
100
                        plt.show()
101
102
                        # Plotting cross-correlated and syncronized GLM
103
       and MMIA non normalized signals
                        plt.figure()
104
                        plt.plot(MMIA_snippets[i][j][:,0],
105
      MMIA_snippets[i][j][:,1], color = 'r', linewidth = 0.5)
                        plt.plot(GLM_snippets[i][j][:,0], GLM_snippets[
106
      i][j][:,1], color = 'black', linewidth = 1)
                        plt.legend(['MMIA','GLM'])
107
                        plt.title('GLM (black) and MMIA (red)
108
      correlated normalized signals for day %d snippet %d' % (int(
      matches[i]), j))
                        plt.xlabel('Time [s]')
109
                        plt.ylabel('Energy')
110
                        plt.grid('on')
111
                        plt.show()
112
113
                    GLM_xcorr[i][j] = GLM_snippets[i][j]
114
                    MMIA_xcorr[i][j] = MMIA_snippets[i][j]
115
116
                    print('Date %d snippet %d cross-correlated and
117
      aligned!' % (int(matches[i]), j))
               elif type(GLM_snippets[i][j]) == np.ndarray and type(
118
      MMIA_snippets[i][j]) != np.ndarray:
```

CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.15 Function get\_GLM\_MMIA\_peaks.py

| 119 | <pre>print('Date %d snippet %d was pre-avoided for lack</pre>    |
|-----|--|
|     | of MMIA data' % (int(matches[i]), j))                            |
| 120 | <pre>elif type(GLM_snippets[i][j]) != np.ndarray and type(</pre> |
|     | <pre>MMIA_snippets[i][j]) == np.ndarray:</pre>                   |
| 121 | <pre>print('Date %d snippet %d was pre-avoided for lack</pre>    |
|     | of GLM data' % (int(matches[i]), j))                             |
| 122 | else:  |
| 123 | <pre>print('Date %d snippet %d was pre-avoided for lack</pre>    |
|     | of GLM and MMIA data' % (int(matches[i]), j))                    |
| 124 |  |
| 125 | <pre>print(' ')</pre>  |
| 126 | <pre>print('All snippets checked!')</pre>                        |
| 127 | <pre>print(' ')</pre>  |
| 128 | <pre>return [GLM_xcorr, MMIA_xcorr, delays]</pre>                |

### B.15 Function get\_GLM\_MMIA\_peaks.py

LISTING B.15: Full code for get\_GLM\_MMIA\_peaks.py function

```
1 def get_GLM_MMIA_peaks(GLM_xcorr, MMIA_xcorr, matches, show_plots):
      1.1.1
2
      This function gets the cross-correlated vector snippets from
3
     GLM and MMIA
      and finds their indexes for every prominent peak in their
4
     signals.
      It returns a list of lists of index vectors for every snippet.
5
6
7
      Parameters
       _ _ _ _ _ _ _ _ _ _ _
8
      GLM_xcorr : list
9
          List of daily lists of synchronized GLM data.
10
      MMIA_xcorr : list
11
          List of daily lists of synchronized MMIA data.
12
      matches : list
13
          List of existing GLM and MMIA dates
14
      show_plots : bool
15
          Boolean variable for plotting. Not ploting makes the
16
     program faster.
17
      Returns
18
       _____
19
      GLM_peaks : list
20
          List of daily lists with vectors of GLM_xcorr indexes for
21
     peaks in the
          signal.
22
      GLM_peaks : list
23
          List of daily lists with vectors of MMIA_xcorr indexes for
24
     peaks in the
          signal.
25
      1.1.1
26
27
      print('Finding peaks in GLM and MMIA cross-correlated signals')
28
     print('This process can take a while...')
29
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.15 Function get\_GLM\_MMIA\_peaks.py

```
GLM_peaks = [None] * len(GLM_xcorr)
30
      MMIA_peaks = [None] * len(MMIA_xcorr)
31
32
      for i in range(len(GLM_xcorr)):
33
           snips = [None] * len(GLM_xcorr[i])
           GLM_peaks[i] = snips
34
      for i in range(len(MMIA_xcorr)):
35
           snips = [None] * len(MMIA_xcorr[i])
36
           MMIA_peaks[i] = snips
37
38
      for i in range(len(GLM_xcorr)):
39
           print('Date %s, %d / %d...' % (matches[i], i+1, len(
40
     GLM_xcorr)))
          for j in range(len(GLM_xcorr[i])):
41
42
               if type(GLM_xcorr[i][j]) == np.ndarray and type(
43
     MMIA_xcorr[i][j]) == np.ndarray:
44
                   print('Finding peaks for date %s snippet %d / %d' %
45
      (matches[i], j, len(GLM_xcorr[i])))
46
                   # Cropping in order to have the same time to
47
     compare
48
                   # Not overlapping conditions
49
                   GLM_left_cond = GLM_xcorr[i][j][-1,0]<=MMIA_xcorr[i</pre>
50
     ][j][0,0]
                   GLM_right_cond = GLM_xcorr[i][j][0,0]>=MMIA_xcorr[i
51
     ][j][-1,0]
52
                   if GLM_left_cond == True or GLM_right_cond == True:
53
                       print('Correlated snippets for date %s snippet
54
     %d do not overlap at all' % (matches[i], j))
                   else:
55
56
                       # Finding the starting position
57
                       GLM_first = 0
58
                       if GLM_xcorr[i][j][0,0] < MMIA_xcorr[i][j</pre>
59
     ][0,0]: # GLM starts first
                            start_pos = np.where(GLM_xcorr[i][j][:,0]
60
     <= MMIA_xcorr[i][j][0,0])[0][-1]
                            GLM_first = 1
61
                       elif GLM_xcorr[i][j][0,0] > MMIA_xcorr[i][j
62
     ][0,0]: # MMIA starts first
                            start_pos = np.where(MMIA_xcorr[i][j][:,0]
63
     <= GLM_xcorr[i][j][0,0])[0][-1]
                       else: # Both start at the sime timestep
64
                            start_pos = 0
65
66
                       # Finding the end position
67
                       GLM_last = 0
68
                        if GLM_xcorr[i][j][-1,0] < MMIA_xcorr[i][j</pre>
69
     ][-1,0]: # GLM ends first
                            end_pos = np.where(MMIA_xcorr[i][j][:,0] <=</pre>
70
      GLM_xcorr[i][j][-1,0])[0][-1]
                        elif GLM_xcorr[i][j][-1,0] > MMIA_xcorr[i][j
71
     ][-1,0]: # MMIA ends first
```

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.15 Function get\_GLM\_MMIA\_peaks.py

| 72  | MMIA_xcorr[i][j][-1,0           | <pre>end_pos = np.where(GLM_xcorr[i][j][:,0] &lt;= ])[0][-1] GLM_last = 1</pre> |
|-----|---------------------------------|---|
| 73  |                                 | —   |
| 74  |                                 | : # Both end at the sime timestep   |
| 75  |                                 | end_pos = -1  |
| 76  |                                 |   |
| 77  | # Cr                            | opping vectors accordingly  |
| 78  | if G                            | LM_first == 1 and GLM_last == 1:  |
| 79  |                                 | GLM_vector = GLM_xcorr[i][j][start_pos:   |
|     | end_pos,1]                      |   |
| 80  | - ·                             | <pre>GLM_time_vector = GLM_xcorr[i][j][start_pos</pre>                          |
| 00  | :end_pos,0]                     |   |
| 01  | -                               | MMIA_vector = MMIA_xcorr[i][j][:,1]   |
| 81  |                                 |   |
| 82  |                                 | <pre>MMIA_time_vector = MMIA_xcorr[i][j][:,0]</pre>                             |
| 83  |                                 |   |
| 84  |                                 | GLM_first == 1 and GLM_last != 1:   |
| 85  |                                 | GLM_vector = GLM_xcorr[i][j][start_pos  |
|     | :-1,1]                          |   |
| 86  |                                 | <pre>GLM_time_vector = GLM_xcorr[i][j][start_pos</pre>                          |
|     | :-1,0]                          |   |
| 87  |                                 | <pre>MMIA_vector = MMIA_xcorr[i][j][0:end_pos,1]</pre>                          |
| 88  |                                 | MMIA_time_vector = MMIA_xcorr[i][j][0:  |
|     | end_pos,0]                      |   |
| 89  | -                               |   |
| 90  | elif                            | GLM_first != 1 and GLM_last == 1:   |
| 91  |                                 | GLM_vector = GLM_xcorr[i][j][0:end_pos,1]                                       |
| 92  |                                 | GLM_time_vector = GLM_xcorr[i][j][0:end_pos                                     |
| 52  | ,0]                             |   |
| 93  |                                 | MMIA_vector = MMIA_xcorr[i][j][start_pos  |
| 33  | :-1,1]                          | min_vootoi min_kooii[i][j][budit_pob  |
| 94  |                                 | MMIA_time_vector = MMIA_xcorr[i][j][  |
| 34  | <pre>start_pos:-1,0]</pre>      |   |
| 95  | btdit_pob. 1,0]                 |   |
|     | olif                            | GLM_first != 1 and GLM_last != 1:   |
| 96  |                                 |   |
| 97  |                                 | GLM_vector = GLM_xcorr[i][j][:,1]   |
| 98  |                                 | GLM_time_vector = GLM_xcorr[i][j][:,0]  |
| 99  |                                 | <pre>MMIA_vector = MMIA_xcorr[i][j][start_pos:</pre>                            |
|     | end_pos,1]                      |   |
| 100 |                                 | MMIA_time_vector = MMIA_xcorr[i][j][  |
|     | <pre>start_pos:end_pos,0]</pre> |   |
| 101 |                                 |   |
| 102 |                                 |   |
| 103 |                                 | lculating indexes of peaks in GLM signal  |
| 104 | GLM_                            | <pre>peak_vec, _ = find_peaks(GLM_vector,</pre>                                 |
|     | prominence = $0.3e-14$ ,        |   |
| 105 |                                 |   |
| 106 | # Ca                            | lculating indexes of peaks in MMIA signal                                       |
| 107 |                                 | _noise_level = np.percentile(MMIA_vector  |
|     | ,90, axis=0)                    |   |
| 108 |                                 | _peak_vec, _ = find_peaks(MMIA_vector,  |
|     |                                 | ght = MMIA_noise_level, prominence = 0.4,                                       |
| 100 | distance=400)                   | o, prominon of t,   |
| 109 | OT M                            | nooka[i][i] = CIM nook  |
| 110 |                                 | <pre>peaks[i][j] = GLM_peak_vec peaks[i][i] = MMIA peak vec</pre>               |
| 111 | MMIA                            | _peaks[i][j] = MMIA_peak_vec  |
| 112 |                                 |   |

B CODE FOR IMPORTANT FUNCTIONS IN ORDER OF APPEARANCE IN MAIN\_TFG.PY B.15 Function get\_GLM\_MMIA\_peaks.py

```
# Cropping GLM_xcorr and MMIA x_corr
113
                        GLM_xcorr_new_snippet = np.zeros((len(
114
      GLM_vector),2))
                         GLM_xcorr_new_snippet[:,0] = GLM_time_vector
115
                        GLM_xcorr_new_snippet[:,1] = GLM_vector
116
                        GLM_xcorr[i][j] = GLM_xcorr_new_snippet
117
118
                        MMIA_xcorr_new_snippet = np.zeros((len(
119
      MMIA_vector),2))
                        MMIA_xcorr_new_snippet[:,0] = MMIA_time_vector
120
                        MMIA_xcorr_new_snippet[:,1] = MMIA_vector
121
                        MMIA_xcorr[i][j] = MMIA_xcorr_new_snippet
122
123
124
                        if show_plots == 1:
125
                             plt.figure()
126
                             plt.plot(GLM_vector, color = 'black',
127
      linewidth=0.5)
128
                             plt.plot(GLM_peak_vec, GLM_vector[
      GLM_peak_vec], "*", color='b')
                             plt.title('GLM peaks on day %d, snippet %d'
129
       % (int(matches[i]), j))
                             plt.xlabel('Samples')
130
                             plt.ylabel('Radiance [J]')
131
                             plt.grid('on')
132
                             plt.show()
133
134
                             plt.figure()
135
                             plt.plot(MMIA_vector, color = 'r',
136
      linewidth=0.5)
                             plt.plot(MMIA_peak_vec, MMIA_vector[
137
      MMIA_peak_vec], "*", color='b')
                             plt.title('MMIA peaks on day %d, snippet %d
138
      ' % (int(matches[i]), j))
                             plt.xlabel('Samples')
139
                             plt.ylabel(r'Irradiance $\left[\dfrac{\mu W
140
      }{m^2}\right]$')
                             plt.grid('on')
141
                             plt.show()
142
143
                else:
144
                    print('Date %s snippet %d was not cross correlated'
145
       % (matches[i], j))
           print(' ')
146
147
       print('Done!')
148
       return [GLM_peaks, MMIA_peaks]
149
```