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1	Constraints on global mantle evolution from Ce-Nd-Hf isotope systematics
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15	
16	Abstract
17	Mantle evolution is governed by continuous depletion by partial melting and replenishment by
18	recycling oceanic and continental crust. Several important unknowns remain, however, such as
19	the extent of compositional variability of the residual depleted mantle, the timescale,
20	composition, and mass flux of recycled oceanic and continental crust, and the relative
21	importance of recycling upper versus lower continental crust. Here, we investigate the Ce-Nd-Hf
22	isotope systematics in a globally representative spectrum of mid ocean ridge and ocean island

influence on the slope, scatter, and extent of the array. The model results suggest a relatively young (<1.5 Ga) average depletion age of the depleted mantle, consistent with Nd and Os

basalts. Using a Monte Carlo approach for reproducing the global Ce-Nd-Hf isotope mantle

array shows that the type and age of depleted mantle and recycled crust have the dominant

27 isotope model ages of abyssal peridotites, and an apparent moderate extent of incompatible 28 element depletion. The latter, however, is deceiving, because it reflects a natural sampling bias, 29 resulting from melting an inherently heterogeneous depleted mantle. In principal, recycling of 30 oceanic crust can potentially explain most of the isotopic range of the isotopically enriched end 31 of the global Ce-Nd-Hf mantle array, but only if the entire compositional variability of the 32 recycled crust is preserved during recycling, residence in the mantle, and re-melting. The latter is 33 unlikely, however, because many sources of internal chemical variance average out on the scale 34 of the bulk oceanic crust, during residence in the mantle, and subsequent sampling by partial 35 melting. Moreover, both the slope and limited scatter of the global Ce-Nd-Hf mantle arrays 36 show that recycling of bulk oceanic crust, that is, both the extrusive basalts and intrusive gabbros 37 of the lower oceanic crust must be considered, and are key to better understand crust-mantle 38 cycling in general. The Monte-Carlo simulation also indicates that the return flux from the 39 continental crust into the mantle is dominated by material from the lower continental crust, 40 consistent with current models of continental crust evolution, which all require that a substantial 41 amount of the mafic lower continental crust must be recycled into the mantle to maintain the 42 average andesitic composition of the continental crust.

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44 *Keywords:* OIB, MORB, Cerium isotopes, REE, mantle heterogeneity, depleted mantle

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### 47 **1. Introduction**

The evolution of Earth's mantle is governed by continuous depletion in incompatible elements through partial melting, and re-enrichment in incompatible elements by recycling oceanic and continental crust (e.g., Hofmann, 1997; Stracke, 2012, 2018; White, 2015a; Zindler and Hart, 1986). "The part of Earth's mantle from which basaltic melt has been extracted" and thus has become depleted in incompatible elements is defined as the "depleted mantle" (DM; Stracke, 53 2016). Depleted mantle is the main source of mid ocean ridge basalts (MORB), but also a 54 significant part of most ocean island basalt (OIB) sources, because most are characterized by a 55 relative depletion in the highly incompatible elements (Willbold and Stracke, 2006), and they are, 56 on average, isotopically depleted (εNd ~+5, (Kumari et al., 2016). Therefore, DM constitutes a

significant fraction of Earth's mantle, but its average extent of depletion and mass fraction of the
total mantle remain poorly constrained (e.g., Stracke, 2016; Stracke et al., 2019).

59 The Sr-Nd-Hf-Pb isotope composition of DM is usually inferred from the isotope ratios of 60 MORB. However, the Hf-Nd isotope (e.g., Cipriani et al., 2004; Salters and Dick, 2002; Stracke 61 et al., 2011) and Os isotope ratios of abyssal peridotites (e.g., Alard et al., 2005; Day et al., 2017; 62 Harvey et al., 2006; Liu et al., 2008), as well isotopically ultra-depleted Nd isotope signatures in 63 melt inclusions (Stracke et al., 2019) have revealed the existence of DM that is isotopically more 64 depleted than MORB. Melting of such "ultra-depleted" mantle also explains the increased Hf-65 Nd isotope variability of MORB compared to OIB (Salters et al., 2011; Sanfilippo et al., 2019), 66 suggesting that ultra-depleted mantle components could be a ubiquitous part of Earth's mantle. 67 The latter would have far-reaching implications for interpreting the isotopic variation in mantle-68 derived melts, because it would imply that the rate of mantle depletion and thereby also the rate 69 of mantle-crust exchange could be higher than previously thought (Stracke, 2018; Stracke et al., 70 2019; 2011). A higher rate of mantle-crust cycling would also decrease the lifetime of different 71 mantle materials, such as depleted mantle and recycled oceanic and continental crust, and thus 72 affect their isotopic evolution.

Here, we use the La-Ce isotope systematics in MORB and OIB to obtain new constraints on the composition and age of DM, but also on the different recycled crustal materials in Earth's mantle. <sup>138</sup>La undergoes branched decay to <sup>138</sup>Ba (66%) by electron capture and to <sup>138</sup>Ce (34%) by  $\beta^{-}$  decay. The total half-life of <sup>138</sup>La is ~1.03 × 10<sup>11</sup> years (Sato and Hirose, 1981), and is similar to the half-life of the <sup>147</sup>Sm-<sup>144</sup>Nd decay system (1.06 × 10<sup>11</sup> years; (Begemann et al., 2001). Cerium isotope ratios in oceanic basalts reflect the time-integrated La/Ce of their mantle source.
Because the Nd and Hf isotope ratios reflect the time-integrated Sm/Nd and Lu/Hf of the
sampled mantle source, combined Ce-Nd-Hf isotope data reflect the time-integrated rare earth
element (REE) patterns of their mantle source. Note that Sm/Hf ratios in oceanic basalts are
relatively invariable (Blichert-Toft et al., 1999; Salters and Stracke, 2004), so that La-Ce-Nd-Sm/Hf-Lu accurately define the curvature of the entire REE pattern.

84 The elements of the three radioactive decay systems La-Ce, Sm-Nd, and Lu-Hf, become 85 increasingly less incompatible in the following order: La, Ce, Nd, Sm~Hf, and Lu. Thus, a 86 negative correlation results between La/Ce and Sm/Nd or Lu/Hf for variable products of 87 magmatic differentiation. With time, this behavior also leads to negative correlations in Ce-Nd 88 and Ce-Hf isotope space (Fig. 1, section 2). Moreover, because Ce is more incompatible than Nd 89 and Hf, Ce/Nd and Ce/Hf are highly variable in different mantle materials, resulting in strongly 90 curved mixing arrays in Ce-Nd and Ce-Hf isotope diagrams for mixture between melts from 91 strongly depleted mantle (DM) and any more light rare earth element (LREE) enriched 92 components (e.g., recycled oceanic and continental crust, Willig and Stracke, 2019). The variable 93 extent of LREE depletion, age, and proportion of depleted mantle sources thus strongly 94 influences the slope of the MORB and OIB isotope arrays in Ce-Nd-Hf isotope space, on both 95 local (i.e., at an ocean island or individual ridge segment) and global scales. In contrast, for other 96 combinations of isotope ratios, for example Sr-Nd, mixing arrays are more linear, and their slope 97 is primarily influenced by the isotope composition of the enriched source components.

98 Cerium isotope ratios, in combination with Nd and Hf isotope ratios, are therefore a unique 99 tracer for identifying variable DM components in the sources of MORB and OIB. But the 100 variable light REE patterns of different recycled oceanic and continental crust components also 101 lead to distinct Ce-Nd-Hf isotope compositions, especially for recycled lower versus upper 102 continental crust. The Ce-Nd-Hf isotope arrays of oceanic basalts are thus also highly sensitive to the nature and relative proportion of the different recycled crustal materials in their mantlesource, on a local and global scale (e.g., Bellot et al., 2015; Willig and Stracke, 2019).

105 Here, we use a Monte Carlo approach to reproduce the global Ce-Nd-Hf mantle array defined 106 by a globally representative, and internally consistent suite of MORB and OIB (Willig and 107 Stracke, 2019). We investigate how different mantle components (DM, recycled oceanic and 108 continental crust) influence the extent and slope of the global Ce-Nd-Hf array. The different 109 components are allowed to vary in composition and age, resulting in a spectrum of mantle 110 compositions that is sampled by partial melting. Each melting event samples a specific 111 combination of sources, and repeated sampling of the compositional spectrum produces a 112 unique modeled Ce-Nd-Hf array. By quantifying how well the modeled arrays fit the observed 113 Ce-Nd-Hf array, we constrain the parameter space that best reproduces the observations with 114 respect to age, composition, and relative proportion of the various mantle materials, and identify 115 the parameters that have the greatest leverage on the extent and slope of the global Ce-Nd-Hf 116 array.

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#### 119 2. Defining the Ce-Nd-Hf "mantle arrays"

High-precision isotope data for MORB and OIB have previously been reported for samples
from the Mid-Atlantic and Pacific Ridges and for Iceland, Hawaii, Gough, Tristan da Cunha, St.
Helena, Mangaia island, and McDonald seamount (Willig and Stracke, 2019). In addition, new
Ce-Nd-Hf data for eight additional OIB samples from Gough, St. Helena, and Tristan da Cunha
are reported in supplementary table 1.

125 The samples for which Ce isotope data exist (supplementary table 1) cover most of the global 126 Sr-Nd-Pb-Hf isotopic spectrum of MORB and OIB. Their Ce isotope ratios range from about 127  $\epsilon$ Ce = -1.5 ( $\epsilon$ Nd = 12,  $\epsilon$ Hf = 20) at the depleted end of the spectrum, which is defined by

128 MORB and Icelandic basalts, to  $\varepsilon Ce = 0.2$  ( $\varepsilon Nd = -3$ ,  $\varepsilon Hf = -3$ ) on the enriched side, represented by OIB from Gough island, and Tristan da Cunha (Fig. 1, supplementary table 1). 129 Samples from Mangaia, St. Helena and McDonald seamount are clustered around  $\varepsilon Ce = 0.7$ 130  $(\epsilon Nd = 5, \epsilon Hf = 5)$ . These HIMU-type OIB have slightly lower  $\epsilon Hf$  for a given  $\epsilon Nd$  and  $\epsilon Ce$ 131 132 relative to the rest of the samples (Fig. 1). However, despite this small offset of the HIMU 133 basalts, all Ce-Nd-Hf data define a narrow linear array in Ce-Nd-Hf isotope space, particularly 134 relative to their distribution in Nd-Pb and Nd-Sr isotopic space (Stracke et al., 2005). 135 For the combined Ce-Nd-Hf data determined by Willig and Stracke (2019, and this study), 136 which are listed in supplementary table 1, ordinary least square bisector fits (Isobe et al., 1990) for the ECe-ENd-EHf data define the slope and intercept of the ENd-ECe, EHf-ECe, and EHf-ENd 137 138 arrays. The ordinary least square bisector defines the line that mathematically bisects the 139 regression lines of variable Y against X, and X against Y. When both X and Y variables have 140 analytical or inherent uncertainty, the ordinary least square bisector fit is likely the most reliable 141 fitting method for determining the underlying relationship between two variables (Isobe et al., 142 1990). This approach results in the following relationships: 143 144  $\varepsilon^{143}$ Nd= $\varepsilon^{138}$ Ce × -9.06 ± 0.35 (2 S.E.) - 0.71 ± 0.60 (2 S.E.)

145 
$$\mathbf{\epsilon}^{176}$$
Hf= $\mathbf{\epsilon}^{138}$ Ce × -14.84 ± 0.72 (2 S.E.) - 0.62 ± 1.81 (2 S.E.)

146  $\mathbf{\epsilon}^{176}$ Hf =  $\mathbf{\epsilon}^{143}$ Nd × 1.59 ± 0.09 (2 S.E.) + 0.59 ± 1.22 (2 S.E.)

148 For calculating the  $\varepsilon$ -values, the CHUR values are from Bouvier et al. (2008) for Nd and Hf  $(^{143}Nd/^{144}Nd = 0.512630 \text{ and } ^{176}Hf/^{177}Hf = 0.282785)$ , and from Willig and Stracke (2019) for Ce 149  $(^{138}\text{Ce}/^{136}\text{Ce} = 1.336897).$ 150 151 Including the few existing Ce-Nd isotope data for MORB and OIB (Bellot et al., 2015; Boyet 152 et al., 2019; Dickin, 1988; Dickin, 1987; Makishima and Masuda, 1994; Tanaka et al., 1987), does 153 not significantly change the slope of the correlation lines defined above. The abundant Nd-Hf 154 literature data (taken from the compilation in Stracke, 2012) produce a tighter fit, but within 155 error of our more limited combined Ce-Nd-Hf data, and almost identical to the equation defined 156 by Vervoort and Patchett (1996): 157  $\epsilon^{176}$ Hf= $\epsilon^{143}$ Nd × 1.50 ± 0.03 (2 S.E.) + 1.51 ± 0.24 (2 S.E.). 158 159 160 161 3. Ce-Nd-Hf mantle geochemistry 162 3.1. Ce-Nd-Hf isotope evolution of different mantle components 163 Evolution of different mantle materials in Ce-Nd-Hf isotope space is a function of their La/Ce, 164 Sm/Nd and Lu/Hf and the time available for isotopic decay. 165 Depleted mantle (DM, Salters and Stracke, 2004; Workman and Hart, 2005) for example, evolves high  $\epsilon$ Nd and  $\epsilon$ Hf, but low  $\epsilon$ Ce. The DM estimates of Salters and Stracke (2004), and 166 the three different estimates of Workman and Hart (2005) evolve along significantly different 167 168 trends in Ce-Nd-Hf isotopic space (Fig. 2). These differences result from the different strategies 169 for estimating the incompatible element content of the DM. Salters and Stracke (2004) base their 170 estimate on the composition of LREE-depleted MORB, whereas Workman and Hart (2005) use 171 abyssal peridotites to estimate the REE composition of the DM. Because most abyssal

172 peridotites are variably affected by partial melting and refertilization by melt-rock interaction 173 (e.g., Brunelli et al., 2014; 2006; Hellebrand and Snow, 2003; Stracke et al., 2011; Warren, 2016), 174 which re-enriches the LREE, the La/Ce (and Ce/Nd,Hf) are comparatively high for given 175 Sm/Nd in the Workman and Hart (2005) estimate. Hence, the Workman and Hart (2005) DM 176 evolves to higher Ce for given Nd and Hf isotope ratios than the Salters and Stracke (2004) DM 177 estimate, resulting in an evolution curve that is steeper than the mantle array for the Workman 178 and Hart (2005), and shallower than the mantle array for the Salters and Stracke (2004) DM 179 estimate.

Bulk recycled oceanic crust (OC) consists of extrusive MORB and intrusive rocks such as the oceanic gabbros of the lower oceanic crust. Recycled bulk OC, as estimated by White and Klein (2014), generally evolves along the εNd-εCe mantle arrays to isotopic compositions between DM and recycled upper and lower continental crust (UCC-LCC; Fig. 2; see more detailed discussion in section 5.1), but plots below the mantle array in εHf-εCe and εHf-εNd diagrams (Fig. 2a, c;

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185 Salters and White, 1998; Stracke et al., 2003).
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186 Recycled UCC and LCC (Rudnick and Gao, 2014) evolve high  $\varepsilon$ Ce, but low  $\varepsilon$ Nd and  $\varepsilon$ Hf.

188 εNd and εHf, with UCC evolving to more extreme isotope ratios than LCC for the same age. In

Continental crust components (UCC and LCC) extend the Nd-Hf mantle array towards lower

189 the Ce-Hf and Ce-Nd isotope diagrams (Fig. 2a, b), UCC and LCC evolve at an angle to the

190 mantle array towards lower  $\varepsilon$ Nd and  $\varepsilon$ Hf for given  $\varepsilon$ Ce relative to the mantle array defined in

191 section 2 (Fig. 2). Although the evolution trends for LCC and recycled bulk OC (White and

192 Klein, 2014) have similar slopes in Ce-Hf isotopic space, bulk OC generally has more moderate

193  $\epsilon$ Ce and  $\epsilon$ Nd values than LCC for any given age. In Ce-Nd isotope space UCC, and especially

194 LCC evolve along trends with shallower slope than bulk OC, which evolves along a trend

195 parallel to the mantle array. Especially recycled LCC develops along a vector that is significantly 196 steeper than the Ce-Nd(Hf) mantle array, which may allow distinguishing recycled LCC and 197 UCC materials in different mantle sources.

198

## 199 3.2 Generating the Ce-Nd-Hf mantle array by sampling heterogeneous mantle

For interpreting the Ce-Nd-Hf mantle array it is important to understand not only how mantle components evolve, but also how partial melting samples these components and how the derivative melts mix to form the erupted melts.

203 Mixing arrays for melts from DM and more incompatible element enriched source materials 204 (OC, UCC, LCC; Fig. 3) are generally curved in Ce-Nd and Ce-Hf isotope space. The degree of 205 curvature increases with increasing depletion of the DM (i.e., lower La/Ce and Ce content; Fig. 206 3). Although the nature of the enriched source component also has a significant influence, 207 variable curvature mostly results from different DM source components, because the La/Ce and 208 Ce contents in various enriched end-members are much less variable than those of different 209 depleted end-members. Variably depleted mantle can thus be identified by arrays with different 210 position and slopes in Ce-Nd(-Hf) isotope diagrams (Fig. 3), on both a local and global scale.

211 On a local scale, the involvement of strongly depleted mantle causes highly curved mixing 212 arrays with melts from an enriched source component. That is, steep slopes on the high  $\varepsilon$ Ce–low

213 ENd-EHf side (reflecting low proportions of DM), but shallow, almost horizontal slopes on the

- 214 low ECe-high ENd-EHf side of the mixing curve (reflecting high proportions of DM, Fig. 3).
- Hence, both different DM compositions, but also variable proportions of DM and enrichedcomponents can be resolved on a local scale.
- On a global scale, the nature of the average DM involved in MORB and OIB genesis controlsthe steepness of the global mantle array, and could influence whether it intersects the chondritic

reference value (Willig and Stracke, 2019). The slope of the global mantle array therefore reflects
the average composition of DM involved in global MORB and OIB genesis, and thus constrains
the average age and extent of depletion of the DM on a global scale.

222 The shape and extent of the global mantle array is also greatly affected by sampling variably 223 depleted and enriched source components in different proportions. For the mixing curves in Fig. 224 3, for example, the proportion of melt from the DM source component varies along each mixing 225 line. In contrast, in Fig. 4, the proportion of melts from DM to melts from a recycled 226 component (here consisting of OC and UCC) is kept constant for each individual mixing curve, 227 but the ratio of DM/(OC+UCC) varies between the different mixing curves. Increasing the 228 proportion of melts from the DM elevates the **E**Nd and **E**Hf relative to **E**Ce, and limits the extent 229 of the mixing curve towards the enriched source component (OC+UCC). Hence, for a similar 230 enriched source component, different proportions of melts from a single DM component can 231 lead to parallel trends in Ce-Nd and Ce-Hf space, which may lead to increasing isotopic variance 232 at the DM end of the arrays, on both local and global scales. Moreover, a relatively large 233 proportion of DM limits the extent of the global Ce-Nd-Hf array, and little variance in the 234 proportion of DM limits the variability perpendicular to the array (Fig. 4).

235 In summary, the geometry of the global Ce-Nd-Hf mantle array, and that of any local array 236 for a given ocean island or ridge segment, is highly sensitive to the average composition and age 237 of the DM involved. Although the Nd-Hf isotope data in MORB also indicate melting of 238 variably depleted mantle sources (Salters et al., 2011), the Ce isotope data, in combination with 239 Nd and Hf are much more sensitive to variably depleted mantle and thus provide higher 240 resolution for distinguishing different DM components. Combining Ce with Nd and Hf isotope 241 data is therefore a unique tool for identifying the role of variably depleted mantle in MORB and 242 OIB genesis. Moreover, other than for Nd-Hf data alone, combined Ce-Nd-Hf isotope data allow discriminating between different types continental material, and thus add new constraintson the relative importance of recycled LCC and UCC for crust-mantle evolution.

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#### 247 4. Modeling the Ce-Nd-Hf mantle array

Although the mantle is a complex assembly of many different incompatible element depleted 248 249 and enriched materials, most of the isotopic variation in oceanic basalts is captured within a 250 simple conceptual framework: depletion by partial melting and replenishment by recycling of the 251 generated oceanic and continental crust back into the mantle (e.g., Hofmann, 1997; Stracke, 252 2012, 2018; White, 2015a; 2015b; Willbold and Stracke, 2010; Zindler and Hart, 1986). Although 253 the mantle is certainly more complex, mixtures between melts from DM, recycled OC, UCC and 254 LCC can reproduce the general distribution and pattern of the observed Sr-Nd-Hf-Pb isotope 255 ratios (e.g., Stracke, 2012, 2018; White, 1985; 2015a; Willbold and Stracke, 2010; Zindler and 256 Hart, 1986).

257 Even when simplifying the problem to sampling of only four generic mantle components 258 (DM, OC, LCC, UCC), however, their abundance, compositional variance, and age are 259 unknown. Moreover, the distribution within the mantle and the sampling of the different mantle 260 components during partial melting is a, more or less, random process. Hence, even generating 261 the global mantle array by random sampling of only four variable components defines a system 262 with many coupled degrees of freedom. Such systems can suitably be investigated with Monte 263 Carlo methods, resulting in reasonable bounds on the range of input parameters, and therefore 264 valuable constraints on the isotopic evolution of Earth's mantle.

265

266 4.1 Model framework

#### 267 4.1.1 Principal model set-up

268 The Ce-Nd-Hf mantle array is modeled by mixing melts from four generic mantle components: 269 DM, recycled bulk OC, recycled UCC and LCC. Repeated sampling of the compositional 270 spectrum of the four components creates a discrete number of model points; each point 271 reflecting a different mixture of melts from these four components. Using 60 model points 272 proved to be a good number for computational efficiency and for having a large-enough number 273 to define a modeled mantle array in Ce-Nd-Hf isotope space. The modeled Ce-Nd-Hf array is 274 thus typically defined by 60 modeled points and is compared to the observed data (section 2, 275 supplementary table 1).

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#### 277 4.1.2 Composition and variability of mantle components

The composition of the DM varies in the model, with equal probability, between the estimates ofSalters and Stracke (2004), and the three different estimates of Workman and Hart (2005).

To model recycling of bulk OC, we use the median composition by White and Klein (2014). The assigned compositional variance reflects differences of  $\pm 1\%$  (S.D.) in the average degree of partial melting of a uniform bulk OC source (cf. Langmuir et al., 1992). As such the model captures geochemical variability between segments of oceanic crust produced at mid oceanic ridges with variable thermal profiles, or similar degrees of melting of compositionally variable bulk OC sources.

286 The UCC and LCC are the average estimates given by Rudnick and Gao (2014), and have no 287 compositional variance assigned. Upper continental crust is recycled into the mantle after 288 continental erosion, transportation, sedimentation, and subduction. All these processes average, 289 and thus minimize the enormous compositional spectrum of the recycled UCC, as evidenced by 290 the striking similarity of average subducted sediment (GLOSS, Plank, 2014) and UCC (Rudnick 291 and Gao, 2014). Rather than assigning an arbitrary compositional variance to the UCC, 292 compositional variance in the recycled continental crust is thus captured by varying the 293 proportion of UCC and LCC.

294

295 4.1.3 Age of the mantle components

All mantle end-members are allowed to vary in age between 0 and 4.56 Ga. At the time of their formation, all four components (DM, bulk OC, LCC, and UCC) derive from, and thus inherit the isotopic composition of the upper mantle. The isotope evolution of the upper mantle is approximated by evolving linearly from bulk silicate earth at 4.56 Ga to an average MORB isotope ratio at the present time (i.e.,  $\varepsilon Ce = \varepsilon Nd = \varepsilon Hf = 0$  to  $\varepsilon Ce = -1.2$ ,  $\varepsilon Nd = 9.2$ ,  $\varepsilon Hf = 18.4$ ).

301 The age of the DM component is the age of its last depletion, i.e., the time at which this 302 component has acquired its assumed composition by melt extraction from the upper mantle.

303 The age of recycled bulk OC is the recycling or mantle residence time, which in this model 304 equals the recycling time of the continental crust (UCC and LCC). However, both continental 305 crust components reside, and thus develop isotopically in the continental crust, before being 306 recycled into the mantle. The age of the crustal materials thus is the sum of their residence time 307 in the crust and mantle. Implicit in the model therefore is that the average age of the recycled OC is lower than the formation age of the recycled UCC and LCC. This is an emergent feature 308 309 of decomposing the UCC-LCC and OC ages into a crustal residence time (only experienced by 310 UCC and LCC) and mantle residence or recycling time (cf. Stracke et al., 2003). A simplifying 311 assumption is that the mean ages for the recycled UCC and LCC are equal. Although this is 312 unlikely in nature, varying the age of the recycled UCC relative to LCC has a similar effect on the 313 isotope ratios of the total continental crust (UCC and LLC) as varying the proportion of UCC 314 and LCC. It is therefore not explored separately.

315

#### 316 4.1.4 Melting and mixing of mantle components

317 During sampling by oceanic volcanism, the DM and the recycled crustal components (OC, LCC318 and UCC) melt to a large-enough extent that Ce, Nd and Hf are quantitatively extracted into the

319 melt. About 90% of the Hf is incorporated in the melt after 6% fractional melting of DM and 99% is extracted after melting to 10%. For Ce and Nd, which are more incompatible than Hf, 320 321 quantitative extraction from source to melt occurs at even lower degrees of melting. In our 322 model the DM melts to a large extent (maximum of 20%), and thus the relative abundance of 323 Ce-Nd-Hf is very similar to those in the DM components used in the model (section 4.1.3). More incompatible element enriched recycled components (OC, UCC, LCC) melt to an even 324 325 larger extent than the peridotitic DM (e.g., Lambart et al., 2016; Pertermann and Hirschmann, 326 2003), which also leads to quantitative extraction of Ce, Nd and Hf into the produced melts. A 327 simplifying assumption in the model therefore is that OC, UCC, LCC melt completely. The 328 melts from each component (DM, OC, UCC, LCC) then mix to form a basalt that is plotted as 329 one of the 60 total points of the modeled Ce-Nd-Hf mantle array.

330

### 331 4.1.5 The modeled Ce-Nd-Hf mantle array

Any specific modeled mantle array, that is, one model solution, is calculated assuming a mean, standard deviation (S.D.), minimum and maximum value for the following parameters: the age of the DM, the recycling time of crustal components in the mantle, the continental residence time (age of the recycled UCC, LCC = continental residence time + mantle residence time), the sampling ratio of OC/DM, OC/(UCC+LCC), and the proportion of LCC in the recycled total continental crust, LCC/(LCC+UCC).

Each of the parameters have minimum and maximum values assigned. The ages of the mantle components are between 4.56 and 0 Ga. The sampling proportions are  $\geq 0$ , and the ratio of recycled OC to DM (OC/DM) in the source is limited to (3/7) (Sobolev et al., 2007), with a mean  $\leq 1/4$ . This restriction is reasonable when compared to estimates of how much oceanic crust has been produced at oceanic ridges over Earth's history (e.g., Salters and Stracke, 2004; Stracke et al., 2003; Tackley, 2015). In the simplest case, assuming that the entire mantle has melted in Earth's history and that oceanic crust production reflects about 10% of partial melting, 345 OC/DM = 1/9, which is significantly less than the maximum OC/DM ~ 1/4 used here. The 346 model thus allows for a considerable range of OC/DM, acknowledging that OC/DM in 347 individual mantle sources could be larger or smaller than expected, on average, for the entire 348 mantle.

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#### 351 4.2 Exploration of model parameter space

One single of many possible model solutions is shown in Fig. 5. To evaluate what constraints the model provides on the input parameters (Table 1), and thus on global mantle evolution, requires identifying which of the modeled arrays reproduce the observed Ce-Nd-Hf array better than others. The latter is done by quantifying the goodness of fit between the modeled and observed Ce-Nd-Hf mantle array.

The goodness of fit algorithm used here is a measure for the average squared distance in isotopic space from modeled basalts to the closest observed oceanic basalt and vice versa. Using this algorithm, the best possible score is 0, which is obtained when the modeled mantle array is identical to the observed mantle array. Similar results are obtained using an alternative scoring algorithm based on the isotopic composition of, and variance in the mantle array end-members. The working, suitability, and validation of different scoring algorithms is further elaborated on in the supplementary materials.

Having a metric for the goodness of fit allows exploration of the parameter space systematically. This is done by selecting two controlled variables of interest and assigning unique values to them, for example the age of the DM and the age of the bulk OC. Other variables (Table 1) are allowed to vary and are adjusted to optimize the goodness of fit score using a gradient descent algorithm. The goodness of fit scores obtained after optimization represent the best fit obtainable for these two chosen controlled variables. Performing this operation for different combinations of values for the controlled variables allows identifying combinations of input parameters that produce a better fit than others, and illustrating these results in diagramssuch as Fig. 6-8.

373 Parameters of interest that are explored are the age of the DM, the recycling time of crustal 374 materials (OC, LCC, UCC) in the mantle and the sampling ratio of recycled LCC/(LCC+UCC); 375 Table 1). To investigate the former two, all possible combinations  $(7 \times 7)$  for the following ages 376 are taken for the DM and recycled crust (OC, LCC, UCC):  $0.5 \pm 0.25$ ,  $1 \pm 0.5$ ,  $1.5 \pm 0.75$ ,  $2 \pm 1$ , 377  $2.5 \pm 1.25$ ,  $3 \pm 1.5$ ,  $3.5 \pm 1.75$ Ga (1 S.D.). For each combination the other parameters (sampling 378 proportions, continental crust residence time) are tuned by gradient descent. The optimized 379 scores reflect how consistent each modeled array is with the observed Ce-Nd-Hf data, for a 380 given average DM age and crustal recycling time (Fig. 6).

To investigate the sensitivity of the mantle array to the ratio of recycled LCC/(LCC+UCC) sampled in the mantle, a similar routine is performed where the age of the DM and the sampling ratio of recycled LCC/(LCC+UCC) are controlled and the other parameters are optimized by gradient descent (Fig. 8).

385

#### 386 4.3. Model Results

## 387 4.3.1 Age of the DM, recycled OC, and recycling of LCC vs. UCC

Optimized goodness of fit scores as a function of the age of the DM and age of the recycled crust (OC, UCC, LCC) are contoured in Fig. 6. Best fit models have mean DM ages <1.5 Ga and recycled crust ages >1 Ga. Outside of these bounds, the model fits become increasingly poor; although admittedly the exact cut-off for scores that provide a good fit is somewhat subjective, and dependent on the selection of controlled parameters.

393 Several observations can be made with respect to the proportion of recycled LCC to total 394 continental crust (LCC+UCC; Fig. 7). The optimal solutions, found on convergence with the 395 gradient decent algorithm, have a relatively high amount of recycled LCC in the melt relative to 396 UCC (ca. 7/3, i.e., LCC/(LCC+UCC) = 0.7; Fig. 7). This fine-tuned proportion of 397 LCC/(LCC+UCC) is a function of the crustal age and the age of the DM. In general, when 398 longer recycling times are imposed on the optimization algorithm, higher LCC/(LCC+UCC) are 399 required (Fig. 7). This effect arises because as the continental crust components become older 400 and isotopically more extreme, less recycled UCC is required to reproduce the most isotopically 401 enriched OIB (Fig. 4). In addition, older, isotopically more extreme DM leads to pronounced 402 curvature in the mixing trends between DM and the recycled components (OC, LCC, UCC), and 403 to mixing over a larger isotopic range, which offsets the mantle array towards higher  $\varepsilon$ Nd and

404  $\epsilon$ Hf at given  $\epsilon$ Ce. In part, the models compensate for this effect during optimization by 405 increasing the ratios of LCC/(LCC+UCC) (Fig. 3, 4 and Fig. 7) and OC/DM (see 406 supplementary material).

407

#### 408 4.3.2 Age of the DM and defining LCC/UCC recycling ratio

409 Better constraints on the sampling of recycled LCC in oceanic basalts can be obtained if the ratio 410 of LCC/(LCC+UCC) is a controlled variable (Fig. 8), which is systematically explored rather 411 than a free parameter optimized by gradient descent (as done for Fig. 7). Optimized goodness of 412 fit scores are mapped out in Fig. 8 as a function of the age of the DM and LCC/(LCC+UCC). 413 Depleted mantle (DM) older than 1.5-2 Ga yields poor fits with the observed Ce-Nd-Hf mantle 414 array, as in Fig. 6. Models devoid of recycled UCC (right side of the diagram) fail to replicate the 415 mantle array. Good fits (score < 0.008) are attained for recycled LCC to UCC sampling ratios of 416 4:6 to 8:2 (i.e., LCC/(LCC+UCC) = 0.4 - 0.8).

417

418

419 **5.** Discussion

## 420 5.1 Recycling oceanic crust (OC)

421 Recycled OC has a large influence on the slope of the modeled Ce-Nd-Hf array, because with 422 age, the recycled bulk OC develops isotope ratios that increasingly deviate from the observed 423 Ce-Hf and Nd-Hf array (Fig. 2). However, mixtures of melts from DM and recycled bulk OC in 424 our model can only account for isotope ratios in the upper left quadrant in Fig. 2a,b, 3a,b and 425 4a,b; that is,  $\varepsilon$ Nd>0,  $\varepsilon$ Hf>0, and  $\varepsilon$ Ce<0. For reproducing isotope ratios that extend the array to

426 εNd<0, εHf<0, and εCe>0, the presence of recycled UCC or LCC is required. This effect makes

427 it difficult to extract tight constraints on the age of the recycled OC. Nevertheless, for the 428 recycled bulk OC and associated variance assumed, the best-fit models have mean ages >1 Ga 429 (and DM ages of < 1.5 Ga, Fig. 6 and 7). The residence time of recycled OC in the mantle 430 therefore exceeds the turn-around time of the mantle, which is on the order of several 100 Myr 431 for convective mantle velocities of cm/yr. Hence, the >1 Ga residence time of OC implies that 432 it is stored for considerable time in Earth's mantle before being reprocessed by partial melting in 433 MORB and OIB sources.

The model results, however, depend strongly on the assumed composition of the recycled 434 435 OC. We use bulk OC (White and Klein, 2014), which consists of the upper extrusive crust 436 (MORB) and the lower oceanic gabbros, rather than solely MORB. To account for differences in 437 bulk OC composition on a global scale, we assign compositional variance as described in section 438 5.1. The limited data for the lower OC (i.e., gabbros) show that it is highly heterogeneous on a 439 local and global scale (e.g., Coogan, 2014; White and Klein, 2014; Fig. 9). This limited knowledge 440 about the average composition of the lower OC, in addition to large variability of MORB 441 (Arevalo and McDonough, 2010; Gale et al., 2013; Jenner and O'Neill, 2012; White and Klein, 442 2014; Yang et al., 2018), imposes considerable uncertainty on the average composition and 443 variability of the bulk recycled OC (e.g., Stracke et al., 2003; White and Klein, 2014). Hence, it is 444 difficult to judge whether the model input parameter, the bulk OC of White and Klein (2014), 445 under-or overestimates the natural variance.

446 For evaluating the uncertainty associated with the recycled bulk OC it is important to consider 447 that the OC is compositionally heterogeneous on different scales. On a local scale, e.g., the scale 448 of ridge segments, compositional variability is caused by differences in mantle composition and 449 extent of mantle melting due to variable mantle temperature. At the scale of an individual cross 450 section of OC, the upper OC (i.e., MORB) differs from the lower crustal gabbros, mostly due to 451 fractional crystallization (e.g., Coogan, 2014). If the compositional heterogeneity within the bulk 452 OC, that is, the upper and lower OC together, is preserved during recycling, residence in the 453 mantle, and subsequent sampling by oceanic volcanism, ancient OC develops considerable 454 isotopic heterogeneity (e.g., Chauvel and Hémond, 2000; Koornneef et al., 2012; Stracke et al., 455 2003; 2005). The expected range of isotope ratios of recycled 2 Ga upper MORB alone, for 456 example, covers the entire Ce-Nd-Hf mantle array (Fig. 9). In principal, therefore, a large part of 457 the mantle array might be explained solely by recycling MORB if its entire compositional 458 variability is considered.

However, the region in Fig. 9 where ancient recycled MORB plot most densely is offset from the mantle array towards lower  $^{138}$ Ce/ $^{136}$ Ce. Moreover, the observed mantle array in Ce-Nd-Hf isotopic space is very narrow, which is difficult to maintain if the entire range of observed MORB compositions would be sampled with equal probability, given the large isotopic variance of recycled MORB perpendicular to the mantle array (Fig. 9). A moderating element is thus required to shift recycled MORB towards the mantle array. This is achieved by the additional presence of lower oceanic gabbros, which develop lower  $\varepsilon$ Ce, but higher  $\varepsilon$ Nd and  $\varepsilon$ Hf compared

to recycled MORB (Fig. 9).

467 Consequently, considering that recycled OC is not just MORB, but rather a package of upper 468 and lower OC (i.e., MORB + gabbros), the isotope ratios of bulk recycled OC fall closer to the 469 mantle array. Moreover, owing to the well defined Ce-Nd-Hf mantle array (Fig. 1), it is apparent 470 that not the entire compositional variance of the recycled OC is transferred into oceanic basalt 471 sources. This is plausible, because many sources of internal chemical variance average out on the 472 scale of the bulk OC, and during subsequent sampling of recycled OC by partial melting. Thus, 473 the geochemical variance between subducting packages of bulk OC is expected to be less than 474 the internal variance within each lithological unit (MORB vs. gabbors). In addition, the predicted 475 global variance of recycled OC is likely reduced by stirring and mingling during subduction and 476 mantle processing, and subsequently, during partial melting of heterogeneous mantle sources by 477 oceanic volcanism.

Hence, for evaluating the influence of recycled OC on mantle evolution it is crucial to
consider the composition of the bulk OC, rather than only the extrusive basalts (cf. Stracke et al.,
2003). Nevertheless, the composition and variance of recycled bulk OC remains a major
unknown; not only in our model, but also for our general understanding of global geochemical
fluxes between the crust and mantle.

483

484 **5.2 Recycling of continental crust** 

485 Although recycling OC can produce a large range of isotope ratios, the majority concentrates at 486  $\epsilon$ Nd>0,  $\epsilon$ Hf>0, and  $\epsilon$ Ce<0 (Fig. 9). Oceanic basalts with  $\epsilon$ Nd<0,  $\epsilon$ Hf<0, and  $\epsilon$ Ce>0 thus

487 require at least one additional, incompatible element enriched, high ECe but low ENd-EHf source 488 component, that is, a continental crust component. Even though the compositional variability of 489 the recycled bulk OC might be larger than accounted for by the composition in our model 490 (White and Klein, 2014, see discussion above), this conclusion appears robust. Chauvel et al. 491 (2008) reached a similar conclusion on basis of the global distribution of Nd-Hf isotope ratios in 492 oceanic basalts, which is re-affirmed by the Ce-Hf and Nd-Hf isotope systematics modeled here. 493 In Ce-Nd and Ce-Hf isotope space, however, LCC and UCC evolution trends are more 494 divergent than in Nd-Hf isotope space (Fig. 2). Thus, considering Ce in addition to Nd-Hf 495 isotope ratios provides valuable additional constraints on the type of recycled continental crust.

496 The sensitivity of the Ce isotope system to the ratio of LCC/UCC sampled in oceanic basalts has 497 a significant impact on the extent and slope of the modeled Ce-Nd-Hf mantle array (Fig. 2). 498 Good model fits result for LCC/UCC = 4/6 to 8/2 (i.e., LCC/(UCC+LCC) = 0.4 - 0.8; Fig. 8), 499 consistent with the high proportion of recycled LCC suggested from Fig. 6. While this range is 500 relatively broad, it does suggest that recycled LCC, whether introduced into the mantle by delamination or subduction erosion (e.g., Arndt and Goldstein, 1989; McKenzie and O'Nions, 501 502 1983; Stracke, 2012, 2018; Willbold and Stracke, 2006, 2010), dominates the return flux of 503 continental crust into Earth's mantle (e.g., Willbold and Stracke, 2006, 2010). The latter is 504 consistent with recent models for developing and maintaining an average andesitic composition of the continental crust (Kelemen et al., 2014; Rudnick and Gao, 2014). 505

506

#### 507 5.3 Age of the depleted mantle

The best-fit DM ages inferred from our model (Fig. 6-8) have mean depletion ages between 0.5 Ga ( $\pm$  0.25, 1 S.D.) to 1.5 Ga ( $\pm$  0.75, 1 S.D.). This range of depletion ages is similar to Sm-Nd model ages of abyssal peridotites, which are generally <1 Ga, but with sporadic ages up to 2.7 Ga (Brunelli et al., 2018; Cipriani et al., 2004; Mallick et al., 2014; Stracke et al., 2011; Warren et al., 2009). Rhenium-Os model ages ( $T_{RD}$ ) also have a similar distribution, with an average of 0.77 Ga  $\pm$  0.93 (2 S.D.) (Day et al., 2017). The Sm-Nd and Re-Os ( $T_{RD}$ ) model ages of abyssal peridotites are thus consistent with the young DM ages inferred from our model (Fig. 6-8).

The latter rely on the assumption that the existing DM estimates used for the modeling (Salters and Stracke, 2004; Workman and Hart, 2005) accurately capture the average composition of the DM. As discussed below (section 5.4), these estimates are expected to somewhat underestimate the average incompatible element depletion of the DM. Using a more incompatible element depleted DM end-member, the model presented here yields younger, less variable ages for the DM. The best-fit age for the DM (<1.5 Ga, Fig. 5-7) with the parameters 521 used in our model (Salters and Stracke, 2004; Workman and Hart, 2005) should thus be 522 considered an upper bound.

523 If the depletion ages inferred from our model (<1.5 Ga) reflect the time of the last melting 524 event the mantle has experienced, they constrain the rate of mantle processing through melting 525 regions in the shallow mantle. This so-called mantle processing rate is proportional to the mass 526 fraction of mantle that has become depleted by partial melting and the number of melting events 527 experienced by its individual parts. The mantle processing rate therefore determines the present 528 extent of mantle depletion, and is a proxy for the overall rate of silicate earth differentiation (e.g., 529 Stracke, 2018; Tackley, 2015). Hence, if the young DM ages of our model indicate that the 530 mantle re-melts within <1.5 Ga, the processing rate of the mantle would be considerably faster 531 than inferred from oceanic crust production rates (e.g., Stracke, 2018; Stracke et al., 2003; 532 Tackley, 2015). If so, the resulting large extent of mantle depletion must be counteracted, on a 533 global scale, by a relatively large return flux of subducted oceanic and continental crust.

534

#### 535

## 5.4 The nature of the depleted mantle

536 If the fast mantle processing rate inferred from our modeled DM ages is correct, most of Earth's 537 mantle should have experienced more than one melting event over Earth's history, and should 538 thus be highly depleted. Mixing arrays with melts from highly depleted mantle and incompatible 539 element enriched sources (OC, UCC, LCC), however, have a pronounced curvature (Fig. 3). On 540 a global scale, this scenario would produce a Ce-Nd-Hf mantle array with a steeper slope than 541 observed (Fig. 3, cf. Willig and Stracke, 2019). Apparently therefore, a DM with moderate 542 isotopic, and thus time-integrated average incompatible element depletion is required to 543 reproduce the global Ce-Nd-Hf mantle array (Fig. 3, 4).

544 For evaluating the composition of the average DM involved in our model -- and in MORB and 545 OIB generation in general- it must be considered that DM is intrinsically heterogeneous. 546 Generating DM by partial melting at mid ocean ridges, for example, strongly depletes the mantle by high degree melting under the ridge axis, but only slightly depletes the mantle by low degree melting off-axis (Langmuir et al., 1992; Fig. 10). Strongly incompatible elements (D < 0.01) are quantitatively extracted at low degrees ( $\leq$ 3-5%) of melting and thus retain significant concentrations only in the lower parts of the residual mantle, i.e., those that have experienced the smallest degree of melting (Fig. 10). Formation of DM therefore invariably produces a range of intrinsically heterogeneous mantle materials.

553 Incompatible element abundances and ratios in aggregate melts from a heterogeneous DM, 554 including isotope ratios of the lithophile radiogenic isotope systems (Sm-Nd, Rb-Sr, Pb-U, Th-U, 555 and here La-Ce), thus reflect the average incompatible element inventory of a range of highly to 556 slightly incompatible element depleted materials. However, although highly depleted materials 557 constitute most of the DM (Fig. 10), they contain only small amounts of incompatible elements 558 (e.g., Ce, Nd, Hf). The total incompatible element inventory of a heterogeneous DM is thus 559 dominated by its least depleted parts. Incompatible element ratios (and concentrations) of this 560 average DM are therefore similar to the more moderately depleted materials of the underlying 561 spectrum of variably depleted mantle. The average La/Ce for the residual mantle shown in Fig. 562 10, for example, is  $0.63 \times \text{La/Ce}_0$ , where  $\text{La/Ce}_0$  is the La/Ce of the mantle before partial 563 melting. This apparent La/Ce corresponds to the residual mantle produced by 1.6% single-stage 564 fractional melting, compared to the average of 6% for the example shown in Fig. 10. Hence, if 565 the apparent moderate extent of depletion reflected in melts from a heterogeneous DM is taken 566 to reflect bulk source composition, rather than the average of a range of heterogeneous 567 materials, the extent of melt extraction and incompatible element depletion of the DM source is 568 underestimated.

569 Consequently, the apparent moderate depletion of the DM indicated by our model is 570 deceiving, because the moderately depleted melts required to reproduce the slope of the global 571 Ce-Nd-Hf reflect a biased average of an intrinsically heterogeneous DM, rather than a 572 moderately incompatible element depleted and compositionally homogeneous DM. The total 573 compositional range of DM involved in generating the observed Ce-Nd-Hf mantle array is574 therefore probably larger than inferred from our model.

Note that underestimating the extent of depletion and compositional variability of DM from incompatible element and isotope ratios in MORB also presents an obstacle for estimating DM composition (Salters and Stracke, 2004; Workman and Hart, 2005), and for inferring the correct extent of depletion and mass of DM from geochemical and isotopic mass balance (e.g., Allègre et al., 1979; Hofmann, 1986).

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#### 584 6. Conclusions

585 Cerium isotope ratios, in addition to the conventional lithophile Nd-Hf (and Sr-Pb) isotope 586 ratios are a unique tool for identifying the role of variably depleted mantle in MORB and OIB 587 generation, and provide valuable constraints on the nature of recycled oceanic and different 588 types of continental crust (i.e., upper vs. lower continental crust) components in Earth's mantle.

589 The combined Ce-Nd-Hf isotope ratios, for example, provide better constraints on the nature 590 and abundance of DM in MORB and OIB sources. On a local scale, the pronounced curvature 591 of mixing arrays towards different DM components allows discerning variably depleted mantle 592 (Fig. 3, 4). The slope and dispersion of the global Ce-Nd-Hf array is also heavily influenced by 593 the average DM involved in MORB and OIB generation (Fig. 2-5). Reproducing the slope of the 594 global Ce-Nd-Hf array by repeated sampling of an underlying variable population of mantle 595 components (DM, recycled OC, UCC, LCC) in a "Monte Carlo" approach requires melts from 596 average DM that are only moderately depleted. However, because these melts do not reflect 597 melting of a single homogeneous DM, but rather the total incompatible element inventory of a 598 much greater range of inherently heterogeneous DM, the extent of depletion and variability of599 the DM is larger than indicated by the modeled melts.

600 Our model results further show that it is crucial to consider the composition of recycled bulk 601 OC, i.e., both the extrusive basalts and intrusive gabbroic rocks of the lower oceanic crust, for 602 evaluating the role recycled OC on mantle compositions and crust-mantle cycling in general (cf. 603 Stracke et al., 2003). The low dispersion of the global Ce-Nd-Hf arrays suggests that the total 604 compositional variance of the oceanic crust is not reflected in MORB and OIB sources. One 605 reason for this effect is that individual variability averages out on the scale of the bulk oceanic 606 crust, and during subsequent sampling of recycled oceanic crust by partial melting. Another 607 reason is that the ubiquitous presence of recycled continental crust, which is required to account 608 for the full range of Ce-Nd-Hf isotope ratios in oceanic basalts (cf. Chauvel et al., 2008), masks 609 the signatures of recycled oceanic crust. Nevertheless, uncertainties in bulk composition of 610 recycled oceanic crust are a major source of uncertainty for evaluating the isotopic evolution of 611 Earth's mantle, and crust mantle cycling in general, especially considering that subducted oceanic 612 crust is by far the greatest mass flux into the mantle.

Owing to the distinct LREE patterns of upper and lower continental crust, combined Ce-Nd-Hf systematics can also distinguish between different recycled continental crust components in MORB and OIB sources, on both local and global scale. Our model result indicate that the return flux from the continental crustal into Earth's mantle is dominated by lower continental crust (e.g., Willbold and Stracke, 2006, 2010), consistent with recent models for developing and maintaining an average andesitic composition of the continental crust (Kelemen et al., 2014; Rudnick and Gao, 2014).

Overall, therefore, Ce in addition to other lithophile isotope ratios (Sr-Nd-Hf-Pb) provide
valuable constraints on the nature of MORB and OIB source components, on crust-mantle
exchange, and silicate Earth evolution in general.

624

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## 633 Tables

**Table 1:** Input parameters used to generate Fig. 5

attribute	mean	S.D.	min	max
Continental crust residence time [Ga]	0.18	0.3	0	4
Crustal material in mantle residence time [Ga]	2	1	0	4
Age DM [Ga]	1	0.5	0	4
Source LCC/(UCC+LCC)	0.72	0.02	0	1
Source CC/OC	0.07	0.1	0	1
Source OC/DM	0.19	0.01	0.01	3/7
Number basalts modeled	150	-	-	-

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Figure 1: (a) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>138</sup>Ce/<sup>136</sup>Ce, (b) <sup>143</sup>Nd/<sup>144</sup>Nd versus <sup>138</sup>Ce/<sup>136</sup>Ce, and (c) 644 <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>143</sup>Nd/<sup>144</sup>Nd diagram for OIB (red dots) and MORB (green squares). Panel (a) 645 and (b) show only data for which combined Ce, Nd and Hf isotopic data exists (supplementary 646 647 table 1), panel (c) includes literature data from the compilation provided in Stracke (2012, gray 648 dots). The vertical and horizontal line denote the isotope composition of the chondritic uniform 649 reservoir (CHUR) (Bouvier et al., 2008; Willig and Stracke 2019; Workman and Hart, 2005). An 650 ordinary least square (OLS) bisector fit (c.f. Isobe et al., 1990) is shown as a black line with a 2 651 S.E. envelope in gray. In panel (c) the OLS bisector fit for the literature data (Stracke, 2012, grey 652 dots) is shown in dark grey.

653

**Figure 2:** Isotopic evolution trends for depleted mantle (DM, green), lower continental crust (LCC, orange), upper continental crust (UCC, red) and bulk oceanic crust (OC, blue) in (a)  $^{176}$ Hf/<sup>177</sup>Hf versus  $^{138}$ Ce/<sup>136</sup>Ce, (b)  $^{143}$ Nd/<sup>144</sup>Nd versus  $^{138}$ Ce/<sup>136</sup>Ce, and (c)  $^{176}$ Hf/<sup>177</sup>Hf versus  $^{143}$ Nd/<sup>144</sup>Nd diagrams, including the OIB and MORB data shown in Fig. 1. The initial isotopic composition of the various end-members is the isotopic composition of the upper mantle at the time of formation, which is approximated by evolving linearly from BSE at 4.56 Ga (i.e.,  $\epsilon$ Ce=

660  $\epsilon$ Nd=  $\epsilon$ Hf=0) to  $\epsilon$ Ce=-1.2,  $\epsilon$ Nd=9.2,  $\epsilon$ Hf=18.4 to present time. The decay constants used are

- 661 given in supplementary table 2.
- 662

**Figure 3:** Diagrams of (a) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>138</sup>Ce/<sup>136</sup>Ce, (b) <sup>143</sup>Nd/<sup>144</sup>Nd versus <sup>138</sup>Ce/<sup>136</sup>Ce, and (c) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>143</sup>Nd/<sup>144</sup>Nd show mixing curves between melts from different depleted mantle components in green and a recycled crustal component, which is a 9:1 mixture of recycled bulk oceanic crust (OC; White and Klein, 2014, blue) and recycled upper continental 667 crust (UCC, Rudnick and Gao, 2014, red). The recycling time of oceanic and continental crust is
668 1.5 Ga. Prior to recycling, the continental crust had a crustal residence time of 1 Ga, hence was
669 formed at 2.5 Ga. The depletion age of the DM is assumed to be 1 Ga, i.e., the recycled
670 components are incorporated during mantle processing into depleted mantle. Tick marks on the
671 mixing curves represent the fraction of crustal melt involved in melt mixing.

The depleted components have compositions taken from Salters and Stracke (2004, S&S) or Workman and Hart (2005, W&H). Also shown is a modeled depleted mantle corresponding to a residual mantle column produced by partial melting of DM composition to 12% (DM, Salters and Stracke, 2004). The latter thus corresponds to highly refractory mantle.

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Figure 4: Diagrams of (a)  ${}^{176}$ Hf/ ${}^{177}$ Hf versus  ${}^{138}$ Ce/ ${}^{136}$ Ce, (b)  ${}^{143}$ Nd/ ${}^{144}$ Nd versus  ${}^{138}$ Ce/ ${}^{136}$ Ce, and 677 (c) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>143</sup>Nd/<sup>144</sup>Nd show mixing curves between melts from recycled oceanic 678 crust (blue) and upper continental crust (red) and DM (green). The proportion of melts from the 679 680 DM to melts from a recycled component (here consisting of OC and UCC) is kept constant for 681 each individual mixing curve, but varies between the different mixing curves. The variability 682 along each individual mixing trend is caused by varying the proportion of OC and UCC within the enriched source component. Increasing the proportion of melts from the DM elevates the 683 684  $\epsilon$ Nd and  $\epsilon$ Hf relative to  $\epsilon$ Ce, and limits the extent of the mixing curve towards the enriched

685 source component. Different proportions of melts from a single DM component thus result in 686 parallel trends in Ce-Nd and Ce-Hf space, which may lead to increasing isotopic variance at the 687 depleted (DM) end of the global isotope arrays on these diagrams.

688

Figure 5: Diagrams of (a) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>138</sup>Ce/<sup>136</sup>Ce, (b) <sup>143</sup>Nd/<sup>144</sup>Nd versus <sup>138</sup>Ce/<sup>136</sup>Ce, and
(c) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>143</sup>Nd/<sup>144</sup>Nd show an exemplary model output for the input parameters
given in Table 1. Round grey dots (150 in total) represent the modeled oceanic basalts; white-

692 filled black squares and dots are the MORB and OIB data given in supplementary table 1.693 Individual mantle end-member isotope compositions used in the modeling are shown in color.

694

695 Figure 6: The diagram shows (contoured) a goodness of fit after optimization for different 696 values for the age of the depleted mantle (mean  $\pm$  1S.D.) = [0.5  $\pm$ 0.25, 1  $\pm$ 0.5, 1.5  $\pm$ 0.75, 2  $\pm$ 1, 697 2.5  $\pm$ 1.25, 3  $\pm$ 1.5, 3.5  $\pm$ 1.75] and recycling time of enriched components in the mantle [0.5 698  $\pm 0.25$ , 1  $\pm 0.5$ , 1.5  $\pm 0.75$ , 2  $\pm 1$ , 2.5  $\pm 1.25$ , 3  $\pm 1.5$ , 3.5  $\pm 1.75$ ]. Free model parameters are the 699 sampled OC/DM, sampled CC/OC, sampled LCC/(LCC+UCC), and the crustal storage time. 700 The goodness of fit score is optimized using a gradient descent algorithm that minimizes the 701 average squared distance between a model point and the closest data point (for details see 702 supplementary material), using 60 modeled oceanic basalts for each model run.

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704 Figure 7: The diagram shows optimized values for sampled LCC/(LCC+UCC) for different 705 values for the age of the depleted mantle (mean  $\pm$  1S.D.) = [0.5  $\pm$ 0.25, 1  $\pm$ 0.5, 1.5  $\pm$ 0.75, 2  $\pm$ 1, 706 2.5 ±1.25, 3 ±1.5, 3.5 ±1.75] and recycling time of enriched components in the mantle [0.5 707  $\pm 0.25$ , 1  $\pm 0.5$ , 1.5  $\pm 0.75$ , 2  $\pm 1$ , 2.5  $\pm 1.25$ , 3  $\pm 1.5$ , 3.5  $\pm 1.75$ ]. Free parameters are the sampled 708 OC/DM, sampled CC/OC, sampled LCC/(LCC+UCC), and the crustal storage time. The 709 goodness of fit score is optimized using a gradient descent algorithm that minimizes the average 710 squared distance between a model point and the closest data point (supplementary material), 711 using 60 modeled oceanic basalts for each model run. Also shown in gray is the parameter area 712 for which sufficiently good goodness of fit scores resulted in Fig. 6.

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Figure 8: The diagram shows a goodness of fit after optimization for different values for the age
of the depleted mantle (mean ± 1S.D.) = [0.5 ±0.25, 1 ±0.5, 1.5 ±0.75, 2 ±1, 2.5 ±1.25, 3 ±1.5,
3.5 ±1.75] and the recycled LCC/(UCC+LCC) sampled by oceanic basalts = [0 ±0, 0.1 ±0, 0.2
±0, 0.3 ±0, 0.4 ±0, 0.5 ±0, 0.6 ±0, 0.7 ±0, 0.8 ±0, 0.9 ±0, 1 ±0]. Free parameters are the

718 sampled OC/DM, sampled CC/OC, crustal storage time, and the recycling time of crustal 719 material in the mantle. The goodness of fit score is optimized using a gradient descent algorithm 720 that minimizes the average squared distance between a model point and the closest data point 721 (supplementary material), using 60 modeled oceanic basalts for each model run.

722

Figure 9: Diagrams of (a)  ${}^{176}$ Hf/ ${}^{177}$ Hf versus  ${}^{138}$ Ce/ ${}^{136}$ Ce, (b)  ${}^{143}$ Nd/ ${}^{144}$ Nd versus  ${}^{138}$ Ce/ ${}^{136}$ Ce, and 723 (c) <sup>176</sup>Hf/<sup>177</sup>Hf versus <sup>143</sup>Nd/<sup>144</sup>Nd show the isotopic distribution of 2Ga MORB (green) and 724 725 lower oceanic crustal gabbro (blue). The modeled MORB (n = 2459) and gabbro (n = 286) 726 isotope ratios are obtained by isotopic evolution from DM at 2Ga (section 5.2) using the La/Ce, 727 Sm/Nd and Lu/Hf from the MORB compilation by Gale et al. (2013), and Gabbro trace 728 element data obtained from the PetDB (supplementary table 3). The large compositional 729 variability of MORB (and gabbro) results in a large range of isotope compositions that covers 730 the entire Ce-Nd-Hf mantle array (white-filled black squares: MORB, white-filled black dots: 731 OIB).

732

733 Figure 10: The diagram on the left shows a schematic triangular melting area beneath a mid 734 ocean ridge, assuming a maximum extent of melting of 12% directly underneath the ridge axis. 735 Assuming that incompatible elements are completely exhausted after 5% of melting (i.e., at about 736 half the height) 58% of the mantle that returns into the deeper mantle is highly depleted (for 737 exhaustion of incompatible elements at 3% of melting it becomes ca. 75%). Effectively, this 738 means that highly depleted mantle, almost devoid of incompatible elements is the volumetrically 739 dominant component of depleted mantle. The diagram on the right shows how the La/Ce, 740 Sm/Nd and Lu/Hf ratios change from bottom to top of the residual mantle, i.e., with increasing 741 degree of melting. La/Ce decreases from La/Ce ~  $0.97 \times La/Ce_0$  (where La/Ce<sub>0</sub> is the La/Ce of the unmelted mantle) in the lower part of the residual mantle (0.1% melt) to La/Ce  $\sim 0.02 \times$ 742

La/Ce<sub>0</sub> in the upper part of the residual mantle (12% melt, black line). For a uniform distribution 743 744 of the degree of melt in the residual mantle from 0.1-12%, the average La/Ce for the residual mantle is  $\sim 0.63 \times La/Ce_0$  (black dashed line). This average extent of depletion corresponds to 745 the residue of 1.6% single-stage fractional melting, rather than the actual 6% produced. Sm and 746 747 Nd are more compatible and yield average Sm/Nd of  $1.40 \times \text{Sm/Nd}_0$  corresponding to ~4% 748 melt depletion (blue lines). For Lu/Hf, the average is  $1.51 \times Lu/Hf_0$  corresponding to ~4.9% 749 fractional melting (red lines). Hence, the average of a heterogeneous DM is increasingly biased 750 towards the least depleted DM components for increasingly more incompatible elements ratios. 751 This bias is transposed to partial melts, i.e., incompatible element ratios in partial melts from a 752 heterogeneous DM are biased towards the least depleted parts of the DM. Bulk partition 753 coefficients for melting of spinel peridotite are  $D_{La} = 0.013$ ,  $D_{Ce} = 0.021$ ,  $D_{Nd} = 0.046$ ,  $D_{Sm} =$ 754 0.075,  $D_{Hf} = 0.073$ ,  $D_{Lu} = 0.186$  (Stracke and Bourdon, 2009).

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**LCC** 2.5 Ga

1.3369

<sup>138</sup>Ce/<sup>136</sup>Ce

1.3370

1.3368

0.5120

0.5115

1.3367













