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### Renewed activity from the magnetar CXOU J164710.2-455216

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**Publication date**

2018

**Document Version**

Final published version

**Published in**

The astronomer's telegram

**License**

Unspecified

[Link to publication](#)

**Citation for published version (APA):**

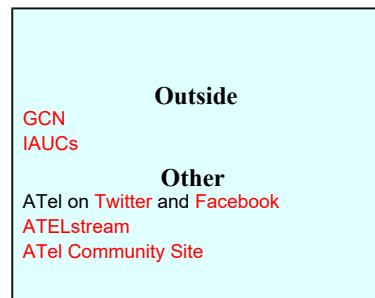
Borghese, A., Coti Zelati, F., Rea, N., Esposito, P., Mereghetti, S., & Pintore, F. (2018). Renewed activity from the magnetar CXOU J164710.2-455216. *The astronomer's telegram*, 11264. <http://www.astronomerstelegram.org/?read=11264>

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## Renewed activity from the magnetar CXOU J164710.2-455216

ATel #11264; *Alice Borghese (U. Amsterdam), Francesco Coti Zelati (IEEC-CSIC), Nanda Rea (IEEC-CSIC, U. Amsterdam), Paolo Esposito (U. Amsterdam), Sandro Mereghetti, Fabio Pintore (INAF/IASF-Mi), on behalf of a larger collaboration*  
on 6 Feb 2018; 16:02 UT

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Subjects: X-ray, Transient, Magnetar

Referred to by ATel #: [11270](#)

On 2018 February 5 at 19:27:11 UT, Swift BAT detected a new burst from a direction consistent with the magnetar CXOU J164710.2-455216 (trigger=808755; Barthelmy et al. 2018, GCN #[22389](#)). Moreover, INTEGRAL detected two short bursts from CXOU J1647 on 2018 February 6 at 07:25:58 UT and 12:33:36 UT. The source was recovering from a major outburst, started on 2017 May 16 (trigger=753085; D'Ai et al. 2017, GCN #[21095](#)), when it reached a 0.5-10 keV absorbed X-ray flux of  $8(+2)(-4)\text{e-}12 \text{ erg s}^{-1} \text{ cm}^{-2}$ , a factor of about 10 larger than the quiescent value.

The XRT began observing the field about 40 s after the BAT trigger for an exposure time of 1.2 ks. The time-averaged spectrum could be well described by an absorbed power law with a photon index of 1.8(3) and a column density of  $3.1(7)\text{e}22 \text{ cm}^{-2}$  (errors are quoted at 1sigma confidence level; see also Barthelmy et al. 2018, GCN #[22389](#)). The source had an averaged count rate of about 0.3 cts/s in the energy band 0.5-10 keV, corresponding to an absorbed flux of  $2.3(2)\text{e-}11 \text{ erg s}^{-1} \text{ cm}^{-2}$ .

We note that the light curve was highly variable and at the end of the observation (about 10 minutes after the trigger) the absorbed 0.5-10 keV flux was approximately  $1.2(+0.1)(-0.4)\text{e-}11 \text{ erg s}^{-1} \text{ cm}^{-2}$ , substantially higher than in the last Swift observation performed before the event, on 2018 January 28 when it was around  $2\text{e-}12 \text{ erg s}^{-1} \text{ cm}^{-2}$ .

We are monitoring the source since its outburst onset in 2017 May with Swift, Chandra and NuSTAR observations (ATEL #[10878](#)). Swift observations are already planned till the end of February, and new simultaneous Chandra and NuSTAR observations will be performed within the next two weeks.

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