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## INTEGRAL observation of MAXI J0911-655

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on 26 Apr 2016; 05:59 UT

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Subjects: X-ray, Binary, Transient

Referred to by ATel #: [9738](#), [10425](#)

The recently discovered transient MAXI J0911-655 (Atels #[8872](#), #[8884](#), #[8914](#), #[8971](#)) was observed during a dedicated INTEGRAL campaign performed from 2016 April 23 at 23:43 to April 25 at 06:51 UTC.

The source was detected by IBIS/ISGRI at a significance of 16 sigma in the 20-40 keV energy range and 11 sigma in the 40-80 keV energy range (effective exposure time 46 ks). The estimated fluxes from the mosaics were of  $9.3^{+/-0.6}$  mCrab (i.e.  $\sim 7.0 \times 10^{-11}$  erg/cm $^2$ /s) and  $10.7^{+/-0.9}$  mCrab (i.e.  $\sim 7.4 \times 10^{-11}$  erg/cm $^2$ /s), respectively.

The source was also detected by the two JEM-X instruments at a significance of 11 sigma in the 3-10 keV energy range and about 4 sigma in the 10-20 keV energy range (effective exposure time 58 ks). The corresponding fluxes estimated from the JEM-X mosaics were  $6.8^{+/-0.6}$  mCrab (i.e.  $\sim 9.6 \times 10^{-11}$  erg/cm $^2$ /s) and  $5.5^{+/-1.4}$  mCrab (i.e.  $\sim 5.1 \times 10^{-11}$  erg/cm $^2$ /s), respectively. These values might be significantly different once the JEM-X off-line energy calibration will be carried out.

By using the INTEGRAL near real time data, we performed a preliminarily broad-band spectral fits of the ISGRI and JEM-X data together. The spectrum of MAXI J0911-655 could be well fit (reduced  $\chi^2/\text{d.o.f.} = 0.7/36$ ) by using a simple absorbed power-law with photon index of  $2.4^{+/-0.2}$  (we fixed the absorption column density to  $4 \times 10^{21}$  cm $^{-2}$ ), see Atel #[8884](#)). The flux measured from the spectral fit is  $3.0 \times 10^{-10}$  erg/cm $^2$ /s in the 3-100 keV energy range.

A more detailed spectral analysis will be performed on the consolidated INTEGRAL data including refined calibrations especially for the JEM-X units.

No thermonuclear bursts have been detected in the JEM-X lightcurve, and thus the nature of the

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accreting compact object in MAXI J0911-655 remains debated.

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