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	Hakobyan, A. A.; Karapetyan, A. G.; Barkhudaryan, L. V.; Mamon, G. A.; Kunth, D.; et al.
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Properties of 500 SNe and their 419 hosts (Hakobyan+, 2016) J/MNRAS/456/2848

Supernovae and their host galaxies. III. The impact of bars and bulges on the radial distribution of supernovae in disc galaxies. Hakobyan A.A., Karapetyan A.G., Barkhudaryan L.V., Mamon G.A., Kunth D.,

Petrosian A.R., Adibekyan V., Aramyan L.S., Turatto M. <Mon. Not. R. Astron. Soc. 456, 2848 (2016)> =<u>2016MNRAS.456.2848H</u> (SIMBAD/NED BibCode)

ADC\_Keywords: Supernovae; Galaxies, optical; Galaxy catalogs;

Morphology; Positional data; Surveys

Keywords: supernovae: general - galaxies: spiral - galaxies: stellar content -

galaxies: structure

#### Abstract:

We present an analysis of the impact of bars and bulges on the radial distributions of the different types of supernovae (SNe) in the stellar discs of host galaxies with various morphologies. We use a well-defined sample of 500 nearby (≤100Mpc) SNe and their low-inclined (i≤60deg) and morphologically non-disturbed S0-Sm host galaxies from the Sloan Digital Sky Survey. We find that in Sa-Sm galaxies, all core-collapse (CC) and vast majority of SNe Ia belong to the disc, rather than the bulge component. The radial distribution of SNe Ia in SO-SO/a galaxies is inconsistent with their distribution in Sa-Sm hosts, which is probably due to the contribution of the outer bulge SNe Ia in SO-SO/a galaxies. In Sa-Sbc galaxies, the radial distribution of CC SNe in barred hosts is inconsistent with that in unbarred ones, while the distributions of SNe Ia are not significantly different. At the same time, the radial distributions of both types of SNe in Sc-Sm galaxies are not affected by bars. We propose that the additional mechanism shaping the distributions of Type Ia and CC SNe can be explained within the framework of substantial suppression of massive star formation in the radial range swept by strong bars, particularly in early-type spirals. The radial distribution of CC SNe in unbarred Sa-Sbc galaxies is more centrally peaked and inconsistent with that in unbarred Sc-Sm hosts, while the distribution of SNe Ia in unbarred galaxies is not affected by host morphology. These results can be explained by the distinct distributions of massive stars in the discs of early- and late-type spirals.

## Description:

The full database of 500 individual SNe (SN designation, type, and offset from host galaxy nucleus) and their 419 host galaxies (galaxy SDSS designation, distance, morphological type, bar, corrected g-band  $D_{25}$ , a/b, PA, and inclination).

## File Summary:

FileName Lrecl Records Explanations ReadMe 20 This file table.dat 115 500 Properties of 500 SNe and their 419 hosts

## See also:

B/sn : Asiago Supernova Catalogue (Barbon et al., 1999-) J/A+A/544/A81 : Supernovae and their hosts in SDSS DR8 (Hakobyan+, 2012) J/MNRAS/444/2428 : Disturbance levels of SNe host galaxies (Hakobyan+, 2014) http://www.sdss.org : SDSS Home Page

# Byte-by-byte Description of file: table.dat

Bytes Format Units Label Explanations 6 Α6 SN Supernova (SN) designation 9- 16 SN type (from spectroscopy) Α8 ---Type 21- 26 F6.2 <u>arcsec</u> oRA SN offset right ascension from host galaxy nucleus, in the  ${\sf E}/{\sf W}$  direction [E/W] E/W direction of SN offset oRA 28 Δ1 n oRA 32- 37 SN offset declination from host galaxy F6.2 arcsec oDE nucleus, in the N/S direction 39 Δ1 --n oDE [N/S] N/S direction of SN offset oDE 41- 59 Δ19 ---Gal SN host galaxy identification (1) 66- 71 74- 77 F6.3 <u>Mpc</u> Dist Host galaxy distance Α4 MType Host galaxy morphological type 81 Α1 [:?] Uncertainty flag on morphology (2)--u\_MType 84 Δ1 ---Bar [B] Presence of bar in the host galaxy u Bar 87 [:?] Uncertainty flag on presence of bar (3) Δ1 ---91-97 F7.2 arcsec Diam Host galaxy diameter (4)100-104 Host galaxy axial ratio F5.3 a/b 107-109 13 <u>deg</u> PΑ Host galaxy position angle 112-115 F4.1 Tnc Host galaxy inclination <u>deg</u>

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Note (1): JHHMMSS.ss+DDMMSS.s for SDSS JHHMMSS.ss+DDMMSS.s

An alternative name is mentioned when SDSS identification was not possible to obtain. Unidentified (anonymous) galaxies are listed with the letter A.

Note (2): Uncertainties in host galaxy morphological type are marked by ":" and "?" (highly uncertain).

Note (3): Uncertainties in host galaxy bar are marked by ":" and "?" (highly uncertain).

Note (4): Corrected major axis in arcsec (isophotal level of 25mag/arcsec² in the SDSS g-band) of the host galaxy.
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## Acknowledgements:

Artur Hakobyan, hakobyan(at)bao.sci.am

#### References:

(End) Artur Hakobyan [BAO], Patricia Vannier [CDS] 26-Feb-2016

The document above follows the rules of the <u>Standard Description for Astronomical Catalogues</u>; from this documentation it is possible to generate **f**77 program to load files <u>into arrays</u> or <u>line by line</u>

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