

Publication Year	2016
Acceptance in OA@INAF	2020-05-13T12:24:27Z
Title	VizieR Online Data Catalog: Hi-GAL. inner Milky Way: +68>=l>=70 (Molinari+, 2016)
Authors	MOLINARI, Sergio; Schisano, E.; ELIA, Davide Quintino; Pestalozzi, M.; TRAFICANTE, ALESSIO; et al.
Handle	http://hdl.handle.net/20.500.12386/24792
Journal	VizieR Online Data Catalog



Portal Simbad VizieR Aladin X-Match Other Help

J/A+A/591/A149 Hi-GAL. inner Milkv Wav: +68≥1≥70 (Molinari+, 2016)

Hi-GAL, the Herschel infrared Galactic Plane Survey: photometric maps and compact source catalogues. First data release for the inner Milky Way: +68°≥ 1 ≥70°. Molinari S., Schisano E., Elia D., Pestalozzi M., Traficante A., Pezzuto S., Swinyard B.M., Noriega-Crespo A., Bally J., Moore T.J.T., Plume R., Zavagno A., Di Giorgio A.M., Liu S.J., Pilbratt G.L., Mottram J.C., Russeil D., Piazzo L., Veneziani M., Benedettini M., Calzoletti L., Faustini F., Natoli P., Piacentini F., Merello M., Palmese A., Del Grande R., Polychroni D., Rygl K.L.J., Polenta G. Barlow M.J., Bernard J.-P., Martin P.G., Testi L., Ali B., Andre P Beltran M.T., Billot N., Carey S., Cesaroni R., Compiegne M., Eden D., Fukui Y., Garcia-Lario P., Hoare M.G., Huang M., Joncas G., Lim T.L., Lord S.D., Martinavarro-Armengol S., Motte F., Paladini R., Paradis D., Peretto N., Robitaille T., Schilke P., Schneider N., Schulz B., Sibthorpe B., Strafella F., Thompson M.A., Umana G., Ward-Thompson D., Wyrowski F. <Astron. Astrophys., 591, A149 (2016)>
=2016A&A...591A.149M (SIMBAD/NED Bi (SIMBAD/NED BibCode)

ADC_Keywords: Surveys; Milky Way; Infrared sources

Keywords: dust, extinction - infrared: ISM - stars: formation - Galaxy: disk methods: data analysis - techniques: photometric

Abstract:

We present the first public release of high-quality data products (DR1) from Hi-GAL, the Herschel infrared Galactic Plane Survey. Hi-GAL is the keystone of a suite of continuum Galactic plane surveys from the near-IR to the radio and covers five wavebands at 70, 160, 250, 350 and 500μm, encompassing the peak of the spectral energy distribution of cold dust for 8≲T≲50K. This first Hi-GAL data release covers the inner Milky Way in the longitude range 68° ≳1≳-70° in a |b|≤1° latitude strip.

Photometric maps have been produced with the ROMAGAL pipeline, which optimally capitalizes on the excellent sensitivity and stability of the bolometer arrays of the Herschel PACS and SPIRE photometric cameras. It delivers images of exquisite quality and dynamical range, absolutely calibrated with Planck and IRAS, and recovers extended emission at all wavelengths and all spatial scales, from the point-spread function to the size of an entire 2°x2° "tile" that is the unit observing block of the survey. The compact source catalogues were generated with the CuTEx algorithm, which was specifically developed to optimise source detection and extraction in the extreme conditions of intense and spatially varying background that are found in the Galactic plane in the thermal infrared.

 $\mbox{Hi-GAL}$ DR1 images are cirrus noise limited and reach the $1\sigma\mbox{-rms}$ predicted by the Herschel Time Estimators for parallel-mode observations at 60"/s scanning speed in relatively low cirrus emission regions. Hi-GAL DR1 images will be accessible through a dedicated web-based image cutout service. The DR1 Compact Source Catalogues are delivered as single-band photometric lists containing, in addition to source position, peak, and integrated flux and source sizes, a variety of parameters useful to assess the quality and reliability of the extracted sources. Caveats and hints to help in this assessment are provided. Flux completeness limits in all bands are determined from extensive synthetic source experiments and greatly depend on the specific line of sight along the Galactic plane because the background strongly varies as a function of Galactic longitude. Hi-GAL DR1 catalogues contain 120581, 291858, 280143, 161946, and 85811 compact sources in the five bands.

Description:

This is the first public data release of high-quality products from the Herschel Hi-GAL survey. The release comes two years after the end of the Herschel observing campaign and is the result of extensive testing of the data reduction and extraction procedures created by members of the Hi-GAL consortium. The complexity and the large variation of the background conditions in all Herschel wavelength bands makes source extraction on the Galactic plane a challenging task. With Hi-GAL DR1, we provide access (http://vialactea.iaps.inaf.it) through a cutout service to high-quality images and compact source catalogues for the Galactic plane at 70, 160, 250, 350, and 500um in the region $68^{\circ} \ge 1 \ge -70^{\circ}$ and $|b| \le 1^{\circ}$.

File Summary:

higalred.dat

higalpsw.dat

578

558

291858

280143

FileName Lrecl Records Explanations ReadMe 80 120581 higalblu.dat 578 Blue (PACS 70um) band HIGAL Herschel catalog

Red (PACS 160um) band HIGAL Herschel catalog

PSW (SPIRE 250um) band HIGAL Herschel catalog

higalpmw.dat higalplw.dat bigalplw.dat tableb1.dat
558 161946 PMW (SPIRE 350um) band HIGAL Herschel catalog PLW (SPIRE 500um) band HIGAL Herschel catalog Clumps or clusters of saturated pixels, crossmatched with IRAS point sources and RMS sources

See also:

J/A+A/526/A151	Hi-Gal sources distance determination (Russeil+, 2011)
J/A+A/549/A130	Cores in IR Dark Clouds for 300≤l≤330 (Wilcock+, 2012)
J/A+A/549/A130	YSOs in Herschel-Hi-GAL survey (Veneziani+, 2013)
J/ApJ/772/45	Hi-GAL: star formation in the third quadrant (Elia+, 2013)
J/A+A/579/A71	Infrared emission of young HII regions (Cesaroni+, 2015)
J/A+A/487/253	RMS survey: 13CO observations of YSOs (Urquhart+ 2008)
J/A+A/501/539	RMS survey. 6cm observations of YSOs (Urquhart+, 2009)
J/ApJ5/208/11	Red MSX Source Survey: massive protostars (Lumsden+, 2013)

http://vialactea.iaps.inaf.it : Via Lactea Home Page

Byte-by-byte Description of file: higalred.dat

Bytes	Format	Units	Label	Explanations
1- 6 8- 29	I6 A22		rowid Name	Sequential number Designation of the source
31- 40	F10.6	dea	GLON	(HIGALPXLLL.llll+B.bbbb) <u>(G1)</u> Galactic longitude
42- 50	F9.6	deg	GLAT	Galactic latitude
52- 56	F5.2	deg	e GLON	rms uncertainty on GLON
58- 62	F5.2	deg	e_GLAT	rms uncertainty on GLAT
64- 72	F9.5	deg	RAdeg	Right ascension (J2000)
74- 83	F10.6	<u>deg</u>	DEdeg	Declination (J2000)
85-109			Image	Atlas image name (ATLAS_IMAGE)
111-118		<u>pix</u>	Xpos	X position
120-127		<u>pix</u>	Ypos	Y position
129-134 136-141		<u>pix</u> pix	e_Xpos e_Ypos	rms uncertainty on Xpos rms uncertainty on Ypos
143-151		 <u>htv</u>	Source	Source identification (SOURCE_ID)
153-161	F9.3	<u>Jγ</u>	Fint	Source-integrated flux measured from the
133 101	13.5	<i>≥y</i> .	. 1	fitting process, uncorrected (FINT UNCORR)
163-175	F13.7	<u>Ју</u>	Fintc	Source-integrated flux measured from the
				fitting process after applying
				photometric corrections as a function of
				the source size, to account for source
477 404	50 3	_		non-Gaussianity and for scan speed (FINT)
177-184	F8.3	<u> 77</u>	e_Fintc	Uncertainty on the integrated flux
				computed by multiplicating the fitted source residual rms (RMS_TOTAL) by the
				fitted source area as estimated by FWHMA
				and FWHMB (ERR_FINT)
186-196	F11.3	MJy/sr	Fpeak	Source peak flux measured from the fitting
			•	process (FPEAK)
198-205	F8.3	<u>MJy/sr</u>	e_Fpeak	Uncertainty on the peak flux (ERR_FPEAK)
207-212	F6.2	arcsec	FWHMA	Full width at half maximum of the source
				along axis a of the elliptical Gaussian
214-221	F8.2	20000	FWHMB	as determined by fitting engine Full width at half maximum of the source
214-221	F0.2	<u>arcsec</u>	FWHMD	along axis b of the elliptical Gaussian
				as determined by fitting engine
223-229	F7.2	arcsec	e FWHMA	rms uncertainty on FWHMA (ERR_FWHMA)
231-237	F7.2	arcsec	e FWHMB	rms uncertainty on FWHMB (ERR_FWHMB)
239-244	F6.1	deg	PĀ	[] Position angle of the elliptical
				Gaussian (N E)
246-250	F5.1	<u>deg</u>	e_PA	rms uncertainty on PA
252-263	F12.3	<u>MJy/sr</u>	Bckg	Background value determined at the source
265 277	543 5			peak position (BACKGROUND)
265-277	F13.5	<u>MJy/sr</u>	Bckga	Coefficient a of the zero-order term of
				the background obtained by the fit at the source peak position (BACK ACOEFF)
279-290	F12 5	MJy/sr	Bckgb	Coefficient b of the first-order term x of
2,5 250	. 12.3	113 97 31	DERED	the background obtained by the fit at the
				source peak position (BACK_BCOEFF)
292-303	F12.5	MJy/sr/pix	Bckgc	Coefficient c of the first-order term y of
				the background obtained by the fit at the
				source peak position (BACK_CCOEFF)
305-315	F11.5	MJy/sr/pix2	Bckgd	Coefficient d of the second-order term x2
				of the background obtained by the fit at the source peak position (BACK_DCOEFF)
317-327	E11 5	MJy/sr/pix2	Rekae	Coefficient e of the second-order term y2
317-327	111.5	<u>1113/731 7 PIXZ</u>	DCKge	of the background obtained by the fit at
				the source peak position (BACK_ECOEFF)
329-339	F11.5	MJy/sr/pix2	Bckgf	Coefficient f of the second-order term xy
			•	of the background obtained by the fit at
				the source peak position (BACK_FCOEFF)
341-353	F13.5	<u>MJy/sr</u>	rms	Standard deviation, loc, of the residuals
				computed within the source area defined
				by FWHMA and FWHMB after subtracting the

355-367	F13.5	M]y/sr	rmss	best fit (RMS_TOTAL) Standard deviation, loc, of the residuals computed within the fitting window after subtracting the best fit, excluding both the pixels that belong to the source and the pixels belonging to other sources that fall within the fitting window
369-377			SNR	(RMS_SURROUND) Signal-to-noise ratio obtained by dividing FPEAK by the residual rms over a source area with FWHMA and FWHMB as semi-axes (SNR)
		Basic detect		nation
379-387			detX	Relevance of the source in the
389-397	F9.3		detY	second-derivative map along the x-axis defined as the ratio between the measured second derivative at source peak position and the adopted local threshold value (DET_X) Relevance of the source in the second-derivative map along the y-axis defined as the ratio between the measured second derivative at source peak position
				and the adopted local threshold value (DET_Y)
399-407	F9.3		detX45	Relevance of the source in the second-derivative map along the bisector of the xy-axis defined as the ratio between the measured second derivative at source peak position and the adopted
409-417	F9.3		detY45	local threshold value (DET_X45) Relevance of the source in the
				second-derivative map along the bisector of the yx-axis defined as the ratio between the measured second derivative at source peak position and the adopted local threshold value (DET Y45)
419-427	F9.3	MJy/sr/pix2	detlimX	Absolute value for the local detection limit threshold adopted for the second derivative along the x-axis coordinate (DETLIM X)
429-437	F9.3	MJy/sr/pix2	detlimY	Absolute value for the local detection limit threshold adopted for the second derivative along the y-axis coordinate
439-447	F9.3	MJy/sr/pix2	detlimx45	(DETLIM_Y) Absolute value for the local detection limit threshold adopted for the second derivative along the bisector of the
449-457	F9.3	MJy/sr/pix2	detlimy45	first and third quadrant (DETLIM_X45) Absolute value for the local detection limit threshold adopted for the second derivative along the bisector of the
459	I1		Clump	second and fourth quadrant (DETLIM_Y45) Flag for confusion at detection level
461-462	12		Ncomp	(CLUMP_FLAG) <u>(G3)</u> Number of Gaussian components used simultaneously in the fitting process
464-467	14	<u>pix</u>	Xcent	(NCOMP) (G4) The x-pixel coordinate of the centre of
469-472	14	<u>pix</u>	Ycent	the source fitting window in the original image (XCENT) The y-pixel coordinate of the centre of the source fitting window in the
474-475	Т2	niv	Xwin	original image (YCENT) Half-width size of the source fitting
7/4=4/3	14	<u>pix</u>	VMTII	window along x coordinate and centred at XCENT (XWINDOW)
477-478	12	<u>pix</u>	Ywin	Half-width size of the source fitting window along y coordinate and centred
480-481	I2		NCont	at YCENT (YWINDOW) Number of other sources falling inside the fitting window whose presence is
				taken into account at fitting stage. Not all those other sources might have
483-486	F4.2	pix	Cent	been fitted at the same time (NCONTAM) Maximum variation in pixels for adjustment of the fit centre with respect to the position of detection, measured as the distance between the latter and the brightest local (within three pixels) pixel in the fitting window (CENT_TOL)
488-491			DOF	Degrees of freedom of the source Gaussian fit
		Quality flag	gs	
493-507	F15.1		Chi2	chi ² determined by the fitting engine.
509-520	F12.2		Chi2O	Estimator of the fidelity between the fit and the data computed as

				phi= $(0(i)-F(i))^2/F(i)$, where $0(i)$ is the observed data in the i pixel of the fitting window and $F(i)$ is the fitted value in the same position (CHI2OPP)
522	I1		fitst	<pre>[0/4] Flag returned from the fitting engine (FIT_STATUS) (G5)</pre>
524-525	A2		Guess	[ABC 0123456789] Flag on quality of guessed source parameters as determined at the detection stage (GUESS_FLAG) (G6)
527-533	Α7		Group	[ABC 0123456789] Flag on quality of guessed source parameters as determined at the detection stage (GROUP_FLAG) (G6)
535	I1		Constr	Flag indicating the number of parameters that reached the tolerance limits allowed to the fit process (CONSTRAINS) (G8)
		D:-		

Racic	avtraction	information

		pasic extra	CCION INIC	or macton
537-543	F7.3		rd2dx	Ratio between the second-derivative value along x direction expected by the fitted model of the source and the second derivative derivative measured at the detection stage (RDETP2DX) (G9)
545-552	F8.3		rd2dy	Ratio between the second-derivative value along y direction expected by the fitted model of the source and the second derivative measured at the detection stage. (RDETP2DY)
554-560	F7.3		rd2dx45	Ratio between the second-derivative value along the bisector of the xy direction expected by the fitted model of the source and the second derivative measured at the detection stage. (RDETP2DX45)
562-568	F7.3		rd2dy45	Ratio between the second-derivative value along the bisector of the yx direction expected by the fitted model of the source and the second derivative measured at the detection stage. (RDETP2DY45)
570-575	A6		ovlap	Flag to indicate whether the source has been detected and extracted in one or more adjacent tiles (OVERLAP_FLAG) (G10)
577-578	A2		ovflux	Flag to indicate which flux values were adopted if detected and extracted in two adjacent tiles (OVFLUX_FLAG) (G11)

Byte-by-byte Description of file: $\underline{\text{higalpsw.dat}} \ \underline{\text{higalpmw.dat}} \ \underline{\text{higalplw.dat}}$

Bytes	Format	Units	Label	Explanations
1 6	T.C.			Consential comban
1- 6 8- 29			rowid	Sequential number
8- 29	AZZ		Name	Designation of the source (HIGALPXLLL.1111+B.bbbb) (G1)
31- 40	F10.6	doa	GLON	Galactic longitude
42- 50			GLON	Galactic longitude Galactic latitude
52- 57			e GLON	rms uncertainty on GLON
59- 64			e GLAT	rms uncertainty on GLAT
66- 74		<u>deg</u> deg	RAdeg	Right ascension (J2000)
	F10.6		DEdeg	Declination (J2000)
87-111		<u>ueg</u>	Image	Atlas image name (ATLAS IMAGE)
113-121			Xpos	X position
123-121			Ypos	Y position
133-131		pix	e Xpos	rms uncertainty on Xpos
140-145		pix	e_xpos e Ypos	rms uncertainty on Ypos
147-155		<u> </u>	Source	Source identification (SOURCE ID)
157-165			Fint	Source-integrated flux measured from the
137 103	13.3	2 <i>y</i> .	11110	fitting process, uncorrected (FINT_UNCORR)
167-174	F8.3	<u>Jy.</u>	e_Fint	Uncertainty on the integrated flux computed by multiplicating the fitted source residual rms (RMS_TOTAL) by the fitted source area as estimated by FWHMA and FWHMB (ERR_FINT)
176-185	F10.3	MJy/sr	Fpeak	Source peak flux measured from the fitting process (FPEAK)
187-198	F12.3	MJy/sr	e_Fpeak	Uncertainty on the peak flux (ERR_FPEAK)
200-206	F7.2	<u>arcsec</u>	FWHMA	Full width at half maximum of the source along axis a of the elliptical Gaussian as determined by fitting engine
208-214		<u>arcsec</u>	FWHMB	Full width at half maximum of the source along axis b of the elliptical Gaussian as determined by fitting engine
216-224	F9.2	arcsec	e_FWHMA	rms uncertainty on FWHMA (ERR_FWHMA)
226-234	F9.2	arcsec	e_FWHMB	rms uncertainty on FWHMB (ERR_FWHMB)
236-241	F6.1	<u>deg</u>	PA	[] Position angle of the elliptical Gaussian (N E)
243-247	F5.1	<u>deg</u>	e_PA	rms uncertainty on PA
249-258	F10.3	MJy/sr	Bckg	Background value determined at the source peak position (BACKGROUND)
260-271	F12.5	MJy/sr	Bckga	Coefficient a of the zero-order term of

13/5/2020				J/A+A/591/A14
				the background obtained by the fit at the
273-283	F11.5	MJy/sr	Bckgb	source peak position (BACK_ACOEFF) Coefficient b of the first-order term x of the background obtained by the fit at the
285-296	F12.5	MJy/sr/pix	Bckgc	source peak position (BACK_BCOEFF) Coefficient c of the first-order term y of the background obtained by the fit at the
298-307	F10.5	MJy/sr/pix2	Bckgd	source peak position (BACK_CCOEFF) Coefficient d of the second-order term x2 of the background obtained by the fit at
309-318	F10.5	MJy/sr/pix2	Bckge	the source peak position (BACK_DCOEFF) Coefficient e of the second-order term y2 of the background obtained by the fit at
320-330	F11.5	MJy/sr/pix2	Bckgf	the source peak position (BACK_ECOEFF) Coefficient f of the second-order term xy of the background obtained by the fit at
332-342	F11.5	MJy/sr	rms	the source peak position (BACK_FCOEFF) Standard deviation, loc, of the residuals computed within the source area defined by FWHMA and FWHMB after subtracting the
344-354	F11.5	MJy/sr	rmss	best fit (RMS_TOTAL) Standard deviation, loc, of the residuals computed within the fitting window after subtracting the best fit, excluding both the pixels that belong to the source and the pixels belonging to other sources that fall within the fitting window
356-364	F9.5		SNR	(RMS_SURROUND) Signal-to-noise ratio obtained by dividing FPEAK by the residual rms over a source area with FWHMA and FWHMB as semi-axes
		Basic detect	ion inform	(SNR) mation
366-373	F8.3		detX	Relevance of the source in the second-derivative map along the x-axis defined as the ratio between the measured second derivative at source peak position
375-382	F8.3		detY	and the adopted local threshold value (DET_X) Relevance of the source in the second-derivative map along the y-axis defined as the ratio between the measured second derivative at source peak position
384-391	F8.3		detX45	and the adopted local threshold value (DET_Y) Relevance of the source in the second-derivative map along the bisector
393-400	F8.3		detY45	of the xy-axis defined as the ratio between the measured second derivative at source peak position and the adopted local threshold value (DET_X45) Relevance of the source in the second-derivative map along the bisector of the yx-axis defined as the ratio between the measured second derivative at source peak position and the adopted
402-408	F7.3	MJy/sr/pix2	detlimX	local threshold value (DET_Y45) Absolute value for the local detection limit threshold adopted for the second derivative along the x-axis coordinate (DETLIM X)
410-416	F7.3	MJy/sr/pix2	detlimY	Absolute value for the local detection limit threshold adopted for the second derivative along the y-axis coordinate
418-425	F8.3	MJy/sr/pix2	detlimx45	limit threshold adopted for the second derivative along the bisector of the
427-433	F7.3	MJy/sr/pix2	detlimy45	limit threshold adopted for the second derivative along the bisector of the
435	I1		Clump	second and fourth quadrant (DETLIM_Y45) Flag for confusion at detection level (CLUMP_FLAG) (G3)
437	I1		Ncomp	Number of Gaussian components used simultaneously in the fitting process
439-442	14	<u>pix</u>	Xcent	(NCOMP) (G4) The x-pixel coordinate of the centre of the source fitting window in the original image (XCENT)
444-447	14	<u>pix</u>	Ycent	The y-pixel coordinate of the centre of the source fitting window in the
449-450	12	<u>pix</u>	Xwin	original image (YCENT) Half-width size of the source fitting window along x coordinate and centred at YCENT (YUTNDOW)
452-453	12	<u>pix</u>	Ywin	at XCENT (XWINDOW) Half-width size of the source fitting window along y coordinate and centred
455-456	12		NCont	at YCENT (YWINDOW) Number of other sources falling inside
https://cdsa	rc.unis	tra.fr/viz-bin/R	ReadMe/J/A	+A/591/A149?format=html&tex=true

458-461	F4.2	p <u>ix</u>	Cent	the fitting window whose presence is taken into account at fitting stage. Not all those other sources might have been fitted at the same time (NCONTAM) Maximum variation in pixels for adjustment of the fit centre with respect to the position of detection, measured as the distance between the latter and the brightest local (within three pixels) pixel in the fitting
463-466			DOF	window (CENT_TOL) Degrees of freedom of the source Gaussian fit
		Quality fl	ags	
				chi ² determined by the fitting engine. Estimator of the fidelity between the fit and the data computed as $phi=(O(i)-F(i))^2/F(i)$, where $O(i)$ is the observed data in the i pixel of the
499	I1		fitst	fitting window and F(i) is the fitted value in the same position (CHI2OPP) [0/4] Flag returned from the fitting engine (FIT_STATUS) (G5)
501-502	A2		Guess	[ABC 0123456789] Flag on quality of guessed source parameters as determined
504-510	A7		Group	at the detection stage (GUESS_FLAG) (G6) [ABC 0123456789] Flag on quality of guessed source parameters as determined
512			Constr	at the detection stage (GROUP_FLAG) (G6) Flag indicating the number of parameters that reached the tolerance limits allowed to the fit process (CONSTRAINS) (G8)
		Basic extr	action info	rmation
514-521	F8.3		rd2dx	Ratio between the second-derivative value along x direction expected by the fitted model of the source and the second derivative derivative measured at the
523-531	F9.3		rd2dy	detection stage (RDETP2DX) (G9) Ratio between the second-derivative value along y direction expected by the fitted model of the source and the second derivative measured at the detection
533-539	F7.3		rd2dx45	stage. (RDETP2DY) Ratio between the second-derivative value along the bisector of the xy direction expected by the fitted model of the source and the second derivative measured at the detection stage. (RDETP2DX45)
541-548	F8.3		rd2dy45	
550-555	A6		ovlap	Flag to indicate whether the source has been detected and extracted in one or more adjacent tiles (OVERLAP_FLAG) (G10)
557-558	A2		ovflux	Flag to indicate which flux values were adopted if detected and extracted in two adjacent tiles (OVFLUX_FLAG) (G11)

Byte-by-byte Description of file: tableb1.dat

Bytes	Format	Units	Label	Explanations
1- 9	F9.5	<u>deg</u>	GLON	Galactic longitude (1)
11- 18	F8.5	<u>deg</u>	GLAT	Galactic latitude (1)
20- 23	A4		Мар	Map designation (1NNN)
25- 26	I2		PixRed	?=- Number of saturated pixels in Red band
28- 29	12	arcsec	ReRed	<pre>?=- Radius of the circularised area of the saturated pixels cluster in Red band</pre>
31- 33	I3		PixPSW	?=- Number of saturated pixels in PSW band
35- 36	12	arcsec	RePSW	?=- Radius of the circularised area of the saturated pixels cluster in PSW band
38- 39	I2		PixPMW	?=- Number of saturated pixels in PWW band
41- 42	12	arcsec	RePMW	?=- Radius of the circularised area of the saturated pixels cluster in PMW band
44- 45	12		PixPLW	?=- Number of saturated pixels in PLW band
47- 48	12	arcsec	RePLW	?=- Radius of the circularised area of the saturated pixels cluster in PLW band
50- 60	A11		IRAS	IRAS source <60° from the cluster barycentre
62-159	A98		RMS	RMS sources <40° from the cluster barycentre

Note (1): Centroid position of the cluster at the shortest wavelength where the

saturation conditions exists.

Global notes:

```
Note (G1): Designation of the source based on its Galactic position in the form
  LLL.1111+b.bbbb. The naming convention for the Hi-GAL catalogue has the form
  HIGALPXLLL.1111+b.bbbb, where HIGALP stands for the preliminary catalogue,
 X stands for the band where the source has been identified among the
 possible choices:
  B - blue band; R - red band; S - PSW band; M - PMW band; L - PLW band.
Note (G3): A value equal to 0 means that the source was identified from an
  isolated group of pixels above the threshold in all the four derivative
  directions. Sources belonging to the extraction of the same atlas image having
  the same value of this flag belong to the same group of pixels above the
 threshold.
Note (G4): The number of Gaussian components includes the source, so the
  minimum value is 1, this number is greater than 1 if the source is fit with
  other nearby detections.
Note (G5): Possible values of the fitst flag are as follows:
 0 = fit convergence failed
  1 = convergence reached
 2 = convergence reached despite the initial accuracy requested to fitting
      engine was set too low
 3 = maximum number of iterations in the fitting process reached
  4 = problems in fitting due to the initial guess
Note (G6): The form of the Guess and group flags are GN, where
```

G is a letter defined as:

A = optimal number of positions to estimate the size B = sufficient number of positions to estimate the size

C = low number of positions to estimate the size and N is a number defining the quality of initial guess size:

0 = initial estimate failed

1 = good initial estimate for sizes

- 2 = one of the two guessed sizes was initially estimate as smaller than the PSF
- 3 = initial estimates of source sizes were higher than three times the PSF. Note (G8): Values of 4 indicate that the source flux has higher unreliability since either the centre and its sizes have reached the maximum

(or the minimum) allowed for the fit engine. Note (G9): Values closer to one indicate a higher reliability of the source

Note (G10): Overlap flag:

- H indicates that the source has been detected in the tile named in column ATLAS IMAGE
- E,W indicate that the source is detected only in the eastern or western adjacent tile, respectively (east is higher Galactic longitude); if the source has been detected in both H and E or W, then the name of the adjacent tile is also listed (e.g. H_1060). In these cases, the entry in the catalogue is the one with the highest S/N.

Note (G11): Ovflux flag:

- 0 indicates that the source has been detected only once and therefore all fluxes refer to this detection
- -1 indicates that the two fluxes differ by more than 15%; the one listed is that with the highest S/N
- 1 indicates that both integrated fluxes lie within 15%, the one in the catalogue is that with the highest S/N
- 2 indicates that the integrated fluxes differ by more than 15% but FPEAK are within 15%; the one listed is that with the highest S/N

History:

Copied at http://tools.asdc.asi.it/HiGAL.jsp

Patricia Vannier [CDS] (End) 31-Aug-2016

The document above follows the rules of the Standard Description for Astronomical Catalogues; from this documentation it is possible to generate f77 program to load files into arrays or line by line

© Université de Strasbourg/CNRS

f □ y ○ · Contact ▷