



TITLE:

Chandrasekhar polynomials - A brief review (Analysis of inverse problems through partial differential equations and related topics)

AUTHOR(S):

Machida, Manabu

CITATION:

Machida, Manabu. Chandrasekhar polynomials - A brief review (Analysis of inverse problems through partial differential equations and related topics). 数理解析研究所講義録 2021, 2174: 88-92

ISSUE DATE:

2021-02

URL:

<http://hdl.handle.net/2433/263958>

RIGHT:

Chandrasekhar polynomials – A brief review

Manabu Machida

Institute for Medical Photonics Research,
Hamamatsu University School of Medicine, Hamamatsu 431-3192, Japan

E-mail address: machida@hama-med.ac.jp

Abstract. A review on the Chandrasekhar polynomials is given. The polynomials often appear in transport theory. The relation to the method of rotated reference frames for the three-dimensional radiative transport equation is clarified.

1. INTRODUCTION

The Chandrasekhar polynomials play an important role in one-dimensional transport theory (see [5, 9] and references therein). Recently, the appearance of the polynomials has been recognized even for the three-dimensional radiative transport equation [10, 11].

We begin with the one-dimensional transport equation. Let μ_t, μ_s be constants such that $\mu_t > \mu_s \geq 0$. Let μ be the third component of vector $\theta \in \mathbb{S}^2$, i.e., μ is the cosine of the polar angle of θ . Let Ω be an interval on the real axis. We write the transport equation as

$$\left(\mu \frac{\partial}{\partial z} + \mu_t \right) I(z, \theta) = \mu_s \int_{\mathbb{S}^2} p(\theta, \theta') I(z, \theta') d\theta', \quad (z, \theta) \in \Omega \times \mathbb{S}^2.$$

The solution $I(z, \theta)$ will be uniquely determined if suitable boundary conditions are imposed. We assume that the scattering phase function $p(\theta, \theta')$ is given by

$$p(\theta, \theta') = \frac{1}{4\pi} \sum_{l=0}^L \beta_l P_l(\theta \cdot \theta') = \sum_{l=0}^L \sum_{m=-l}^l \frac{\beta_l}{2l+1} Y_{lm}(\theta) Y_{lm}^*(\theta'),$$

where P_l are Legendre polynomials, Y_{lm} are spherical harmonics, and the symbol $*$ means complex conjugate. The coefficient $\beta_0 = 1$ and for $1 \leq l \leq L$, $|\beta_l| < 2l+1$. Using associated Legendre polynomials $P_l^m(\mu)$, spherical harmonics are given by

$$Y_{lm}(\theta) = \sqrt{\frac{2l+1}{4\pi} \frac{(l-m)!}{(l+m)!}} P_l^m(\mu) e^{im\varphi},$$

where $\varphi \in [0, 2\pi)$ is the azimuthal angle of θ . We note that this $p(\theta, \theta')$ implies scatterers are spherically symmetric. In optics, coefficients β_l are often given by $\beta_l = (2l+1)g^l$ with the anisotropy factor $g \in (-1, 1)$ [3].

By changing the spatial variable as $x = \mu_t z$, we can rewrite the transport equation as

$$\left(\mu \frac{\partial}{\partial x} + 1 \right) \psi(x, \theta) = \varpi \int_{\mathbb{S}^2} p(\theta, \theta') \psi(x, \theta') d\theta', \quad (x, \theta) \in \Omega \times \mathbb{S}^2,$$

where $\varpi = \mu_s/\mu_t \in [0, 1)$ is called the albedo for single scattering and $\psi(x, \theta) = I(x/\mu_t, \theta)$. Chandrasekhar's polynomials appear when the solution $\psi(x, \theta)$ to the

homogeneous equation is sought assuming the form

$$\psi(x, \theta) = \sum_{l=0}^{\infty} \sum_{m=-l}^l f_{lm}(\nu) Y_{lm}(\theta) e^{-x/\nu},$$

where $\nu \in \mathbb{R}$ is a parameter and $f_{lm}(\nu)$ are coefficients which will be later related to Chandrasekhar's polynomials. See [1] for the equivalence between the method of discrete ordinates and the spherical-harmonic expansion.

Let us introduce h_l as

$$h_l = \begin{cases} 2l + 1 - \varpi\beta_l, & 0 \leq l \leq L, \\ 2l + 1, & l \geq L + 1. \end{cases}$$

We note the relation

$$\int_{-1}^1 \mu P_l^m(\mu) P_{l'}^m(\mu) d\mu = \frac{2}{4(l+1)^2 - 1} \frac{(l+1+m)!}{(l-m)!} \delta_{l+1, l'} + \frac{2}{4l^2 - 1} \frac{(l+m)!}{(l-1-m)!} \delta_{l-1, l'}.$$

By substituting the assumed form of $\psi(x, \theta)$ into the homogeneous transport equation, we obtain

$$\mu \sum_{l'=0}^{\infty} \sum_{m'=-l'}^{l'} f_{l'm'}(\nu) Y_{l'm'}(\theta) - \nu \sum_{l'=0}^{\infty} \sum_{m'=-l'}^{l'} \frac{h_{l'}}{2l'+1} f_{l'm'}(\nu) Y_{l'm'}(\theta) = 0.$$

Then by multiplying $Y_{lm}^*(\theta)$ and integrating over θ , we obtain

$$\sum_{l'=|m|}^{\infty} \sqrt{\frac{l'^2 - m^2}{4l'^2 - 1}} \delta_{l-1, l'} f_{l'm}(\nu) - \frac{\nu h_l}{2l+1} f_{lm}(\nu) + \sum_{l'=|m|}^{\infty} \sqrt{\frac{(l+1)^2 - m^2}{4(l+1)^2 - 1}} \delta_{l+1, l'} f_{l'm}(\nu) = 0.$$

Let us define $f_{lm}(\nu) = 0$ for $l < |m|$. If we multiply $\sqrt{2l+1}$ in the above equation, we obtain

$$\sqrt{\frac{l^2 - m^2}{2l-1}} f_{l-1, m}(\nu) - \frac{\nu h_l}{\sqrt{2l+1}} f_{lm}(\nu) + \sqrt{\frac{(l+1)^2 - m^2}{2(l+1)+1}} f_{l+1, m}(\nu) = 0.$$

2. CHANDRASEKHAR POLYNOMIALS

Let $x \in \mathbb{R}$. Chandrasekhar introduced polynomials $G_l^m(x)$ which satisfy the following three-term recurrence relation [2, 4].

$$(l+m)G_{l-1}^m(x) - h_l x G_l^m(x) + (l-m+1)G_{l+1}^m(x) = 0, \quad l \geq m \geq 0,$$

with $G_m^m(x) = (2m-1)!!$. See [8] for the case $m=0$.

Then the normalized Chandrasekhar polynomials $g_l^m(x)$ ($l \geq |m|$) were introduced [6, 7]. By setting [15]

$$g_l^m(x) = \sqrt{\frac{(l-m)!}{(l+m)!}} G_l^m(x),$$

we see that g_l^m satisfy the following three-term recurrence relation.

$$\sqrt{l^2 - m^2} g_{l-1}^m(x) - h_l x g_l^m(x) + \sqrt{(l+1)^2 - m^2} g_{l+1}^m(x) = 0, \quad l \geq |m|.$$

Indeed, the three-term recurrence relation for f_{lm} is recovered if we put $g_l^m = f_{lm}/\sqrt{2l+1}$. We set the initial term as

$$g_m^m(x) = \frac{(2m-1)!!}{\sqrt{(2m)!}} = \frac{\sqrt{(2m)!}}{2^m m!}, \quad m \geq 0.$$

Moreover, $g_l^{-m}(x)$ and $g_l^m(-x)$ are related to $g_l^m(x)$ as

$$g_l^{-m}(x) = (-1)^m g_l^m(x), \quad g_l^m(-x) = (-1)^{l+m} g_l^m(x).$$

3. EIGENPROBLEM

To avoid tedious calculations, in this section we assume m is nonnegative: $m = 0, 1, \dots$. It is straightforward to extend results below to the case of negative m . It is also possible to write $p(\theta, \theta')$ only with $m \geq 0$ making use of the formula $P_l^{-m}(\mu) = (-1)^m \frac{(l-m)!}{(l+m)!} P_l^m(\mu)$. Let us introduce

$$\sigma_l = \frac{\mu_t h_l}{2l+1},$$

and

$$y_l^m(x) = \sqrt{(2l+1)\sigma_l} g_l^m(x).$$

Using the new notation, the three-term recurrence relation for g_l^m becomes

$$b_l(m) y_{l-1}^m(x) - \frac{x}{\mu_t} y_l^m(x) + b_{l+1}(m) y_{l+1}^m(x) = 0,$$

where

$$b_l(m) = \sqrt{\frac{l^2 - m^2}{(4l^2 - 1)\sigma_l \sigma_{l-1}}}.$$

By imposing the truncation condition

$$g_{M+1}^m(\xi) = 0, \quad M = l_{\max} + m \quad \text{or} \quad l_{\max},$$

where M determines the highest degree of P_l^m used to express $\psi(x, \theta)$, we arrive at the eigenproblem

$$B(m) Y_\xi(m) = \frac{\xi}{\mu_t} Y_\xi(m),$$

where $Y_\xi(m) = (y_m^m(\xi), y_{m+1}^m(\xi), \dots, y_M^m(\xi))^T$. The tridiagonal matrix $B(m)$ is given by [12, 14]

$$\{B(m)\}_{ll'} = b_l(m) \delta_{l', l-1} + b_{l'}(m) \delta_{l', l+1}.$$

In the method of rotated reference frames [12, 14], eigenmodes are labeled by eigenvalues of $B(m)$. The number of rows and columns of $B(m)$ is $l_{\max} + 1$ when $M = l_{\max} + m$ and is $l_{\max} - m + 1$ when $M = l_{\max}$. Since $B(m)$ is a symmetric tridiagonal matrix with nonzero off-diagonal elements, its eigenvalues are distinct. Also if ξ/μ_t is an eigenvalue for $y_l^m(\xi)$, then $-\xi/\mu_t$ is another eigenvalue and $y_l^m(-\xi) = (-1)^l y_l^m(\xi)$ [12]. Essentially the same tridiagonal matrix W was introduced in [15]. Elements of W are given by $\{W\}_{ll'} = w_l(m) \delta_{l', l-1} + w_{l'}(m) \delta_{l', l+1}$, where $w_l(m) = \sqrt{(l^2 - m^2)/(h_l h_{l-1})}$. Let ξ_j/μ_t ($j = 1, \dots, l_{\max} + 1$) denote eigenvalues of $B(m)$. We note that $\{\xi_j\}$ are eigenvalues of W .

For simplicity, hereafter, we suppose $M = l_{\max} + m$ and $l_{\max} \geq 1$ is an odd integer. There are $(l_{\max} + 1)/2$ positive eigenvalues and $(l_{\max} + 1)/2$ negative eigenvalues for each m . Then we can write eigenvalues as

$$\xi_1 > \xi_2 > \cdots > \xi_{\frac{l_{\max}+1}{2}} > 0 > \xi_{\frac{l_{\max}+1}{2}+1} > \cdots > \xi_{l_{\max}+1},$$

and $\xi_{l_{\max}+2-j} = -\xi_j$ ($j = 1, \dots, (l_{\max} + 1)/2$).

The following lemmas hold.

Lemma 3.1 (Orthogonality [13, 15]). *We have*

$$\frac{1}{Z_j} \sum_{l=m}^{l_{\max}+m} y_l^m(\xi_i) y_l^m(\xi_j) = \delta_{ij}, \quad i, j = 1, 2, \dots, l_{\max} + 1,$$

where $Z_j = \sum_{l=m}^{l_{\max}+m} [y_l^m(\xi_j)]^2$.

Proof. Eigenvectors corresponding to two distinct eigenvalues of a symmetric real matrix are orthogonal. \square

Lemma 3.2 (Completeness [13, 15]). *We have*

$$\sum_{j=1}^{l_{\max}+1} \frac{1}{Z_j} y_l^m(\xi_j) y_{l'}^m(\xi_j) = \delta_{ll'}, \quad l, l' = m, m+1, \dots, m+l_{\max},$$

where $Z_j = \sum_{l=m}^{l_{\max}+m} [y_l^m(\xi_j)]^2$.

Proof. Let us introduce vectors $X_j = Y_{\xi_j}(m)/\sqrt{Z_j}$ and matrix $X = (X_1, \dots, X_{l_{\max}+1})$. Then X is an orthogonal matrix: $X^{-1} = X^T$. Next we introduce matrices $Z = \text{diag}(\sqrt{Z_1}, \dots, \sqrt{Z_{l_{\max}+1}})$ and $Y = (Y_{\xi_1}(m), \dots, Y_{\xi_{l_{\max}+1}}(m))$. The matrix Y is expressed as $Y = XZ$.

Let us consider

$$\sum_{j=1}^{l_{\max}+1} D_j y_{l'}^m(\xi_j) = \delta_{ll'}, \quad l, l' = m, m+1, \dots, m+l_{\max}.$$

To find D_j ($j = 1, \dots, l_{\max} + 1$), we introduce vectors $D = (D_1, \dots, D_{l_{\max}+1})^T$ and $F = (\delta_{m,l}, \delta_{m+1,l}, \dots, \delta_{m+l_{\max},l})^T$, and write the relation as $YD = F$. Since $D = Z^{-1}X^T F$, we obtain

$$D_j = \frac{1}{Z_j} y_l^m(\xi_j).$$

This completes the proof. \square

Remark 3.3. If we define $|y_{\xi_j}(m)\rangle = Y_{\xi_j}(m)/\sqrt{Z_j}$ and $\langle l|y_{\xi_j}(m)\rangle = y_l^m(\xi_j)/\sqrt{Z_j}$, then the orthogonality and completeness in Lemma 3.1 and Lemma 3.2 are equivalently expressed as

$$\langle y_{\xi_i}(m)|y_{\xi_j}(m)\rangle = \delta_{ij}, \quad \sum_{j=1}^{l_{\max}+1} |y_{\xi_j}(m)\rangle \langle y_{\xi_j}(m)| = 1.$$

4. CONCLUDING REMARKS

In this paper, we focused on the case $\mu_a > 0$, i.e., $\varpi < 1$. It is possible to consider the conservative (nonabsorbing) case $\mu_a = 0$ but it must be done separately. When $\varpi = 1$, $\sigma_0 = 0$ and the element $b_1(0)$ becomes infinity. In this case, we need to remove the top left part of $B(0)$.

In application, the numerical evaluation of the Chandrasekhar polynomials is important. Various numerical techniques have been developed [5, 6, 7].

ACKNOWLEDGEMENTS

The author acknowledges support from JSPS KAKENHI Grant No. 17K05572, 18K03438.

REFERENCES

- [1] Barichello L B and Siewert C E 1998 On the equivalence between the discrete ordinates and the spherical harmonics methods in radiative transfer *Nucl. Sci. Eng.* **130** 79–84
- [2] Benassi M, Garcia R D M, Karp A H and Siewert C E 1984 A high-order spherical harmonics solution to the standard problem in radiative transfer *Astrophys. J.* **280** 853–864
- [3] Bigio I and Fantini S 2016 *Quantitative Biomedical Optics* (Cambridge: Cambridge University Press)
- [4] Chandrasekhar S 1960 *Radiative Transfer* (New York: Dover)
- [5] Ganapol B D 2014 Chandrasekhar polynomials and the solution to the transport equation in an infinite medium *J. Comp. Theor. Trans.* **43** 433–473
- [6] Garcia R D M and Siewert C E 1989 On discrete spectrum calculations in radiative transfer *J. Quant. Spec. Rad. Trans.* **42** 385–394
- [7] Garcia R D M and Siewert C E 1990 On computing the Chandrasekhar polynomials in high order and high degree *J. Quant. Spec. Rad. Trans.* **43** 201–205
- [8] Inönü E 1970 Orthogonality of a set of polynomials encountered in neutron transport and radiative transfer theories *J. Math. Phys.* **11** 568–577
- [9] Kuščer I and McCoormick N J 1991 Some analytical results for radiative transfer in thick atmospheres *Trans. Theo. Stat. Phys.* **20** 351–381
- [10] Machida M 2014 Singular eigenfunctions for the three-dimensional radiative transport equation *J. Opt. Soc. Am. A* **31** 67–74
- [11] Machida M 2015 An F_N method for the radiative transport equation in three dimensions *J. Phys. A: Math. Theor.* **48** 325001
- [12] Markel V A 2004 Modified spherical harmonics method for solving the radiative transport equation *Waves Random Media* **14** L13–L19
- [13] McCormick N J 1995 Identities for the Chandrasekhar polynomials in the spherical harmonics radiative transfer method *J. Quant. Spec. Rad. Trans.* **54** 851–856
- [14] Panasyuk G, Schotland J C and Markel V A 2006 Radiative transport equation in rotated reference frames *J. Phys. A: Math. Gen.* **39** 115–137
- [15] Siewert C E and McCormick N J 1997 Some identities for Chandrasekhar polynomials *J. Quant. Spec. Rad. Trans.* **57** 399–404