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ARTICLE

Determination of the microscopic mineralogy of inclusion in an amygdaloidal pillow basalt by fs-LIMS

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We present chemical depth profiling studies on mineralogical inclusions embedded in amygdale calcium carbonate by our Laser Ablation Ionisation Mass Spectrometer designed for in situ space research. An IR femtosecond laser ablation is employed to generate ions that are recorded by a miniature time-of-flight mass spectrometer. The mass spectra were measured at several locations on the sample surface and yield chemical depth profiles along the depth length of about 30 μm . The presence of oxides and sulphides within inclusion material allows us to derive elemental abundance calibration factors (relative sensitivity coefficients, RSCs) for major and minor elements. These are obtained from the atomic intensity correlations performed on the depth profiling data. With the RSCs corrections the quantitative analysis of more complex mineralogical phases within the inclusion is conducted by correlating atomic abundance fractions in ternary diagrams, typically used in geology. The spatial resolution of the depth profiles was sufficient to study chemically distinct micrometre-sized objects, such as mineralogical grains and thin layers of minerals including micrometre-sized filamentous structures. The method presented here is well-suited for the quantitative chemical analyses of highly heterogeneous materials where the ablation condition can vary locally with the material composition making the application of standard reference materials less accurate. The presented method is developed to distinguish between abiotic and biological material while searching for micrometre-sized extinct or extant life forms on the surfaces of Solar System bodies.

Introduction

One of the major objectives of current space research and planetology is the search for signatures of extant and extinct life on other planetary bodies. With the assumption that life on other planets should be based on the same physical and chemical rules as life we know from Earth and evolved along similar pathways, these searches should follow known terrestrial practice. Ancient extinct life can be identified by finding remains of organic material, stromatolitic layers, and individual microfossils within planetary rocks and soils¹. Identification of metabolic products together with the

deciphering of fossilization mechanisms can be essential in such searches. Among available analytical instrumentation, only these instruments, which can deliver selective and sensitive chemical measurements of microscopic structures can be employed for this task^{2, 3}.

For conducting analyses on planetary bodies only small, light, low-powered and sufficiently robust instrumentation can be used. Several space-borne instruments including Laser-Induced Breakdown Spectroscopy (LIBS), X-Ray Fluorescence (XRF), and Laser Raman Spectroscopy have proved their capabilities in

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delivering composition measurements of the sample surface and important biosignatures in numerous performance studies. However, their measurement sensitivity might not be sufficient to investigate the chemical composition of microscopic structures of sparse life^{2, 4}. An example of the instrument combining Raman, LIBS and IR techniques is the SuperCam instrument designed on the NASA Mars2020 with the performance figures of each of applied technique described in recent publication and references therein^{5, 6}. Compared to these techniques, laser-based mass spectrometry adds complementary information. With current LIMS spatial resolution (lateral and vertical) and sensitivity, detailed analyses of grain-size objects can be conducted with the determination of context mineralogy and detection of micro-sized fossils⁷⁻⁹.

Laser ablation ionisation mass spectrometry (LIMS) based on time-of-flight mass analysers is a well-known technique with its origin in the mid-1970s. In the beginning, the technique suffered from various drawbacks, which mostly were affecting the quantification and the measurement procedure. With the continuous advances in electronics, vacuum, and laser systems, this technology has re-emerged for laboratory and field application for chemical analysis of solid-state samples¹⁰⁻¹². Over the years, a high degree of miniaturisation of the LIMS instrument has been achieved with the objective to become a space-borne instrument^{8, 13-15}. Robustness of the time-of-flight mass analyser, flexible operation of the laser ablation ion source, the control over the ion optics settings and detector gain allow direct and sensitive chemical analysis of raw samples providing efficient ion production and nearly 100% ion transmission from the measurement spot to ion detector^{15, 16}. For the analysis of micrometre-sized objects such as grains or microfossils, the integration of a microscope camera system into the system helps in an initial characterisation of micro-sized objects and improving overlap of the laser spot with the object of interest^{8, 17}.

LIMS with an fs laser ablation ion source shows an increasingly higher performance. It delivers improved quantitative chemical composition information with

considerably reduced element fractionation¹⁸⁻²⁰. Several recent studies show that LIMS can be considered as a standardless technique delivering semi-quantitative results¹⁸⁻²². Further improvements to the quantitative analysis can be achieved by adding to atomic abundance derived from the single-charged atomic mass peaks also multiple ionised atomic mass peaks²³. An improved efficiency of the atomic ion production and quantification of atomic ions can be achieved applying a double pulse laser ablation/ionisation ion source²⁴. In double pulse system, the first pulse is ablating the sample surface and the second, shortly delayed pulse interacts with freshly formed plasma plume resulting in additional plasma heating, atomisation and ionisation. These processes influence also post-plasma chemistry efficiency. The presence of isobaric multiple-charged atomic ions and clusters in ion beam can introduce uncertainty in the quantification of the atomic signal measurement; hence, by applying double pulse laser ion source, the isobaric interferences can be minimized. It has been demonstrating also that the isobaric interferences due to clusters could be reduced by implementing collisional cell which allows partial temporal separation of the cluster and atomic ion distributions²⁵. High-resolution time-of-flight mass spectrometric measurements of atomic ions produced in laser ablation are typically difficult due to wide kinetic energy distributions of produced ions. Only, recently a high-resolution ($M/\Delta M \sim 10^4$) laboratory LIMS system with a double reflectron system was developed to resolve most of mass peaks of cluster ions, hence, the isobaric interferences resulting from polyatomic species can be avoided²⁶. In miniature LIMS systems designed for in situ applications not all these implementations can be made. Here, we propose alternative method of the abundance calibration. The applied method uses depth-profiling data and requires presence of simple chemical compounds such as oxides and sulphides within the sample material.

In recent years, fs-LIMS proved its capabilities for chemical analysis with high spatial (lateral, vertical) resolution. High lateral resolution is achieved by focusing laser radiation onto the surface to a spot size of

a few micrometres and by tuning over the pulse energy, ablation rate and thickness of the ablated layer can be well controlled so that the measurements can be conducted with the depth resolution down to a few nanometres^{27–32}. In our recent studies on the Ni/Cr NIST standard, we have accomplished a depth resolution of about 30 nm applying UV femtosecond laser radiation arranged in a double pulse configuration^{24, 32}. 3D chemical mapping analysis with LIMS were recently presented^{33, 34}.

There are other laboratory analytical techniques, which offer high spatial resolution depth profiling capabilities such as Glow Discharge Mass Spectrometry (GD-MS), Laser Ablation Inductively Coupled Mass Spectrometry (LA-ICP-MS) and Secondary Ion Mass Spectrometry (SIMS)^{35–40}. These instruments in current state of development are less suitable for application to space research but can be useful in applications in terrestrial environments. Current LIMS systems fulfil the requirements of size, weight or power consumption necessary available aboard a spacecraft, either a rover or a lander⁷.

In the current contribution, we introduce a procedure for the in-situ atomic abundance calibration using micrometre-sized inclusion embedded in a calcium carbonate host mineral. The procedure uses the depth profiling data e.g. the measurements of atomic intensities in a function of ablation layer. By correlating the intensities of atomic species measured at specific ablation layers and observation of linear correlation, we can deduce the presence of the molecular compounds and determine corrections to the measured atomic intensities. These corrections are obtained by comparison of the measured slopes of the correlations and the expected values from the actual chemical formula. From the measurements of various oxides and sulphides we obtain the correction coefficients for most of major and minor elements. Determination of these corrections on simple compounds allows for a more detailed analysis of complex mineralogical phases present in the sample at microscopic scales. Having the obtained corrected elemental abundances, atomic abundance fractions are plotted in ternary diagrams

allowing the investigation of more complex mineralogical phases. A similar procedure which is known as LRS-linear regression slope method was applied in earlier MC-ICP-MS studies to improve accuracy and precision of isotope ratios^{41–43}. This data reduction protocol used the simultaneous responses of all isotopes measured. In the current study, the simultaneous response of elements is expected while ablating specific molecular compound.

Similar materials to that studied here are found on Mars and can be considered as a distinct astropaleontological target. The calcium carbonate phase in basalts can host veins and voids, which are recognised as habitats for endolithic microorganisms and form secondary infilling mineralisation in volcanic environments^{44–47}. Thus, extended characterisation of these materials and understanding their heterogeneities are important to account for their biological or abiotic origin.

Experimental

Materials

The sample of an amygdaloidal pillow basalt from Kinghorn, Fife, Scotland⁴⁸ was processed to a 30 μm thick thin-section, mounted on a stainless-steel sample holder and introduced into the vacuum chamber for its mass spectrometric analysis (Fig. 1). The left panel of Fig. 1 shows the thin-section of basalt with an inclusion of calcium carbonate. The right panel shows a close-up image of the dark inclusion, which contains a number of needle-shaped microscopic structures. The enclosure into the calcium carbonate host happened some 360–320 Ma ago⁴⁸.

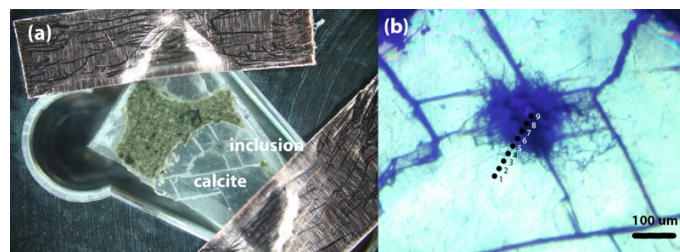


Fig. 1. (a): The sample prior to the introduction into the vacuum chamber. The sample is placed within milled depression in the sample holder made of stainless steel and is attached by a copper tape. (b): Close up microscope image of the

amygdale calcium carbonate sample with a dark inclusion in the centre. A network of needle-like structures spreading out from a denser, darker central part of the inclusion can be readily identified. The ten black spots with numbers indicate the location at which laser ablation mass spectrometric analyses were conducted.

The LIMS instrument

The mass spectrometric investigations were conducted with our miniature laser ablation ionisation mass spectrometer designed for in situ space research applications. The instrument combines a femtosecond-laser ablation ion source with a miniature time-of-flight mass analyser²². Current studies are conducted with a NIR-fs-laser radiation (pulse width, $\Delta t \sim 190$ fs, $\lambda = 775$ nm, laser pulse repetition rate of 1 kHz) focused to a circular spot of about 10 μm in diameter. In all measurements, a pulse energy of about 8 μJ on the sample surface was applied. The resulting craters were inspected with an optical microscope¹⁷. The detailed crater morphology analyses can be found in our previous publications³¹. The measurements were computer-controlled using our custom-designed operating system¹⁵. The sample was attached to the xyz-micro-translation stages. The stage allows for about 2 μm positioning accuracy of the location of interest on the sample by using prior measurements size calibration with an in-situ microscope camera system¹⁷.

Mass spectrometric measurements

Mass spectra were collected at ten locations on the sample, starting from the pure calcium carbonate mineral, across the interface of the calcium carbonate host all the way into the dark inclusion, and on the inclusion's surface seen as a dark area in the centre of Fig. 1b. The dark inclusion consists of needle-shaped micrometre-sized structures that are embedded in complex mineralogical phases other than calcium carbonate. For each location 300 mass spectra were acquired corresponding to 300 ablated layers. The individual layer spectrum is measured by accumulation mass spectra of 200 laser shots on board of the acquisition system. Additionally, mass spectra on the calcium carbonate host were collected by drilling through the 30 μm thin-section, which corresponds to about 366 layers. The mass calibration of the spectra, mass peak integration and principles of depth profiling

analysis were discussed in detail in our previous publications^{49, 50}.

Relative sensitivity coefficients

The Relative Sensitivity Coefficient (RSC) for an element in the sample is obtained by dividing measured element concentration, X_m by the expected concentration, X :

$$(1) \text{RSC}_X = \frac{X_m}{X}$$

The LIMS instrument delivers measurement of atomic intensities. These intensities are calculated by integrating mass peaks of this element in the mass spectrum. Thus, the atomic concentrations are presented as the atomic fractions. To derive the RSC of an element, one compares the measured and expected atomic fractions:

$$(2) \text{RSC}_X = \frac{A_X/A_{\text{tot}}}{P_X}$$

where A_X stands for the atomic ion intensity obtained from the spectrum and A_{tot} is the sum of all atomic intensities measured in the sample. Hence, A_X/A_{tot} is the measured fraction of given element in the sample. P_X denotes the stated fraction of the X atoms to all atoms in the sample. When the sample is a pure mineralogical compound such as calcium carbonate, CaCO_3 , the atomic fractions of P_{Ca} , P_{C} , P_{O} are 0.2, 0.2 and 0.6, respectively.

Knowing concentration ratio between elements X_i and reference element Y in the sample, we can derive another formula for calculating RSCs:

$$(3) \text{RSC}_{X_i} = \text{RSC}_Y \frac{X_{mi} Y}{Y_m X_i}$$

where RSC_Y is the RSC of reference element Y, X_{mi}/Y_m is the ratio of the measured concentrations (slope in the linear intensity correlation), Y/X is the stated ratio of element Y to element X_i (see ESI, Section A for more details).

Using depth profiling data and the signal intensity correlation, we can identify elements, which form simple chemical compounds even they are distributed within sample volume irregularly. In addition to their intensity relationship, the linear intensity correlation of the measuring signals also yields the location where the

signals are measured in the sample i.e., where these minerals are present in the sample. Hence, we can identify these compounds at specific ablation layers. The determination of RSCs from the signal intensity correlation between species in a simple chemical compound can be made this way, e.g., two or three-element compounds such as oxides or sulphides present in the sample as layers or grains can be made this way. For the correlation, we select the atomic signals of a compound, which originate from various ablation layers. Atomic signals emerging from the same compound will always create a linear correlation, since the element ratio will be constant in all different layers, independently of the amount of the compound. The slope of a linearly

Results and discussion

Mass spectra of the calcium carbonate phase and the dark inclusion

Figure 2a shows the mass spectrum measured on the pure calcium carbonate with C, O, and Ca as the major elements that are characteristic for the calcium carbonate mineral composition, CaCO_3 . From purely mass spectrometric perspective, it is hard to identify polymorphic modifications. The calcium carbonate could be calcite, aragonite, or vaterite, but earlier it was determined to be calcite⁷. Additionally, we also observe mass peaks of the elements Na and K. The dynamic range of the spectrum is close to 10^4 , which allowed for simultaneous detection of the major elements and trace elements with a bulk concentration of at least 100 ppm

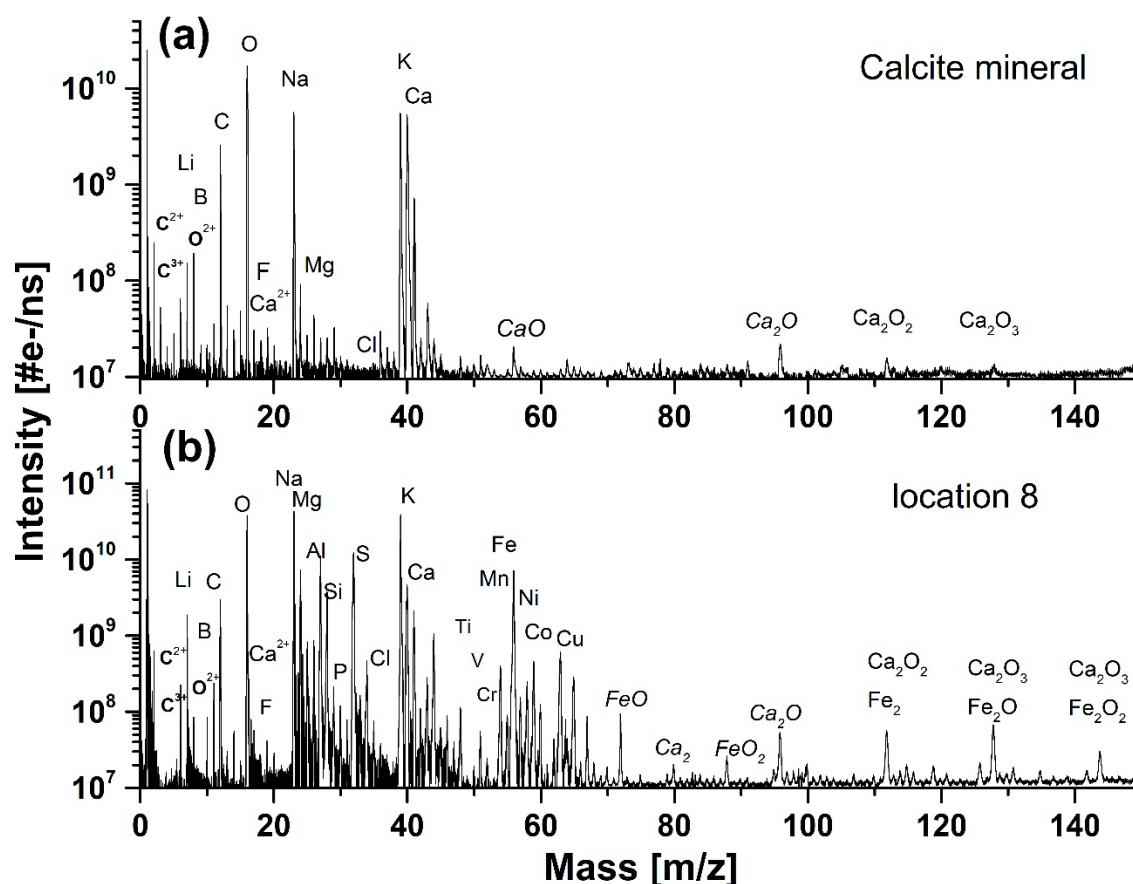


Fig. 2 (a) The mass spectrum recorded on a pure calcium carbonate mineral on location 4, (b) and on the inclusion on location 8.

correlated data set is then the effective abundance ratio of the two correlating atomic species.

(atomic fractions). In addition to C, O and Ca also Li, B, F, Mg, S, Cl, and Sr elements are identified in the spectrum.

These elements are known impurities in the calcium carbonate minerals. Other mass peaks within the low atomic mass range are double- and triple-charged ions of C, O, and Ca. Non-equilibrium chemistry in the expanding and cooling ablation plume leads to the formation of hydrocarbons, C_xH_y , and oxides, Ca_xO_y (where $x = 1-3$ and $y = 1-4$), which ions are also observed in the mass spectrum. Similar mass spectra are recorded at locations 1–4 (see also Fig. 1b), indicating chemical homogeneity of the host mineral.

The mass spectrum displayed in Fig. 2b was measured on the inclusion at location 9 (see also Fig. 1). The major and minor elements observed are O, Na, K, Fe and C, Mg, Al, Si, S, Ca, Ni, Co, respectively, including also the trace elements, Li, N, P, Ti, V, Cr, Cu, Zn and Sr. The dynamic range in the spectra is larger than 10^4 . In addition, multiple charged ions, hydrocarbons and oxides of the main constituents C, O, Ca, and Fe can be identified in the mass spectra. The resemblance of the mass spectra at location 9 and at locations 6, 7 and 8 is large, indicating negligible chemical heterogeneity in the dark inclusion. However, the depth profiling analysis shows that the inclusion material is heterogeneous.

Chemical depth profiling on calcium carbonate mineral

Figure 3a shows the chemical depth profiles measured for the elements C, O, Na, and Si along the 30 micrometre-thick calcium carbonate sample as a function of the ablation layer number. The measured mass peak intensities are observed to decrease fast for about first 30 ablation layers and for the remaining ablation layers - approximately monotonically with the ablation layer number. Because of the mass peak shape distortions and peak intensity fluctuations in the spectra recorded during an initial ablation phase (here for the first 30 ablation layers) these spectra are removed from the analysis. Above the ablation layer number 366 we observe a sudden signal increase for all elements,

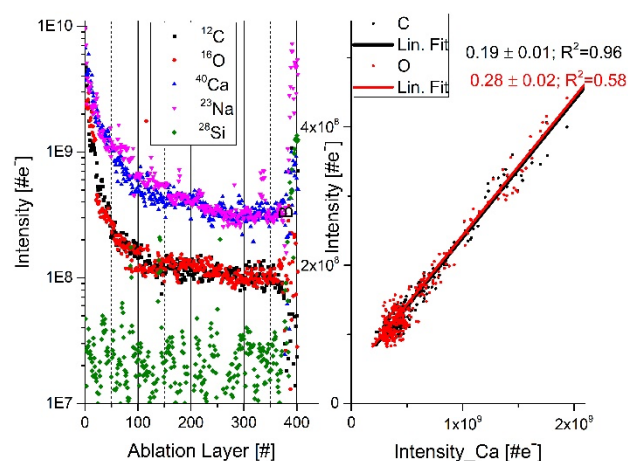


Fig. 3 (a): Depth profile of C, O, Ca, Na, and Si as a function of the ablation number. After 366 ablation layers we observe an increase of the Si and Na atomic intensities indicating transition to soda lime silica glass substrate material. (b): Correlation between C and Ca, and O and Ca atomic intensities, respectively, derived from the depth profiling data.

including Na and Si. The Na and Si signals show the transition to the soda lime glass substrate, which indicates that the laser has drilled a crater through approximately $30\mu\text{m}$ thick thin-section sample. Approximately circular craters with a diameter of about $\phi 10\mu\text{m}$ were produced as observed with the internal optical microscope¹⁷. More advanced analyses of the crater characteristic were not conducted in the current study.

Mass peak intensity correlation of C, O and Ca and RSC values

Calcium carbonate mineral

The mass spectrometric analysis of the composite spectrum (see Fig. 2a) recorded on pure calcium carbonate yield C, O and Ca atomic intensities from which we can calculate atomic fractions of each element and compare these results to the expected atomic fraction in $CaCO_3$ using formula 2. These calculation yields RSCs for C, O and Ca of 0.55 ± 0.03 , 0.36 ± 0.05 , and 3.3, respectively (see also ESI, section A, Table 1 for more details). The intensity correlation plots of C and Ca and O and Ca, derived from individual mass spectra measured along the ablation depth are displayed in Fig. 3b. The mass peak intensities correlated in plots are calculated from the peak areas. We observe linear correlations for these data and the slope coefficient as

shown in Fig. 3b together with the R^2 correlation coefficient. The slope coefficients derived from the linear fits are 0.19 ± 0.01 and 0.28 ± 0.02 for the correlation of C with Ca and of O with Ca, respectively. Considering the expected atomic fractions of the elements in the pure calcium carbonate mineral, we derive relative sensitivity coefficients, RSCs from the formula 3 by taking RSC_{Ca} determined using the first method as reference. The RSCs values obtained with this method are 0.63 ± 0.02 , and 0.30 ± 0.03 for C, O, respectively (see also ESI, Section A, Table 2) indicating consistency of the RSCs values determined by the two methods.

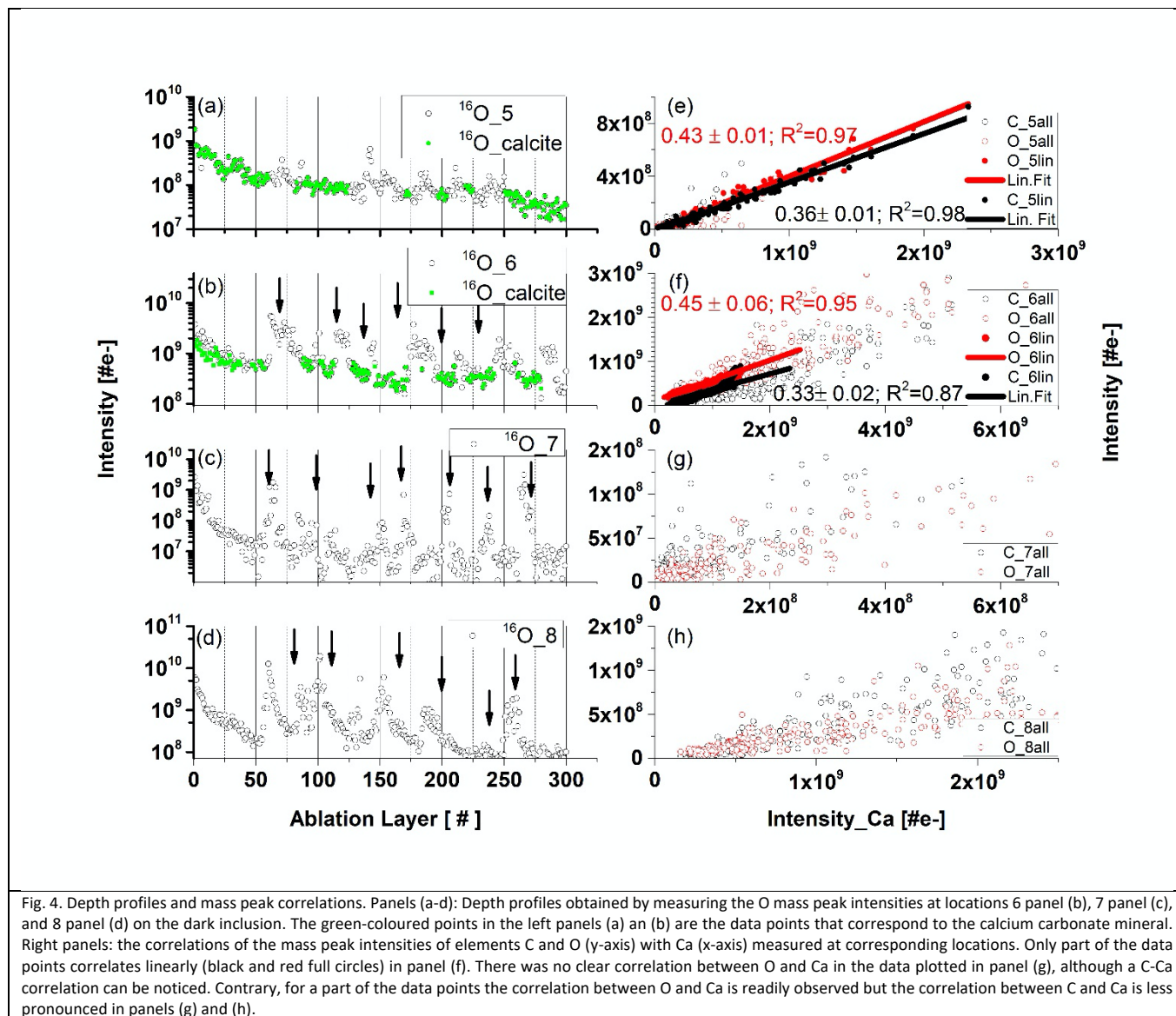
Inclusion material, heterogeneities

Figure 4 compares the O depth profiles in a function of ablation layer number recorded at various locations on the inclusion including also location at the interface between pure calcium carbonate and inclusion material (Fig. 4, panel a, see also location 5 in Fig. 1b). Similar depth profiles are observed for C and Ca although some differences at specific ablation layers in the profile shapes can be noticed. An increase of mass peak intensities at specific ablation layer form characteristic peaks on the depth-profiling curve. A systematic shift of these peaks towards lower ablation layer numbers is observed while the measurement locations become closer to the centre of the inclusion, as indicated by the arrows in Fig. 4, (b, c and d). This shift is likely due to measurement of the same assembly of dark needles embedded in the matrix with a certain angle to the surface normal. The closer one gets to the centre of the dark inclusion, the earlier one measures the needle assembly because the darker area is closer to the surface and the needles were growing in opposite direction, along the direction of depth of the thin section. Figure 4 (panels' e to h) displays the mass peak intensity correlations of C with Ca, and O with Ca, measured at locations from 5 to 8. At location 5, the mass peak intensities correlate linearly well and only a small group of data points lay outside the correlation line. Contrary, at location 6, the linear correlation is observed only for a small part of data points (Fig. 4, panel f). At locations 7 and 8, no obvious linear correlation can be observed; the data points form a diffuse distribution with a large

intensity spread. The data points in the depth profiles measured in locations 5 and 6 (Fig. 4, panels' a and b, respectively), for which the calcium carbonate mineral was identified via the linear correlations (red and black data points on the linear regression in panels e and panel f) are indicated as green circles in panels' a and b, respectively. These are found at the ablation layer numbers between regions of increased peak intensities on the depth-profiling curve. These layers with calcium carbonate, if present, are not easy to detect at locations 7-9. By fitting the fraction of the data that correlates to the calcium carbonate mineral at location 6, one obtains slope coefficient values for O, which are larger compared to that determined at the location of a pure calcium carbonate. The slope coefficient values in correlation of C and Ca are observed to be similar to each other at these locations with the slope coefficient values of 0.41 ± 0.06 and 0.31 ± 0.02 at location 6 and 0.45 and 0.33 at location 7 for C/Ca and O/Ca correlations, respectively (Fig. 4e,f). Considering data at locations 6 we can repeat the procedure applied for pure calcium carbonate phase

and derive RSC coefficients of C, O and Ca at the inclusion location (see ESI Section A, Table 3) using the composite spectrum obtained by summing up only the contributing the data points to linear correlations. These calculations yield RSCs of C, O, and Ca of 0.91

minor elements recorded at location 9. Significant mass peak intensity variations for all elements can be observed along the depth profiles at specific ablation layer numbers. Signal intensity variations indicate either a deficit or an increase of the element concentration at



± 0.05 , 0.44 ± 0.05 and 2.71 ± 0.04 , respectively. Applying again Eq. 3 with RSC of Ca equal 2.71 ± 0.04 as reference and considering slope coefficients C/Ca and O/Ca we can obtain the RSC of C and O equal to 0.97 ± 0.09 and 0.38 ± 0.05 , respectively. The RSC of Ca is chosen as the reference (see Section A, Table 3 and 4 in ESI for more details).

Depth profiles at location 9

Figure 5 shows depth profiles of several major and

these specific ablation layers. Some increase of the ablation efficiency can occur because of the increased absorption of laser radiation at darker spots such as needles. In agreement with the signal intensity correlation studies at locations 7 and 8, we also observe that the Ca, C, and O signals do not correlate linearly. Moreover, at certain depths, a depletion of Ca is observed shown as shaded areas in Fig. 5a. At the ablation layer ranges 53:58, 65:71, 128:132, 156:161, and 165:171, we can observe that the drop of the K

intensity is accompanied by an increase of the intensities of Na, C, and Ca (Fig. 5 a). Considering correlations of the minor elements, an increase of the S, Ni, Co, Mg, Mn, Cu, and B at several locations is typically accompanied by a decrease of the Fe and Ti mass peak intensities (see shaded areas in Fig. 5b and Fig. 5c). Only at the ablation, layer 110 all peaks intensities are observed to correlate. From the microscope images of the needle-like structures we can estimate their size to be in the range of 0.8 – 2 μm (Fig. 1b). The narrowest peaks in the depth profile involve 5 ablation layers (see Fig. 5). Thus, the thickness of the measured individual ablation layers is estimated to be in a range 0.2 – 0.4 μm .

Determination of RSCs from the data collected at locations 6 to 9

If the atomic signals were measured stoichiometric, the slope in the correlation would represent the abundance ratios of the elements of the relevant chemical formula. So far, we have determined RSCs for the elements C and O through their abundance correlation with Ca. Based on our previous studies we know that in the inclusion material other small chemical compounds are present including oxides and sulphides. By correlating the intensities of other elements with oxygen and sulphur intensities, we can gain the information about the RSCs for these elements

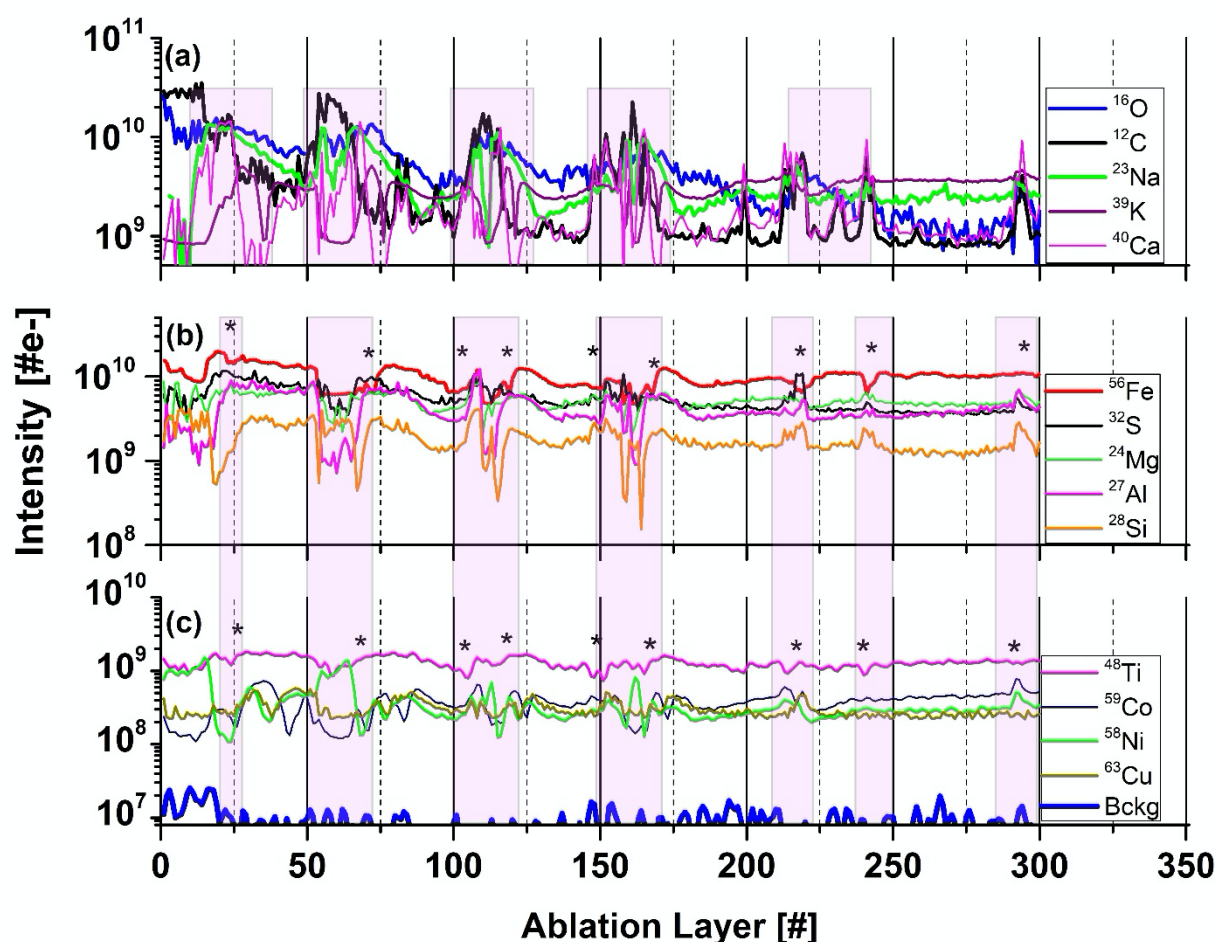


Fig. 4 Variation of the intensities as function of ablation layer number for several elements (isotopes) measured on the dark inclusion at location 8 (Fig. 1b). Panel (a) shows the intensity variations of the C, O, Ca, Na, and K intensities. Panel (b) shows variations of the Fe, S, Mg, Al, and Si intensities. Panel (c) shows variations of Ti, Ni, Co, and Cu intensities. At ablation layers marked by (*) the intensities of Ni and Co correlate well with S, and anticorrelate with Fe. Shaded area highlights layer with specific element correlation, respectively anticorrelations. The traces are shifted to each other vertically for better visualisation of their shapes.

providing that they form oxides and/or sulphides. The details of RSCs calculations are given in ESI, Section A Table's 5,6 and 7. Here we briefly discussed the results obtained with the linear intensity correlation method.

In Figure 6, the intensities of several elements are set in correlation with S (left panel) and O (right panel) intensities, respectively. For some parts of the data set, a linear correlation, either with oxygen or sulphur is observed indicating that these elements may have their origin in oxides and/or sulphides. This agrees with the results of our earlier studies, where identification of oxides and sulphides with other analytical methods were performed⁷. Typically, the fs-LIMS measurements on Standard Research Materials show that RSCs for almost all metallic elements are, generally, close to one within about 15%. An exception are Al, Ti, Cu and Mg. RSCs of non-metallic elements such as C, S, O, Si and Ca various within broader range and depends sensitively on the ablation conditions and material properties^{19, 22}.

Iron is an abundant element in the inclusion. Typically, iron occurs ubiquitously in hydrothermal ore deposits in sulphides, oxides, silicates, and in carbonate minerals. In the correlation of the Fe with O intensities, we observe three different linear slopes (Fig. 6d). The coefficient values for the red and blue slope are 1.4 ± 0.1 and 3.3 ± 0.3 , respectively. With RSC of O derived in the analysis of calcium carbonate on the inclusion equal to 0.38 ± 0.05 , we derive RSC of Fe equal 1.25 ± 0.04 considering chemical formulas FeO for blue slope. From the slope coefficient of the red curve we can derive RSC of Fe equal to 1.05 ± 0.05 assuming that FeO₂ is present in the sample (Fig.6 panel d). From these data we obtain average RSC of Fe equal to 1.15 ± 0.05 . RSC value for Fe is typically close to 1 (within 15% uncertainty) as it can be determined in the analyses of various materials including metallic or rock samples^{19, 22, 51}. The latter compound, FeO₂, however, is unlikely to be present in our sample. The most common oxidation states of Fe are +2 and +3. With these valences Fe can be found in minerals such FeO (ferric oxide), Fe₂O₃ (Hematite) or Fe₃O₄ (Magnetite). Most viable candidate instead can be iron (II) hydroxide or ferrous hydroxide with the formula Fe(OH)₂. Ferrous hydroxide is produced when iron(II)

salts, e.g., iron(II) sulphate is treated with hydroxide ions. This process is likely to occur in aqueous environments including hydrothermal systems⁵². The slope coefficient of the third correlation of Fe with O observed in Fig. 6c is smaller than the one assigned to the oxides. This result implies that those measured atomic signals may originate from more complex mixture of relevant chemical compounds or mineral phases. For a sample consisting of heterogeneities within a volume smaller than the sampled volume, the determination of RSC from the mass spectrum at this location would yield inaccurate results, because an unique element composition is assumed for the entire volume, thus entire data set. Using depth profiling, the mass peak intensity correlation method can be applied for each individual layer and can yield more accurate results. The specific chemical composition of each layer is preserved and the spatial resolution with the profiling depth is expected to be in sub-micrometre range. In this way an isolation between various compound can be achieved and the different mineralogical phases are recognised by means of their unique elemental composition (abundance ratio) within the various ablation layers. Hence, they can be identified from the correlation studies. Nevertheless, the lateral resolution of this study is at least order of magnitude lower and may affect an isolation of the signals if more than one unique mineralogical grains will be ablated simultaneously.

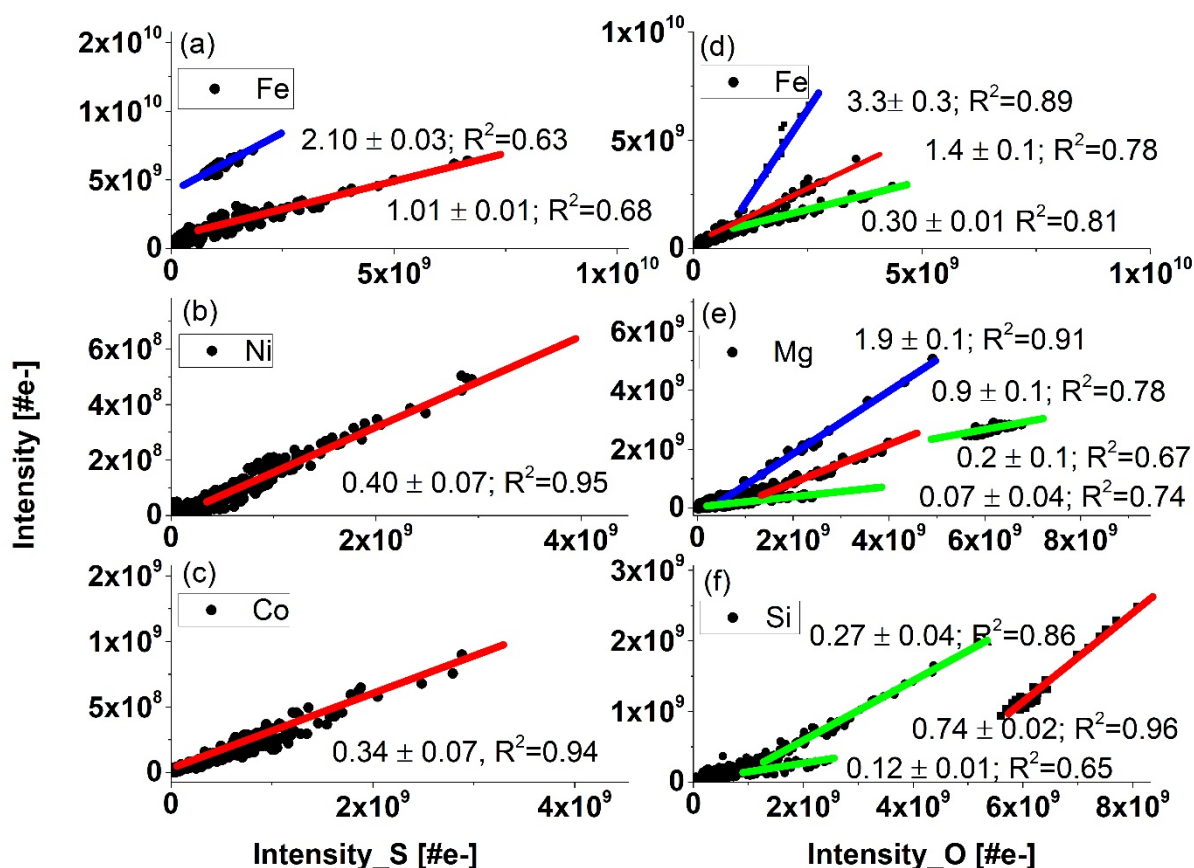


Fig. 6 The correlation of Fe, Ni, and Co with S (panels a-c), and the correlation of Fe, Mg, and Si with O (panels d-f) based on all data measure on locations 6 to 9 (see Fig. 1b). The y-axis represents the intensity of the indicated element in the legend. The slopes (blue) are given for each correlation curve for the oxides and sulphides_RSCs are determined from the blue slopes. Lower slope values than expected for sulphides and oxides indicates that the analysed element was within another compound or a more complex.

In similar analysis conducted for the correlation of Fe with S, we observe two linearly correlated populations (Fig. 6a). Taking the RSC value of Fe as 1.15 ± 0.05 , we can determine the RSC value of S. The two correlation curves give the slope coefficients 2.10 ± 0.03 (blue) and 1.01 ± 0.01 (red), respectively. Assigning the blue and red slopes as being due to the compounds FeS and FeS₂, we can derive RSC of S equal to 0.54 ± 0.06 and 0.57 ± 0.04 , respectively. The FeS and FeS₂ are the most common forms of iron sulphides known in natural environments, such as Mackinawite (metastable form of FeS, pyrrhotite (thermodynamically more stable form of FeS), cubic FeS, (formed in presence of foreign ions e.g. chloride ions), and pyrite (FeS₂). These compounds are formed frequently also in hydrothermal environments. In our previous study, no pyrite could be identified in the

inclusion though this compound is the most important sulphide gangue mineral and occurs in virtually all major hydrothermal base and precious metal mineral deposits. Nevertheless, if the deposits are subjected to high-grade metamorphism, pyrrhotite (FeS) replaces pyrite. This occurs typically by dissolution of pyrite from aqueous solutions and presence of Fe(SH)₂ instead FeS₂ in inclusion material^{53, 54, 52}.

Similar analyses of RSCs of other elements including Mg, Al, Si, and Ca are conducting assuming that these elements can be present as sulphides and oxides. For the Mg and O intensity correlation, we observe four slopes and with S one slope. The slope with the coefficient 1.9 ± 0.1 corresponds likely to MgO. The correlation slope coefficient 0.91 ± 0.06 indicates that MgO₂ could be

present in the inclusion. Again, in this case similarly to other above cases, it is more likely that $\text{Mg}(\text{OH})_2$ is present instead. The slope value in correlation Mg and S is consistent with the measurement on MgS. Three linear correlations of Si with O are observed. Quartz (SiO_2) is typically, the gangue mineral, most commonly encountered in ore-forming hydrothermal systems. Slope coefficient 0.74 in correlation with O is taken to derive the RSC of this element. It is likely that again $\text{Si}(\text{OH})_2$ instead of SiO_2 is present in the sample. The correlation curves with smaller slope coefficients can reflect the presence of other mineralogical phases with a lower Si fraction in the compound. Similarly, for Al correlation with O, the measured slope coefficient is consistent with presence $\text{Al}(\text{OH})_3$ rather than common compound Al_2O_3 . The analysis of Mg and Ca correlation curves with S yield RSCs similar to that obtained for correlation with O or from the analysis of calcium carbonate. The correlation of Ca with S with measured slope coefficient indicates presence of $\text{Ca}(\text{SH})_2$ and the the analysis of correlation Al with S is consistent with Al_2S_3 . The results of calculations of RSCs are given in Table 7.

The slope coefficients of the Ni S, and Co and S correlation curves are 0.34 ± 0.07 and 0.4 ± 0.07 , respectively, yield however much lower values of RSCs as expected for these metals (Table 7). These coefficients are somewhat too small to yield the expected chemical formula NiS and CoS, assuming that RSCs of these elements are typically close to one^{22, 51}. One of the reason for these RSCs being lower than 1 can be that the compounds NiS and CoS are well mixed and effectively a (Ni, Co)S compound is sampled in our experiment. The determination of RSC for S (from Fe-sulphides) also allows to conclude that the size of the other Ni and Co sulphides inclusions are smaller than the ablated layer thickness ($< \sim 150$ nm). The ablation layers where NiS and CoS are measured cannot be resolved to see separate compounds. The optical imaging of the inclusion area shows large density of needles characteristic of millerite mineral⁷. In addition to Co, also Cu, Mn, and Fe are known impurities, which are identified in millerite. Figure 5 shows that at the ablation

layers at which an increase of the Ni, Co and S signals is observed also signals of Mn, Cu or Mg increases. An exception are Fe and Ti signals, which decrease for these specific ablation layers.

A small part of data points of Ti are observed to linearly correlate with the slope coefficient 0.22 ± 0.03 indicating that Ti may be a part of more complex mineralogical phase.(Fig. 5, panel c). In the analysed locations on the inclusion there is also no correlation of Ti with O observed. TiO_2 (anatase mineral) was identified, however, in the previous studies in other inclusion locations. For the other trace elements including B, V, Cr, Mn, Zn and Cu measured here the correlation curves are too diffuse and the linear fit is not applied. However, their RSCs are expected to be close to one^{18, 20, 51}. The RSCs for non-metals such as Si, C, O, S or P are typically smaller than one. RSCs of Na and K are typically larger than 1, because it is relatively easy to ionise alkali elements comparing to other elements. RSCs of Na and K could not be derived from the correlation of these elements with O or S. For the minor and trace elements, the measurement intensities have larger statistical fluctuations resulting in less accurate determination of the slope coefficients. There are ablation layers in which the relevant atomic signals are weak or not detectable. For too large signals, the mass peak intensities can be under-measured due to the detector saturation effects. For large differences in material properties from one ablation layer to the other, the ablation process and efficiency of ion production may change affecting linearity of the atomic intensity correlation. The effects affecting the mass peak intensity measurements are discussed in more detail in our recent publication⁵⁵. The RSCs of investigated elements are summarised in Table 1.

For a sample consisting of heterogeneities within a volume smaller than the sampled volume, the determination of RSC from the mass spectrum at this location would yield inaccurate results, because an unique element composition is assumed for the entire volume, thus entire data set. Using depth profiling, the mass peak intensity correlation method can be applied for each individual layer, which yields more accurate

⁵⁵results. The specific chemical composition of each layer is preserved and the spatial resolution with the profiling depth is expected to be in sub-micrometre range. In this way an isolation between various compounds can be achieved and the different mineralogical phases are recognised by means of their unique elemental composition (abundance ratio) within the various ablation layers. Hence, they can be identified from the correlation studies. Nevertheless, the lateral resolution of this study is at least order of magnitude lower and may affect an isolation of the signals if more than one unique mineralogical grains will be ablated simultaneously.

Only, RSCs of major and some minor elements can be obtained using present method (see also ESI). The corrections were applied to the atomic intensities of the major species Ca, O, S, Mg and Si. Such corrections can be important when the analysis of mineralogical phases are conducted as the rock chemistry depends on the concentrations of these elements. In ESI, Section B, we show also the atomic intensities (background level value 10^7) and atomic fractions of all elements identified in this study.

Mineralogical phases

Calcium carbonate

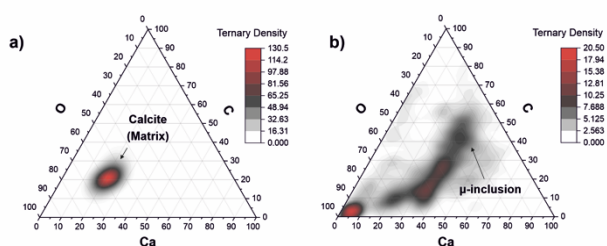


Fig. 5 Ternary diagram of the C, O, and Ca concentration fractions with applied RSCs that were derived on a pure calcium carbonate phase at location 1-4. (a) and on the dark inclusion (b) at locations from 6 to 9. Only a part of the data points in panel (b) can be assigned to calcium carbonate mineral. On the O axes the 3 times O concentration is applied thus the fractions of C, Ca and 3O are the same for carbonate mineral and experimental points are in the middle area in the triangle. The other remaining points indicate presence of other mineralogical compounds.

Ternary plots are frequently used in the mineralogical analysis in geology. They can be helpful in

identification of the mineralogical context of the analysed sample just by applying a minimum number of chemical species, which are necessary to describe the composition. In the conventional analysis of carbonate phases, ternary diagram correlating abundance fractions (Ciderite-Magnesite-Calcite) is applied. For pure calcium carbonate, the ternary plot of the C, O, and Ca abundance fractions is shown in Fig.7, panel a. The experimental data points are observed to be located close to each other, with element abundance ratios (atomic fractions, 20:20:60:) that agrees with the chemical formula CaCO_3 . Contrary, the plot of the data representing the measurements on locations 6 to 9 shows very diffuse distribution with large spread from the area expected for the calcite mineral (Fig. 7, panel b). This may indicate that the other minerals instead a calcium carbonate are present at the inclusion location. This goes along with our previous investigations with spectroscopic methods, which identified the presence of mineralogical phases such as clinochlore $[(\text{Mg}_5\text{Al})(\text{AlSi}_3)\text{O}_{10}(\text{OH})_8]$, clinochlore-transformed augite $(\text{Ca, Na})(\text{Mg, Fe, Al, Ti})(\text{Si, Al})_2\text{O}_8$ and plagioclase (feldspar group) $[\text{NaAlSi}_3\text{O}_8\text{-CaAl}_2\text{Si}_2\text{O}_8]$.

Clinochlore, Pyroxene-Augite, Plagioclase

Clinochlore is a member of chlorite group of phyllosilicate minerals and often contains considerable amounts of Fe^{2+} , grading into chamosite. Clinochlore is described by the chemical formula $\text{Mg}_5\text{Al}(\text{AlSi}_3\text{O}_{10})(\text{OH})_8$, and Fe, Mn, Zn, Ca, and Cr are common impurities. Figure 8a shows the ternary plot of the Mg, Al, and Fe element abundances to illustrate their relative proportions measured in the dark inclusion (Fig. 1b). These proportions indicate that the clinochlore mineral is a major mineral in the dark inclusion with the average $\text{Mg}/(\text{Mg}+\text{Fe})$ abundance ratio of $\sim 0.34 \pm 0.14$. Within ultramafic rocks, metamorphism produces predominantly clinochlore chlorite, an Mg-rich end member, which could be identified as the inclusion.

Clinochlore-transformed augite mineral is a member of pyroxene group namely calcic clinopyroxenes (diopside, hedenbergite, augite). Pure diopside of $\text{CaMgSi}_2\text{O}_6$ chemical composition can undergo as a

complete solid solution to hedenbergite $\text{CaFeSi}_2\text{O}_6$ or augite $(\text{Ca},\text{Mg},\text{Fe}^{2+},\text{Fe}^{3+},\text{Al})_2(\text{Si},\text{Al})_2\text{O}_6$. In another chemical formula proposed for pyroxene group $(\text{Ca}_x\text{Mg}_y\text{Fe}_z)(\text{Mg}_{y1}\text{Fe}_{z1})\text{Si}_2\text{O}_6$, for augite, the relative composition of Ca can vary within the range $0.4 \leq x \leq 0.9$ where $x+y+z=1$ and $y1+z1=1$ ⁵⁶. The ternary plot of Mg, Fe, and Ca proportions indicates presence of magnesium augite in the inclusion material (Fig. 8b).

Figure 8c shows the ternary diagram of the abundances of the elements K, Na, and Ca, which is applied for the classification of feldspar minerals. One can identify the distribution of data along the base of the K coordinate, and the miscibility gap in the centre. The sequence of minerals along the base of the triangle indicate the presence of the plagioclase mineral or the

In our previous studies the analyses conducted by various analytical techniques showed that the dark inclusion investigated here is highly weathered basalt with olivine, and pyroxene (augite) phenocrysts with a dominating amygdales phase⁷. Typically, amygdales in basaltic lavas are formed after lava cools down, and hydrothermal fluids pass through the cavities. Their precipitation in the bubbles form secondary minerals such as calcium carbonate, zeolites, and quartz, depending on the fluid composition. The depth profiles indicate also presence of layers of various mineralogical phases around the millerite needles. Oxides and sulphides of several metals are present as separate entities and some of the metals fill the sites of complex mineralogical compounds⁵².

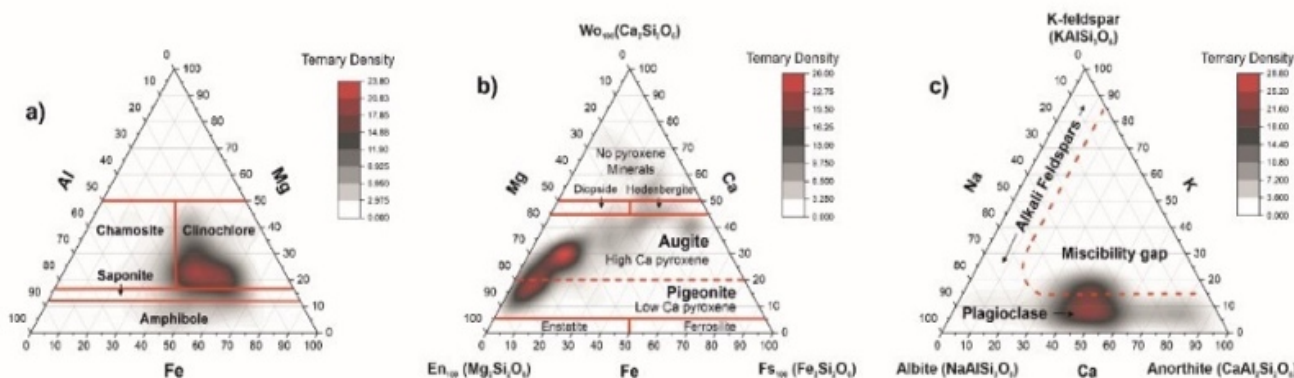


Fig. 8 Ternary plots of the elements relevant to chlorite and feldspar groups. (a): The correlation of the Al, Mg, and Fe abundances showing characteristic regions for various minerals of the chlorite group; (b): The correlation between the Mg, Fe, and Ca abundance ratios indicating the presence of magnesium augite and pigeonite in the sample material; (c): The correlation diagram of the Ca, K, and Na elements indicate the presence of the plagioclase feldspars in the inclusion material.

solid solution series of plagioclase being a mixture between albite (sodium aluminosilicate: $[\text{Na}(\text{AlSi}_3\text{O}_8)]$) and anorthite (calcium aluminosilicate: $[\text{Ca}(\text{Al}_2\text{Si}_2\text{O}_8)]$) and the notable absence of alkali feldspars⁵⁷. The red line denoting location of the feldspars is schematic. The RSC values of K and Na are chosen 1, which can be the reason that we observe the plagioclase on ternary diagram shifted more in the alkali region. The RSCs of K and Na are expected to be larger than 1 but they cannot here be derived accurately using the abundance correlation method applied to simple compounds.

Summary

We describe the mass spectrometric measurements conducted by fs-laser ablation/ionisation time-of-flight mass spectrometry and introduce principles of data analysis, which yields to true elemental abundances. The corrections to measured atomic abundances are obtained by analysing atomic intensity correlations using depth-profiling data. In heterogeneous sample such as analysed inclusion embedded in a calcium carbonate mineral, chemical compounds are present at specific depths and can be identified in specific ablation layers. By correlating the measured atomic intensities,

one can select the data points for which atomic intensities, correlate linearly. With the slope coefficients of these correlations, the relative abundances of the correlated elements are derived. The method was tested successfully on a pure calcium carbonate phase followed by locations at the interface between inclusion and calcite mineral and on the inclusion. RSCs for major and minor elements were derived by correlating atomic intensity of these elements with either oxygen or sulphur. Subsequently, with the corrected elemental abundances more complex mineralogical phases were analysed by applying the ternary atomic abundance fraction correlations on specific elements typically chosen in geological analysis of the mineralogical context.

The mass spectrometric results are consistent with the results presented in our former studies of this sample conducted with several other instruments. However, LIMS studies cannot differentiate between minerals with the same chemical formula such as calcium carbonate mineral, which can occur as aragonite, calcite or vaterite. To conduct such analysis the Raman measurements will be necessary. Similarly, with the IR analysis one would approve our ambiguous assignment of presence of hydroxides or hydrosulphides in the sample. The presented method can be important in studies of highly heterogeneous materials, because matrix-matched reference materials are often not available for complex natural samples. The in situ determined RSCs can overcome the problems associated with the local heterogeneity, which can occur in the micrometre range similar to the sample probing laser spot. Thus, the in situ atomic abundance calibration introduced here can allow quantitative chemical composition investigation of heterogeneous structures. With the detection of heterogeneity in the micrometre scale, one can analyse grain-size objects such as micro-sized minerals or microfossils in detail. The presented method can be also of interest in searches for ancient fossilised life forms on planetary surfaces because in addition to context mineralogy, high depth profiling resolution provides means for isolation and

analyses of microfossils embedded in mineralogical phases.

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References

1. K. E. Peters, J. M. Walters and J. M. e. Moldowan, *The Biomarker Guide*, Cambridge University Press, 2005.
2. A. H. Stevens, A. McDonald, C. de Koning, A. Riedo, L. J. Preston, P. Ehrenfreund, P. Wurz and C. S. Cockell, *Sci Rep-Uk*, 2019, **9**.
3. A. Riedo, C. de Koning, A. H. Stevens, C. S. Cockell, A. McDonald, A. C. Lopez, V. Grimaudo, M. Tulej, P. Wurz and P. Ehrenfreund, *Astrobiology*, 2020, **20**, 1224-1235.
4. A. Riedo, C. de Koning, A. Stevens, A. C. o. L. p. McDonald, A, M. Tulej, P. Wurz, C. S. Cockell and P. Ehrenfreund, *Astrobiology*, 2020, **accepted**.
5. J. M. Reess, M. Bonafous, L. Lapauw, O. Humeau, T. Fouchet, P. Bernardi, P. Cais, M. Deleuze, O. Forni, S. Maurice, S. Robinson, R. C. Wiens and S. Team, *Proc Spie*, 2018, **11180**.
6. P. Sarrazin, D. Blake, M. Gailhanou, F. Marchis, C. Chalumeau, S. Webb, P. Walter, E. Schyns, K. Thompson and T. Bristow, *J Instrum*, 2018, **13**.
7. A. Neubeck, M. Tulej, M. Ivarsson, C. Broman, A. Riedo, S. McMahon, P. Wurz and S. Bengtson, *International Journal of Astrobiology*, 2016, **15**, 133-146.
8. M. Tulej, A. Riedo, M. B. Neuland, S. Meyer, P. Wurz, N. Thomas, V. Grimaudo, P. Moreno-Garcia, P. Broekmann, A. Neubeck and M. Ivarsson, *Geostand Geoanal Res*, 2014, **38**, 441-466.
9. M. Tulej, A. Neubeck, M. Ivarsson, A. Riedo, M. B. Neuland, S. Meyer and P. Wurz, *Astrobiology*, 2015, **15**, 669-682.
10. Y. M. Lin, Q. A. Yu, W. Hang and B. L. Huang, *Spectrochim Acta B*, 2010, **65**, 871-883.
11. A. A. Sysoev and A. A. Sysoev, *Eur J Mass Spectrom*, 2002, **8**, 213-232.
12. A. A. Sysoev, A. V. Karpov, V. V. Milyaeva and A. A. Sysoev, *Eur J Mass Spectrom*, 2018, **24**, 96-107.
13. U. Rohner, J. A. Whitby and P. Wurz, *Meas Sci Technol*, 2003, **14**, 2159-2164.
14. M. Tulej, M. Iakovleva, I. Leya and P. Wurz, *Anal Bioanal Chem*, 2011, **399**, 2185-2200.
15. A. Riedo, A. Bieler, M. Neuland, M. Tulej and P. Wurz, *J Mass Spectrom*, 2013, **48**, 1-15.
16. R. Wiesendanger, M. Tulej, A. Riedo, S. Frey, H. Shea and P. Wurz, *J Anal Atom Spectrom*, 2017, **32**, 2182-2188.
17. R. Wiesendanger, D. Wacey, M. Tulej, A. Neubeck, M. Ivarsson, V. Grimaudo, P. Moreno-Garcia, A. Cedeno-Lopez, A. Riedo and P. Wurz, *Astrobiology*, 2018, **18**, 1071-1080.
18. J. He, R. F. Huang, Q. Yu, W. Hang and B. L. Huang, *Chem J Chinese U*, 2009, **30**, 57-59.
19. B. C. Zhang, M. H. He, W. Hang and B. L. Huang, *Anal Chem*, 2013, **85**, 4507-4511.

20. S. D. Zhang, B. C. Zhang, W. Hang and B. L. Huang, *Spectrochim Acta B*, 2015, **107**, 17-24.
21. Z. B. Yin, L. Hang, R. Liu, W. Hang and B. L. Huang, *J Mass Spectrom*, 2018, **53**, 435-443.
22. A. Riedo, M. Neuland, S. Meyer, M. Tulej and P. Wurz, *J Anal Atom Spectrom*, 2013, **28**, 1256-1269.
23. V. V. Milyaeva, E. E. Sil'nikov, A. M. Mikhailov and A. A. Sysoev, *Eur J Mass Spectrom*, 2017, **23**, 167-173.
24. M. Tulej, R. Wiesendanger, A. Riedo, G. Knopp and P. Wurz, *J Anal Atom Spectrom*, 2018, **33**, 1292-1303.
25. R. F. Huang, Y. M. Lin, L. F. Li, W. Hang, J. He and B. L. Huang, *Anal Chem*, 2010, **82**, 3077-3080.
26. R. Wiesendanger, V. Grimaudo, M. Tulej, A. Riedo, R. Lukmanov, N. Ligterink, R. Fausch, H. Shea and P. Wurz, *J Anal Atom Spectrom*, 2019, **34**, 2061-2073.
27. M. B. Neuland, S. Meyer, K. Mezger, A. Riedo, M. Tulej and P. Wurz, *Planet Space Sci*, 2014, **101**, 196-209.
28. S. Frey, R. Wiesendanger, M. Tulej, M. Neuland, A. Riedo, V. Grimaudo, P. Moreno-Garcia, A. C. Lopez, M. Mohos, B. Hofmann, K. Mezger, P. Broekmann and P. Wurz, *Planet Space Sci*, 2020, **182**.
29. A. Riedo, M. Tulej, U. Rohner and P. Wurz, *Rev Sci Instrum*, 2017, **88**.
30. V. Grimaudo, P. Moreno-Garcia, A. C. Lopez, A. Riedo, R. Wiesendanger, M. Tulej, C. Gruber, E. Lortscher, P. Wurz and P. Broekmann, *Anal Chem*, 2018, **90**, 5179-5186.
31. V. Grimaudo, P. Moreno-Garcia, A. Riedo, M. B. Neuland, M. Tulej, P. Broekmann and P. Wurz, *Anal Chem*, 2015, **87**, 2037-2041.
32. V. Grimaudo, M. Tulej, A. Riedo, R. Lukmanov, N. F. W. Ligterink, C. de Koning and P. Wurz, *Rapid Commun Mass Sp*, 2020, **34**.
33. V. Grimaudo, P. Moreno-Garcia, A. Riedo, S. Meyer, M. Tulej, M. B. Neuland, M. Mohos, C. Gutz, S. R. Waldvogel, P. Wurz and P. Broekmann, *Anal Chem*, 2017, **89**, 1632-1641.
34. M. H. He, Y. F. Meng, S. S. Yan, W. Hang, W. G. Zhou and B. L. Huang, *Anal Chem*, 2017, **89**, 565-570.
35. B. Fernandez and R. Pereiro, *Anal Bioanal Chem*, 2010, **396**, 2723-2724.
36. B. Fernandez, R. Pereiro and A. Sanz-Medel, *Anal Chim Acta*, 2010, **679**, 7-16.
37. B. Hattendorf, J. Pisonero, D. Gunther and N. Bordel, *Anal Chem*, 2012, **84**, 8771-8776.
38. A. Pelster, M. Korsgen, T. Kurosawa, H. Morita and H. F. Arlinghaus, *Anal Chem*, 2016, **88**, 9638-9646.
39. J. Pisonero, *Anal Bioanal Chem*, 2006, **384**, 47-49.
40. J. Pisonero, J. Koch, M. Walle, W. Hartung, N. D. Spencer and D. Gunther, *Anal Chem*, 2007, **79**, 2325-2333.
41. J. Fietzke, V. Liebetrau, D. Guenther, K. Gurs, K. Hametner, K. Zumholz, T. H. Hansteen and A. Eisenhauer, *J Anal Atom Spectrom*, 2008, **23**, 955-961.
42. J. Fietzke, V. Liebetrau, D. Gunther, A. Frische, K. Zumholz, T. H. Hansteen and A. Eisenhauer, *Geochim Cosmochim Acta*, 2008, **72**, A267-A267.
43. V. N. Epov, S. Berail, M. Jimenez-Moreno, V. Perrot, C. Pecheyran, D. Amouroux and O. F. X. Donard, *Anal Chem*, 2010, **82**, 5652-5662.
44. M. Ivarsson, *Biogeosciences*, 2012, **9**, 3625-3635.
45. M. Ivarsson, W. Bach, C. Broman, A. Neubeck and S. Bengtson, *Geomicrobiol J*, 2018, **35**, 460-467.
46. M. Ivarsson, S. Bengtson, V. Belivanova, M. Stampanoni, F. Marone and A. Tehler, *Geology*, 2012, **40**, 163-166.
47. S. McMahon, T. Bosak, J. P. Grotzinger, R. E. Milliken, R. E. Summons, M. Daye, S. A. Newman, A. Fraeman, K. H. Williford and D. E. G. Briggs, *J Geophys Res-Planet*, 2018, **123**, 1012-1040.
48. S. McMahon, J. Parnell and N. J. F. Blamey, *International Journal of Astrobiology*, 2012, **11**, 163-167.
49. R. Wiesendanger, M. Tulej, V. Grimaudo, A. C. Lopez, R. Lukmanov, A. Riedo and P. Wurz, *Journal of Chemometrics*, 2018, DOI: 10.1002/cem.3081, 1-10.
50. S. Meyer, A. Riedo, M. B. Neuland, M. Tulej and P. Wurz, *J Mass Spectrom*, 2017, **52**, 580-590.
51. M. B. Neuland, V. Grimaudo, K. Mezger, P. Moreno-Garcia, A. Riedo, M. Tulej and P. Wurz, *Meas Sci Technol*, 2016, **27**.
52. T. M. Seward, A. Williams-Jones and A. Migdisov, in *Treatise on Geochemistry*, Elsevier Ltd., 2013, vol. 13.
53. U. M. Graham and H. Ohmoto, *Geochim Cosmochim Acta*, 1994, **58**, 2187-2202.
54. H. Ohmoto, K. Hayashi and Y. Kajisa, *Geochim Cosmochim Acta*, 1994, **58**, 2169-2185.
55. M. Tulej, A. Neubeck, A. Riedo, R. Lukmanov, V. Grimaudo, N. Ligterink, M. Ivarsson, W. Bach and C. de Koning, *J Mass Spectrom*, 2020, **55**, e4660.
56. N. Morimoto, *Mineral Mag*, 1988, **52**, 535-550.
57. <http://geology.com/minerals/plagioclase.shtml>).